

# Do Neutron Stars give us Nuclear Matter properties?



## Machine Learning answers

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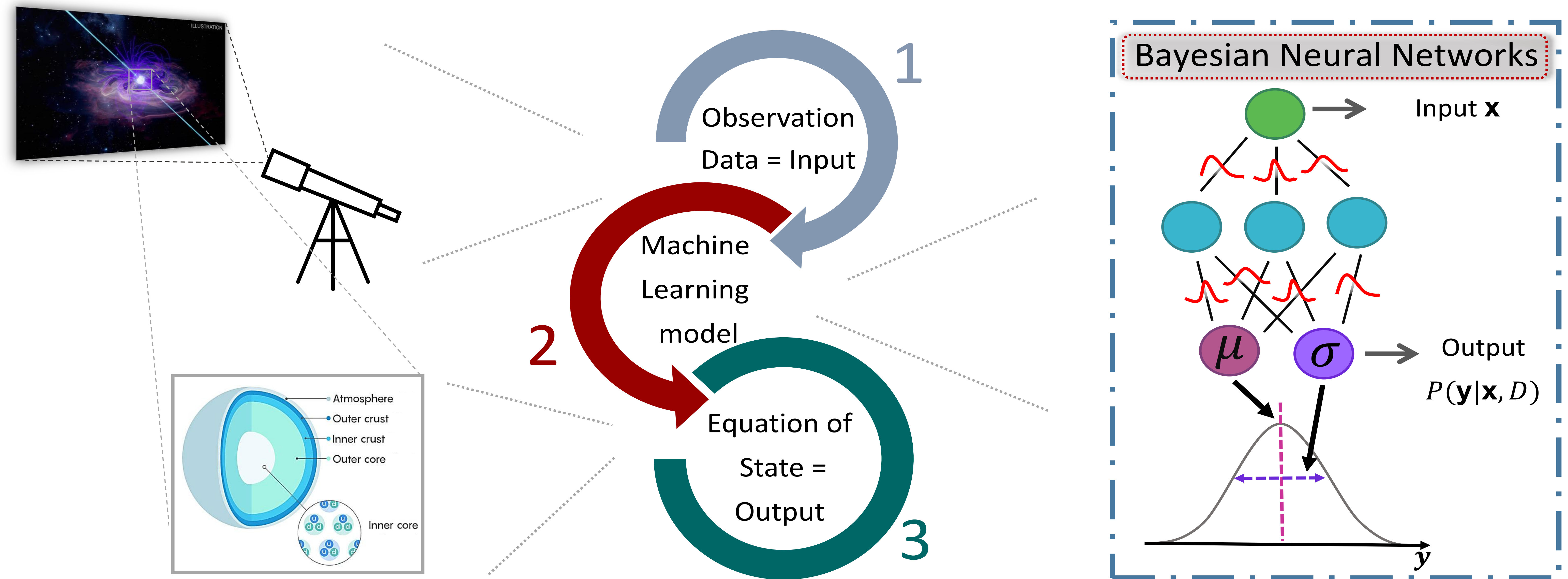
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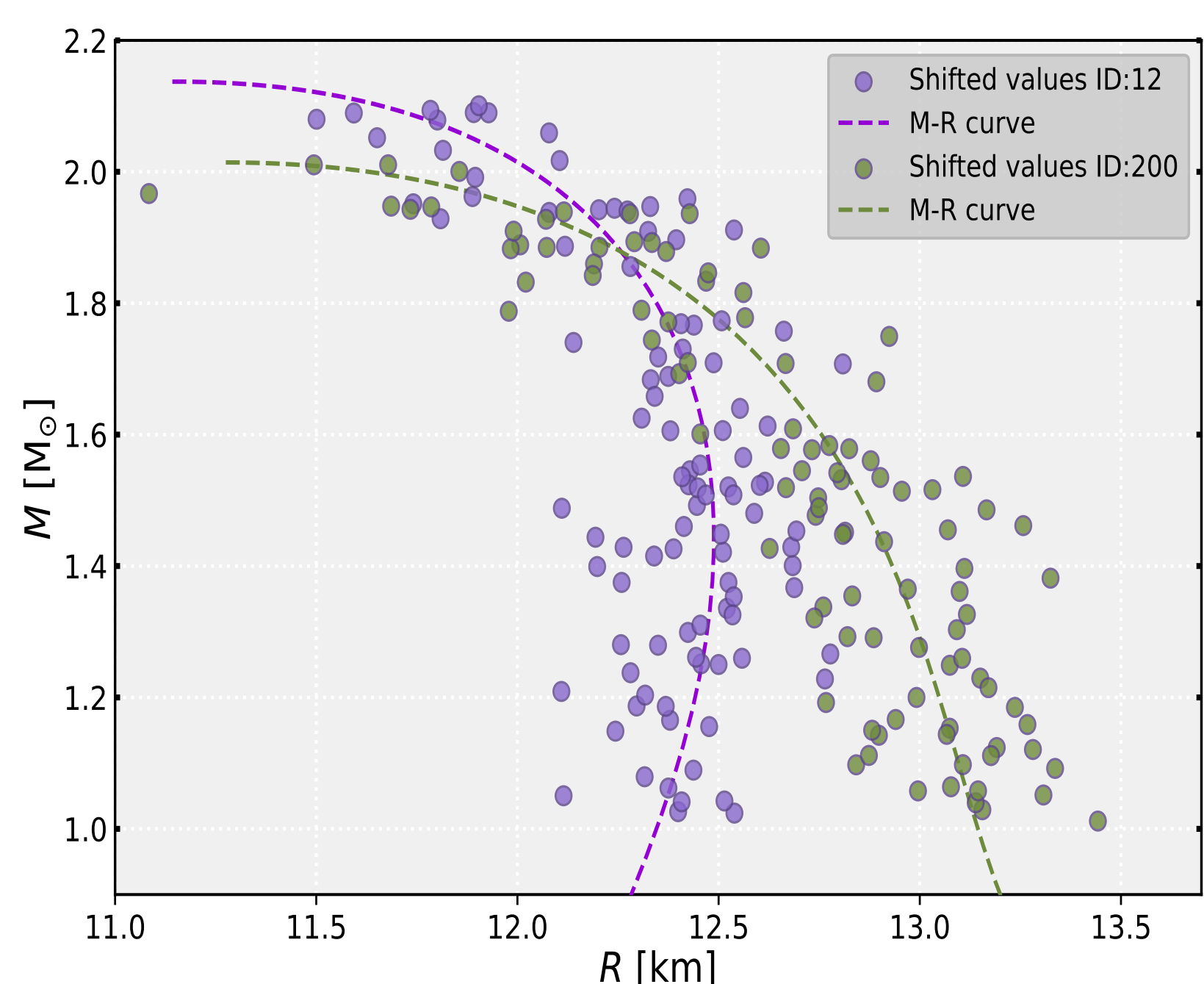
### Overview

Neutron stars are a remarkable object towards the study of nuclear matter properties under extreme conditions. In this work we extract some properties contained in the Equation of State (EoS) **3** with the help of Bayesian Neural Networks (BNN) **2** using the mass and the radius from observations **1**.



### 1 Input features

Input vector is:  $\mathbf{x}_i = [M_1, M_2, M_3, M_4, M_5, R_1, R_2, R_3, R_4, R_5]$



How we do data augmentation in the input:

$$M_i^{(0)} = \mathcal{U}[M_{\odot}, M_{Max}],$$

$$M_i = \mathcal{N}(M_i^{(0)}, \sigma_M),$$

$$R_i = \mathcal{N}(\mathbf{R}(M_i^{(0)}), \sigma_R)$$

( $i = 1, \dots, 5$ )

The process is repeated  $n_s$  times for the same EoS.

### 2 What is a Bayesian Neural Network?

Is the combination between Neural Networks (NN) and Bayesian Inference (BI).

NN: Mapping of Input space  $\mathbf{x}$  onto the output space  $\mathbf{y}$  by several successive layers of linear transformations (given by the weights  $\mathbf{w}$ ) interleaved with element-wise non-linear transforms. The probabilistic view is  $P(\mathbf{y}|\mathbf{x}, \mathbf{w})$ .

BI: The calculus of the posterior distribution of the weights, given the training data,  $D = (\mathbf{x}, \mathbf{y})$ :

$$P(\mathbf{w}|D) = \frac{\text{Prior Likelihood}}{\text{Evidence}} = \frac{P(\mathbf{w}) P(D|\mathbf{w})}{P(D)}$$

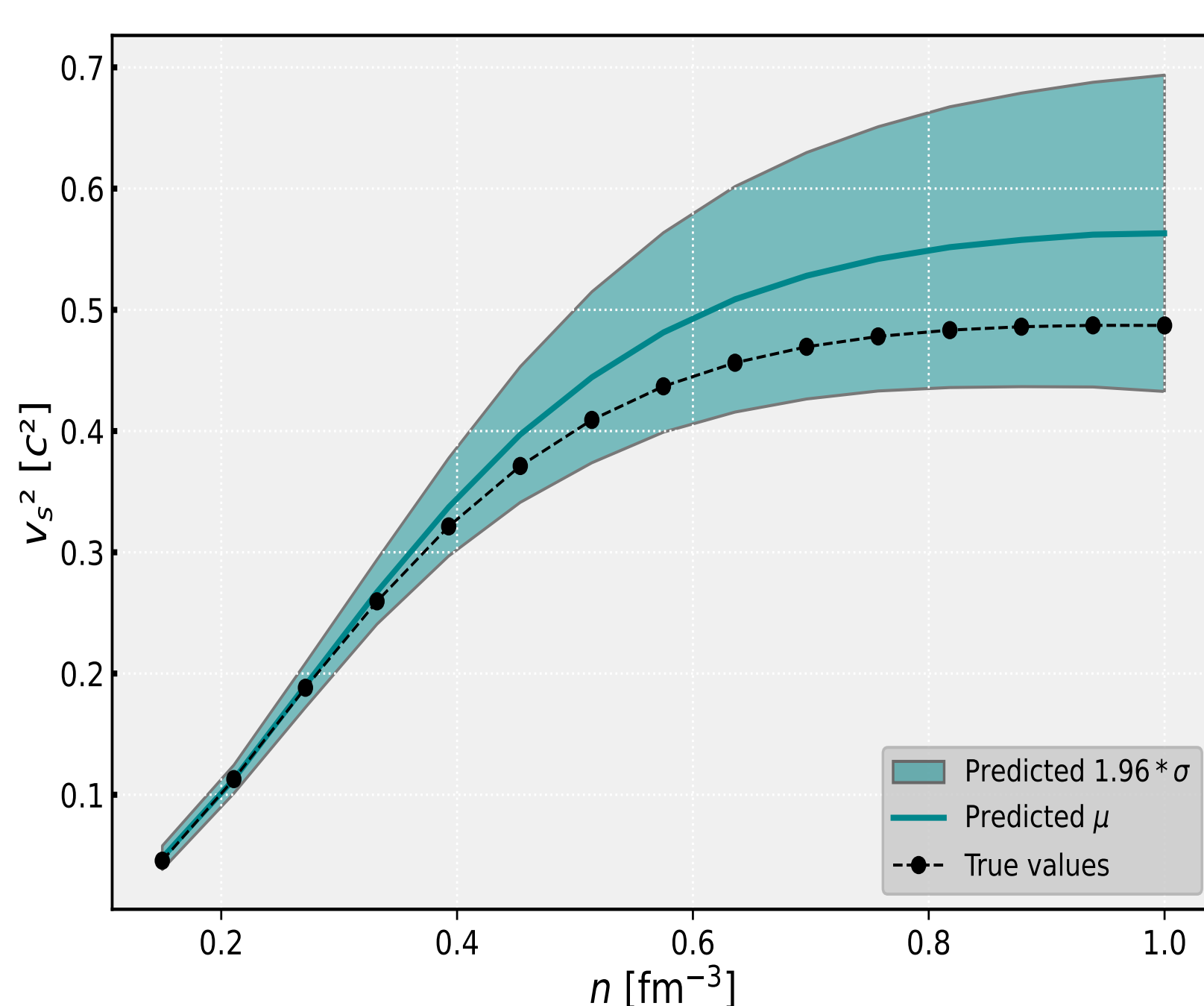
The prediction distribution of a BNN is then for an unknown  $\hat{\mathbf{y}}$  of a test data item  $\hat{\mathbf{x}}$ :

$$P(\hat{\mathbf{y}}|\hat{\mathbf{x}}, D) = \int P(\hat{\mathbf{y}}|\hat{\mathbf{x}}, \mathbf{w}) P(\mathbf{w}|D) d\mathbf{w}$$

### 3 Output predictions

Output vector is:  $\mathbf{y}_i = [v_s^2(n_1), \dots, v_s^2(n_{15})]$ ,  $n_i$  = baryonic density.

Our 15 true values for one EoS of the test set are all contained in the 95% Confidence Interval predicted by the model.



### Ongoing work

- Working with other outputs, such as pressure and proton fraction;
- Using tidal deformability in the input;
- Testing how data augmentation affects our results;
- And so on ...