

# Theory uncertainties in the extraction of $\alpha_S$ from the $Z p_T$ spectrum

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Freiburg, Germany

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arXiv:24xx.yyyy  
with T. Cridge and F. Tackmann



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# Plan of the talk

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- » Theory uncertainty with Theory Nuisance Parameters (TNPs)
- » Theory uncertainties in the extraction of  $\alpha_S$  from the  $Z p_T$  spectrum
  - Perturbative uncertainties
  - Nonperturbative uncertainties
- » Conclusions

# But why?

A precise determination of the strong-coupling constant from the recoil of Z bosons with the ATLAS experiment at  $\sqrt{s} = 8$  TeV

A precise measurement of the Z-boson double-differential transverse momentum and rapidity distributions in the full phase space of the decay leptons with the ATLAS experiment at  $\sqrt{s} = 8$  TeV

[arXiv:2309.12986, 2309.09318]

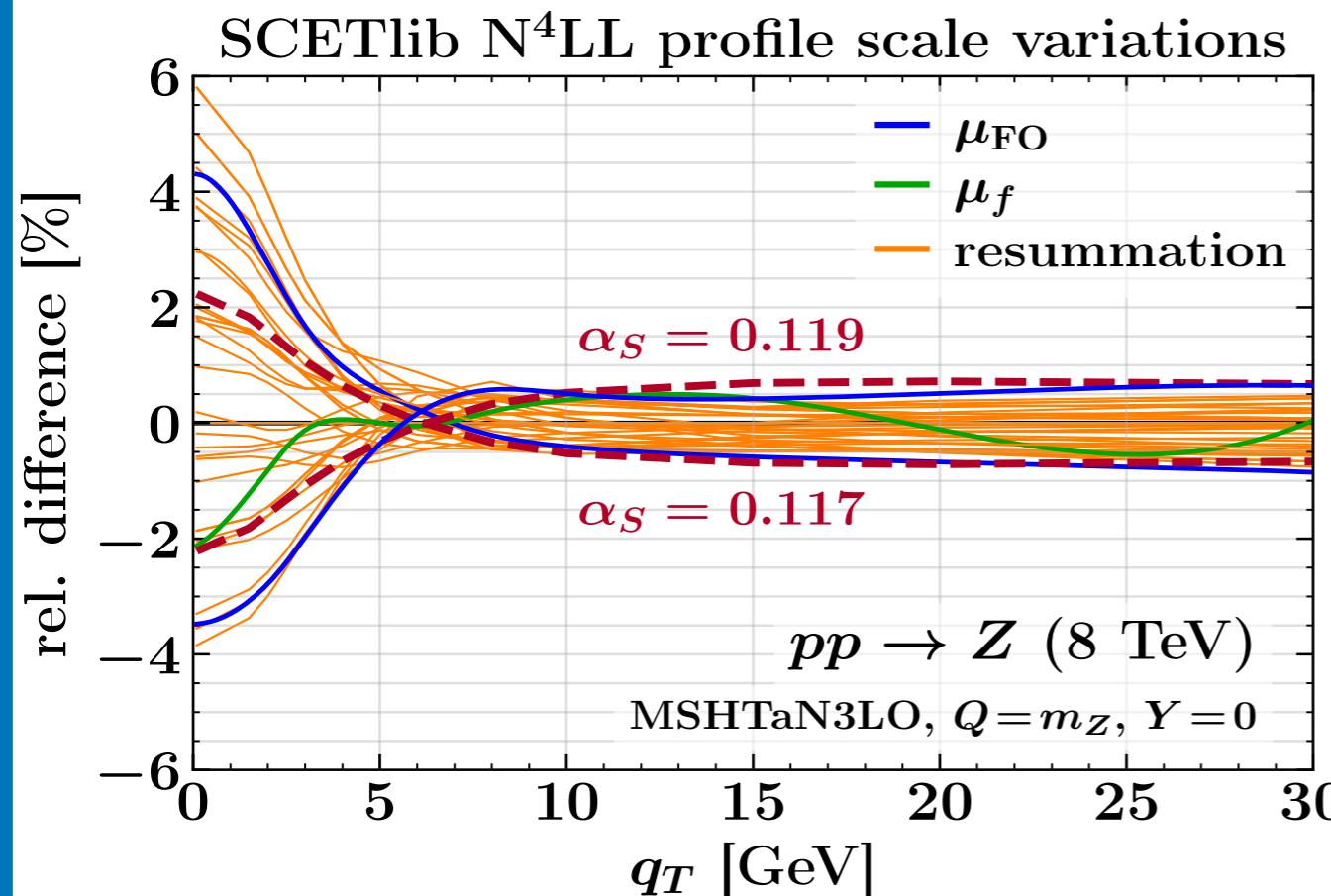
***super* precise ATLAS measurement of  $Z p_T$  spectrum**

no fiducial cuts, cross-section from  $Z \rightarrow ll$  full phase space

$$\alpha_S(m_Z) = 0.1183 \pm 0.0009$$

	In units of $10^{-3}$	
Experimental uncertainty	$\pm 0.44$	
PDF uncertainty	$\pm 0.51$	
Scale variation uncertainties	$\pm 0.42$	
Matching to fixed order	0	-0.08
Non-perturbative model	+0.12	-0.20
Flavour model	+0.40	-0.29
QED ISR		$\pm 0.14$
$N^4LL$ approximation		$\pm 0.04$
Total	+0.91	-0.88

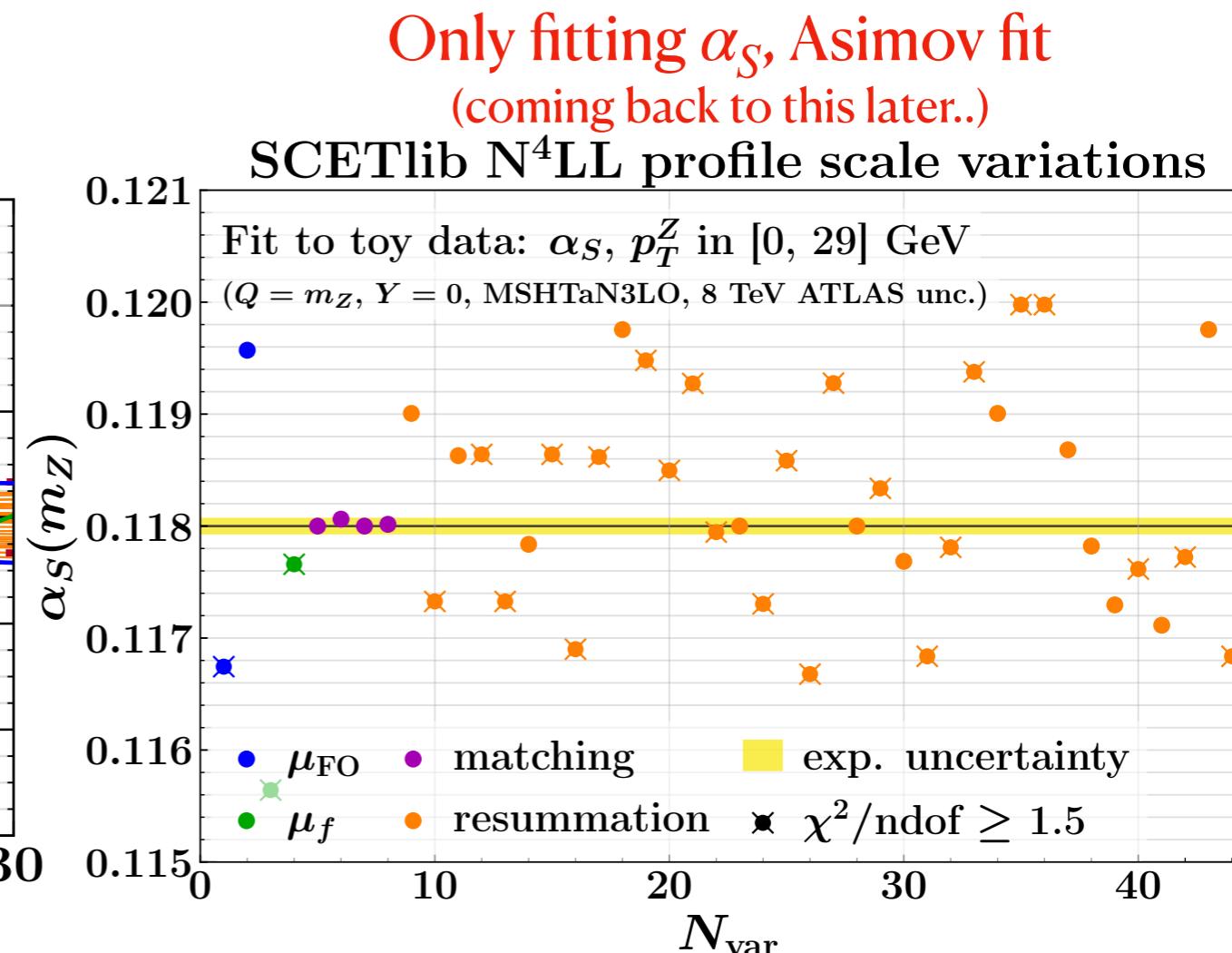
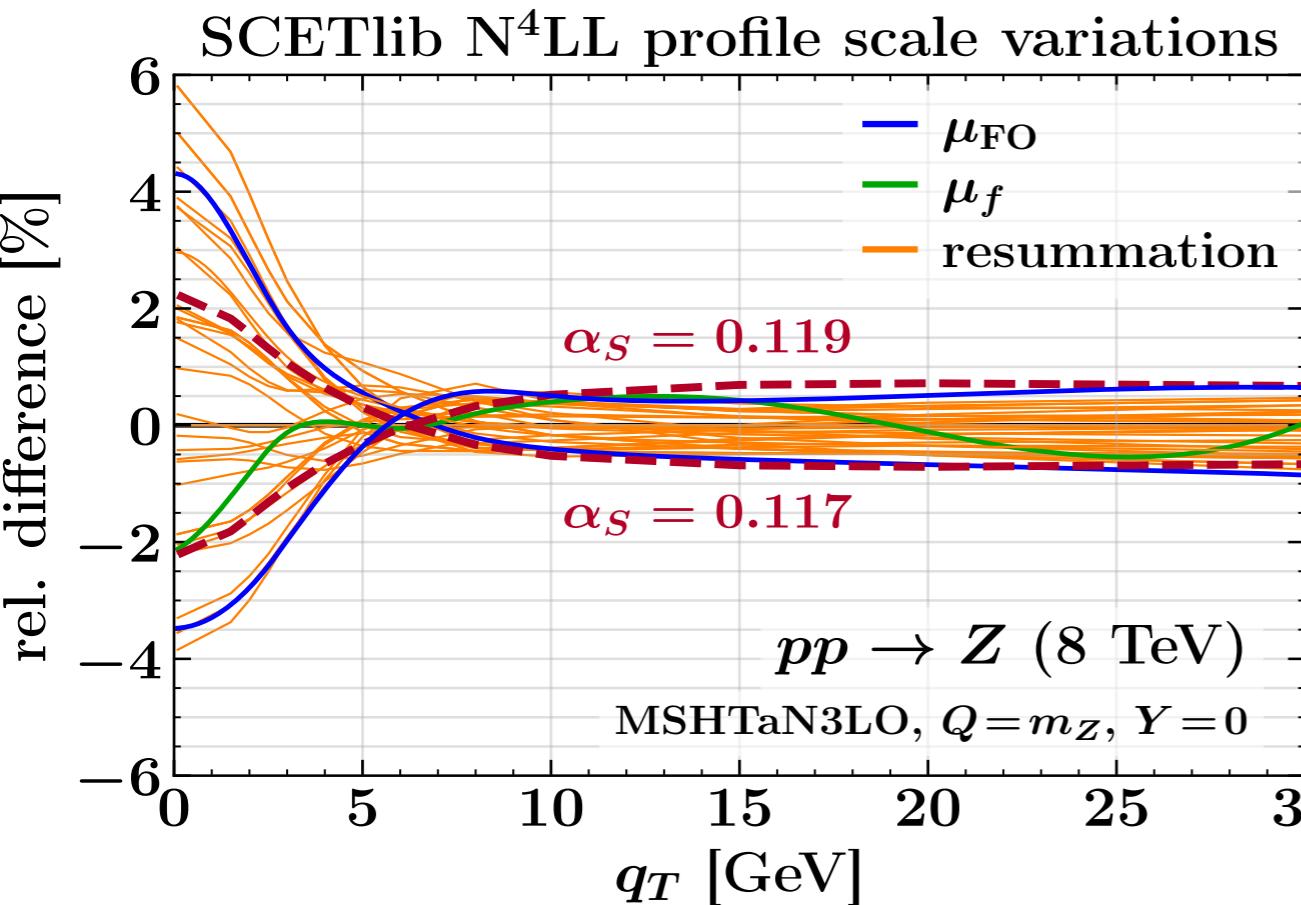
# Naive approach: scale variations



➤  $\alpha_s$  sensitivity in the  $q_T$  spectrum is a shape effect → theory correlations are crucial

\* uncertainties in units of  $10^{-3}$

# Naive approach: scale variations and scanning



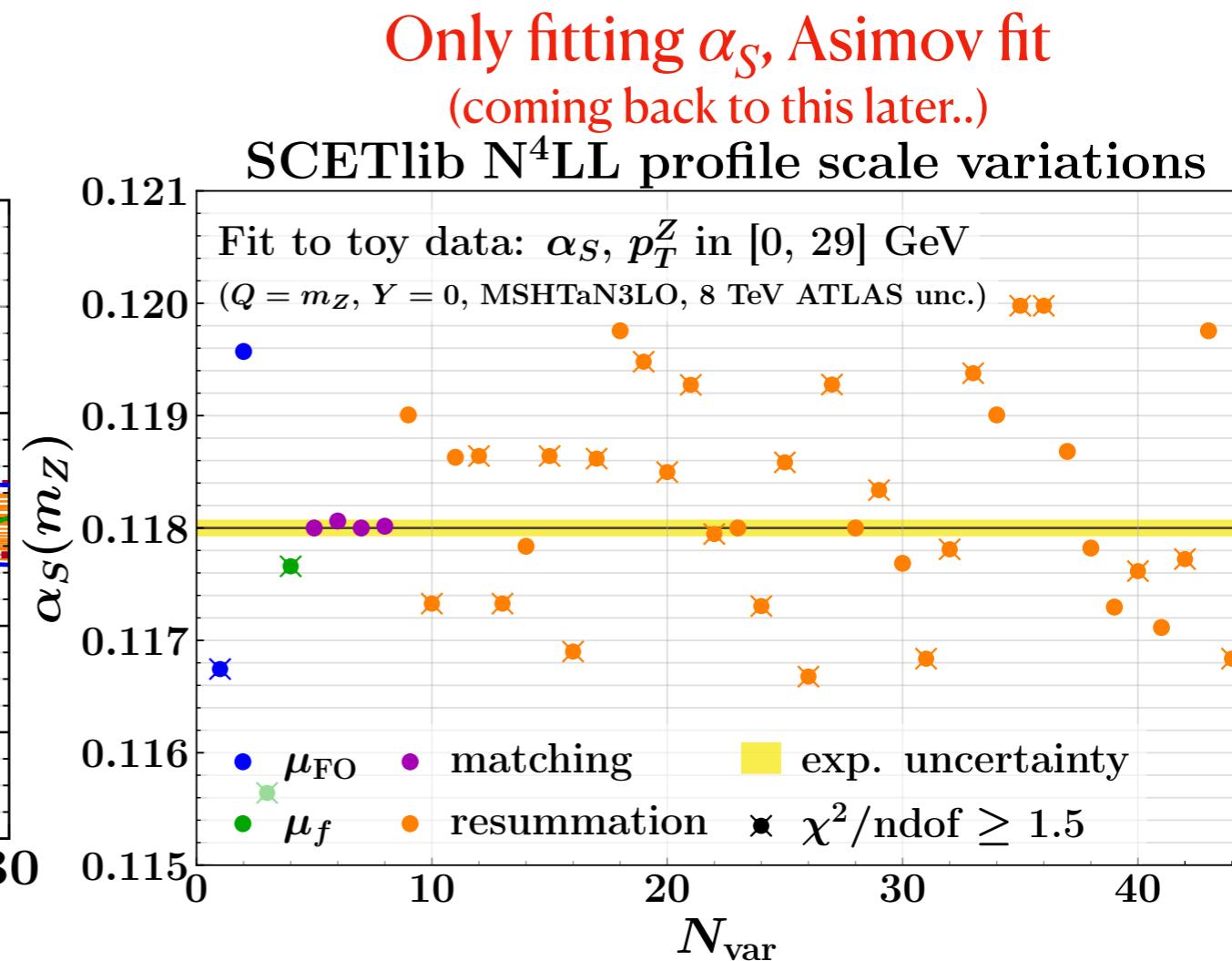
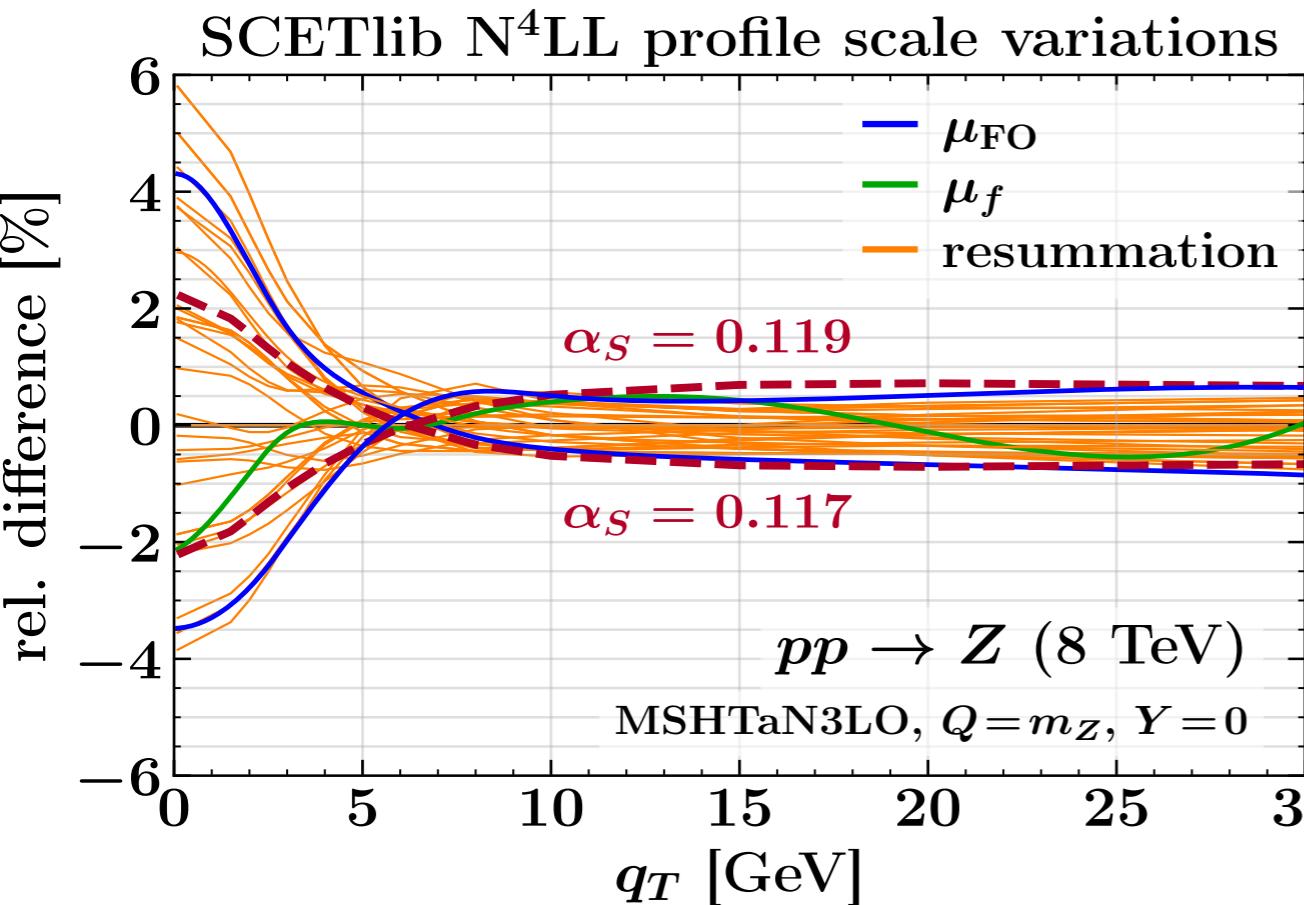
- $\alpha_S$  sensitivity in the  $q_T$  spectrum is a shape effect → theory correlations are crucial
- Each variation is a 100 % (anti)correlated correlation model, strongly impacts the result:

Sum in quadrature:  $\Delta_{\text{total}} = \sqrt{\Delta_{\text{FO}}^2 + \Delta_{\text{resum}}^2 + \Delta_{\text{match}}^2} \sim 2.6$  [neglecting  $\mu_f$ ]

Envelope:  $\Delta_{\text{total}} \sim 2.1$

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Envelope:  $\Delta_{\text{total}} \sim 2.1$

scale variations are not sufficient! can we do better?

\* uncertainties in units of  $10^{-3}$

# Theory Nuisance Parameters (TNPs)

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Consider a series expansion in a small parameter  $\alpha$ :

$$f(\alpha) = \underbrace{f_0 + \alpha f_1 + \alpha^2 f_2}_{\text{NNLO}} + \alpha^3 f_3 + \alpha^4 f_4 + \mathcal{O}(\alpha^5)$$

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- 2 Parametrized and include the leading source of uncertainty:

$$f^{\text{pred}}(\alpha) = f_0 + \alpha f_1 + \alpha^2 f_2 + \alpha^3 f_3(\theta_3) + \mathcal{O}(\alpha^4) \rightarrow \text{named N}^{2+1}\text{LO}$$

using theory nuisance parameters  $\theta_n$ ;

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using theory nuisance parameters  $\theta_n$ ;

- the expansion converges, so the next order [ $f_4$ ] is not yet relevant;
- once  $f_3$  is known (or strongly constrained), include the next order;
- $\theta_n$  have physical value, true parameters

# Theory Nuisance Parameters (TNPs)

---

3 How to *define* these  $\theta_n$ ?

- simplest case:  $f_3(\theta_3) \equiv \theta_3$
- better: account for the internal structure of  $f_3$   
(given the process: partonic channels, color, ... )

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Consider the  $p_T$  spectrum, leading power  $p_T$  dependence is known to all orders:

$$p_T \frac{d\sigma}{dp_T} = \left[ H \times B_a \otimes B_b \otimes S \right] (\alpha_S, L \equiv \ln p_T/m_Z) + \mathcal{O}\left(\frac{p_T^2}{m_Z^2}\right)$$

$F = \{H, B, S\}$  solution to RGE equations

$$F(\alpha_S, L) = F(\alpha_S) \exp \int_0^L dL' \left\{ \Gamma[\alpha_S(L')] L' + \gamma_F[\alpha_S(L')] \right\}$$

boundary conditions      anomalous dimensions

# TNPs in the $p_T$ spectrum

Have three independent scalar perturbative series, for  $N^{2+1}LL$ :

$$F(\alpha_S) = 1 + \sum_{n=1} \left( \frac{\alpha_S}{4\pi} \right)^n F_n \longrightarrow F(\alpha_S) = 1 + \frac{\alpha_S}{4\pi} F_1 + \left( \frac{\alpha_S}{4\pi} \right)^2 F_2(\theta_2^F)$$

$$\gamma(\alpha_S) = \sum_{n=0} \left( \frac{\alpha_S}{4\pi} \right)^{n+1} \gamma_n \quad \begin{array}{c} \nearrow \\ \searrow \end{array} \quad \begin{aligned} \gamma_F(\alpha_S) &= \frac{\alpha_S}{4\pi} \gamma_{F0} + \left( \frac{\alpha_S}{4\pi} \right)^2 \gamma_{F1} + \left( \frac{\alpha_S}{4\pi} \right)^3 \gamma_{F,2}(\theta_3^\gamma) \\ \Gamma(\alpha_S) &= \frac{\alpha_S}{4\pi} \Gamma_0 + \left( \frac{\alpha_S}{4\pi} \right)^2 \Gamma_1 + \left( \frac{\alpha_S}{4\pi} \right)^3 \Gamma_2 + \left( \frac{\alpha_S}{4\pi} \right)^4 \Gamma_3(\theta_3^\gamma) \end{aligned}$$

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➤ Pulling out known color factor:

$$F_n(\theta_n) = 4C_r(4C_A)^{n-1}(n-1)! \theta_n^F \qquad C_r \text{ leading color factor}$$

$$\gamma_n(\theta_n) = 2C_r(4C_A)^n \theta_n^\gamma \qquad C_A^{n-1} \text{ leading } n\text{-loop color factor}$$

# TNPs in the $p_T$ spectrum

4 How to vary  $\theta_n$ ?

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factorizing out  $\left(\frac{\alpha_S}{4\pi}\right)^n$     1    +4.9    -24.0    -4065.5    -123979.0     $C_{gg}$

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factorizing out  $(n-1)!$       1      +0.4      -0.2      -1.2      -1.0

$\rightarrow \theta_n \sim \mathcal{O}(1)$

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factorizing out $\left(\frac{\alpha_S}{4\pi}\right)^n$	1	+4.9	-24.0	-4065.5	-123979.0	$C_{gg}$
	1	-8.5	-48.6	-1386.7	-42014.9	$C_{q\bar{q}}^V$
factorizing out $4^n$	1	+1.2	-1.5	-63.5	-484.3	
	1	-2.1	-3.0	-21.7	-164.1	
factorizing out $C_r C_A^{n-1}$	1	+0.4	-0.2	-2.4	-5.9	
	1	-1.6	-0.8	-1.8	-4.6	
factorizing out $(n-1)!$	1	+0.4	-0.2	-1.2	-1.0	
	1	-1.6	-0.8	-0.9	-0.8	

$\rightarrow \theta_n = 0 \pm \mathcal{O}(1)$

# TNPs in the $p_T$ spectrum

4 How to vary  $\theta_n$ ?

$$\gamma_n(\theta_n) = 2C_r(4C_A)^n \theta_n^r$$

AD:  $\gamma(\alpha_S) = \frac{\alpha_S}{4\pi} \gamma_1 + \left(\frac{\alpha_S}{4\pi}\right)^2 \gamma_2 + \left(\frac{\alpha_S}{4\pi}\right)^3 \gamma_3 + \left(\frac{\alpha_S}{4\pi}\right)^4 \gamma_4 + \left(\frac{\alpha_S}{4\pi}\right)^5 \gamma_5 + \mathcal{O}(\alpha_S^6)$

	$\left(\frac{\alpha_S}{4\pi}\right)^n$	4.0	56.2	474.9	2824.8	42824.1	$-\gamma_m/2$
factorizing out $\left(\frac{\alpha_S}{4\pi}\right)^n$		5.3	36.8	239.2	141.2	70000.0	$\Gamma_{\text{cusp}}$
		2.0	7.0	14.8	22.1	83.6	
factorizing out $2 \cdot 4^n$		2.7	4.6	7.5	1.1	136.7	
		1.5	1.8	1.2	0.6	0.8	
factorizing out $C_r C_A^n$		2.0	1.5	0.6	0.03	1.3	

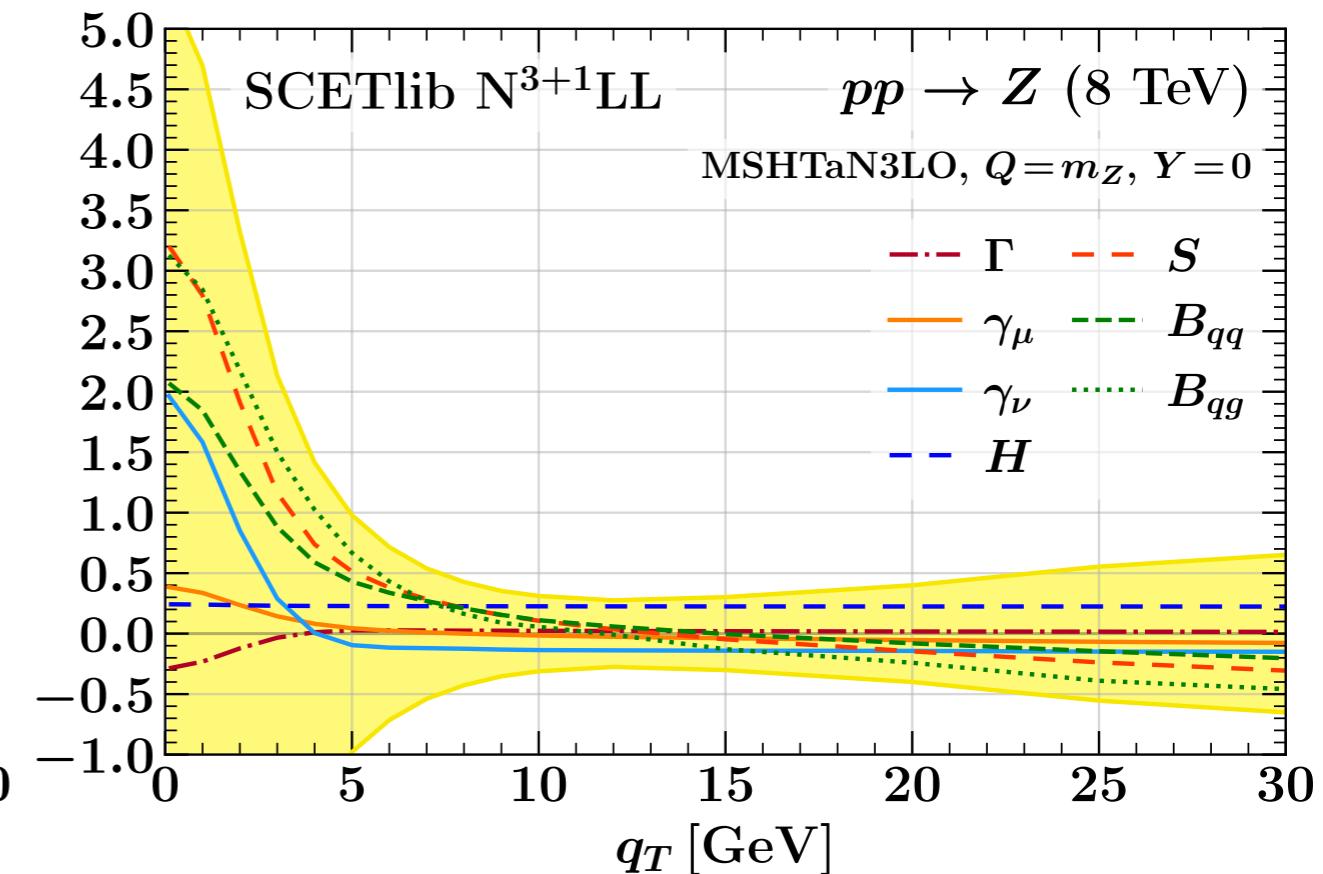
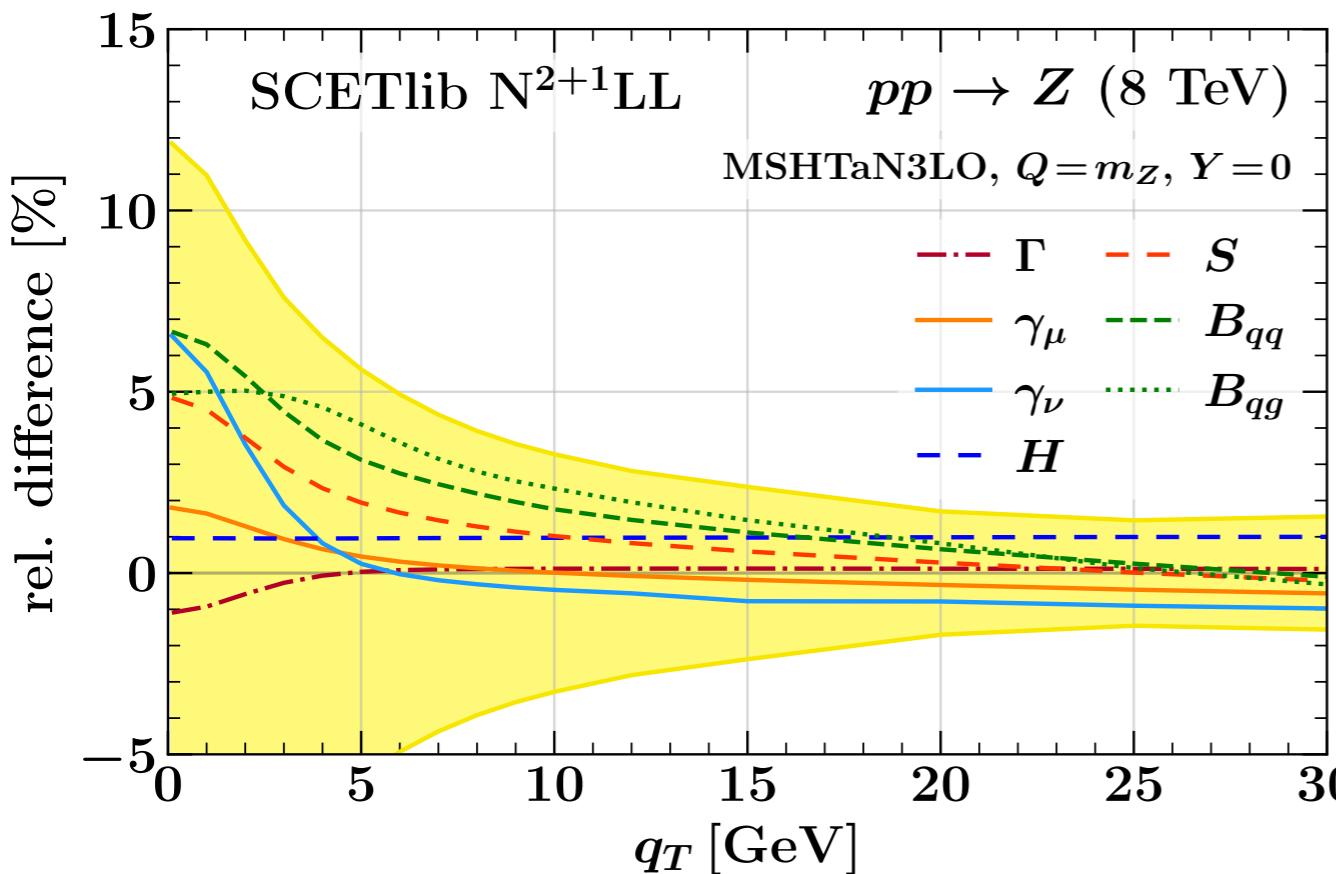
➤ Iterate this game with different functions and do some “statistics”:

→  $\theta_n = 0 \pm 1$

look at known  $n$ -loop coefficients from population sample [here](#)

# Application of TNPs to $Z p_T$ spectrum

Remember  $N^{k+1}\text{LL}$ :  $N^{k+1}\text{LL}$  resummation structure + highest-order boundary conditions/ anomalous dim. as TNPs



- Varying each  $\theta_i$  independently: correctly describe correlation across  $q_T$
- Add in quadrature for the total uncertainty
- For the beams  $B_{qj}$ :  $f_n = (0 \pm 1.5) \times f_n^{\text{true}}$ , DGLAP splitting functions not varied

many other interesting plots [here](#)

**Going towards  $\alpha_S$**

# Asimov test fitting $\alpha_S(m_Z)$ from $Z p_T$

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Use TNPs to study the expected uncertainty/sensitivity on  $\alpha_S$  on toy data (**Asimov test**)

## Our theory inputs:

- SCETlib only  $N^{3+1}LL$  resummed contribution  
[default central scales and variations, no mass corrections and nonsingular power corrections\*]

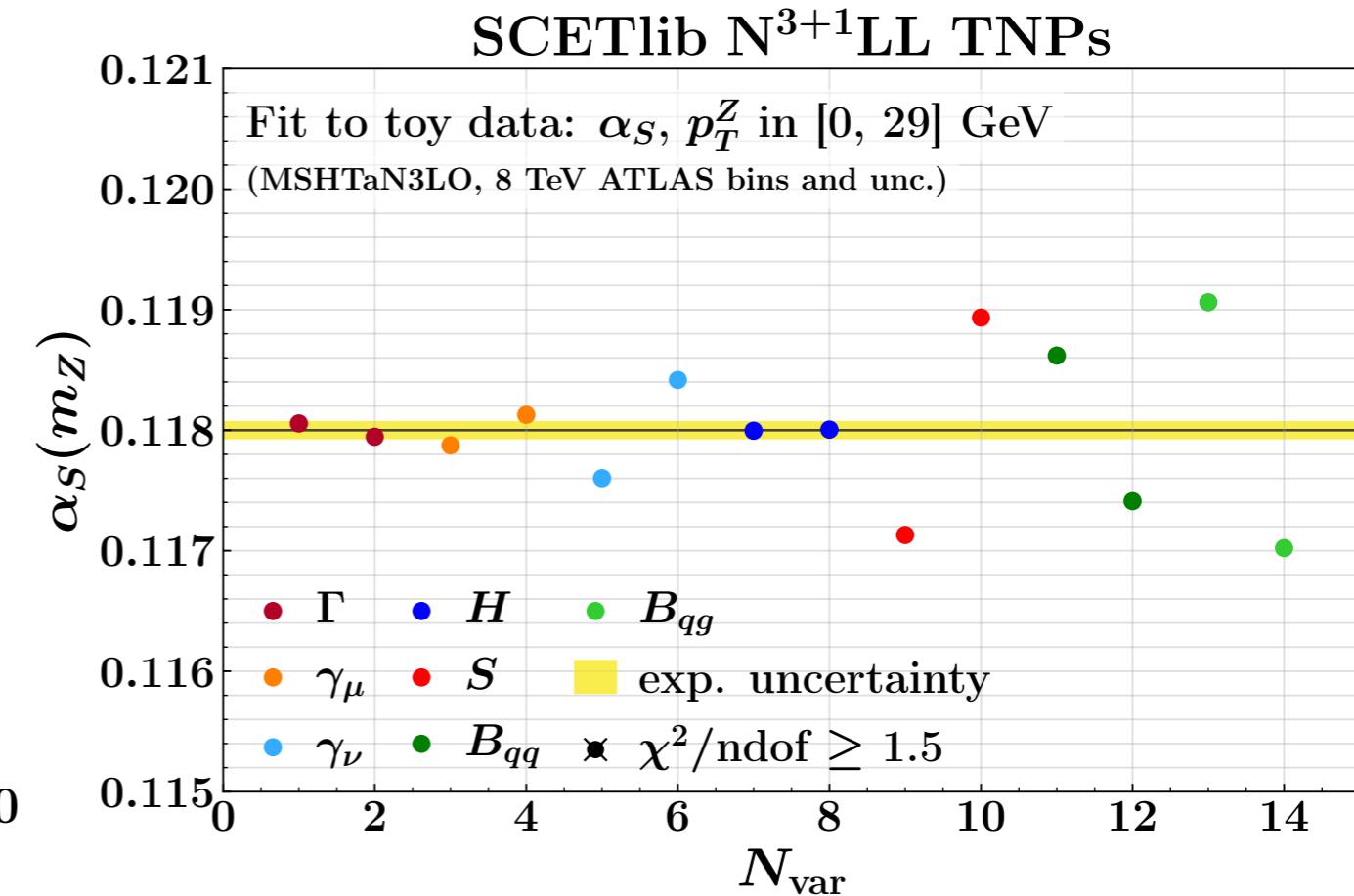
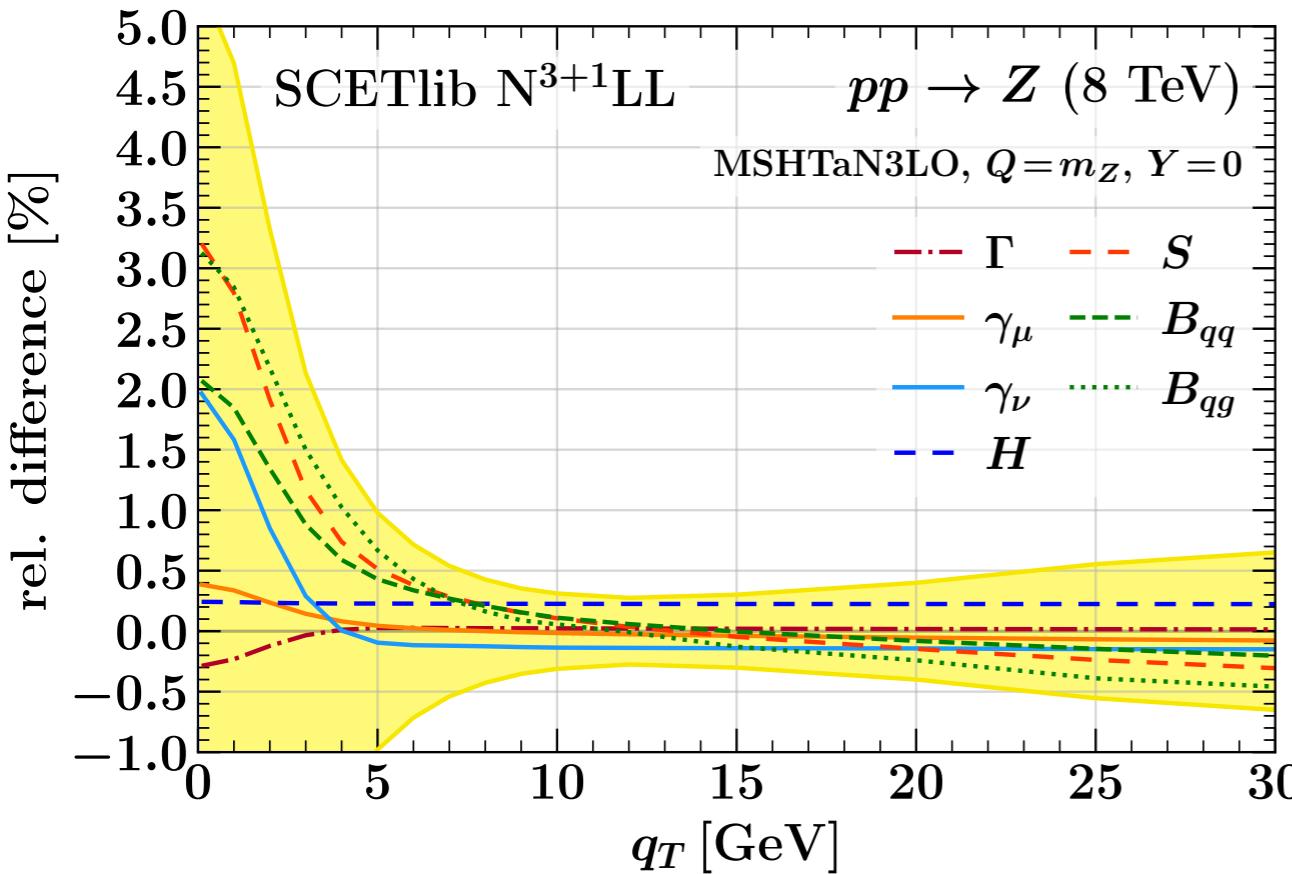
## Our toy data:

- Data defined as central theory prediction [ $\alpha_S = 0.118$ ]  
[fixed nonp. params, [MSHT20aN3LO](#) PDF set]
- 72 data points in ATLAS binning,  
9  $q_T$  bins in [0,29] GeV for each 8  $Y$  bin in [0.0,3.6]  
[integrated in  $q_T$ ,  $Y$  and  $Q$ ]
- Using ATLAS exp. uncertainties and complete correlations for all 72 bins
- Using Minuit as minimizer for the fit

[\*nonsingular p.c. don't affect our conclusions, obviously necessary when fitting against real data]

# Scanning TNPs

Only fitting  $\alpha_S$



Repeat fit for each TNP variation, using TNPs at N<sup>3+1</sup>LL;  
still does not let the fit decide between moving  $\alpha_S$  or theory

TNPs correctly account for their correlations  
being an independent source of uncertainty

→ sum in quadrature:  $\Delta_{\text{total}} = 1.6$

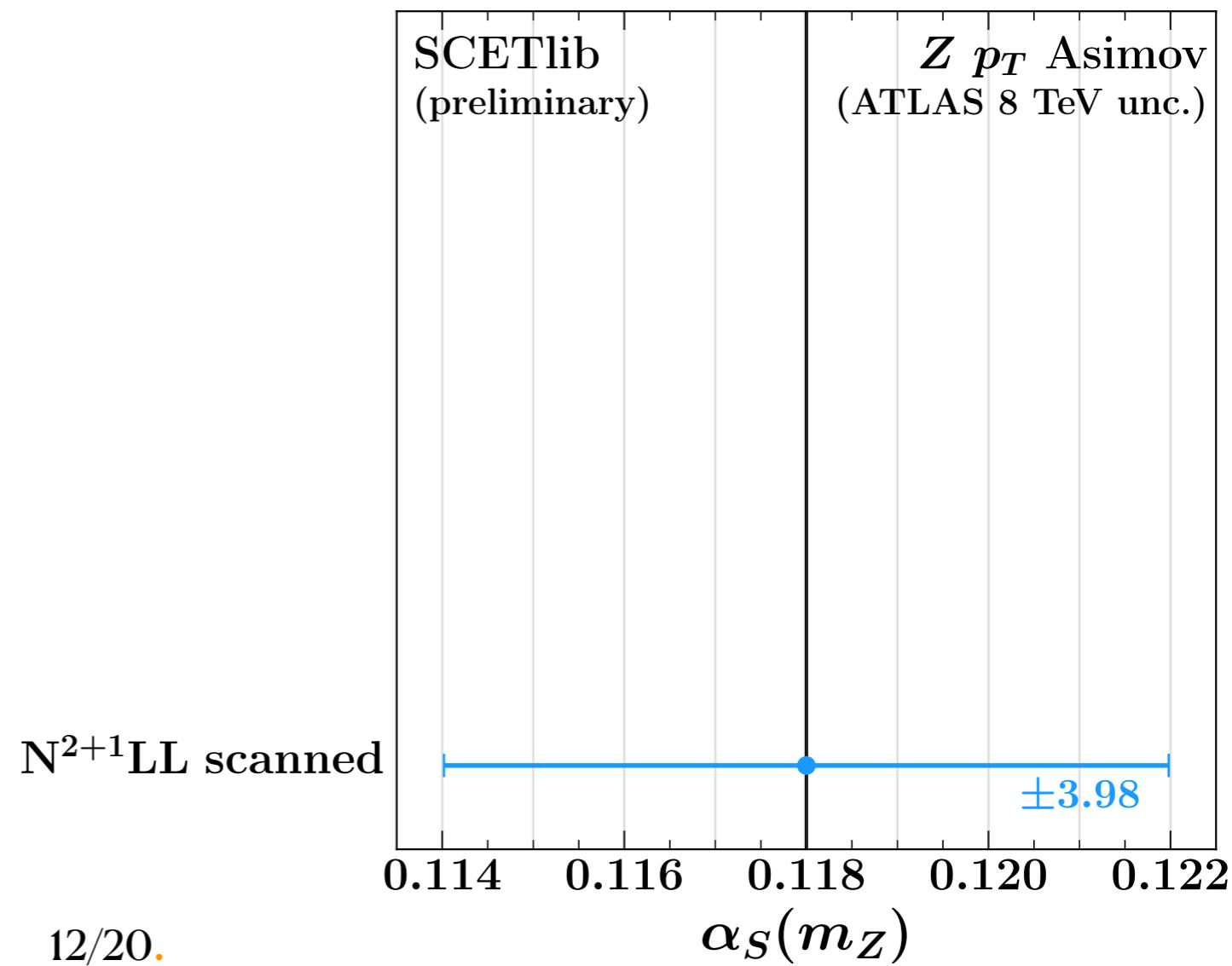
\* uncertainties in units of  $10^{-3}$

# Perturbative uncertainty in Asimov fit

**Scanning:** vary one TNP at a time and re-fit  $\alpha_S$

**Profiling:** fitting  $\alpha_S$  *together* with all TNPs (allow the fit to decide what to do)

➤ Fit  $N^{2+1}LL$  against  $N^{2+1}LL$  data



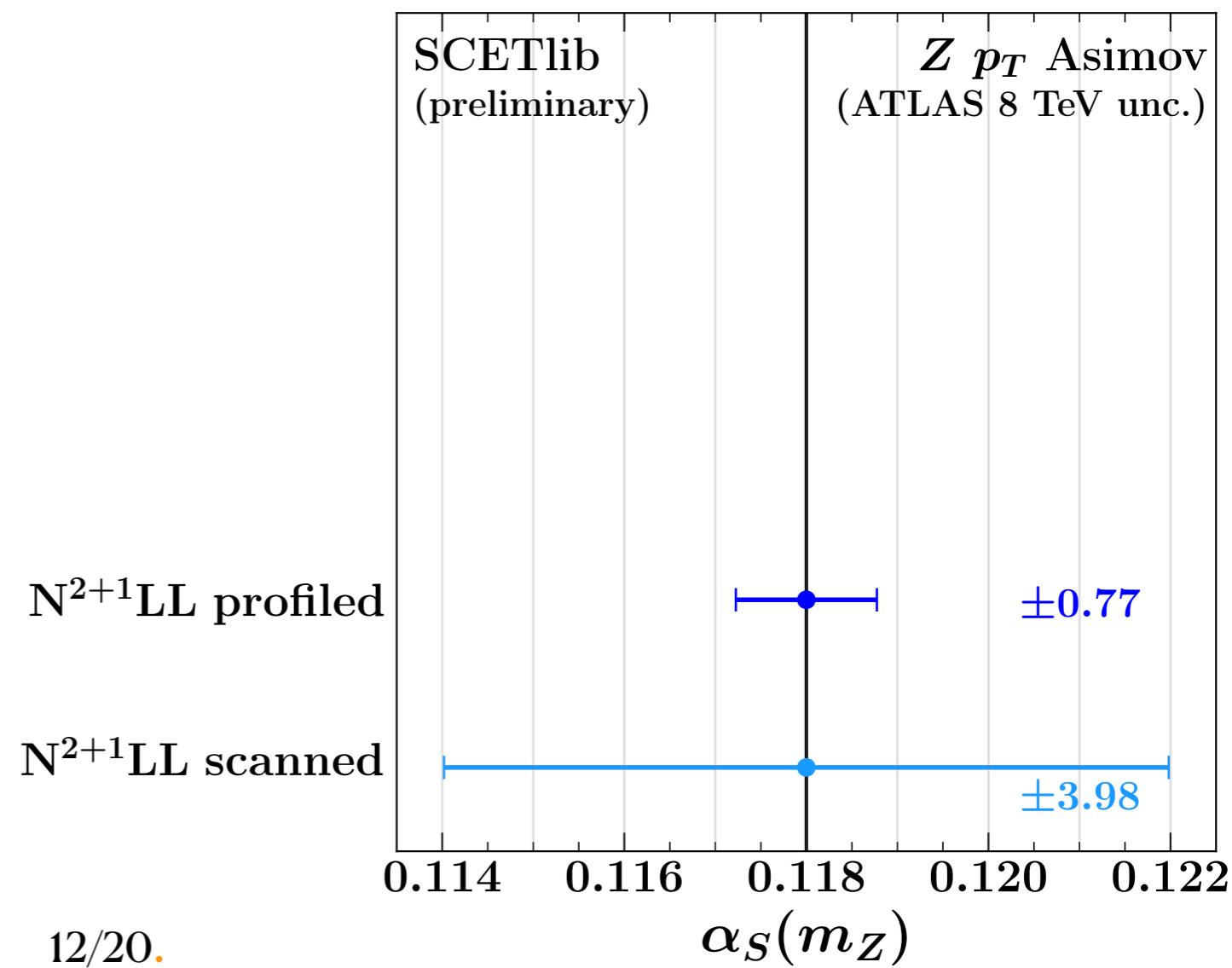
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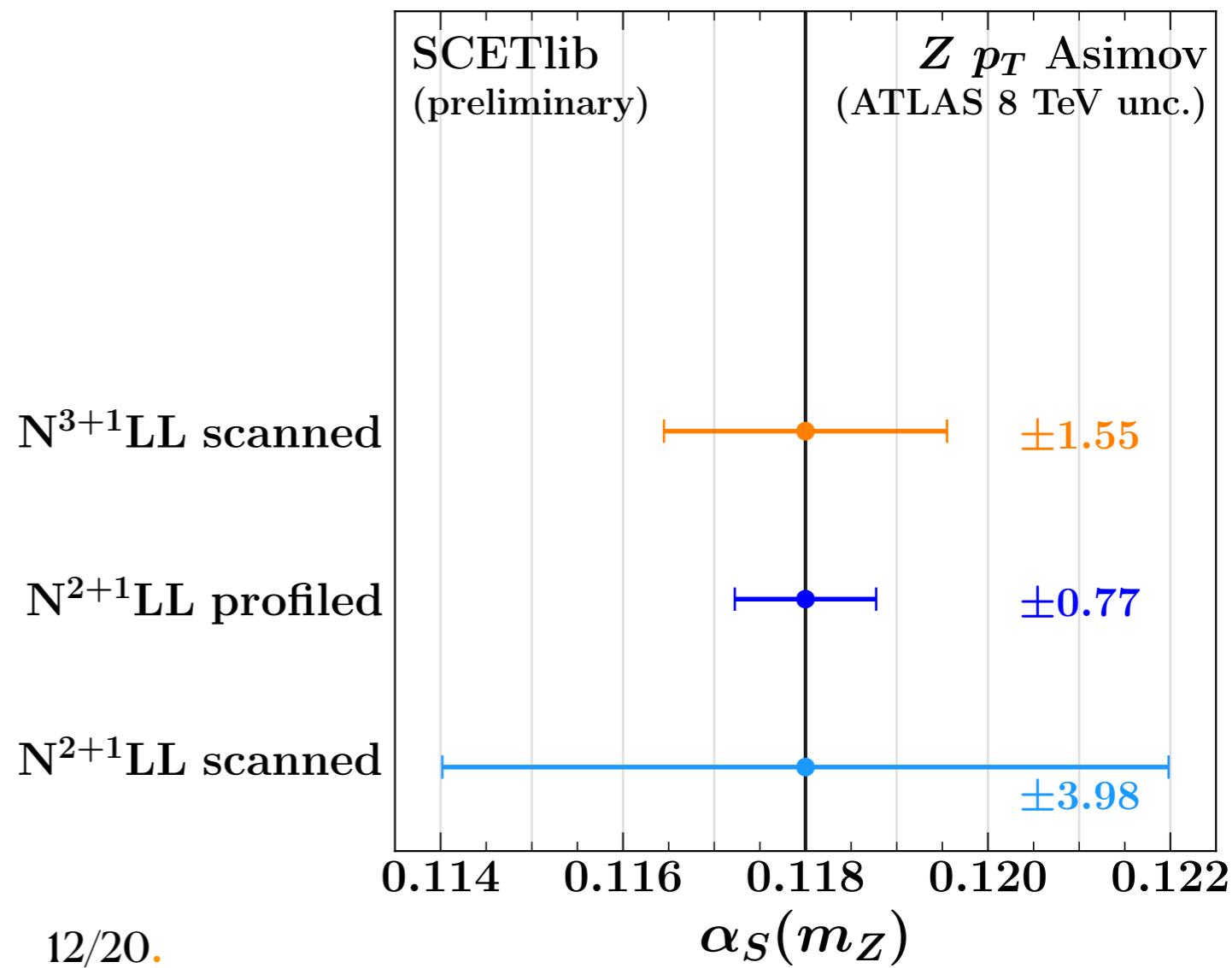
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- Fit  $N^{3+1}\text{LL}$  against  $N^{3+1}\text{LL}$  data



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# Perturbative uncertainty in Asimov fit

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**Profiling:** fitting  $\alpha_S$  *together* with all TNPs (allow the fit to decide what to do)

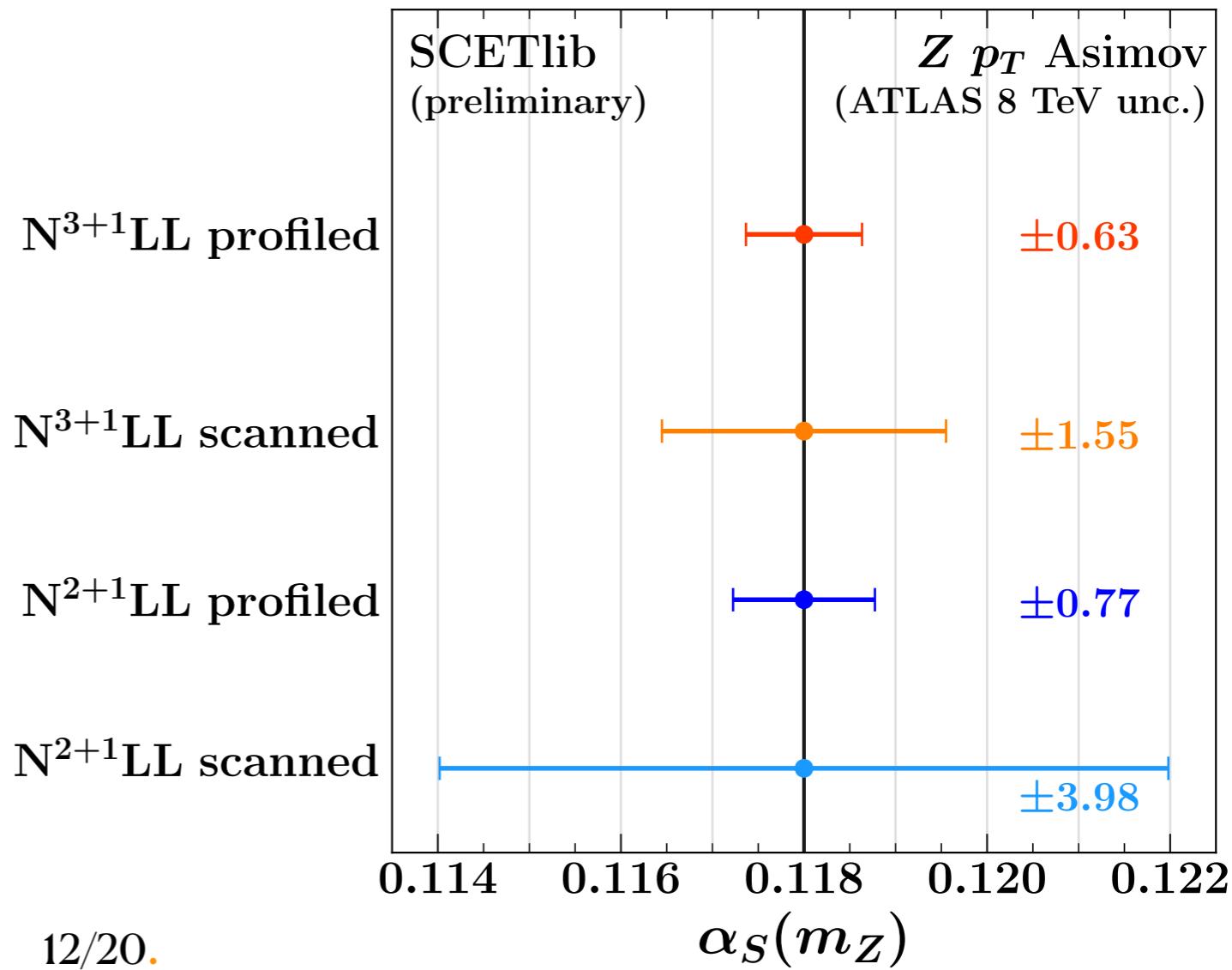
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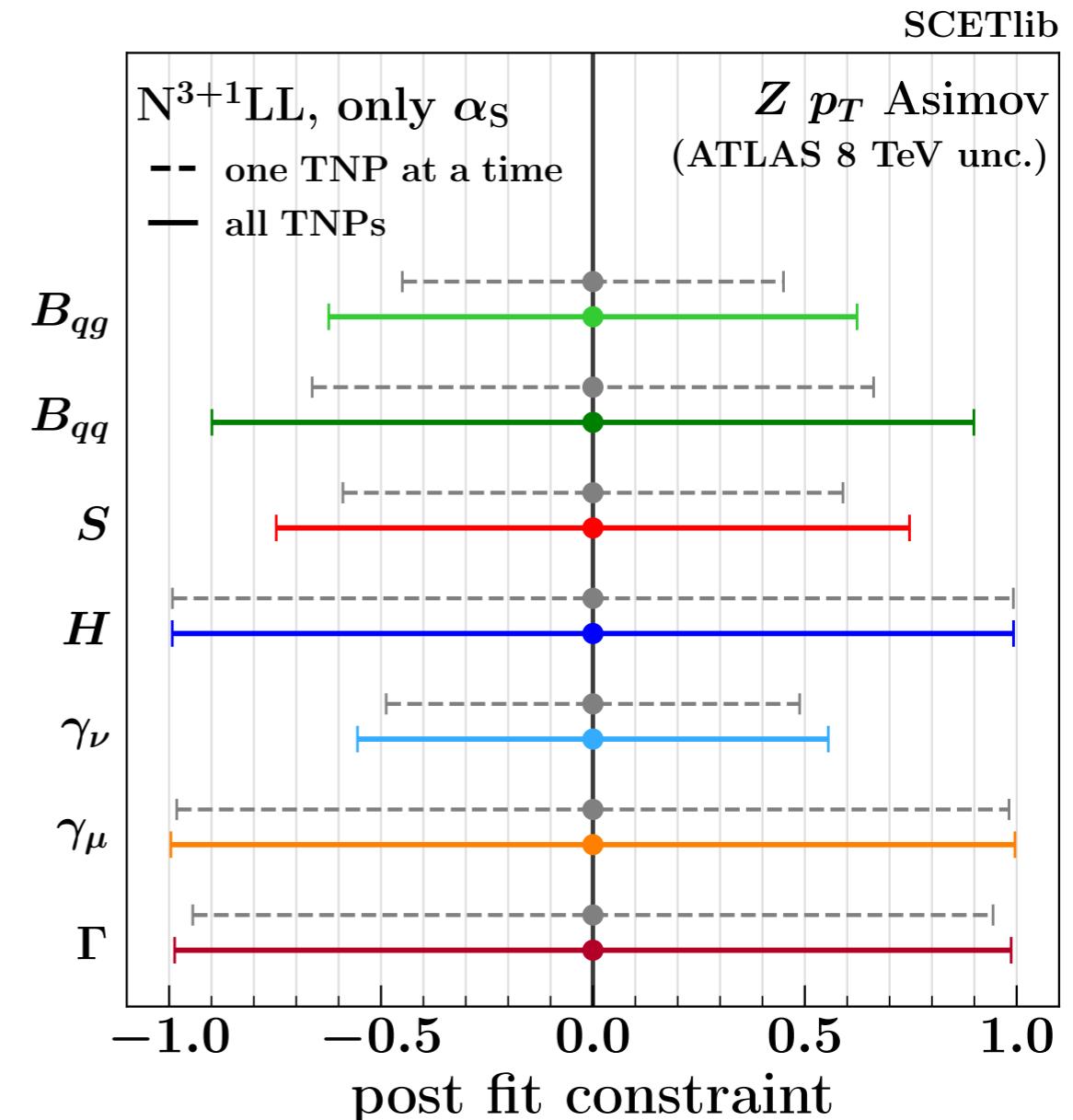
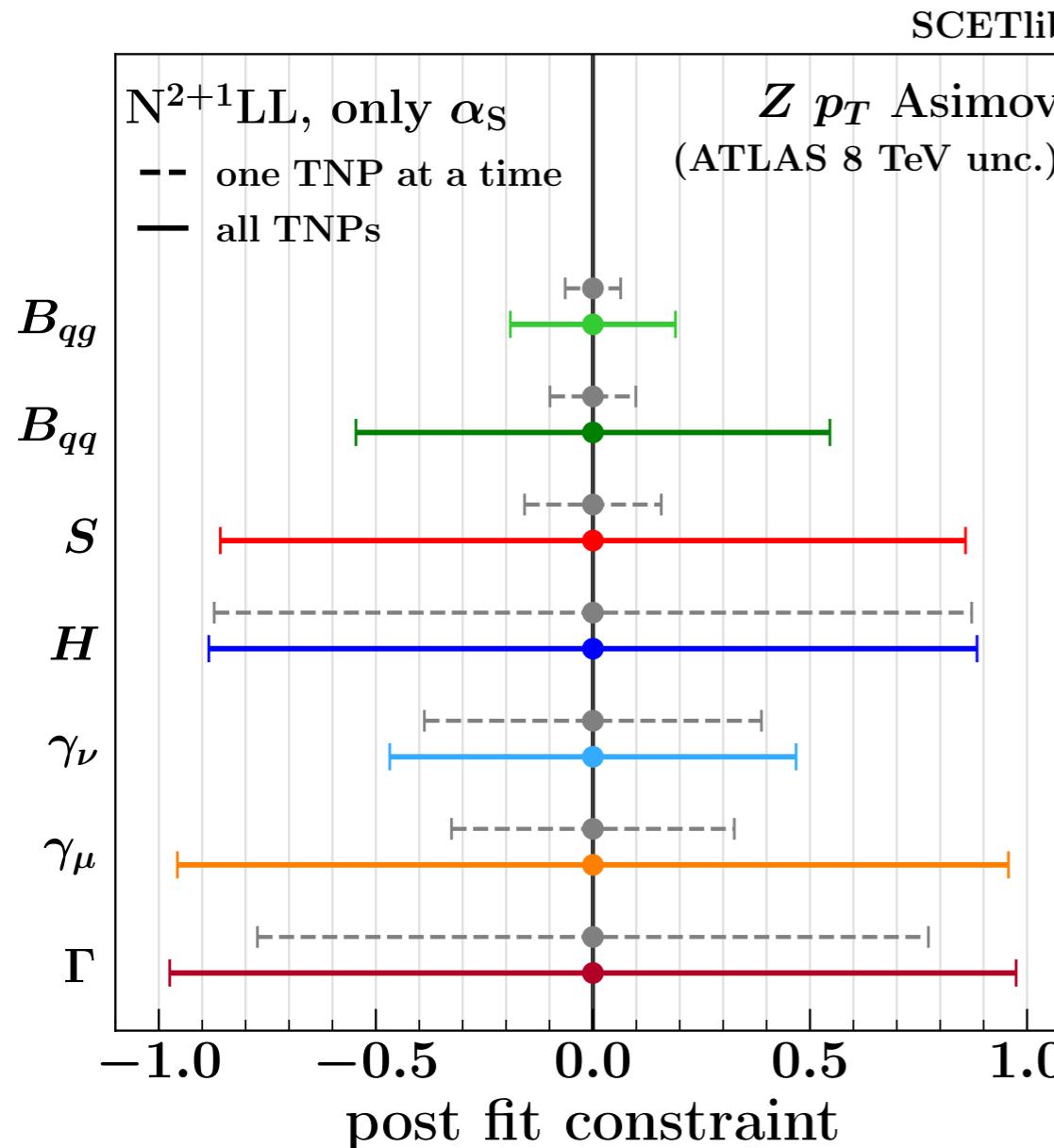
**profiling** constraints the TNPs allowing data to reduce the theory uncertainty!

still need to look at the TNPs pull plot to understand the post-fit uncertainty

\* uncertainties in units of  $10^{-3}$



# Constraints on TNPs



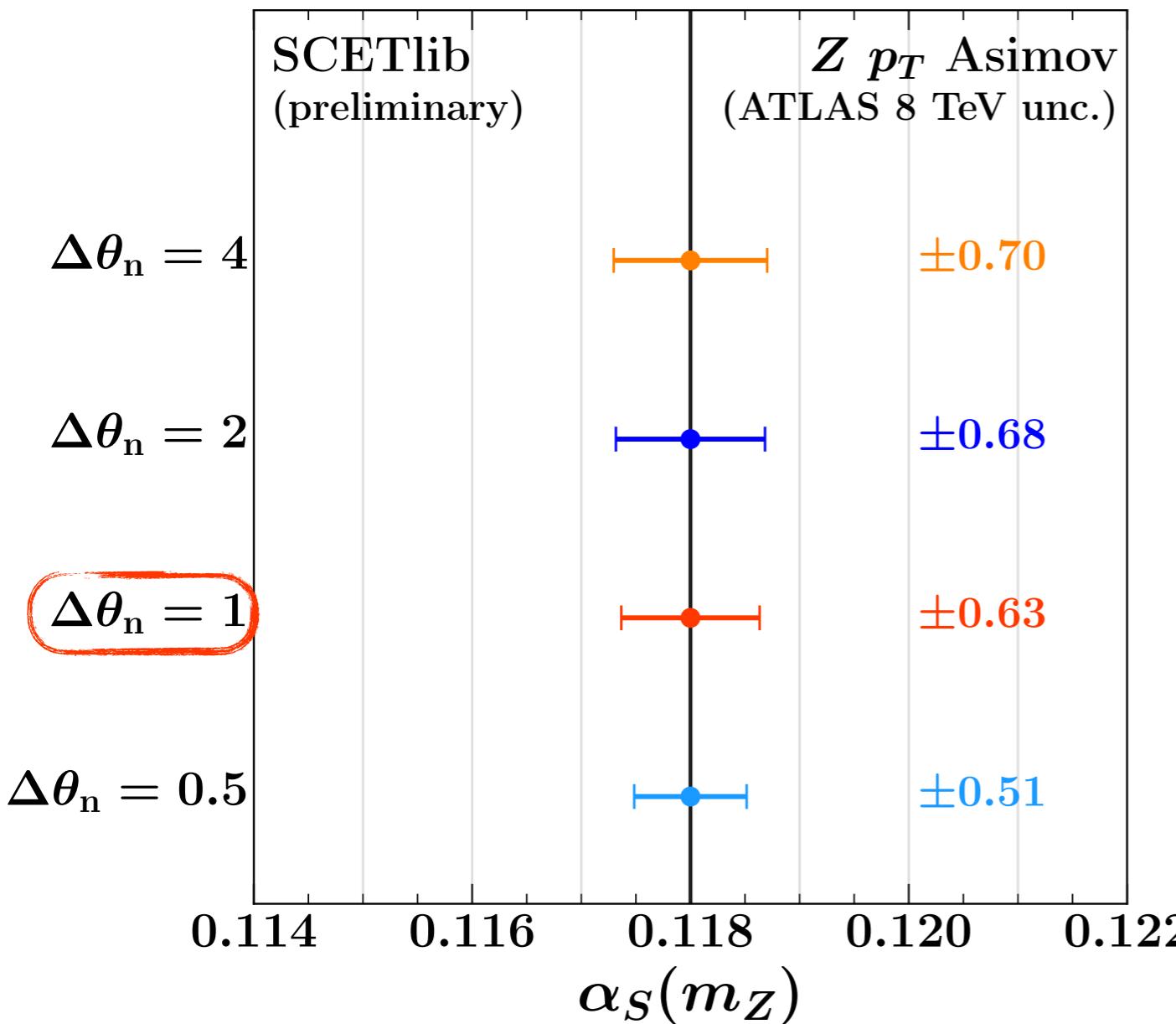
- $N^{2+1}\text{LL}$ : TNPs much more constrained than at  $N^{3+1}\text{LL}$
- If TNPs get strongly constrained, the next order becomes relevant for the uncertainty correlations!

# Different constraints on TNPs

What happens by changing the prior theory constraint?

Using now  $\theta_n = 0 \pm \Delta\theta_n$  with  $\Delta\theta_n = 0.5, 1, 2, 4$

Fit  $N^{3+1}\text{LL}$  against  $N^{3+1}\text{LL}$  data



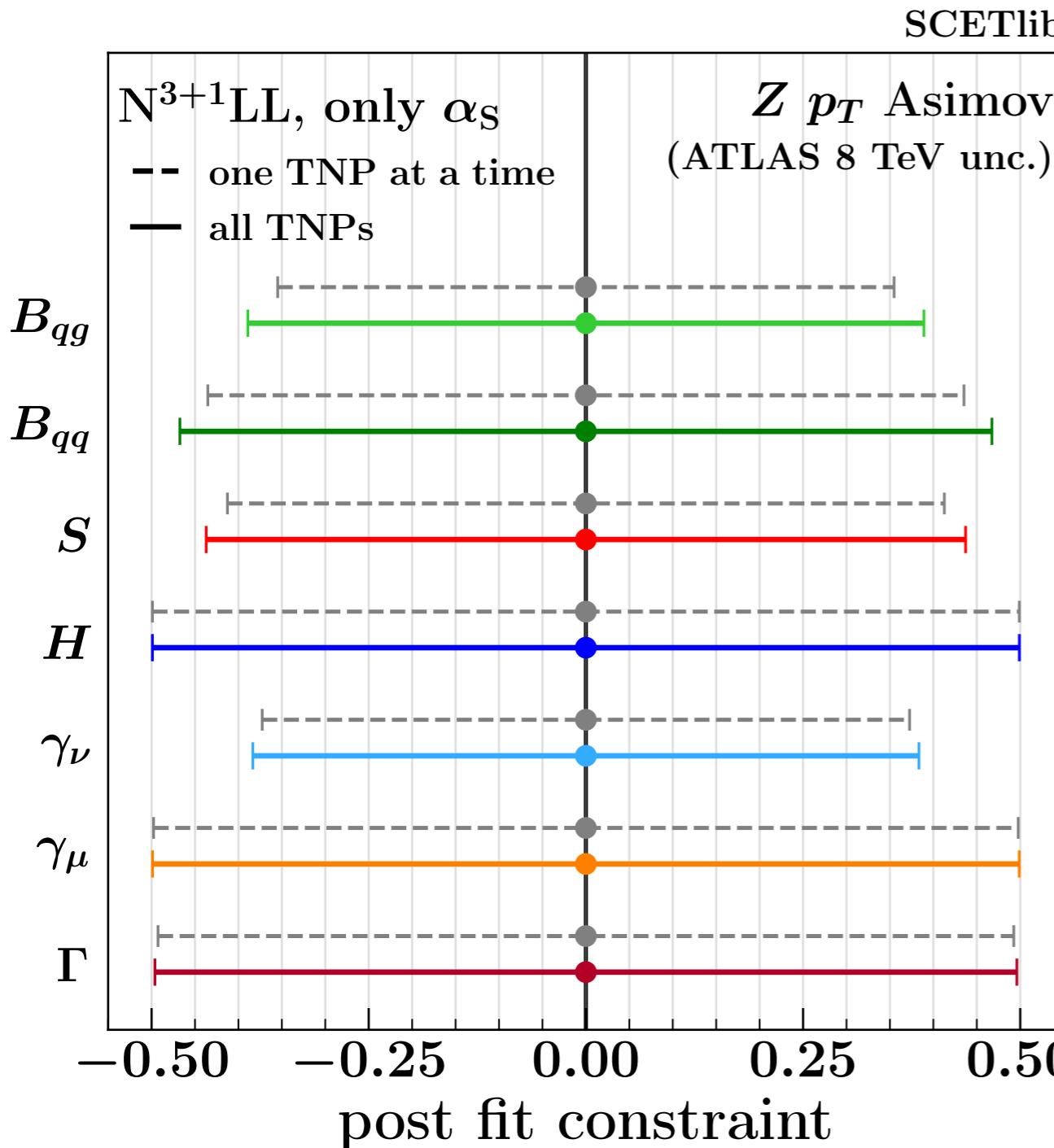
- The effect strongly depends on the power of the experimental constraint
- Only slight increase in the uncertainties when relaxing the TNPs constraint
- Again, need to look at the TNPs pull plot to understand the post-fit uncertainty

\* uncertainties in units of  $10^{-3}$

# Different constraints on TNPs

Using now  $\theta_n = 0 \pm \Delta\theta_n$  with  $\Delta\theta_n = 0.5$

Fit  $N^{3+1}LL$  against  $N^{3+1}LL$  data



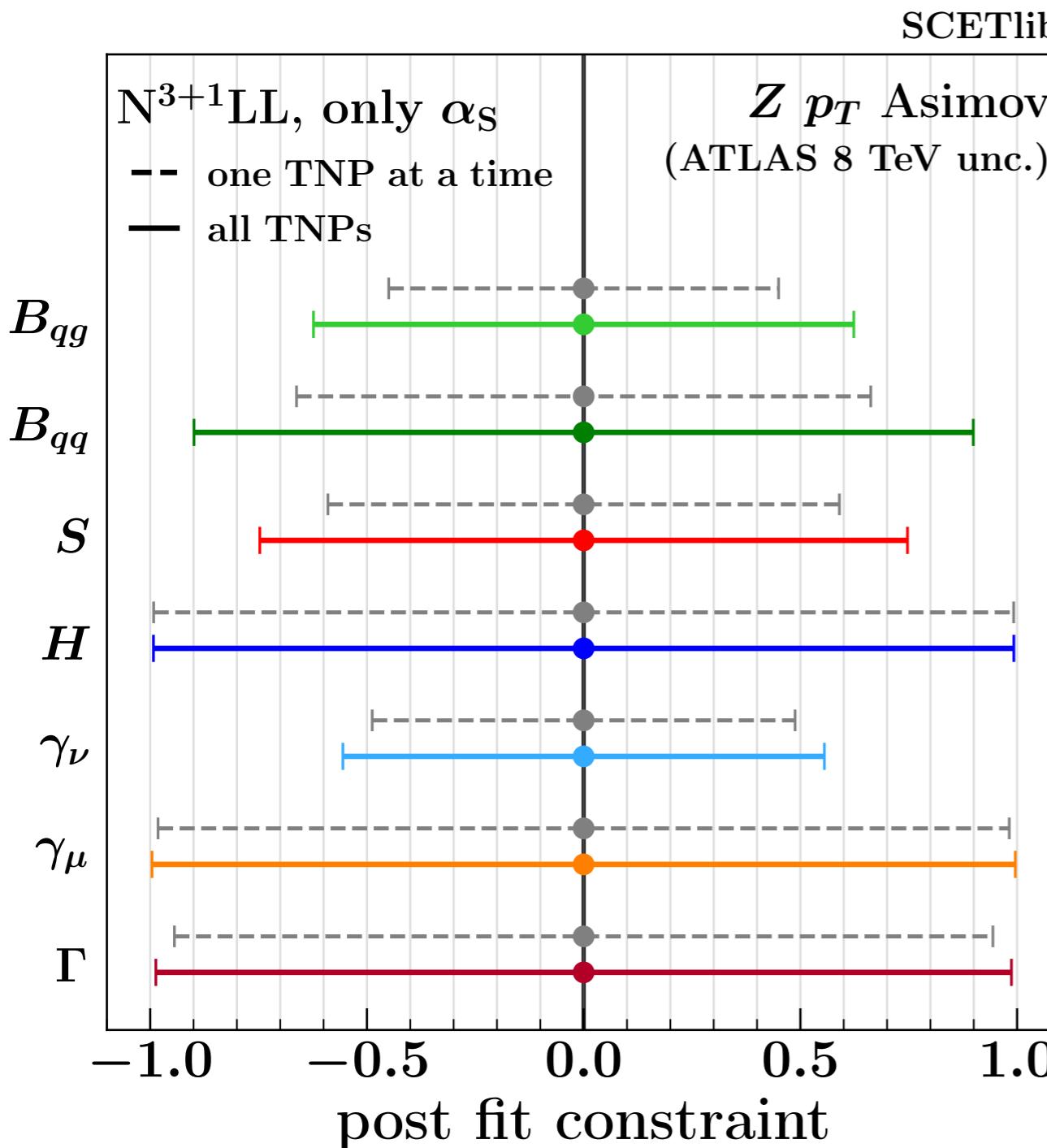
[don't be fooled by the different range!]

- 1  $\Delta\theta_n = 0.5$  not really constrained by exp.,  
but very tight theory constraint for TNPs  
[exp. uncert.  $\gtrsim$  theory uncert.]

# Different constraints on TNPs

Using now  $\theta_n = 0 \pm \Delta\theta_n$  with  $\Delta\theta_n = 1$

Fit  $N^{3+1}\text{LL}$  against  $N^{3+1}\text{LL}$  data



[don't be fooled by the different range!]

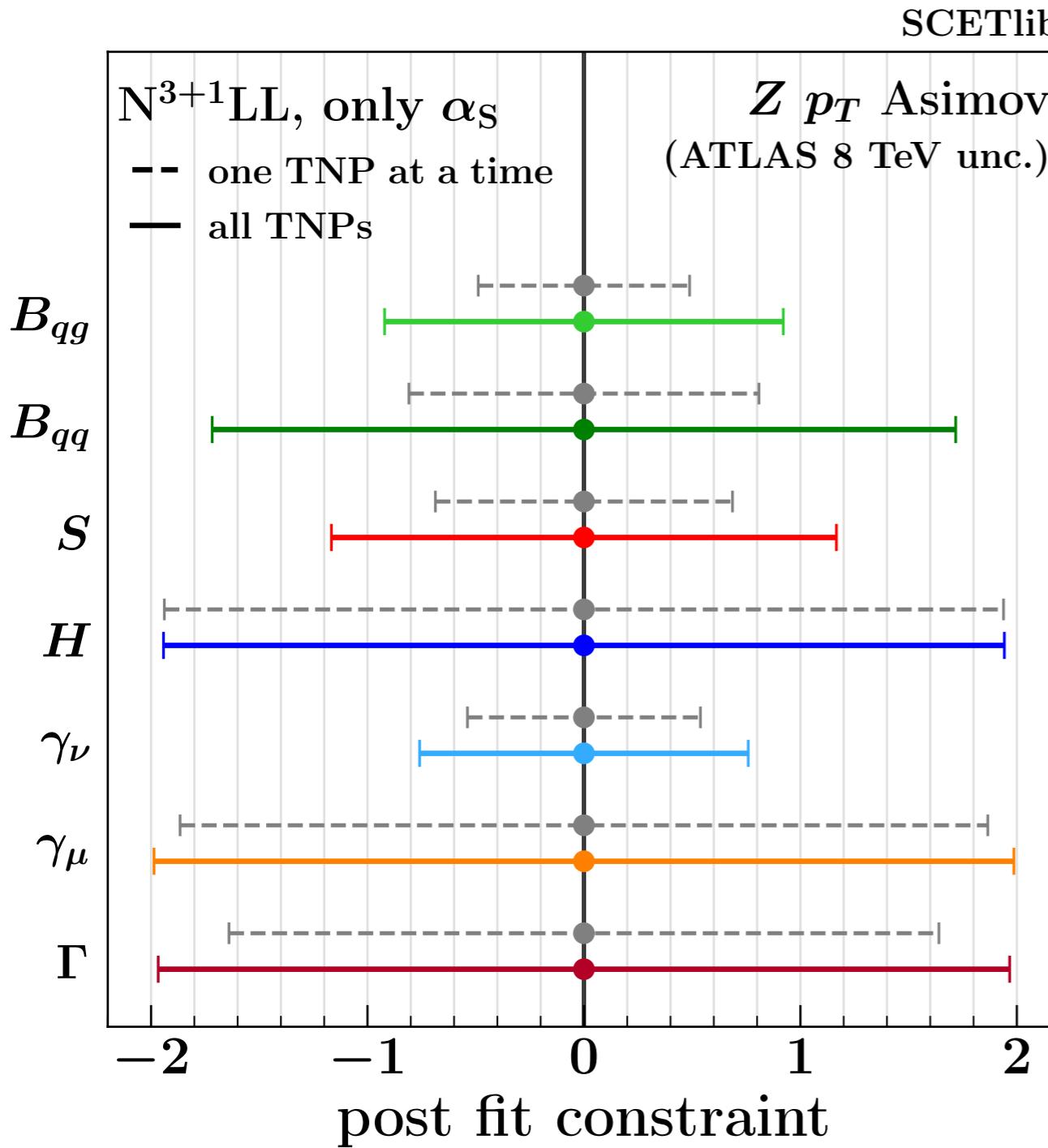
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[exp. uncert.  $\gtrsim$  theory uncert.]

2  $\Delta\theta_n = 1$  start seeing the exp. constraint

# Different constraints on TNPs

Using now  $\theta_n = 0 \pm \Delta\theta_n$  with  $\Delta\theta_n = 2$

Fit  $N^{3+1}LL$  against  $N^{3+1}LL$  data



[don't be fooled by the different range!]

①  $\Delta\theta_n = 0.5$  not really constrained by exp.,  
but very tight theory constraint for TNPs  
[exp. uncert.  $\gtrsim$  theory uncert.]

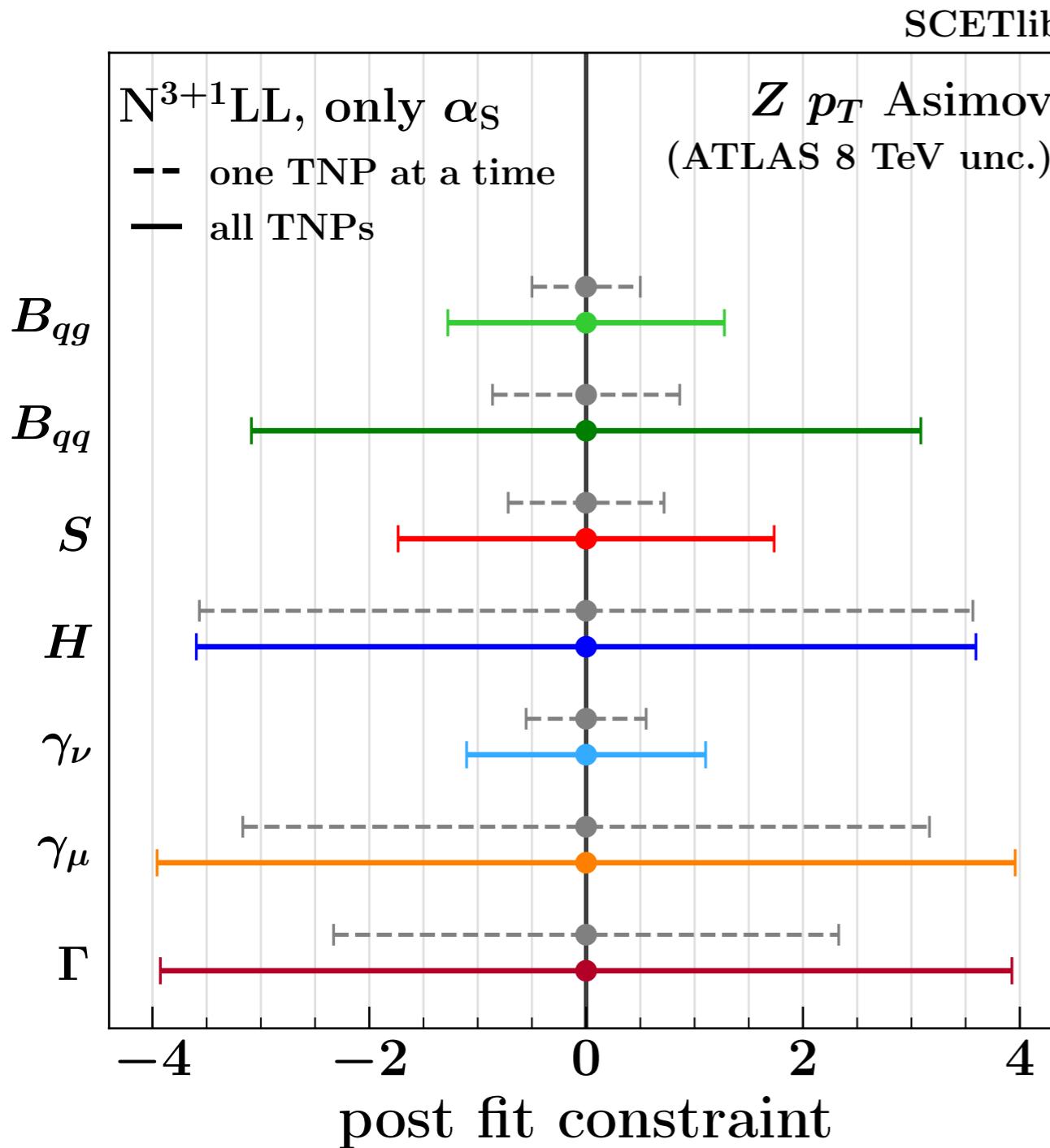
②  $\Delta\theta_n = 1$  start seeing the exp. constraint

③  $\Delta\theta_n = 2$  it's basically a factor 2 w.r.t  $\Delta\theta_n = 1$

# Different constraints on TNPs

Using now  $\theta_n = 0 \pm \Delta\theta_n$  with  $\Delta\theta_n = 4$

Fit  $N^{3+1}\text{LL}$  against  $N^{3+1}\text{LL}$  data



[don't be fooled by the different range!]

- 1  $\Delta\theta_n = 0.5$  not really constrained by exp., but very tight theory constraint for TNPs [exp. uncert.  $\gtrsim$  theory uncert.]
- 2  $\Delta\theta_n = 1$  start seeing the exp. constraint
- 3  $\Delta\theta_n = 2$  it's basically a factor 2 w.r.t  $\Delta\theta_n = 1$
- 4 with  $\Delta\theta_n = 4$  data can constrain TNPs more

# Nonperturbative uncertainty in Asimov fit

1 Collins-Soper (CS) kernel [ $\sim$  rapidity anomalous dimensions]:

$$\tilde{\gamma}_\nu(b_T) = \tilde{\gamma}_\nu^{\text{pert}} \left( b_6^*(b_T) \right) + \tilde{\gamma}_\nu^{\text{nonp}}(b_T) \quad \tilde{\gamma}_\nu^{\text{nonp}}(b_T) = -\lambda_\infty f_\nu \left( \frac{\lambda_2}{\lambda_\infty} b_T^2 + \frac{\lambda_4}{\lambda_\infty} b_T^4 \right)$$

2 Transverse Momentum Distributions (TMDs)[ $\sim$  intrinsic  $k_T$  of the partons inside the protons]:

$$\tilde{f}(b_T) = \tilde{f}_{\text{pert}}(b_T) \tilde{f}_{\text{nonp}}(b_T) \quad \ln \left( \tilde{f}_{\text{nonp}}(b_T) \right) = -\Lambda_\infty b_T f \left( \frac{\Lambda_2}{\Lambda_\infty} b_T + \frac{\Lambda_4}{\Lambda_\infty} b_T^3 \right)$$

$\lambda_2, \lambda_4$  and  $\Lambda_2, \Lambda_4$  quadratic/quartic small  $b_T$  coefficients

$\lambda_\infty, \Lambda_\infty$  determine  $b_T \rightarrow \infty$  behavior

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From Collins and Rogers '14,  $f_\nu(x)$  and  $f(x)$  behavior

$$f_\nu(x \rightarrow 0) \sim x^2, \quad f_\nu(x \rightarrow \infty) \sim \text{const}$$
$$\log(f(x \rightarrow 0)) \sim x^2, \quad \log(f(x \rightarrow \infty)) \sim x$$

# Nonperturbative uncertainty in Asimov fit

---

What is used in our fits:

$$f_\nu \left( \frac{\lambda_2}{\lambda_\infty} b_T^2 + \frac{\lambda_4}{\lambda_\infty} b_T^4 \right) = \tanh \left( \frac{\lambda_2}{\lambda_\infty} b_T^2 + \frac{\lambda_4}{\lambda_\infty} b_T^4 \right)$$

$$f \left( \frac{\Lambda_2}{\Lambda_\infty} b_T + \frac{\Lambda_4}{\Lambda_\infty} b_T^3 \right) = 2 \tanh \left( \frac{\Lambda_2}{\Lambda_\infty} b_T + \frac{\Lambda_4}{\Lambda_\infty} b_T^3 \right)$$

# Nonperturbative uncertainty in Asimov fit

What is used in our fits:

$$f_\nu \left( \frac{\lambda_2}{\lambda_\infty} b_T^2 + \frac{\lambda_4}{\lambda_\infty} b_T^4 \right) = \tanh \left( \frac{\lambda_2}{\lambda_\infty} b_T^2 + \frac{\lambda_4}{\lambda_\infty} b_T^4 \right)$$

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Also using inputs from **lattice QCD** for the CS kernel [some details [here](#)]:

- exploit lattice QCD calculations of the CS kernel to obtain good constraints on  $\lambda_\infty$ ,  $\lambda_2$  and  $\lambda_4$

representative values:

$$\lambda_\infty = 1.7 \pm 0.5$$

$$\lambda_2 = 0.09 \pm 0.03$$

$$\lambda_4 = 0.007 \pm 0.007$$

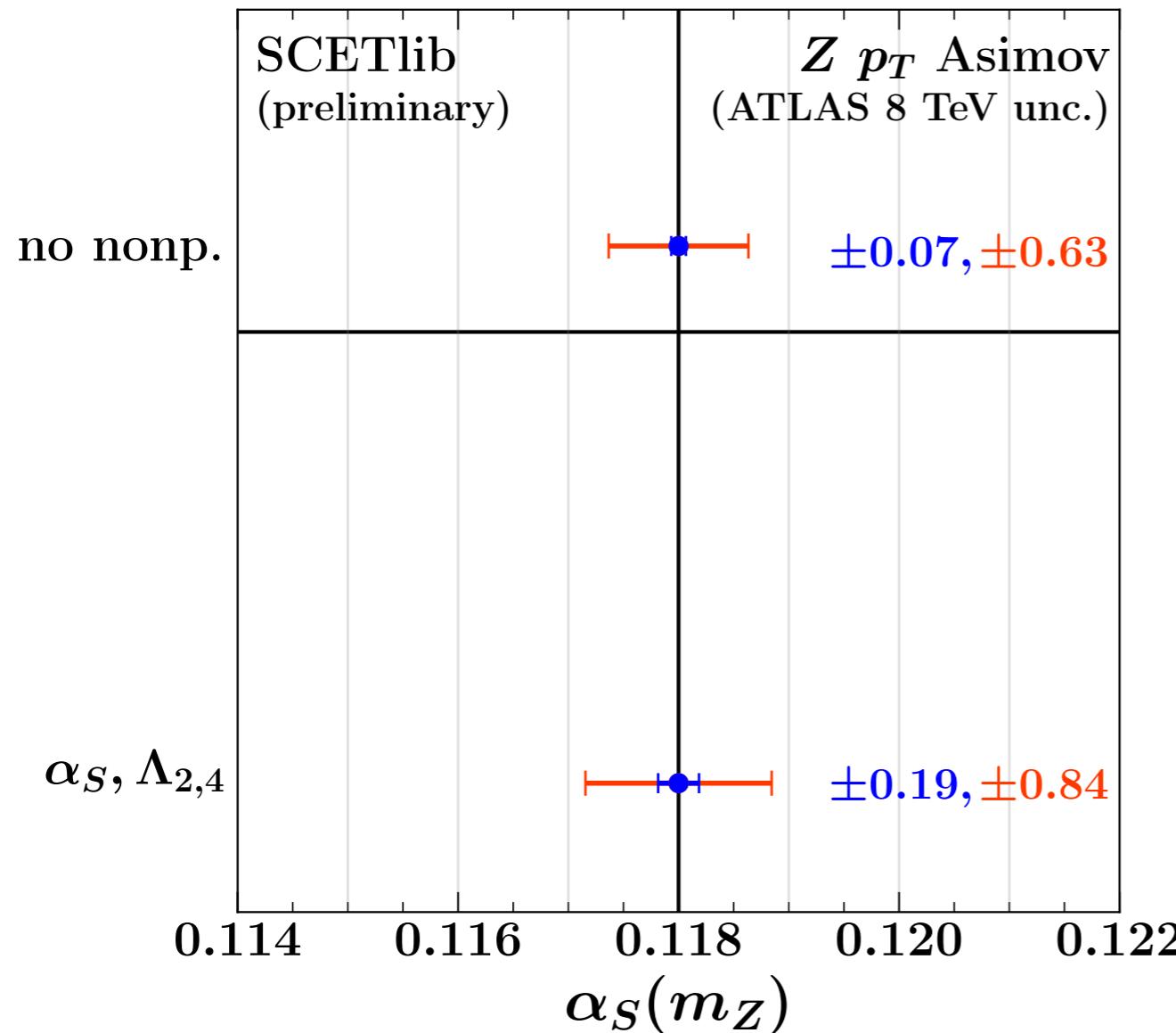
+ full covariance matrix from  
lattice fit

# Nonperturbative uncertainty in Asimov fit

fit unc. only: fitting *only*  $\alpha_S$  and nonp.

$N^{3+1}$ LL profiled: including TNPs

Fit  $N^{3+1}$ LL against  $N^{3+1}$ LL data



➤ Fit only  $\alpha_S$ ,  $\Lambda_2$  and  $\Lambda_4$  (fixed  $\tilde{\gamma}_\nu^{\text{nonp}}$ )

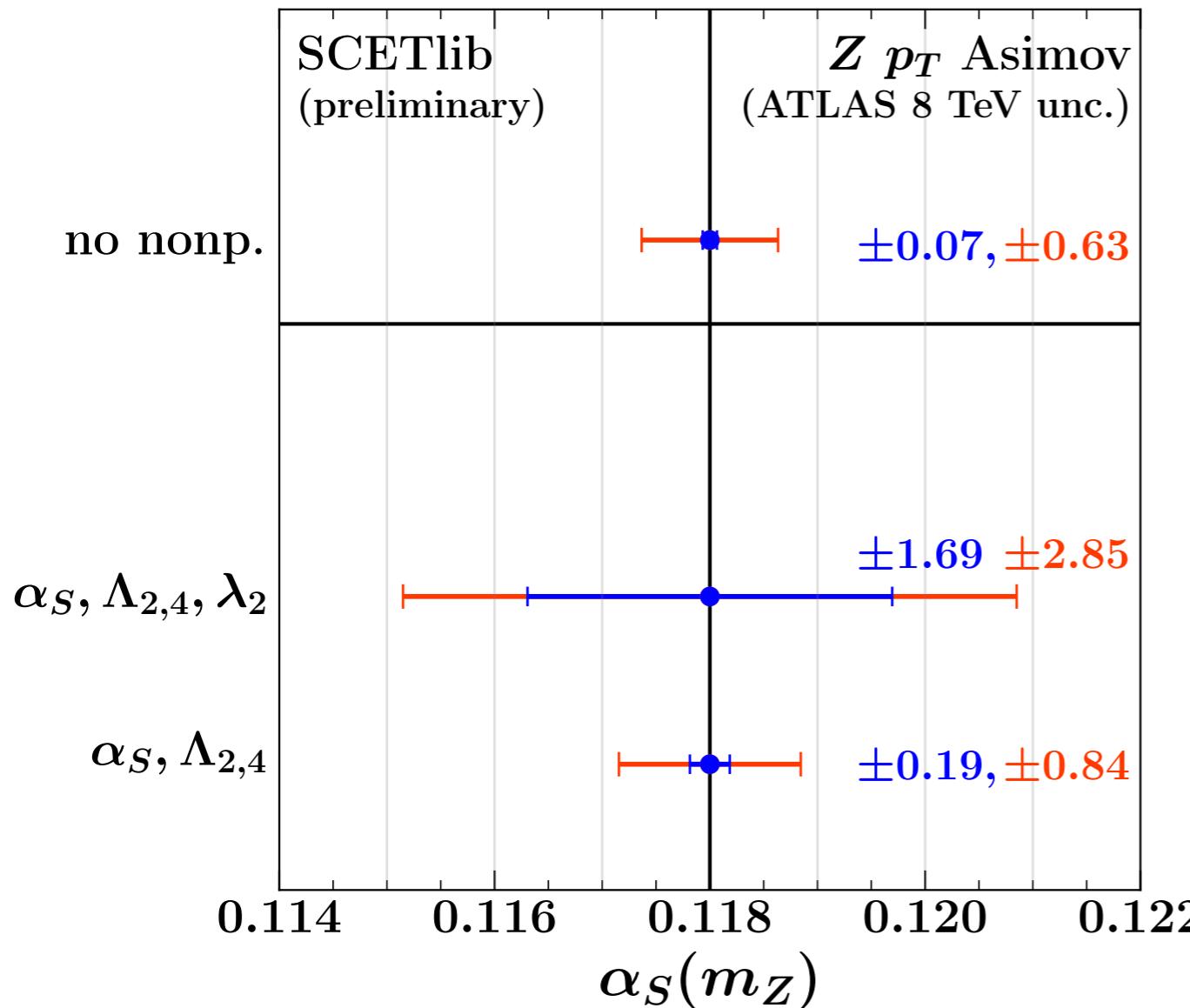
\* uncertainties in units of  $10^{-3}$

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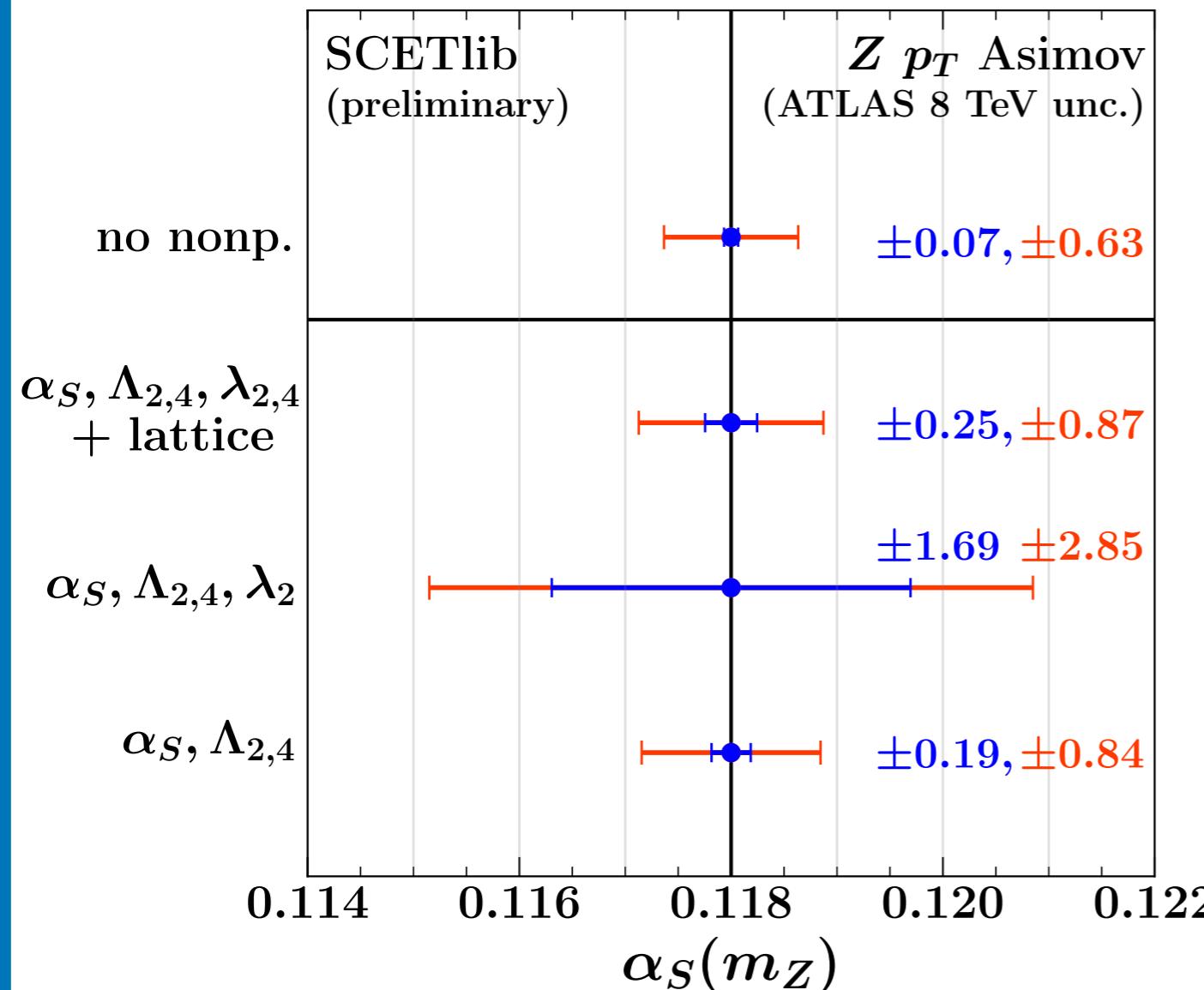
➤ Not using lattice constraints  
parameters fitted:  $\lambda_2, \Lambda_2, \Lambda_4$

\* uncertainties in units of  $10^{-3}$

# Nonperturbative uncertainty in Asimov fit

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**N<sup>3+1</sup>LL profiled:** including TNPs

Fit N<sup>3+1</sup>LL against N<sup>3+1</sup>LL data

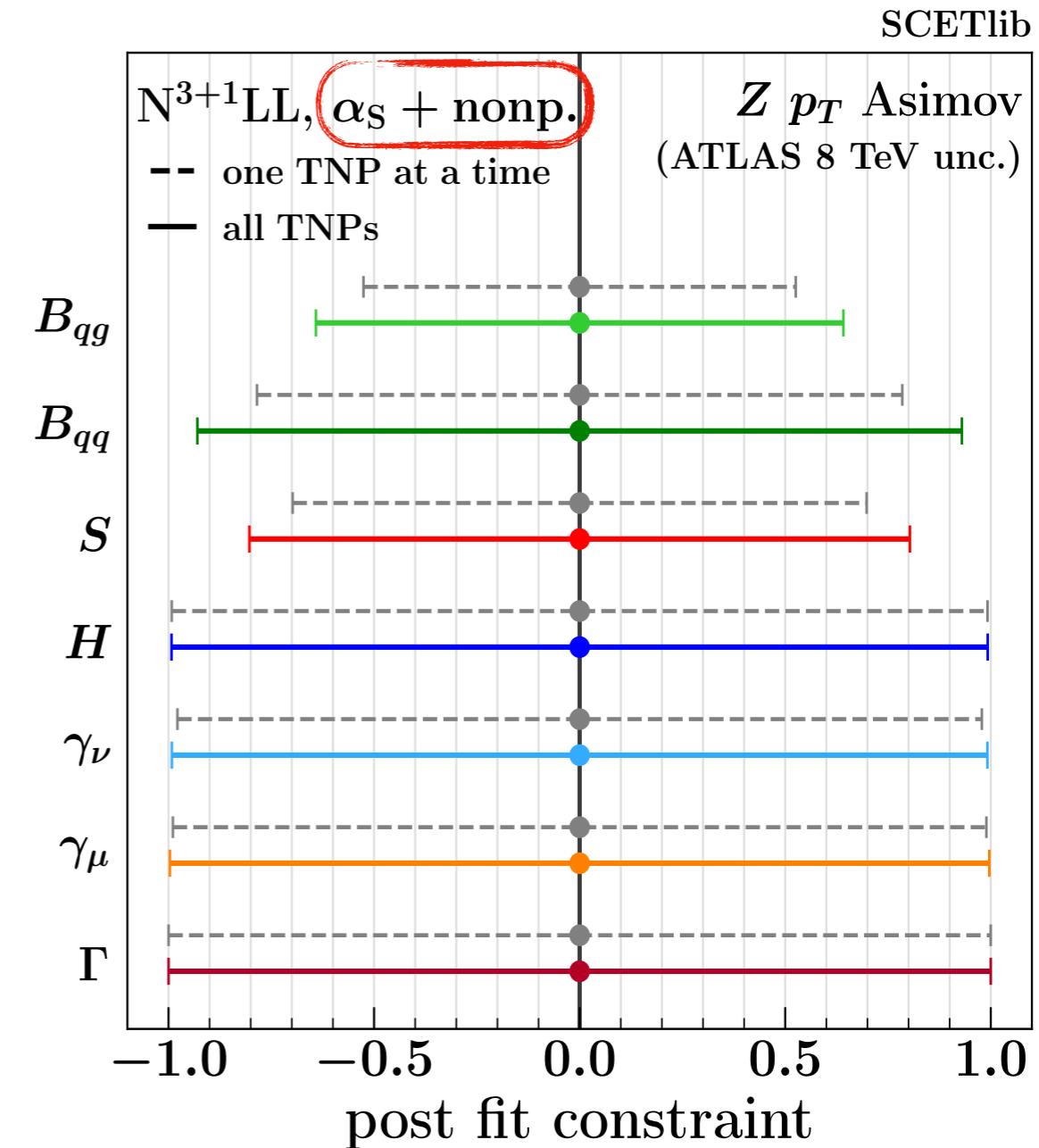
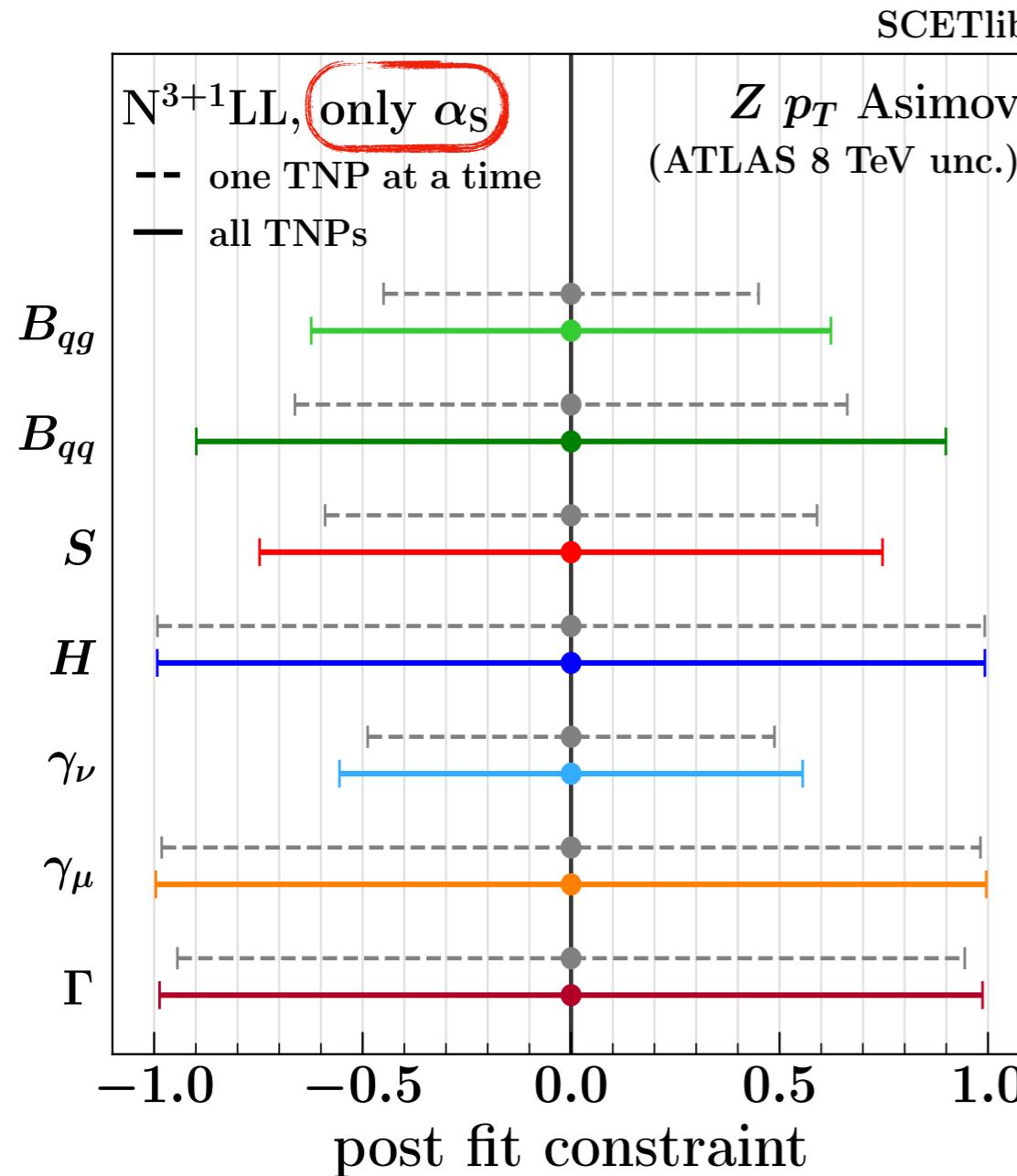


- Fit only  $\alpha_S$ ,  $\Lambda_2$  and  $\Lambda_4$  (fixed  $\tilde{\gamma}_\nu^{\text{nonp}}$ )
- Not using lattice constraints  
parameters fitted:  $\lambda_2, \Lambda_2, \Lambda_4$
- Using lattice constraints  
parameters fitted:  $\lambda_2, \lambda_4, \Lambda_2, \Lambda_4$

\* uncertainties in units of 10<sup>-3</sup>

# Nonperturbative uncertainty in Asimov fit

► parameters fitted  $\lambda_2$ ,  $\lambda_4$ ,  $\Lambda_2$ ,  $\Lambda_4$  + lattice QCD constraints



Data now also constraint nonperturbative parameters, therefore less constraint on TNPs

plots with different  $\Delta\theta_n$  for TNPs here

# Conclusions

---

Need for theoretical predictions including correlations for interpretation of LHC precision measurements:

**1 Theory Nuisance Parameters** perfect candidate

- » include correct correlations across the  $p_T$  spectrum
- » can be constrained by data reducing theory uncertainty
- » work as advertised for Asimov tests

**2 Nonperturbative model**

- » importance of fitting CS kernel
- » can be improved with lattice constraints

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**2 Nonperturbative model**

- » importance of fitting CS kernel
- » can be improved with lattice constraints

**3 Under investigation:**

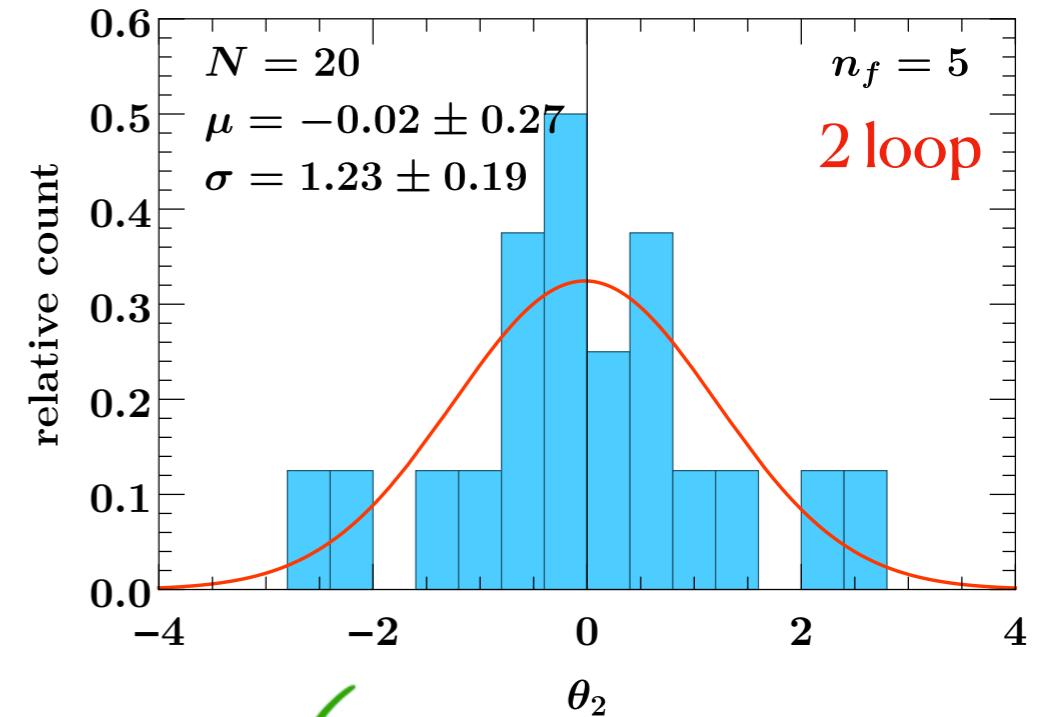
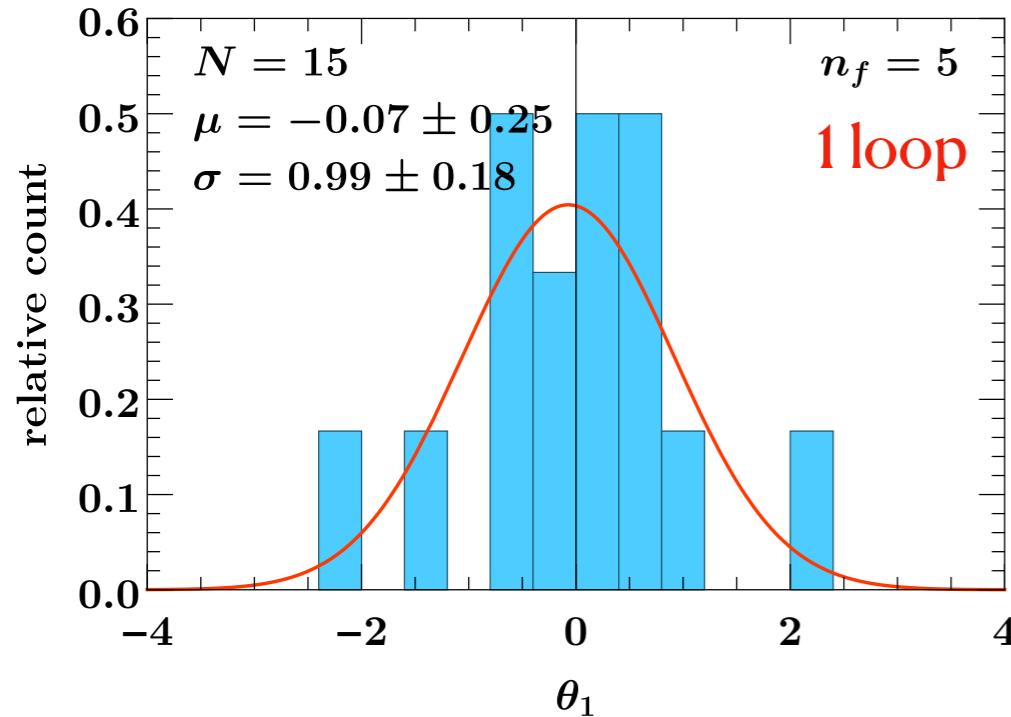
- » PDFs: scanning and/or profiling
- » Quark mass effects
- » Fits against real data

THANK YOU!

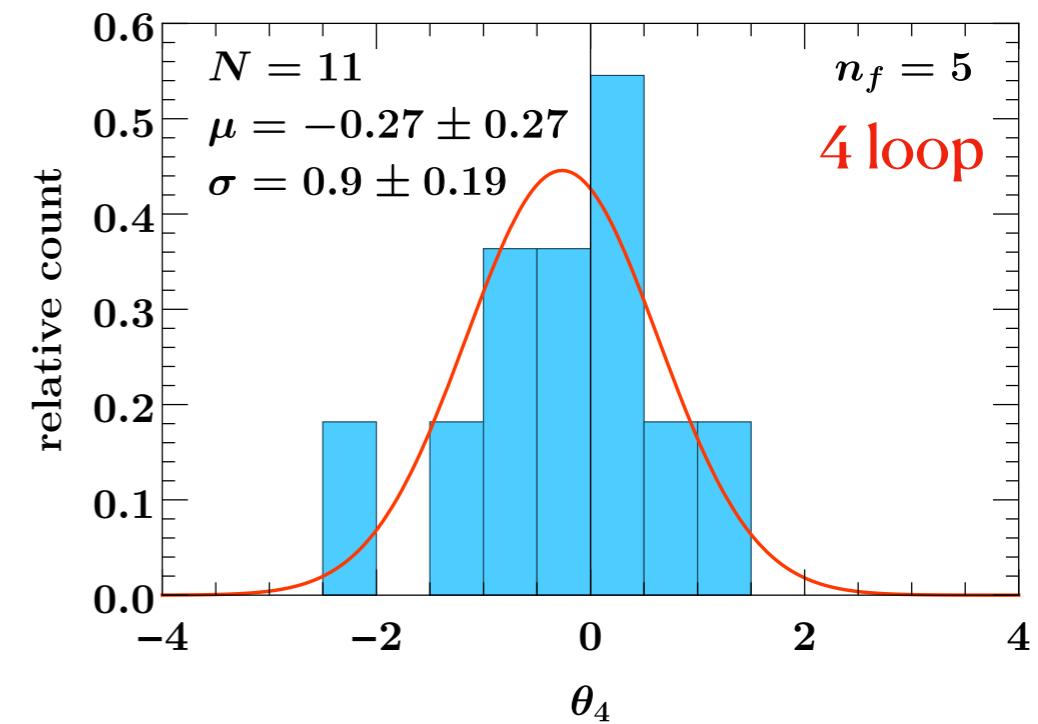
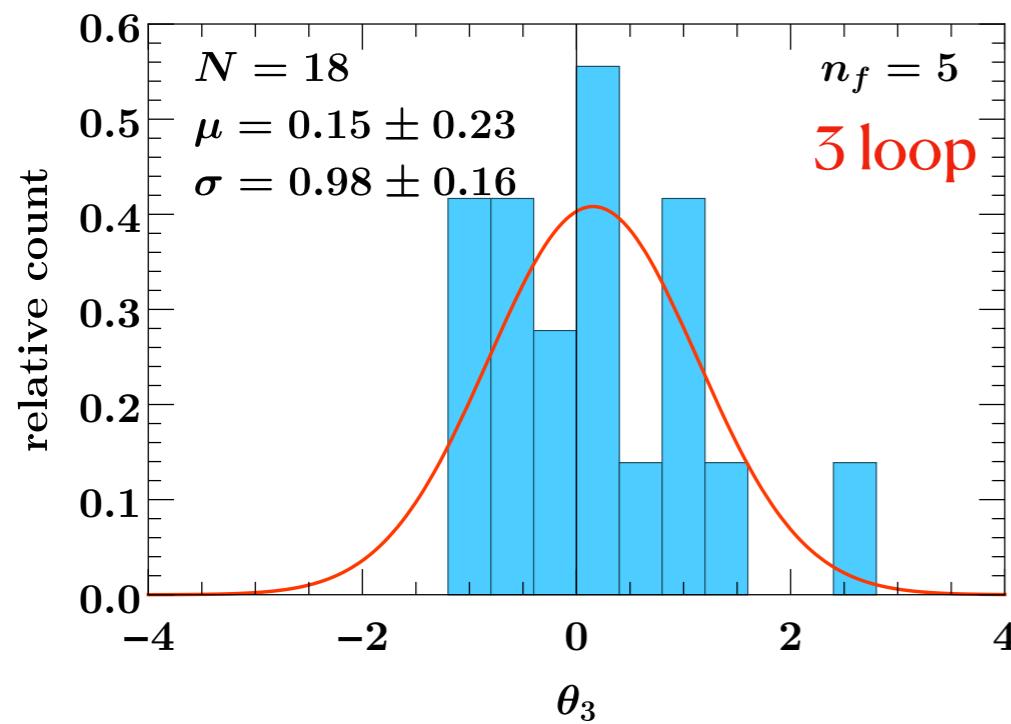
# Backup slides

# TNPs for Boundary Conditions

$$F_n(\theta_n) = 4C_r(4C_A)^{n-1}(n-1)! \theta_n^F$$

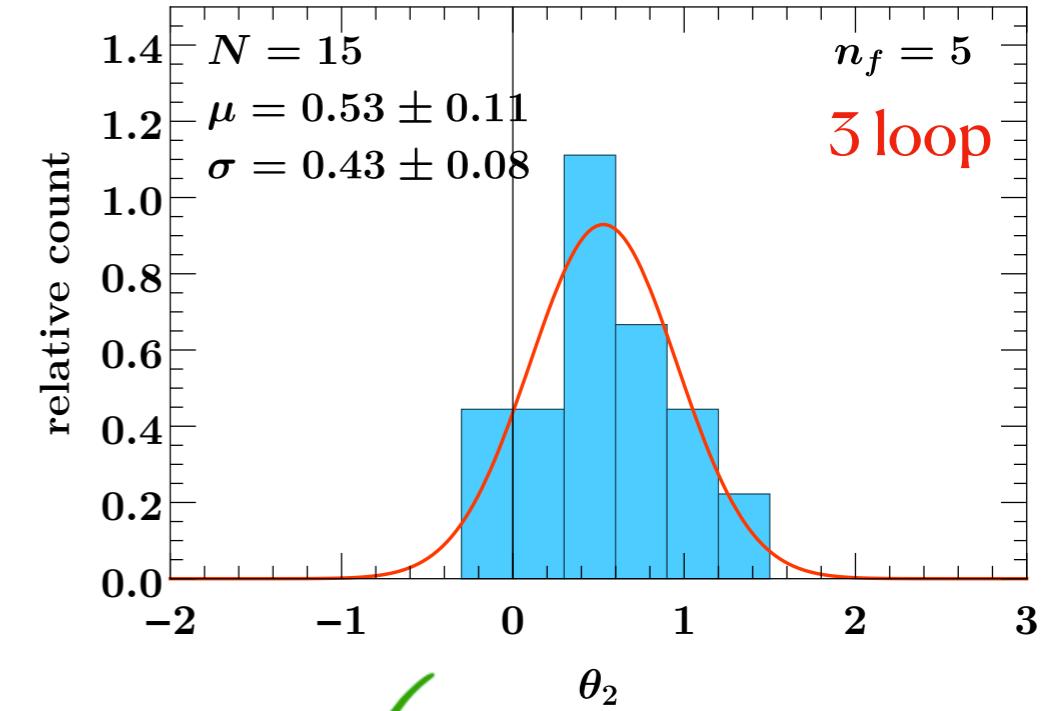
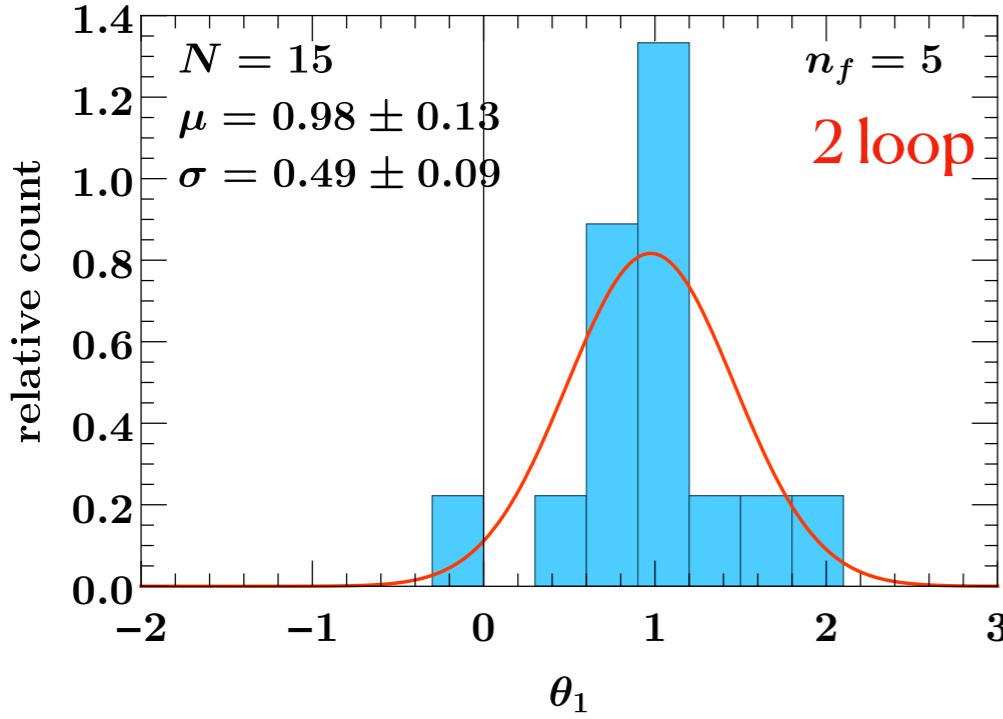


Fit to a Gaussian with  $\theta = 0$  and  $\Delta\theta_n = 1$  ✓

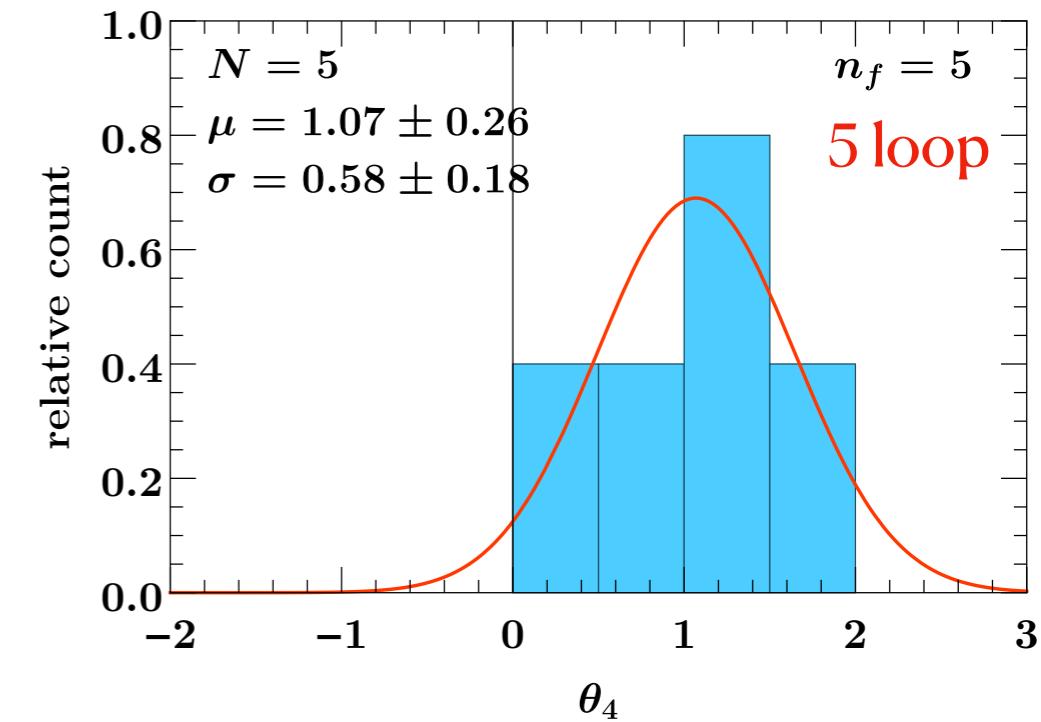
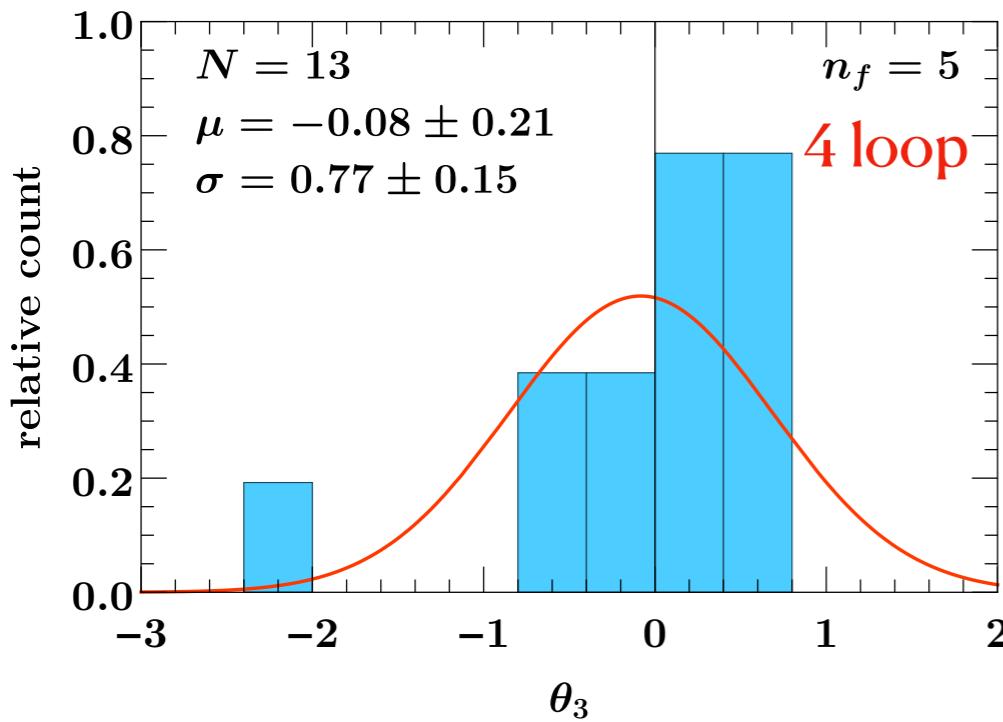


# TNPs for Anomalous Dimensions

$$\gamma_n(\theta_n) = 2C_r(4C_A)^n \theta_n^r$$

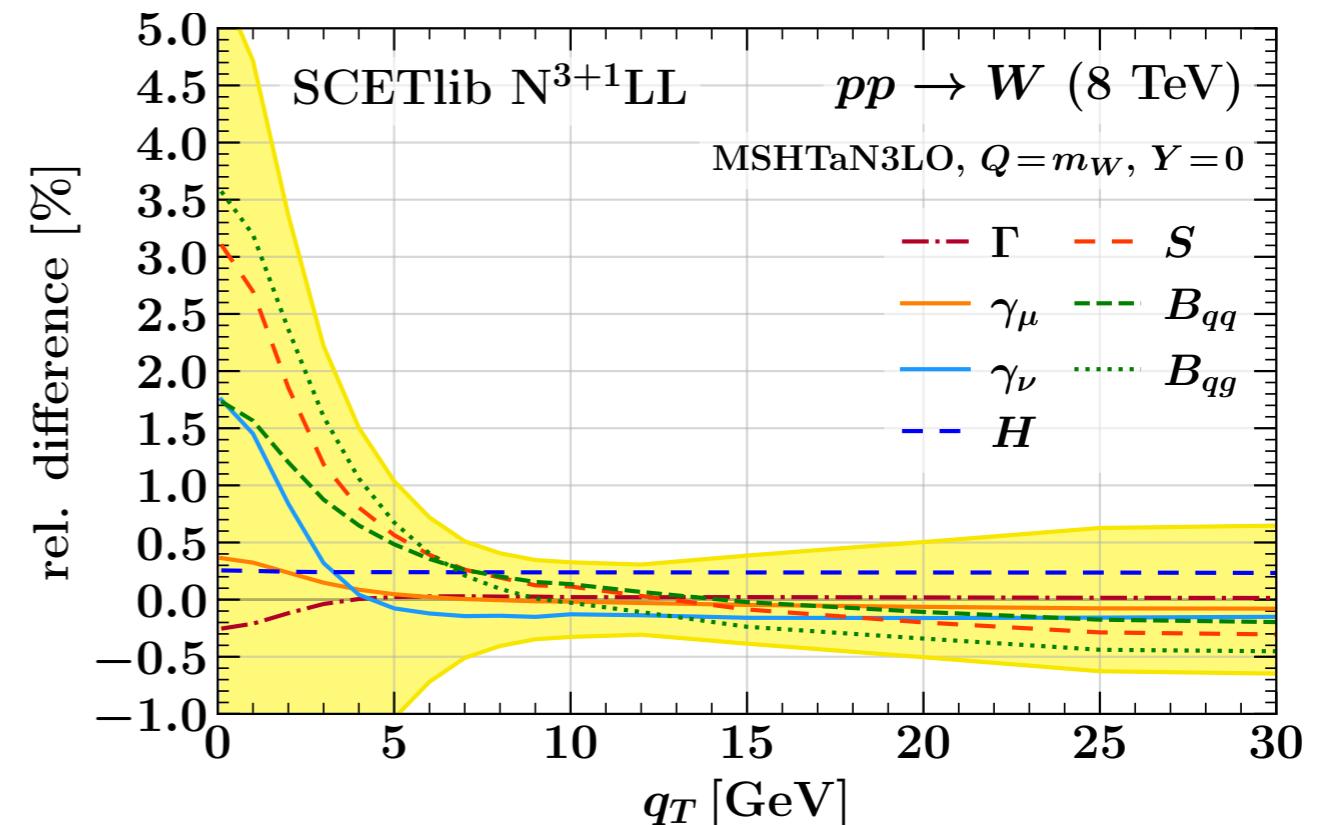
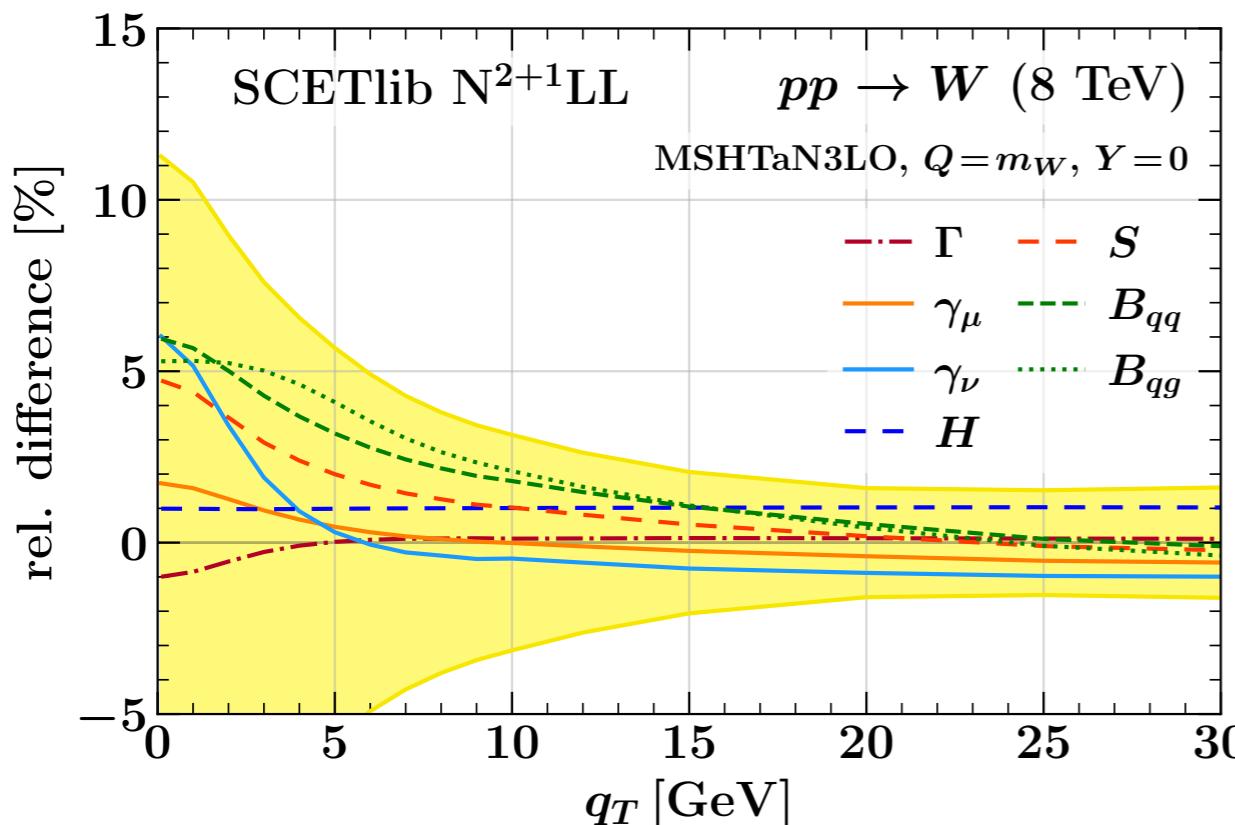
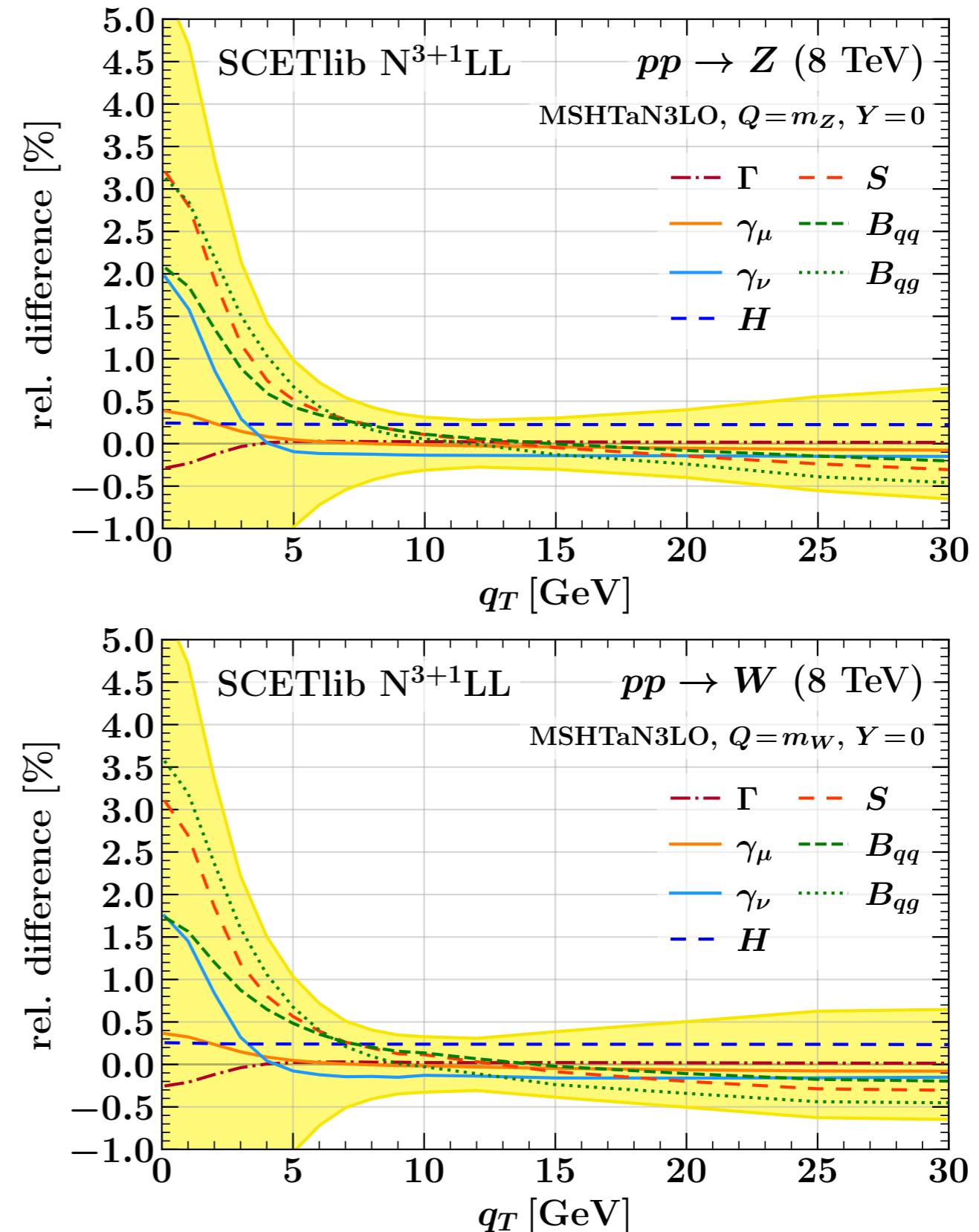
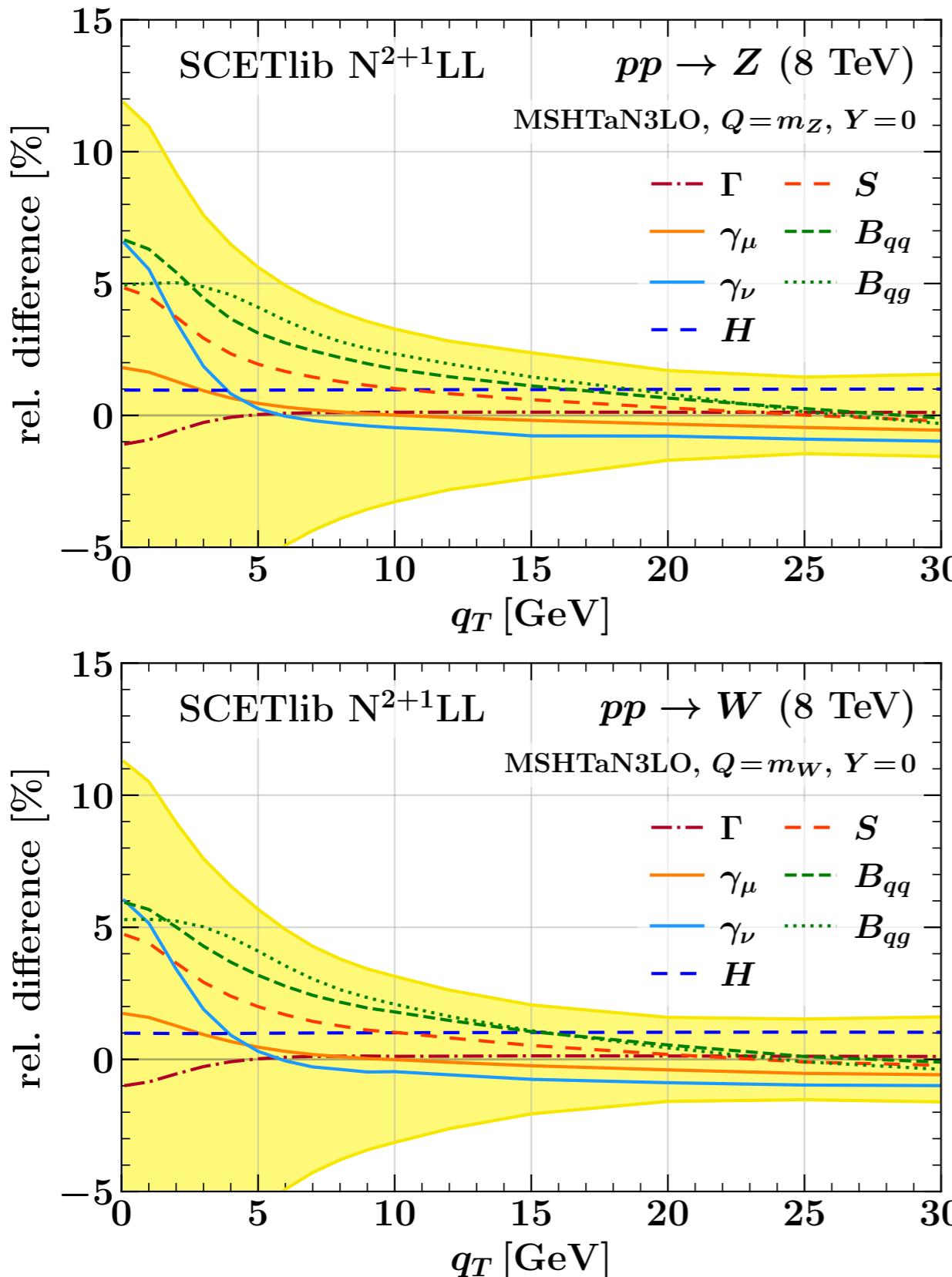


Fit to a Gaussian with  $\theta \neq 0$  and  $\Delta\theta_n = 0.5$  ✓



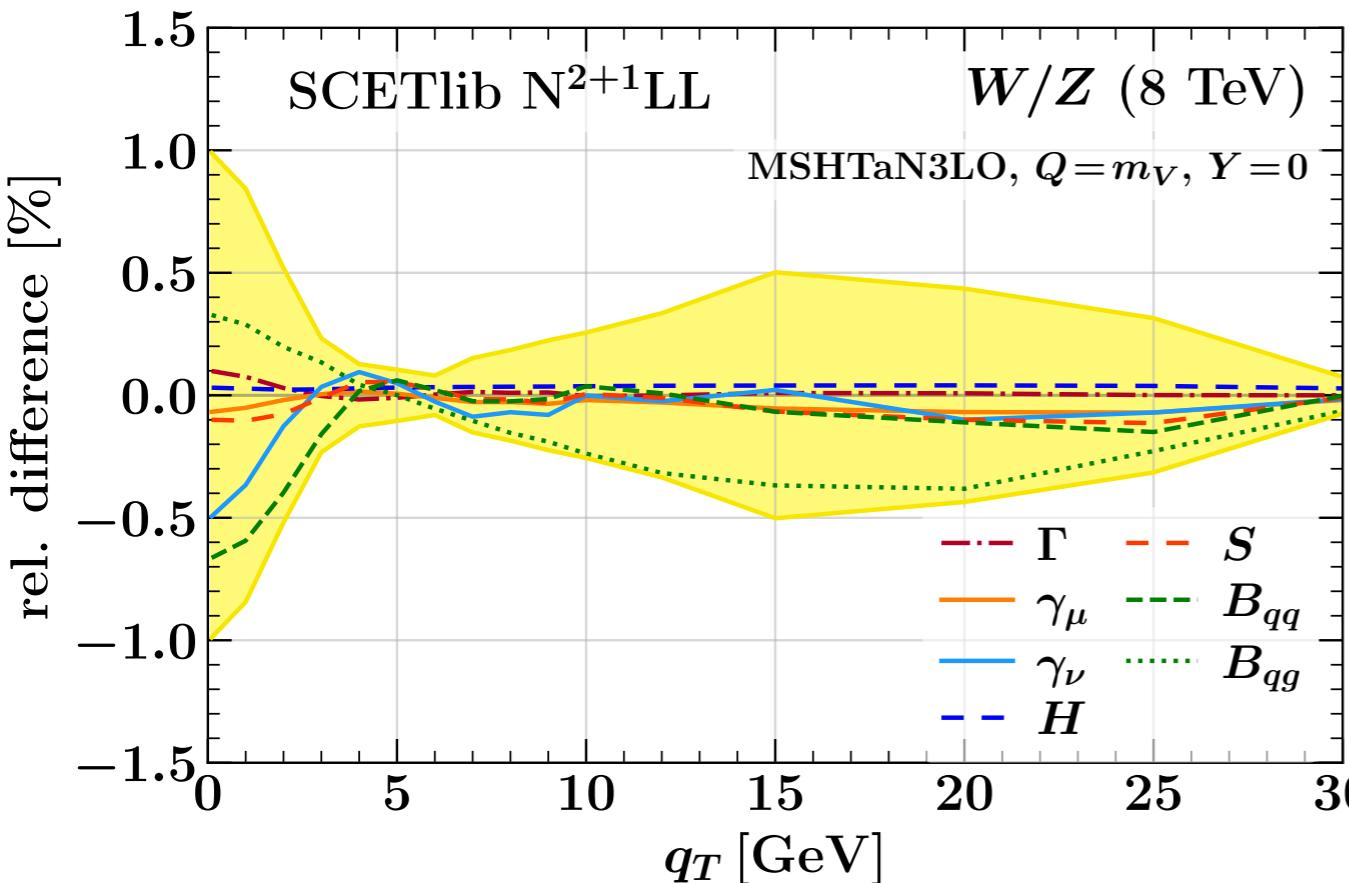
go back slide

# Application to Drell-Yan $p_T$ spectrum

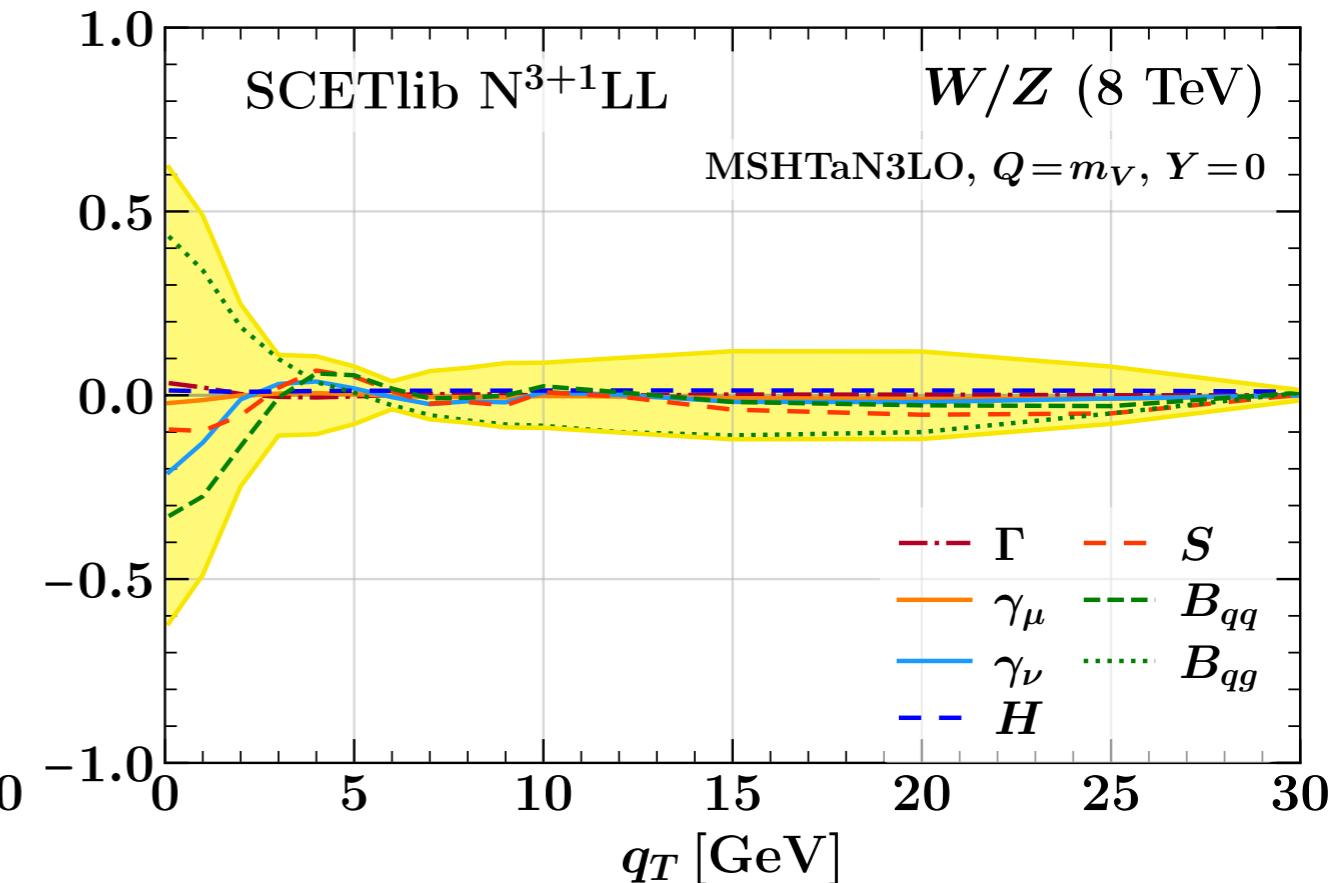


# Application to Drell-Yan $p_T$ spectrum

relative impact for  $W/Z$



relative impact for  $W/Z$

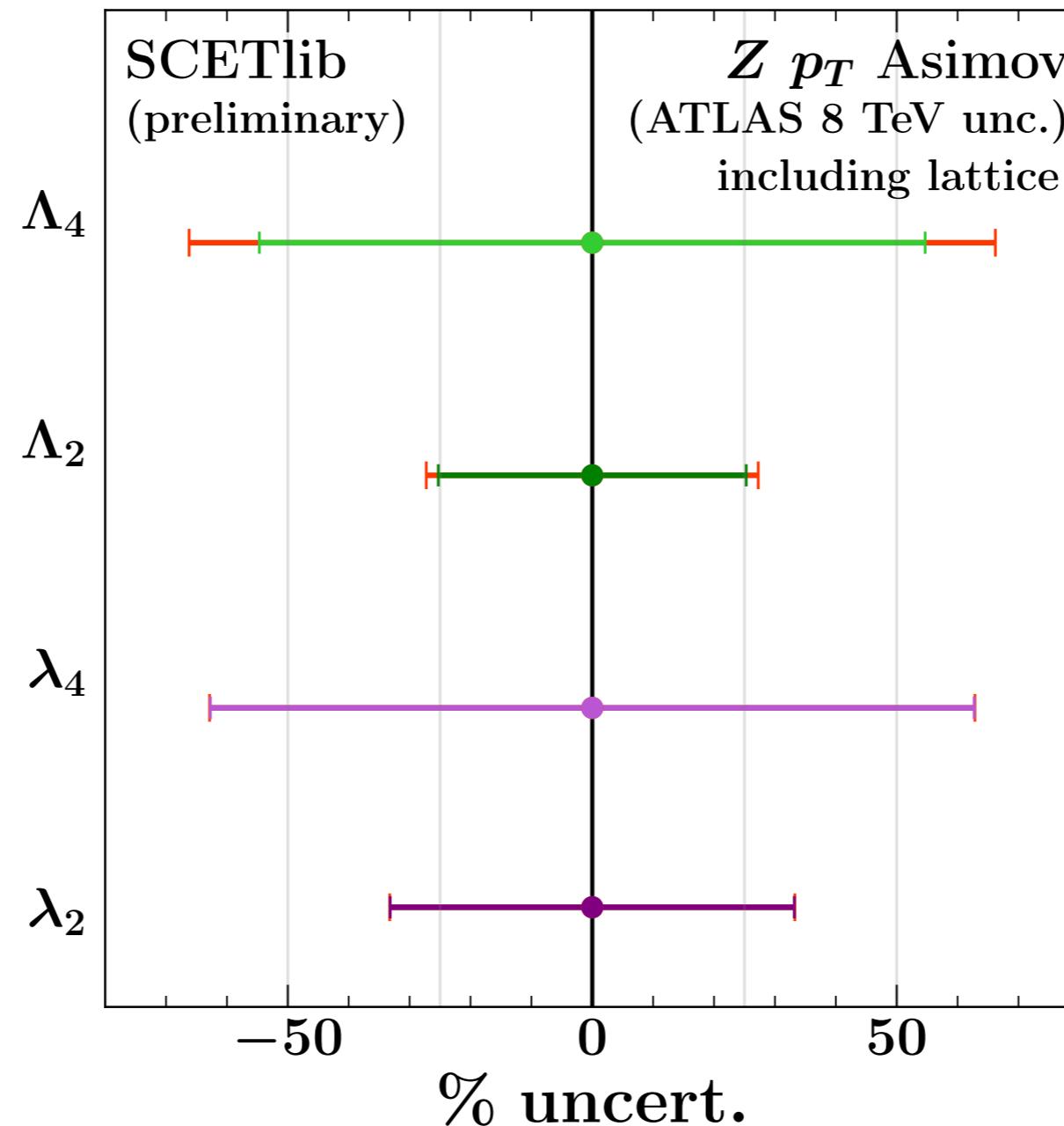


Correlation in  $p_T$  between  $W$  and  $Z$  captured ✓

go back slide

# Constraints on nonp. parameters

- parameters fitted  $\lambda_2$ ,  $\lambda_4$ ,  $\Lambda_2$ ,  $\Lambda_4$  + lattice QCD constraints

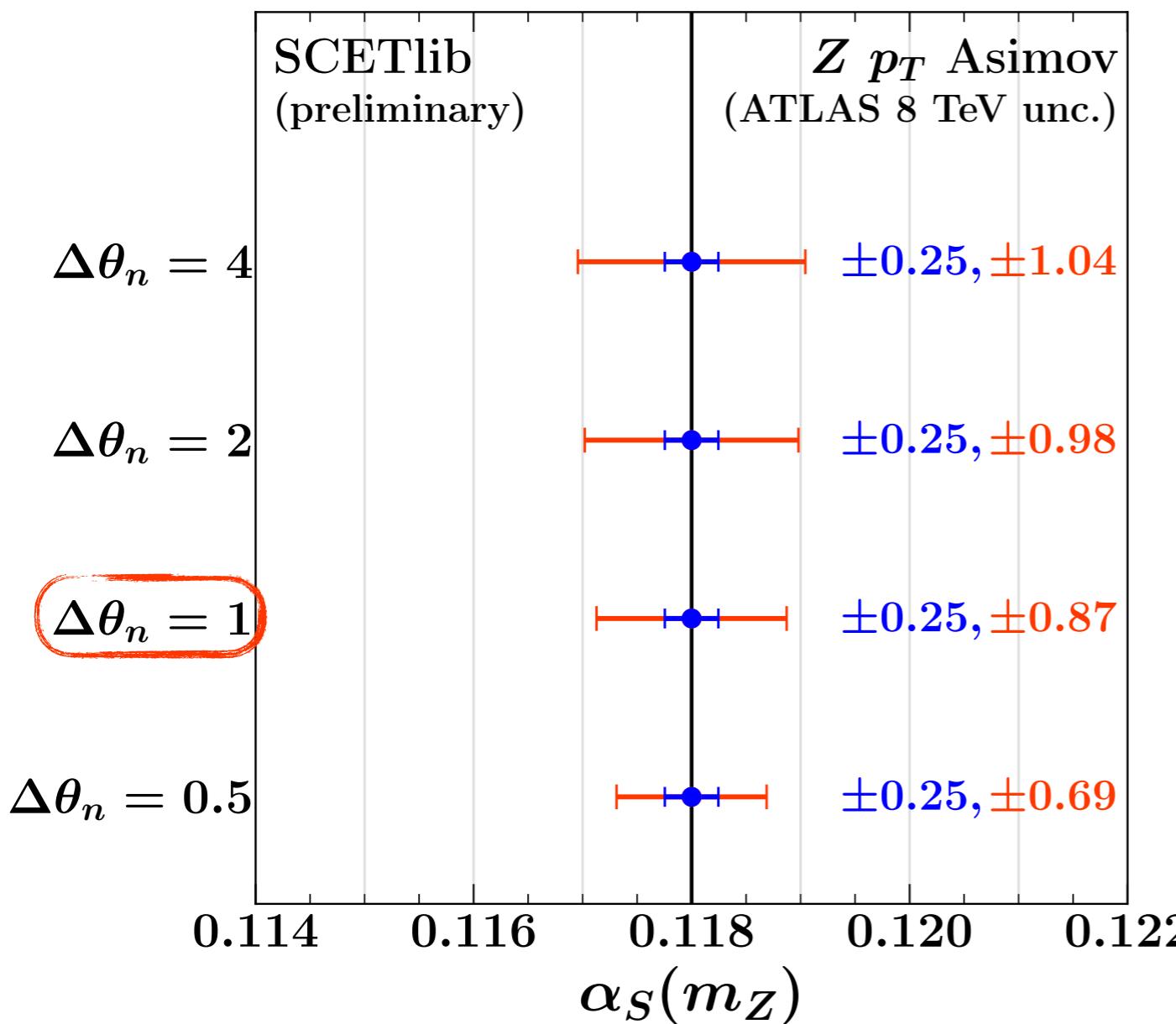


# Different constraints on TNPs including nonp.

What happens by changing the constraint?

Using now  $\theta_n = 0 \pm \Delta\theta_n$  with  $\Delta\theta_n = 0.5, 1, 2, 4$

Fit  $N^{3+1}LL$  against  $N^{3+1}LL$  data



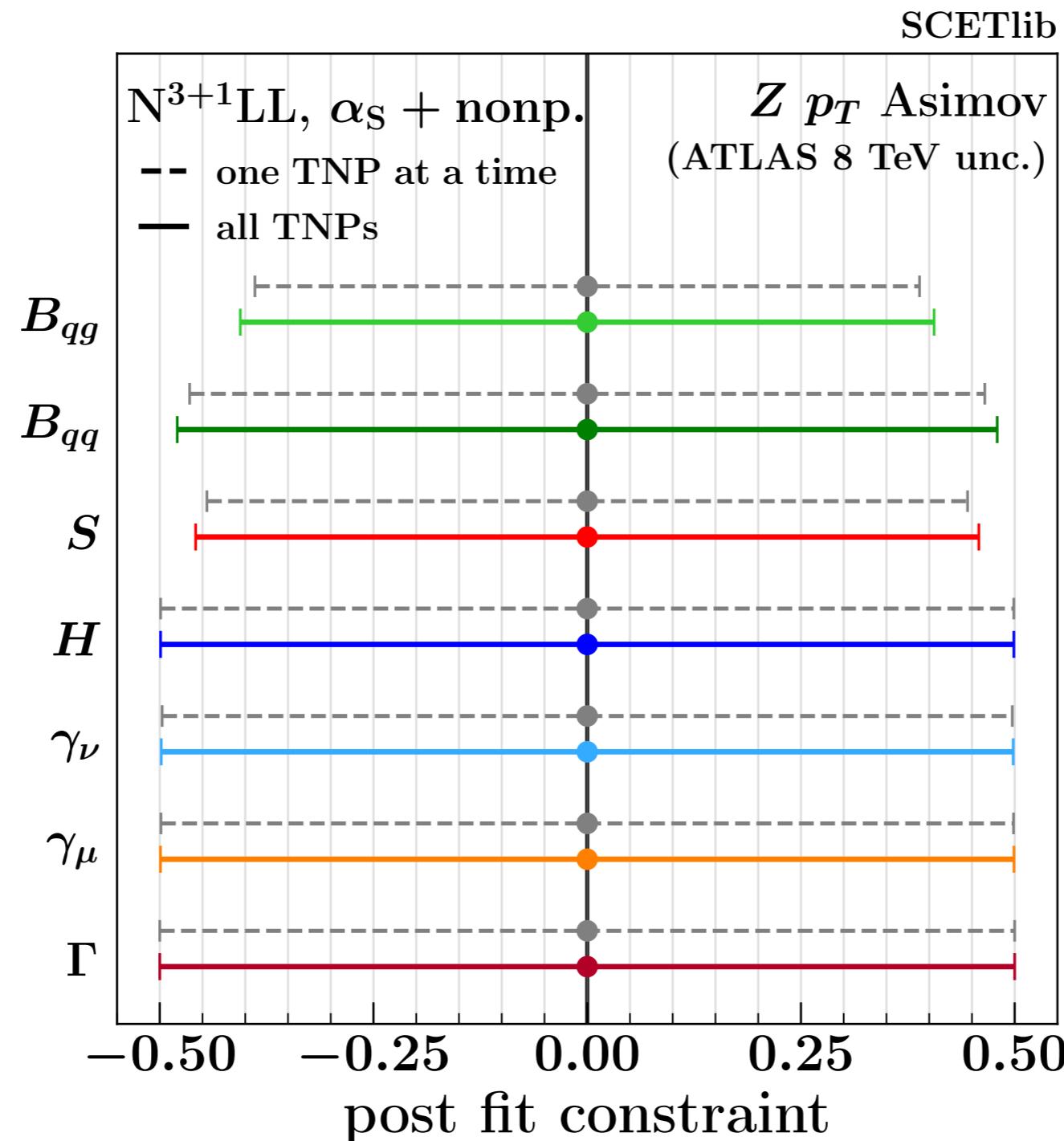
fit unc. only: fitting *only*  $\alpha_S$  and nonp.  
 $N^{3+1}LL$  profiled: including TNPs

\* uncertainties in units of  $10^{-3}$

# Different constraints on TNPs including nonp.

Using now  $\theta_n = 0 \pm \Delta\theta_n$  with  $\Delta\theta_n = 0.5$

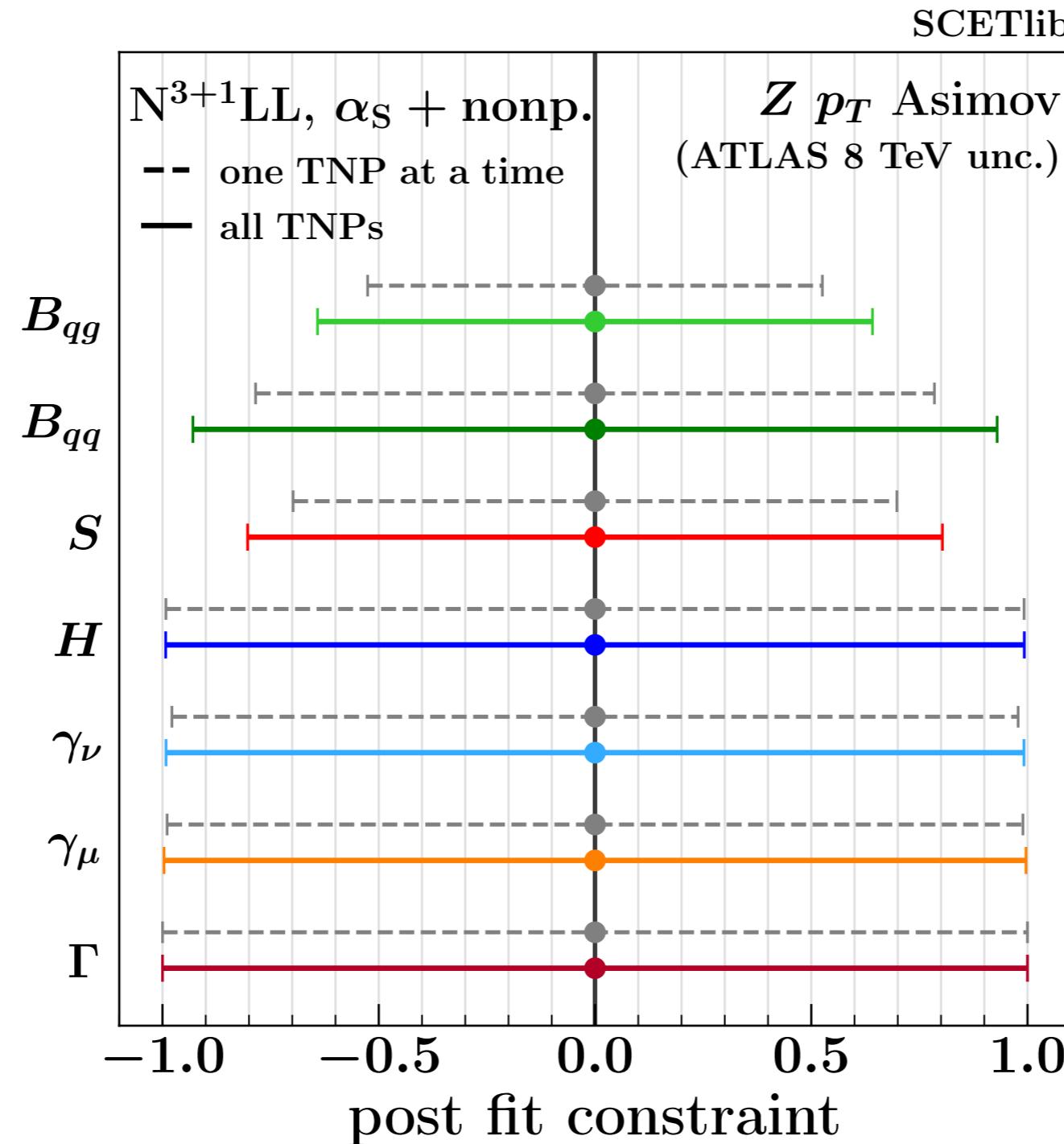
Fit  $N^{3+1}\text{LL}$  against  $N^{3+1}\text{LL}$  data



# Different constraints on TNPs including nonp.

Using now  $\theta_n = 0 \pm \Delta\theta_n$  with  $\Delta\theta_n = 1$

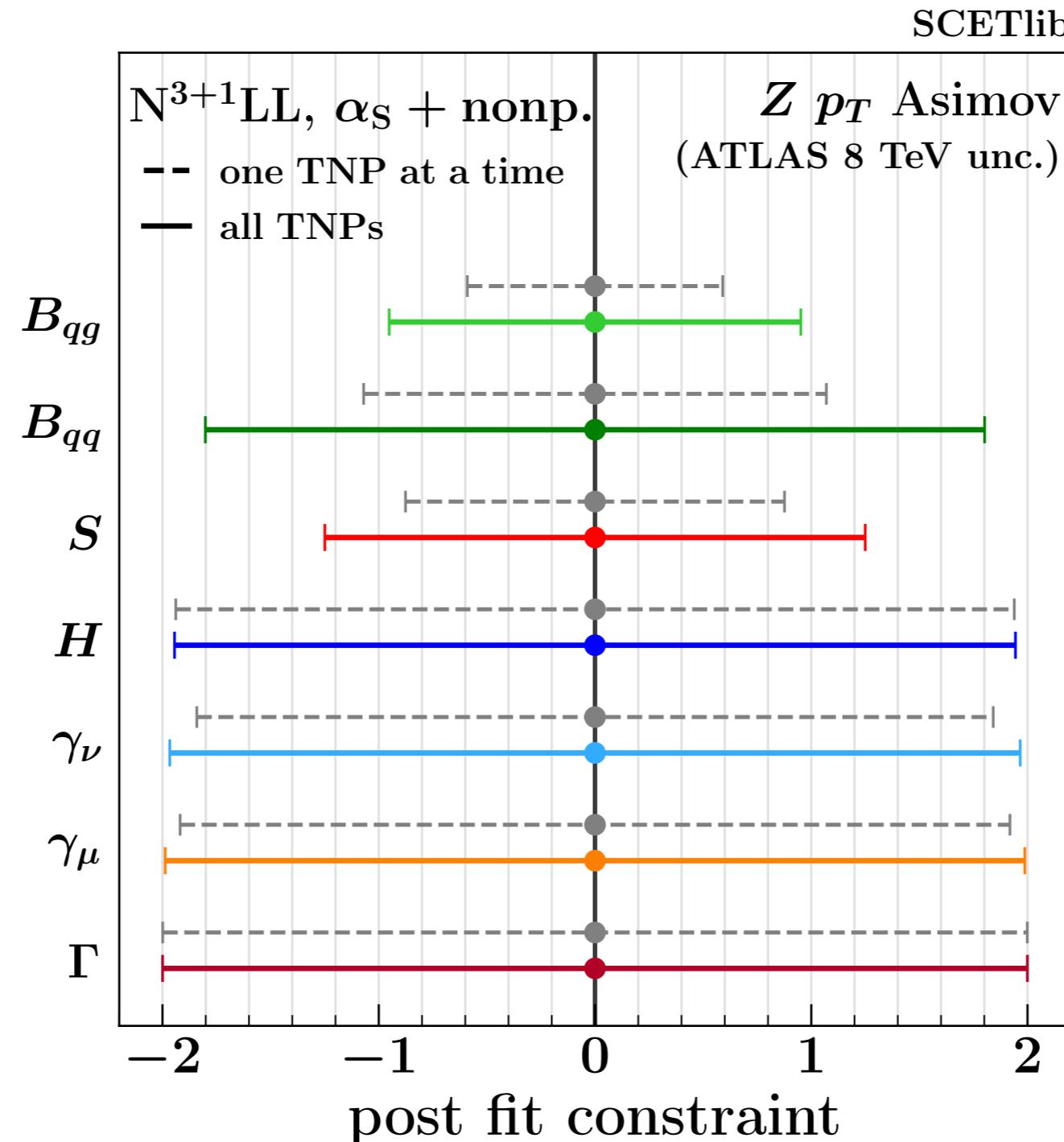
Fit  $N^{3+1}\text{LL}$  against  $N^{3+1}\text{LL}$  data



# Different constraints on TNPs including nonp.

Using now  $\theta_n = 0 \pm \Delta\theta_n$  with  $\Delta\theta_n = 2$

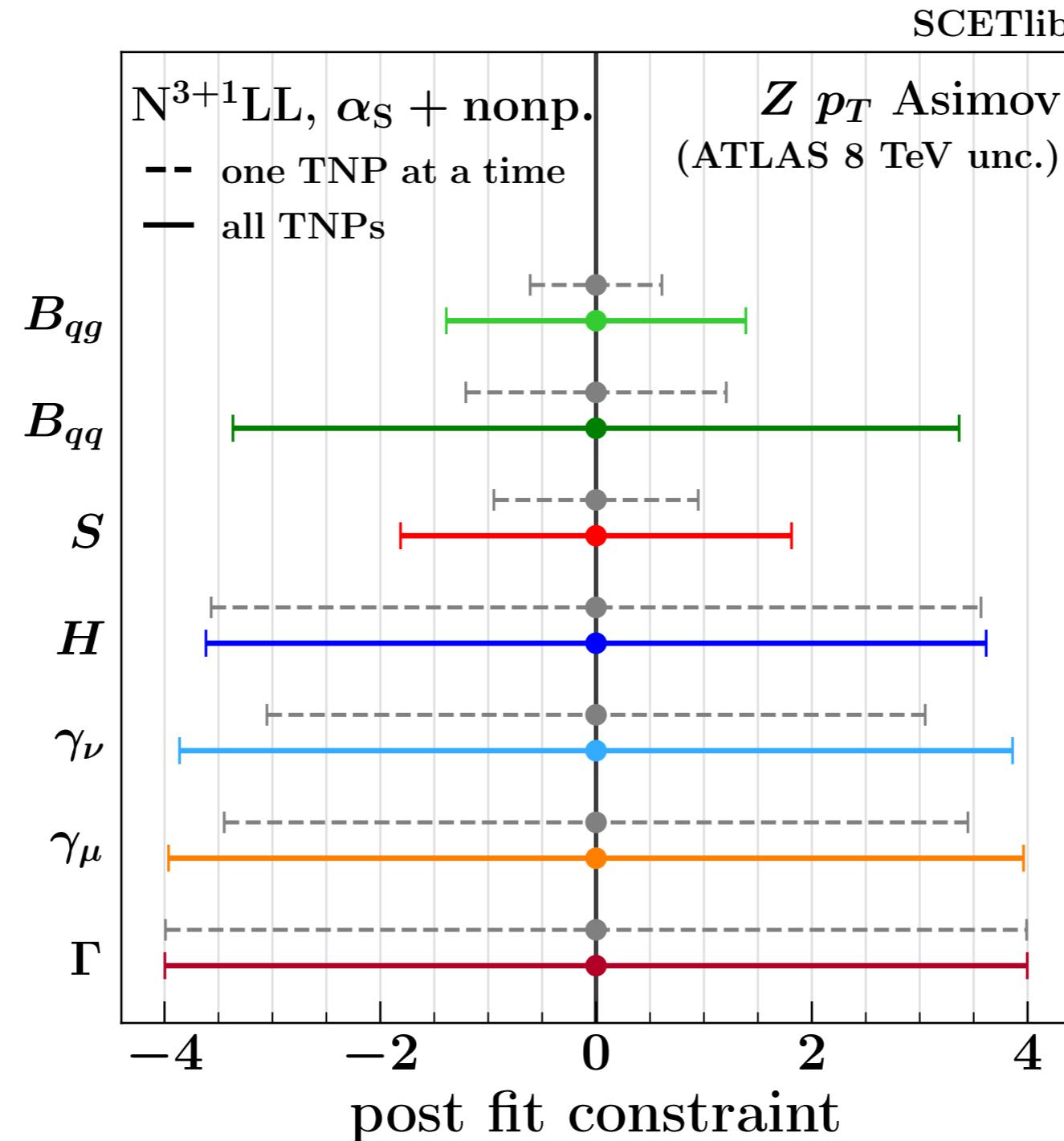
Fit  $N^{3+1}\text{LL}$  against  $N^{3+1}\text{LL}$  data



# Different constraints on TNPs including nonp.

Using now  $\theta_n = 0 \pm \Delta\theta_n$  with  $\Delta\theta_n = 4$

Fit  $N^{3+1}\text{LL}$  against  $N^{3+1}\text{LL}$  data



# Acknowledgments

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**European Research Council**

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