Soft factorization on AdS spacetimes

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Overview and Motivation

Derivation from Classical soft theorems

Derivation from CFT₃ Ward identity

Conclusion and future directions

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Single soft theorems on flat spacetime

Factorization of scattering amplitudes involving soft particles

$$\lim_{k \to 0} \mathcal{M}_{n+1}(\varepsilon, k; \{p_a\}) = \mathsf{Soft factor} \ \times \ \mathcal{M}_n(\{p_a\})$$

 ε : polarization; k: soft momentum; p_a : hard momenta \mathcal{M}_n : Amplitude with n external particles

The soft factor admits a frequency expansion

$$S_{\text{em}} = \sum_{i} S_{\text{em}}^{(i)} \qquad S_{\text{em}}^{(0)} = \sum_{a=1}^{n} Q_{(a)} \eta_{(a)} \frac{\varepsilon \cdot p_{(a)}}{p_{(a)} \cdot k} \text{ (leading)}$$
$$= \pm 1(-1) \text{ for outgoing (incoming)} \quad Q_{(a)} \text{ charges}$$

 $\eta_{(a)}=+1(-1)$ for outgoing (incoming), $Q_{(a)}$ charges and ε^{μ} the polarization

 Renewed interest in relation to asymptotic symmetries and holography on asymptotically flat spacetimes [see talk by Laura Donnay]

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Soft theorem = Large gauge Ward identity

The soft factor satisfies

 $\lim_{\omega \to 0} \langle \mathsf{out} | a^{\mathsf{out}}_+(\omega \hat{x}) \mathcal{S} | \mathsf{in} \rangle = S^{(0)}_{\mathsf{em}} \langle \mathsf{out} | \mathcal{S} | \mathsf{in} \rangle = -\lim_{\omega \to 0} \langle \mathsf{out} | \mathcal{S} a^{\mathsf{in}\dagger}_-(\omega \hat{x}) | \mathsf{in} \rangle$

with S the S-matrix for the hard process, \hat{x} and ω the soft photon orientation and frequency resp. $[k^{\mu} = \omega(1, \hat{x})]$

► Equivalent to large gauge Ward identity satisfied by soft modes, with gauge parameter e(z_k, z
_k)

$$\lim_{\omega \to 0} \frac{\sqrt{2}\omega}{1+z\bar{z}} \langle \mathsf{out} | a^{\mathsf{out}}_{+}(\omega \hat{x}) \mathcal{S} | \mathsf{in} \rangle = \sum_{k=1}^{n} Q_{(k)} \eta_{(k)} \epsilon(z_k \,, \bar{z}_k) \langle \mathsf{out} | \mathcal{S} | \mathsf{in} \rangle$$

$$\epsilon(z_k, \bar{z}_k) = rac{1}{z - z_k} \quad (+ {\sf ve helicity})$$

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 Also realized to subleading orders [see talks by Silvia Nagy, Sangmin Choi and Shreyansh Agrawal]

Motivation

- Broad Goal: Generalizations of soft factorization for scattering on non-asymptotically flat spacetimes?
- On spacetimes with a cosmological constant, we lack scattering amplitudes. They can be defined in appropriate flat spacetime limits [see talks by Luis Fernando Alday, Paul McFadden, Bin Zhu, and this session]
- If we can address soft factorization in AdS, we might better understand flat spacetime holography

Thus far for bulk scattering processes:

- ▶ Weinberg's soft theorems derivable from the flat limit $(L \rightarrow \infty)$ limit of a CFT₃ Ward identity [E. Hijano, D. Neuenfeld, JHEP **11** 009 (2020)]
- ► We can find AdS radius dependent (L⁻²ⁿ) corrections of soft factors for flat spacetime amplitudes about the flat limit of AdS.
- These were derived using "classical soft theorems" [N. Banerjee, KF,
 A. Mitra, JHEP 08, 105 (2021); N. Banerjee, A. Bhatacharjee, A. Mitra, JHEP
 01, 038(2020)] and from a CFT₃ Ward identity [N. Banerjee, KF, A. Mitra,
 JHEP 04, 055 (2023)]

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Overview and Motivation

Derivation from Classical soft theorems

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Classical soft theorems (CST)

- Classical soft theorems [A. Laddha, A. Sen, JHEP 10 056 (2018); A. P. Saha, B. Sahoo, A. Sen, JHEP 06 153 (2020)] : Classical limits of soft photon and soft graviton factors derivable from ω → 0 limit of radiative solutions from classical scattering.
- Applicability: $\Delta E_{\text{scatterer}} \ll 1$; $\lambda_{\text{radiation}} \gg b$ (impact parameter)
- Derivable from energy flux in $\omega \to 0$

$$\lim_{\omega \to 0} \varepsilon^{\mu} \tilde{a}_{\mu} \left(\omega, \vec{x} \right) = -i \frac{e^{i\omega R}}{4\pi R} S_{\text{em}}$$
$$\lim_{\omega \to 0} \varepsilon^{\mu\nu} \tilde{e}_{\mu\nu}(\omega, \vec{x}) = -i \frac{e^{i\omega R}}{4\pi R} S_{gr}$$

where \tilde{a}_{μ} and $\tilde{e}_{\mu\nu}(\omega, \vec{x})$ are electromagnetic and gravitational perturbations in frequency space, ε^{μ} and $\varepsilon^{\mu\nu}$ are photon and graviton polarizations; R: distance to scatterer.

CST on AdS spacetimes about the flat limit

Hierarchy of length scales for probe scattering

GM << r << L

Leads to $1/L^{2n}$ corrections of flat spacetime results.

• Double scaling limit : $\omega \to 0$ as $L \to \infty$, with $\gamma := \omega L >> 1$

Produces $1/\gamma^{2n}$ corrections of flat spacetime soft factors

- ► $1/L^n$ Trajectory corrections of flat asymptotic trajectories exist Relevant for $\mathcal{O}(\gamma^{-4})$ corrected soft factors
- Derived leading (ω⁻¹) and subleading (ln ω⁻¹) soft photon and graviton factors, with their respective 1/γ² corrections. Derived γ⁻² corrected large gauge Ward identity in the leading soft photon case [N. Banerjee, KF, A. Mitra, JHEP 08, 105 (2021)]

$1/\gamma^2$ corrected soft photon theorem on AdS $_4$

► Leading soft photon factor result $S_{\text{em}}^{(0)} = S_{\text{em}}^{\text{f}(0)} + S_{\text{em}}^{\text{L}(0)} + \mathcal{O}\left(\gamma^{-4}\right)$

$$S_{\rm em}^{\mathsf{f}(0)} = \sum_{a=1}^{n} Q_{(a)} \eta_{(a)} \frac{\epsilon_{\mu} p_{(a)}^{\mu}}{p_{(a)} \cdot k} , \qquad S_{\rm em}^{L(0)} = \frac{\omega^2}{4\gamma^2} \sum_{a=1}^{n} Q_{(a)} \eta_{(a)} \frac{\epsilon_{\mu} p_{(a)}^{\mu}}{p_{(a)} \cdot k} \frac{\vec{p}_{(a)}^2}{\left(p_{(a)} \cdot k\right)^2}$$

From leading (universal) term, we infer the single soft theorem

$$\begin{split} \lim_{\omega \to 0} \omega \left(\langle \mathsf{out} | a_+(\omega \hat{x}) \mathcal{S} | \mathsf{in} \rangle + \langle \mathsf{out} | \tilde{a}_+(\omega \hat{x}) \mathcal{S} | \mathsf{in} \rangle \right) \\ &= \left(S_{\mathsf{em}}^{\mathsf{f}(0)} + S_{\mathsf{em}}^{L(0)} \right) \left\langle \mathsf{out} | \mathcal{S} | \mathsf{in} \right\rangle \end{split}$$

Parametrizing for massless scattering, we have

$$\begin{split} \lim_{\omega \to 0} \frac{\sqrt{2\omega} n_L}{(1+z\bar{z})} \langle \mathsf{out} | \tilde{a}_+ \mathcal{S} | \mathsf{in} \rangle &= \sum_{k=1}^n Q_{(k)} \eta_{(k)} \epsilon_L^{\mathsf{cst}}(z_k \,, \bar{z}_k) \langle \mathsf{out} | \mathcal{S} | \mathsf{in} \rangle \\ \epsilon_L^{\mathsf{cst}}(z_k \,, \bar{z}_k) &= \frac{1}{16\gamma^2} \frac{(1+z_k \bar{z}_k)^2 \, (1+z\bar{z})^2}{(\bar{z}-\bar{z}_k)^2 (z-z_k)^3} \quad (\text{+ve helicity}) \end{split}$$

Subtlety : Overall numerical constant n_L is not fixed by CST

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Flat spacetime from Lorentzian AdS/CFT

• Analysis in global AdS_{d+1} with metric

$$ds^{2} = \frac{L^{2}}{\cos^{2}\rho} \left(-d\tau^{2} + d\rho^{2} + \sin^{2}\rho \, d\Omega_{d-1}^{2} \right)$$

AdS boundary at $\rho = \frac{\pi}{2}$ with coordinates $x = \{\tau, z, \bar{z}\}$ Define

$$\frac{r}{L} = \tan(\rho) \approx \rho \ , \qquad \frac{t}{L} = \tau$$

 \blacktriangleright AdS metric in the $L \rightarrow \infty$ limit goes to the flat spacetime metric

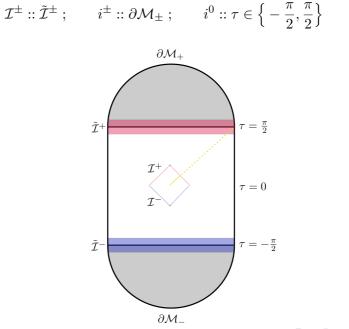
$$\lim_{L \to \infty} ds^2 = -dt^2 + dr^2 + r^2 d\Omega_2^2$$

Asymptotically flat patch centered about $\tau = 0$ in global AdS

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Bulk field modes from boundary operators by HKLL reconstruction
 [A. Hamilton, D. Kabat, G. Lifschytz, D. Lowe, PRD 74 066009 (2006)]

Asymptotic mapping



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Maxwell field in AdS from boundary sources

Bulk field with magnetic boundary conditions (no Coulombic sources)

$$\hat{\mathcal{A}}_{\mu}(X) = \mathcal{A}_{\mu}(\rho, x) \xrightarrow[\rho \to \frac{\pi}{2}]{} (\cos \rho) j_{\mu}(x)$$

Bulk field has HKLL reconstruction from boundary currents

$$\hat{\mathcal{A}}_{\mu}(X) = \int d^3x' \left[K^V_{\mu}(X;x') \epsilon^{a'b'}_{\tau'} \nabla_{a'} j^+_{b'} + K^S_{\mu}(X;x') \gamma^{a'b'} \nabla_{a'} j^+_{b'} + c.c. \right]$$

 $\epsilon^{a'b'c'}$ and $\nabla_{a'}$: bndy Levi-Civita tensor and covariant derivative; $K^V_\mu(X;x')$ and $K^S_\mu(X;x')$: HKLL kernels from "vector" ($\Delta=2$) and "scalar" ($\Delta=1$) type solutions of Maxwell's equations.

 \blacktriangleright We can expand the kernels about $L \rightarrow \infty$ to find

$$\hat{\mathcal{A}}_{\mu}(X) = \hat{\mathcal{A}}_{\mu}^{\mathsf{flat}}(X) + \hat{\mathcal{A}}_{\mu}^{L^{-2}}(X) + \mathcal{O}\left(L^{-4}\right)$$

Bulk fields about flat limit

The HKLL kernels have a sum over discrete energies. These are large and scale like L to get continuous flat spacetime frequencies

$$2\kappa + l + \Delta = \omega_{\kappa} \to \omega L \qquad \Rightarrow \qquad \sum_{\kappa} \to \frac{1}{2} \int d\omega L$$

We find

$$\hat{\mathcal{A}}_{z}^{\text{out; flat}}(X) = \frac{1}{4\pi} \int d^{3}x' \int d\omega \, r \, j_{l}(r\omega)[[\cdots]]$$
$$\hat{\mathcal{A}}_{z}^{\text{out};L^{-2}}(X) = \frac{1}{16\pi} \int d^{3}x' \int d\omega \, \frac{r}{\gamma^{2}} \, j_{l}(r\omega)l(l+1)[[\cdots]]$$

with

$$[[\cdots]] = \left[\sum_{l,m} \frac{Y_{lm}^* \left(\Omega'\right)}{-l(l+1)} \partial_z Y_{lm} \left(\Omega\right) (i)^{-l} e^{i\omega t} e^{-i\omega L \left(\tau' - \frac{\pi}{2}\right)} D^{\bar{z}'} j_{\bar{z}'}^+ + c.c. \right]$$

[Likewise for ingoing modes, and other components]

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Soft modes from bulk fields

The bulk field solutions are about the flat limit, and we can define flat spacetime modes in the usual way

$$\hat{a}_{(\lambda)}^{\mathsf{out}} = \lim_{t \to \infty} i \int d^3 \vec{x} \, \varepsilon^{(\lambda)\mu} e^{-iq \cdot x} \overleftrightarrow{\partial_0} \hat{\mathcal{A}}_{\mu}^{\mathsf{out}}(X)$$

 \blacktriangleright The $\hat{\mathcal{A}}_z^{\mathrm{out;\;flat}}(X)$ and $\hat{\mathcal{A}}_z^{\mathrm{out;}L^{-2}}(X)$ pieces give

$$\hat{a}_{+}^{\text{out;flat}}(\omega \hat{x}) = \frac{1}{4} \frac{1 + z\bar{z}}{\sqrt{2}\omega} \int d\tau' \, e^{i\omega L\left(\frac{\pi}{2} - \tau'\right)} \int d^2 z' \epsilon(z') D^{\bar{z}'} j_{\bar{z}'}^-(x')$$
$$\hat{a}_{+}^{\text{out;}L^{-2}}(\omega \hat{x}) = \frac{1}{4} \frac{1 + z\bar{z}}{\sqrt{2}\omega} \int d\tau' \, e^{i\omega L\left(\frac{\pi}{2} - \tau'\right)} \int d^2 z' \epsilon_L(z') D^{\bar{z}'} j_{\bar{z}'}^-(x')$$

with

$$\epsilon(z') = \frac{1}{(z-z')}, \quad \epsilon_L(x') = \frac{1}{8\pi\gamma^2} \int d\Omega_w \left[\frac{(1+z'\bar{z}')^2 (1+z_w\bar{z}_w)^2}{(\bar{z}'-\bar{z}_w)^2 (z-z_w)^3} \right]$$

• Soft limit: Phase dominated by $au = \frac{\pi}{2}$ around $\tilde{\mathcal{I}}^+$

$$\lim_{\omega \to 0} \lim_{L \to \infty} \int d\tau' \, e^{i\omega L\left(\frac{\pi}{2} - \tau'\right)} \int d^2 z' \equiv \int_{\tilde{\mathcal{I}}^+} d^3 x'$$

Recovering classical soft theorem gauge parameter

Consider a "collinear expansion"

$$|z_w - z'| \approx |z - z'| + \delta \approx |z_w - z| \qquad \text{with } \delta << 1$$

This assumes $\{z_w, \bar{z}_w\}$ can be expanded about $\{z, \bar{z}\}$ or $\{z', \bar{z}'\}$, with |z - z'| as a minimal distance.

Helps realize properties of the leading soft photon factor:

diverges as ω^{-1} and collinearly diverges as $z \to z'$

• We can now formally integrate $\epsilon_L(x')$ to find

$$\epsilon_L(x') = 8\epsilon_L^{\mathsf{cst}}(x') + \text{`` corrections''}$$

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with "corrections" involving derivatives of $\epsilon_L^{\text{cst}}(x')$

Ward identity to Soft theorem

• The integrated U(1) Ward identity on the boundary

$$\int d^3x' \,\alpha(x')\partial'_{\mu}\langle 0|T\{j^{\mu}(x')X\}|0\rangle = \left(\sum_{i=1}^n Q_i\alpha(x'_i) - \sum_{j=1}^m Q_j\alpha(x'_j)\right)\langle 0|T\{X\}|0\rangle$$

 $T{X}: n (\tau > 0)$ and $m (\tau < 0)$ time ordered operators with charges Q_i $(i = 1, \dots, n)$ and $Q_j (j = 1, \dots, m)$

- ▶ $\langle 0|T\{X\}|0\rangle$ goes to the S-matrix on replacing operators with corresponding modes in the flat limit.
- α(x')|_{Ĩ[±]} = ε(x') gives the Weinberg soft photon theorem
 [E. Hijano, D. Neuenfeld, JHEP 11 009 (2020)]

▶ With $\alpha(x')|_{\tilde{I}^{\pm}} = 8\epsilon_L^{\text{cst}}(x')$ we recover the γ^{-2} corrected soft photon theorem and fixes $n_L = \frac{1}{8}$ [N. Banerjee, KF, A. Mitra, JHEP **04**, 055 (2023)]

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- ▶ Soft theorems about the flat spacetime limit of AdS involve corrections in powers of γ^{-2} where $\gamma = \omega L >> 1$
- \blacktriangleright Two independent derivations give a consistent result for γ^{-2} corrected soft photon factor
- Future: Derivation using Witten diagrams with a L^{-1} expansion
- ▶ Implications on CCFT sector within CFT about $\tau = \pm \frac{\pi}{2}$? [L. P. de Gioia, A-M Raclariu, 2303.10037 [hep-th]]
- Soft and flat limits of boundary correlators do not commute on AdS [C. Chowdhury, A. Lipstein, J. Mei, Y. Mo, 2407.16052 [hep-th]]
- \blacktriangleright Could a resummation of γ^{-2n} corrections relate the two limits?

Thank You

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