

Simulation of Galaxy Density Profiles with Velocity-Dependent Self-Interaction

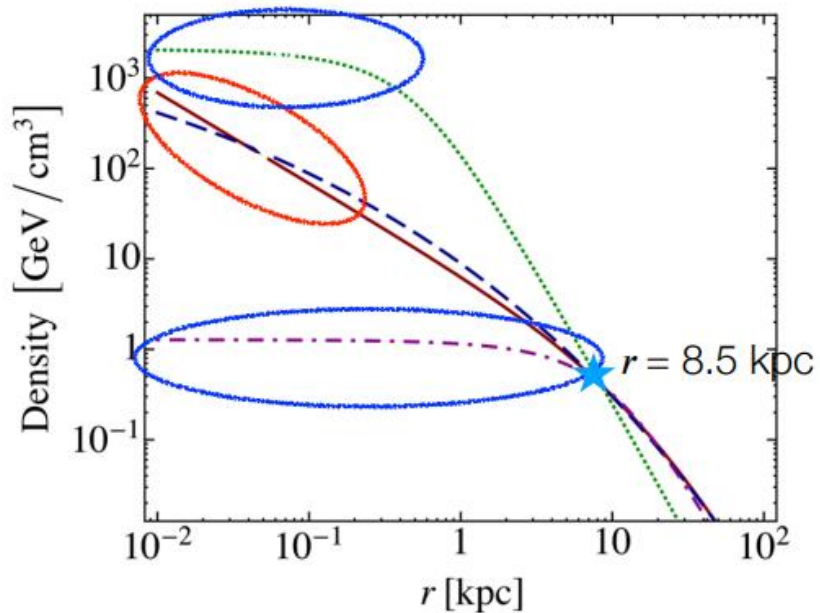
CAU HEP Workshop

THEP

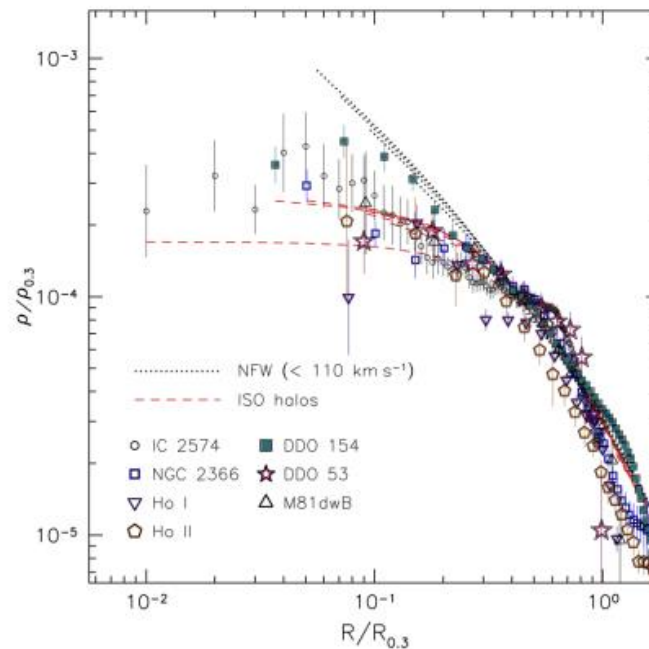
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27 Dec 2023

Core - Cusp Problem



Cohen, Lisanti, Pierce, Slatyer, JCAP 1310, 061 (2013)



Oh, de Blok, Brinks, Walter, Kennicutt, Astrophys. J. 141, 6 (2011)

- ✓ Red solid: NFW
- ✓ Blue dashed: Einasto
- ✓ Green dotted: Burkert $r_s = 0.5$ kpc
- ✓ Purple dot-dashed: Burkert $r_s = 10$ kpc

Cuspy

Cored

$$\rho_{\text{pseudo-iso}} = \frac{\rho_C}{1 + (r/r_s)^2}$$

ρ_C : core density

r_s : core radius

Velocity-dependent Dark Matter

In the low energy collision, effective range theory allows us a model-independent description for velocity-dependent cross-section. It is given by

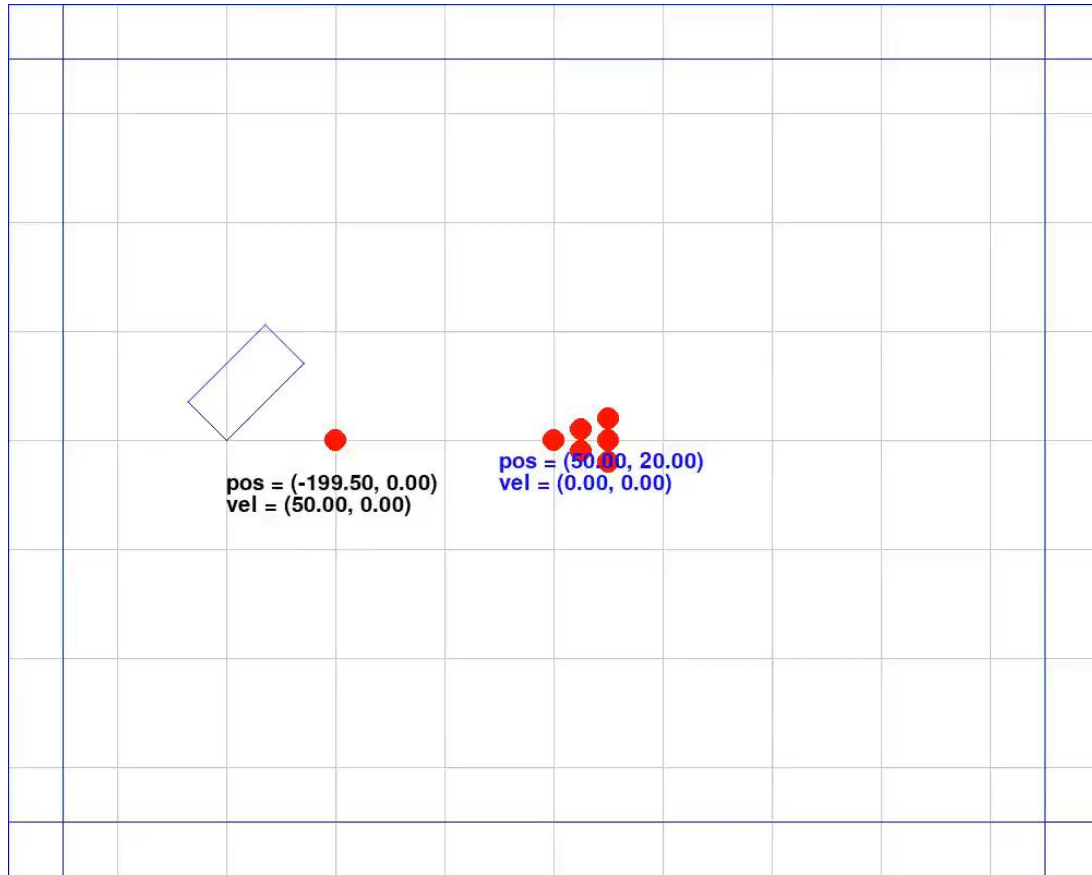
$$\sigma_0 = \frac{4\pi}{k^2} \sin^2 \delta_0 \approx \frac{4\pi a^2}{1 + k^2 \left(a^2 - ar_e \right) + \frac{1}{4} a^2 r_e^2 k^4}$$

(k : momentum, δ_0 : phase shift, a : scattering length, r_e : effective range)

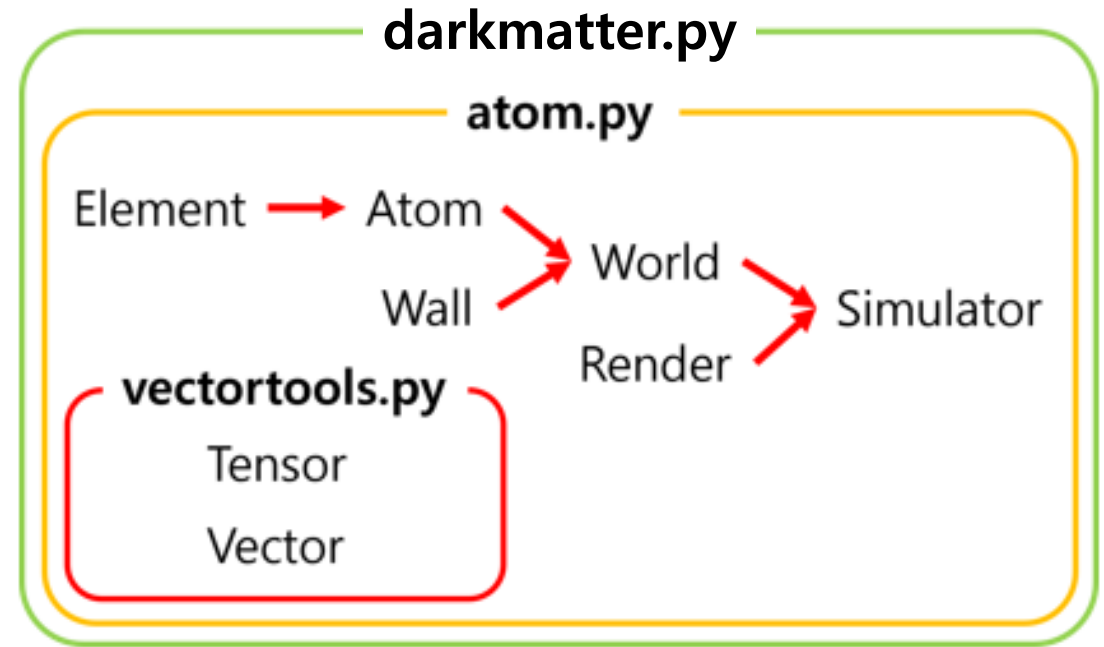
We consider special case $r_e = \frac{a}{2}$ in this work.

Then we get an expression of the following form, $\sigma(v) = \sigma_0 \left(1 + \left(\frac{v}{w} \right)^2 \right)^{-2}$

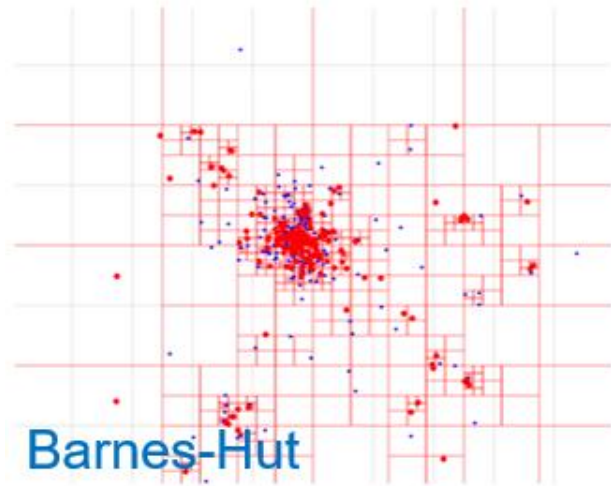
Implement of Dark Matter Simulation



atom.py



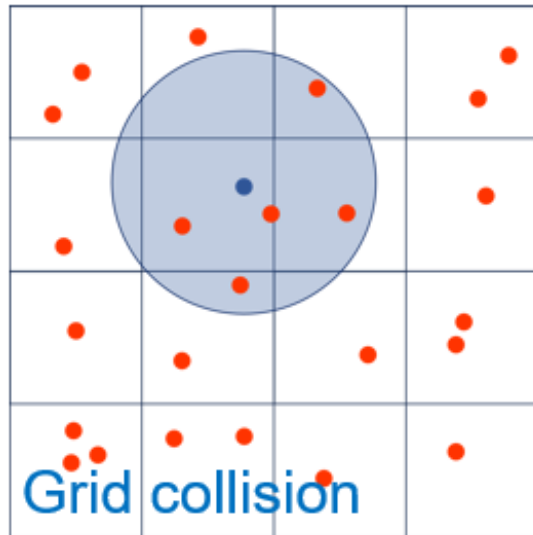
Implement of Dark Matter Simulation



- N-body Simulation algorithm

Direct N-body method: $t \sim O(N^2)$

Barnes-Hut algorithm: $t \sim O(N \log N)$ ✓

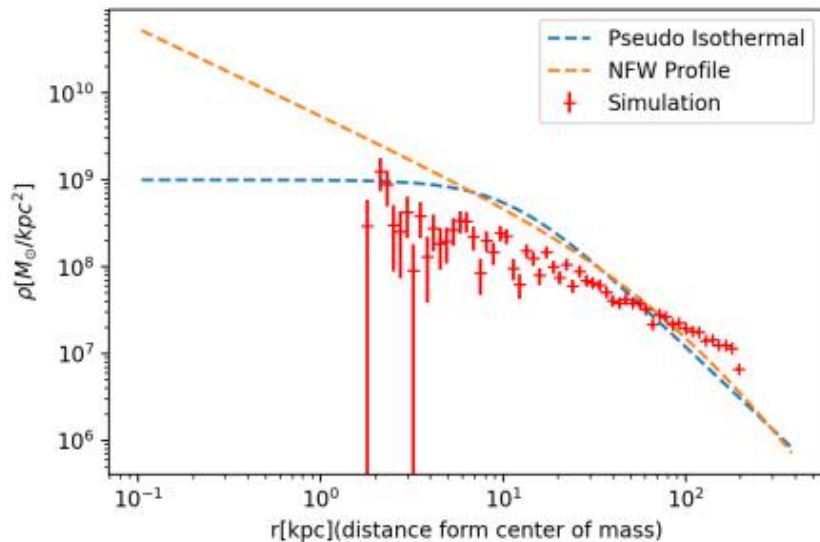
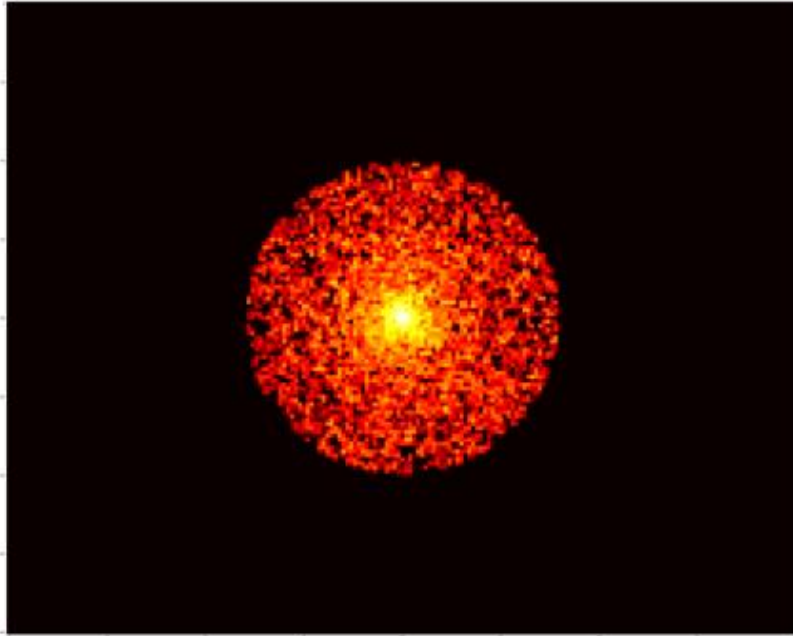


- Collision detection algorithm

Direct N collision detection: $t \sim O(N^2)$

Grid collision detection: $t \sim O(N^2/nm)$ ✓

Simulation Setup



	CDM	v-indep SIDM	v-dep SIDM
$\frac{\sigma_0}{m}$	$0 \text{ cm}^2/\text{g}$	$2.17 \text{ cm}^2/\text{g}$	$2.17 \text{ cm}^2/\text{g}$
w	-	∞	180 km/s

$$\sigma(v) = \sigma_0 \left(1 + \left(\frac{v}{w} \right)^2 \right)^{-2}$$

Disk radius: 200 kpc

Initial Density profile: $\rho \propto \frac{1}{r}$

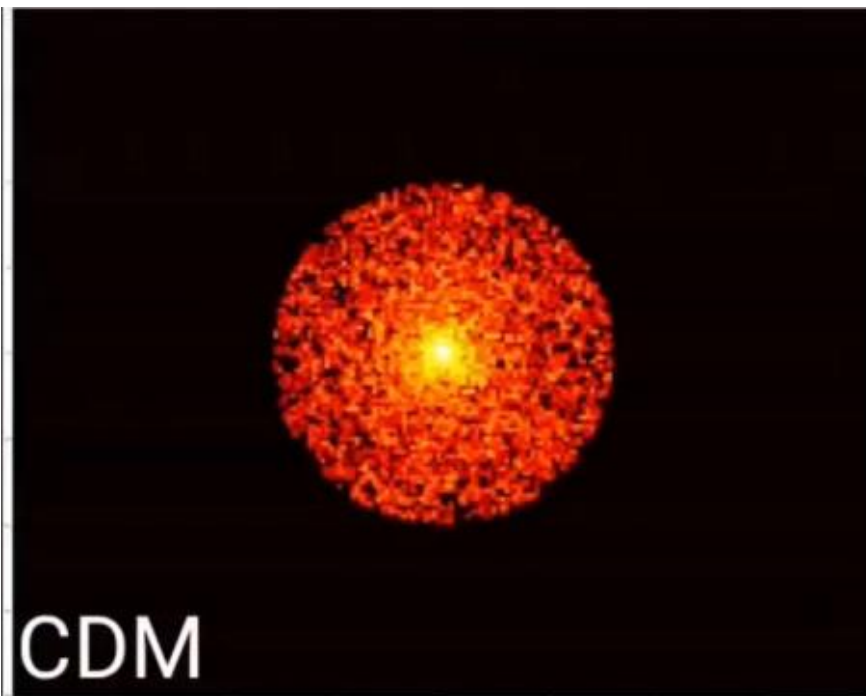
Initial Velocity: 5 kpc/Gyr & Random direction

Total mass: $2.5 \times 10^{12} M_{\odot}$

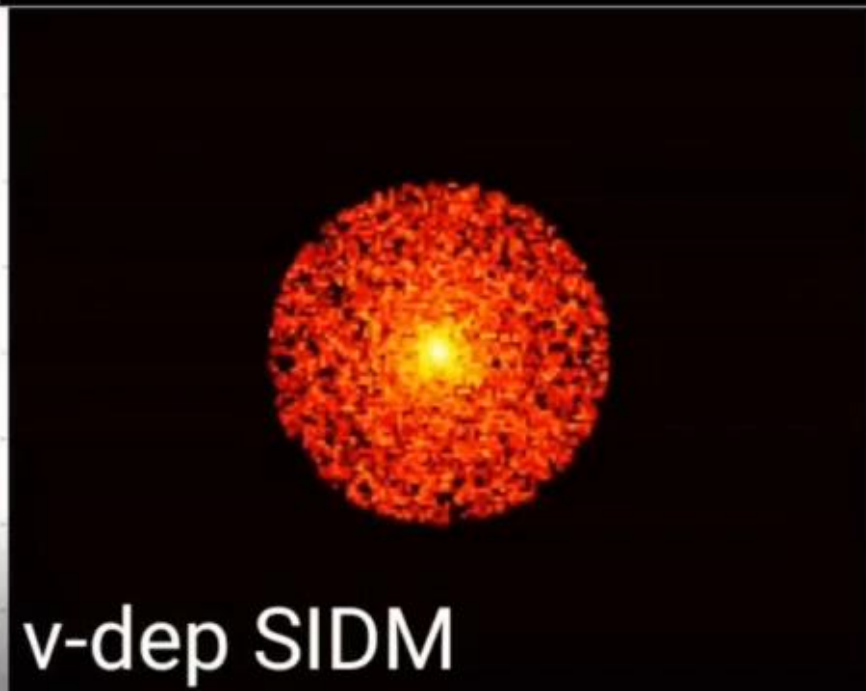
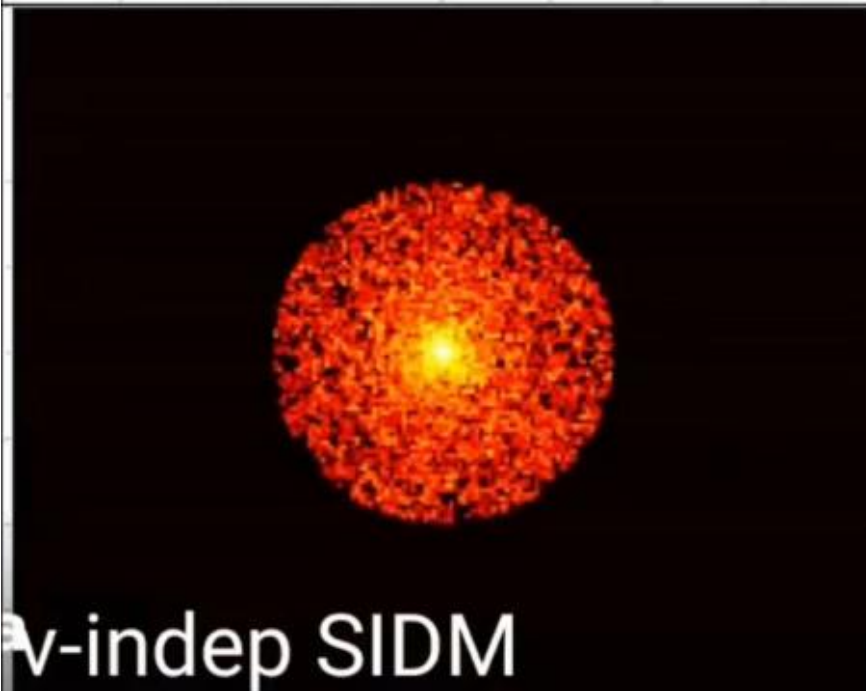
Particle Number: 5000

Unit time step: 0.01 Gyr

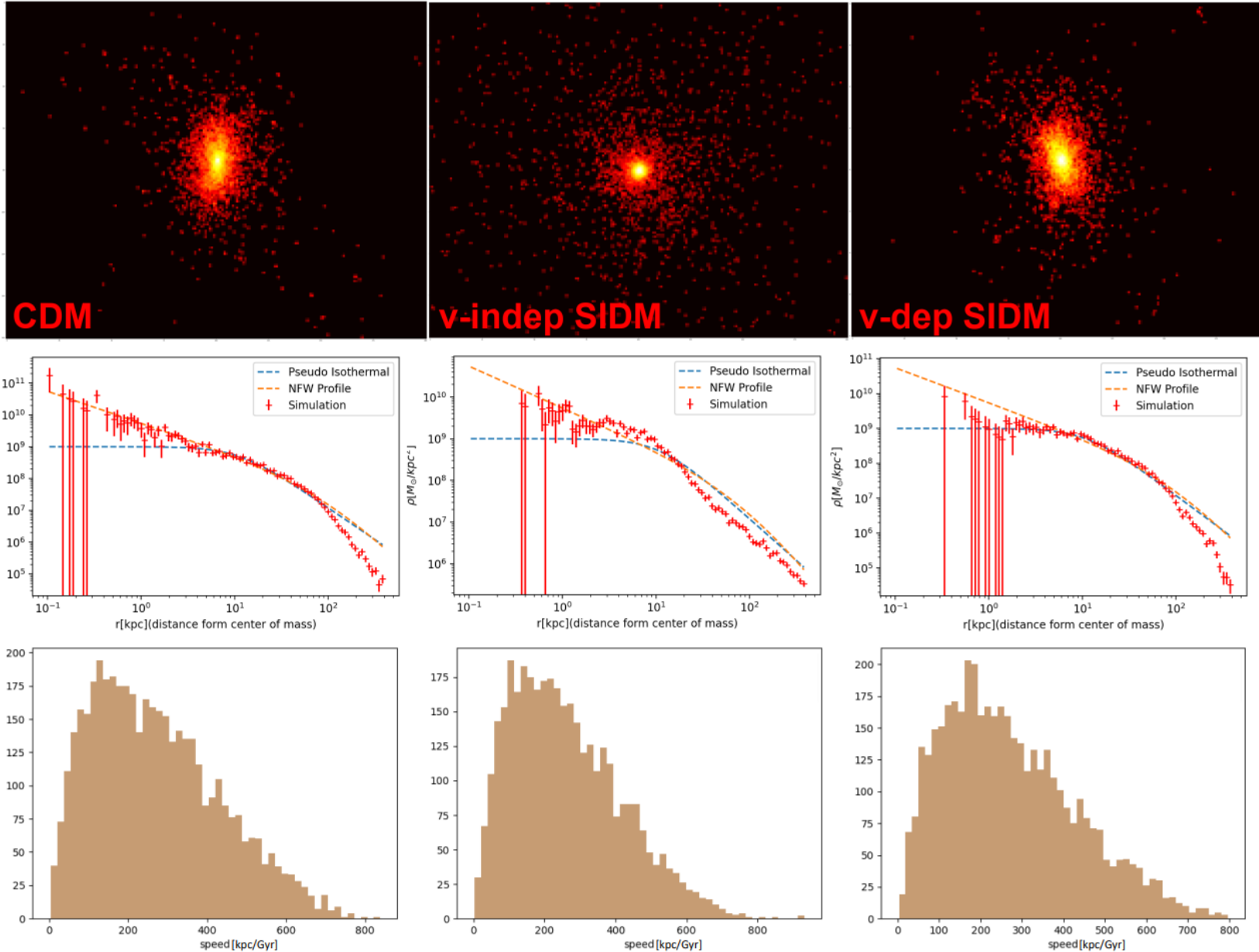
Total simulation time: 5 Gyr



00.01Gyr



Simulation Results



References

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- Sean Tulin, & Hai-Bo Yu (2018). Dark matter self-interactions and small scale structure. *Physics Reports*, 730, 1–57.
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