Predictions from scoto-seesaw with A₄ modular symmetry

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Flavor structure: Hint for new physics

Modular symmetry as a predictive framework

Mass hierarchy of neutrinos shaped by the scoto-seesaw framework

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Flavour Structure of Standard Model

• In Standard Model (SM) there are three families of five fermions:

 $Q_{Li}, u_{Ri}^c, d_{Ri}^c, I_{Li}, e_{Ri}$

Elementary Particles



I II III Three Families of Matter

• Flavor problem occurs when three generations have to live together.

Flavour Structure of Standard Model

• This problem arises from Mass terms:

$$e_{Ri}^{c}M_{ij}^{e}l_{Lj}$$
 : Charged leptons (1)

$$I_i M_{ij}^{\nu} I_j$$
 : Neutrinos (2)

$$U_{PMNS} = V_e^{\dagger} V_{\nu} \tag{3}$$





- This peculiar pattern consists flavor puzzle.
- Leads to new physics : Symmetry approach.





There are various ways of doing this:

- Froggatt Nielsen Mechanism: Solves hierarchy problem but no explanation for large mixing angles in lepton sector.
- **Discrete Symmetry:** Explains mixing pattern but symmetry breaking mechanism is quite complicated.

In Modular symmetry approach flavor symmetry is realised in a non-linear way hence making it special.

Mass hierarchy of neutrinos shaped by the scoto-seesaw framework

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Model Framework

• We extend the SM symmetry by including A_4 modular symmetry.

•
$$\Delta m_{atm}^2 = 2.5 \times 10^{-3} \text{ eV}^2$$
 and $\Delta m_{sol}^2 = 7.41 \times 10^{-5} \text{ eV}^2$.

• origins of these two scales may stem from separate mechanisms : Scoto-seesaw.

	Fermions					Scalars			Yukawa couplings						
Fields	L_ℓ	ℓ_R^c	N_{R_1}	N_{R_2}	f	$H_{u,d}$	η	η'	χ	$Y_{1}^{(4)}$	$Y_{1'}^{(4)}$	$Y_{1}^{(8)}$	$Y_{1'}^{(8)}$	$Y_{1''}^{(8)}$	$Y_{1}^{(10)}$
$SU(2)_L$	2	1	1	1	1	2	2	2	1	-	-	-	-	-	-
$U(1)_Y$	-1/2	1	0	0	0	$\pm 1/2$	1/2	-1/2	0	-	-	-	-	-	-
A_4	$1,1^{\prime},1^{\prime\prime}$	$1,1^{\prime\prime},1^{\prime}$	1	1′	1	1	1	1	1	1	1′	1	1′	1″	1
kı	0	0	4	4	5	0	3	3	5	4	4	8	8	8	10

Table: Particle content and modular Yukawa couplings of the model and their charges under $SU(2)_L \times U(1)_Y \times A_4$, where k_l is modular weight.

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Atmospheric Neutrino Mass Scale

• Generating the atmospheric scale through type-I seesaw.



Figure: Feynman diagram at tree level

• The Superpotential at tree level :

$$\begin{aligned} \mathcal{W}_{\nu}^{\mathcal{T}} &= \alpha_{\mathcal{T}} \left(Y_{1}^{(4)} L_{e} H_{u} N_{R_{1}} + Y_{1'}^{(4)} L_{\tau} H_{u} N_{R_{1}} + Y_{1'}^{(4)} L_{\mu} H_{u} N_{R_{2}} + Y_{1}^{(4)} L_{\tau} H_{u} N_{R_{2}} \right) \\ &+ \kappa_{1} Y_{1}^{(8)} N_{R_{1}} N_{R_{1}} + \kappa_{2} Y_{1''}^{(8)} N_{R_{2}} N_{R_{2}}. \end{aligned}$$

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• Neutrino mass matrix:

$$M_D = \begin{pmatrix} Y_1^{(4)} & 0\\ 0 & Y_{1'}^{(4)}\\ Y_{1'}^{(4)} & Y_1^{(4)} \end{pmatrix} \alpha_T v_u, \qquad M_R = \begin{pmatrix} \kappa_1 Y_1^{(8)} & 0\\ 0 & \kappa_2 Y_{1''}^{(8)} \end{pmatrix}$$
(4)

$$(M_{\nu})_{\rm tree} = -M_D M_R^{-1} M_D^T \tag{5}$$

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Solar Neutrino mass scale

The Superpotential at loop level:

$$\mathcal{W}_{\nu}^{\mathcal{L}} = \beta_{L} \left(Y_{1}^{(8)} L_{e} \eta f + Y_{1''}^{(8)} L_{\mu} \eta f + Y_{1'}^{(8)} L_{\tau} \eta f \right) + \kappa_{S} Y_{1}^{(10)} f f + \lambda_{1} Y_{1}^{(8)} H_{d} \eta \chi .$$



Figure: Radiative neutrino mass generation

Neutrino mass matrices at loop level:

$$(M_{\nu})_{\text{loop}} = \beta_{L}^{2} M_{f} \begin{pmatrix} \left(Y_{1}^{(8)}\right)^{2} & \left(Y_{1}^{(8)}Y_{1''}^{(8)}\right) & \left(Y_{1}^{(8)}Y_{1'}^{(8)}\right) \\ * & \left(Y_{1''}^{(8)}\right)^{2} & \left(Y_{1''}^{(8)}Y_{1'}^{(8)}\right) \\ * & * & \left(Y_{1''}^{(8)}\right)^{2} \end{pmatrix} \mathcal{F}(m_{\eta_{R}}, m_{\eta_{I}}, M_{f})$$

Total neutrino mass:

$$M_{\nu} = (M_{\nu})_{\rm tree} + (M_{\nu})_{\rm loop} \tag{6}$$

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Input Parameters	Range				
$\operatorname{Re}[\tau]$	$\pm [0.0, 0.5]$				
$\operatorname{Im}[\tau]$	[0.8, 1.5]				
$lpha_T$	$[10^{-3}, 10^{-2}]$				
β_L	$[10^{-1}, 2]$				
M_1 (GeV)	$[1, 10] \times 10^{11}$				
M_2 (GeV)	$[1, 10] \times 10^{11}$				
M_f (GeV)	$[1, 10^4]$				
m_{η_R} (GeV)	[1, 400]				
m_{η_I} (GeV)	[1, 400]				

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Model Predictions



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Results from Neutrino phenomenology

Model favors normal ordering of neutrino mass



Figure: Predictions of neutrino oscillation parameters from model

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Results from Neutrino phenomenology



Figure: Predicted range for δ_{CP}

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Results from Neutrino phenomenology



Figure: Neutrinoless double beta decay parameter and Majorana phases

Lepton Flavor Violating Decay mode $\mu ightarrow e \gamma$



Figure: Expected branching ratio (BR) of $\mu \rightarrow e\gamma$ with experimental upper limit as 4.2×10^{-13}

Lepton Flavor Violating Decay mode $\mu \rightarrow 3e$



Figure: Expected BR of $\mu \rightarrow 3e$ with experimental upper limit as 1×10^{-12}

- Modular symmetry is highly predictive in terms of mixing parameters in lepton sector.
- Neutrino mass is hierarchical by scoto-seesaw.
- Our model predicts normal ordering for neutrino masses and $m_{
 m lightest} \in (9.2, 20.0) \times 10^{-3}$ eV.
- The neutrinoless double beta decay parameter $|m_{ee}| \in (3.15, 6.66) \times 10^{-3}$ eV, which is within the potential reach of upcoming experiments.
- A₄ modular symmetry within the scoto-seesaw framework leads to a highly predictive model whose predictions can be tested in various experiments.

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Thankyou



$$\Gamma^{0
u}_{etaeta} = rac{1}{T^{0
u}_{etaeta}} = G^{0
u} \cdot \left| M^{0
u} \right|^2 \cdot \langle m_{etaeta}
angle^2$$
 Effective Majorana mass

$$\langle m_{etaeta}
angle = \sum_i U_{ei}^2 m_i$$
 ,

a larger $< m_{ee} >$ increases the chances of detecting the decay.

$$f(\tau) = (-1)^k f(\tau) \tag{7}$$

Thus for odd weights, modular forms vanishes.



I. q-Series: an example

The generating function, the infinite *q*-series

$$\sum_{n=0}^{\infty} p(n)q^n = \frac{1}{\prod_{k=1}^{\infty} (1-q^k)}$$

= 1 + q + 2 q^2 + 3 q^3 + 5 q^4 + ...,
 $\in \mathbb{Z}[[q]]$



Leonhard Euler

▶ ★ E ▶ E = 9QQ

is modular!

$$q^{-\frac{1}{24}} \sum_{n=0}^{\infty} p(n)q^n = \frac{q^{-\frac{1}{24}}}{\prod_{k=1}^{\infty} (1-q^k)} = \frac{1}{\eta(\tau)}$$

$$\Rightarrow \text{ a shift by } q^{\frac{-1}{24}}$$

$$\Rightarrow q \to \tau$$

• a shift by
$$q^{rac{-1}{24}}$$

• $q
ightarrow au$

I. q-Series: as modular forms

The Dedekind eta function:

$$\eta(au) = rac{q^{rac{1}{24}}}{\sum_{n=0}^{\infty} p(n)q^n} = q^{rac{1}{24}} \prod_{k=1}^{\infty} (1-q^k)$$

is modular!

II. Modular Forms: the partition example

$$\eta^{24} : \mathbb{H} \to \mathbb{C}$$
$$\eta^{24}(\tau) = (c\tau + d)^{-12} \eta^{24} \left(\frac{a\tau + b}{c\tau + d}\right)$$
$$= q \prod_{k=1}^{\infty} (1 - q^k)^{24}$$

is a modular form of weight 12.

II. Modular Forms: as in 2d conformal field theory and string theory

A string moving in time = a cylinder.



The partition functions are computed by identifying the initial and final time. This turns the cylinder into a torus. As a result the string partition functions are modular forms!

Modularity is very helpful in studying these physical theories.

Neutrino Phenomenology

Priya Mishra

II. Modular Forms: as in 2d conformal field theory and string theory String theory/CFT is very helpful to understand modularity.

Example 1: free chiral boson/Heisenberg algebra

$$Z(\tau) = q^{-\frac{1}{24}} \sum_{n=0}^{\infty} p(n)q^n = \frac{1}{\eta(\tau)}$$

ground state energy

MODULAR SYMMETRY



► The modular symmetry is a geometrical symmetry of the two-dimensional torus, *T*².

(16)

- The two-dimensional torus is constructed as division of the two-dimensional Euclidean space R^2 by a lattice Λ , $T^2 = \frac{R^2}{\Lambda}$.
- ▶ Instead of *R*², one can use the one-dimensional complex plane.
- The lattice is spanned by two basis vectors, e_1 and e_2 as $m_1e_1 + m_2e_2$, where m_1 and m_2 are integer.
- There ratio is

$$\tau = \frac{e_2}{e_1}$$



in the complex plane, represents the shape of the Torus T^2 and parameter τ is called the modulus.

▶ The same lattice can be spanned by other basis vectors such as

$$\begin{pmatrix} e_1' \\ e_2' \end{pmatrix} = \begin{pmatrix} a & b \\ c & d \end{pmatrix} \begin{pmatrix} e_1 \\ e_2 \end{pmatrix}$$
(17)

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where a, b, c, d are integer satisfying ad - bc = 1. That is the SL(2, Z).



- One interesting thing about finite modular groups ⁷ are: they are isomorphic to discrete symmetry groups like $\Gamma_2 \simeq S_3$, $\Gamma_3 \simeq A_4$, $\Gamma_4 \simeq S_4$, $\Gamma_5 \simeq A_5$, $\Gamma'_3 \simeq A'_4$, $\Gamma'_5 \simeq A'_5$ etc.
- ▶ The modular group is defined as a group of 2 × 2 matrices having integer entries and determinant 1.

$$\Gamma(N) = \left\{ \begin{pmatrix} a & b \\ c & d \end{pmatrix} \in SL(2, Z); \begin{pmatrix} a & b \\ c & d \end{pmatrix} = \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} (mod N) \right\}.$$

Therefore, the group $\Gamma(N)$ acts on the complex variable τ , varying in the upper-half $\mathcal{H} = \text{Im}(\tau) > 0$, as linear fractional transformation given by

$$\gamma(\tau) = \frac{a\tau + b}{c\tau + d}, \quad \mathcal{H} = \{\tau \in \mathbb{C}, \text{ Im}(\tau) > 0\}, \quad \gamma = \begin{bmatrix} a & b \\ c & d \end{bmatrix} \in \text{SL}(2, \mathbb{Z}), \tag{18}$$

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⁷Feruglio, Ferruccio, Are neutrino masses modular forms?, From My Vast Repertoire ...

► The generators of modular group being

$$S = \begin{bmatrix} 0 & -1 \\ 1 & 0 \end{bmatrix} \text{ and } T = \begin{bmatrix} 1 & 1 \\ 0 & 1 \end{bmatrix}.$$
 (19)
$$S : \tau \to -\frac{1}{\tau} \qquad T : \tau \to \tau + 1.$$
 (20)

- A function F is said to be entire modular of weight k if it satisfies these below conditions:
 - 1. \mathcal{F} is analytic in the upper plane H, H =(x + iy | y > 0; $x, y \in R$). 2. $\mathcal{F}\left(\frac{a\tau+b}{c\tau+d}\right) = (c\tau+d)^k \mathcal{F}(\tau)$.

3. The fourier series of ${\cal F}$ is given by the form of (called q expansion form).

$$\mathcal{F}(\tau) = \sum_{n=0}^{\infty} a_n q^n \qquad q = e^{2\pi i \tau}$$
(21)

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Under the modular transformation, chiral superfields ψ_i (i denotes flavors) with weight -k transform as

$$\psi_i \to (c\tau + d)^{-k_i} \rho(\gamma) i j \psi_j.$$
 (8)

Weight (<i>k</i>)	d_k	A_4 representations
2	3	3
4	5	3 + 1 + 1 ′
6	7	3 + 3 + 1
8	9	3+3+1+1'+1"
10	11	3+3+3+1 +1'

Table: A_4 representations for different weight k.

UV completion of SUSY scotogenic loop

$$\mathcal{W}_{S} = \mu H_{u} H_{d} + \mu_{\eta} Y_{1}^{(6)} \eta \eta' + \frac{1}{2} \mu_{\chi} Y_{1}^{(10)} \chi \chi + \lambda_{1} Y_{1}^{(8)} H_{d} \eta \chi + \lambda_{2} Y_{1}^{(8)} H_{u} \eta' \chi.$$

$$\left(M_{\nu}^{ij}\right)_{\text{loop}} = \sum_{l=1}^{3} \mathcal{F}_{1l} M_{f} \mathbf{h}^{i} \mathbf{h}^{j} + \sum_{l=1}^{3} \mathcal{F}_{2l} m_{\tilde{\eta}_{l}} \mathbf{h}^{i} \mathbf{h}^{j} , \qquad (9)$$

where \mathcal{F}_{1l} and \mathcal{F}_{2l} are the loop function given by:

$$\mathcal{F}_{1I} = \frac{1}{32\pi^2} \left[\left[U_R(2,I) \right]^2 \frac{m_{\eta_{RI}}^2}{M_f^2 - m_{\eta_{RI}}^2} \ln\left(\frac{M_f^2}{m_{\eta_{RI}}^2}\right) - \left[U_I(2,I) \right]^2 \frac{m_{\eta_{II}}^2}{M_f^2 - m_{\eta_{II}}^2} \ln\left(\frac{M_f^2}{m_{\eta_{II}}^2}\right) \right], \tag{10}$$

$$\mathcal{F}_{2I} = \left[U_\eta(2,I) \right]^2 \frac{1}{32\pi^2} \left[\frac{m_{fR}^2}{m_{\tilde{\eta}_I}^2 - m_{fR}^2} \ln\left(\frac{m_{\tilde{\eta}_I}^2}{m_{fR}^2}\right) - \frac{m_{fI}^2}{m_{\tilde{\eta}_I}^2 - m_{fI}^2} \ln\left(\frac{m_{\tilde{\eta}_I}^2}{m_{\tilde{\eta}_I}^2}\right) \right]. \tag{11}$$

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$$\mathcal{F}_{1l} \to \mathcal{F} = \frac{1}{32\pi^2} \left[\frac{m_{\eta_R}^2}{M_f^2 - m_{\eta_R}^2} \ln\left(\frac{M_f^2}{m_{\eta_R}^2}\right) - \frac{m_{\eta_l}^2}{M_f^2 - m_{\eta_l}^2} \ln\left(\frac{M_f^2}{m_{\eta_l}^2}\right) \right].$$
(12)

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