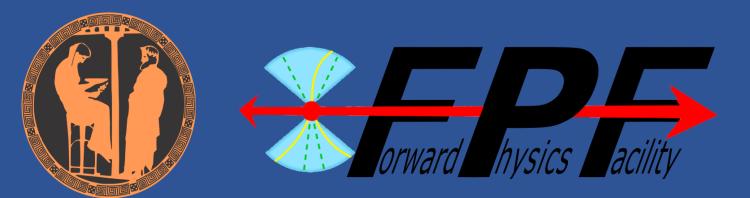
Tuning Pythia for Forward Physics Experiments Max Fieg 2309.08604 + PRD

In collaboration with Felix Kling, Holger Schulz, Torbjörn Sjöstrand





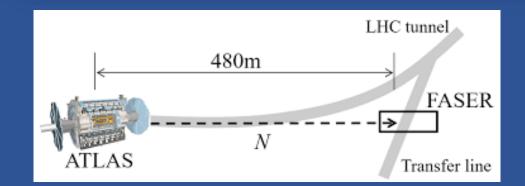
FASER and the Forward Physics Facility (FPF)

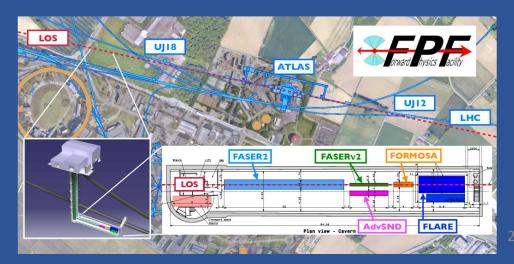
In the forward, large η , region at the LHC there is an intense flux of hadrons which can decay to, e.g., neutrinos or BSM states

The Forward Search Experiment (FASER) sits 480m downstream from the ATLAS IP and is looking for the decays of long-lived particles

The proposed Forward Physics Facility program, would carve out a cavern along the beamline to host a suite of experiments with different technologies







FASER and the Forward Physics Facility (FPF)

First FASER results already in! Dark photon bounds and collider neutrino discovery

> CERN-FASER-CONF-2023-001 29 March 2023

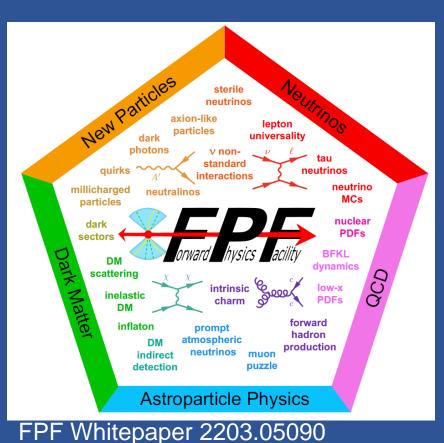
First Results from the Search for Dark Photons with the FASER Detector at the LHC

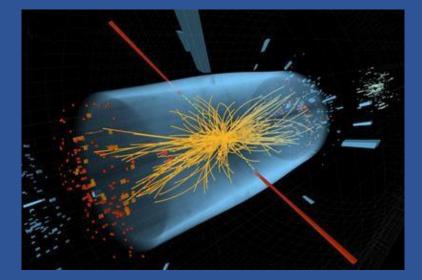
FASER Collaboration

First Direct Observation of Collider Neutrinos with FASER at the LHC

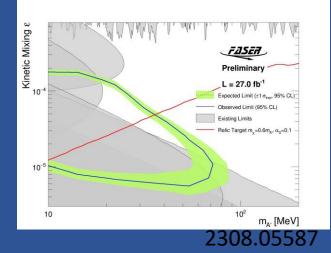
FASER Collaboration

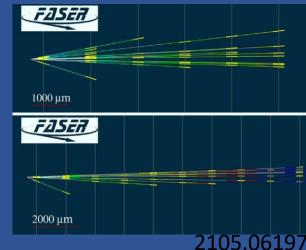
 If the FPF is approved, we are in for a broad physics program that requires careful study of forward physics





 Monte Carlo event generators used for LHC are tuned to central physics and have excellent agreement





 Forward physics studies require an understanding of forward (light) hadron production

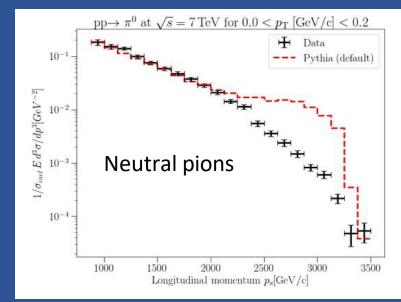


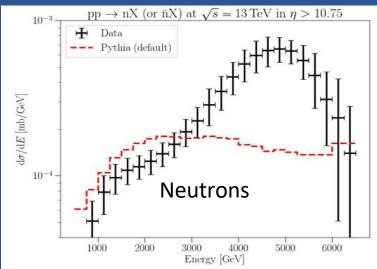


Let's tune Pythia for forward physics without spoiling the success in the central region

Main problem

- LHCf has measured neutral pions, neutrons, and photons (aka pions) at $\sqrt{s} = 7,13$ TeV.
 - Expect similar hadronization mechanism at each energy
 - π^{\pm} important for u_{μ} production
 - Charm decay important for all flavors at high energies – see 2309.12793
- Central Pythia tunes do not describe forward particle fluxes measured by LHCf
 - Other generators don't do very well either
- Use forward measurements from LHCf as our target and tune hadronization parameters
 - bonus if we can minimize the impact on central predictions







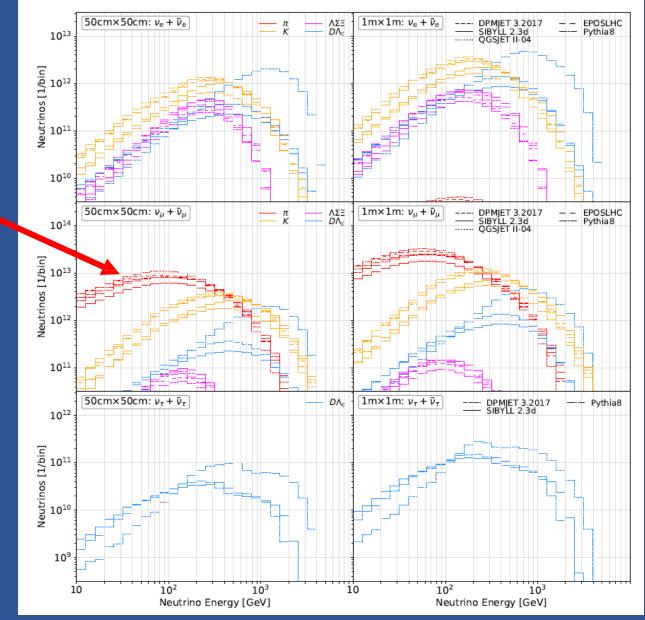
Second problem

Different generators can give very different hadron / neutrino fluxes

How can we get a handle on flux uncertainties?

One method sometimes taken is to take the spread of generators' predictions

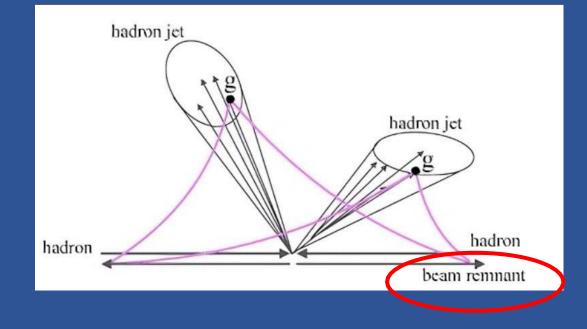
 But this is too dependent on the weakest generator... Need something more robust



1. Pythia tuning methods

- Maximize success in fitting forward production while minimizing impact on central physics
- 2. Tuning uncertainties
- Provide a tuning uncertainty which translates to a flux uncertainty
- 3. Applications at FASER
- Demonstrate tune for applications

Tuning methods: beam remnant After a coarse scan through many parameters find a subset of tuning parameters which are important for forward physics. Those that are associated with the <u>beam remnant</u>



We tune parameters relating to:

• <u>Primordial kT</u> of incoming partons to tune overall normalization

BeamRemnants:primordialKTremnant

BeamRemnants:primordialKTsoft

<u>Remnant → baryon fragmentation</u>
 <u>function</u> to produce more hard neutrons

BeamRemnants:hardRemnantBaryon

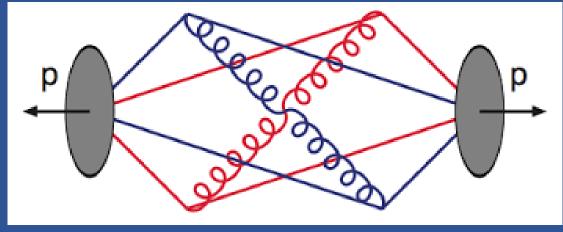
BeamRemnants:bRemnantBaryon

 Reduce "<u>Popcorn production</u>" to produce fewer hard mesons from remnant diquarks

Tuning Methods: Color Reconnection (CR)

As baseline tunes, we compare the Monash tune vs. a central tune based on QCD Color Reconnection (1505.01681)

Here, explicit colors are assigned to partons in an MPI and string reconnections can occur if they reduce the total string length



We find that that using the QCD CR tune as our baseline is some improvement over our Monash based tune

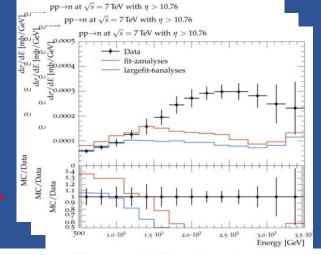
Tuning methods

With parameters identified, we generate and fit the parameters to data



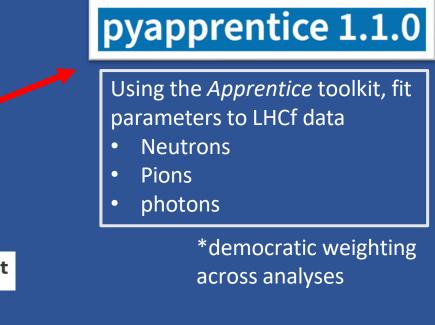


Generate events in Pythia across tuning space



Rivet — the particle-physics MC analysis toolkit

Fill out LHCf histograms for pion, neutron and photon analyses at 7 and 13 TeV

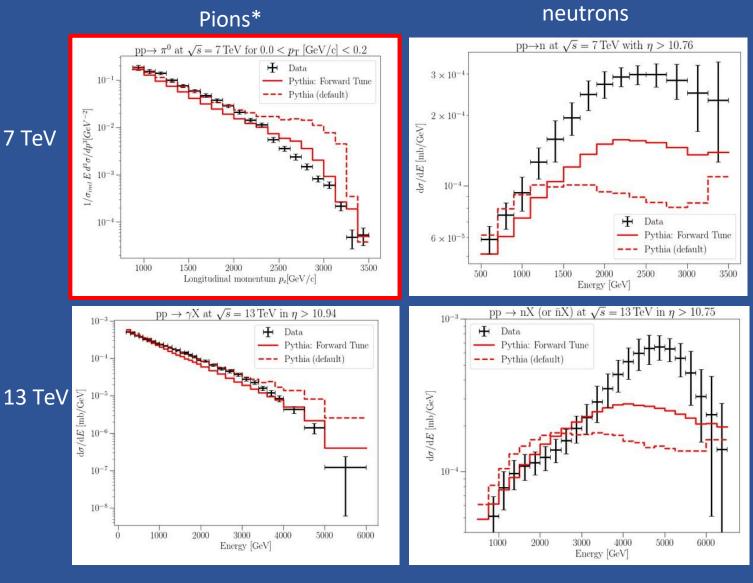


Tuning results

Excess hard pions reduced by disabling the "popcorn mechanism": forces a remnant diquark to form a baryon

Independent handle on baryons by modifying diquark → baryon fragmentation function

Flux normalization controlled by fitting primordial parton p_T: "kT_{remn}"

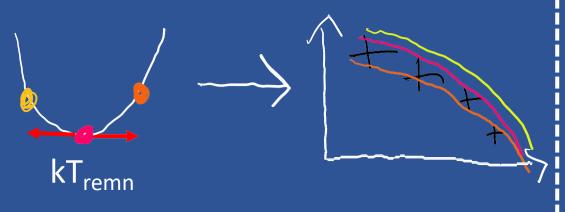


Can we define an uncertainty that captures imperfections in our tune? a naïve $\Delta\chi^2$ returns an unreasonable underestimate of uncertainties ¹¹

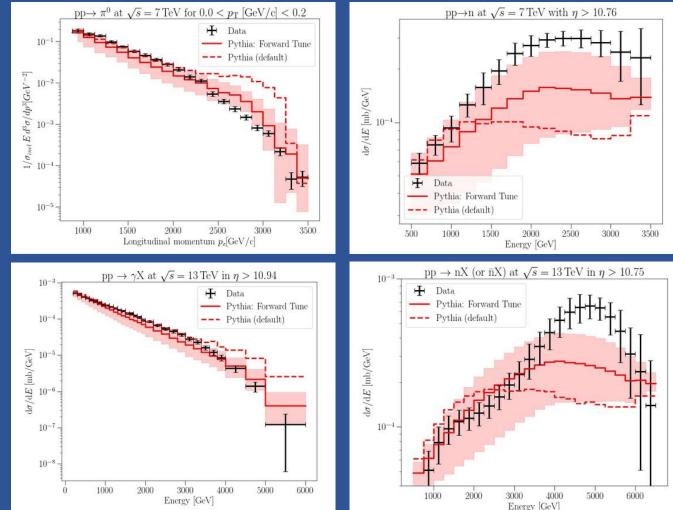
Tuning uncertainties

We reduce to the most sensitive tuning parameter (kT_{remn}) and take a pragmatic data-driven approach

- Define a band specified by $kT_{remn} \pm \Delta$
- Increase ∆ from best fit until 68% of the datapoints are contained in the band



By construction, result is a band enveloping 68% of data, resembling 1σ



How does Monash Compare?

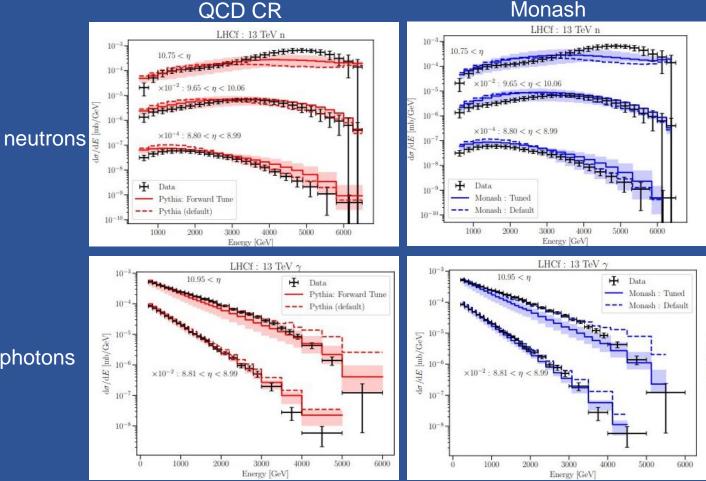
Tuning Results: Monash vs. QCD CR

Monash tune is comparable but with some notable deficiencies

QCD CR better predicts the shape of the forward neutron spectra, Monash predicts more soft neutrons

Monash also underpredicts the photon spectra

~ 20% overall improvement of QCD CR over Monash



Tuning Results

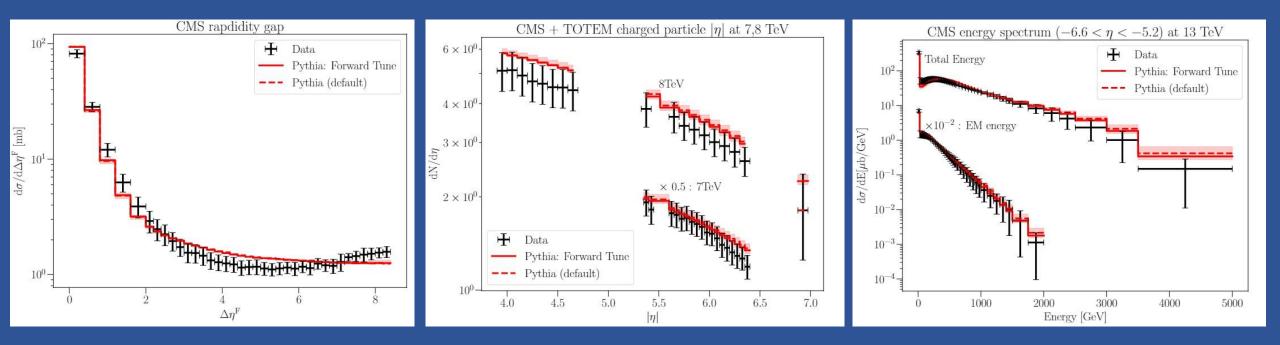
| Full name | Shorthand | Baseline (QCDCR) | Forward Tune | Uncertainty |
|--|--------------------|------------------|--------------|-------------------|
| BeamRemnants:dampPopcorn | $d_{ m pop}$ | 1 | 0 | |
| ${\tt BeamRemnants:hardRemnantBaryon}$ | $f_{ m remn}$ | off | on | |
| ${\tt BeamRemnants:aRemnantBaryon}$ | $a_{ m remn}$ | - | 0.36 | |
| ${\tt BeamRemnants: bRemnantBaryon}$ | $b_{ m remn}$ | - | 1.69 | |
| BeamRemnants:primordialKTsoft | $\sigma_{ m soft}$ | 0.9 | 0.58 | $0.26 \dots 1.27$ |
| BeamRemnants:primordialKThard | $\sigma_{ m hard}$ | 1.8 | 1.8 | |
| BeamRemnants:halfScaleForKT | $Q_{ m half}$ | 1.5 | 10 | |
| BeamRemnants:halfMassForKT | $m_{ m half}$ | 1 | 1 | |
| BeamRemnants:primordialKTremnant | $\sigma_{ m remn}$ | 0.4 | 0.58 | $0.26 \dots 1.27$ |
| | | | | |

*Some details skipped over here, see paper or ask me for details

Did we spoil success in the central region , at CMS, ATLAS or even TOTEM?

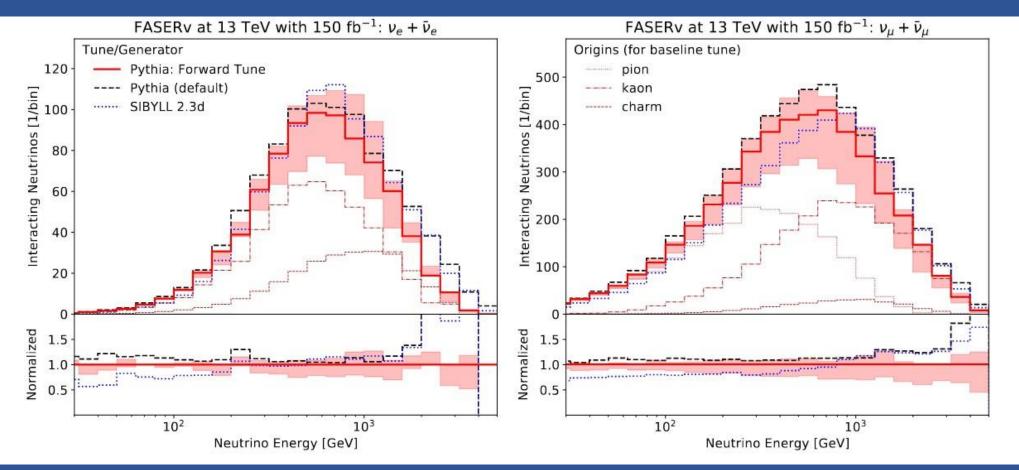
Impact on central physics

Some "central" analyses where we would most likely see effect of tuning



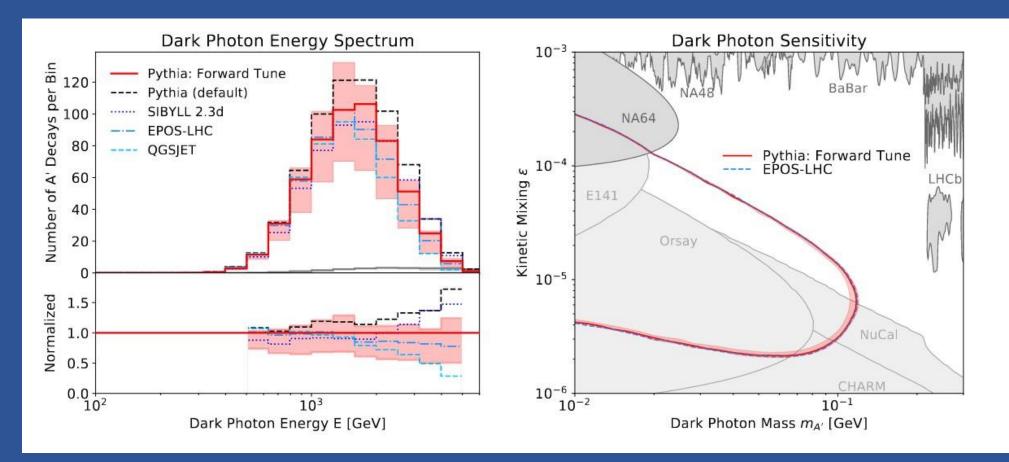
Applications for forward physics - Neutrinos

 Interacting electron and muon neutrino spectrum at FASER. Our improved tune predicts ~10% fewer neutrinos as compared to the default Pythia configuration, and we find a ~20% uncertainty band



Applications for forward physics – Dark Photons Dark photon spectra for fixed $m_{A'}$, ε and dark photon reach plot

-About 50% uncertainty in number of dark photon decays. Reach is largely unaffected due to large ε suppression

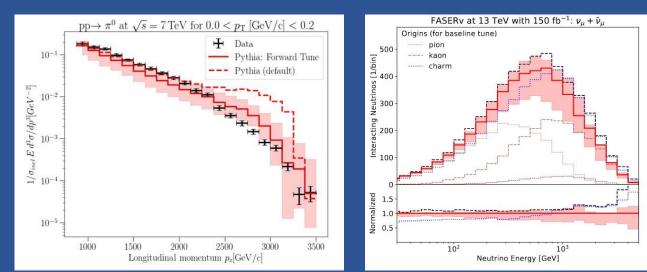


Summary

- We tune Pythia for forward physics purposes at the LHC, by fitting beam remnant parameters which have negligible impact on central physics
- We provide a data-driven uncertainty estimate
- We demonstrate an application of our tune by showing its impact on neutrino and dark photon measurements at FASER

Thank you for listening!

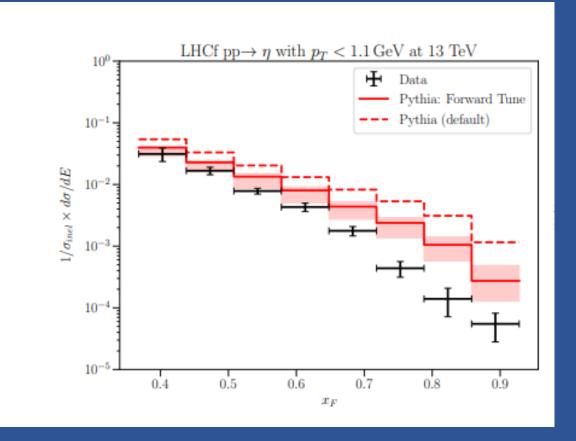
| Full name | Shorthand | Baseline (QCDCR) | Forward Tune | Uncertainty |
|----------------------------------|---------------------|------------------|--------------|-------------|
| BeamRemnants:dampPopcorn | $d_{ m pop}$ | 1 | 0 | |
| BeamRemnants:hardRemnantBaryon | fremn | llo | on | |
| BeamRemnants:aRemnantBaryon | $a_{ m remn}$ | 122 | 0.36 | |
| BeamRemnants:bRemnantBaryon | bremn | - | 1.69 | |
| BeamRemnants:primordialKTsoft | $\sigma_{\rm soft}$ | 0.9 | 0.58 | 0.261.27 |
| BeamRemnants:primordialKThard | $\sigma_{ m hard}$ | 1.8 | 1.8 | |
| BeamRemnants:halfScaleForKT | $Q_{\rm half}$ | 1.5 | 10 | |
| BeamRemnants:halfMassForKT | $m_{ m half}$ | 1 | 1 | |
| BeamRemnants:primordialKTremnant | $\sigma_{\rm remn}$ | 0.4 | 0.58 | 0.261.27 |

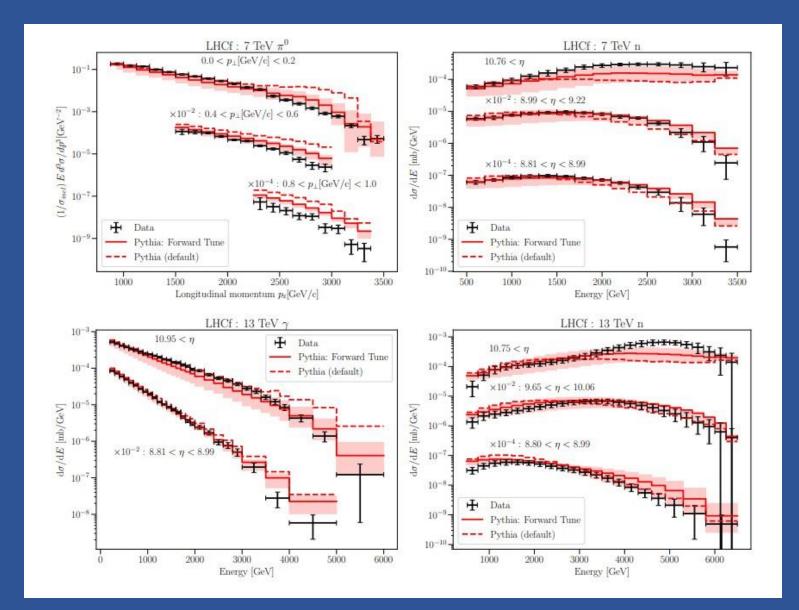




Back up

Eta analysis





Variation of MultipartonInteractions 3.1 MultipartonInteractions:alphaSvalue 3.4 MultipartonInteractions:ecmPow 3.6 MultipartonInteractions:enhanceScreening Variation of ColourReconnection and BeamRemnants 4.7 BeamRemnants:primordialKTremnant Variation of TimeShower and SpaceShower

Variation of StringPT and StringZ

| 6.1 | StringPT:sigma | 8 | | 2 | 28 | | | | | 4 | • | |
|-----|---------------------------|---|--|---|----|--|--|--|--|---|---|--|
| 6.2 | StringPT:enhancedFraction | | | | | | | | | | | |
| | StringPT:enhancedWidth . | | | | | | | | | | | |
| 6.4 | StringDT:closeDecking | | | _ | | | | | | _ | | |

| Vari | iation of StringFlav | |
|------|-----------------------------|----------------|
| 7.1 | StringFlav:probStoUD | |
| 7.2 | StringFlav:probQQtoQ | |
| 7.3 | StringFlav:probSQtoQQ | |
| 7.4 | | |
| 7.5 | StringFlav:mesonUDvector | |
| 7.6 | StringFlav:mesonSvector | and the second |
| 7.7 | StringFlav:mesonCvector | |
| 7.8 | | |
| 7.9 | StringFlav:etaSup | |
| 7.10 | 0 | |
| 7.11 | | |
| 7.12 | StringFlav:popcornRate | |
| 7.13 | StringFlav:popcornSpair | |
| 7.14 | StringFlav:popcornSmeson | |
| | StringFlav:suppressLeadingB | |

8 Variation of Diffraction

7

| 8.1 | Diffraction:mMinPert | • |
|------|--|---|
| 8.2 | Diffraction:mWidthPert | |
| 8.3 | Diffraction:probMaxPert | |
| 8.4 | Diffraction:pickQuarkNorm | • |
| 8.5 | Diffraction:pickQuarkPower | |
| 8.6 | Diffraction:primKTwidth | |
| 8.7 | $Diffraction: large Mass Suppress \ . \ . \ . \ . \ .$ | • |
| 8.8 | | |
| 8.9 | | |
| 8.10 | $Diffraction:mPowPomP \ \ \ldots $ | |
| 8.11 | Diffraction:bProfile | |
| 8.12 | Diffraction:doHard | 1 |
| | | |

9 Variation of Diffraction (SaS Model)

| 9.1 | PDF:PomSet | |
|-----|------------------------------|--|
| 9.2 | SigmaDiffractive:mMin | |
| | SigmaDiffractive:lowMEnhance | |
| 9.4 | SigmaDiffractive:mResMax | |
| 9.5 | SigmaDiffractive:dampen | |
| | SigmaDiffractive:SaSepsilon | |
| | | |

10 Variation of Diffraction (ABMST Model)

| 10.1 | ${\bf SigmaDiffractive: ABMSTmodeSD}$ | | • | | | | | • | 5 | • |
|------|---|---|---|---|---|----|----|----|---|----------|
| 10.2 | SigmaDiffractive:ABMSTmultSD | | | | ÷ | | 8 | | | æ |
| | SigmaDiffractive:ABMSTpowSD | | | | | | | | | |
| 10.4 | SigmaDiffractive:ABMSTmultDD | | • | ÷ | | | | a. | | |
| 10.5 | SigmaDiffractive:ABMSTpowDD | | | | | | | | | |
| 10.6 | SigmaDiffractive:ABMSTygap | | | 5 | | ्र | i. | | | <u>.</u> |
| 10.7 | ${\small SigmaDiffractive:} ABMSTypow .$ | • | | 5 | | e. | 2 | | | |

Monash

