

Probing pseudo-Nambu-Goldstone dark matter evading direct detection bounds

Takashi Toma (Kanazawa U.)

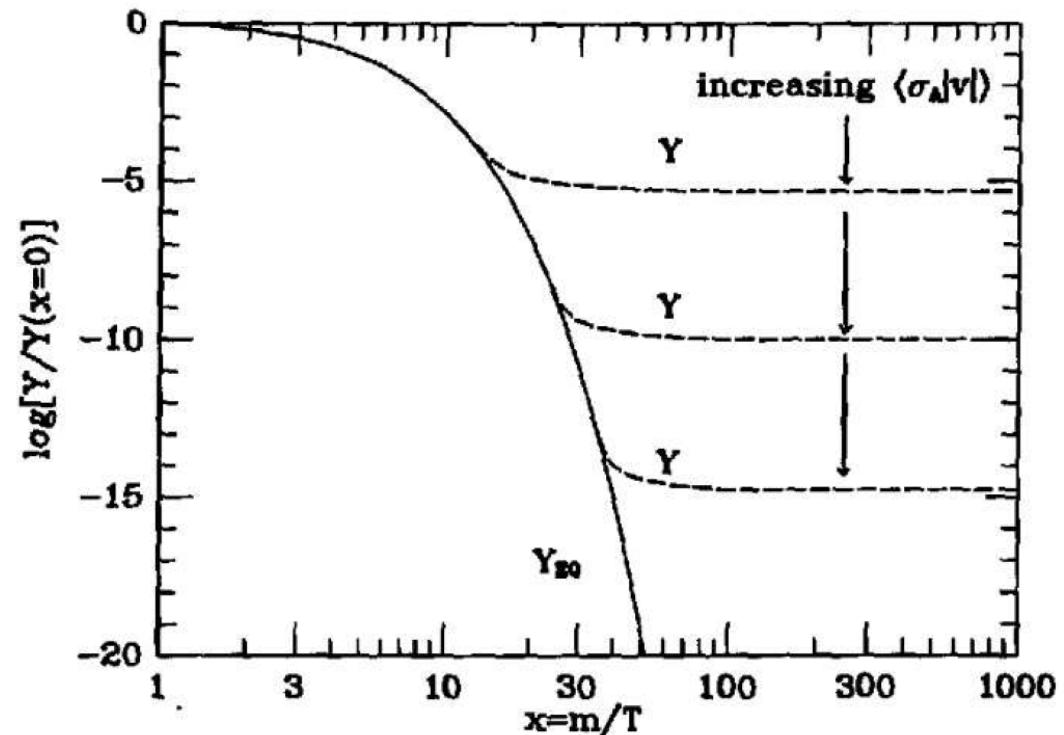
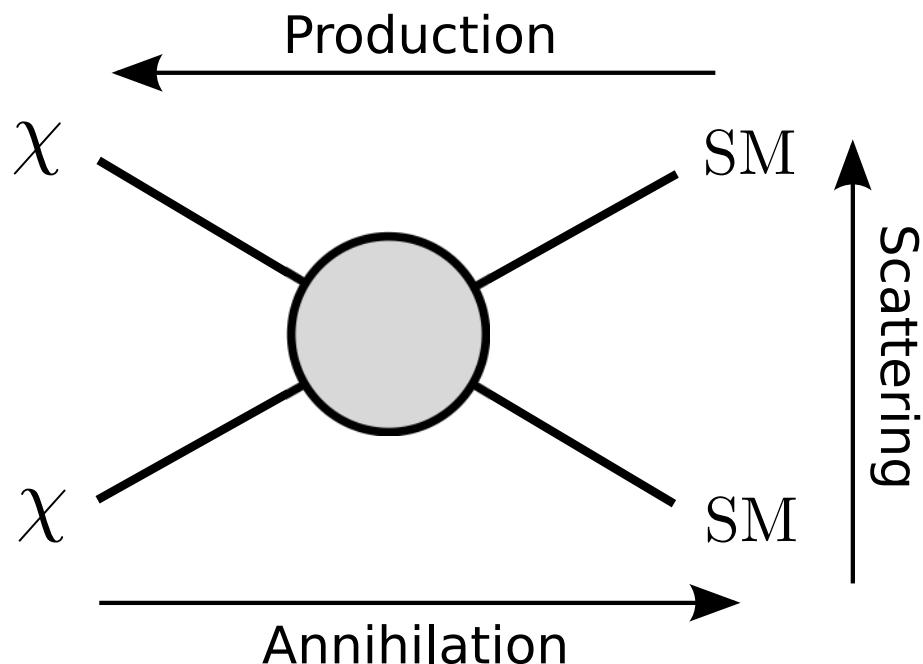
NuDM-2024 @ Cairo, Egypt
11th–14th Dec. 2024

Based on
Phys. Rev. Lett. **119**, no.19, 191801 (2017)
JHEP **12**, 089 (2018), JHEP **05**, 057 (2020)
Phys. Rev. D **104**, no.3, 035011 (2021)
and work in progress



Collaborators: Gross, Lebedev, Ishiwata, Abe, Tsumura, Yamatsu, and more

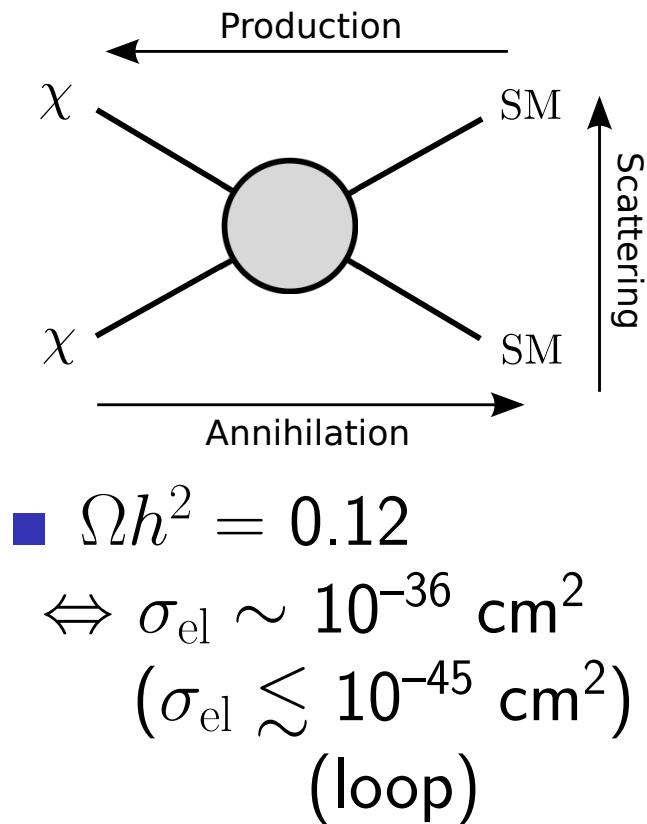
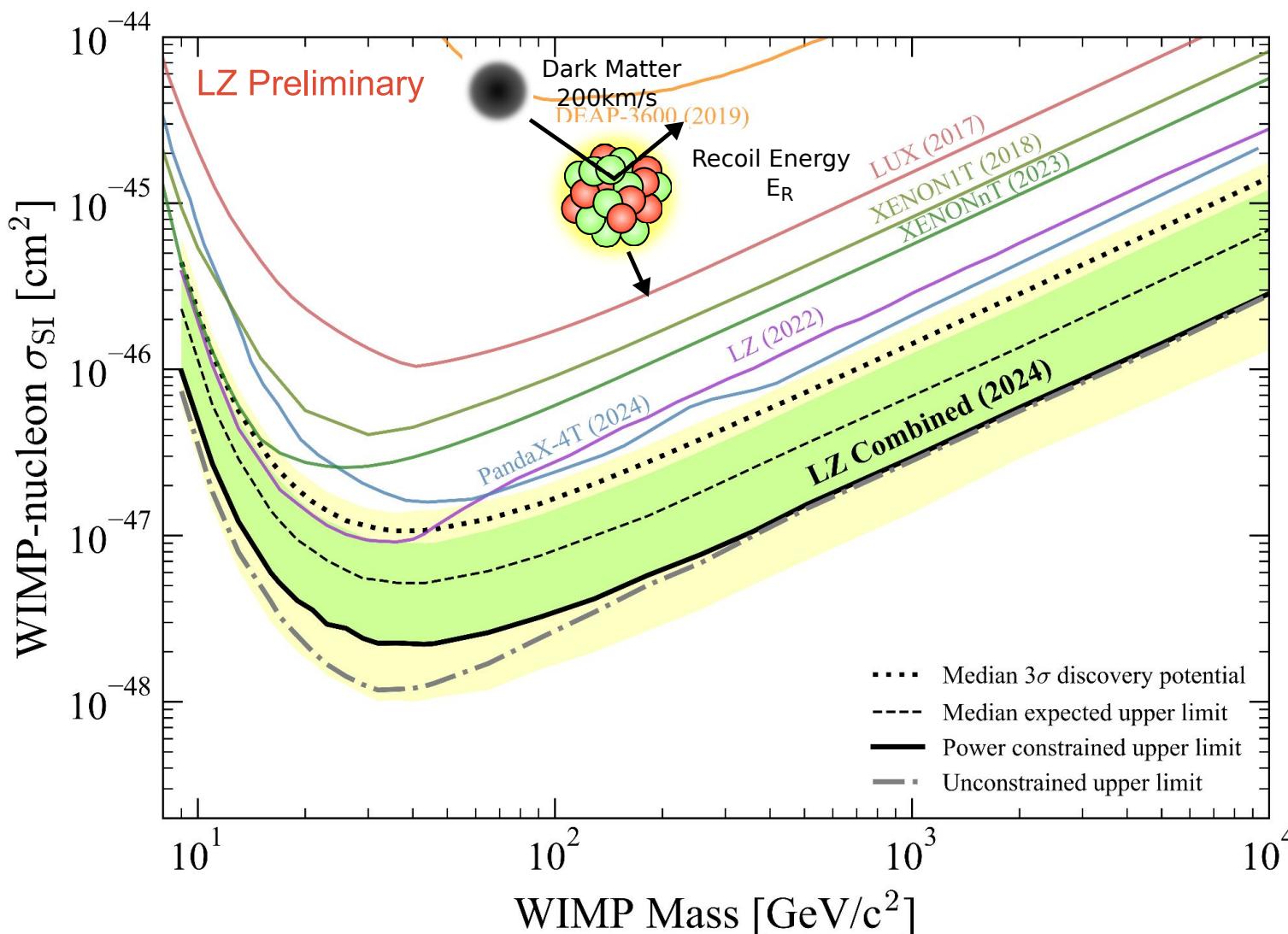
Thermal dark matter



$$\frac{dn_\chi}{dt} + 3Hn_\chi = -\langle \sigma v \rangle (n_\chi^2 - n_\chi^{\text{eq}2})$$

- WIMPs are thermalized with SM particles in early universe.
- To get $\Omega_\chi h^2 = 0.12$, roughly $\sigma \sim 1\text{pb} \sim 10^{-26}\text{cm}^3/\text{s} \sim 10^{-36}\text{cm}^2$
- Almost independent on DM mass

Status of direct detection experiments



LZ talk @ TeVPA2024

- LZ gives the strongest bound $2.2 \times 10^{-48} \text{ cm}^2$ at 43 GeV.
- Need a way to evade the bound

Wayout

- v_χ dependent cross section ($v_\chi \sim 10^{-3}$)

Ex.1 pNG DM ($i\mathcal{M} \propto v_\chi^2$)

C. Gross, O. Lebedev, TT, PRL (2017) [arXiv:1708.02253]

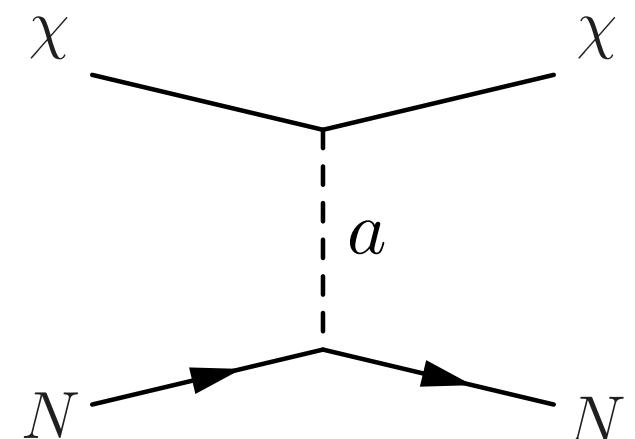
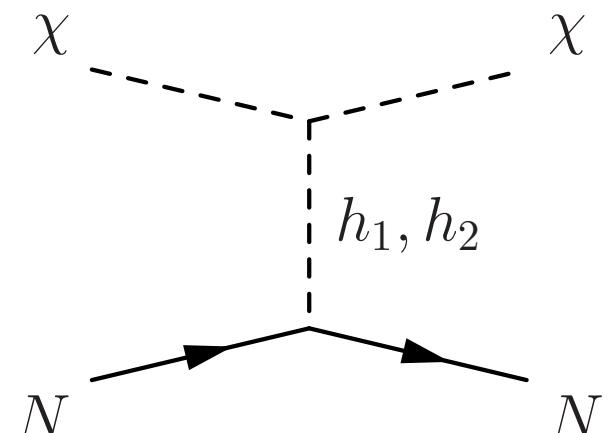
Ex.2 Fermionic DM with Pseudo-scalar int.

$$\mathcal{L} = a \bar{\chi} \gamma_5 \chi$$

T. Abe, M. Fujiwara, J. Hisano, JHEP (2019) [arXiv:1810.01039]

- Challenging to explore with standard way of direct detection experiments

⇒ These DM could be detected if boosted.



The pNG DM model

The pNG DM model

C. Gross, O. Lebedev, TT, PRL (2017) [arXiv:1708.02253]

- Introduce a complex scalar field $S = (s + i\chi)/\sqrt{2}$ (DM: χ)
- Global $U(1)$ symmetry is assumed (invariant under $S \rightarrow e^{i\alpha}S$)

$$\mathcal{V} = -\frac{\mu_H^2}{2}|H|^2 - \frac{\mu_S^2}{2}|S|^2 + \frac{\lambda_H}{2}|H|^4 + \lambda_{HS}|H|^2|S|^2 + \frac{\lambda_S}{2}|S|^4$$

$$- \left(\frac{\mu_S'^2}{4} S^2 + \text{H.c.} \right) \quad \leftarrow \text{soft breaking mass term}$$

- After H and S get VEVs, ϕ and s mix

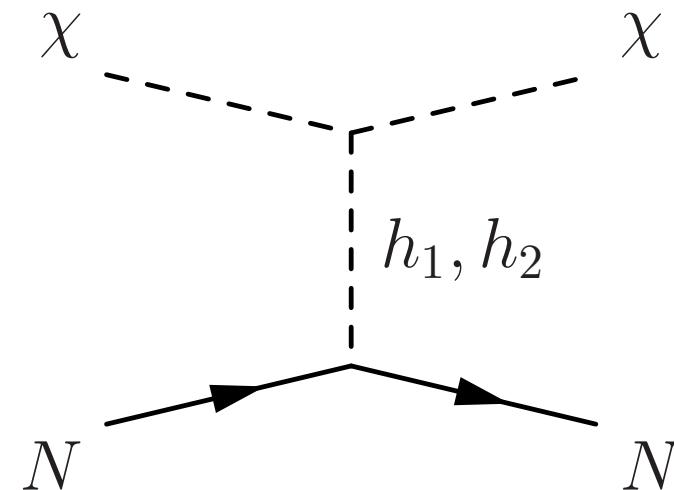
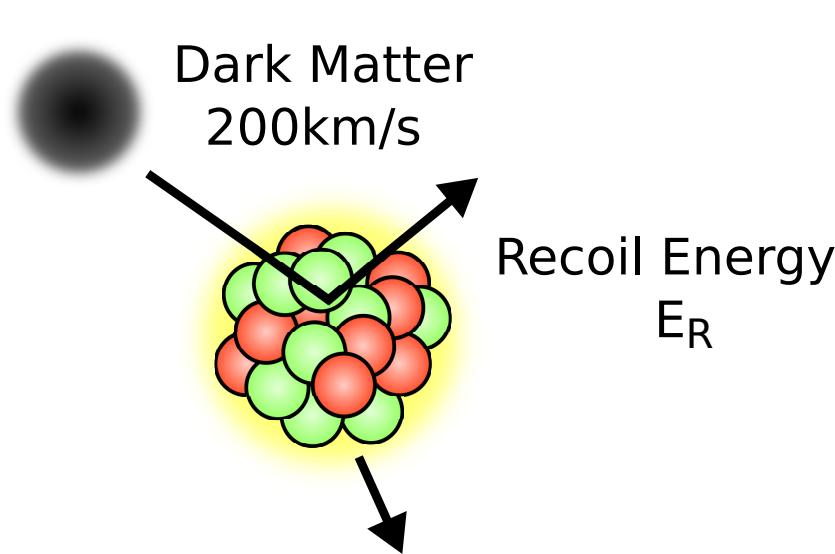
$$H = \begin{pmatrix} 0 \\ (v + \phi)/\sqrt{2} \end{pmatrix}, \quad S = \frac{v_s + s + i\chi}{\sqrt{2}}$$

$$\begin{pmatrix} \phi \\ s \end{pmatrix} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} h_1 \\ h_2 \end{pmatrix}$$

- $\sin \theta \lesssim 0.3 \quad \leftarrow \text{Constrained by EWPT, } h_2 \text{ direct search at LHC}$

Direct detection (tree level)

C. Gross, O. Lebedev, TT, PRL (2017) [arXiv:1708.02253]



- Scattering amplitude cancels between h_1, h_2 mediated diagrams

$$i\mathcal{M} \sim i \left(\frac{m_{h_1}^2}{q^2 - m_{h_1}^2} - \frac{m_{h_2}^2}{q^2 - m_{h_2}^2} \right) \sim i \frac{q^2(m_{h_1}^2 - m_{h_2}^2)}{m_{h_1}^2 m_{h_2}^2} \rightarrow 0$$

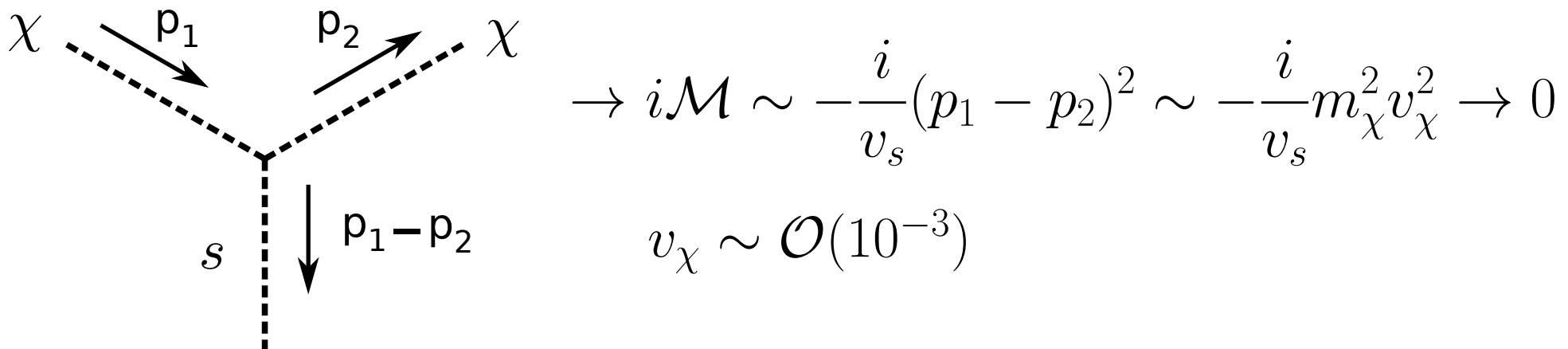
\Rightarrow Naturally evade the direct detection bounds

- The cancellation is due to nature of Goldstone boson

Direct detection (tree level)

C. Gross, O. Lebedev, TT, PRL (2017) [arXiv:1708.02253]

$$\text{Rewrite with } S = \frac{(v_s + s)}{\sqrt{2}} e^{i\chi/v_s} \Rightarrow \mathcal{L} \supset \frac{1}{v_s} s \left[(\partial_\mu \chi)^2 - m_\chi^2 \chi^2 \right]$$

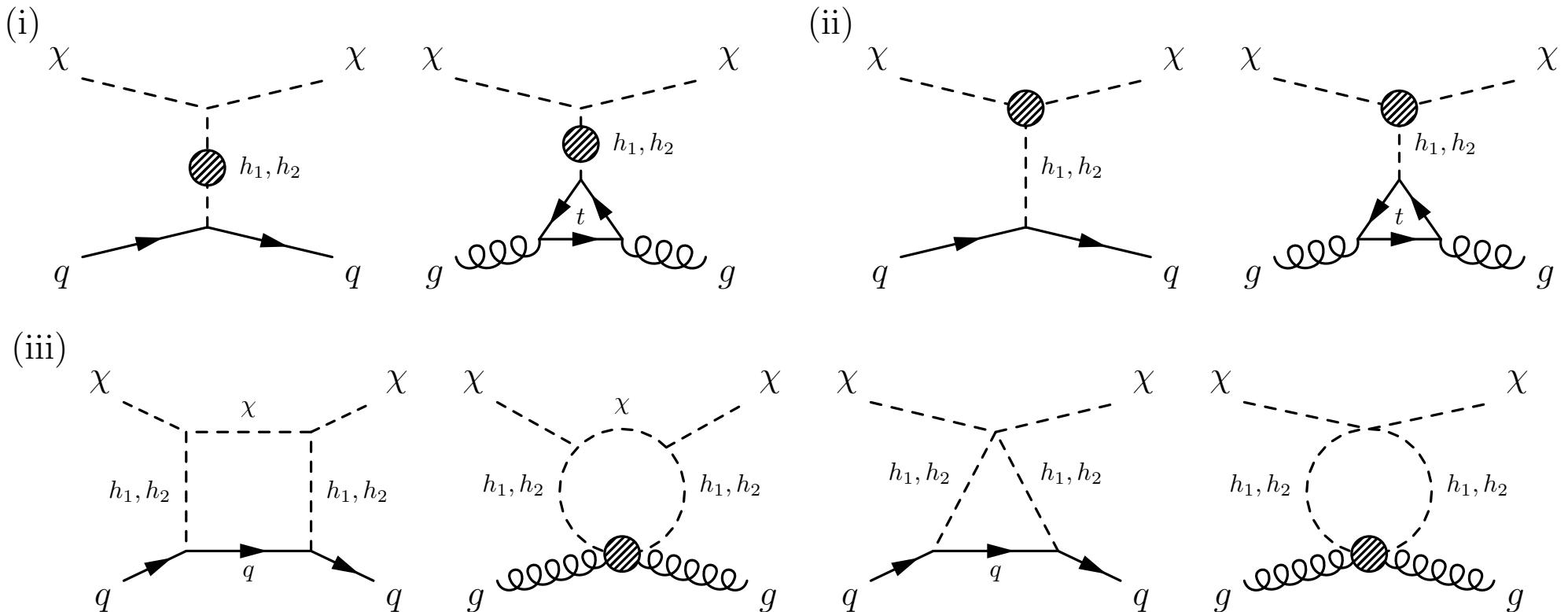


- The cancellation is due to nature of Goldstone boson
- All interactions are written with derivative couplings $\mathcal{L}_{\text{int}} = \mathcal{L}_{\text{int}}(\partial_\mu \chi)$
- Only 4 independent parameters ($m_\chi, m_{h_2}, \sin \theta, v_s (\lambda_S)$)

Direct detection (1-loop level)

D. Azevedo et al., JHEP [arXiv:1810.06105]
 K. Ishiwata, TT, JHEP [arXiv:1810.08139]
 S. Glaus et al., JHEP [arXiv:2008.12985]

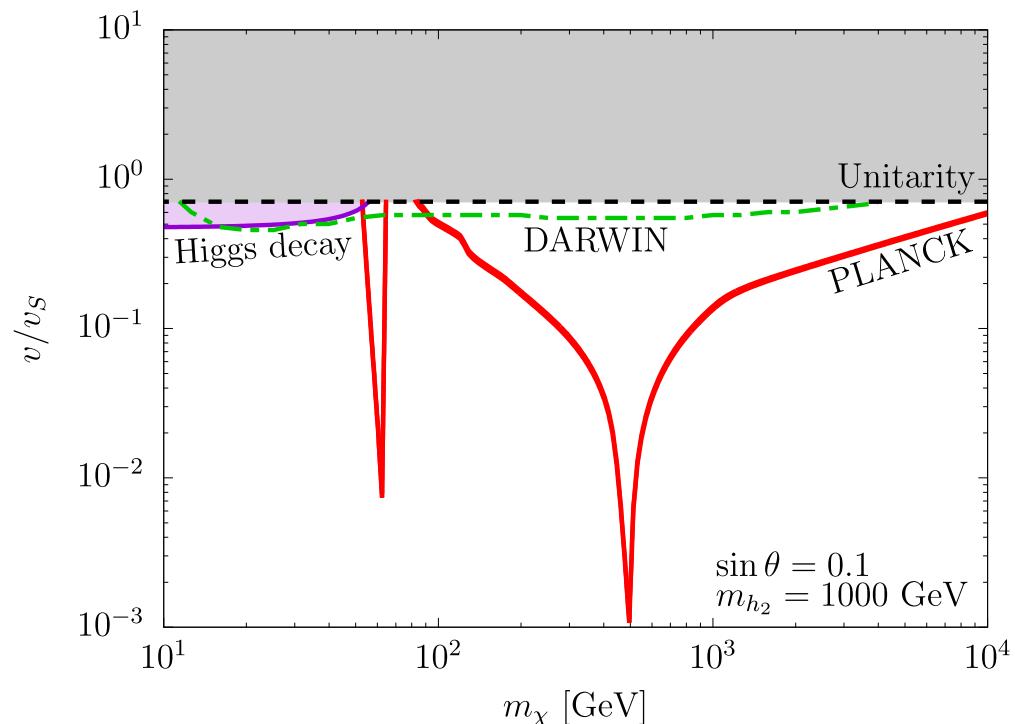
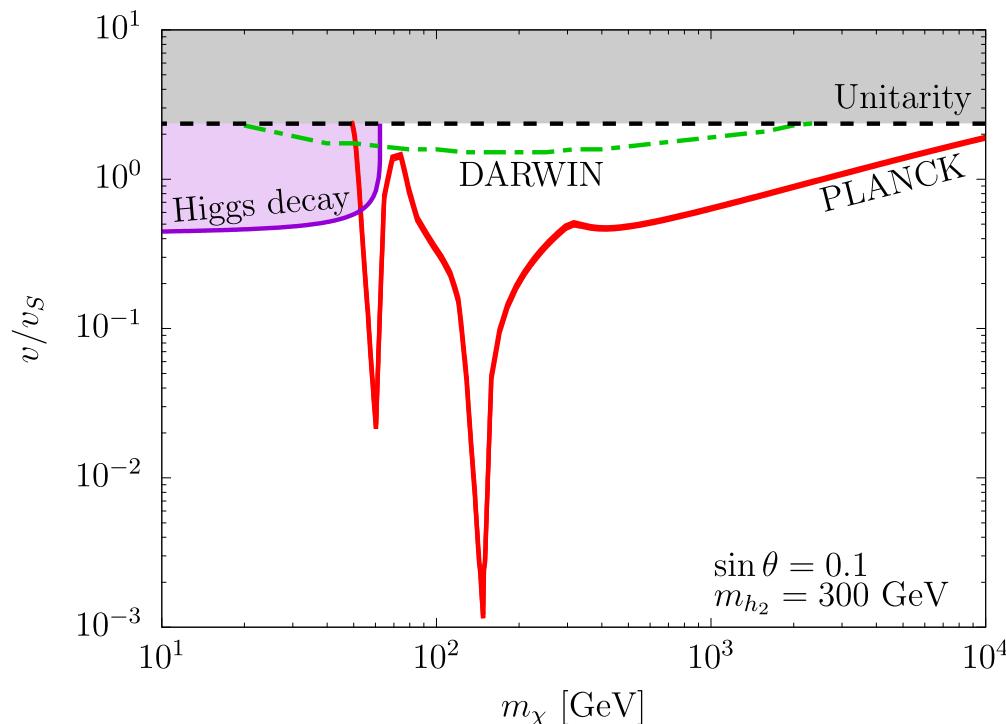
Constant amplitude at 1-loop level



- (i) self-energy correction
- (ii) vertex correction
- (iii) box and triangle → two Yukawa couplings
 (sub-dominant)

Numerical calculations

K. Ishiwata, TT, JHEP [arXiv:1810.08139]

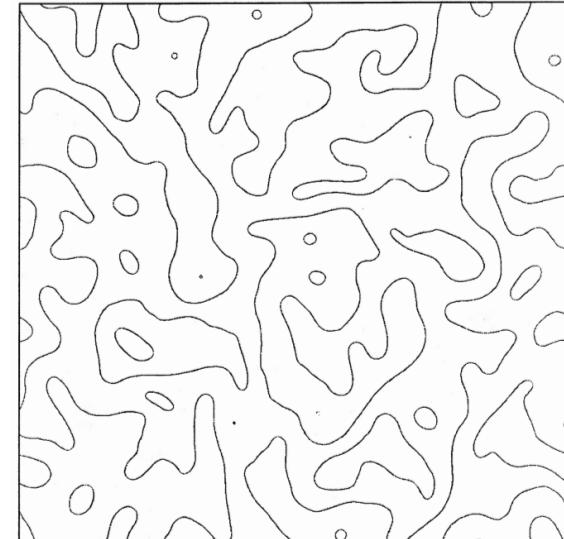
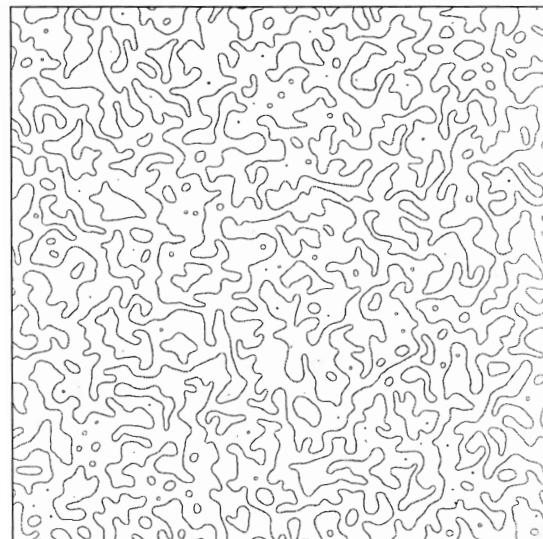
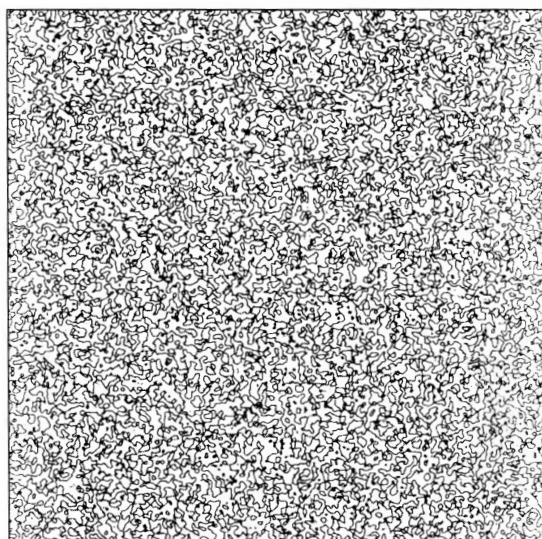


- $\sin \theta = 0.1$
- Unitarity bound: $\lambda_S \leq 8\pi/3$
- Higgs Invisible decay $\text{Br}(h_1 \rightarrow \text{inv}) \lesssim 20\%$ at LHC
- $\Omega h^2 = 0.12$ (PLANCK)
- $\sigma_{\text{SI}}^p = \mathcal{O}(10^{-48}) \text{ cm}^2$ at most

Extension of the model

Domain wall problem

- Domain walls due to spontaneous breaking of \mathbb{Z}_2 symmetry
⇒ distort CMB
- Solutions:
 - UV completion
 - low energy inflation after the \mathbb{Z}_2 breaking
 - decay before BBN (making the domain wall unstable)



time evolution →→→
Press, Ryden, Spergel ApJ (1989)

UV completion

Y. Abe, TT, K. Tsumura, JHEP (2020) [arXiv:2001.03954]

- Origin of the soft breaking term? $\frac{m_\chi^2}{4} S^2 + \text{H.c.}$

	Q_L	u_R^c	d_R^c	L	e_R^c	H	ν_R^c	S	Φ
$SU(3)_c$	3	$\bar{3}$	$\bar{3}$	1	1	1	1	1	1
$SU(2)_L$	2	1	1	2	1	2	1	1	1
$U(1)_Y$	+1/6	-2/3	+1/3	-1/2	+1	+1/2	0	0	0
$U(1)_{B-L}$	+1/3	-1/3	-1/3	-1	+1	0	+1	+1	+2

- Gauged $U(1)_{B-L}$ extension (New fields: ν_R , Φ)
- Potential $\mathcal{V} \supset \mu_c \Phi^* S^2 + \text{h.c.} \rightarrow m_\chi^2 S^2$ at low energy
The soft breaking term is induced.
- 3 ν_R for anomaly cancellation
- Seesaw for ν mass: $\mathcal{L} \supset \Phi \nu_R \nu_R$

GUT embedding

Y. Abe, TT, K. Tsumura, N. Yamatsu, PRD (2021) [arXiv:2104.13523]

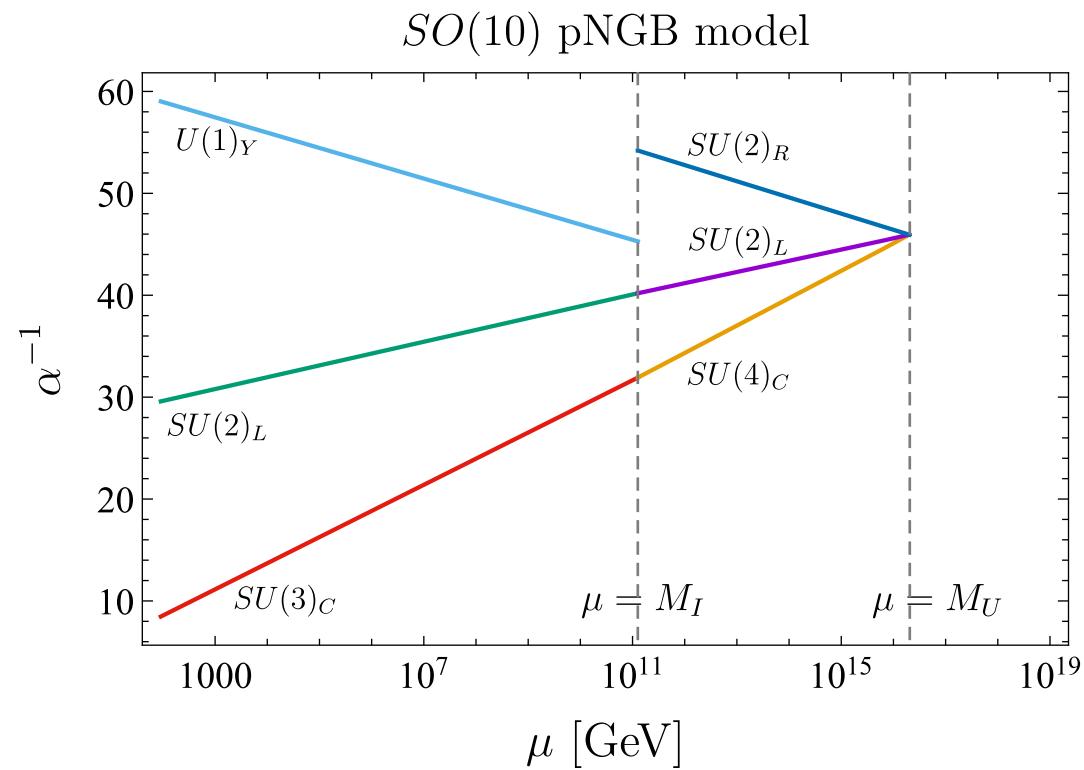
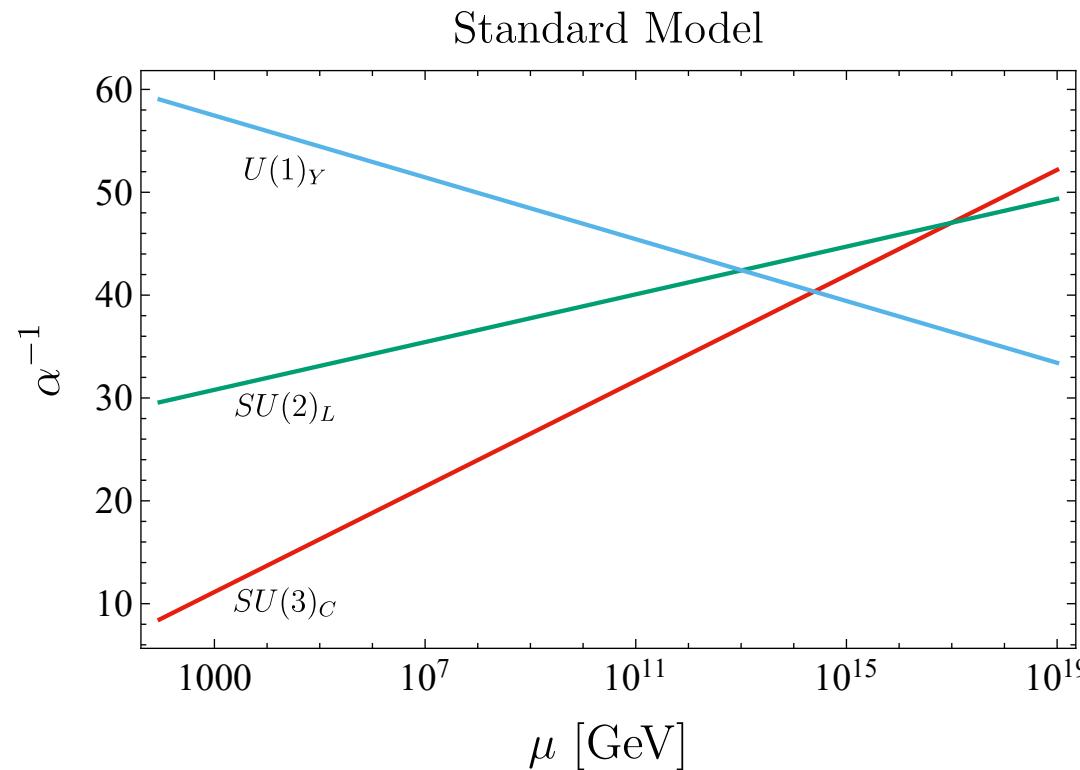
	fermions	H	S	Φ	SO(10)	
	A_μ	$\Psi_{\mathbf{16}}$	$\Phi_{\mathbf{10}}$	$\Phi_{\mathbf{16}}$	$\Phi_{\overline{\mathbf{126}}}$	$\Phi_{\mathbf{210}}$
$SO(10)$	45	16	10	16	$\overline{126}$	210

- We embed the UV complete model in $SO(10)$ GUT.
- The pNG model is reproduced at low energy.
- Breaking pattern: $SO(10) \rightarrow G_{\text{PS}} \rightarrow G_{\text{SM}}$
at $\mu = M_U$ at $\mu = M_I$

Pati-Salam symmetry: $G_{\text{PS}} = SU(4)_C \times SU(2)_L \times SU(2)_R$

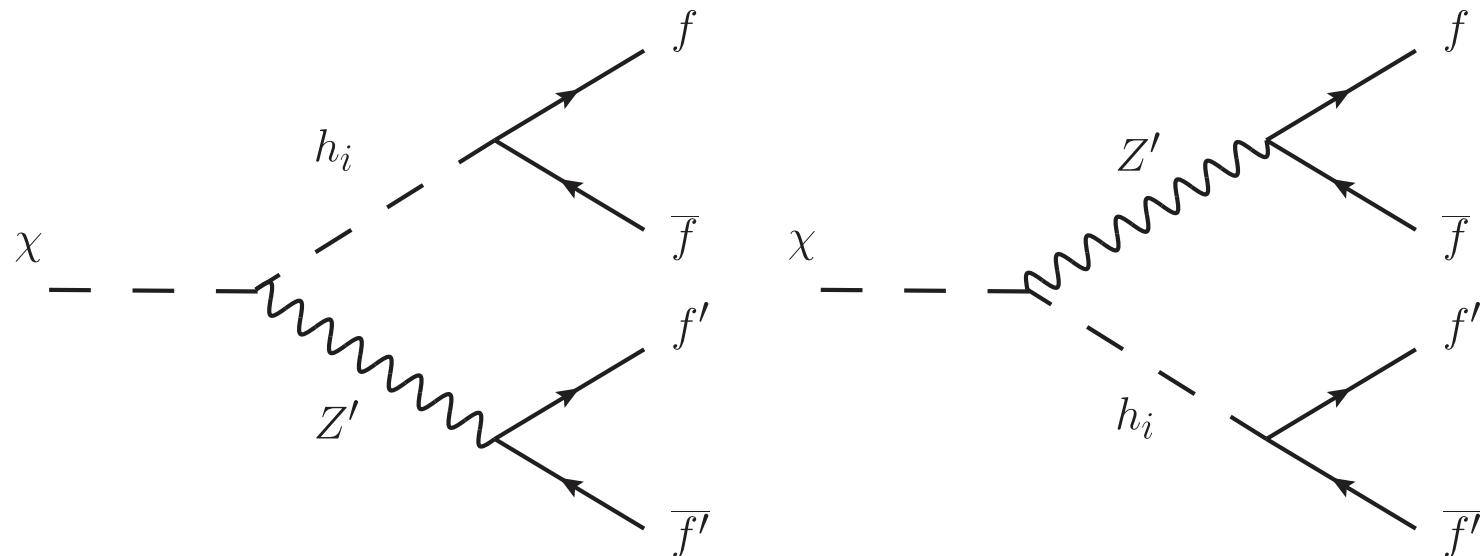
- GUT scale (M_U)
Intermediate scale (M_I) = breaking scale of $U(1)_{B-L}$,
Determined by matching conditions of gauge couplings

Gauge coupling unification



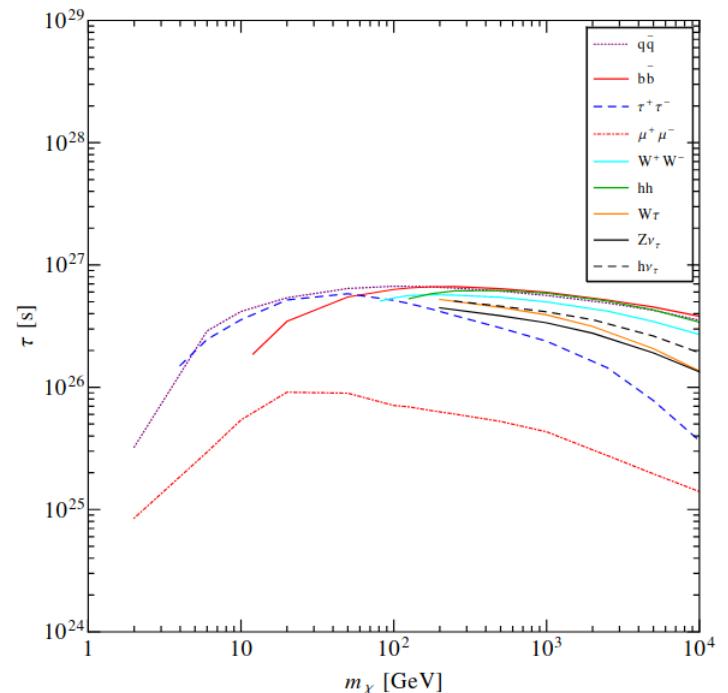
- Intermediate scale, GUT scale: determined by matching conditions of gauge couplings
- $v_\phi = M_I = 1.26 \times 10^{11} \text{ GeV}$, $M_U = 2.06 \times 10^{16} \text{ GeV}$
- $g_{B-L} = 0.38$ at $\mu = M_I$

DM decay

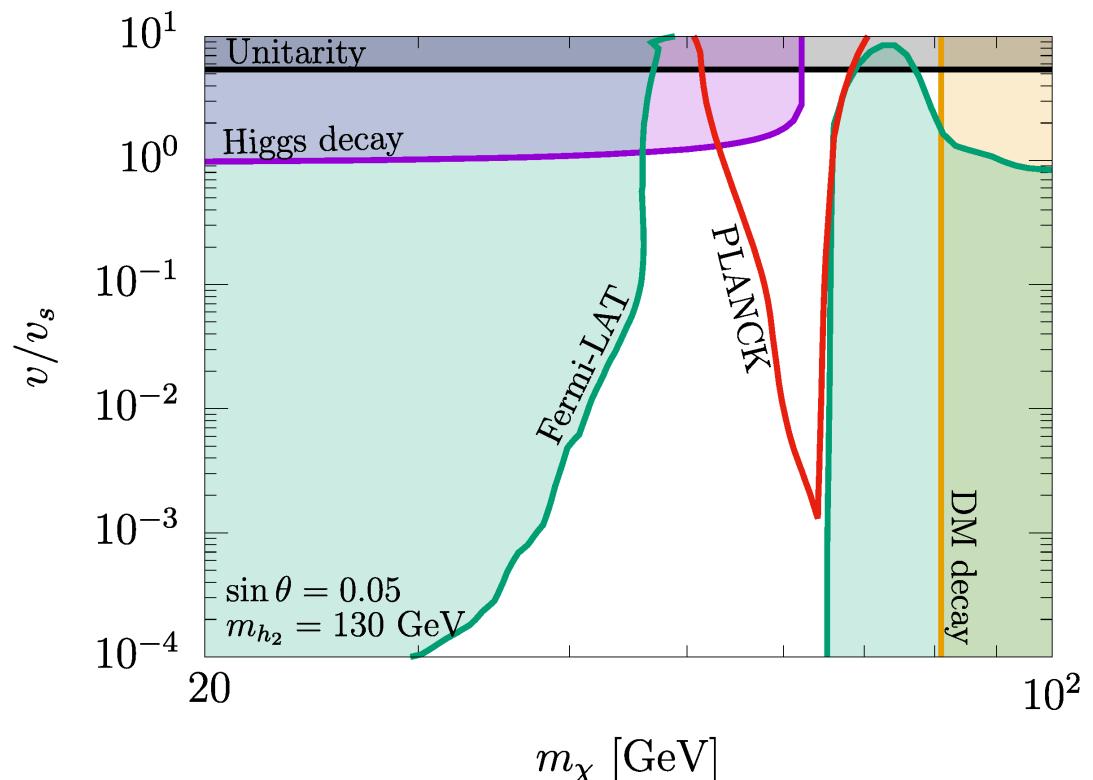
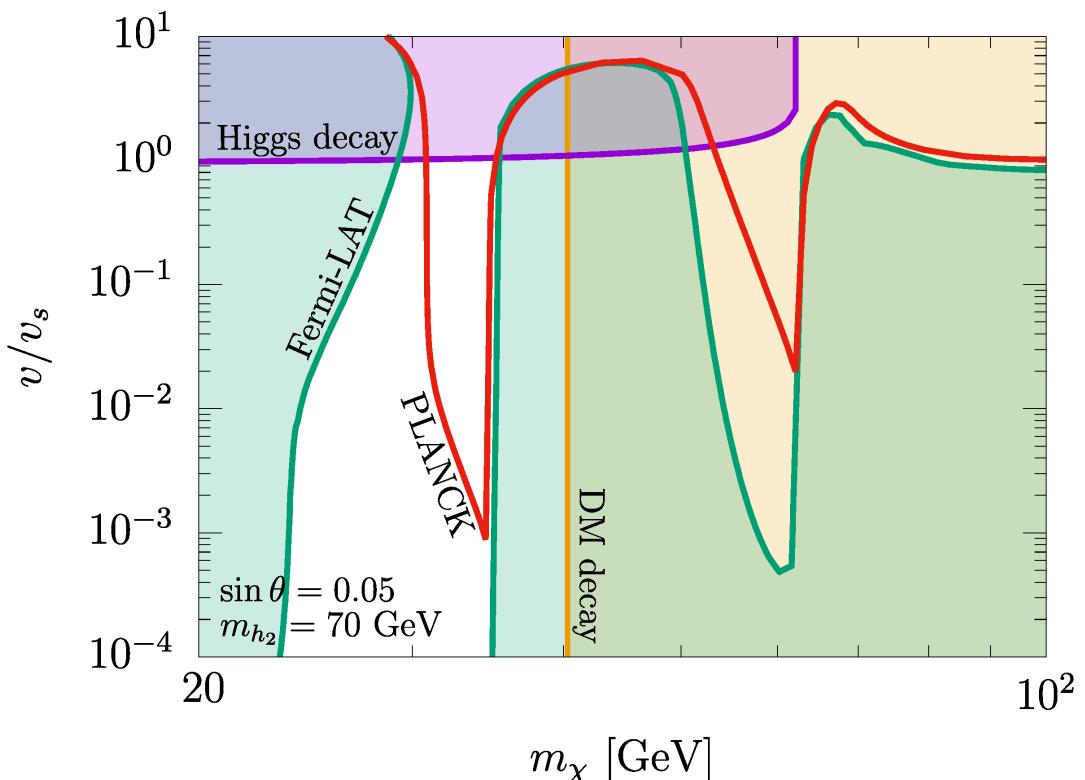


- DM lifetime: $\tau_{\text{DM}} \gtrsim 10^{17}$ sec at least (the age of the universe).
- Cosmic ray observations give stronger limits: $\tau_{\text{DM}} \gtrsim 10^{27}$ sec.
- 3-body decays $\chi \rightarrow f\bar{f}h_i, f\bar{f}Z$ if $m_\chi \gtrsim m_{h_i}, m_Z \rightarrow$ excluded

Baring et al. (2015)



Parameter space



- $v/v_s \sim \sqrt{\lambda_S}$
- $\tau_\chi \gtrsim 10^{27}$ sec
- Fermi-LAT: $\chi\chi \rightarrow b\bar{b}, WW \rightarrow$ gamma-ray production
- close to the h_2 resonance

Boosting mechanism

Other variations of pNG DM models

- THDM + S with global U(1) [Zhang, Cai, Jiang, Tang, Yu, Zhang, JHEP 05 \(2021\) 160](#)
⇒ Gravitational waves from strong 1st order phase transition
- Various pNG DM models

	DW problem	stable/decaying	real/complex	Landau pole
Global U(1)	✗	stable	real	✓
Gauged U(1) _{B-L}	✓	decaying	real	✓
SO(10) GUT	✓	decaying	real	✓
Hidden gauged U(1)	✓	decaying	real	✓
Global SU(2) with gauged U(1)	✓	stable	complex	✗
Gauged SU(2)	✓	stable	complex	✓
Gauged U(1) with 2 scalars	✓	stable	real	✓

[Abe, Hamada, Tsumura, JHEP 05 076 \(2024\)](#)

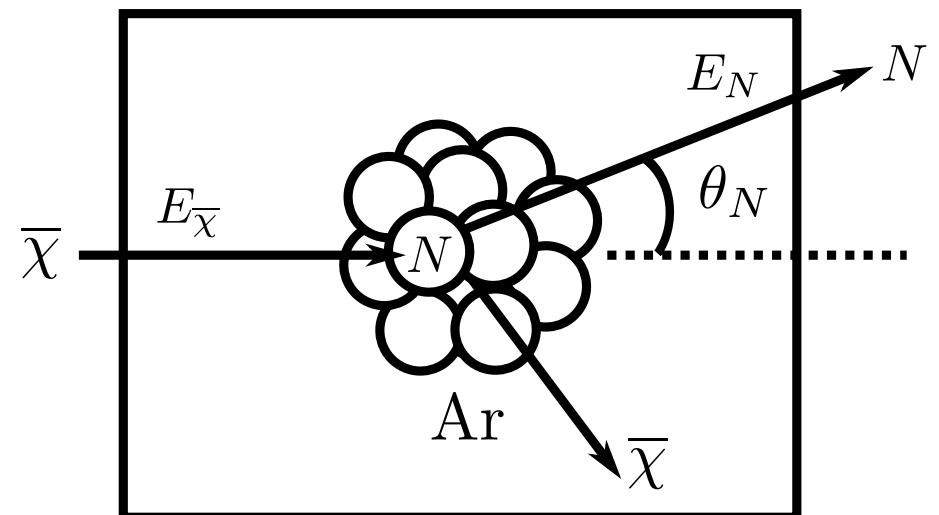
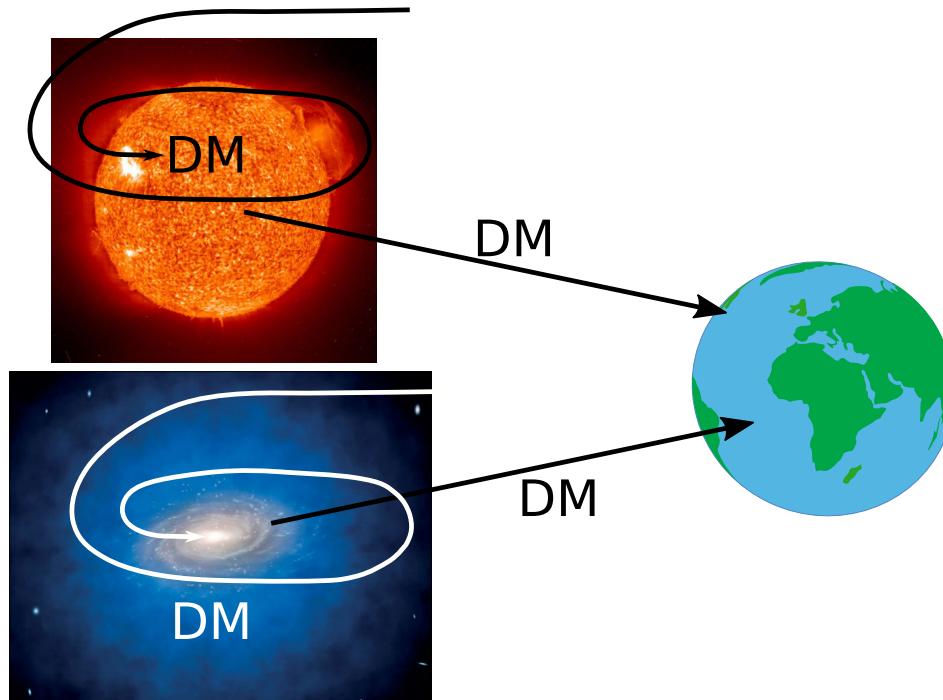
- No domain wall

Two component DM $\chi_2\chi_2 \rightarrow \chi_1\chi_1$

Semi-annihilating DM $\chi\chi \rightarrow \chi\phi \Rightarrow$ DM is accelerated

How to probe pNG DM

- PNG DM can be detectable if boosted by some mechanism.
 $\sigma_{\text{el}} \propto Q^4 \Rightarrow$ strong enhancement of signals $\Rightarrow \nu$ detectors (DUNE)
- Boosting mechanisms
 - Semi-annihilations
 - Decay or annihlations of heavy particles
 - Up-scattering with high energy cosmic-rays

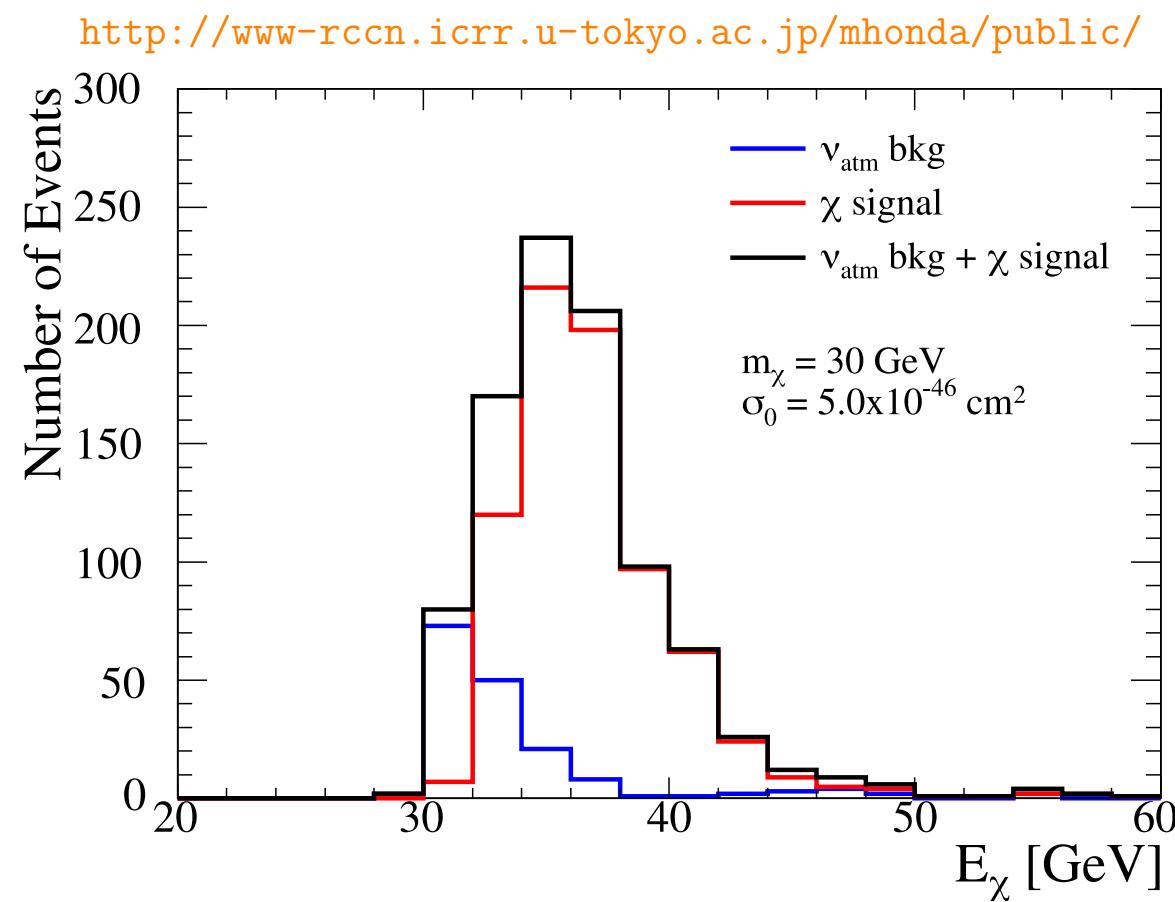
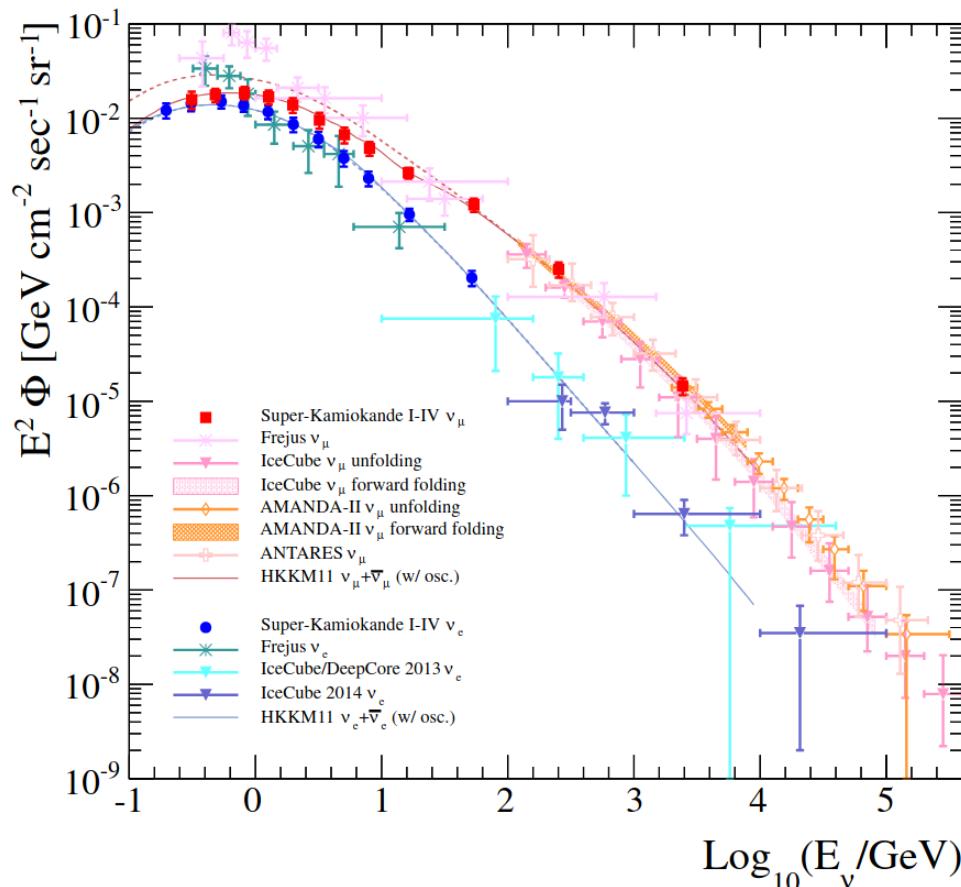


How to probe pNG DM 2

- Simplest option: \mathbb{Z}_3 symmetry \Rightarrow semi-annihilation ($\chi\chi \rightarrow \chi Z'$)

$$E_\chi = \frac{5m_\chi^2 - m_{Z'}^2}{4m_\chi}, \quad \gamma_\chi \equiv \frac{E_\chi}{m_\chi} = \frac{5}{4} - \frac{m_{Z'}^2/m_\chi^2}{4} \leq 1.25$$

- Atmospheric ν background and signals



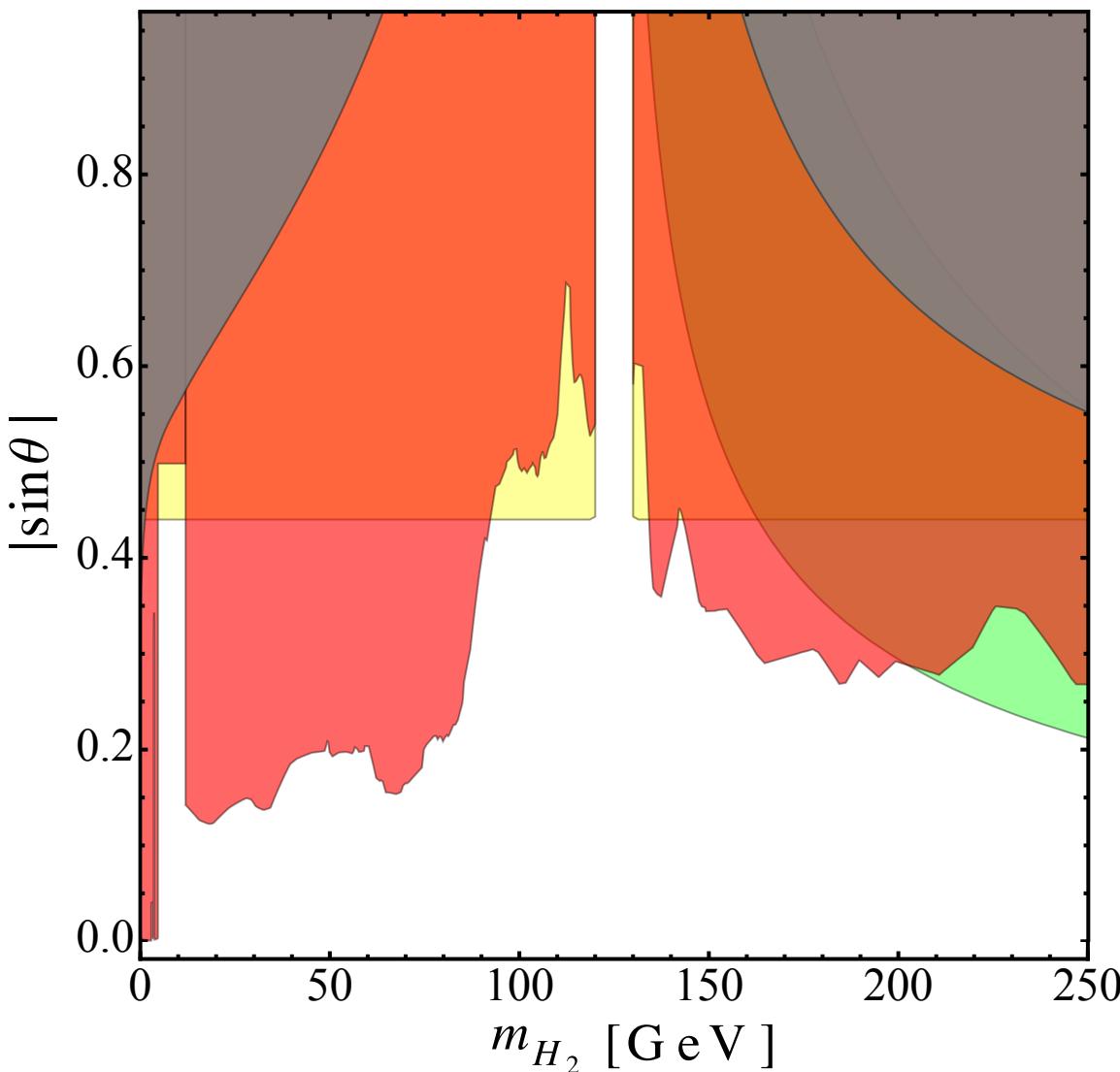
Summary

- 1 Thermal DM is strongly constrained by direct detection experiments.
⇒ pNG DM $\sigma_{\text{el}} \propto v_\chi^4$.
- 2 Further extensions of the pNG DM model have been proposed.
⇒ Solution for the domain wall problem and origin of S^2 term
- 3 Strong pNG DM signals are expected if it is boosted.
⇒ Extended models with semi-annihilations

Backup

Bound on $\sin \theta$

Falkowski, Gross, Lebedev, JHEP 1505 (2015) [arxiv:1502.01361]



- **Red:** h_2 direct search at LHC
- **Yellow:** h_1 coupling measurements
- **Green:** Favored region from stability of scalar potential
- **Gray:** Electroweak precision tests
- $|\sin \theta| \lesssim 0.3$ if $m_{h_2} \gtrsim m_{h_1}$
- $m_\chi \lesssim m_{h_2}$ (above EW scale)

The pNG DM model

C. Gross, O. Lebedev, TT, PRL (2017) [arXiv:1708.02253]

- χ is mass eigenstate itself $m_\chi^2 = \mu_S'^2$
Invariant under $S \rightarrow S^*$, $\Rightarrow \chi$ can be a DM candidate
- Higgs portal DM
- Scalar potential $\mathcal{V} = \mu_{h_1\chi\chi} h_1 \chi^2 + \mu_{h_2\chi\chi} h_2 \chi^2 + \dots$

$$\mu_{h_1\chi\chi} = -\frac{m_{h_1}^2 \sin \theta}{v_s}, \quad \mu_{h_2\chi\chi} = \frac{m_{h_2}^2 \cos \theta}{v_s},$$

SM Yukawa int. $\mathcal{L} \supset y_q (\cos \theta h_1 + \sin \theta h_2) \bar{q} q$

$$\lambda_H = \frac{\cos^2 \theta m_{h_1}^2 + \sin^2 \theta m_{h_2}^2}{v^2}, \quad \lambda_S = \frac{\sin^2 \theta m_{h_1}^2 + \cos^2 \theta m_{h_2}^2}{v_s^2},$$

$$\lambda_{HS} = \frac{\sin \theta \cos \theta (m_{h_2}^2 - m_{h_1}^2)}{v v_s}$$