Exploring DSNB boosted sub-GeV dark matter: insights from XENONnT and LZ experiments

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Introduction

Motivation for searching the Dark Matter (DM)

- Cold DM: It is a non-luminous matter which occupies 27% of the mass and energy in the observable Universe. It does not interact with photons and interacts only "weakly" with ordinary matter.
- Astronomical and cosmological observations at various scales:
	- (i) Rotation curves of spiral galaxies and galaxy clusters
	- (ii) Gravitational lensing
	- (iii) Cosmic Microwave background (CMB) fluctuations

- **Direct Detection Experiments:** *XENONnT, LUX-ZEPLIN, Super-CDMS, Dark-Side, PandaX-4T, etc.*
- **Indirect Detection Experiment:** *IceCube, HESS, MAGIC, etc.*
- **Accelerator searches:** *ATLAS, CMS at CERN*

DM landscape: a wide mass range

DM landscape: a wide mass range

Thermal relic DM vs boosted DM

The maximum recoil energy of the target:

 $T_r^{\text{max}} \approx \frac{Q^2}{2m}$ $\frac{Q^2}{2m_T} \approx \frac{2m_\chi^2m_Tv_\chi^2}{(m_\chi+m_T)}$ $(m_\chi + m_T)^2$

Thermal relic DM vs boosted DM

The maximum recoil energy of the target:

Thermal relic DM vs boosted DM

The maximum recoil energy of the target:

[DSNB boosted dark matter](#page-9-0)

Diffuse Supernova Neutrino Background @ IISER Bhopal

Right after the first star formation event, the Universe has been surrounded by an isotropic flux of MeV-energy neutrinos and antineutrinos of all flavors, produced from all supernovae events from the core-collapse explosions of huge stars throughout the Universe. This cumulative and isotropic flux of MeV neutrinos form DSNB.

BDM Flux At The Underground Detectors ONIER Bhopal

The DSNB-boosted DM differential flux,

$$
\frac{d\Phi_{\chi}}{dT_{\chi}} = D_{\text{halo}} \sum_{\alpha} \int_{E_{\nu}^{\text{min}}}^{E_{\nu}^{\text{max}}} dE_{\nu} \frac{1}{m_{\chi}} \frac{d\sigma_{\nu\chi}}{dT_{\chi}} \frac{d\Phi_{\nu\alpha}^{\text{DSNB}}}{dE_{\nu}}
$$

*D*_{halo} encodes the line of sight integral of DM density within our galactic halo,

$$
D_{\rm halo} = \int_{\Delta\Omega} \frac{d\Omega}{4\pi} \int_0^{\ell_{\rm max}} \rho_{\rm NFW}[r(\ell,\psi)]d\ell
$$

■ We consider Navarro-Frenk-White (NFW) profile for galactic DM density,

$$
\rho_{\text{NFW}}(r) = \rho_{\odot} \left[\frac{r}{r_{\odot}} \right]^{-1} \left[\frac{1 + \frac{r_{\odot}}{r_{\rm s}}}{1 + \frac{r}{r_{\rm s}}} \right]^2
$$

For the differential cross section, we assume the constant cross section approximation, i.e.

$$
\frac{d\sigma_{\nu\chi}}{dT_{\chi}}=\frac{\sigma_{\nu\chi}}{T_{\chi}^{\rm max}}
$$

BDM Flux At The Underground Detectors ONIER Bhopal

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$$
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$$

DM flux gets attenuated by the elements of the atmosphere and Earth before reaching to the underground detector. However as $\eta^{\rm atm}_i << \eta^{\rm Earth}_i$, we have **neglected the attenuation due to elements of the atmosphere**.

$$
\frac{d\sigma_{\chi i}}{dT_i} = \frac{\sigma_{\chi i}}{T_i^{\max}}
$$

$$
i \equiv \{e, \mathcal{N}\}
$$

BDM Flux At The Underground Detectors O IISER Bhopal

The DSNB-boosted DM differential flux,

$$
\frac{d\Phi_{\chi}}{dT_{\chi}} = D_{\text{halo}} \sum_{\alpha} \int_{E_{\nu}^{\text{min}}}^{E_{\nu}^{\text{max}}} dE_{\nu} \frac{1}{m_{\chi}} \frac{d\sigma_{\nu_{\chi}}}{dT_{\chi}} \frac{d\Phi_{\nu_{\alpha}}^{\text{DSNB}}}{dE_{\nu}}
$$

DM flux gets attenuated by the elements of the atmosphere and Earth before reaching to the underground detector. However as $\eta^{\rm atm}_i << \eta^{\rm Earth}_i$, we have **neglected the attenuation due to elements of the atmosphere**.

Implications of nuclear form factor

101

The spin independent DM-nuclei scattering cross section can be written as, 10-¹ Ge $\sigma_{\chi \cal N}^{\rm SI} (q^2) = \frac{\mu_{\chi \cal N}^2}{\mu^2}$ Xe $\frac{\mu_{\chi N}^2}{\mu_{\chi n}^2}$ Α²σ_{χn}F²(q²) 10^{-3} $F^2(E_{\rm m})$ 10^{-5} 10-⁷ $z = 1.4$ km 10⁻⁹ $=\sigma_{\chi\rm e} = \sigma_{\chi n} = 10^{-29} \; {\rm cm}^2$ 10⁶ $\sigma_{\nu\chi} = \sigma_{\chi e} = \sigma_{\chi n} = 10^{-29}$ cm² 0 50 100 150 200 250 300 10−³ E_n (keV_{nr}) χ^0 = 10⁵ keV ο. 10⁵ *T* (Helm, 1956) $^{-2}$ s⁻¹) Unattenuated 10^{-4} $\frac{0}{\chi}$ = 10⁴ keV $\mathbf{0}$. Attenuated $[\chi - \mathcal{N}]$ (*F* =Helm) *T mχ* = 300 MeV $10⁴$ Attenuated $[\chi - \mathcal{N}]$ ($F = 1$) $\sum_{r=1}^{10^2}$ *T* $\frac{0}{\chi}$ = 10³ keV 0. $\left| \frac{5}{10^{-5}} \right|$ 10^3 $\frac{0}{\chi}$ = 10² keV $\mathbf{0}$. *T T^zχ* · *^d* Φ*χ dT ^zχ* 10^2 10−⁶ *F* = Helm $10¹$ $F=1$ 10^{0} - 10^{-7} $\sigma_{\nu\chi}$ 10^{-1} 10^{0} 10^{1} 10^{2} 10^{3} 10^{4} 10^{5} 10^{-2} 10⁻¹ 10⁰ 10¹ 10² 10³ *T z χ* (keV) *z* (km) [JCAP 03 (2024) 028]

DM signal at the underground detectors

After reaching the underground detector, the DSNB-boosted DM can scatter off both the electrons and nuclei of the target material, triggering both electronic and nuclear recoils. The differential event rate with respect to the recoil energy *Tⁱ* can be written as,

$$
\frac{dR}{dT_i} = t_{\rm run} N_{\rm target}^i \mathcal{A} \int dT_x^z \frac{d\Phi_x}{dT_x^z} \frac{d\sigma_{xi}}{dT_i}
$$

 $m_x = 300$ MeV

Effect of Earth attenuation in the resulting limits

Anirban Majumdar [DSNB Boosted Dark Matter](#page-0-0) 15th October 2024 9 / 13

Resulting Limits

V. De Romeri, A. Majumdar et al., JCAP 03 (2024) 028

Conclusions

- ☞ DSNB Boosted DM produces a subdominant, semi-relativistic component of Galactic DM.
- ☞ Consideration of Earth attenuation is crucial for accurate interpretation of experimental results.
- ☞ *Although a significant part of our constraints lie in a region of parameter space already probed by other searches, these results highlight the complementarity and significance of the LZ and XENONnT data in probing the sub-GeV DM parameter space.*

THANK YOU

χ 2 **function utilized**

For the analysis of LZ data, we have performed a spectral analysis using the following Poissonian χ^2 function

$$
\chi^2(\overrightarrow{S};\alpha,\beta,\delta)=2\sum_{i=1}^{51}\left[R_{\text{pred}}^i(\overrightarrow{S};\alpha,\beta,\delta)-R_{\text{exp}}^i+R_{\text{exp}}^i\ln\left(\frac{R_{\text{exp}}^i}{R_{\text{pred}}^i(\overrightarrow{S};\alpha,\beta,\delta)}\right)\right]+\\\left(\frac{\alpha}{\sigma_\alpha}\right)^2+\left(\frac{\beta}{\sigma_\beta}\right)^2+\left(\frac{\delta}{\sigma_\delta}\right)^2,
$$

The following Gaussian χ^2 function is used for the analysis of XENONnT data

$$
\chi^2(\overrightarrow{\mathcal{S}};\beta) = \sum_{i=1}^{30} \left(\frac{R_{\text{pred}}^i(\overrightarrow{\mathcal{S}};\beta) - R_{\text{exp}}^i}{\sigma^i} \right)^2 + \left(\frac{\beta}{\sigma_{\beta}} \right)^2
$$

Geophysical properties of Earth

We model the Earth's interior as a sphere of constant electron and nuclear densities ($n_e = 8 \times 10^{23}~\rm cm^{-3}$ and $n_\mathcal{N}=3.44\times 10^{22}~\rm cm^{-3}$), based on the abundances of the main elements as shown in following table.

