## **DM Mass from Angular Dependence**

with D. Kim [In preparation]





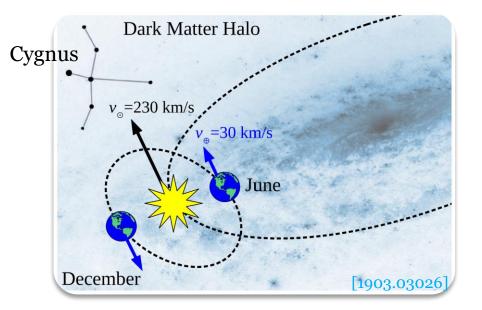
Jong-Chul Park Chungnam National Univ. PPC 2024, October 16 (2024)

#### **DM Direct Detection: Basics**

$$\Phi_{\chi} = n_{\chi} v_{\rm rel} \& n_{\chi} = \rho_{\chi} / m_{\chi}$$

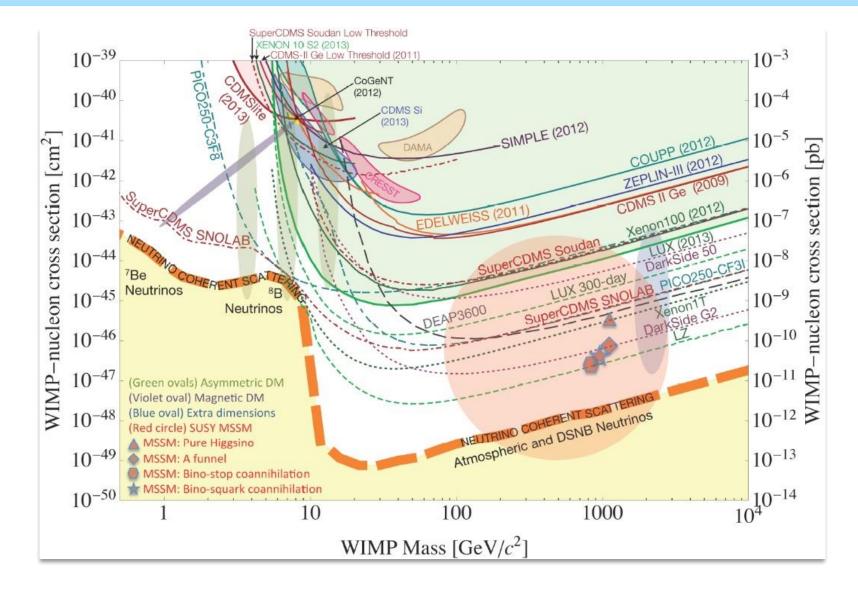
$$\frac{dN}{dE_R}(t) \propto N_T \frac{\rho_{\chi}}{m_{\chi}} \int_{v > v_{\min}} dv^3 \frac{d\sigma}{dE_R} v f_{\text{Earth}}(\vec{v}, t)$$

$$v_{\rm min} = \sqrt{m_T E_R / 2\mu_{\chi T}^2}$$



$$f_{\text{Earth}}(\vec{v}, t) = f_{\text{Galaxy}}(\vec{v} + \vec{v}_{\odot} + \vec{v}_{\oplus}(t))$$

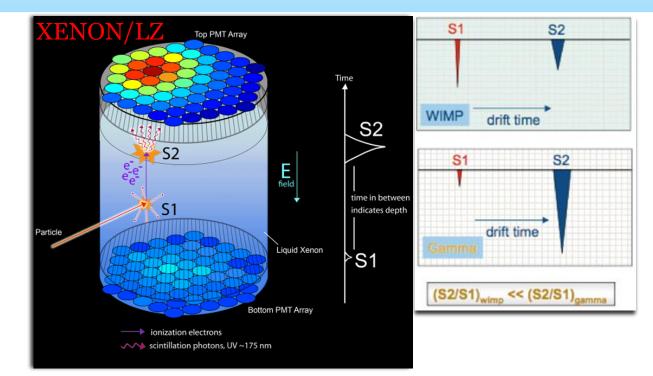
#### **DM (WIMP) Direct Detection: Results**

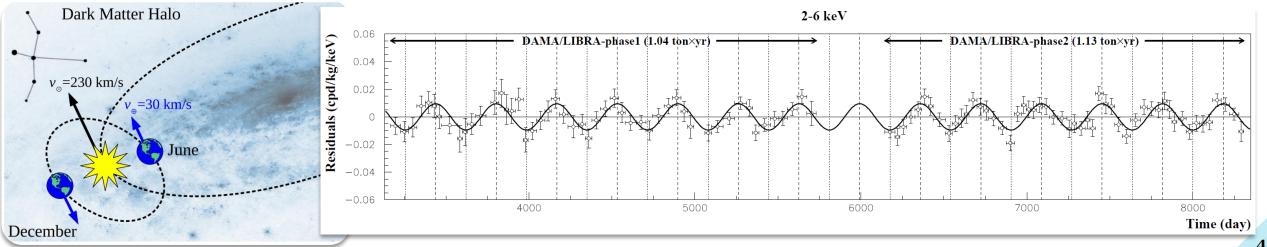


#### **DM Direct Detection: Some Issues**

#### 1. DM signals vs Backgrounds

- ✓ Event discrimination via signal characteristics: most of experiments
- ✓ Earth's motion around the Sun →
   Annual modulation in event rate (e.g. DAMA, COSINE), Directional
   detection (e.g. DRIFT, NEWSdm)





### **DM Direct Detection: Some Issues**

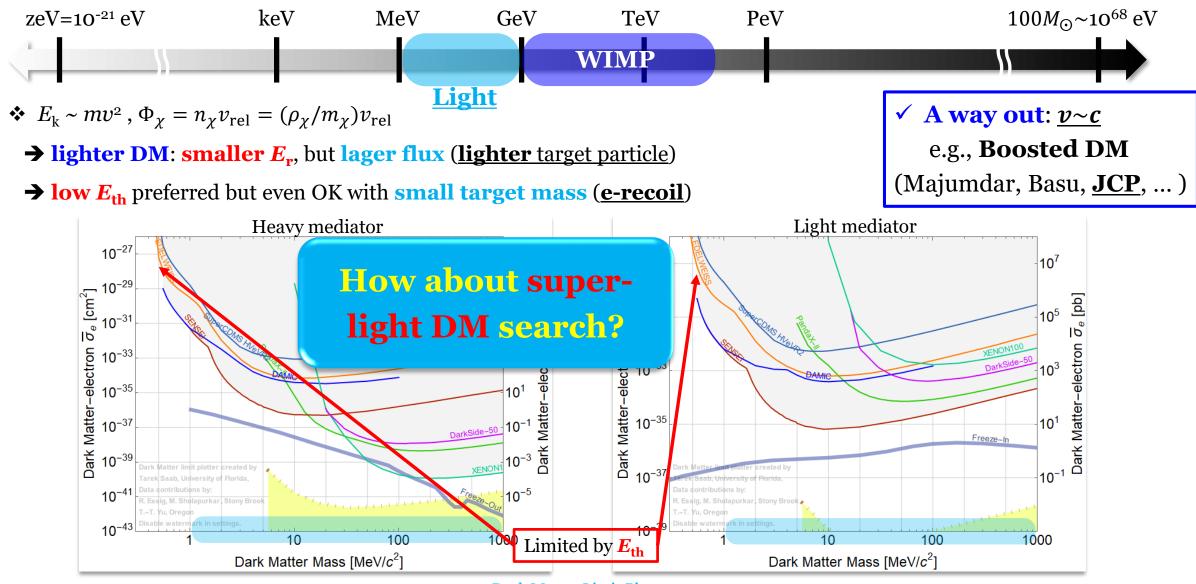
#### 2. Mass & interaction of DM

T. Schwetz, PPC11 CERN

 $10^{-39}$ **MS-II Ge Low Threshold (20**  $10^{-3}$ CoGeN'  $\checkmark$  Differential recoil rate: 2012)  $10^{-40}$  $10^{-4}$ CDMS Si SIMPLE (2012)  $10^{-41}$  $[cm^2]$  $10^{-5}$ Amplitude  $\rightarrow$  Interaction strength nucleon cross section [pb]  $10^{-6}$ Curvature ( $\sim$ distribution)  $\rightarrow$  Mass WIMP-nucleon cross section  $10^{-43}$ SNOL  $10^{-44}$ 10-9 10-45 / keV [arb, Units] Neutrinos 10 m [GeV]  $10^{-10}$ 20 $10^{-46}$ 30  $10^{-11}$  WIN  $10^{-12}$  WIN  $10^{-47}$ Green ovals) Asymmetric DM (Violet oval) Magnetic DM  $10^{-48}$ ue oval) Extra dimensions Atmospheric and DSNB Neu ed circle) SUSY MSSM 3 MSSM: Pure Higgsino  $10^{-49}$  $10^{-13}$ uo MSSM: Bino-stop coannihilation events MSSM: Bino-squark coannihilation  $10^{-50}$  $10^{-14}$ 1000 100 10  $10^{4}$ WIMP Mass  $[\text{GeV}/c^2]$ 10 30 0 20 40 50 E<sub>rr</sub> [keV]

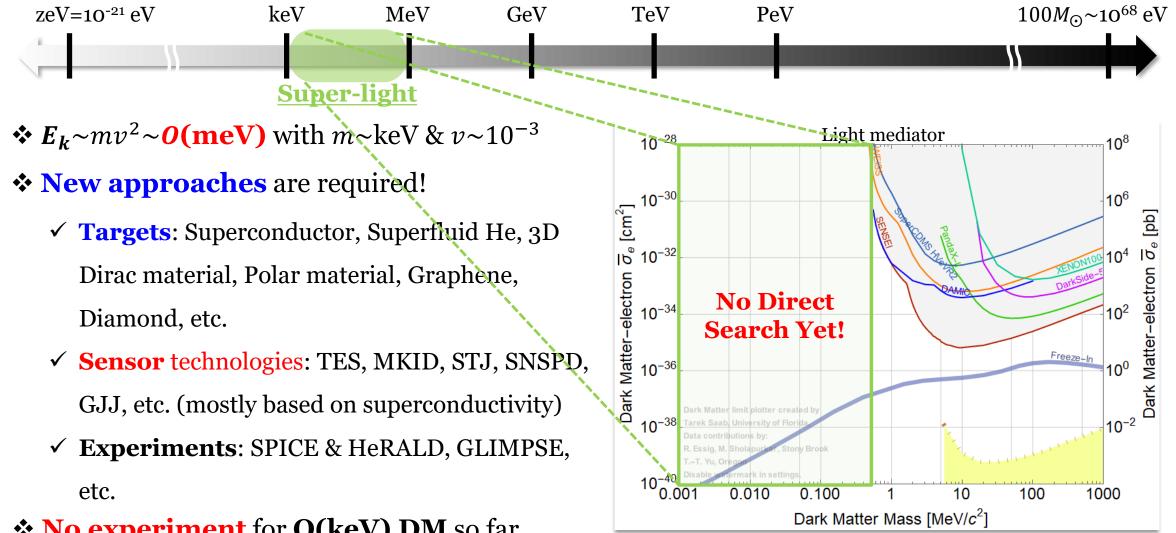
erCDMS Soudan Low Threshol NON 10 S2 (2013

### **Light DM Direct Search**



Dark Matter Limit Plotter

### **Super-Light DM Direct Search**



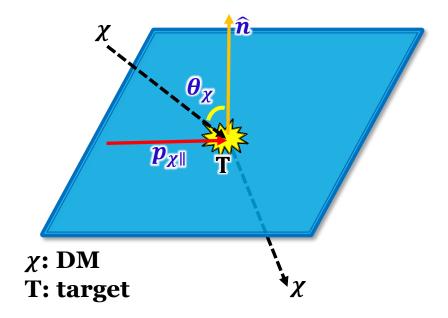
**Dark Matter Limit Plotter** 

### **Potential Questions for LDM Direct Detection**

- Low *E* sensor technologies mostly feature the "on-off" type working principle or relatively poor *E* resolution.
- 1. DM signals vs Backgrounds :
  - ✓ Event discrimination via signal characteristics: difficult!
  - ✓ For better directional detection, higher  $E_R$  is preferred, e.g., longer track.
  - ✓ But, light DM induces lower  $E_R$ : less visible signals (tracks)
    - → Can light DM be connected to directional recoil detection?
- 2. Mass determination:
  - $\checkmark$  We may recognize a DM event occurrence, but utilizing the differential  $E_R$  spectrum is difficult!
    - → Is there any alternative method to determine the mass of DM?

#### **Answers for the Questions**

★ Experiments using (effectively) 2D detectors: the experimental signatures are related to the behavior of targets scattered by DM along the detection plane → the incident angle ( $\theta_{\chi}$ ) of a DM particle affects the resulting event rate.



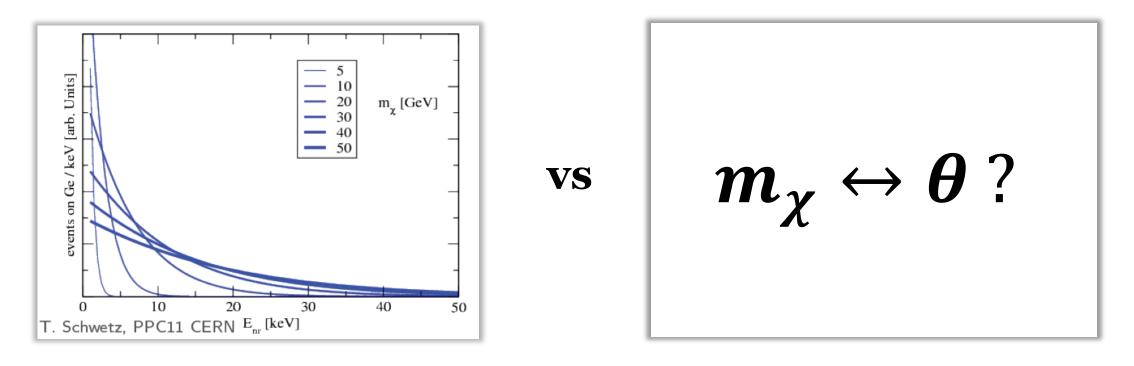
#### **Answers for the Questions**

- ★ Experiments using (effectively) 2D detectors: the experimental signatures are related to the behavior of targets scattered by DM along the detection plane → the incident angle ( $\theta_{\chi}$ ) of a DM particle affects the resulting event rate.
- ◆ Due to the motion of the Sun, the DM flux (DM wind) has a directional preference: CYGNUS!
  - → Non-trivial dependence of event rates on the incident angle  $(\theta_{\chi}^{w})$  of the DM wind.



#### **Answers for the Questions**

- ♦ (effectively) 2D detectors:  $v_{\chi\parallel}$  is more relevant to event rates than  $v_{\chi\perp}$  w.r.t. the detector plane.
- ✤ Heavy DM: a small v is good enough to get over the  $E_{th}$ , leaving a detectable signature + m via  $dR/dE_R$ . vs Light DM: a large v is preferred (+ no or poor  $dR/dE_R$ ).
- The 2D detection plane gets exposed to the DM wind at various angles. The resultant angular distribution of event rates per unit exposure time allows for the determination of the DM mass.

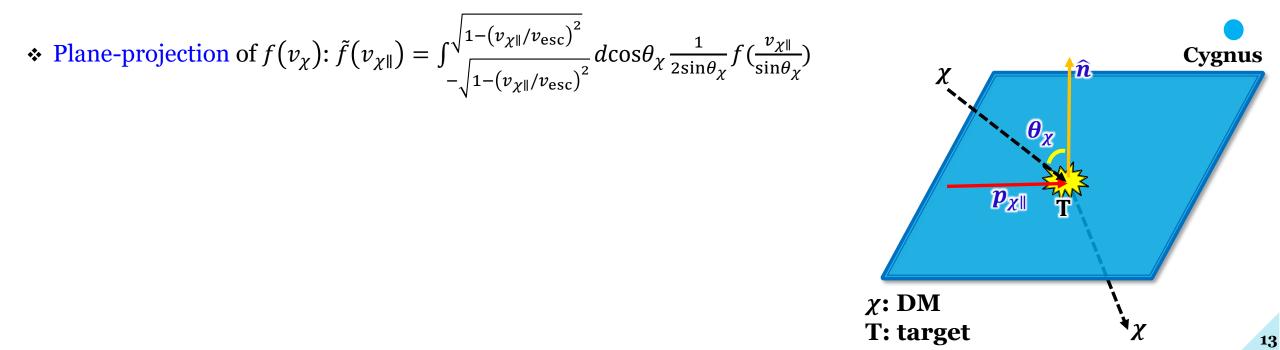


## 2D Detection: Angular Dependence

#### **Angular Dependence of Event Rates**

- \* Number of events/unit detector mass/unit run time:  $n_{\text{eve}} = \int dE_r dv_{\chi} f(v_{\chi}) \frac{d}{dE_r} \left( \overline{N}_T \langle \sigma_{\chi T} v_{\text{rel}} \rangle \frac{\rho_{\chi}}{m_{\chi}} \right)$  with
  - $\overline{N}_{\mathrm{T}} = N_{\mathrm{T}}/M_{\mathrm{T}}.$
- \* If the detector of interest is 2D,  $v_{\chi\parallel}$  (to the detection plane) affects the event rate:

$$n_{\text{eve}} = \frac{\rho_{\chi}}{m_{\chi}} \int dE_r dv_{\chi \parallel} \tilde{f}(v_{\chi \parallel}) \frac{d}{dE_r} \left( \overline{N}_{\text{T}} \langle \sigma_{\chi T} v_{\text{rel} \parallel} \rangle \right)$$



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• Plane-projection of 
$$f(v_{\chi})$$
:  $\tilde{f}(v_{\chi\parallel}) = \int_{-\sqrt{1-(v_{\chi\parallel}/v_{esc})^2}}^{\sqrt{1-(v_{\chi\parallel}/v_{esc})^2}} d\cos\theta_{\chi} \frac{1}{2\sin\theta_{\chi}} f(\frac{v_{\chi\parallel}}{\sin\theta_{\chi}})$ 

★ Revolution of the solar system around the GC:  $f(v_{\chi}) \rightarrow F(V_{\chi})$  with  $V_{\chi} \equiv |\vec{v}_{\chi} + \vec{v}_{\odot}|$ 

$$n_{\text{eve}}(\Theta) = \frac{\rho_{\chi}}{m_{\chi}} \int dE_r dV_{\chi \parallel} \tilde{F}(V_{\chi \parallel}; \Theta) \frac{d}{dE_r} (\bar{N}_{\text{T}} \langle \sigma_{\chi T} V_{\text{rel}\parallel}(\Theta) \rangle)$$

For  $\Theta$  between the Cygnus direction and  $\hat{n}$ , a plane-projection procedure of  $F(V_{\chi})$  should be done individually  $\rightarrow n_{eve}$  depends on  $\Theta$  non-trivially.

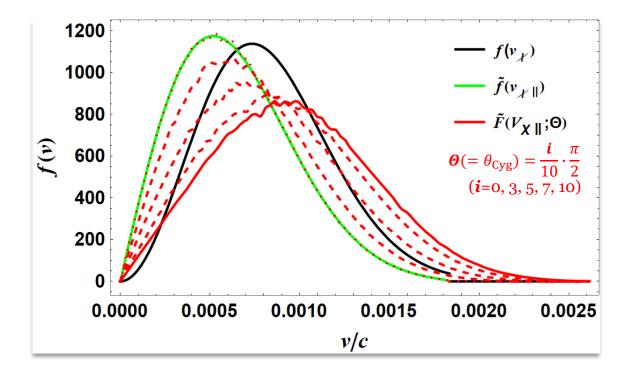
**χ: DM** T: target Cygnus

Θ

#### Angular Dependence -> Angular Modulation

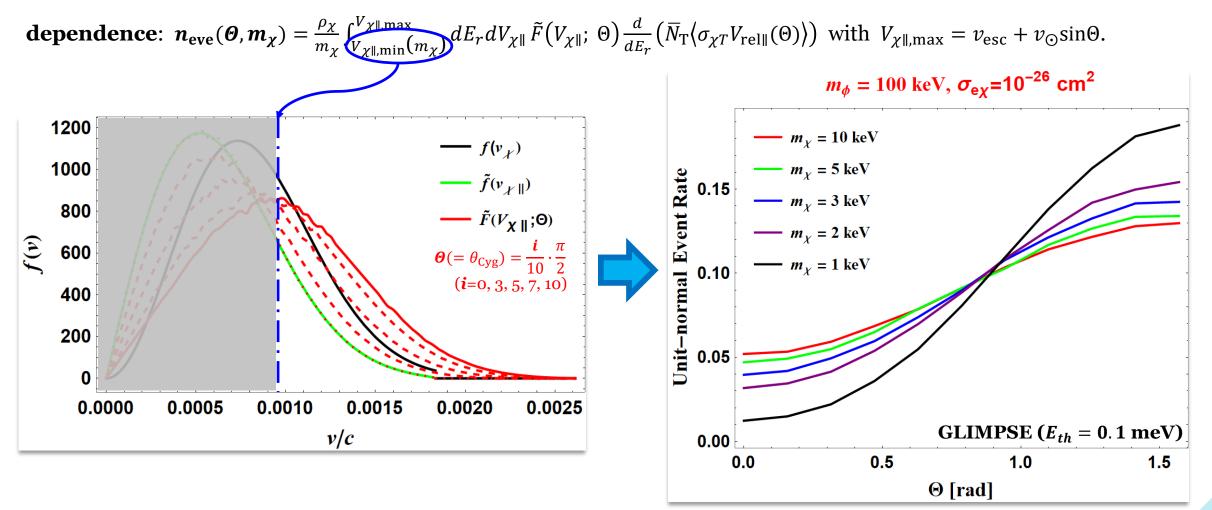
- ♦  $E_{\text{th}} \neq 0 \rightarrow V_{\chi\parallel,\min}$  for DM signal detection.
- \* For smaller  $m_{\chi}$ , larger  $V_{\chi\parallel}$  is required.  $\rightarrow$  A dependence of  $n_{eve}$  on  $m_{\chi}$  through  $V_{\chi\parallel,\min}(m_{\chi}) \rightarrow$  The curvature of the  $\Theta$

**dependence**:  $n_{\text{eve}}(\Theta, m_{\chi}) = \frac{\rho_{\chi}}{m_{\chi}} \int_{V_{\chi\parallel,\min}(m_{\chi})}^{V_{\chi\parallel,\max}} dE_r dV_{\chi\parallel} \tilde{F}(V_{\chi\parallel}; \Theta) \frac{d}{dE_r} (\overline{N}_{\text{T}} \langle \sigma_{\chi T} V_{\text{rel}\parallel}(\Theta) \rangle) \text{ with } V_{\chi\parallel,\max} = v_{\text{esc}} + v_{\odot} \sin\Theta.$ 



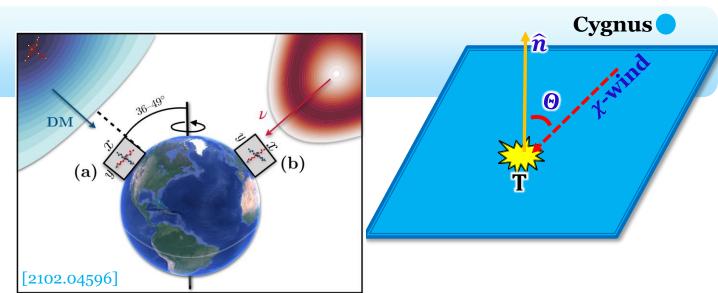
#### Angular Dependence → Angular Modulation

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#### Summary

- > DM flux (DM wind) carries
  - a directional preference: CYGNUS.



**Thank You!** 

➤ Angular modulation → New method for DM mass determination as well as BG rejection!

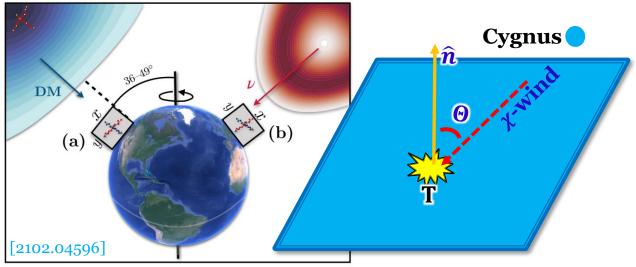
- Senerally applied to the (effectively) 2D or 2D-projectable direct detection experiments allowing for directionality observables
- > Experiments even w/ good  $E_R$ : an additional way to cross-check their results

## Supplemental

### **Angular Modulation vs Annual Modulation**

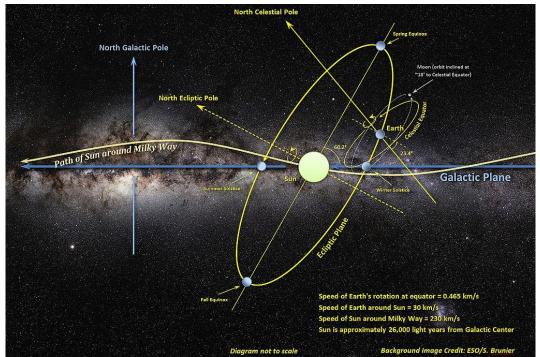
#### ♦ Angular modulation (→ Daily!)

- ✓ Effects from the change of the DM wind direction (𝒫) relative to the plane-normal direction due to Detector's motion
  - → contribution: **revolution**  $\approx$  **rotation**
- $\checkmark N_{\text{event}}(\boldsymbol{\Theta}) \text{ from } N_{\text{event}}(t) \text{ using } \boldsymbol{\Theta}(t)$ 
  - → <u>BG rejection</u> + <u>mass</u> information



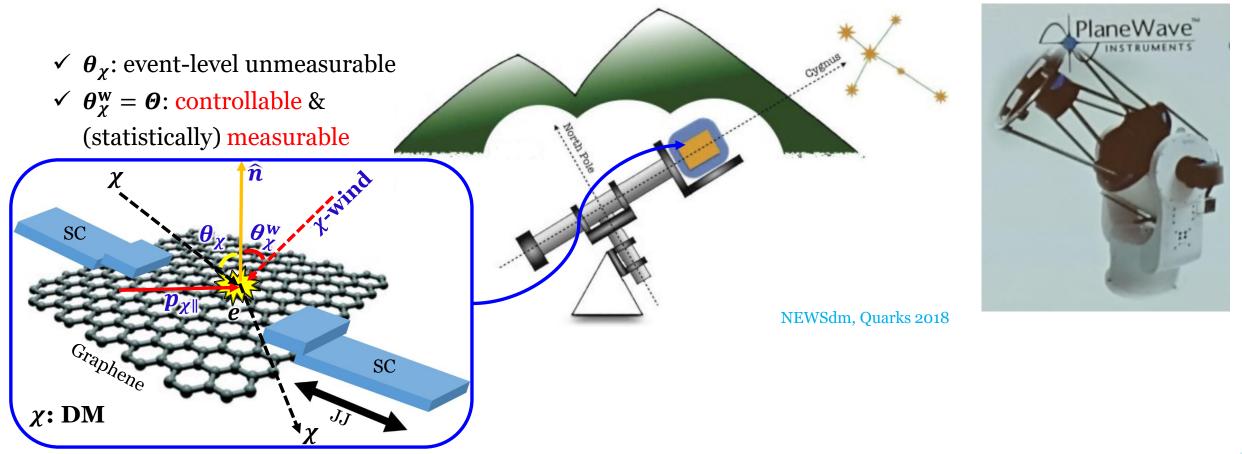
#### **\*** Annual modulation

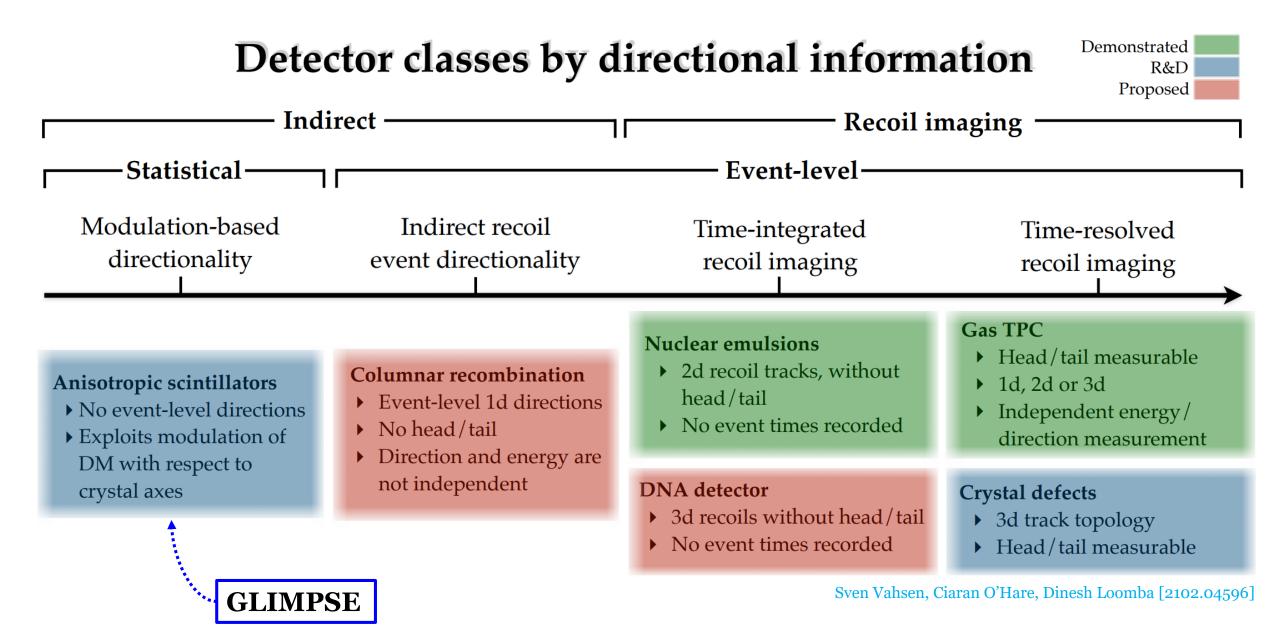
- ✓ Effects from the change of  $|\vec{v}_{rel}|$  due to Earth's motion relative to Sun's motion
  - $\rightarrow$  contribution: **revolution**  $\gg$  rotation
- $\checkmark N_{\text{event}}(t) \rightarrow \text{BG rejection}$



#### **Directional Dependence: Angular Information?**

- $\checkmark$  Actively rotating the detector to run the experiment with a fixed  $\theta_{\chi}^{w} = \Theta$ .
- ✓ **Timing information** of each signal **→ statistically**  $\Theta(t)$ .





## **GLIMPSE** <u>Graphene-based super-Light</u>

Invisible Matter Particle SEarch

[Kim, **JCP**, Lee, Fong, 2002.07821 & in progress]



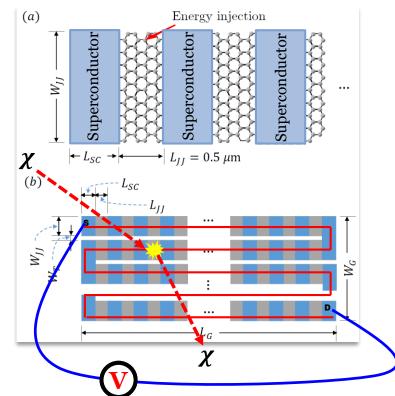
We proposed a new super-light DM direct detection experiment, adopting the Graphene-based Josephson Junction\* (GJJ) microwave single photon detector.

\* A "state-of-the-art" technology:

much lower  $E_{th} \sim O(0.1 \text{ meV})$ 

### **Super-Light DM Direct Search Using GJJs**

[Kim, **JCP**, Lee, Fong, 2002.07821]



I. Single graphene strip (a): the 1D assembly of a long graphene strip &

a number of superconducting material pieces

- → an array of SC-graphene-SC-graphene-SC-… (SGSGS…).
- II. Each sequence of SGS represents a single GJJ device.
- III. 2D detector unit (b): all GJJs are connected in series so that even a single

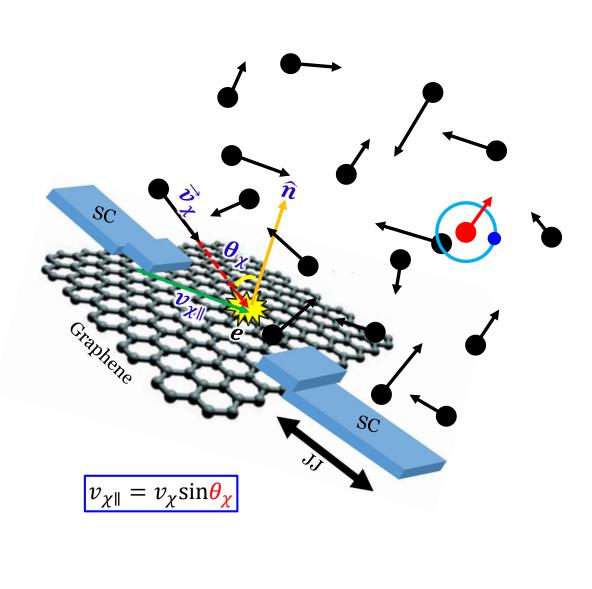
switched GJJ by DM interaction allows the series resistance between S & D

 $\rightarrow$  V changes from 0 to a finite value.

\* A much larger-scale detector can be made of a stack of such detector units (3D).

# **Angular Dependence**

#### **DM** *v* **Distribution: Angle Dependence**

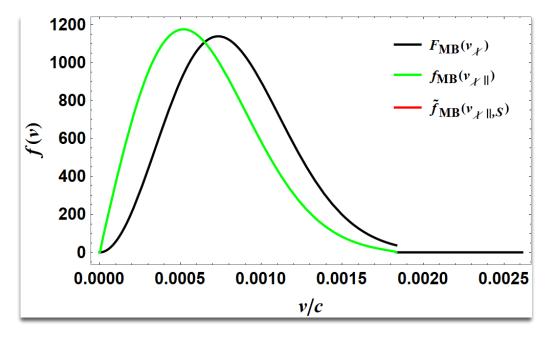


• DM

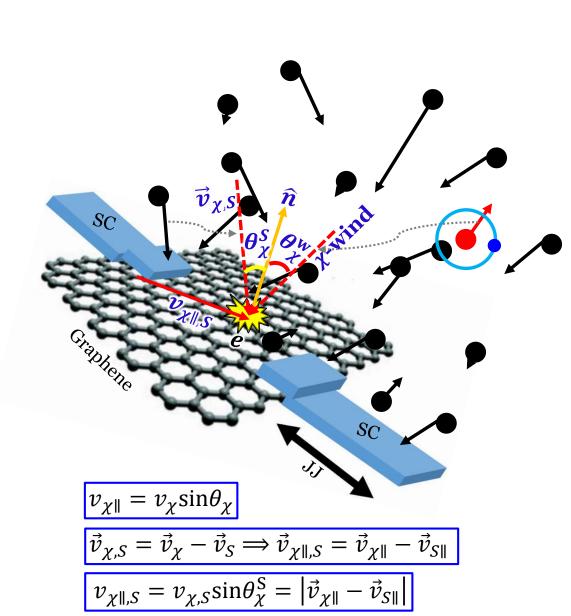
• Sun

• Earth

Cygnus



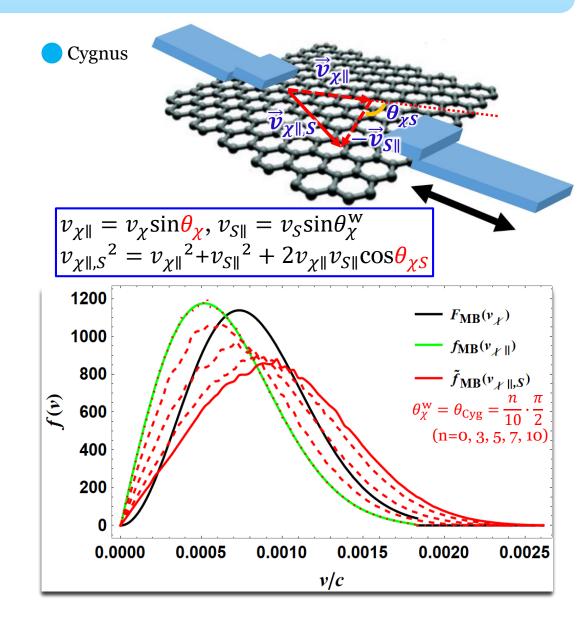
#### **DM** *v* **Distribution: Angle Dependence - Sun**



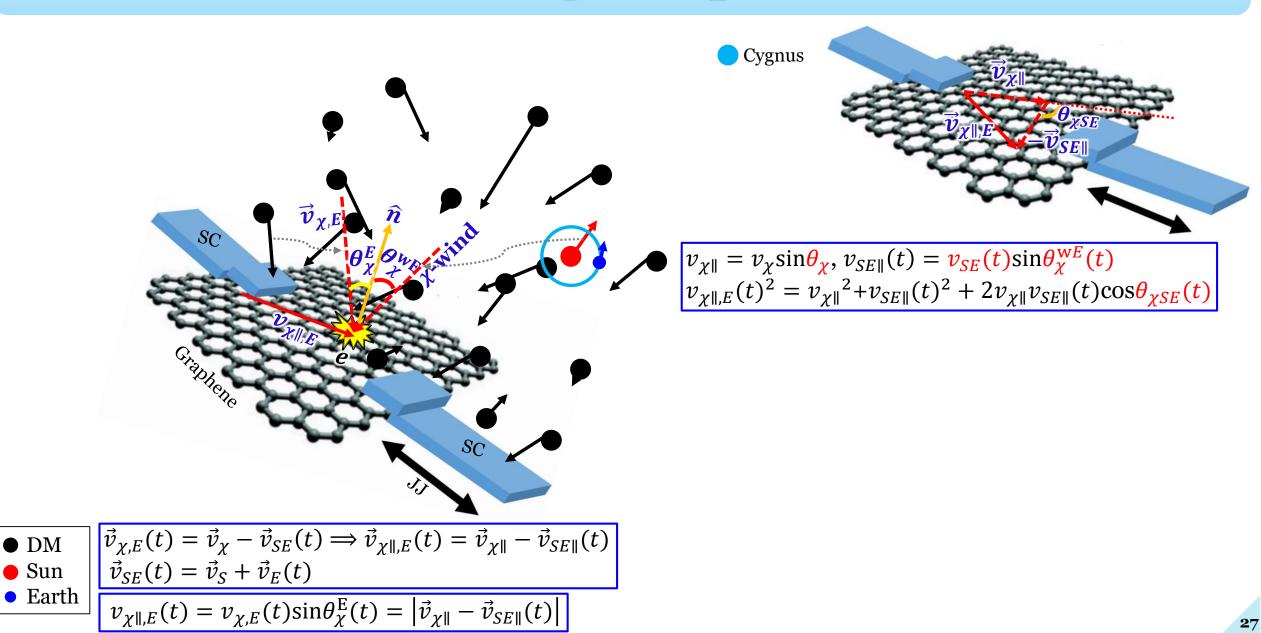
• DM

• Sun

• Earth



#### DM v Distribution: Angle Dependence - Sun-Earth



### **Signal Rate: Directional Dependence**

- **Constrained** elastic scattering: Electron <u>confined in the 2D graphene sheet</u> even after the collision.
- → Momentum transfer: the change of  $p_{\chi\parallel}$  → Signal rate: DM incident direction dependence
- → DM signals: in situ validation by actively rotating the detector or time information of each signal

