

Extended phase space for subⁿ-leading soft theorems

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Why symmetries ?

asymptotic symmetries $\xrightarrow[\text{id.}]{\text{Ward}}$ soft theorems

- Important ingredient for flat space holography

Why symmetries ?

asymptotic symmetries $\xrightarrow[\text{id.}]{\text{Ward}}$ **subleading/loop corrections for** soft theorems

[Campliglia,Ladha,Strominger,Lysov,Pasterski,Peraza,Donnelly,Freidel,Speranza,Geiller,Pranzetti,Ciambelli,Leigh,Pai,Oliveri,Speziale,Raclariu,Zwikel,Sahoo,Sen,Krishna,Pasterski,Donnay,Nguyen,Ruzziconi,Agrawal,Choi,Puhm,Bhatkar,Bianchi,He,Huang,Wen,Mitra,Conde,Mao,Wu,Bern,Davies,Di Vecchia,Nohle.....]

- Important ingredient for flat space holography

Strange symmetries

asymptotic symmetries $\xrightarrow[\text{id.}]{\text{Ward}}$ **subleading** soft theorems

$$\lim_{\omega \rightarrow 0} A_{n+1} = \underbrace{\left(\frac{1}{\omega} S^{(0)} + S^{(1)} + \dots \right) A_n}_{\text{universal}} + \underbrace{\dots}_{\text{non-universal}}$$

- Example: subleading order gauge theory extracted via

$$\lim_{\omega \rightarrow 0} (1 + \omega \partial_\omega) A_{n+1} = S^{(1)} A_n$$

was shown to arise as a Ward identity for an overleading gauge parameter [Campiglia, Laddha]

$$\Lambda(r, u, z, \bar{z}) = r \Lambda_1(u, z, \bar{z}) + \Lambda_0(u, z, \bar{z})$$

- Violates the fall-off of the fields

$$A_z(r, u, z, \bar{z}) = A_z^{(0)}(u, z, \bar{z}) + \dots$$

asymptotic symmetries $\xrightarrow[\text{id.}]{\text{Ward}}$ **subleading** soft theorems

$$\lim_{\omega \rightarrow 0} A_{n+1} = \underbrace{\left(\frac{1}{\omega} S^{(0)} + S^{(1)} + \dots \right) A_n}_{\text{universal}} + \underbrace{\dots}_{\text{non-universal}}$$

- Goal 1: extend phase space such that symmetries act canonically (at all orders)
- Goal 2: consequences for soft theorems

- Work in Yang-Mills

$$D^\mu \mathcal{F}_{\mu\nu} = \mathcal{J}_\nu (= 0)$$

- Arbitrary gauge choice, coordinates $x = (\tau, \vec{y})$, usually Bondi (r, u, z, \bar{z}) .
- Very general expansion for

$$\mathcal{A}_\mu = \sum_{n,k} A_\mu^{(-n;k)}(\vec{y}) \frac{\log^k \tau}{\tau^n},$$

with n, k such that $\lim_{r \rightarrow \infty} \mathcal{A}_\mu$ at most $\mathcal{O}(1)$.

- Phase space

$$\Gamma^0 = \left\{ \mathfrak{A}^0 \mid \text{with } \mathfrak{A}^0 \text{ constrained by e.o.m. and gauge choice} \right\}$$

Divergent gauge parameter

- Standard large gauge transformations give leading order soft theorems

$$\Lambda^{(0)} = \Lambda^{(0)}(\vec{y})$$

- General divergent gauge parameter

$$\Lambda_+(x) = \sum_{n,k} \tau^n \log^k \tau \Lambda^{(n;k)}(\vec{y}) ,$$

- May also be field dependent $\check{\Lambda}_+ = \check{\Lambda}_+(\mathcal{A}_\mu(x), \Lambda_+(x))$
- Violates gauge field fall-off
- Need **new objects** on which these symmetries act canonically.
- Helps to think of problem as a symmetry breaking...

Stueckelberg trick

- Stueckelberg trick: originally introduced to restore *broken local symmetry* in e.g. massive theories
- Promote gauge parameter of symmetry we want to restore to a field

$$\Lambda_+(x) \rightarrow \Psi(x),$$

with

$$\Psi(x) = \sum_{n,k} \tau^n \log^k \tau \Psi^{(n;k)}(\vec{y}).$$

- the phase space is now

$$\Gamma_\infty^{\text{ext}} := \Gamma^0 \times \{\Psi(x)\}$$

- Ψ is Goldstone-type field

Consistency relation

- Ψ comes from the bulk
- Dressed gauge field

$$\tilde{\mathcal{A}}_\mu = e^{i\Psi} \mathcal{A}_\mu e^{-i\Psi} + ie^{i\Psi} \partial_\mu e^{-i\Psi}$$

- Consistency condition:

$$\delta_\Lambda \tilde{\mathcal{A}}_\mu = \tilde{D}_\mu \Lambda$$

where

$$\Lambda = \Lambda^{(0)} + \Lambda_+$$

- The transformation of \mathcal{A}_μ is unchanged, and the overleading transformations only transform Ψ .

Stueckelberg field transformation

- The Stueckelberg (Goldstone) field transforms as

$$\delta_\Lambda \Psi = \mathcal{O}_{-i\Psi}^{-1} (\Lambda - e^{i\Psi} \Lambda^{(0)} e^{-i\Psi})$$

with

$$\mathcal{O}_X := \frac{1 - e^{-ad_X}}{ad_X} = \sum_{k=0}^{\infty} \frac{(-1)^k}{(k+1)!} (ad_X)^k$$

and its inverse

$$(\mathcal{O}_X)^{-1} = \left(\frac{1 - e^{-ad_X}}{ad_X} \right)^{-1} = \sum_{m=0}^{\infty} \frac{B_m^+ ad_X^m}{m!}$$

where B_n^+ are the Bernoulli numbers

$$B_0^+ = 1, \quad B_1^+ = \frac{1}{2}, \quad B_2^+ = \frac{1}{6}, \quad \dots$$

Stueckelberg field transformation

- The Stueckelberg (Goldstone) field transforms as

$$\delta_{\Lambda} \Psi = \mathcal{O}_{-i\Psi}^{-1} (\Lambda - e^{i\Psi} \Lambda^{(0)} e^{-i\Psi})$$

- Perturbatively

$$\delta_{\Lambda}^{[m]} \Psi = \frac{B_m^+}{m!} (ad_{-i\Psi})^m [\Lambda + (-1 + 2\delta_{m,1}) \Lambda^{(0)}]$$

- at 0th order in the field, it transforms via a shift

$$\delta_{\Lambda}^{[0]} \Psi = \Lambda - \Lambda^{(0)} = \Lambda_+$$

i.e. Goldstone modes for the symmetry breaking in the bulk.

- recall that

$$B_{2k+1} = 0, \quad \text{for } k > 0$$

i.e. only even powers of the fields will contribute to the transformation

Charges in the bulk

- The *bulk* action,

$$S[\tilde{\mathcal{A}}_\mu] = \int_D \text{tr} \left(\tilde{\mathcal{F}}_{\mu\nu} \tilde{\mathcal{F}}^{\mu\nu} \right) d\text{vol},$$

where $\tilde{\mathcal{F}}^{\mu\nu}$ is constructed from $\tilde{\mathcal{A}}_\mu$, and takes the form

$$\tilde{\mathcal{F}}_{\mu\nu} \equiv \partial_\mu \tilde{\mathcal{A}}_\nu - \partial_\nu \tilde{\mathcal{A}}_\mu - i[\tilde{\mathcal{A}}_\mu, \tilde{\mathcal{A}}_\nu] = e^{i\Psi} \mathcal{F}_{\mu\nu} e^{-i\Psi}$$

- same as without Ψ , but apply covariant phase space formalism on the tilded objects

$$\Omega^{\text{bulk}}[\delta_1, \delta_2] = - \int_\Sigma \text{tr} \left(\delta_1 \tilde{\mathcal{F}}^{\mu\nu} \wedge \delta_2 \tilde{\mathcal{A}}_\nu \right) dS_\mu - (1 \leftrightarrow 2),$$

Charges at null infinity

- Covariant phase space formalism gives:

$$\tilde{Q}_\Lambda = \int_\Sigma \partial_\nu \text{tr}(\Lambda \tilde{\mathcal{F}}^{\mu\nu}) dS_\mu.$$

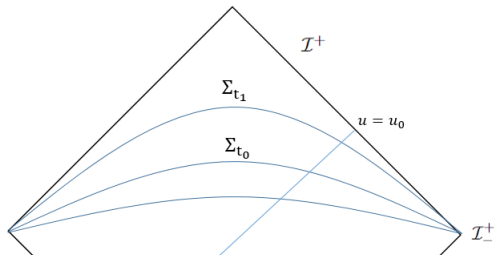


Figure: Cauchy Surfaces $\Sigma_{t_0}, \Sigma_{t_1}$, with $t_0 < t_1$, and constant u ray. Reach \mathcal{I}^+ by taking $r \rightarrow +\infty$ with u fixed.

Dressed charge algebra

- Charge on the celestial sphere

$$\tilde{Q}_\Lambda = \int_{S^2} \text{tr}(\Lambda \tilde{\mathcal{F}}_{ru}) \sqrt{g_{S^2}} dz d\bar{z}.$$

- We can now define the charge density

$$\tilde{q}_\Lambda = \text{tr} \left(\sqrt{g_{S^2}} \Lambda \tilde{\mathcal{F}}_{ru} \right)^{(0)}$$

In terms of the original field strength, the charge density is

$$\begin{aligned} \tilde{q}_\Lambda &= \text{tr} \left(\sqrt{g_{S^2}} \Lambda e^{i\Psi} \mathcal{F}_{ru} e^{-i\Psi} \right)^{(0)} \\ &= \text{tr} \left(\sqrt{g_{S^2}} e^{-i\Psi} \Lambda e^{i\Psi} \mathcal{F}_{ru} \right)^{(0)} \end{aligned}$$

- Charge algebra

$$\{\tilde{q}_{\Lambda_1}, \tilde{q}_{\Lambda_2}\} = \tilde{q}_{-i[\Lambda_1, \Lambda_2]}$$

Recursion relations to all orders

- Assume fall-off

$$\mathcal{A}_\mu(u, r, z, \bar{z}) = \sum_{n \in \mathbb{Z}} A_\mu^{(-n)}(u, z, \bar{z}) \frac{1}{r^n},$$

- Further expand in u (dual of energy)

$$\mathcal{A}_\mu(u, r, z, \bar{z}) = \sum_{n, k \in \mathbb{Z}} A_\mu^{(-n, k)}(z, \bar{z}) \frac{u^k}{r^n},$$

- The overleading gauge parameter and Stueckelberg field also have a simpler expansion in this case

$$\Lambda_+(u, r, z, \bar{z}) = \sum_{k=1}^{\infty} r^k \Lambda_+^{(k)}(u, z, \bar{z})$$

$$\Psi(u, r, z, \bar{z}) = \sum_{k=1}^{\infty} r^k \Psi^{(k)}(u, z, \bar{z})$$

- Work in radial gauge

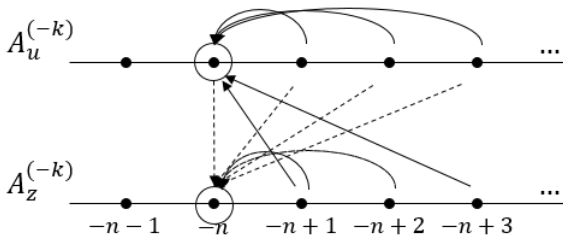
$$\mathcal{A}_r = 0, \quad \text{and} \quad A_u^{(0)} = 0,$$

Recursion relations to all orders

- Using e.o.m. and Bianchi identities, we find

$A_u^{(-n)}$ depends on $\{A_z^{(-k)}, A_{\bar{z}}^{(-k)}\}_{k < n}$ and $A_u^{(-n,0)}$

$A_z^{(-n)}$ depends on $A_u^{(-n)}, \{A_z^{(-k)}, A_{\bar{z}}^{(-k)}, A_u^{(-k)}\}_{k < n}$ and $A_z^{(-n,0)}$.



- the standard $u \rightarrow -\infty$ fall-offs for the radiative data $A_z^{(0)}$,

$$A_z^{(0)}(u, z, \bar{z}) = A_z^{(0,0)}(z, \bar{z}) + o(u^{-\infty}),$$

translates to

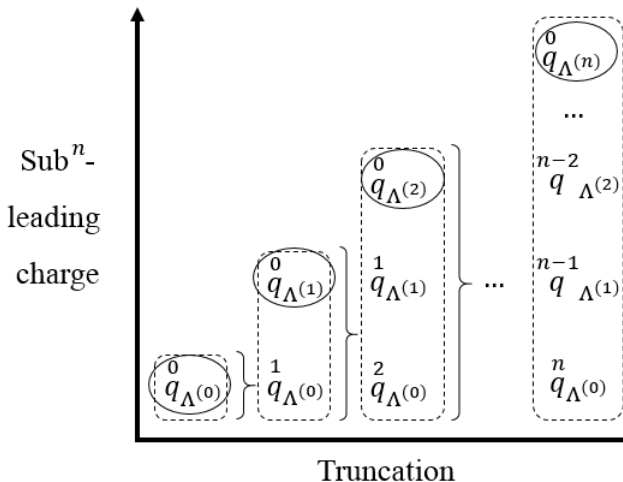
$o(u^{-\infty})$	$F_{uz}^{(0)}$	$F_{u\bar{z}}^{(0)}$				
$O(1)$	$F_{uz}^{(-1)}$	$F_{u\bar{z}}^{(-1)}$	$F_{ru}^{(-2)}$	$F_{z\bar{z}}^{(0)}$		
$O(u)$	$F_{uz}^{(-2)}$	$F_{u\bar{z}}^{(-2)}$	$F_{ru}^{(-3)}$	$F_{z\bar{z}}^{(-1)}$	$F_{rz}^{(-2)}$	$F_{r\bar{z}}^{(-2)}$
$O(u^2)$	$F_{uz}^{(-3)}$	$F_{u\bar{z}}^{(-3)}$	$F_{ru}^{(-4)}$	$F_{z\bar{z}}^{(-2)}$	$F_{rz}^{(-3)}$	$F_{r\bar{z}}^{(-3)}$
...						
$O(u^n)$	$F_{uz}^{(-n-1)}$	$F_{u\bar{z}}^{(-n-1)}$	$F_{ru}^{(-n-2)}$	$F_{z\bar{z}}^{(-n)}$	$F_{rz}^{(-n-1)}$	$F_{r\bar{z}}^{(-n-1)}$

Table: fall-offs in u

Charges revisited

- on the celestial sphere (i.e., \mathcal{I}_-^+), the charge is the limit $r \rightarrow +\infty$ and $u \rightarrow -\infty$ in the quantity

$$\tilde{Q}_\Lambda = \int_{S^2} \text{tr}(\Lambda \tilde{\mathcal{F}}_{ru}) r^2 \gamma_{z\bar{z}} dz d\bar{z}.$$



Treuncated charge algebra and recursion relations

- Explicitly, *recursion for charges*

$$\begin{aligned} \tilde{q}_{\Lambda^{(n)}}^0 &= \text{Tr} \left(\Lambda^{(n)} F_{ru}^{(-2-n)} \right) \\ \tilde{q}_{\Lambda^{(j)}}^{n-j} &= \tilde{q}_{[e^{ad-i\psi}(\Lambda^{(j)})]\{n\}}^0, \end{aligned}$$

- Truncated charge algebra : for $0 \leq j, k \leq n$,

$$\left\{ \tilde{q}_{\Lambda_1^{(k)}}^{n-k}, \tilde{q}_{\Lambda_2^{(j)}}^{n-j} \right\} = \begin{cases} \tilde{q}_{-i[\Lambda^{(k)}, \Lambda^{(j)}]}^{n-k-j} & \text{if } j+k \leq n, \\ 0 & \text{otherwise.} \end{cases}$$

With this result, we have a closed Poisson algebra, \mathfrak{p}_n such that the action of the charges is canonical and we have the following chain

$$\mathfrak{p}_0 \subset \dots \subset \mathfrak{p}_{n-1} \subset \mathfrak{p}_n.$$

- Truncation of

$$\{\tilde{q}_{\Lambda_1}, \tilde{q}_{\Lambda_2}\} = \tilde{q}_{-i[\Lambda_1, \Lambda_2]}$$

- match previous results for $n = 1$ YM and electromagnetism [\[Campiglia, Peraza\]](#)

Infinite algebras

- infinite dimensional YM version of the $w_{1+\infty}$ algebra

$$[S_m^{p,a}(\bar{z}), S_n^{q,b}(\bar{z})] = -i f_c^{ab} S_{m+n}^{p+q-1,c}(\bar{z})$$

- Following procedure from [Freidel,Pranzetti,Raclariu], construct

$$S_{m,n}^{1+\frac{s}{2},a} = \int dzd\bar{z} z^{m+\frac{s}{2}} \bar{z}^{n-\frac{s}{2}} r_s^a$$

where

$$r_s^a T^a \xleftarrow{\text{renormalised}} \mathcal{R}_s$$

and the charges \mathcal{R}_s satisfy some recursion relations which are a subset of out recursion relations, corresponding to the self-dual sector.

Infinite algebras

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$$[S_m^{p,a}(\bar{z}), S_n^{q,b}(\bar{z})] = -i f_c^{ab} S_{m+n}^{p+q-1,c}(\bar{z})$$

- Following procedure from [Freidel,Pranzetti,Raclariu],

$$S_{m,n}^{1+\frac{s}{2},a} \leftarrow r_s^a$$

where we have restricted to the self-dual sector via

$$r_s^1 = r_s^1(A_z^{(0)})$$

$$r_s^2 = r_s^2(A_z^{(0)}, A_z^{(0)})$$

$$r_s^3 = r_s^3(A_z^{(0)}, A_z^{(0)}, A_z^{(0)})$$

...

where A_z corresponds to the positive helicity, and $A_{\bar{z}}$ to the negative helicity modes.

Infinite algebras to boring algebras

Motivation

Set-up

Stueckelberg trick

Charges

Recursion relations

Infinite algebras

Quasi-universal soft limits

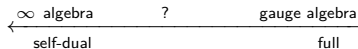
- infinite dimensional YM version of the $w_{1+\infty}$ algebra

$$[S_m^{p,a}(\bar{z}), S_n^{q,b}(\bar{z})] = -i f_c^{ab} S_{m+n}^{p+q-1,c}(\bar{z})$$

- Recall that including all the charges we just get the gauge algebra

$$\{\tilde{q}_{\Lambda_1}, \tilde{q}_{\Lambda_2}\} = \tilde{q}_{-i[\Lambda_1, \Lambda_2]}$$

- Anything in between ?



- Similar for gravity.

Back to soft limits

- Soft theorems

$$\lim_{\omega \rightarrow 0} A_{n+1} = \underbrace{\left(\frac{1}{\omega} S^{(0)} + S^{(1)} + \dots \right) A_n}_{\text{universal}} + \underbrace{\dots}_{\text{non-universal}}$$

- universal terms come from symmetries
- in some simple cases non-universal terms vanish
- Focus on subleading via

$$\lim_{\omega \rightarrow 0} (1 + \omega \partial_\omega) A_{n+1} = S^{(1)} A_n + \dots$$

- take a more interesting scenario

$$\lim_{\omega \rightarrow 0} (1 + \omega \partial_\omega) A_{n+1} = S^{(1)} A_n + \text{quasi-universal}$$

Quasi-universal terms from symmetries

- QED with higher derivative interactions, e.g. ϕF^2 [Elvang,Jones,Naculich]

$$\lim_{\omega \rightarrow 0} (1 + \omega \partial_\omega) A_{n+1} = S^{(1)} A_n + \tilde{S}^{(1)} [A_n]$$

where $S^{(1)}$ is the usual

$$S^{(1)} = -ie \sum_{k=1}^n Q_k \frac{p_s^\mu \varepsilon_s^\nu}{p_k \cdot p_s} \mathcal{J}_{k\mu\nu}$$

and $\tilde{S}^{(1)}$ is an *operator*

$$\tilde{S}_+^{(1)} A_n = \sum_{k=1}^n \frac{[sk]}{\langle sk \rangle} \mathcal{F}_k A_n, \quad \tilde{S}_-^{(1)} = (\tilde{S}_+^{(1)})^\dagger$$

where \mathcal{F}_k is a particle changing operator.

Quasi-universal terms from symmetries

$$\lim_{\omega \rightarrow 0} (1 + \omega \partial_\omega) A_{n+1} = S^{(1)} A_n + \tilde{S}^{(1)} [A_n]$$

- [Laddha,Mitra] showed that both $S^{(1)}$ and \tilde{S}^1 follow as a Ward identity for an gauge symmetry which **violates the fall-off of the fields**

$$\Lambda(u, r, z, \bar{z}) = r\Lambda(z, \bar{z}) + \frac{u}{2}(D^2 + 2)\Lambda(z, z\bar{b})$$

and, schematically, the symplectic potential

$$\theta(\varphi, \delta\varphi) = \theta_{old}(\varphi, \delta\varphi) + \theta_{new}(\varphi, \delta\varphi)$$

- $\theta_{old}(\varphi, \delta\varphi)$ gives the universal soft factor
- $\theta_{new}(\varphi, \delta\varphi)$ is subleading relative to $\theta_{old}(\varphi, \delta\varphi)$ for a standard δ_Λ ,
- it becomes of the same order when we allow $\delta_{r\Lambda}$ and gives the quasi-universal term

Quasi-universal terms in extended phase space

$$\lim_{\omega \rightarrow 0} (1 + \omega \partial_\omega) A_{n+1} = S^{(1)} A_n + \tilde{S}^{(1)} [A_n]$$

- Overleading part of gauge parameter acts on the Stueckelberg field

$$\Lambda(u, r, z, \bar{z}) = r\Lambda(z, \bar{z}) + \frac{u}{2}(D^2 + 2)\Lambda(z, z\bar{b})$$

- the symplectic potential is dressed with the Stueckelberg fields

$$\theta(\tilde{\varphi}, \delta\tilde{\varphi}) = \theta_{old}(\tilde{\varphi}, \delta\tilde{\varphi}) + \theta_{new}(\tilde{\varphi}, \delta\tilde{\varphi})$$

- Universal and (mildly) non-universal terms arising from **symmetries acting canonically** on an extended phase space
- Different gauge choices can simplify things !

Future directions

- More structure from symmetries

$$\sum_{n,k} \tau^n \log^k \tau \Lambda^{(n;k)}(\vec{y}) \xrightarrow{\text{Ward identity}} \text{universal} + (\text{mildly non-universal}) \text{ terms}$$

- Loop corrections
- Procedure is quite general, currently extending to gravity

[Geiller,SN,Peraza,Pizzolo].

$$\tilde{g} = e^{\mathcal{L}_V} g.$$

where we have a Stueckelberg vector coming from the diffeo parameter

$$\xi^\mu \rightarrow V^\mu$$

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Thank You !