

# FuPhy 2024



## Probing the neutron skin variations in isotope pairs by hyperon-antihyperon production in antiproton–nucleus interactions

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- Motivation
- Neutron skin
- Method
- Remarks on feasibility

# JG|U Mass – Radius Relation

- Stellar matter in hydrostatic equilibrium
- Described by GR and Tolman–Oppenheimer–Volkoff (TOV) equation

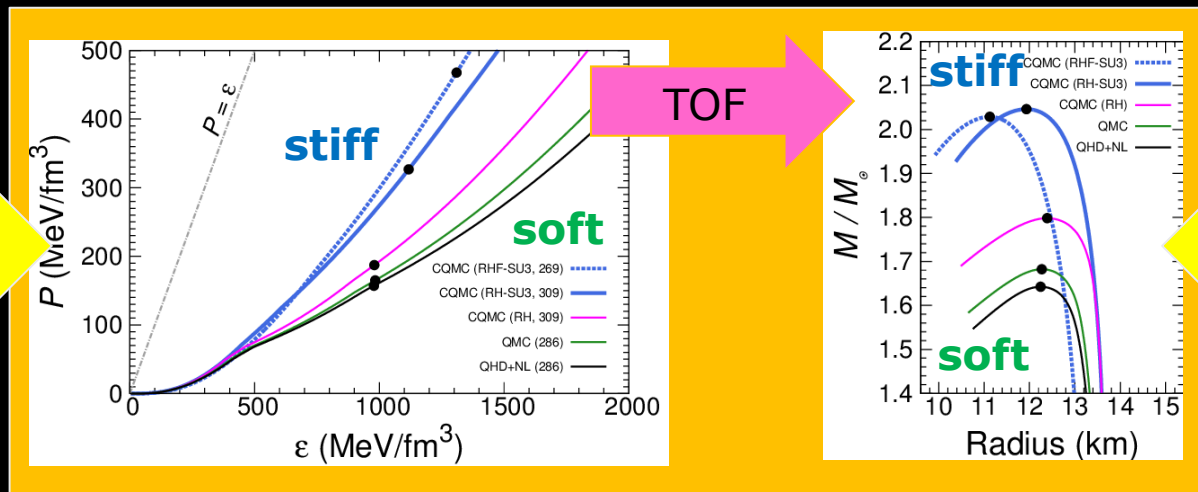
$$\frac{dP}{dr} = -\frac{G \cdot M(r) \cdot \varepsilon(P)}{r^2 c^2} \cdot \left(1 + \frac{P}{\varepsilon(P)}\right) \cdot \left(1 + \frac{4\pi r^3 P}{M(r) \cdot c^2}\right) \left(1 - \frac{2GM(r)}{rc^2}\right)^{-1}$$

- P: pressure
- $\varepsilon$ : energy density
- M(r): mass enclosed within radius r

$$\frac{dM(r) \cdot c^2}{dr} = 4\pi r^2 \varepsilon(P)$$

- Specific solution (M,R) requires Equation of state  $\varepsilon(P)$  as input
  - Soft EoS: low maximum mass and small radii
  - Stiff EoS: high maximum mass and large radii

Nuclear and particle physics



Astronomical observations

# Nuclear Equation of State

➤ Approximated by Taylor expansion around normal nuclear density  $\rho_0$

isoscalar part  $E_{SNM}$

isovector part  $E_{sym}$

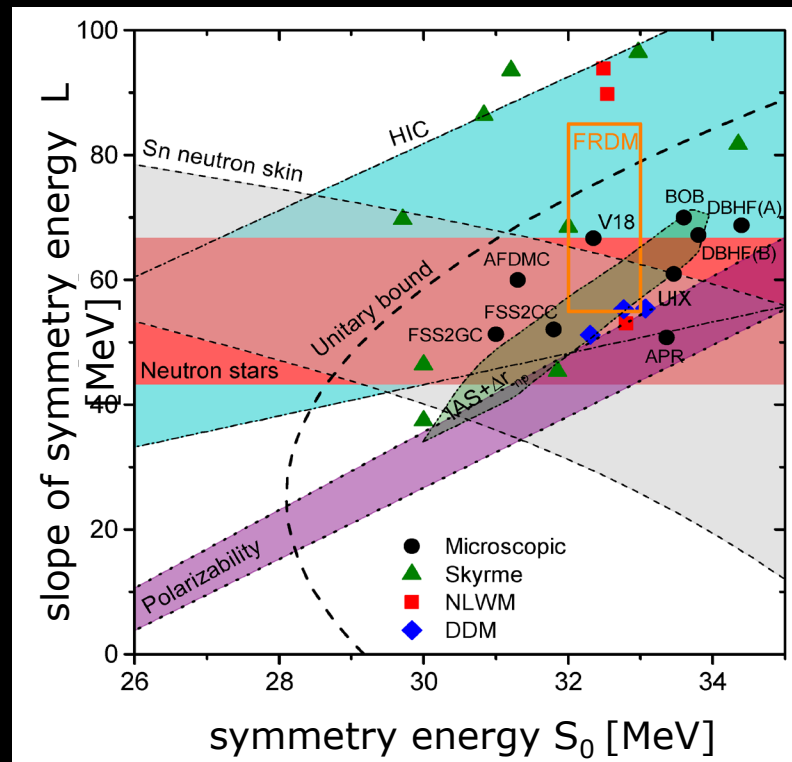
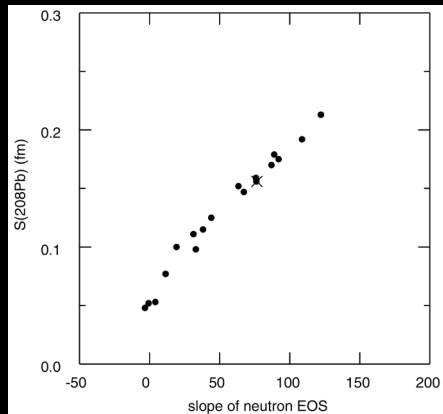
$$E(\rho, \delta) = E_0 + \frac{K_0}{2} \cdot \left( \frac{\rho - \rho_0}{3\rho_0} \right)^2 + \left[ S_0 + L \cdot \left( \frac{\rho - \rho_0}{3\rho_0} \right) + \frac{K_{sym}}{2} \cdot \left( \frac{\rho - \rho_0}{3\rho_0} \right)^2 \right] \cdot \delta^2$$

incompressibility  $K_0 = 9\rho_0^2 \frac{d^2 E_{SNM}}{d\rho^2}(\rho_0)$

symmetry energy  $S_0 = \frac{1}{2} \frac{\partial^2 E}{\partial \delta^2}(\rho_0, 0)$

slope of  $E_{sym}$   $L = 3\rho_0 \frac{dE_{sym}}{d\rho}(\rho_0)$

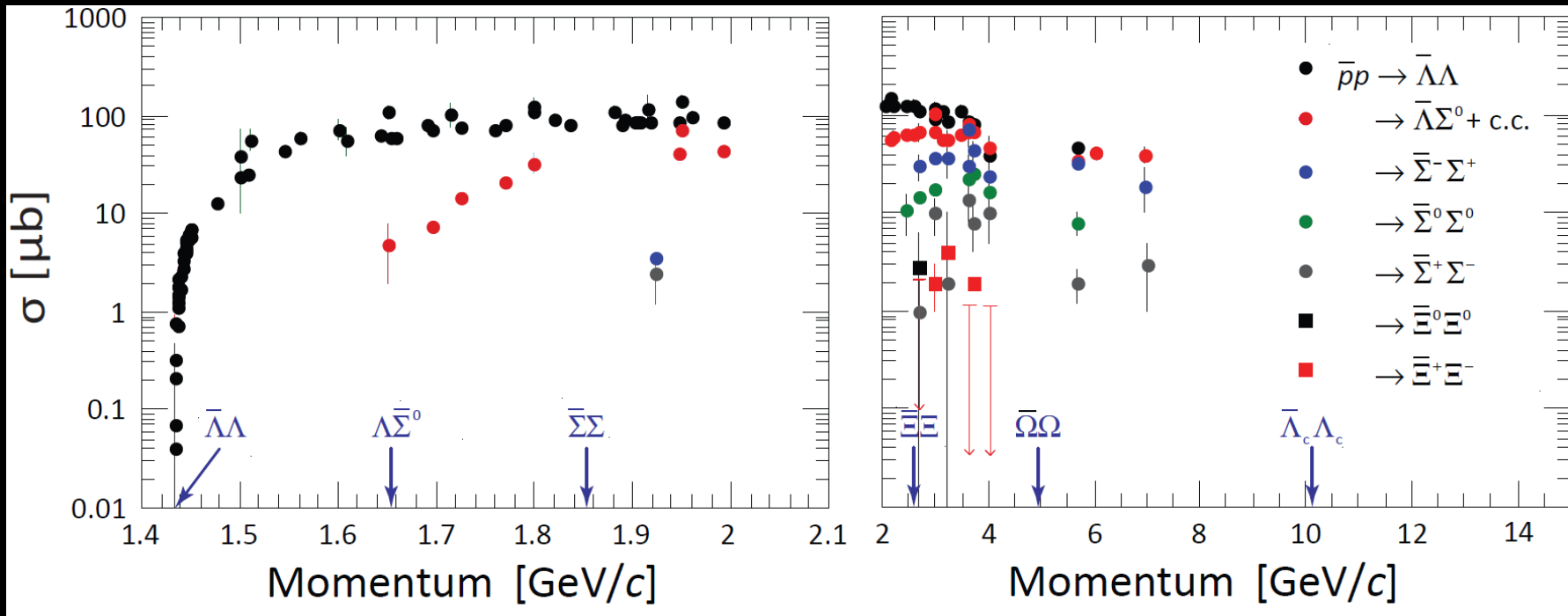
$K_{sym} = 9\rho_0^2 \frac{d^2 E_{sym}}{d\rho^2}(\rho_0)$



# Nuclear Physics with $\bar{p}+A$ interactions

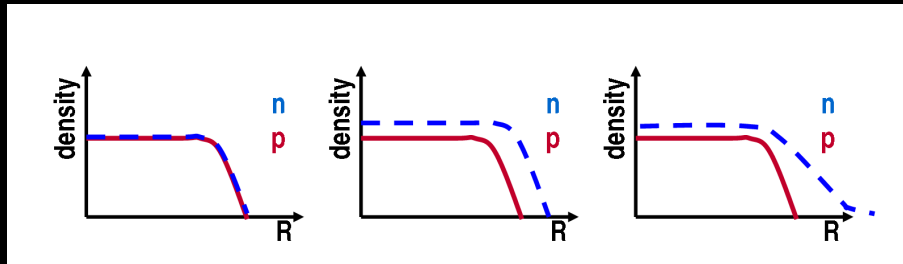
neutron skin  
 anihyperons in nuclei  
 hyper- and antihyper atoms  
 double hypernuclei

$\bar{\Lambda}\Lambda$  and  $\bar{\Lambda}\Sigma$   
 $\bar{\Lambda}\Lambda$   
 tagged negativ (anti)hyperons  
 tagged  $\Xi$

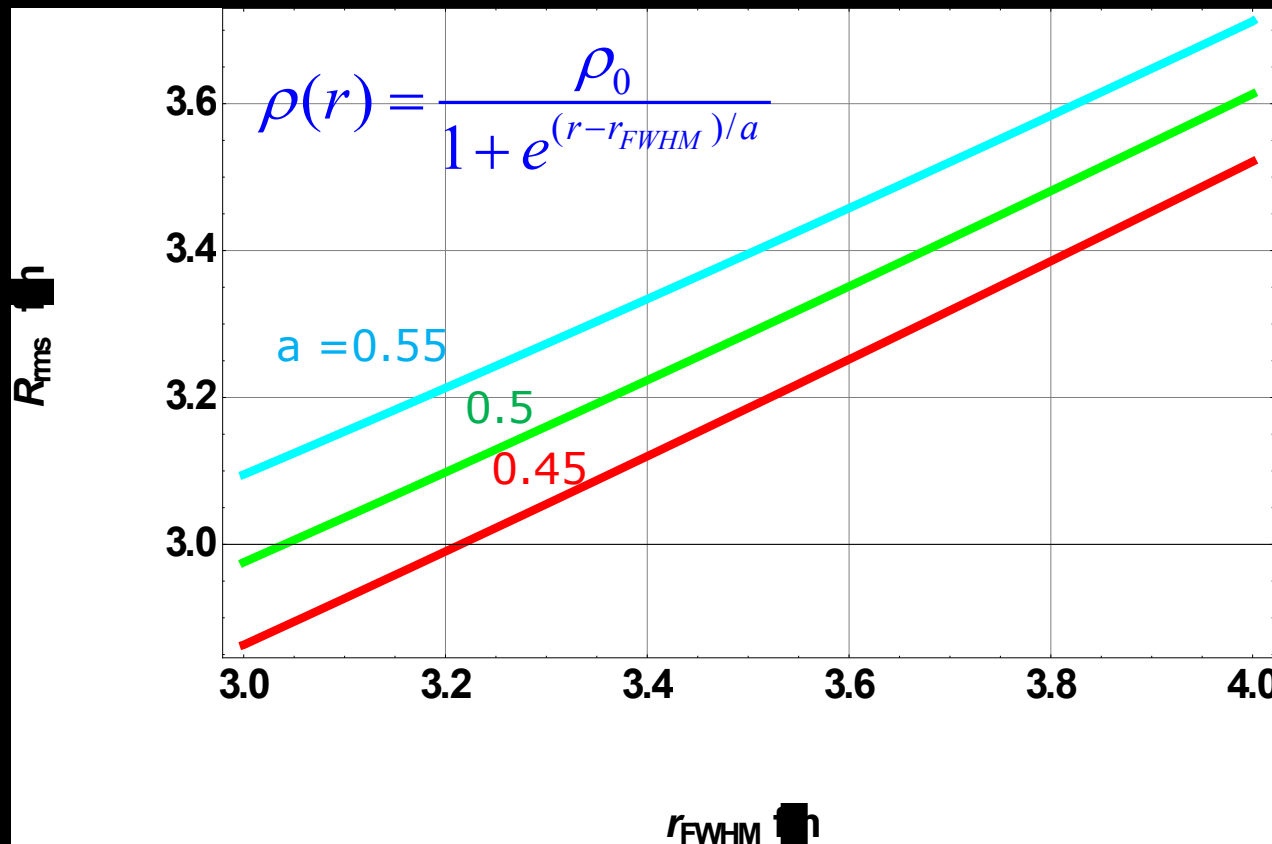


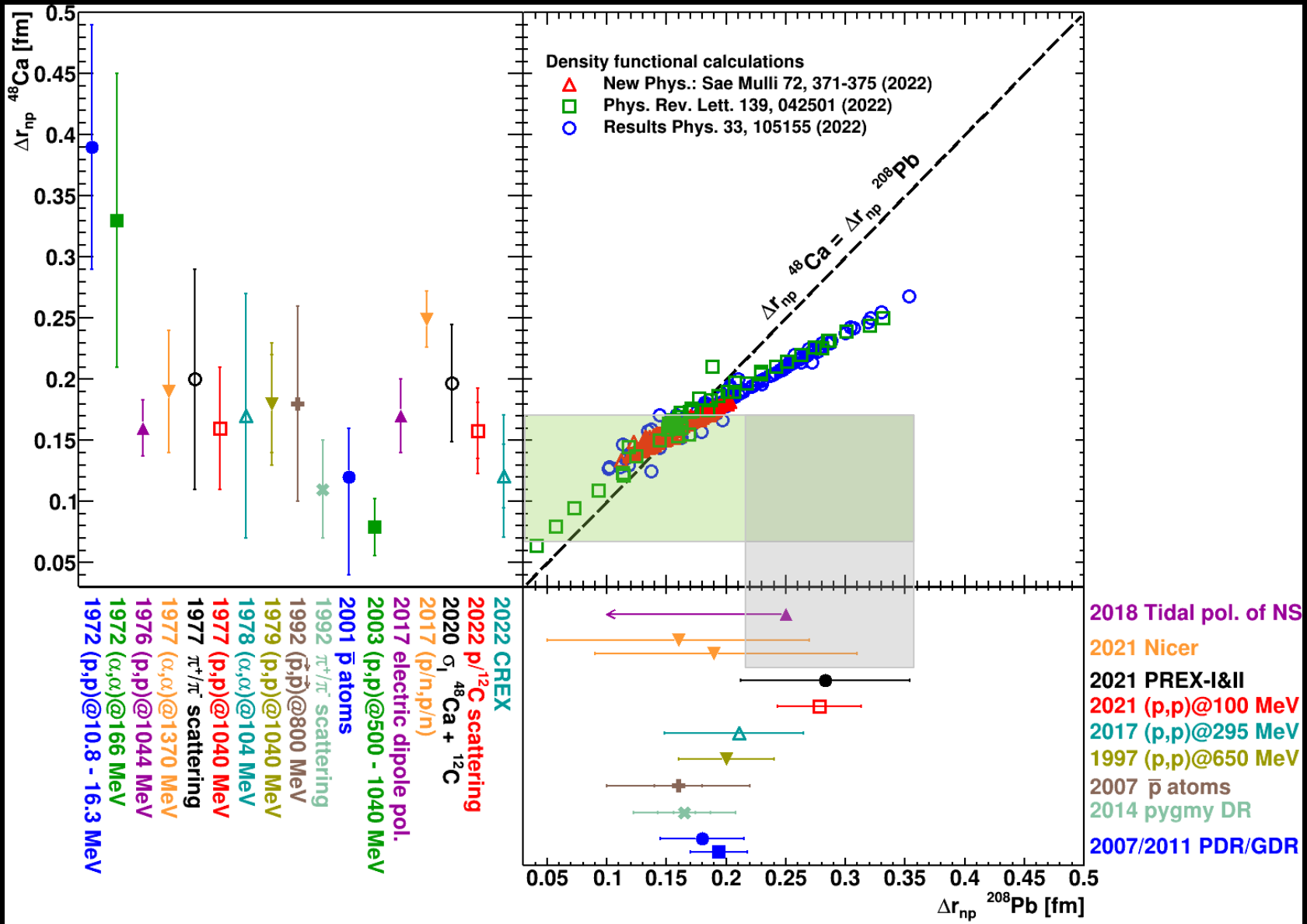
# Neutron skin

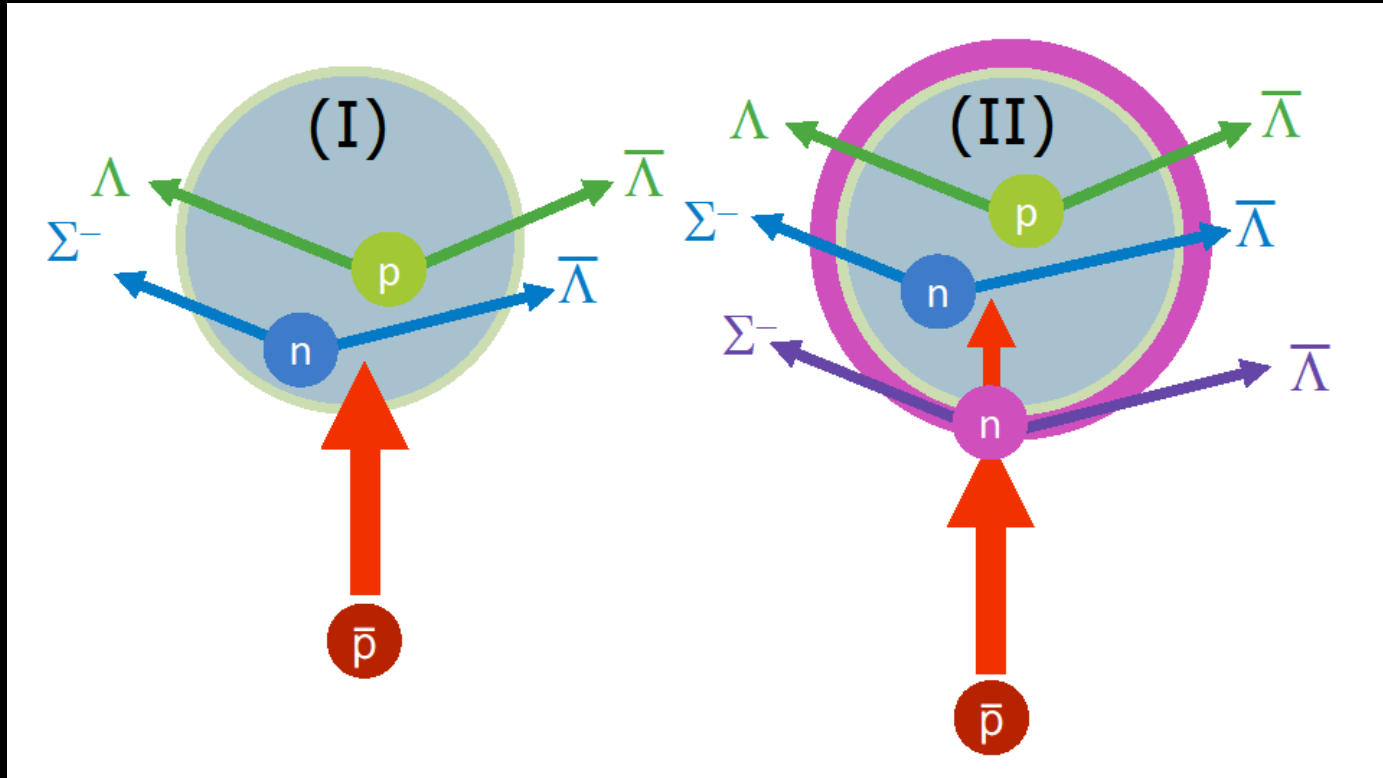
- Neutron skin defined as  $R_{rms}(n) - R_{rms}(p)$
- Definition of neutron skins is not unique



M Thiel et al 2019 J. Phys. G: Nucl. Part. Phys. 46 093003







$$DR = \frac{p_{\Sigma^- \bar{\Lambda}}^a / p_{\Lambda \bar{\Lambda}}^a}{p_{\Sigma^- \bar{\Lambda}}^b / p_{\Lambda \bar{\Lambda}}^b} \approx \frac{1 + p_{abs}}{1 - p_{abs}}$$

$$\text{with } p_{abs} = 1 - \exp\left(-\sigma_{\bar{p}n} \cdot \int_{\Delta_n} \rho_n \cdot dr_n\right)$$

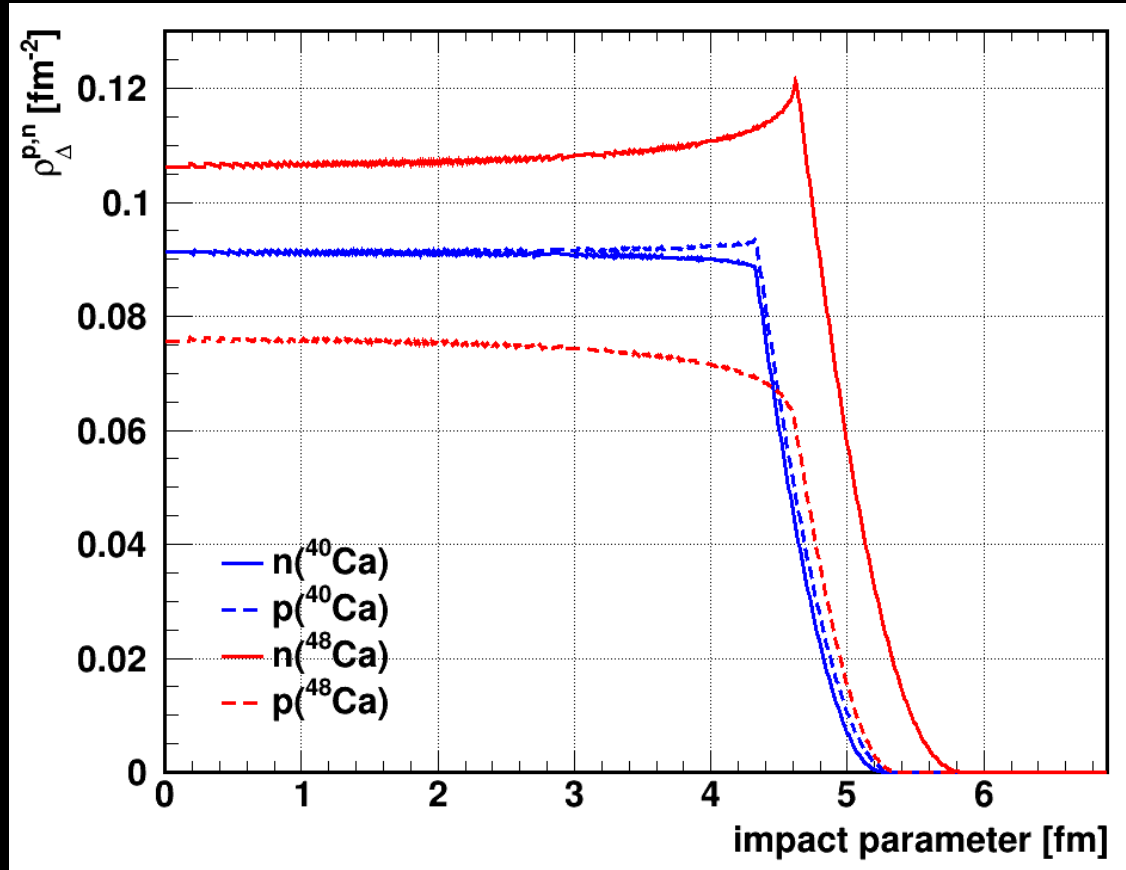
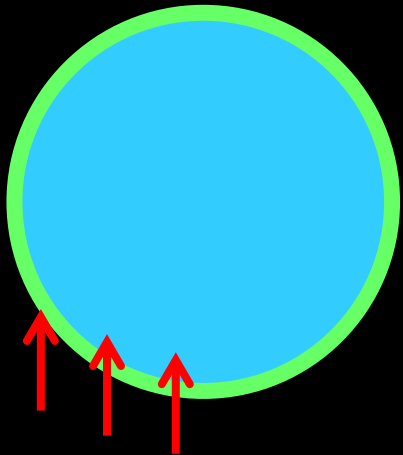
# JG|U Studied isotope pairs

isotope	abundance [%]	exp. proton radius $R_p^{exp}$ [fm]	proton radius $R_p$ [fm]	neutron radius $R_n$ [fm]	neutron skin $\Delta R_{pn}$ [fm]	skin difference $\Delta_n$ [fm]	RMF model
$^{20}\text{Ne}$	90.5	$2.992 \pm 0.008$ [51]	2.782	2.758	-0.024	–	NL3 [52]
$^{22}\text{Ne}$	9.3	$2.986 \pm 0.021$ [51]	2.800	2.887	0.087	0.111	NL3 [52]
$^{40}\text{Ca}$	96.9	$3.4776 \pm 0.0019$ [53]	3.452	3.416	-0.036	–	NL1 [52]
$^{48}\text{Ca}$	0.187	$3.4786 \pm 0.0106$ [53, 54]	3.525	3.731	0.206	0.242	NL1 [52]
$^{40}\text{Ca}$	96.9	$3.4776 \pm 0.0019$ [53]	3.391	3.354	-0.037	–	NL3 [52]
$^{48}\text{Ca}$	0.187	$3.4786 \pm 0.0106$ [53, 54]	3.472	3.659	0.187	0.224	NL3 [52]
$^{40}\text{Ca}$	96.9	$3.4776 \pm 0.0019$ [53]	3.396	3.360	-0.036	–	NL3* [55]
$^{48}\text{Ca}$	0.187	$3.4786 \pm 0.0106$ [53, 54]	3.475	3.666	0.191	0.227	NL3* [55]
$^{40}\text{Ca}$	96.9	$3.4776 \pm 0.0019$ [53]	3.117	3.090	-0.028	–	Set I [56]
$^{48}\text{Ca}$	0.187	$3.4786 \pm 0.0106$ [53, 54]	3.243	3.357	0.1114	0.142	Set I [56]
$^{58}\text{Ni}$	68.1	$3.770 \pm 0.002$ [57]	3.769	3.768	0.000	–	NL3 [52]
$^{64}\text{Ni}$	0.926	$3.854 \pm 0.002$ [57]	3.822	3.947	0.125	0.125	NL3 [52]
$^{129}\text{Xe}$	26.4	$4.7775 \pm 0.0050$ [53]	4.768	4.932	0.164	–	NL3 [52]
$^{130}\text{Xe}$	4.1	$4.7818 \pm 0.0049$ [53]	4.776	4.950	0.174	0.010	NL3 [52]
$^{131}\text{Xe}$	21.2	$4.7808 \pm 0.0049$ [53]	4.784	4.968	0.184	0.020	NL3 [52]
$^{132}\text{Xe}$	26.9	$4.7859 \pm 0.0048$ [53]	4.792	4.986	0.194	0.030	NL3 [52]
$^{134}\text{Xe}$	10.4	$4.7899 \pm 0.0047$ [53]	4.809	5.023	0.213	0.049	NL3 [52]
$^{136}\text{Xe}$	8.9	$4.7964 \pm 0.0047$ [53]	4.826	5.059	0.233	0.069	NL3 [52]



# Finite impact parameter range

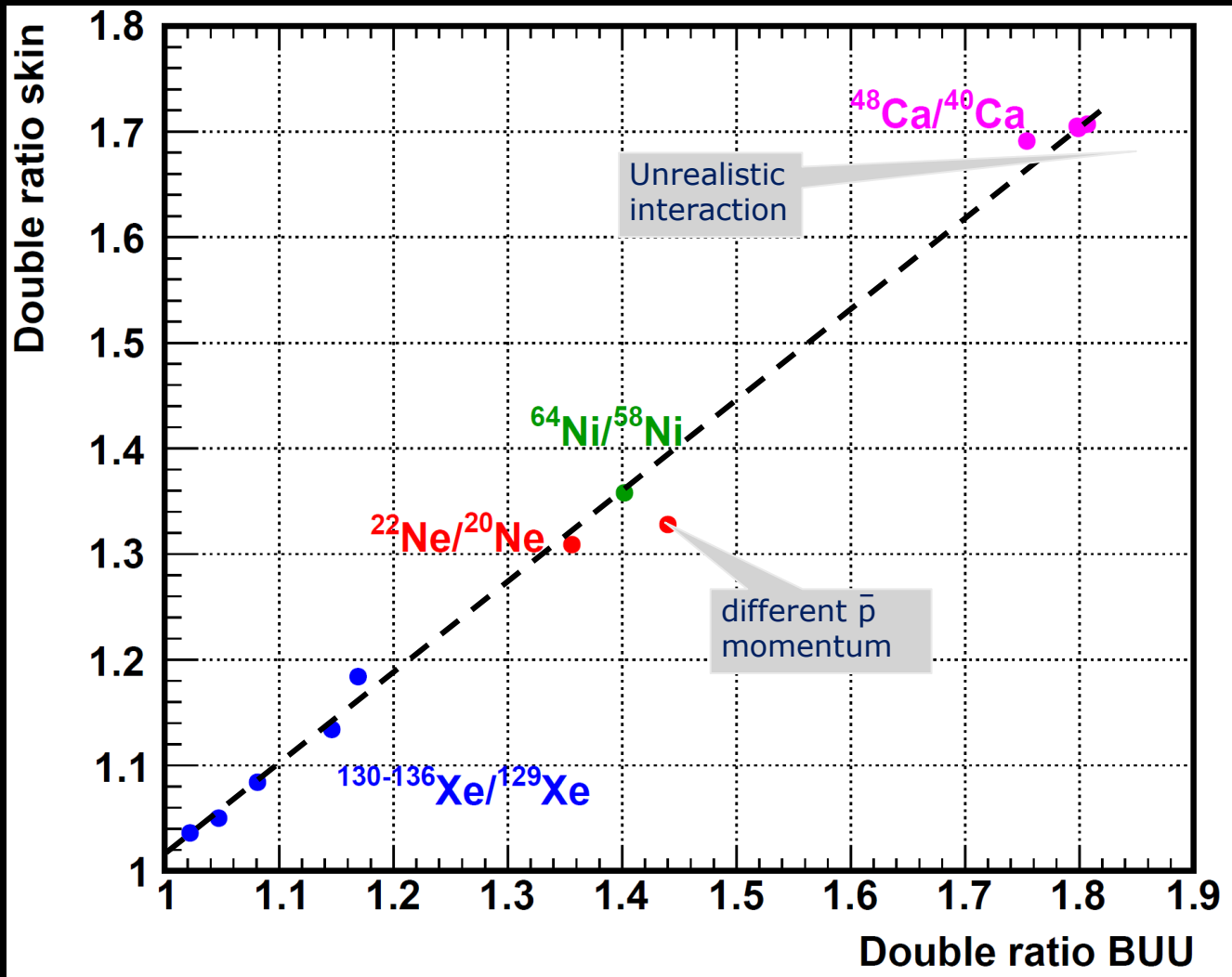
- ▶ Straight line geometry
- ▶ Integrated density within 1 interaction length  $\sim 0.18\text{fm}^2$  at  $2.4\text{GeV}/c$



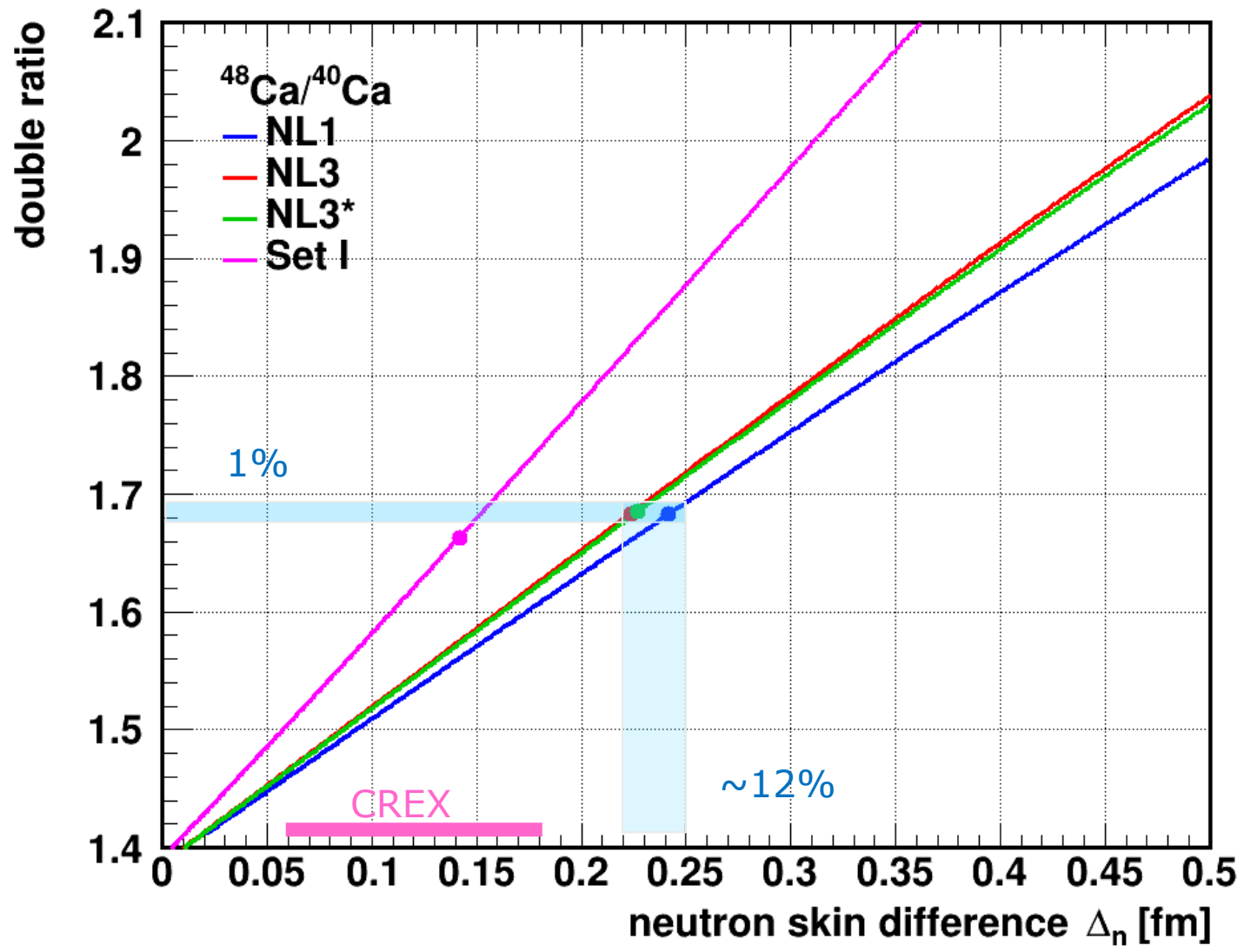
- ▶ Impact parameter averaged neutron-to-proton ratio:  $^{40}\text{Ca}$  0.97 and  $^{48}\text{Ca}$  1.65

target	RMF model	$p(\bar{p})$ [ GeV/c ]	events [10 <sup>6</sup> ]	$\Lambda\bar{\Lambda}$	$\Sigma^-\bar{\Lambda}$	double ratio		
						GiBUU	Equation 9	$\rho_\Delta$
<sup>20</sup> Ne	NL3 [52]	1.7	99	7579	1629	—	—	—
<sup>22</sup> Ne	NL3 [52]	1.7	98	6347	1965	1.440±0.054	1.115	1.328
<sup>20</sup> Ne	NL3 [52]	2.4	167	32387	10870	—	—	—
<sup>22</sup> Ne	NL3 [52]	2.4	171	29227	13297	1.356±0.021	1.100	1.309
<sup>40</sup> Ca	NL1 [52]	2.4	415	76323	22880	—	—	—
<sup>48</sup> Ca	NL1 [52]	2.4	450	66694	36074	1.799±0.018	1.224	1.703
<sup>40</sup> Ca	NL3 [52]	2.4	415	74280	21827	—	—	—
<sup>48</sup> Ca	NL3 [52]	2.4	450	61313	32391	1.798±0.019	1.207	1.705
<sup>40</sup> Ca	NL3* [55]	2.4	415	78753	23438	—	—	—
<sup>48</sup> Ca	NL3* [55]	2.4	450	64212	34523	1.807±0.018	1.210	1.707
<sup>40</sup> Ca	Set I [56]	2.4	415	67000	17446	—	—	—
<sup>48</sup> Ca	Set I [56]	2.4	450	57016	26042	1.754±0.020	1.129	1.691
<sup>58</sup> Ni	NL3 [52]	2.4	100	16811	5230	—	—	—
<sup>64</sup> Ni	NL3 [52]	2.4	108	14978	6534	1.402±0.030	1.113	1.358
<sup>129</sup> Xe	NL3 [52]	2.4	109	13717	6238	—	—	—
<sup>130</sup> Xe	NL3 [52]	2.4	109	13394	6225	1.022±0.022	1.001	1.036
<sup>131</sup> Xe	NL3 [52]	2.4	109	13403	6379	1.047±0.023	1.018	1.050
<sup>132</sup> Xe	NL3 [52]	2.4	109	13335	6556	1.081±0.023	1.026	1.084
<sup>134</sup> Xe	NL3 [52]	2.4	109	12771	6656	1.146±0.025	1.043	1.134
<sup>136</sup> Xe	NL3 [52]	2.4	109	12680	6739	1.169±0.025	1.062	1.184

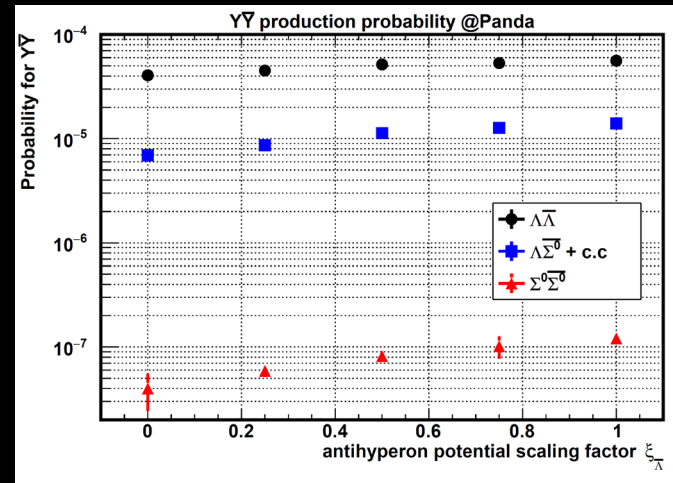
# BUU vs. Schematic model



- ▶ Advantage of schematic model: explore sensitivity between neutron skin and double ratio



- PANDA detector as reference
- Goal: 1% measurement of double ratio
- $2 \cdot 10^6$  interaction per second
- Reconstruction efficiency
  - $\Lambda$  30%
  - $\bar{\Lambda}$  30%
  - $\Sigma^-$  3% (detection of n required!)
- $S/B = 73$
- Statistical uncertainty limited by  $\Sigma^-$
- Required time  $\sim 1$  day at PANDA
  
- There are many interesting topics for energetic antiproton beams



D16.2: Development of dedicated simulation and software tools for the measurement of antihyperons in nuclei.

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 representing the Networking activity THEIA (WP16) within STRONG-2020

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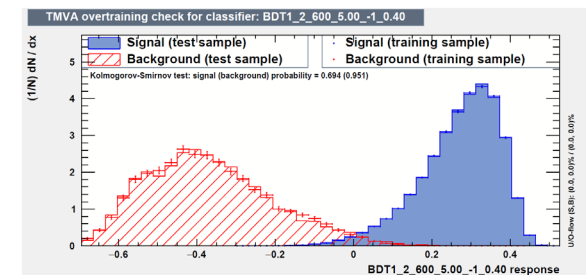


Figure 10: Output of the boosted decision tree (BDT) algorithm for background events (red histogram) and signal events (blue histogram).