

Signatures of nuclear structure changes in neutron-rich Cr isotopes

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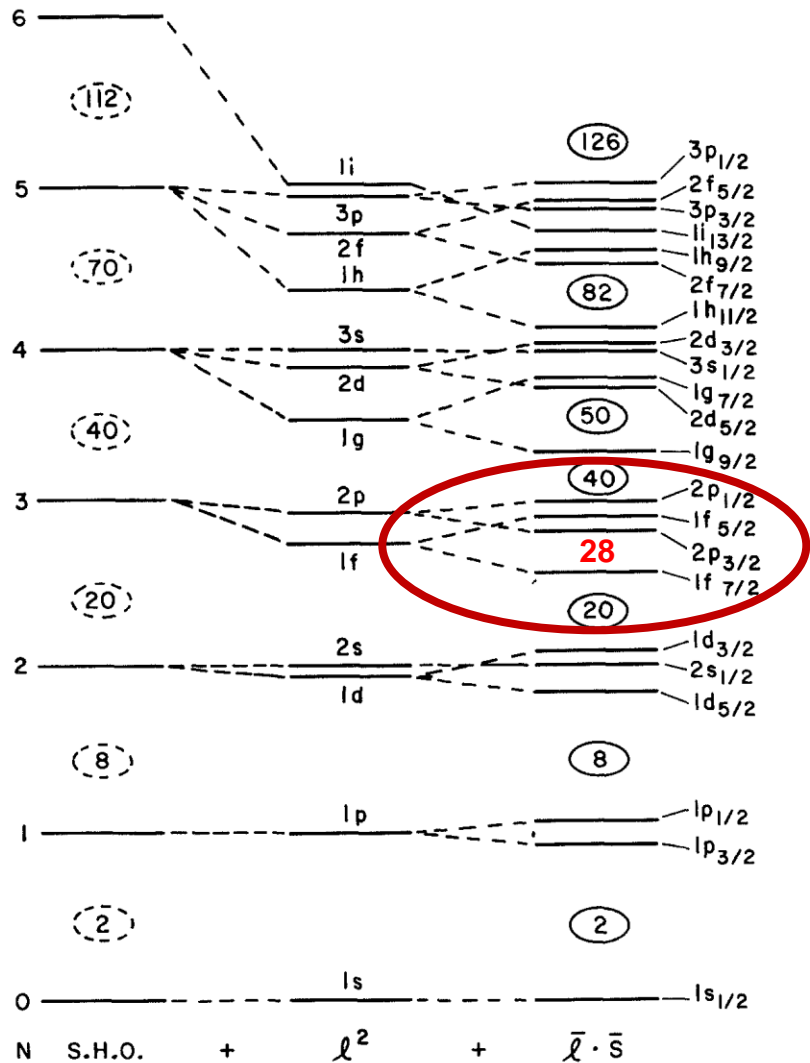
Prof. dr. G. Neyens, KU Leuven

For the CRIS collaboration

LISA Conference 2024



Introduction to the nuclear shell model



fp-shell

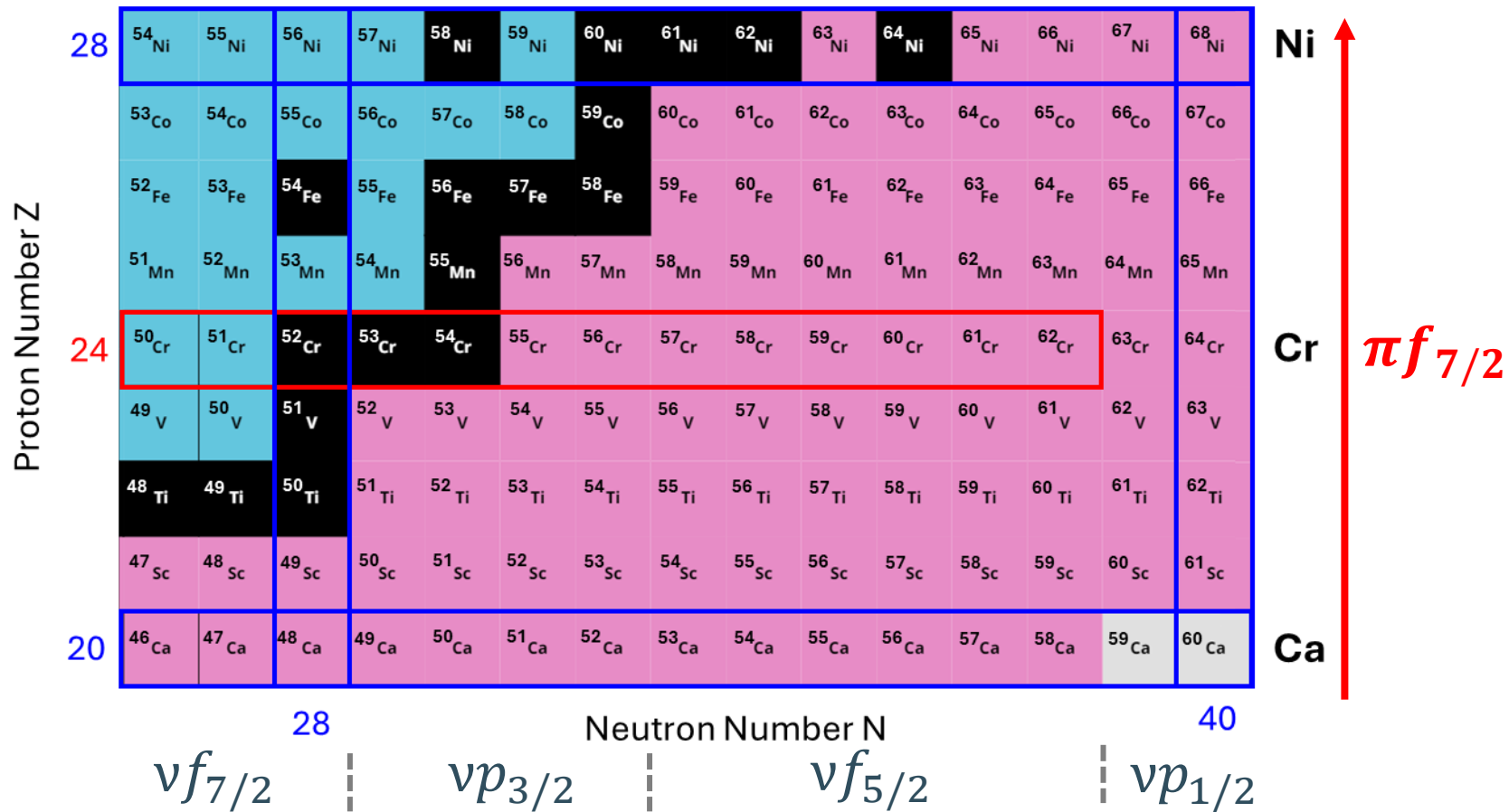
$^{51-61}\text{Cr}$

$Z = 24$

$N = 27 - 37$

All valence neutrons
lie in fp-shell

Cr between the Ni and Ca isotopic chains

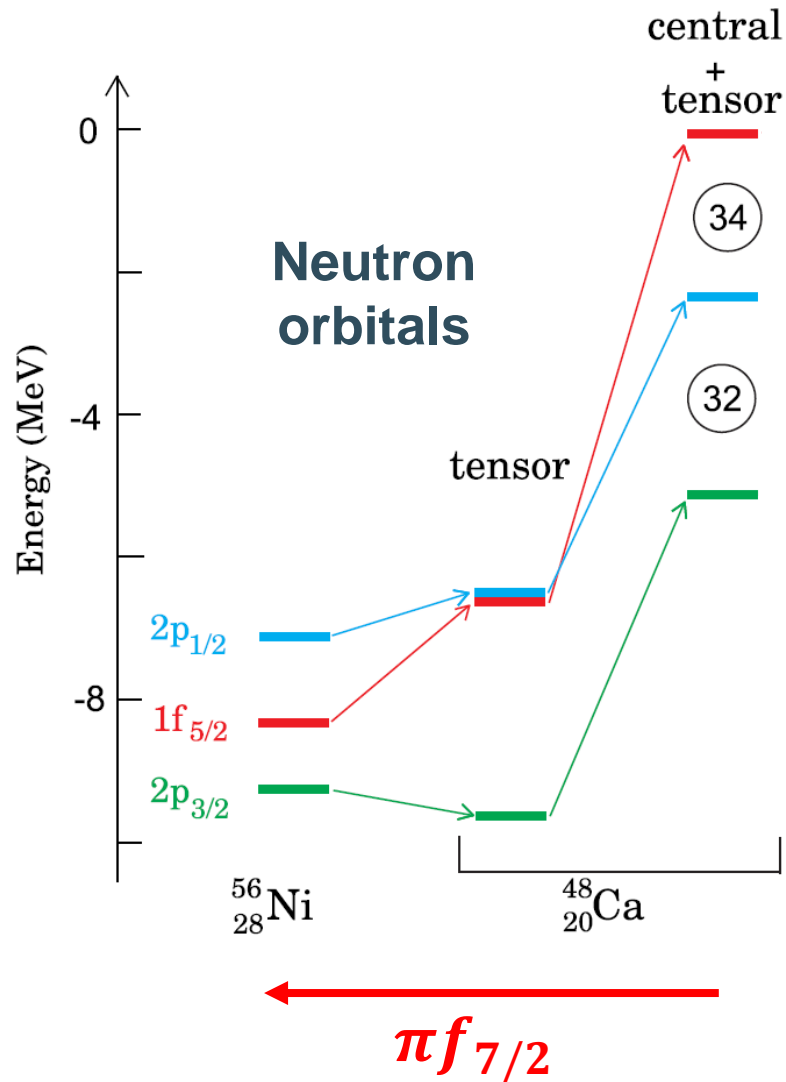


Studied $^{50-62}\text{Cr}$

Cr (Z=24) inbetween
Ca (Z=20) and Ni (Z=28)

Middle of 2 magic proton
shell closures

Shell structure evolution from Ca to Ni



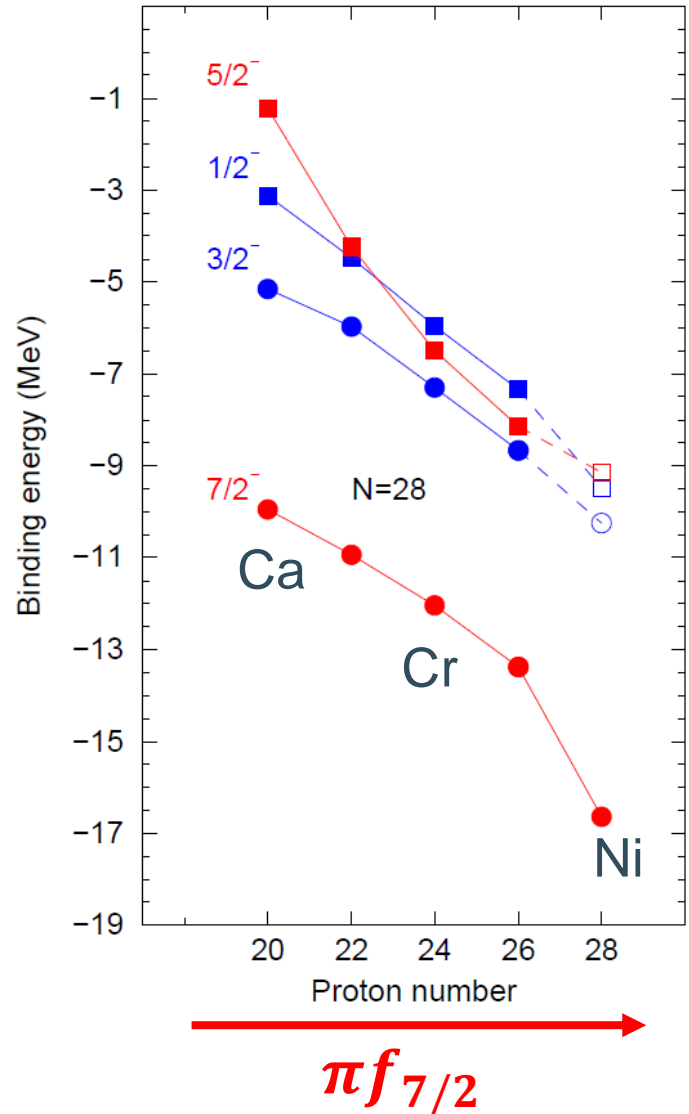
From **theoretical calculations**

Clear crossing of the $\nu f_{5/2}$ and $\nu p_{1/2}$ orbitals

New subshell gap formation at $N=32, 34$ in Ca

T. Otsuka et al. 2020

Shell structure evolution from Ca to Ni



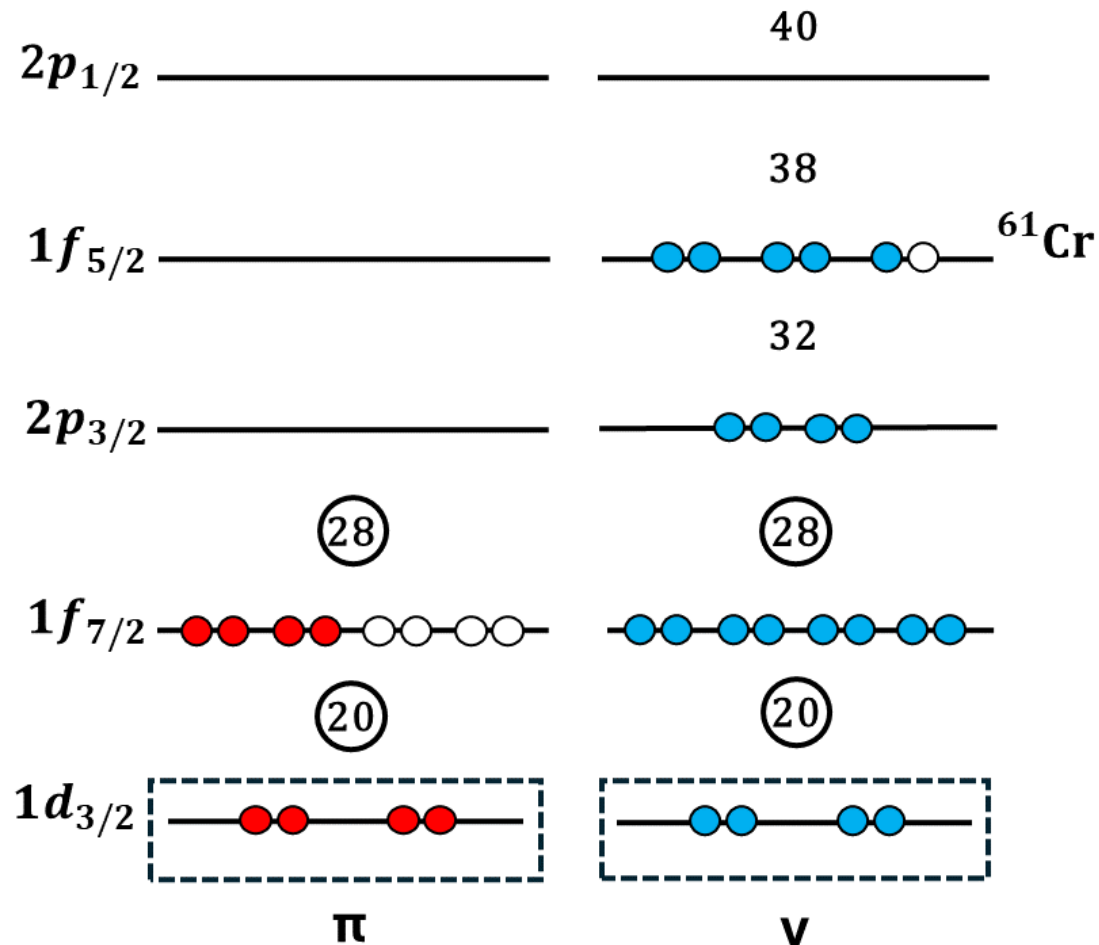
From **experimental results**

Clear crossing of the $\nu f_{5/2}$ and $\nu p_{1/2}$ orbitals

New subshell gap formation at $N=32, 34$ in Ca

O. Sorlin et al. 2008

Nuclear shell structure in Cr



Unnuanced orbital filling and single-particle behaviour

Literature

- $51 \rightarrow 7/2$
 - $53 \rightarrow 3/2$
 - $55 \rightarrow 3/2$
 - $57 \rightarrow 5/2$
 - $59 \rightarrow 5/2$
 - $61 \rightarrow 5/2$
-
- $51 \rightarrow 7/2^-$ measured
 - $53 \rightarrow 3/2^-$ measured
 - $55 \rightarrow 3/2^-$ measured
 - $57 \rightarrow 3/2^-$ tentative
 - $59 \rightarrow 1/2^-$ tentative
 - $61 \rightarrow 5/2^-$ tentative

Observables to investigate

Nuclear ground State Spin I

even-even $\rightarrow 0^+$

even-odd \rightarrow depends on
specific configuration of
neutrons

Nuclear magnetic dipole moment μ

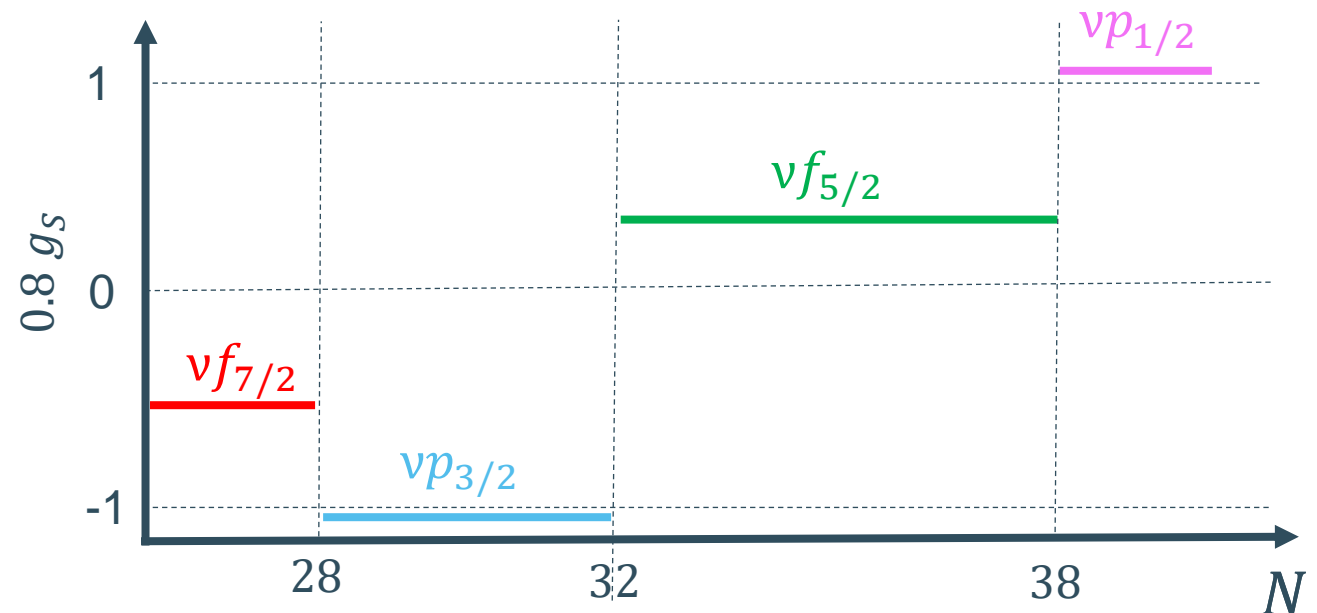
$$\mu = gI\mu_N.$$

Interaction of nucleus with magnetic fields

g-factor sensitive to specific configuration

RMS Charge Radius $\langle r_c^2 \rangle$

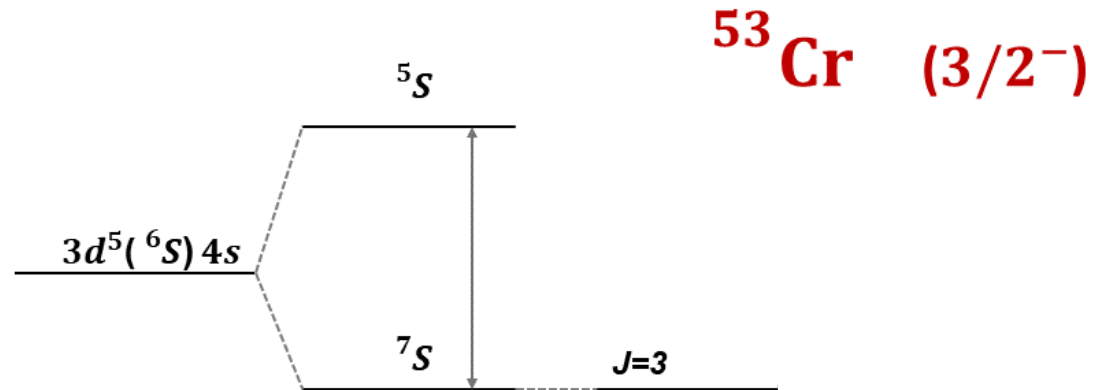
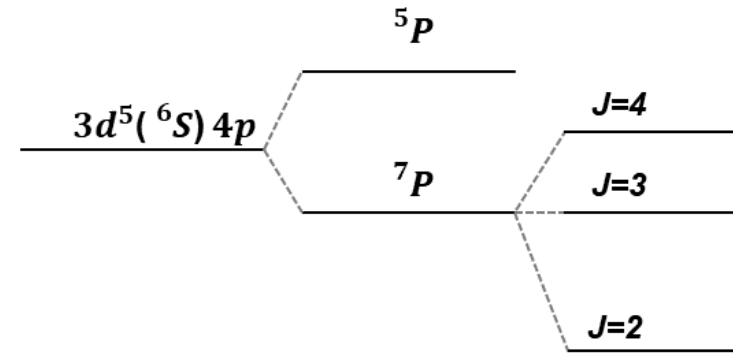
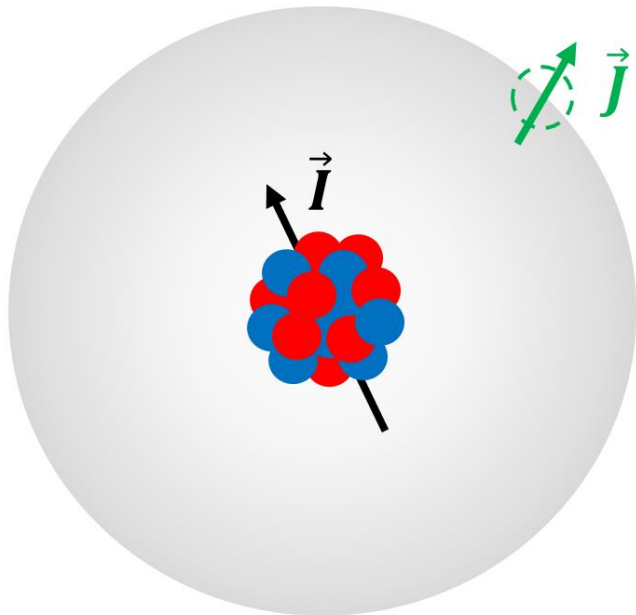
Further interpretation necessary



Experimental Method

Hyperfine Structure (HFS)

Interaction of the nuclear charge and current distribution with the electromagnetic fields of atomic electrons



$^{53}\text{Cr} (3/2^-)$

Configuration	Term $2S+1L$	State $\vec{J} = \vec{L} + \vec{S}$
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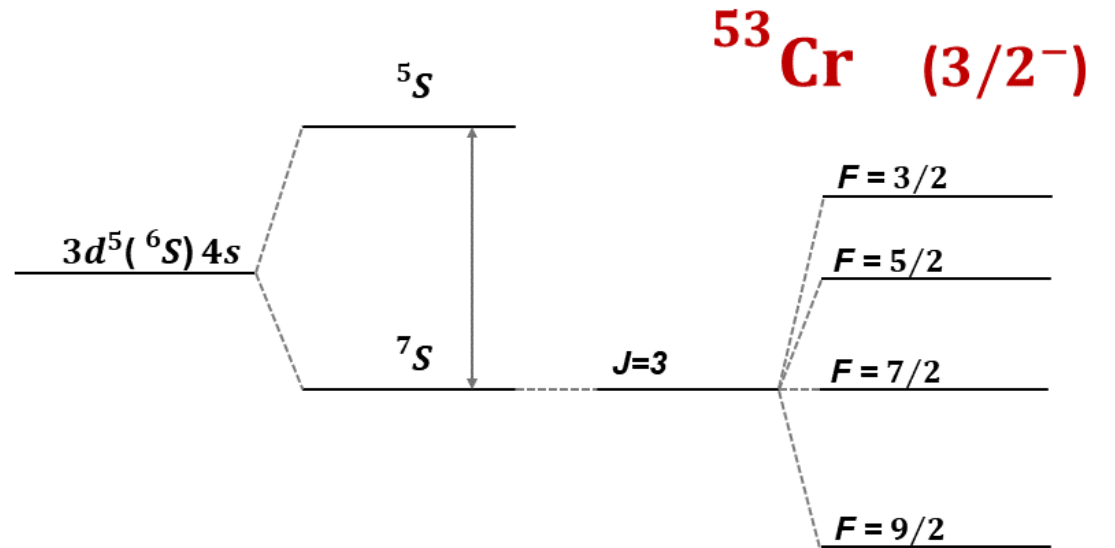
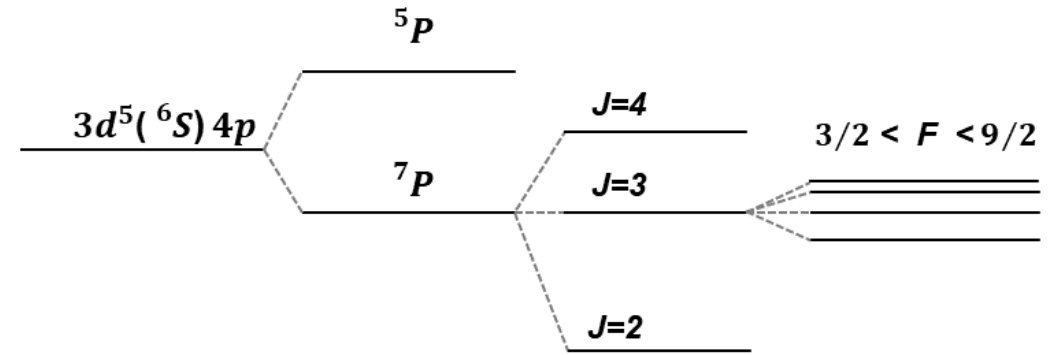
Hyperfine Structure (HFS)

Nuclear Spin

$$\vec{F} = \vec{I} + \vec{J}$$

Number of levels dictated by coupling

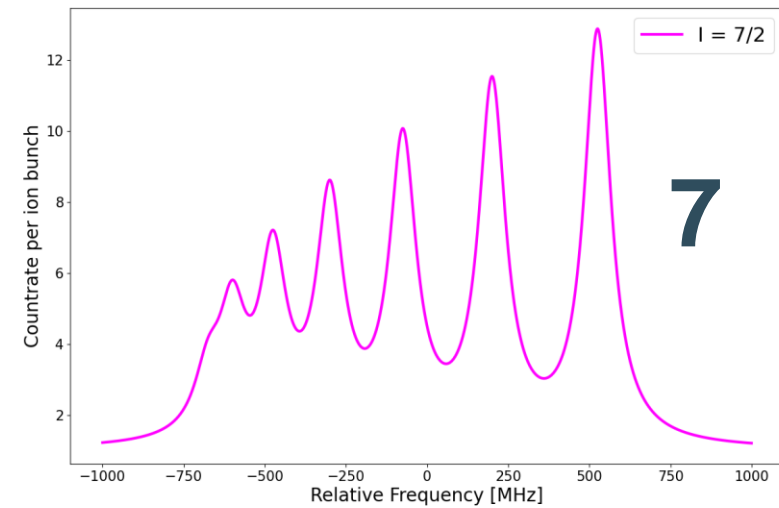
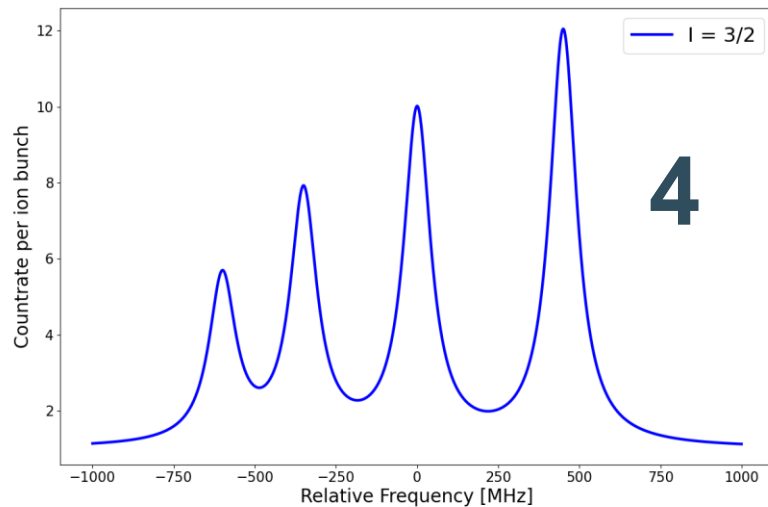
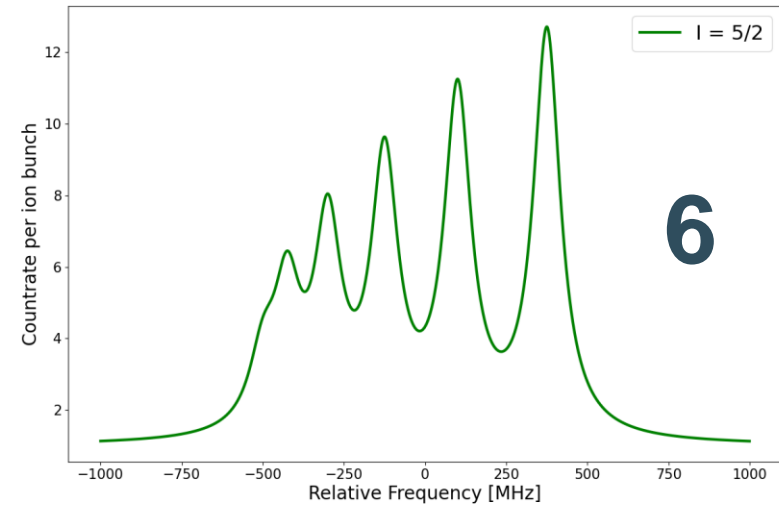
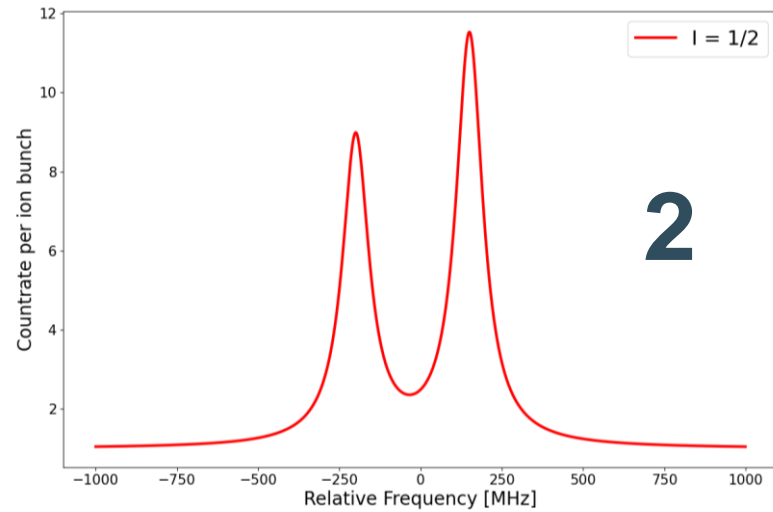
$$|I - J| \leq F \leq I + J$$



^{53}Cr ($3/2^-$)

Configuration	Term $2S+1L$	State $\vec{J} = \vec{L} + \vec{S}$	HFS $\vec{F} = \vec{I} + \vec{J}$
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Hyperfine Structure (HFS)



Hyperfine Structure (HFS)

Magnetic Dipole Moment

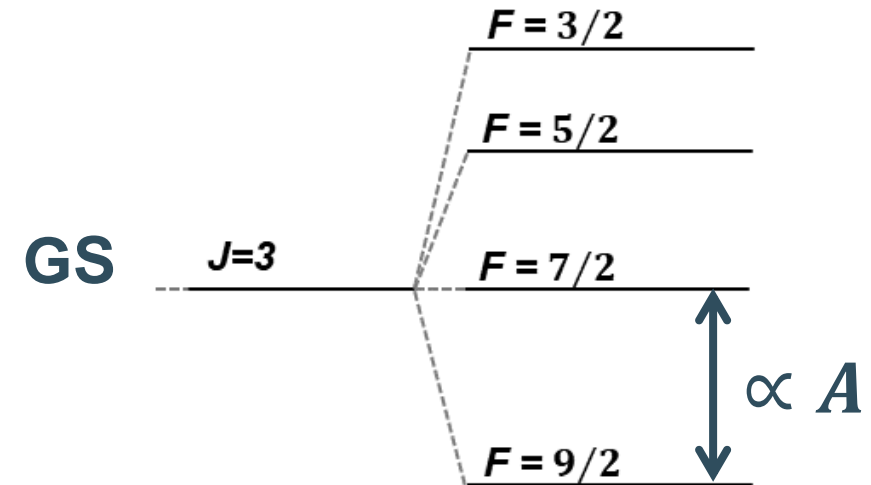
Interaction with magnetic field of electrons at center of nucleus B_0

$$U = -\vec{\mu} \cdot \vec{B}.$$

Splitting proportional to HF A parameter
 → Proportional to g-factor

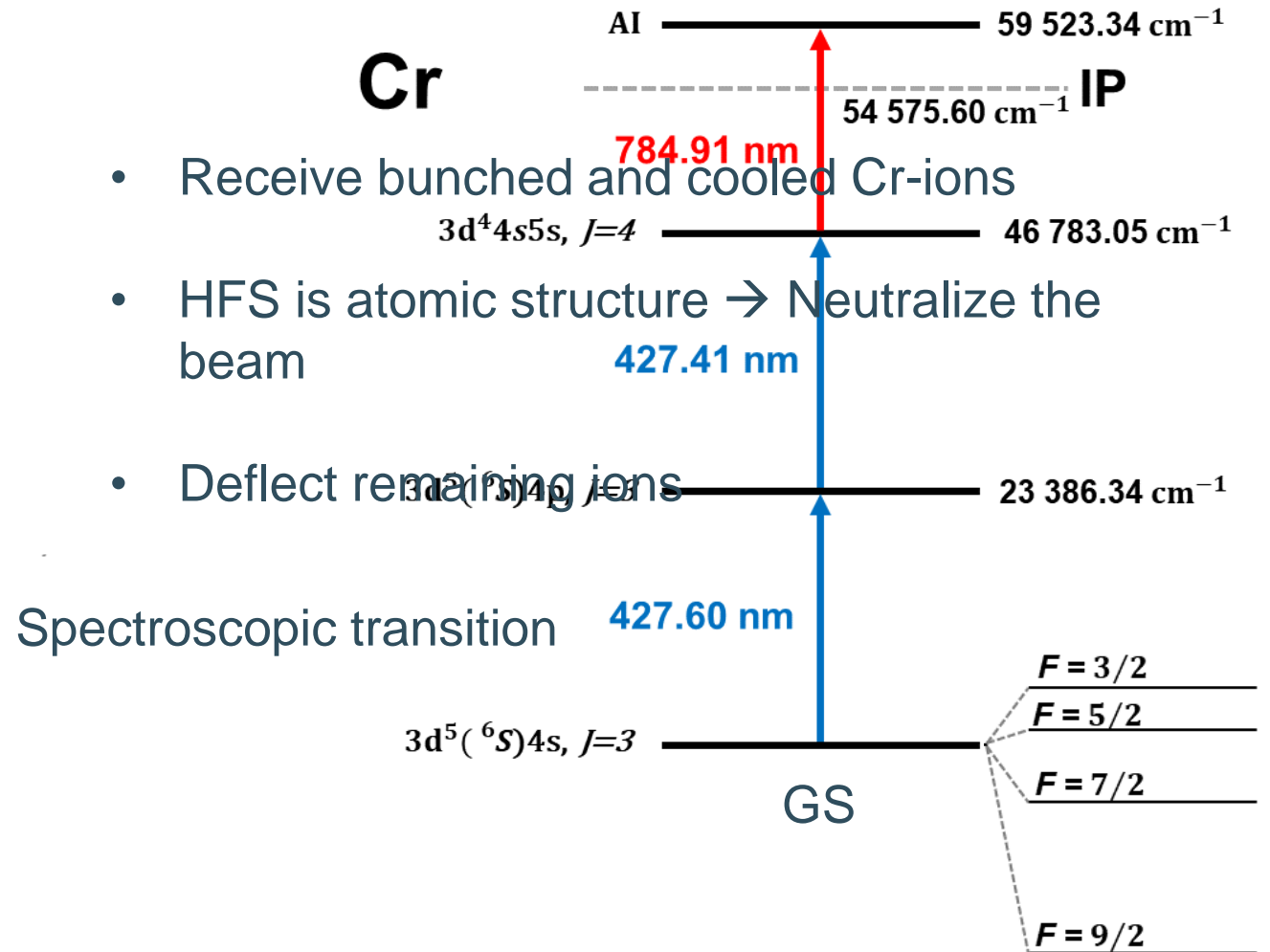
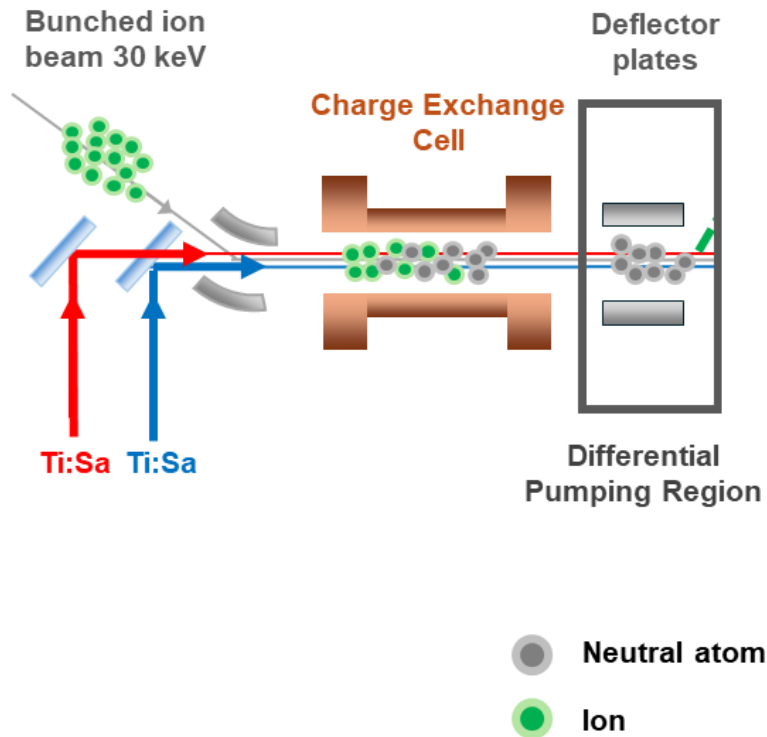
$$\Delta E_{M1}(F) = -\frac{AC}{2} \quad A = \frac{\mu B_0}{\hbar^2 I J}$$

⁵³Cr (3/2⁻)

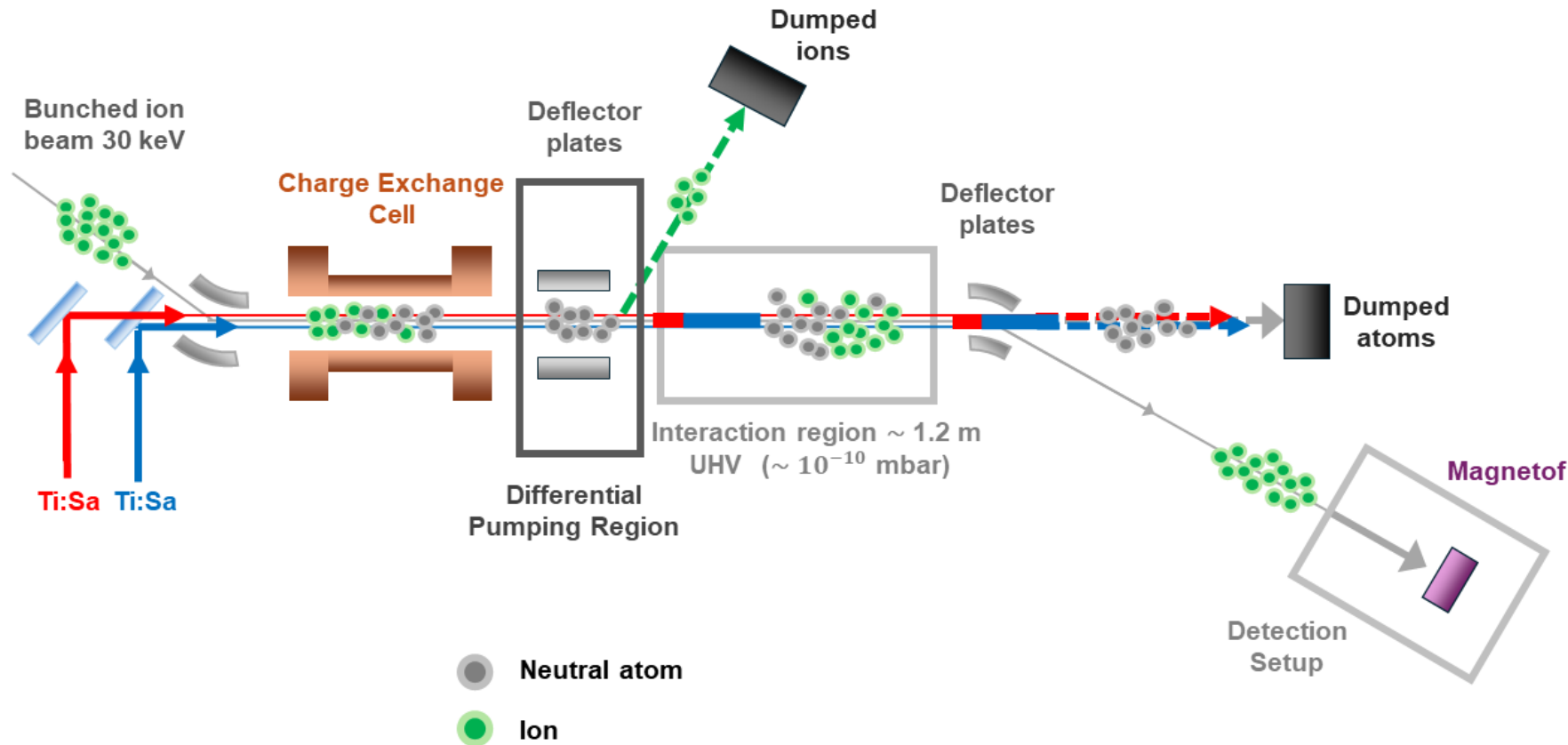


State	HFS
$\vec{J} = \vec{L} + \vec{S}$	$\vec{F} = \vec{I} + \vec{J}$

Collinear Resonance Ionization Spectroscopy (CRIS)

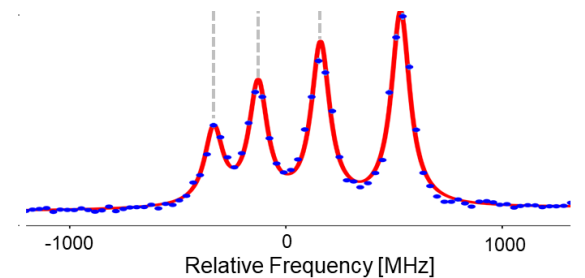


Collinear Resonance Ionization Spectroscopy (CRIS)

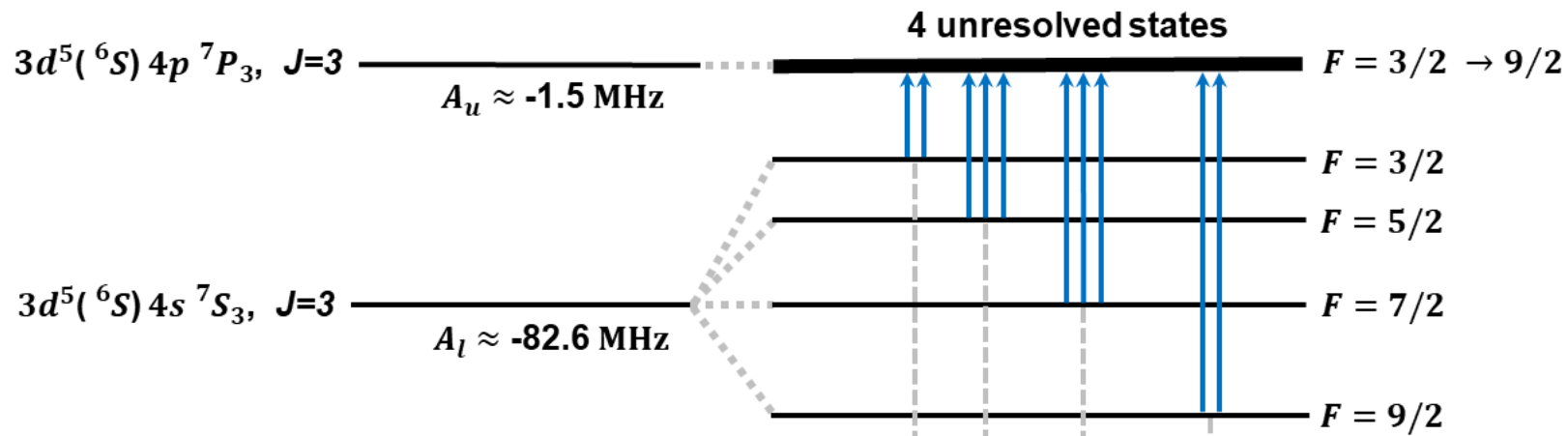


Ions detected when spectroscopic step on resonance

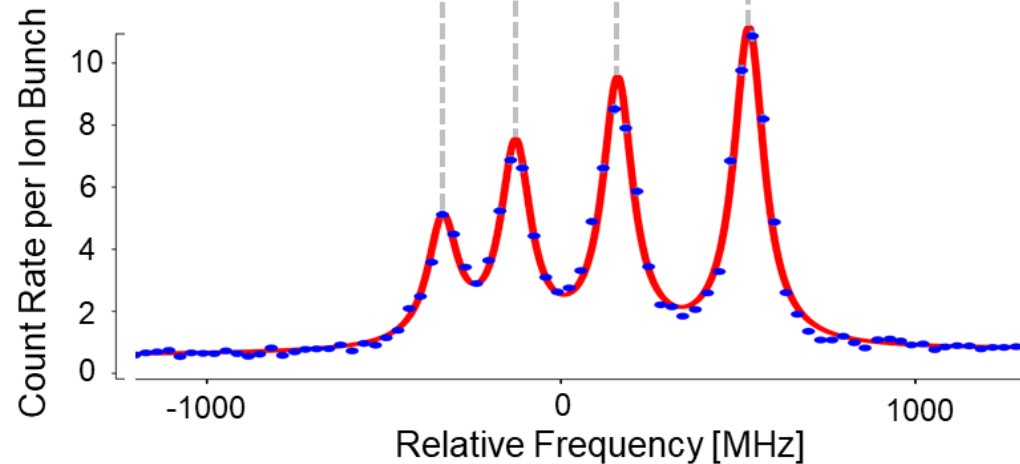
Resolve HFS



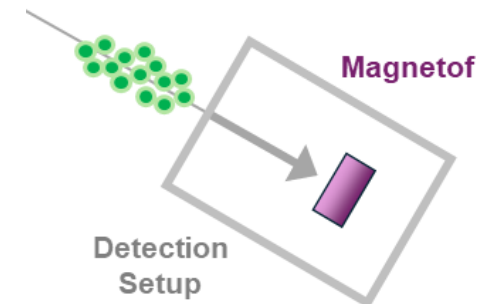
Collinear Resonance Ionization Spectroscopy (CRIS)



^{53}Cr
 $(3/2^-)$



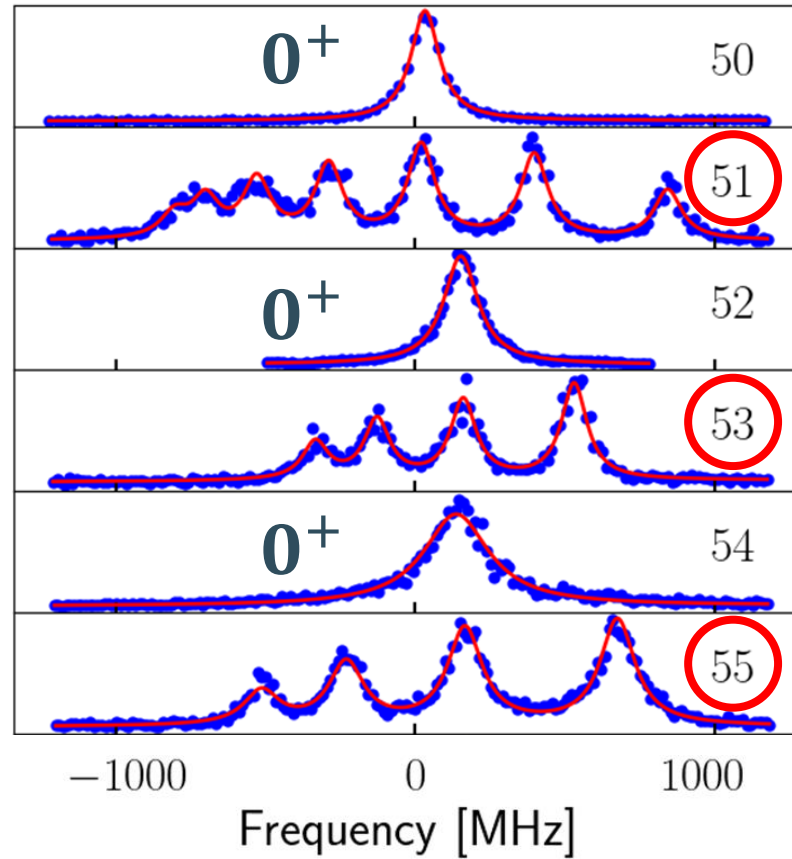
Resolve the HFS of the ground state



Ground state spins

51,53,55Cr

All behave as expected, both according to orbital filling and compared to literature



I Lit

I CRIS

$7/2^-$

$7/2^-$

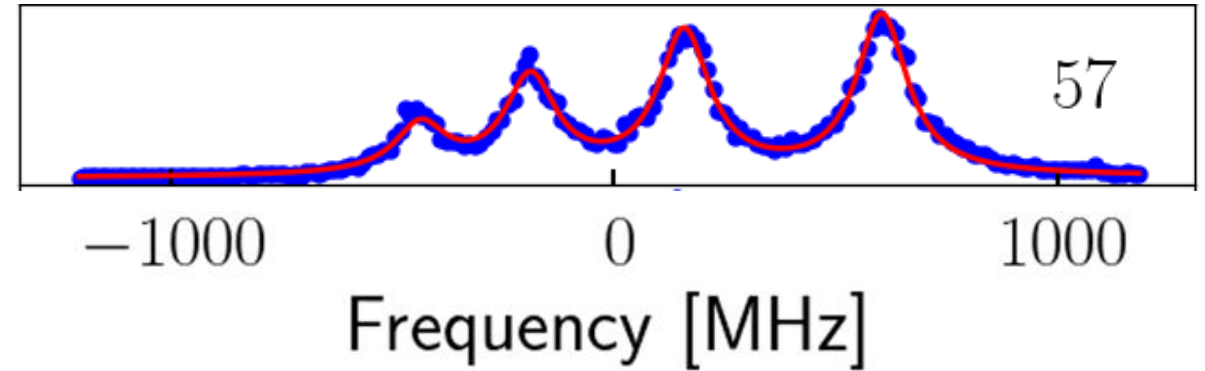
$3/2^-$

$3/2^-$

$3/2^-$

$3/2^-$

^{57}Cr $N = 33$



From 57 onwards expect to fill the $\nu f_{5/2}$ orbital and therefore have a spin 5/2

Not the case, CRIS confirms literature 3/2 spin

I Lit

I CRIS

$(3/2)^-$

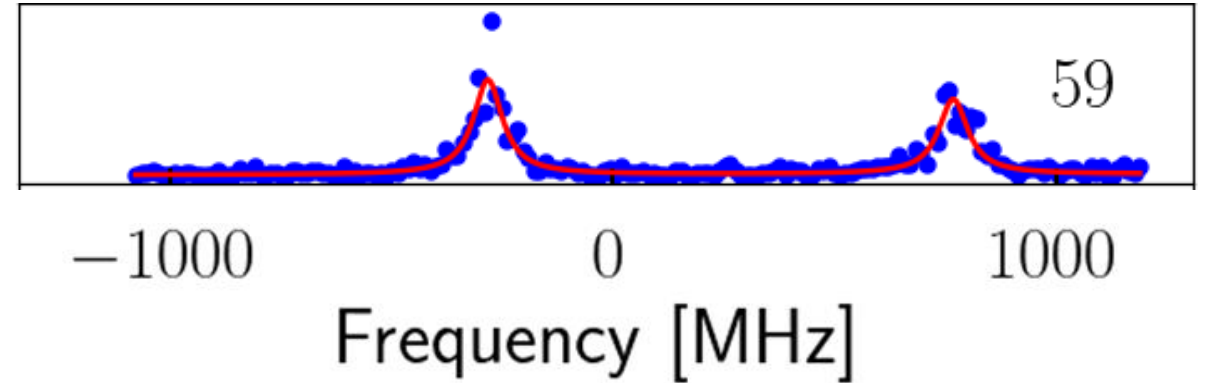
$3/2^-$

^{59}Cr $N = 35$

Naively expect 3 neutrons in $\nu f_{5/2}$ orbital

Correctly predicted spins by shell model calculations GXPF1A and KB3G

$$^{59}\text{Cr} \quad 1/2^- \quad (1f_{7/2})^8, (2p_{3/2})^4, (1f_{5/2})^2, (1p_{1/2})^1$$



I Lit

I CRIS

$(1/2)^-$

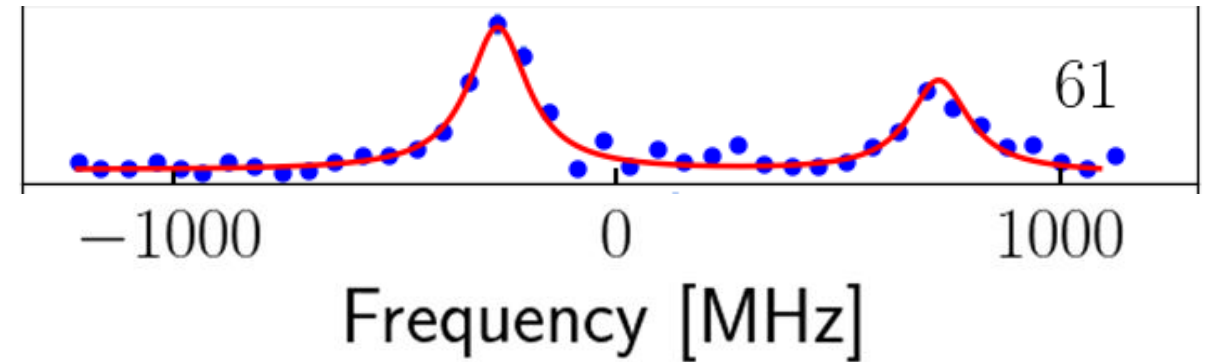
$1/2^-$

^{61}Cr $N = 37$

Spin 5/2 most likely configuration according to three previous publications

H.L. Crawford et al. 2009

All use β - and γ -decay spectroscopy



I Lit

I CRIS

$(5/2)^-$

$1/2^-$

^{61}Cr $N = 37$

H.L. Crawford et al. 2009

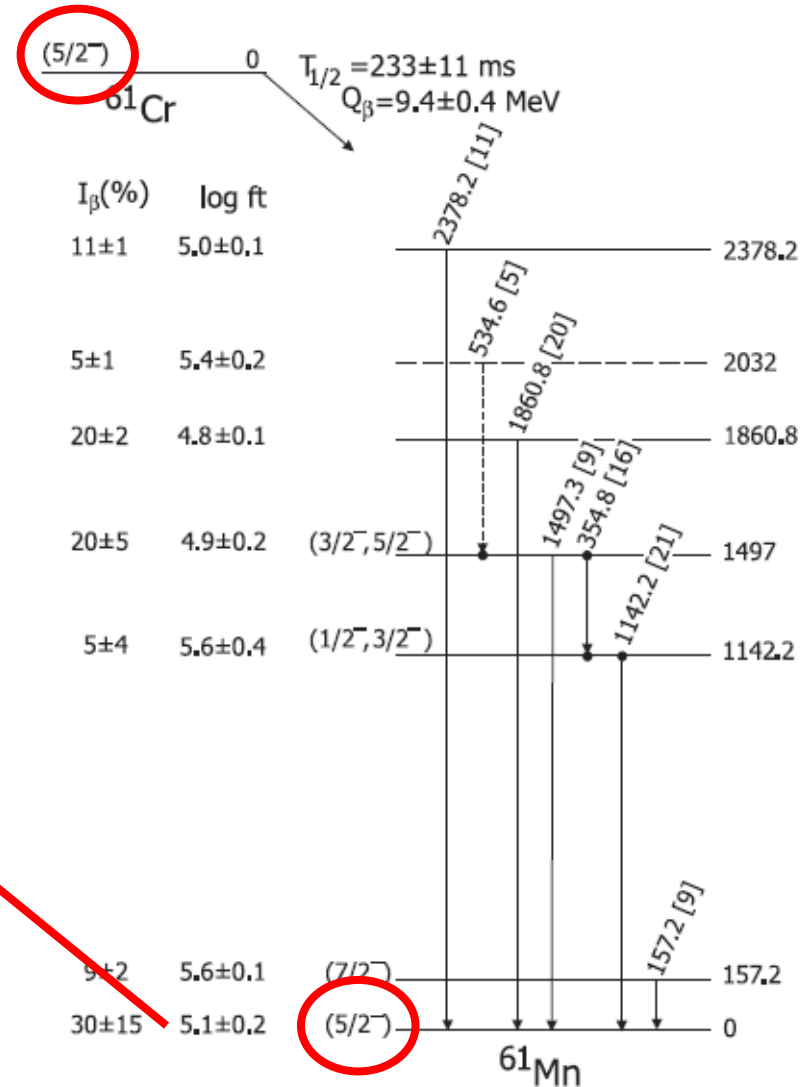
Both ^{61}Mn and ^{61}Cr ground state spins only tentatively assigned

$5/2^-$ spin of ^{61}Mn confirmed at COLLAPS C. Babcock et al. 2015

$\log ft = 5.1(2) \rightarrow$ (super)allowed transition

NOT supporting $\nu p_{1/2} \rightarrow \pi f_{7/2}$

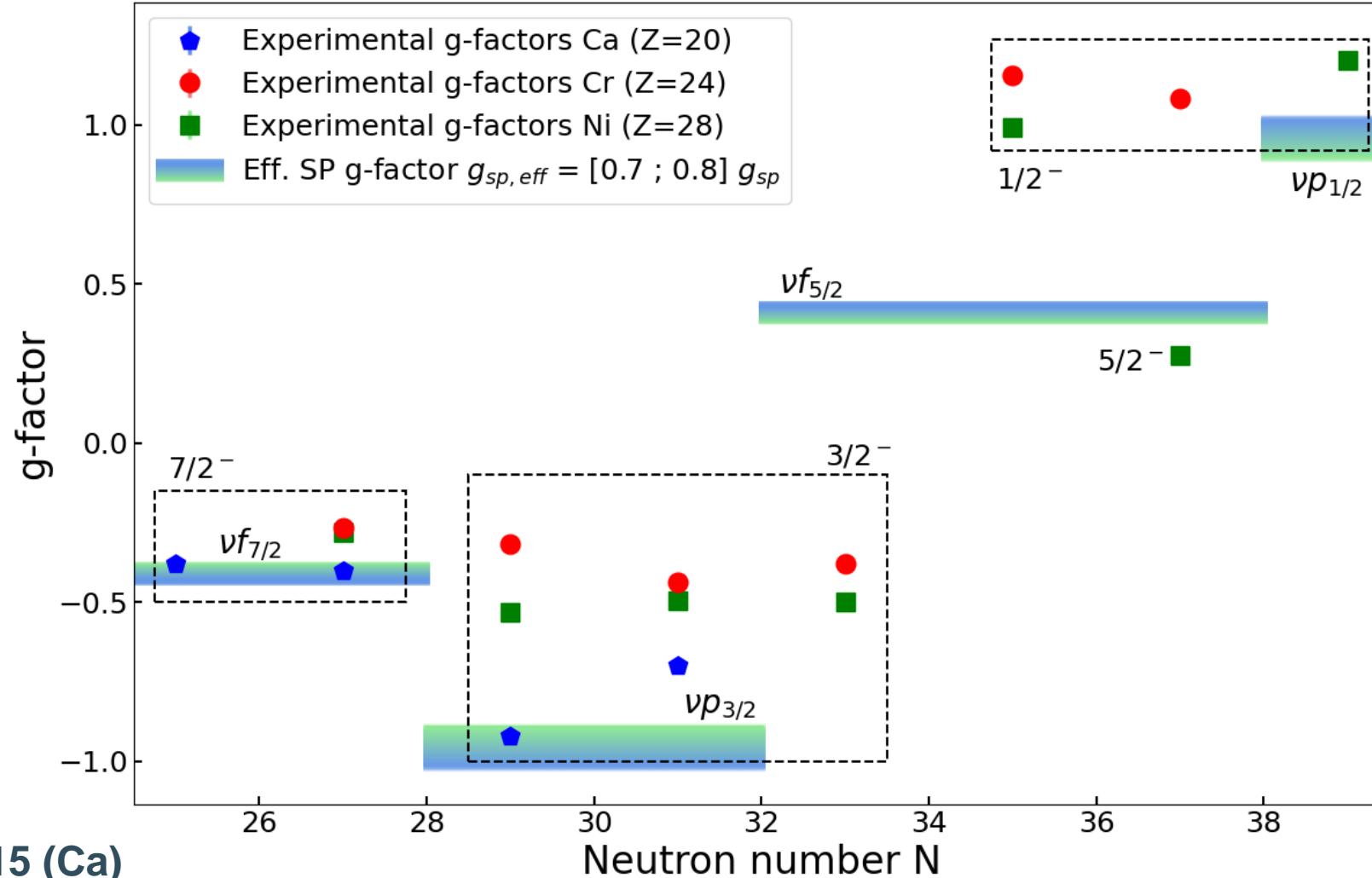
\rightarrow **Pandemonium effect**
overestimating transition strengths



g-factors

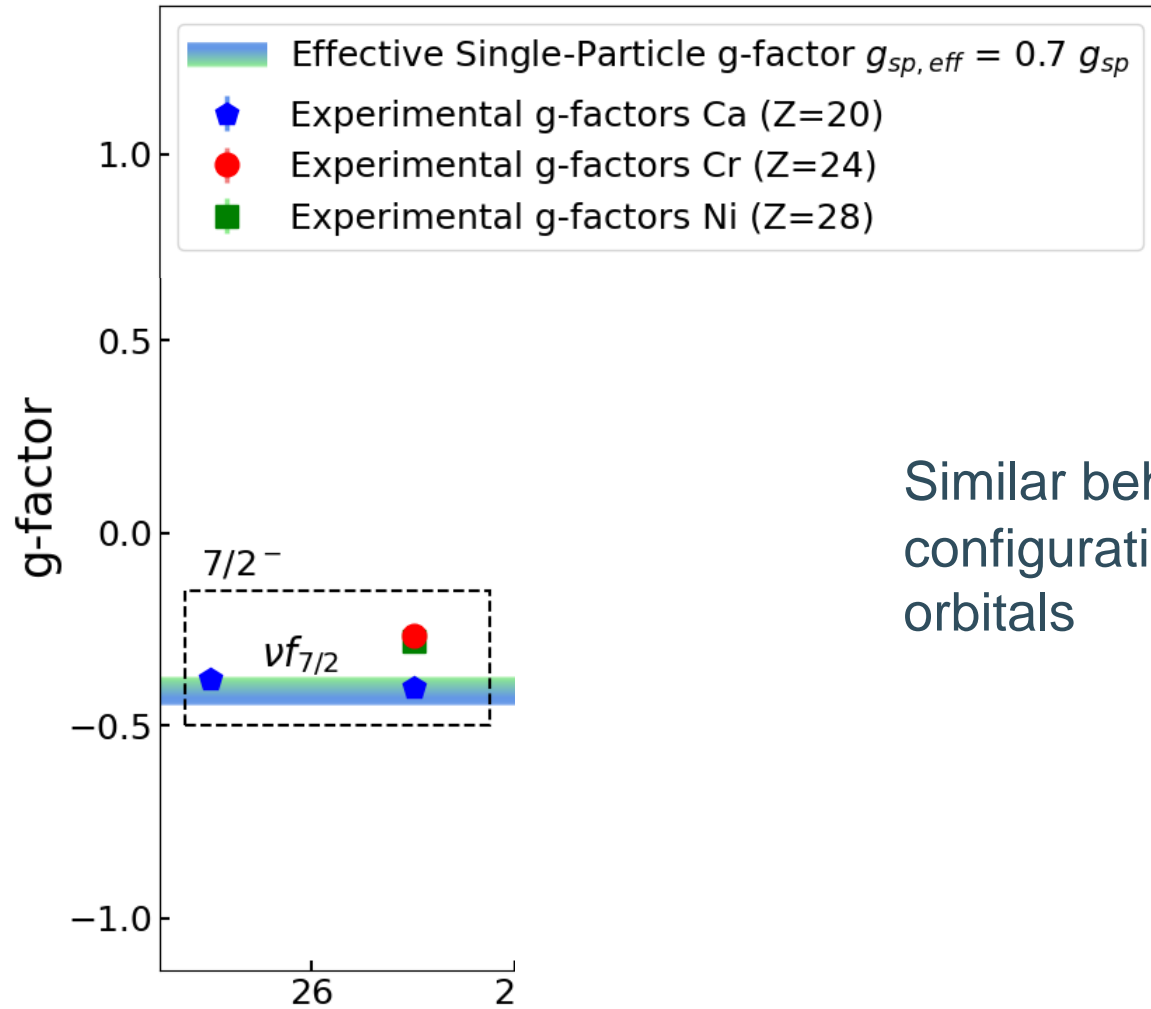
$$g = \frac{\mu}{I}$$

g-factors



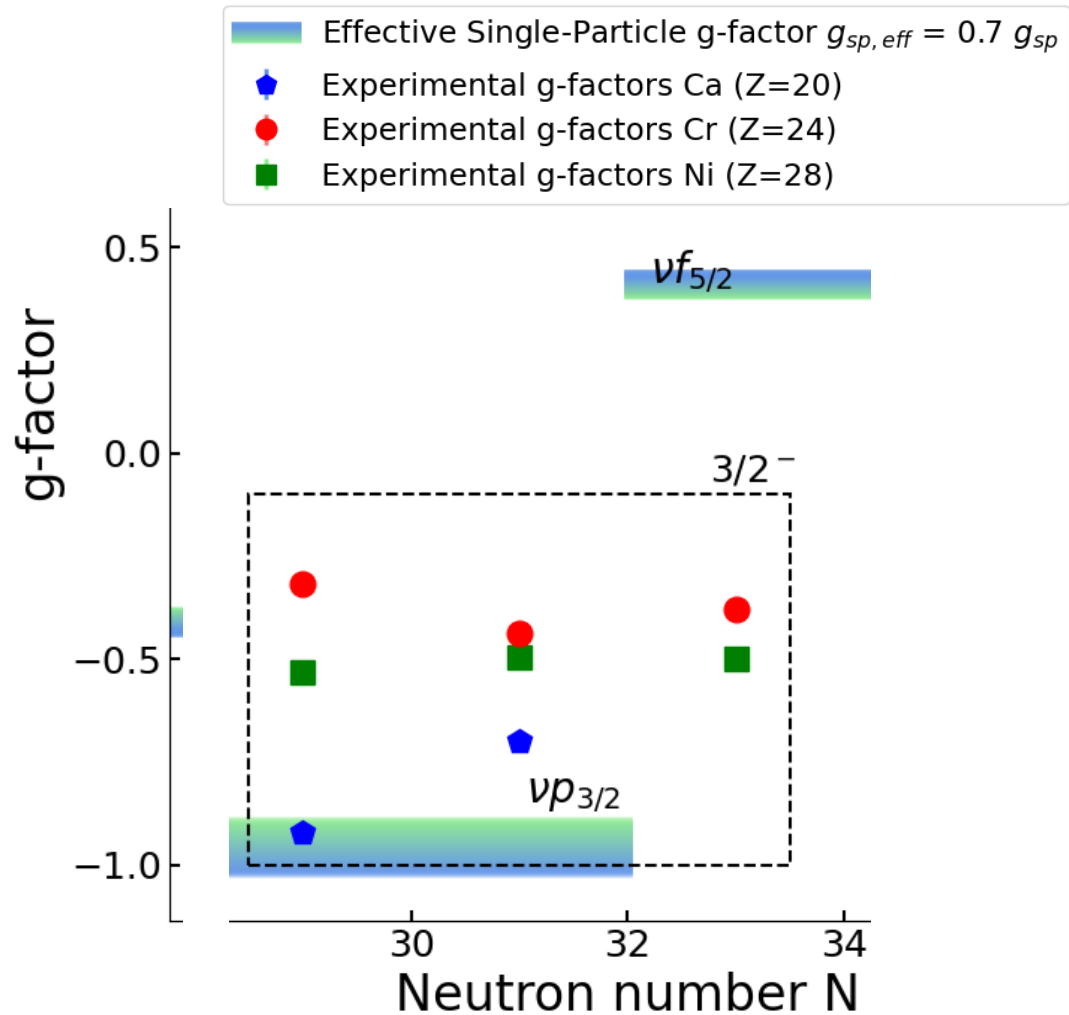
R.F. Ruiz et al. 2015 (Ca)
 P. Müller et al. 2024 (Ni)

^{51}Cr



Similar behaviour to ^{55}Ni , slight configuration mixing with higher neutron orbitals

53,55,57Cr



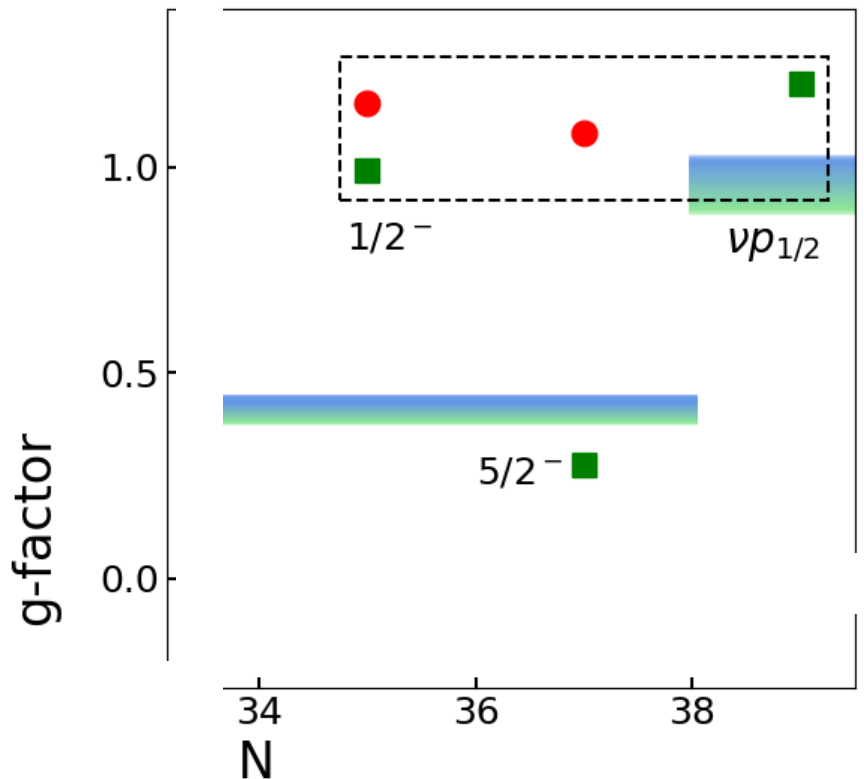
Far from single-particle g-factors

Large amounts of neutron and/or proton excitations, as observed in the Ni isotopes, can be expected.

No solid conclusion can be drawn without further investigation

$^{59,61}\text{Cr}$

- ◆ Experimental g-factors Ca (Z=20)
- Experimental g-factors Cr (Z=24)
- Experimental g-factors Ni (Z=28)



Highly single-particle like behaviour from experimental g-factor and shell model wave function calculations

→ Opposes expected large deformation in this area

Isotope	I^π	Wave function (ν)	GXPF1A Probability	KB3G Probability
^{59}Cr	$1/2^-$	$(1f_{5/2})^2, (1p_{1/2})^1$	0.63	0.60
^{61}Cr	$1/2^-$	$(1f_{5/2})^4, (1p_{1/2})^1$	0.76	0.75

S. Suchyta et al. 2014

Conclusion and Outlook

- Unequivocal spin measurements of $^{51-61}\text{Cr}$
 - Spin $1/2^-$ of ^{61}Cr instead of $5/2^-$ from previous measurements.
 - Reinterpretation and remeasurement of β - and γ -decay spectroscopy necessary
 - Dipole moments of $^{51-61}\text{Cr}$ measured and compared with corresponding Ca and Ni dipole moments.
-
- Compare dipole moments and spins to calculations, shell model/ab initio
 - New laser scheme to allow measurement of quadrupole moments
 - Interpret the measured charge radii of $^{50-62}\text{Cr}$

Thank you for listening

^{61}Cr as a Doorway to the $N = 40$ Island of Inversion

L. Lalanne,^{1,2,3,*} M. Athanasakis-Kaklamanakis,^{1,2} D.D. Dao,³ Á. Koszorús,¹ Y. C. Liu,⁴ R. Mancheva,^{2,1} F. Nowacki,³ J. Reilly,⁵ C. Bernerd,² K. Chrysalidis,² T. E. Cocolios,¹ R. P. de Groote,¹ K. T. Flanagan,⁵ R. F. Garcia Ruiz,⁶ D. Hanstorp,⁷ R. Heinke,¹ M. Heines,¹ P. Lassegues,¹ K. Mack,⁵ B. A. Marsh,² A. McGlone,⁵ K. M. Lynch,⁵ G. Neyens,¹ B. van den Borne,¹ R. Van Duyse,¹ X. F. Yang,⁴ and J. Wessolek^{5,2}

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⁴*School of Physics and State Key Laboratory of Nuclear Physics and Technology, Peking University, Beijing 100871, China*

⁵*Department of Physics and Astronomy, The University of Manchester, Manchester M13 9PL, United Kingdom*

⁶*Massachusetts Institute of Technology, Cambridge, MA 02139, USA*

⁷*Department of Physics, University of Gothenburg, SE-412 96 Gothenburg, Sweden*

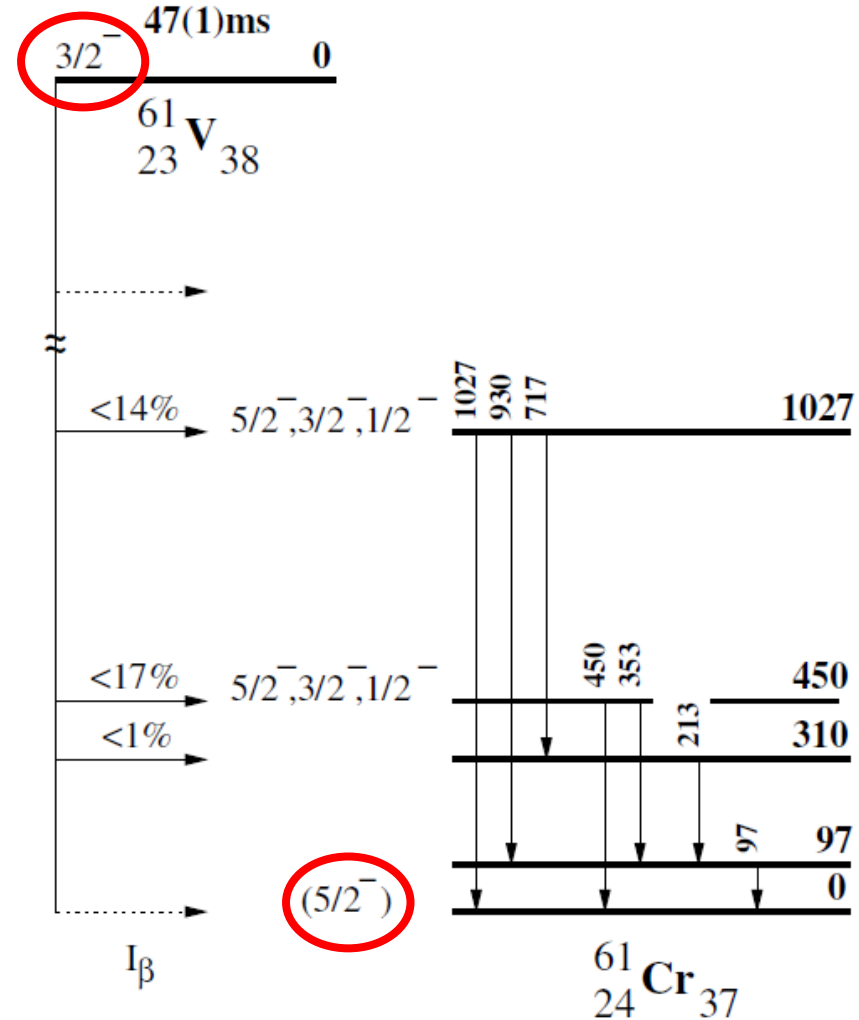
(Dated: July 25, 2024)

^{61}Cr $N = 37$

L. Gaudefroy et al. 2005

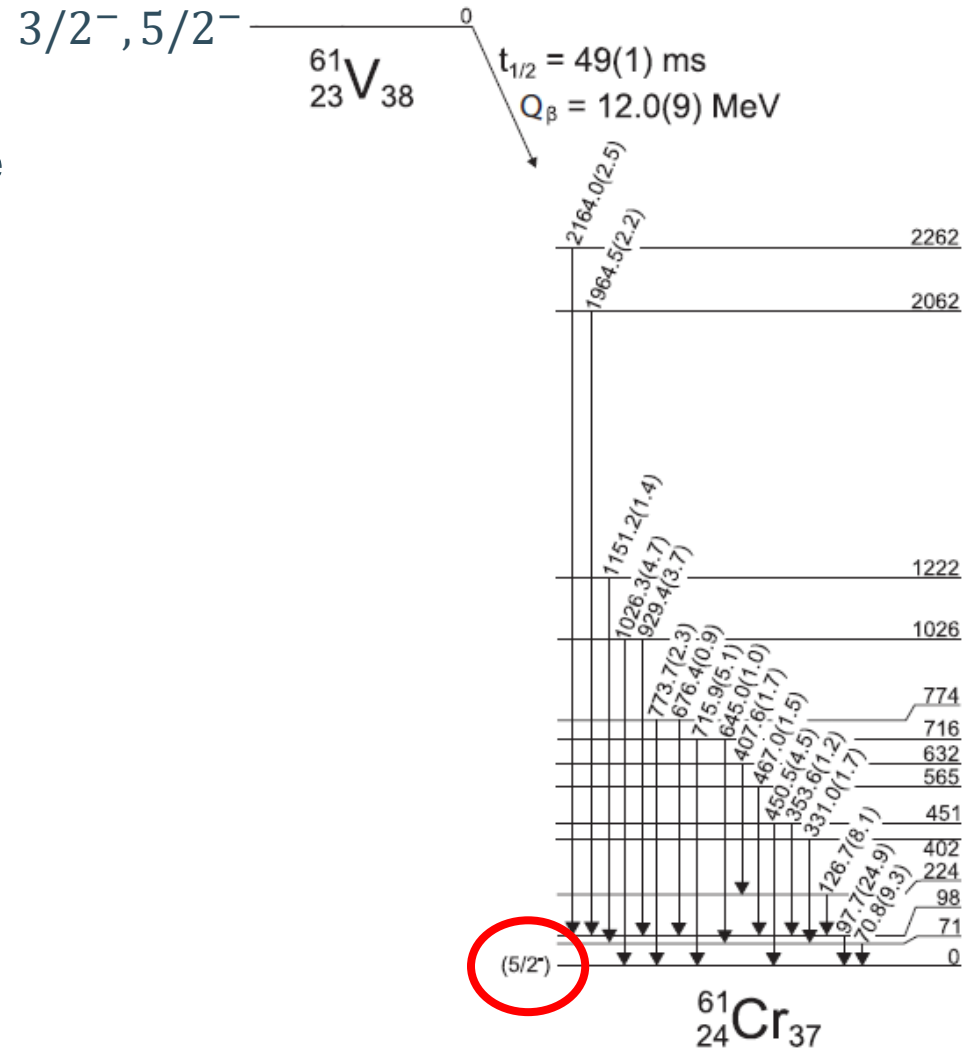
$3/2^-$ spin of $^{61}_{23}\text{V}$ assuming small quadrupole deformation

Concluded $f_{5/2} : 5/2^-$ configuration of ^{61}Cr because GT selection rules favor $\pi f_{7/2} \rightarrow \nu f_{5/2}$ transitions



^{61}Cr $N = 37$

S. Suchyta et al. 2014



$3/2^-$ spin of $^{61}_{23}\text{V}$ most likely for prolate deformation

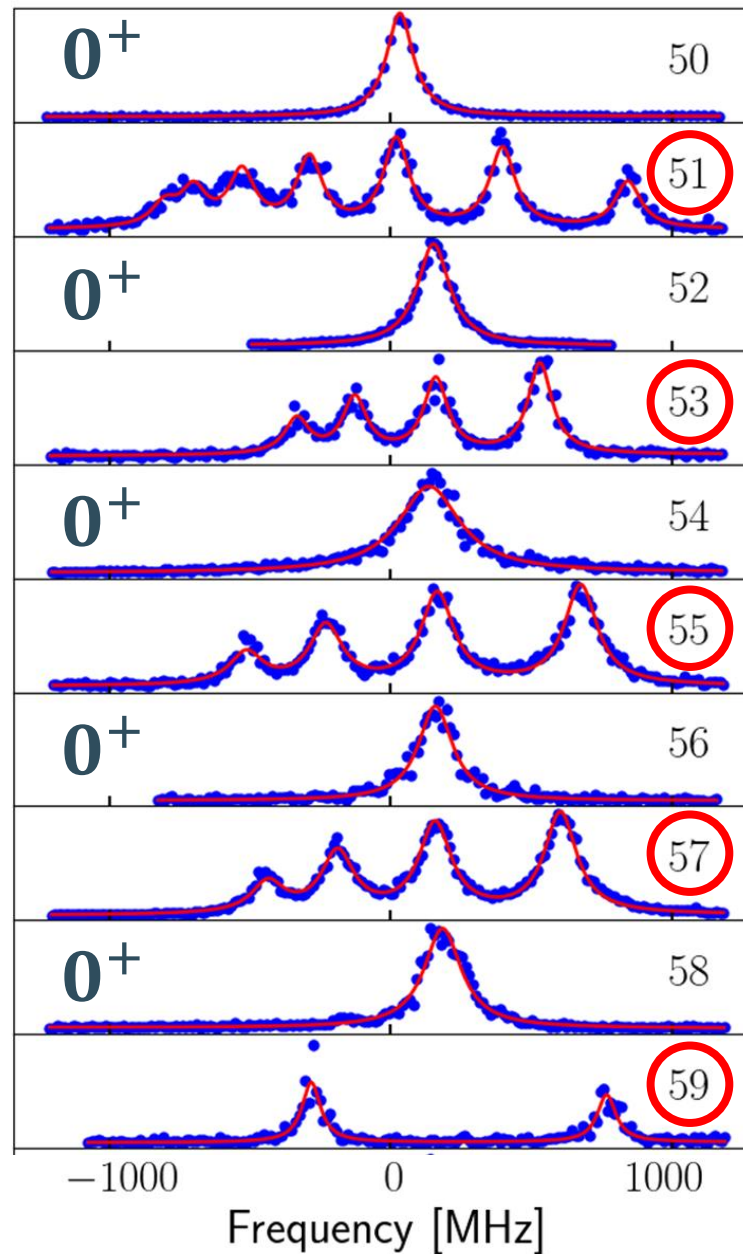
$5/2^-$ spin of $^{61}_{23}\text{V}$ most likely for oblate deformation

S. Suchyta et al. agree with previous spin assignments

But ^{61}Cr $1/2^-$

$^{51-59}\text{Cr}$

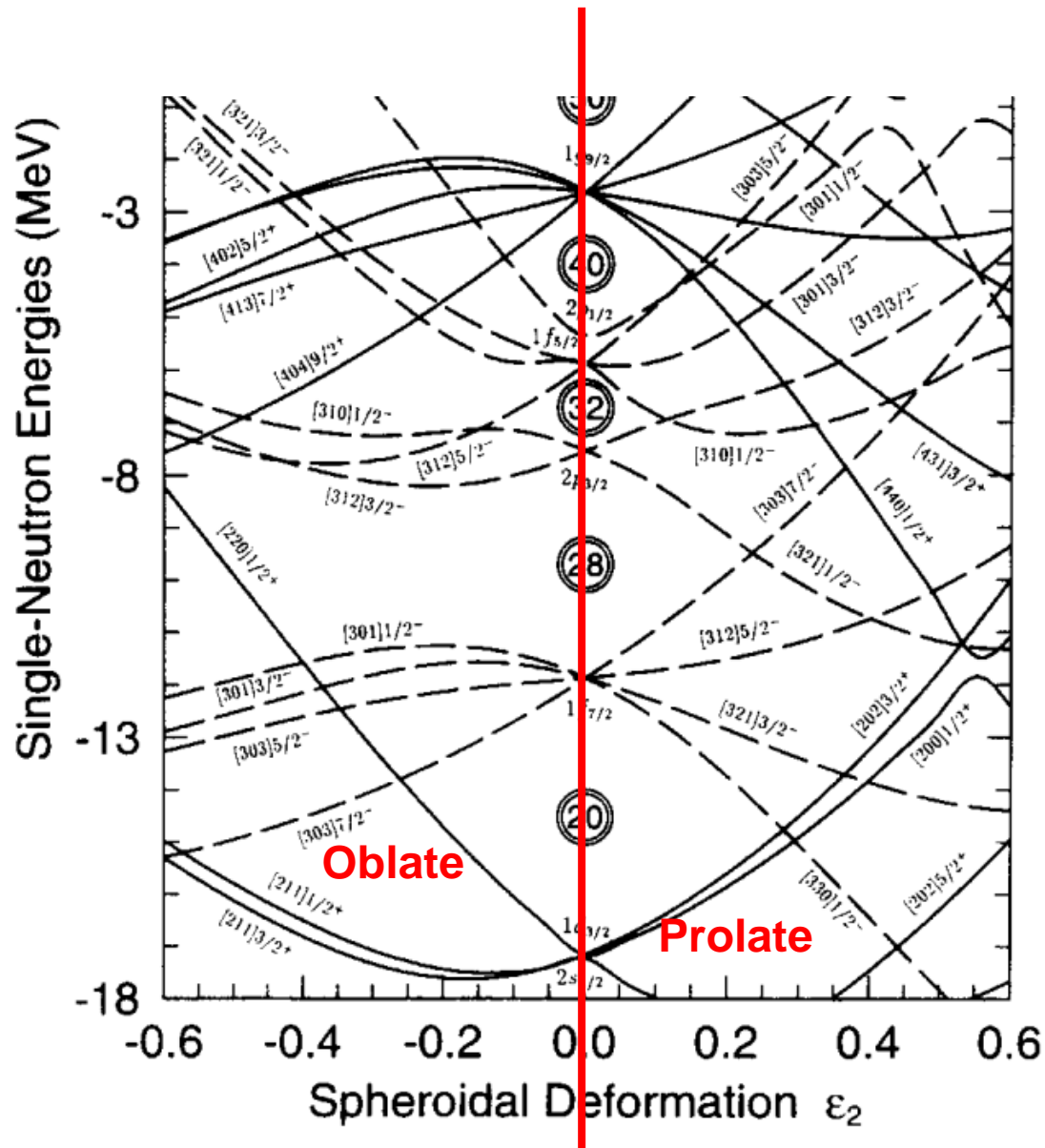
All behave as previously measured/predicted



<i>I</i> Lit	<i>I</i> CRIS
1^+	1^+
$7/2^-$	$7/2^-$
$3/2^-$	$3/2^-$
$3/2^-$	$3/2^-$
$(3/2)^-$	$3/2^-$
$(1/2)^-$	$1/2^-$

Folded Yukawa SP level schemes

O. Sorlin et al. 1998



Literature and CRIS values of I

51, 53, 55 behave as expected, both according to naive orbital filling and compared to literature

55 \rightarrow N=31, clearly after N=32 (filling the $\nu p_{3/2}$ orbital) less SP like behaviour.

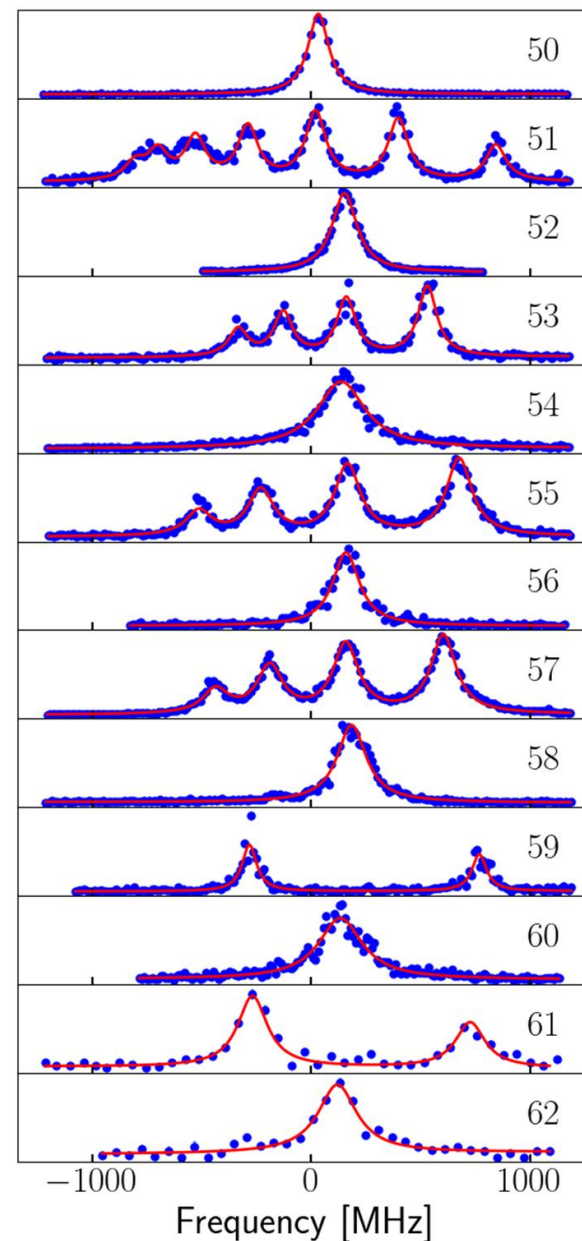
61 = Star of the experiment, wrong spin assignment in multiple published results.

Unnuanced SP suggestions

5/2

5/2

5/2



I Lit I CRIS

$7/2^-$ $7/2^-$

$3/2^-$ $3/2^-$

$3/2^-$ $3/2^-$

$(3/2)^-$ $3/2^-$

$(1/2)^-$ $1/2^-$

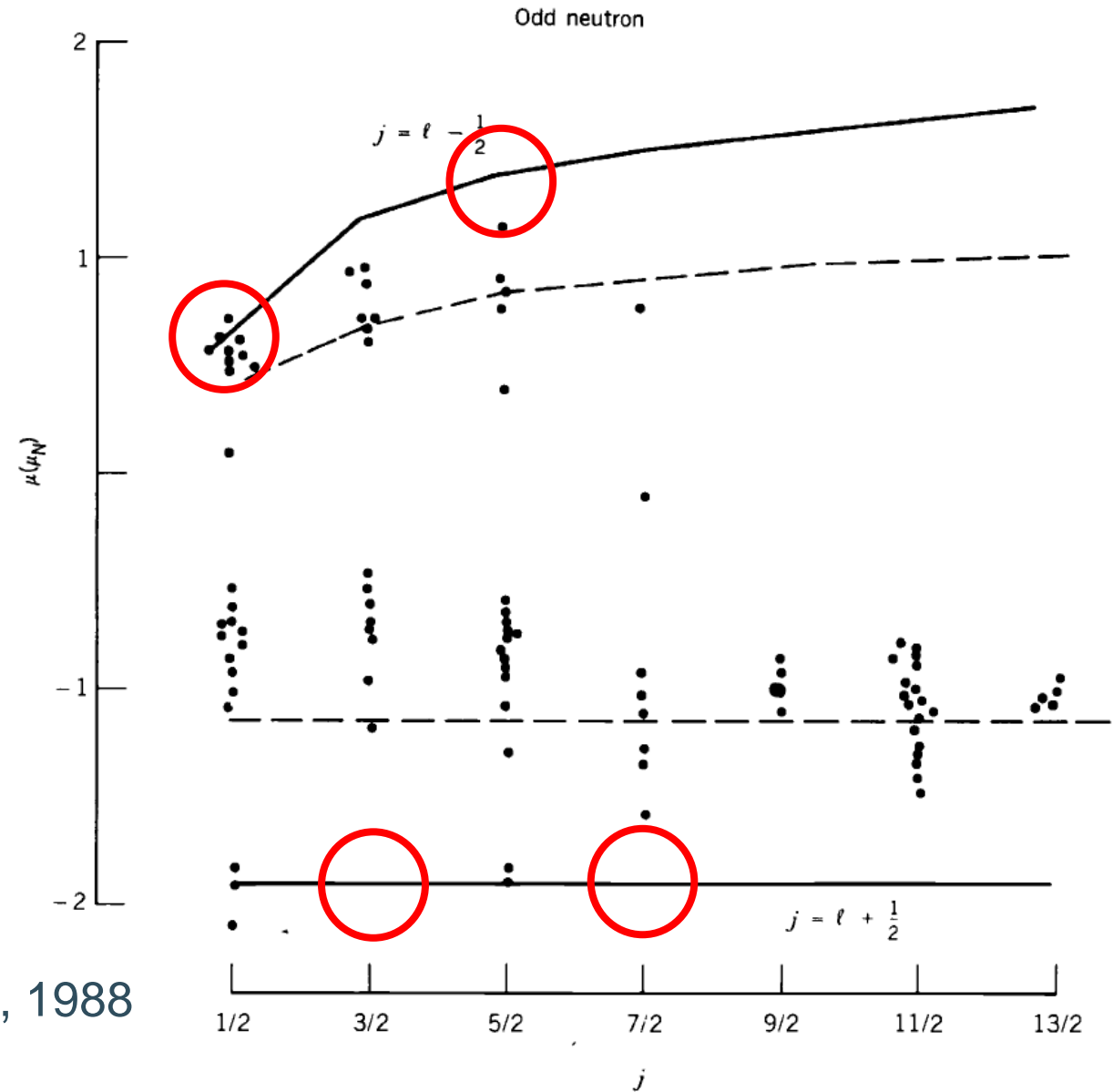
$(5/2)^-$ $1/2^-$

Schmidt moments even-odd

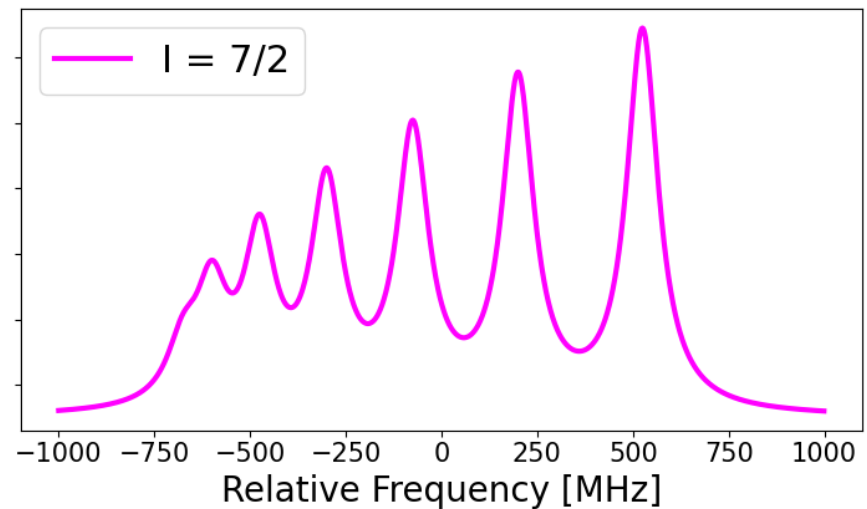
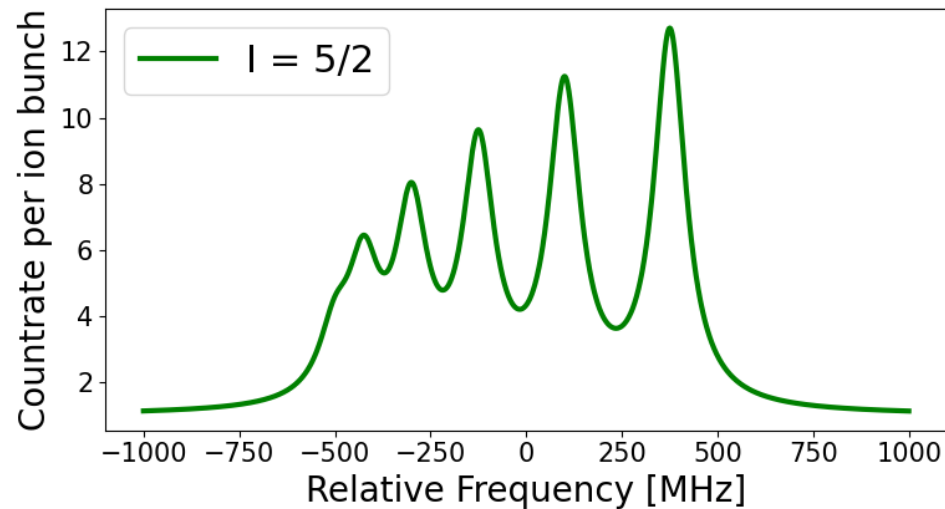
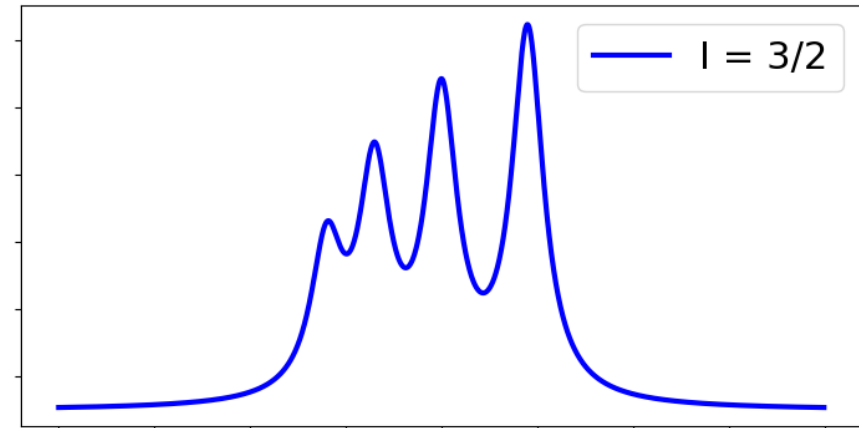
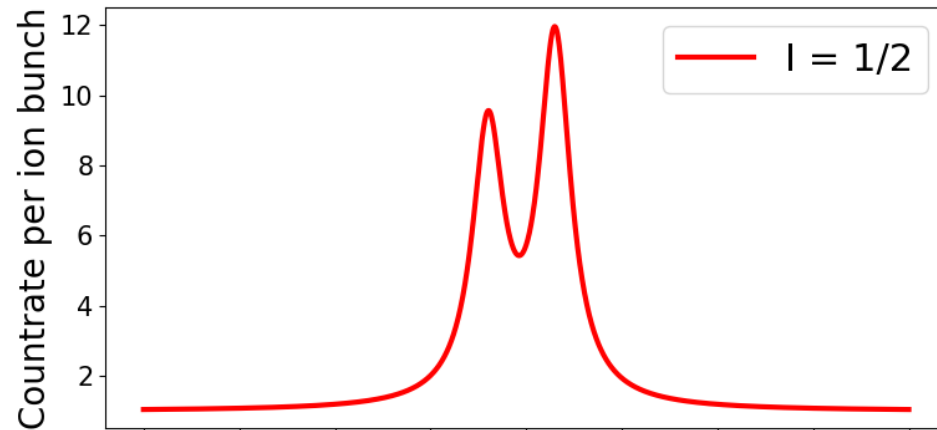
$$\nu p_{3/2}, \nu f_{7/2} \rightarrow j = l + 1/2$$

$$\nu p_{1/2}, \nu f_{5/2} \rightarrow j = l - 1/2$$

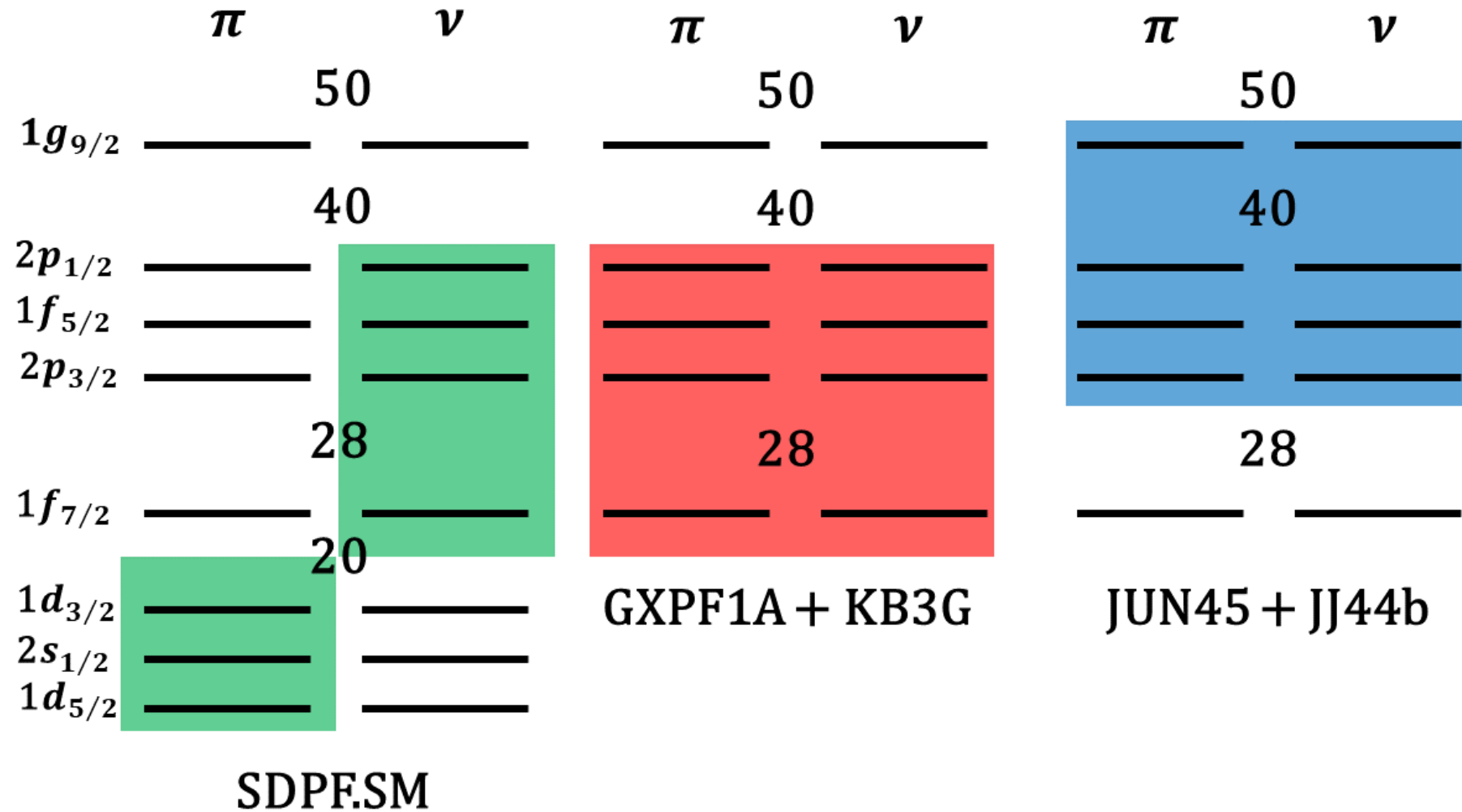
Krane, 1988



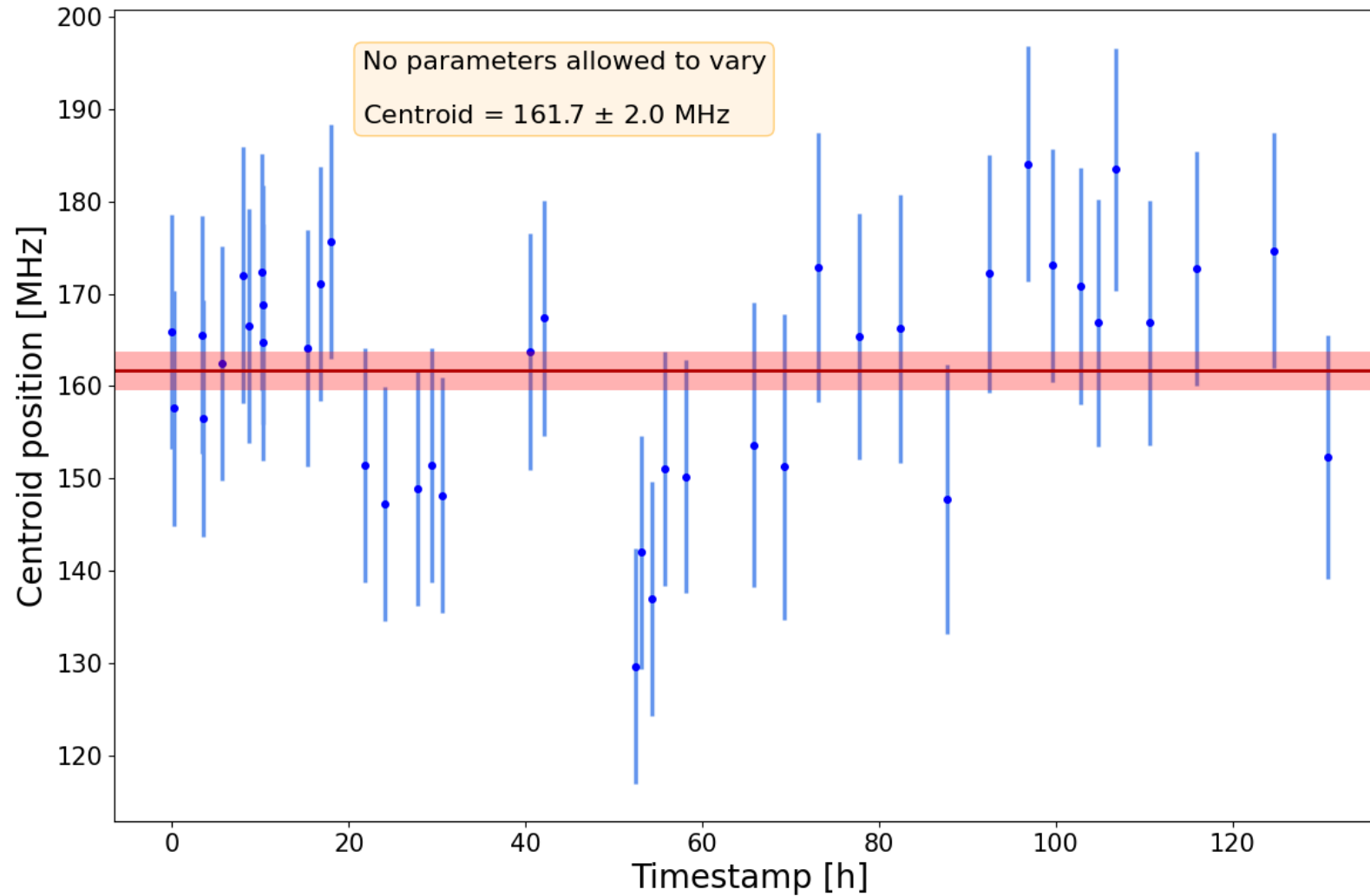
Number of peaks simulations



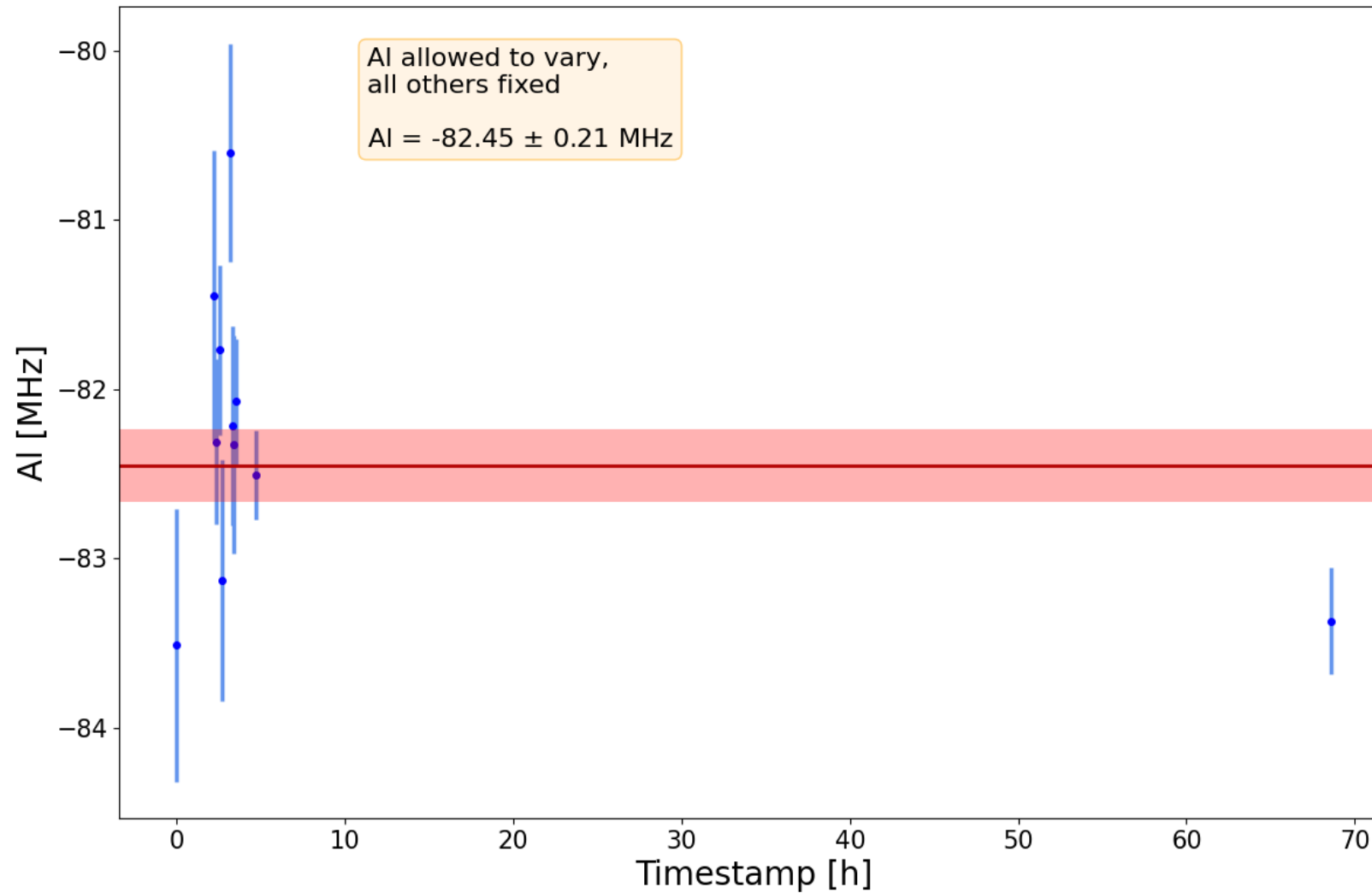
Model spaces shell model interactions



Examples mean and error calculations



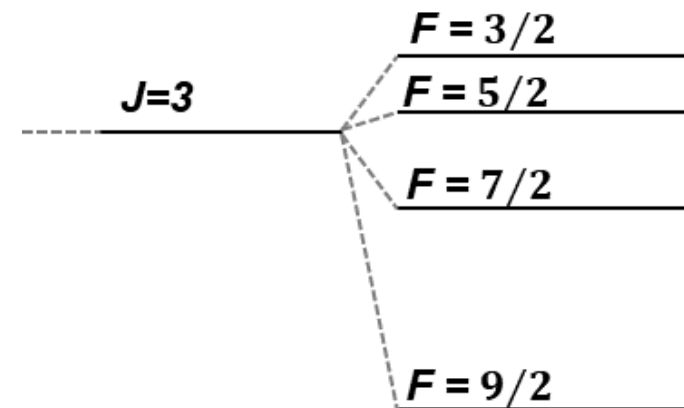
Examples mean and error calculations



Hyperfine Structure (HFS)

^{53}Cr ($3/2^-$)

Spin I	# HF levels
$1/2$	2
$3/2$	4
$5/2$	6
$7/2$	7



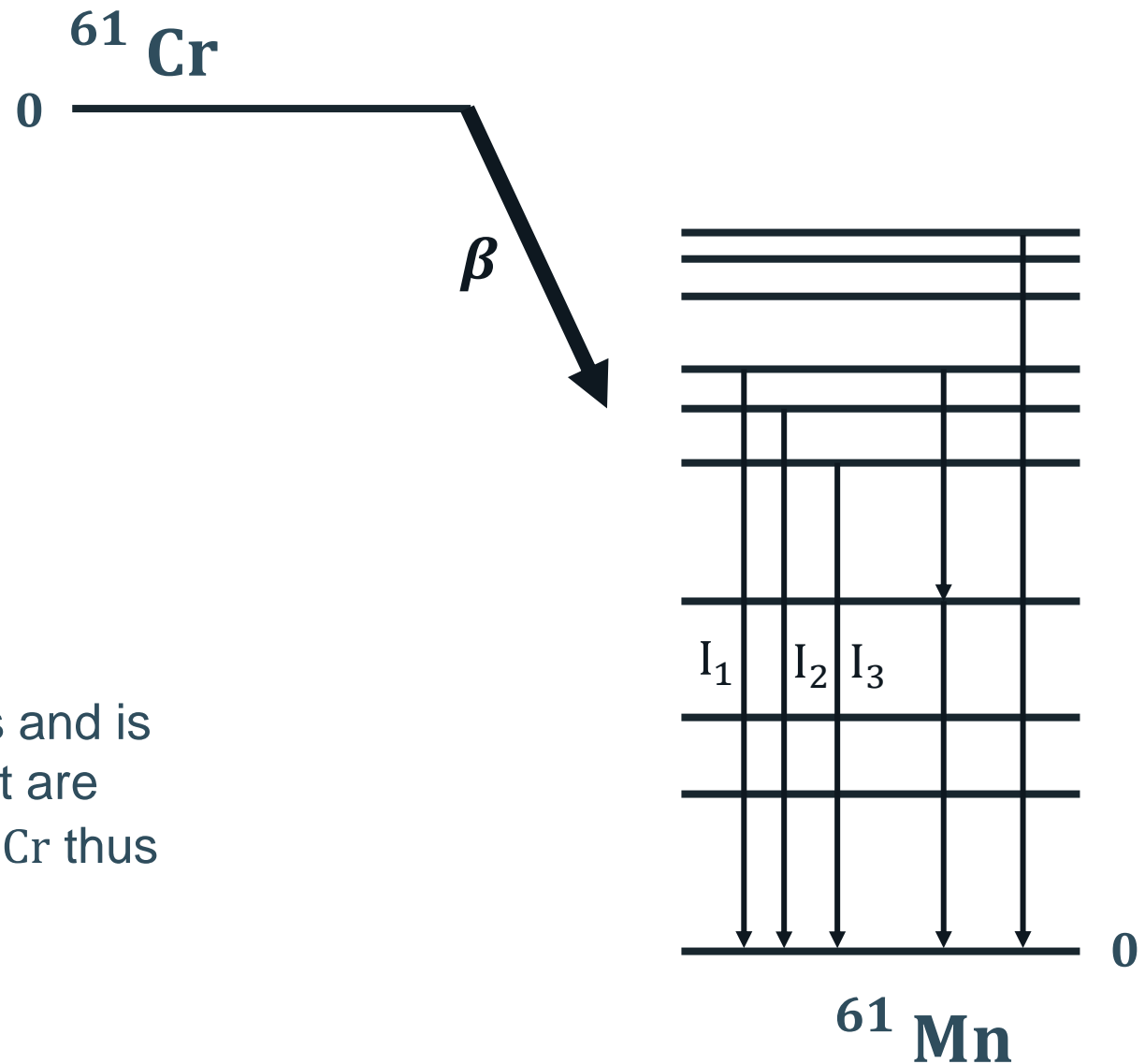
State	HFS
$\vec{J} = \vec{L} + \vec{S}$	$\vec{F} = \vec{I} + \vec{J}$

The Pandemonium effect

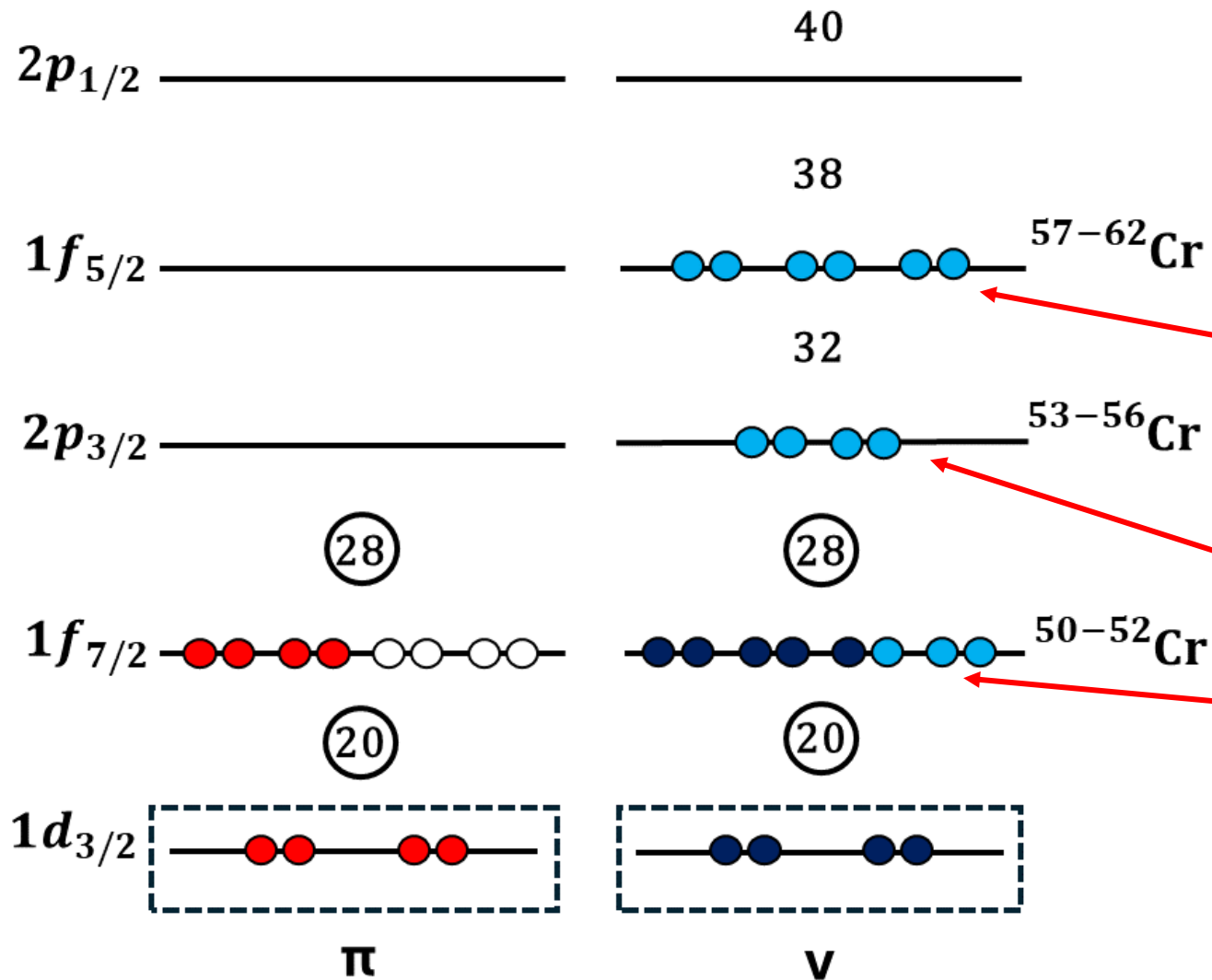
Feeding of the ground state of ^{61}Cr

$$F = 100\% - I_1 - I_2 - \dots$$

Assume we observe all gamma rays and is simply not the case. All gammas that are not observed are assigned to GS ^{61}Cr thus overestimating its strength



Nuclear shell structure in Cr



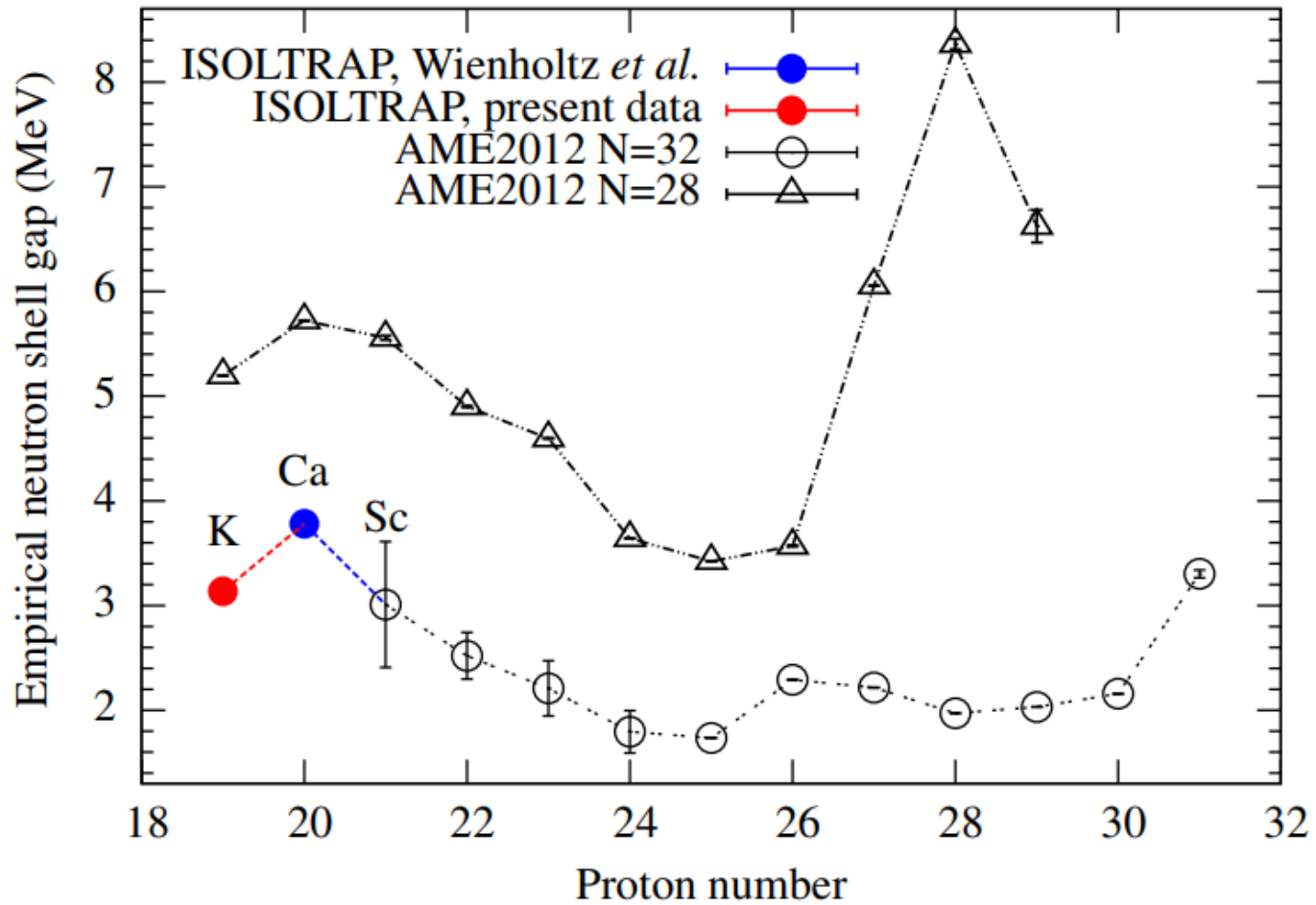
Unnuanced orbital filling
and single-particle
behaviour

Literature

- 61 → 5/2
- 59 → 5/2
- 57 → 5/2
- 55 → 3/2
- 53 → 3/2
- 51 → 7/2

- 61 → 5/2
- 59 → 5/2
- 57 → 5/2
- 55 → 3/2
- 53 → 3/2
- 51 → 7/2

32, 34 neutron shell gaps from mass measurements



Charge Radii calculation procedure

$$\delta\nu^{AA'} = K \frac{m_{A'} - m_A}{m_{A'}(m_A - m_e)} + F \delta\langle r^2 \rangle^{AA'}.$$

F = Field shift

$$\mu^{AA'} = \frac{m_{A'} - m_A}{m_{A'}(m_A - m_e)},$$

can be rewritten as

K = Mass shift

$$\frac{\delta\nu^{AA'}}{\mu^{AA'}} = F \frac{\delta\langle r^2 \rangle^{AA'}}{\mu^{AA'}} + K.$$

$$x^{AA'} = \frac{\delta\langle r^2 \rangle^{AA'}}{\mu^{AA'}} \quad \& \quad y^{AA'} = \frac{\delta\nu^{AA'}}{\mu^{AA'}},$$

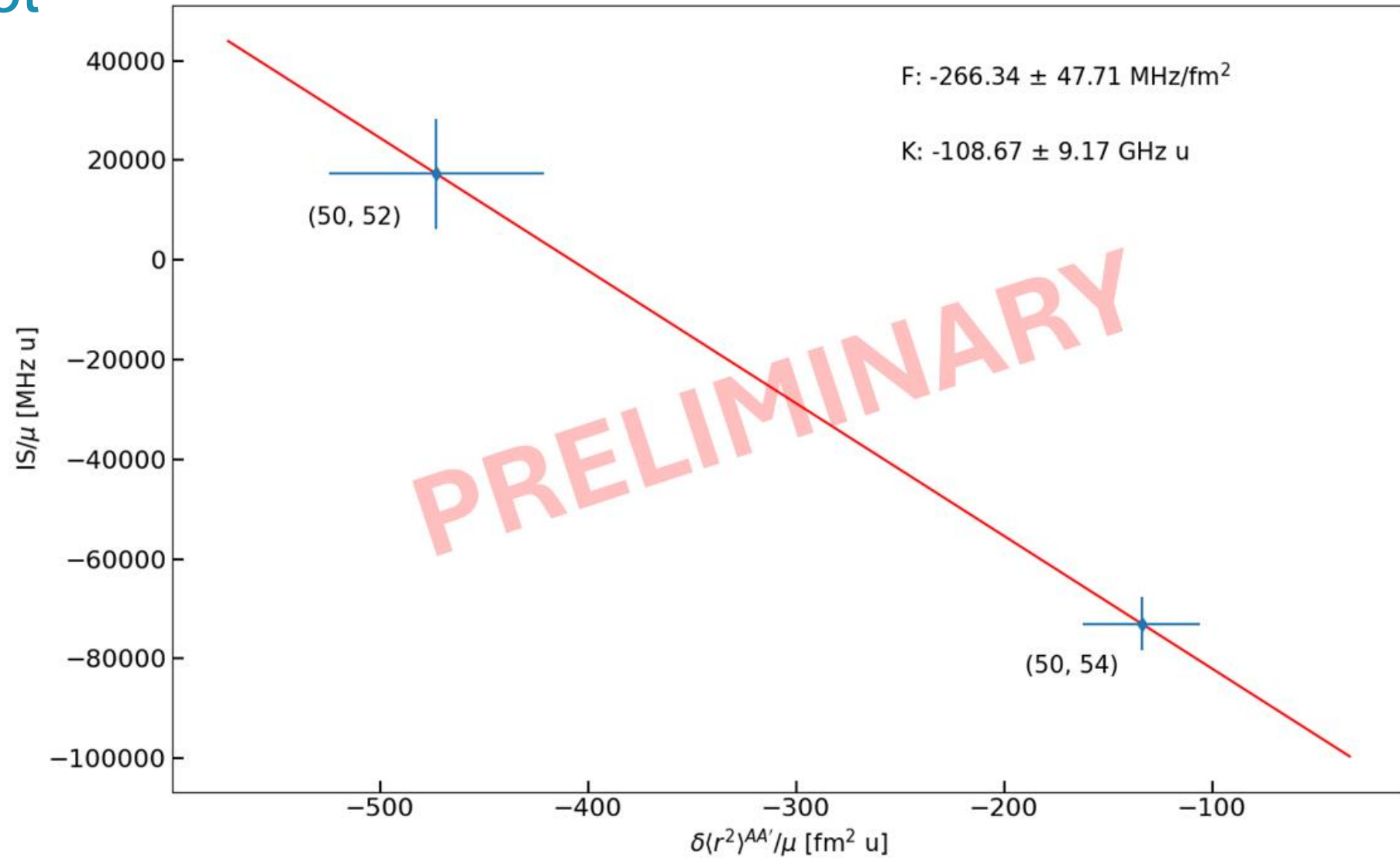
and mass shift K can be calculated as

$$F = \frac{x^{50,54} - x^{50,52}}{y^{50,54} - y^{50,52}} \quad \& \quad K = y^{50,54} - F x^{50,54}.$$

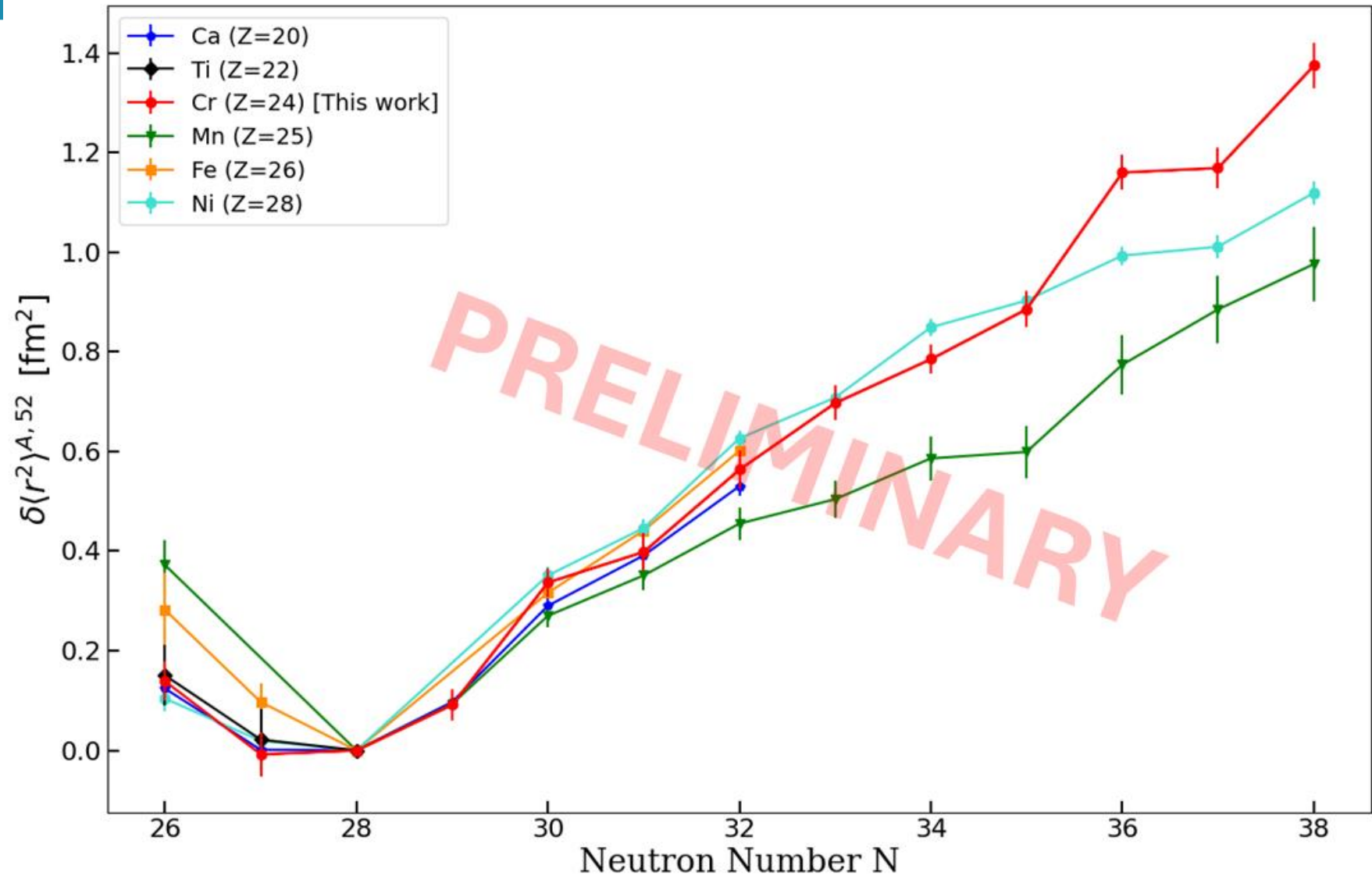
and K with their respective errors are shown in equation

$$F = -266(48)\text{MHz/fm}^2 \quad \& \quad K = -109(12)\text{GHz u}$$

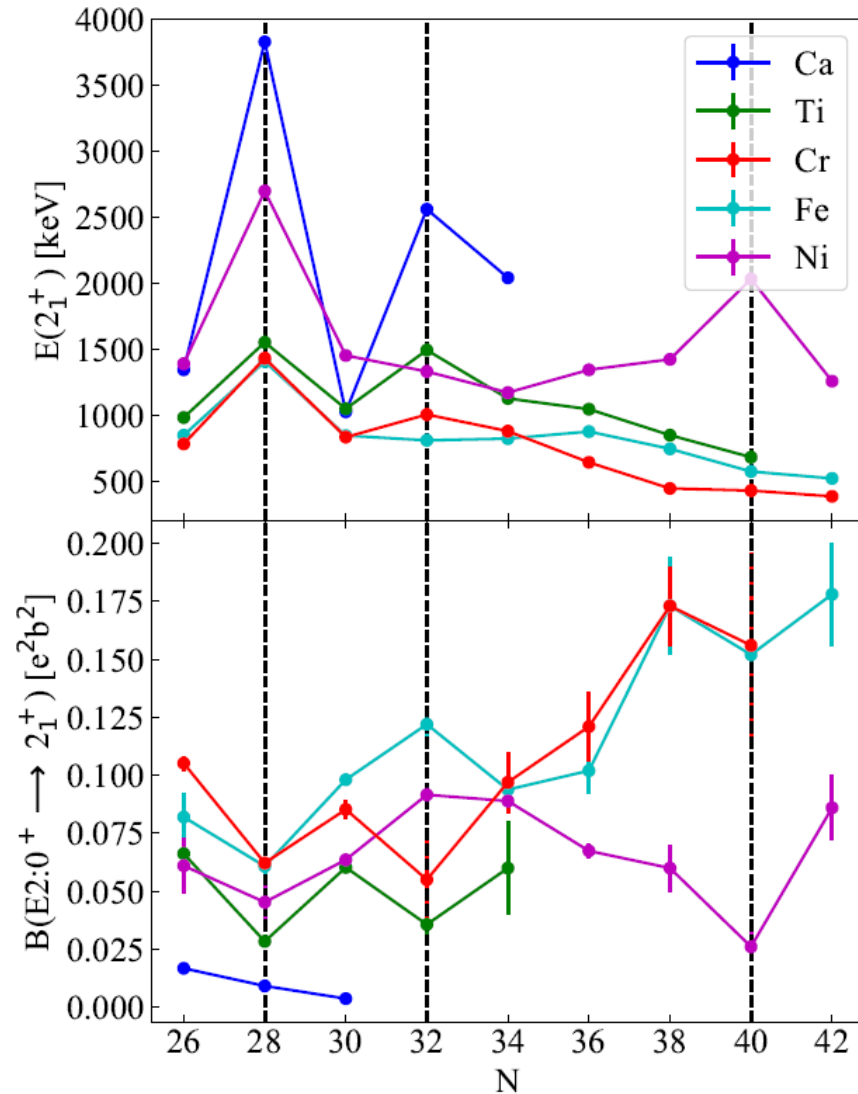
King Plot



Charge Radii



Shell structure evolution from Ca to Ni



Towards $N=40$

Low $E(2_1^+)$ in Cr and Fe, compared to Ni

High $B(E2: 0^+ \rightarrow 2_1^+)$ in Cr and Fe, not in Ni

Collective behaviour and deformation expected in Cr going towards $N=40$

Observables to investigate

Nuclear ground State Spin I

even-even $\rightarrow 0^+$

even-odd \rightarrow depends on
specific configuration of
neutrons

Nuclear magnetic dipole moment μ

$$\mu = gI\mu_N.$$

Interaction of nucleus with magnetic fields

g-factor sensitive to specific configuration

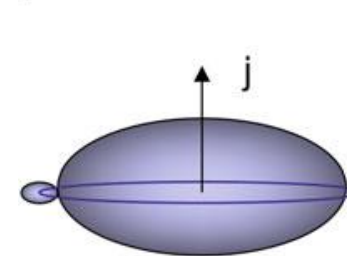
RMS Charge Radius $\langle r_c^2 \rangle$

Expect small radii for stable
nuclei = around magic numbers

Large radii for deformed nuclei

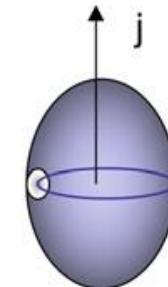
Nuclear electric quadrupole moment Q

particle in an orbit



oblate
 $Q < 0$

hole in an orbit



prolate
 $Q > 0$

Observables to investigate

Nuclear ground State Spin I

even-even $\rightarrow 0^+$

even-odd \rightarrow depends on
specific configuration of
neutrons orbitals

Nuclear magnetic dipole moment μ

even-even $\rightarrow 0^+$

even-odd \rightarrow depends on
specific configuration of
neutrons orbitals

Interaction of nucleus with magnetic fields

g-factor sensitive to specific configuration

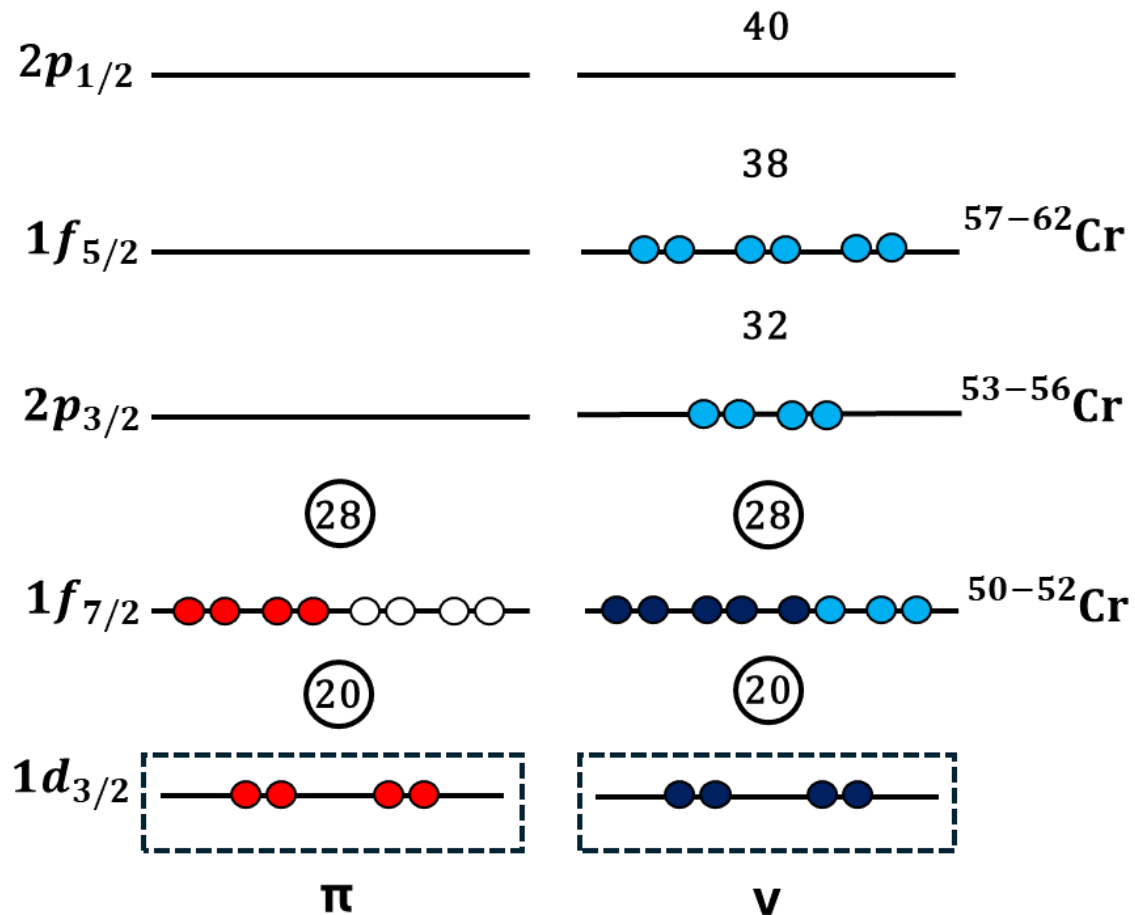
RMS Charge Radius $\langle r_c^2 \rangle$

Further interpretation necessary

Nuclear electric quadrupole moment Q

For Cr impossible to measure within
experimental resolution (~ 100 MHz)

Nuclear shell structure in Cr



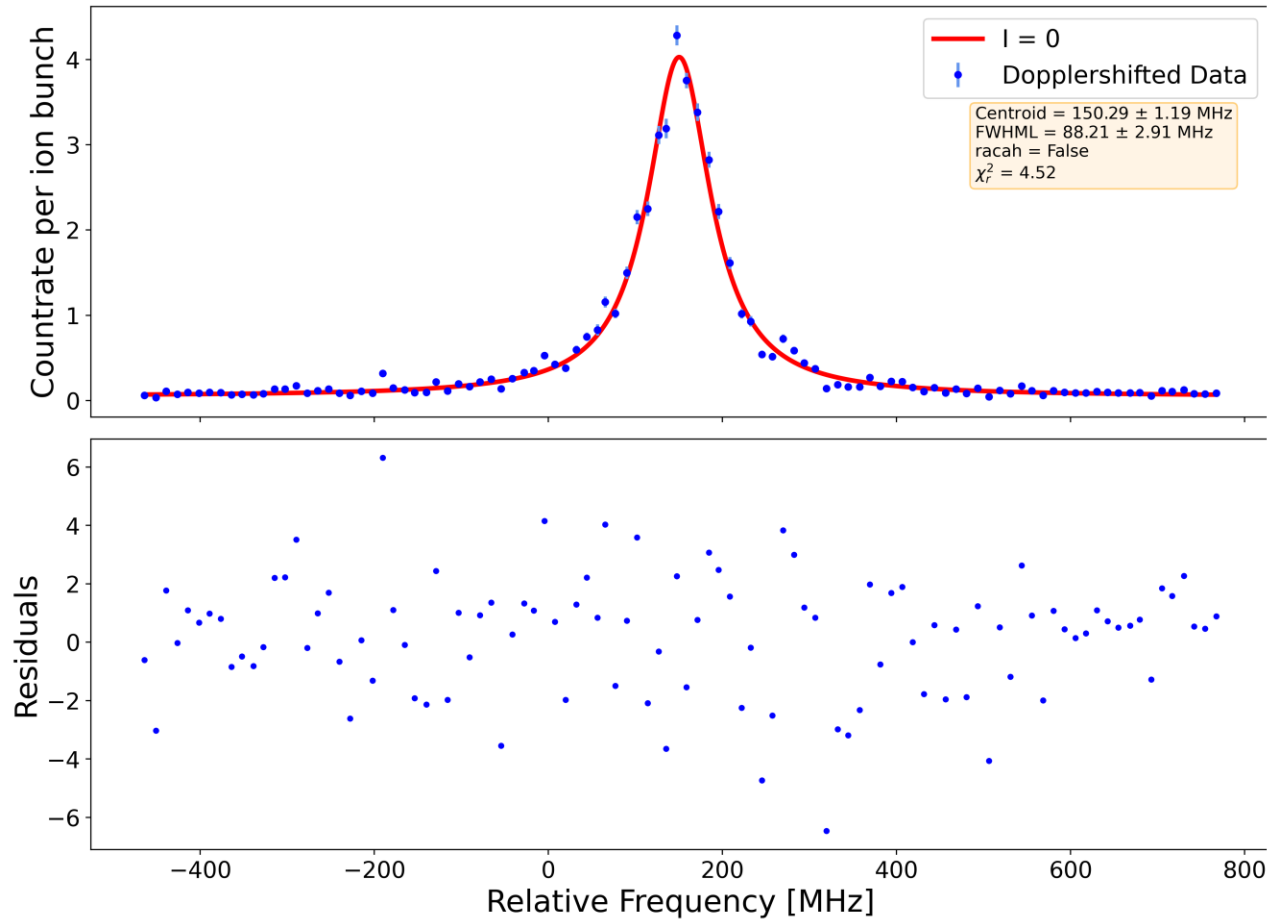
Unnuanced orbital filling and single-particle behaviour

Literature

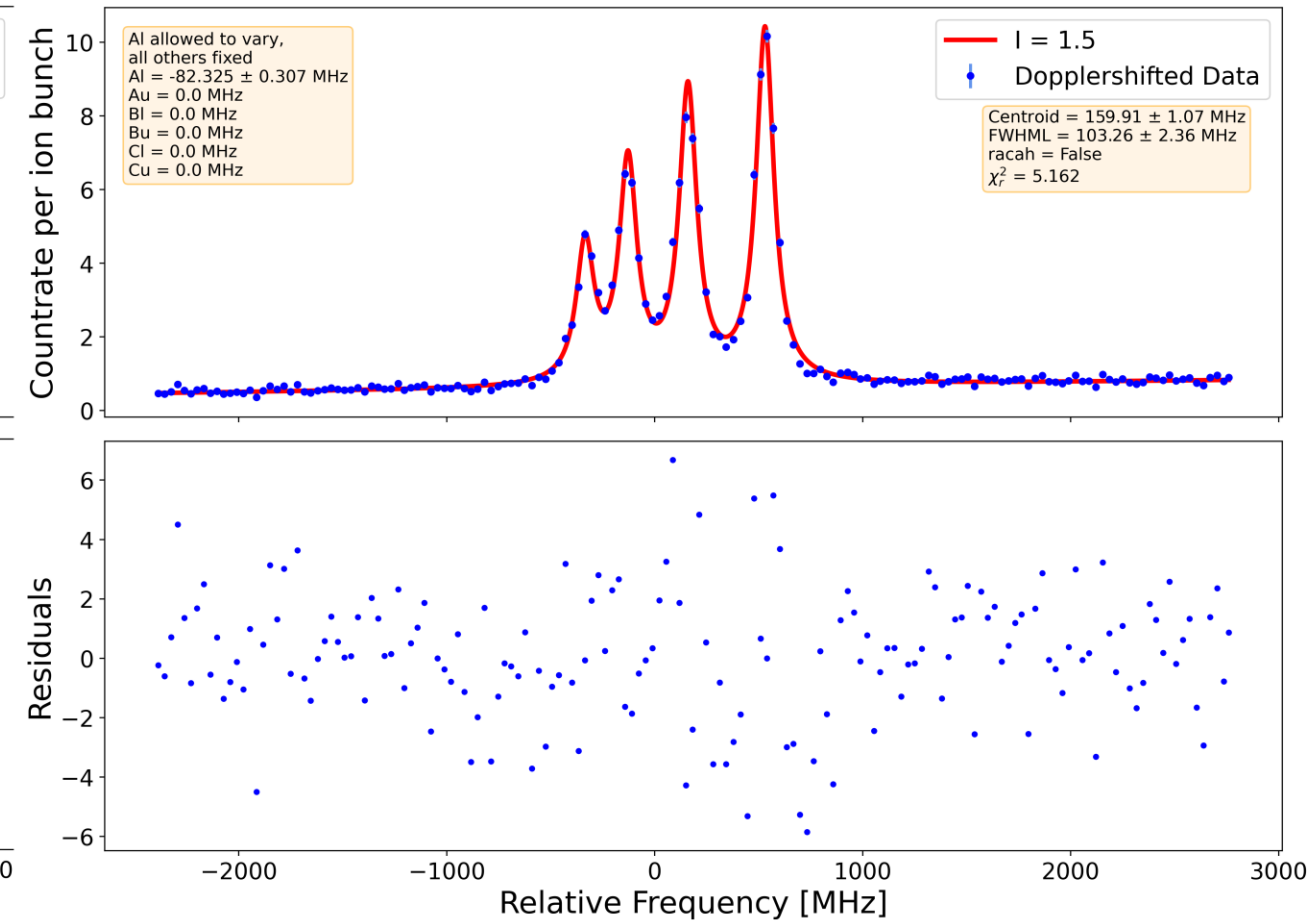
- | | |
|------------------------|------------------------------------|
| • $51 \rightarrow 7/2$ | • $51 \rightarrow 7/2^-$ measured |
| • $53 \rightarrow 3/2$ | • $53 \rightarrow 3/2^-$ measured |
| • $55 \rightarrow 3/2$ | • $55 \rightarrow 3/2^-$ measured |
| • $57 \rightarrow 5/2$ | • $57 \rightarrow 3/2^-$ tentative |
| • $59 \rightarrow 5/2$ | • $59 \rightarrow 1/2^-$ tentative |
| • $61 \rightarrow 5/2$ | • $61 \rightarrow 5/2^-$ tentative |

Isotope	J^π	GXPF1A		KB3G	
		Wave function (neutron)	Probability	Wave function (neutron)	Probability
^{55}Cr	$1/2_1^-$	$(0f_{7/2})^8, (1p_{3/2})^2, (0f_{5/2})^1, (1p_{1/2})^0$	0.24	$(0f_{7/2})^8, (1p_{3/2})^2, (0f_{5/2})^1, (1p_{1/2})^0$	0.26
		$(0f_{7/2})^8, (1p_{3/2})^2, (0f_{5/2})^0, (1p_{1/2})^1$	0.22	$(0f_{7/2})^8, (1p_{3/2})^1, (0f_{5/2})^1, (1p_{1/2})^1$	0.18
		$(0f_{7/2})^8, (1p_{3/2})^1, (0f_{5/2})^1, (1p_{1/2})^1$	0.16	$(0f_{7/2})^8, (1p_{3/2})^2, (0f_{5/2})^0, (1p_{1/2})^1$	0.14
	$3/2_1^-$	$(0f_{7/2})^8, (1p_{3/2})^3, (0f_{5/2})^0, (1p_{1/2})^0$	0.47	$(0f_{7/2})^8, (1p_{3/2})^3, (0f_{5/2})^0, (1p_{1/2})^0$	0.37
		$(0f_{7/2})^8, (1p_{3/2})^2, (0f_{5/2})^0, (1p_{1/2})^1$	0.11		
	$5/2_1^-$	$(0f_{7/2})^8, (1p_{3/2})^2, (0f_{5/2})^1, (1p_{1/2})^0$	0.44	$(0f_{7/2})^8, (1p_{3/2})^2, (0f_{5/2})^1, (1p_{1/2})^0$	0.42
^{57}Cr	$1/2_1^-$	$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^0, (1p_{1/2})^1$	0.51	$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^0, (1p_{1/2})^1$	0.19
		$(0f_{7/2})^8, (1p_{3/2})^2, (0f_{5/2})^2, (1p_{1/2})^1$	0.12	$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^1, (1p_{1/2})^0$	0.18
				$(0f_{7/2})^8, (1p_{3/2})^3, (0f_{5/2})^2, (1p_{1/2})^0$	0.14
	$3/2_1^-$	$(0f_{7/2})^8, (1p_{3/2})^3, (0f_{5/2})^1, (1p_{1/2})^1$	0.39	$(0f_{7/2})^8, (1p_{3/2})^3, (0f_{5/2})^1, (1p_{1/2})^1$	0.10
		$(0f_{7/2})^8, (1p_{3/2})^3, (0f_{5/2})^2, (1p_{1/2})^0$	0.12	$(0f_{7/2})^8, (1p_{3/2})^3, (0f_{5/2})^2, (1p_{1/2})^0$	0.47
	$5/2_1^-$	$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^1, (1p_{1/2})^0$	0.45	$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^1, (1p_{1/2})^0$	0.38
	$(0f_{7/2})^8, (1p_{3/2})^3, (0f_{5/2})^1, (1p_{1/2})^1$	0.15	$(0f_{7/2})^8, (1p_{3/2})^3, (0f_{5/2})^1, (1p_{1/2})^1$	0.15	
^{59}Cr	$1/2_1^-$	$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^2, (1p_{1/2})^1$	0.63	$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^2, (1p_{1/2})^1$	0.10
	$3/2_1^-$	$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^2, (1p_{1/2})^1$	0.44	$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^2, (1p_{1/2})^1$	0.60
		$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^3, (1p_{1/2})^0$	0.12	$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^3, (1p_{1/2})^0$	0.36
		$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^2, (1p_{1/2})^1$	0.11	$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^2, (1p_{1/2})^1$	0.22
	$5/2_1^-$	$(0f_{7/2})^8, (1p_{3/2})^3, (0f_{5/2})^2, (1p_{1/2})^2$	0.11	$(0f_{7/2})^8, (1p_{3/2})^2, (0f_{5/2})^1, (1p_{1/2})^2$	0.10
		$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^1, (1p_{1/2})^2$	0.28	$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^3, (1p_{1/2})^0$	0.47
$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^3, (1p_{1/2})^0$		0.21	$(0f_{7/2})^8, (1p_{3/2})^3, (0f_{5/2})^3, (1p_{1/2})^1$	0.11	
^{61}Cr	$1/2_1^-$	$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^2, (1p_{1/2})^1$	0.15	$(0f_{7/2})^8, (1p_{3/2})^2, (0f_{5/2})^3, (1p_{1/2})^2$	0.10
		$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^4, (1p_{1/2})^1$	0.76	$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^4, (1p_{1/2})^1$	0.75
				$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^4, (1p_{1/2})^1$	0.44
	$3/2_1^-$	$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^3, (1p_{1/2})^2$	0.71	$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^4, (1p_{1/2})^1$	0.44
		$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^4, (1p_{1/2})^1$	0.13	$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^3, (1p_{1/2})^2$	0.33
	$5/2_1^-$	$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^3, (1p_{1/2})^2$	0.78	$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^3, (1p_{1/2})^2$	0.60
			$(0f_{7/2})^8, (1p_{3/2})^4, (0f_{5/2})^4, (1p_{1/2})^1$	0.10	

Lorentzian Fitted Dopplershifted data Cr-58 Scan 5327



Lorentzian Fitted Dopplershifted data Cr-53 Scan 5216



Weighted mean

$$\hat{\mu} = \frac{\sum_{i=1}^n x_i w_i}{\sum_{i=1}^n w_i}. \quad w_i = \frac{1}{\sigma_i^2},$$

Std variation of mean

$$\hat{\sigma}_{\hat{\mu}}^2 = \sigma_{\hat{\mu}}^2 \chi^2 = \frac{1}{n-1} \frac{1}{\sum_{i=1}^n w_i} \sum_{i=1}^n w_i (x_i - \hat{\mu})^2,$$

$$\chi^2 \equiv \frac{1}{n-1} \sum_{i=1}^n w_i (x_i - \hat{\mu})^2.$$

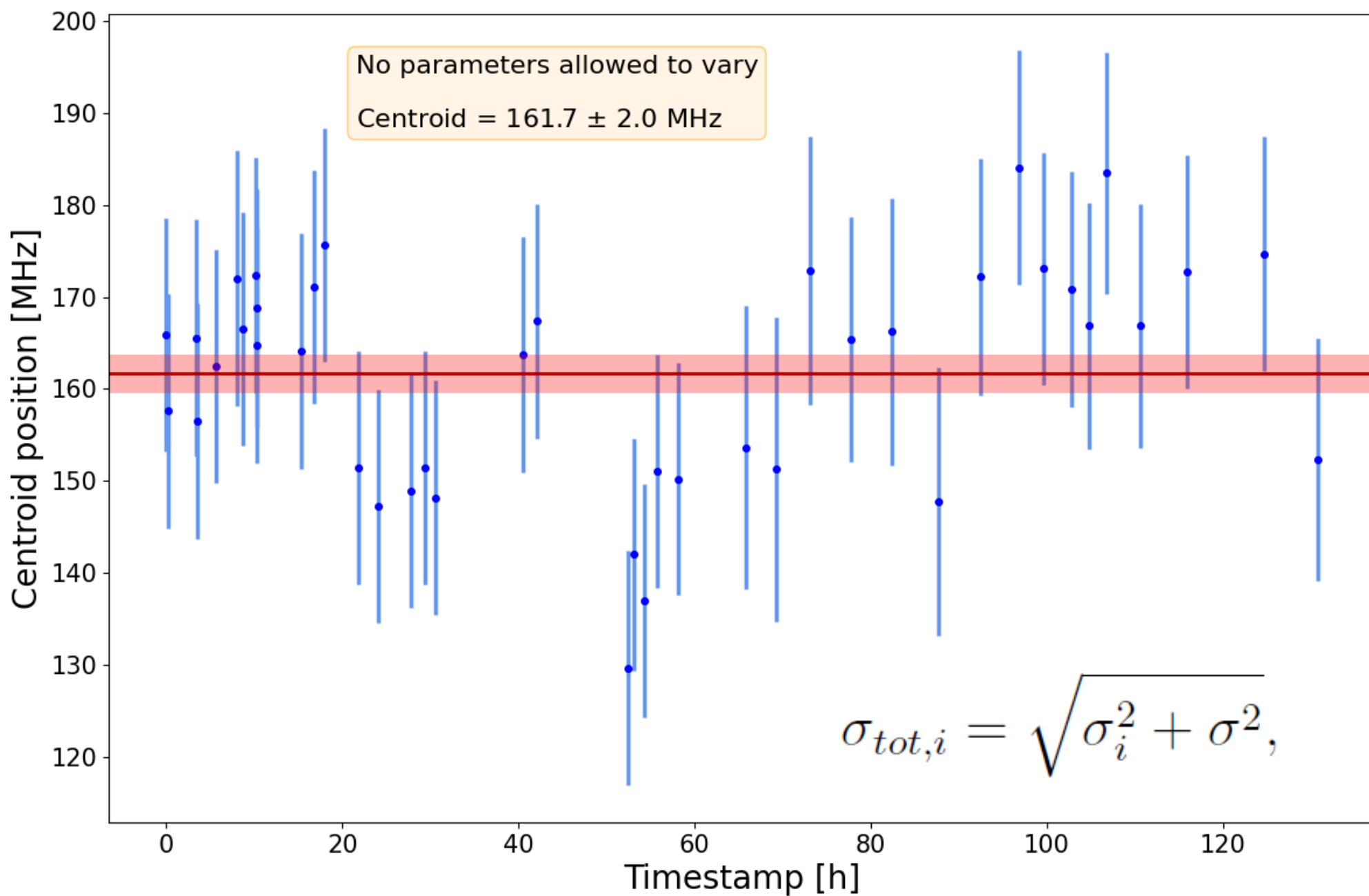
Weighted mean

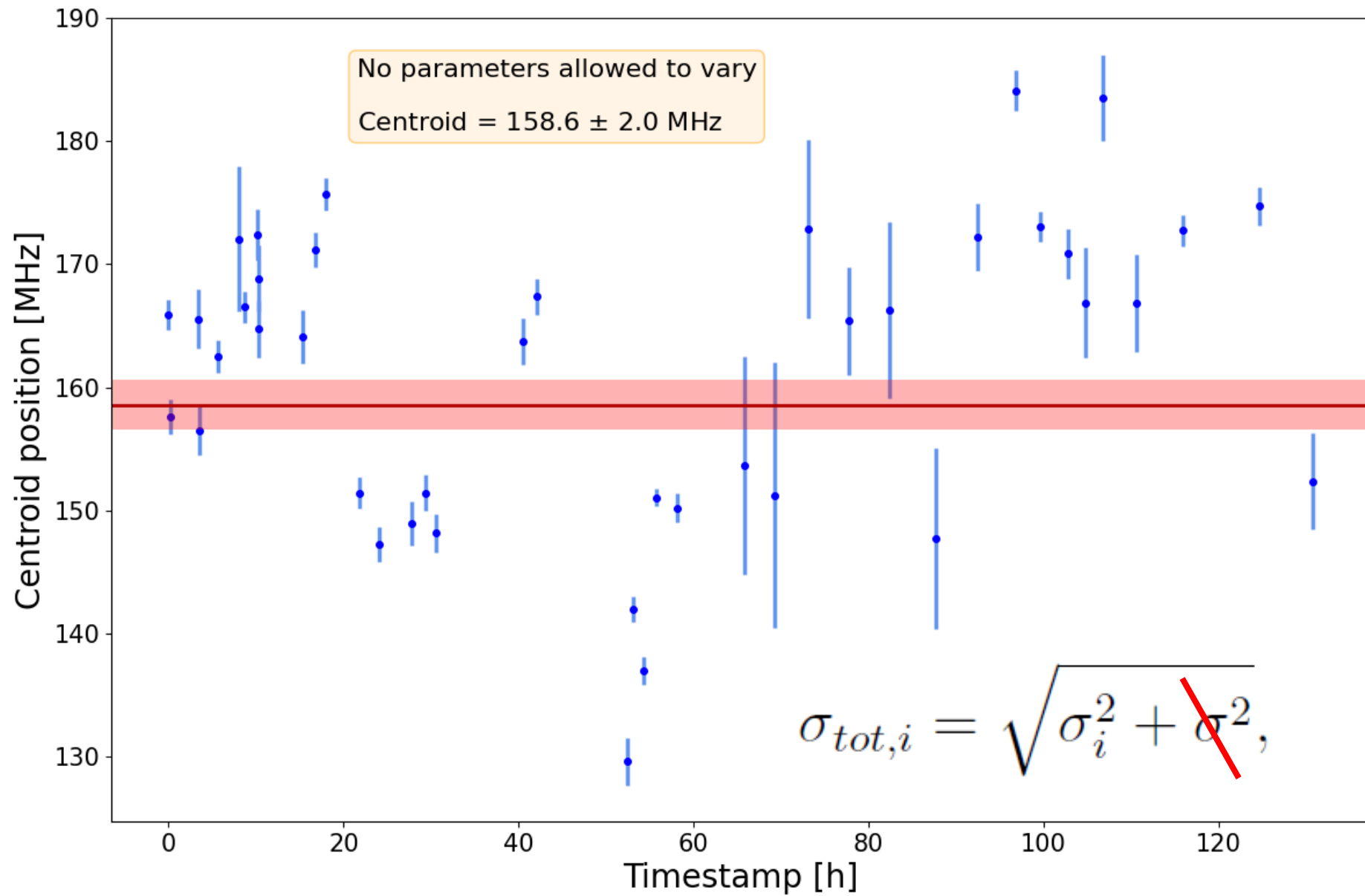
$$\hat{\mu} = \frac{\sum_{i=1}^n x_i w_i}{\sum_{i=1}^n w_i}. \quad w_i = \frac{1}{\sigma_i^2},$$

How to calculate the error on each datapoint

Std variation of the ref. isotope

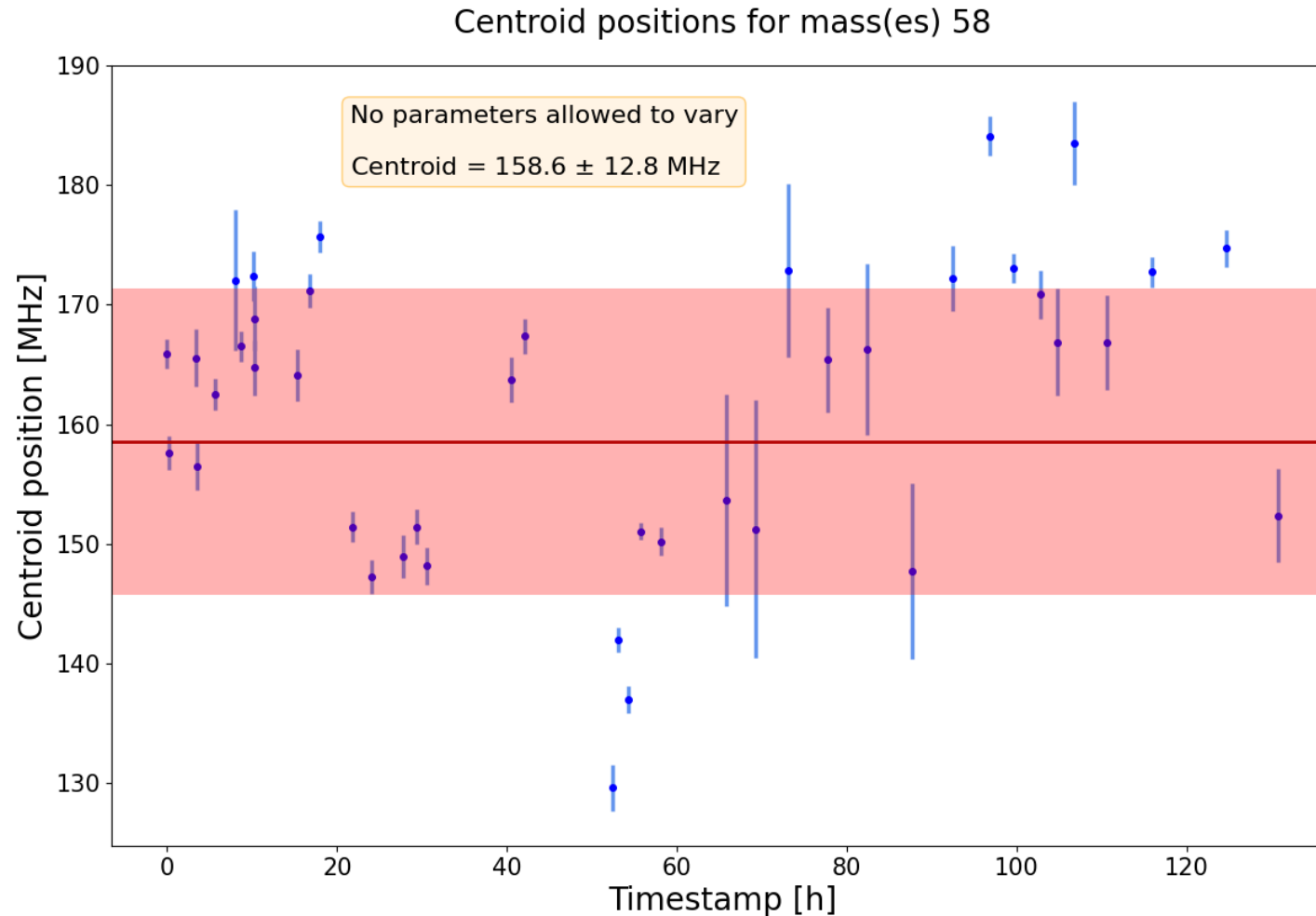
$$\sigma_{tot,i} = \sqrt{\sigma_i^2 + \sigma^2}, \quad \sigma^2 = \frac{\sum_{i=1}^n (x_i - \hat{\mu})^2}{n - 1}.$$



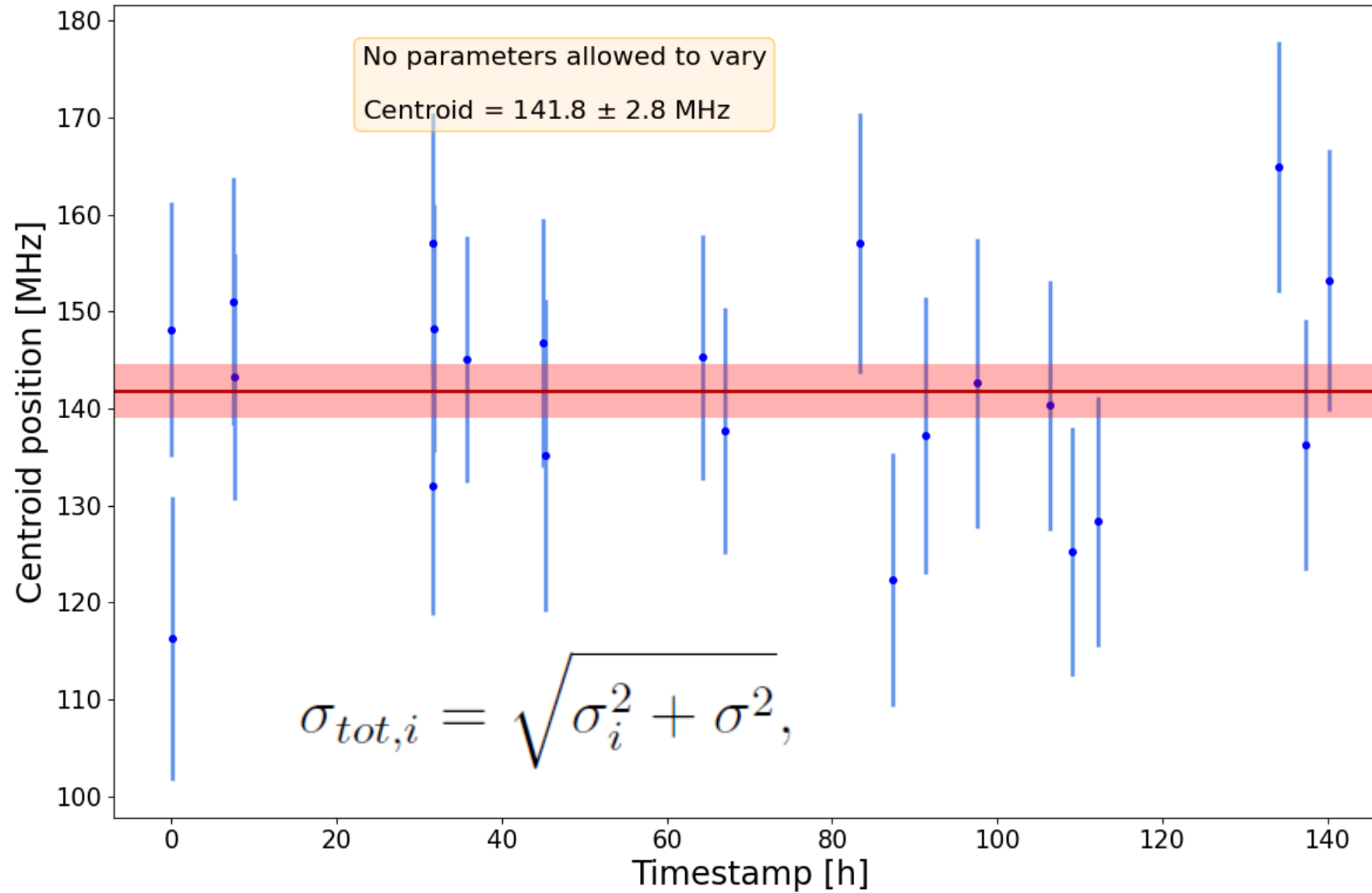


Last option: add σ to final error on weighted mean \rightarrow Very large errors

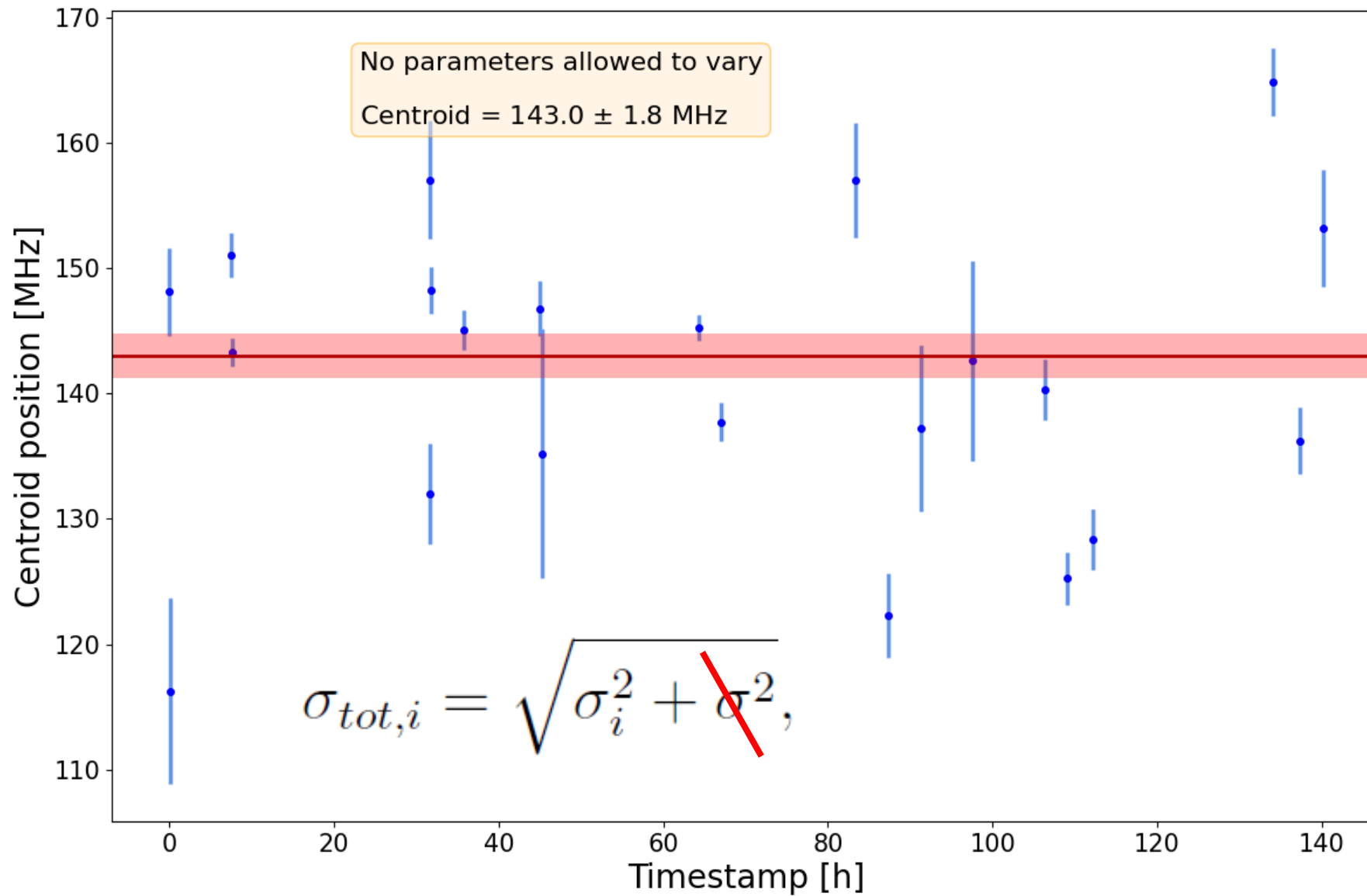
$$\sigma_{tot,i} = \sqrt{\sigma_i^2 + \cancel{\sigma^2}},$$



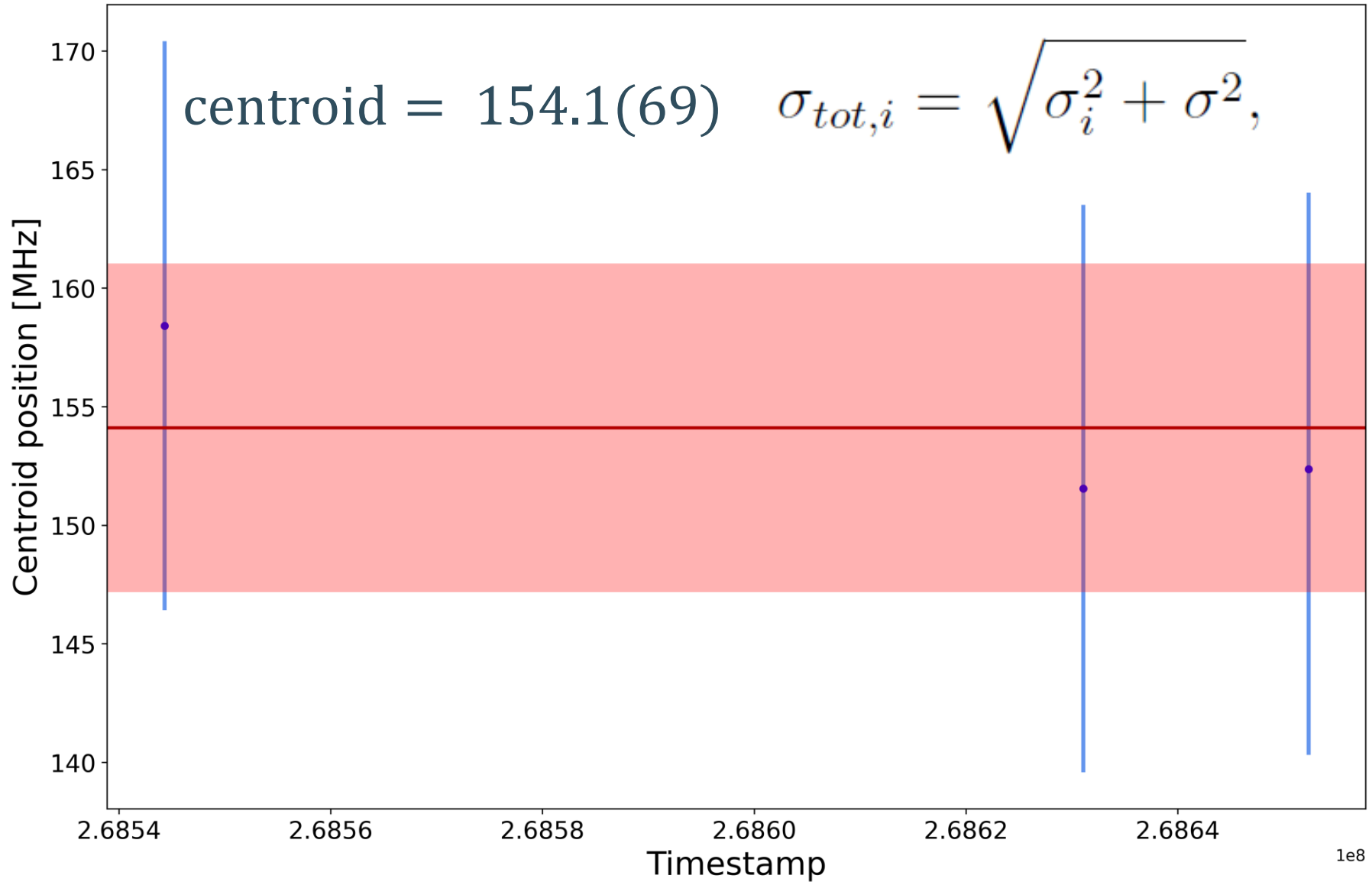
Centroid positions for mass(es) 54



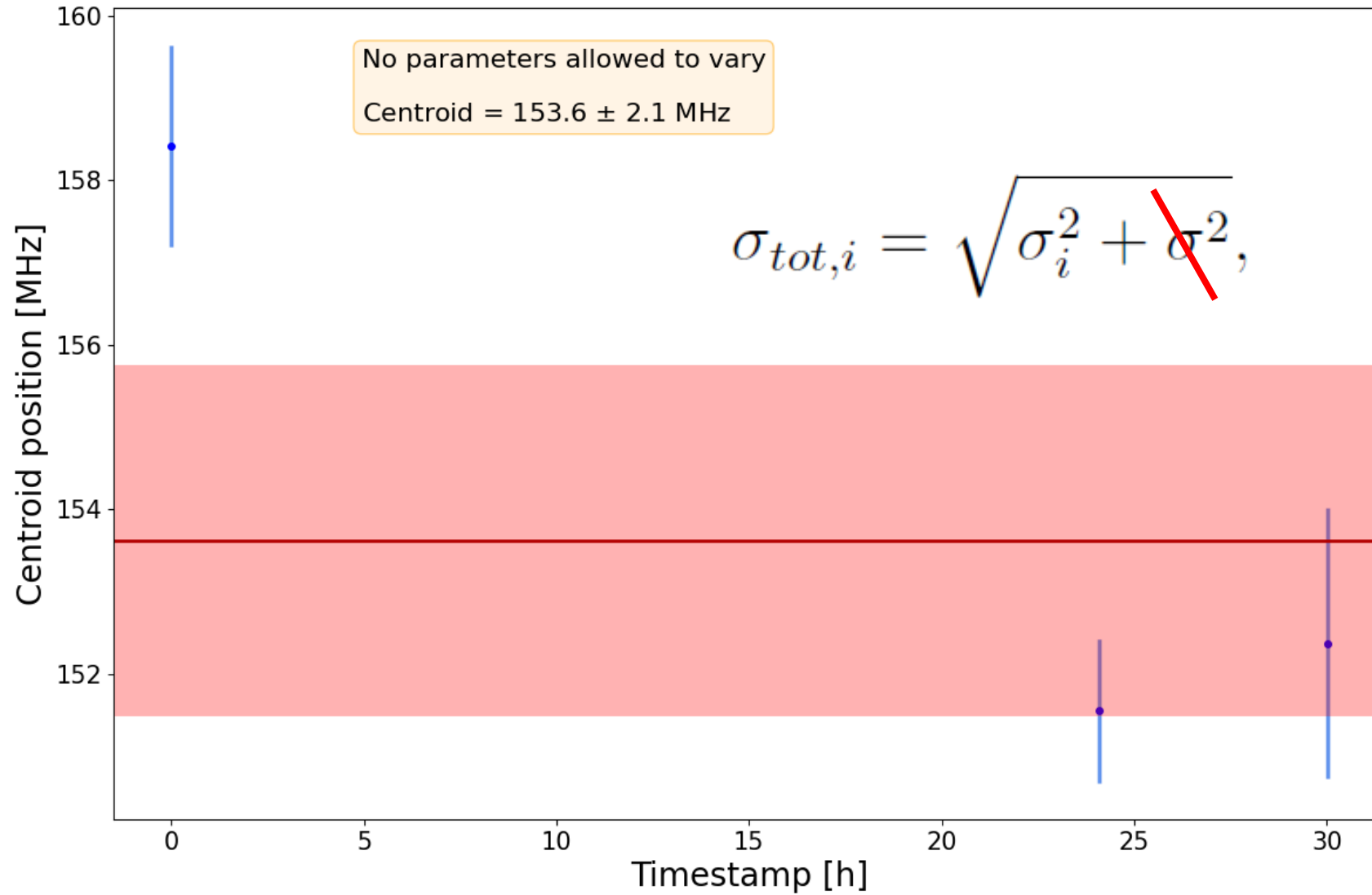
Centroid positions for mass(es) 54



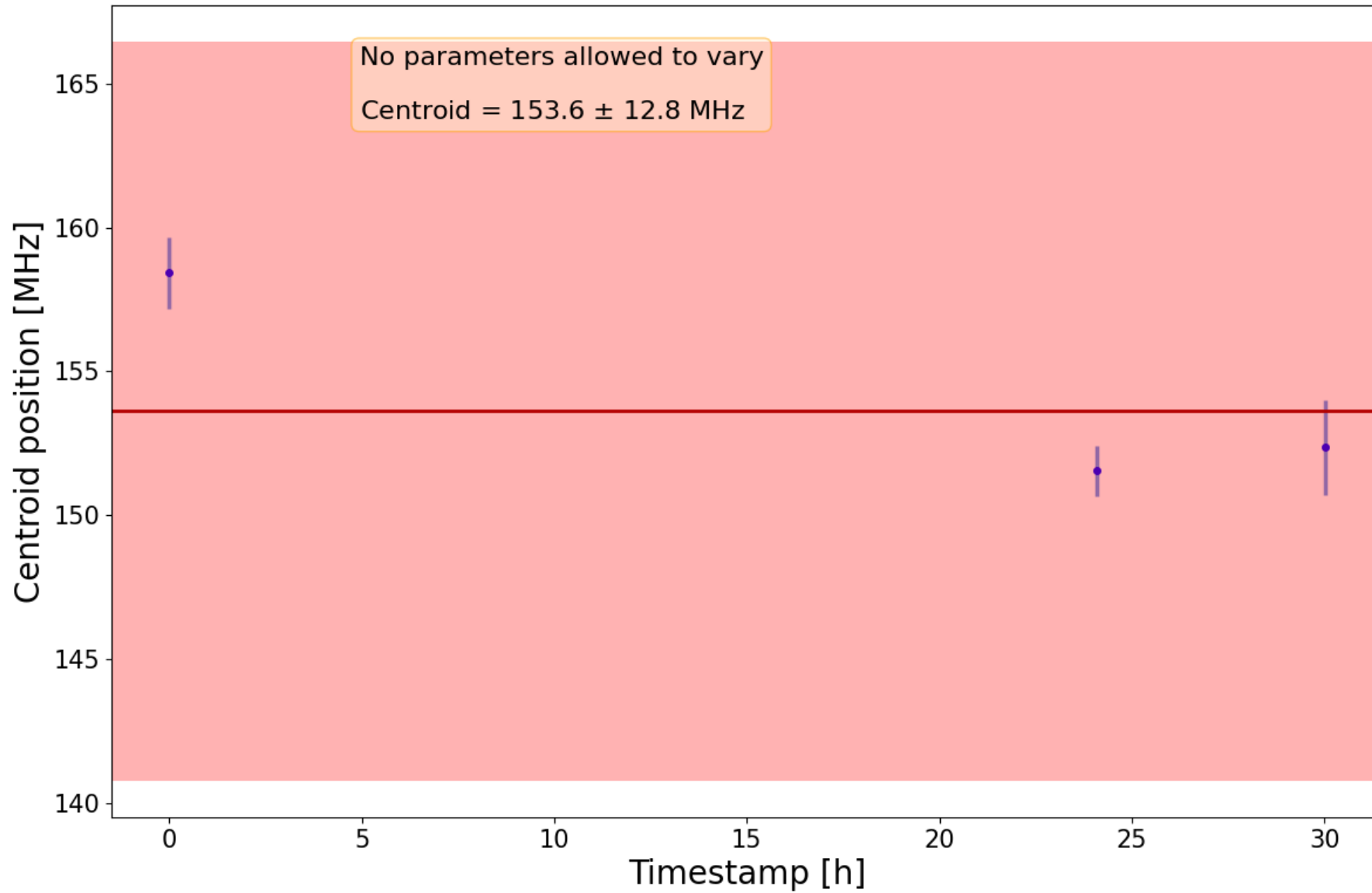
Centroid positions for mass(es) 52



Centroid positions for mass(es) 52



Centroid positions for mass(es) 52



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