Recycling: A New Mechanism for Producing Ultra Heavy Particle Dark Matter

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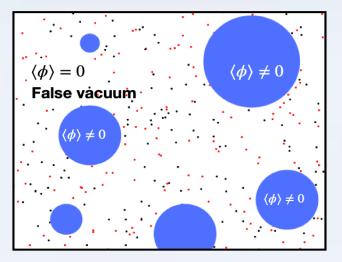
arXiv: 2310.08526

Outline of Talk

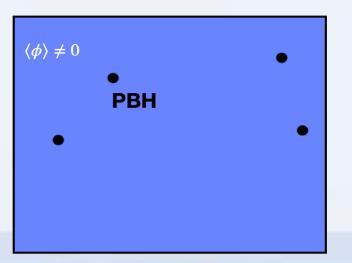
- Overall story of recycled dark matter
- Trapped (Multicomponent Dark Sector)
- PBH formation
- UHDM produced from PBHs

Recycling Mechanism

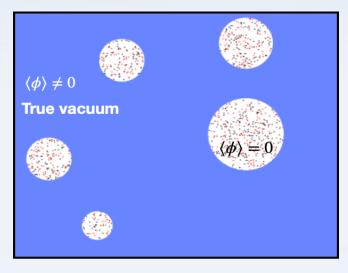
(1) First Order Phase Transition



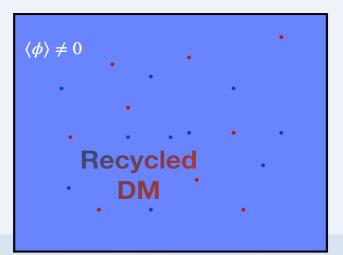
(3) Black hole formation



(2) Dark sector trapped



(4) Hawking radiation



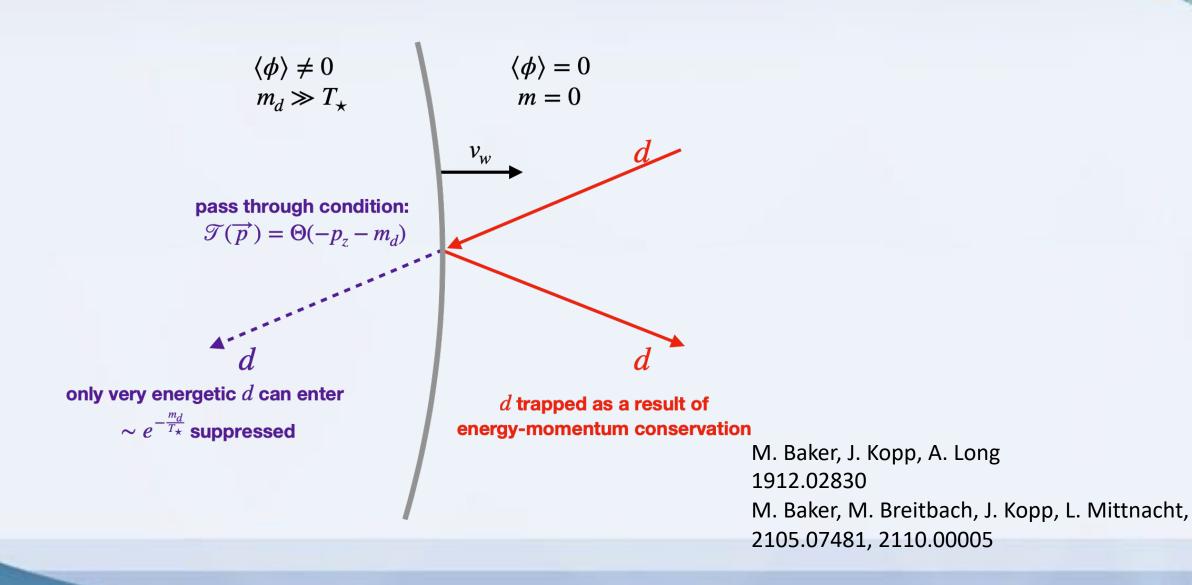
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Dark Sector FOPT

$$\mathscr{L} \supset - \underbrace{y_{\chi} \phi \bar{\chi} \chi}_{\text{VLawa mass}} + \underbrace{\mu^2 \phi^2 - \lambda \phi^4}_{\text{FOPT}} + \underbrace{\mathscr{L}_{\text{SM-DS}}}_{\text{SM-DS}}$$

• $\mathscr{L}_{\rm SM-DS}$ can keep dark sector and SM sector in equilibrium before FOPT,

Trapped Dark Sector



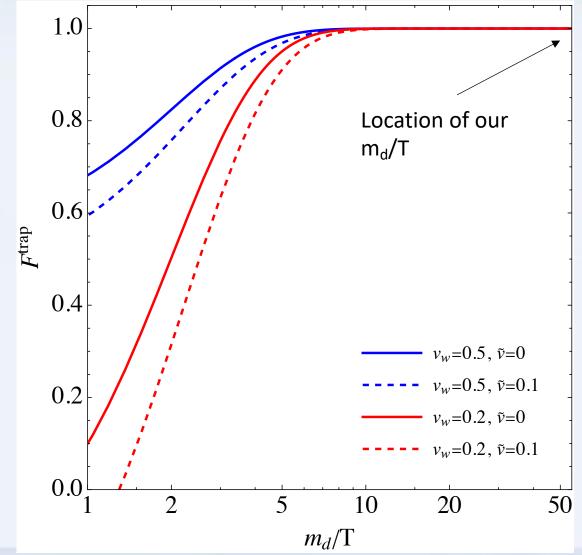
Trapped Dark Sector During FOPT

$$J_w = \frac{g_d T^3 (1 + \tilde{\gamma} m_d (1 - \tilde{v})/T)}{4 \pi^2 \tilde{\gamma}^3 (1 - \tilde{v})^2} e^{-\tilde{\gamma} m_d (1 - \tilde{v})/T}.$$

 $n_d^{ ext{filtered}} = rac{J_w}{\gamma_w \, v_w}$

$$F^{\rm trap} = 1 - n_d^{\rm filtered} / n_d$$

- The probability of a DM particle penetrating through the moving bubble wall is given by $J_{w.}$ The probability of being trapped in false vacuum is given by $F^{trapped}$
- Tildes are in the bubble wall rest frame. Quantities without tildes are in the global plasma rest frame
- The greatest suppression to trapping is in the exponential which is controlled by dark matter mass and particle temperature



Thermal History

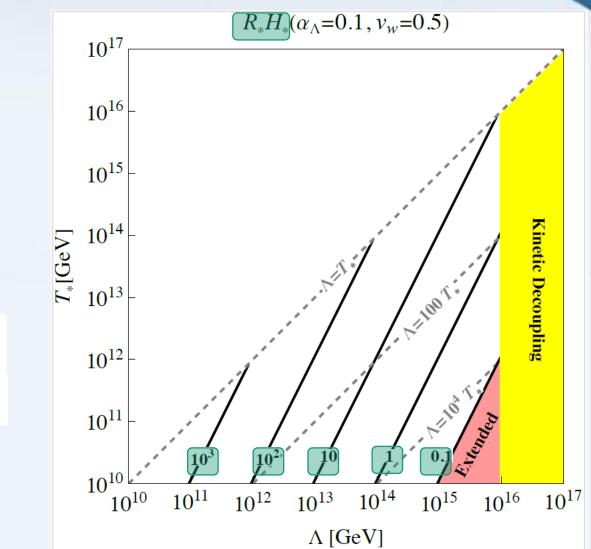
$$\chi + \bar{\chi} \to \mathrm{SM} + \mathrm{SM}$$

$$\mathcal{L}_{\rm SM-DS} = \frac{\alpha_{\Lambda}}{\Lambda} \bar{\chi} \chi H^{\dagger} H$$

$$lpha_\Lambda \gtrsim 0.17 imes \left(rac{g_\star}{106.75 + 4.5}
ight)^{1/4} \left(rac{\Lambda}{10^{16} \ {
m GeV}}
ight)^{1/2}$$

When T < Λ the cross section is a constant $\Gamma_{ann} \times t_c < 1$ Time scale of collapse: $\Gamma_{ann} = n_{\chi} \frac{\alpha_{\Lambda}^2}{\Lambda^2}$ $t_c \simeq R_{\star} / v_w$

$$R_{\star} > 0.27 \, H_{\star}^{-1} \, \left(\frac{\alpha_{\Lambda}}{0.1}\right) \, \left(\frac{0.5}{v_w}\right)^{-1} \, \left(\frac{T_{\star}}{\Lambda}\right) \, \left(\frac{M_{\rm pl}}{T_{\star}}\right)^{-1}$$



PBH Formation

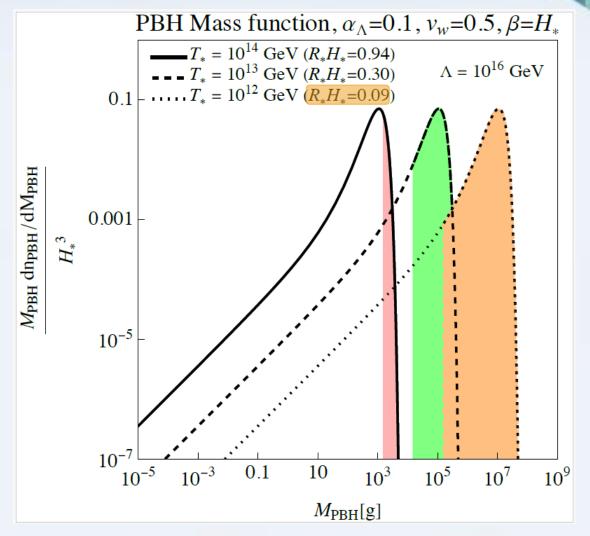
$$M_{\rm PBH} = \left[\frac{2\pi^3}{90} \left(g_\phi + \frac{7}{8}g_\chi\right)\right]^{1/2} R_\star^2 T_\star^2 M_{\rm PH}$$

Include lower bound due to annihilation in small pockets

$$R_{\star} > 0.27 \, H_{\star}^{-1} \, \left(\frac{\alpha_{\Lambda}}{0.1}\right) \, \left(\frac{0.5}{v_w}\right)^{1/2} \, \left(\frac{T_{\star}}{\Lambda}\right) \, \left(\frac{M_{\rm pl}}{T_{\star}}\right)^{1/2}$$

$$\frac{dn_{\rm PBH}}{dM_{\rm PBH}} \simeq \frac{I_{\star}^4 \,\beta^4 M_{\rm PBH}^{-1/2}}{384 \, T_{\star} \, v_w^3 \left(\frac{2\pi^3}{90\mathrm{G}} \left(g_{\phi} + \frac{7}{8}g_{\chi}\right)\right)^{1/4}} \left(e^{4\beta R_{\star}/v_w - I_{\star} \,e^{\beta R_{\star}/v_w}}\right) \left(1 - e^{-I_{\star} e^{\beta R_{\star}/v_w}}\right) \times \Theta\left(M_{\rm PBH} - 6.9 \times 10^{-3} \sqrt{g_{\phi} + \frac{7}{8}g_{\chi}} \,\frac{\alpha_{\Lambda}^2}{v_w} \,\frac{M_{\rm Pl}^4}{T_{\star}\Lambda^2}\right)$$

Lower cutoff of M_{PBH} due energy leaving the pocket



PBH Evaporation & DM Relic Abundance

Total number of DM = (Number of DM from PBHs) x (Total number of PBHs)

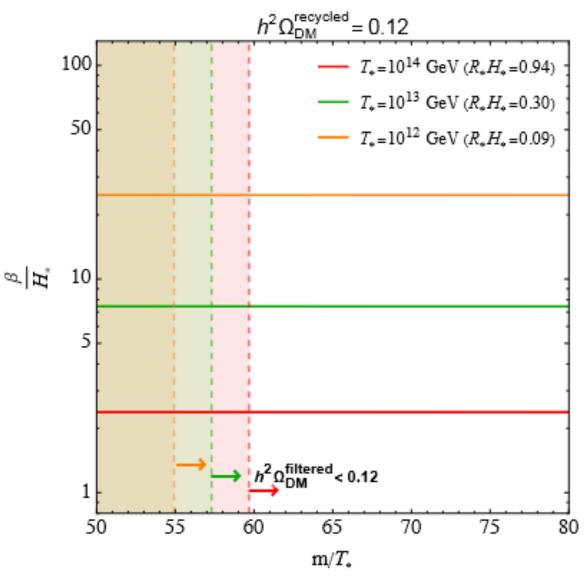
$$N_{i} = \frac{120 \zeta(3)}{\pi^{3}} \frac{g_{i}}{g_{\star}(T_{\text{PBH}})} \frac{M_{\text{PBH}}^{2}(t_{\star})}{M_{\text{Pl}}^{2}}, \qquad T_{\text{PBH}}(t_{\star}) > m_{i},$$

$$N_{i} = \frac{15 \zeta(3)}{8\pi^{5}} \frac{g_{i}}{g_{\star}(T_{\text{PBH}})} \frac{M_{\text{Pl}}^{2}}{m_{i}^{2}} \qquad T_{\text{PBH}}(t_{\star}) < m_{i}.$$

$$Y_{\text{DM}} = \frac{3}{4} \beta_{\text{PBH}} N_{\text{DM}} \frac{T_{\star}(M_{\text{PBH}})}{M_{\text{PBH}}}$$

$$\Omega_{\mathrm{DM}} = rac{
ho_{\mathrm{DM}(t_0)}}{
ho_c} = rac{m_{\mathrm{DM}} r_{\mathrm{DM}}}{
ho_c} s(t_0)$$

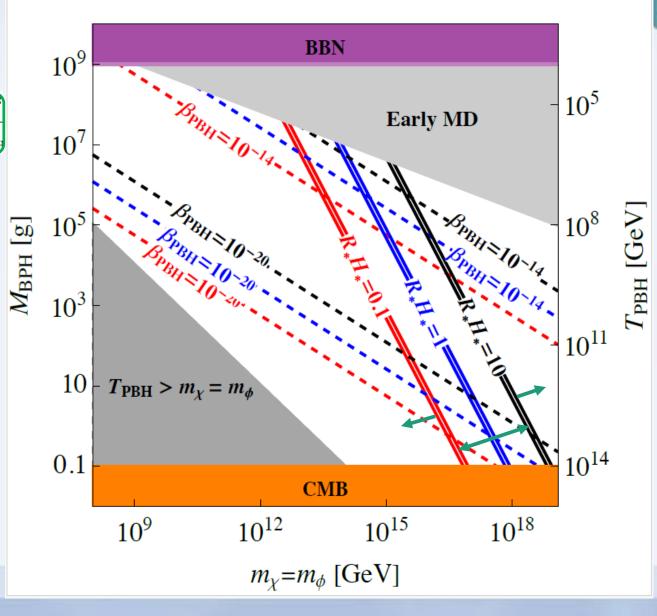
$$\Omega_{DM} = \frac{45\sqrt{3} \times 5^{1/4}\zeta(3)}{16 \times 2^{1/4}\pi^{23/4}} \frac{(3g_{\chi} + 4g_{\phi})(7g_{\chi} + 8g_{\phi})^{1/4}}{(8g_{\star} + 7g_{\chi} + 8g_{\phi})^{3/2}} \beta_{\text{PBH}} \frac{R_{\star}}{H_{\star}^{-1}} \frac{s(t_0)}{m\rho_c} \frac{M_{\text{PBH}}^{7/2}}{M_{\text{PBH}}^{3/2}}$$



Recycling Dark Matter Parameter Space

$$\Omega_{DM} = \frac{45\sqrt{3} \times 5^{1/4}\zeta(3)}{16 \times 2^{1/4}\pi^{23/4}} \frac{(3g_{\chi} + 4g_{\phi})(7g_{\chi} + 8g_{\phi})^{1/4}}{(8g_{\star} + 7g_{\chi} + 8g_{\phi})^{3/2}} \beta_{\text{PBH}} \frac{R_{\star}}{H_{\star}^{-1}} \frac{s(t_0)}{m\rho_c} \frac{M_{\text{Pl}}^{7/2}}{M_{\text{PBH}}^{3/2}}$$

- Left arrow: Everything left is excluded since this leads to an extended mass function which we leave for future study
- Right arrow: Everything right is excluded since our choice of cut off scale leads effective field theory breaking down in this area
- Double arrow: Everything in between is viable since filtering is exponentially suppressed.



Summary

- Multicomponent DS is trapped during FOPT and collapses into PBH. The same UHDM is reproduced (Recycled) by Hawking evaporation of PBHs and matches observed relic abundance.
- New mechanism for generating Ultra Heavy Dark Matter
- Recycling mechanism can be tested GWs in the MHz-GHz range

Recycled Dark Matter: arXiv: 2310.08526

T. Gehrman, B. Es Haghi, K. Sinha, T. Xao **PBH+Baryogensis+HFGWs:** arXiv:2211.08431 **PBH+DM+HFGWs:** arXiv:2305.09194

Thank You!

Annihilation of Dark Matter After Evaporation of PBHS

• Issue may arise where annihilation of dark matter after PBH evaporation may lead to a lower relic abundance

$$\frac{\Gamma_{\chi\chi}(\tau_{PBH}+\epsilon)}{H(\tau_{PBH}+\epsilon)} \simeq \frac{n_{\chi}(\tau_{PBH}+\epsilon)\langle\sigma v\rangle_{\chi\chi}(\tau_{PBH}+\epsilon)}{H(\tau_{PBH}+\epsilon)}$$

$$\langle\sigma v\rangle_{\chi\chi}(\tau_{PBH}+\epsilon) = \frac{\alpha_{\Lambda}^{2}}{\Lambda^{2}} \qquad n_{\chi}(\tau_{PBH}+\epsilon) \simeq N_{\chi}\beta\frac{\rho_{\rm rad}(t_{\star})}{M_{\rm PBH}}\frac{T^{3}(\tau_{\rm PBH})}{T^{3}(t_{\star})}$$

$$\frac{\Gamma_{\chi\chi}(\tau_{PBH}+\epsilon)}{H(\tau_{PBH}+\epsilon)} < \frac{27\sqrt{2}\zeta(3)}{8192\times2^{1/4}\sqrt{5}\pi^{7/2}}\frac{\sqrt{g_{\star}}}{(7g_{\chi}+8g_{\phi})^{1/4}}\frac{\alpha_{\Lambda}^{2}\beta}{C}\frac{M_{\rm Pl}}{M_{\rm PBH}} \ll$$

The Need for a Multicomponent Dark Sector

 Typical FOPT Scenario: One scalar φ is massless in the true vacuum and fermion gains its mass from the symmetry breaking of the scalar. Fermion is the dark matter candidate.

 $y_{\chi}\phi\bar{\chi}\chi \qquad \bar{\chi}\chi
ightarrow \phi\phi$

 The annihilation of a fermion to a massless scalar leads to energy density leaving the pocket. This highly suppresses PBH formation and therefore inhibits recycling.

Recycling: $\langle \phi \rangle \sim \mathcal{O}(55-60) \times T_* \times (1/y_{\chi})$

Possible Questions/ Extra Information

How do ϕ and χ affect PBH formation?

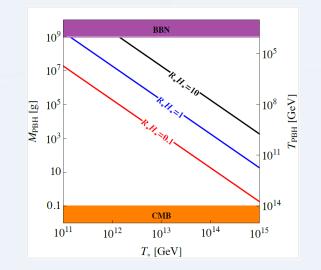
An attractive force caused by ϕ particles (Yukawa force) could enhance PBH formation by causing χ particles to collapse before the degeneracy pressure becomes important. However, we did not pressure the strength of this force in our work, but detailed calculation needs to be done.

$$T>\Lambda \qquad \sigma_{
m ann}=lpha_{\Lambda}^2/T^2 \qquad \Gamma_{
m ann}=n_{\chi/\phi}\,lpha_{\Lambda}^2/T^2\,\propto\,T$$

$$T \lesssim \Lambda$$
 $\Gamma_{
m ann} = n_{\chi/\phi} \, lpha_{\Lambda}^2 / \Lambda^2 \propto T^3$

 $T_\star \lesssim T \lesssim \Lambda$

 $\mathcal{L} \supset -y_\chi \phi ar{\chi} \chi + \mu^2 \phi^2 - \lambda \phi^4 + \mathcal{L}_{\mathrm{SM-DS}}$



Fermi Ball Scenario: A fermion asymmetry in the pocket leads to Fermi ball formation while the scalar is massless leading to a recycling scenario

Black Hole Thermodynamics

$$T_{\rm BH} = \frac{M_{\rm Pl}^2}{8\pi M_{\rm BH}}$$

$$\frac{d^2 u_i(E,t)}{dt dE} = \frac{g_i}{8\pi^2} \frac{E^3}{e^{E/T_{\rm BH}} \pm 1}$$

$$\begin{aligned} \frac{dM_{\rm BH}}{dt} &= -4\pi r_{\rm S}^2 \sum_i \int_0^\infty \frac{d^2 u_i(E,t)}{dt dE} dE = -\frac{g_*(T_{\rm BH})}{30720\pi} \frac{M_{\rm Pl}^4}{M_{\rm BH}^2} \\ M(t) &= M_i \left(1 - \frac{t - t_i}{\tau}\right)^{1/3} \qquad \qquad \tau = \frac{10240\pi}{g_*(T_{\rm BH})} \frac{M_i^3}{M_{\rm Pl}^4} \end{aligned}$$

D. Baumann, P. Steinhardt, and N. Turok 2007 arXiv:hep-th/0703250v1

P. Gondolo, P. Sandick, and B. Shams Es Haghi 2020

$$\label{eq:schwarzschild radius} \begin{split} r_{s} &\equiv Schwarzschild radius\\ T_{BH} &\equiv Black \ Hole \ Temperature\\ M_{i} &\equiv Initial \ Mass \ of \ Black \ Hole\\ Tau &\equiv We \ take \ this \ to \ be \ the \ Evaporation \ time\\ t_{i} &\equiv Formation \ time\\ u_{i} &\equiv Energy \ density\\ M_{Pl} &\equiv \ Reduced \ plank \ mass\\ M_{Pl} &\equiv \ 1.220 \ 890 \ x \ 10^{19} \ GeV \end{split}$$

$$\chi \phi$$
, m _{χ/ϕ} = 0 $\chi \phi$ SM

