

MAIN CONNECTIONS BETWEEN COSMOLOGY & PARTICLE PHYSICS

THE EARLY UNIVERSE

O2 ACDM MODEL

COSMIC MICROWAVE BACKGROUND

04 CERN EXPERIMENTS

THE EARLY UNIVERSE

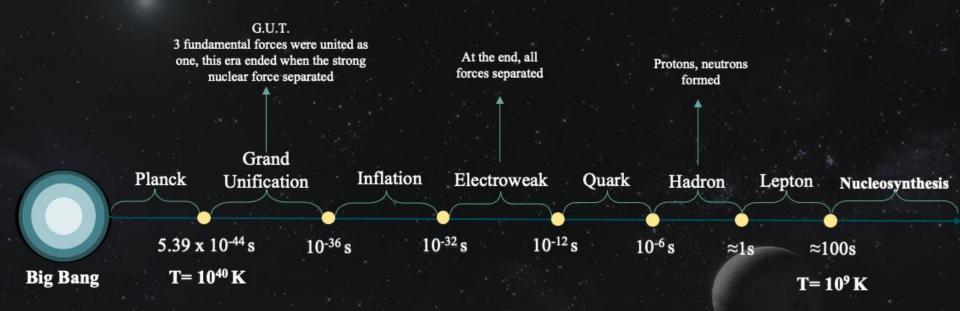


Georges Lemaître 1927 Hypothesis of the Primeval Atom



Edwin Hubble

Radiation Era ≈ 50.000 years



Source: Author's own graph

^{*} The strong nuclear force is responsabile for the strong bond between neutrons and protons in atomic nuclei

^{*} Weak nuclear force is responsible for the radioactive decay processes and the nuclear fusion inside stars

I. PARTICLE BEHAVIOR IN THE EPOCHS

I. ELECTROWEAK

- Quark-gluon plasma
- Large numbers of exotic particles (W, Z, Higgs Bosons)

II. BETWEEN ELECTROWEAK-QUARK

■ Baryogenesis (baryon asymmetry problem—- quarks dominate)

III. QUARK

- Quarks, electrons, neutrinos form
- Quarks not bound in Hadrons!

IV. HADRON

■ Protons, neutrons formed

I. PARTICLE BEHAVIOR IN THE EPOCHS

V. LEPTON

■ Leptogenesis (dominance of leptons over anti-leptons)

VI. NUCLEAR

- Atomic nuclei formed
- Energy dominated by photons

II. FUNDAMENTAL FORCES

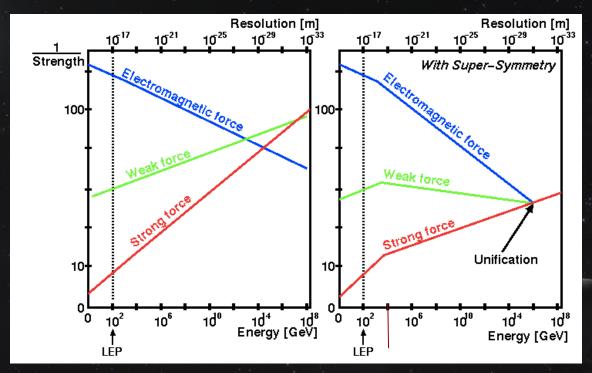


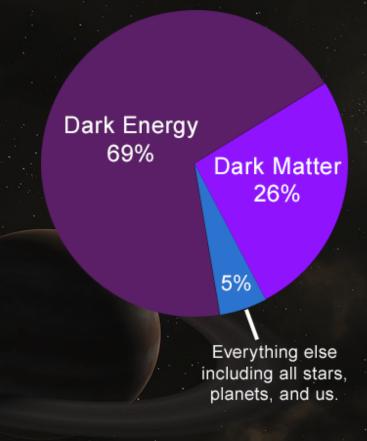
Figure 1: Strength reciprocal vs. energy graph not taking into account SUSY (left) and taking it (right)

Source: CERN, "Particle Physics Education CD-ROM," 1999.

2

ACDM MODEL





How do they work?



Figure 2: Content of the universe

Source: https://chandra.harvard.edu/darkuniverse/

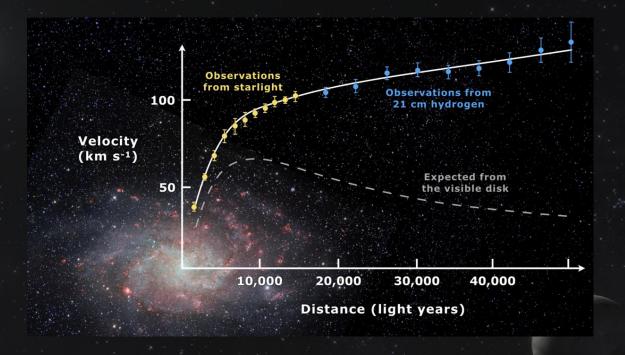


Figure 3: Observed vs expected galaxy velocities Source: https://commons.wikimedia.org/wiki/File:Rotation_curve_of _spiral_galaxy_Messier_33_(Triangulum).png



Vera Rubin



Fritz Zwicki

DARK MATTER POSSIBLE CANDIDATES

I.

WIMPS

(Weakly interacting massive particle)

II.

AXIONS

III.

STERILE NEUTRINOS



Heavy, electromagnetically neutral subatomic particle that is hypothesized to make up most dark matter and therefore some 22% of the universe

Characteristics:

- They don't absorb/emit light
- They don't interact strongly with other particles

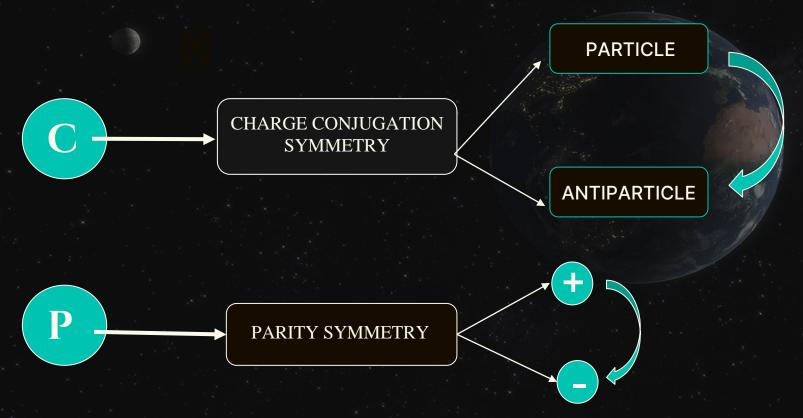
AXIONS

Axions are another well-motivated dark matter candidate. They were proposed to resolve the strong CP problem.

Characteristics:

- No electric charge
- Very small mass
- Very low interaction for strong and weak forces

CP PROBLEM



Source: Author's own graph

STERILE NEUTRINOS

A special kind of neutrino that has been proposed to explain some unexpected experimental results, but they have not been definitively discovered. Scientists are looking hard for them in many different experiments.

Characteristics:

- No electric charge
- Very small mass
- Very low interaction for strong and weak forces

EXPANSION OF THE UNIVERSE



ACDM MODEL OF COSMOLOGY

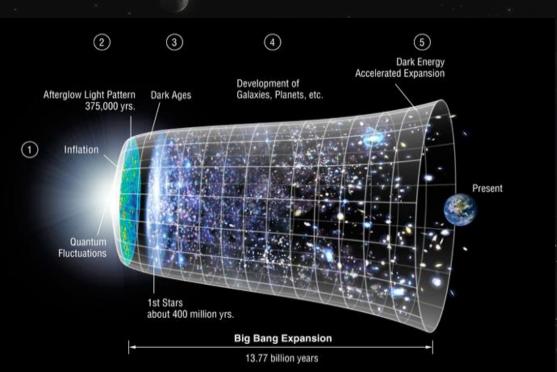


Figure 4: Big Bang Expansion https://lambda.gsfc.nasa.gov/education/graphic_history/univ_evol.html

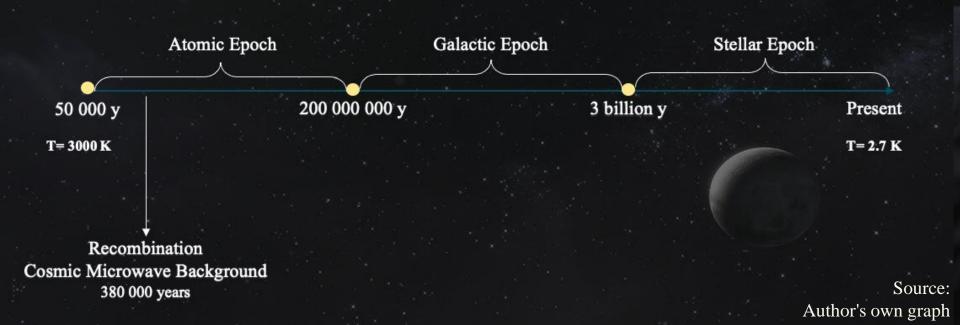
This model proposes specifically cold matter, such as:

- 1. Non baryonic
- 2. Cold
- 3. Dissipationless
- 4. Collisionless

3

COSMIC
MICROWAVE
BACKGROUND

Matter Era – until today



COSMIC MICROWAVE BACKGROUND

- Isotropic photon background with a blackbody spectrum (T=2.725 K)
- Before recombination: hot, dense plasma of particles (photons scattered off, Thomson scattering)
- Proves that photons/baryons **must** have existed in a highly interacting thermal state
- Small irregularities: Random quantum fluctuations before recombination

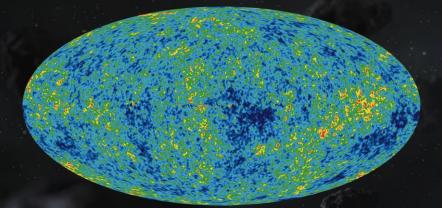


Figure 5: Cosmic Microwave Background (blue spots colder than red spots) Source: *WMAP*

4

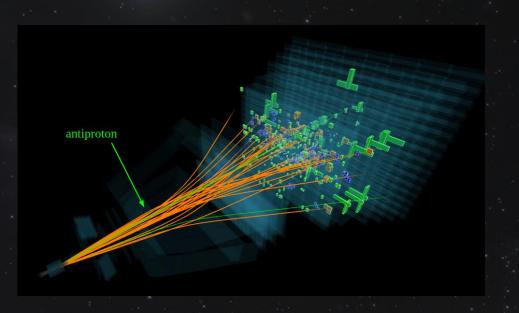
CERN EXPERIMENTS

I. LHCB / LHCF

II. ALICE

OTHERS (ATLAS, CMS, ELENA/AD)

LARGE HADRON COLLIDER BEAUTY LARGE HADRON COLLIDER FORWARD



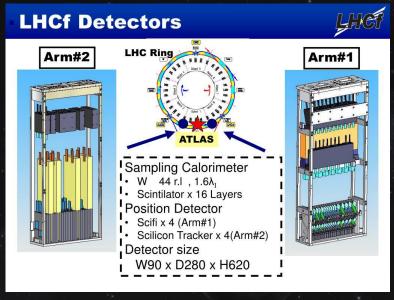
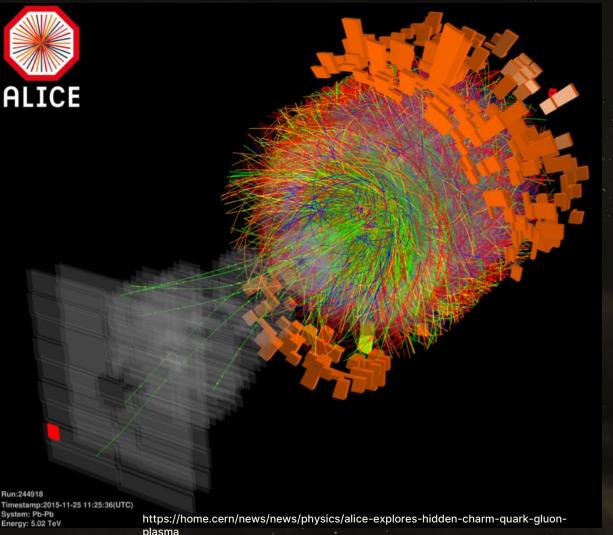


Figure 6: proton-proton collision at LHCb https://home.web.cern.ch/news/news/physics/lhcb-reveals-secret-antimatter-creation-cosmic-collisions

Figure 7: LHCf Detectors https://www.slideserve.com/varick/lhcf-detectors





What is the quark-gluon plasma state?



What happens at ALICE?

RESOURCES

Aretz, S.; Borowski, A.; Schmeling, S. (2017), "Development and Evaluation of a Construct Map for the Understanding of the Expansion of the Universe". *Science Education Review Letters*Auriemma, G. (2014), "LHC, Astrophysics and Cosmology", *Acta Polytechnica CTU Proceedings 1(1): 42–48*

Bambi, C.; Dolgov, A. (2016), "Introduction to Particle Cosmology", *Springer-Verlag Berlin Heidelberg* Bergstrom, L.; Goobar, A. (2004), "Cosmology and Particle Astrophysics", *Springer* Pettini, M.; "Introduction to cosmology", *Cambridge University*

https://cfa.harvard.edu/research/topic/dark-energy-and-dark-matter

https://pages.uoregon.edu/jimbrau/astr123/Notes/Chapter27.html

https://www.osti.gov/servlets/purl/6260191

https://neutrinos.fnal.gov/types/sterile-neutrinos/

https://en.wikipedia.org/wiki/Axion

https://www.astro.caltech.edu/~george/ay20/eaa-wimps-machos.pdf

https://physicsworld.com/a/quintessence/

https://home.web.cern.ch/news/news/physics/lhcb-reveals-secret-antimatter-creation-cosmic-collisions

https://www.slideserve.com/varick/lhcf-detectors

