

# Viability of Boosted Light Dark Matter in a Two-Component Scenario

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IIT Hyderabad, University of Hyderabad**

**Reference:**

“Viability of Boosted Light Dark Matter in a Two-Component Scenario” , A Basu , A Chakraborty, N. Kumar, and S. Sadhukhan (<https://arxiv.org/abs/2310.09349>)

## Motivation:

- The existence of Dark Matter (DM) is proven only through indirect gravitational probes.
- Astrophysical observation predicts the amount of DM (  $\sim 26.8\%$  ) of the total energy of the Universe.
- A Plethora of DM direct and indirect detections for WIMP are only of minimal success.
- We are studying **light DM (MeV-GeV scale)**.
- The light DM can receive sufficient energy for the nuclear recoil if it is **boosted**.
- Detection prospects get better with the **boost**.

# Model Description:

Neutrinophilic Two Higgs doublet model **v2HDM**.  $\Phi_1$  and  $\phi_2$

Neutrinophilic suggests a strong coupling between one doublet and  $N_R$ .

Yukawa interaction of the model

$$\mathcal{L}_Y = Y_{\alpha\beta}^d \bar{Q}_{L,\alpha} \Phi_1 d_{R,\beta} + Y_{\alpha\beta}^u \bar{Q}_{L,\alpha} \tilde{\Phi}_1 u_{R,\beta} + Y_{\alpha\beta}^l \bar{L}_{L,\alpha} \Phi_1 l_{R,\beta} + Y_{\alpha\beta}^\nu \bar{L}_{L,\alpha} \tilde{\Phi}_2 N_{R,\beta} + h.c.$$

$N_R$  is the RHN. Odd under  $Z_2$ .

Scalar singlet  $\phi_3$  gives MeV scale scalar DM

$$V_{DM} = \frac{1}{2} m_{\phi_3}^2 \phi_3^2 + \frac{\lambda_{\phi_3}}{4!} \phi_3^4 + \kappa_1 \Phi_1^\dagger \Phi_1 \phi_3^2 + \kappa_2 \Phi_2^\dagger \Phi_2 \phi_3^2.$$

Vector-like doublet  $\mathbf{N}$   
Vector-like singlet  $\chi$

$$L_{VLL} = m_N \bar{\mathbf{N}} \mathbf{N} + m_\chi \bar{\chi} \chi + y_N \bar{\mathbf{N}} \tilde{\Phi}_1 \chi + h.c.$$

arxiv: 0711.4022  
0802.4353  
1011.6188

# Model Description: Fields & Charge Assignments:

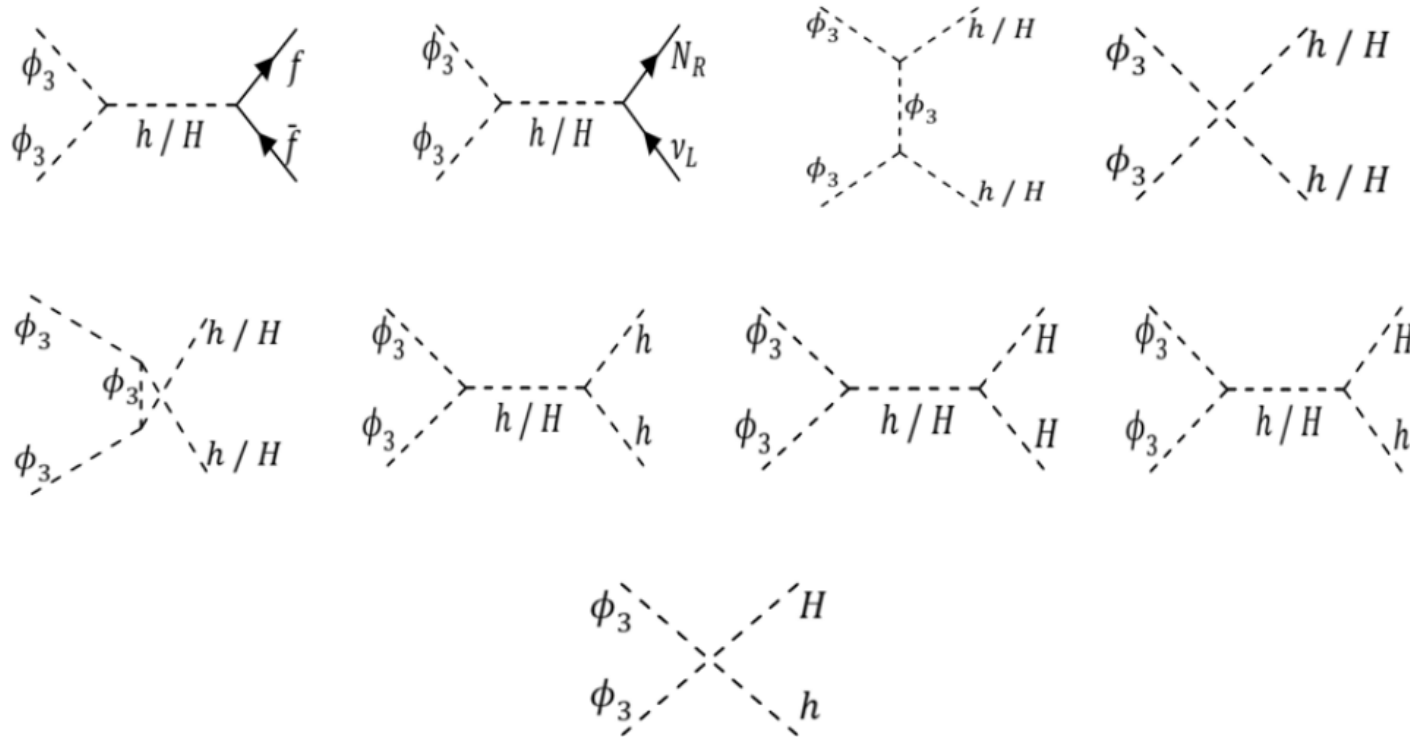
| Particle Name           | $SU(2)_L$ Charges | $U(1)_Y$ Charges | $Z_2$ Charges | $Z_2^{\text{DM}}$ Charges |
|-------------------------|-------------------|------------------|---------------|---------------------------|
| <b>Scalar Fields</b>    |                   |                  |               |                           |
| $\Phi_1$                | 2                 | 1                | 1             | 1                         |
| $\Phi_2$                | 2                 | 1                | -1            | 1                         |
| $\phi_3$                | 1                 | 0                | 1             | -1                        |
| <b>Fermionic Fields</b> |                   |                  |               |                           |
| $N$                     | 2                 | -1               | 1             | -1                        |
| $\chi$                  | 1                 | 0                | 1             | -1                        |

- $\Phi_3$  is the scalar MeV scale DM.
- Mixing of  $N$  and  $\chi$  produces the heavy fermionic DM  $\chi_1$ .  $m_{\chi_1} \sim 100 \text{ GeV}$ .
- $\Phi_2$  is even under  $Z_2^{\text{DM}}$ -> suggests that it provides a portal interaction and does not serve as a DM candidate.
- Odd under  $Z_2^{\text{DM}}$ -> suggests stable DM candidate.

# Dark Matter: Relic Density Aspects: MeV scale Scalar Singlet DM:

## Feynman Diagrams for the scalar DM annihilation

Scalar DM  $\phi_3$ :



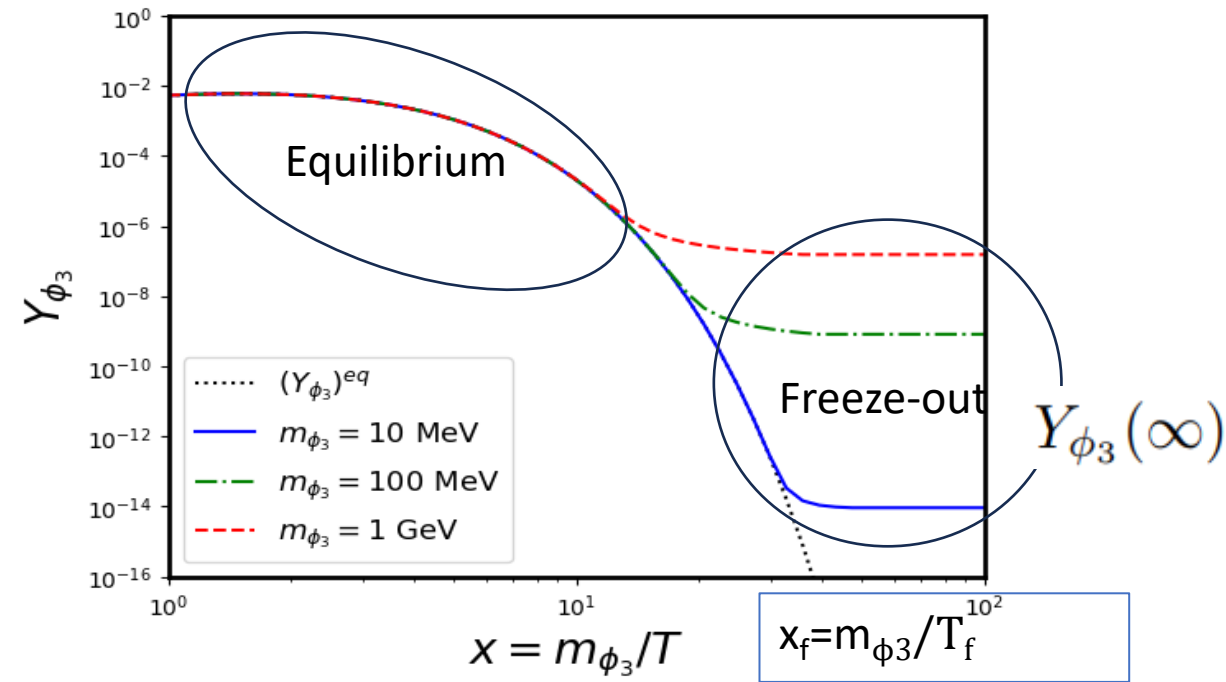
$$\sigma_{\phi_3} = \sigma(s)_{\phi_3\phi_3 \rightarrow f\bar{f}} + \sigma(s)_{\phi_3\phi_3 \rightarrow N_R\nu_L} + \sigma(s)_{\phi_3\phi_3 \rightarrow HH} + \sigma(s)_{\phi_3\phi_3 \rightarrow hh} + \sigma(s)_{\phi_3\phi_3 \rightarrow hH}$$

# Dark Matter: Relic Density Aspects: Scalar Singlet DM:

$$\frac{dY_{\phi_3}}{dx} = -\frac{1}{x^2} \frac{s(m_{\phi_3})}{H(m_{\phi_3})} \langle \sigma v \rangle_{\phi_3 \phi_3 \rightarrow SM} (Y_{\phi_3}^2 - Y_{\phi_3,eq}^2)$$

$$Y_{\phi_3,eq} = 0.145 \frac{g_i}{g^*} x^{3/2} e^{-x}$$

$$\Omega_{\phi_3} h^2 = \frac{m_{\phi_3} s_0 Y_{\phi_3}(\infty)}{\rho_c / h^2}$$



| Scalar DM |                       |
|-----------|-----------------------|
| Mass      | Relic density         |
| 10 MeV    | $2.43 \times 10^{-8}$ |
| 100 MeV   | $2.10 \times 10^{-2}$ |
| 1 GeV     | 36.9                  |

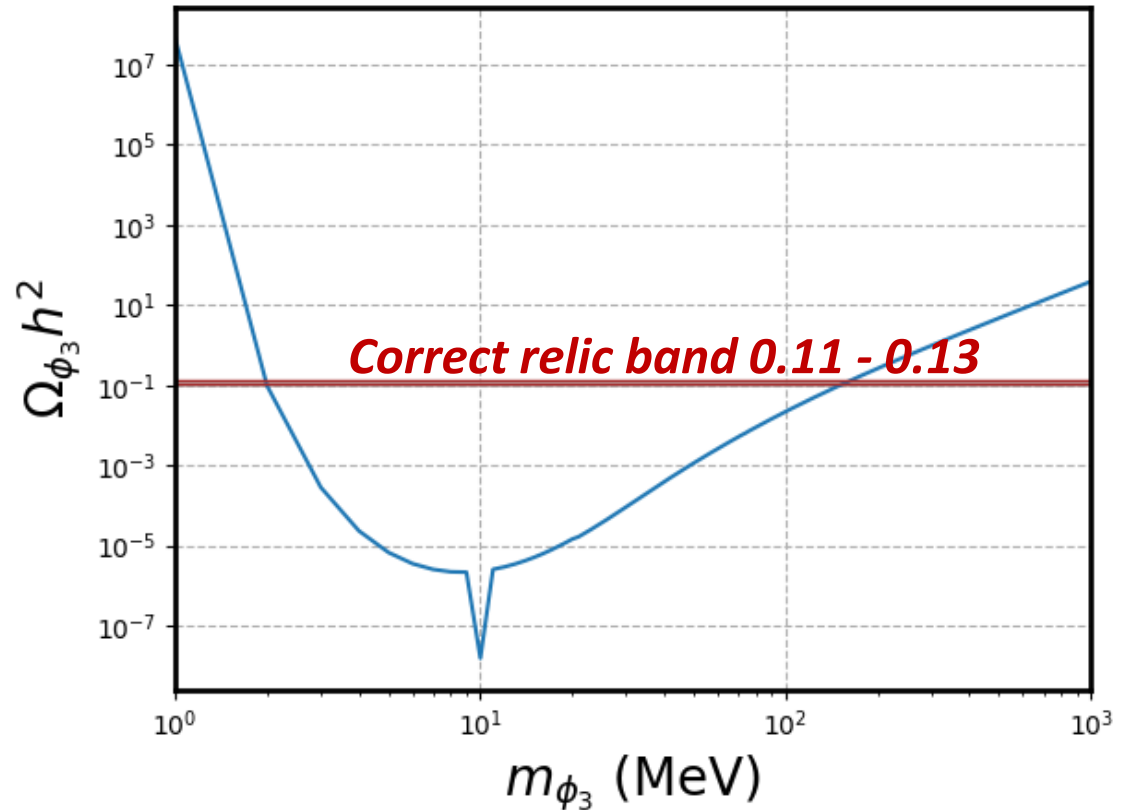
# Dark Matter: Relic Density Aspects: Scalar Singlet DM:

$$\Omega h^2 = \frac{2.14 \times 10^9 \text{ GeV}^{-1}}{\sqrt{g_*} M_{pl}} \frac{1}{J(x_f)}$$

$$J(x_f) = \int_{x_f}^{\infty} \frac{\langle \sigma v \rangle(x)}{x^2} dx$$

$\langle \sigma v \rangle$  is the thermal averaged cross-section

References: <https://arxiv.org/pdf/0804.2741>  
<https://arxiv.org/pdf/2006.09721>  
<https://arxiv.org/pdf/1808.01272>

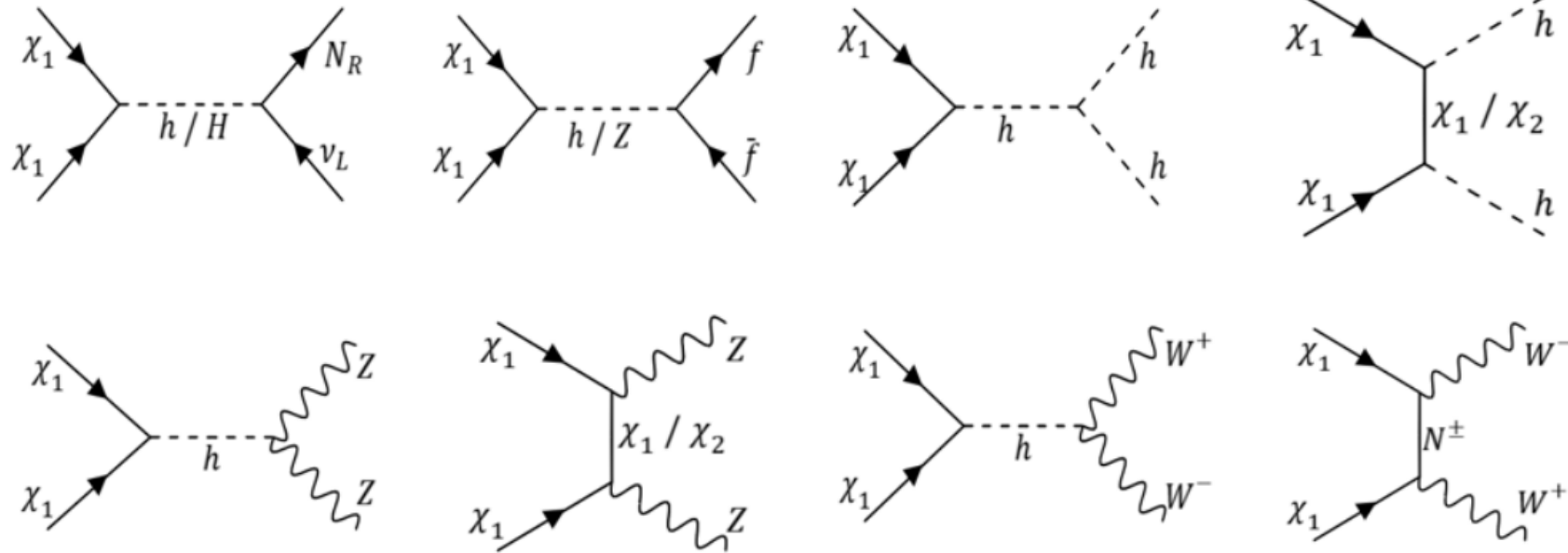


**Allowed window 2-120 MeV**

# Dark Matter: Relic Density Aspects: Fermionic DM:

## Feynman Diagrams for the fermionic DM annihilation

Fermionic DM  $\chi_1$ :



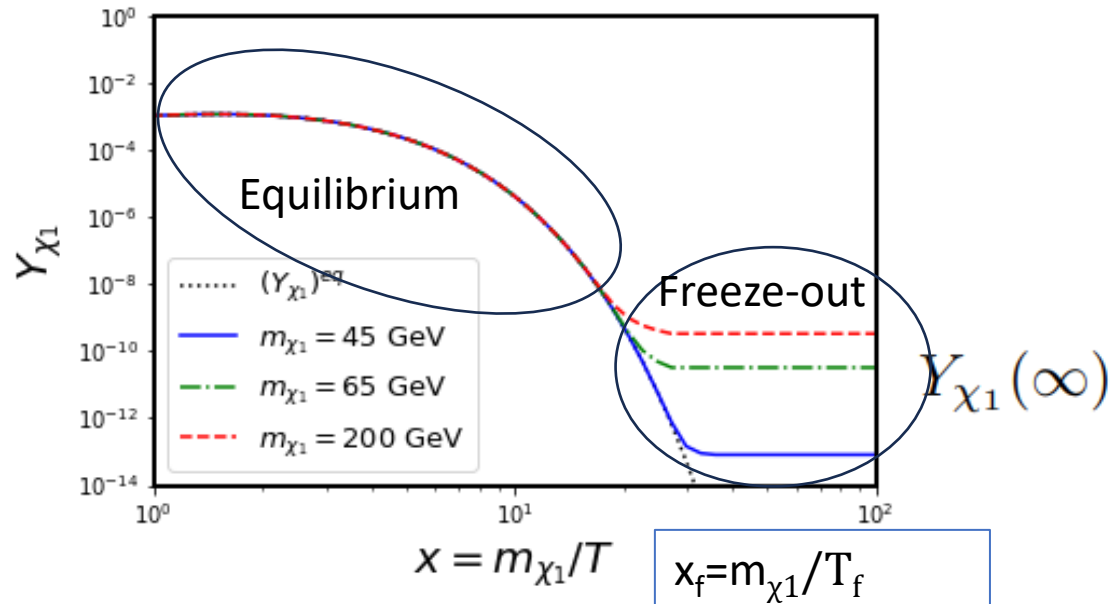
$$\begin{aligned} \sigma_{\chi_1} = & \sigma(s)_{\chi_1\chi_1 \rightarrow f\bar{f}} + \sigma(s)_{\chi_1\chi_1 \rightarrow N_R\nu_L} + \sigma(s)_{\chi_1\chi_1 \rightarrow hh} + \sigma(s)_{\chi_1\chi_1 \rightarrow ZZ} \\ & + \sigma(s)_{\chi_1\chi_1 \rightarrow W^+W^-} \end{aligned}$$



# Dark Matter: Relic Density Aspects: Fermionic DM:

$$\frac{dY_{\chi_1}}{dx} = -\frac{1}{2} \frac{1}{x^2} \frac{s(m_{\chi_1})}{H(m_{\chi_1})} \langle \sigma v \rangle_{\chi_1 \chi_1 \rightarrow SM} (Y_{\chi_1}^2 - Y_{\chi_1,eq}^2)$$

$$Y_{\chi_1,eq} = 0.145 \frac{g_i}{g^*} x^{3/2} e^{-x}$$

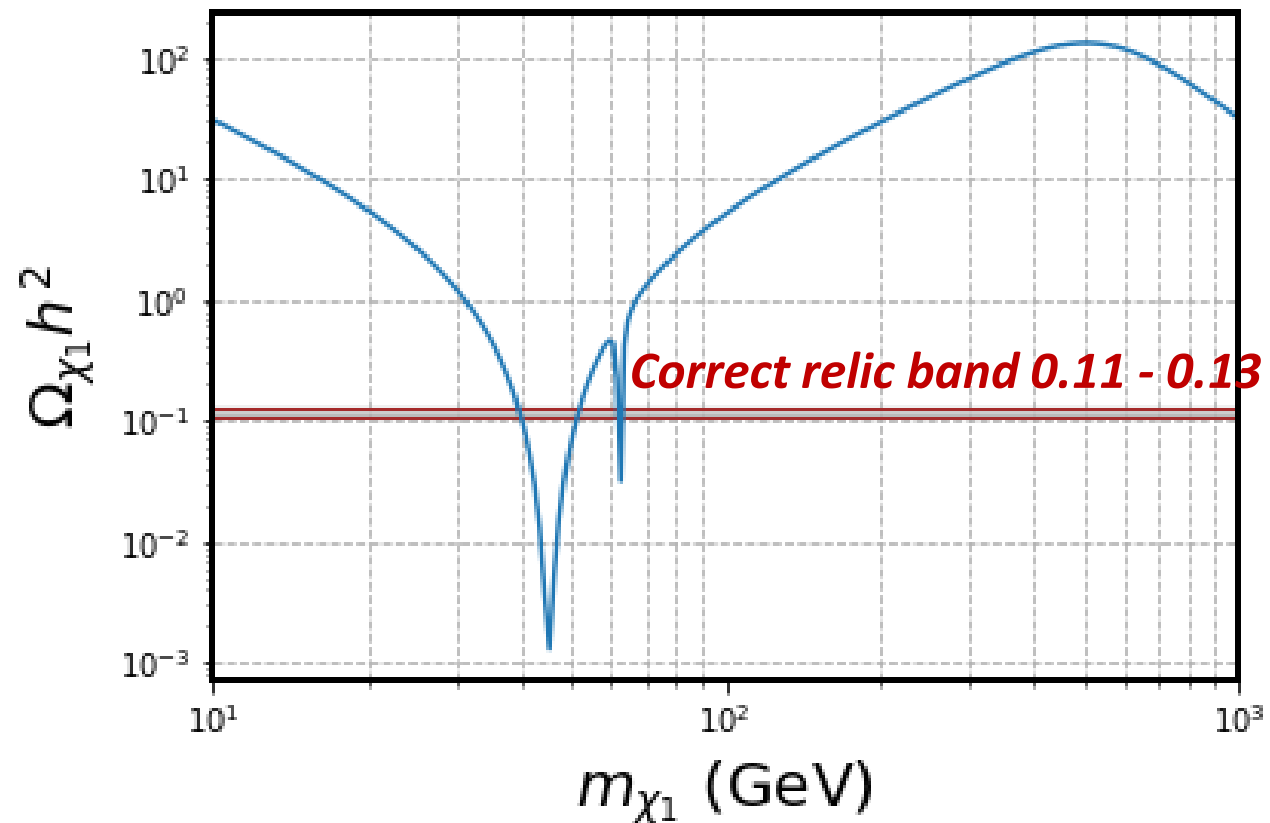


$$\Omega_{\chi_1} h^2 = \frac{m_{\chi_1} s_0 Y_{\chi_1}(\infty)}{\rho_c/h^2}$$

| Fermion DM |                       |
|------------|-----------------------|
| Mass       | Relic density         |
| 45 GeV     | $2.07 \times 10^{-3}$ |
| 65 GeV     | 0.67                  |
| 200 GeV    | 22.16                 |

# Dark Matter: Relic Density Aspects: Fermionic DM:

$$\Omega_{\chi_1} h^2 = \frac{2.14 \times 10^9 x_f}{\sqrt{g^*} M_{pl} \langle \sigma v \rangle}$$

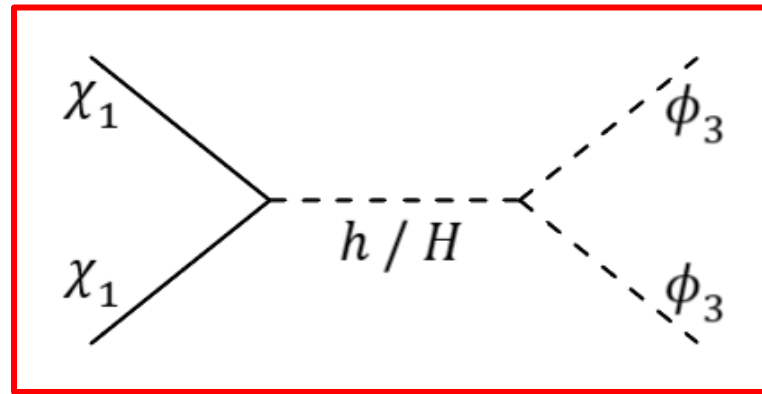


References: <https://arxiv.org/pdf/1812.06505>  
<https://arxiv.org/pdf/1510.02760>

**Allowed window 40 GeV – 50 GeV.**

# Two-component DM: Coupled Boltzmann Equation :

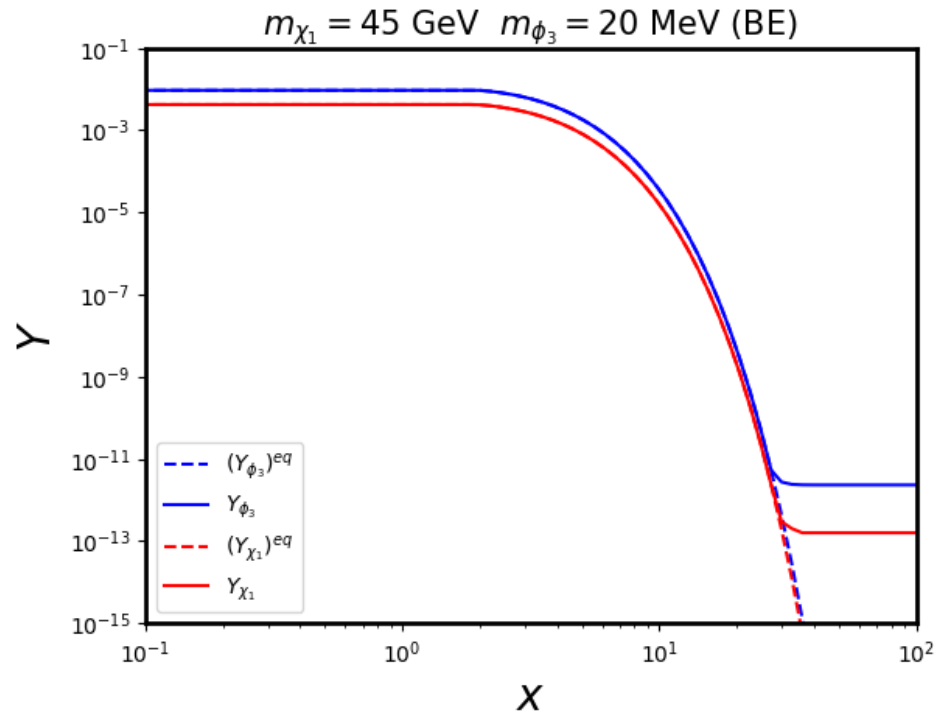
Higgs portal interaction between the two DM candidates serves as the **connector between the two DM candidate**



$$\frac{dY_{\chi_1}}{dx} = -\frac{1}{2} \frac{\lambda_{\chi\phi}}{x^2} \left( Y_{\chi_1}^2 - Y_{\phi_3}^2 \frac{Y_{\chi_1,eq}^2}{Y_{\phi_3,eq}^2} \right) - \frac{1}{2} \frac{\lambda_{\chi}}{x^2} (Y_{\chi_1}^2 - Y_{\chi_1,eq}^2)$$
$$\frac{dY_{\phi_3}}{dx} = -\frac{\lambda_{\phi}}{x^2} (Y_{\phi_3}^2 - Y_{\phi_3,eq}^2) + \frac{\lambda_{\chi\phi}}{x^2} \left( Y_{\chi_1}^2 - Y_{\phi_3}^2 \frac{Y_{\chi_1,eq}^2}{Y_{\phi_3,eq}^2} \right)$$

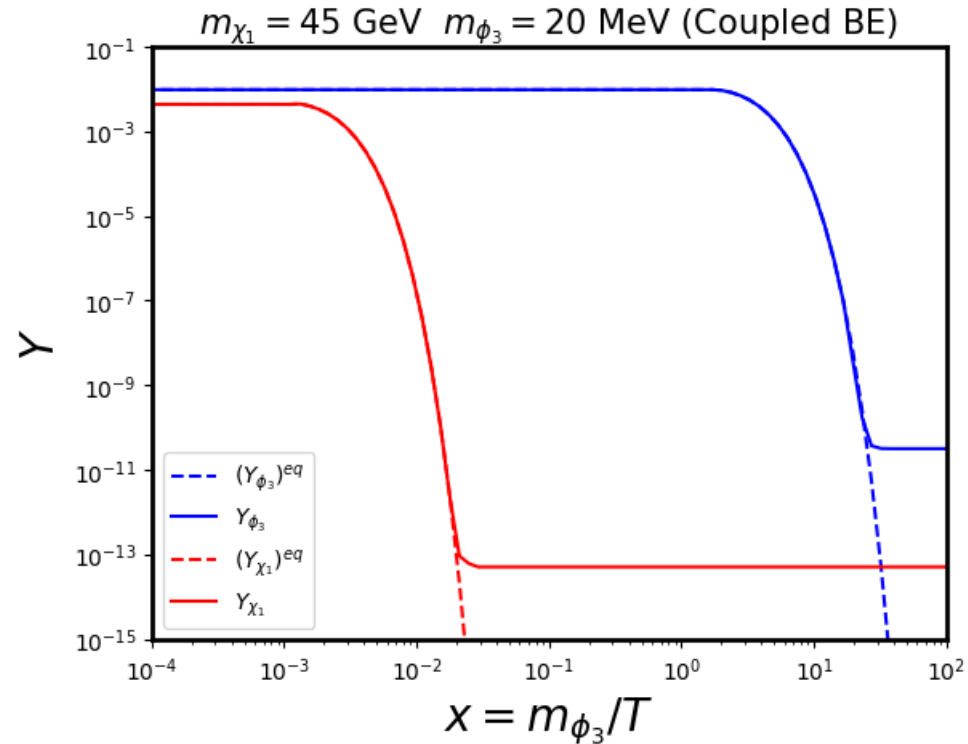
# Two-component DM: Coupled Boltzmann Equation :

$$m_{\chi_1} = 45 \text{ GeV} \quad m_{\phi_3} = 20 \text{ MeV}$$



$$\Omega_{\chi_1} h^2 = 2.07 * 10^{-3}$$

$$\Omega_{\phi_3} h^2 = 5.50 * 10^{-6}$$

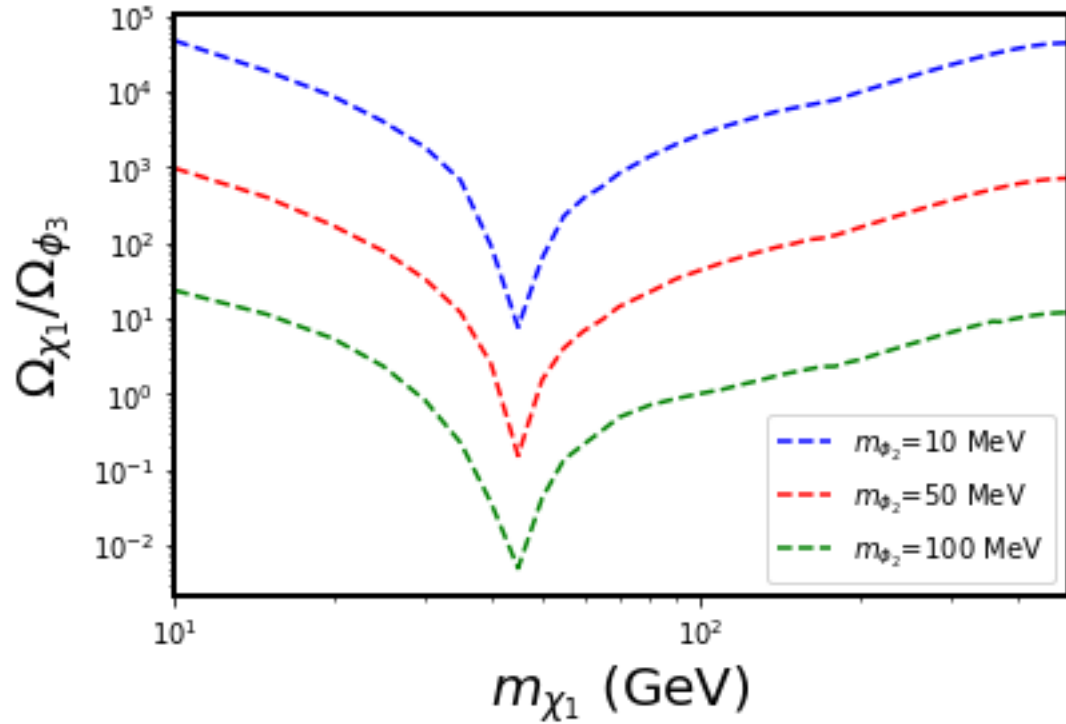


$$\Omega_{\chi_1} h^2 = 6.2 * 10^{-4}$$

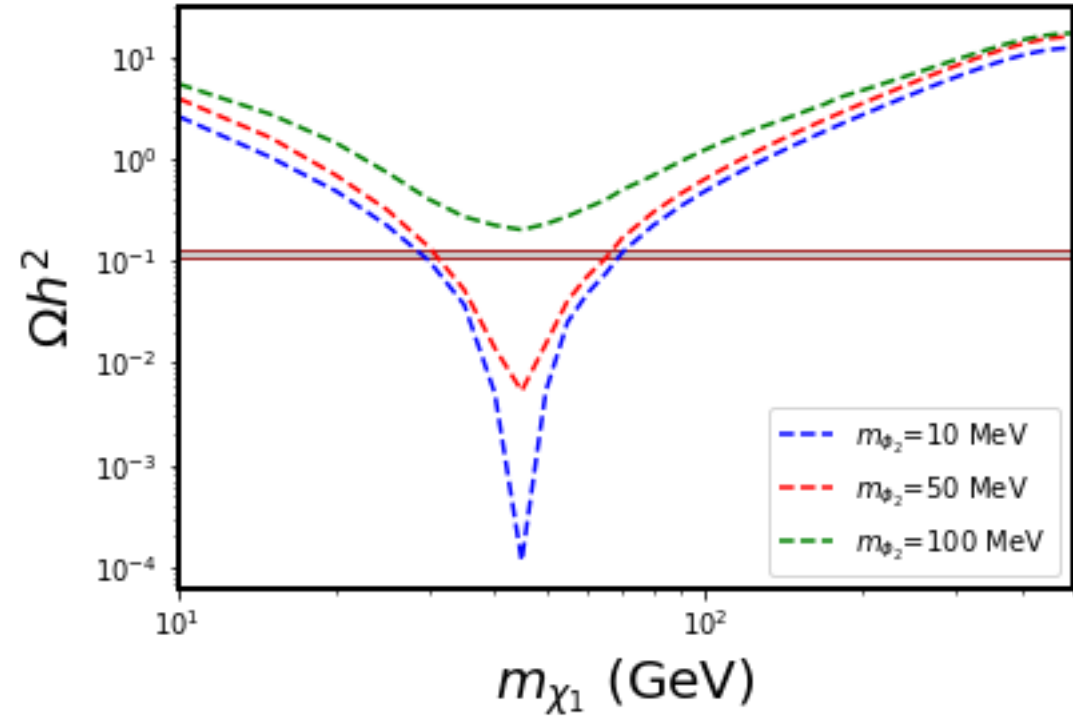
$$\Omega_{\phi_3} h^2 = 1.2 * 10^{-4}$$

$$\Omega h^2 = \Omega_{\chi_1} h^2 + \Omega_{\phi_3} h^2 = 7.4 * 10^{-4}$$

# Two-component DM: Coupled Boltzmann Equation :

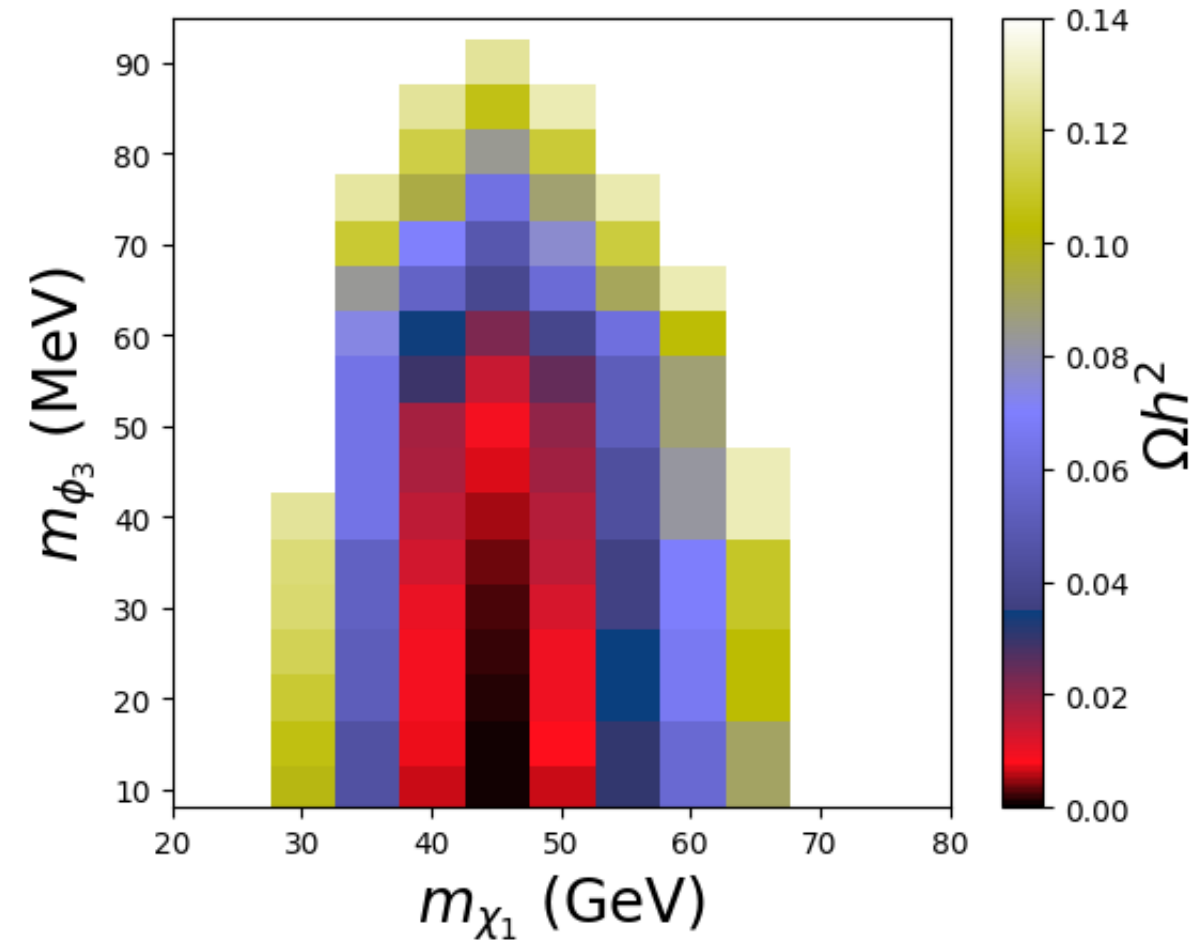


The Scalar DM  $\phi_3$  with 50 MeV mass, contributes  $>50\%$  of total DM.



Allowed mass window for fermionic DM  $\chi_1$  increased up to 30 GeV – 70 GeV.

# Two-component DM: Coupled Boltzmann Equation :



- $\Omega h^2 = 0.1199 \pm 3 \cdot 0.0027$  ( $3\sigma$  upper limit  $\lesssim 0.13$  for the relic density measured by the Planck experiment) is used.
- The yellowish-green makes a boundary on the  $m_{\chi_1} - m_{\phi_3}$  plane for the under-abundant relic density.
- The white region is ruled out due to the over-abundance of the relic density.

# Boosted DM:

$$s = 4m_{\phi_3}^2 + 4m_{\phi_3}^2 \left(1 - \frac{1}{\gamma_{\phi_3}^2}\right)$$

$$\gamma_{\phi_3} \sim \frac{m_{\chi_1}}{m_{\phi_3}}$$

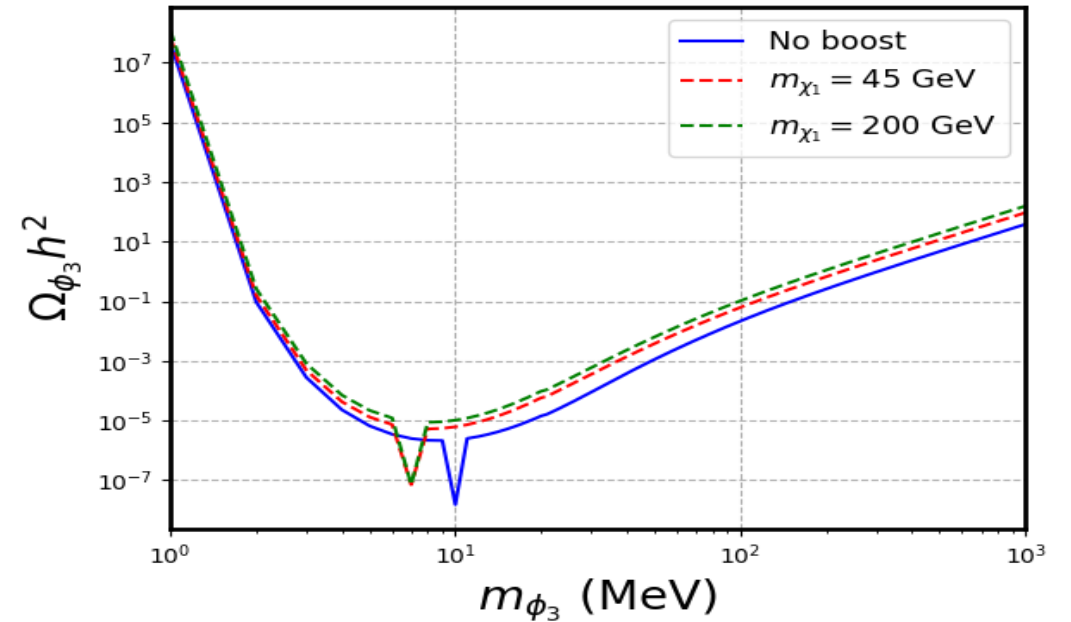
No boost case  $\rightarrow \gamma_{\phi_3} = 1$

resonance condition  $s = m_H^2$

or,  $m_{\phi_3} = m_H/2$

$m_H = 20$  MeV is taken as a benchmark

Resonance is due to the H-dominated “s” channel annihilation.



With boost case, for say  $m_{\chi_1} = 45$  GeV

The COM energy  $s \approx 8m_{\phi_3}^2$

Thus, the resonance condition  $s = m_H^2$

gives,  $m_{\phi_3} = m_H/(2\sqrt{2})$

# Conclusion:

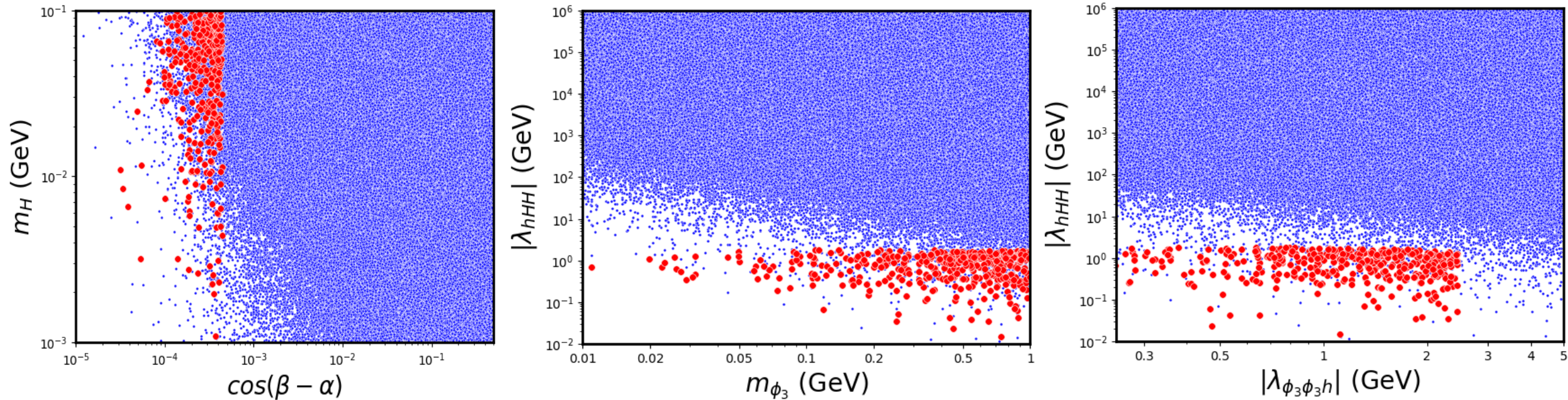
- **Goal:** Relic aspects study of Two-Component DM (Depending on a particular BSM Model).
- **What we are doing:** The DM-DM interaction is allowed through the Higgs portal, which leads to a two-component model. We showed that the relic aspects change in the **Boosted** case due to the DM-DM interaction.
- **Final Results:** We got a mass range for the two DM components and can comment that for some particular choice of masses, the **light scalar DM contributes >50% of total DM.**

# Thank You



# Theoretical constraints and Benchmark Points :

- Theoretical constraints such as stability of vacuum and tree-level perturbative unitarity.
- The Higgs invisible bound. SM Higgs boson ( $h$ ) (125 GeV) invisible BR is  $<10\%$ . [arxiv: 2303.01214](#)
- Z invisible bound.



We chose  $m_H = 20$  MeV. Also checked other values (50,100) MeV. No such variation in relic density observed, other than the shift of resonance drop.

# Parameter Space :

The chosen parameter space for our work is,

| Benchmark point |         |          |                      |            |            |         |           |
|-----------------|---------|----------|----------------------|------------|------------|---------|-----------|
| $m_H$           | $m_h$   | $\alpha$ | $\tan \beta$         | $\kappa_1$ | $\kappa_2$ | $Y_\nu$ | $m_{N_R}$ |
| 20 MeV          | 125 GeV | 89.998   | $1.8 \times 10^{-4}$ | 0.01       | 0.002      | 1       | 10 MeV    |

| A few selected couplings |                          |                          |                                    |                                    |                         |                         |
|--------------------------|--------------------------|--------------------------|------------------------------------|------------------------------------|-------------------------|-------------------------|
| $\lambda_{hHH}$<br>(GeV) | $\lambda_{hhH}$<br>(GeV) | $\lambda_{hhh}$<br>(GeV) | $\lambda_{\phi_3\phi_3h}$<br>(GeV) | $\lambda_{\phi_3\phi_3H}$<br>(GeV) | $\lambda_{\nu_L N_R h}$ | $\lambda_{\nu_L N_R H}$ |
| -1.04                    | 0.001                    | 63.5                     | 2.45                               | $3.3 \times 10^{-4}$               | $-2.83 \times 10^{-5}$  | 0.70                    |