

Cooking nuclear pasta at zero temperature: from semi-classical to fully microscopic recipes

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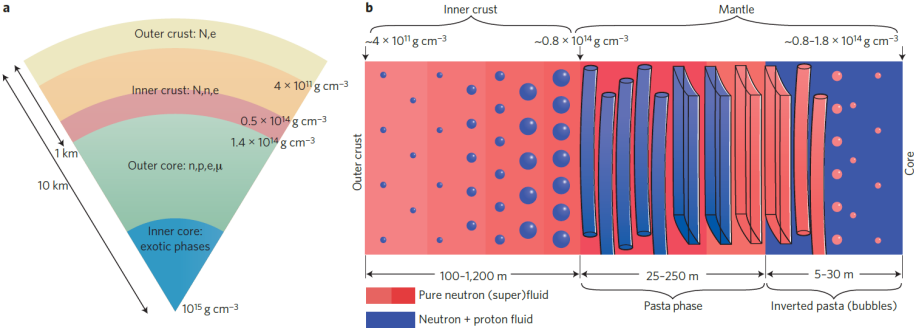
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Karpacz school on theoretical physics, June 6, 2024

Neutron star schematic structure



[Newton+'13]



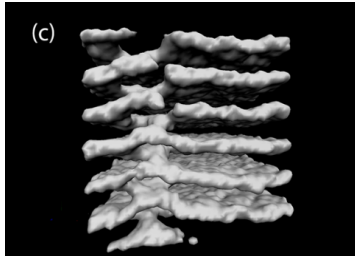
Role of pasta phases

► Transport properties

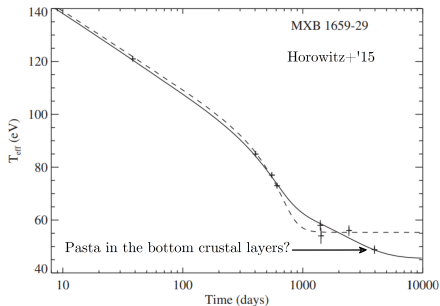
Thermal and electrical conductivities are determined by electron scattering on pasta structures.

Possible manifestations:

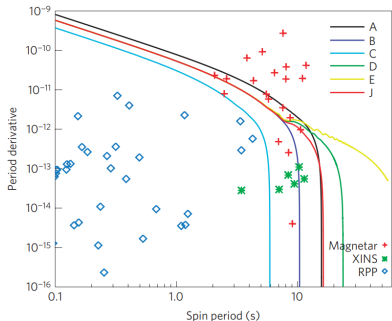
- NS thermal evolution
- Magnetic field evolution



Topological defects in pasta [Horowitz+'15]



[Horowitz+'15], [Pons+'13]



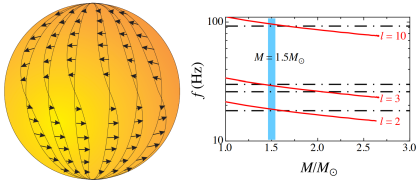
Role of pasta phases

► **Mechanical properties**

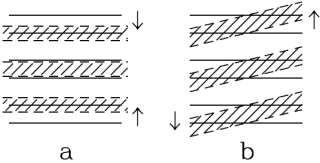
In response to deformations pasta behave like liquid crystals.

Possible manifestations:

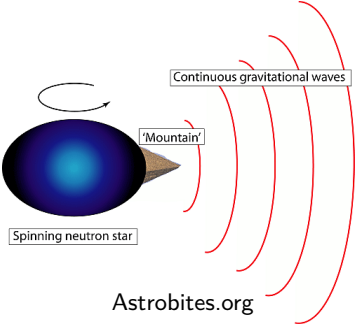
- NS oscillation frequencies
- Gravitational waves emission



[Bastrukov+'09], [Sotani'11]



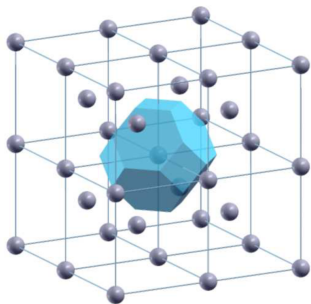
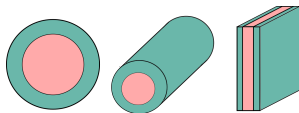
Pasta deformations [Pethick&Potekhin'98]



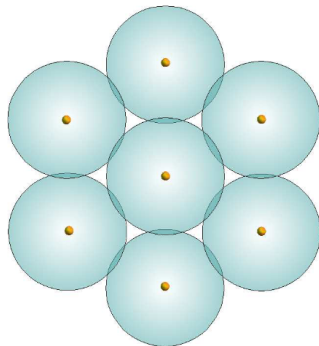
Astrobit.es.org

Modelling pasta

- ▶ It is assumed that matter is in its ground state at $T = 0 \Rightarrow$ cold catalyzed matter
- ▶ Matter is divided into identical cells
- ▶ Energy per particle is minimized at given density

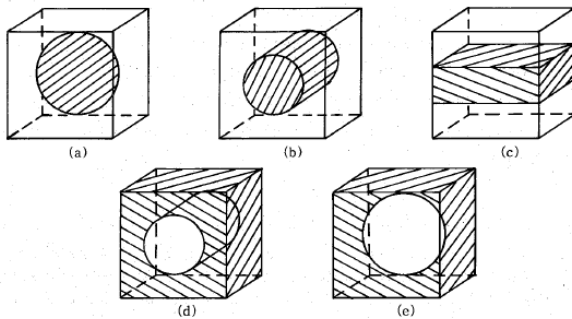


Body-centered cubic lattice of quasi-spherical nuclei [Chamel&Haensel'08]



Spherical Wigner-Seitz approximation [Chamel&Haensel'08]

Recipe I. Liquid drop models



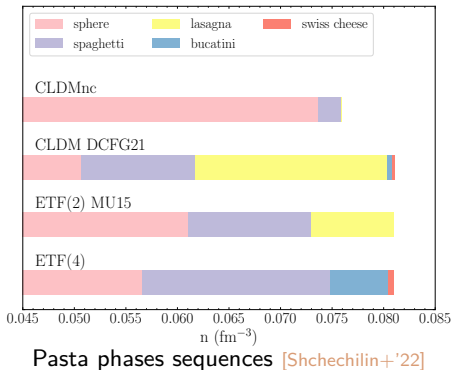
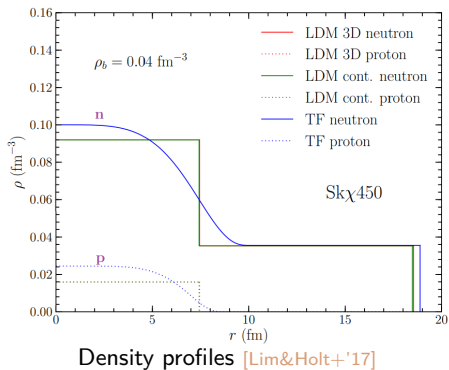
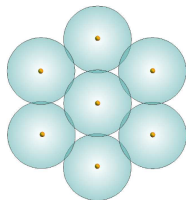
Pasta shapes [Hashimoto+'84]

- ▶ First predicted in liquid-drop models [Ravenhall+'83], [Hashimoto+'84]
- ▶ $E_{\text{tot}} = E_{\text{bulk}} + E_{\text{surf}} + E_{\text{coul}}$
- ▶ At certain filling factor $u = V_{\text{cl}}/V$, E_{surf} is reduced by shape rearrangements

Recipe II. Semi-classical models

Extended Thomas-Fermi (ETF) method

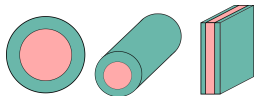
- Describes smooth nucleons distributions ρ
- Surface energy is included self-consistently
- **lack of shell effects**



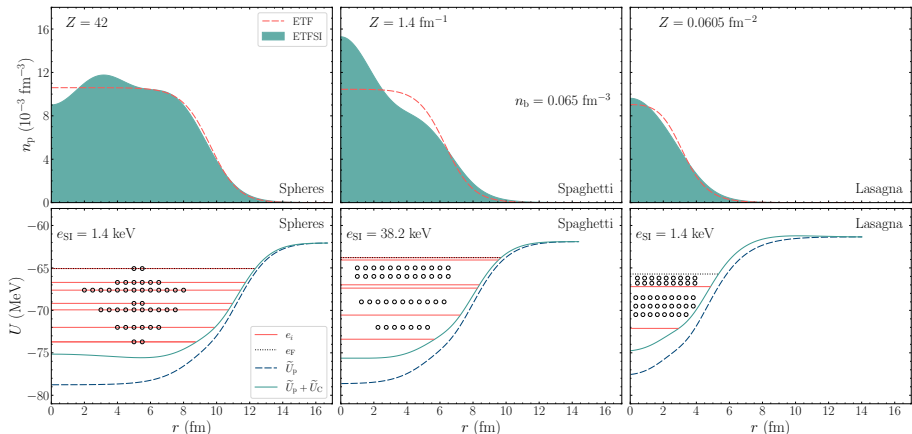
Recipe II*. Semi-classical with microscopic corrections

Extended Thomas-Fermi + Strutinsky Integral (ETFSI)

- Construct central and Coulomb fields U_{ETF} from the ETF density profiles ρ
- Solve Hartree-Fock equations for ψ with U_{ETF}
- Add microscopic (shell) effects perturbatively

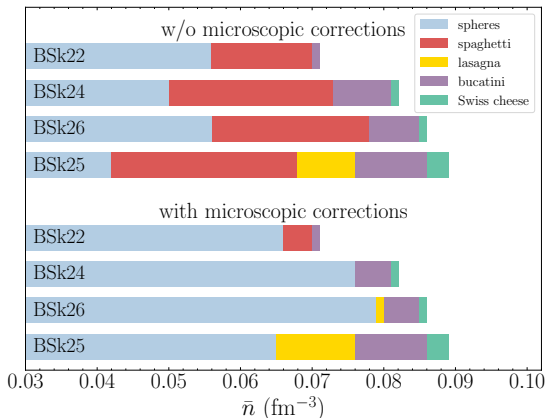


[Pearson&Chamel'22], [Shchechilin+'24]



Semi-classical models with microscopic corrections

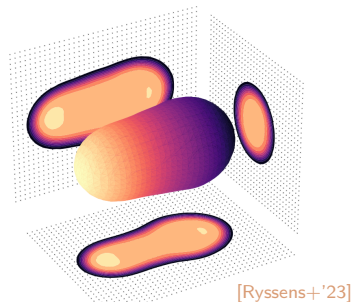
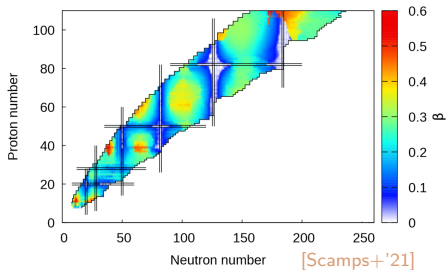
- ▶ Allows fast estimation of the role of the nuclear interaction employed
- ▶ Microscopic effects are important \Rightarrow **call for the fully quantum calculations**



Pasta sequences with nuclear functionals differing in symmetry energy [Shchepochin+'23]

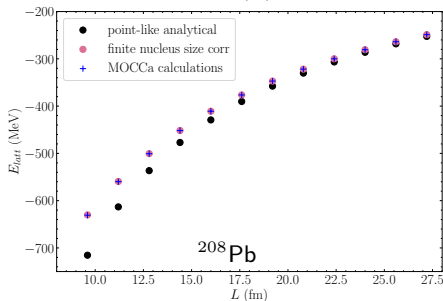
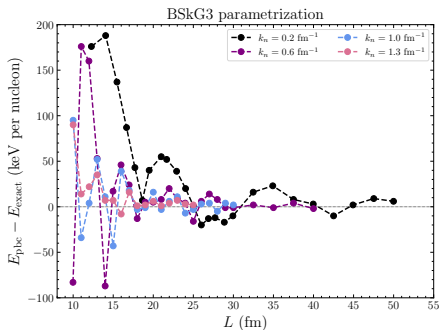
Recipe III. Microscopic models

- ▶ Hartree-Fock-Bogoliubov method
 - Mean-field approach for finding self-consistent solution for ψ
 - Pairing correlations are included
- ▶ 3D HFB code MOCCa [Ryssens+'15]
 - tool of choice for deformed nuclei
- ▶ BSkG3 parametrization of generalized Skyrme force [Grams+'23]
 - rms (exp masses): 0.631 MeV
 - Max NS mass: $2.3 M_{\odot}$
 - \Rightarrow suitable for unified EOS



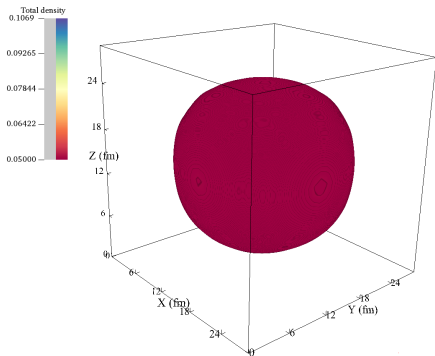
MOCCa code generalization

- ▶ We implement periodic boundary conditions
 - Pure neutron matter energy tends to analytical result at large box sizes
 - Lattice energy agrees with analytic expression for simple cubic lattice
- ▶ We scale the problem for the large number of particles $\mathcal{N} \sim 5000$

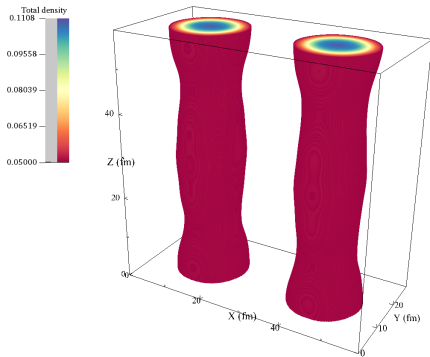


Microscopic models for pasta

- ▶ **Proof of principle** calculations with maximal symmetries and no pairing (Hartree-Fock in 1/8 of the full box).



(Spherical shape) $Z = 40$, $A = 1318$,
 $\bar{\rho} = 0.05 \text{ fm}^{-3}$



(Spaghetti shapes) $Z = 160$, $A = 5240$,
 $\bar{\rho} = 0.05 \text{ fm}^{-3}$

Conclusions

► Motivation

- Pasta in neutron stars can impact: cooling, gravitational waves, etc.
- Reliable calculations of pasta are needed

► Main results

- We show within the perturbative approach that microscopic effects are important
- To take them into account fully self-consistently we adapt the 3D Hartree-Fock-Bogoliubov code MOCCa for pasta calculations

► Perspectives

- Calculation of the mantle composition and equation of state
- Investigation of dynamics and stability of pasta structures

