

Scattering amplitudes for Kerr black holes and higher-spin symmetry



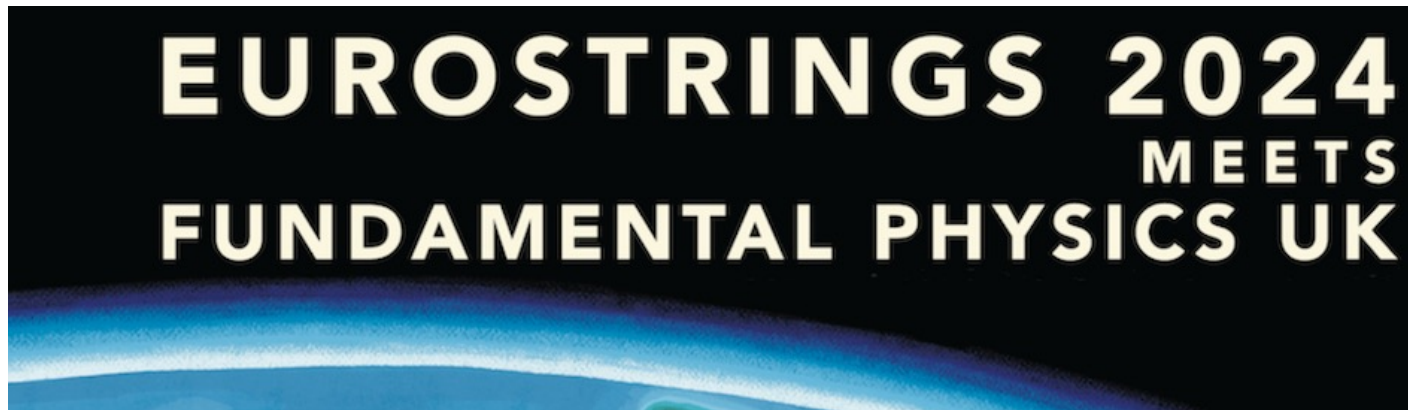
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Based on refs: Chiodaroli, HJ, Pichini [2107.14779];
Cangemi, Chiodaroli, HJ, Ochirov, Pichini, Skvortsov
[2212.06120], [2311.14668], [2312.14913]

Classical BH dynamics – PN, PM, spin

Post-Newtonian (PN) expansion:

Bound systems:



expand in G and v

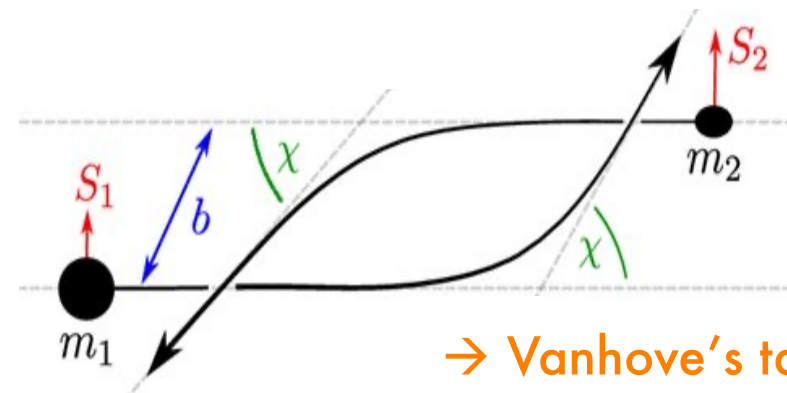
$$v^2 \sim \frac{GM}{r}$$

(virial theorem)

Post-Minkowskian (PM) expansion:

Gravitational scattering:

expand in $G \rightarrow$ loop expansion

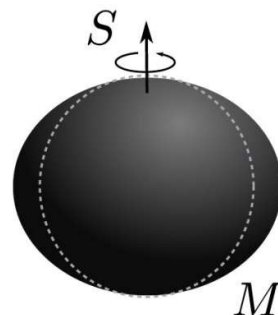


\rightarrow Vanhove's talk

Spin-multipole expansion:

Rotating black holes

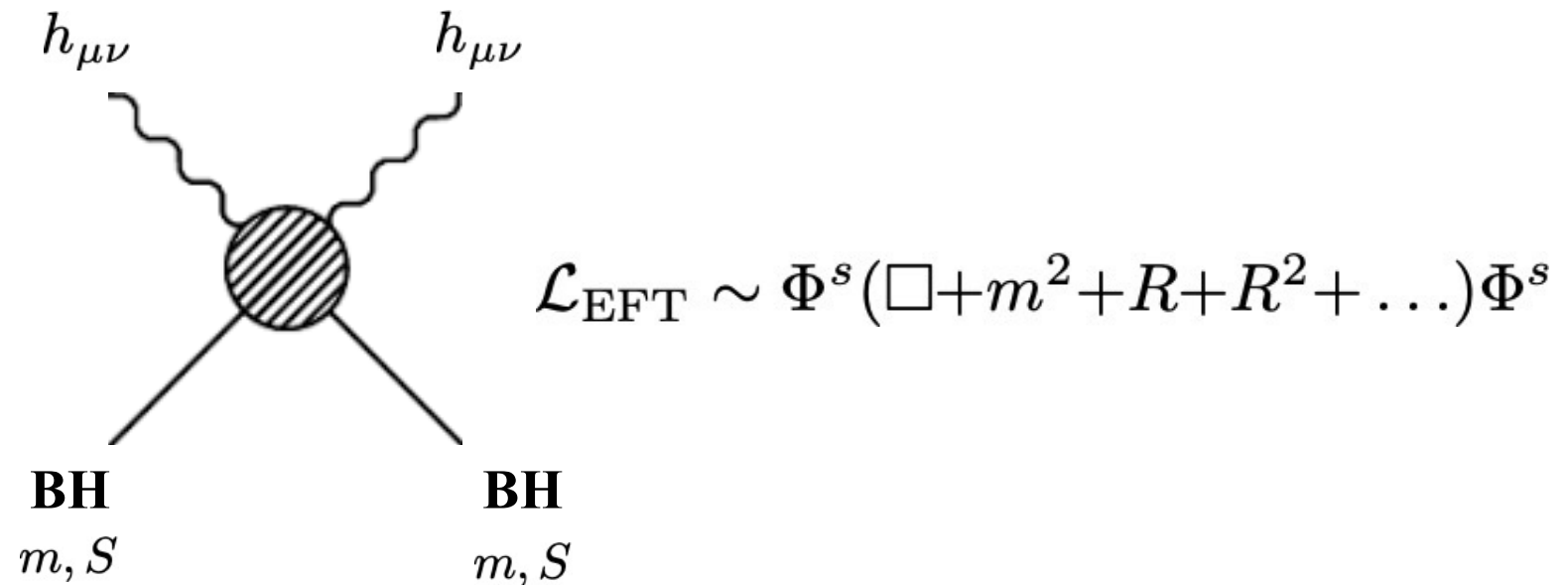
expand in S_1 and S_2



Methods

- BH perturbation theory
- Worldline EFTs
- Quantum scattering ampl's
- *higher-spin QFTs*

Kerr BH Compton scattering



- Eternal BHs = asymptotic states $(m, S, Q = 0)$
- Loops probe finite-size effects
(horizon, tidal effects, QNM, etc.) $r_S = 2Gm$
- Tree-level = superextremal Kerr $Gm \ll S/m$
- Point-particle approximation valid
- Compton \rightarrow BH dynamics \rightarrow BH EFT

Outline

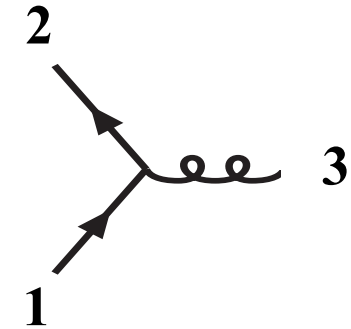
- **Motivation**
- **The AHH higher-spin amplitudes**
- **The problem of Compton scattering**
- **Higher-spin gauge symmetry and EFTs**
- **Chiral HS fields and Compton spin-s result**
- **Conclusion**

Higher-spin 3pt amplitudes & Kerr BH

Natural higher-spin gravitational 3pt amplitudes:

$$M(1\Phi^s, 2\Phi^s, 3h^+) = (\varepsilon_3^+ \cdot p_1)^2 \frac{\langle \mathbf{12} \rangle^{2s}}{m^{2s}}$$

$$M(1\Phi^s, 2\Phi^s, 3h^-) = (\varepsilon_3^- \cdot p_1)^2 \frac{[\mathbf{12}]^{2s}}{m^{2s}}$$



Arkani-Hamed, Huang, Huang ('17)

Linearized energy-momentum tensor for Kerr source

Vines ('17)

$$T^{\mu\nu}(-k) = 2\pi \delta(p \cdot k) p^{(\mu} \exp(m^{-1} S * ik)^{\nu)}{}_{\rho} p^{\rho}$$

Non-minimal worldline action for Kerr:

Levi, Steinhoff ('15)

$$L_{\text{SI}} = \sum_{n=1}^{\infty} \frac{(-1)^n}{(2n)!} \frac{C_{ES^{2n}}}{m^{2n-1}} D_{\mu_{2n}} \cdots D_{\mu_3} \frac{E_{\mu_1\mu_2}}{\sqrt{u^2}} S^{\mu_1} S^{\mu_2} \cdots S^{\mu_{2n-1}} S^{\mu_{2n}}$$

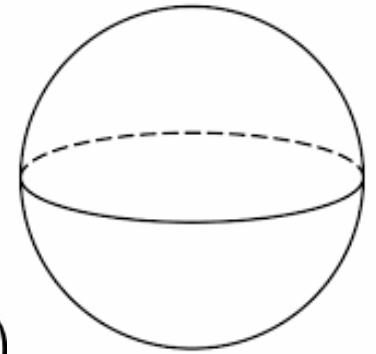
$$+ \sum_{n=1}^{\infty} \frac{(-1)^n}{(2n+1)!} \frac{C_{BS^{2n+1}}}{m^{2n}} D_{\mu_{2n+1}} \cdots D_{\mu_3} \frac{B_{\mu_1\mu_2}}{\sqrt{u^2}} S^{\mu_1} S^{\mu_2} \cdots S^{\mu_{2n-1}} S^{\mu_{2n}} S^{\mu_{2n+1}}$$

(spin-multipole expansion)

Spin operator (QM)

Introduce projective 3-sphere coordinates

$$z^a = (x_1 + ix_2, x_3 + ix_4) \rightarrow 1 = z^a \bar{z}_a = |x|^2$$



parametrizes $SU(2) \leftrightarrow$ spin wavefn $z^a \sim (|\uparrow\rangle, |\downarrow\rangle)$

Classical and quantum spin related as:

$$S^\mu = \frac{s}{2m} (\bar{z}^a z_a)^{2s-1} (\langle \bar{\mathbf{1}} | \sigma^\mu | \mathbf{1} \rangle + \langle \mathbf{1} | \sigma^\mu | \bar{\mathbf{1}} \rangle)$$

massive
spinor-helicity
formalism

Properties:

Transversality of spin vector: $p_1 \cdot S = 0$

Equals an expectation value: $S^\mu = \langle \hat{S}^\mu \rangle \equiv (\bar{z})^{2s} \cdot \hat{S}^\mu \cdot (z)^{2s}$

Gives spin operator: $[\hat{S}^\mu, \hat{S}^\nu] = i\epsilon^{\mu\nu\rho} \hat{S}_\rho \quad \hat{S}^2 = s(s+1)\mathbb{1}$

Massive spinor helicity

Following AHH bold massive spinors \leftrightarrow symmetrized little group indices

$$|\mathbf{i}\rangle \equiv |i^a\rangle z_{i,a}, \quad |\mathbf{i}] \equiv |i^a] z_{i,a} \quad \text{AHH}$$

(spinors define maps: $SL(2, \mathbb{C}) \rightarrow SU(2)$)

Analytic functions of spinors now possible:

$$\langle \mathbf{12} \rangle^{2s} = \text{degree-}4s \text{ polynomial in } (z_1^a, z_2^a)$$

Massive polarizations are null vectors Chiodaroli, HJ, Pichini

$$\epsilon_i^\mu = \frac{\langle \mathbf{i} | \sigma^\mu | \mathbf{i} \rangle}{\sqrt{2} m_i} = \frac{[\mathbf{i} | \bar{\sigma}^\mu | \mathbf{i} \rangle}{\sqrt{2} m_i} = (z_i^1)^2 \epsilon_{i,-}^\mu - \sqrt{2} z_i^1 z_i^2 \epsilon_{i,L}^\mu - (z_i^2)^2 \epsilon_{i,+}^\mu$$

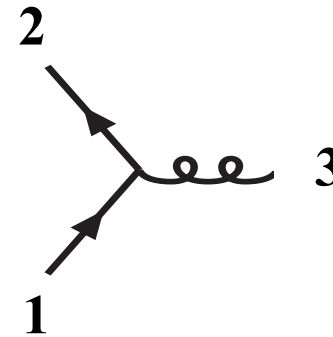
Higher-spin states automatically symmetric, transverse, traceless

$$\epsilon_i^{\mu_1 \mu_2 \dots \mu_s} \equiv \epsilon_i^{\mu_1} \epsilon_i^{\mu_2} \dots \epsilon_i^{\mu_s} = \text{degree-}2s \text{ polynomial in } z_i^a$$

AHH amplitudes = Kerr BHs

Relate in/out states by Lorentz transf.

$$|\mathbf{2}\rangle := |\bar{\mathbf{1}}\rangle + p_3 \cdot \sigma |\bar{\mathbf{1}}\rangle / (2m).$$



AHH factor \rightarrow exponential of spin operator:

$$\frac{\langle \mathbf{12} \rangle^{2s}}{m^{2s}} = \left\langle \sum_{n=0}^{2s} \frac{1}{n!} \left(\frac{p_3 \cdot \hat{S}}{m} \right)^n \right\rangle = \langle e^{p_3 \cdot \hat{a}} \rangle$$

Quantum Kerr and root Kerr 3pt \rightarrow Quantum Newman-Janis shift

$$M_{3,\pm}^{\text{Kerr}} = \langle e^{\pm p_3 \cdot \hat{a}} \rangle M_{3,\pm}^{\text{Schwarzchild}}$$

$$A_{3,\pm}^{\sqrt{\text{Kerr}}} = \langle e^{\pm p_3 \cdot \hat{a}} \rangle A_{3,\pm}^{\text{Coulomb}}$$

with ring-radius operator: $\hat{a}^\mu = \frac{\hat{S}^\mu}{m}$

(original argument: Guevara, Ochirov, Vines; see also Chung, Huang, Kim, Lee)

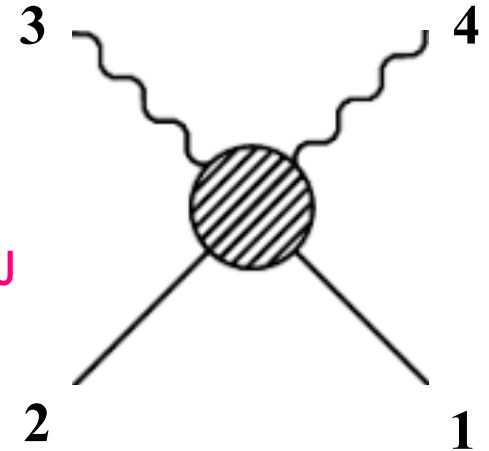
Kerr Compton amplitudes

Candidate Compton amplitudes via BCFW:

same helicity case:

$$M(1\phi^s, 2\bar{\phi}^s, 3h^+, 4h^+) = i \frac{\langle \mathbf{12} \rangle^{2s} [34]^4}{m^{2s-4} s_{12} t_{13} t_{14}}$$

Ochirov, HJ



opposite helicity case:

$$M(1\phi^s, 2\bar{\phi}^s, 3h^-, 4h^+) = i \frac{[4|p_1|3\rangle^{4-2s} ([41]\langle 32\rangle + [42]\langle 31\rangle)^{2s}}{s_{12} t_{13} t_{14}}, \quad \text{AHH}$$

Needed for NLO calculations:



However, for $s > 2$ there is a spurious pole \rightarrow need corrections

$$\frac{1}{[4|p_1|3\rangle^{2s-4}}$$

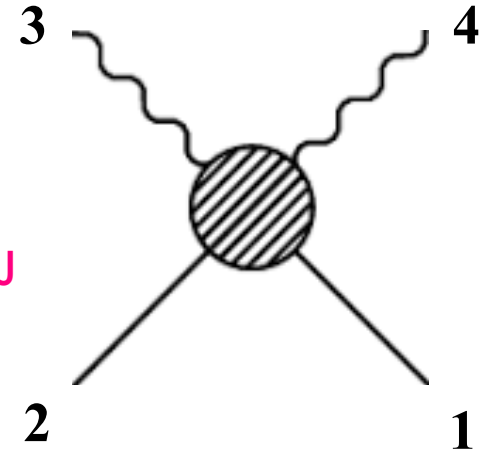
Gauge theory root-Kerr

Higher-spin gauge-theory: root-Kerr amplitudes

same helicity case:

$$A(1\phi^s, 2\bar{\phi}^s, 3A^+, 4A^+) = i \frac{\langle \mathbf{12} \rangle^{2s} [34]^2}{m^{2s-2} t_{13} t_{14}}$$

Ochirov, HJ



opposite helicity case:

$$A(1\phi^s, 2\bar{\phi}^s, 3A^-, 4A^+) = -i \frac{[4|p_1|3\rangle^{2-2s} ([41]\langle 32\rangle + [42]\langle 31\rangle)^{2s}}{t_{13} t_{14}} \quad \text{AHH}$$

Not needed for physics purposes, but provide useful toy model!

Again, for $s > 1$ spurious pole $\frac{1}{[4|p_1|3\rangle^{2s-2}} \rightarrow$ need corrections

Which quantum EFTs give Kerr amplitudes ?

EFTs behind root-Kerr

Identify EFTs from covariant formulas:

$$A(1\phi^s, 2\bar{\phi}^s, 3A^+) = mx \frac{\langle \mathbf{12} \rangle^{2s}}{m^{2s}}$$

spin-0: $A(1\phi^0, 2\bar{\phi}^0, 3A) = \varepsilon_3 \cdot (p_1 - p_2) \equiv A_{\phi\phi A}$ (scalar)

spin-1/2: $A(1\phi^{1/2}, 2\bar{\phi}^{1/2}, 3A) = \bar{u}_2 \not{\varepsilon}_3 u_1 \equiv A_{\lambda\lambda A}$ (fermion)

spin-1: $A(1\phi^1, 2\bar{\phi}^1, 3A) = 2(\varepsilon_1 \cdot \varepsilon_2 \varepsilon_3 \cdot p_2 + \varepsilon_2 \cdot \varepsilon_3 \varepsilon_1 \cdot p_3 + \varepsilon_3 \cdot \varepsilon_1 \varepsilon_2 \cdot p_1)$
 $\equiv A_{WWA}$ (W-boson)

spin-3/2: $A(1\phi^{3/2}, 2\bar{\phi}^{3/2}, 3A) = \bar{u}_2^\mu \not{\varepsilon}_3 u_{1\mu} - \frac{2}{m} \bar{u}_{2\mu} f_3^{\mu\nu} u_{1\nu} - \frac{1}{2m} \bar{u}_2^\mu f_3^{\rho\sigma} \gamma_\rho \gamma_\sigma u_{1\mu} \equiv A_{\psi\psi A}$
 (gravitino)

general spin- s given as a generating function:

$$\sum_{s=0}^{\infty} A(1\phi^s, 2\bar{\phi}^s, 3A) = A_{\phi\phi A} + \frac{A_{WWA} - (\varepsilon_1 \cdot \varepsilon_2)^2 A_{\phi\phi A}}{(1 + \varepsilon_1 \cdot \varepsilon_2)^2 + \frac{2}{m^2} \varepsilon_1 \cdot p_2 \varepsilon_2 \cdot p_1}$$

Chiodaroli,
HJ, Pichini

For $s > 1 \rightarrow$ higher-derivative HS effective theories (no massless limit)

Kerr/root-Kerr double copy

Kerr amplitudes related to gauge th. via double copy

Chiodaroli,
HJ, Pichini

$$M(1\phi^s, 2\bar{\phi}^s, 3h^\pm) = iA(1\phi^{s_L}, 2\bar{\phi}^{s_L}, 3A^\pm)A(1\phi^{s_R}, 2\bar{\phi}^{s_R}, 3A^\pm)$$

The general spin- s 3pt amplitude \rightarrow generating fn

$$\sum_{2s=0}^{\infty} M(1\phi^s, 2\bar{\phi}^s, 3h) = M_{0\oplus 1/2} + A_{WWA} \left(A_{0\oplus 1/2} + \frac{A_{1\oplus 3/2} - (\epsilon_1 \cdot \epsilon_2)^2 A_{0\oplus 1/2}}{(1 + \epsilon_1 \cdot \epsilon_2)^2 + \frac{2}{m^2} \epsilon_1 \cdot p_2 \epsilon_2 \cdot p_1} \right)$$

From double-copy structure, we can infer:

EFTs	$s = 1/2$	$s = 1$	$s = 3/2$	$s = 2$	$s = 5/2$	$s \geq 3$
Kerr	Major.	Proca	Rar.-Sch.	KK grav.	HS	HS
$\sqrt{\text{Kerr}}$	Dirac	W -boson	gravitino	HS	HS	HS

Cangemi,
Chiodaroli,HJ,
Ochirov,
Pichini,
Skvortsov

For $s > 2$ Kerr \rightarrow higher-derivative HS EFTs (no massless limit)

Higher-spin (HS) theories

What special about the low-spin EFTs ?

Kerr (root-Kerr) EFTs for $s \leq 2$ ($s \leq 1$)

Chiodaroli,
HJ, Pichini

→ well-behaved massless limit

→ exhibits gauge symmetry (SSB)

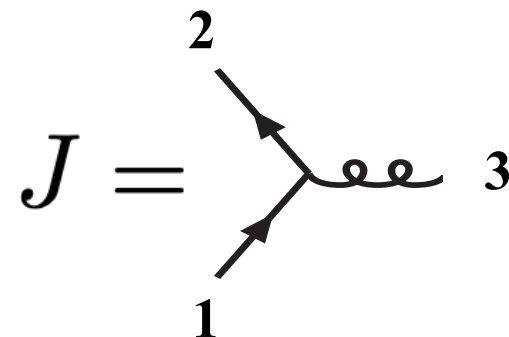
$s = 1$ (YM + W-boson) → non-abelian gauge symmetry

$s = 3/2$ (GR + massive gravitino) → supersymmetry

$s = 2$ (GR + massive KK graviton) → General covariance

Furthermore: satisfy a current constraint

$$p_1 \cdot J = \mathcal{O}(m)$$



Connected to tree-level unitarity constraint;

Porrati et al.

longitudinal modes suppressed in low-mass (high-energy) limit

Using HS gauge invariance

Cangemi, Chiodaroli, HJ,
Ochirov, Pichini, Skvortsov

Consider spin-2 root-Kerr case:

physical field: $\Phi_{\mu\nu}$

Stückelberg fields: $\{B_\mu, \varphi\}$

Imposing a linearized massive higher-spin gauge transformation:

$$\begin{aligned}\delta\Phi_{\mu\nu} &= \frac{1}{2}\partial_\mu\xi_\nu + \frac{1}{2}\partial_\nu\xi_\mu + \frac{m}{\sqrt{2}}\eta_{\mu\nu}\xi, \\ \delta B_\mu &= \partial_\mu\xi + \frac{m}{\sqrt{2}}\xi_\mu, \\ \delta\varphi &= \sqrt{3}m\xi,\end{aligned}$$

← gauge parameter

Makes sure that:

→ DOFs are correct,

→ small-mass limit better behaved than naively expected

Massive Ward identities

We write down ansatz for off-shell interactions:

Cangemi, Chiodaroli, HJ,
Ochirov, Pichini, Skvortsov

$$V_{\Phi\bar{\Phi}A} \sim m (\epsilon_1)^2 (\epsilon_2)^2 \epsilon_3 \left(\frac{p^3}{m^3} + \frac{p}{m} \right),$$

$$V_{B\bar{\Phi}A} \sim m (\epsilon_1) (\epsilon_2)^2 \epsilon_3 \left(\frac{p^2}{m^2} + 1 \right),$$

$$V_{\varphi\bar{\Phi}A} \sim m (\epsilon_2)^2 \epsilon_3 \left(\frac{p}{m} \right),$$

and constrain them using Ward identities

$$V_{\xi\bar{\Phi}A}|_{(2,3)} = V_{\zeta\bar{\Phi}A}|_{(2,3)} = 0$$

where the vertices corresponding to gauge parameters are:

$$V_{\xi\bar{\Phi}A} := \frac{m}{\sqrt{2}} V_{B\bar{\Phi}A} - \frac{i}{2} p_1 \cdot \frac{\partial}{\partial \epsilon_1} V_{\Phi\bar{\Phi}A},$$

$$V_{\zeta\bar{\Phi}A} := \sqrt{3} m V_{\varphi\bar{\Phi}A} - i p_1 \cdot \frac{\partial}{\partial \epsilon_1} V_{B\bar{\Phi}A} + \frac{m}{2\sqrt{2}} \left(\frac{\partial}{\partial \epsilon_1} \right)^2 V_{\Phi\bar{\Phi}A}.$$

→ 3pt amplitude: $A(\Phi_1^2 \bar{\Phi}_2^2 A_3^+) = A_0 \frac{\langle \mathbf{12} \rangle^3}{m^4} (c_1 [\mathbf{12}] + (1 - c_1) \langle \mathbf{12} \rangle)$

unique after current constraint: $c_1 = 0$

General spin-s EFTs

Consider tower $k = 0, 1, 2, \dots, s$ of HS fields and gauge parameters:

$$\Phi^k := \Phi^{\mu_1 \mu_2 \dots \mu_k}, \quad \xi^k := \xi^{\mu_1 \mu_2 \dots \mu_k} \quad \text{Zinoviev (2001)}$$

(double-traceless) (traceless)

Gauge transformation: $\delta\Phi^k = \partial^{(1}\xi^{k-1)} + m\alpha_k\xi^k + m\beta_k\eta^{(2}\xi^{k-2)}$

$$\alpha_k = \frac{1}{k+1} \sqrt{\frac{(s-k)(s+k+1)}{2}}, \quad \beta_k = \frac{1}{2} \frac{k}{k-1} \alpha_{k-1}$$

Minimal Lagrangian:

$$\mathcal{L}_0 = \mathcal{L}_F + \frac{1}{2} \sum_{k=0}^{s-1} (-1)^k (k+1) G^k G^k$$

Gauge-fixing fn:

$$G^k = \partial \cdot \Phi^{k+1} - \frac{k}{2} \partial^{(1} \tilde{\Phi}^{k+1)} + m (\alpha_k \Phi^k - \gamma_k \tilde{\Phi}^{k+2} - \delta_k \eta^{(2} \tilde{\Phi}^k))$$

Feynman-gauge Lagr:

$$\mathcal{L}_F = \sum_{k=0}^s \frac{(-1)^k}{2} \left[\Phi^k (\square + m^2) \Phi^k - \frac{k(k-1)}{4} \tilde{\Phi}^k (\square + m^2) \tilde{\Phi}^k \right]$$

Cangemi, Chiodaroli, HJ,
Ochirov, Pichini, Skvortsov

Non-minimal interactions

Cangemi, Chiodaroli, HJ,
Ochirov, Pichini, Skvortsov

3pt vertex: $V_{\Phi^k \Phi^s A^{\mathfrak{h}}} = V_{\Phi^k \Phi^s A^{\mathfrak{h}}}^{\text{min.}} + V_{\Phi^k \Phi^s A^{\mathfrak{h}}}^{\text{non-min.}}$

Ward identities:
$$V_{\xi^k \Phi^s A^{\mathfrak{h}}} := m\alpha_k V_{\Phi^k \Phi^s A^{\mathfrak{h}}} - \frac{i}{k+1} p_1 \cdot \frac{\partial}{\partial \epsilon_1} V_{\Phi^{k+1} \Phi^s A^{\mathfrak{h}}} + \frac{m\beta_{k+2}}{(k+2)(k+1)} \frac{\partial}{\partial \epsilon_1} \cdot \frac{\partial}{\partial \epsilon_1} V_{\Phi^{k+2} \Phi^s A^{\mathfrak{h}}}$$

Constraints imposed:

(WI) Ward identities $V_{\xi^k \Phi^s A^{\mathfrak{h}}} \Big|_{(2,3), \epsilon_1^2 \rightarrow 0} = 0;$

(CC) Current constraint $p_1 \cdot \frac{\partial}{\partial \epsilon_1} V_{\Phi^s \Phi^s A^{\mathfrak{h}}} \Big|_{(2,3), \epsilon_1^2 \rightarrow 0} = \mathcal{O}(m).$

(PC) Power-counting bound on derivatives in non-minimal vertices: $V_{\Phi^{s_1} \Phi^{s_2} A^{\mathfrak{h}}}^{\text{non-min.}} \sim \partial^{s_1+s_2-2\mathfrak{h}} (F_{\mu\nu})^{\mathfrak{h}};$

(ND) Near-diagonal interactions: if $|s_1 - s_2| > \mathfrak{h}$ then $V_{\Phi^{s_1} \Phi^{s_2} A^{\mathfrak{h}}} = 0.$

Gives unique Kerr and root-Kerr 3pt amplitudes (matching AHH)

HS perturbation theory

Calculations expected to simplify in Feynman gauge: Cangemi, Chiodaroli, HJ, Ochirov, Pichini, Skvortsov

Feynman-gauge propagator for any field obtained as generating fn:

$$\Delta(\epsilon, \bar{\epsilon}) = \sum_{s=0}^{\infty} (\epsilon)^s \cdot \Delta^{(s)} \cdot (\bar{\epsilon})^s = \frac{1}{p^2 - m^2 + i0} \frac{1 - \frac{1}{4}\epsilon^2 \bar{\epsilon}^2}{1 + \epsilon \cdot \bar{\epsilon} + \frac{1}{4}\epsilon^2 \bar{\epsilon}^2}$$

e.g. for root-Kerr Compton amplitude, we obtain

$$A(\Phi_1^s \Phi_2^s A_3^- A_4^+) = \frac{\langle 3|1|4 \rangle^2 (U + V)^{2s}}{m^{4s} t_{13} t_{14}} + \frac{\langle 3|1|4 \rangle \langle \mathbf{13} \rangle [\mathbf{24}] P^{(2s)}}{m^{4s} t_{13}} + \langle \mathbf{13} \rangle \langle \mathbf{32} \rangle [\mathbf{14}] [\mathbf{42}] \frac{P^{(2s-1)}}{m^{4s}} + C_s,$$

contact term
 $C_{s < 2} = 0$

with a polynomial:

$$P^{(k)} = \frac{\varsigma_1^k - \varsigma_2^k}{\varsigma_1 - \varsigma_2}$$

and variables

$$\varsigma_1 = \langle \mathbf{1} | 1+4 | \mathbf{2} \rangle, \quad \varsigma_2 = \langle \mathbf{2} | 2+3 | \mathbf{1} \rangle$$

Chiral fields $(2s,0)$

Chiral higher-spin approach

Easier way to get correct DOFs:

Ochirov, Skvortsov;
Cangemi, et al.

Change Lorentz rep. $(s, s) \longrightarrow (2s, 0)$

Chiral fields $|\Phi\rangle := \Phi_{\alpha_1 \dots \alpha_{2s}}$ $SL(2, \mathbb{C})$ indices

Minimal Lagrangian $\mathcal{L}_{\min.}^{(s)} = \langle D_\mu \Phi | D^\mu \Phi \rangle - m^2 \langle \Phi | \Phi \rangle$

Gives "correct" all-plus helicity amplitudes:

$$A_n(1^s, 2^s, 3^+, 4^+, \dots, n^+) = \langle \mathbf{12} \rangle^{2s} A_n^{\text{scalar}}$$

However, breaks parity badly, and also naive renormalizability...

W-bosons in SM: $\mathcal{L}^{(1)} = \langle \Phi | \left\{ |\overleftarrow{D}| \overrightarrow{D} | \otimes \frac{1}{1 - \frac{ig}{m^2} |F_-|} \right\} | \Phi \rangle - m^2 \langle \Phi | \Phi \rangle + \mathcal{O}(\Phi^4)$

Chalmers, Siegel

Non-minimal chiral interactions

Cangemi, Chiodaroli, HJ,
Ochirov, Pichini, Skvortsov

Restore parity at 3pts \rightarrow AHH 3pt amplitudes:

Root-Kerr non-minimal interactions:

$$\mathcal{L}^{(s)} = \langle D_\mu \Phi | D^\mu \Phi \rangle - m^2 \langle \Phi | \Phi \rangle + \sum_{k=0}^{2s-1} \frac{ig}{m^{2k}} \langle \Phi | \left\{ |\overleftarrow{D} \overrightarrow{D}|^{\odot k} \otimes |F_-| \right\} | \Phi \rangle + \mathcal{O}(F^2)$$

Kerr non-minimal interactions:

$$\mathcal{L}_{\text{Kerr}} = \sqrt{-g} \left\{ \frac{1}{2} \langle \nabla_\mu \Phi | \nabla^\mu \Phi \rangle - \frac{m^2}{2} \langle \Phi | \Phi \rangle - \frac{1}{4} \sum_{k=0}^{2s-2} \frac{2s-k-1}{m^{2k}} \langle \Phi | \left\{ (|\overleftarrow{\nabla} \overrightarrow{\nabla}|)^{\odot k} \odot |R_-| \right\} | \Phi \rangle \right\} + \mathcal{O}(R^2)$$

Interactions behave as geometric series $\sim \frac{1}{1 - |\overleftarrow{D} \overrightarrow{D}|} \odot |F_-|$

Omnipresent polynomials

Consider geometric sums:

$$\begin{array}{c}
 \begin{array}{ccccccc}
 \bullet & \xrightarrow{\quad} & \bullet & \xrightarrow{\quad} & \bullet & \xrightarrow{\quad} & \bullet & \xrightarrow{\quad} & \bullet \\
 \varsigma_1^4 & & \varsigma_1^3 \varsigma_2 & & \varsigma_1^2 \varsigma_2^2 & & \varsigma_1 \varsigma_2^3 & & \varsigma_2^4
 \end{array} \\
 = \frac{\varsigma_1^5 - \varsigma_2^5}{\varsigma_1 - \varsigma_2}
 \end{array}$$

$$= \frac{\varsigma_1^6}{(\varsigma_1 - \varsigma_2)(\varsigma_1 - \varsigma_3)} + \text{perm}(\varsigma_1, \varsigma_2, \varsigma_3)$$

In general:

Complete homogenous symmetric polynomials:

Cangemi, Chiodaroli, HJ,
Ochirov, Pichini, Skvortsov

$$P_n^{(k)} = \frac{\varsigma_1^k}{(\varsigma_1 - \varsigma_2)(\varsigma_1 - \varsigma_3) \dots (\varsigma_1 - \varsigma_n)} + \text{perm}(\varsigma_1, \varsigma_2, \dots, \varsigma_n)$$

Compton spin variables:

$$\varsigma_1 = \langle \mathbf{1} | 1+4 | \mathbf{2} \rangle, \quad \varsigma_2 = \langle \mathbf{2} | 2+3 | \mathbf{1} \rangle, \quad \varsigma_3 = m \langle \mathbf{21} \rangle, \quad \varsigma_4 = m [\mathbf{21}]$$

Constraints for fixing R^2 contact term

Assumptions: contact terms depend only on $C^{(s)} = C^{(s)}[P_n^{(k)}]$

- well-behaved classical limit $s \rightarrow \infty$;
- compatible with massive higher-spin gauge invariance;
- s -independent numerical coefficients;
- parity invariance
- all contact terms have spinor-helicity structure $\sim (\langle \mathbf{13} \rangle \langle \mathbf{32} \rangle [\mathbf{14}] [\mathbf{42}])^2$
- classical spin hexadecapole S^4 is fixed by $s = 2$ amplitude
- improved behavior in $m \rightarrow 0$ limit: $M(1^s, 2^s, 3^-, 4^+) \sim m^{-4s+4}$

Kerr amplitude from chiral fields + contact

Cangemi, Chiodaroli, HJ,
Ochirov, Pichini, Skvortsov

Final Kerr Compton amplitude (quantum spin):

$$\begin{aligned} M(\mathbf{1}^s, \mathbf{2}^s, \mathbf{3}^-, \mathbf{4}^+) &= \frac{\langle 3|1|4 \rangle^4 P_1^{(2s)}}{m^{4s} s_{12} t_{13} t_{14}} - \frac{\langle \mathbf{13} \rangle [\mathbf{42}] \langle 3|1|4 \rangle^3}{m^{4s} s_{12} t_{13}} P_2^{(2s)} + \frac{\langle \mathbf{13} \rangle \langle \mathbf{32} \rangle [\mathbf{14}] [\mathbf{42}]}{m^{4s} s_{12}} (\langle 3|1|4 \rangle^2 P_2^{(2s-1)} + m^4 \langle 3|\rho|4 \rangle^2 P_4^{(2s-1)}) \\ &+ \frac{\langle \mathbf{13} \rangle \langle \mathbf{32} \rangle [\mathbf{14}] [\mathbf{42}]}{m^{4s-2} s_{12}} \langle 3|1|4 \rangle \langle 3|\rho|4 \rangle (P_2^{(2s-2)} - m^2 \langle \mathbf{12} \rangle [\mathbf{12}] P_4^{(2s-2)}) \\ &+ \frac{\langle \mathbf{13} \rangle^2 \langle \mathbf{32} \rangle^2 [\mathbf{14}]^2 [\mathbf{42}]^2}{2m^{4s-4}} \langle \mathbf{12} \rangle [\mathbf{12}] \left[(1 + \eta) P_{5|\zeta_1}^{(2s-2)} + (1 - \eta) P_{5|\zeta_2}^{(2s-2)} \right] + \alpha C_\alpha^{(s)}. \end{aligned}$$

Includes some dissipative effects after matching to BHPT

Bautista, Guevara,
Kavanagh, Vines, et al

Does not include: near-zone contributions or loop corrections

(similar expression for root-Kerr gauge theory)

Root-Kerr Lagrangian and classical amplitude

Chiral spin-s Lagrangian (gauge theory)

Cangemi, Chiodaroli, HJ,
Ochirov, Pichini, Skvortsov

$$\begin{aligned} \mathcal{L} = & \langle D_\mu \Phi | D^\mu \Phi \rangle - m^2 \langle \Phi | \Phi \rangle + \sum_{k=0}^{2s-1} \frac{ig}{m^{2k}} \langle \Phi | \left\{ |\overleftarrow{D}| \overrightarrow{D} |^{\odot k} \odot |F_-| \right\} | \Phi \rangle + \mathcal{O}(|F_-|^2) \\ & + \sum_{k \leq l=0}^{2s-4} \sum_{j=0}^{2s-3-l} \frac{g^2}{m^{2(j+l)+6}} \langle \Phi | \left\{ (|\overleftarrow{D}| \overrightarrow{D} | + m^2) \odot |\overleftarrow{D}| \overrightarrow{D} |^{\odot j} \odot |\overleftarrow{D}| \overrightarrow{D}_+ |^{\odot k} \odot |\overleftarrow{D}_+| \overrightarrow{D} |^{\odot (l-k)} \odot \mathfrak{F}_6 \right\} | \Phi \rangle. \end{aligned}$$

Field-strength dependence: $\mathfrak{F}_6 = \frac{1}{4} \{T^c, T^{c'}\} |F_-^c| \odot |\overleftarrow{D}| F_+^{c'} | \overrightarrow{D}$

Classical root-Kerr amplitude: $\lim_{s \rightarrow \infty, \hbar \rightarrow 0} \mathcal{A}(1, 2, 3^-, 4^+)$

$$\begin{aligned} = & -2g^2 (p \cdot \chi)^2 \left\{ \left(\frac{[T^{c_3}, T^{c_4}]}{q^2 (p \cdot q_\perp)} + \frac{1}{2} \frac{\{T^{c_3}, T^{c_4}\}}{(p \cdot q_\perp)^2} \right) \left(e^x \cosh z - w e^x \operatorname{sinhc} z + \frac{w^2 - z^2}{2} E(x, y, z) \right) \right. \\ & \left. - \frac{[T^{c_3}, T^{c_4}]}{q^2 (p \cdot q_\perp)} \left(x(w^2 + z^2) - w(x^2 - y^2 + z^2) \right) \tilde{E}(x, y, z) \right\}. \end{aligned}$$

$$E(x, y, z) = \frac{e^y - e^x \cosh z + (x - y) e^x \operatorname{sinhc} z}{(x - y)^2 - z^2} + (y \rightarrow -y)$$

$$x = a \cdot q_\perp,$$

$$y = a \cdot q,$$

$$z = |a| \frac{p \cdot q_\perp}{m},$$

$$w = \frac{a \cdot \chi p \cdot q_\perp}{p \cdot \chi}$$

Final classical results – Kerr BH

Classical root-Kerr amplitude: $\lim_{s \rightarrow \infty, \hbar \rightarrow 0} \mathcal{A}(1, 2, 3^-, 4^+)$

$$= -2g^2(p \cdot \chi)^2 \left\{ \left(\frac{[T^{c3}, T^{c4}]}{q^2(p \cdot q_\perp)} + \frac{1}{2} \frac{\{T^{c3}, T^{c4}\}}{(p \cdot q_\perp)^2} \right) \left(e^x \cosh z - w e^x \operatorname{sinhc} z + \frac{w^2 - z^2}{2} E(x, y, z) \right) - \frac{[T^{c3}, T^{c4}]}{q^2(p \cdot q_\perp)} \left(x(w^2 + z^2) - w(x^2 - y^2 + z^2) \right) \tilde{E}(x, y, z) \right\}.$$

Classical Kerr BH amplitude: Cangemi, Chiodaroli, HJ, Ochirov, Pichini, Skvortsov

$$\begin{aligned} \mathcal{M}(1, 2, 3^-, 4^+) &= \frac{(p \cdot \chi)^4}{q^2(p \cdot q_\perp)^2} \left(e^x \cosh z - w e^x \operatorname{sinhc} z + \frac{w^2 - z^2}{2} E(x, y, z) \right) \\ &+ \frac{(p \cdot \chi)^4}{q^2(p \cdot q_\perp)^2} \frac{w^2 - z^2}{2} (w - x) \tilde{E}(x, y, z) \\ &- \frac{(p \cdot \chi)^4}{(p \cdot q_\perp)^4} \frac{(w^2 - z^2)^2}{2} \left(\frac{\partial \tilde{E}}{\partial x} + \eta \frac{\partial \tilde{E}}{\partial z} \right) \quad + \alpha z \text{ (polygamma terms)} \end{aligned}$$

Matches explicit BH perturbation theory from GR \rightarrow Teukolsky eqn.

up to spin S^7 (ignoring polygamma terms) Bautista, Guevara, Kavanagh, Vines

Conclusion: Kerr dynamics from HS

Kerr dynamics is non-trivially constrained by

- massive higher-spin gauge symmetry
- power counting, current constraint, ...

Checks: → uniquely predicts previously known Kerr 3-4pt amplitudes
→ constrains $s \geq 2$ 4pt contact terms, but not unique...

Additional constraints imposed:

- chiral Lagrangian,
- symmetric homogeneous polynomials,...
- classical limit consistency,
- matching to Teukolsky BHPT (mod. polygamma terms)

Outlook: → classical loop corrections to Compton
→ implications for quantum BHs,
→ including absorption and emission effects