COSMALLOGY

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MAIN CONNECTIONS BETWEEN COSMOLOGY & PARTICLE PHYSICS

THE EARLY UNIVERSE





COSMIC MICROWAVE BACKGROUND



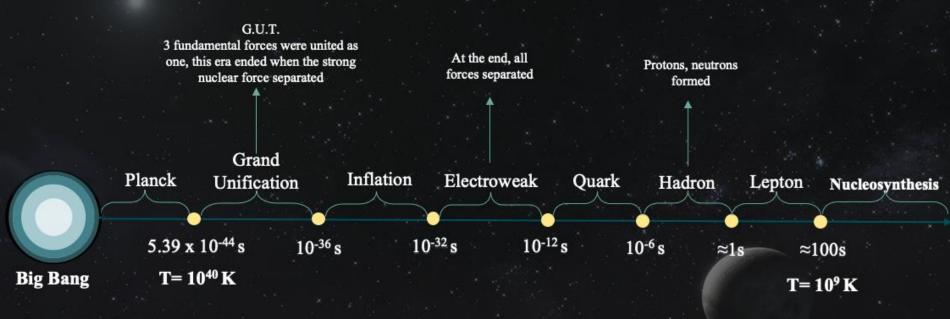
THE EARLY UNIVERSE





Georges Lemaître 1927 Hypothesis of the Primeval Atom Edwin Hubble

Radiation Era ≈ 50.000 years



* The strong nuclear force is responsabile for the strong bond between neutrons and protons in atomic nuclei

* Weak nuclear force is responsible for the radioactive decay processes and the nuclear fusion inside stars

Source: Author's own graph

I. PARTICLE BEHAVIOR IN THE EPOCHS

I. ELECTROWEAK

Quark-gluon plasma

■ Large numbers of exotic particles (W, Z, Higgs Bosons)

II. BETWEEN ELECTROWEAK-QUARK

Baryogenesis (baryon asymmetry problem—- quarks dominate)

III. QUARK

- Quarks, electrons, neutrinos form
- Quarks not bound in Hadrons!

IV. Hadron

Protons, neutrons formed

I. PARTICLE BEHAVIOR IN THE EPOCHS

V. LEPTON

■ Leptogenesis (dominance of leptons over anti-leptons)

VI. nuclear

- Atomic nuclei formed
- Energy dominated by photons

II. FUNDAMENTAL FORCES

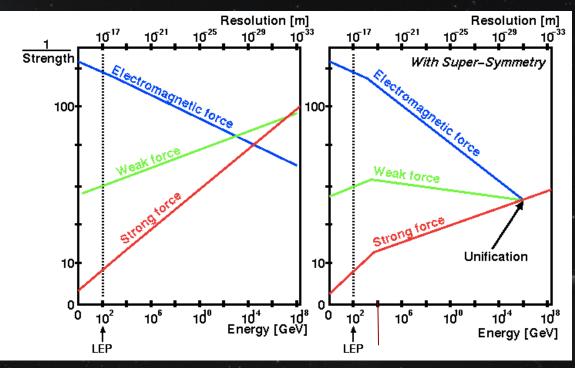


Figure 1: Strength reciprocal vs. energy graph not taking into account SUSY (left) and taking it (right) Source: CERN, "Particle Physics Education CD-ROM," 1999.

V ΛCDM MODEL



What are these particles?

Dark Energy 69%

Dark Matter 26% How do they work?



5%

Everything else including all stars, planets, and us.

Figure 2: Content of the universe Source: <u>https://chandra.harvard.edu/darkuniverse/</u>

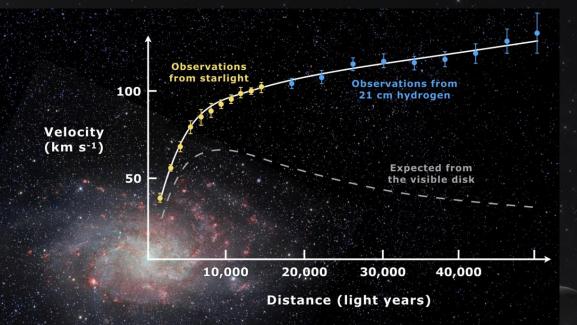


Figure 3: Observed vs expected galaxy velocities Source:https://commons.wikimedia.org/wiki/File:Rotation_curve_of _spiral_galaxy_Messier_33_(Triangulum).png



Vera Rubin



Fritz Zwicki

DARK MATTER **POSSIBLE CANDIDATES**



WIMPS

(Weakly interacting massive particle)



AXIONS

III.

STERILE NEUTRINOS

THE WIMPS

Heavy, electromagnetically neutral subatomic particle that is hypothesized to make up most dark matter and therefore some 22% of the universe

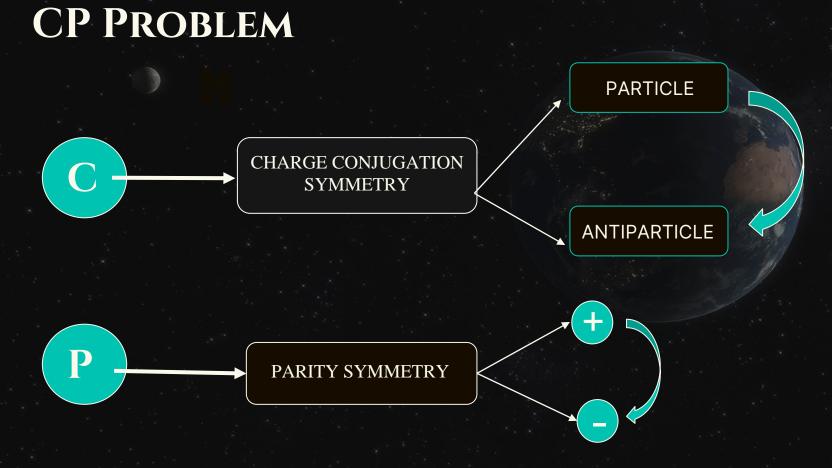
Characteristics:

- They don't absorb/emit light
- They don't interact strongly with other particles

AXIONS

Axions are another well-motivated dark matter candidate. They were proposed to resolve the strong CP problem. Characteristics:

- No electric charge
- Very small mass
- Very low interaction for strong and weak forces



Source: Author's own graph

STERILE

NEUTRINOS

A special kind of neutrino that has been proposed to explain some unexpected experimental results, but they have not been definitively discovered. Scientists are looking hard for them in many different experiments.

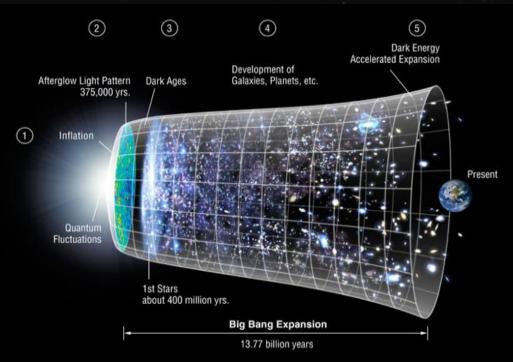
Characteristics:

- No electric charge
- Very small mass
- Very low interaction for strong and weak forces

EXPANSION OF THE UNIVERSE

QUINTESSENCE

ACDM MODEL OF COSMOLOGY



This model proposes specifically cold matter, such as:

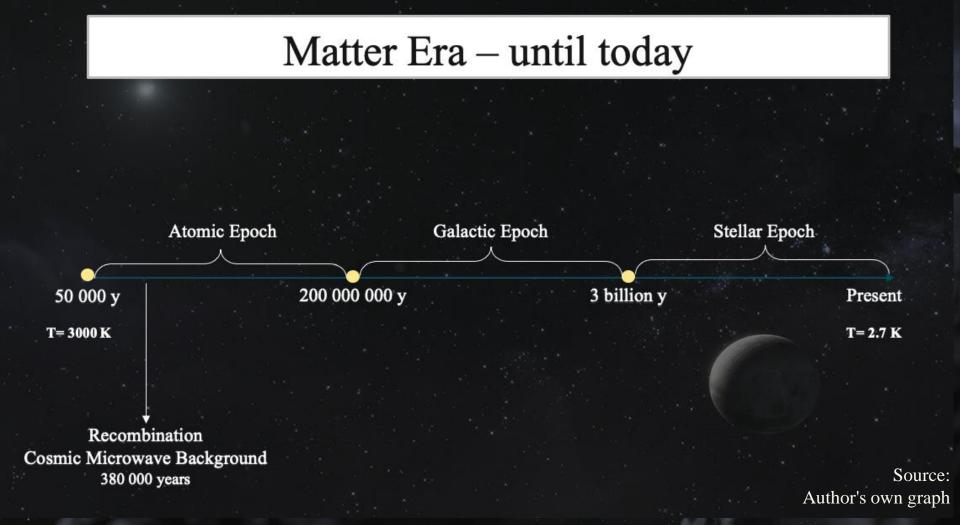
- 1. Non baryonic
- 2. Cold
- 3. Dissipationless
- 4. Collisionless

Figure 4: Big Bang Expansion

 $https://lambda.gsfc.nasa.gov/education/graphic_history/univ_evol.html$

NASA/WMAP Science Team

COSMIC MICROWAVE BACKGROUND



COSMIC MICROWAVE BACKGROUND

- Isotropic photon background with a blackbody spectrum (T=2.725 K)
- Before recombination: hot, dense plasma of particles (photons scattered off, Thomson scattering)
- Proves that photons/baryons **must** have existed in a highly interacting thermal state
- Small irregularities: Random *quantum fluctuations* before recombination

Figure 5: Cosmic Microwave Background (blue spots colder than red spots) Source: *WMAP*

CERN EXPERIMENTS

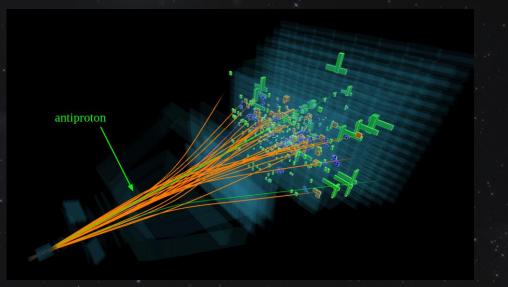
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III. OTHERS (ATLAS, CMS, ELENA/AD)

LARGE HADRON COLLIDER BEAUTY LARGE HADRON COLLIDER FORWARD



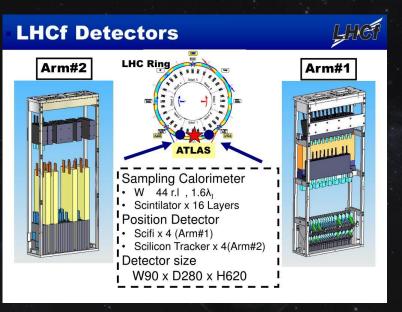
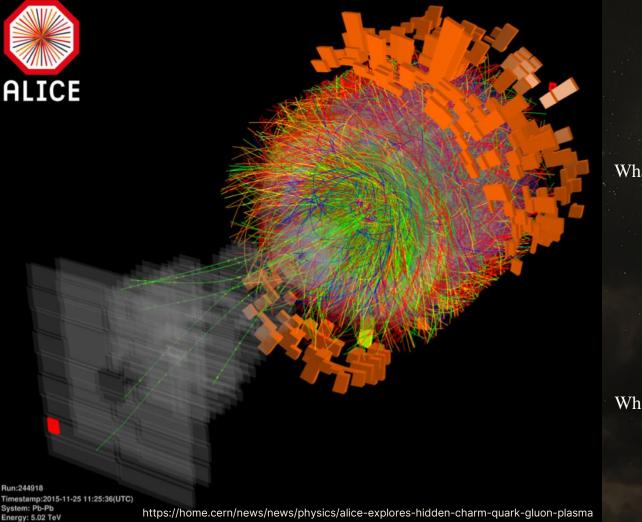


Figure 6: proton-proton collision at LHCb <u>https://home.web.cern.ch/news/news/physics/lhcb-reveals-secret-</u> <u>antimatter-creation-cosmic-collisions</u> **Figure 7**: LHCf Detectors <u>https://www.slideserve.com/varick/lhcf-detectors</u>





What is the quark-gluon plasma state?



What happens at ALICE?

RESOURCES

Aretz, S.; Borowski, A.; Schmeling, S. (2017), "Development and Evaluation of a Construct Map for the Understanding of the Expansion of the Universe". *Science Education Review Letters* Auriemma, G. (2014), "LHC, Astrophysics and Cosmology", *Acta Polytechnica CTU Proceedings 1(1): 42–48*

Bambi, C.; Dolgov, A. (2016), "Introduction to Particle Cosmology", *Springer-Verlag Berlin Heidelberg* Bergstrom, L.; Goobar, A. (2004), "Cosmology and Particle Astrophysics", *Springer* Pettini, M.; "Introduction to cosmology", *Cambridge University*

https://cfa.harvard.edu/research/topic/dark-energy-and-dark-matter https://pages.uoregon.edu/jimbrau/astr123/Notes/Chapter27.html https://www.osti.gov/servlets/purl/6260191 https://neutrinos.fnal.gov/types/sterile-neutrinos/ https://en.wikipedia.org/wiki/Axion https://www.astro.caltech.edu/~george/ay20/eaa-wimps-machos.pdf https://physicsworld.com/a/quintessence/ https://home.web.cern.ch/news/news/physics/lhcb-reveals-secret-antimatter-creation-cosmic-collisions https://www.slideserve.com/varick/lhcf-detectors https://www.physicsoftheuniverse.com/topics_bigbang_timeline.html

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