


BAYESIAN ESTIMATES FOR THE UNCERTAINTIES

Alexander Huss



WHAT IS THE **UNCERTAINTY** Δ_{TH} OF MY RESULT?

- increasingly urgent to address with $\Delta_{\text{EXP}} \searrow$ (\leftrightarrow HL-LHC)
 - ▶ what does 5σ mean if Δ_{TH} non-negligible?
 - ▶ interpretation of data in need for robust Δ_{TH} : PDF fits, χ^2 in ATLAS jets, ...
- various sources that contribute to Δ_{TH} :
 - ▶ Δ_{α_s} , Δ_{param} : parametric uncertainties \leftrightarrow exp. extraction
 - ▶ Δ_{PDF} : parton distribution functions (PDFs) \leftrightarrow fits
 - ▶ $\Delta_{\text{non pert.}}$: hadronisation, UE, ... \leftrightarrow parton showers [e.g. HERWIG vs. PYTHIA]
 - ▶ Δ_{MHO} : *missing higher-order (MHO)* corrections 

CONVENTIONAL APPROACH FOR Δ_{MHO} — SCALE VARIATION

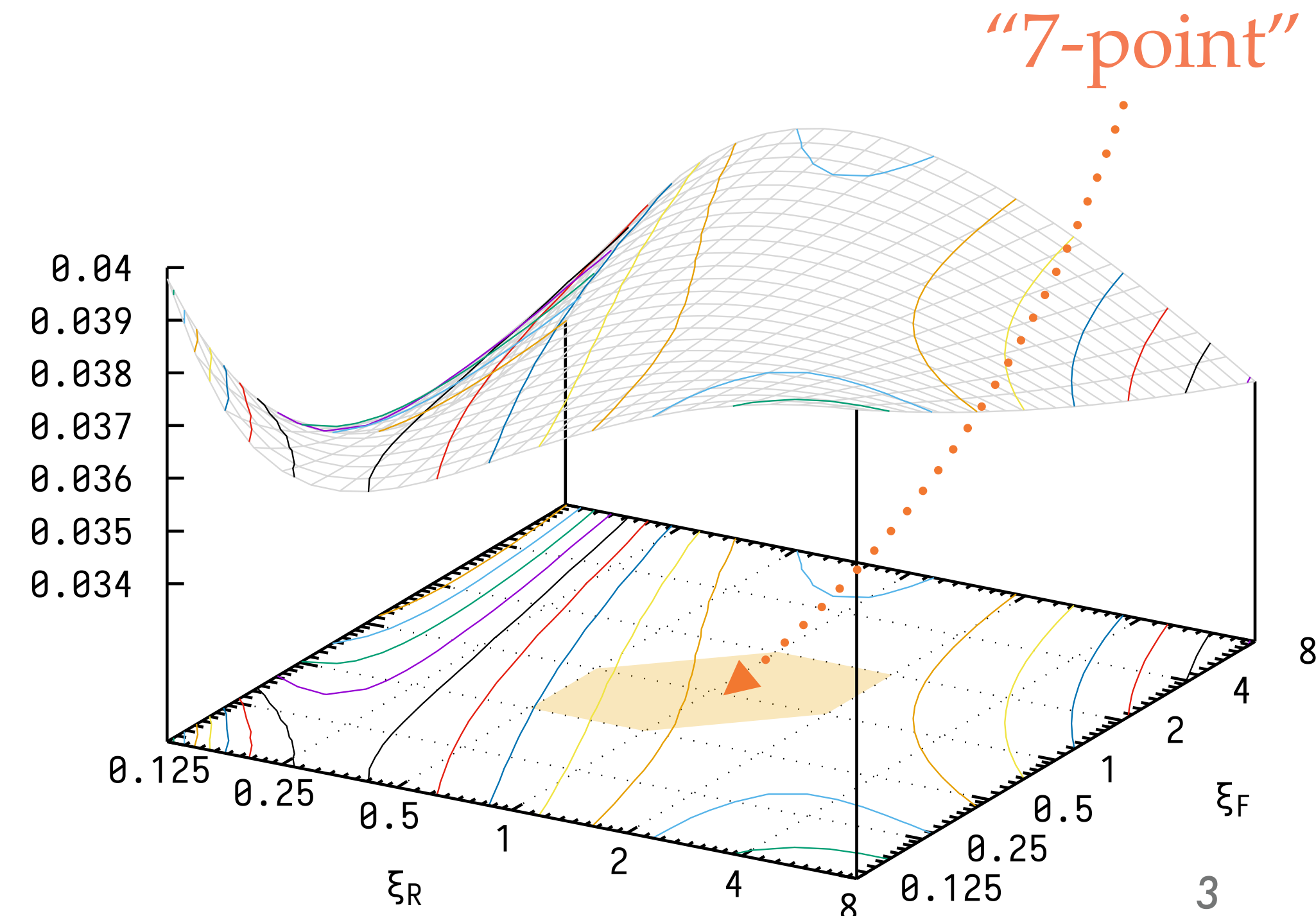
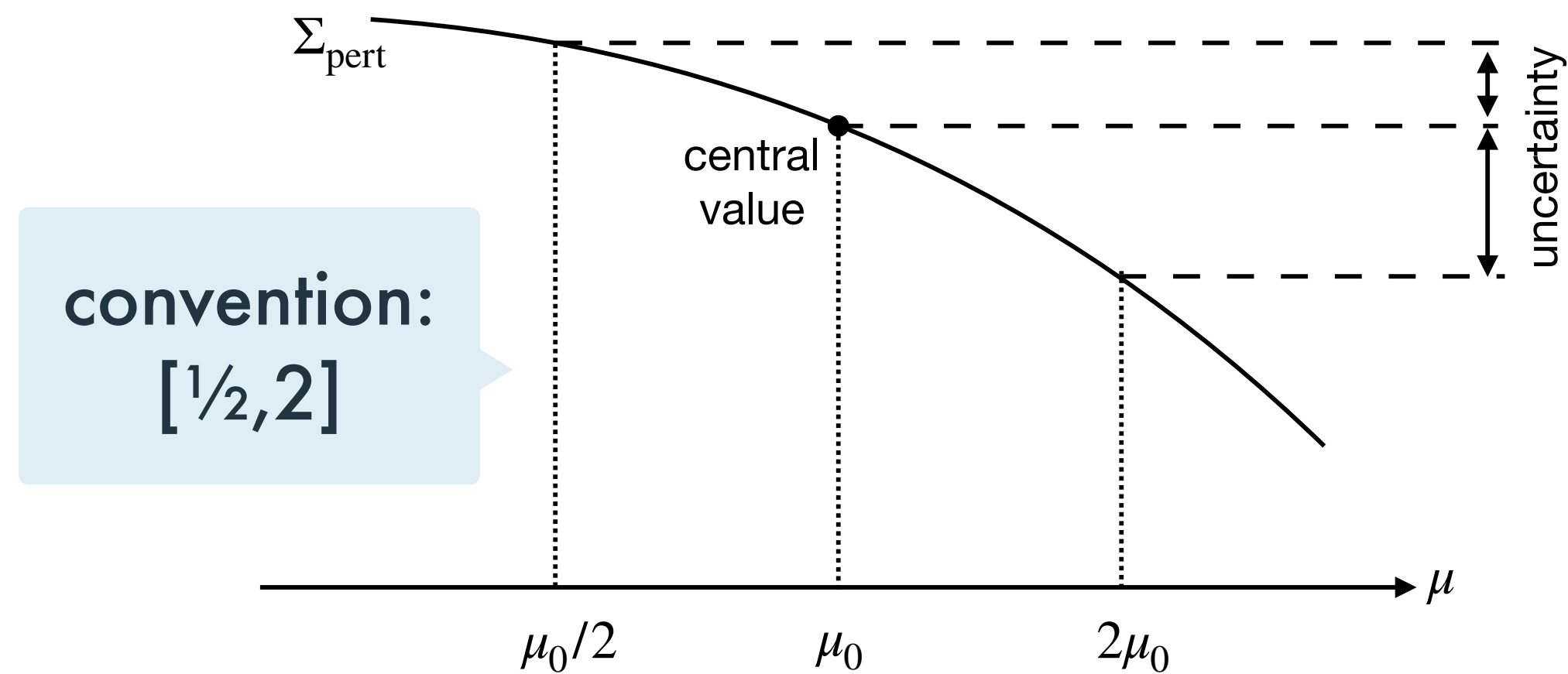
- approximation for an observable @ (next-to-)ⁿ leading order:

▶ NⁿLO: $\Sigma \simeq \Sigma_n(\mu) = \sum_{k=0}^n \Sigma^{(k)}(\mu) \propto \alpha_s^{n_0+k}$

electroweak (EW):
 ↪ scheme dependence
 ↪ $\alpha \ll \alpha_s$

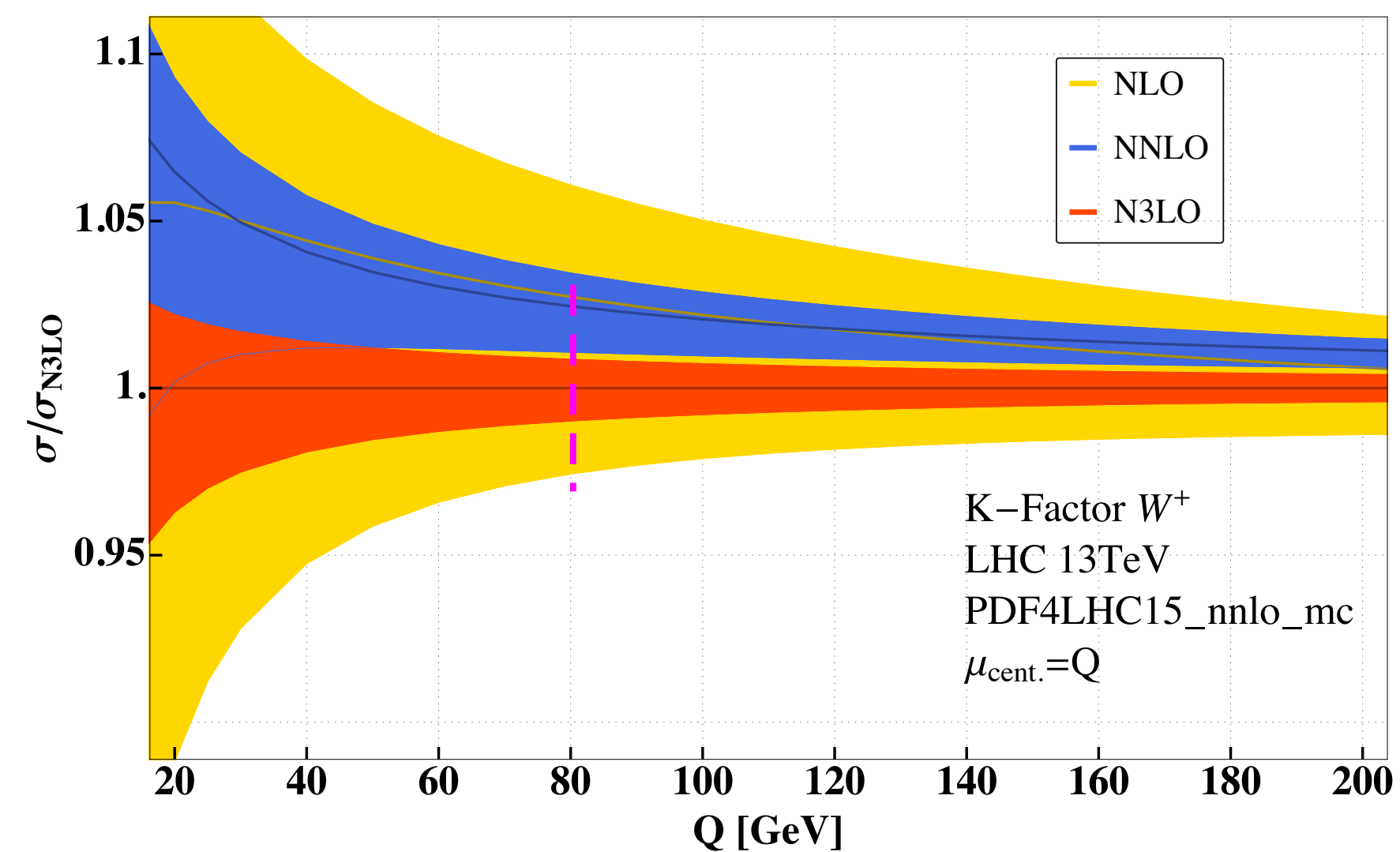
- truncation of series induces a sensitivity to terms of the next order

$$\mu \frac{d}{d\mu} \Sigma_n(\mu) = \mathcal{O}(\alpha_s^{n_0+n+1}) = \mathcal{O}(\Delta_{\text{MHO}})$$



ISSUES WITH STANDARD SCALE VARIATIONS

- known to be insufficient:
 - exclusive jet(s) (veto)
 - ratios (correlation?)
 - cancellations (e.g. $q\bar{q}$ vs. qg in DY)



[Duhr, Dulat, Mistlberger '20]

- choice of the central scale
 - fastest apparent convergence (FAC)
 - $\hookrightarrow \Sigma^{(n)}(\mu_{\text{FAC}}) = 0$
 - principle of minimal sensitivity (PMS)
 - $\hookrightarrow \left. \frac{\partial}{\partial \mu} \Sigma^{(n)}(\mu) \right|_{\mu_{\text{PMS}}} = 0$
 - BLM/PMC
 - [Brodsky, Lepage, Mackenzie '83]; [Brodsky, Di Giustino '12]
 - ...
- crucially:** *no probabilistic interpretation!*
 - \rightsquigarrow can we do better?

PROBABILITY DISTRIBUTIONS FOR Δ_{MHO}

[Cacciari, Houdeau '11]

- Sequence of perturbative *corrections* δ_k normalised w.r.t. LO (dimensionless)

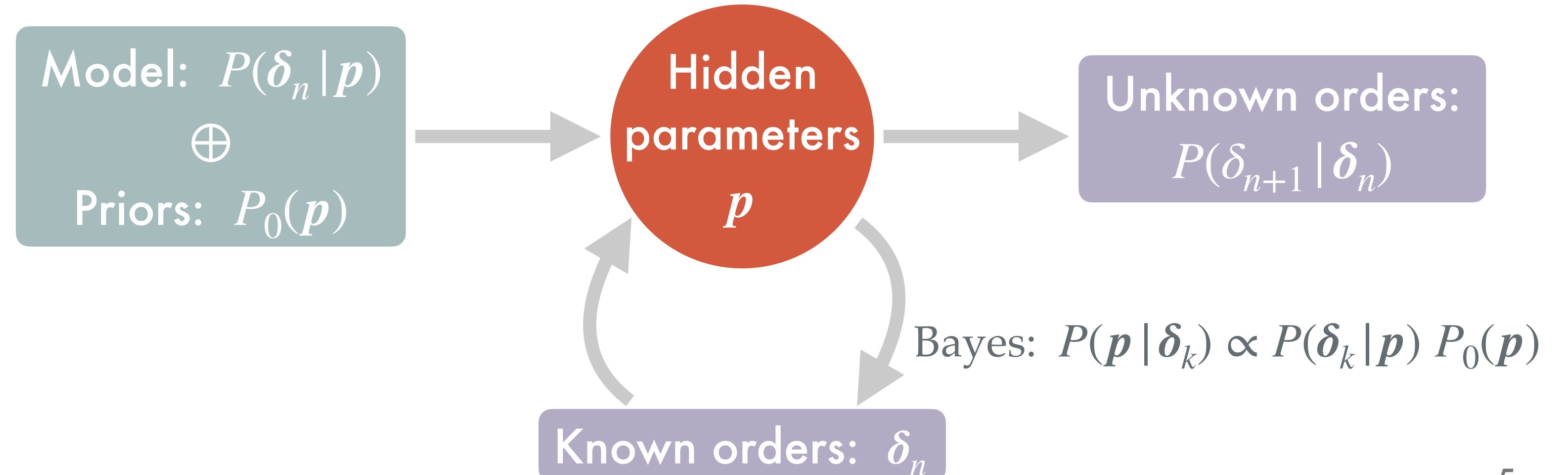
$$\Sigma_n = \Sigma^{(0)} (1 + \delta_1 + \dots + \delta_n) \quad \rightsquigarrow \quad \delta_k = \mathcal{O}(\alpha_s^k)$$

- Probability distribution for δ_{n+1} , given $\delta_n = (\delta_0, \delta_1, \dots, \delta_n)$

$$P(\delta_{n+1} | \delta_n) = \frac{P(\delta_{n+1})}{P(\delta_n)} = \frac{\int d^m \mathbf{p} P(\delta_{n+1} | \mathbf{p}) P_0(\mathbf{p})}{\int d^m \mathbf{p} P(\delta_n | \mathbf{p}) P_0(\mathbf{p})}$$

$$P(A, B) = P(A | B) P(B)$$

$$P(A) = \int dB P(A, B)$$



THE CH MODEL

[Cacciari, Houdeau '11]

- perturbative expansion $\delta_k = c_k \alpha_s^k$ bounded by a geometric series: $|c_k| \leq \bar{c} \quad \forall k$

$$\left| \sum_k \delta_k \right| \leq \sum_k |c_k| \alpha_s^k \leq \sum_k \bar{c} \alpha_s^k$$

- ▶ one hidden parameter: \bar{c}
- ▶ constrain upper bound \bar{c} from known orders
 \rightsquigarrow constraint on unknown coefficients c_{n+1}

- limitations:

- ▶ α_s at what scale? why not: $\frac{\alpha_s}{\pi}, \frac{\alpha_s}{2\pi}, \alpha_s \ln^2(v), \alpha_s \ln(v), \dots$?

- why not let the model figure out the expansion parameter itself?

$$c_k \sim \eta^k$$

\hookrightarrow **survey of observables**

[Bagnaschi, Cacciari, Guffanti, Jenniches '14]

\hookrightarrow **fitting**

[Forte, Isgro, Vita '13]

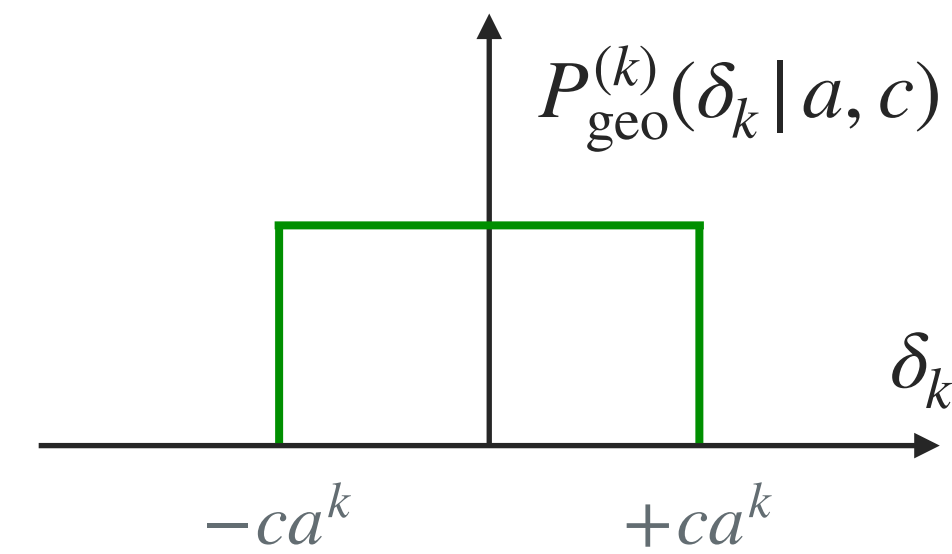
THE GEOMETRIC MODEL

[Bonvini '20]

- bounded by a geometric series with expansion parameter a :

$$|\delta_k| \leq c a^k \quad \forall k \quad \Leftrightarrow \text{two model parameters: } a, c$$

- model:**
$$P_{\text{geo}}^{(k)}(\delta_k | a, c) = \frac{1}{2c a^k} \Theta\left(c - \frac{|\delta_k|}{a^k}\right)$$

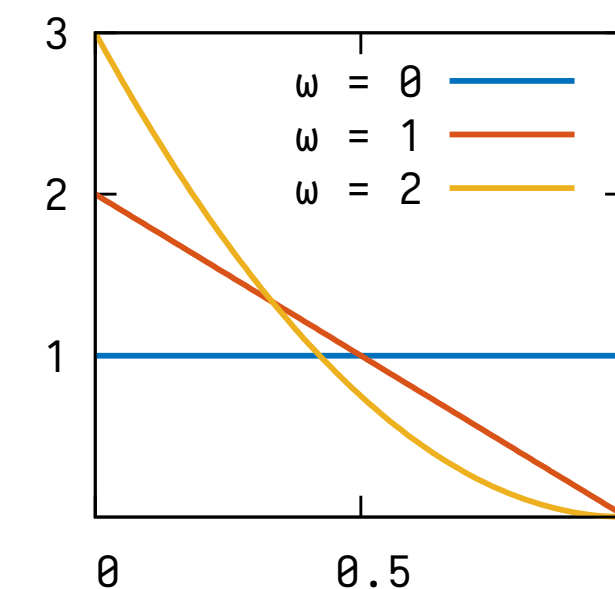


- priors:**
$$P_0(a, c) = P_0(a) P_0(c)$$

$$P_0(a) = (1 + \omega) (1 - a)^\omega \Theta(a) \Theta(1 - a)$$

$$P_0(c) = \frac{\varepsilon}{c^{1+\varepsilon}} \Theta(c - 1)$$

$$\Leftrightarrow dc/c \sim d \ln(c) \quad (\varepsilon: \text{regulator})$$

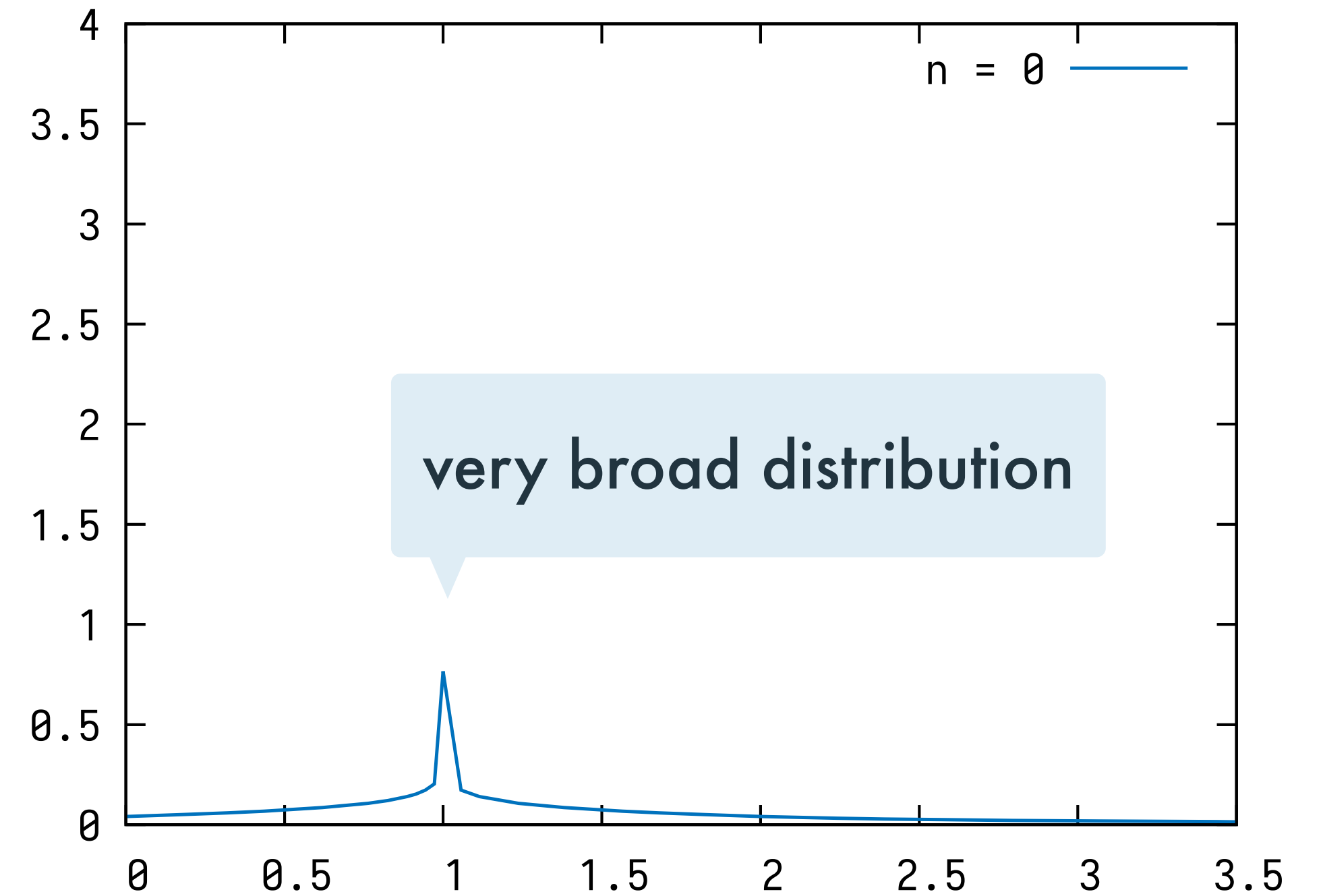
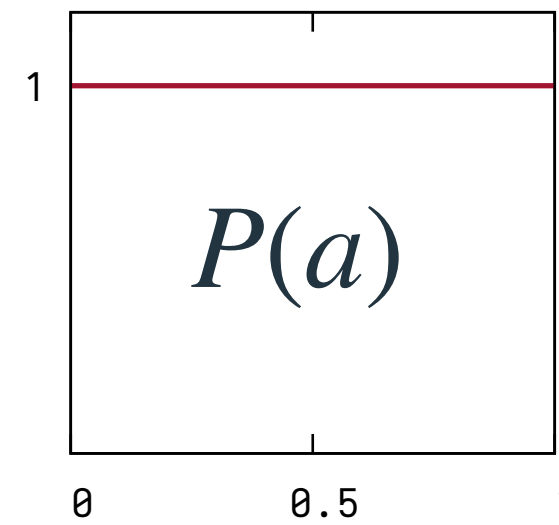


THE INFERENCE STEP — GEOMETRIC SERIES: $\delta_k = (0.7)^k$

LO $\delta_0 \equiv 1$

$$P_0(a, c) = \Theta(a) \Theta(1 - a) P_0(c)$$

chose $\omega = 0$ for flat prior in a



no inference yet!
 $P(\delta_1)$ entirely determined by the model & priors

$$P(\delta_1) = \int da \int dc P_{\text{geo}}^{(1)}(\delta_1 | a, c) P_0(a, c)$$

THE INFERENCE STEP — GEOMETRIC SERIES: $\delta_k = (0.7)^k$

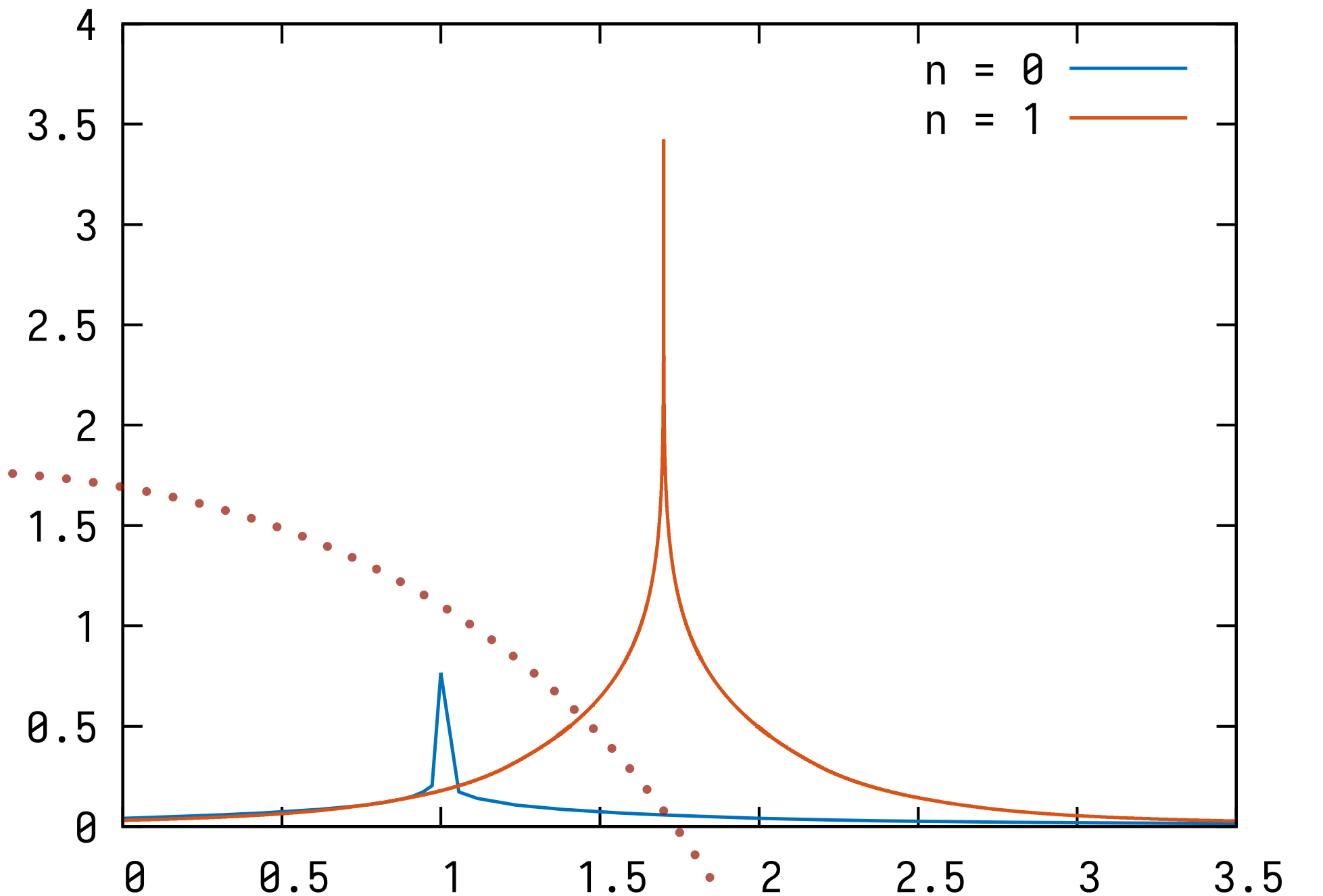
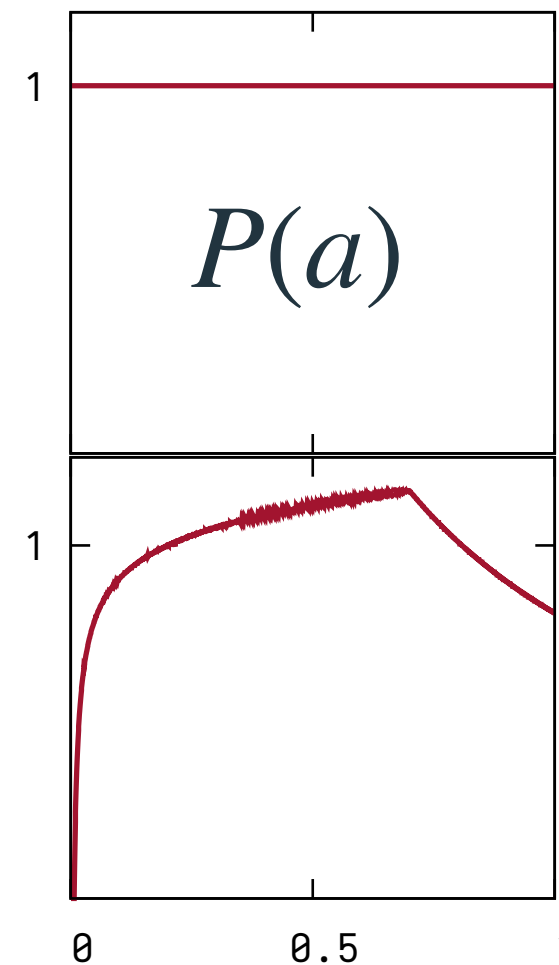
- LO $\delta_0 \equiv 1$

$$P_0(a, c) = \Theta(a) \Theta(1 - a) P_0(c)$$

- NLO $\delta_1 = 0.7$

$$P(a, c | \delta_1) \propto P_{\text{geo}}^{(1)}(\delta_1 | a, c) P_0(a, c)$$

}
}
}
 posterior likelihood prior



Bayes' theorem:

$$P(A | B) = \frac{P(B | A) P(A)}{P(B)}$$

δ_k independent:

$$P(\delta_2 | \delta_1) = P(\delta_2)$$

$$P(\delta_2 | \delta_1) = \int da \int dc P(\delta_2 | \delta_1, a, c) P(a, c | \delta_1)$$

$$\propto \int da \int dc P_{\text{geo}}^{(2)}(\delta_2 | a, c) P_{\text{geo}}^{(1)}(\delta_1 | a, c) P_0(a, c)$$

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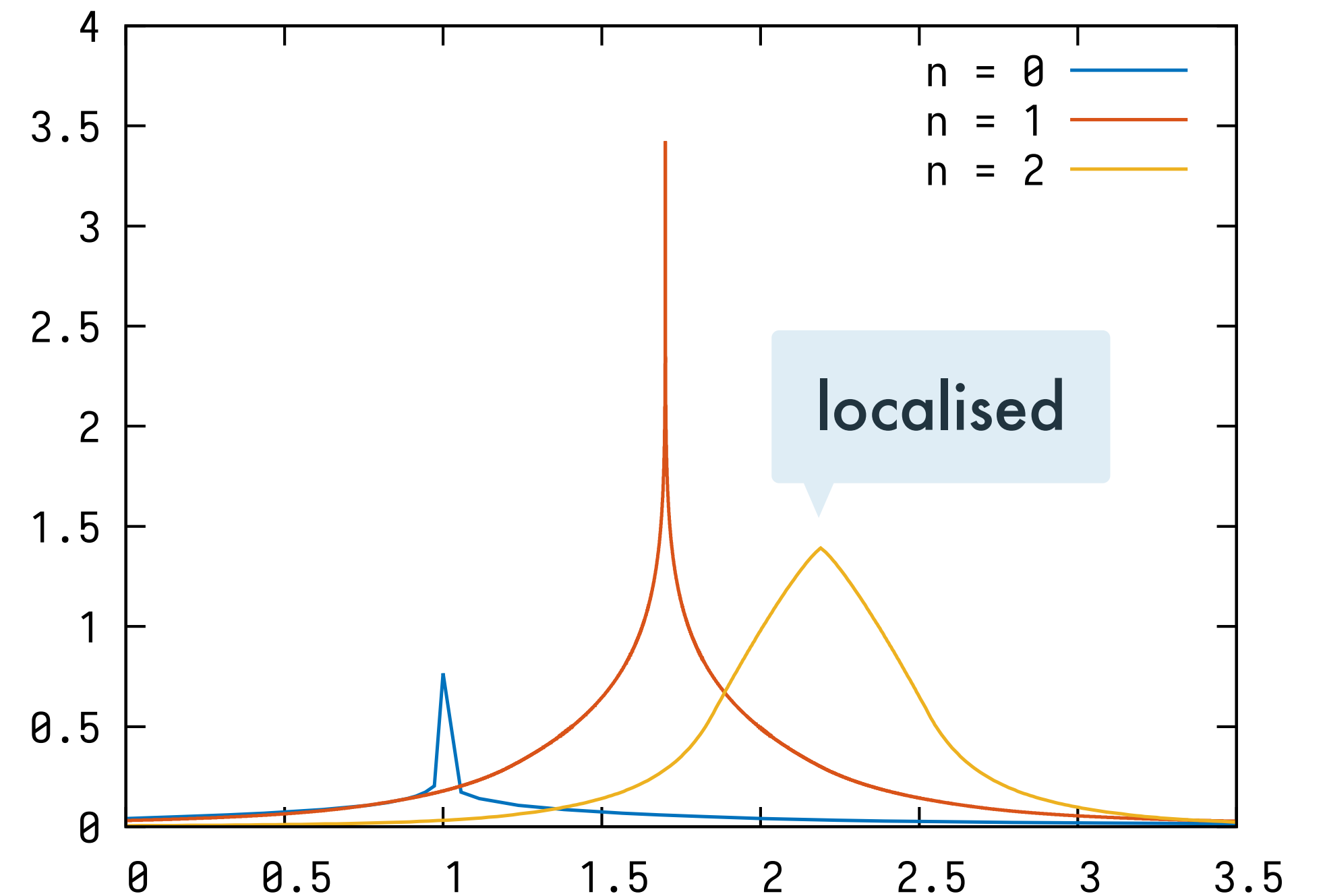
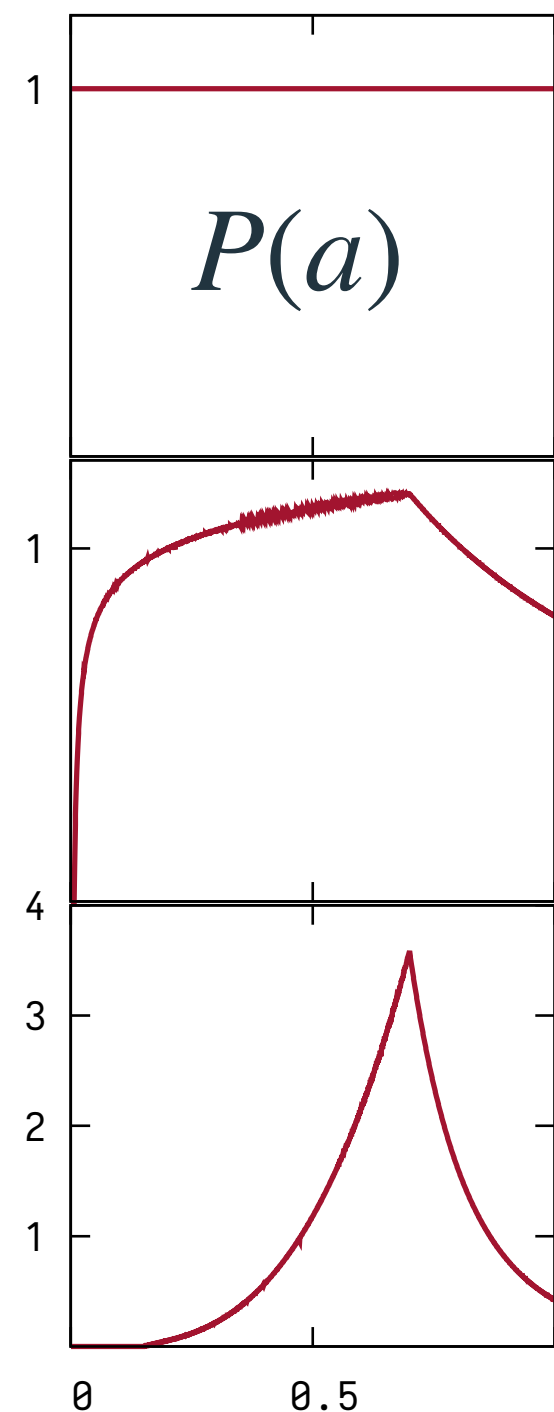
- N²LO $\delta_2 = 0.7^2$

$$P(a, c | \delta_1, \delta_2) \propto P(\delta_2 | \delta_1, a, c) P(a, c | \delta_1)$$

$$\propto P_{\text{geo}}^{(2)}(\delta_2 | a, c) P_{\text{geo}}^{(1)}(\delta_1 | a, c) P_0(a, c)$$

Bayes' theorem
& independence

$a \sim 0.7$
also: $c \sim 1$



$$P(\delta_3 | \delta_1, \delta_2) \propto \int da \int dc \prod_{k=1}^3 \left[P_{\text{geo}}^{(k)}(\delta_k | a, c) \right] P_0(a, c)$$

THE INFERENCE STEP — GEOMETRIC SERIES: $\delta_k = (0.7)^k$

- LO $\delta_0 \equiv 1$

$$P_0(a, c) = \Theta(a) \Theta(1 - a) P_0(c)$$

- NLO $\delta_1 = 0.7$

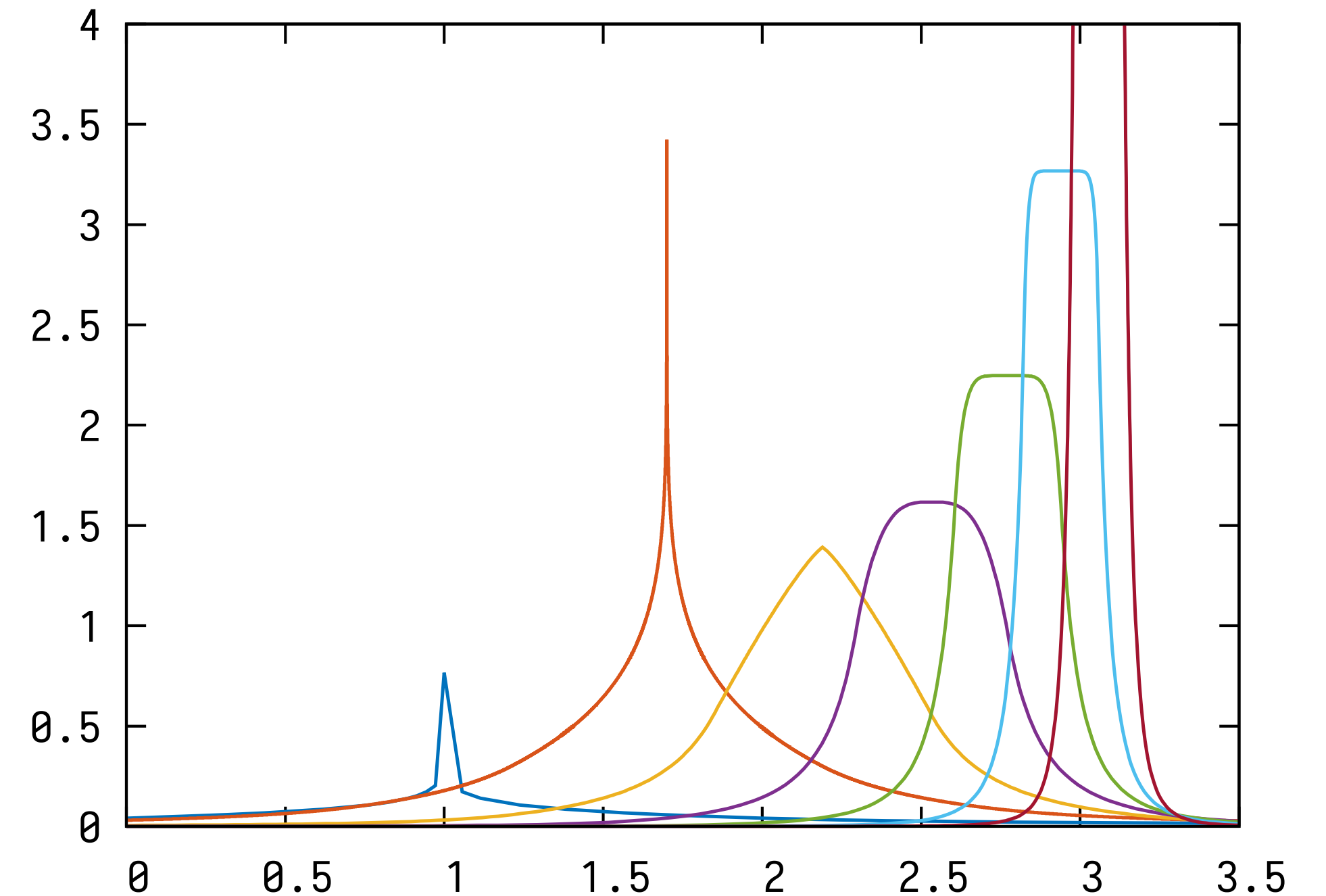
$$P(a, c | \delta_1) \propto P_{\text{geo}}^{(1)}(\delta_1 | a, c) P_0(a, c)$$

- N²LO $\delta_2 = 0.7^2$

$$P(a, c | \delta_1, \delta_2) \propto P(\delta_2 | \delta_1, a, c) P(a, c | \delta_1)$$

$$\propto P_{\text{geo}}^{(2)}(\delta_2 | a, c) P_{\text{geo}}^{(1)}(\delta_1 | a, c) P_0(a, c)$$

- ...



$$P(\delta_{n+1} | \delta_n) \propto \int da \int dc \prod_{k=1}^n \left[P_{\text{geo}}^{(k)}(\delta_k | a, c) \right] P_0(a, c)$$

can be solved analytically

THE *abc* MODEL — ASYMMETRIC GEOMETRIC MODEL

[Duhr, AH, Mazeliauskas, Szafron '21]

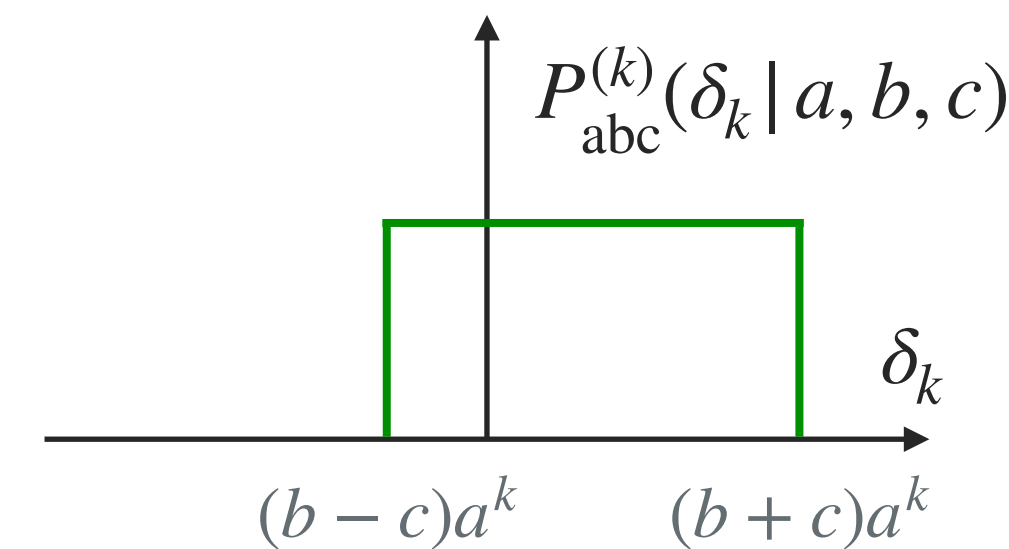
geometric model is symmetric: $P(\delta_0, \dots, \delta_n) = P(|\delta_0|, \dots, |\delta_n|) \rightsquigarrow \langle \delta_{n+1} \rangle_{\text{geo}} = 0$

allow for different lower & upper bound:

$$b - c \leq \frac{\delta_k}{a^k} \leq b + c \quad \forall k \quad \Leftrightarrow \text{three model parameters: } a, b, c$$

bias/offset

model: $P_{abc}^{(k)}(\delta_k | a, b, c) = \frac{1}{2c|a|^k} \Theta\left(c - \left|\frac{\delta_k}{a^k} - b\right|\right)$

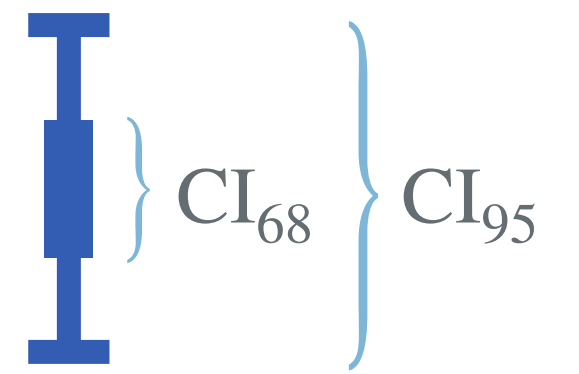


priors: $P_0(a, b, c) = P_0(a) P_0(b, c)$

$$P_0(a) = \frac{1}{2} (1 + \omega) (1 - |a|)^\omega \Theta(1 - |a|) \quad \Leftrightarrow \text{support: } [-1, +1] \text{ (alternating } \checkmark)$$

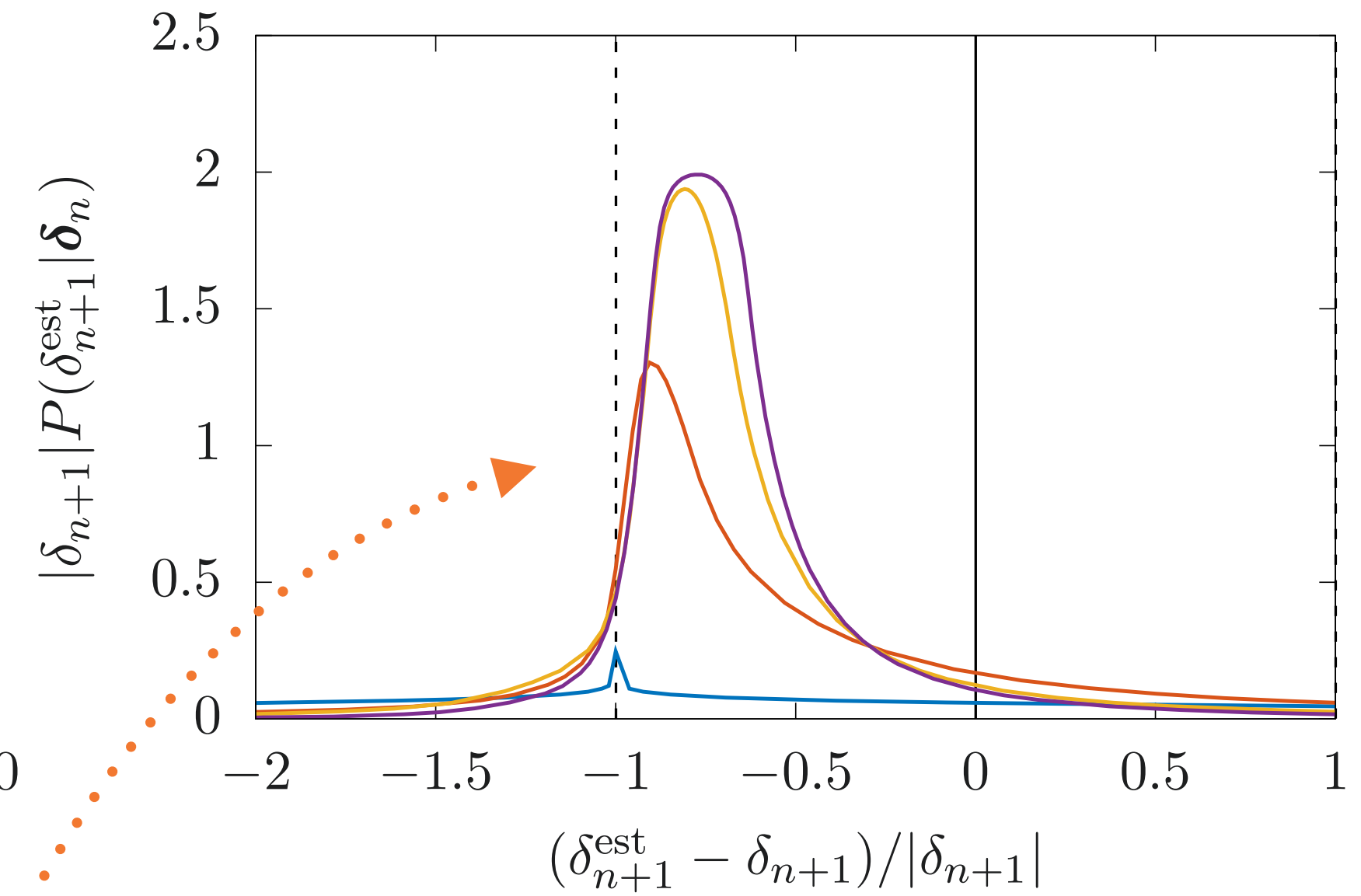
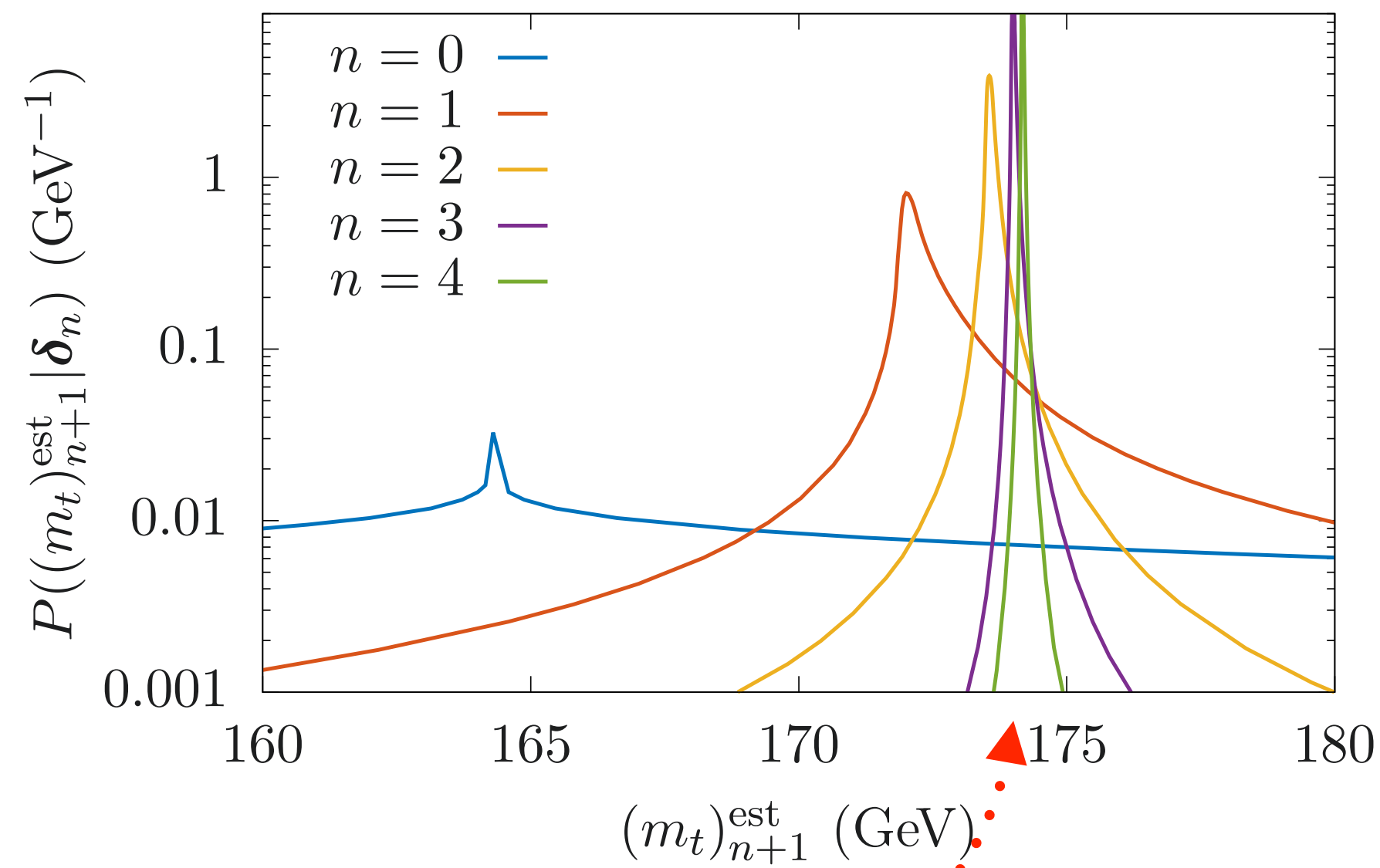
$$P_0(b, c) = \frac{\varepsilon \eta^\varepsilon}{c^{1+\varepsilon}} \Theta(c - \eta) \frac{1}{2\xi c} \Theta(\xi c - b)$$

A REAL-WORLD EXAMPLE — m_t (OS \leftrightarrow $\overline{\text{MS}}$)



$$m_t = \frac{Z_m^{\text{OS}}}{Z_m^{\overline{\text{MS}}}(\mu_R)} \overline{m}_t(\mu_R)$$

$$= \sum_k m_t^{(k)}$$

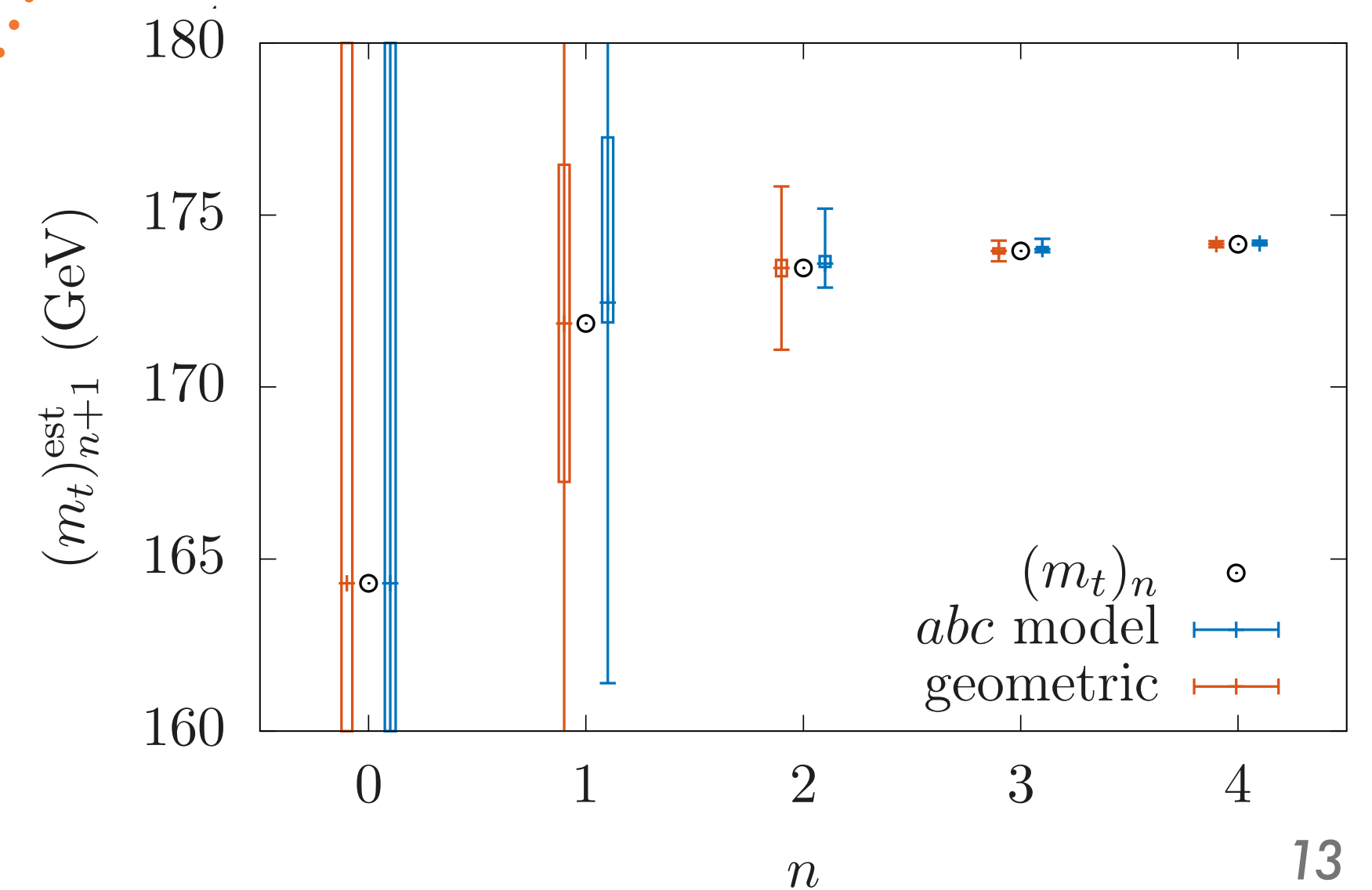


strongly peaked $n \nearrow$

positive corrections
anticipated

estimate for $m_t - (m_t)_4$

- ▶ $\text{CI}_{68} = [0.008, 0.046] \text{ GeV}$
- ▶ $\text{CI}_{95} = [-0.027, 0.112] \text{ GeV}$



WHAT TO DO WITH THE THE SCALE μ ?

• $\forall \mu \rightsquigarrow P(\delta_3 | \delta_0, \delta_1, \delta_2; \mu)$

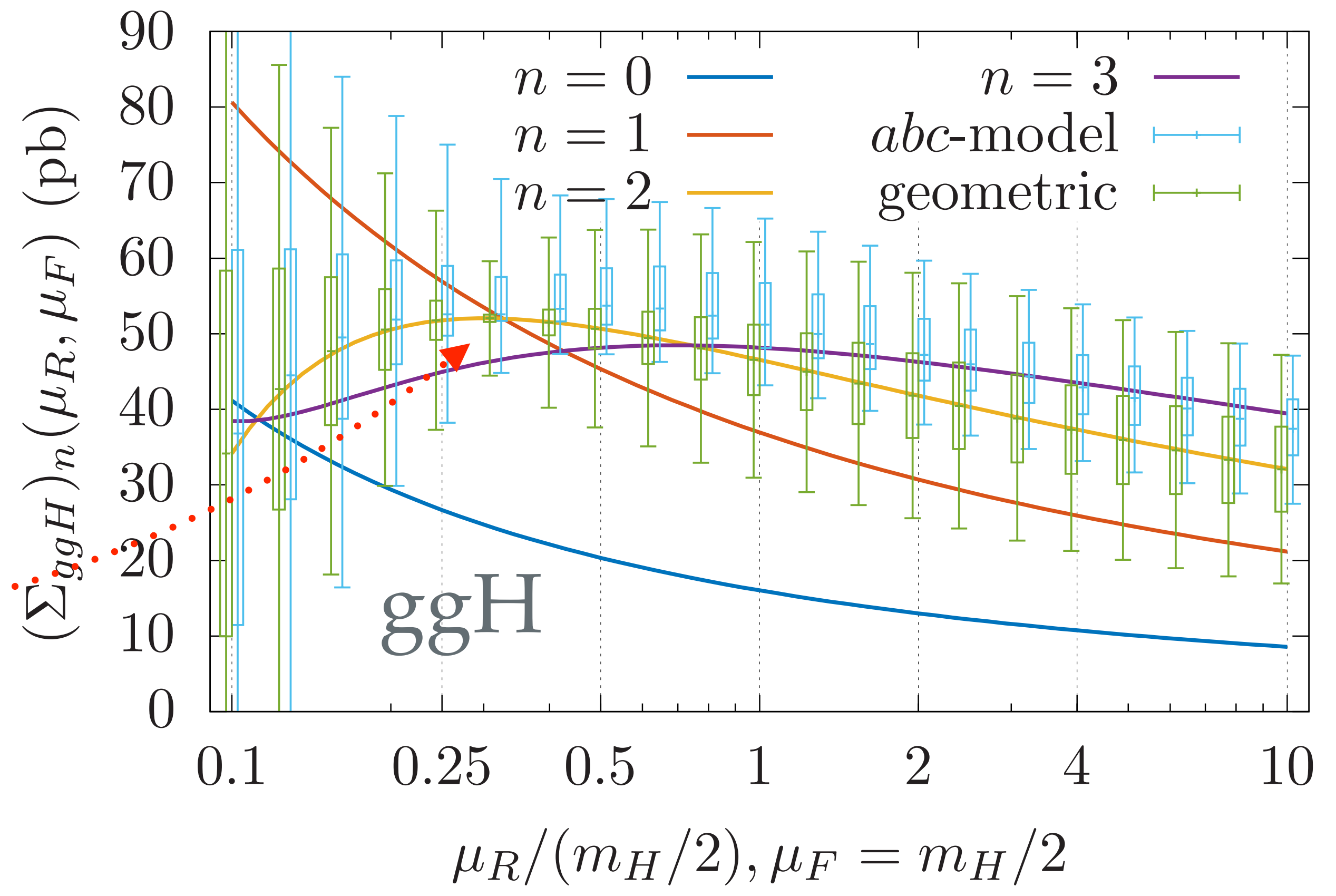
▶ $CI_{68/95}$  (geo)  (abc)

• geo

- ▶ always entered around NNLO
- ▶ very narrow peak

• abc

- ▶ $\mu/\mu_0 \gtrsim 1 \rightsquigarrow$ anticipate pos. N3LO
- ▶ $\mu/\mu_0 \lesssim 1 \rightsquigarrow$ bias slowly disappears



WHAT TO DO WITH THE THE SCALE μ ?

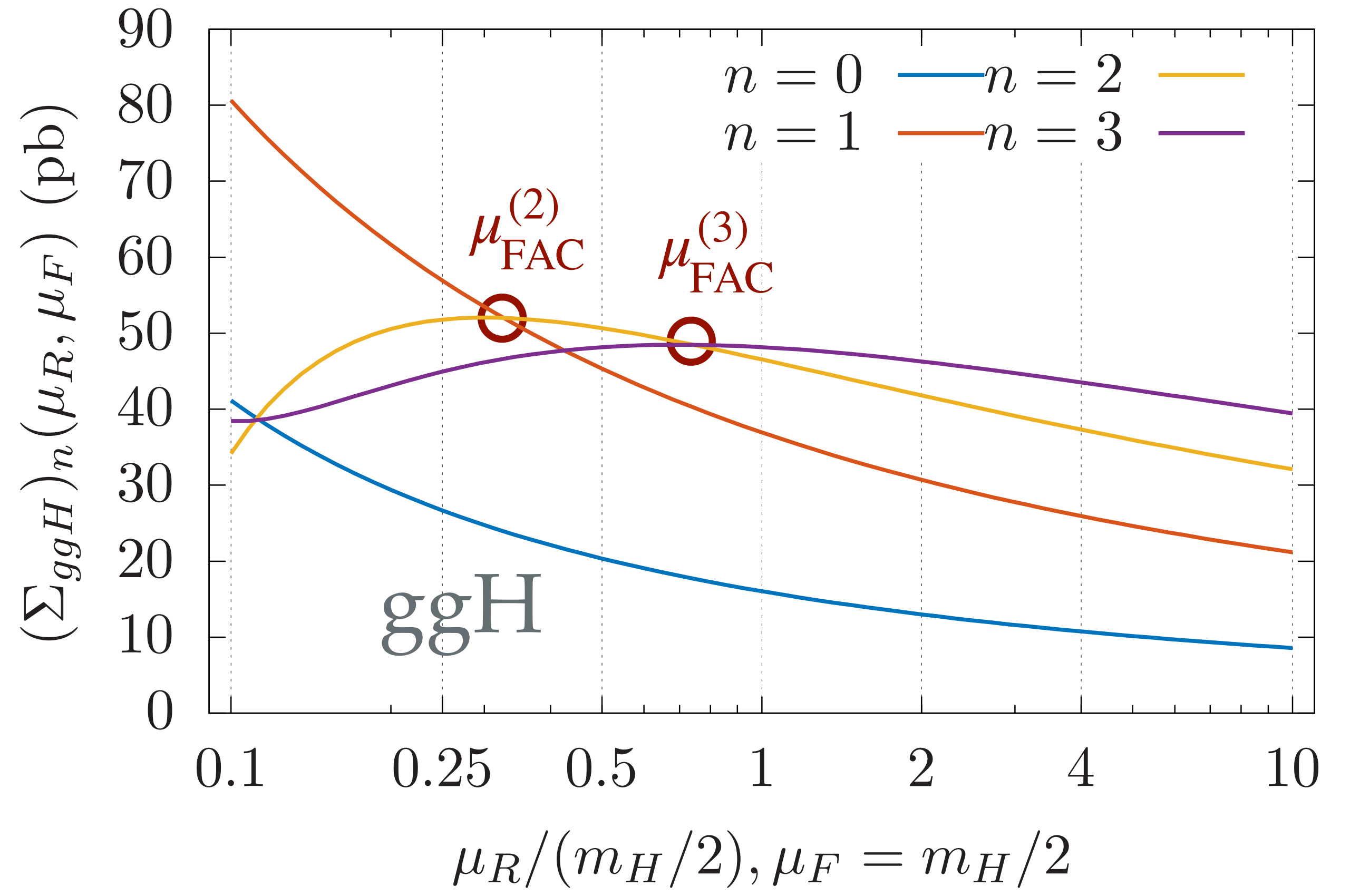
• $\forall \mu \rightsquigarrow P(\delta_3 | \delta_0, \delta_1, \delta_2; \mu)$

▶ $CI_{68/95}$  (geo)  (abc)

• two options:

1. invoke some *principle* to pick the “*optimal*” scale

▶ FAC, PMS, PMC, ...



Fastest Apparent Convergence
 $\Sigma_n(\mu_{FAC}) = \Sigma_{n-1}(\mu_{FAC})$

depends on order
 might not be unique

WHAT TO DO WITH THE THE SCALE μ ?

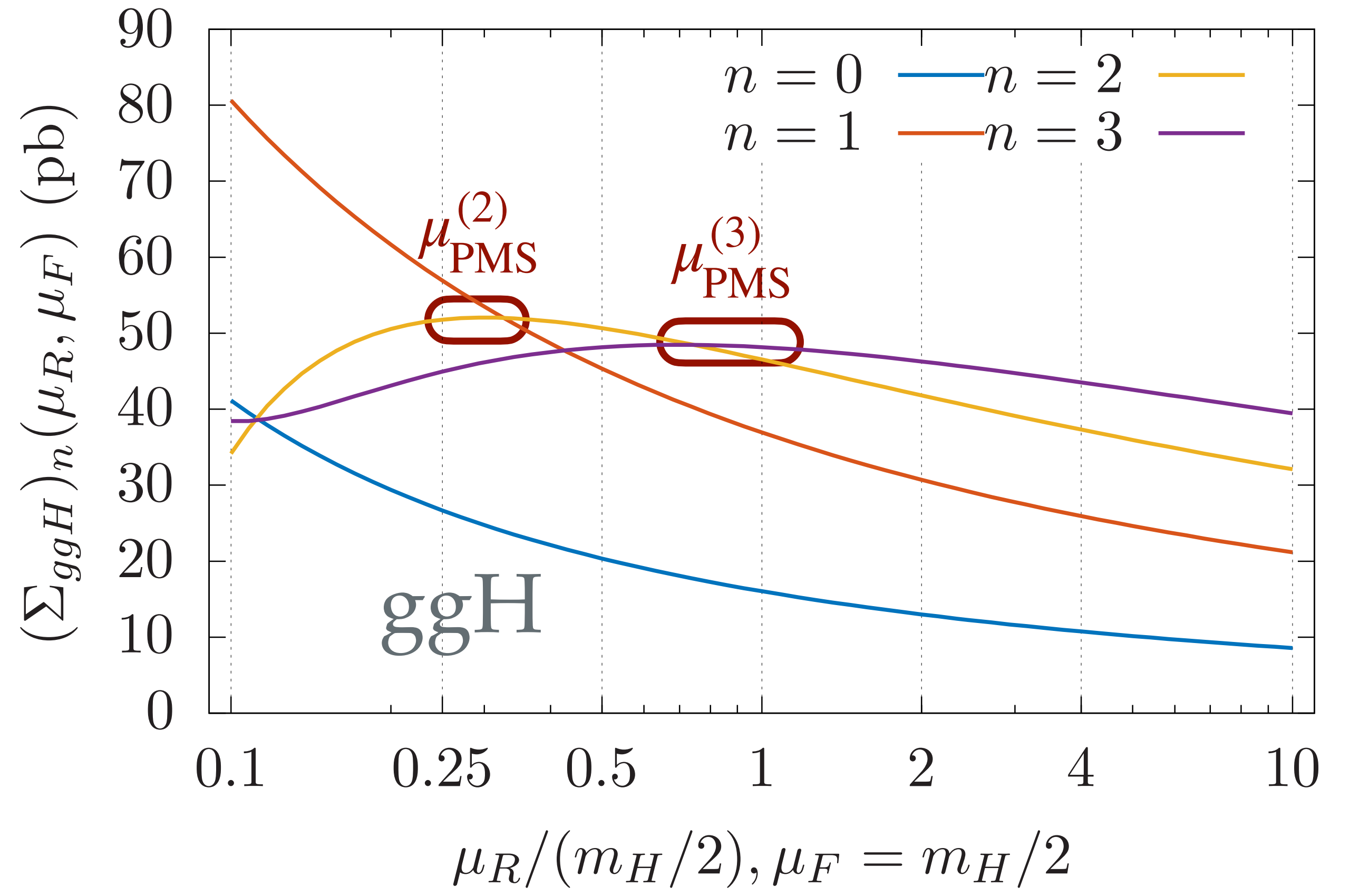
• $\forall \mu \rightsquigarrow P(\delta_3 | \delta_0, \delta_1, \delta_2; \mu)$

▶ $CI_{68/95}$  (geo)  (abc)

• two options:

1. invoke some *principle* to pick the “*optimal*” scale

▶ FAC, PMS, PMC, ...



Principle of Minimal Sensitivity
 $\frac{\partial}{\partial \mu} \Sigma_n(\mu) \Big|_{\mu_{\text{PMS}}} = 0$

depends on order
 might not be unique

WHAT TO DO WITH THE THE SCALE μ ?

• $\forall \mu \rightsquigarrow P(\delta_3 | \delta_0, \delta_1, \delta_2; \mu)$

▶ $CI_{68/95}$  (geo)  (abc)

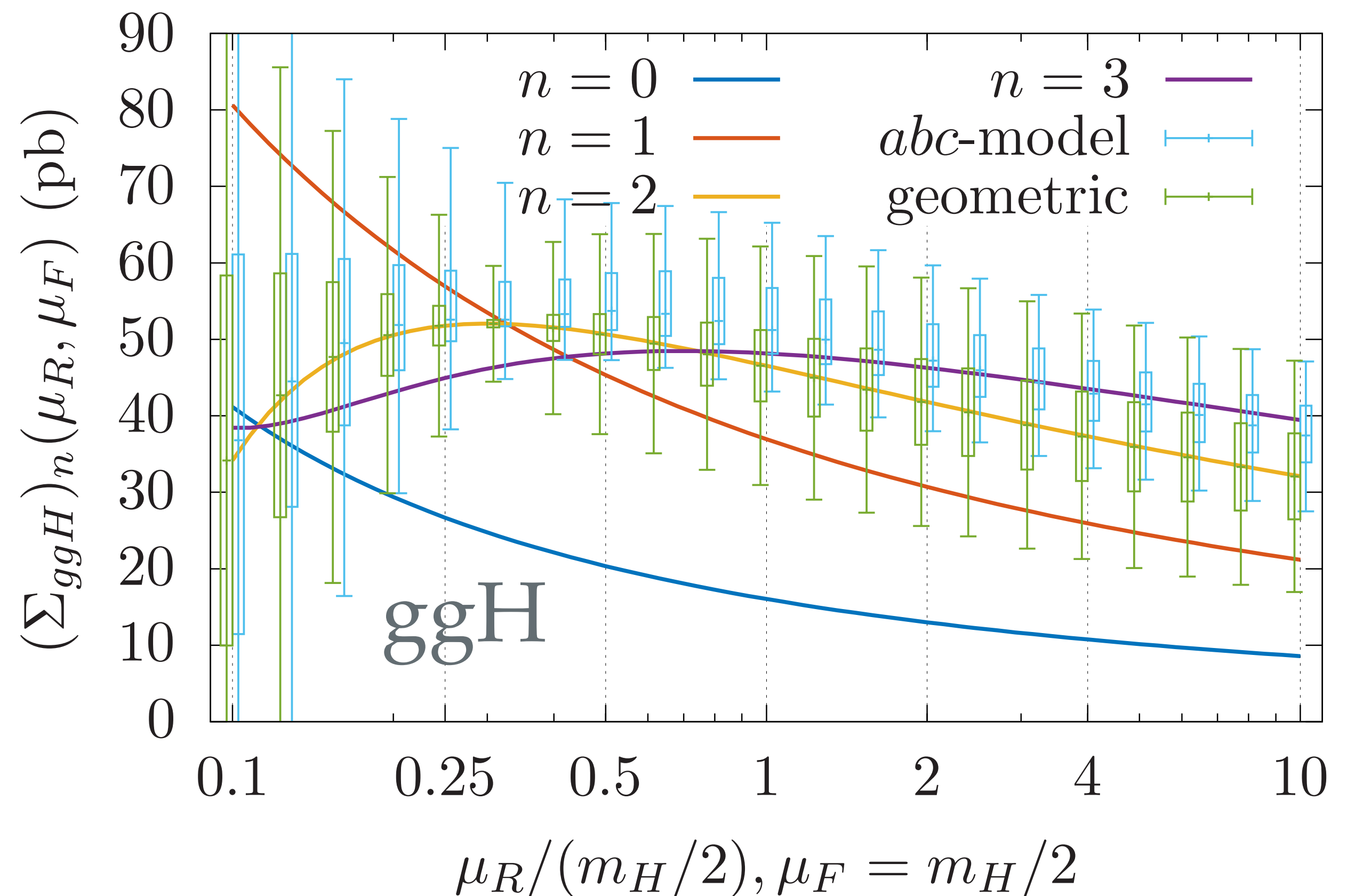
• two options:

1. invoke some *principle* to pick the “*optimal*” scale

▶ FAC, PMS, PMC, ...

2. combine different $P(\delta_{n+1} | \delta_n; \mu)$

pursued in the following



PRESCRIPTIONS FOR SCALES

Scale Marginalisation (sm):

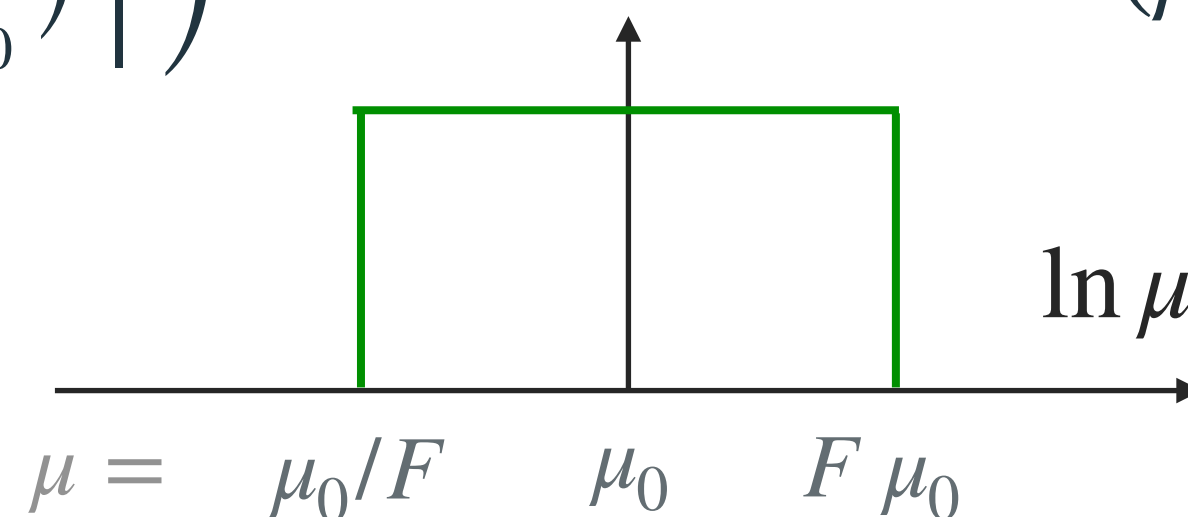
[Bonvini '20]

- treat μ as a hidden model parameter & marginalise over it:

$$\begin{aligned} P_{\text{sm}}(\delta_{n+1} | \delta_n) &= \int d\mu P(\delta_{n+1}, \mu | \delta_n) \\ &= \int d\mu P(\delta_{n+1} | \delta_n; \mu) P(\mu | \delta_n) \end{aligned}$$

- $P(\mu | \delta_n) \propto P(\delta_n; \mu) P_0(\mu)$ with prior:

$$P_0(\mu) = \frac{1}{2\mu \ln F} \Theta\left(\ln F - \left|\ln\left(\frac{\mu}{\mu_0}\right)\right|\right)$$



Scale Average (sa):

[Duhr, AH, Mazeliauskas, Szafron '21]

- μ has no probabilistic interpretation \rightsquigarrow average over it:

$$P_{\text{sa}}(\delta_{n+1} | \delta_n) = \int d\mu w(\mu) P(\delta_{n+1} | \delta_n; \mu)$$

- weight function:

$$w(\mu) = \frac{1}{2\mu \ln F} \Theta\left(\ln F - \left|\ln\left(\frac{\mu}{\mu_0}\right)\right|\right)$$

PEAK OF THE DISTRIBUTIONS*

[Duhr, AH, Mazeliauskas, Szafron '21]

Scale Marginalisation (sm):

- if $\mu_{\text{FAC}} \in [\mu_0/F, F\mu_0]$ then $P_{\text{sm}}(\delta_{n+1} | \delta_n)$ peaks at $\Sigma_n(\mu_{\text{FAC}})$
 - ▶ $P(\delta_n | \mu)$ dominated by ($k = n$) term
 - ▶ symmetric model
 - ↔ $\delta_n(\mu) = 0$ enhanced

Scale Average (sa):

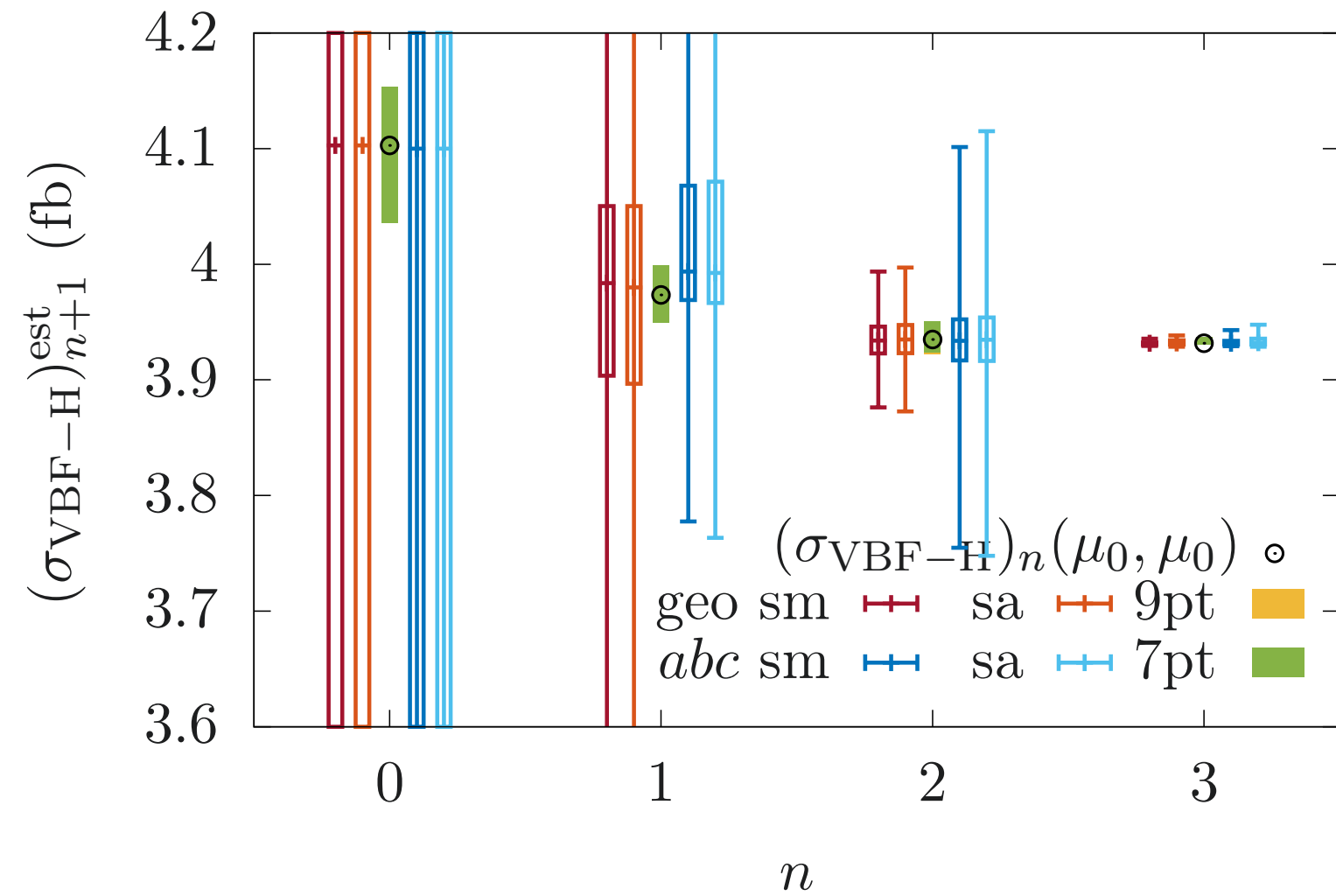
- if $\mu_{\text{PMS}} \in [\mu_0/F, F\mu_0]$ then $P_{\text{sa}}(\delta_{n+1} | \delta_n)$ peaks at $\Sigma_n(\mu_{\text{PMS}})$
 - ▶ overlap between $P(\delta_{n+1} | \delta_n; \mu)$ enhanced at stationary point
 - ↔ $\Sigma'_n(\mu_{\text{PMS}}) \approx 0$

Choice of how to interpret the scale has consequences for predictions!

* for symmetric models, a convergent series, and reasonable assumptions

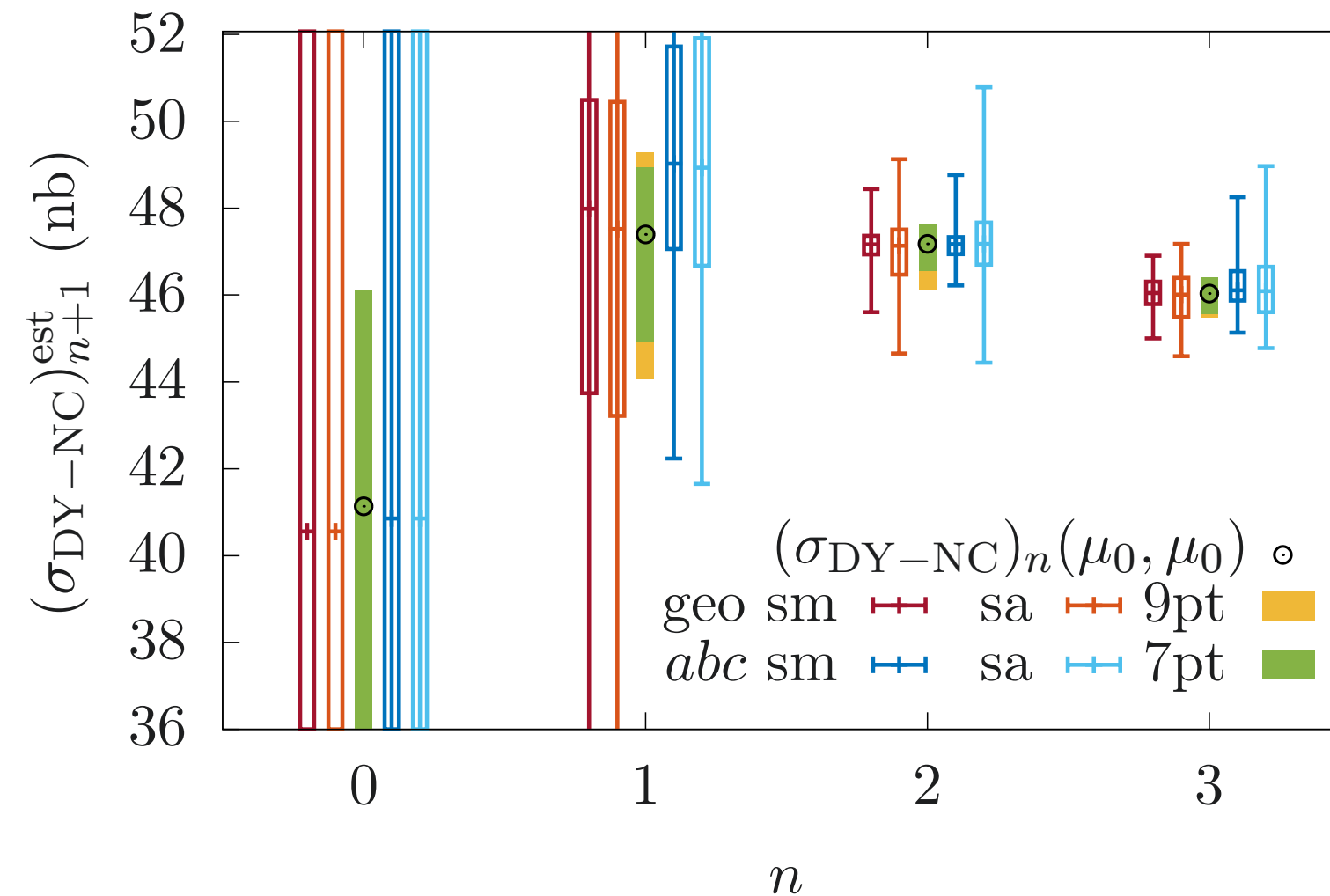
INCLUSIVE CROSS SECTIONS UP TO N³LO

VBF-H



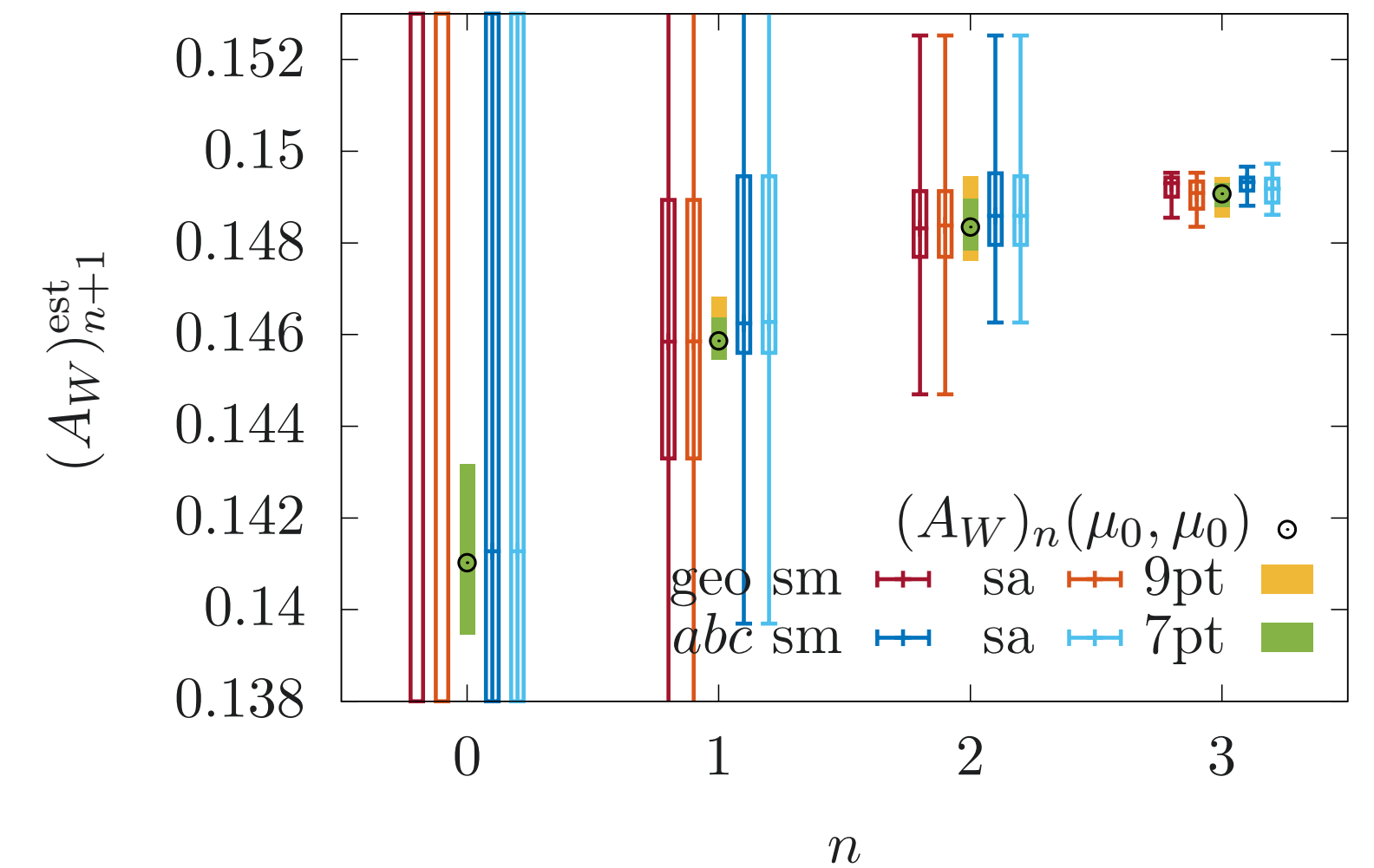
- $n < 2$: CI_{68} bigger than 9pt
- $\delta_1 < 0 \rightsquigarrow abc$ alternating
- $n > 2$: all prescriptions similar

DY-NC



- δ_3 is large and outside of 9pt!
- similar unc.: $sa \simeq 9pt$
- $n = 2$: $sm \ll$ others (μ_{FAC})
- $n = 3$: all prescriptions similar

$$A_W = \frac{W^+ - W^-}{W^+ + W^-}$$



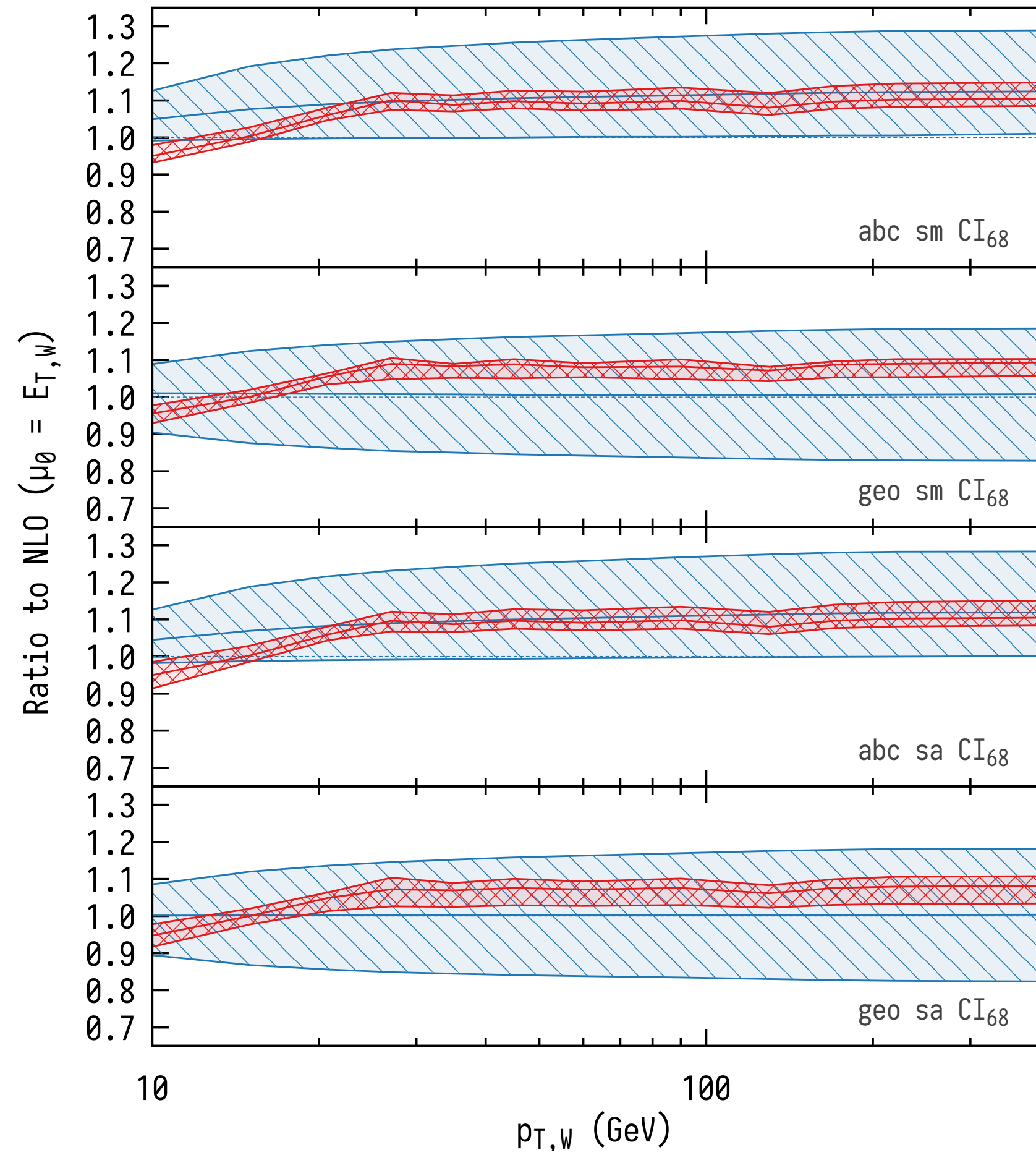
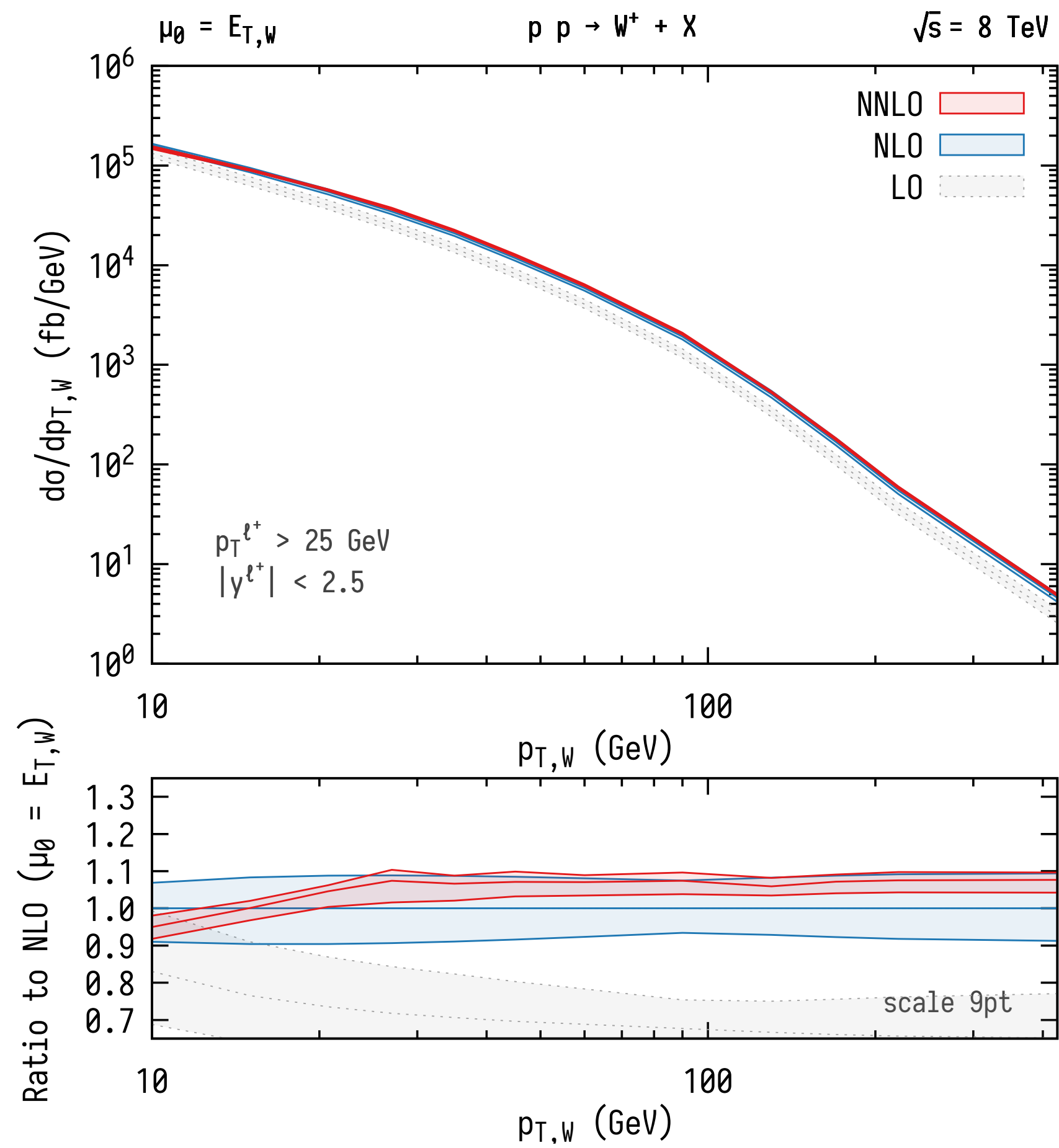
- large cancellations in the ratio
- $n < 2$: 9pt performs poorly
- $(A_W)_n \nearrow$ (anticipated by abc)
- size: $abc \lesssim$ others

overall: not radically different estimates for Δ_{MHO} ($n \geq 2$)

DIFFERENTIAL DISTRIBUTIONS

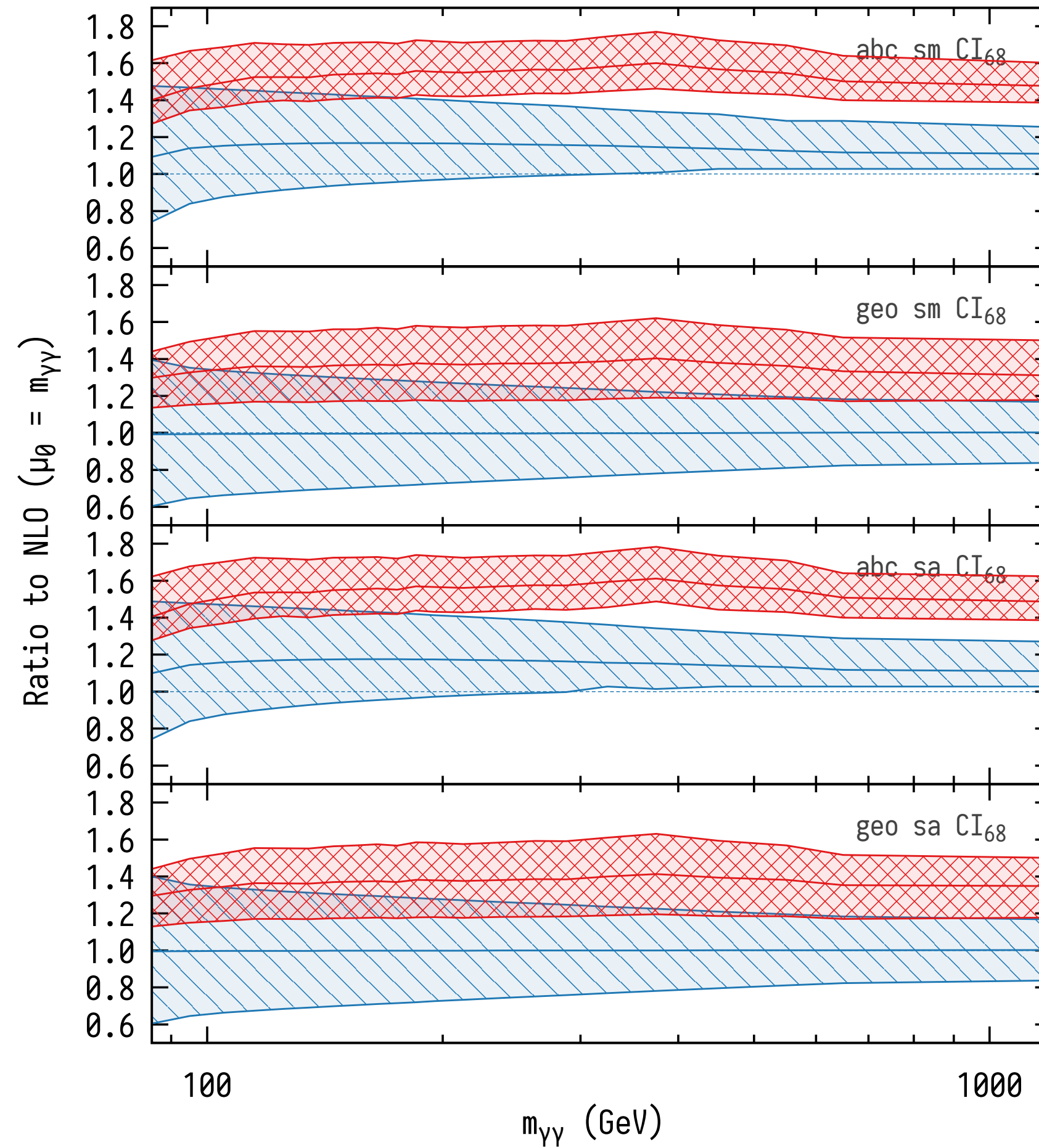
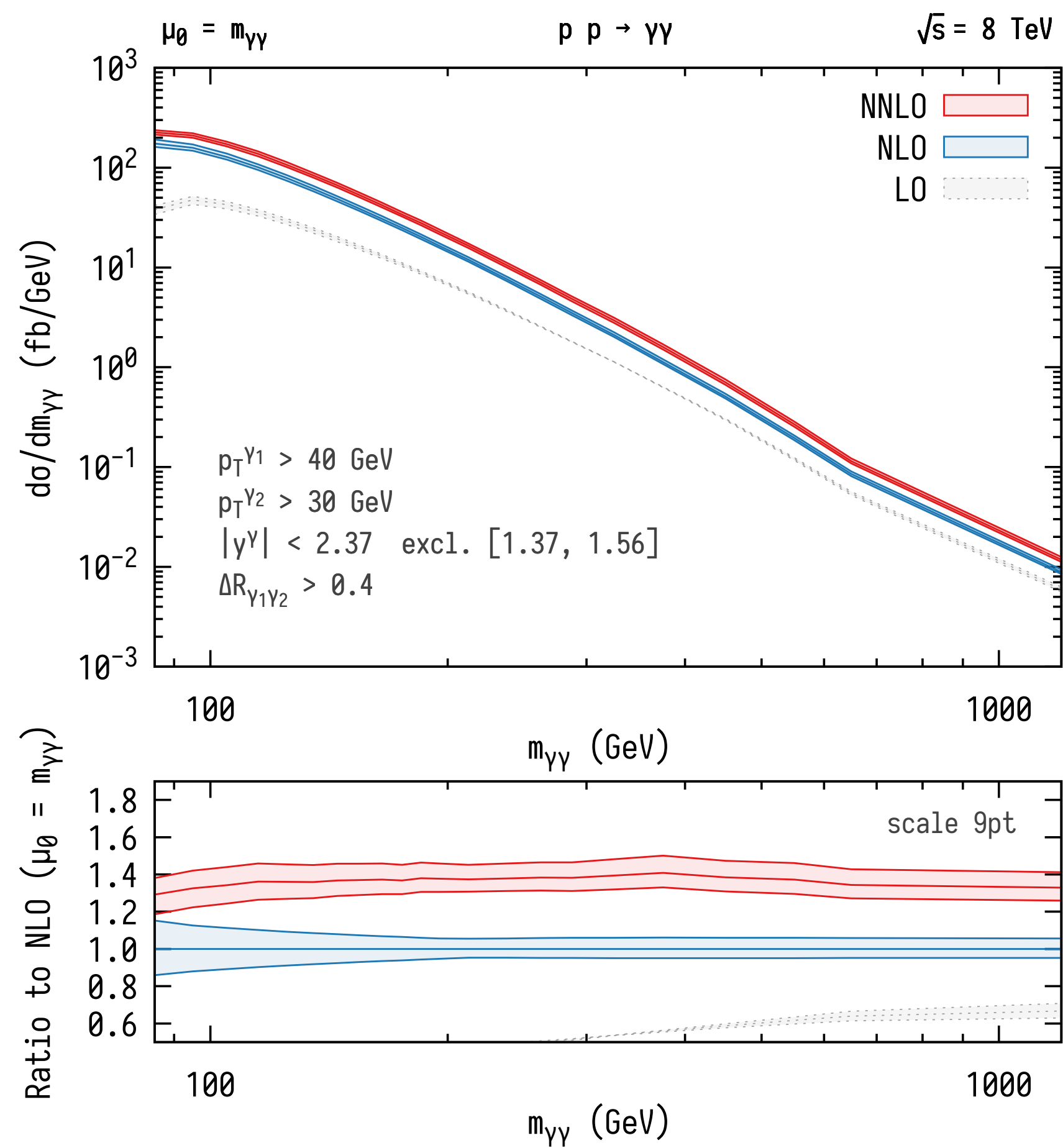
- Bayesian approach also applicable to distributions
 - ↔ treat each bin individually ↔ **will not include correlations!**
- new challenges
 - ▶ no longer “easy” to identify an appropriate hard scale μ_0 (up to rescaling)
 - ↔ inclusive ggH: M_H vs. $\frac{1}{2} M_H$? Just let the model figure it out.
 - ▶ differential distributions can probe different kinematic regimes
 - ↔ dynamical scale choice ↔ *many choices!*
 - ↔ e.g. in jet production: p_T^j , $p_T^{j_1}$, $\langle p_T^j \rangle_{\text{avg}}$, $H_T \equiv \sum_{i \in \text{jets}} p_T^i$, $\hat{H}_T \equiv \sum_{i \in \text{partons}} p_T^i$, ...
 - ▶ re-cycling via quadrature limited ↔ **ideally interpolation grids**

W-BOSON + JET PRODUCTION



- $n < 2$:
 - ▶ CI₆₈ bigger than 9pt
 - ▶ *abc* captures pos. shift
- $n = 2$:
 - ▶ almost identical bands
 - ▶ Δ_{MHO} very robust
- sm vs. sa
 - ▶ almost identical CI

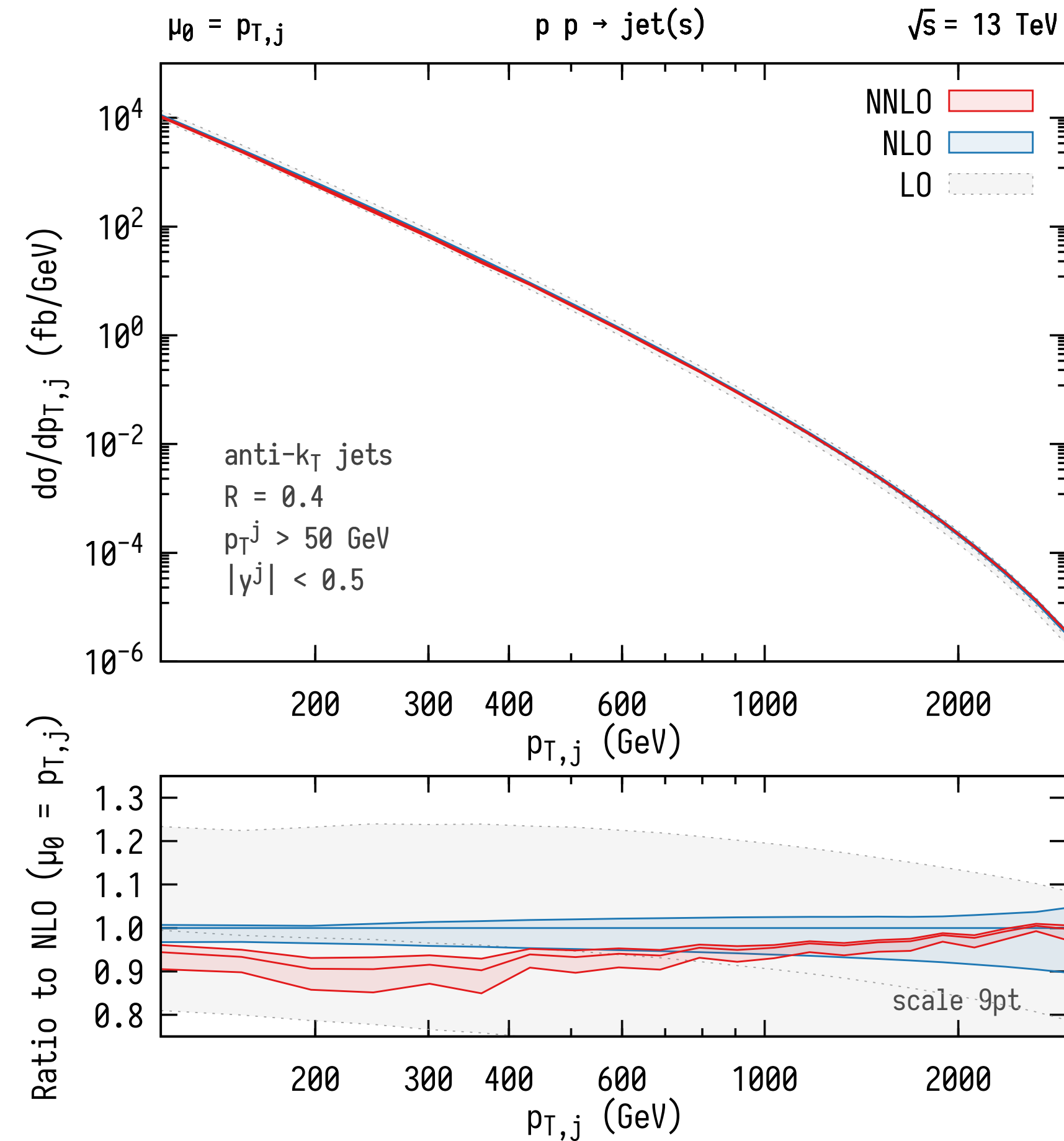
DI-PHOTON PRODUCTION



- example where 9pt fails
 - large corrections
 - $\Delta_{MHO}^{NNLO} \gtrsim \Delta_{MHO}^{NLO}$
 - no sign of convergence
- $n < 2$:
 - $CI_{68} \sim 2-3 \times 9pt$
- $n = 2$:
 - marginal overlap for geo
 - differences in *size & position*
 - ideally N3LO for robust Δ_{MHO}
- $sm \simeq sa$
 - large corrections prohibit FAC points

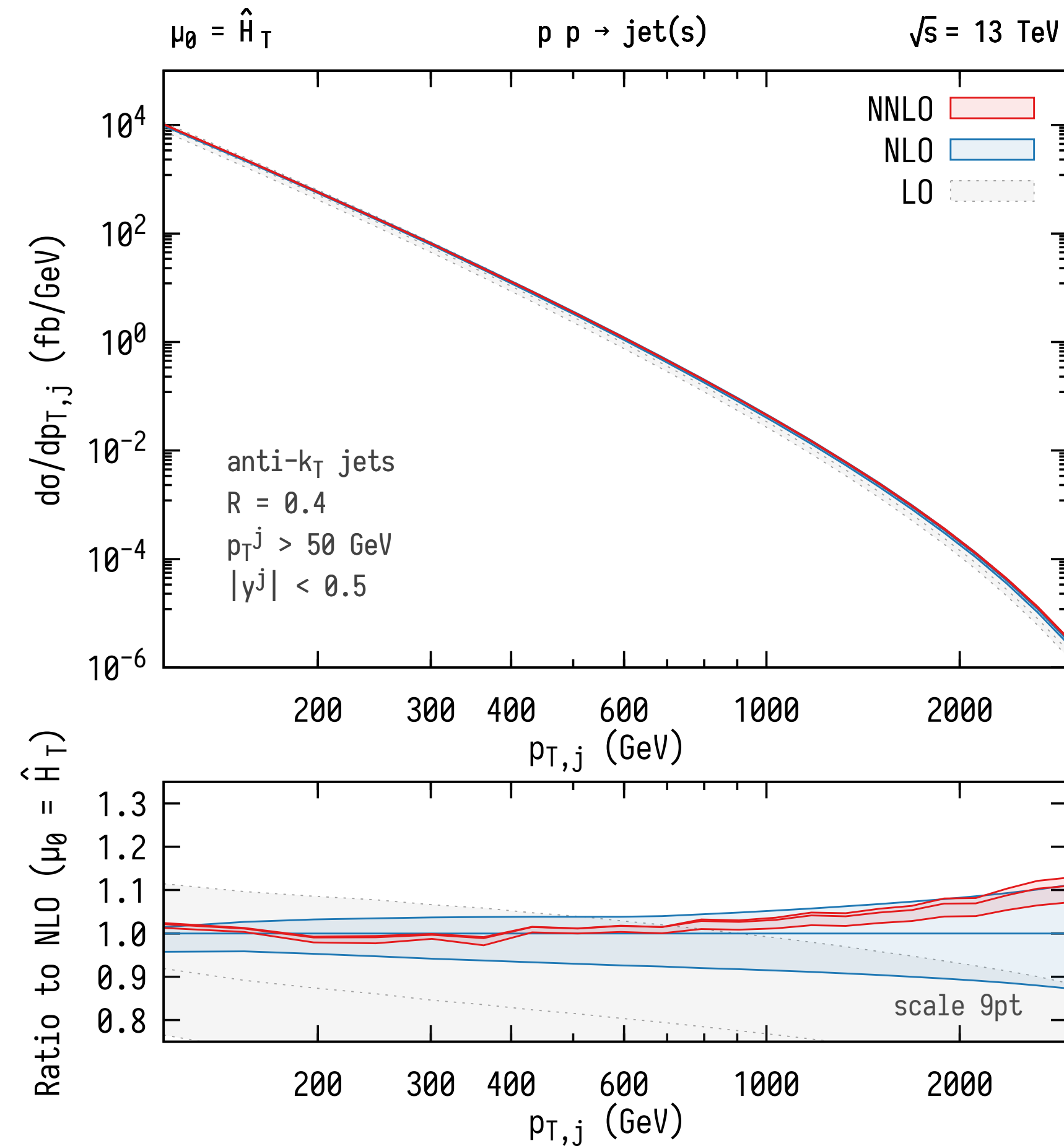
THE PROBLEM WITH JETS...

$\mu_0 = p_T^j$: infrared sensitivity



instability
at low p_T

$\mu_0 = \hat{H}_T$: recommendation [Currie et al. '18]

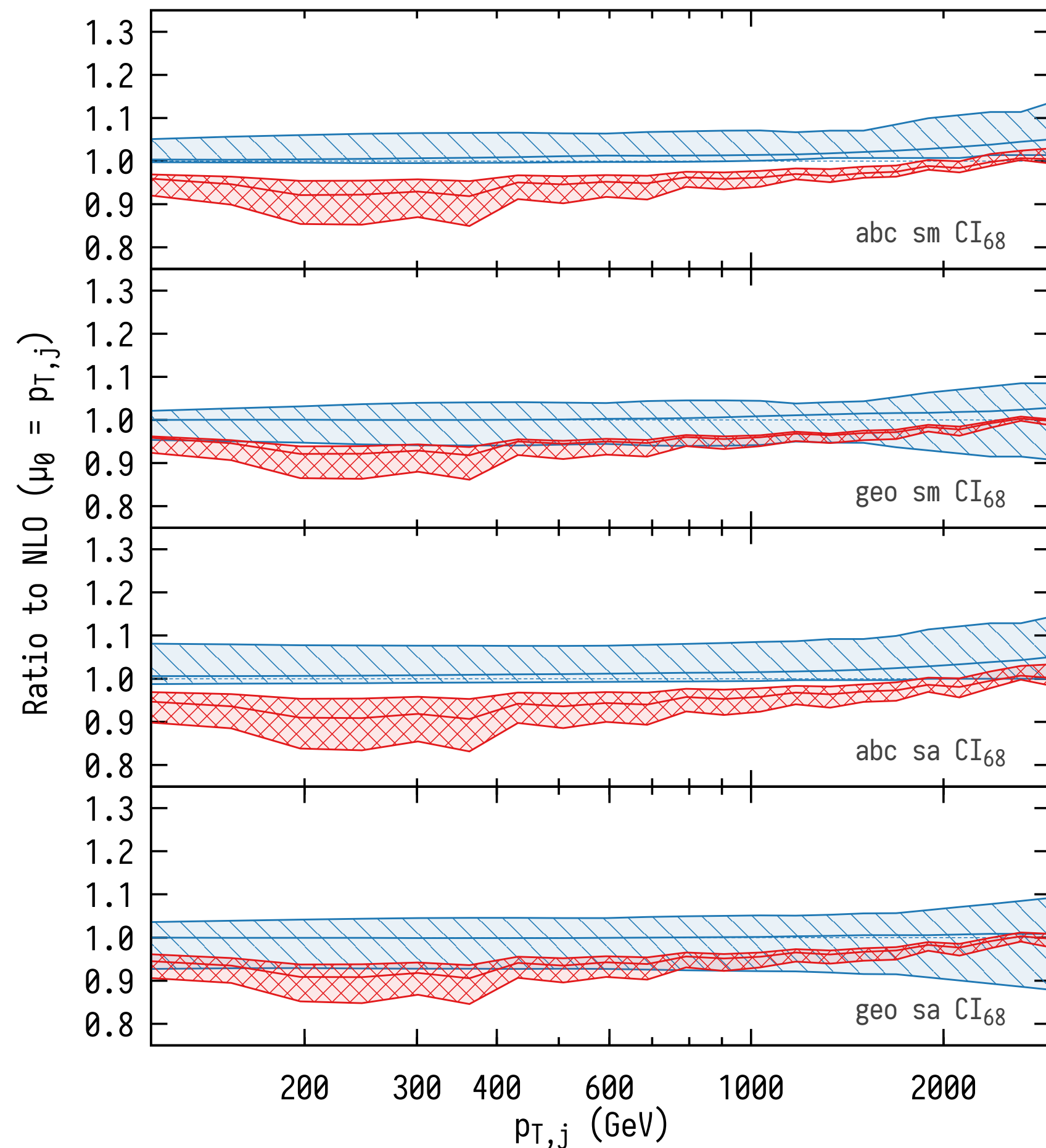


THE PROBLEM WITH JETS... PERSISTS

non-trivial change of dynamical scales
cannot be captured by a simple re-scaling

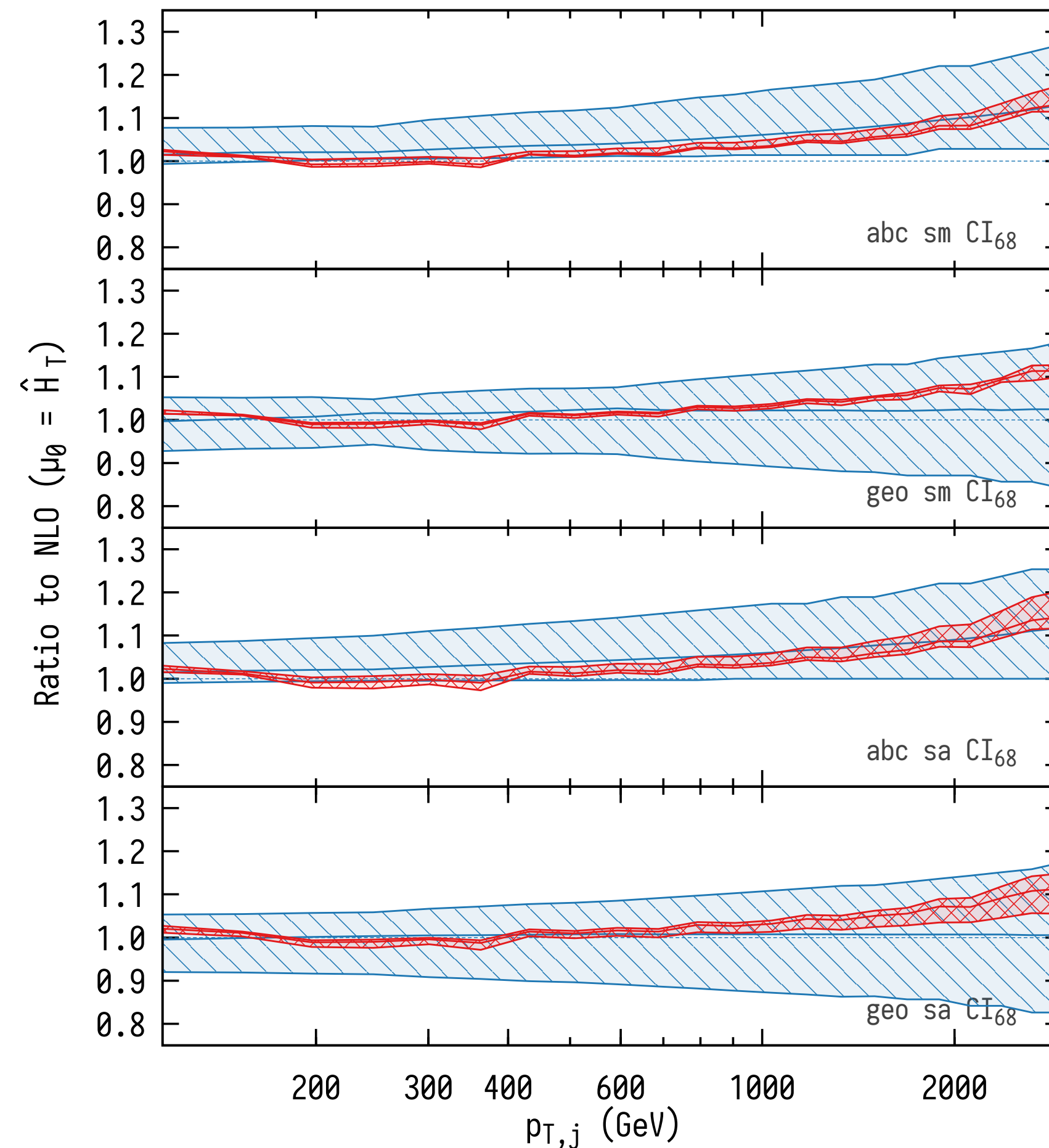
$\mu_0 = p_T^j$: infrared sensitivity

$\mu_0 = \hat{H}_T$: recommendation [Currie et al. '18]



larger NLO
 Δ_{MHO}

barely any
difference
at NNLO



abc captures
positive
corrections

WORK IN PROGRESS — CORRELATIONS

[AH, Mazeliauskas w.i.p]

- idea: if two bins show similar (opposite) perturbative behaviour
↪ two bins should be partially (anti-)correlated.

- we want: joint probability distribution $P(x, y)$ for two bins x & y
↪ preserve projections for compatibility:

$$P(x) = \int dy P(x, y) = \int dz P(x, z)$$

- ↪ hidden parameter $-1 < c < +1$ to smoothly implement the correlation
- possibilities: algorithmic “earth movers distance”; map $P(x)$ onto $P(y)$, ...
↪ can be done much simpler

WORK IN PROGRESS — CORRELATION MODEL IN miho

[AH, Mazeliauskas w.i.p]

- projections of multi-dim. Gaussians (+ correlation matrix) are again Gaussian
 \hookrightarrow map P_i onto Gaussians, implement correlations, map back

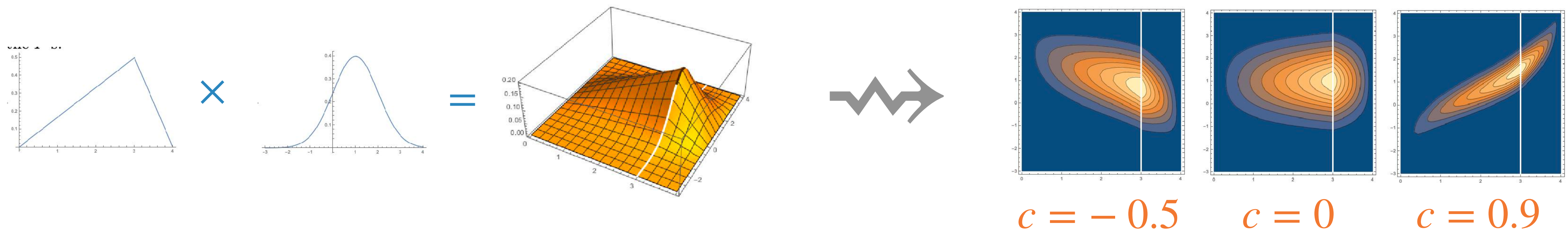
$$\begin{aligned}
 P(x, y) &= P_1(x)P_2(y) \\
 &\times \left. \frac{d\Phi^{-1}(\alpha)}{d\alpha} \right|_{\alpha=\Sigma_1(x)} \left. \frac{d\Phi^{-1}(\beta)}{d\beta} \right|_{\beta=\Sigma_2(y)} \\
 &\times \frac{1}{2\pi\sqrt{1-c^2}} \exp\left(-\frac{1}{2(1-c^2)} [\xi(x)^2 + \eta(y)^2 - c2\xi(x)\eta(y)]\right)
 \end{aligned}$$

$$\Sigma_i(x) = \int_{-\infty}^x dx' P_i(x')$$

$$\Phi^{-1}(p) = \sqrt{2}\text{Erf}^{-1}(-1 + 2p)$$

$$\xi(x) = \Phi^{-1}(\Sigma_1(x))$$

$$\eta(y) = \Phi^{-1}(\Sigma_2(y))$$



use inference to constrain c

CONCLUSIONS & OUTLOOK

- Bayesian inference is a powerful framework to estimate Δ_{MHO}
 - ▶ probabilistic interpretation $\Leftrightarrow P(\delta_{n+1} | \delta_n)$
 - ▶ exposes our *assumptions & biases* clearly \Leftrightarrow model & priors
 - ▶ but: it is not more reliable than scale variation \rightsquigarrow careful analysis required
- typically for $n < 2$: $\text{CI}_{68} > 9\text{pt}$; $n \geq 2$: $\text{CI}_{68} \simeq 9\text{pt}$
- *public code*: ミホ (miho) \rightsquigarrow <https://github.com/aykhuss/miho>
- future directions
 - ▶ correlations (p_T^W / p_T^Z , p_T^Z vs. p_T^ℓ , PDF fits & data interpretation, ...) [AH, Mazeliauskas w.i.p]
 - ▶ marginalisation over models, ...

relying on a single prescription for TH unc. in *precision measurements* that does not admit a probabilistic interpretation is potentially dangerous!

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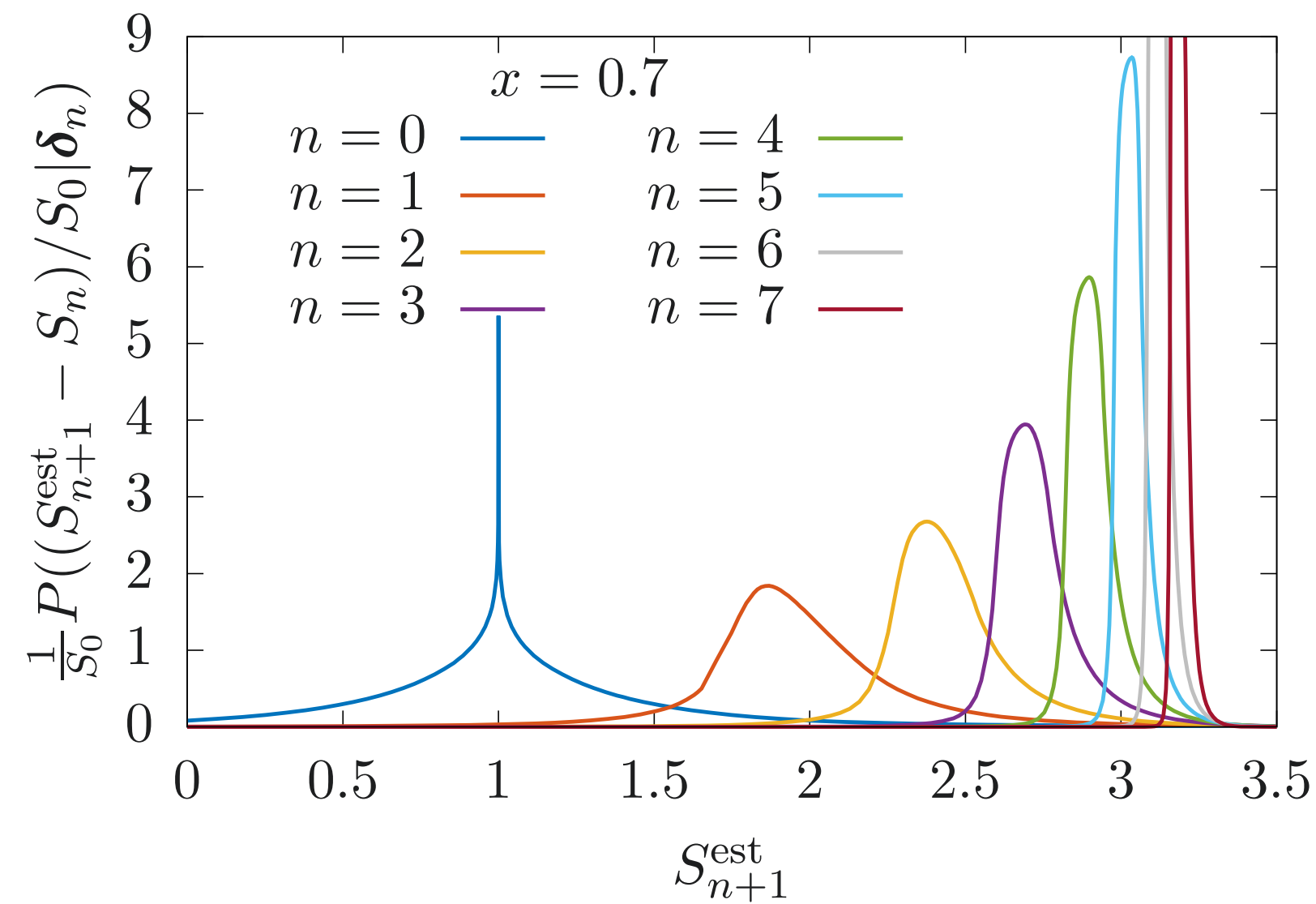
▶ marginalisation over models, ...

Thank you!

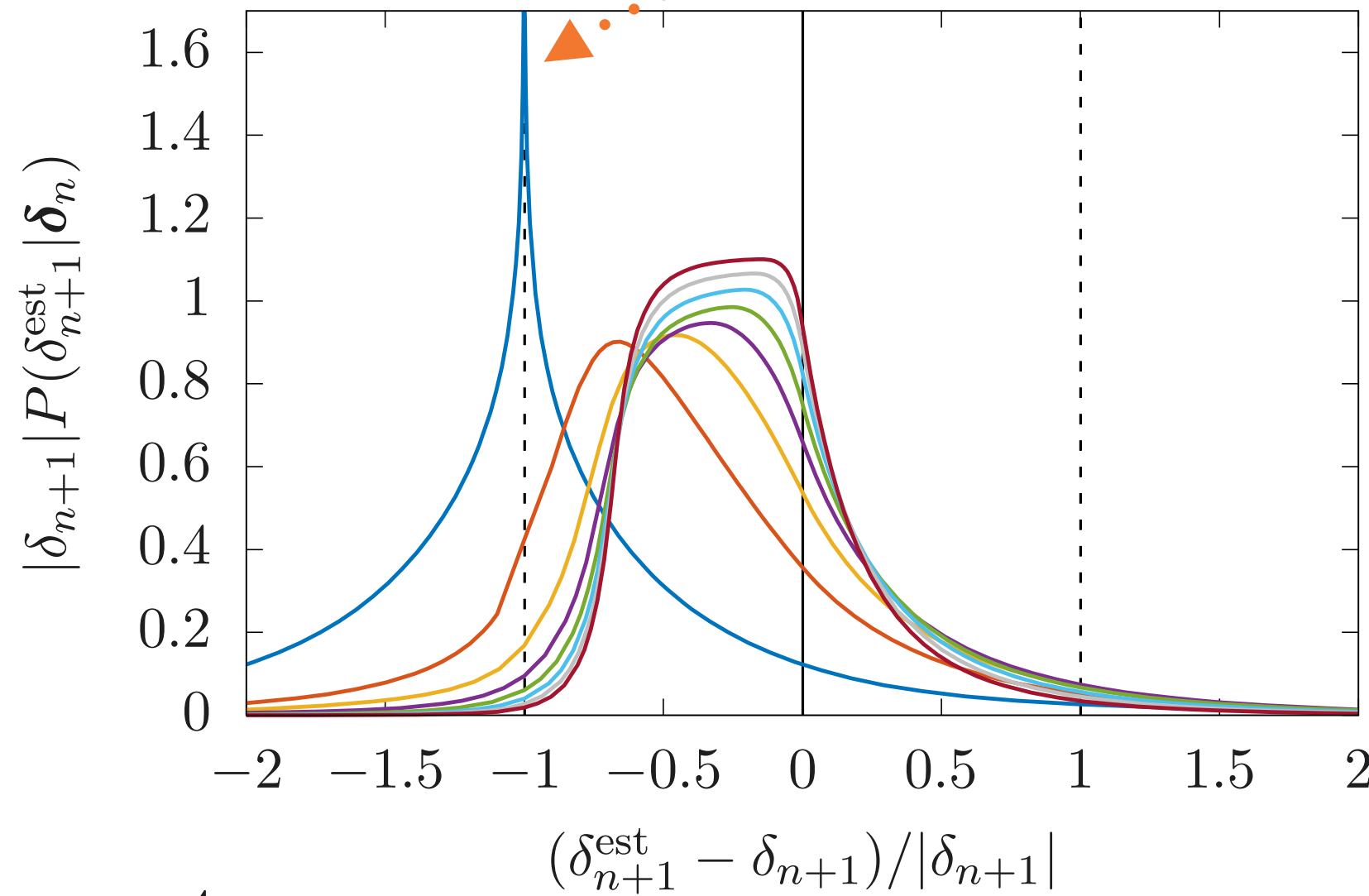
BACKUP

BACKUP

TOY EXAMPLE — $\delta_k = (0.7)^k$



..... geometric would be localised here

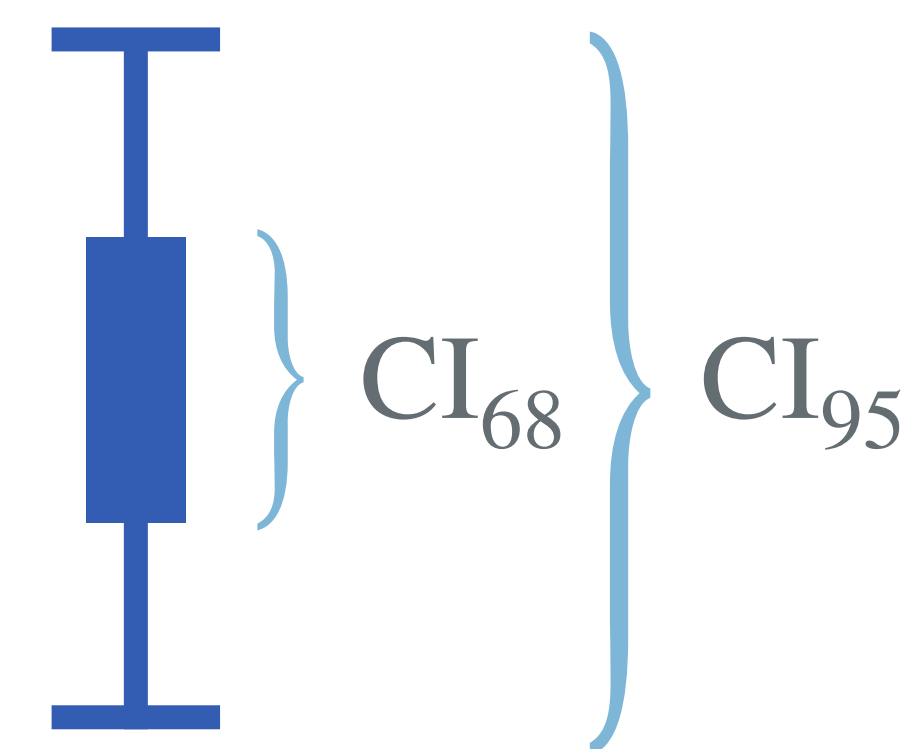
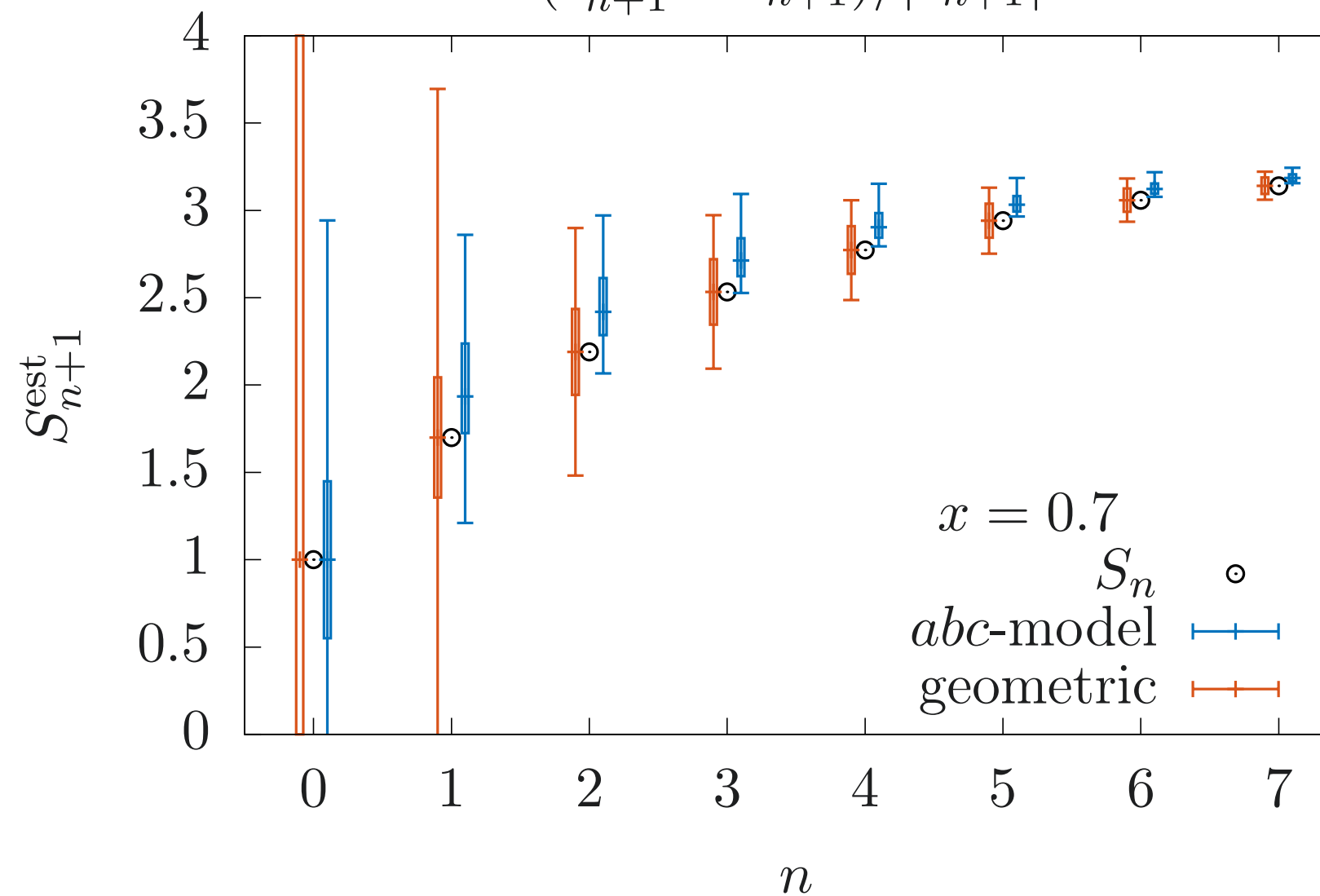


correctly anticipates positive δ_{n+1}

Confidence Intervals

$$CI_x = [\Sigma_x^{\text{low}}, \Sigma_x^{\text{upp}}]$$

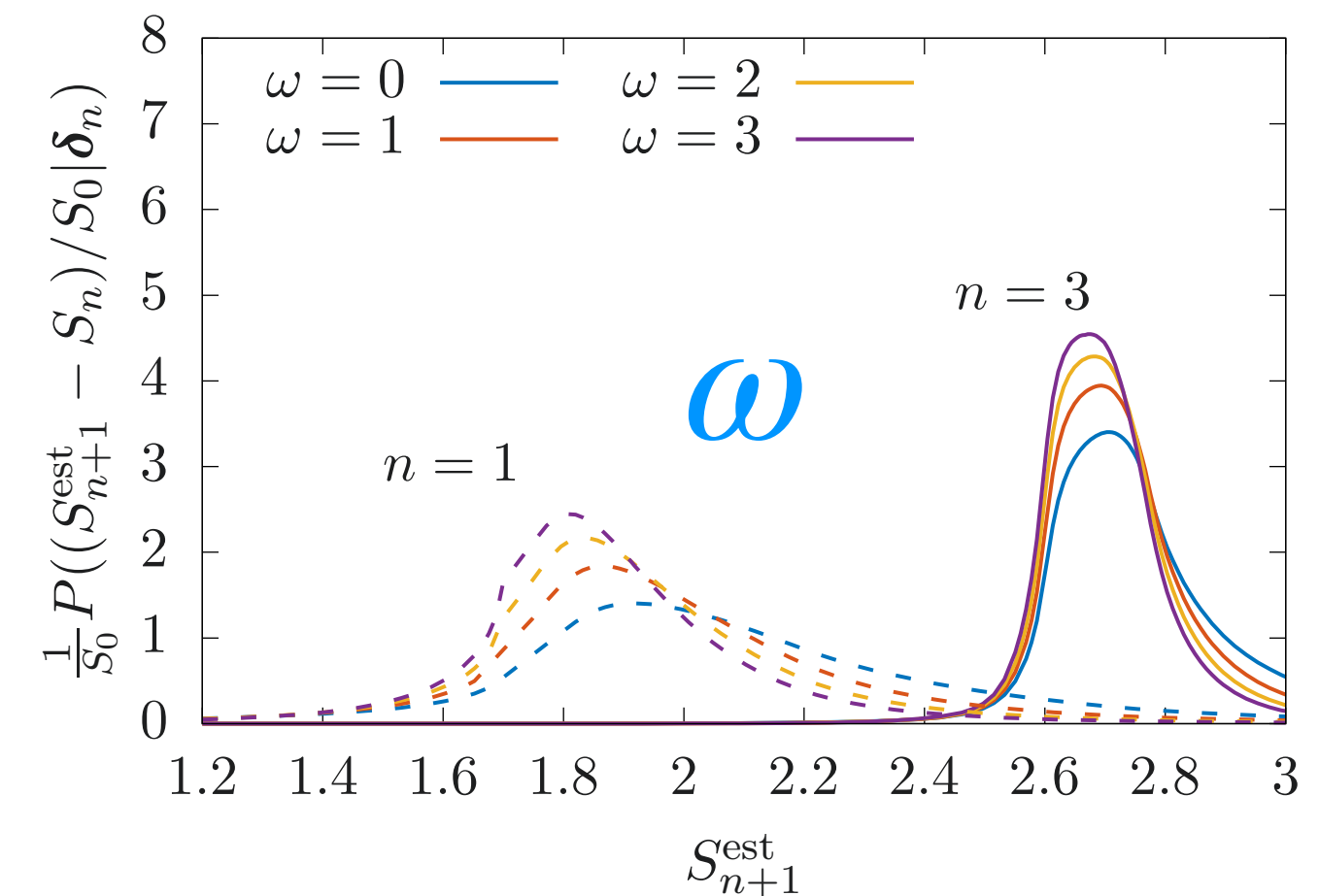
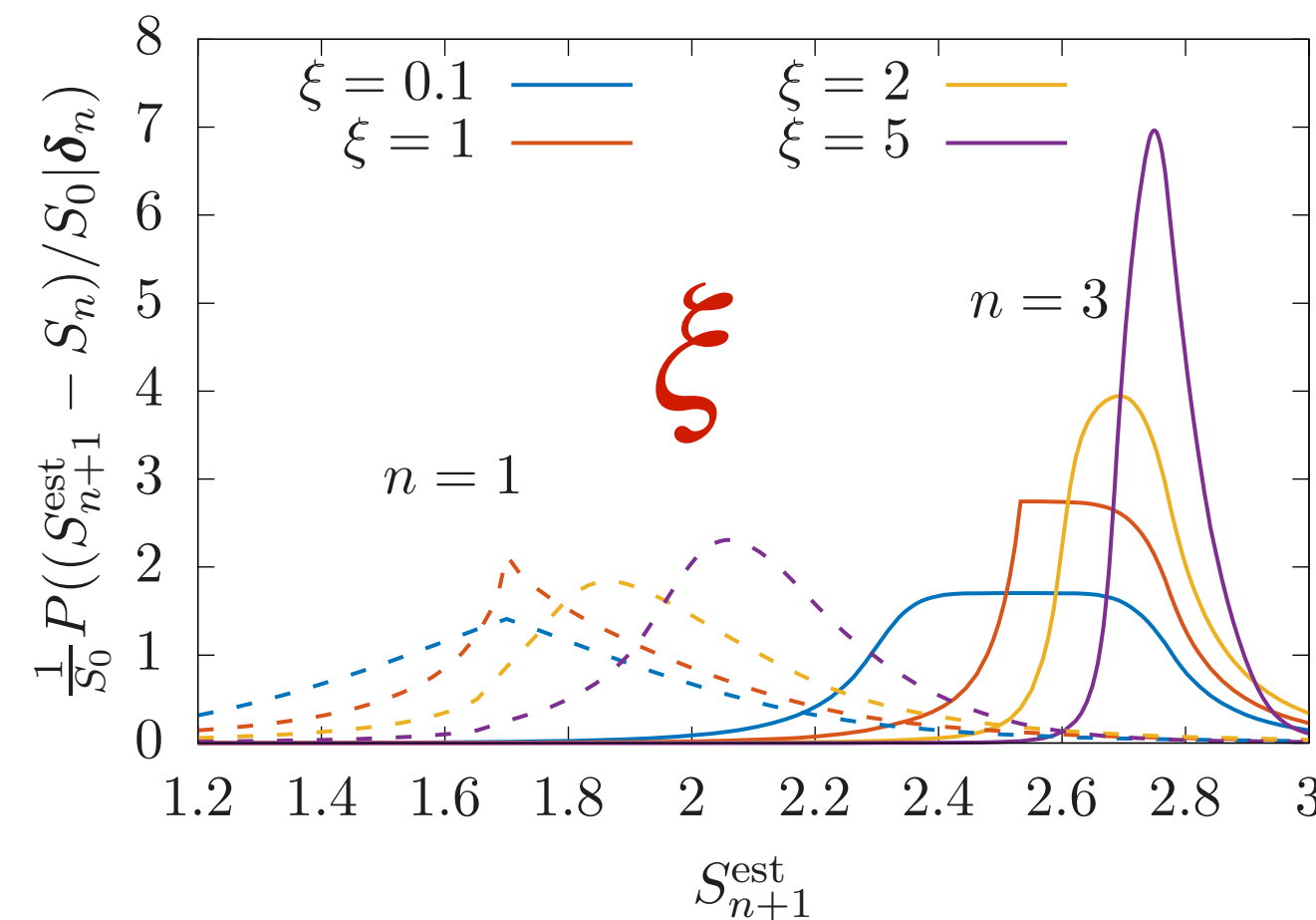
containing $x\%$ of the probability



SENSITIVITY ON THE PRIORS — $(\epsilon, \omega, \eta, \xi) \simeq (0.1, 1, 0.1, 2)$

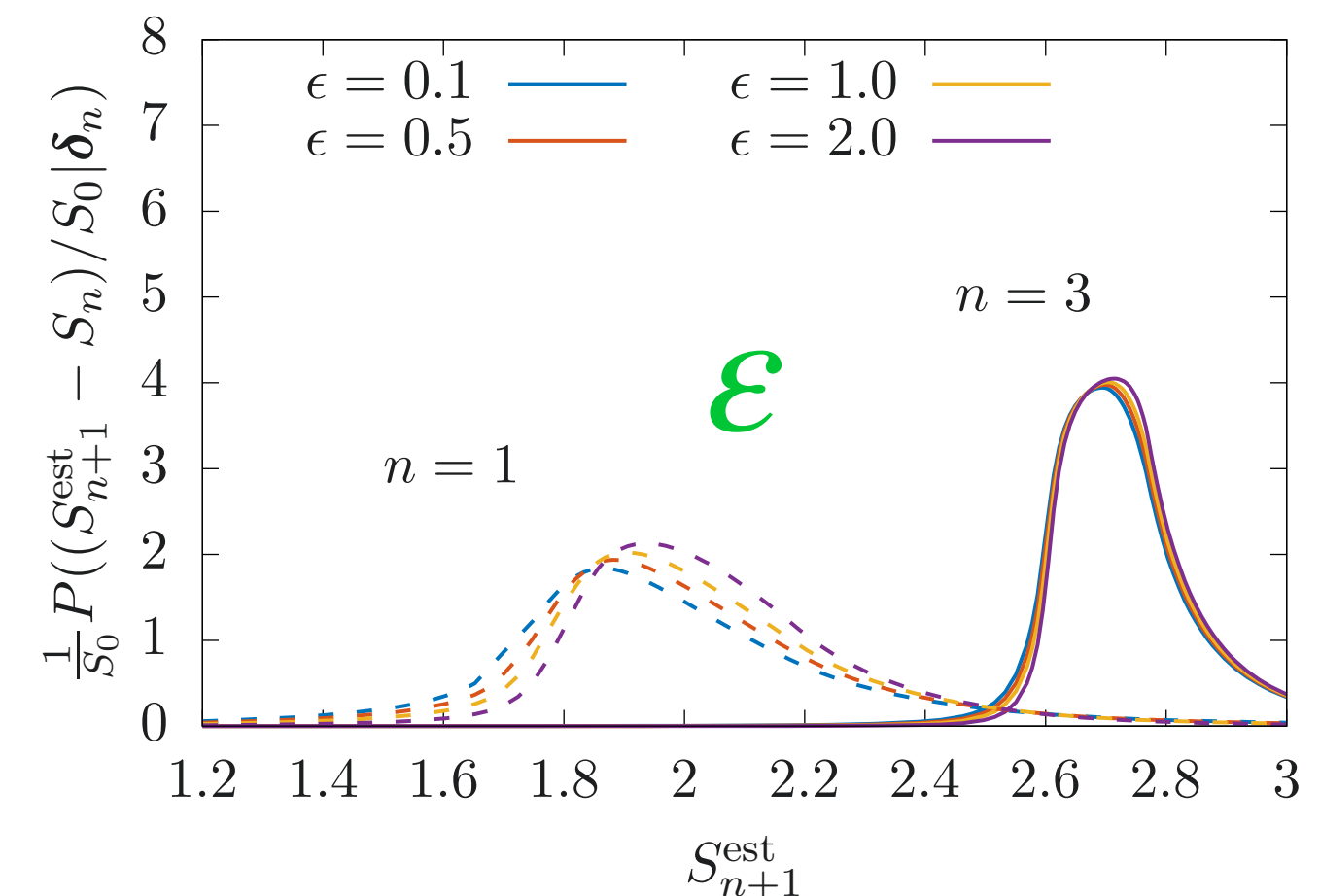
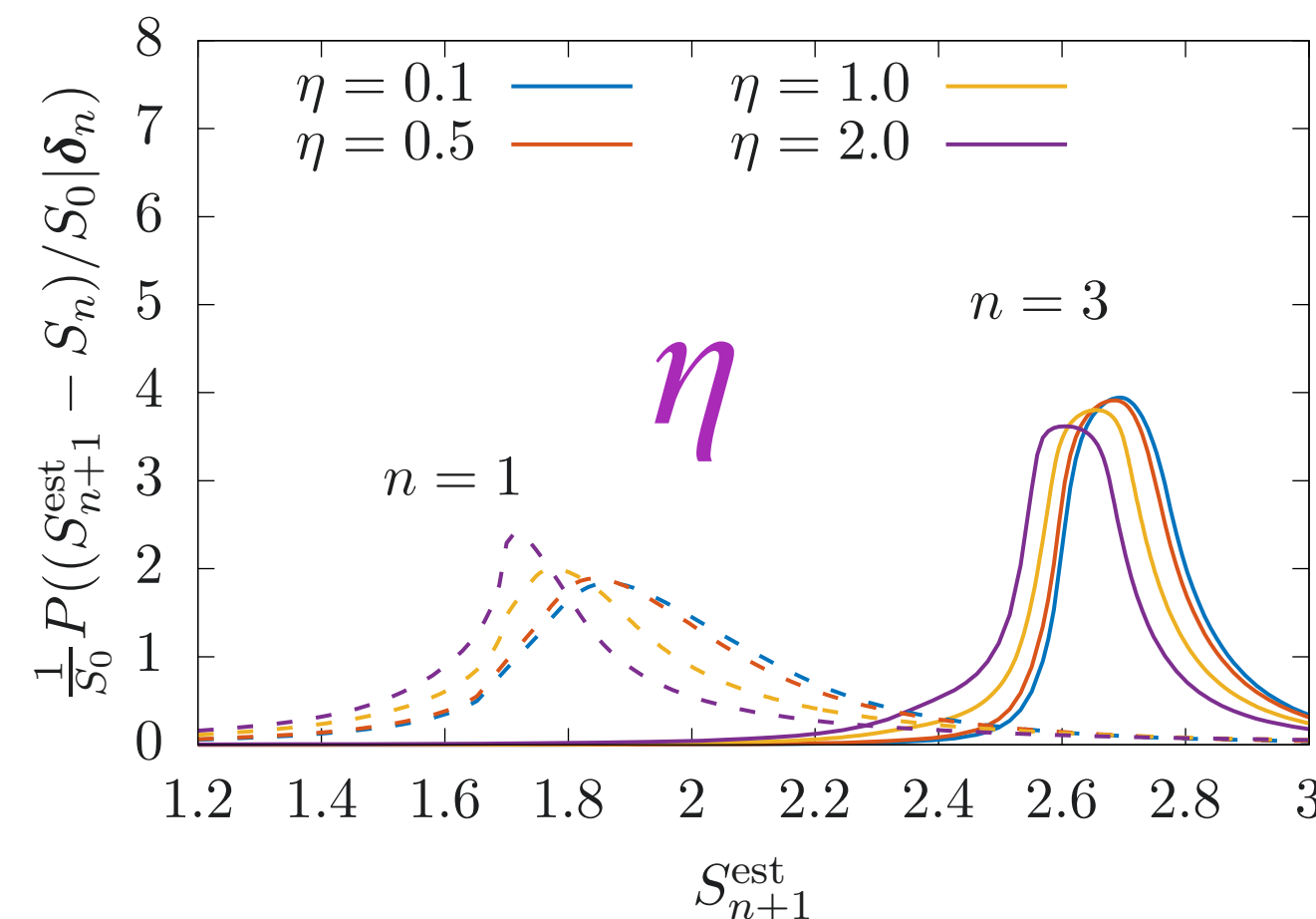
$$P_0(a) = \frac{1}{2} (1 + \omega) (1 - |a|)^\omega \Theta(1 - |a|) \quad P_0(b, c) = \frac{\epsilon \eta^\epsilon}{2 \xi c^{2+\epsilon}} \Theta(c - \eta) \Theta(\xi c - b)$$

- dependence on priors decreases as n increases
- less for ξ as it controls the asymmetry of the distribution



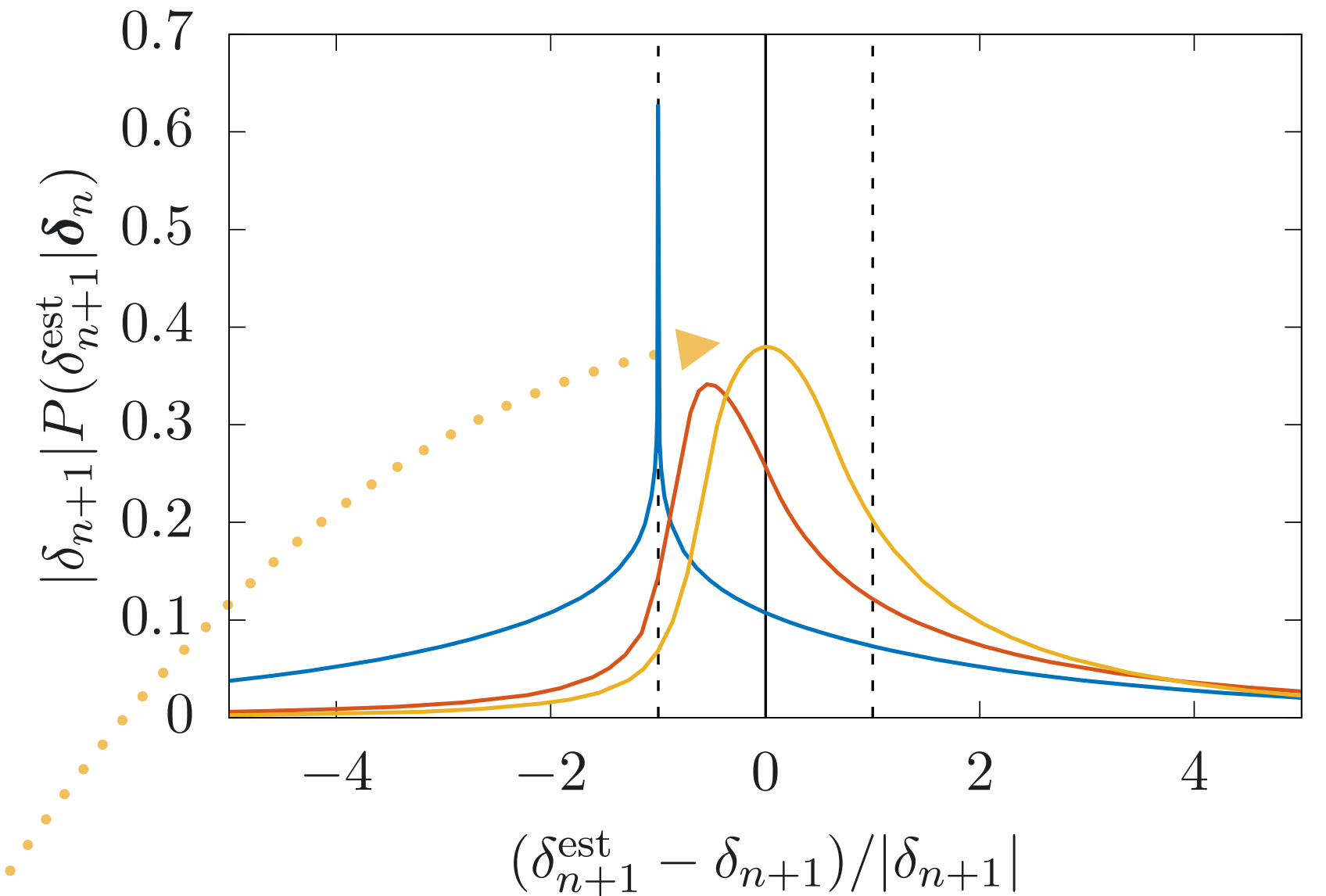
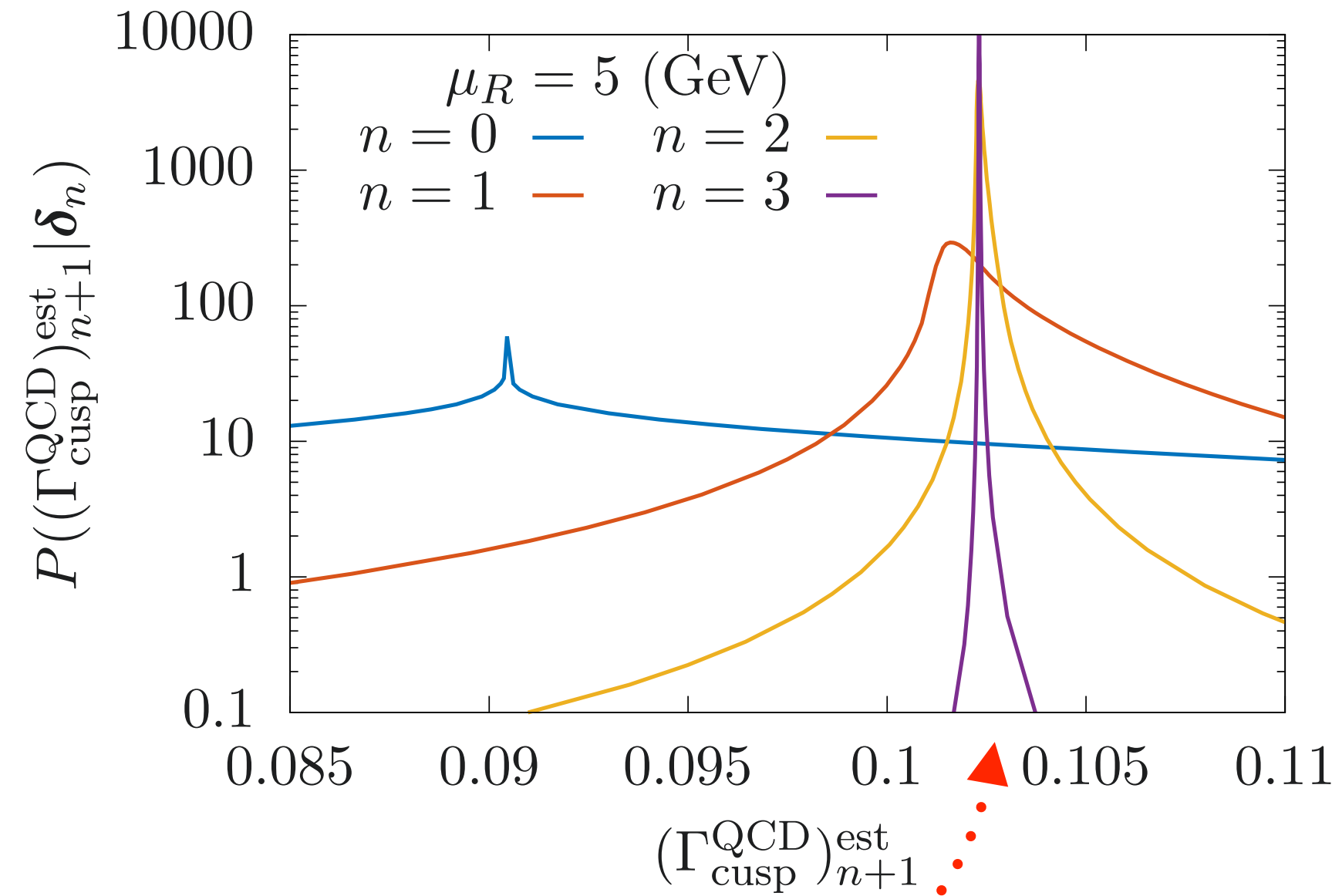
- for geometric series:

$$\left. \begin{array}{l} \eta \rightarrow 0 \\ \xi \rightarrow \infty \end{array} \right\} \rightsquigarrow S_{n+1}^{\text{est}} \rightarrow S_{n+1}$$



A REAL-WORLD EXAMPLE — CUSP ANOMALOUS DIMENSION IN QCD

$$\Gamma_{\text{cusp}}^{\text{QCD}} = C_F \sum_{k=1}^{\infty} \left(\frac{\alpha_s}{\pi}\right)^k \gamma_{\text{QCD}}^{(k)}$$



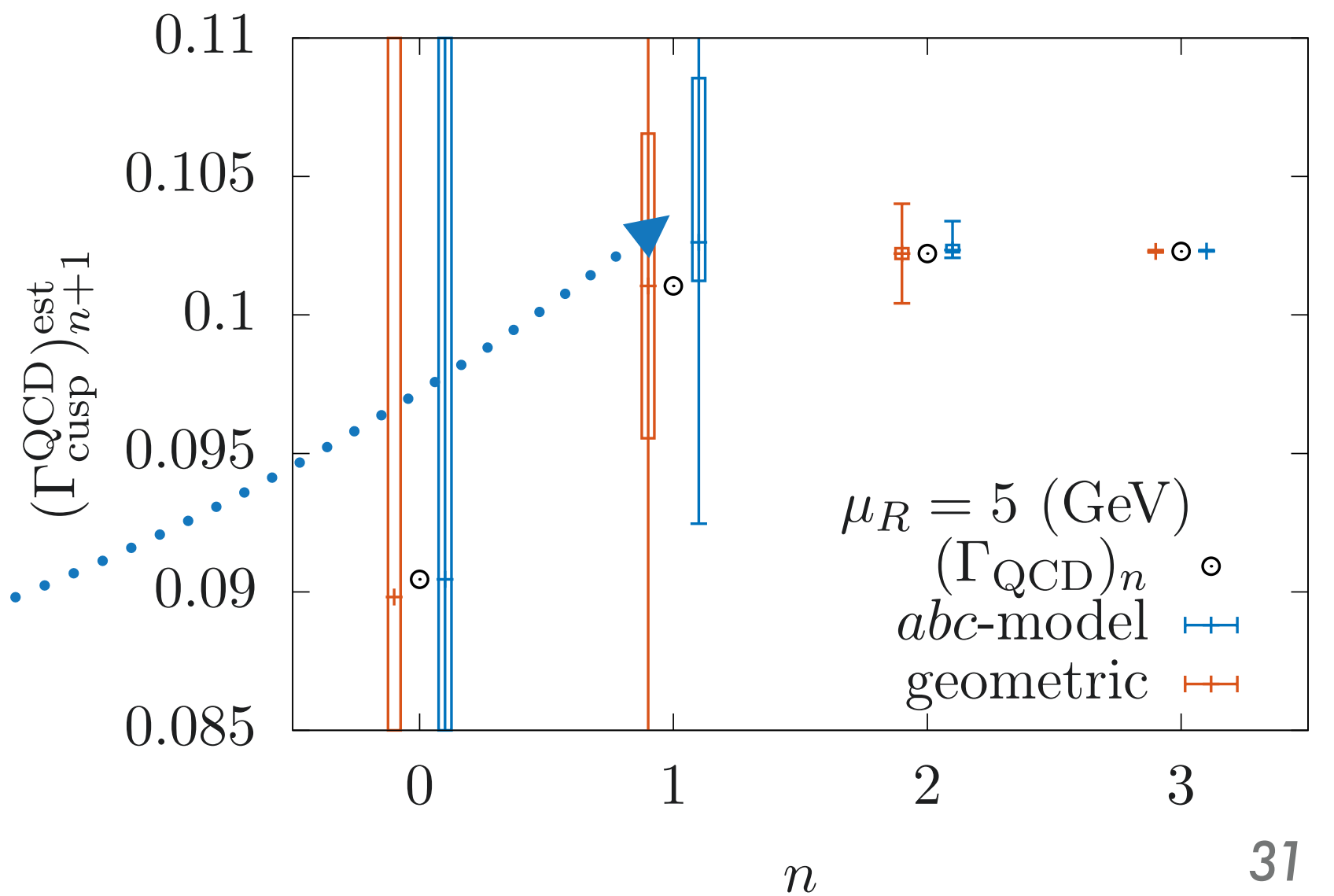
estimate for $\Gamma_{\text{cusp}}^{\text{QCD}} - (\Gamma_{\text{cusp}}^{\text{QCD}})_4$

- ▶ $\text{CI}_{68} = C_F \left(\frac{\alpha_s}{\pi}\right)^5 [2.1, 9.5]$
- ▶ $\text{CI}_{95} = C_F \left(\frac{\alpha_s}{\pi}\right)^5 [-0.38, 21]$

strongly peaked $n \nearrow$

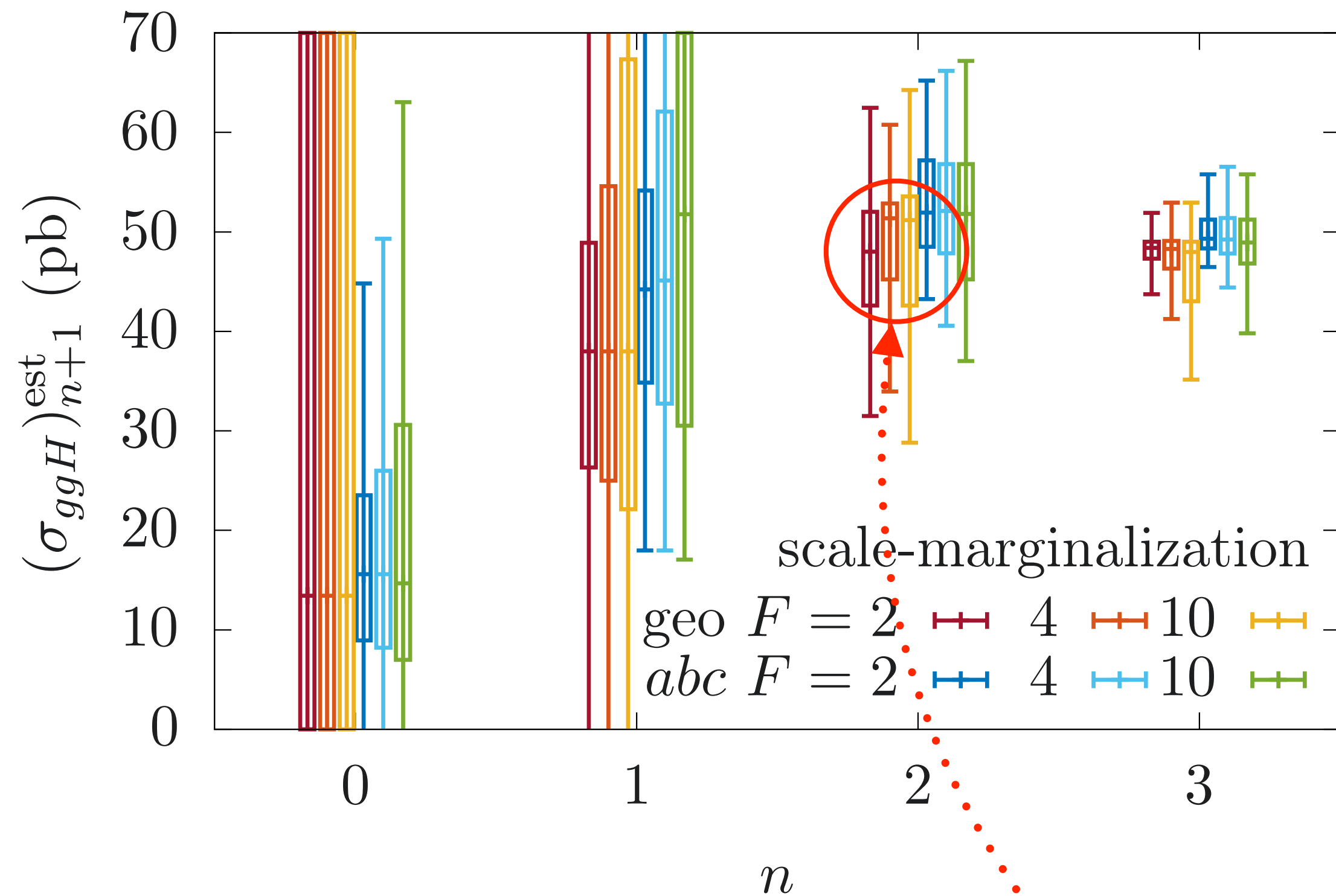
$n = 2$: anticipates the exact known result

CI_{68} always captures next term



SENSITIVITY ON THE RANGE F

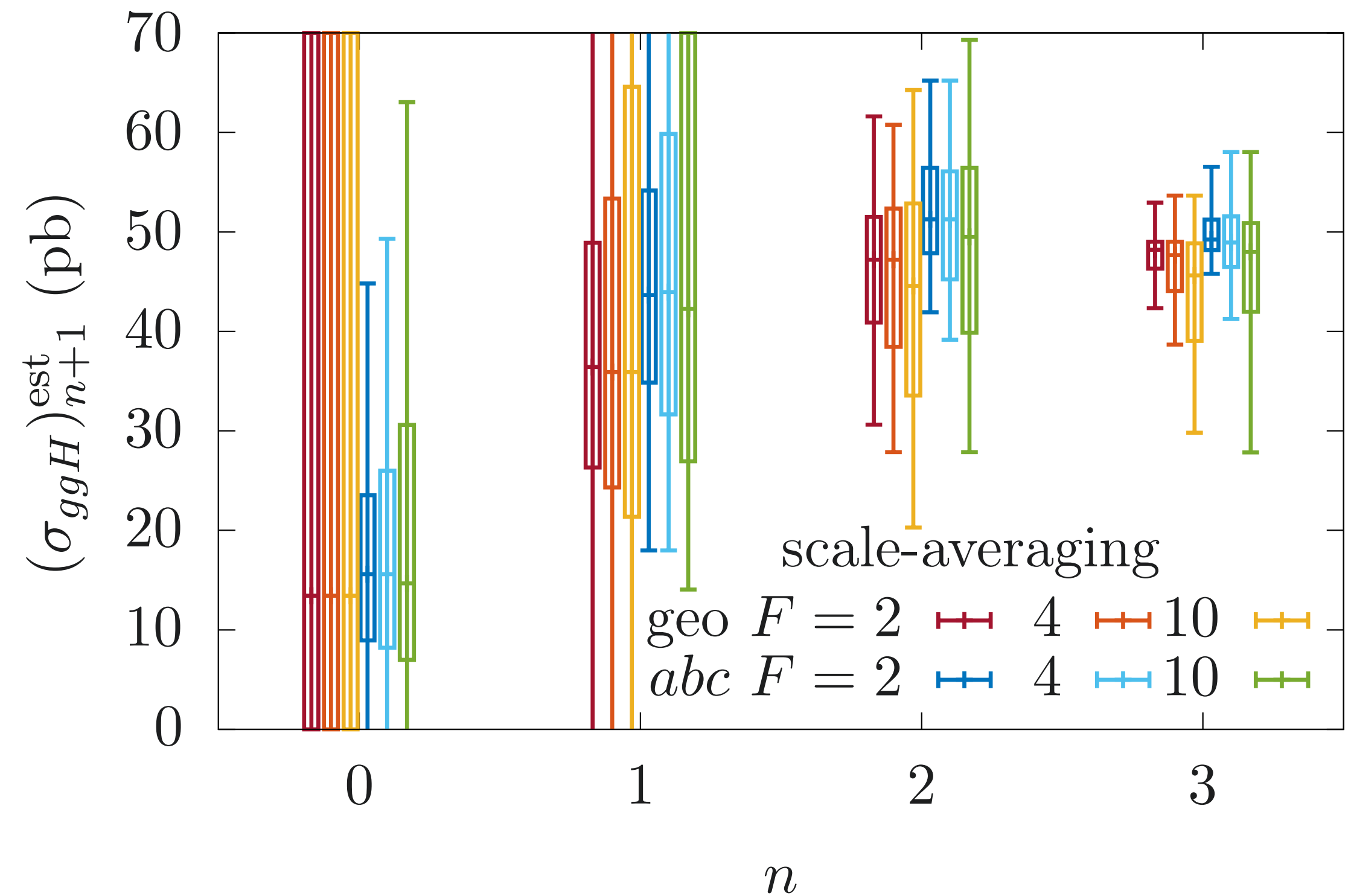
Scale Marginalisation (sm):



milder dependence on F in sm procedure

transition to $\mu_{\text{FAC}} \in [\mu_0/F, F\mu_0]$

Scale Average (sa):

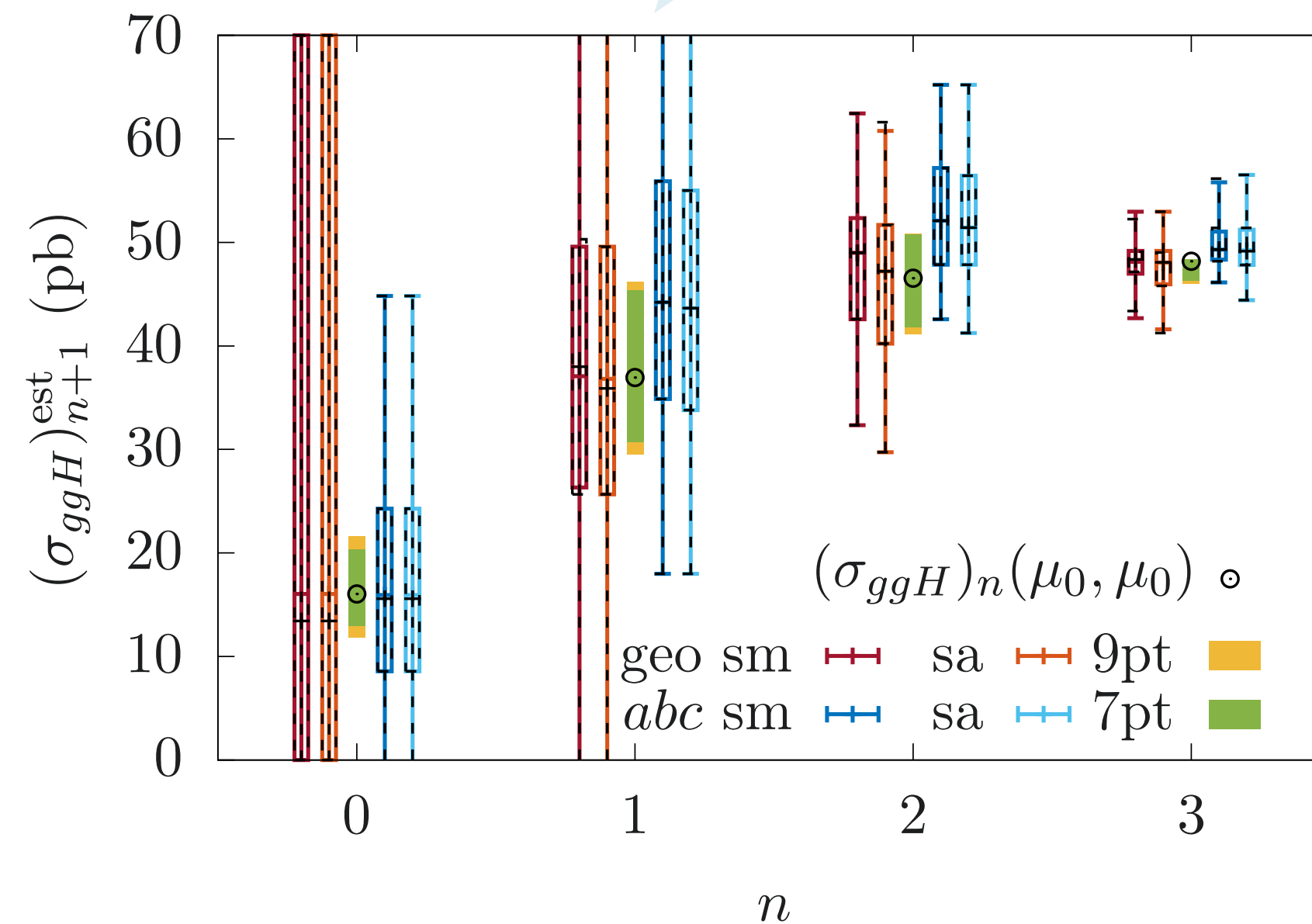
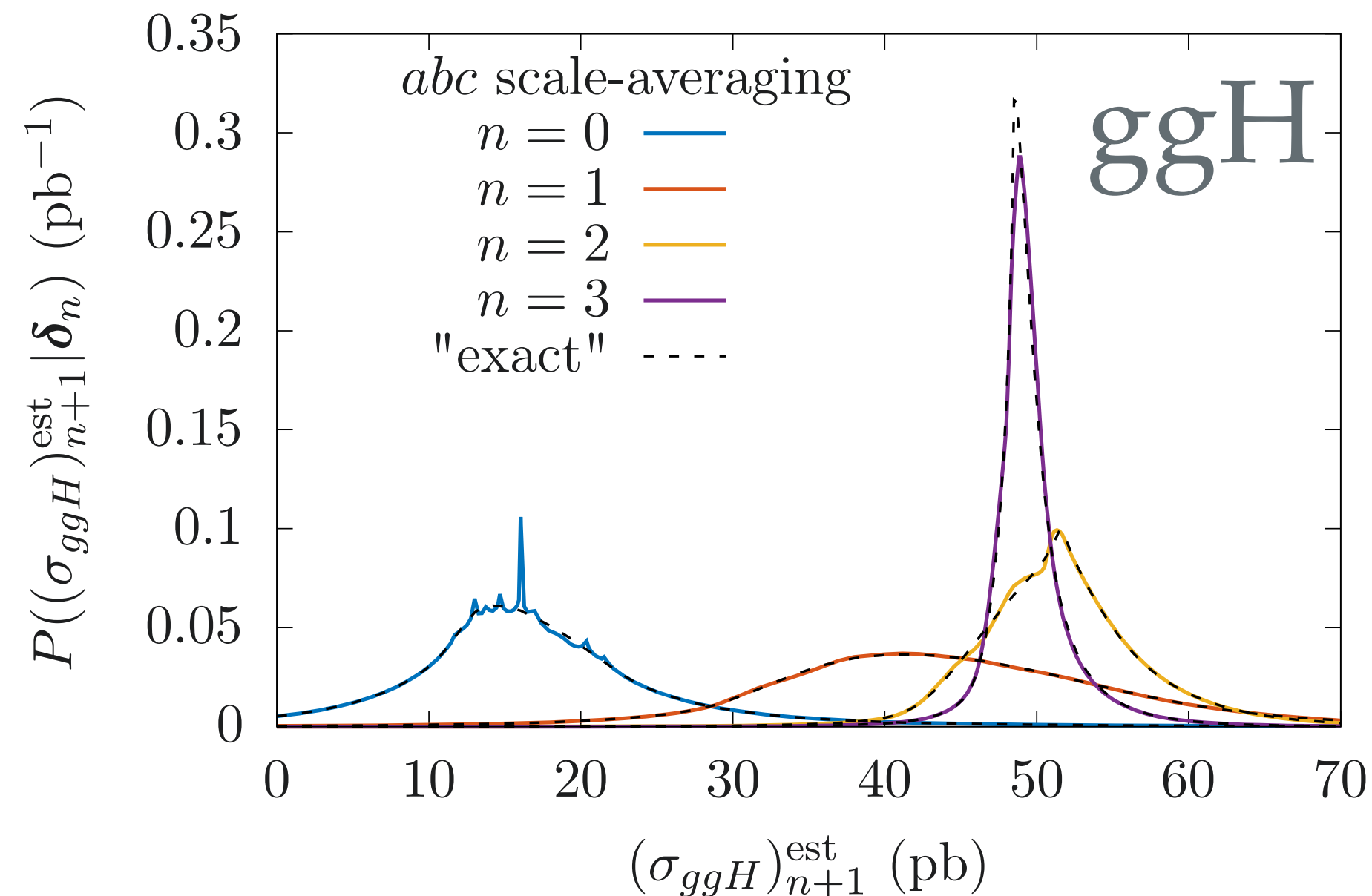


INTERMEZZO — INTEGRATION OVER μ

- the integration over μ is in general very costly (numerical)
 - \rightsquigarrow approximate it using *quadrature rule* (works well for $CI_{68/95}$)
 - \rightsquigarrow recycle existing calculations done for $\{\mu_0/2, 2\mu_0\}$
- Gauss-Legendre ($w_0 = \frac{8}{18}$; $w_{\pm 1} = \frac{5}{18}$) \rightsquigarrow corresponds to $F \simeq 2.45$

$$\int d\mu w(\mu) f(\mu) \simeq w_{-1} f(\mu_0/2) + w_0 f(\mu_0) + w_{+1} f(2\mu_0)$$

more conservative
up to NLO



from NNLO on
similar to 7/9-pt

N3LO 7/9-pt
asymmetric