



(proposal INFN + ENEA in APRENDE)



CERN-INTC-2024-006 / INTC-P-689
10/01/2024

Study of $n + {}^{63,65}\text{Cu}$ reactions and their relevance for nuclear technologies and Astrophysics

M. Bacak¹, D. M. Castelluccio^{2,3}, S. Cristallo^{4,3}, P. Console Camprini^{2,3}, M. Diakaki⁵
G. Grasso², A. Guglielmelli⁶, C. Massimi^{7,3}, M. Mastromarco^{3,8}, A. Mengoni^{2,3},
A. Musumarra^{9,3}, M. P. Pellegriti³, M. Pignatari¹⁰, E. Pirovano¹¹, N. Terranova^{2,3},
R. N. Sahoo³, D. Vescovi^{4,3}, and the n_TOF Collaboration

¹ CERN

² ENEA – Agency for New Technologies, Energy and Sustainable Economic Development, Italy

³ INFN – National Institute for Nuclear Physics, Italy

⁴ INAF – National Institute for Astrophysics, Italy

⁵ NTUA – National Technical University of Athens, Greece

⁶ European Commission, Joint Research Centre, Ispra, Italy

⁷ Department of Physics and Astronomy, University of Bologna, Italy

⁸ Department of Physics, University of Bari, Italy

⁹ Department of Physics and Astronomy, University of Catania, Italy

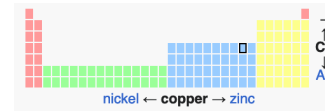
¹⁰ Konkoly Observatory, Research Centre for Astronomy and Earth Sciences, HUN-REN & CSFK, MTA Centre of Excellence, Budapest, Hungary

¹¹ PTB – National Metrology Institute, Germany



ALMA MATER STUDIORUM
UNIVERSITÀ DI BOLOGNA

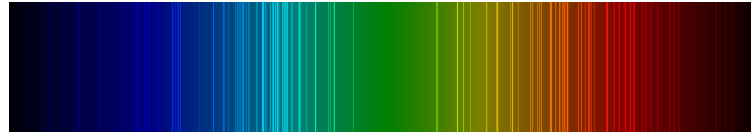
Spokesperson: Cristian Massimi (massimi@bo.infn.it)
Technical coordinator: Olivier Aberle (oliver.aberle@cern.ch)



Native_Copper_Macro_Digon3.jpg: "Jonathan Zander (Digon3)" derivative work: MaterialsScientist (talk) - Native_Copper_Macro_Digon3.jpg

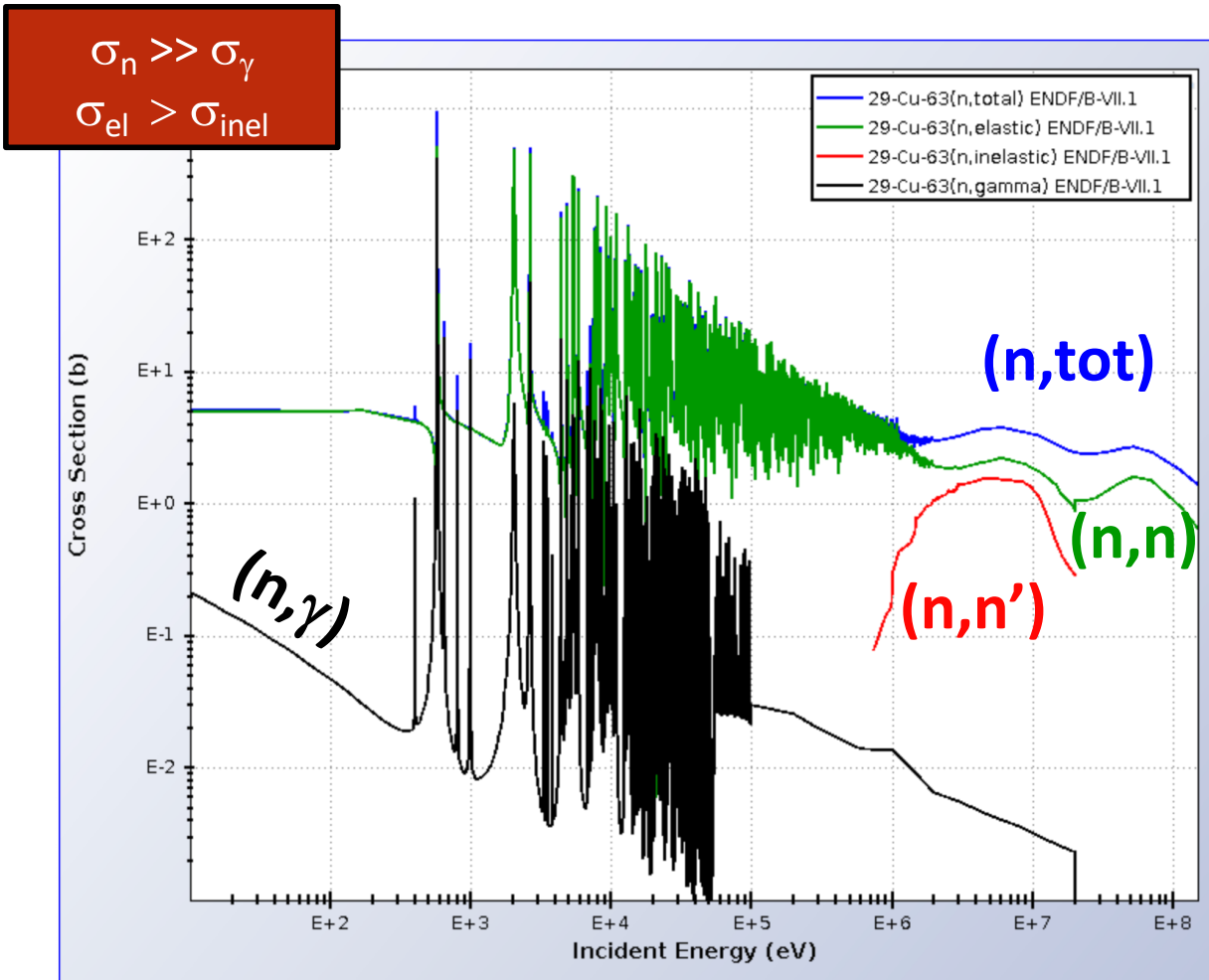


CERN-INTC-2024-006 / INTC-P-689
10/01/2024



- Introduction to scientific motivations
- Data in the literature
- Proposed measurements
- Conclusions

Why $^{63,65}\text{Cu}$?



Natural abundance:

^{63}Cu 69.2 %

^{65}Cu 30.8 %

Prodotti della reazione	Q -valore (keV)
$^{63}\text{Cu} + n$	0
$^{64}\text{Cu} + \gamma$	7915.9 ± 0.6
$^{60}\text{Co} + \alpha$	1717.0 ± 0.6
$^{63}\text{Ni} + p$	715.4 ± 0.6

Prodotti della reazione	Q -valore (keV)
$^{65}\text{Cu} + n$	0
$^{66}\text{Cu} + \gamma$	7065.9 ± 0.9
$^{62}\text{Co} + \alpha$	-193 ± 18
$^{65}\text{Ni} + p$	-1355.5 ± 0.8

Why $^{63,65}\text{Cu}$?

➤ Nuclear Technology

Nucl. Techn.

➤ Nuclear Astrophysics

Nucl. Astro.

^{63}Cu and ^{65}Cu very similar

- ^{63}Cu presented here
- ^{65}Cu in backup slides

Why $^{63,65}\text{Cu}$?

Nucl. Techn.

TAratura Pila Rapida Potenza ZerO (TAPIRO) research reactor:

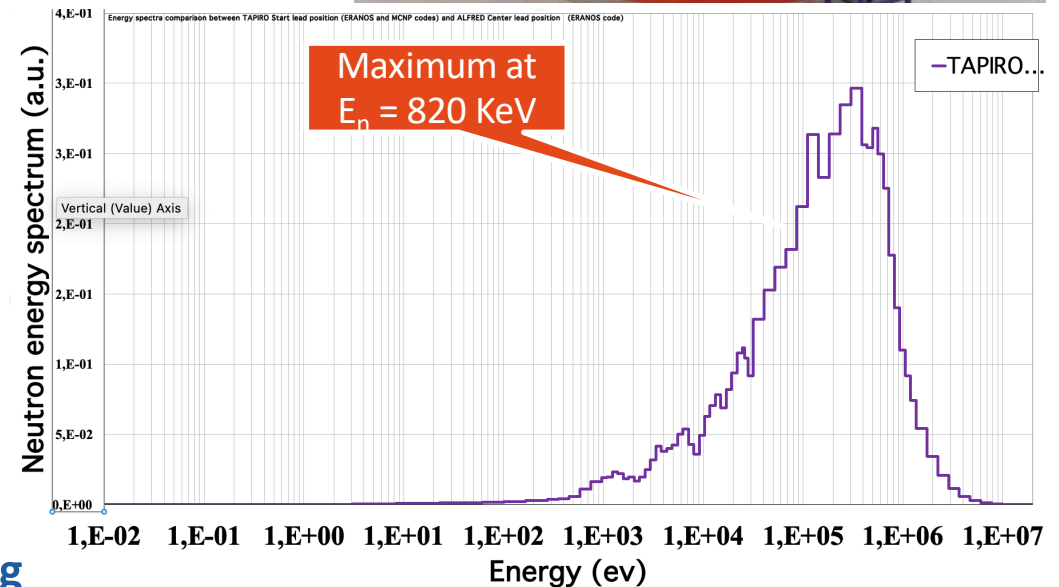
- 5 kW power
- U-Mo fuel
- Core = 12 cm cylinder
- ^{235}U enrichment = 93.5%
- 4×10^{12} n/s

➤ FAST SPECTRUM

Evaluation benchmark

Material test

TAPIRO can play a **pivotal** role in supporting the development of fast reactor projects

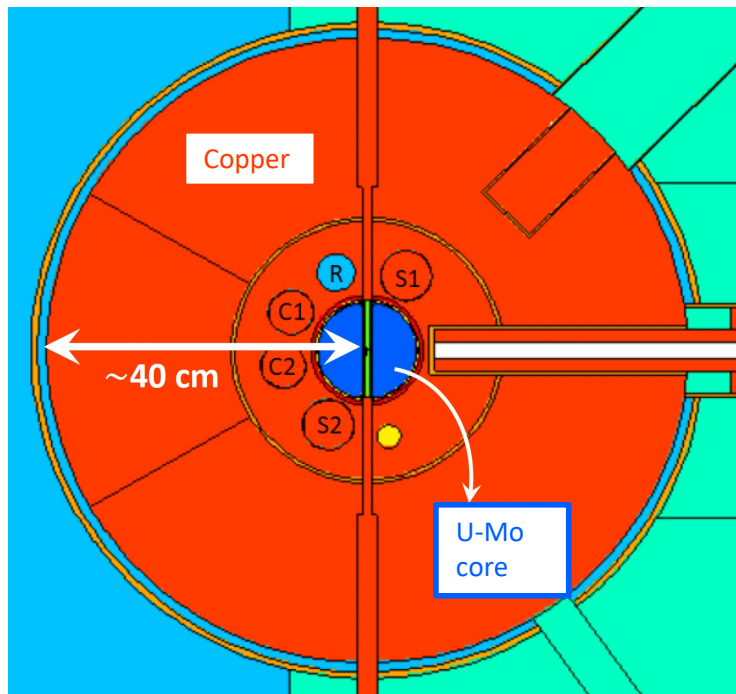


Why $^{63,65}\text{Cu}$?

Nucl. Techn.

TAPIRO: k_{eff} study by MCNP

ENDF/B-VIII.0 $\rightarrow k_{\text{eff}} = 1.00000$



^{63}Cu	Evaluation	k_{eff}
	JEFF3.3	1.00637 ± 0.00001
JENDL-5	1.00147 ± 0.00001	
TENDL-2021	1.00102 ± 0.00001	

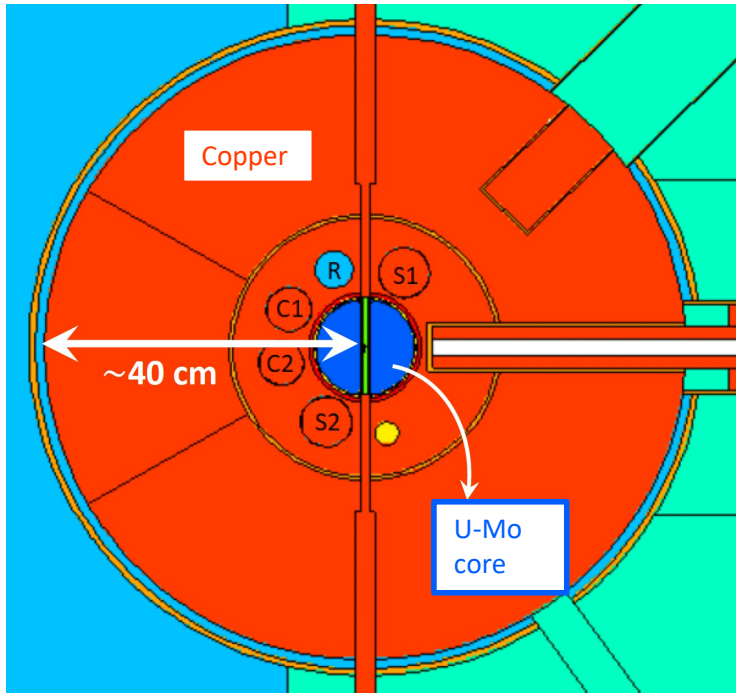
^{65}Cu	Evaluation	k_{eff}
	JEFF3.3	0.99980 ± 0.00001
JENDL-5	0.99782 ± 0.00001	
TENDL-2021	1.00017 ± 0.00001	

Why $^{63,65}\text{Cu}$?

Nucl. Techn.

TAPIRO: k_{eff} study by MCNP

ENDF/B-VIII.0 $\rightarrow k_{\text{eff}} = 1.00000$



^{63}Cu	Evaluation	k_{eff}
	JEFF3.3	1.00637 ± 0.00001
	JENDL-5	1.00147 ± 0.00001
	TENDL-2021	1.00102 ± 0.00001

^{65}Cu	Evaluation	k_{eff}
	JEFF3.3	0.99980 ± 0.00001
	JENDL-5	0.99782 ± 0.00001
	TENDL-2021	1.00017 ± 0.00001

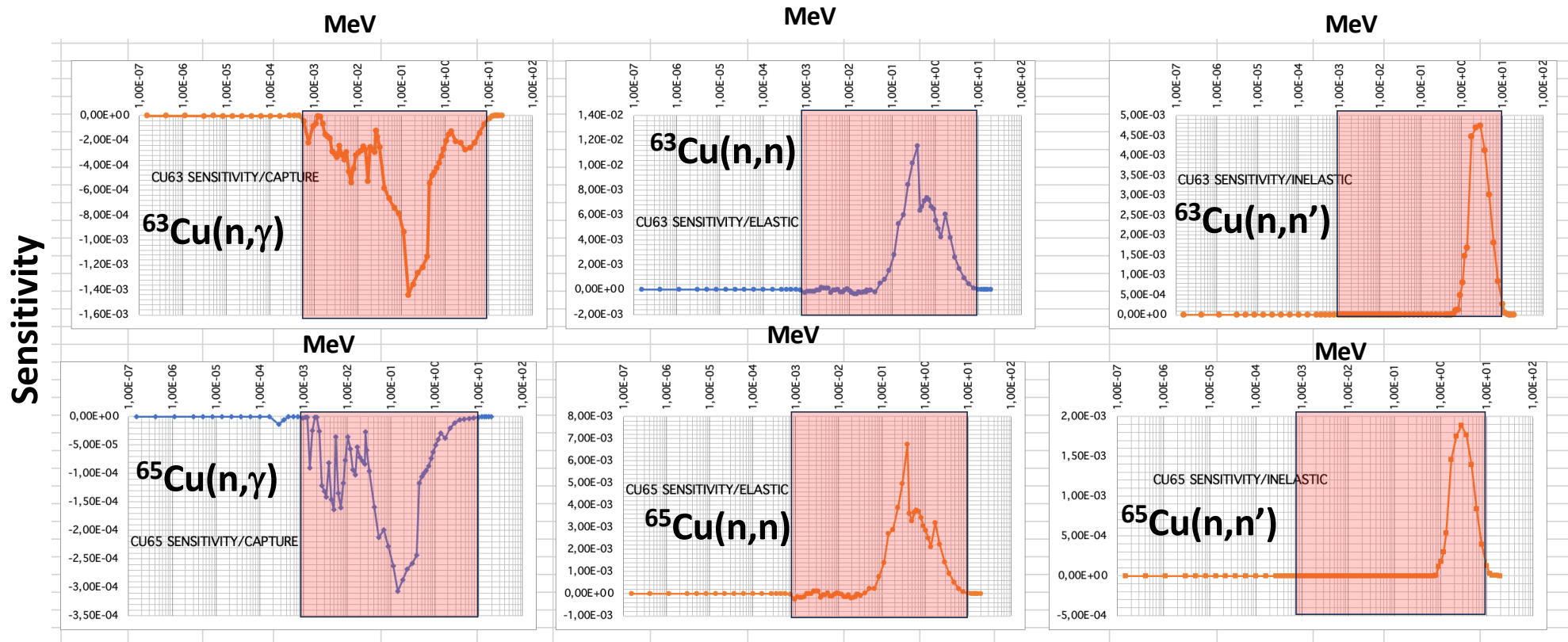
Library effect =
1.5 x regulation rod

Why $^{63,65}\text{Cu}$?



Nucl. Techn.

TAPIRO: spectral parameters sensitivity & uncertainty study by ERANOS 2.3



$10 \text{ keV} < E_n < 300 \text{ keV}$

$50 \text{ keV} < E_n < 5 \text{ MeV}$

$1 \text{ MeV} < E_n < 5 \text{ MeV}$

Why $^{63,65}\text{Cu}$?

Fusion – study at FNG

2016

Fusion Engineering and Design 109–111 (2016) 843–847
Contents lists available at ScienceDirect
Fusion Engineering and Design
journal homepage: www.elsevier.com/locate/fusengdes

Copper benchmark experiment at the Frascati Neutron Generator for nuclear data validation
M. Angelone*, D. Flammini, S. Loreti, F. Moro, M. Pillon, R. Villari
*ENEA Dipartimento Fusione e Tecnologie per la Sicurezza Nucleare, C.R. Frascati, via E. Fermi 45, 00044 Frascati, Italy

HIGHLIGHTS

- A benchmark experiment was performed using pure copper with 14 MeV neutrons.
- The experiment was performed at the Frascati Neutron Generator (FNG).
- Activation foils, thermoluminescent dosimeters and scintillators were used to measure reactions rates (RR), nuclear heating and neutron spectra.
- The paper presents the RR measurements and the post analysis using MCNP5 and JEFF-3.1.1, JEFF-3.2 and FENDL-3.1 libraries.
- C/Es are presented showing the need for deep revision of Cu cross sections.

ARTICLE INFO

Article history:
Received 27 July 2015
Received in revised form 16 December 2015
Accepted 26 January 2016
Available online 6 February 2016

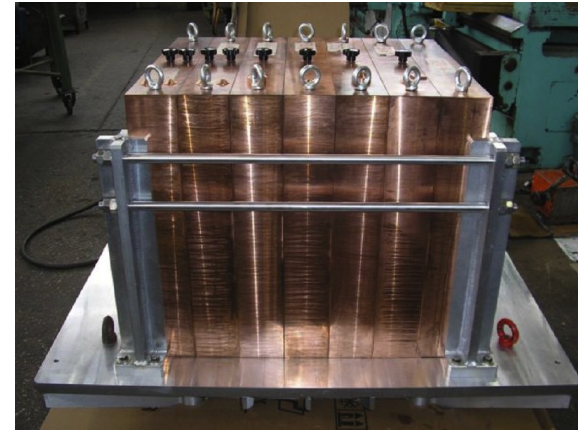
Keywords:
Benchmark experiment
14 MeV neutrons
Copper cross sections
Frascati Neutron Generator
Activation technique
MCNP Monte Carlo code

ABSTRACT

A neutronics benchmark experiment on a pure Copper block (dimensions 60 × 70 × 60 cm³), aimed at testing and validating the recent nuclear data libraries for fusion applications, was performed at the 14-MeV Frascati Neutron Generator (FNG) as part of a F4E specific grant (F4E FPA-395 01) assigned to the European Consortium on Nuclear Data and Experimental Techniques. The relevant neutronics quantities (e.g., reaction rates, neutron flux spectra, doses, etc.) were measured using different experimental techniques and the results were compared to the calculated quantities using fusion relevant nuclear data libraries.

This paper focuses on the analyses carried-out by ENEA through the activation foils techniques. $^{197}\text{Au}(n,\gamma)^{198}\text{Au}$, $^{186}\text{W}(n,\gamma)^{187}\text{W}$, $^{115}\text{In}(n,\gamma)^{116}\text{In}$, $^{19}\text{Ni}(n,p)^{19}\text{Co}$, $^{27}\text{Al}(n,\alpha)^{24}\text{Na}$, $^{109}\text{Ag}(n,2n)^{108}\text{Ag}$ activation reactions were used. The foils were placed at eight different positions along the Cu block and irradiated with 14 MeV neutrons. Activation measurements were performed by means of High Purity Germanium (HPGe) detector. Detailed simulation of the experiment was carried-out using MCNP5 Monte Carlo code and the European JEFF-3.1.1 and 3.2 nuclear cross-sections data files for neutron transport. The calculated reaction rates (C) were compared to the experimental quantities (E) and the C/E ratio with relative uncertainties was assessed, compared to the experimental quantities (E) and the C/E ratio with relative uncertainties was assessed.

© 2016 Elsevier B.V. All rights reserved.



Conclusions

...

The present results call for a deep revision/re-evaluation of the copper cross sections. The new release JEFF-3.2 for Cu provided the highest disagreement in the C/E analysis and must be revised. To this end the results of the companion sensitivity/uncertainty post-analysis will help in identifying the main causes of uncertainty in the Cu cross sections. It worth to note that the largest discrepancy among the C/E values was observed for the thermal (capture) reactions suggesting problems and uncertainties in the $^{63,65}\text{Cu}$ capture and elastic cross sections at lower energy rather than at high energy.

Why $^{63,65}\text{Cu}$?

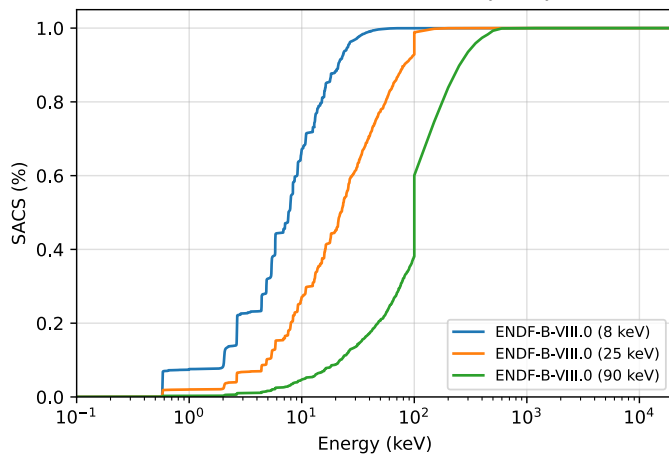
How copper was produced?

Not clear! Candidates:

1. Weak *s* process (Massive stars)
2. Main *s* process (AGB)
3. SNe Ia
4. SNe II

^{62}Ga 116.00 ms β^+	^{63}Ga 32.40 s β^+	^{64}Ga 2.63 m β^+	^{65}Ga 15.20 m β^+	^{66}Ga 9.49 h β^+	^{67}Ga 3.26 d β^+	^{68}Ga 1.13 h β^+
^{61}Zn 1.48 m β^+	^{62}Zn 9.19 h β^+	^{63}Zn 38.47 m β^+	^{64}Zn 48.63 59 mb	^{65}Zn 243.63 d 162 mb, β^+	^{66}Zn 27.9 35 mb	^{67}Zn 4.1 153 mb
^{60}Cu 23.70 m β^+	^{61}Cu 3.33 h β^+	^{62}Cu 9.67 m β^+	^{63}Cu 69.17 94 mb	^{64}Cu 12.70 h β^+	^{65}Cu 30.83 41 mb	^{66}Cu 5.12 m β^-
^{59}Ni 75.99 ka 87 mb, β^+	^{60}Ni 26.223 30 mb	^{61}Ni 1.14 82 mb	^{62}Ni 3.634 22.3 mb	^{63}Ni 100.11 a 31 mb, β^-	^{64}Ni 0.926 8.7 mb	^{65}Ni 2.52 h β^-
^{58}Co 70.86 d β^+	^{59}Co 100 38 mb	^{60}Co 5.27 a β^-	^{61}Co 1.65 h β^-	^{62}Co 1.50 m β^-	^{63}Co 27.40 s β^-	^{64}Co 300.00 ms β^-
^{57}Fe 2.119 40 mb	^{58}Fe 0.282 12.1 mb	^{59}Fe 44.50 d β^-	^{60}Fe 1.50 Ma β^-	^{61}Fe 5.98 m β^-	^{62}Fe 1.13 m β^-	^{63}Fe 6.01 s β^-

The *s* process requires
MACS @ $kT = 8, 25, 90$ keV

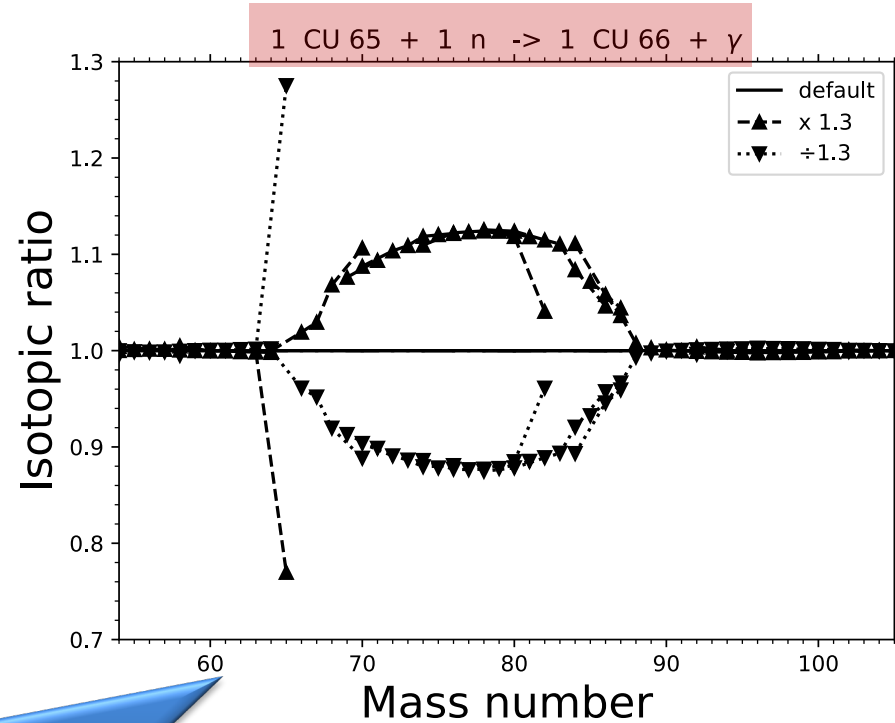
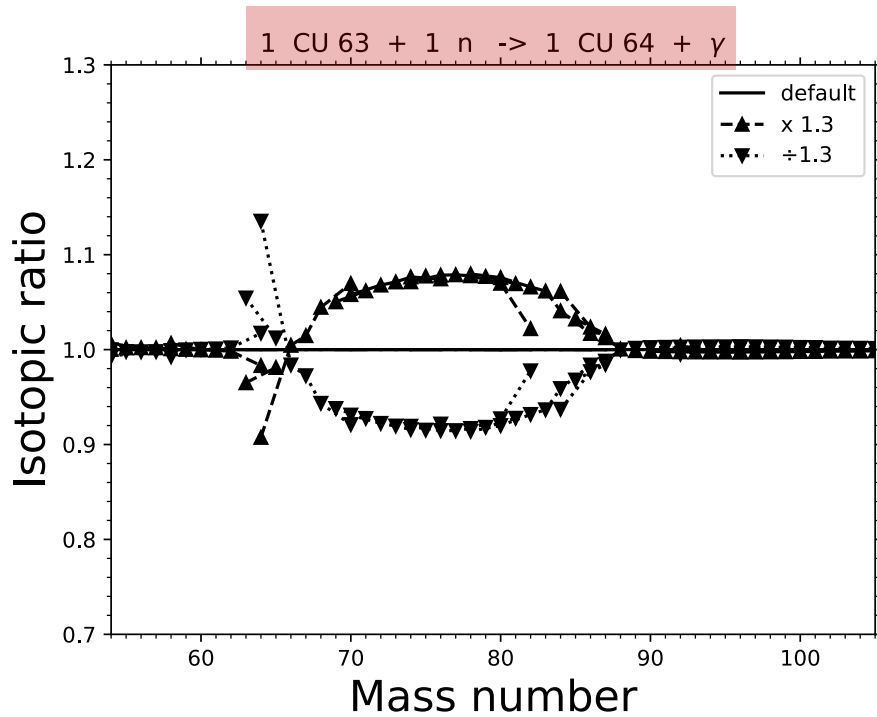


With accurate determination of Cu MACS, it is possible to clarify what is the *s*-process contribution to Cu. Once this is done, it will be possible to constrain the Cu production by other nucleosynthesis processes, where stellar and nuclear uncertainties are much larger.

Why $^{63,65}\text{Cu}$?

Propagation effect in the weak s-process nucleosynthesis

^{62}Ga 116.00 ms β^+	^{63}Ga 32.40 s β^+	^{64}Ga 2.63 m β^+	^{65}Ga 15.20 m β^+	^{66}Ga 9.49 h β^+	^{67}Ga 3.26 d β^+	^{68}Ga 1.13 h β^+
^{61}Zn 1.48 m β^+	M. Pignatari, et al., <i>The s process in massive stars, a benchmark for neutron capture reaction rates</i> , European Physical Journal A 59 (2023) 12, 10.1140/epja/s10050-023-01206-1				^{66}Zn 27.9 35 mb	^{67}Zn 4.1 153 mb
^{60}Cu 23.70 m β^+	^{61}Cu 3.33 h β^+	^{62}Cu 9.67 m β^+	^{63}Cu 69.17 94 mb	^{64}Cu 12.70 h β^+	^{65}Cu 30.83 41 mb	^{66}Cu 5.12 m β^-



Impact of $^{63,65}\text{Cu}(n,\gamma)$ cross sections on the efficiency for the production of elements heavier than Cu

Data in literature

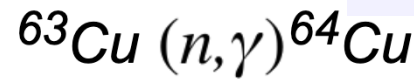


View Maxwellian-Averaged (n,g)
Cross Section

Isotope

(Examples: Ba138, Ta180m, Se.)

▼ Recommended MACS30 (Maxwellian Averaged Cross Section @ 30keV)



Total MACS at 30keV: 60.1 ± 6.2 mb

Cross sections do not include stellar enhancement factors!

▼ History

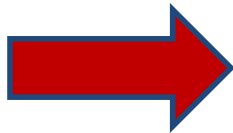
Version	Total MACS [mb]	Partial to gs [mb]	Partial to isomer [mb]
1.0	60.1 ± 6.2	-	-
0.3	55.6 ± 2.2	-	-
0.0	94 ± 10	-	-

(Version 0.0 corresponds to Bao et al.)

▼ Comment

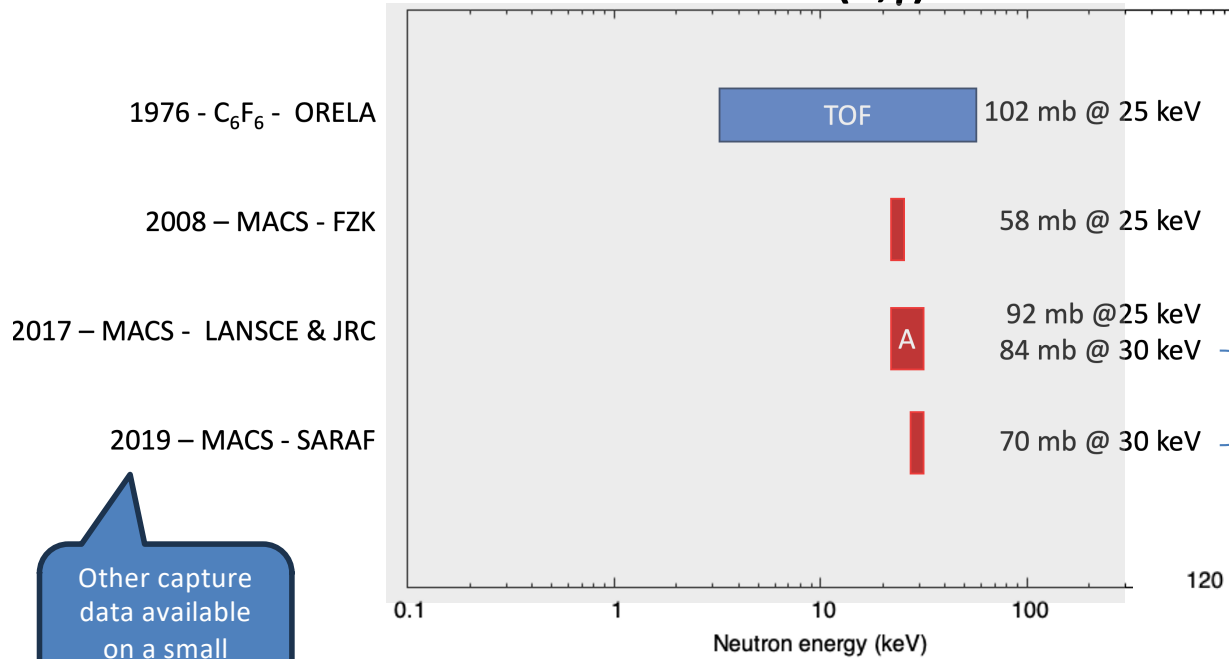
New rec. value is from [HKU08](#), renormalized by 632 mb/586 mb = 1.0785, and recalculated with normalized energy dependencies of [tendl15](#), [endfb71](#), [jendl40](#). Uncertainty is the deviation between different evaluations plus 4% exp. uncertainty from [HKU08](#). **Note the large deviation between the activation measurement and the TOF measurements. More investigation needed!**

Last review: April 2017



Data in literature

$^{63}\text{Cu}(n,\gamma)$

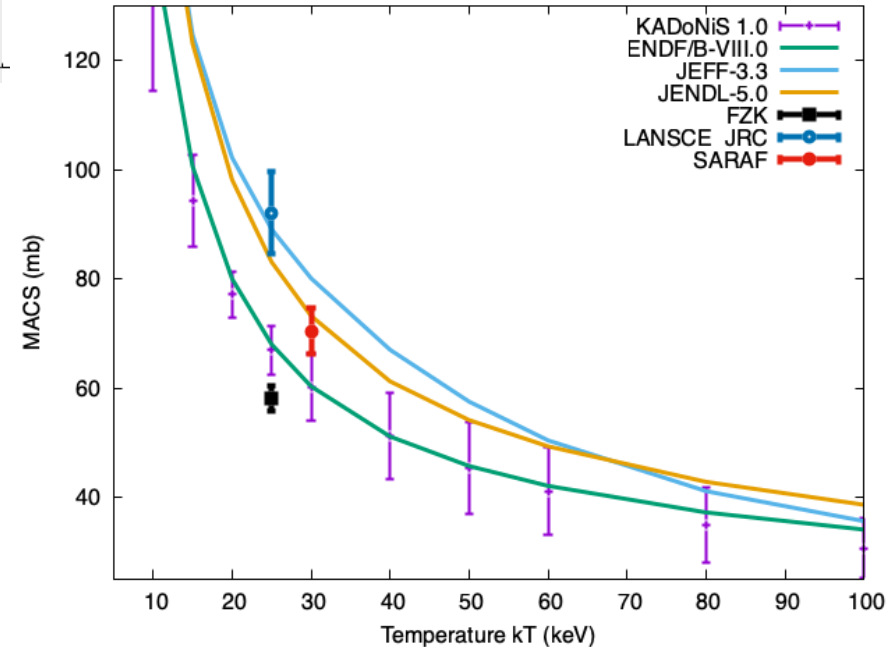


Factor 2 !!!

20%

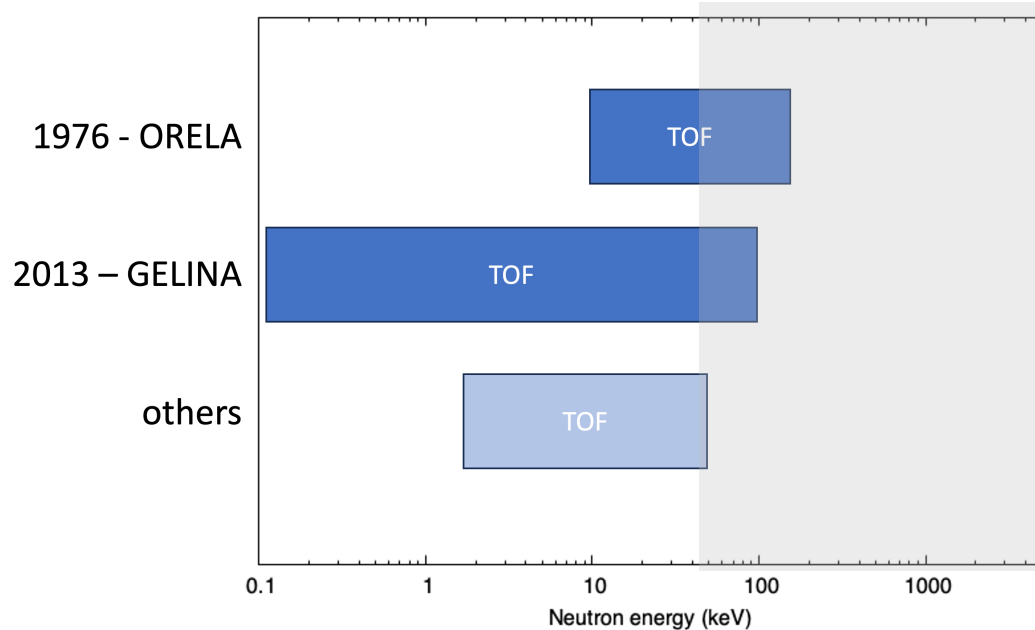
Other capture data available on a small energy range

- Time-of-flight data needed for $E_n < 300$ keV
- n_TOF can accurately cover this energy region



Data in literature

⁶³Cu(n, tot)



- Time-of-flight data needed for $50 \text{ KeV} < E_n < 5 \text{ MeV}$
- n_TOF can accurately cover this energy region with a transmission experiment;
- $\sigma_{\text{tot}} = \sigma_n + \sigma_\gamma + \sigma_{n'}$
- $\sigma_{\text{tot}} \cong \sigma_n$ for $E_n < 2 \text{ MeV}$

[Link1](#)
[Link2](#)
[Link3](#)

Nuclear Physics A396 (1982) 509-540
© North-Holland Publishing Company

ELASTIC AND INELASTIC SCATTERING OF NEUTRONS FROM ^{54,56}Fe AND ^{63,65}Cu
(I). Measurements from 8 to 14 MeV and a spherical optical model analysis

S. M. EL-KADI, C. E. NELSON, F. O. PURSER and R. L. WALTER
Department of Physics, Duke University, Durham, NC 27706, USA
and
Triangle Universities Nuclear Laboratory, Duke Station, NC 27706, USA
and
A. BEYERLE, C. R. GOULD and L. W. SEAGONDOLLAR
Department of Physics, North Carolina State University, Raleigh, NC 27650
and
Triangle Universities Nuclear Laboratory, Duke Station, NC 27706, USA
Received 31 March 1982

Abstract: Differential cross sections were measured at 8, 10, 12 and 14 MeV for elastic scattering of neutrons from enriched samples of ⁵⁶Fe, ⁵⁴Fe, ⁶³Cu and ⁶⁵Cu. Inelastic scattering to the first excited state in ^{54,56}Fe was also observed. For the ^{63,65}Cu isotopes, the inelastic cross sections for scattering to the combined group of the three (five) states were determined at 8 and 10 MeV (12 and 14 MeV). The elastic scattering data are compared to predictions of earlier global optical models. New spherical optical model representations were obtained. These data were combined with data for nickel, tin and lead to generate a new global parametrisation. Comparisons of derived volume integrals for the potentials, total cross sections and potential radii are made to available information.

⁶³Cu(n, n)

⁶³Cu(n, n')

Angular distribution

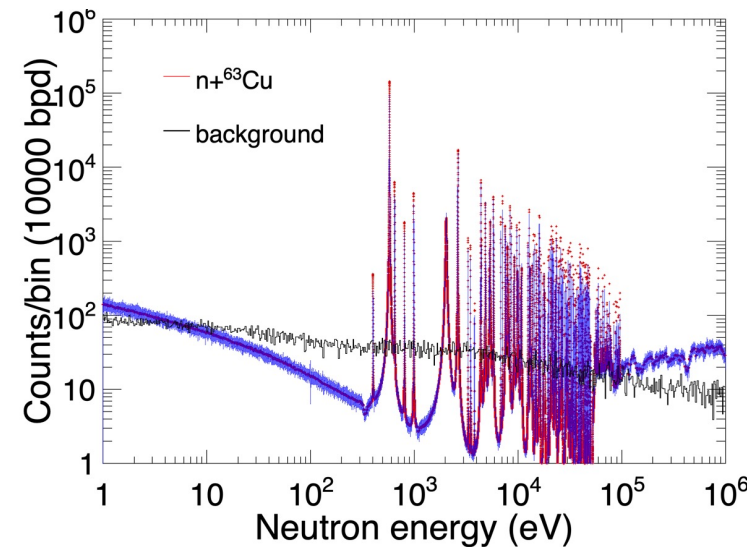
- Currently, detector R&D, to study these channels at n_TOF
- Possible synergy with JRC-Geel

Addendum in the future?

Proposed measurements

Consist of **capture** and **transmission** experiments at **EAR1** using highly enriched samples (**99.8%** in ^{63}Cu and **99.0%** in ^{65}Cu)

- Capture at EAR1 using 4 C_6D_6

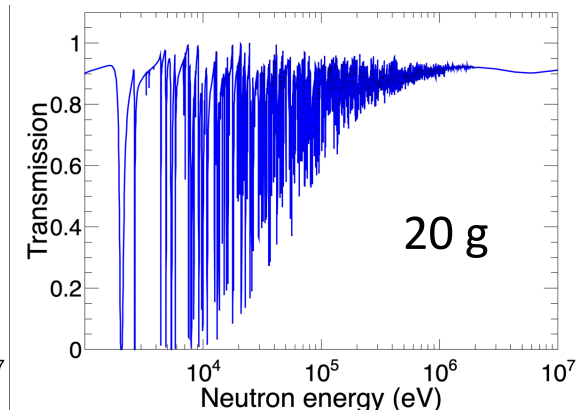
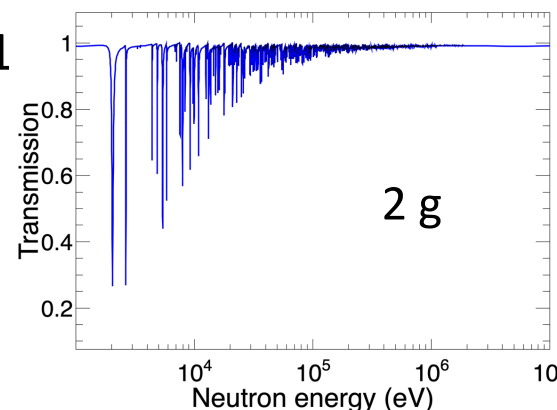


Example:

2 grams of ^{63}Cu and cylindrical sample of 1.5 cm radius (2.7×10^{-3} at/b)

2×10^{18} protons

- Transmission at EAR1 using ^{235}U fission chamber from PTB



Proposed measurements

Consist of **capture** and **transmission** experiments at **EAR1** using highly enriched samples (**99.8%** in ^{63}Cu and **99.0%** in ^{65}Cu)

Experiment	Sample	Protons	Comments
Capture	^{63}Cu	2.0×10^{18}	
Capture	^{65}Cu	2.0×10^{18}	
Capture	^{nat}Cu	0.3×10^{18}	EAR1 or EAR2
Capture	Empty-sample	0.2×10^{18}	background study
Capture	Pb	0.2×10^{18}	background study
Capture	C	0.2×10^{18}	background study
Capture	^{197}Au	0.1×10^{18}	normalization
Transmission	^{63}Cu	1.0×10^{18}	"Sample-in"
Transmission	^{65}Cu	1.0×10^{18}	"Sample-in"
Transmission	Empty-sample	1.0×10^{18}	"Sample-out"
		8.0×10^{18}	

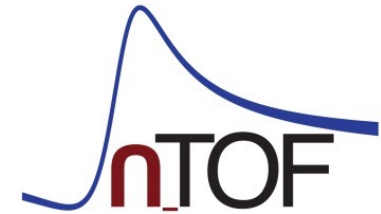
Target:

- Observe resonances up to 50-200 keV
- σ_γ uncertainty below 3-5%

Target:

- σ_{tot} uncertainty below 5% for $E_\gamma > 100$ keV with 100 bpd

Conclusion



- $n+^{63,65}\text{Cu}$ is an intriguing physics case! (new entry in HPRL?)
- Capture and transmission @ EAR1
 - 5×10^{18} Protons for $^{63}\text{Cu}(n,\gamma)$ and $^{65}\text{Cu}(n,\gamma)$
 - 3×10^{18} Protons for $^{63}\text{Cu}(n,\text{tot})$ and $^{65}\text{Cu}(n,\text{tot})$
- Proton request split into 2 runs over 2 years
 - e.g., (n,γ) in 2024 and (n,tot) in 2025
- Other reactions in future dedicated addendum (if feasible)

Total: 8×10^{18} protons on target

Acknowledgement:



CERN-INTC-2024-006 / INTC-P-689
10/01/2024



- ERANOS Sensitivity studies by **Donato Maurizio Castelluccio**
- MCNP calculations by **Patrizio Console Camprini**
- Massive stars sensitivity study by **Marco Pignatari**



ALMA MATER STUDIORUM
UNIVERSITÀ DI BOLOGNA

Cristian Massimi

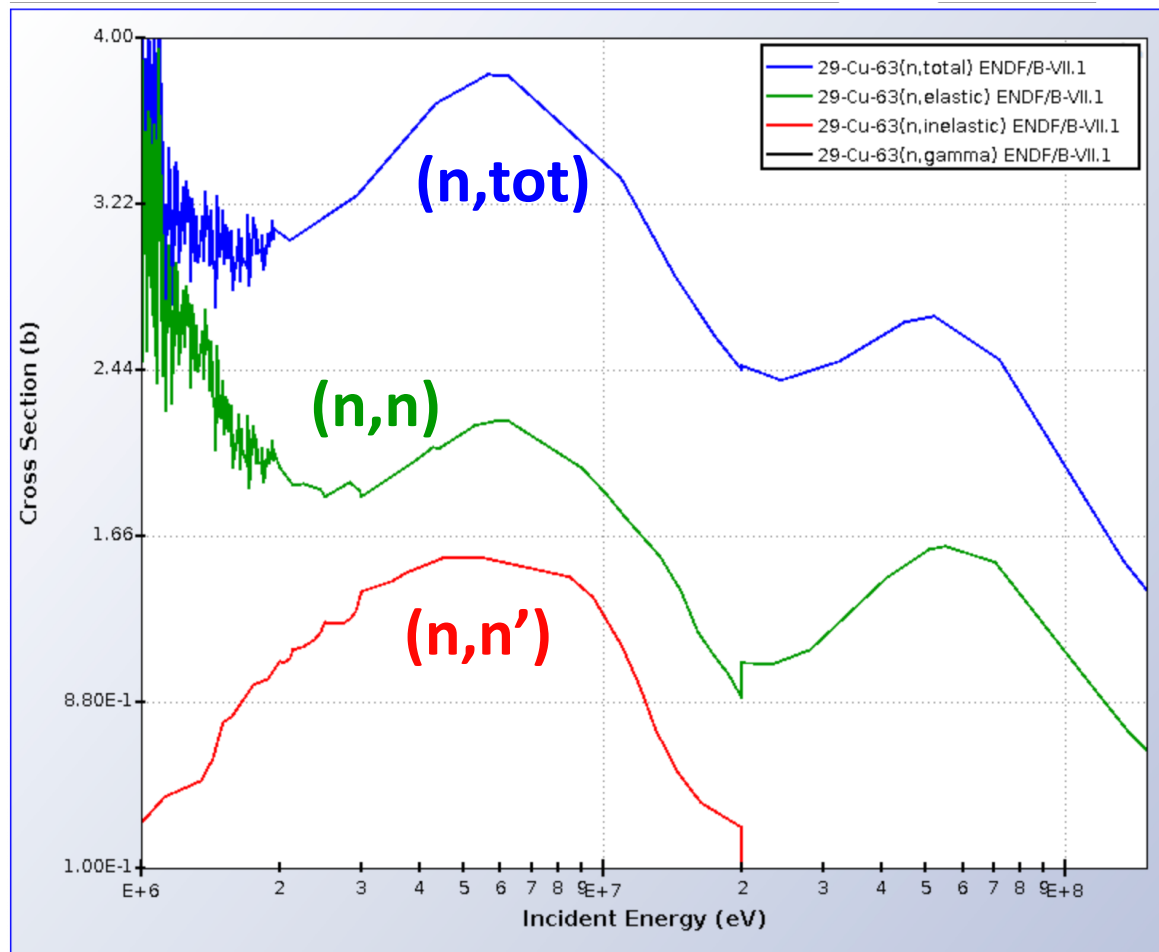
Department of Physics and Astronomy

cristian.massimi@unibo.it

www.unibo.it

Why $^{63,65}\text{Cu}$?

Natural abundance:
 ^{63}Cu 69.2 %



$$\sigma_n \gg \sigma_\gamma$$

$$\sigma_{el} > \sigma_{inel}$$

Prodotti della reazione	Q -valore (keV)
$^{63}\text{Cu} + n$	0
$^{64}\text{Cu} + \gamma$	7915.9 ± 0.6
$^{60}\text{Co} + \alpha$	1717.0 ± 0.6
$^{63}\text{Ni} + p$	715.4 ± 0.6

Why $^{63,65}\text{Cu}$?



Nucl. Techn.

TAPIRO: ERANOS \rightarrow S/U analysis

Sensitivity and Uncertainty (S/U) analysis was performed using the deterministic code ERANOS 2.3, for:

1. k_{eff} parameter $\left\{ \begin{array}{l} (n,n) \\ (n,n') \\ (n,\gamma) \end{array} \right.$
2. ratio of fission reaction rates U238/U235 $\left\{ \begin{array}{l} (n,n') \\ (n,n) \\ (n,\gamma) \end{array} \right.$
3. ratio of fission r.r. Np237/U235 $\left\{ \begin{array}{l} (n,n) \\ (n,n') \\ (n,\gamma) \end{array} \right.$

Relevant energy region

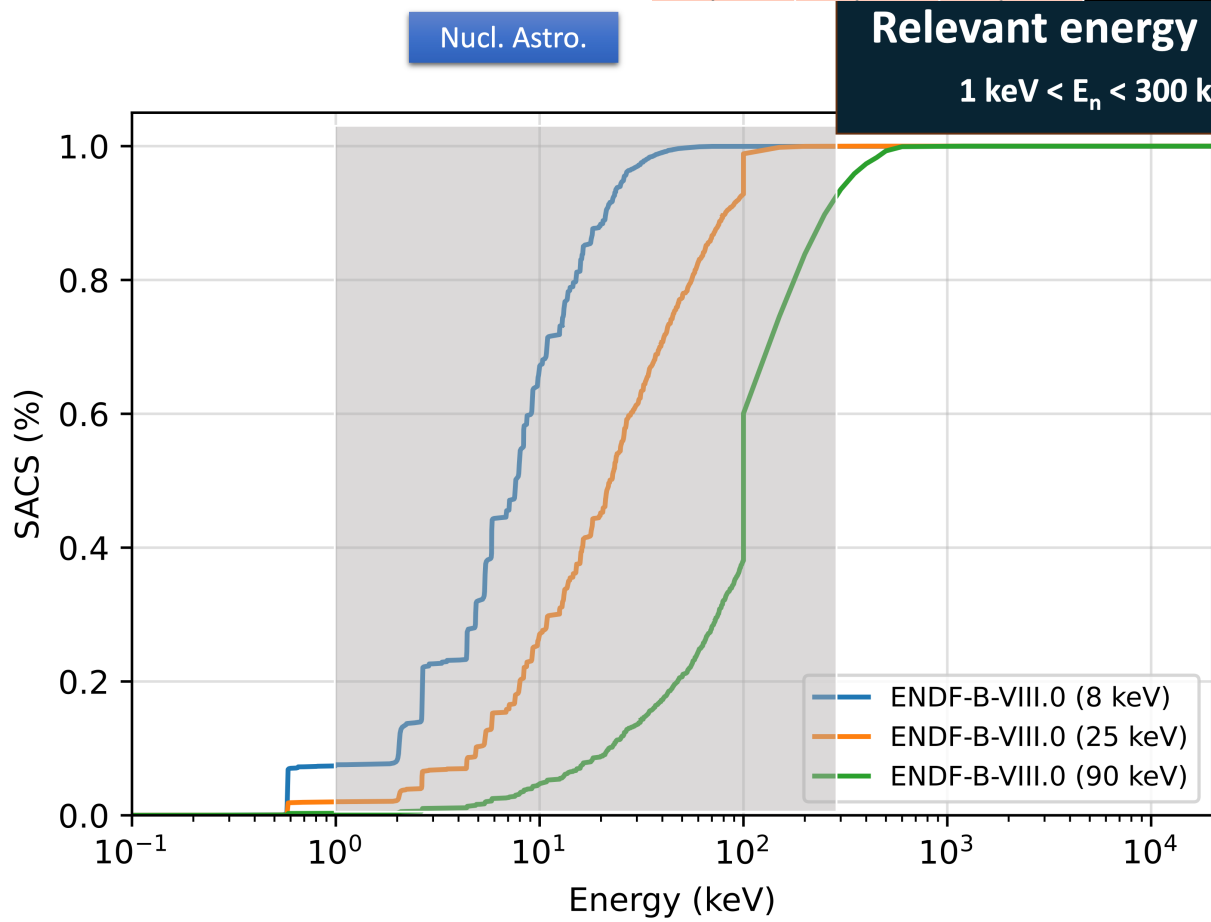
- Elastic: $50 \text{ keV} < E_n < 5 \text{ MeV}$
- Inelastic: $1 \text{ MeV} < E_n < 5 \text{ MeV}$
- Rad. capture: $10 \text{ keV} < E_n < 300 \text{ keV}$

Why $^{63,65}\text{Cu}$?

^{62}Ga 116.00 ms β^+	^{63}Ga 32.40 s β^+	^{64}Ga 2.63 m β^+	^{65}Ga 15.20 m β^+	^{66}Ga 9.49 h β^+	^{67}Ga 3.26 d β^+	^{68}Ga 1.13 h β^+	
^{61}Zn 1.48 m β^+	^{62}Zn 9.19 h β^+	^{63}Zn 38.47 m β^+	^{64}Zn 48.63 59 mb	^{65}Zn 243.63 d 162 mb, β^+	^{66}Zn 27.9 35 mb	^{67}Zn 4.1 153 mb	
^{60}Cu 23.70 m β^+	^{61}Cu 3.33 h β^+	^{62}Cu 9.67 m β^+	^{63}Cu 69.17 94 mb	^{64}Cu 12.70 h β^+	^{65}Cu 30.83 41 mb	^{66}Cu 5.12 m β^-	
						^{64}Ni 0.926 8.7 mb	^{65}Ni 2.52 h β^-
						^{63}Co 27.40 s β^-	^{64}Co 300.00 ms β^-
						^{62}Fe 1.13 m β^-	^{63}Fe 6.01 s β^-

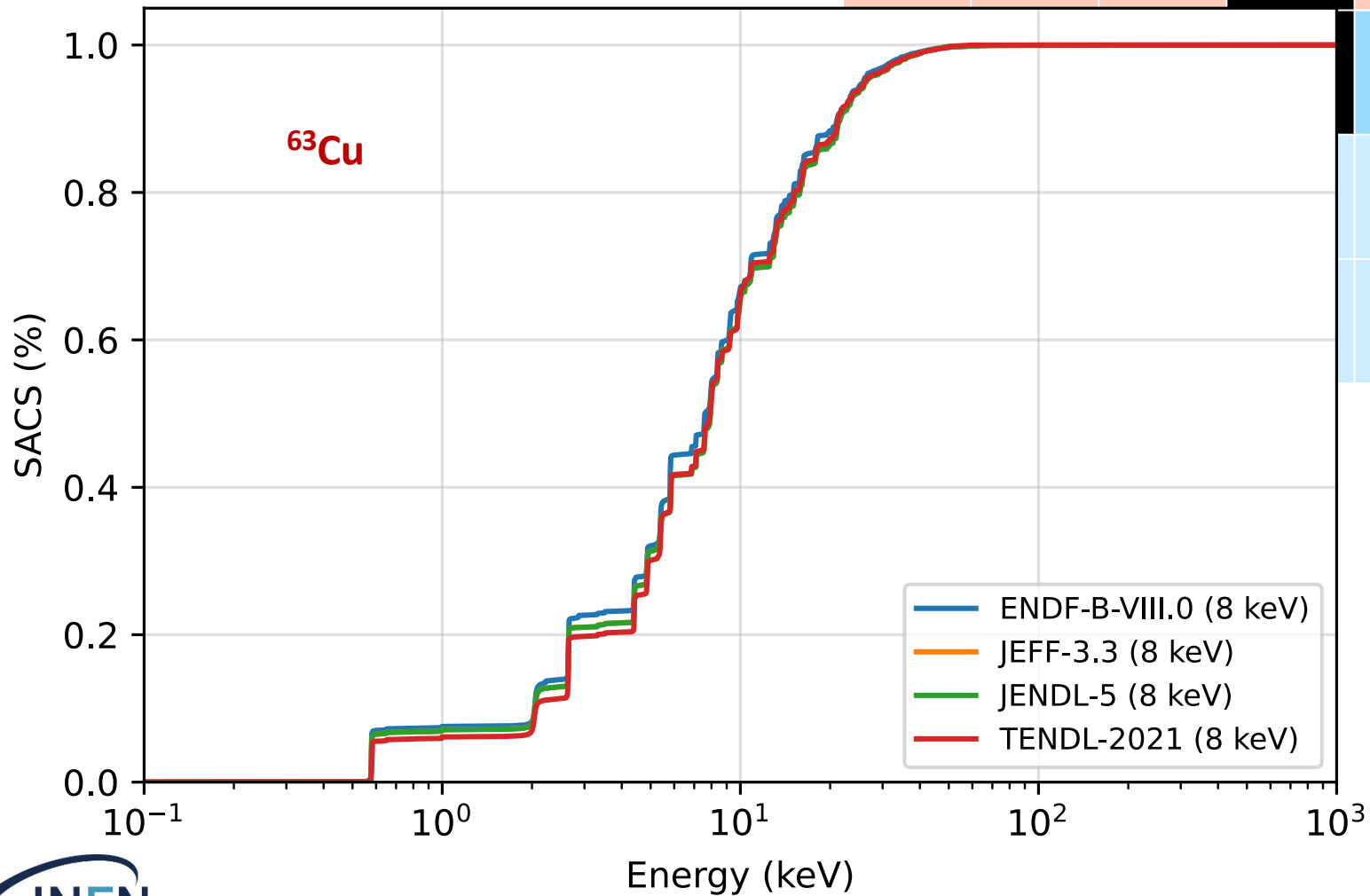
$^{63}\text{Cu}(n,\gamma)$

% MACS



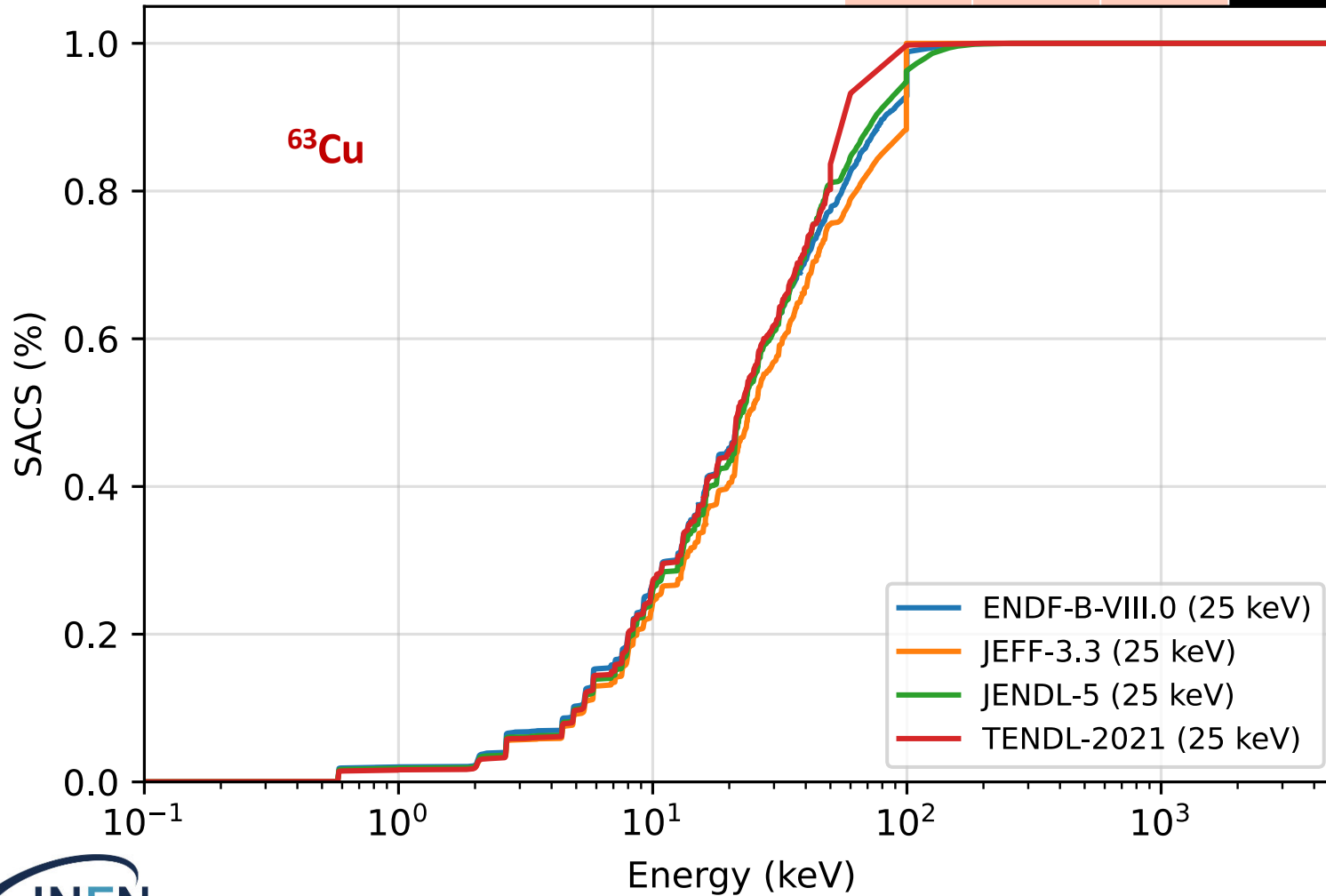
Why $^{63,65}\text{Cu}$?

^{62}Ga 116.00 ms β^+	^{63}Ga 32.40 s β^+	^{64}Ga 2.63 m β^+	^{65}Ga 15.20 m β^+	^{66}Ga 9.49 h β^+	^{67}Ga 3.26 d β^+	^{68}Ga 1.13 h β^+
^{61}Zn 1.48 m β^+	^{62}Zn 9.19 h β^+	^{63}Zn 38.47 m β^+	^{64}Zn 48.63 59 mb	^{65}Zn 243.63 d 162 mb, β^+	^{66}Zn 27.9 35 mb	^{67}Zn 4.1 153 mb
^{60}Cu 23.70 m β^+	^{61}Cu 3.33 h β^+	^{62}Cu 9.67 m β^+	^{63}Cu 69.17 94 mb	^{64}Cu 12.70 h β^+	^{65}Cu 30.83 41 mb	^{66}Cu 5.12 m β^-
				^{63}Ni 100.11 a 31 mb, β^-	^{64}Ni 0.926 8.7 mb	^{65}Ni 2.52 h β^-
				^{62}Co 1.50 m β^-	^{63}Co 27.40 s β^-	^{64}Co 300.00 ms β^-
				^{61}Fe 5.98 m β^-	^{62}Fe 1.13 m β^-	^{63}Fe 6.01 s β^-



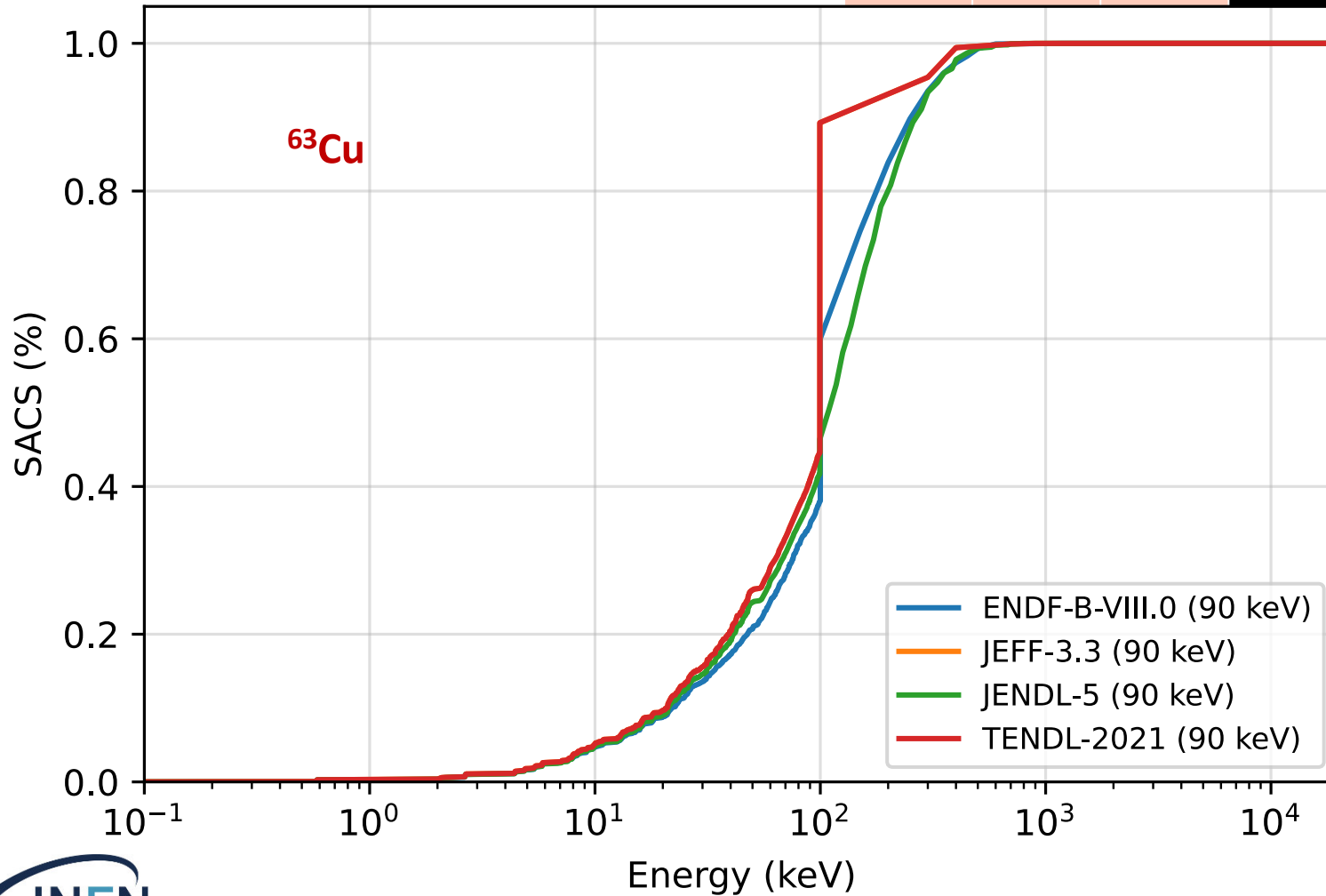
Why $^{63,65}\text{Cu}$?

^{62}Ga 116.00 ms β^+	^{63}Ga 32.40 s β^+	^{64}Ga 2.63 m β^+	^{65}Ga 15.20 m β^+	^{66}Ga 9.49 h β^+	^{67}Ga 3.26 d β^+	^{68}Ga 1.13 h β^+
^{61}Zn 1.48 m β^+	^{62}Zn 9.19 h β^+	^{63}Zn 38.47 m β^+	^{64}Zn 48.63 59 mb	^{65}Zn 243.63 d 162 mb, β^+	^{66}Zn 27.9 35 mb	^{67}Zn 4.1 153 mb
^{60}Cu 23.70 m β^+	^{61}Cu 3.33 h β^+	^{62}Cu 9.67 m β^+	^{63}Cu 69.17 94 mb	^{64}Cu 12.70 h β^+	^{65}Cu 30.83 41 mb	^{66}Cu 5.12 m β^-
				^{63}Ni 100.11 a 31 mb, β^-	^{64}Ni 0.926 8.7 mb	^{65}Ni 2.52 h β^-
				^{62}Co 1.50 m β^-	^{63}Co 27.40 s β^-	^{64}Co 300.00 ms β^-
				^{61}Fe 5.98 m β^-	^{62}Fe 1.13 m β^-	^{63}Fe 6.01 s β^-



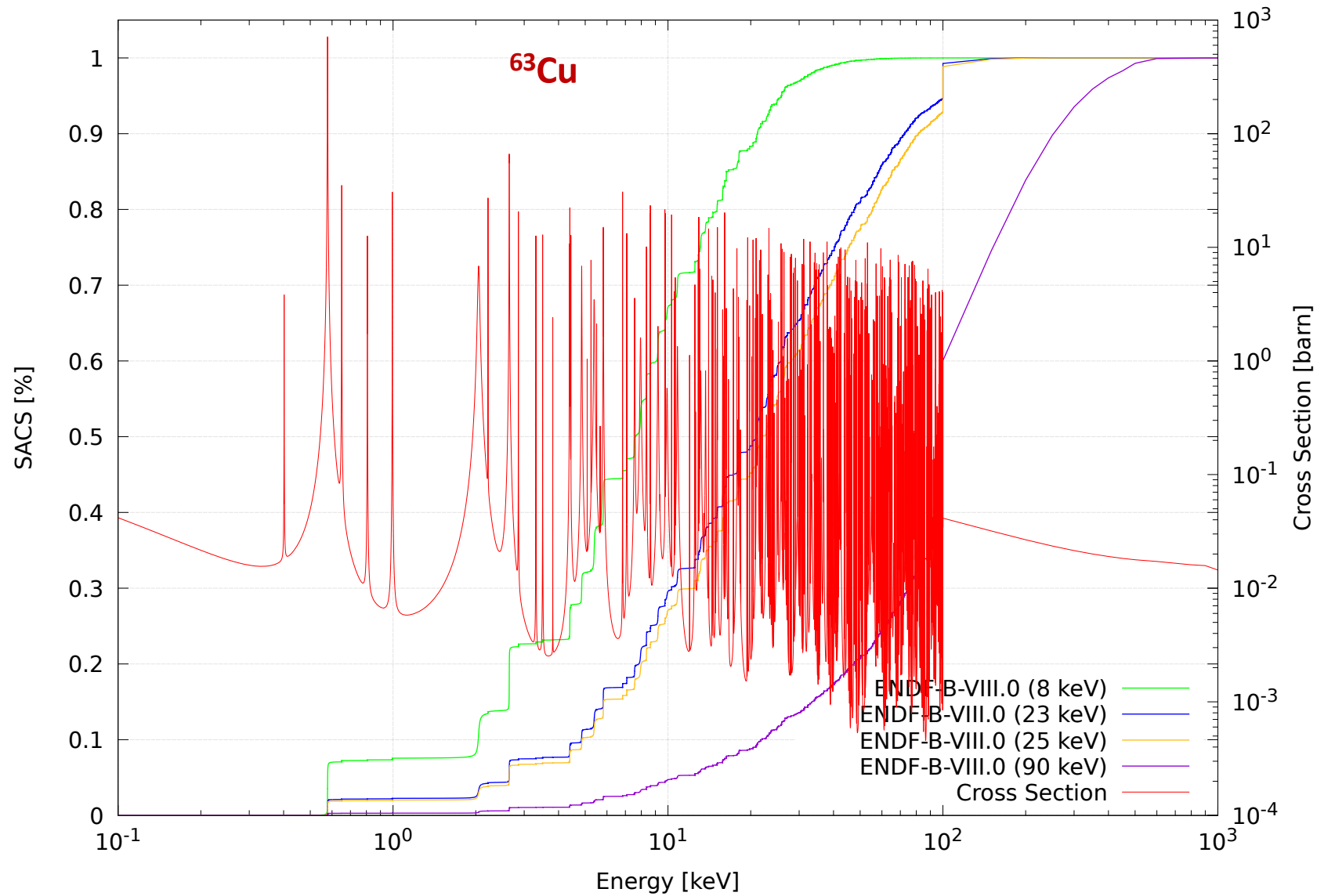
Why $^{63,65}\text{Cu}$?

^{62}Ga 116.00 ms β^+	^{63}Ga 32.40 s β^+	^{64}Ga 2.63 m β^+	^{65}Ga 15.20 m β^+	^{66}Ga 9.49 h β^+	^{67}Ga 3.26 d β^+	^{68}Ga 1.13 h β^+
^{61}Zn 1.48 m β^+	^{62}Zn 9.19 h β^+	^{63}Zn 38.47 m β^+	^{64}Zn 48.63 59 mb	^{65}Zn 243.63 d 162 mb, β^+	^{66}Zn 27.9 35 mb	^{67}Zn 4.1 153 mb
^{60}Cu 23.70 m β^+	^{61}Cu 3.33 h β^+	^{62}Cu 9.67 m β^+	^{63}Cu 69.17 94 mb	^{64}Cu 12.70 h β^+	^{65}Cu 30.83 41 mb	^{66}Cu 5.12 m β^-

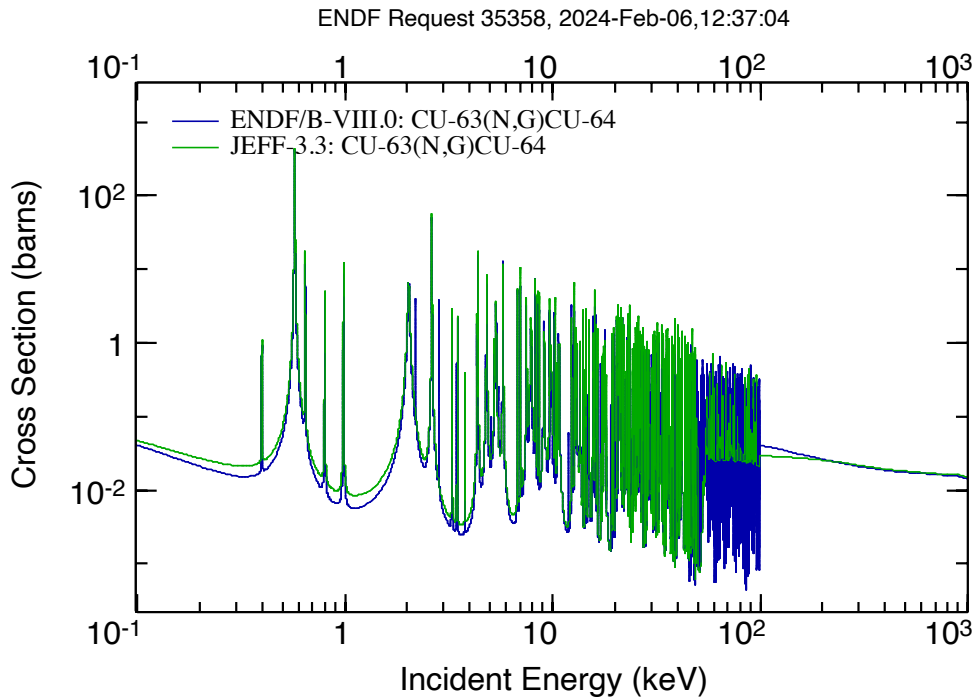


^{63}Ni 100.11 a 31 mb, β^-	^{64}Ni 0.926 8.7 mb	^{65}Ni 2.52 h β^-
^{62}Co 1.50 m β^-	^{63}Co 27.40 s β^-	^{64}Co 300.00 ms β^-
^{61}Fe 5.98 m β^-	^{62}Fe 1.13 m β^-	^{63}Fe 6.01 s β^-

Why $^{63,65}\text{Cu}$?

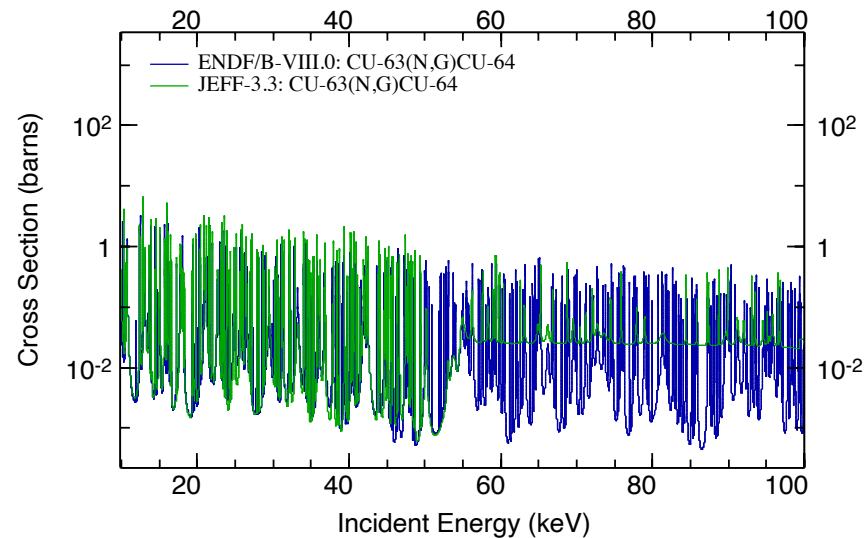


Why $^{63,65}\text{Cu}$?



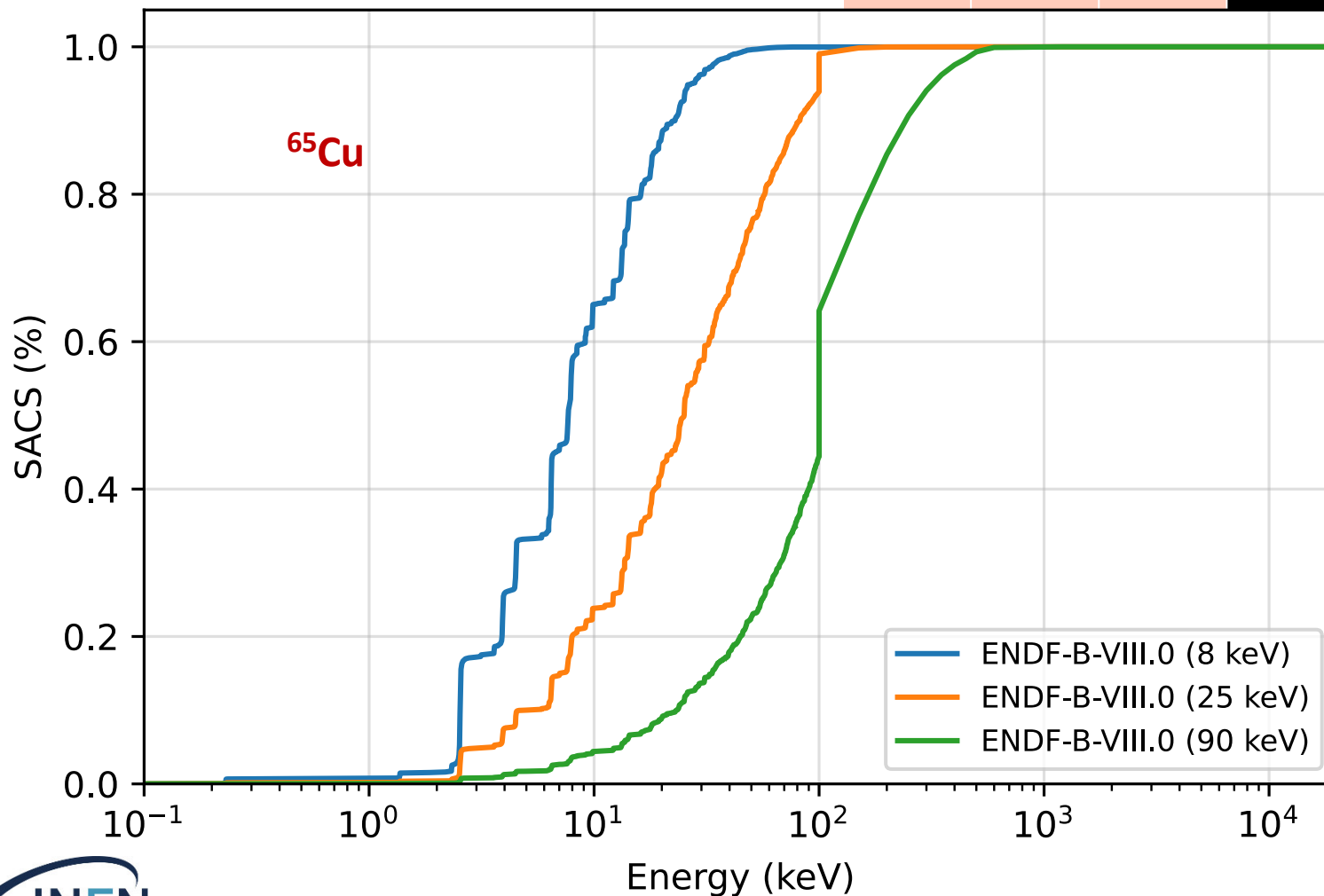
^{62}Ga 116.00 ms β^+	^{63}Ga 32.40 s β^+	^{64}Ga 2.63 m β^+	^{65}Ga 15.20 m β^+	^{66}Ga 9.49 h β^+	^{67}Ga 3.26 d β^+	^{68}Ga 1.13 h β^+
^{61}Zn 1.48 m β^+	^{62}Zn 9.19 h β^+	^{63}Zn 38.47 m β^+	^{64}Zn 48.63 59 mb	^{65}Zn 243.63 d 162 mb, β^+	^{66}Zn 27.9 35 mb	^{67}Zn 4.1 153 mb
^{60}Cu 23.70 m β^+	^{61}Cu 3.33 h β^+	^{62}Cu 9.67 m β^+	^{63}Cu 69.17 94 mb	^{64}Cu 12.70 h β^+	^{65}Cu 30.83 41 mb	^{66}Cu 5.12 m β^-
^{59}Ni 75.99 ka 87 mb, β^+	^{60}Ni 26.223 30 mb	^{61}Ni 1.14 82 mb	^{62}Ni 3.634 22.3 mb	^{63}Ni 100.11 a 31 mb, β^-	^{64}Ni 0.926 8.7 mb	^{65}Ni 2.52 h β^-
^{58}Co 70.86 d β^+	^{59}Co 100 38 mb	^{60}Co 5.27 a β^-	^{61}Co 1.65 h β^-	^{62}Co 1.50 m β^-	^{63}Co 27.40 s β^-	^{64}Co 300.00 ms β^-
^{57}Fe 2.119 40 mb	^{58}Fe 0.282 12.1 mb	^{59}Fe 44.50 d β^-	^{60}Fe 1.50 Ma β^-	^{61}Fe 5.98 m β^-	^{62}Fe 1.13 m β^-	^{63}Fe 6.01 s β^-

ENDF Request 35358, 2024-Feb-06,12:31:43



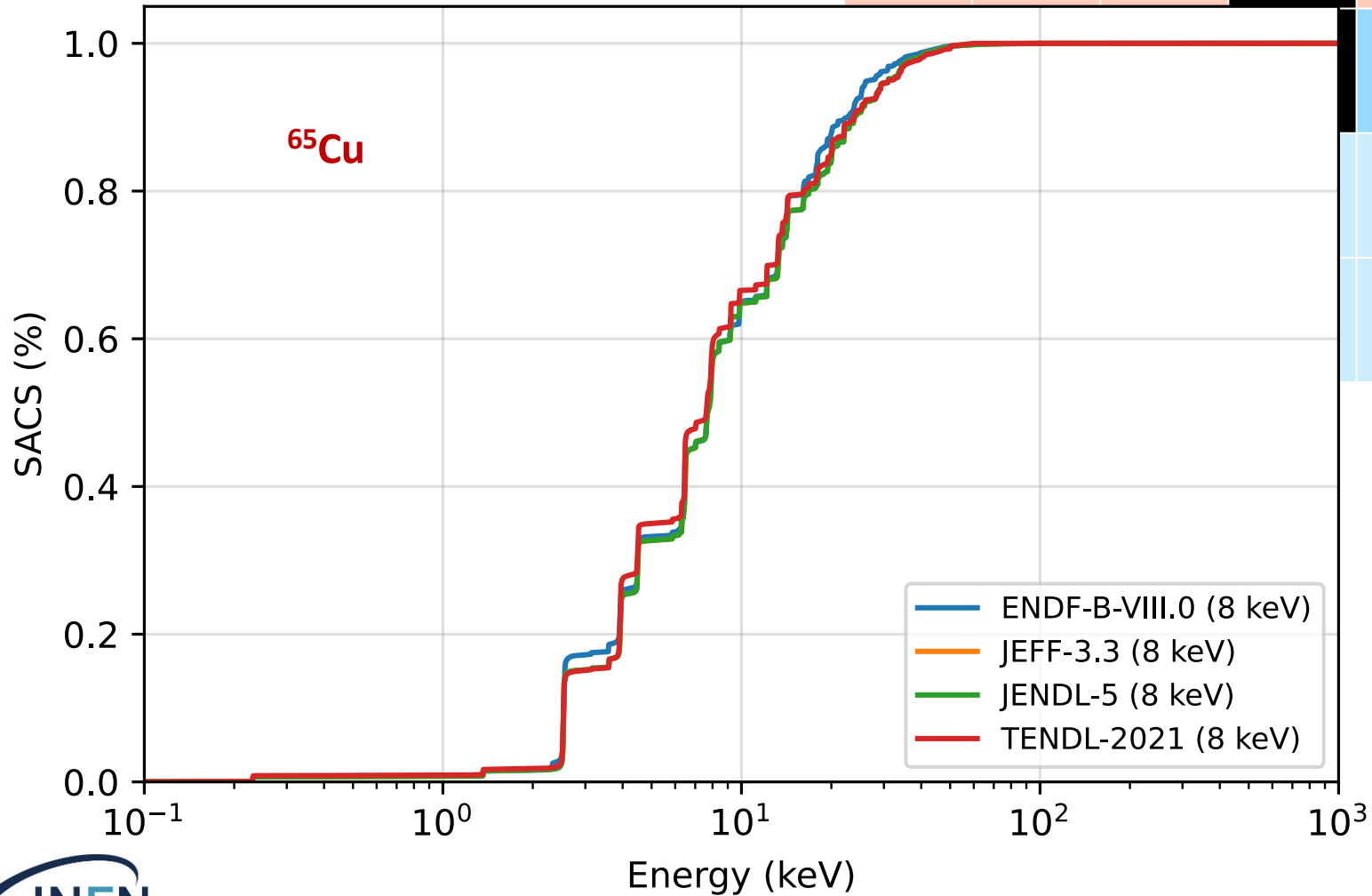
Why $^{63,65}\text{Cu}$?

^{62}Ga 116.00 ms β^+	^{63}Ga 32.40 s β^+	^{64}Ga 2.63 m β^+	^{65}Ga 15.20 m β^+	^{66}Ga 9.49 h β^+	^{67}Ga 3.26 d β^+	^{68}Ga 1.13 h β^+
^{61}Zn 1.48 m β^+	^{62}Zn 9.19 h β^+	^{63}Zn 38.47 m β^+	^{64}Zn 48.63 59 mb	^{65}Zn 243.63 d 162 mb, β^+	^{66}Zn 27.9 35 mb	^{67}Zn 4.1 153 mb
^{60}Cu 23.70 m β^+	^{61}Cu 3.33 h β^+	^{62}Cu 9.67 m β^+	^{63}Cu 69.17 94 mb	^{64}Cu 12.70 h β^+	^{65}Cu 30.83 41 mb	^{66}Cu 5.12 m β^-
				^{63}Ni 100.11 a 31 mb, β^-	^{64}Ni 0.926 8.7 mb	^{65}Ni 2.52 h β^-
				^{62}Co 1.50 m β^-	^{63}Co 27.40 s β^-	^{64}Co 300.00 ms β^-
				^{61}Fe 5.98 m β^-	^{62}Fe 1.13 m β^-	^{63}Fe 6.01 s β^-



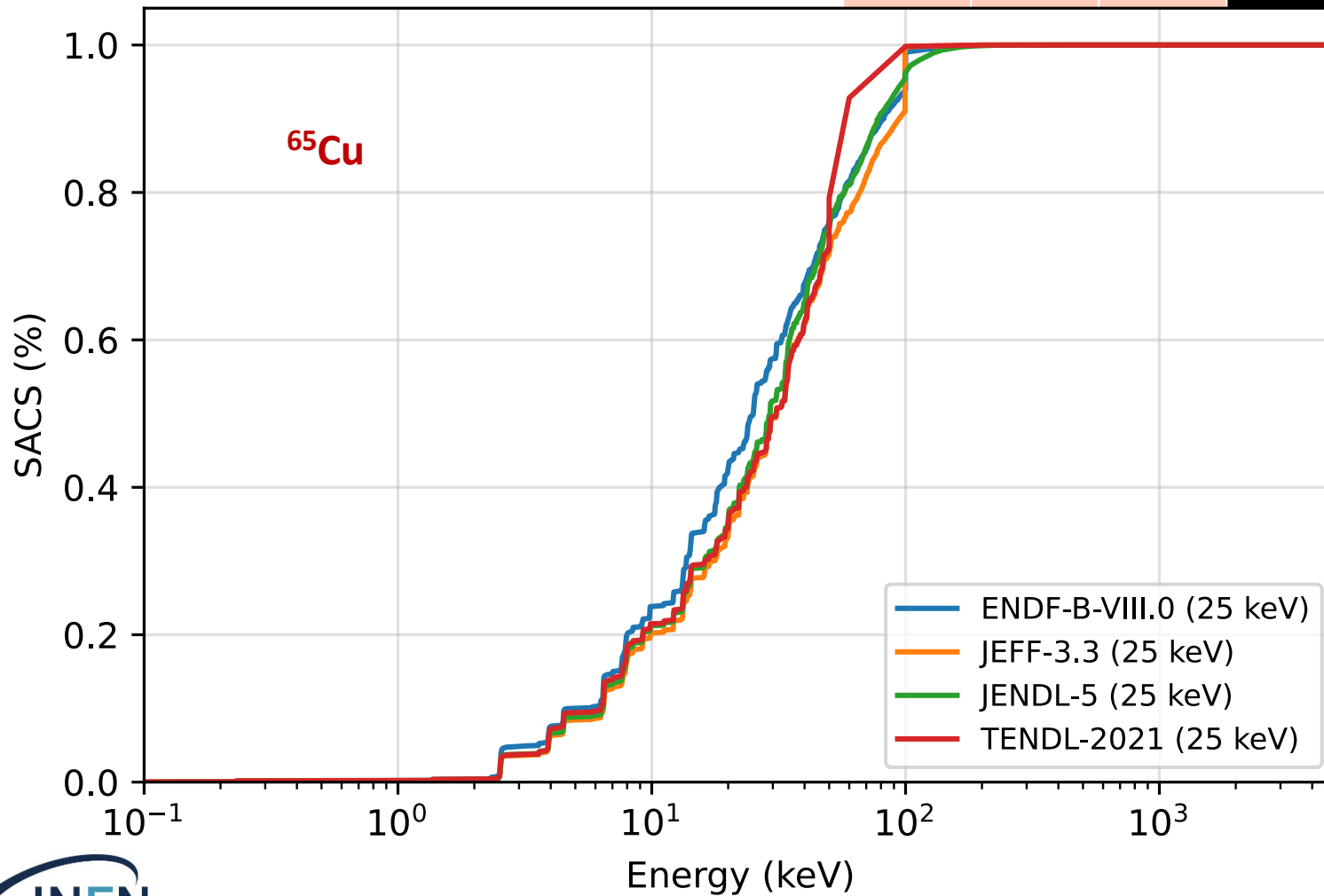
Why $^{63,65}\text{Cu}$?

^{62}Ga 116.00 ms β^+	^{63}Ga 32.40 s β^+	^{64}Ga 2.63 m β^+	^{65}Ga 15.20 m β^+	^{66}Ga 9.49 h β^+	^{67}Ga 3.26 d β^+	^{68}Ga 1.13 h β^+
^{61}Zn 1.48 m β^+	^{62}Zn 9.19 h β^+	^{63}Zn 38.47 m β^+	^{64}Zn 48.63 59 mb	^{65}Zn 243.63 d 162 mb, β^+	^{66}Zn 27.9 35 mb	^{67}Zn 4.1 153 mb
^{60}Cu 23.70 m β^+	^{61}Cu 3.33 h β^+	^{62}Cu 9.67 m β^+	^{63}Cu 69.17 94 mb	^{64}Cu 12.70 h β^+	^{65}Cu 30.83 41 mb	^{66}Cu 5.12 m β^-
	^{63}Ni 100.11 a 31 mb, β^-	^{64}Ni 0.926 8.7 mb	^{65}Ni 2.52 h β^-			
	^{62}Co 1.50 m β^-	^{63}Co 27.40 s β^-	^{64}Co 300.00 ms β^-			
	^{61}Fe 5.98 m β^-	^{62}Fe 1.13 m β^-	^{63}Fe 6.01 s β^-			



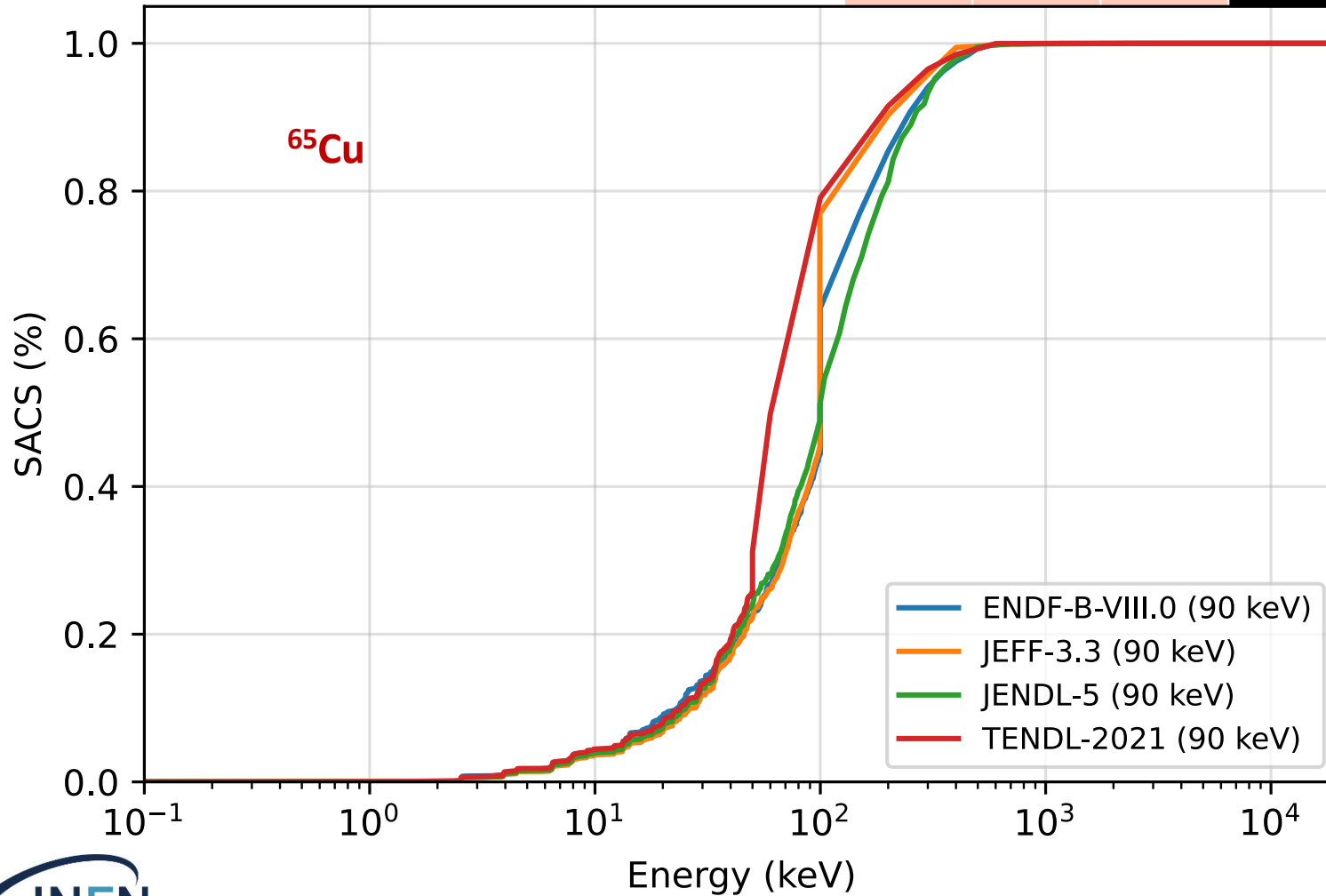
Why $^{63,65}\text{Cu}$?

^{62}Ga 116.00 ms β^+	^{63}Ga 32.40 s β^+	^{64}Ga 2.63 m β^+	^{65}Ga 15.20 m β^+	^{66}Ga 9.49 h β^+	^{67}Ga 3.26 d β^+	^{68}Ga 1.13 h β^+
^{61}Zn 1.48 m β^+	^{62}Zn 9.19 h β^+	^{63}Zn 38.47 m β^+	^{64}Zn 48.63 59 mb	^{65}Zn 243.63 d 162 mb, β^+	^{66}Zn 27.9 35 mb	^{67}Zn 4.1 153 mb
^{60}Cu 23.70 m β^+	^{61}Cu 3.33 h β^+	^{62}Cu 9.67 m β^+	^{63}Cu 69.17 94 mb	^{64}Cu 12.70 h β^+	^{65}Cu 30.83 41 mb	^{66}Cu 5.12 m β^-
				^{63}Ni 100.11 a 31 mb, β^-	^{64}Ni 0.926 8.7 mb	^{65}Ni 2.52 h β^-
				^{62}Co 1.50 m β^-	^{63}Co 27.40 s β^-	^{64}Co 300.00 ms β^-
				^{61}Fe 5.98 m β^-	^{62}Fe 1.13 m β^-	^{63}Fe 6.01 s β^-



Why $^{63,65}\text{Cu}$?

^{62}Ga 116.00 ms β^+	^{63}Ga 32.40 s β^+	^{64}Ga 2.63 m β^+	^{65}Ga 15.20 m β^+	^{66}Ga 9.49 h β^+	^{67}Ga 3.26 d β^+	^{68}Ga 1.13 h β^+
^{61}Zn 1.48 m β^+	^{62}Zn 9.19 h β^+	^{63}Zn 38.47 m β^+	^{64}Zn 48.63 59 mb	^{65}Zn 243.63 d 162 mb, β^+	^{66}Zn 27.9 35 mb	^{67}Zn 4.1 153 mb
^{60}Cu 23.70 m β^+	^{61}Cu 3.33 h β^+	^{62}Cu 9.67 m β^+	^{63}Cu 69.17 94 mb	^{64}Cu 12.70 h β^+	^{65}Cu 30.83 41 mb	^{66}Cu 5.12 m β^-



^{63}Ni 100.11 a 31 mb, β^-	^{64}Ni 0.926 8.7 mb	^{65}Ni 2.52 h β^-
^{62}Co 1.50 m β^-	^{63}Co 27.40 s β^-	^{64}Co 300.00 ms β^-
^{61}Fe 5.98 m β^-	^{62}Fe 1.13 m β^-	^{63}Fe 6.01 s β^-

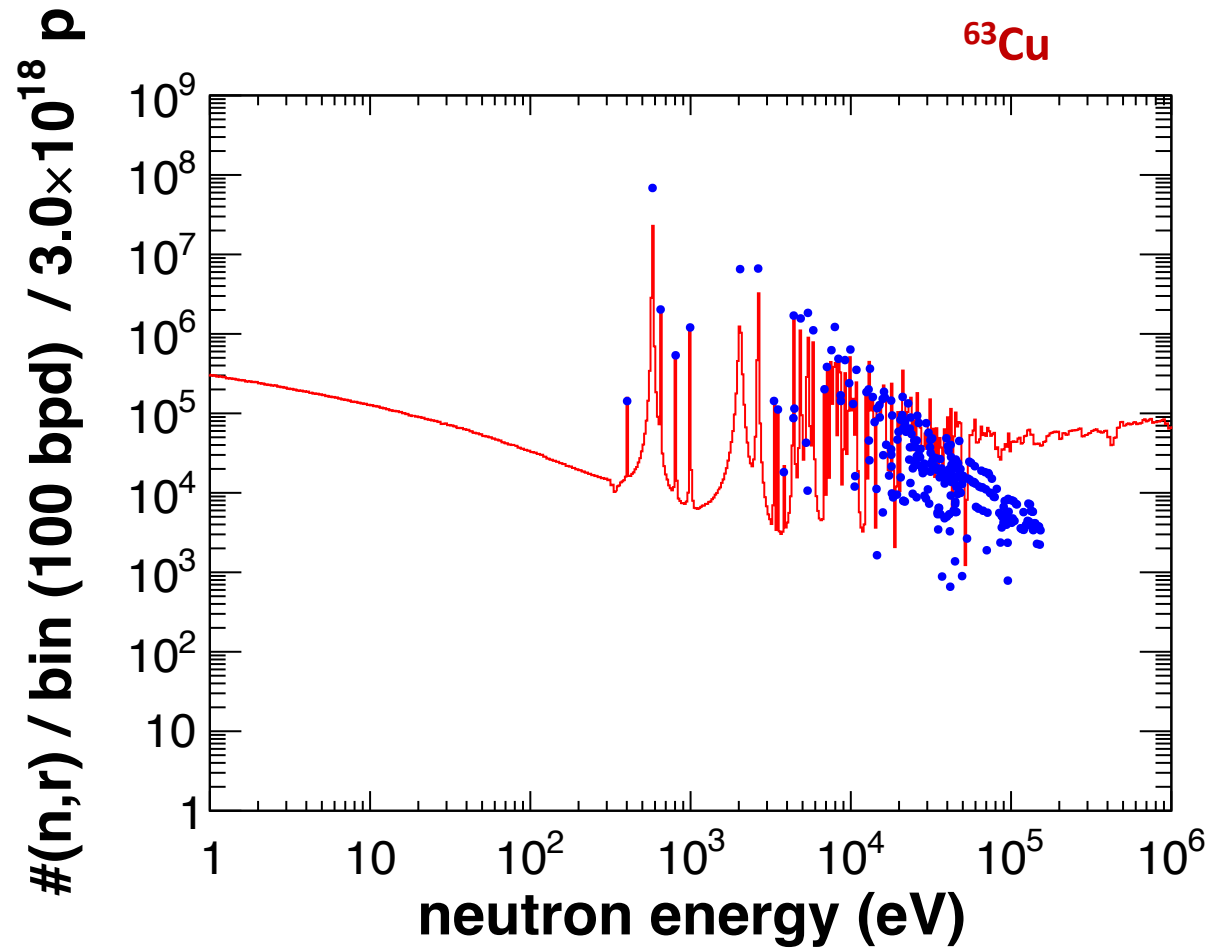
Uncertainties in a capture experiment

Mastromarco et al., Eur. Phys. J. A (2019) 55: 9

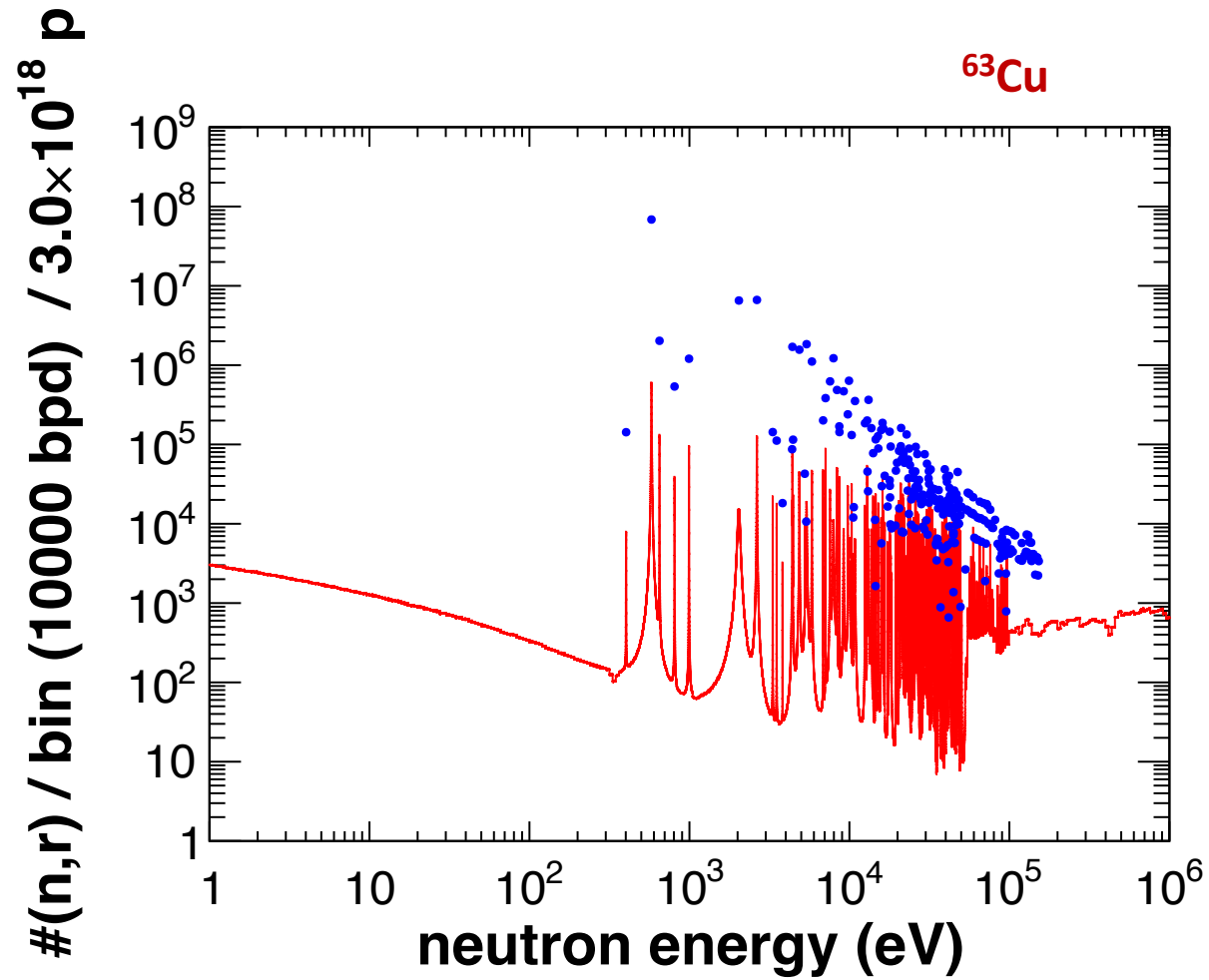
Table 3. Summary of the correlated uncertainties in the $^{155}\text{Gd}(n, \gamma)$ and $^{157}\text{Gd}(n, \gamma)$ cross section measurements.

Source of uncertainty	$^{155}\text{Gd}(n, \gamma)$		$^{157}\text{Gd}(n, \gamma)$	
	near thermal	resonance region	near thermal	resonance region
Normalization	1.2%	1.2%	1.2%	1.2%
PHWT	1.5%	1.5%	1.5%	1.5%
Background	1.4%	$\approx 1\%$	1.0%	$\approx 1\%$
BIF	1.5%		1.5%	
Flux	1.0%	1.0%	1.0%	1.0%
Sample mass	1.0%	$< 0.2\%$	2.1%	$< 0.2\%$
Temperature		1%		1%
Total	3.2%	2.6%	3.5%	2.6%

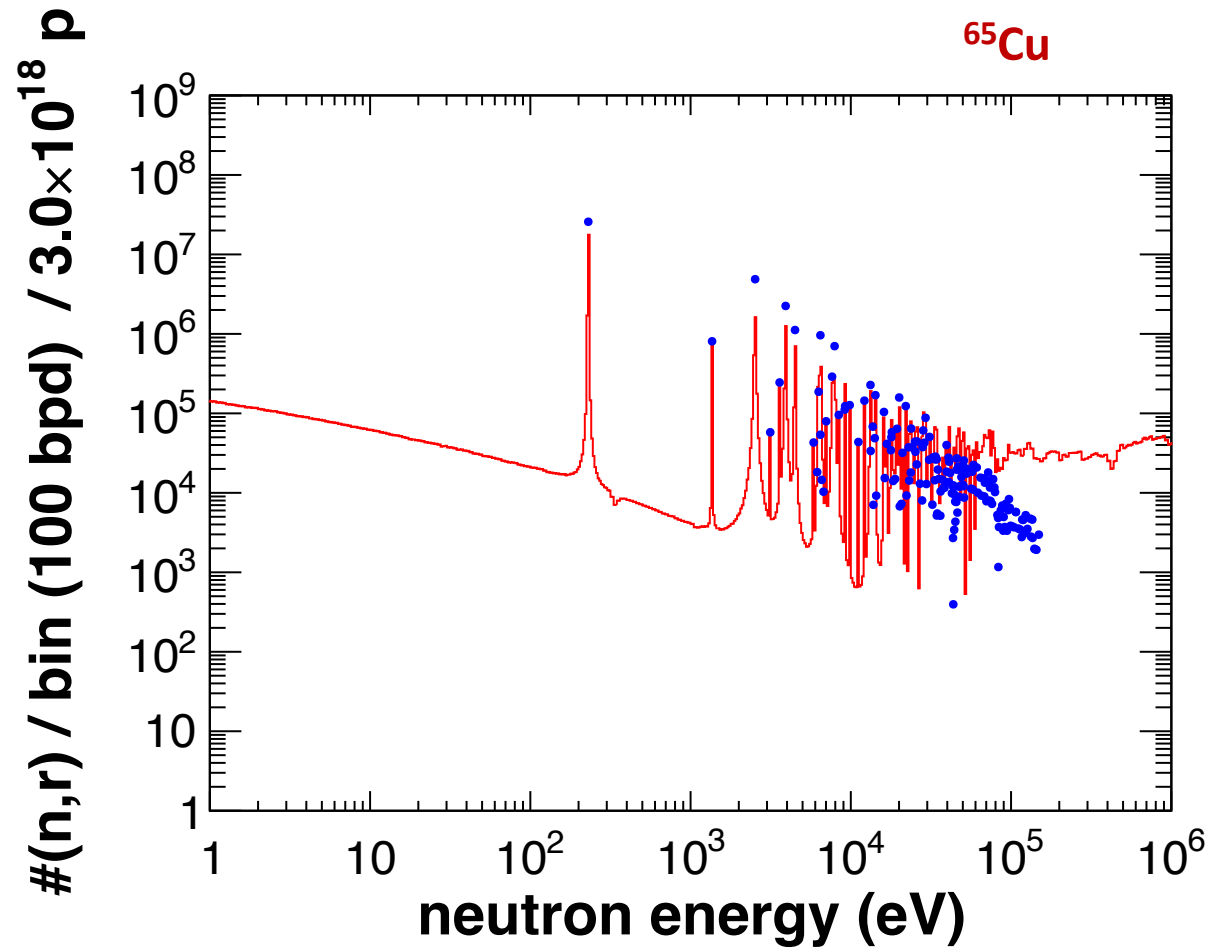
Count rate estimate



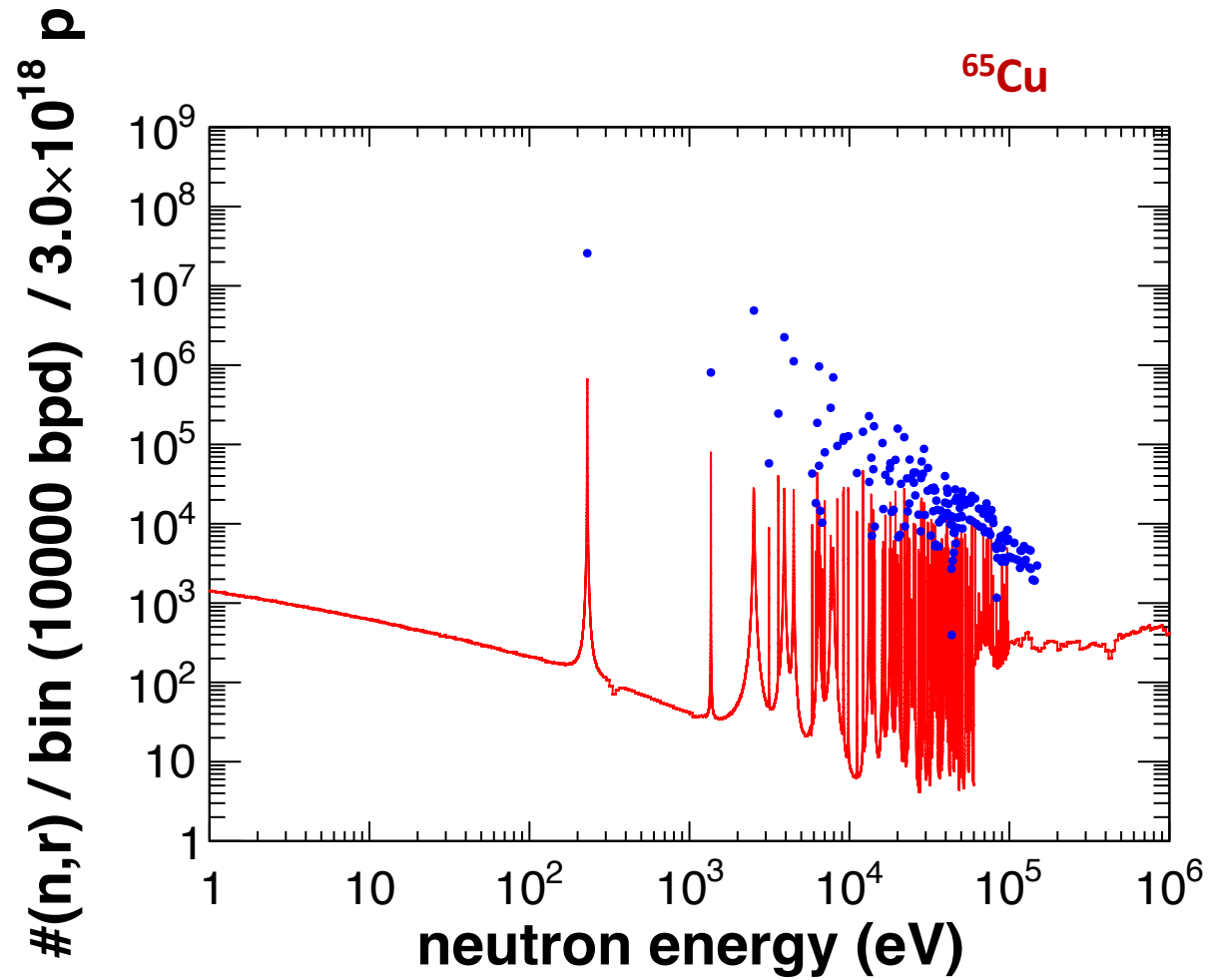
Countrate



Count rate estimate

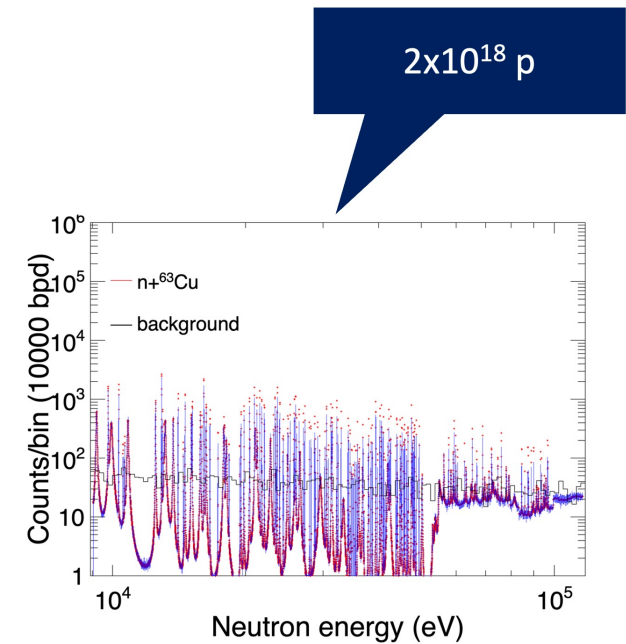
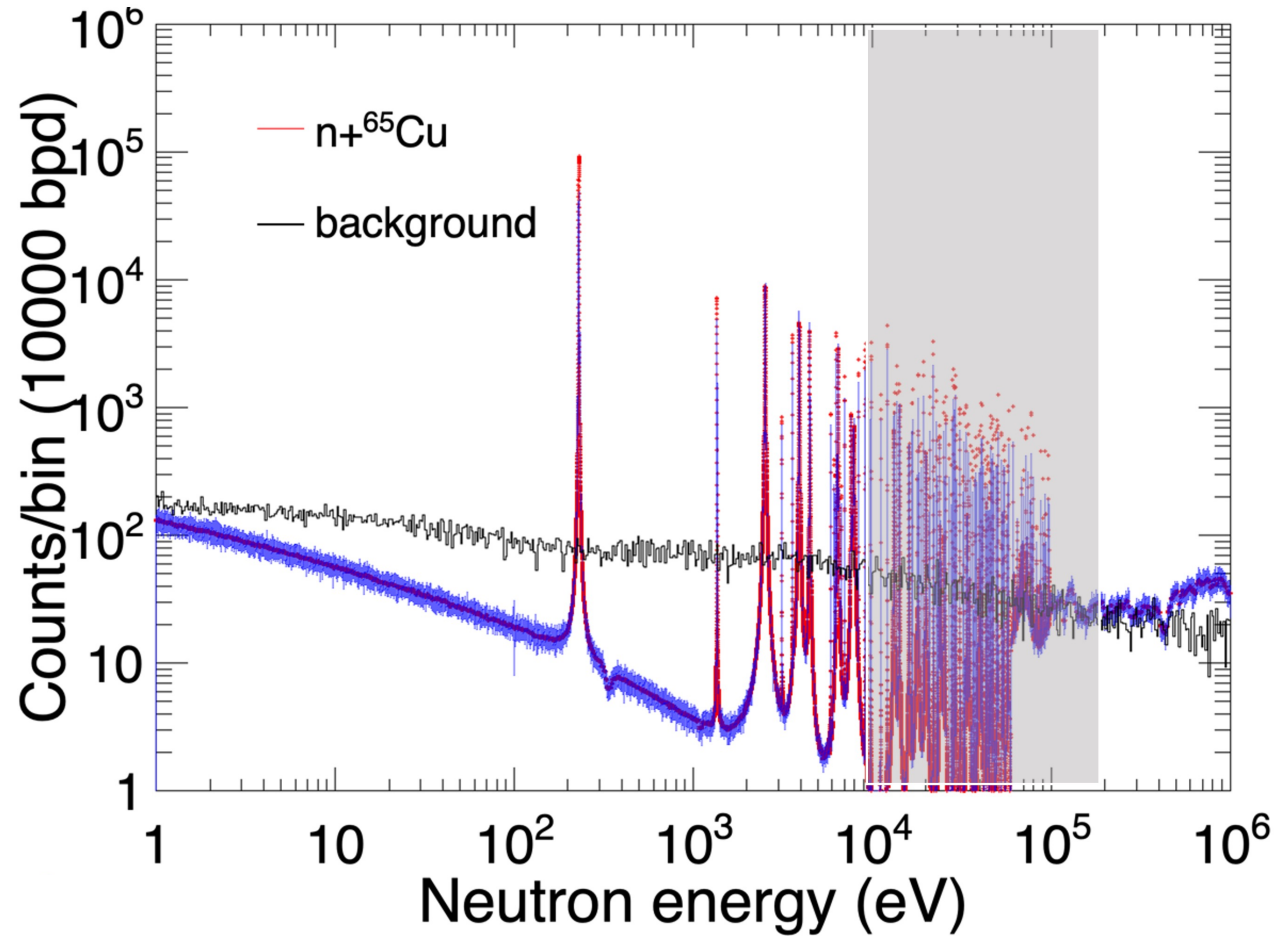


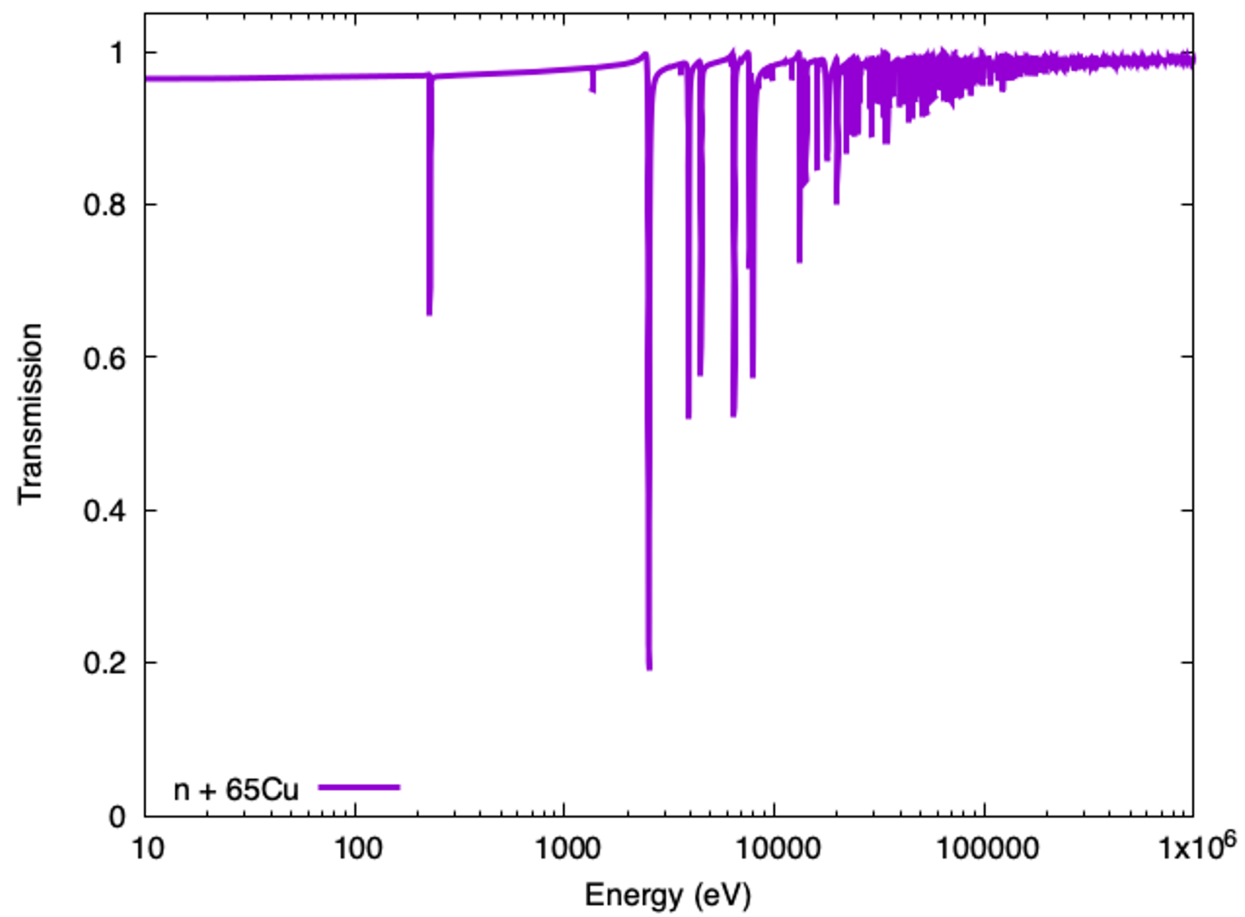
Count rate estimate



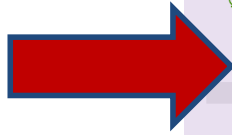
Count rate estimate

(n,γ) @ EAR 1, 4 C_6D_6 , $\emptyset = 3$ cm, mass = 2 g





Transmission



n	Display	Year	Author-1	Energy range, eV	Points	Reference	Subentry#P	NSR-Key	Info+
1)	29-CU-63 (N,TOT),,SIG C4: MF=3 MT=1 Op=0								
Quantity: [CS] Cross section									
1	<input type="checkbox"/> + i X4 X4± CSV + T4 Cov	1977	M.S.Pandey+	5.00e1	1	+ J, PR/C, 15, 600, 197702	10725002 [6]	1977PA04	
2	<input type="checkbox"/> + i X4 X4± CSV + T4 Cov			3.83e3	1.12e6	12165		1977PA04	
3	<input type="checkbox"/> + i X4 X4± CSV + T4 Cov			3.16e1	1.85e5	6208		1977PA04	
4	<input type="checkbox"/> + i X4 X4± CSV + T4 Cov	1972	A.I.Dyumin+	1.42e7	1	+ J, IZV, 36, 852, 1972	40149011 [6]	1972DY02 #2:1972dy02 #3:pdf	
5	<input type="checkbox"/> + i X4 X4± CSV + T4 Cov	1968	G.Rohr+	3.98e4	1.43e5	880	+ P, EANDC (E)-89, 1, 6802	20151002 [5]	
g*	<input type="checkbox"/> + i X4 X4± CSV + T4 Cov	1966	W.M.Good+	2.06e3	5.95e4	299	+ J, PR, 151, 912, 1966	11626024 [5]	1966G038
7	<input type="checkbox"/> + i X4 X4± CSV + T4 Cov	1957	H.Marshak+	1.10e3	9.72e4	164	+ J, PR, 106, 110, 57	11640010 [4]	1957MA59
2)	29-CU-63 (N,TOT),,SIG,,AV C4: MF=3 MT=1 Op=0								
Quantity: [CS] Cross section									
8	<input type="checkbox"/> + i X4 X4± CSV + T4 Cov	1975	V.V.Filippov	5.00e4	9.50e5	11	[pdf]+ C, 75KIEV, 2, 53, 197505	41301002 [6]	
3)	29-CU-63 (N,TOT),,SIG,,RES C4: MF=3 MT=1 Op=0								
Quantity: [CS] Cross section at resonance									
9	<input type="checkbox"/> + i X4 X4± CSV + T4 Cov	1958	R.E.Cote+	5.77e2	5.39e3	3	+ J, PR, 111, 288, 58	116410182 [4]	1958CO69
4)	29-CU-63 (N,TOT),,TRN C4: MF=3 MT=? Op=0								
Quantity: [CS] Transmission									
10	<input type="checkbox"/> + i X4 X4± CSV + T4 Cov	2013	K.Kauwenberghs+	1.50e2	9.00e4	21528	+ R, EUR-26479, 2013	23325002 [1]	
11	<input type="checkbox"/> + i X4 X4± CSV + T4 Cov			1.50e2	9.00e4	21528		23325003 [1]	
12	<input type="checkbox"/> + i X4 X4± CSV + T4 Cov			1.50e2	9.00e4	21528		23325004 [1]	
5)	29-CU-63 (N,TOT),,WID C4: MF=402 MT=6010Op=0								
Quantity: [RP] Resonance width									
13	<input type="checkbox"/> + i X4 X4± CSV + T4 Cov	1958	R.E.Cote+	5.77e2	5.39e3	3	+ J, PR, 111, 288, 58	116410182 [4]	1958CO69

PHYSICAL REVIEW C VOLUME 15, NUMBER 2 FEBRUARY 1977

Neutron total cross sections and resonance parameters of ^{63}Cu and ^{65}Cu . I*

M. S. Pandey[†] and J. B. Garg[†]

State University of New York at Albany, Albany, New York 12222

J. A. Harvey

Oak Ridge National Laboratory, Oak Ridge, Tennessee 37830

(Received 8 April 1976)

High resolution neutron total cross sections of the isotopes of copper have been measured from about 10 to 150 keV using 5 nsec electron pulses and a flight path of 78.2 m. From the area and shape analysis of the transmission and total cross section data, precise values of the resonance parameters, such as E_0 , Γ_n^0 , Γ_n^1 , J^π , etc., have been determined. For example, for ^{63}Cu many *s*-wave resonances have been observed from 10 to 150 keV giving values of $\langle D \rangle_{J=1} = (2.7 \pm 0.3)$ keV, $\langle D \rangle_{J=2} = (4.0 \pm 0.5)$ keV, $\langle D \rangle_{J=1+2} = (1.63 \pm 0.13)$ keV, $S_{0J=1} = (3.0 \pm 0.6) \times 10^{-4} \text{ eV}^{-1/2}$, $S_{0J=2} = (2.0 \pm 0.5) \times 10^{-4} \text{ eV}^{-1/2}$, $S_{0J=1,2} = (2.5 \pm 0.4) \times 10^{-4} \text{ eV}^{-1/2}$. For ^{65}Cu *s*-wave resonances were observed giving values of $\langle D \rangle_{J=1} = (3.6 \pm 0.4)$ keV, $\langle D \rangle_{J=2} = (5.0 \pm 0.7)$ keV, $\langle D \rangle_{J=1+2} = (2.12 \pm 0.19)$ keV, $S_{0J=1} = (2.9 \pm 0.6) \times 10^{-4} \text{ eV}^{-1/2}$, $S_{0J=2} = (1.8 \pm 0.5) \times 10^{-4} \text{ eV}^{-1/2}$, $S_{0J=1,2} = (2.3 \pm 0.4) \times 10^{-4} \text{ eV}^{-1/2}$.