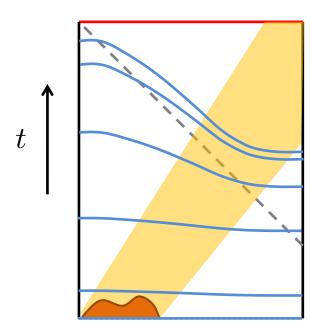
On the quantum mechanics of old black holes



Márk Mezei (Oxford)

Eurostrings 2024 meets FPUK Southampton, 5/9/2024

Renaissance of Euclidean gravity

Euclidean gravitational path integral

- (Mesoscopic) derivation of black hole entropy [Gibbons-Hawking, ...]
- No boundary wavefunction of the Universe [Hartle-Hawking, ...]
- Problems: relation to Lorentzian approaches, wormholes, conformal mode problem, ...

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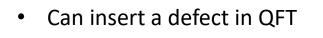
Renaissance in holographic context

- Derivation of holographic entanglement entropy [Ryu-Takayanagi, Lewkowycz-Maldacena, ...]
- Wormholes encode spectral statistic of dual quantum system [Saad-Shenker-Stanford, ...]
- (Mesoscopic) resolution of the black hole information paradox [Penington, Almheiri et al., East Coast & West Coast papers, ..., Kumar's talk]

Relating path integral and Hilbert space descriptions • Cut open QFT on a sphere $\phi(x)$ $\psi^*[\phi(x)]$ $\psi[\phi(x)]$

Relating path integral and Hilbert space descriptions

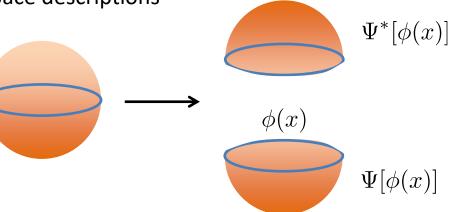
• Cut open QFT on a sphere



$$\mathbf{L} = \exp\left(-h\int dx\,\phi(x)\right)$$

• Two ways of computing

$$\langle L \rangle = \int D\Phi \ L[\phi] \ e^{-S[\Phi]}$$
 Euclidean
= $\int D\phi \ L[\phi] \left| \Psi[\phi] \right|^2$ Hilbert space



Relating path integral and Hilbert space descriptions

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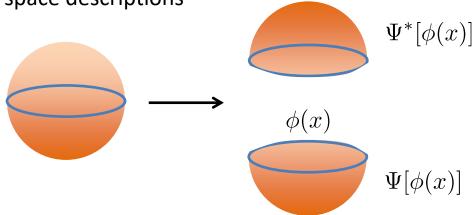
• Can insert a defect in QFT

$$\boldsymbol{L} = \exp\left(-h\int dx\,\phi(x)\right)$$

$$egin{aligned} \langle m{L}
angle &= \int D\Phi \; m{L}[\phi] \, e^{-S[\Phi]} & & ext{Euclidean} \ &= \int D\phi \; m{L}[\phi] \, |\Psi[\phi]|^2 & & ext{Hilbert space} \end{aligned}$$

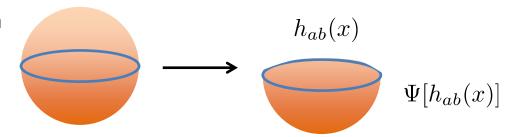
• Complementary advantages:

Euclidean:saddle points, analytic continuation to LorentzianHilbert space:unitarity, probabilistic interpretation, integer degeneracies



Can repeat in gravity

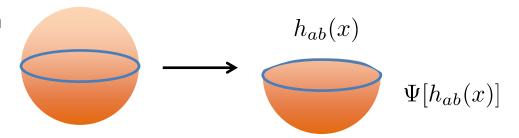
 Cut open sphere gravity path integral along a geodesic (minimal volume slice) [Wheeler-DeWitt, ..., Witten]



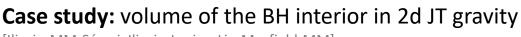
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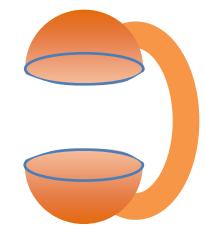
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- Need to take care of diffeomorphism invariance
- Can insert (diff. invariant) operators, can compute both Euclidean and Hilbert space way
- What to do with (bra-ket) wormholes? Should we cut them open too? [Marolf-Maxfield, Gorbenko's talk]
- Geometry now fluctuating



[Iliesiu-MM-Sárosi, Iliesiu-Levine-Lin-Maxfield-MM]

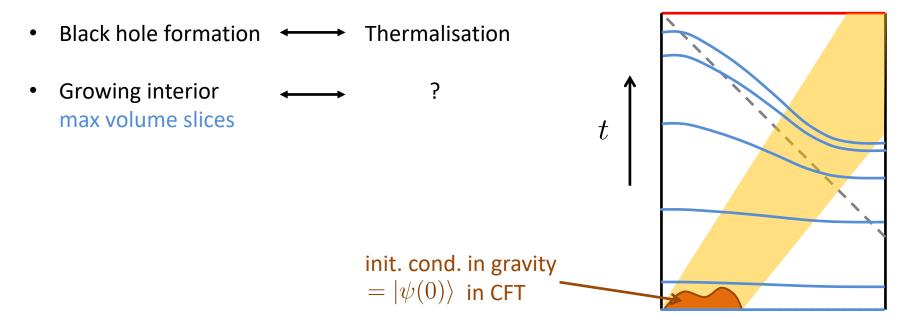


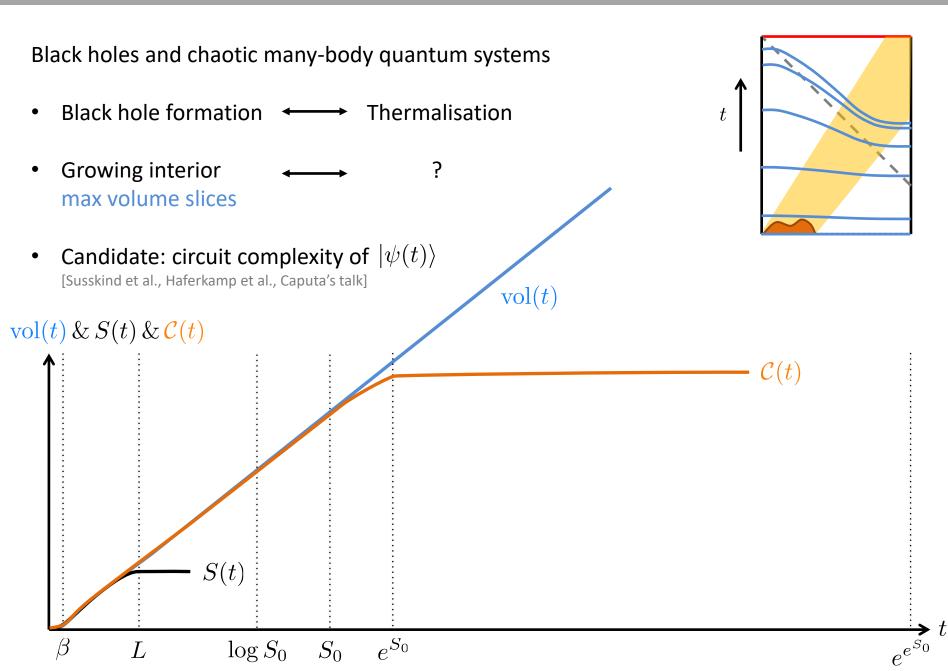
Cross-fertilisation between black holes and chaotic many-body quantum systems [Sonner's talk]

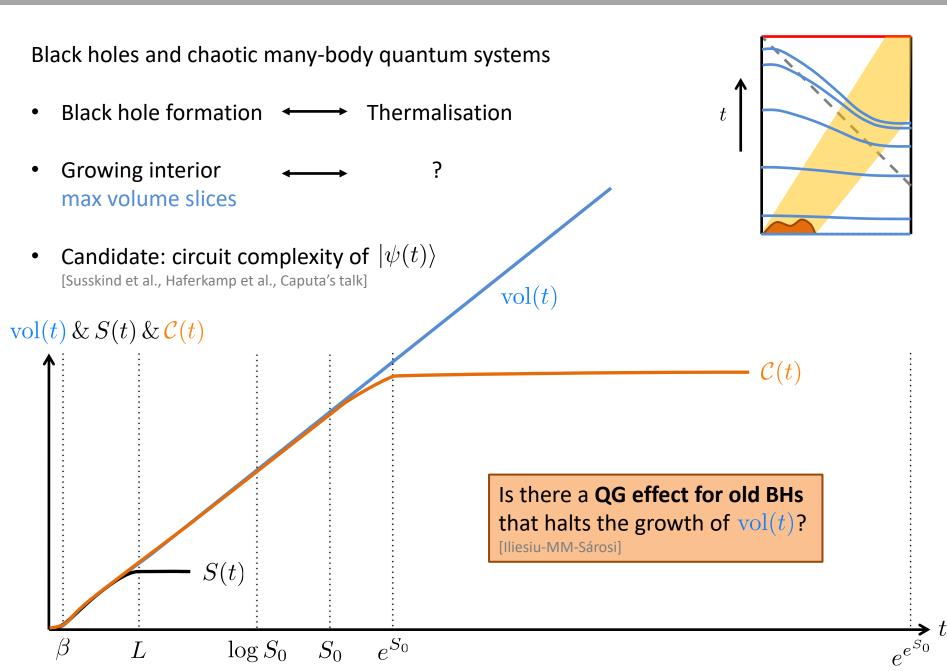
•

Black hole formation \longleftrightarrow Thermalisation $t \uparrow$ init. cond. in gravity $= |\psi(0)\rangle$ in CFT

Cross-fertilisation between black holes and chaotic many-body quantum systems [Sonner's talk]







Outline

Volume in Euclidean gravity

Hilbert space view

Summary

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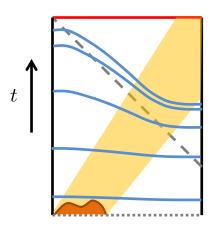
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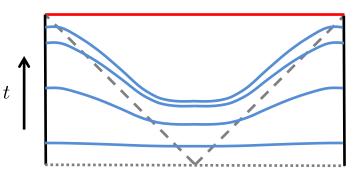
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Summary

In search for the simplest model

• Analogue setup in pure gravity [Hartman-Maldacena]

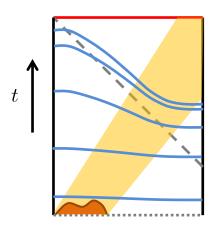


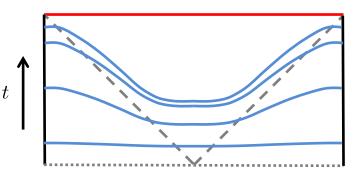


Eternal BH = $|TFD\rangle$ in CFT \otimes CFT

In search for the simplest model

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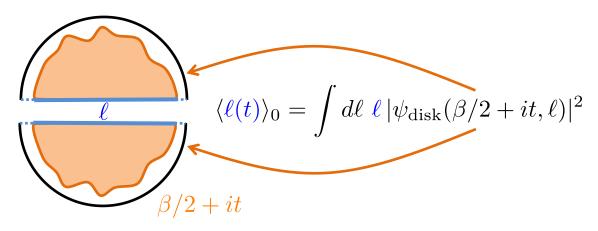
Eternal BH = $|TFD\rangle$ in CFT \otimes CFT

 Lowest dimension: 2d JT (dilaton) gravity Also effective theory for near-extremal BHs, UV complete on its own

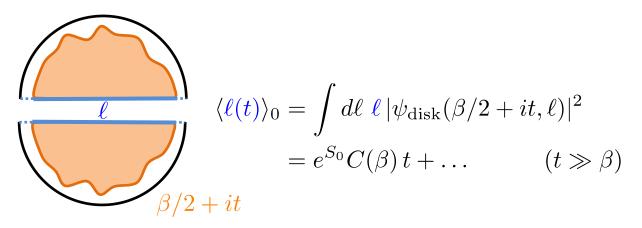
[Teitelboim, Jackiw, Almheiri-Polchinski, Jensen, Maldacena-Stanford-Yang]

$$Z(\beta) = \int Dg D\phi \ e^{-S_0 \chi - I_{\rm JT}}$$
$$I_{\rm JT} = \int \phi \left(R + 2\right) + \text{bdy term}$$

• Genus 0 [Yang]

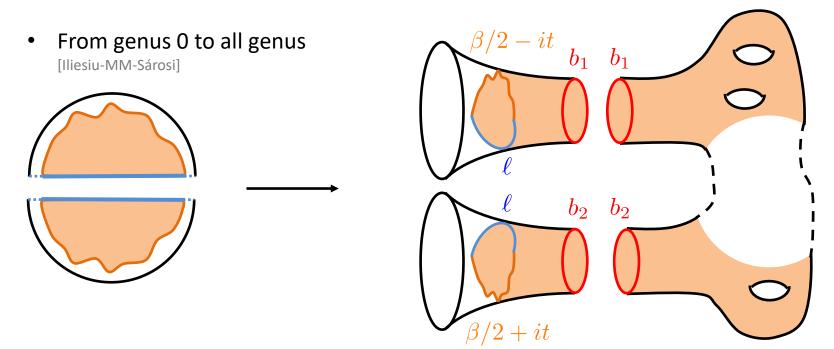


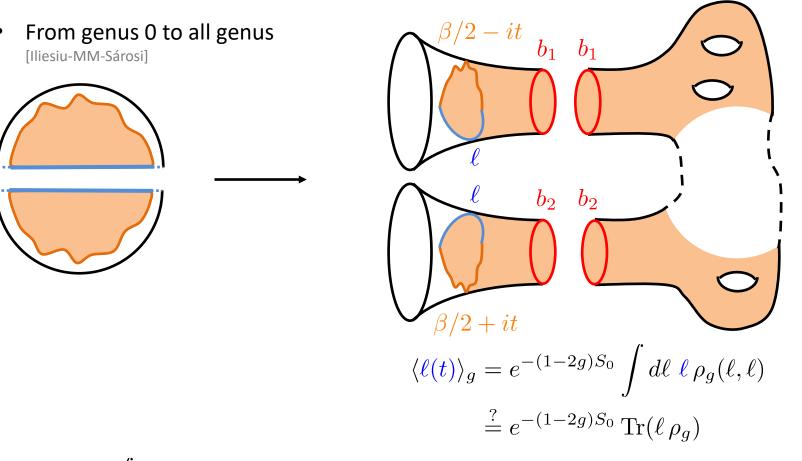
• Genus 0 [Yang]



- Organise path integral computation in way suggestive of Hilbert space view
- Perturbative gravity insensitive to the Heisenberg time e^{S_0}

Need nonperturbative effects for old BHs





• Puzzle:
$$\int d\ell \, \rho_g(\ell,\ell) = \infty$$

Unclear how to cut open QG path integral

Evaluation

• Genus expansion resummed through RMT, JT is dual to an ensemble of QMs

[Saad-Shenker-Stanford, Mirzakhani, Eynard-Orantin, ..., de Boer's talk]

$$\left\langle \ell(t) \right\rangle \approx \frac{1}{Z(\beta)} \int dE \, d\omega \, \frac{\left\langle \rho(E + \omega/2)\rho(E - \omega/2) \right\rangle}{\omega^2 \left\langle \rho(E) \right\rangle} \, e^{-\beta E} \left[1 - \cos(\omega t) \right] \qquad (t \gg \beta)$$

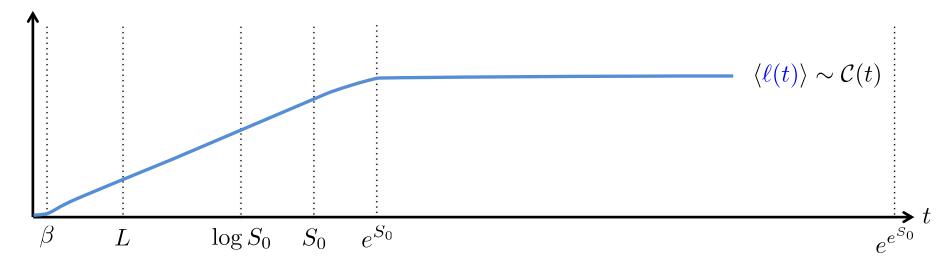
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Linear growth and saturation [Iliesiu-MM-Sárosi]



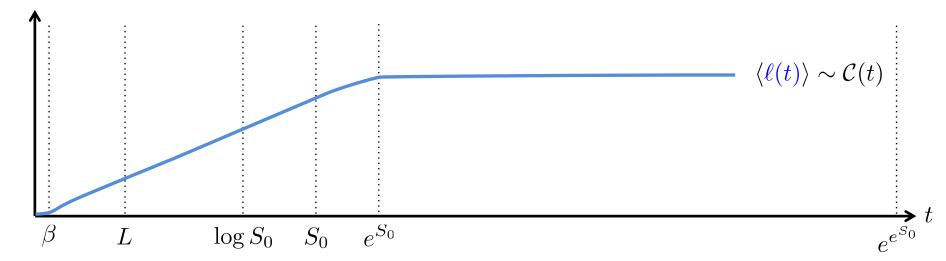
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• Linear growth and saturation [Iliesiu-MM-Sárosi]



• Motivates spectral complexity for an individual Hamiltonian

$$\mathcal{C}(t) = \sum_{E_1 \neq E_2} \frac{e^{-\beta(E_1 + E_2)/2}}{(E_1 - E_2)^2} \left[1 - \cos\left((E_1 - E_2)t\right)\right]$$

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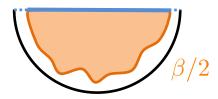
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Goal: understand and go beyond Euclidean result using Hilbert space perspective [Iliesiu-Levine-Lin-Maxfield-MM]

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Idea: semiclassical space of states with modified inner product

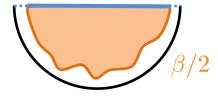
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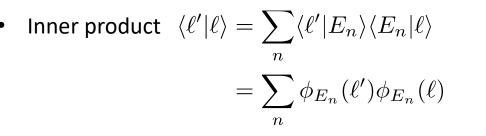
• Inner product $\langle \ell' | \ell \rangle = \sum_n \langle \ell' | E_n \rangle \langle E_n | \ell \rangle$

$$=\sum_{n}\phi_{E_{n}}(\ell')\phi_{E_{n}}(\ell)$$

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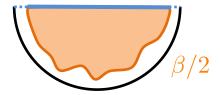
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• Ensemble average $\overline{\langle \ell' | \ell \rangle} = \int dE \ \overline{\rho(E)} \phi_E(\ell') \phi_E(\ell)$

$$\overline{\langle \ell' | \ell \rangle} = \delta(\ell - \ell') + \frac{\ell'}{\ell} + O(e^{-4S_0})$$

• Reproduces all Euclidean gravity results, resolves $\int d\ell \,
ho_g(\ell,\ell) = \infty$ puzzle



Construction of Hilbert space

• Construct null states for one member of the ensemble

$$\begin{split} |\psi\rangle &= \int d\ell \,\psi(\ell) |\ell\rangle \,, \quad \psi(\ell) \equiv \int dE \,\rho_0(E) \chi(E) \phi_E(\ell) \,, \\ \langle\psi|\psi\rangle &= \sum_n |\chi(E_n)|^2 \end{split}$$

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Reduction in dimension of Hilbert space to e^{S_0}

Hamiltonian is unmodified [Harlow-Jafferis]

$$H|\psi\rangle = \int d\ell \,\left(-\frac{1}{2}\partial_{\ell}^2 + 2e^{-\ell}\right)\psi(\ell)|\ell\rangle$$

Length operators

A family of length operators

$$\hat{\ell}_{\Delta} = -\frac{1}{\Delta} \log \left(\int d\ell \, e^{-\Delta \ell} |\ell\rangle \langle \ell| \right)$$

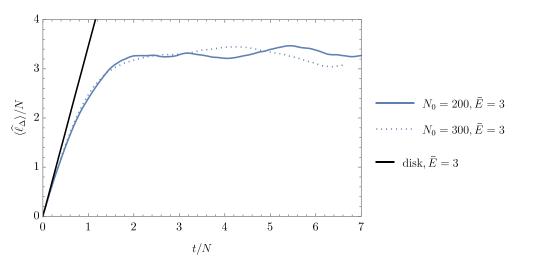
- If $|\ell\rangle$'s were all orthogonal $\hat{\ell}_{\Delta} = \int d\ell \, \ell |\ell\rangle \langle \ell |$
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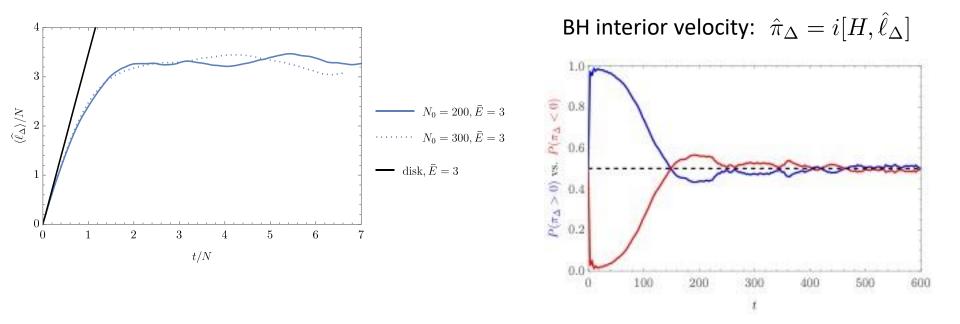


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Path integral and Hilbert space perspectives are complementary both conceptually and in computations

- In QFT they are related by cutting open the path integral In QG topology change hinders cutting
- Can absorb topology change into inner product on state space (at least in JT) Gives rise to plethora of null states
- Studied the volume of black hole interior from both the path integral and Hilbert space perspectives, cross-fertilisation with many-body quantum chaos

Summary

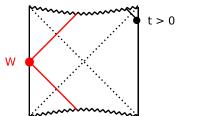
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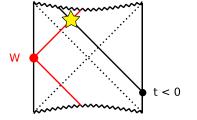
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Thank y

Outlook

- More analytic control in JT, extension to higher dimensions
- Experience of infalling observer? Firewall? [Stanford-Yang, Mathur, AMPS(S), Saad, Blommaert-Chen-Nomura]





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Appendix

Spectral complexity

Related quantity:
$$\langle \ell(t) \rangle \approx \frac{1}{Z(\beta)} \int dE \, d\omega \, \frac{\langle \rho(E+\omega/2)\rho(E-\omega/2) \rangle}{\omega^2 \langle \rho(E) \rangle} \, e^{-\beta E} \left[1 - \cos(\omega t)\right]$$

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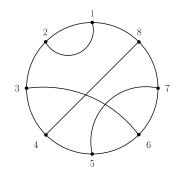
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$$C(t)$$

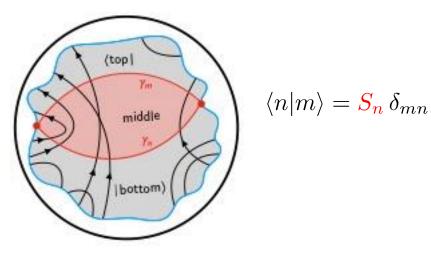
Hilbert space for DSSYK

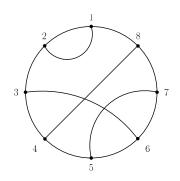
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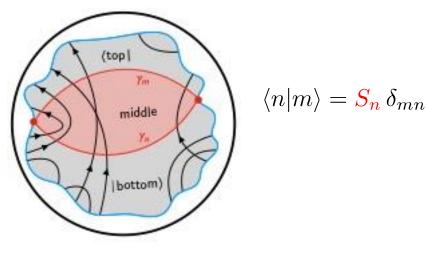
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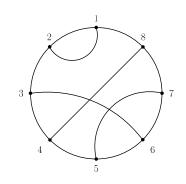




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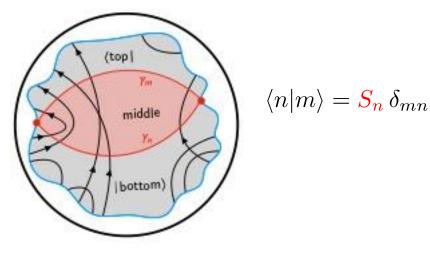


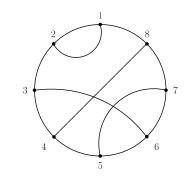


• $|m\rangle$ form a Krylov basis for $\{T^m|0\rangle\}$ and chord number is Krylov state complexity [Lin, Rabinovici-Sánchez-Garrido-Shir-Sonner]

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- Partition function is a transition amplitude $Z(\beta) = \langle 0|e^{-\beta T}|0\rangle$

• Limit $\lambda = \frac{p^2}{N} \to 0$, $\ell = \lambda m = \text{fixed gives JT at genus 0, length is Krylov complexity}$

Is there a genus expansion for DSSYK?