



de Sitter vacua, swampland and supergravity

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Plan:

- Introduction
- Restrictions on gaugings
- Aspects of non-linear SUSY
- Outlook

Introduction

What if string theory has no de Sitter vacua?

See e.g. Dine, Seiberg '85, Danielsson, Van Riet '18

- We observe an expanding universe, and the simplest explanation is to have some dark energy dominating.
- On top of that, long-lived meta-stable de Sitter vacua would be the simplest (but maybe misleading?) solution.
- Extremely hard to find in controlled string theory setups - **maybe impossible!**

How can we uncover this difficulty in 4D supergravity?

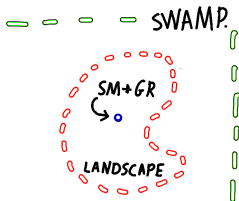
→ Scalar potentials of 4D supergravity typically arise from:

- ▶ Superpotentials.
- ▶ Gaugings.
- ▶ Non-linear supersymmetry.

The vast freedom in choosing the above ingredients makes it impossible to exclude de Sitter vacua in 4D supergravity.

→ Can we find restrictions / implications from the **swampland**?

- ✗ Swamp: EFTs that **do not** arise from S.T.
- ✓ Landscape: EFTs that **do** arise from S.T.



We distinguish swamp from landscape with so-called **conjectures**: *General properties that we guess and test on compactifications, black holes, quantum gravity, etc.*

**We are talking about EFTs coupled to gravity.*

Which conjectures can restrict the ingredients of supergravity?

- ▶ Weak gravity conjecture (magnetic) gives UV cut-off

$$qg M_P > \Lambda_{UV} .$$

Arkani-Hamed, Motl, Nicolis, Vafa '06

- ▶ Non-supersymmetric (AdS) vacua (supported from flux) always decay.

Ooguri, Vafa '16

- ▶ Non-existence of stable de Sitter has been deemed as conjecture too; but we don't want to use it here.

Obied, Ooguri, Spodyneiko, Vafa '18, Andriot '18

Restrictions on de Sitter vacua from gaugings

Example: 4D N=1 Lagrangian of SG + FI term is

$$e^{-1}\mathcal{L} = -\frac{1}{2}M_P^2 R + \frac{1}{2}\epsilon^{\kappa\lambda\mu\nu} (\bar{\psi}_\kappa \bar{\sigma}_\lambda D_\mu \psi_\nu - \psi_\kappa \sigma_\lambda D_\mu \bar{\psi}_\nu) \\ - \frac{1}{4g^2} F_{\mu\nu} F^{\mu\nu} + i\xi \epsilon^{\kappa\lambda\mu\nu} \bar{\psi}_\kappa \bar{\sigma}_\lambda \psi_\mu A_\nu - 4g^2 \xi^2 M_P^4 .$$

- ▶ Since ξ is the gravitino charge, the WGC gives

$$\Lambda_{UV} \sim g \xi M_P .$$

- ▶ The vacuum energy is $4g^2 \xi^2 M_P^2$ and so the Hubble is

$$H \sim g \xi M_P \sim \Lambda_{UV} ,$$

therefore such simple de Sitter is not a good 4D EFT!

Example: FI-gauged 4D N=2 SG + N=2 vect.mult. has potential

$$V = \frac{1}{1 - |z|^2} \left[\xi_0^2 (|z|^2 - 3) - 2\xi_0\xi_1(z + \bar{z}) + \xi_1^2(1 - 3|z|^2) \right],$$

where $\xi_{0,1}$ are FI-gaugings and z a complex scalar.

- ▶ If we set $\xi_0 = 0$ and $\xi_1 = 2gq_{3/2}$ we find $H \sim \Lambda_{UV}$ since

$$\langle V \rangle = 4g^2 q_{3/2}^2,$$

albeit tachyonic.

- ▶ Notably for $\xi_0 \neq 0$ and $\xi_1 = 0$ one can show that such SUSY AdS are not scale-separated. *Cribiori, Dall'Agata '22*

See e.g. *Cribiori, FF '23*

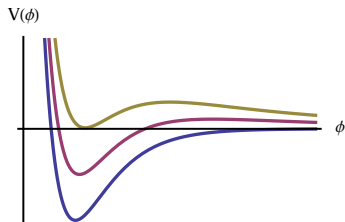
Further results:

- ▶ We have shown that **all known stable de Sitter vacua of gauged N=2** suffer from the same issue. *Cribiori, Dall'Agata, FF '20, Dall'Agata, Emelin, FF, Moritsu '21*
- ▶ All these examples have $m_{3/2} = 0$, $q_{3/2} \neq 0$, so **also in tension with the “festina lente” bound**. *Montero, Vafa, Van Riet, Venken '21*
- ▶ SG has a plethora of other couplings that could indirectly restrict de Sitter, even without gaugings (Yukawa couplings, magnetic moments, etc.).

Decay of vacua from non-linear realizations

KKLT adapted to supergravity with non-linear realizations:

1. No-scale Kähler with
 $W = W_0 + Ae^{-a\Phi}$.
2. + nilpotent chiral goldstino
 $K + X\bar{X}$ and $W + fX$.
3. Lift to metastable de Sitter.



Ferrara, Kallosh, Linde '14

We have broken supersymmetry, which according to the swampland, **should always have a decay channel to SUSY**.

- ▶ SUSY broken with a chiral superfield

$$X = A + \sqrt{2}\theta^\alpha G_\alpha + \theta^\alpha\theta_\alpha F,$$

that is constrained to satisfy

$$X^2 = 0.$$

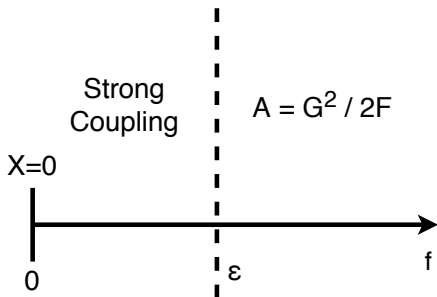
Rocek '78, Casalbuoni, De Curtis, Dominici, Feruglio, Gatto '89

- ▶ The equation $X^2 = 0$ is solved as

I) $A \equiv \frac{G^2}{2F}$ assuming $\langle F \rangle \sim f \neq 0 \rightarrow$ *Volkov – Akulov*.

II) If $\langle F \rangle = 0$ then $X^2 = 0 \rightarrow X = 0$.

- ▶ Solution II) neutralizes all fluctuations of X ('evaporates').



PUZZLE: To get SUSY the EFT passes from small f , but we cannot pass from there because the EFT breaks down.

Decay channel:

- ▶ Trade the auxiliary field of X with a 3-form:

$$F \sim \star dC_3$$

- ▶ Couple the 3-form to a super-membrane with charge Q , and assuming the decay to SUSY vacuum in one go

$$Q \sim f,$$

because $\Delta F \sim Q$.

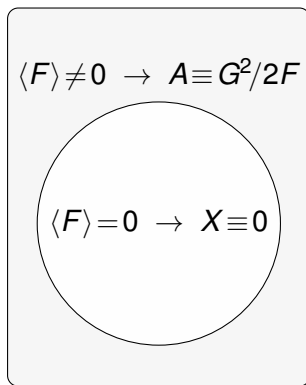
- ▶ The super-membrane tension T_3 also implies non-linear supersymmetry on its world-volume with

$$m \sim T_3^{1/3}.$$

We can solve $X^2 = 0$
depending on the background:

1. $\star dC_3 \neq 0 \rightarrow A = \frac{G^2}{2F}$.

2. $\star dC_3 = 0 \rightarrow X = 0$.



Decay time:

- ▶ Generically these can still be long-lived vacua depending on the properties of T_3 and Q .
- ▶ If we **assume** the two non-linear supersymmetries have the same strong coupling scale

$$f \sim m^2,$$

then we get from BT **very short life-time for dS** because

$$P \sim e^{-B} \sim e^{-\frac{27\pi^2}{2}}.$$

- ▶ Incidentally this also realizes the WGC condition $T_3 < Q$.

Outlook

What if string theory has no de Sitter vacua?

Then we may be able to see this manifest in 4D supergravity equipped with established swampland conjectures:

- ▶ We can uplift SUSY AdS with pure **NL SUSY** → then it can decay via BT bubbles; constraints on the decay time?
- ▶ We can break SUSY and uplift with **gaugings** → then it can violate WGC or FL.
- ▶ We can break it with N=1 **superpotentials** → restrictions?

Thank you