# Gauged Global Strings

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#### BSM in Particle Physics and Cosmology



arXiv: 2311.07639 with Xuce Niu and Fengwei Yang

#### Outline

- Introduction global strings and gauge strings
- Gauge  $U(1)_Z \times \text{global } U(1)_{PQ}$ and string solutions
- Cosmological implication

  rich string structure
  opening up QCD axion window
  gauge string radiating axions?
- Conclusion

•

global U(1) symmetry breaking 
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• tension

gradient term 
$$\mu \simeq 2\pi \int_{m^{-1}}^{L} dr \frac{1}{r} |\partial_{\theta} \Phi(r,\theta)|^2 = \pi f_a^2 \ln(mL)$$

## Gauge strings



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• gauge string solution

$$\Phi(r,\theta) = \frac{1}{\sqrt{2}} f_a e^{i\theta}$$
$$Z_{\mu} = \frac{1}{e} \partial_{\mu} \theta \qquad r \to \infty$$

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$$\mu \simeq 2\pi \int_{m^{-1}}^{L} dr \left| \left( \frac{1}{r} \partial_{\theta} - ieZ_{\mu} \right) \Phi(r, \theta) \right|^{2} = 0$$
  
core  $\mu \simeq \mathcal{O}(1)\pi f_{a}^{2}$ 

### Motivation of cosmic strings

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• theoretically interesting classical field solutions

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• theoretically interesting classical field solutions



 phenomenological rich cosmology (Kibble mechanism) axion dark matter abundance new observables (CMB, ...)

# $U(1)_{Z} \times U(1)_{PQ}$

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• Lagrangian

$$\begin{aligned} \mathscr{L} &= -\frac{1}{4} Z_{\mu\nu} Z^{\mu\nu} + D_{\mu} \Phi_{1}^{\dagger} D^{\mu} \Phi_{1} - \frac{\lambda_{1}}{4} \left( |\Phi_{1}|^{2} - \frac{v_{1}^{2}}{2} \right)^{2} + D_{\mu} \Phi_{2}^{\dagger} D^{\mu} \Phi_{2} - \frac{\lambda_{2}}{4} \left( |\Phi_{2}|^{2} - \frac{v_{2}^{2}}{2} \right)^{2} \\ D_{\mu} &= \partial_{\mu} - ieZ_{\mu} \\ \text{assume that } v_{1} > v_{2} \end{aligned}$$

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• 
$$\Phi_1 \rightarrow \Phi_1 e^{i\alpha_z + i\alpha_{PQ}}$$
  
 $\Phi_2 \rightarrow \Phi_2 e^{i\alpha_z - i\alpha_{PQ}}$ 
 $U(1)_Z$ 
 $1$ 
 $1$ 
 $U(1)_{PQ}$ 
 $1$ 
 $-1$ 

• cross section of vacuum manifold



• cross section of vacuum manifold



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- axion direction is orthogonal to the longitudinal mode of  $Z^{\mu}$ 

$$a(x) = v_a \alpha_{PQ}, \quad v_a = \frac{2v_1 v_2}{\sqrt{v_1^2 + v_2^2}} \sim 2v_2$$

• KSVZ-like model

introduce  $Q_L$  and  $Q_R$  with color charge and U(1)<sub>PQ</sub> charge

$$\mathcal{L} = -\frac{y}{\Lambda} \left( \Phi_1 \Phi_2^* \bar{Q}_L Q_R + h \cdot c \cdot \right)$$

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• Barr and Seckel's model  $Q_{1L} Q_{1R} Q_{2L} Q_{2R}$ color, U(1)<sub>Z</sub> and U(1)<sub>PQ</sub> charges

$$\mathcal{L} = \Phi_1 \bar{Q}_{1L} Q_{1R} + \Phi_2 \bar{Q}_{2L} Q_{2R} + h \cdot c \,.$$

## String Solutions



• (1,0) string  $\Phi_1 = \frac{1}{\sqrt{2}} v_1 e^{i\theta}, \quad \Phi_2 = \frac{1}{\sqrt{2}} v_2, \quad Z_\mu = c \partial_\mu \theta, \quad r \to \infty$  ф,

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• gradient energy

$$\mu_{k,(1,0)} = \int_{0}^{2\pi} d\theta \int_{\delta}^{L} dr \, r \left( \left| \left( \frac{1}{r} \partial_{\theta} - ieZ_{\theta} \right) \Phi_{1} \right|^{2} + \left| \left( -ieZ_{\theta} \right) \Phi_{2} \right|^{2} \right) \right)$$
$$= \pi \ln(\frac{L}{\delta}) \left[ v_{1}^{2} (1 - ec)^{2} + v_{2}^{2} (ec)^{2} \right]$$

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• outside core ( minimize it by varying c )

$$\mu_{k,(1,0)} = \pi \frac{v_1^2 v_2^2}{v_1^2 + v_2^2} \ln(\frac{L}{\delta}) = \pi f_a^2 \ln(\frac{L}{\delta})$$

• (0,1) string  

$$\Phi_1 = \frac{1}{\sqrt{2}} v_1, \quad \Phi_2 = \frac{1}{\sqrt{2}} v_2 e^{i\theta}, \quad Z_\mu = c \,\partial_\mu \theta, \quad r \to \infty$$

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(1,0) string is equivalent to (0,-1) string through a gauge transformation

$$\left(\Phi_1 = \frac{1}{\sqrt{2}} v_1 e^{i\theta}, \Phi_2 = \frac{1}{\sqrt{2}} v_2\right) \xrightarrow{\alpha_Z \to \alpha_Z - \theta} \left(\Phi_1 = \frac{1}{\sqrt{2}} v_1, \Phi_2 = \frac{1}{\sqrt{2}} v_2 e^{-i\theta}\right)$$

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(1,0) can be viewed as an anti-string of (0,1)

# $(1,0) + (0,1) \rightarrow ?$



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# (1,1) strings

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ф

 $\phi_{z}$ 

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- the profile of  $Z_{\theta}$  can simultaneously cancel the gradient energy of  $\Phi_1$  and  $\Phi_2$ 

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• (1,1) gauge string

(1,0) and (0,1) global strings

• magnetic self-energy, scalar potential energy, and gradient energy

- magnetic self-energy, scalar potential energy, and gradient energy
- (1,0) string

$$\mu_{(1,0)} \simeq \pi v_1^2 + \pi v_1^2 \ln\left(\frac{m_1}{m_Z}\right) + \pi v_2^2 \ln\left(\frac{m_Z L}{2}\right)$$

(0,1) string

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(1,1) string

$$\mu_{(1,1)} = \pi v_1^2 + \pi v_1^2 \ln\left(\frac{m_1}{m_Z}\right) + 0$$

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- heavy core of (1,0) string  $\mu_{(1,0)} > \mu_{(0,1)}$
- binding energy of (1,1) string  $\mu_{(1,0)} + \mu_{(0,1)} - \mu_{(1,1)} = \pi v_2^2 \left[ 2 \ln \left( \frac{m_Z L}{2} \right) - 1 \right]$



• consider  $v_1 \gg v_2$ 

first phase transition  $\langle \Phi_1(x) \rangle = \frac{v_1}{\sqrt{2}}$  and  $\langle \Phi_2(x) \rangle = 0$ 



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U(1) gauge strings form
 (1, n) string



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  - (0,1) strings form via Kibble mechanism

• (1, n) string  $\rightarrow ?$ 

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•  $(1, \mathbf{n})$  string  $\rightarrow ?$ 



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•  $(1, \mathbf{n})$  string  $\rightarrow ?$ 



• (1, n) string  $\rightarrow$  (1, 0) string to minimize the energy



• (1,0) string encounters (1,0) string

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 Y-junctions

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 Other works on simulations of Y-junctions found 1) some fraction of Y-junctions remain 2) scaling solution

Urrestilla, Vilenkin JHEP(2008) Rajentie, Skellariodou, Stoica, JCAP (2007) Copeland, Saffin JHEP (2005)

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• dark matter abundance

misalignment + string radiation + domain wall collapse

 $\rho_{a,0} = \rho_a^{\text{vac}}(t_0) + m_a n_a^{\text{str}}(t_0) + m_a n_a^{\text{DW}}(t_0)$ 



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# Gauged global string

## Gauged global string

- (0,1) string tension is the same as a standard QCD axion string
- heavy core of (1,0) string

$$\mu_{(1,0)}(t) \simeq \pi v_1^2 \ln\left(\frac{m_1}{m_Z}\right) + \pi f_a^2 \ln\left(\frac{m_Z t}{2}\right)$$








#### QCD axion window



#### Future explorations



# Gauged global string



• gauge strings radiate gravitons

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$$\mathcal{L} = -\frac{1}{4} Z_{\mu\nu} Z^{\mu\nu} + \frac{1}{2} e^2 (\phi_1^2 + \phi_2^2) Z_{\mu}^2 - g(\phi_1, \phi_2) e Z^{\mu} \partial_{\mu} a + \frac{1}{2} f(\phi_1, \phi_2) (\partial_{\mu} a)^2$$

$$g(\phi_1, \phi_2) = f_a \frac{\phi_1^2}{v_1^2} - f_a \frac{\phi_2^2}{v_2^2}$$

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• Kalb-Ramond field  $B^{\mu\nu}$ ,  $\partial_{\mu}a = \frac{1}{2} \epsilon_{\mu\nu\alpha\beta} \partial^{\nu} B^{\alpha\beta}$ 

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• radiation power  $\frac{\mathrm{d}P_a}{\mathrm{d}\Omega} \sim e^2 f_a^2$ 

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# Conclusion





• U(1)<sub>Z</sub> × U(1)<sub>PQ</sub> (1,0), (0,1) and (1,1) strings

Cosmology Y-Junctions opening QCD axion mass windows

• (1,1) gauge string radiating axions and gravitons