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Latest Three-Flavor Neutrino Oscillation Results from NOvA

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On behalf of the NOvA Collaboration

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Neutrino Mixing and Oscillations

- Neutrinos are fundamental particles
 *** tiny non-zero masses
 - * comes in three flavors ν_e , ν_μ , ν_τ
- Flavor eigenstates are mixed with the mass eigenstates by a unitary matrix
- Neutrinos oscillate between flavor eigenstates
- The oscillation probability is given by

$$P(\nu_{\alpha} \to \nu_{\beta}) = \left| \sum_{i} U_{\alpha i}^{*} e^{-i \frac{m_{i}^{2}L}{2E}} U_{\beta i} \right|^{2}$$

• Neutrino oscillations let us probe the elements of the mixing matrix



Sensitive to Δm_{32}^2 and mixing angle θ_{23} (resolution of octant degeneracy)



Sensitive to the CP-violating phase δ_{CP} , mixing angle θ_{13} (although mild) and mass ordering

The NOvA Experiment

- NuMI Off-axis ν_e Appearance Experiment
 - NuMI: Neutrinos at the Main Injector
 - Off-axis: Detectors situated 14.6 mrad off-axis to beam direction
 - $\nu_{\mu}(\bar{\nu}_{\mu})$ dis-appearance and $\nu_{e}(\bar{\nu}_{e})$ appearance
 - Functionally identical liquid scintillation detectors. ND and FD located at 1km and 810km from the beam source

• Primary Goals:

- Measure neutrino oscillation parameters
- Resolve neutrino mass ordering
- Resolve octant degeneracy
- Measure δ_{CP} , the CP-violating phase



Check out Prof. Bipul Bhuyan's talk from Plenary-I for the details

Beyond Neutrino Oscillations

- Non-standard interactions
- Neutrino cross-sections
- Sterile neutrinos
- Magnetic monopoles
- Dark matter
- And many more!

The NuMI beam line at Fermilab provides an intense $\nu/\bar{\nu}$ beam

How to Measure Neutrino Oscillations?

Observe how flavor changes with energy over a long distance, while mitigating uncertainties related to neutrino flux, interaction cross sections, and detector performance.



Credits: Alex Himmel

Near Detector Spectra



- The observed un-oscillated $\nu_{\mu}/\bar{\nu}_{\mu}$ candidates at the Near Detector
- We use this sample in predicting both the ν_{μ} and ν_{e} signal events at the Far Detector

Near Detector Spectra



- Dominant background: beam $\nu_e/\bar{\nu}_e$ events
- We use these samples in predicting the background events for the ν_e appearance analysis



Extrapolation: Mitigating Corrections



- Correct FD simulations by observing the differences in the ND data and simulations
 - Takes into account Far/Near transformation, oscillations, and detector acceptance
- Significantly reduces the impact of systematic uncertainties
 - e.g. uncertainty on neutrino cross-sections goes down from $\sim 15\%$ to $\sim 4-5\%$

Enhancing Sensitivity to Oscillations

 $\nu_{\mu}/\bar{\nu}_{\mu}$ Sample

 $\nu_e/\bar{\nu}_e$ Sample





New Low Energy ν_e Sample

- Developed a new selection to retain ν_e events in the low energy region where neutrino-anti neutrino asymmetry is maximal
- Improves sensitivity to mass orderings by ~few percent (depending on the oscillation parameters)
- No low energy events for the anti-neutrino beam mode



Far Detector $\nu_{\mu}(\bar{\nu}_{\mu})$ Observations

• Observed $\nu_{\mu}(\bar{\nu}_{\mu})$ candidates from 10 years of NOvA Data (neutrino beam exposure of 26.6 × 10²⁰ POT and anti-neutrino beam exposure of 12.5 × 10²⁰ POT)



Far Detector $\nu_e(\bar{\nu}_e)$ Observations

• Observed $\nu_e(\bar{\nu}_e)$ candidates from 10 years of NOvA Data (neutrino beam exposure of 26.6 × 10²⁰ POT and anti-neutrino beam exposure of 12.5 × 10²⁰ POT)





Fitting Procedure

- We perform a joint fit to $\nu_{\mu}/\bar{\nu}_{\mu}$ disappearance and $\nu_{e}/\bar{\nu}_{e}$ appearance data to extract oscillation parameters
- External constraints on solar parameters
- Reactor constraints on θ_{13} :
 - Unconstrained
 - Daya Bay 1D reactor constraint: $\sin^2 2\theta_{13} = 0.0851 \pm 0.0024$
 - Daya Bay 2D $(\Delta m_{32}^2, \theta_{13})$ constraint





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Results: I



	Frequentist results (w/ Daya Bay 1D θ ₁₃ constraint)				
	Normal MO		Inverted MO		
Δm_{32}^2 / 10 ⁻³ eV ²	+2.433	+0.035 -0.036	-2.473	+0.035 -0.035	
sin ² θ_{23}	0.546	+0.032 -0.075	0.539	+0.028 -0.075	
δ _{CP}	0.88 π		1.51 π		
Rejection significance (σ)			1.36		

- NOvA's measurements consistent with the rest of the accelerator and atmospheric experiments
- Δm_{32}^2 best-fit lies in the normal mass ordering (NO)
- $\sin^2(\theta_{23})$ best-fit value lies in the upper octant

Results: II



• NOvA data disfavors $\delta_{CP} = 3\pi/2$ in NO and $\delta_{CP} = \pi/2$ in IO

- The new NOvA measurements of δ_{CP} are consistent with our previous (2020) analysis
- The T2K, joint NOvA+T2K results favor different δ_{CP} regions in NO, same in IO

Synergy With Reactor Measurements

• NOvA data has a mild preference for the normal mass ordering

• Preference enhances with 1D and 2D reactor constraints



	No Constraint		1D Constraint		2D Constraint	
	Prob	BF	Prob	BF	Prob	BF
Normal Ordering Preference	69%	2.2	76%	3.2	87%	6.8

Results Contd.



NOvA produced the most precise (~1.5%) measurement of Δm_{32}^2 .

Conclusions

- ★ Latest three-flavor neutrino oscillation results from 10 years of NOvA data with doubled neutrino beam dataset (compared to 2020)
- * NOvA data prefers upper octant with reactor constraints on θ_{13} (prob=69%)
- * Mild preference to normal mass ordering (posterior prob. = 87%)
- * The most precise single experiment measurement of Δm_{32}^2 (precision=1.5%)
- * Frequentist best-fit values

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The NOvA Collaboration



Thank you for your attention!

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Back Up

Near-to-Far Extrapolation

- Functionally identical detectors cancel out systematic uncertainties on the best fit neutrino oscillation parameters
- The near detector (ND) data-MC differences are extrapolated in true energy bins to provide datadriven predictions of un-oscillated ν_{μ} ($\bar{\nu}_{\mu}$) and oscillated ν_{e} ($\bar{\nu}_{e}$) events at the far detector (FD)
- The ν_{μ} ($\bar{\nu}_{\mu}$) extrapolation is divided into 4 hadronic energy fraction quartiles to improve the sensitivity of the experiment
- Extrapolation is further divided into 3 bins of final state lepton transverse momentum (p_t) which takes into account the neutrino interaction mis-modeling and the differences in ND and FD



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Uncertainties on FD Predictions



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Selection

20

Uncertainties on Oscillation Parameters





Source of Uncertainty	$\sin^2\theta_{23}$	δ_{CP}/π	$ \Delta m^2_{32} $ (×10 ⁻³ eV ²)
Beam Flux	+0.00042 / -0.00069	+0.0012 / -0.011	+0.00053 / -0.0012
Detector Calibration	+0.0033 / -0.017	+0.014 / -0.17	+0.013 / -0.016
Detector Response	+0.00031 / -0.0043	+0.004 / -0.037	+0.0016 / -0.0026
Lepton Reconstruction	+0.0027 / -0.0046	+0.007 / -0.034	+0.0083 / -0.014
Near-Far Uncor.	+0.0025 / -0.0024	+0.0072 / -0.043	+0.0022 / -0.0034
Neutrino Cross Sections	+0.0031 / -0.0051	+0.018 / -0.11	+0.0058 / -0.011
Neutron Uncertainty	+0.0028 / -0.00075	+0.0056 / -0.011	+0.0022 / -0.0041
Systematic Uncertainty	+0.0067 / -0.019	+0.027 / -0.21	+0.017 / -0.024
Statistical Uncertainty	+0.023 / -0.083	+0.081 / -0.76	+0.032 / -0.044

Table: Summary of uncertainties on Ana2024 frequentist joint best-fit point, evaluated at the NOUO best-fit values i.e. $\sin^2\theta_{23} = 0.55$, $\delta_{CP}/\pi = 0.88$, and $|\Delta m_{32}^2|$ (×10⁻³ eV²) = 2.43.

FD $\nu_{\mu}(\bar{\nu}_{\mu})$ Events By Quartiles





Ratios to No Oscillations

