High precision 209 Bi(n, γ) cross section measurement at nTOF EAR2

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Gen-IV & ADS nuclear systems



Figure adapted from https://www.psi.ch/en/fast



New generation of advanced nuclear power plants (GenIV) and ADS is under design and development aiming to:

- Excel safety and reliability during operation.
- **Sustainability** energy generation and minimization of nuclear waste in long-term.
- Comply with **Non proliferation** agreements.

Development on nuclear technology requires **improvement** on **nuclear data** accuracy:

- Delayed neutron emission probabilities.
- Neutron induced cross sections.



MYRRHA and ²¹⁰Po production



Figure adapted from https://www.sckcen.be/en/expertises-0



MYRRHA is a multipurpose Accelerator Driven System (ADS) nuclear reactor able to operate in subcritical and critical mode. It is made of:

- 600 MeV Proton accelerator.
- Spallation target.
- A core with MOX fuel **cooled** by liquid **lead-bismuth** (Pb-Bi).



Radiological Burden associated with the use of Pb-Bi coolant is because of ²⁰⁹Bi(n,y) populating ²¹⁰Bi and ^{210m}Bi: ²¹⁰Po (t_{1/2}~5 days)/²⁰⁶Tl (t_{1/2}~3My)

Calculations for ²¹⁰Po concentration for MYHRRA in one irradiation ranges from 5 to 20% depending on the evaluation.

²⁰⁹Bi(n, γ) XS and challenges





Figure adapted from L. Fiorito et al. EPJ Nuclear Sci. Technol. (2018) 4, 48



- 5-10% accuracy in RRR covering neutron energies range from thermal up to 35 keV.
- <15% accuracy in the neutron energy range from 35-100 keV.





The nTOF facility @ CERN







Experimental setup @ nTOF EAR2





The experiment took place in **April 2024** at the **nTOF EAR2** using the **state-of-the-art (n,y)** experimental setup:

- Lightweight sample holder design to minimize dead material around the target under study.
 - Reduce background from scattered neutrons
- 9 sTEDs in compact configuration @90° and 4.5 cm from the target position:
 - High sensitivity (n, γ) detection setup.
- 2 C₆D₆ detectors @135° and 17 cm and 20 cm from target:
 - Angular anisotropies + other possible systematics.
- **SiMOn2** : neutron beam monitoring/rel. normalization.

Experimental setup especially well suited for small (n,y) cross sections and/or high radioactive targets



High precision measurement







Preliminary ²⁰⁹Bi(n,y) data







After a very preliminar background subtraction:

- Different neutron resonance shape between thin and thick target data
 - We can account for MS and other systematics thanks to both datasets!
- Expected useful thin and thick datasets up to E_n~30 keV
- For **E**_n>30 keV, only the thick target will be of use.





- ²⁰⁹Bi(n,y) cross section is important for both, nuclear energy applications and nuclear astrophysics:
 - Critical to assess the radiological burden associated to ²¹⁰Po for GenIV and ADS systems cooled by Bi/Pb mixtures.
 - ²⁰⁹Bi(n,y) is **important** for **s-process** and for **U/Th clocks**.
- Measurement very **challenging** because:
 - Small (n,γ).
 - (n,ɣ)/(n,el)<<1.
- Present <u>high precision ²⁰⁹Bi(n,γ) cross section measurement</u> goals:
 - 5-10% uncertainty in RRR En<35 keV.
 - o <15% uncertainty in NRR 35 keV<En< 100 keV.</p>
- It took place in April 2024 at the high luminosity nTOF EAR2:
 - Experimental setup design to enhance sensitivity
 - Several efforts over background estimation and normalization
 - **2 targets** with different **thicknesses** to control MS and Self-Shielding
- Very promising preliminary results from both thin and thick targets:
 - Analysis just started!

Experiment design to overcome previous measurements.



G. de la Fuente phD candidate





Thank you very much for your attention!







• Introduction

• The nTOF facility and the experimental setup

• Very preliminary results

• Summary & conclusions



s-process & Th/U cosmic clocks



Figure adapted from U. Ratzel et al Phys. Rev. C 70 065803 (2004)

From the **astrophysical** point of view, $^{209}Bi(n,y)$ is **important** for understanding the chemical evolution of the Universe via s-process and for U/Th cosmic clocks:

- ²⁰⁹Bi is the last stable s-process nucleus before the α -unstable region feeding to the region above ²⁰⁶Pb recycling heavy material.
- ²⁰⁹Bi(n,y) also contribute to ²⁰⁷TI branching point towards production of ²⁰⁷Pb.
- Th/U ratios, used as cosmic clocks for stars, needs accurate data on 209 Bi(n, χ) to disentangle radiogenic contributions from **Th** and **U** α -chain decays.





XS & neutron reaction yield



There is a catch: the microscopic $\sigma_{\gamma}(En)$ cross-section is not an experimental observable; the reaction yield $Y_{\gamma}(En)$ is:

$$Y(E_n) = Y_0(E_n) + Y_1(E_n) + Y_2(E_n) + \mathcal{O}(\Sigma^4)$$

Y₀= capture Y₁= scattering+capture Y₂=scattering+scattering+capture

$$\mathscr{L}_{n}(E_{n}) = \prod_{i=0}^{n-1} \left[\int_{0}^{l_{max,i}} \Sigma e^{-\Sigma\sigma_{T}(E_{n,i})l_{i}} dl_{i} \int \frac{d\sigma_{nn}(E_{n})}{d\Omega} d\Omega_{i} \right] \int_{0}^{l_{max,n}} \Sigma\sigma_{\gamma}(E_{n,n}) e^{-\Sigma\sigma_{T}(E_{n,n})l_{n}} dl_{n}$$

Is this important?

Yes, for measurements with the following characteristics:

• Small $\sigma_{\gamma}(E_n)$

•
$$\sigma_{\gamma}(E_n)/\sigma_n(E_n) << 1$$

• Thick targets

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4	²⁰⁹ Bi(n,ɣ)	Ì
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²⁰⁹Bi(n, γ) **ORNL** using **C**₆**F**₆ detectors in **1976**

²⁰⁹Bi(n, γ) **nTOF** using **C**₆**D**₆ detectors @ EAR1 in 2006

TABLE I. Capture parameters for resolved resonances in 203 Bi (n, γ) . Many resonances above 30 keV were not fitted individually and those analyzed correspond to those reported in neutron transmission measurements (Refs. 1 and 2).

E_{\circ} (eV)	1	J	Γ_n (meV)	Γ_{γ} (meV)	$g\Gamma_{\gamma}\Gamma_{n}/\Gamma$ (meV
801.6(1)	0	5	4309(145)	33.3(12)	18.2(6)
2323.8(6)	0	4	17888(333)	26.8(17)	12.0(8)
3350.83(4)	1	5	87(9)	18.2(3)	9.5(2)
4458.74(2)	1	5	173(13)	23.2(22)	11.3(11)
5114.0(3)	0	5	5640(270)	65(2)	35.3(11)
6288.59(2)	1	4	116(18)	17.0(17)	6.7(7)
6525.0(3)	1	3	957(100)	25.3(14)	8.6(5)
9016.8(4)	1	6	408(77)	21.1(14)	13.0(9)
9159.20(7)	1	5	259(45)	21.4(21)	10.9(11)
9718.910(1)	1	4	104(22)	74(7)	19.5(21)
9767.2(3)	1	3	900(114)	90(8)	28.7(26)
12098					65(4) ^c
15649.8(1.0)	1	5	1000	47(4)	20.2(17)
7440.0(1.3)	1	6	1538(300)	32(3)	20.4(18)
17839.5(9)	1	5	464(181)	43(4)	21.7(20)
20870	1	5	954(227)	34.4(33)	18.3(17)
21050	1	4	7444(778)	33(3)	14.8(13)
22286.0(9)	1	5	181(91)	33.6(32)	15.1(15)
23149.1(1.3)	1	6	208(154)	25.3(25)	14.7(15)

E	Cnergy (keV) ^a	$g\Gamma_n\Gamma_\gamma/\Gamma$ (meV)	Energy (keV) ^a	$g\Gamma_n\Gamma_\gamma/\Gamma$ (meV)
	0.8000 ^b	25.0± 3.0 ^b	27.45 ^e	156.5 ± 21.6
	2.310 ^b	20.0 ± 10.0 °	28.79	11.5 ± 2.0
	3.351	10.9 ± 0.1^{d}	29.01	8.9± 1.9
	4,458	10.8± 0.2	29,20	12.7 ± 2.2
	5.113	40.8± 0.6	29,52	10.5 ± 2.0
	6.289	6.4± 0.2	30.48	18.0 ± 4.0
	6.527	9.1 ± 0.2	32.73	5.1 ± 2.7
	9.018	10.6± 0.4	32.90	16.0 ± 6.5
	9.159	9.6± 0.4	33,31	243.0 ± 18.0
	9.375	0.8± 0.3	34.68	29.0 ± 4.0
	9.718	21.7± 0.6	37.25	11.2 ± 3.9
	9.766	21.2 ± 0.6	38.10	38.1± 3.9
	12.09	7.7± 0.9	39.17	22.5 ± 4.1
<u> </u>	12.10	49.0± 7.8	42,40	19,2 ± 4,5
	12,24	3.6 ± 0.7	43.60	13.7 ± 4.0
1	14.88	10,1 ± 0,9	44.09	5.1 ± 3.8
	15,51	68.0± 9.3	44.60	30.2 ± 4.2
	15.65	21.8 ± 1.3	45,18	28.8± 5.0
	17.44	16.7 ± 1.0	45.56	129.0 ± 12.2
	17.83	21.7 ± 1.0	46.49	17.8 ± 5.0
	20.86	12.8± 0.9	49.85	28.6± 5.2
	21.06	15.8 ± 1.2	51.74	8.4 ± 5.4
	22.27	18.2 ± 1.1	52.77	5.1 ± 5.6
	23.13	11.2 ± 1.1	53.70	33.2 ± 7.3
	23.85	5.0 ± 1.1	54,22	23.4 ± 5.8
_!	24,20	5.0 ± 1.1	55,42	15.1 ± 5.8
	25.27	30.7± 2.2	57.76	5.1 ± 5.8
	27.05	70.4± 3.7	61.57	64.4 ± 9.1
- 1	27.29	12.8 ± 2.2	69.14	216.0 ± 25.1
			R	Macklin Phys. Rev. C. 14 4 (1976)

Might be affected by **systematics** associated with experimental **W.F.** and **n-sensitivity** (C_cF_c)

Goals of the present $^{209}Bi(n, \gamma)$ measurement:

- 5-10% accuracy in RRR covering neutron energies range from thermal up to 35 keV.
- <15% accuracy in the neutron energy range from 35-100 keV.