

Probing Muon ($g-2$) at Forward Detectors at LHC

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UC Irvine

FLASY 2024

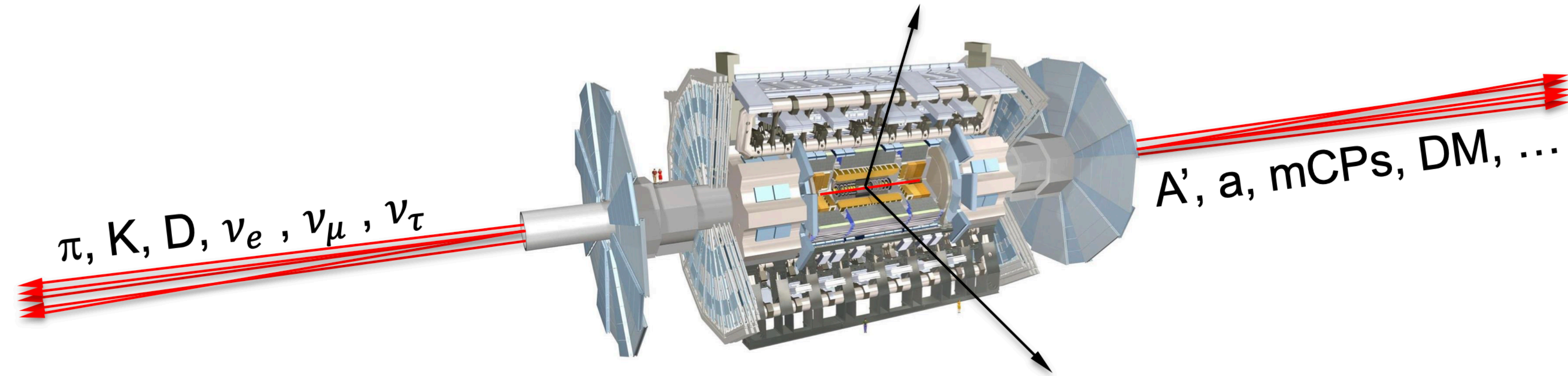
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June 28th, 2024

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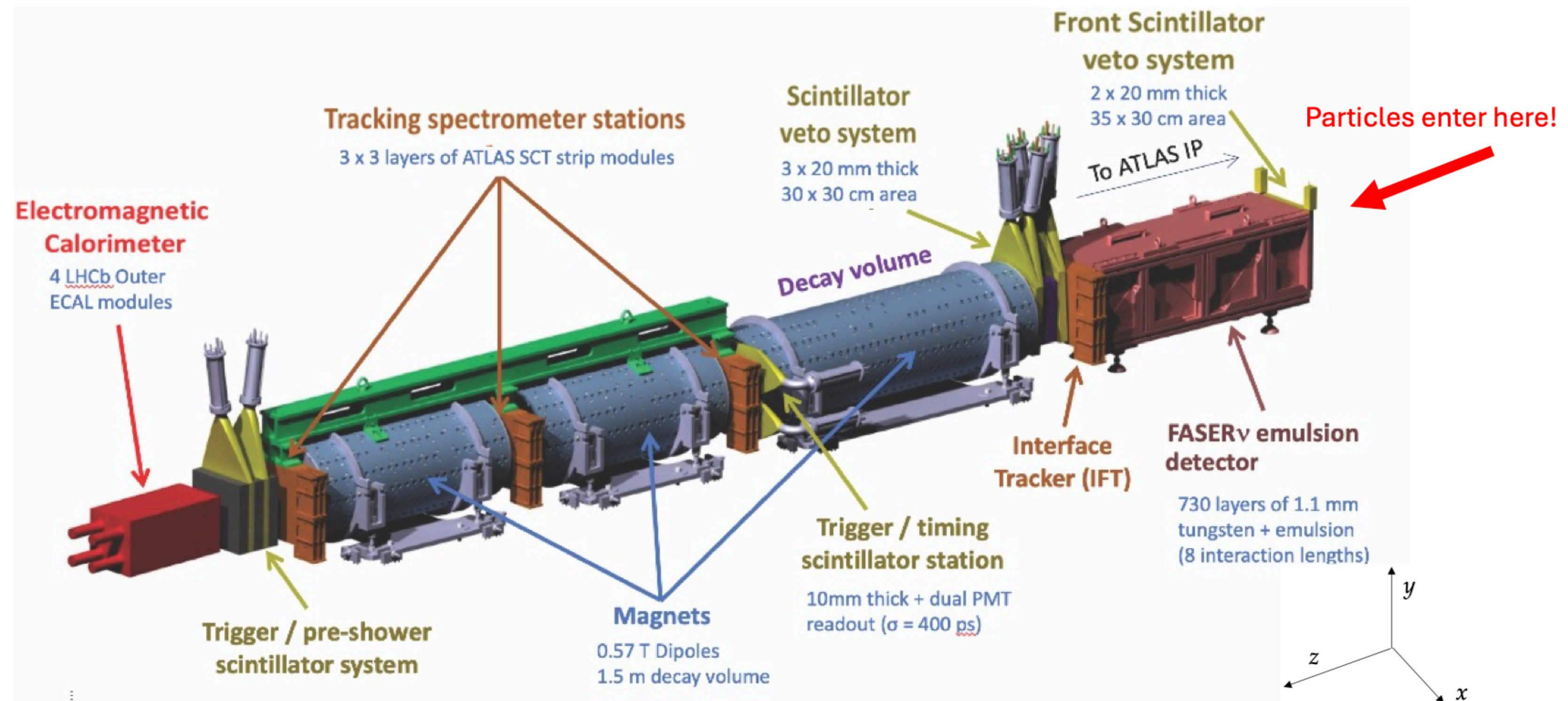
Forward Region at the LHC



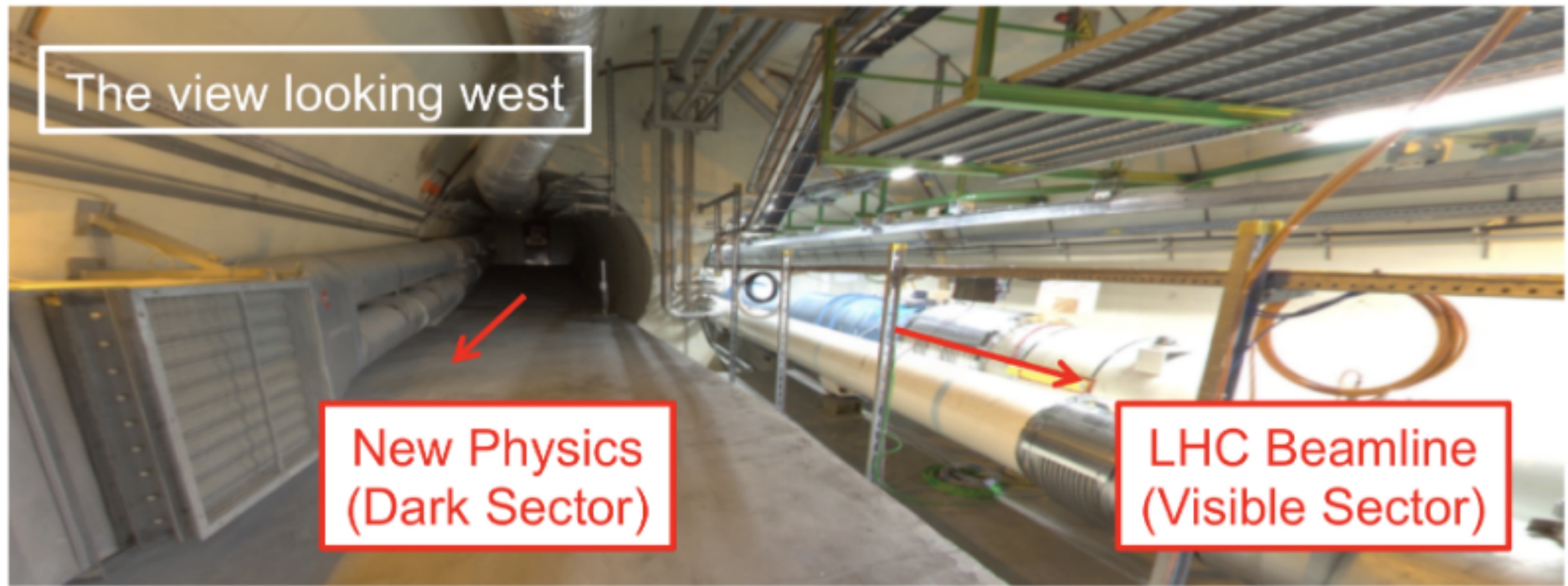
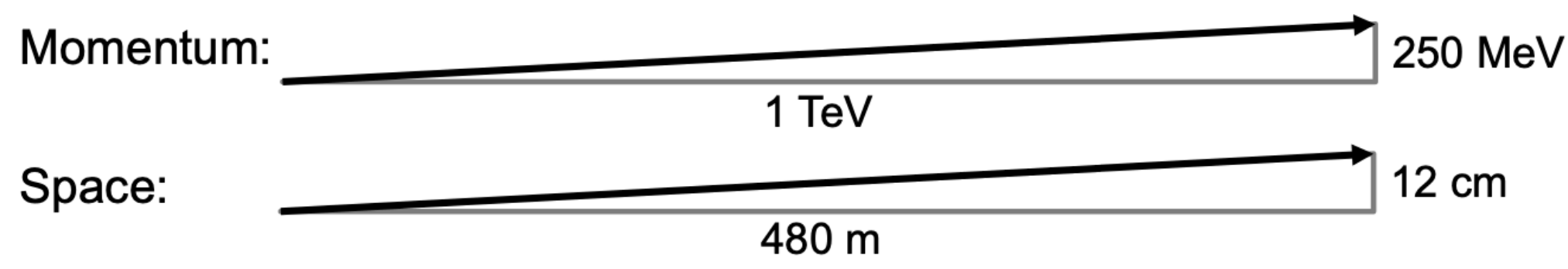
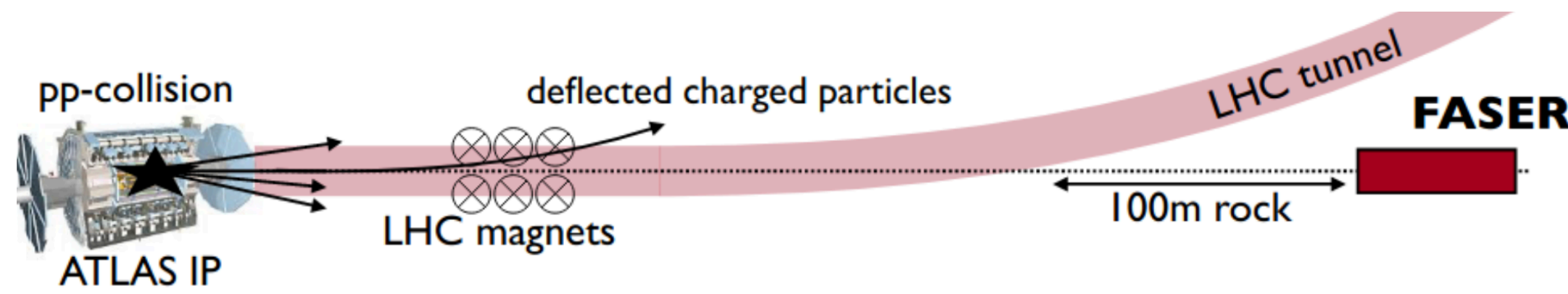
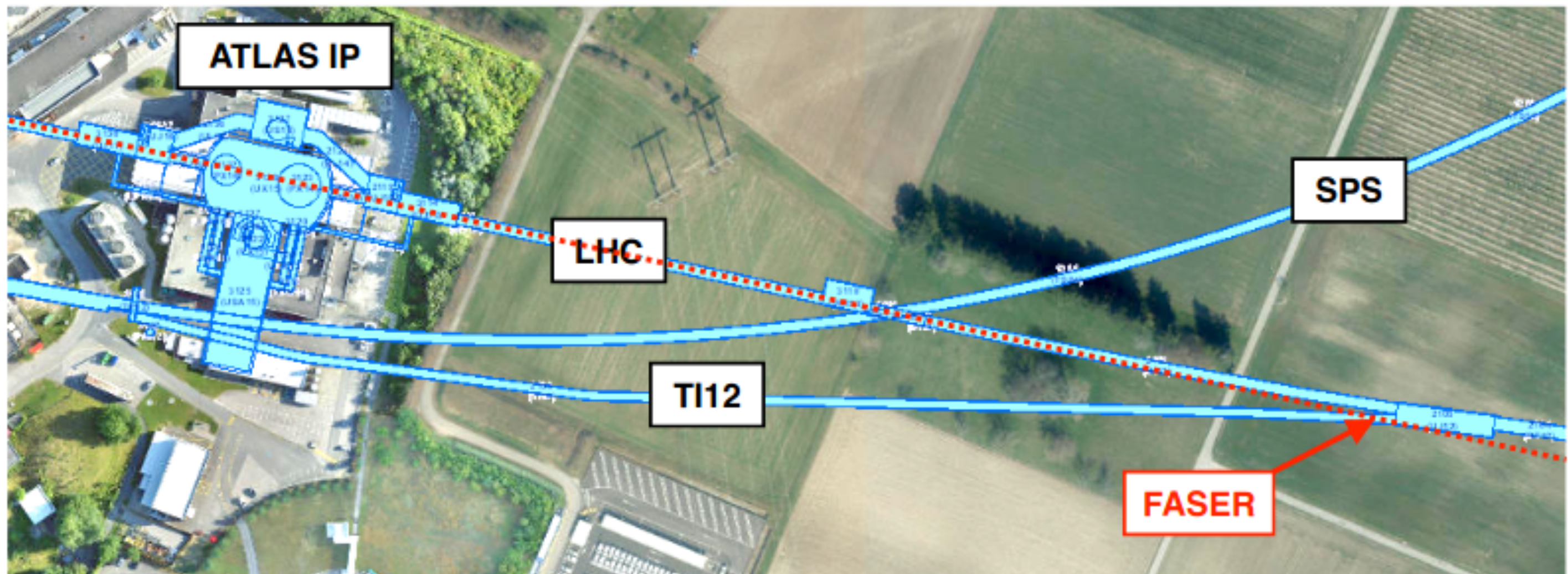
- pp collisions at the LHC produce an intense flux of particles in the forward direction
- These particles are light and weakly coupling:
 - SM (ν, μ, \dots) and BSM (ALPs, dark photon, DM, ...)
- Conventional transverse detectors will miss these particles

ForwArd Search Experiment(ν) - FASER(ν)

- FASER: 25cm x 25cm x 1.5m decay volume
 - 1708.09389 (first paper), 1811.10243 (LOI), 1812.09139
- FASER ν : 25cm x 25cm x 1m tungsten emulsion detector
 - 1908.02310, 2001.03073
- $\eta \gtrsim 8.5$ coverage.



Location for forward detectors at LHC



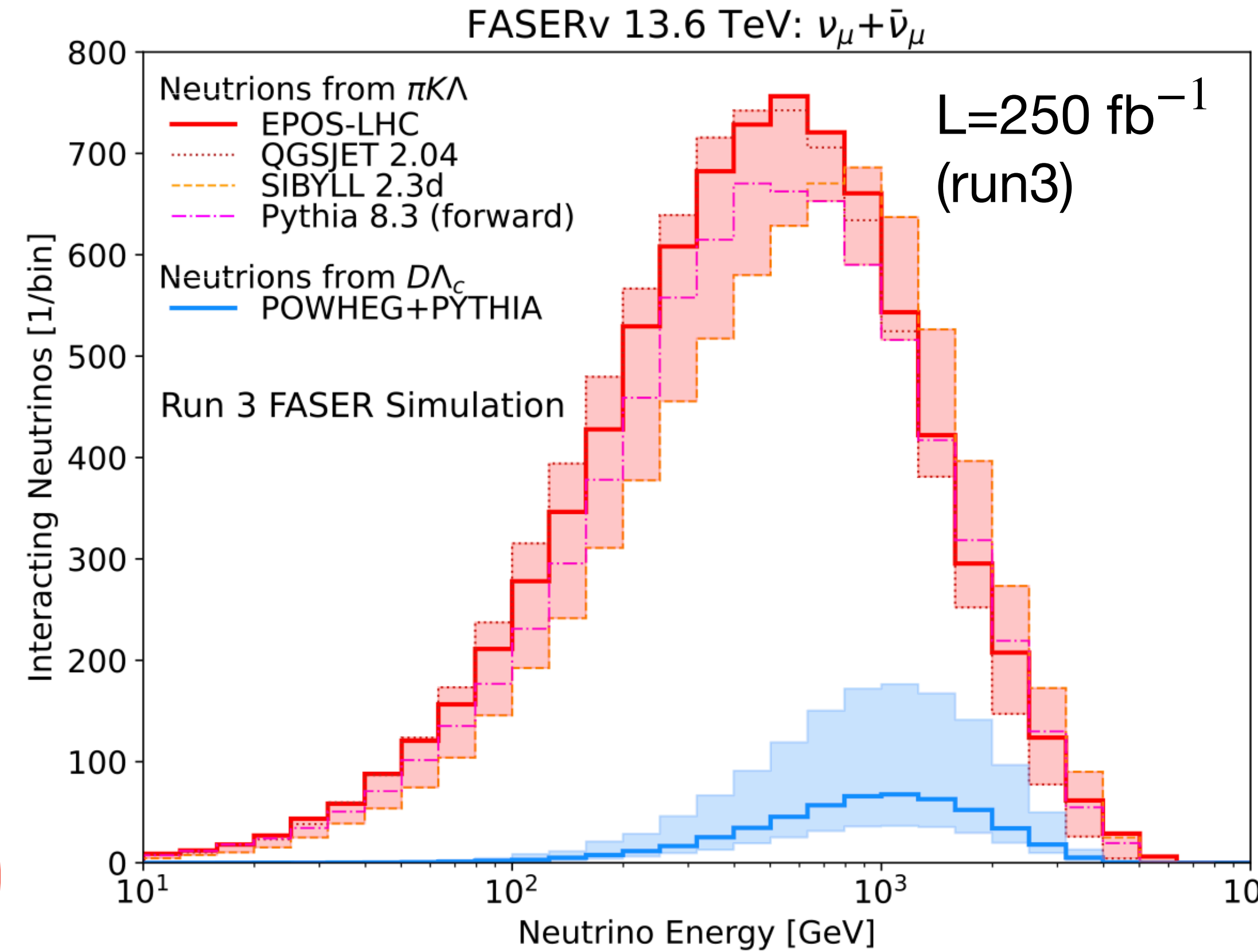
Neutrino Flux at FASER

$$\nu_e: K \longrightarrow \pi e \nu_e, D \longrightarrow K e \nu_e$$

$$\nu_\mu: \pi^\pm \longrightarrow \mu \nu_\mu, K^\pm \longrightarrow \mu \nu_\mu$$

Generators		FASER ν at Run 3			FASER ν at Run 4		
light hadrons	charm hadrons	$\nu_e + \bar{\nu}_e$	$\nu_\mu + \bar{\nu}_\mu$	$\nu_\tau + \bar{\nu}_\tau$	$\nu_e + \bar{\nu}_e$	$\nu_\mu + \bar{\nu}_\mu$	$\nu_\tau + \bar{\nu}_\tau$
EPOS-LHC	–	1149	7996	–	3382	23054	–
SIBYLL 2.3d	–	1126	7261	–	3404	21532	–
QGSJET 2.04	–	1181	8126	–	3379	22501	–
PYTHIAforward	–	1008	7418	–	2925	20508	–
–	POWHEG Max	1405	1373	76	4264	4068	255
–	POWHEG	527	511	28	1537	1499	91
–	POWHEG Min	294	284	16	853	826	51
Combination		1675^{+911}_{-372}	8507^{+992}_{-962}	28^{+48}_{-12}	4919^{+2748}_{-1141}	24553^{+2568}_{-3219}	91^{+163}_{-41}

CC events



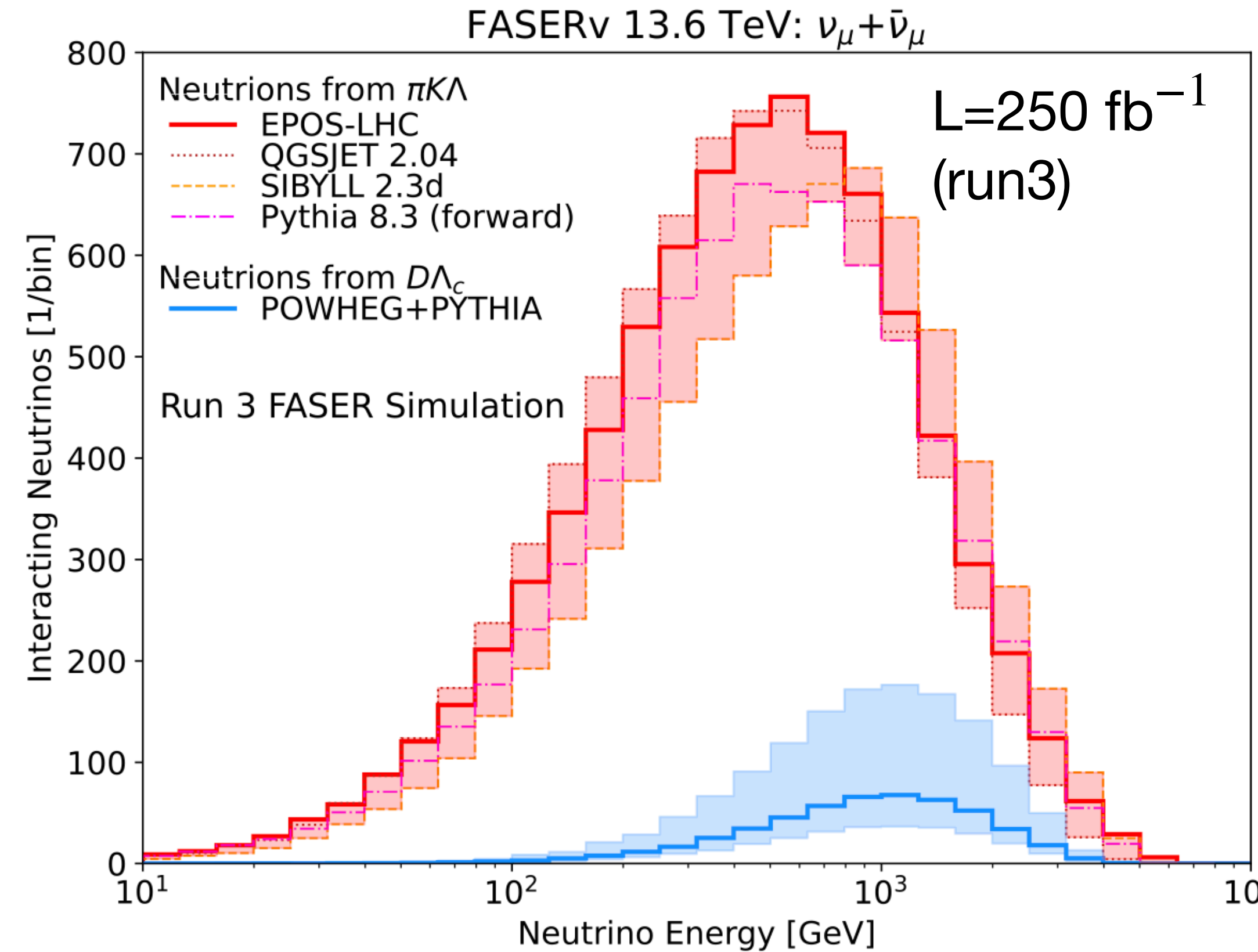
Neutrino Rate Predictions for FASER;
2402.13318

Neutrino Flux at FASER

$$\nu_e: K \longrightarrow \pi e \nu_e, D \longrightarrow K e \nu_e$$

$$\nu_\mu: \pi^\pm \longrightarrow \mu \nu_\mu, K^\pm \longrightarrow \mu \nu_\mu$$

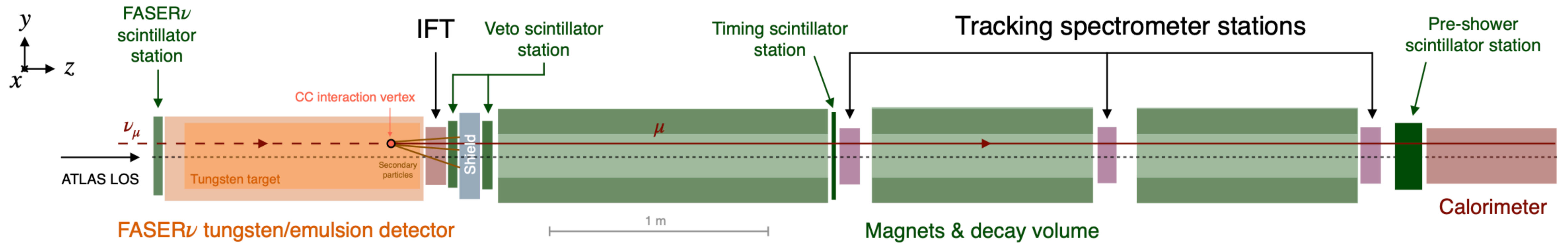
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light hadrons	charm hadrons	$\nu_e + \bar{\nu}_e$	$\nu_\mu + \bar{\nu}_\mu$	$\nu_\tau + \bar{\nu}_\tau$	$\nu_e + \bar{\nu}_e$	$\nu_\mu + \bar{\nu}_\mu$	$\nu_\tau + \bar{\nu}_\tau$
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Already many new exciting results!!!

Neutrino Rate Predictions for FASER,
2402.13318

First Observation of Collider Neutrinos

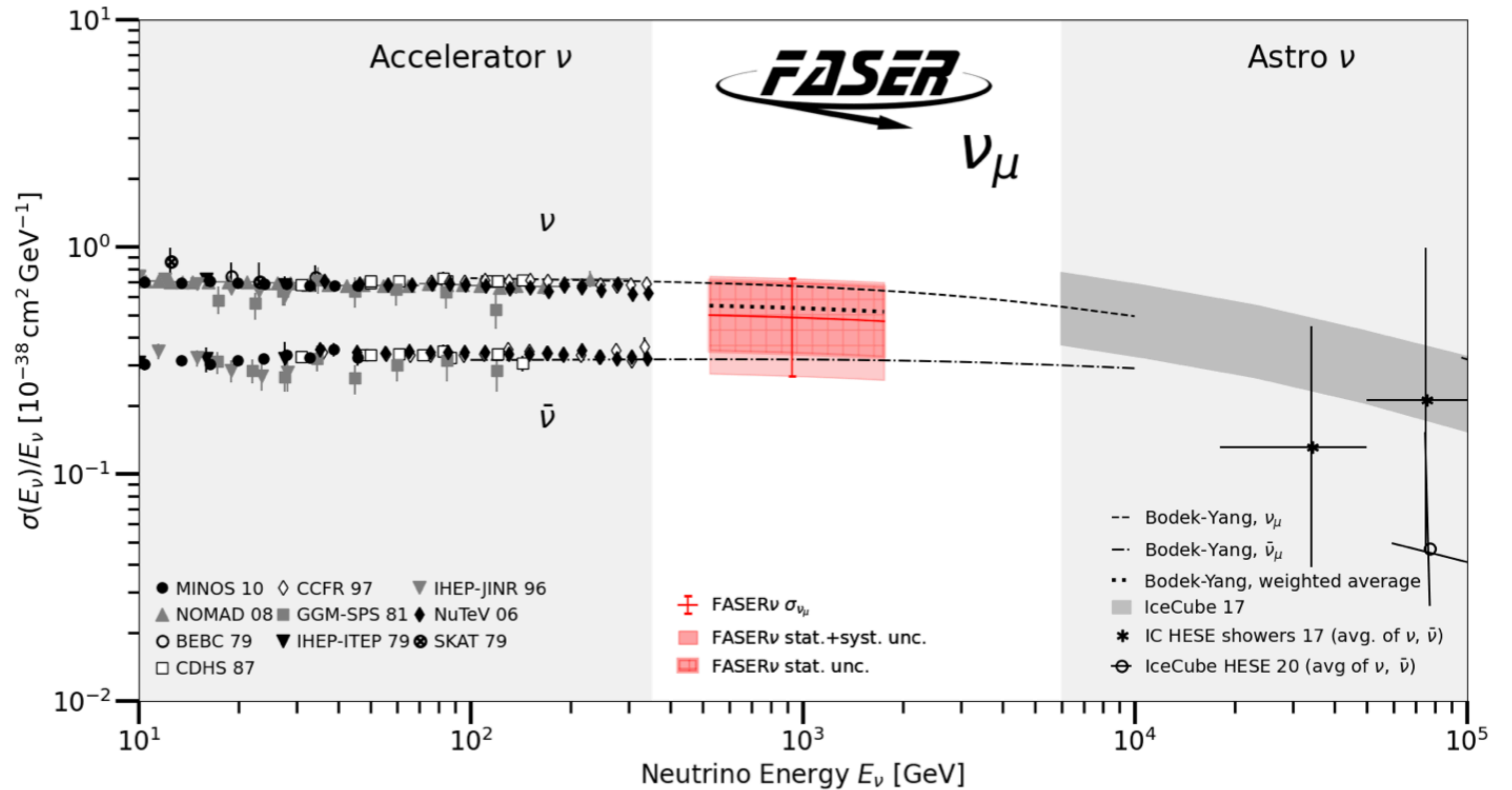
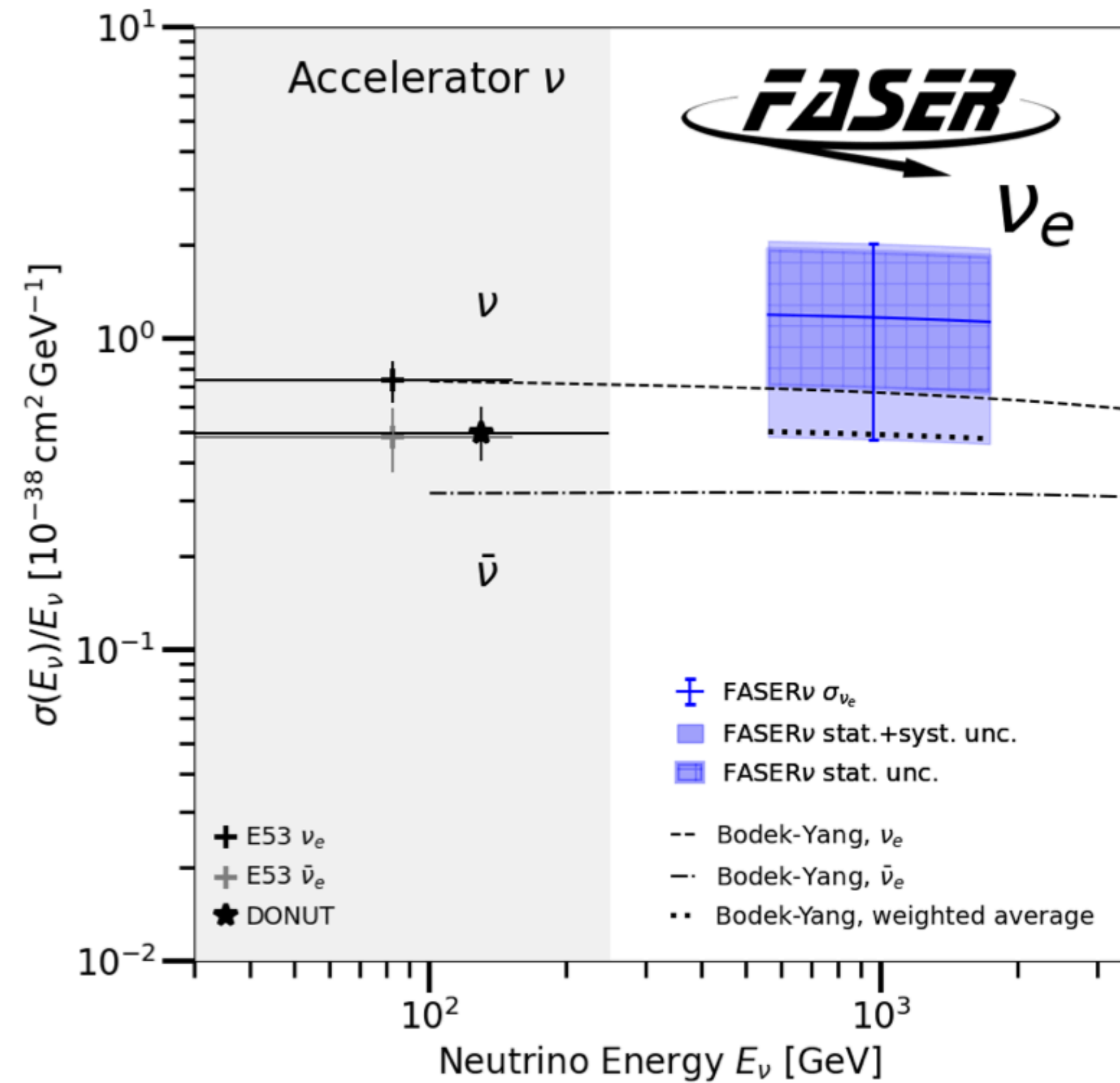


At FASER

**$\sim 150 \nu_{\mu}$ CC events with
 35.4 fb^{-1} of data.**

First Direct Observation of Collider Neutrinos with FASER at the LHC; 2303.14185

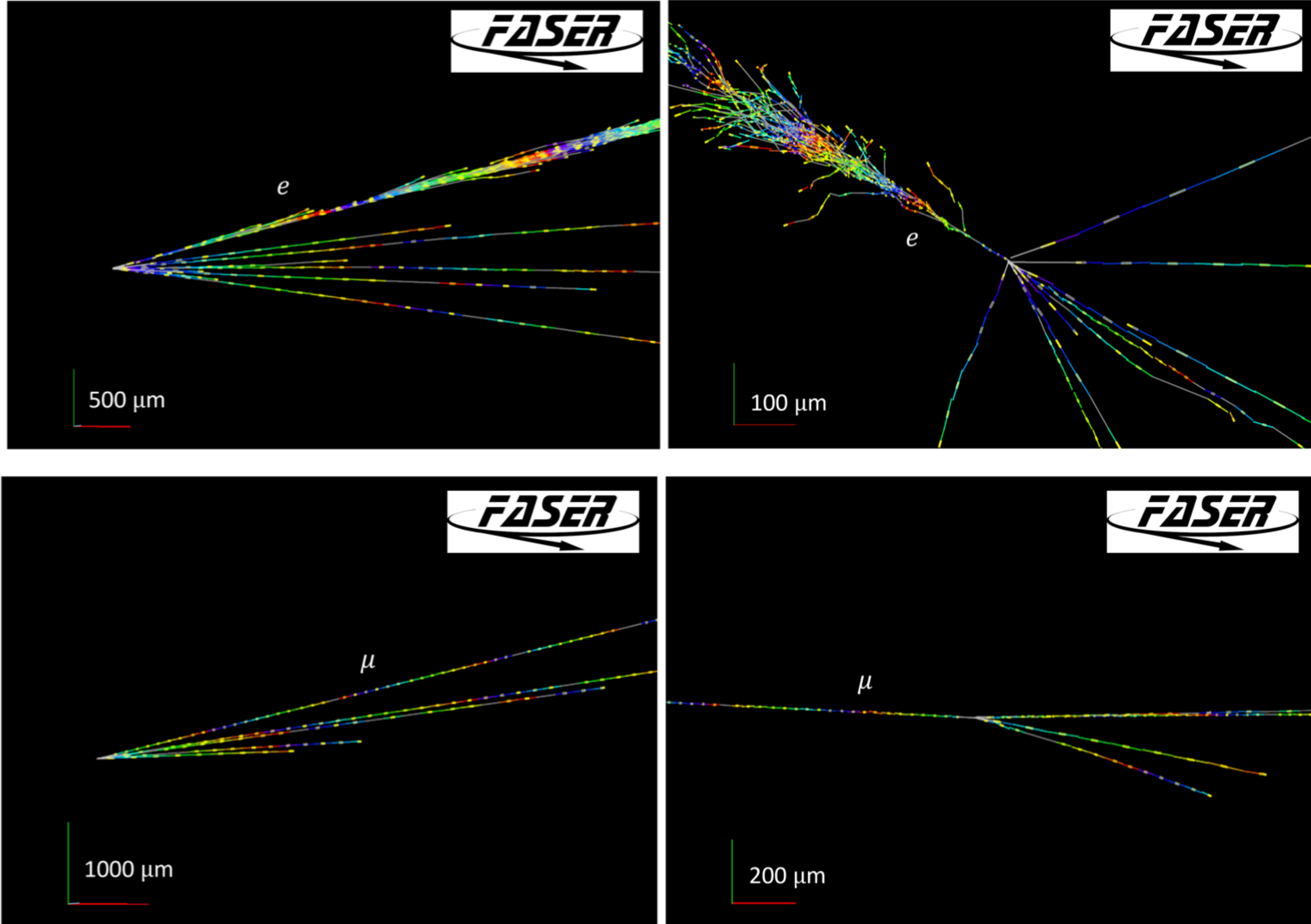
First Neutrino Cross-Section Measurements at LHC



4 ν_e and 8 ν_μ events with 9.5 fb^{-1} of data.

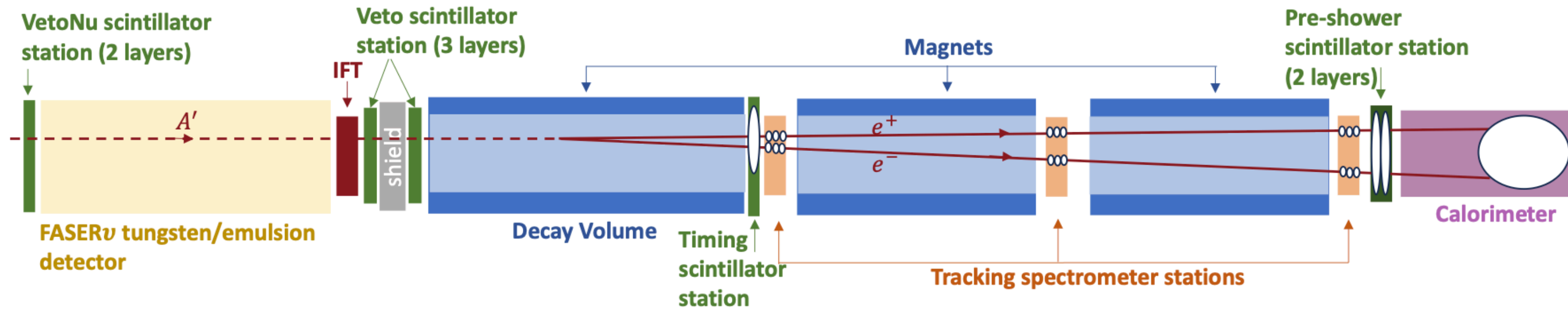
First Measurement of the ν_e and ν_μ Interaction Cross Sections at the LHC with FASER's Emulsion Detector; 2403.12520

ν_e and ν_μ events at FASER ν

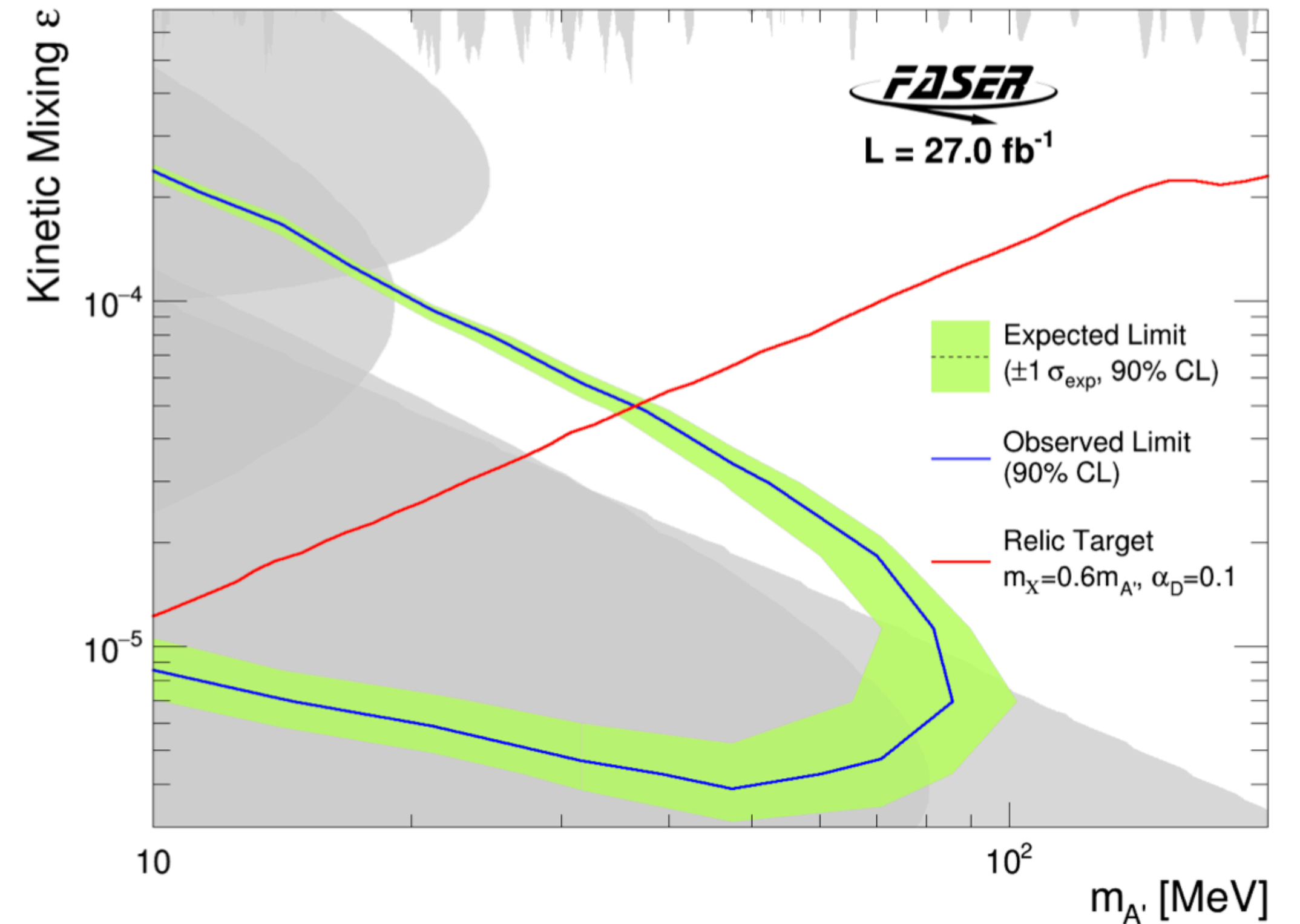
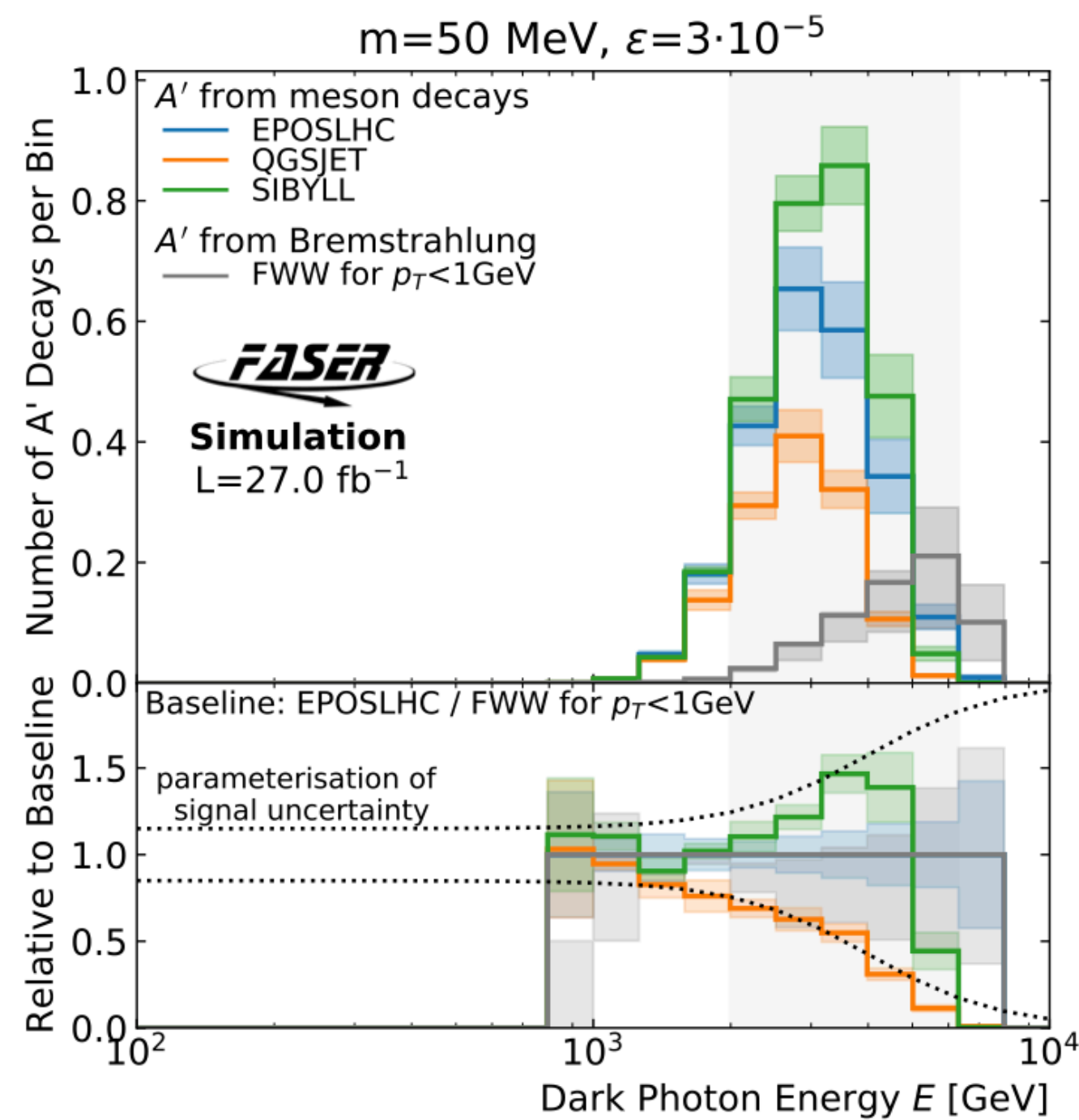


First Measurement of the ν_e and ν_μ Interaction Cross Sections at the LHC with FASER's Emulsion Detector; 2403.12520

Dark Photon Searches at FASER

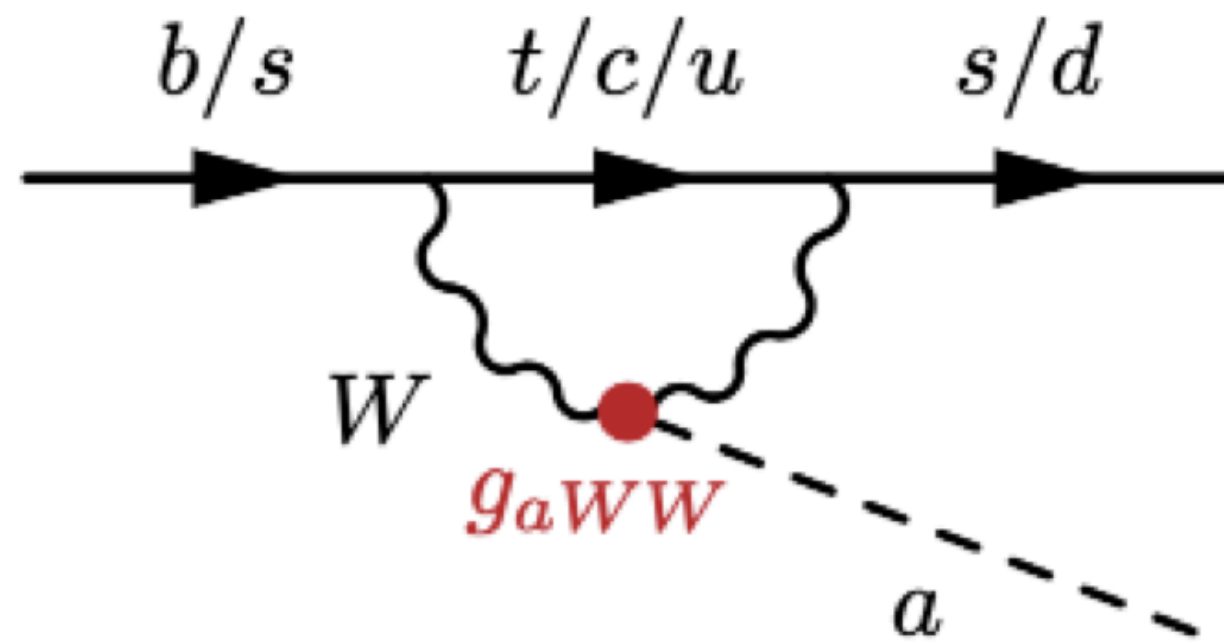


Search for Dark Photons with the FASER detector at the LHC;
2308.05587

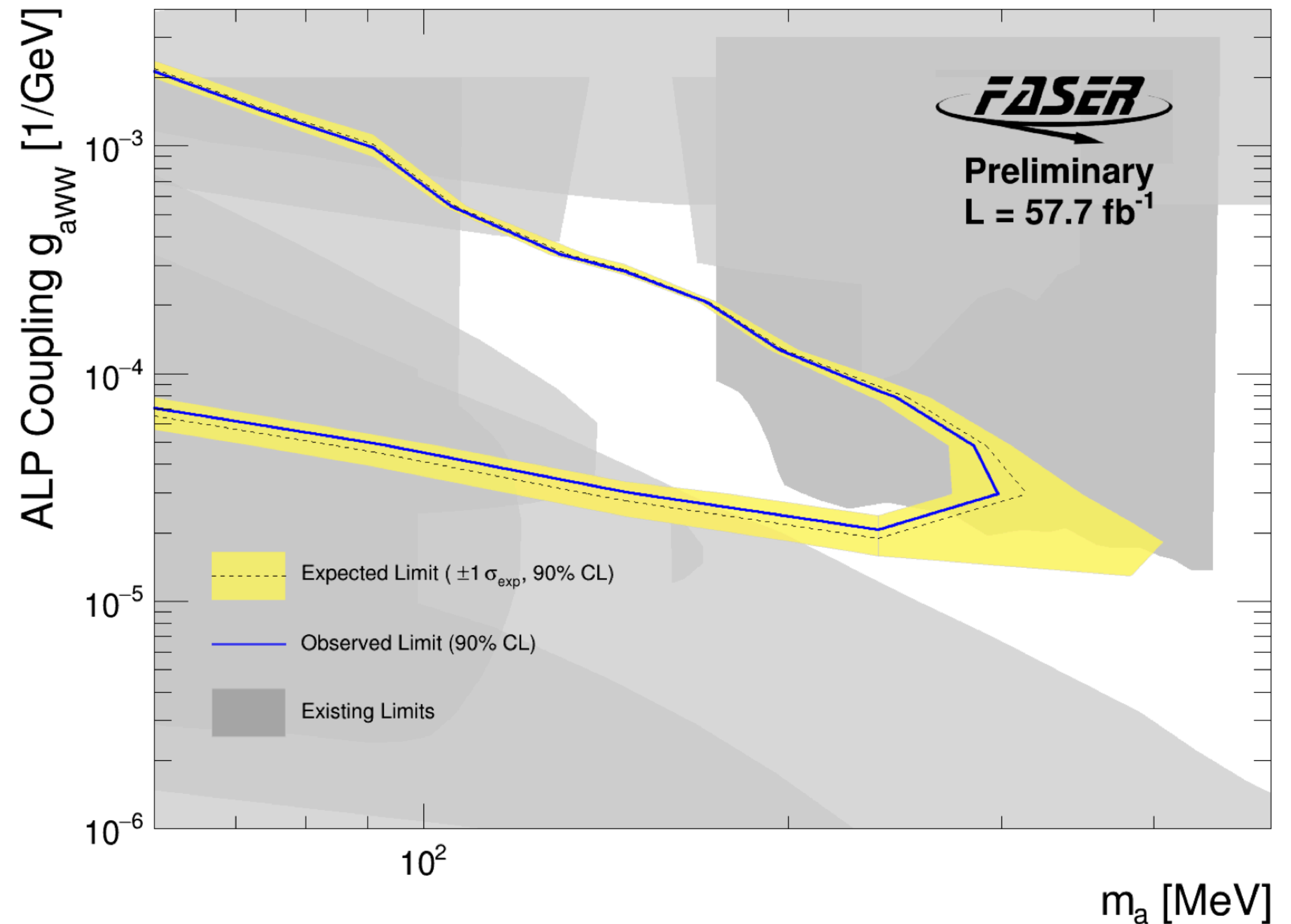


ALP Searches at FASER

$$\mathcal{L} \supset -\frac{1}{2}m_a^2 a^2 - \frac{1}{4}g_{aWW} a W^{a,\mu\nu} \tilde{W}_{\mu\nu}^a$$



Search for Axion-Like Particles in Photonic Final States with the FASER Detector at the LHC; [Conf note](#)



Proposed Expansion for HL-LHC: Forward Physics Facility

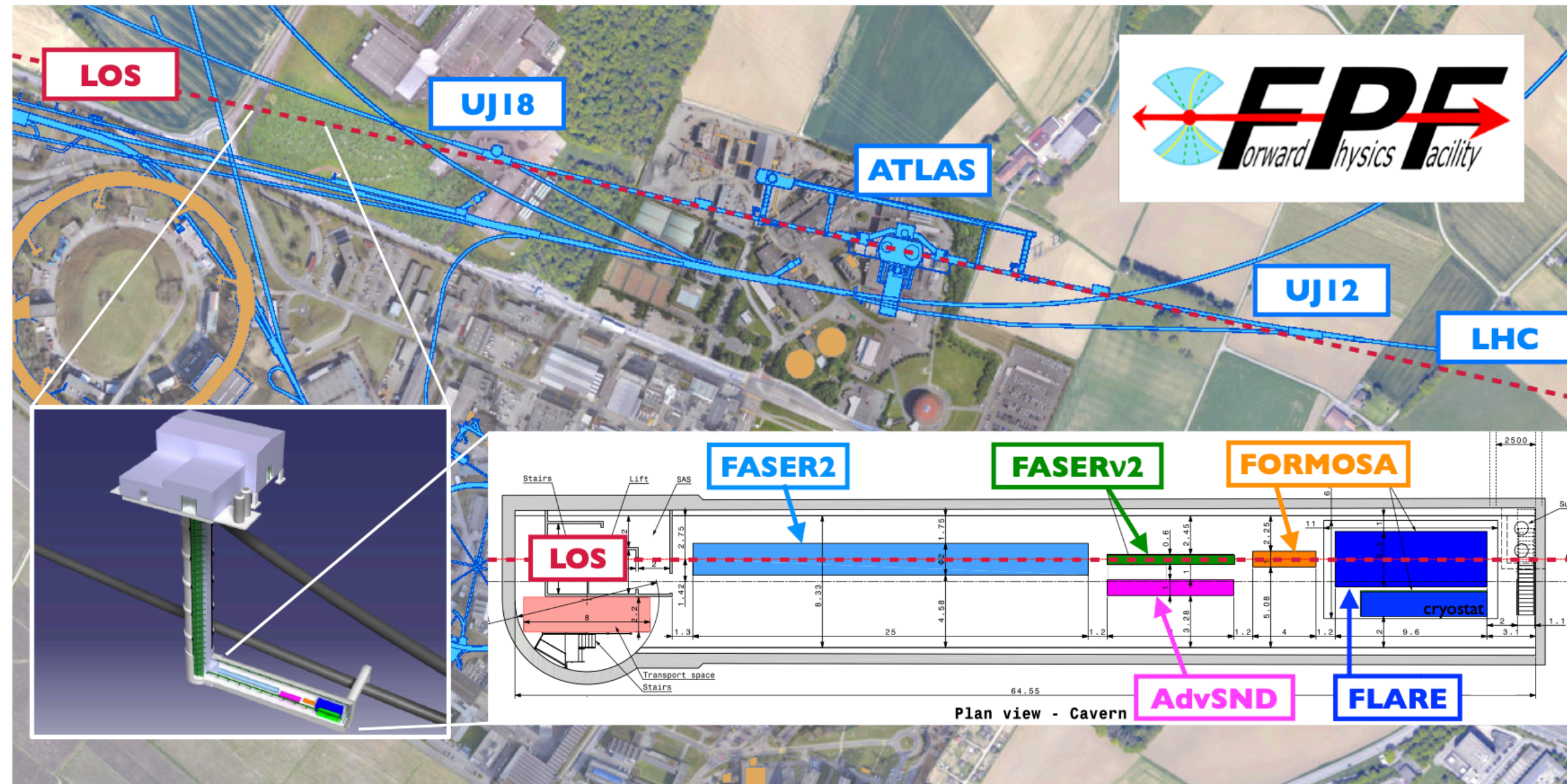


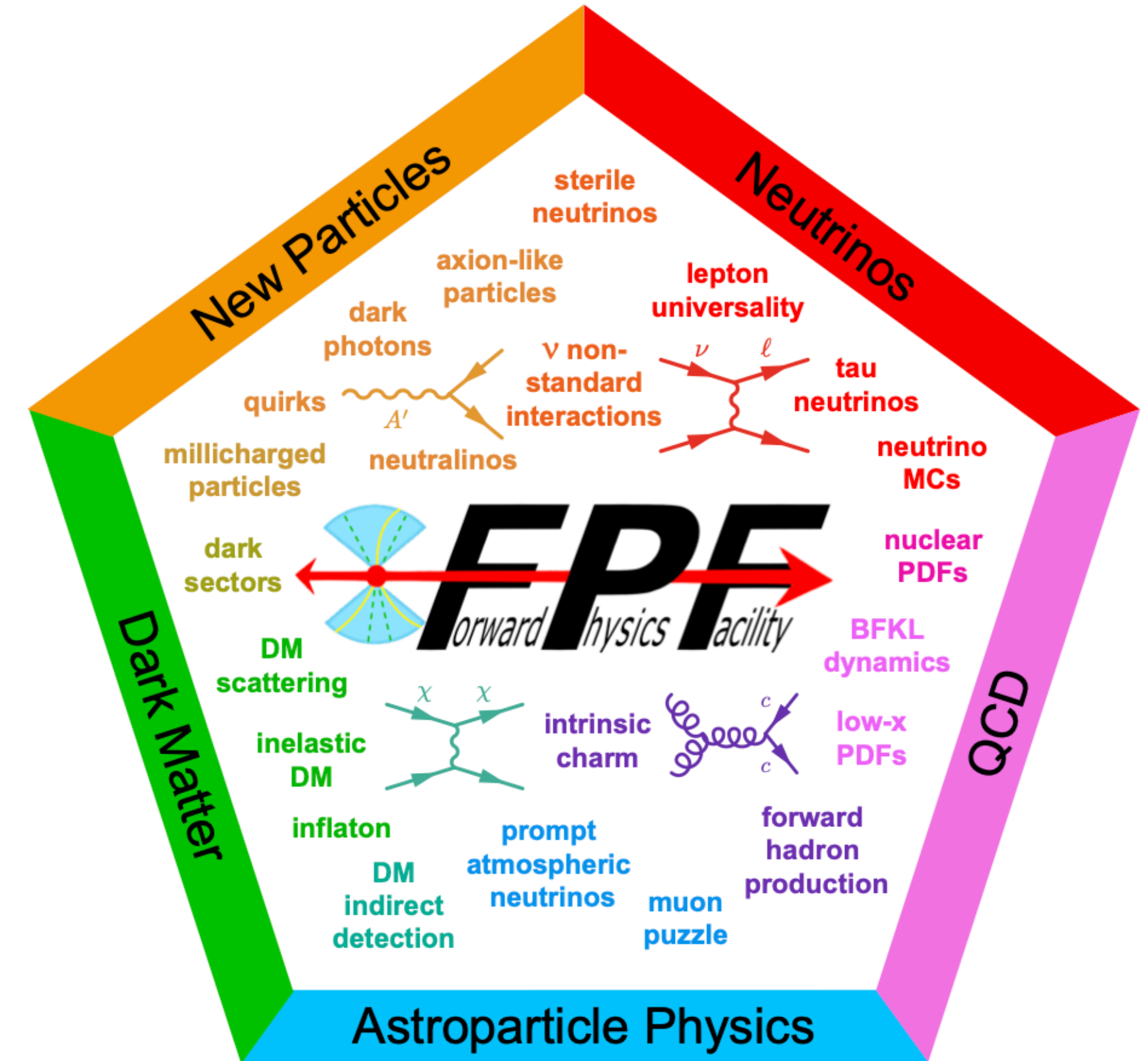
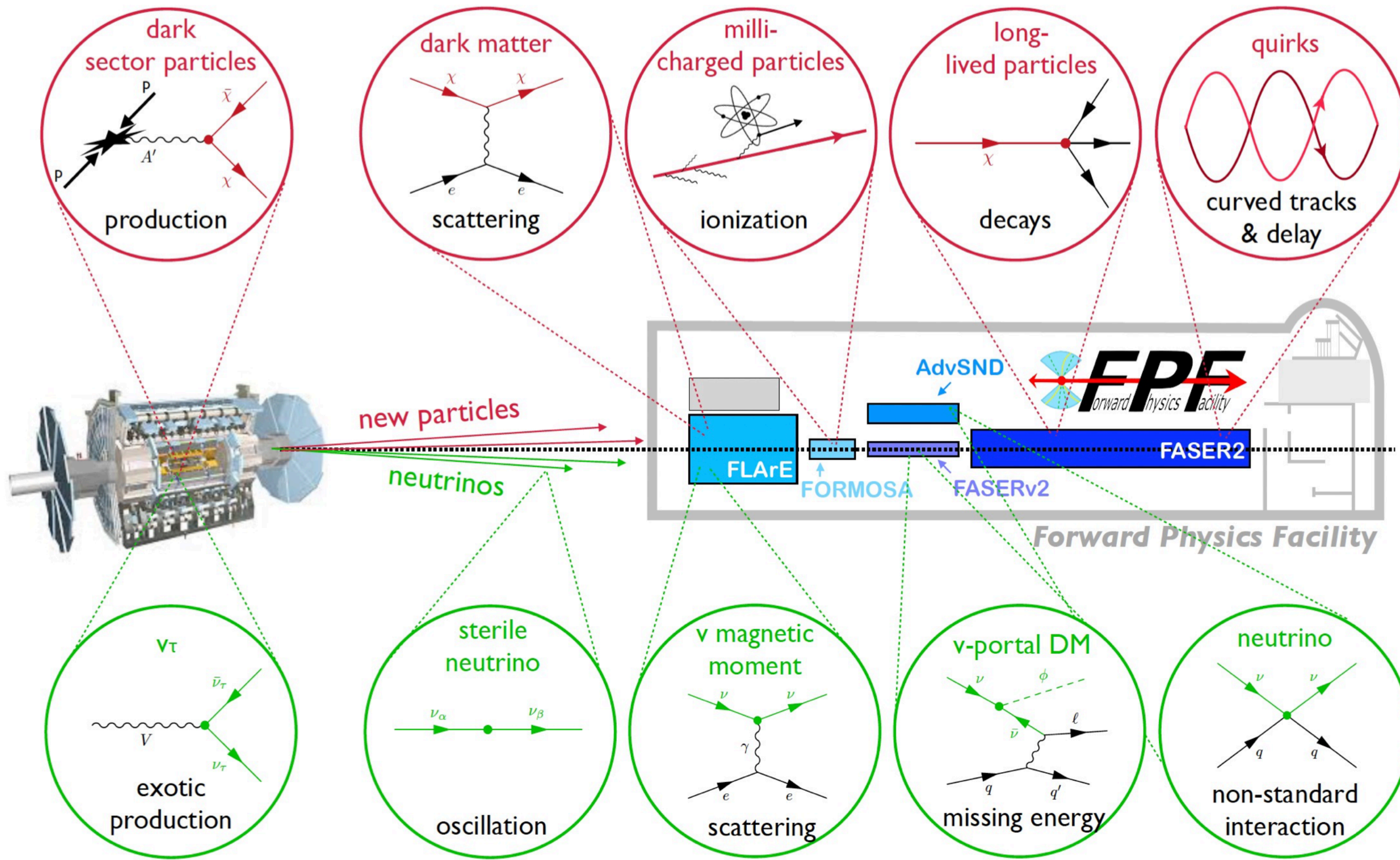
Figure 1: The preferred location for the Forward Physics Facility, a proposed new cavern for the High-Luminosity era. The FPF will be 65 m-long and 8.5 m-wide and will house a diverse set of experiments to explore the many physics opportunities in the far-forward region.

FPF is proposed to house 5 detectors in the forward direction to study SM and BSM physics.

The Forward Physics Facility: Sites, Experiments, and Physics Potential; 2109.10905

The Forward Physics Facility at the High-Luminosity LHC; 2203.05090

Many Physics opportunities at FPF



Muons at Forward Detectors

$$\nu_e: K \longrightarrow \pi e \nu_e, D \longrightarrow K e \nu_e$$

$$\nu_\mu: \pi^\pm \longrightarrow \mu \nu_\mu, K^\pm \longrightarrow \mu \nu_\mu$$

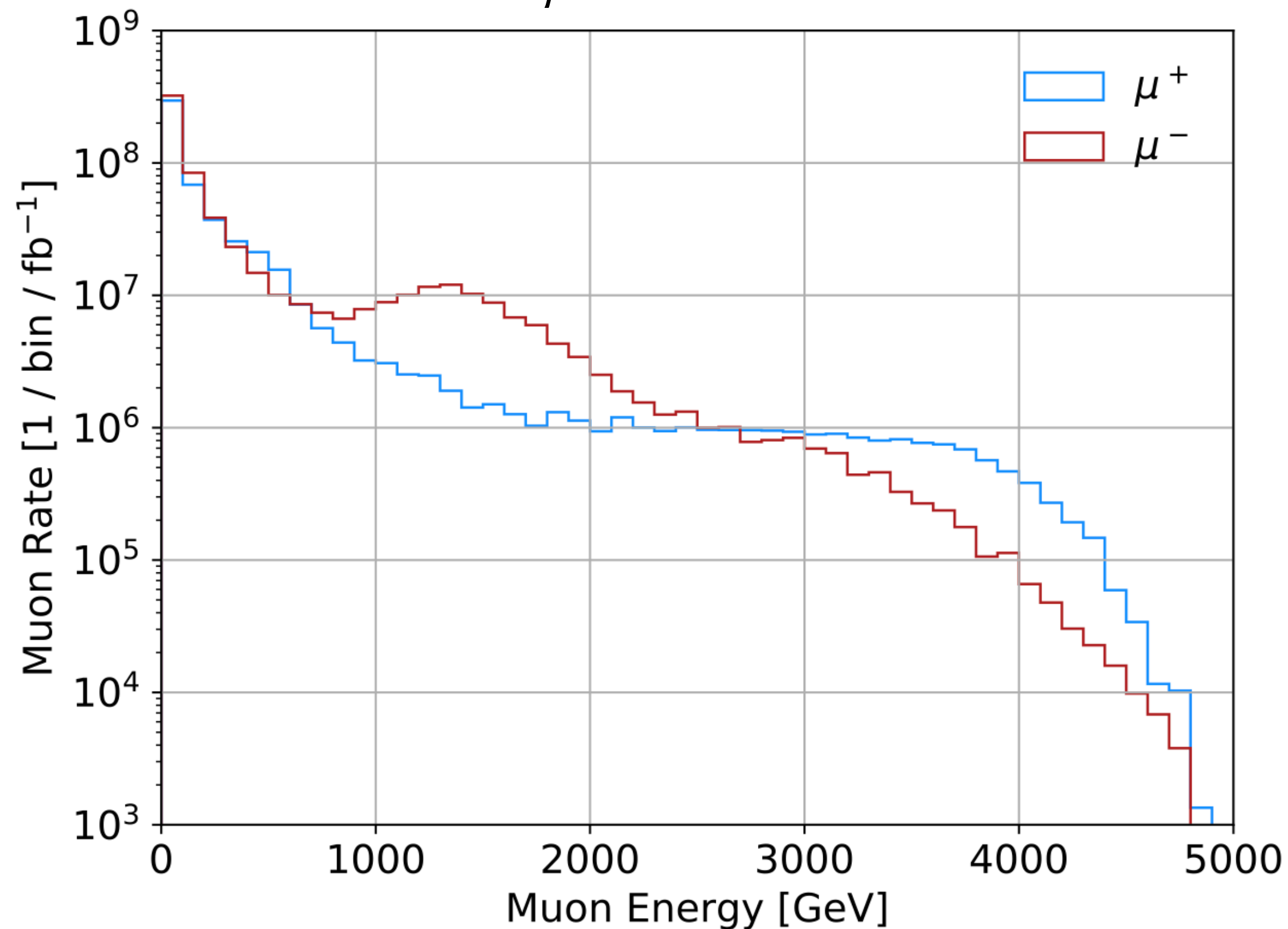
$$\nu_\tau: D_s \longrightarrow \tau \nu_\tau$$

But what about all these muons?

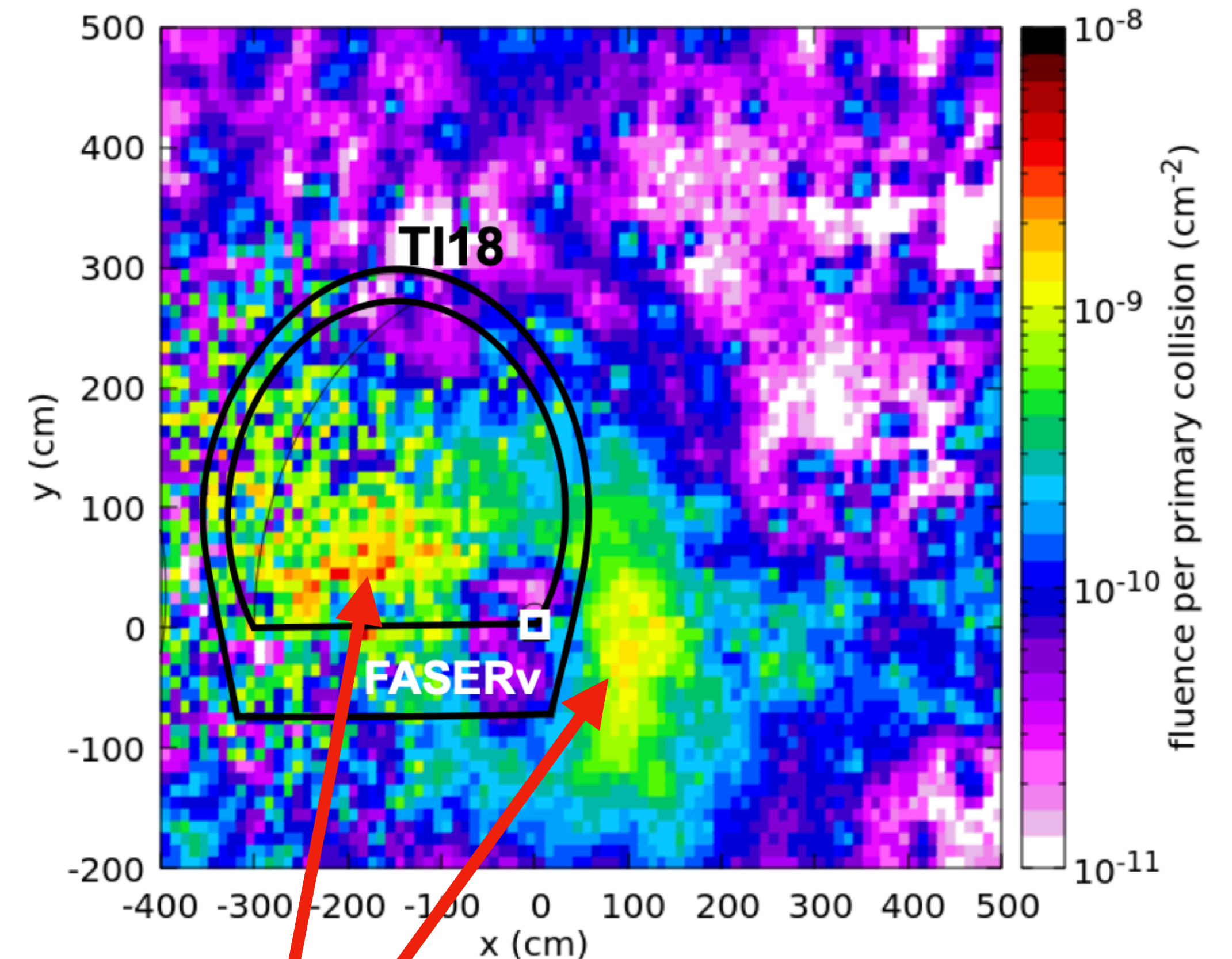
Are they just backgrounds or can we do some physics with them?

One Scientist's Background is Another's Signal

$N_{\mu} \sim 2 * 10^9$, through FASER during Run3!!!



First neutrino interaction candidates at the LHC; 2105.06197



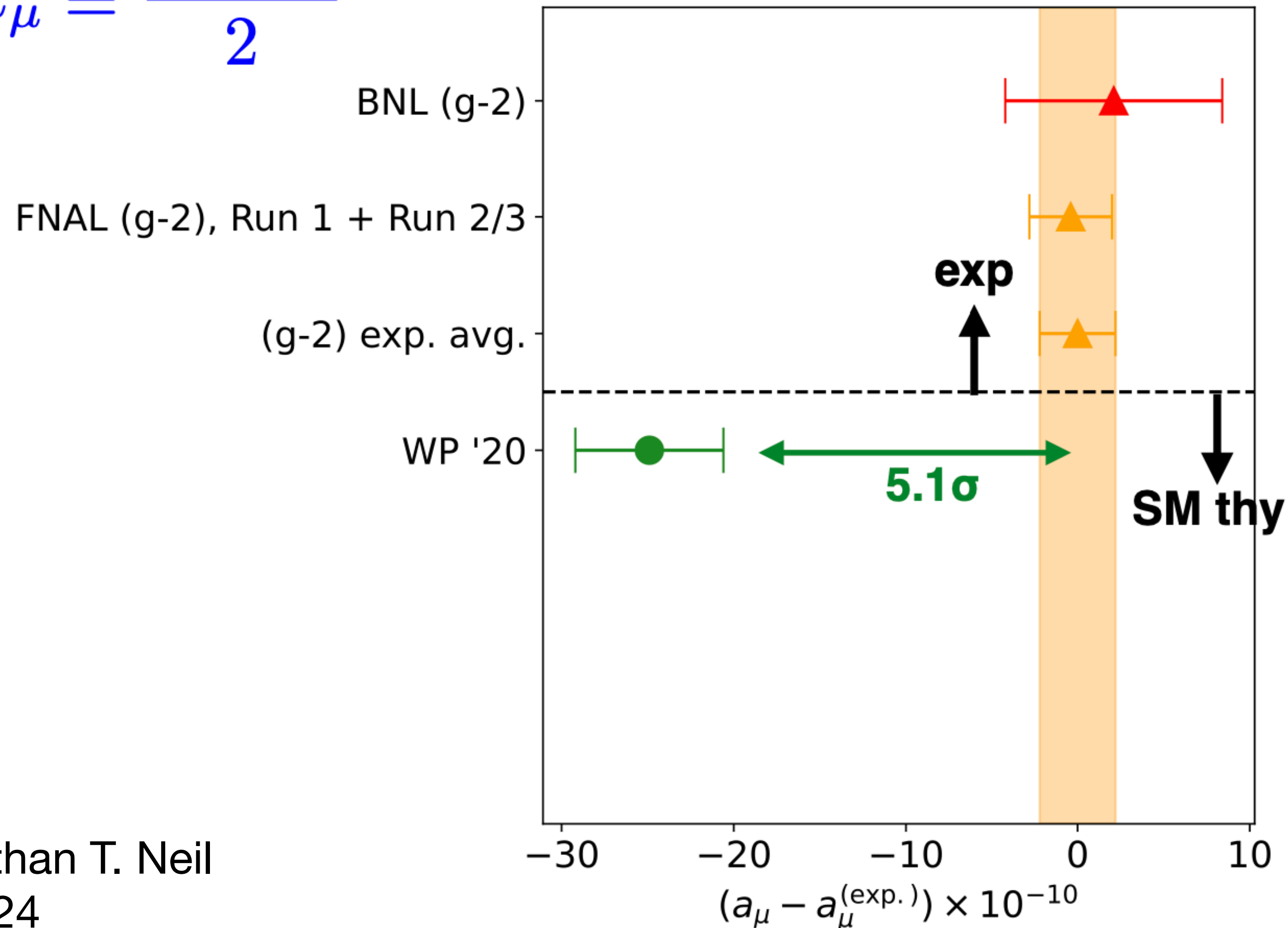
FASER ν T.P, 2001.03073

FASER intentionally avoids the maximum flux of muons

Muon (g-2) puzzle

$$a_\mu \equiv \frac{g_\mu - 2}{2}$$

(Latest experiment: FNAL Muon (g-2) Collaboration, PRL **131**, 161802 (2023); arXiv:2308.06230)
(WP '20: T. Aoyama et al (Muon (g-2) Theory Initiative), arXiv:2006.04822)



- Latest FNAL (g-2) results in $\sim 5\sigma$ tension with “SM theory” prediction from Theory Initiative whitepaper!

Simple model with a muonphilic scalar

- A SM singlet scalar, S , that couples only to the muons.

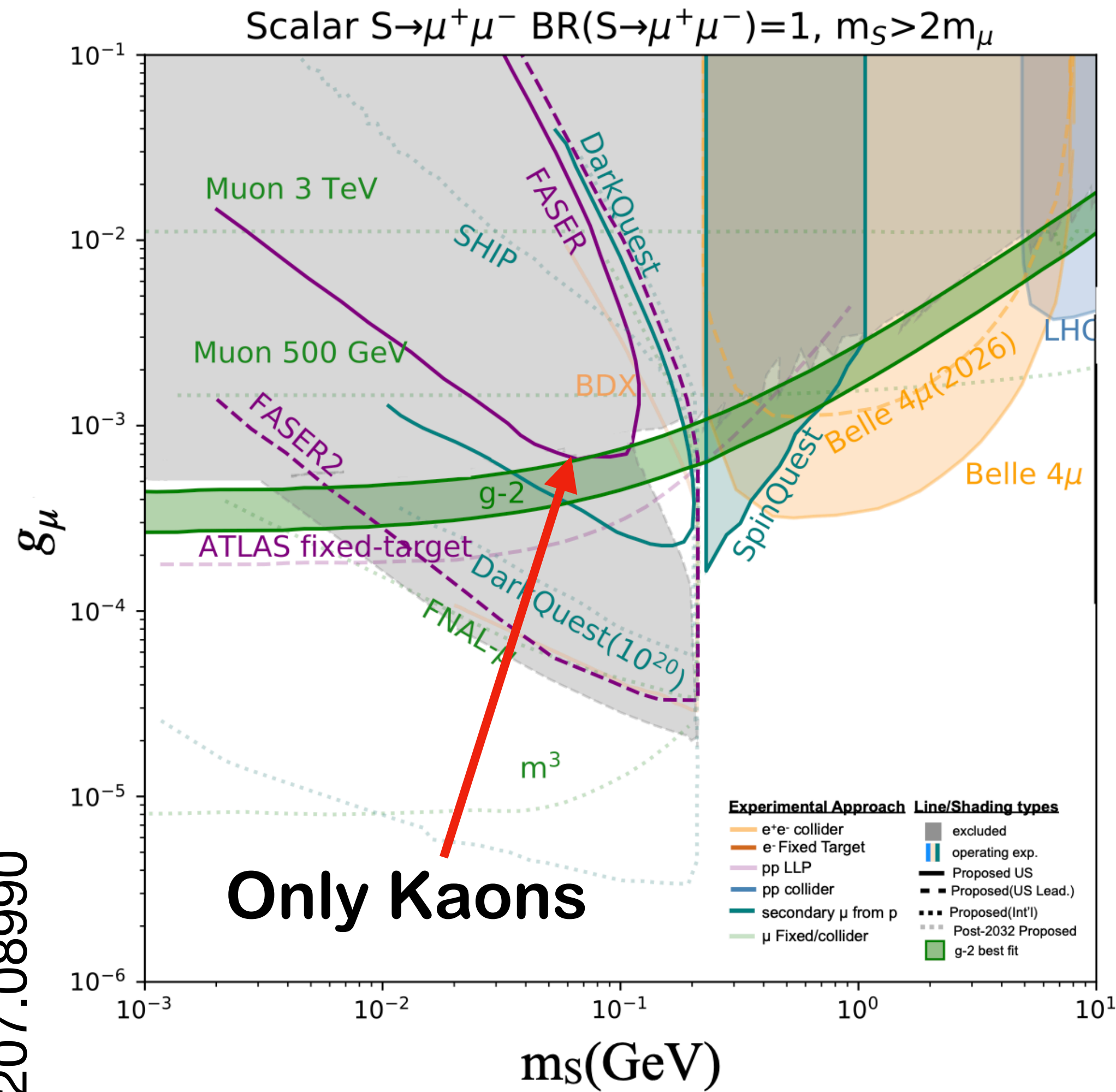
- $\mathcal{L} \supset \frac{1}{2} (\partial_\nu S)^2 - \frac{1}{2} m_S^2 S^2 - g_S S \bar{\mu} \mu$

- Contribution to $\Delta a_\mu = (g - 2)_\mu / 2$ is given by

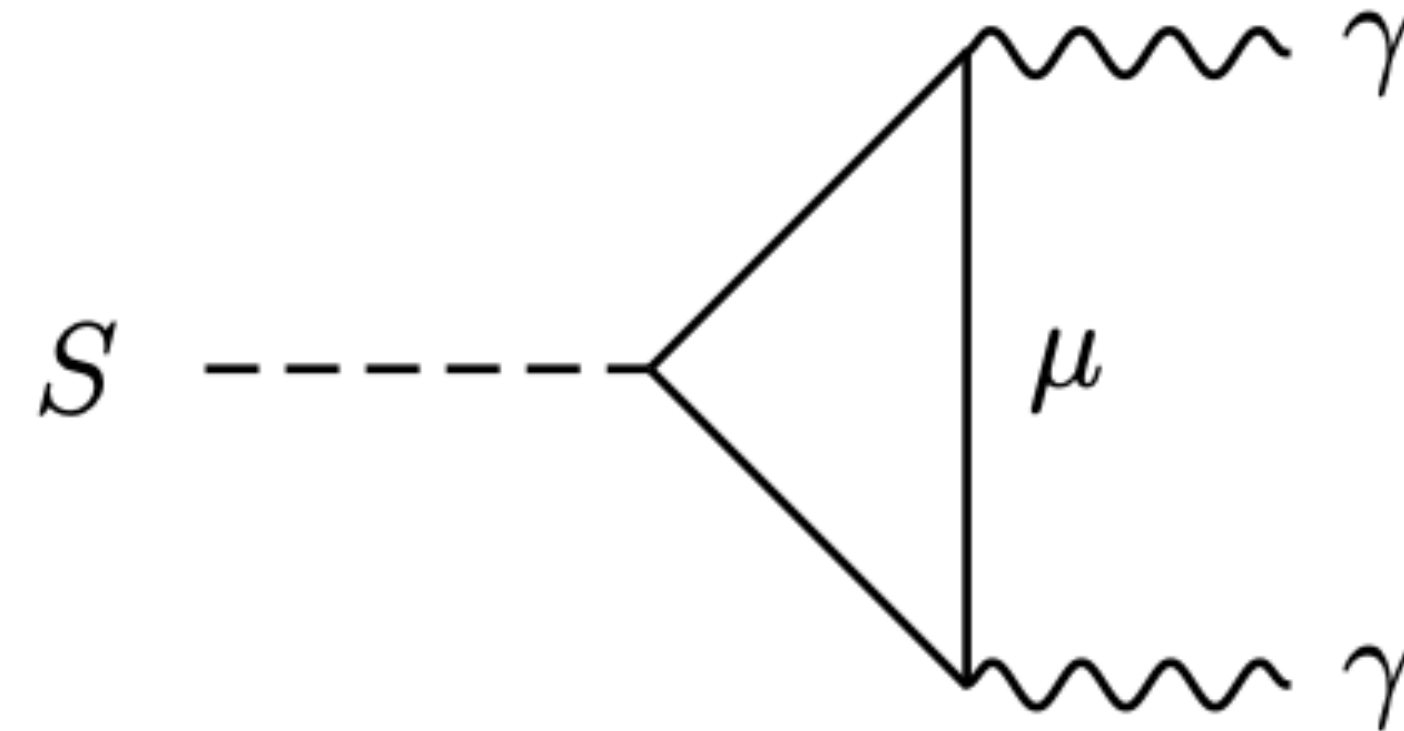
$$\Delta a_\mu = \frac{g_\mu^2}{8\pi^2} \int_0^1 dz \frac{(1-z)^2(1+z)}{(1-z)^2 + z(m_S/m_\mu)^2}$$

Simple model with a muonphilic scalar (Cont.)

Snowmass White Paper: New flavors and rich structures in dark sectors
 Philip Harris, Philip Schuster, Jure Zupan;
 2207.08990



- Current constraints miss a small region near $m_S \lesssim 2 * m_\mu$.
- For $m_S \lesssim 2 * m_\mu$, S decays to 2 photons.

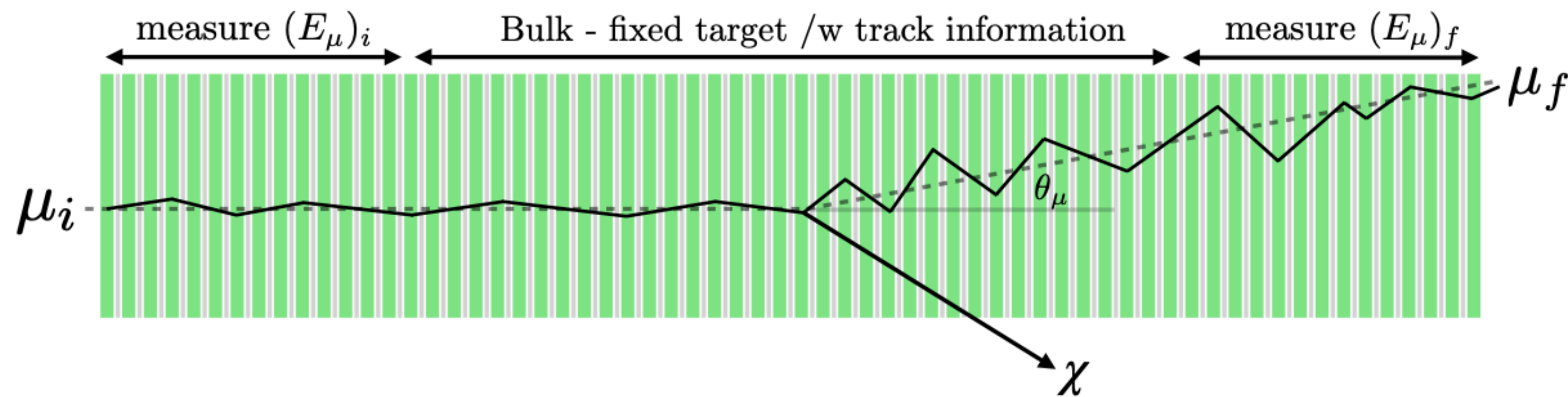


See also,
 Brian Batell, Ayres Freitas, Ahmed Ismail, David McKeen;
 1712.10022

Using *FASER ν* emulsion detector

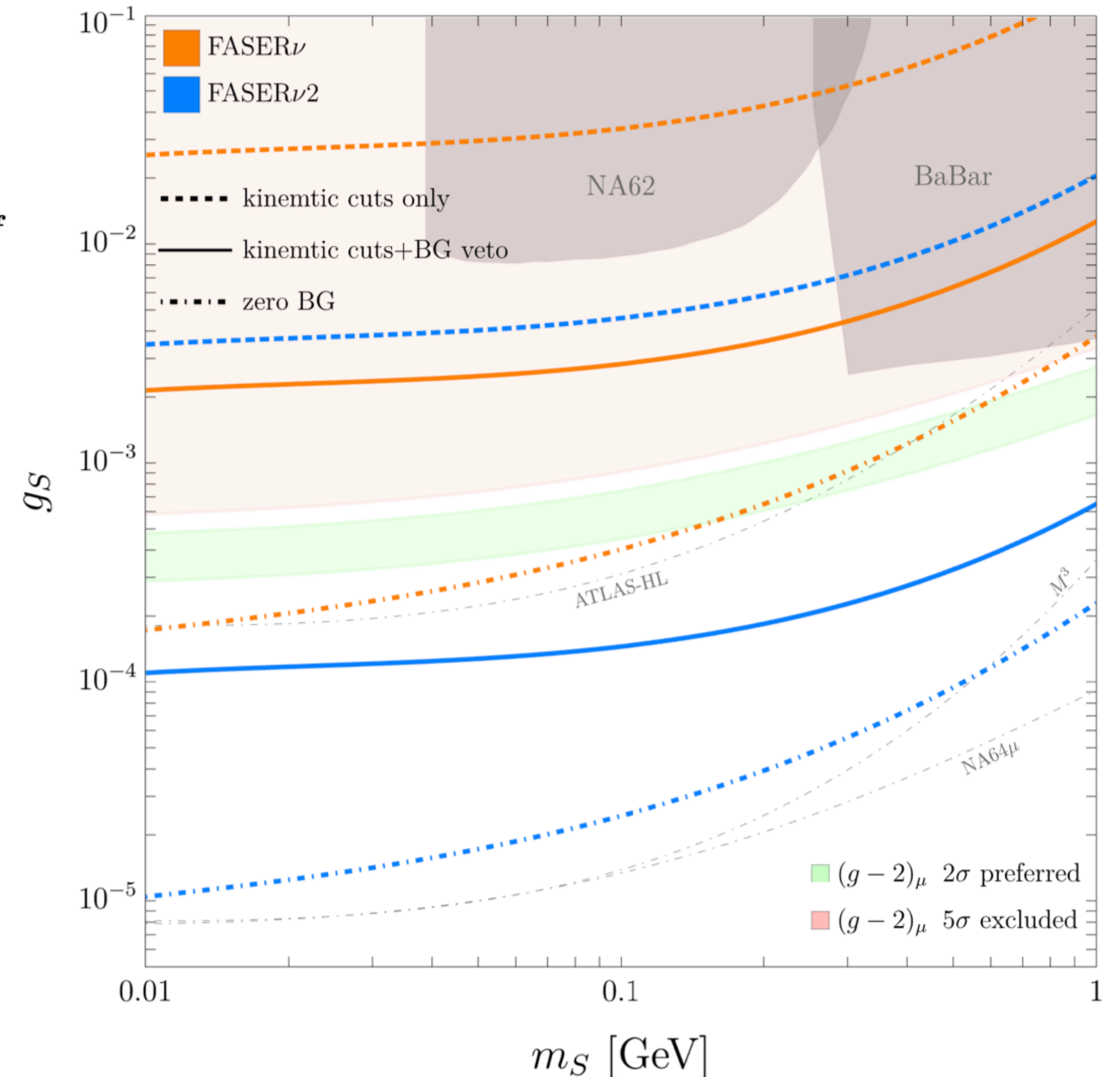
Akitaka Ariga, Reuven Balkin, Iftah Galon, Enrique Kajomovitz, Yotam Soreq
2305.03102

Using missing momentum technique

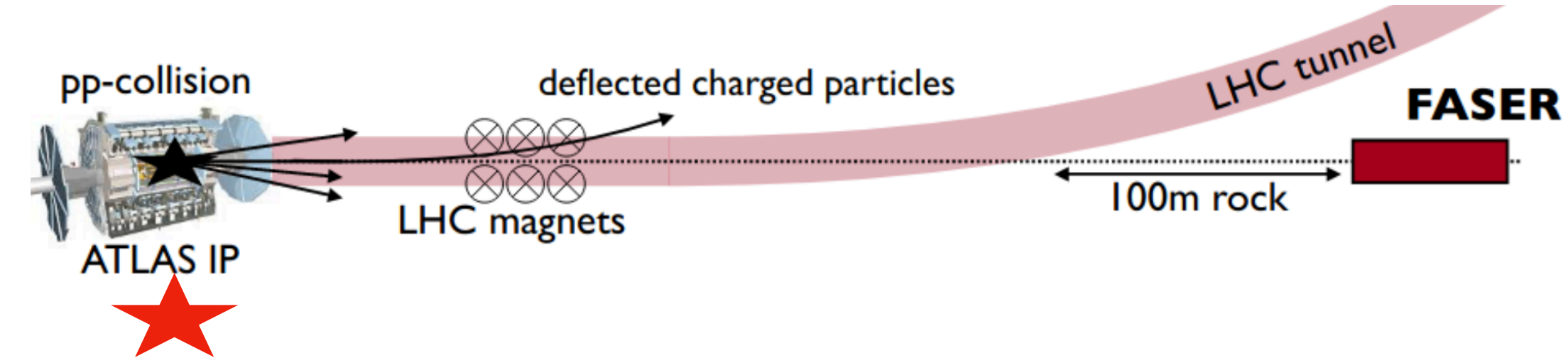


Backgrounds:

$$N + \mu \rightarrow \begin{cases} N + \mu + e^- + e^+ & \text{(Pair production)} \\ N^+ + \mu + e^- & \text{(Ionization)} \\ N + \mu + \gamma & \text{(Bremsstrahlung)} \\ N^* + \mu + \dots & \text{(Nuclear)} \end{cases}$$

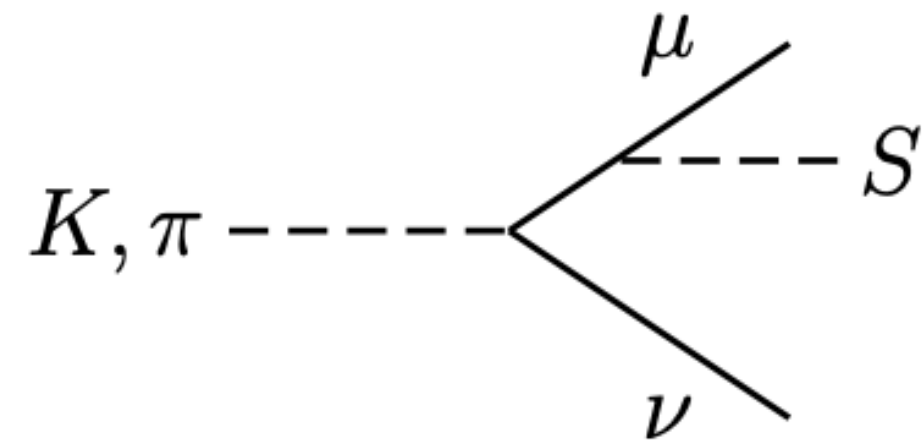


Production from 3 body decays near ATLAS IP



- **Scalar decays via W** Carl E. Carlson, Benjamin C. Rislow; 1206.3587

$$\frac{dBR(K \rightarrow \mu\nu S)}{dE_S dQ^2} = \frac{m_K y^2 \times BR(K \rightarrow \mu\nu)}{8\pi^2 m_\mu^2 (m_K^2 - m_\mu^2)^2 (Q^2 - m_\mu^2)^2} \times \left((m_K^2 - 2m_K E_S + Q^2) Q^2 (Q^2 - m_\mu^2) - (Q^4 - m_\mu^2 m_K^2) (Q^2 + m_\mu^2 - m_S^2) + 2m_\mu^2 Q^2 (m_K^2 - Q^2) \right)$$

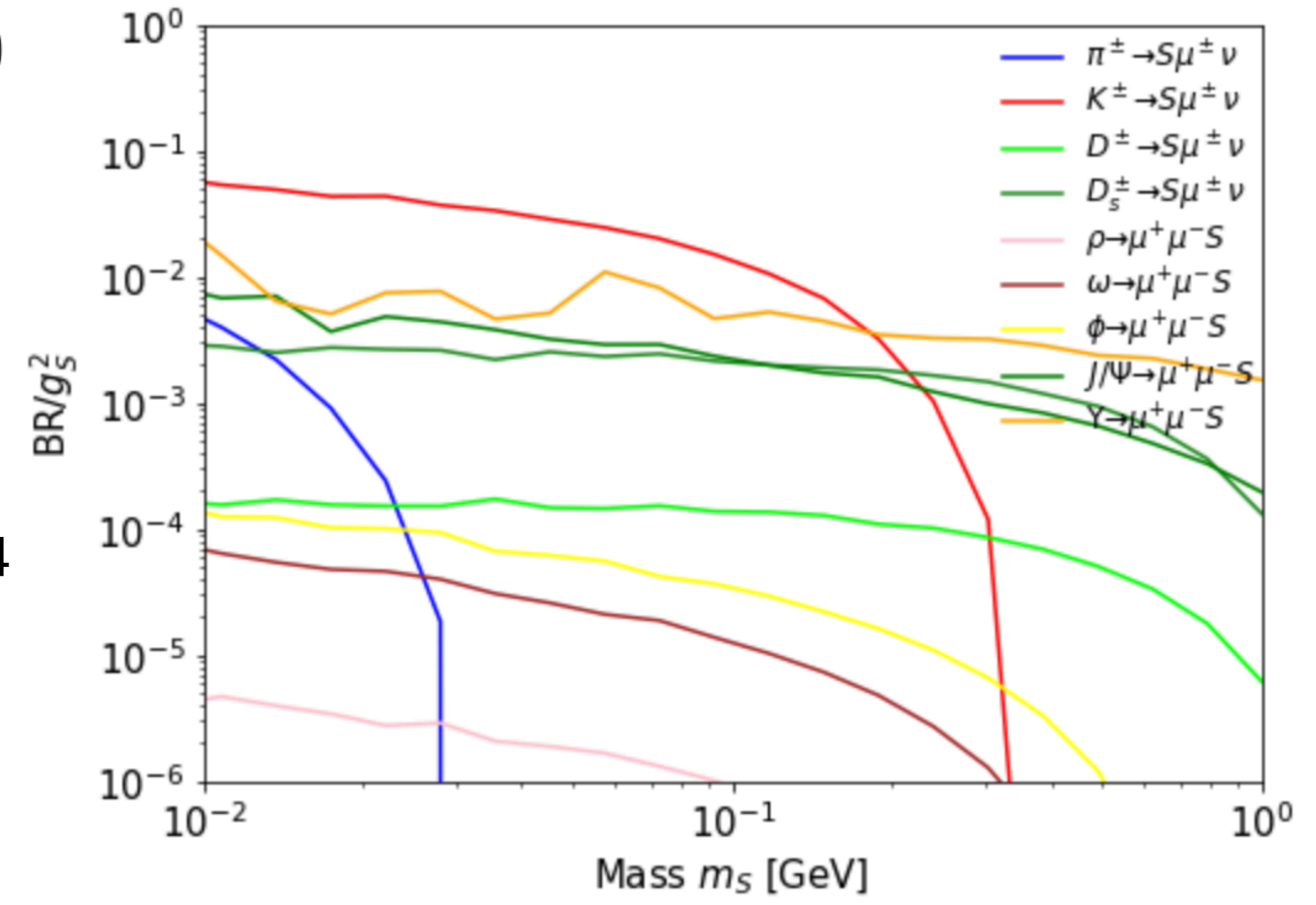


- **Vector decays via γ** Manimala Mitra, Dibyakrupa Sahoo; 2103.08284

$$\frac{d^2\Gamma_{s^\pm}}{dt du} \equiv \frac{d^2\Gamma(J/\psi \rightarrow \mu^- \mu^+ X_{s^\pm})}{dt du} = \frac{\alpha^2 g_{s^\pm}^2 f_J^2}{27 \pi m_J^5 Y} |A_{s^\pm}|^2$$

$$Y = (t - m_\mu^2)^2 (u - m_\mu^2)^2$$

Decay constant **Squared amplitude**

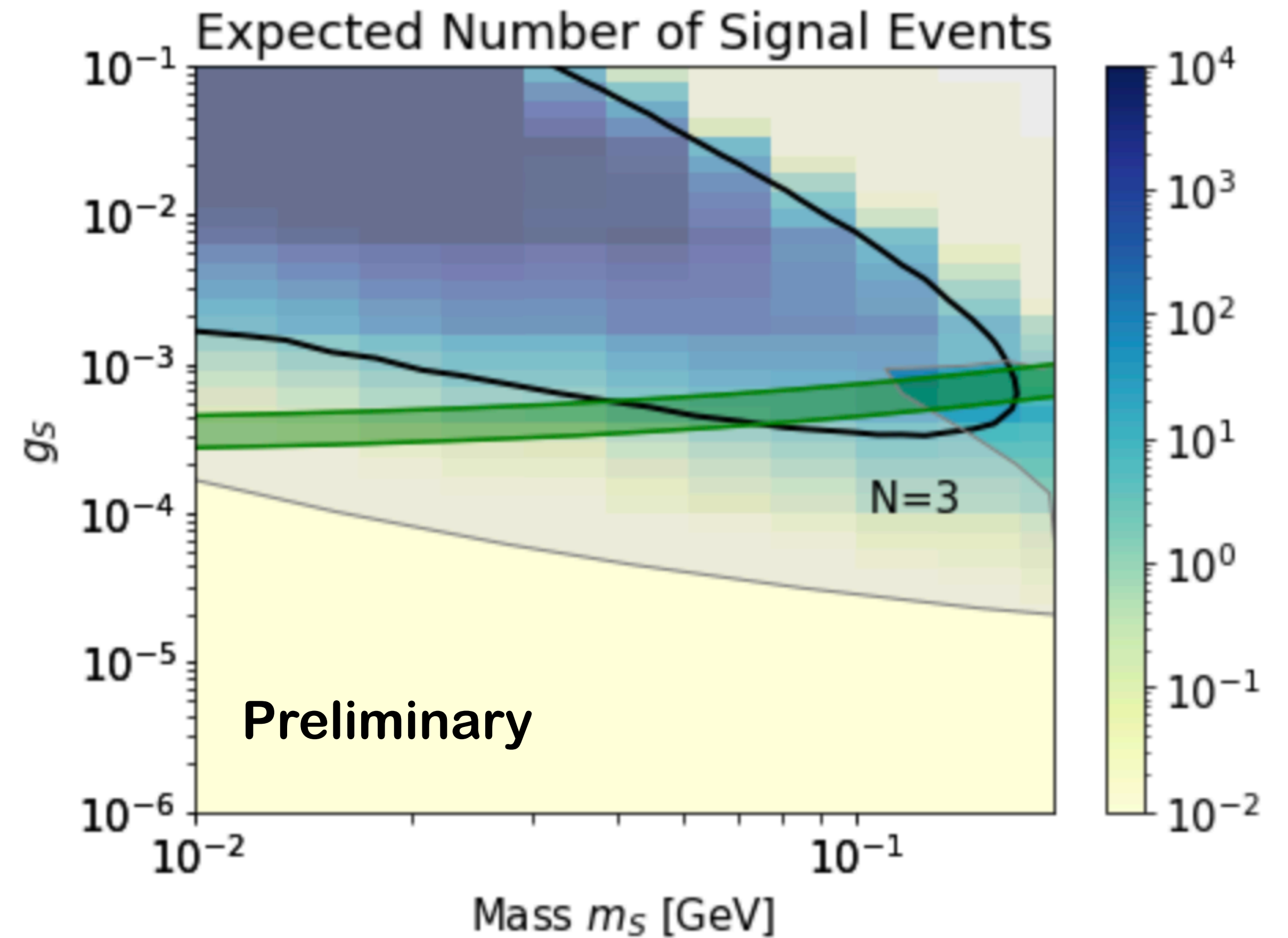


Implemented via FORESEE
 Felix Kling, Sebastian Trojanowski,
 2105.07077

Production from 3 body decays near ATLAS IP

Significant event rates expected at FASER during Run 3.

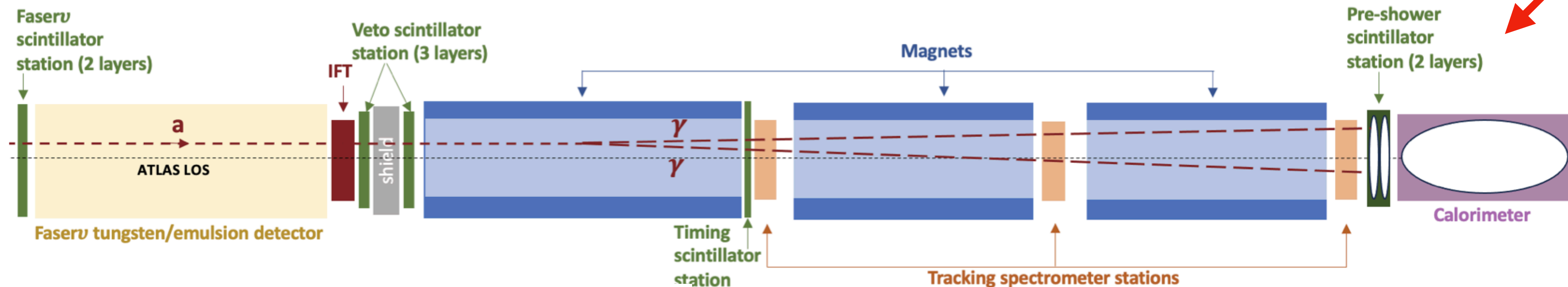
But what about backgrounds?



3 body decays (cont.)

- The signal we expect is “no activity” with some energy deposition in the calorimeter.
- Very similar to FASER’s ALPs analysis. (Conf note)
- The dominant background there is from neutrino interactions.

ALP search at FASER

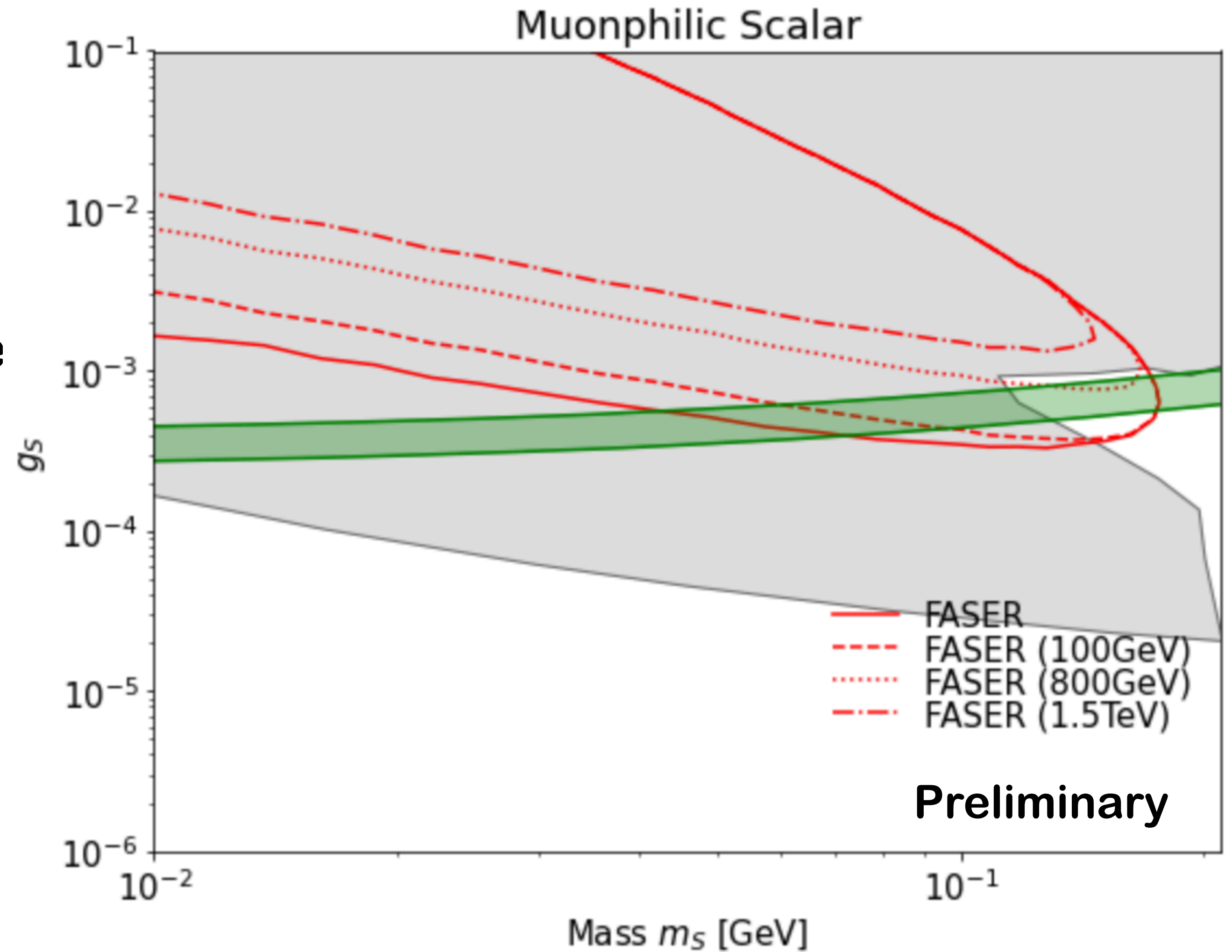


> 1.5 TeV signal region	
Light	$0.23^{+0.01}_{-0.11}$ (flux) ± 0.11 (exp.) ± 0.04 (stat.)
Charm	$0.19^{+0.32}_{-0.09}$ (flux) ± 0.06 (exp.) ± 0.03 (stat.)
Total	0.42 ± 0.38 (90.6%)

- $E_{calo} > 1.5$ TeV reduces the neutrino backgrounds to $\sim 0.42/50 fb^{-1}$.

3 body decays (cont.)

Applying the same energy cut, we do not probe the (g-2) band at FASER.

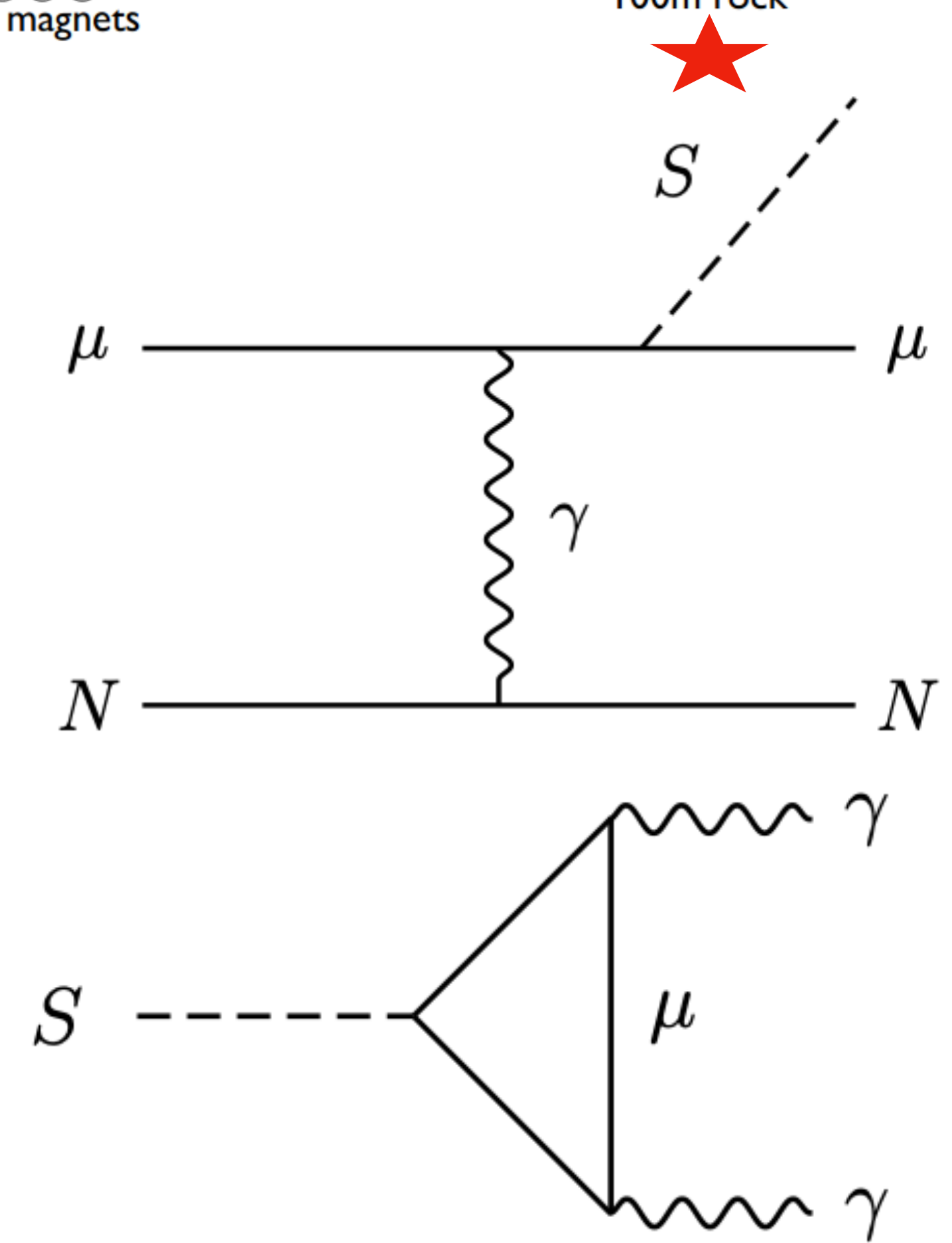
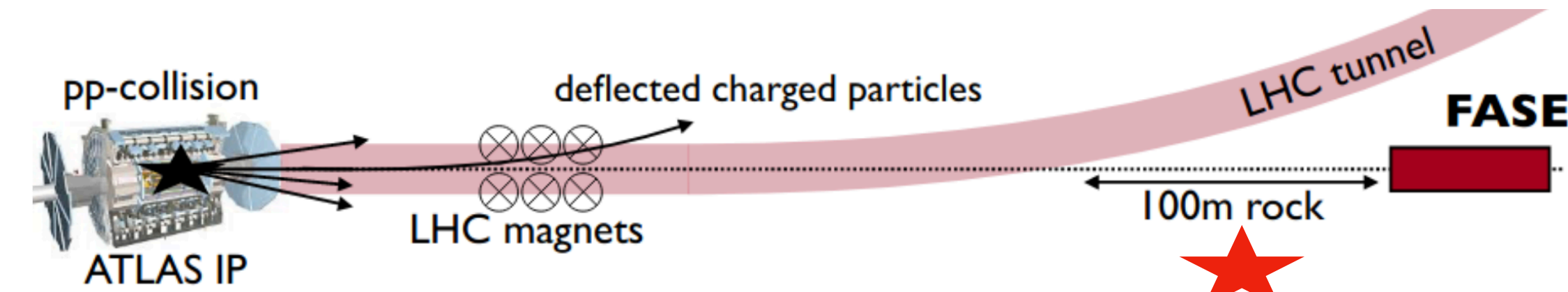


Production from Bremsstrahlung in Rock

- Incoming muon brems off S within the rock in front of $FASER\nu$ detector.
- Muon undergoes multiple scattering in the rock, away from the $FASER$ detector.
- Decay length is given by

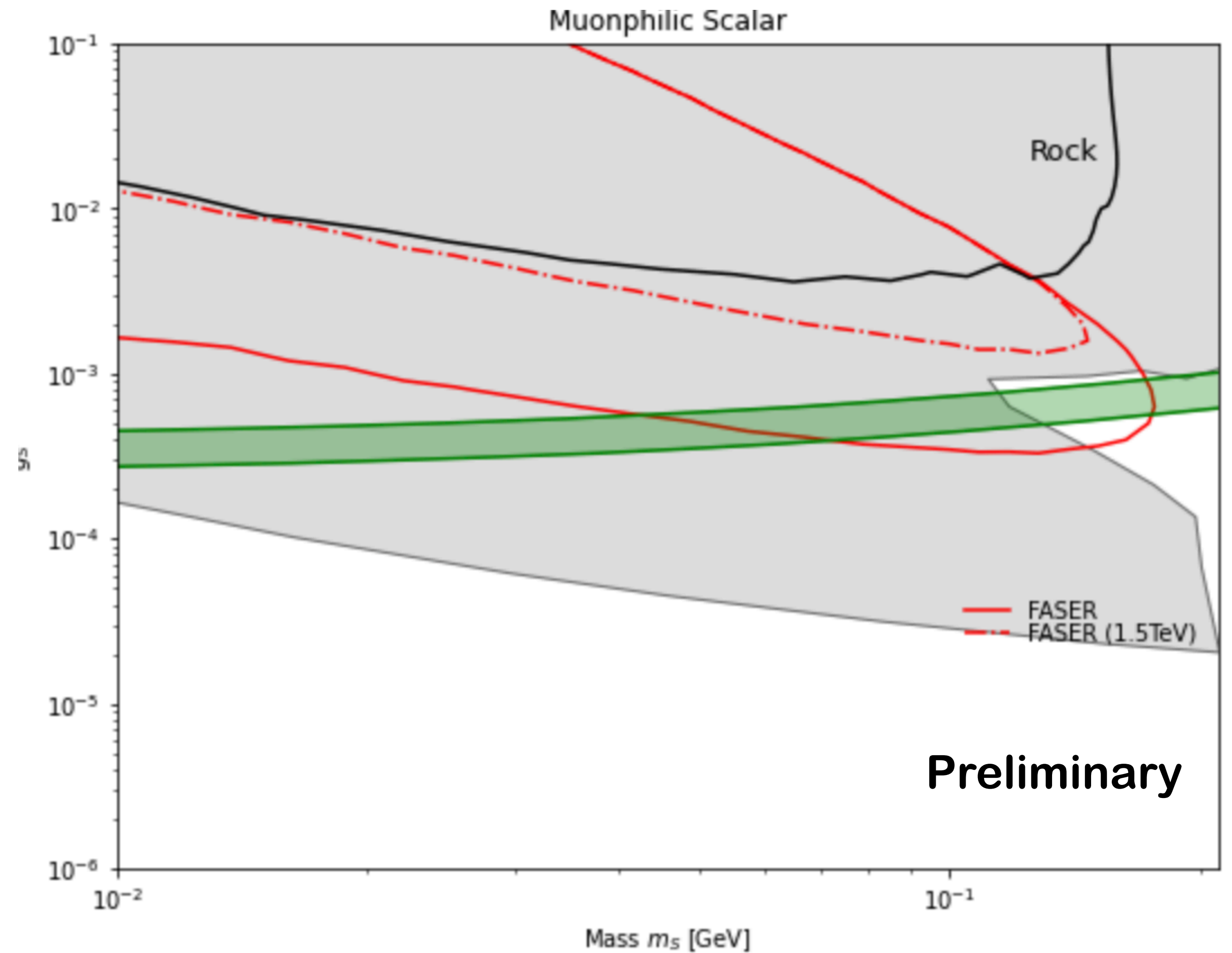
$$L_S = 200 \text{ m} \times \left(\frac{E_s}{30 \text{ GeV}} \right) \times \left(\frac{5 \times 10^{-4}}{g_S} \right)^2 \times \left(\frac{100 \text{ MeV}}{m_S} \right)^4$$

Chien-Yi Chen, Maxim
Pospelov, Yi-Ming
Zhong; 1701.07437



Production from Bremsstrahlung in Rock (cont.)

- Even without imposing an energy cut, bremsstrahlung production in rock is not sensitive to the (g-2) band.

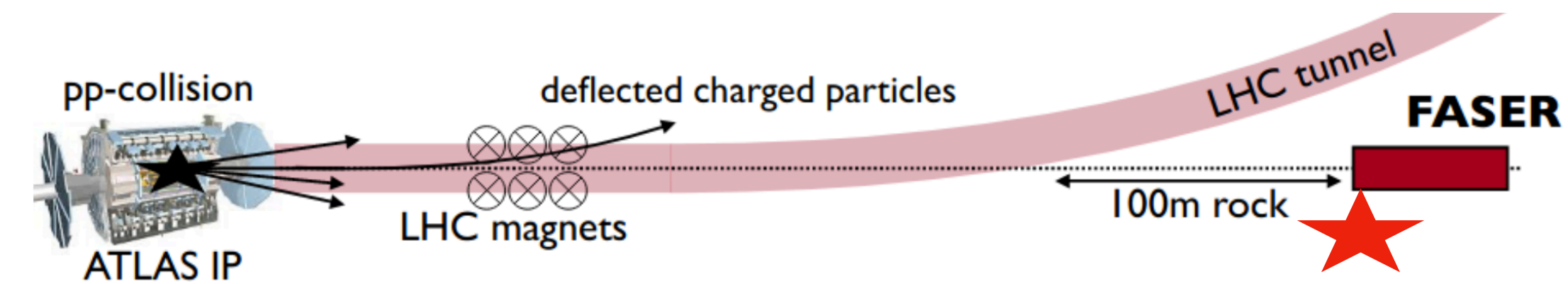


Production from Bremsstrahlung in $FASER\nu$

- Incoming muon brems off S within the $FASER\nu$ detector.
- The target is Tungsten.
- Decay length is given by

$$L_S = 20 \text{ m} \times \left(\frac{E_S}{3 \text{ GeV}} \right) \times \left(\frac{5 \times 10^{-4}}{g_S} \right)^2 \times \left(\frac{100 \text{ MeV}}{m_S} \right)^4$$

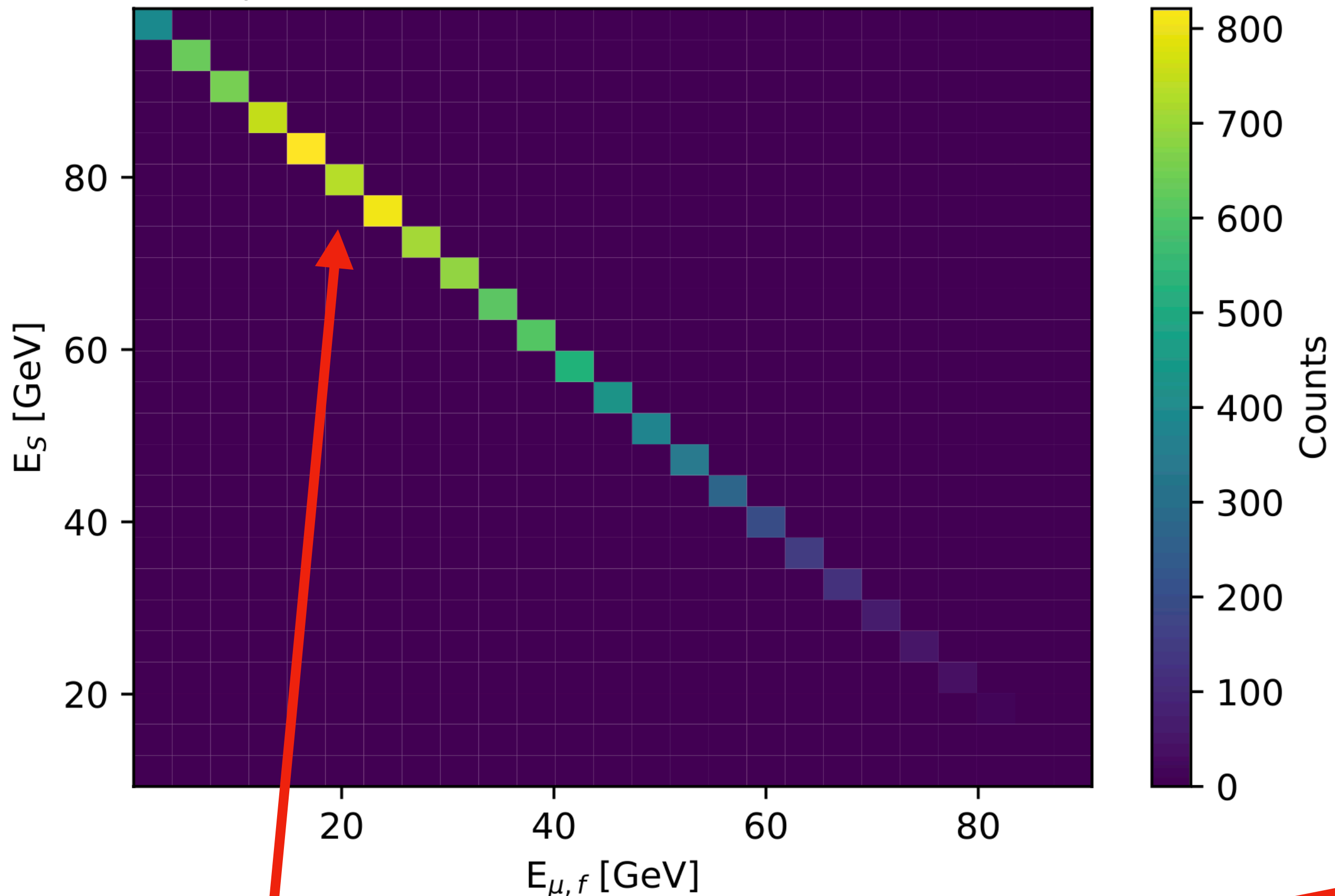
Signal mainly from low energy S



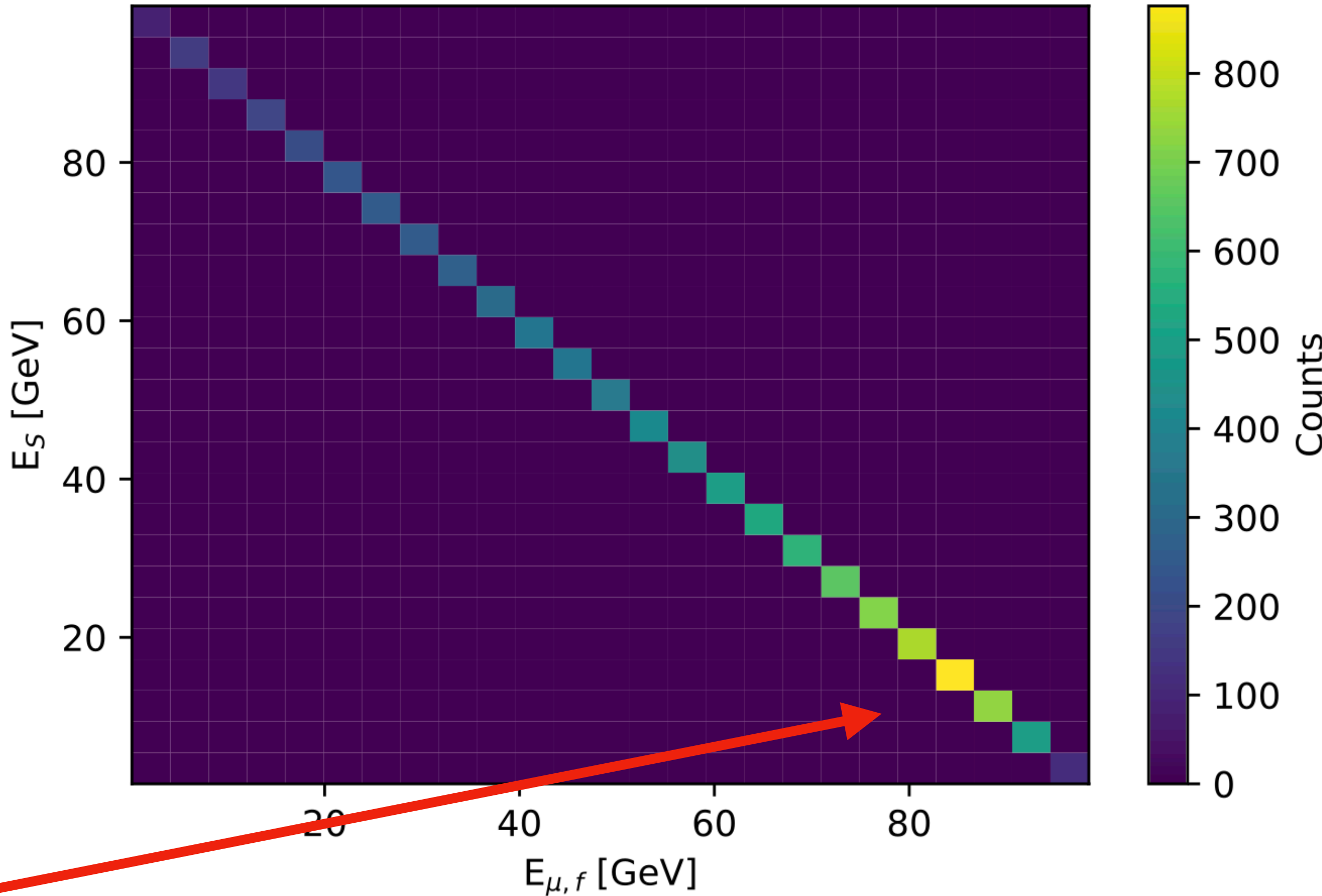
Chien-Yi Chen, Maxim Pospelov, Yi-Ming Zhong; 1701.07437

Production from Bremsstrahlung in FASER ν (cont.)

$E_{\mu,i} = 100 \text{ GeV}, m_S = 0.10 \text{ GeV}, g_S = 1$

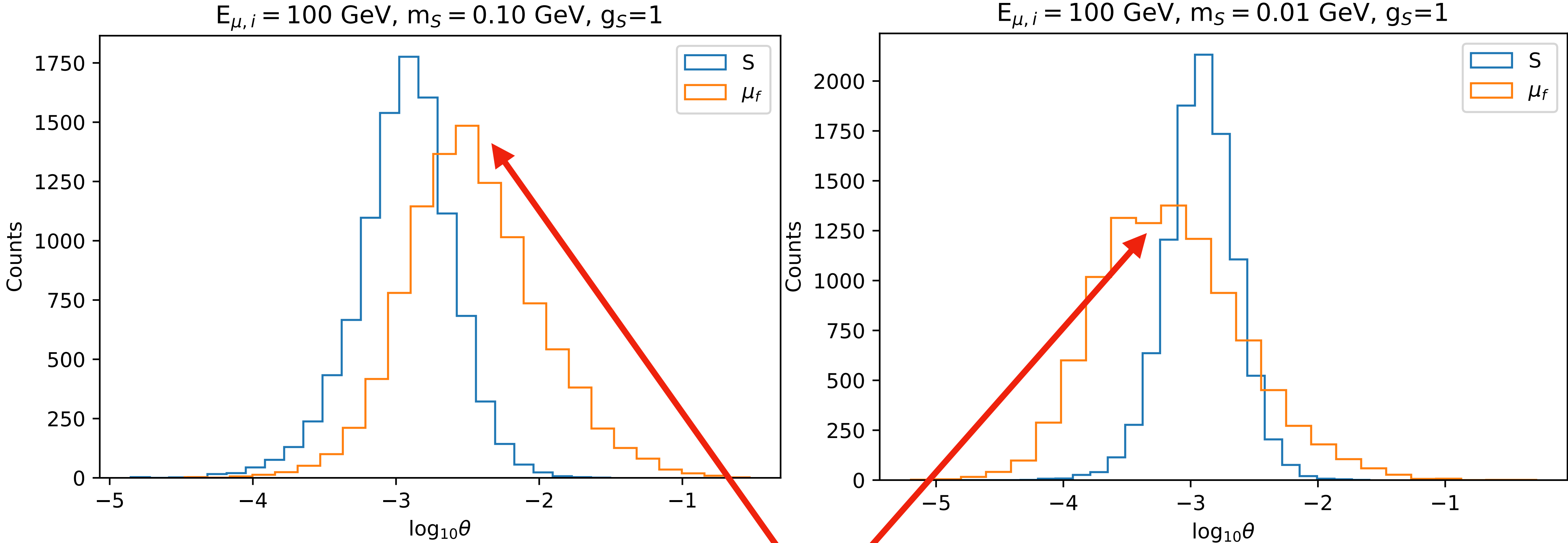


$E_{\mu,i} = 100 \text{ GeV}, m_S = 0.01 \text{ GeV}, g_S = 1$



Heavier (lighter) S takes away more (less) energy

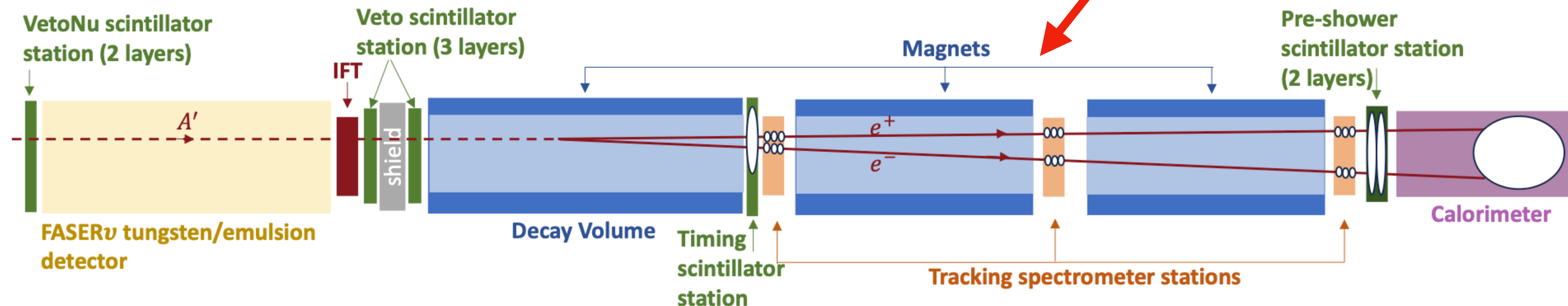
Production from Bremsstrahlung in FASER ν (cont.)



We expect the muon to enter the detector, not scatter away.

Production from Bremsstrahlung in FASER ν (cont.)

- The signal we expect is 1 muon track with some energy deposition in the calorimeter.
- But this was an important background to FASER's dark photon search.

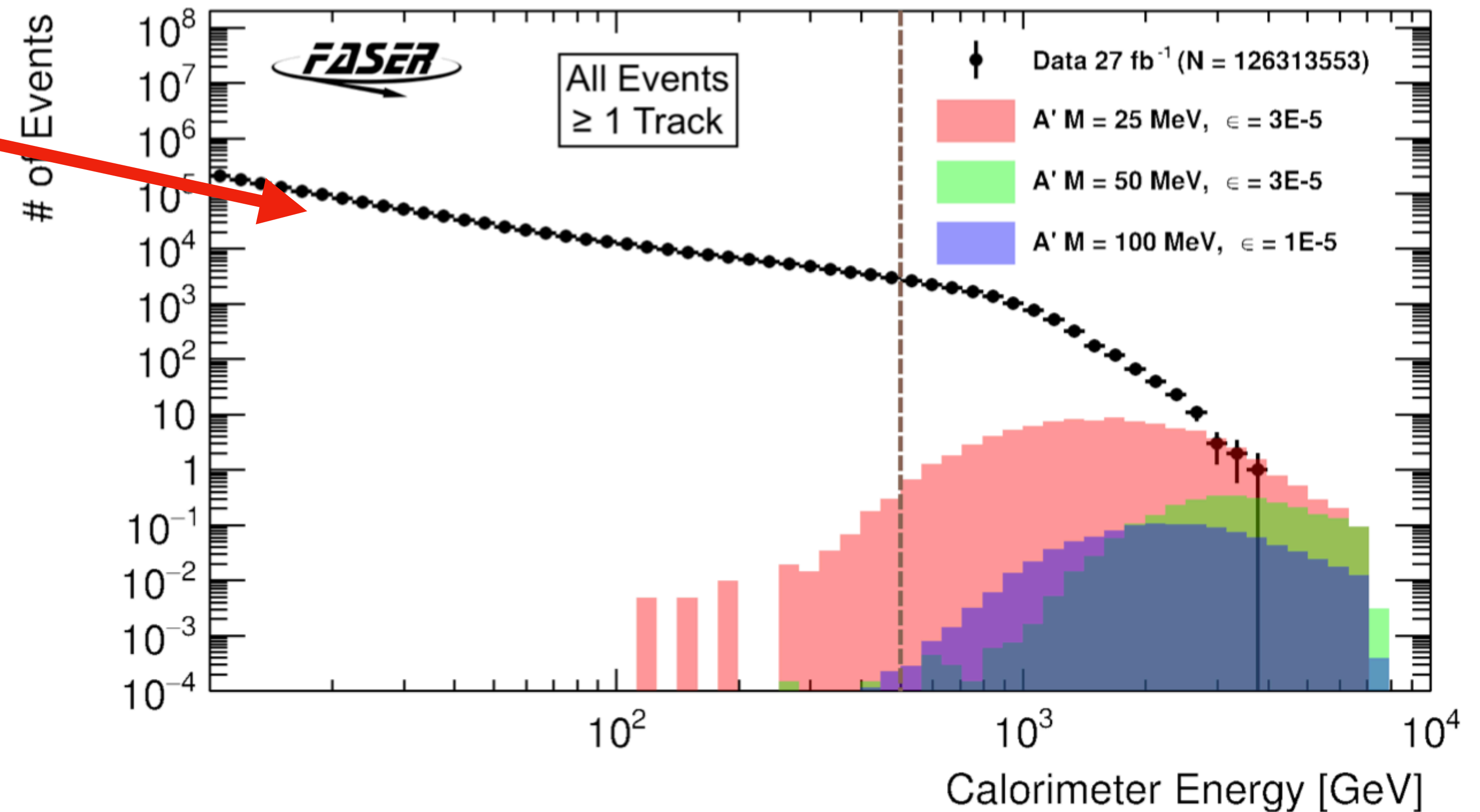


Search for Dark Photons with the
FASER detector at the LHC;
2308.05587

Production from Bremsstrahlung in FASER ν (cont.)

There is an overwhelming number of background events that can mimic our signal.

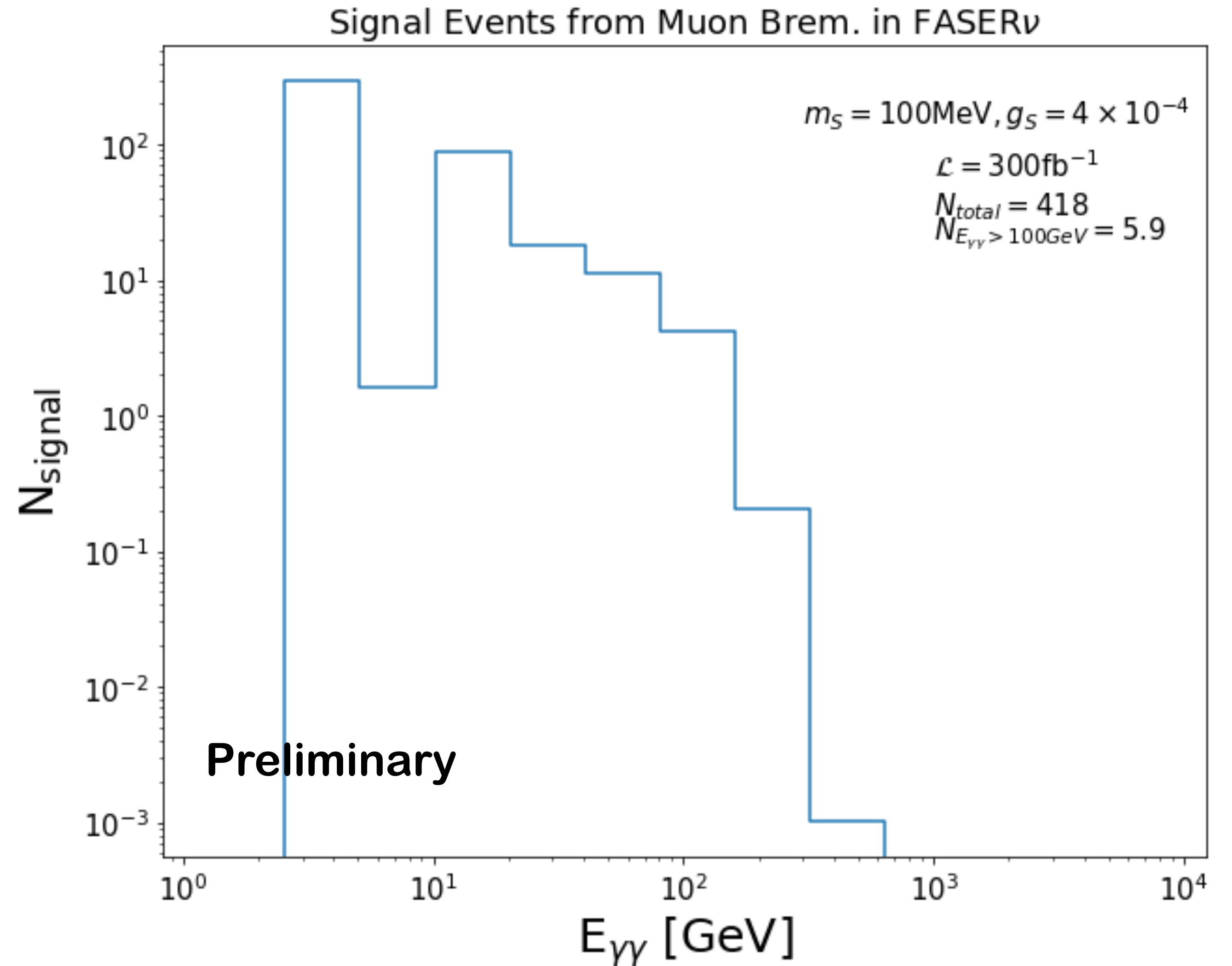
Can we use the fact that low energy S from soft muons dominate our signal reach?



Search for Dark Photons with the
FASER detector at the LHC;
2308.05587

Di-photon Energy Spectrum

- Most of the signal events have low $E_{\gamma\gamma}$.
- This is due to the short decay length requirement.

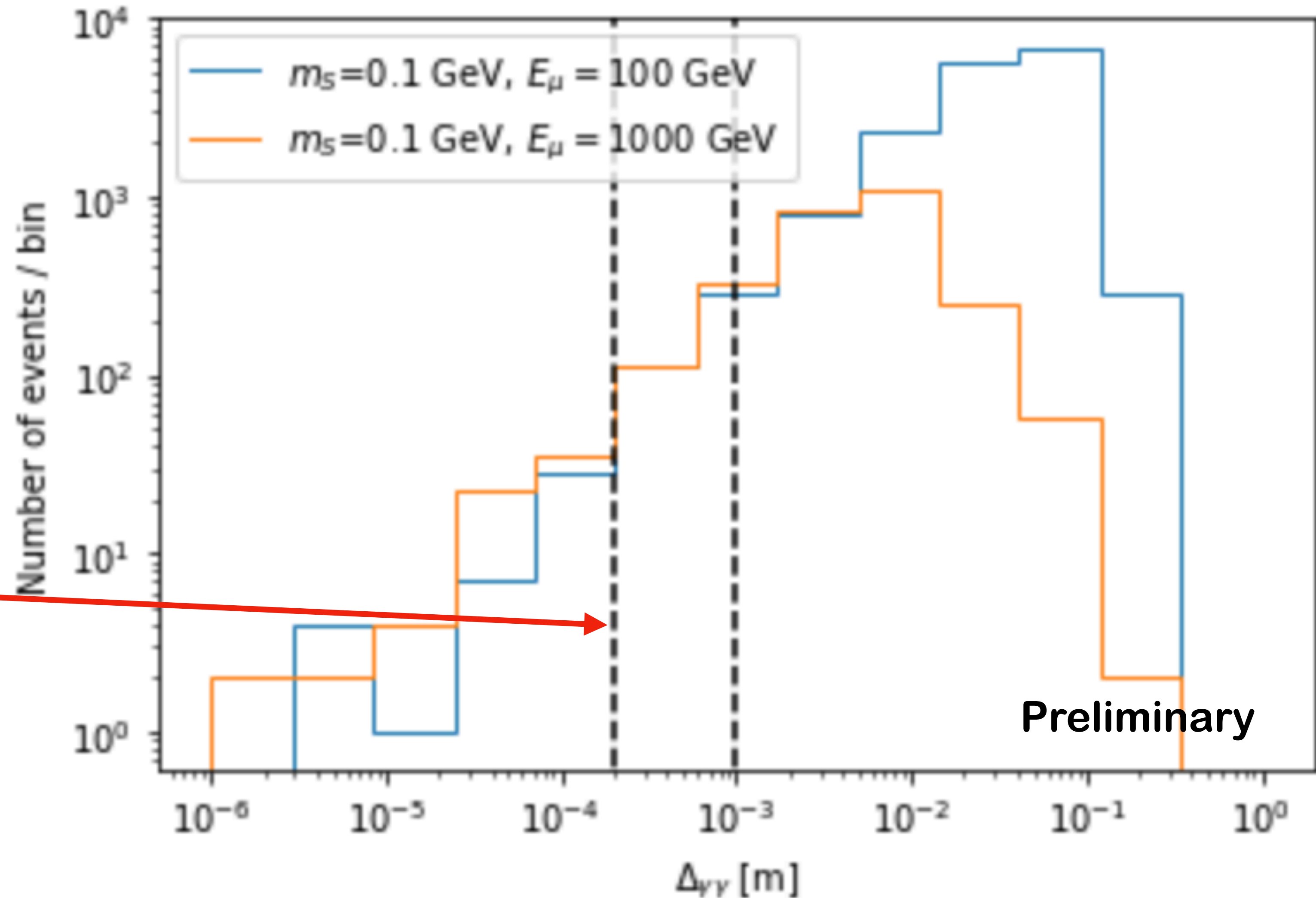


Di-photon Separation Spectrum

- Low energy S tend to decay into 2 photons with greater separation.

- $\Delta_{\gamma\gamma} \sim \frac{m_S}{E_S} * \Delta_z$

- Can we see such small spatial separation between 2 photons?



High Precision Preshower

- The FASER collaboration is working on a High Precision Preshower.

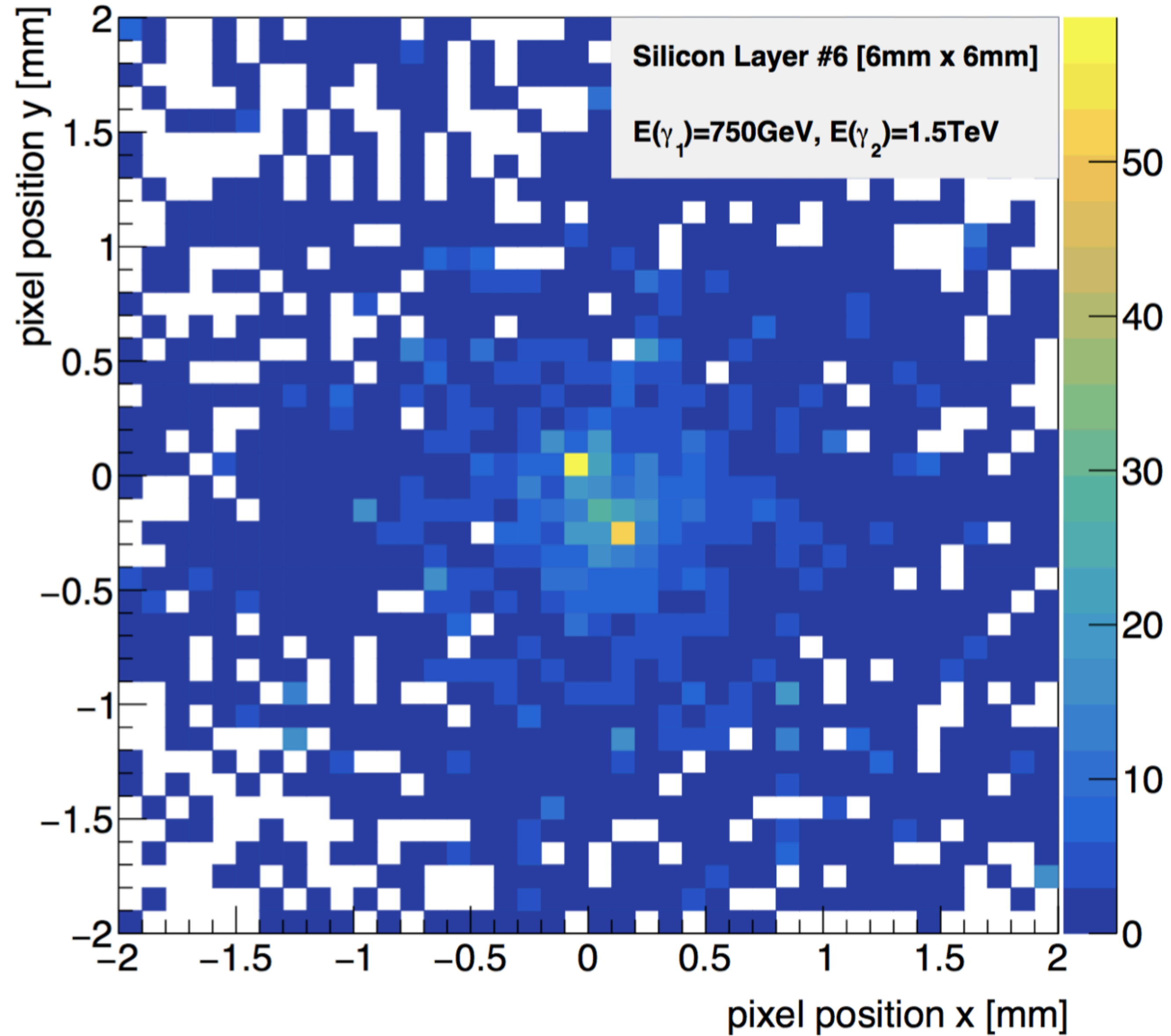
Preshower TP

ABSTRACT: The FASER detector is designed to search for light weakly interacting new particles decaying into charged final states at the LHC. While the first physics data will be taken at the start of Run 3 of the LHC program, an upgrade is already foreseen to enhance the sensitivity to long-lived particles decaying into photons. A high-precision preshower detector will be constructed within the next two years allowing to distinguish the predicted axion-like particles signature of two very closely spaced highly energetic photons. Profiting from recent developments in monolithic pixel silicon detectors, the FASER Collaboration plans to build instrumented silicon pixel detector planes with a granularity of $100 \mu\text{m}$ interleaved with tungsten absorber planes. The addition of the new pre-shower detector will expand the physics search capability of FASER.

Expected to be installed before 2025 data taking!!!

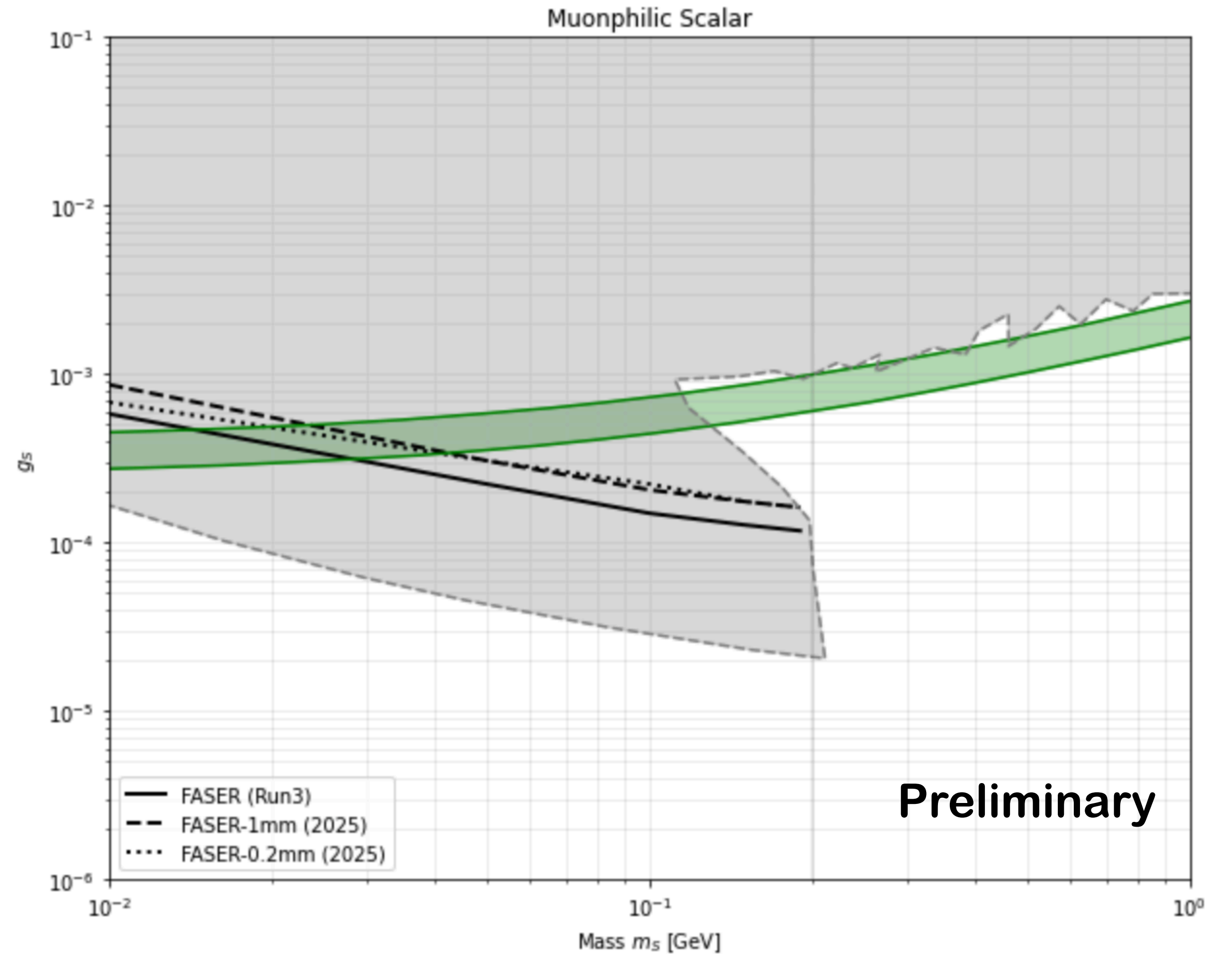
High Precision Preshower

Preshower TP



Bremsstrahlung in FASER ν With High Precision Preshower

- Requiring $\Delta_{\gamma\gamma} > 0.2$ mm suppresses most of the backgrounds.
- In 2025, FASER expects $\sim 90 \text{ fb}^{-1}$ with preshower. Run 4 proposal for FASER
- This is a reduction in luminosity ($300 \text{ fb}^{-1} \rightarrow 90 \text{ fb}^{-1}$).
- But even with only 2025 data, FASER can probe the unconstrained (g-2) band below $2 * m_{\mu}$!!!!



Summary

- There is a lot of physics to be studied in the forward region at LHC.
 - Neutrinos, **Muons**, QCD, PDFs, DM, ALPs,.....
- Neutrinos were the prime target so far, with muons an annoying background.
- We propose a simple detection scheme to probe muon (g-2) band at FASER using just 2025 data at LHC.
- It is the era of **Multimessenger Collider Physics**.

“These sources are complicated... Unless you have many ways to *look* at them, you’re not going to figure them out”

-Francis Halzen on Multimessenger Astronomy
Scientific American

These collisions are complicated... Unless you have many ways to *look* at them, you’re not going to figure them out

Multimessenger Collider Physics

Thank You

Borrowed from Max Fieg