



UNIVERSITÄT  
HEIDELBERG  
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SEIT 1386

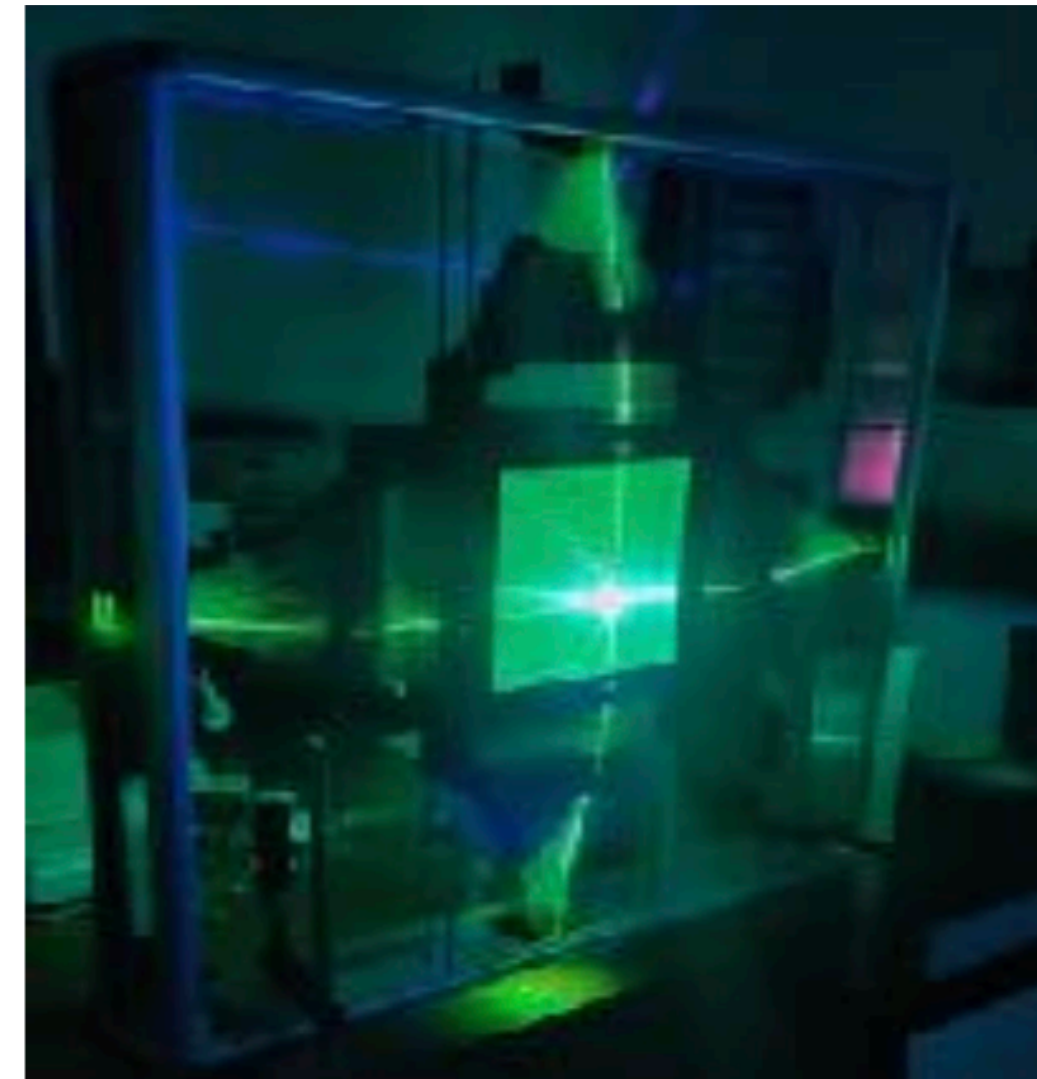
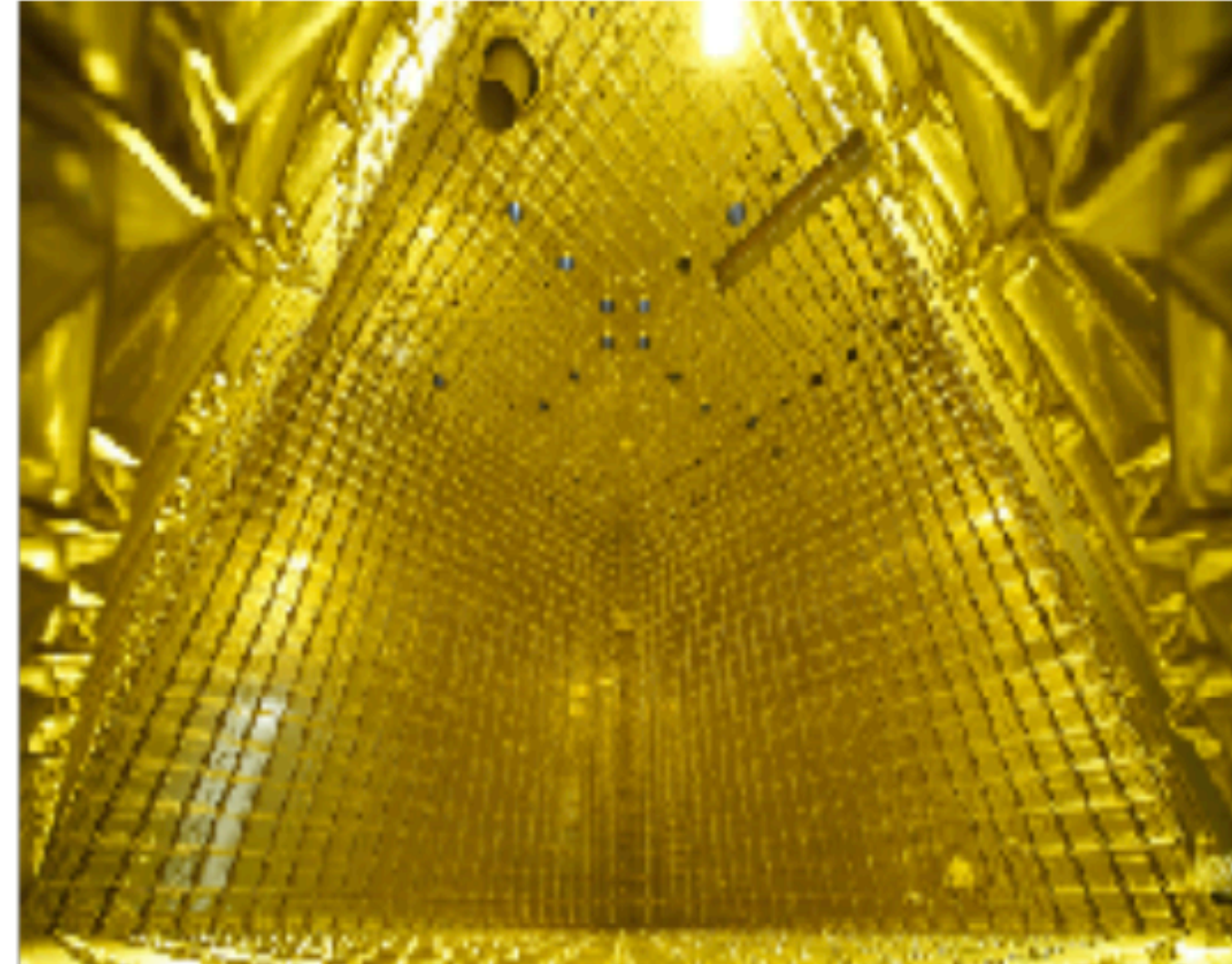
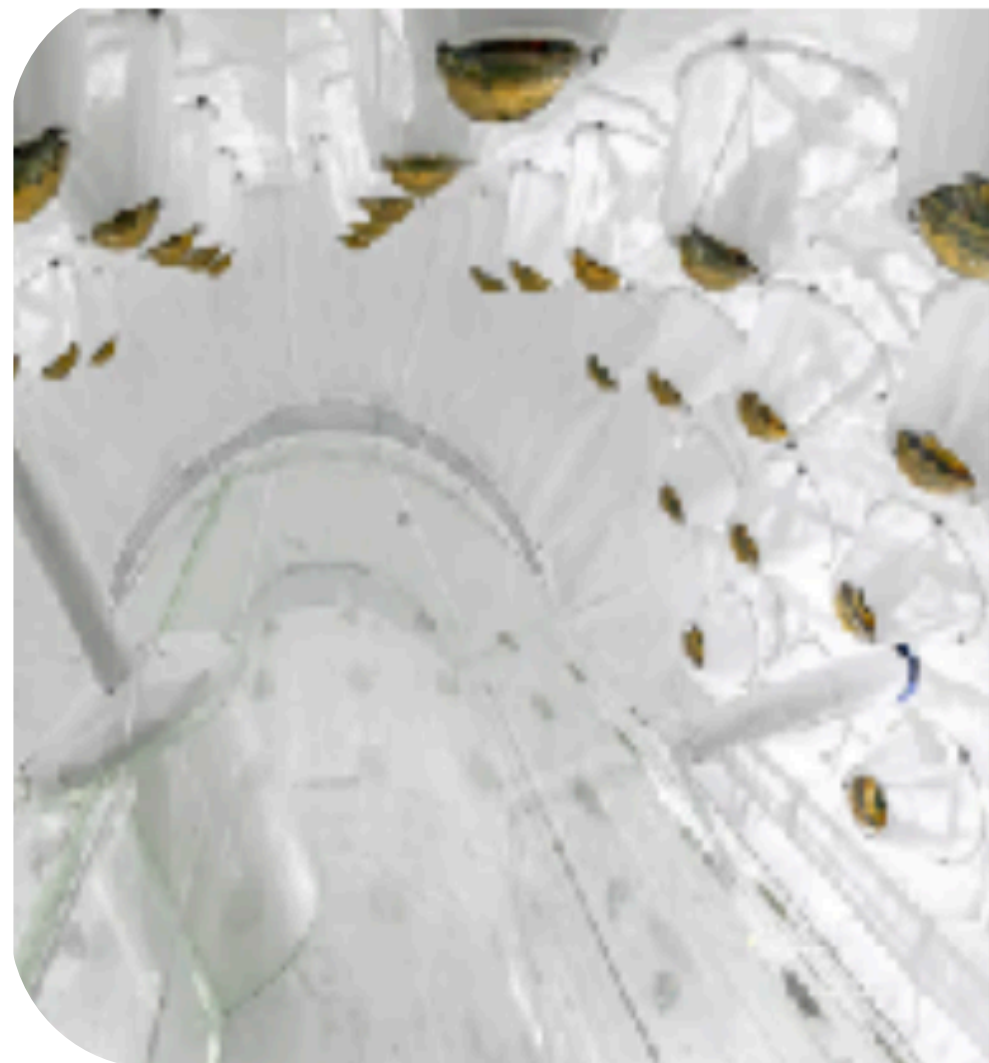


KIRCHHOFF-  
INSTITUT  
FÜR PHYSIK

# Liquid (non-HEP) Detectors - What it needs for Future Research

HighRR Lecture Week, Bergen, 13.06.2024

Belina VON KROSIGK (bkrosigk@kip.uni-heidelberg.de)



Credit: Roxanne Guenette & Jocelyn Monroe



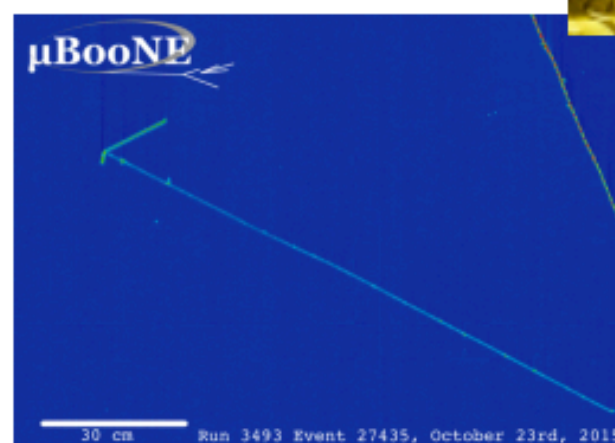
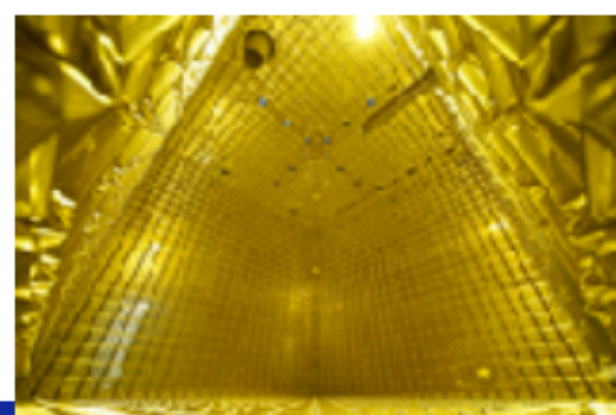
# The science of interest: rare event searches

Slide from:

R. Guenette, J. Monroe  
DRD2 Collaboration Meeting  
5-7 February 2024

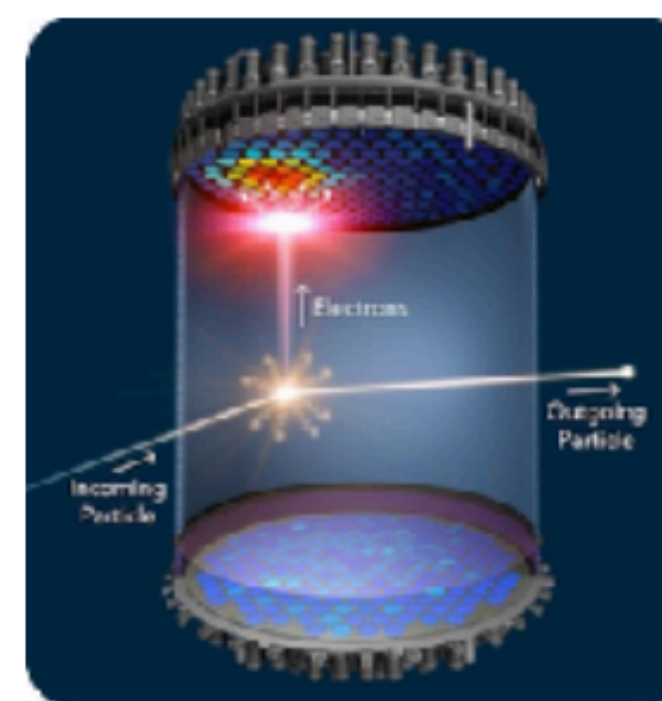
## Neutrinos

- Oscillation precision measurements ( $\delta_{CP}$ , mass ordering,  $\theta_{23}$  octant, sterile vs)
- Neutrino interactions (from CEvNS to DIS)
- Astro neutrinos



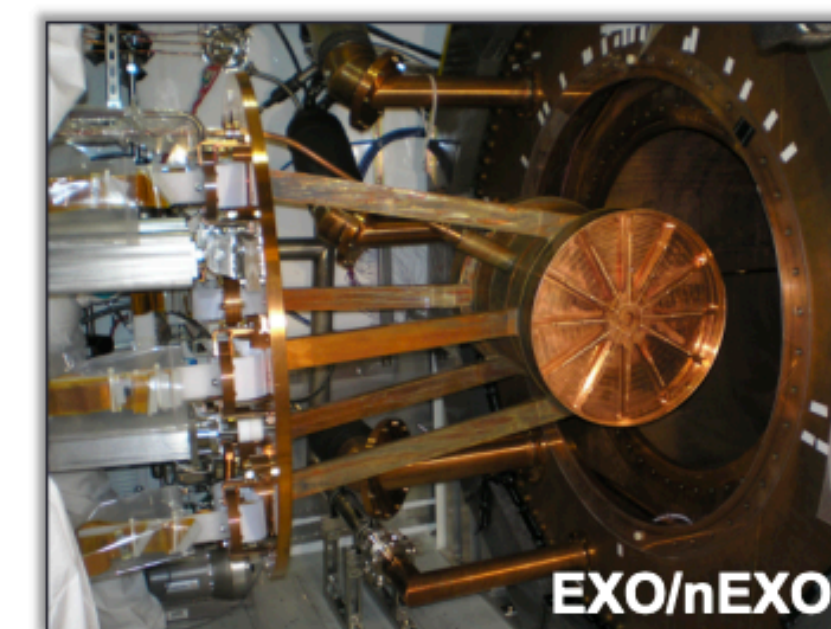
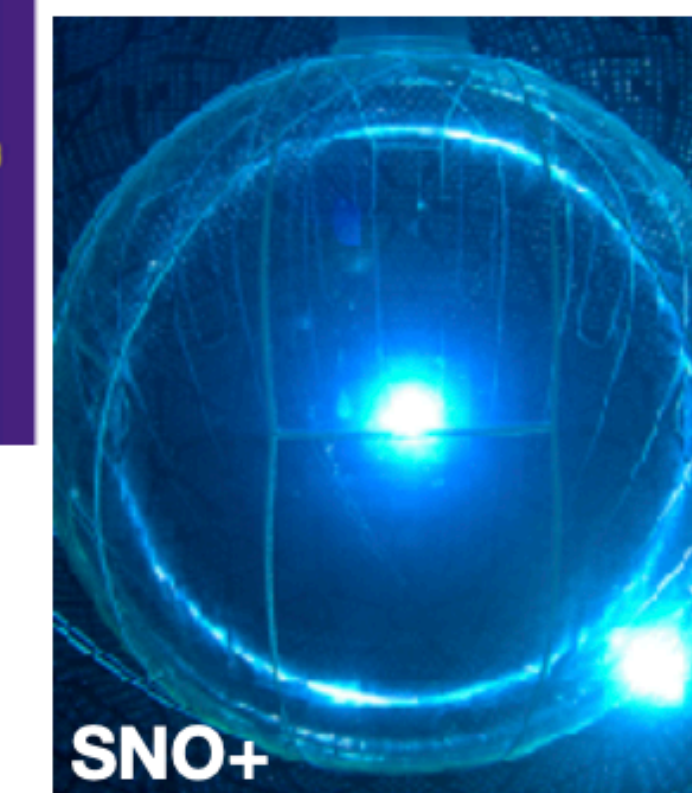
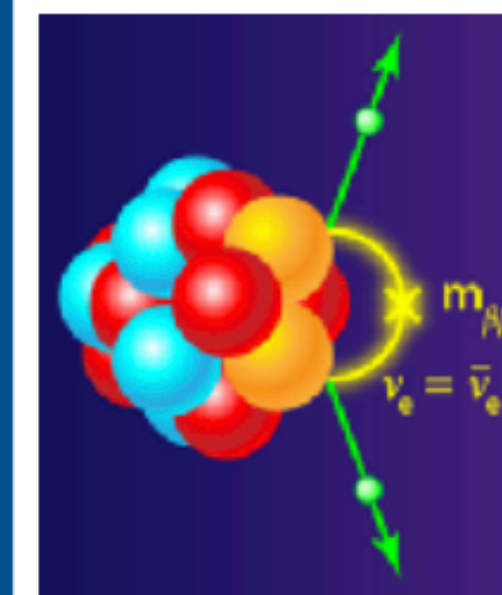
## Dark Matter

- Direct detection (WIMPs, ...)



## $0\nu\beta\beta$

- Search for Majorana neutrinos







# The physics needs (high-level overview)

Slide from:

R. Guenette, J. Monroe  
DRD2 Collaboration Meeting  
5-7 February 2024

## Neutrinos

- **Push Energy thresholds down** to  $\sim 1\text{MeV}$  to enhance oscillation physics, supernovae vs study, to enable solar vs ...
- **Unambiguous readout**
- **Scalability**

## Dark Matter

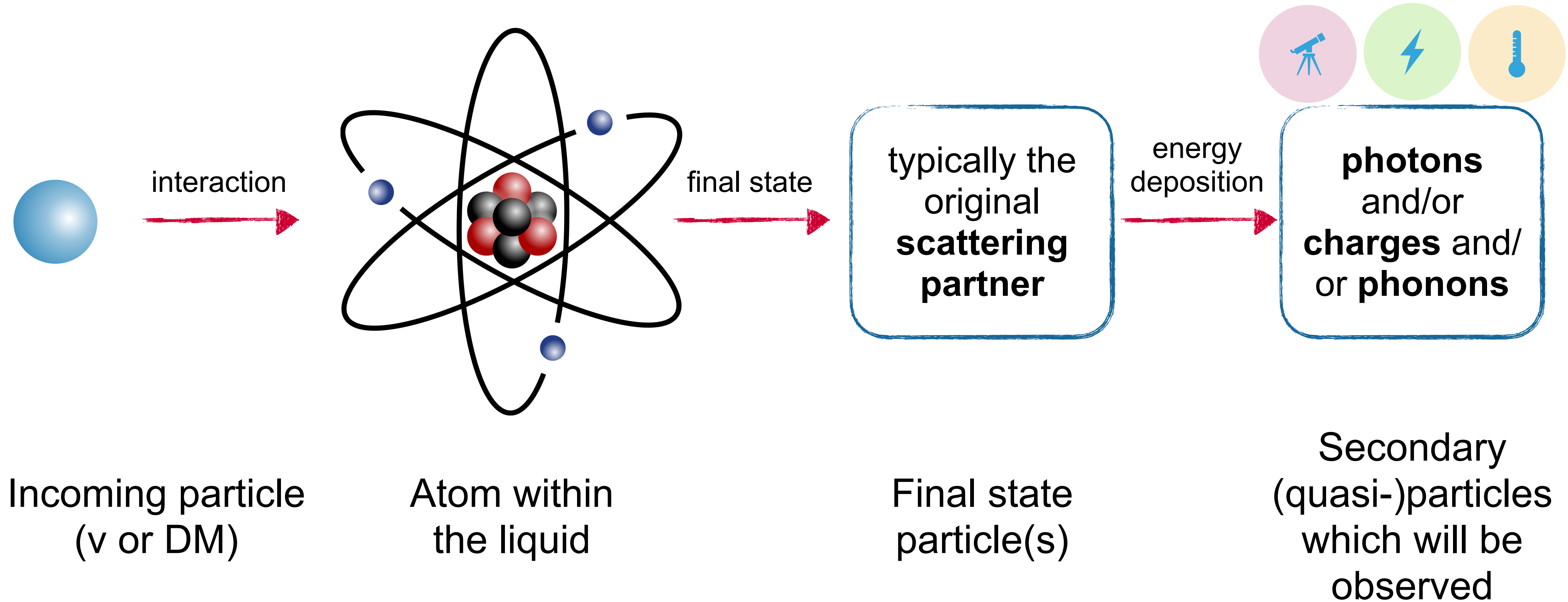
- **Push Energy thresholds down** to  $1\text{ meV}/10\text{ eV}/1\text{ keV}$  to enable low mass DM/ $1\text{ GeV DM}/\text{WIMPs}$ .
- **Reduce background rates**
- **Scalability**

## $0\nu\beta\beta$

- **Improve Energy Resolution** to sub-% FWHM
- **Reduce background rates**
- **Scalability**



# The basic detection principle







# The basic detection principle

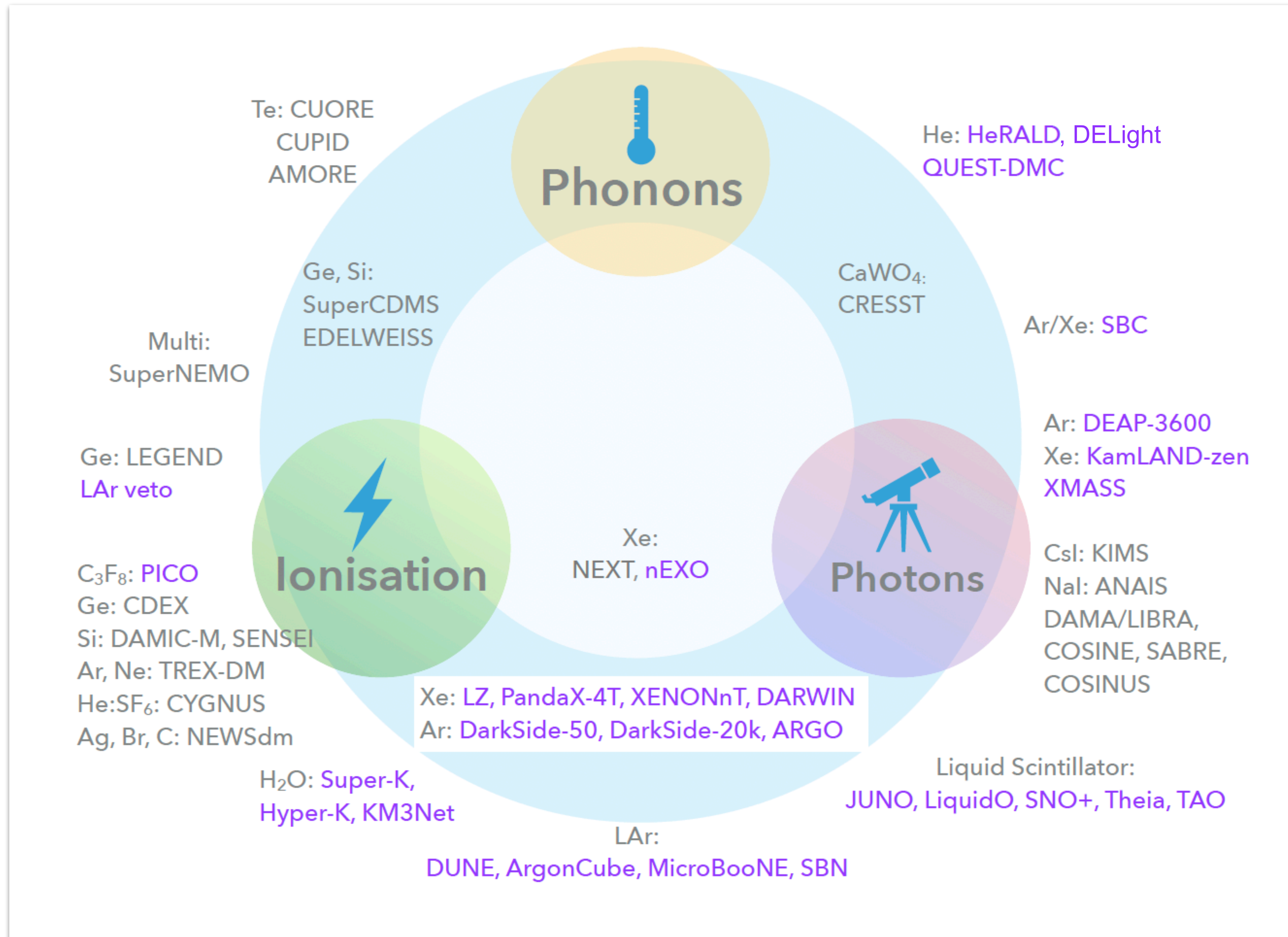
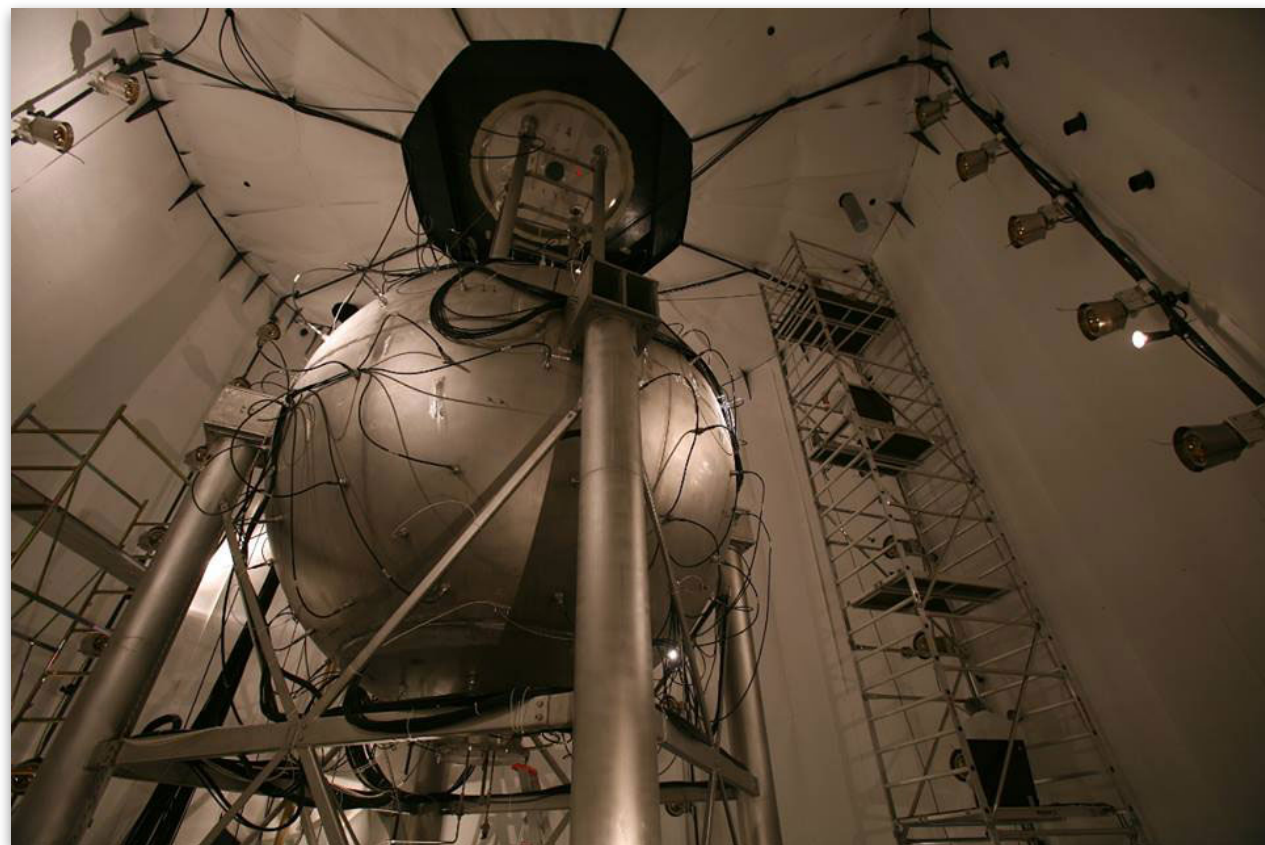
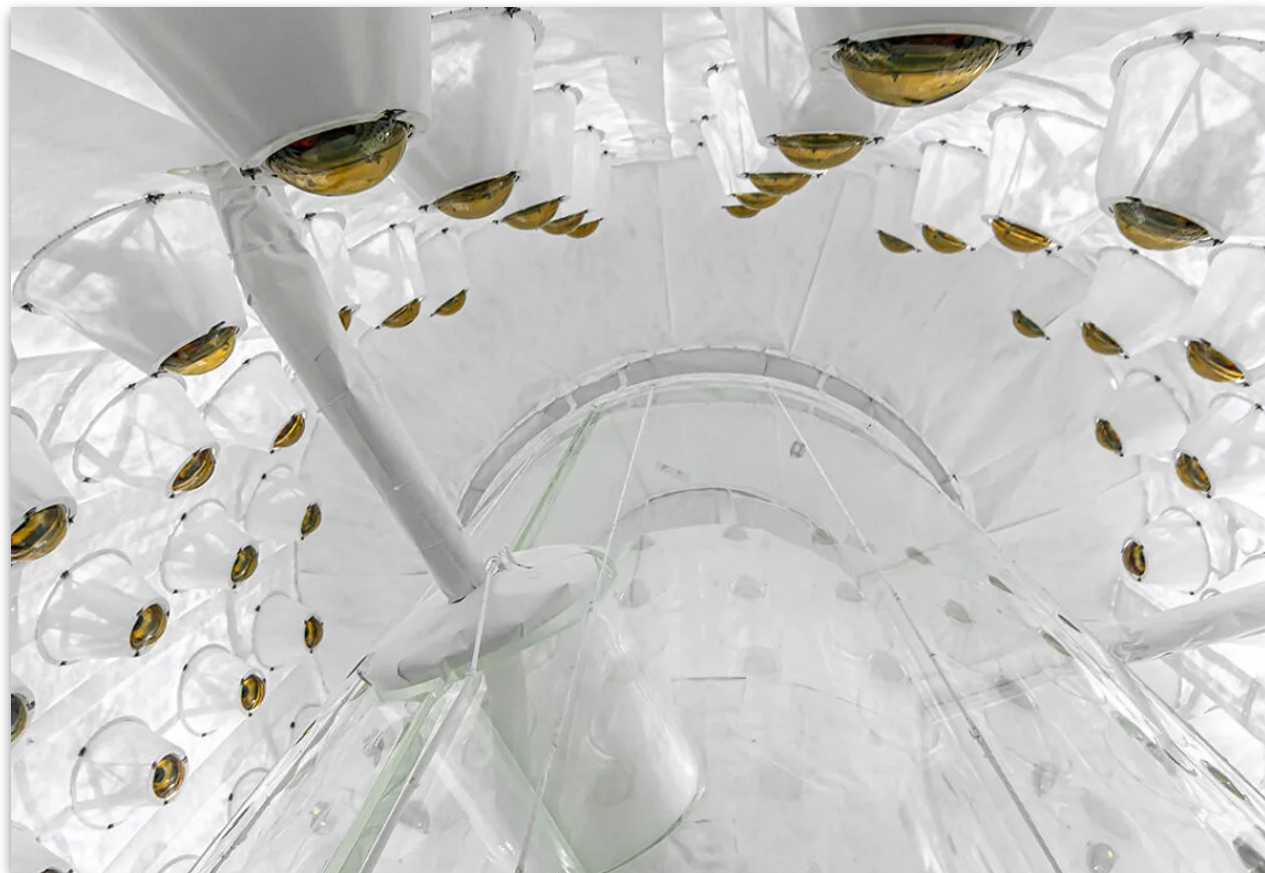


Figure modified from L. Baudis, ECFA, Plenary Input Session

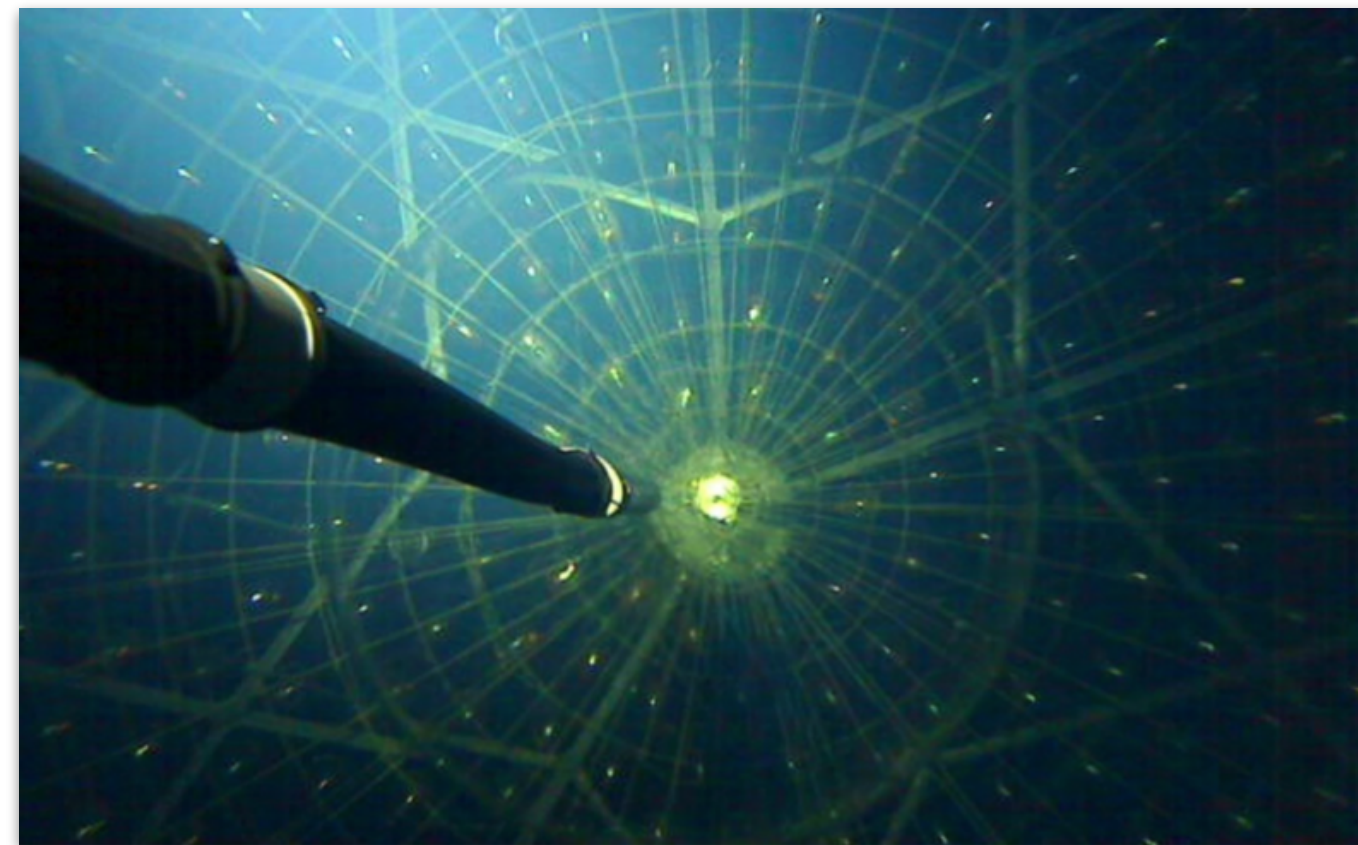
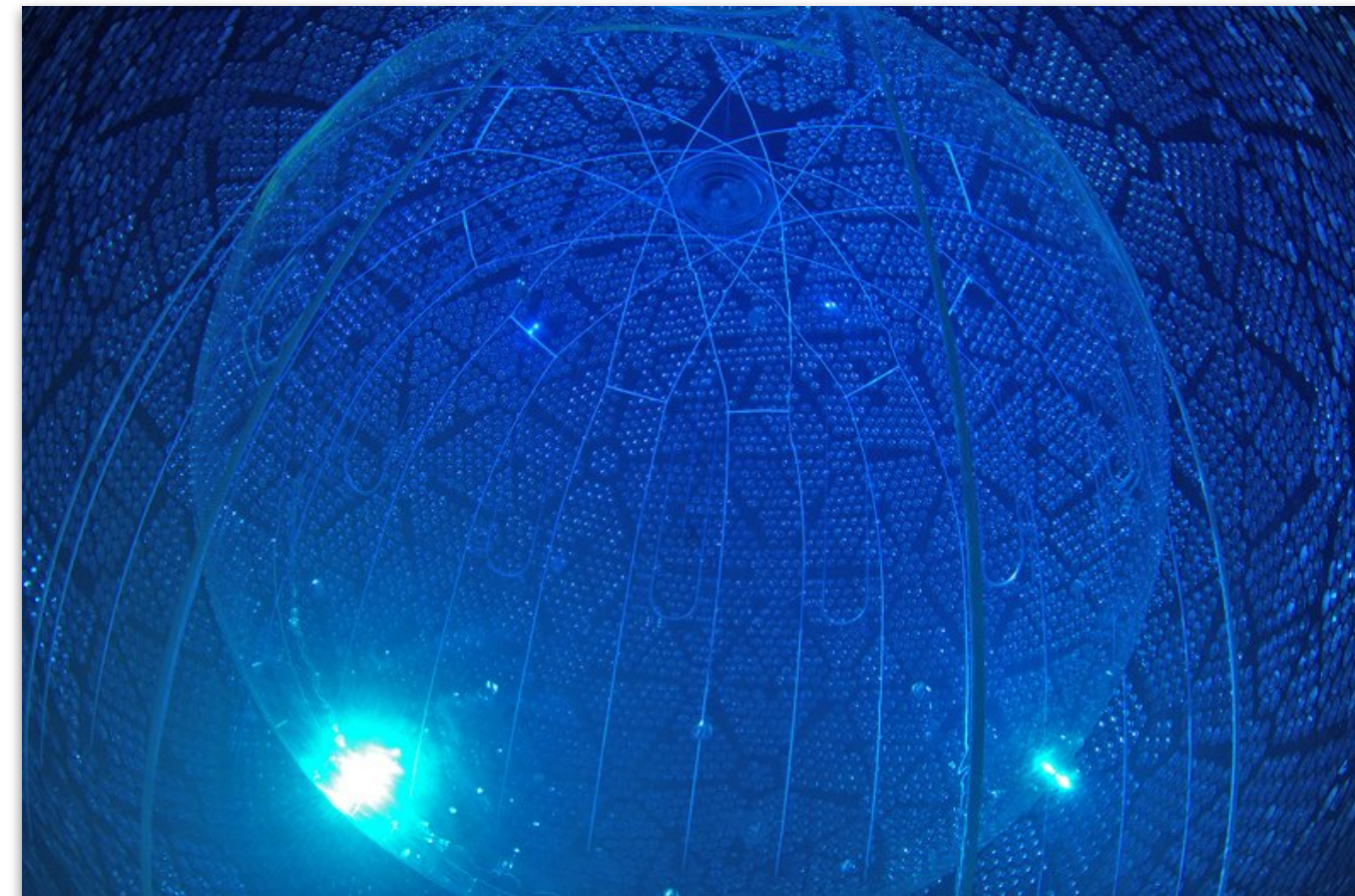


# The liquids

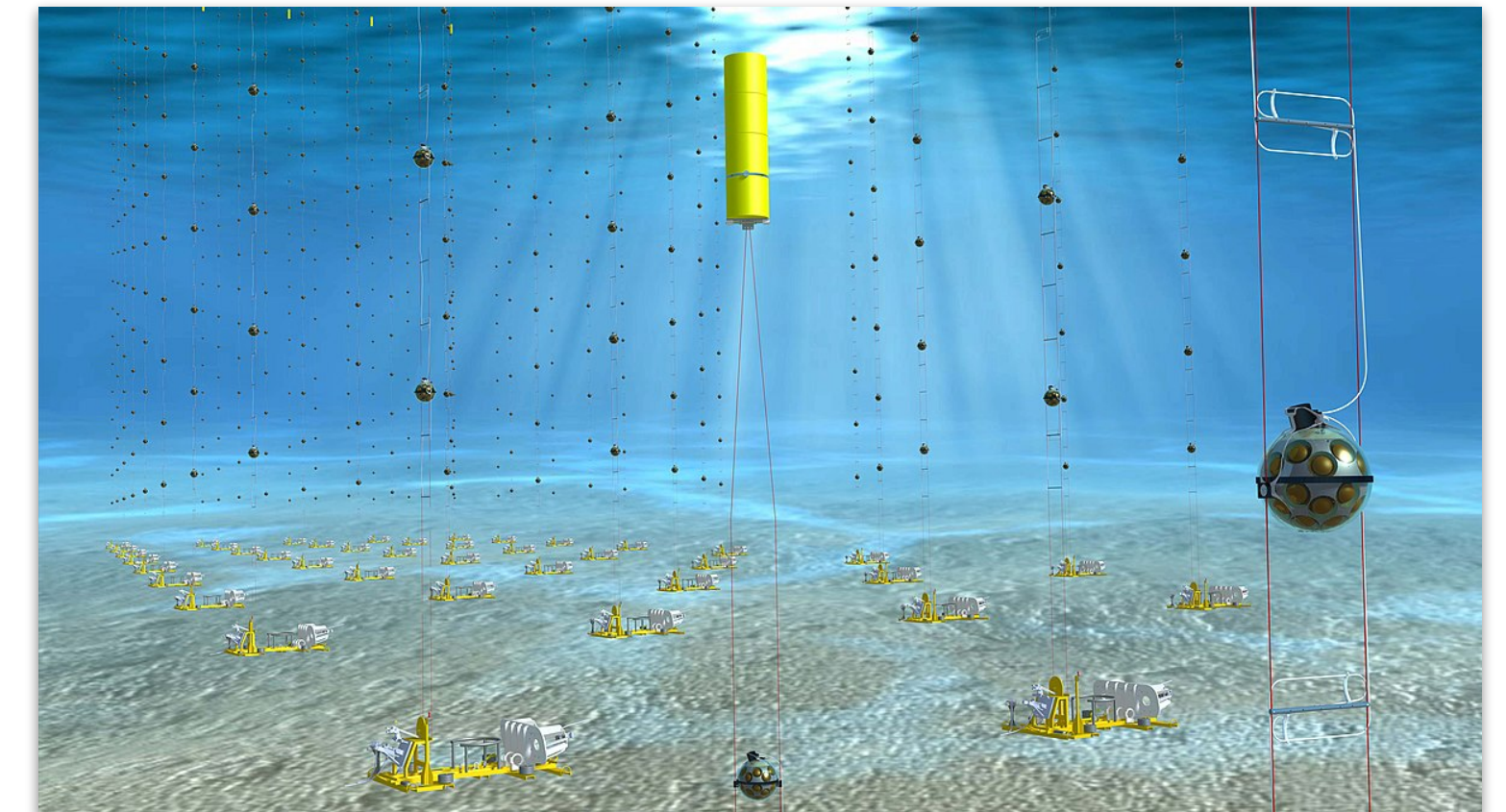
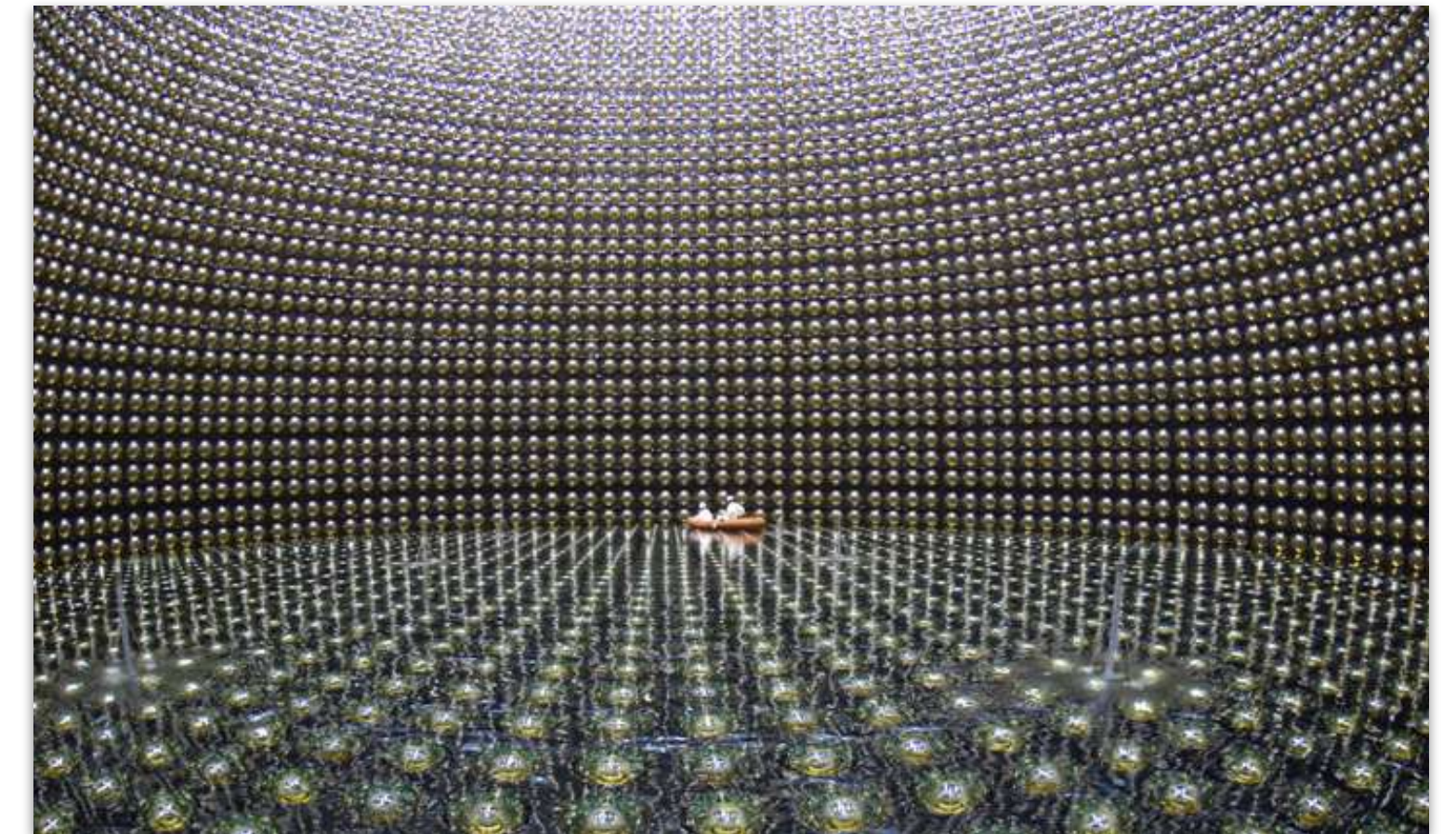
## Liquid Noble Gases



## Liquid Scintillator



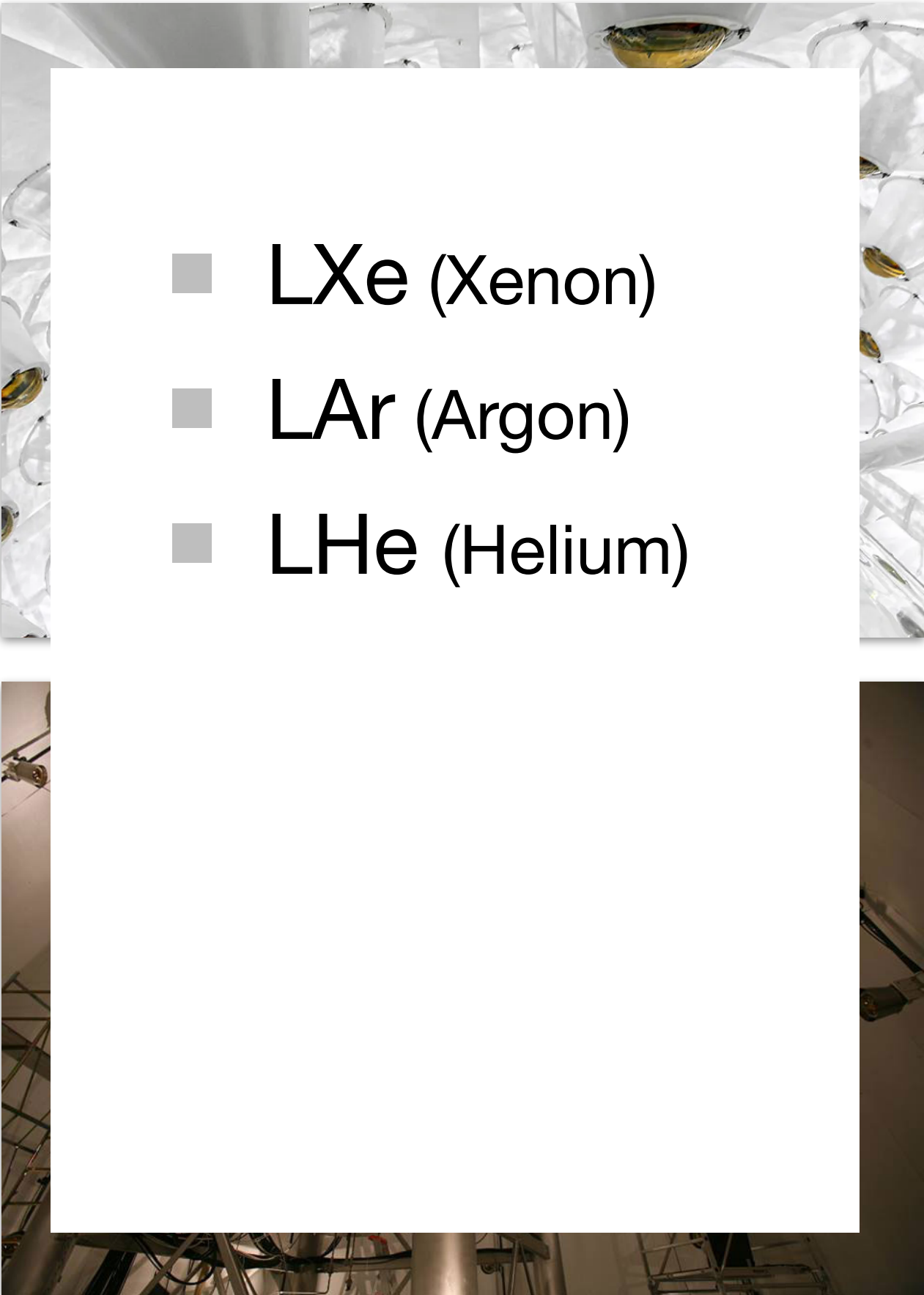
## Water



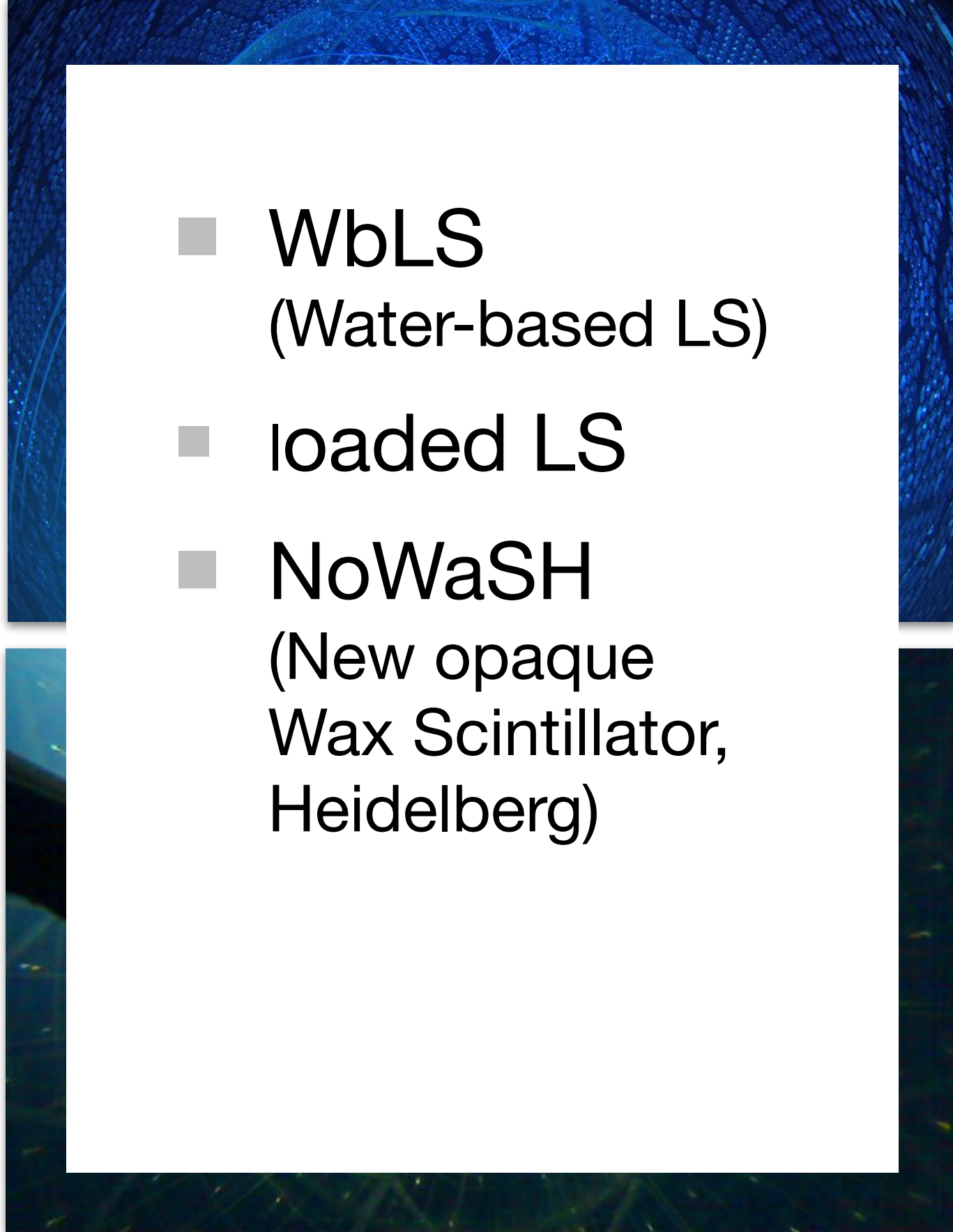


# The liquids

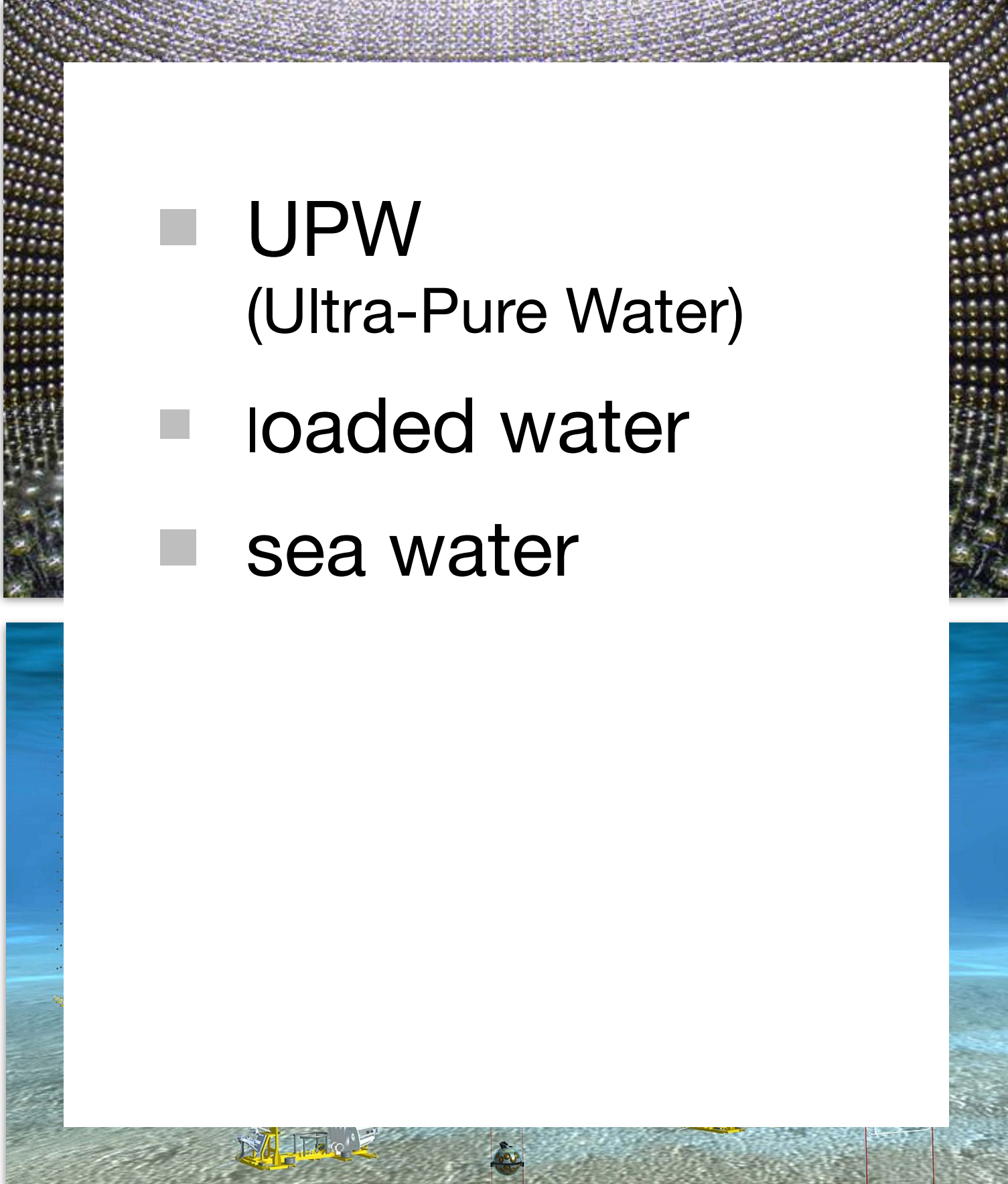
## Liquid Noble Gases

- 
- LXe (Xenon)
  - LAr (Argon)
  - LHe (Helium)

## Liquid Scintillator

- 
- WbLS  
(Water-based LS)
  - loaded LS
  - NoWaSH  
(New opaque  
Wax Scintillator,  
Heidelberg)

## Water

- 
- UPW  
(Ultra-Pure Water)
  - loaded water
  - sea water



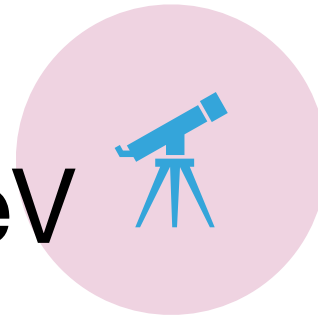


# Primary signals and their energy and time scales

## Liquid Noble Gases

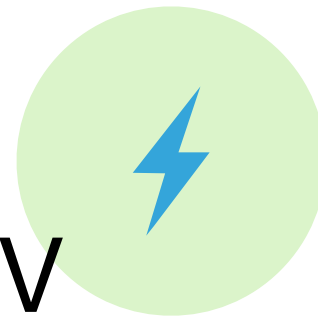
### ■ Atomic excitation → dimer formation → dimer decay

- after  $O(\text{ns} - \mu\text{s})$
- $O(10^4)$  photons / MeV
- VUV - UV



### ■ Ionization

- instantaneous
- $O(10^4)$  charges / MeV



### ■ Quasi-particle creation

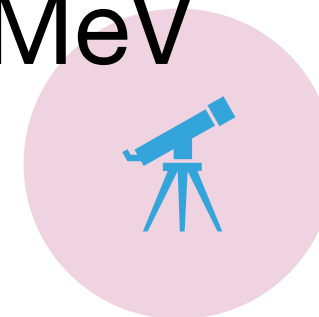
- instantaneous
- $O(10^{5-6})$  quasi-particles / keV



## Liquid Scintillator

### ■ Atomic excitation / ionization → fluorescence and phosphorescence

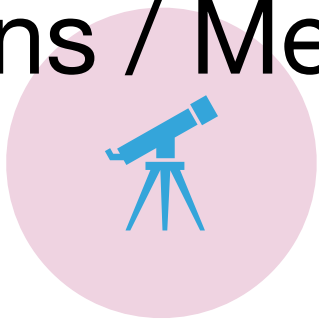
- after  $O(\text{ns} - \mu\text{s})$
- $O(10^4)$  photons / MeV
- UV - visible light



## Water

### ■ Cherenkov light

- instantaneous
- $O(10^2)$  photons / MeV
- visible light



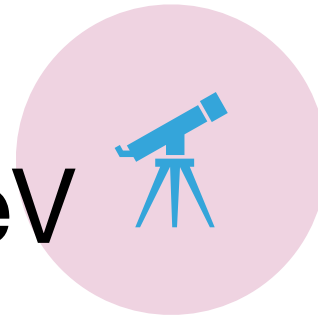


# Primary signals and their energy and time scales

## Liquid Noble Gases

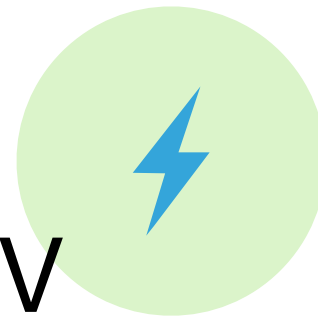
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### ■ Quasi-particle creation

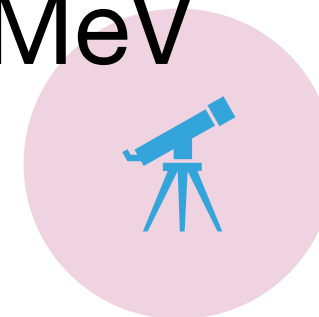
- instantaneous
- $O(10^{5-6})$  quasi-particles / keV



## Liquid Scintillator

### ■ Atomic excitation / ionization → fluorescence and phosphorescence

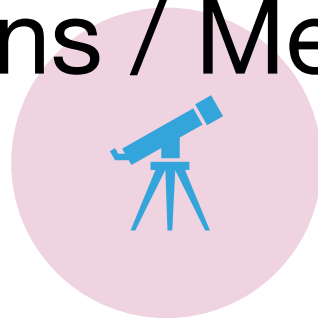
- after  $O(\text{ns} - \mu\text{s})$
- $O(10^4)$  photons / MeV
- UV - visible light



## Water

### ■ Cherenkov light

- instantaneous
- $O(10^2)$  photons / MeV
- visible light



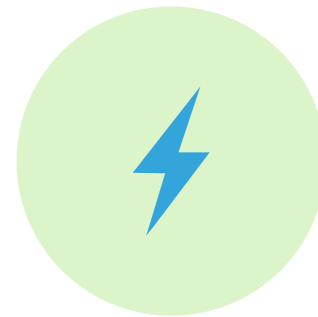
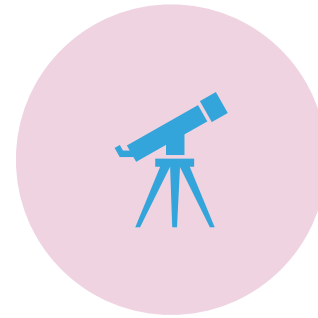
NB: signal collection times can be as slow as ms scale!



# Primary signals and their energy and time scales

## Liquid Noble Gases

light + charge  
or  
light + phonon  
readout



## Liquid Scintillator

## Water

light readout

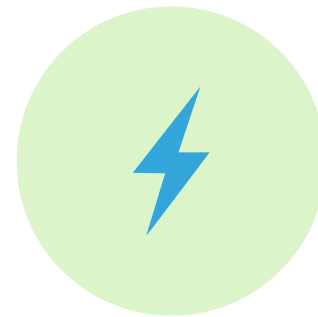
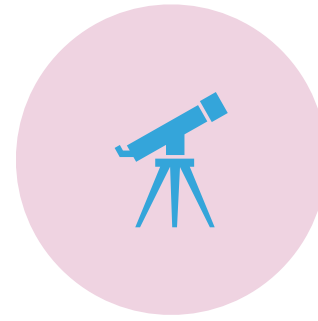


NB: signal collection times  
can be as slow as ms scale!

# Primary signals and their energy and time scales

## Liquid Noble Gases

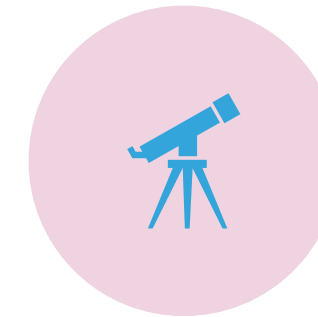
light + charge  
or  
light + phonon  
readout



## Liquid Scintillator

## Water

light readout

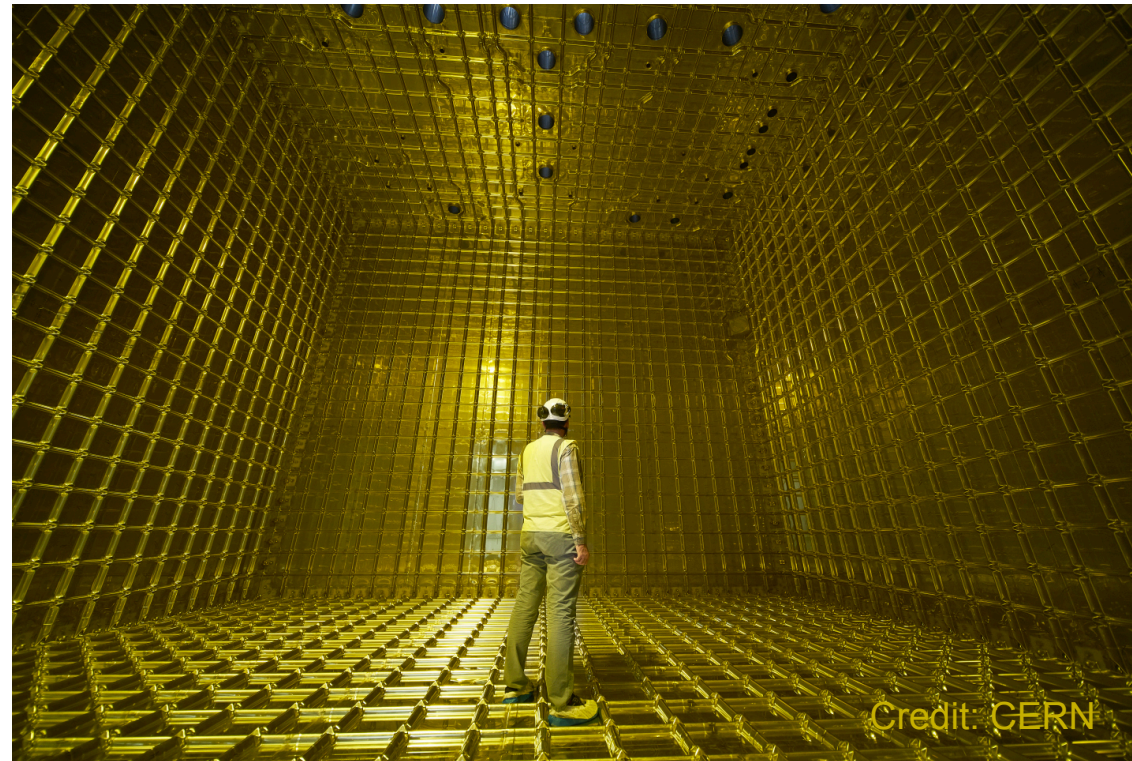


The liquids are used both  
as main target and for  
veto detectors.

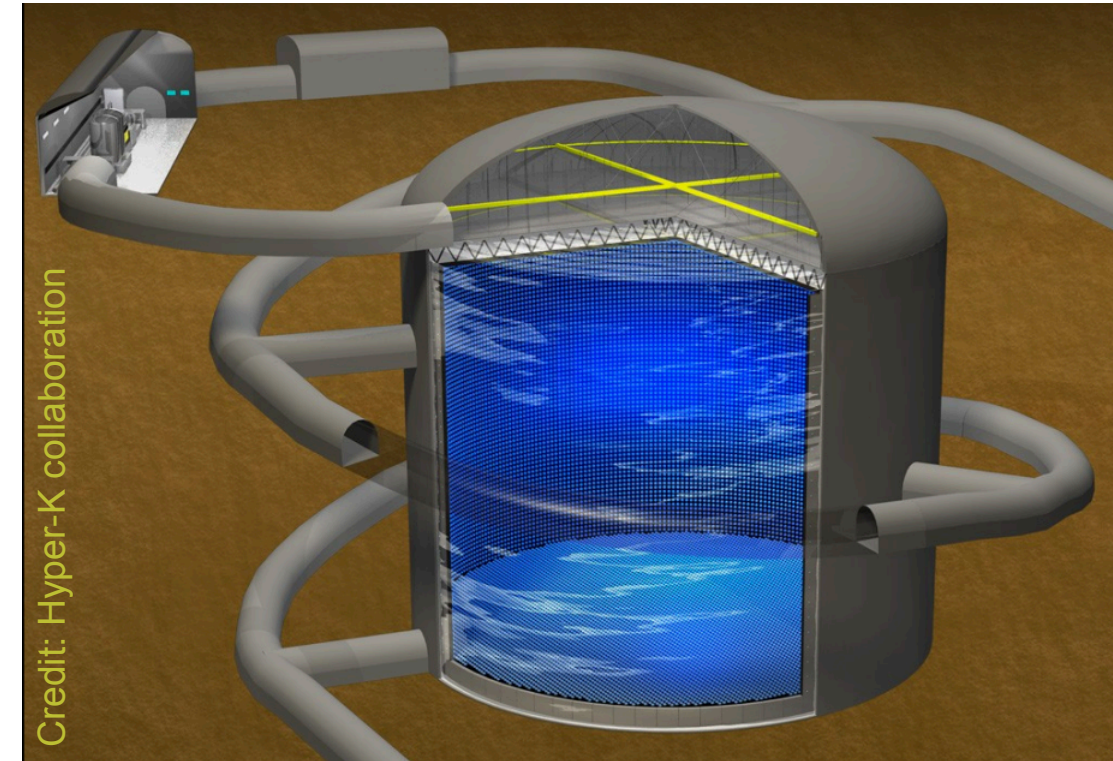


# Categorization of main areas for R&D streamlining

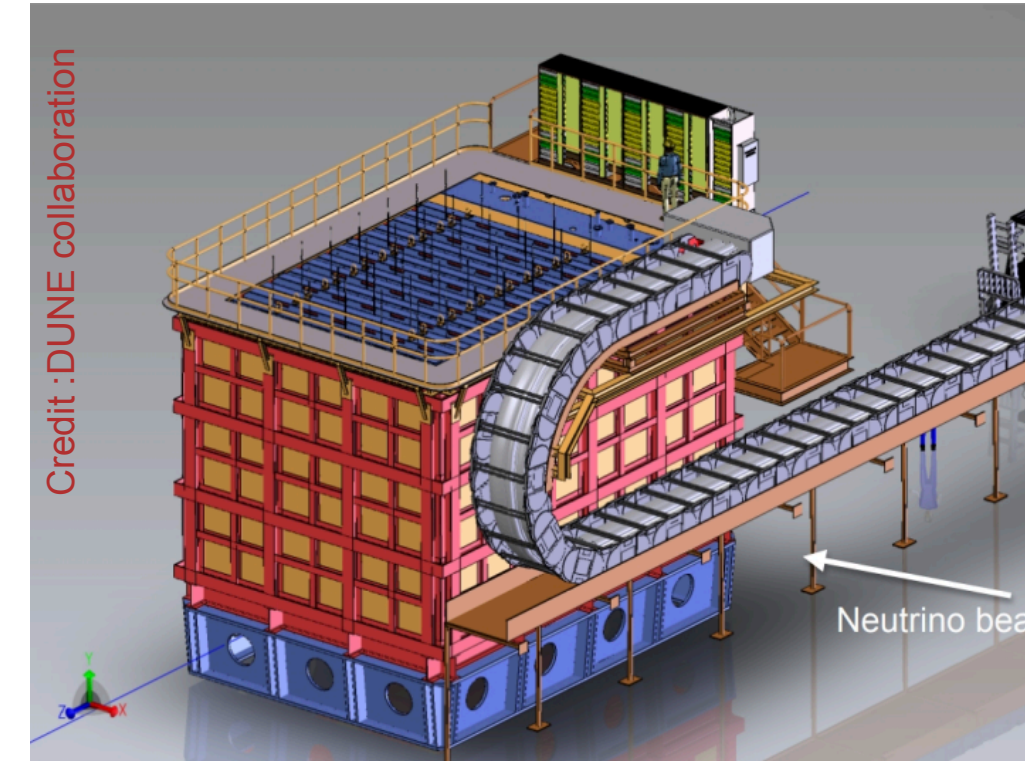
DUNE programme



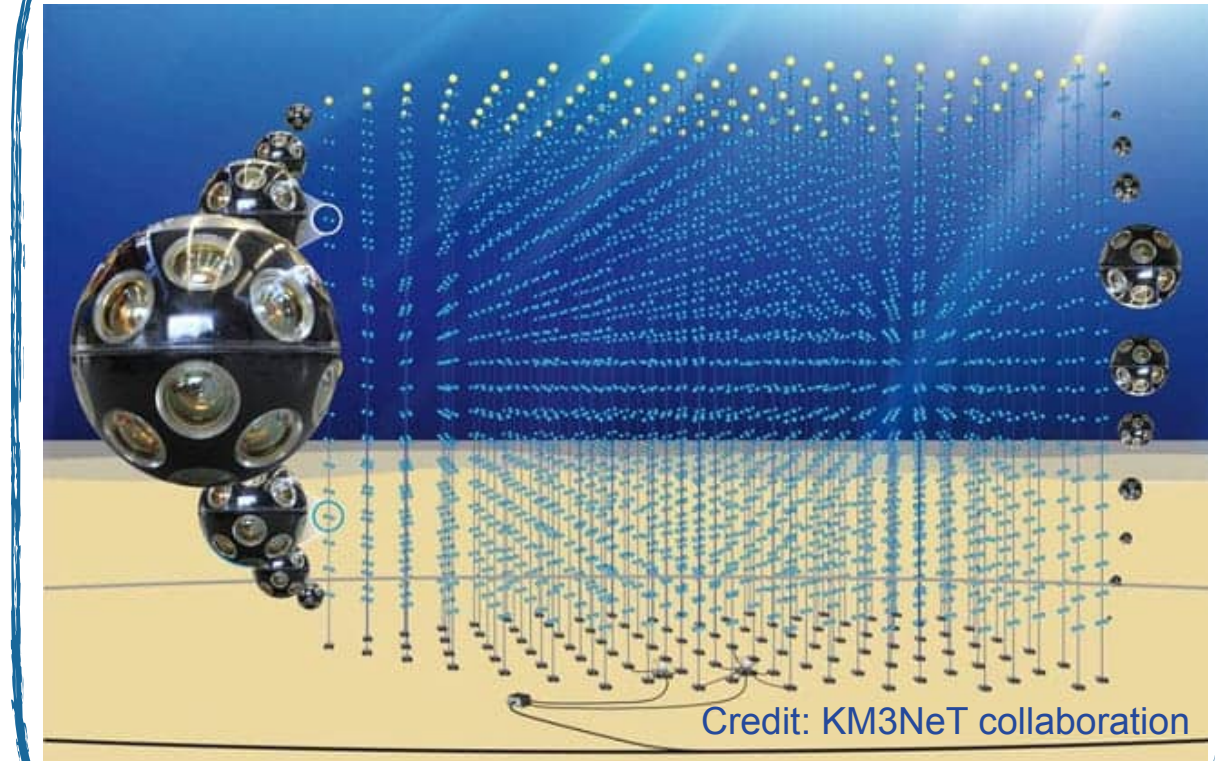
Hyper-K programme



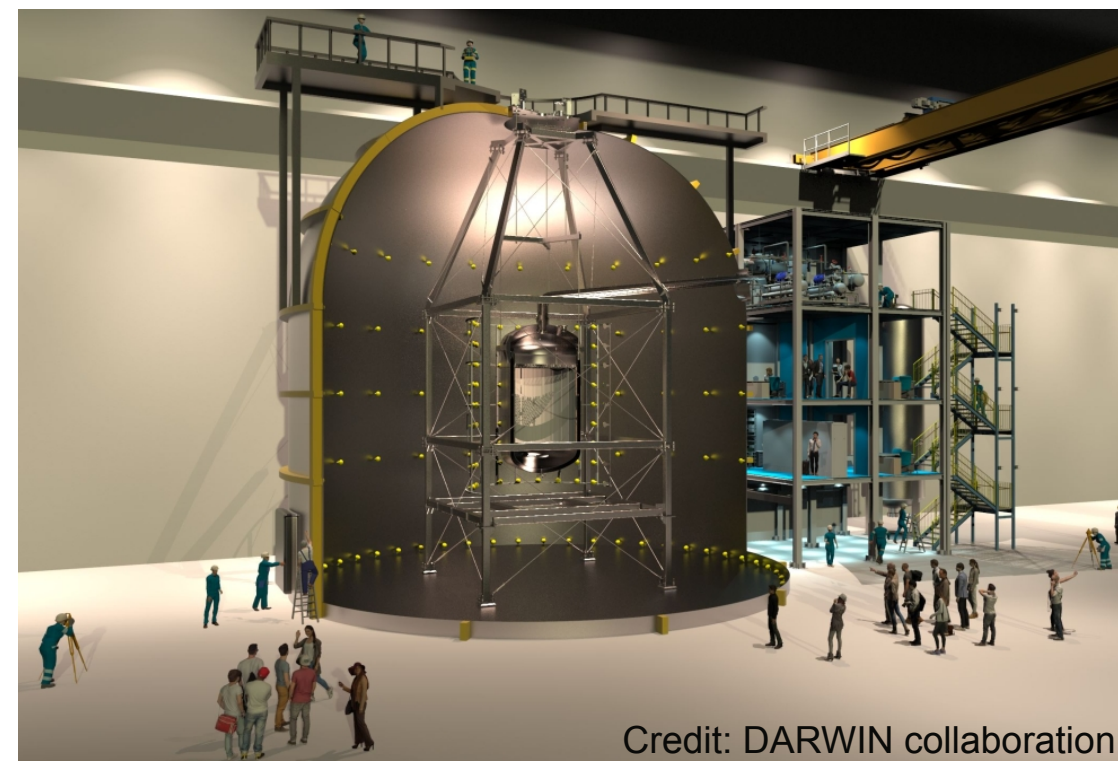
neutrino near detectors



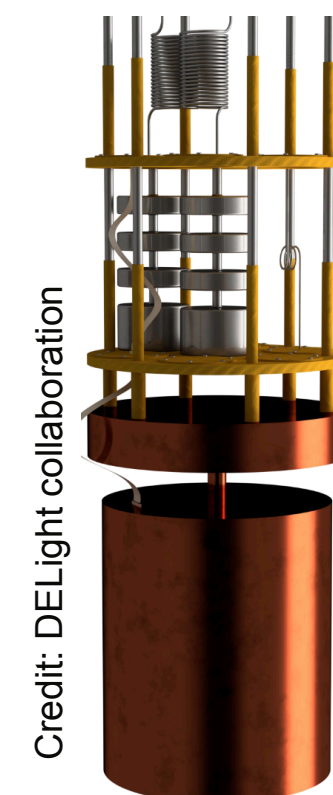
neutrino telescopes



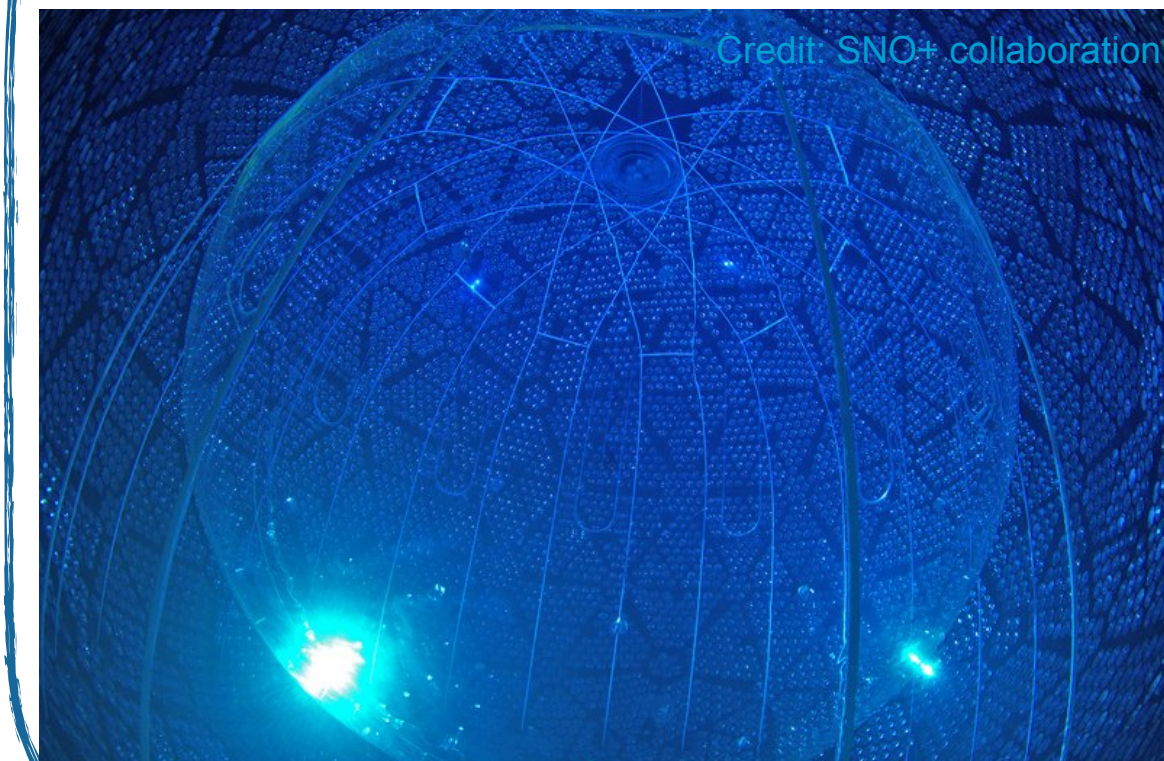
multi-tonne scale  
dark matter detectors



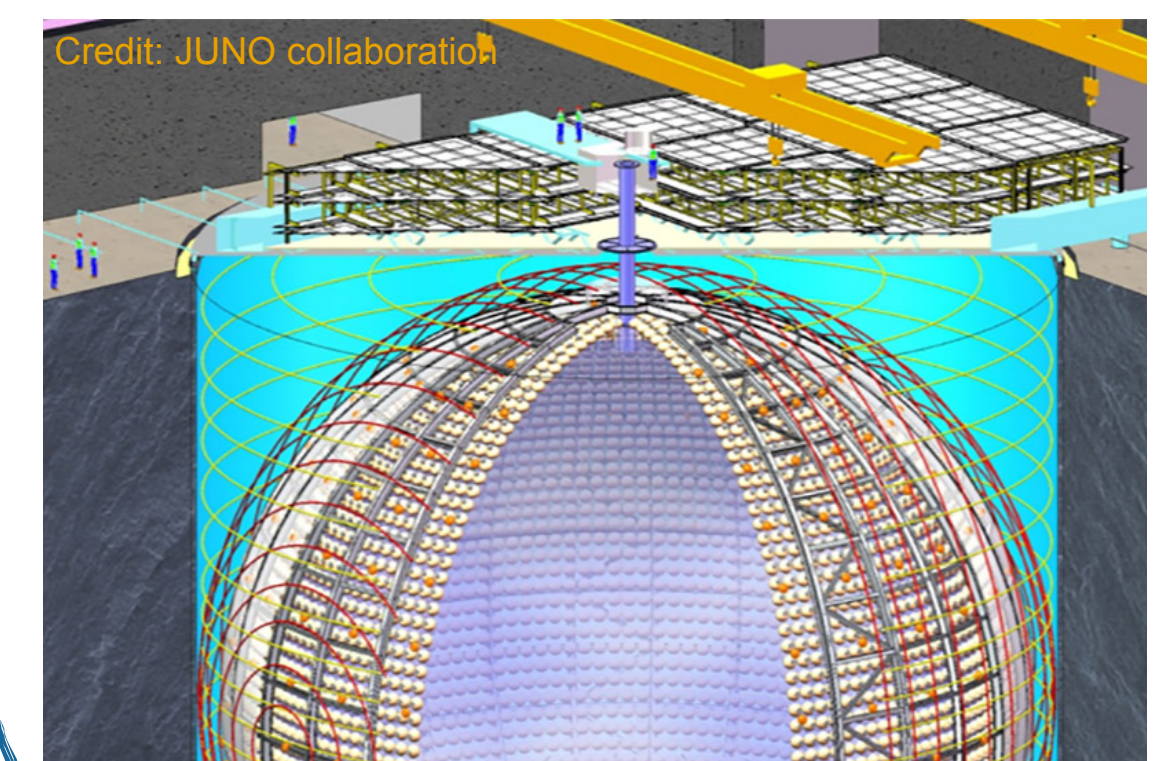
light dark matter  
detectors (g- to kg-scale)



tonne-scale  
 $0\nu\beta\beta$  experiments

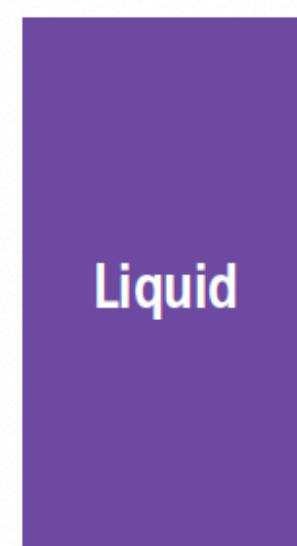
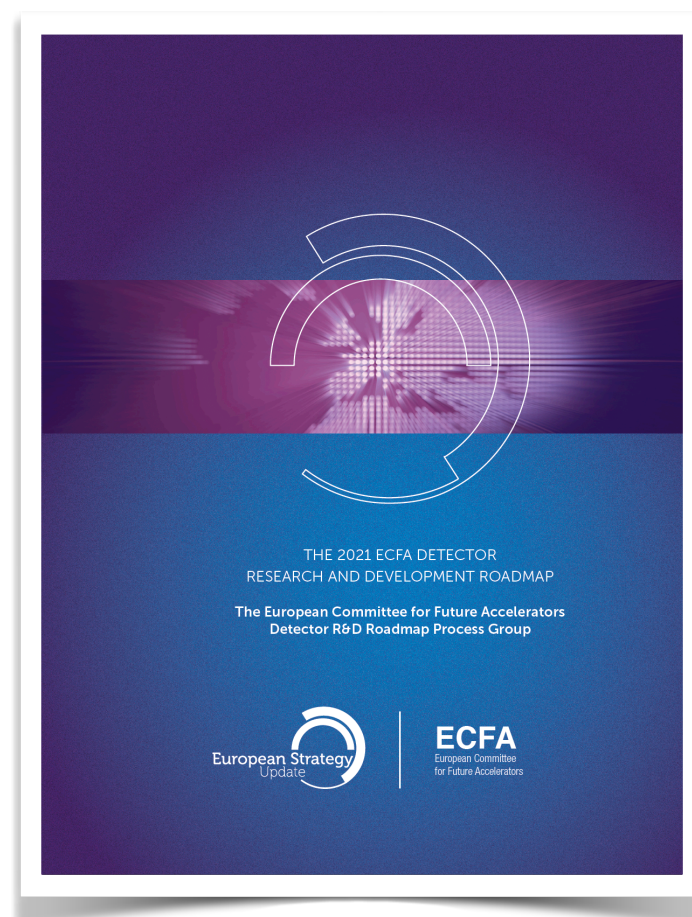


low-energy scintillator  
neutrino detectors

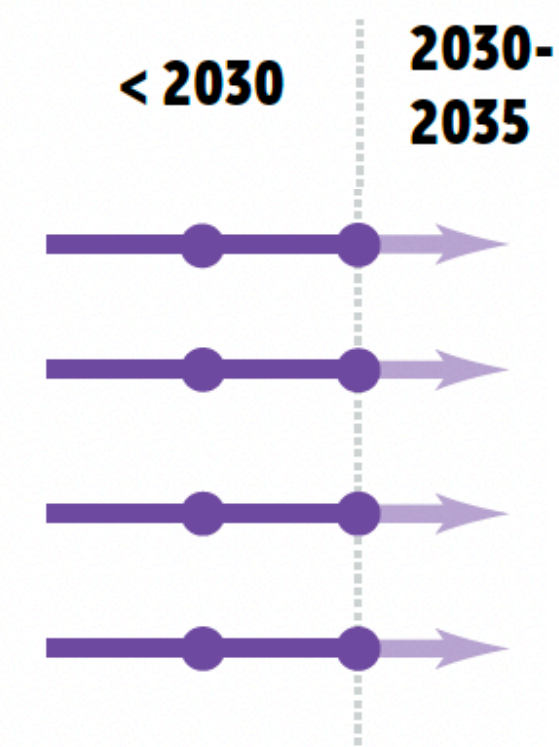




# Major detector R&D themes for liquid detectors



- DRDT 2.1** Develop readout technology to increase spatial and energy resolution for liquid detectors
- DRDT 2.2** Advance noise reduction in liquid detectors to lower signal energy thresholds
- DRDT 2.3** Improve the material properties of target and detector components in liquid detectors
- DRDT 2.4** Realise liquid detector technologies scalable for integration in large systems



## ■ Readout development

- Develop readout technology to increase **spatial and energy resolution**

## ■ Measurement strategy

- Advance **noise** reduction to lower **energy thresholds**

## ■ Target properties

- Improve the **material properties** of target and detector components

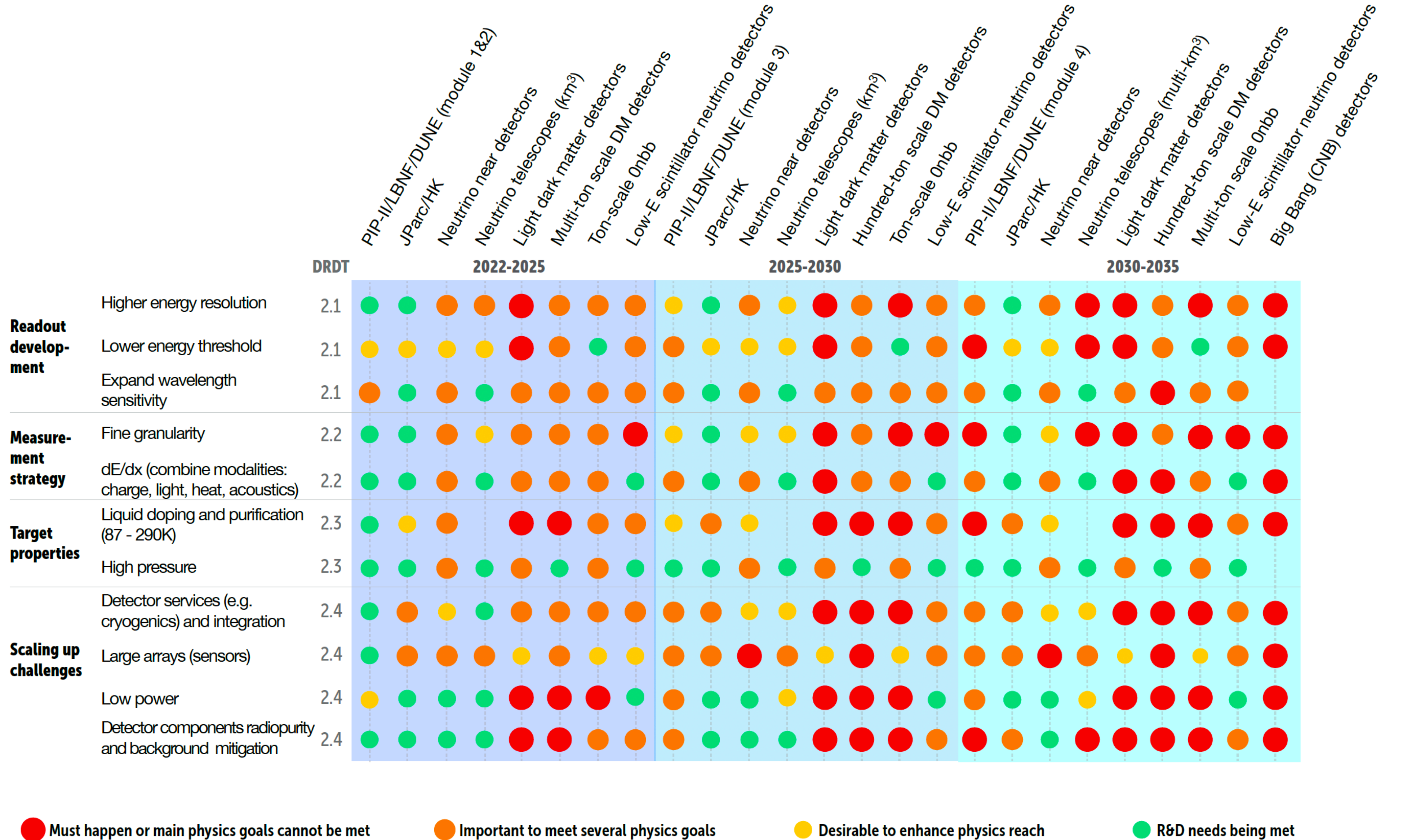
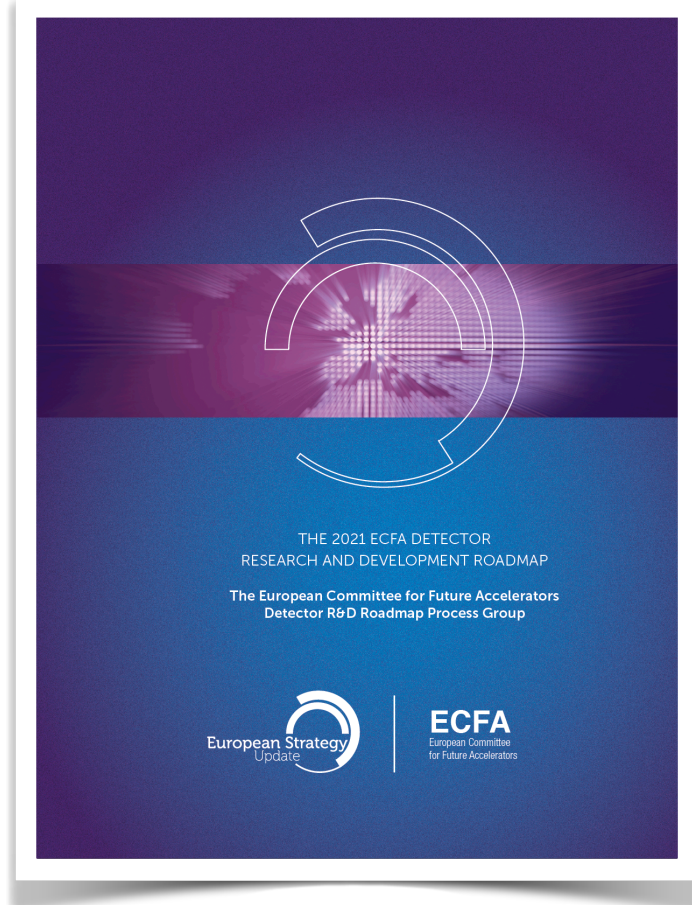
## ■ Scaling up challenges

- Realize technologies **scalable** for integration in large systems





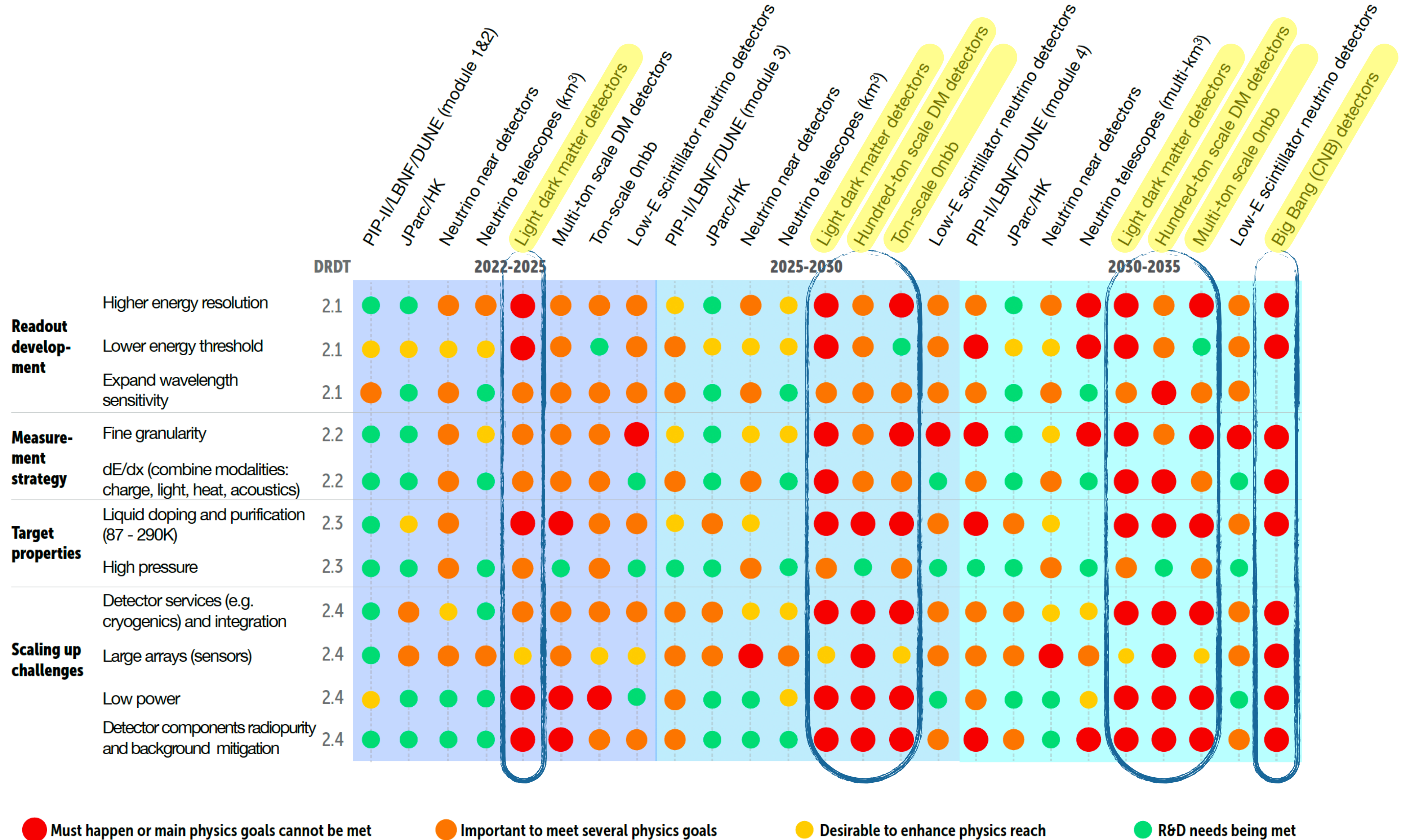
# Major detector R&D themes for liquid detectors





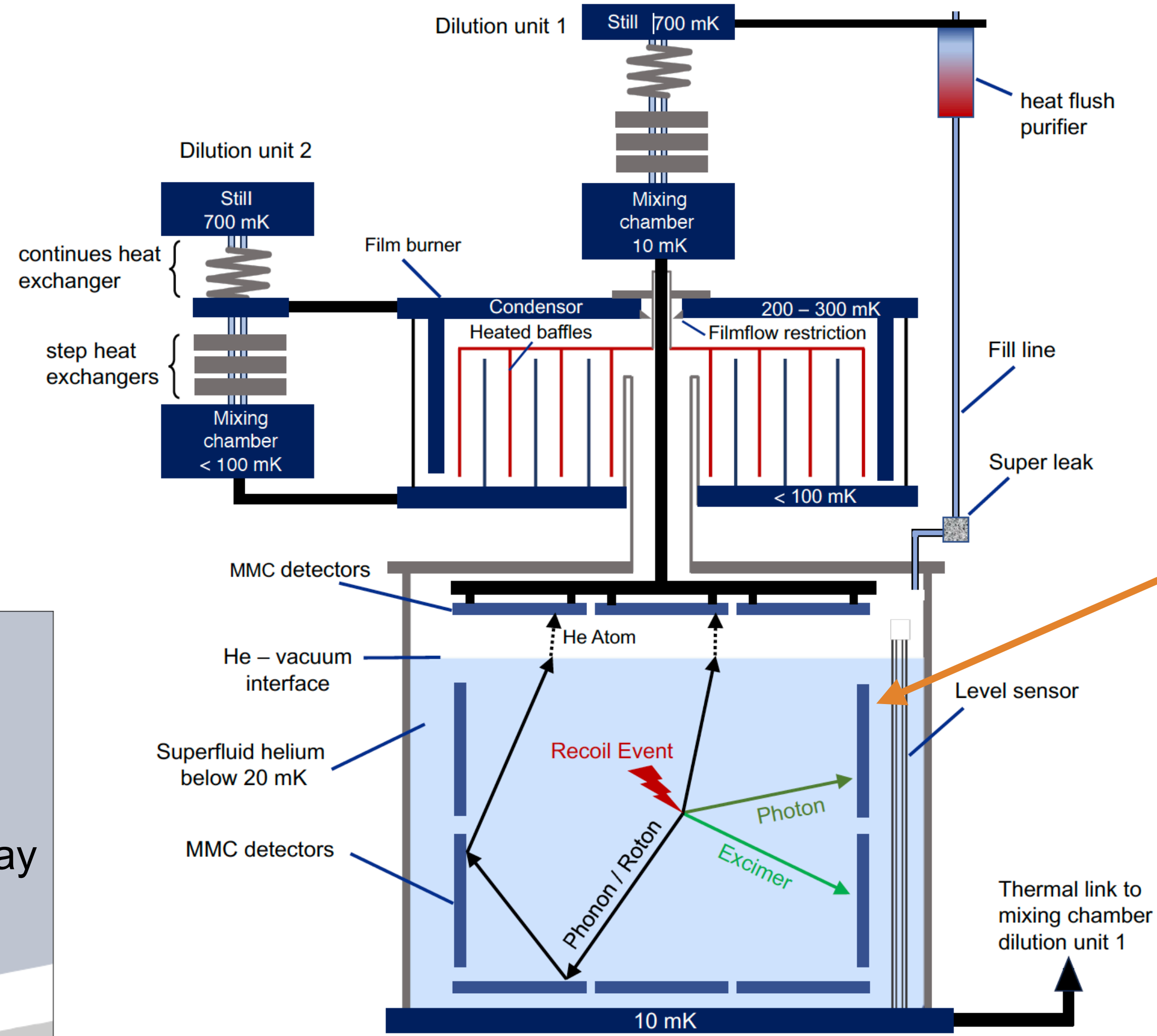
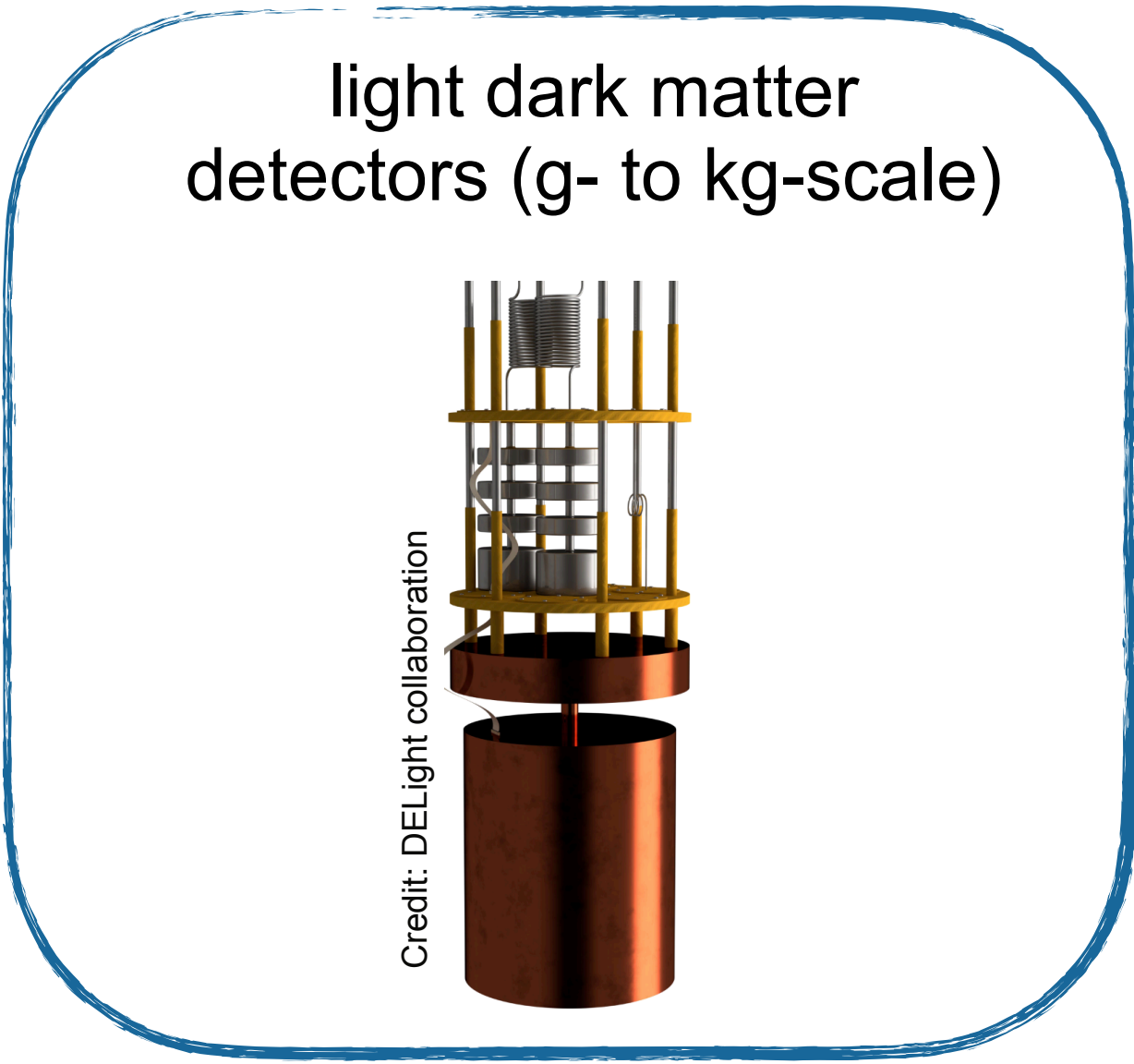


# Major detector R&D themes for liquid detectors

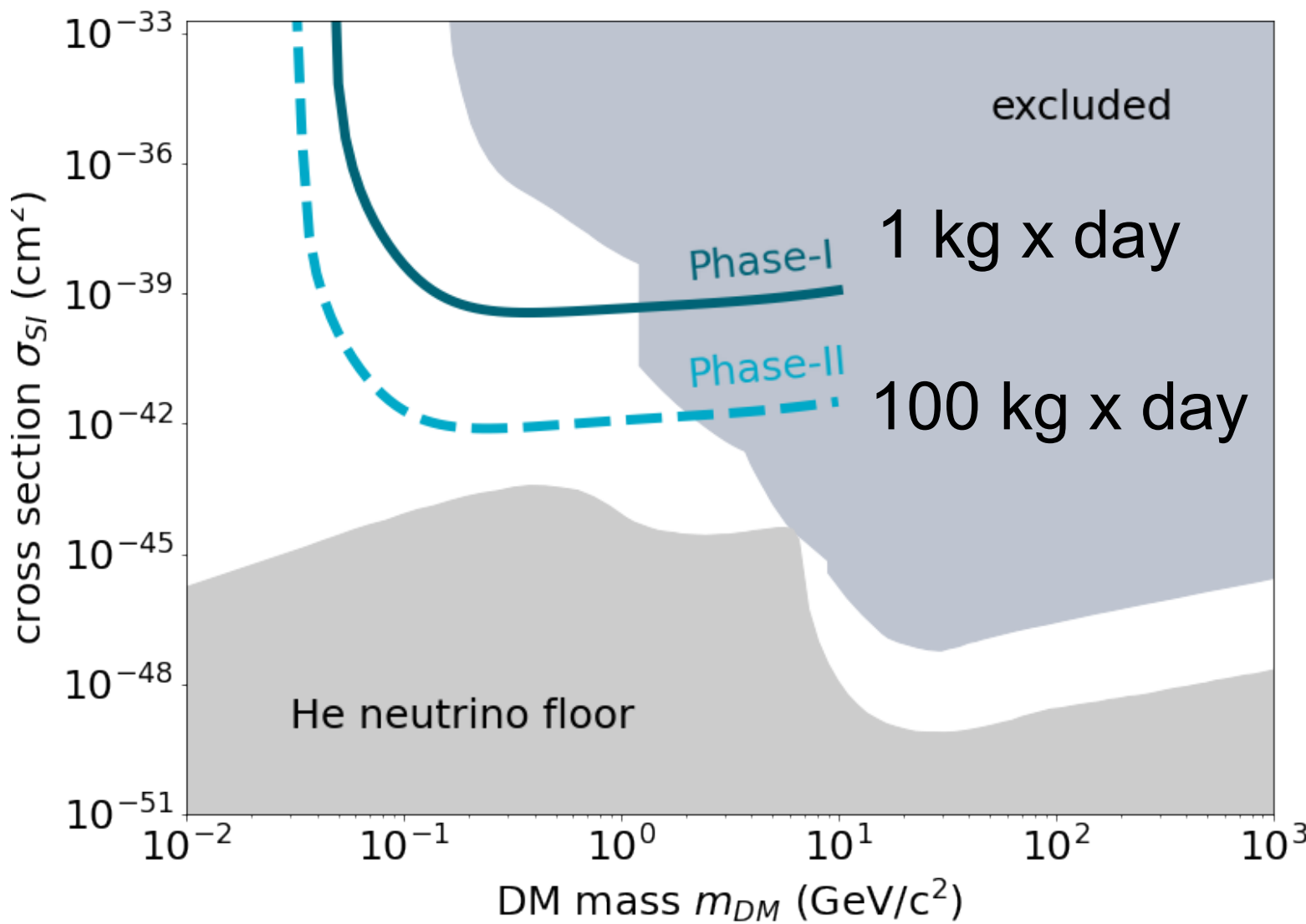
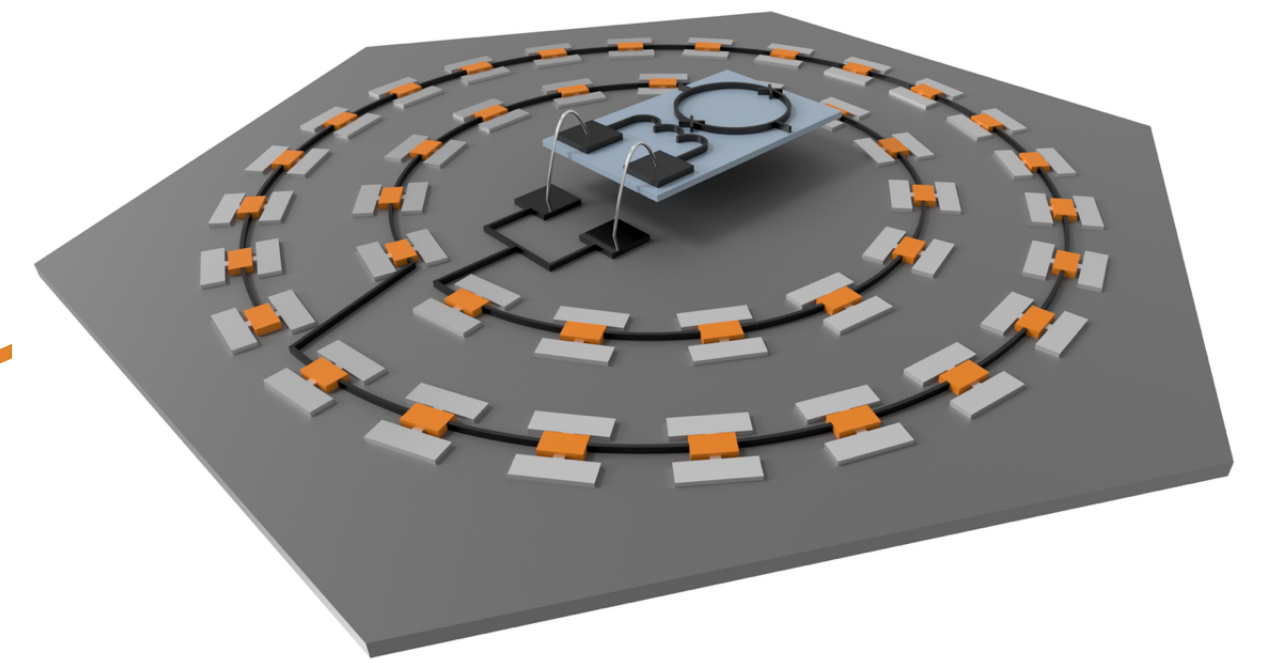




# Light dark matter detectors



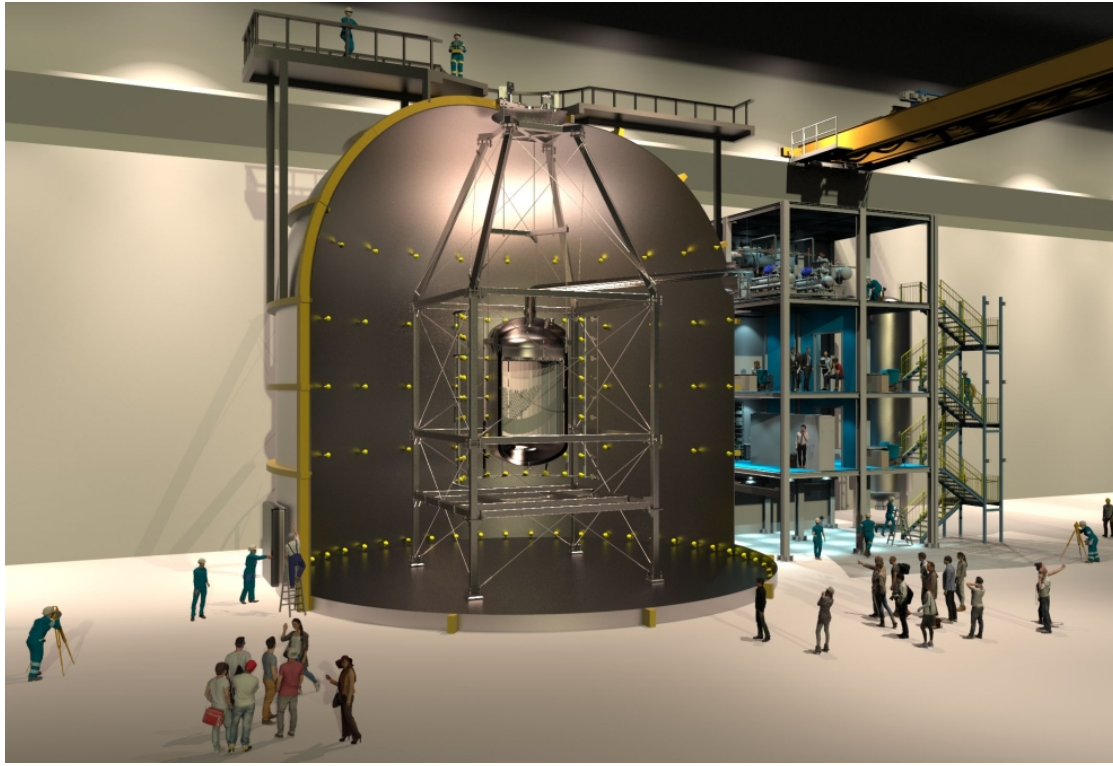
exploration of sub-GeV dark matter mass region with superfluid  $^4\text{He}$  (LHe)





# Hundred-ton scale dark matter detectors

multi-tonne scale  
dark matter detectors



## Dark Matter

- Dark photons
- Axion-like particles
- Planck mass

## WIMPs

- Spin-independent
- Spin-dependent
- Sub-GeV
- Inelastic

## Sun

- pp neutrinos
- Solar metallicity
- ${}^7\text{Be}$ ,  ${}^8\text{B}$ , hep

## Neutrino Nature

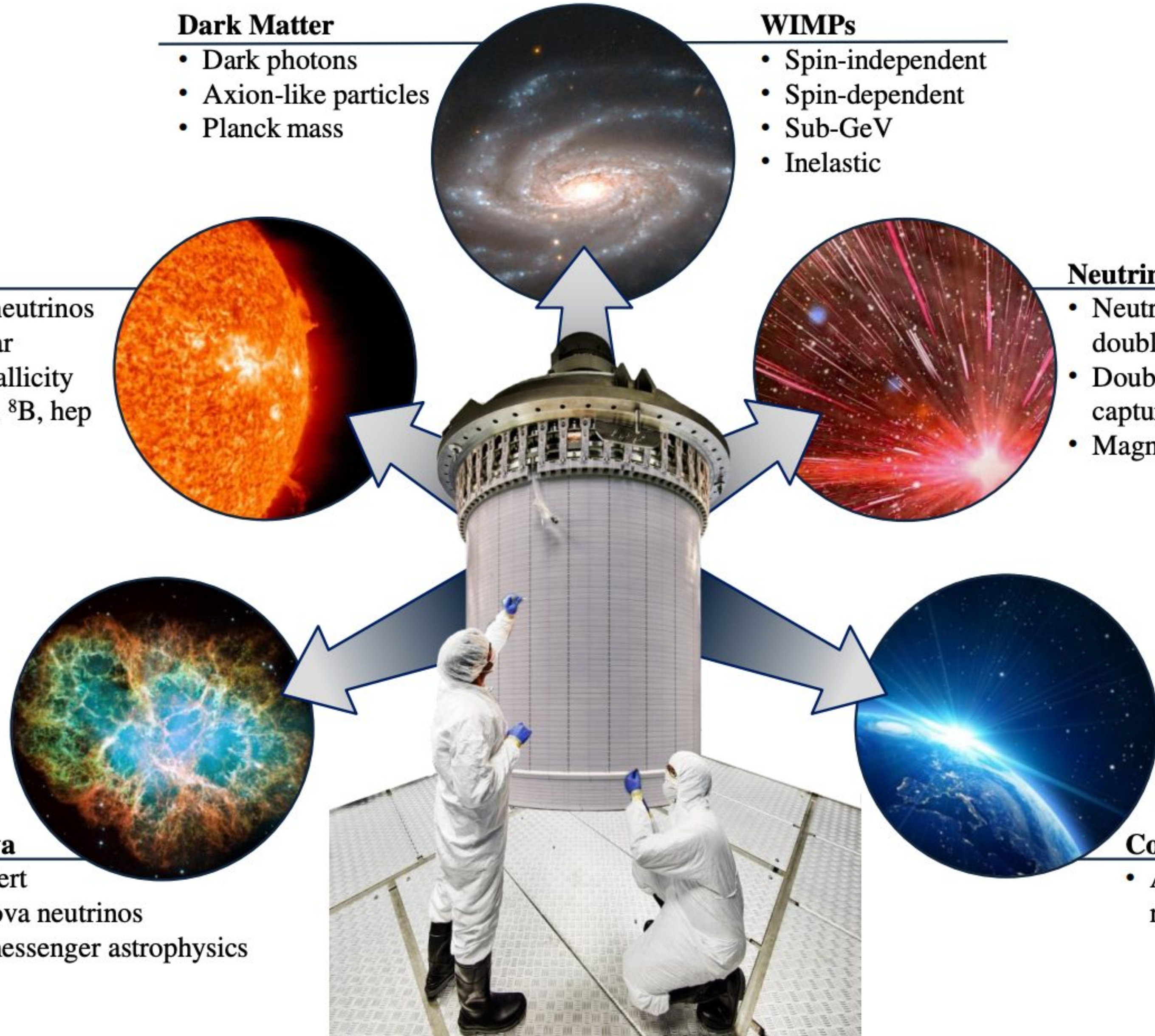
- Neutrinoless double beta decay
- Double electron capture
- Magnetic Moment

## Supernova

- Early alert
- Supernova neutrinos
- Multi-messenger astrophysics

## Cosmic Rays

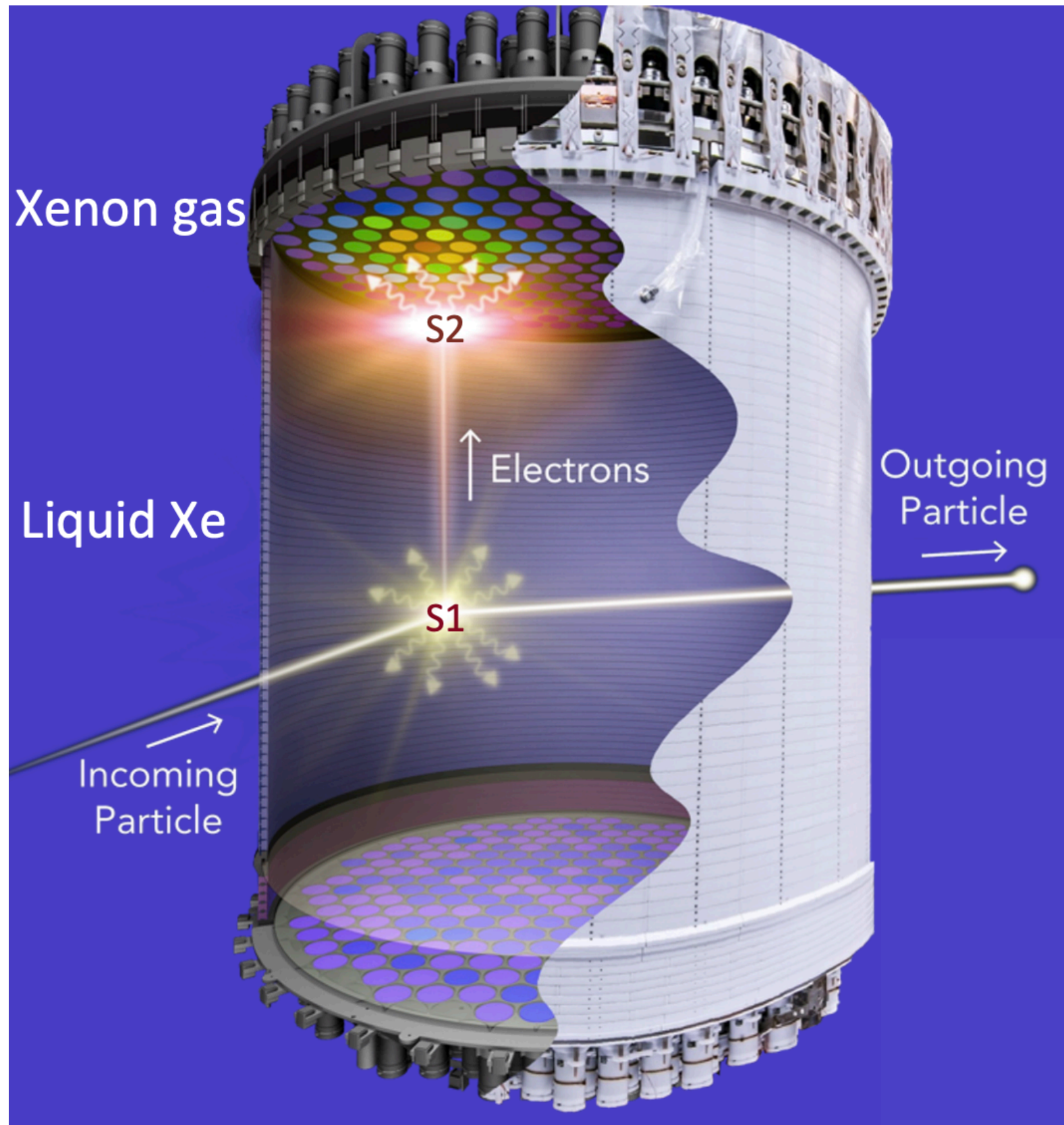
- Atmospheric neutrinos



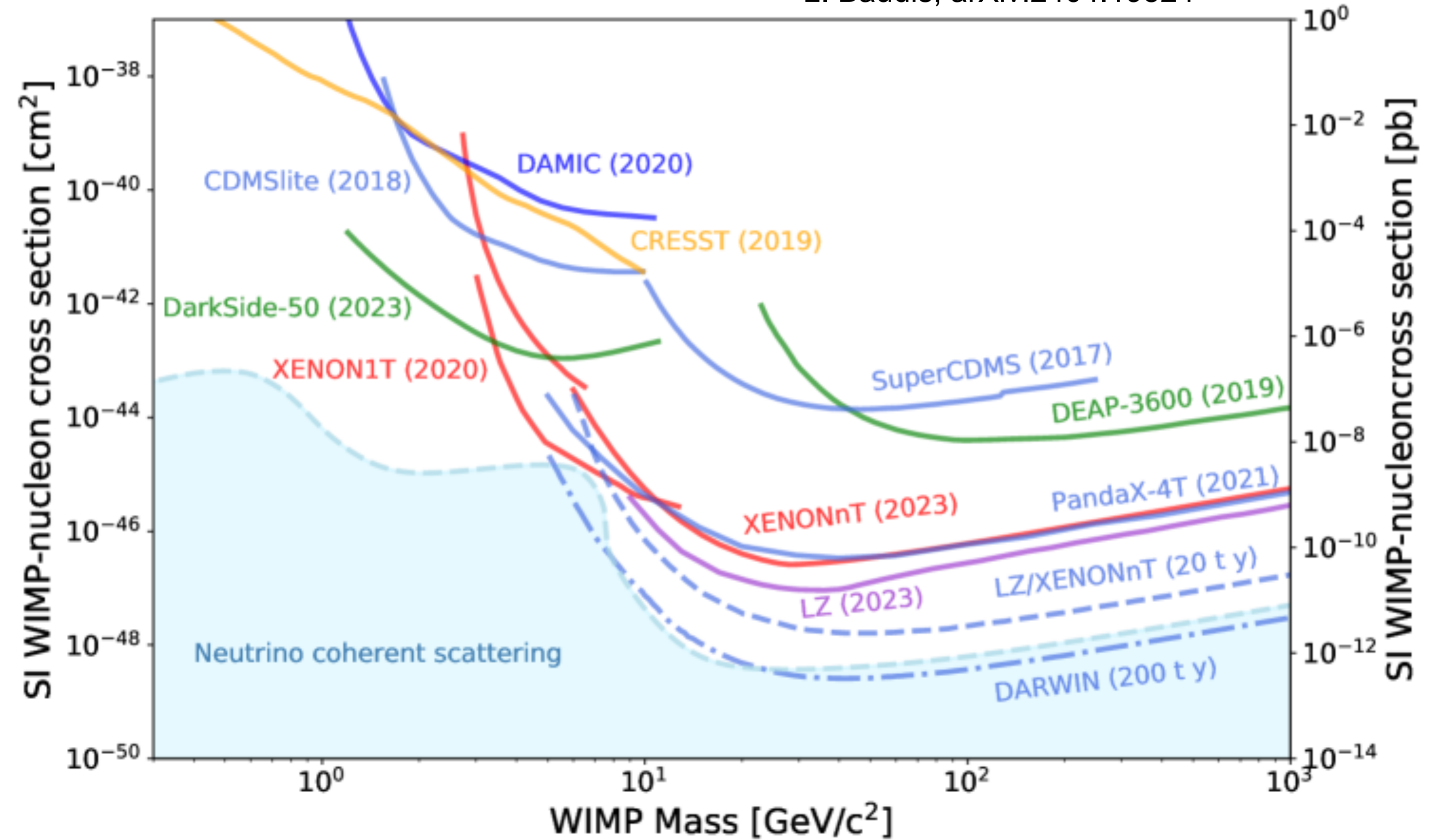


# Hundred-ton scale dark matter detectors

## Dual-phase xenon and argon TPCs



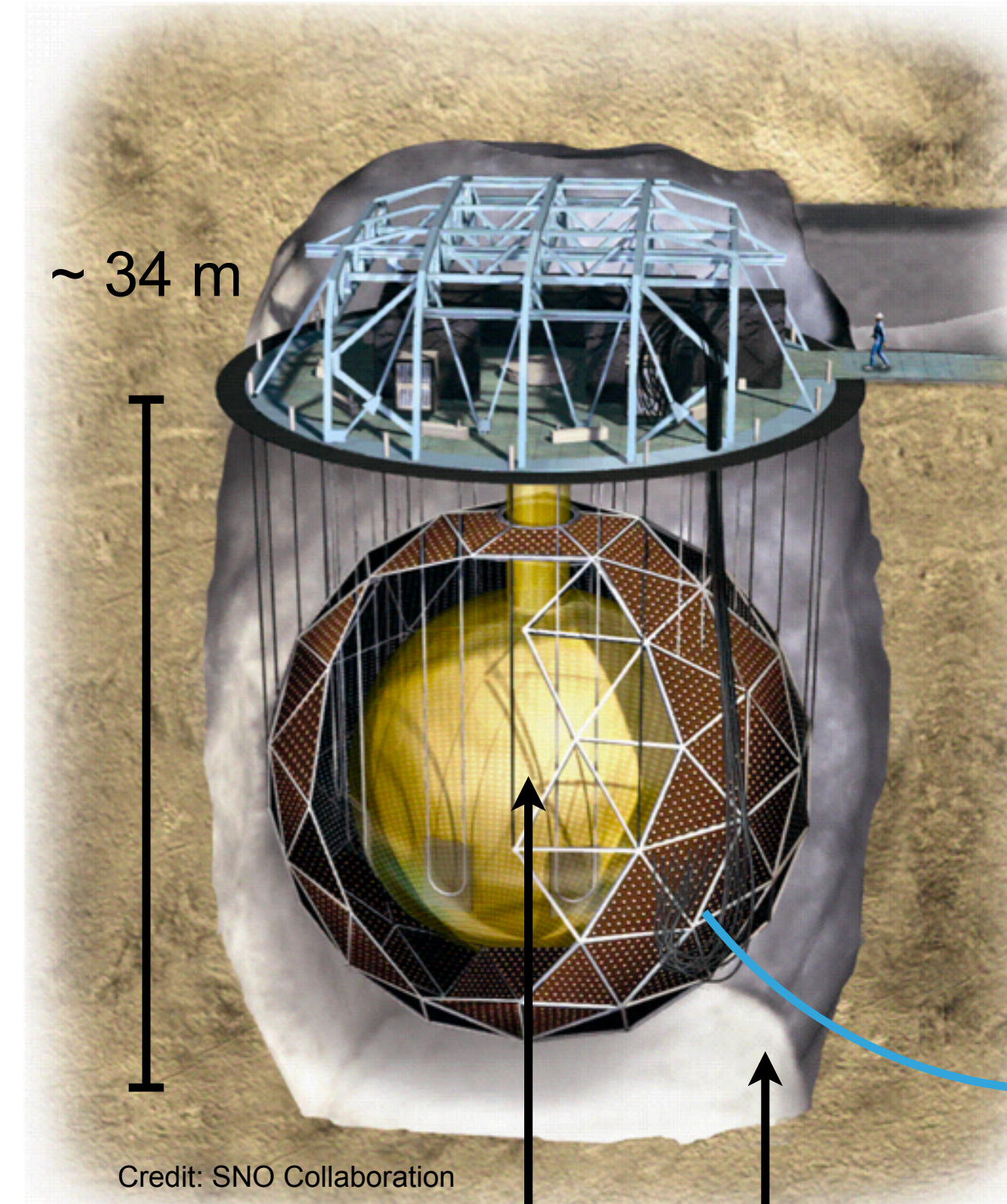
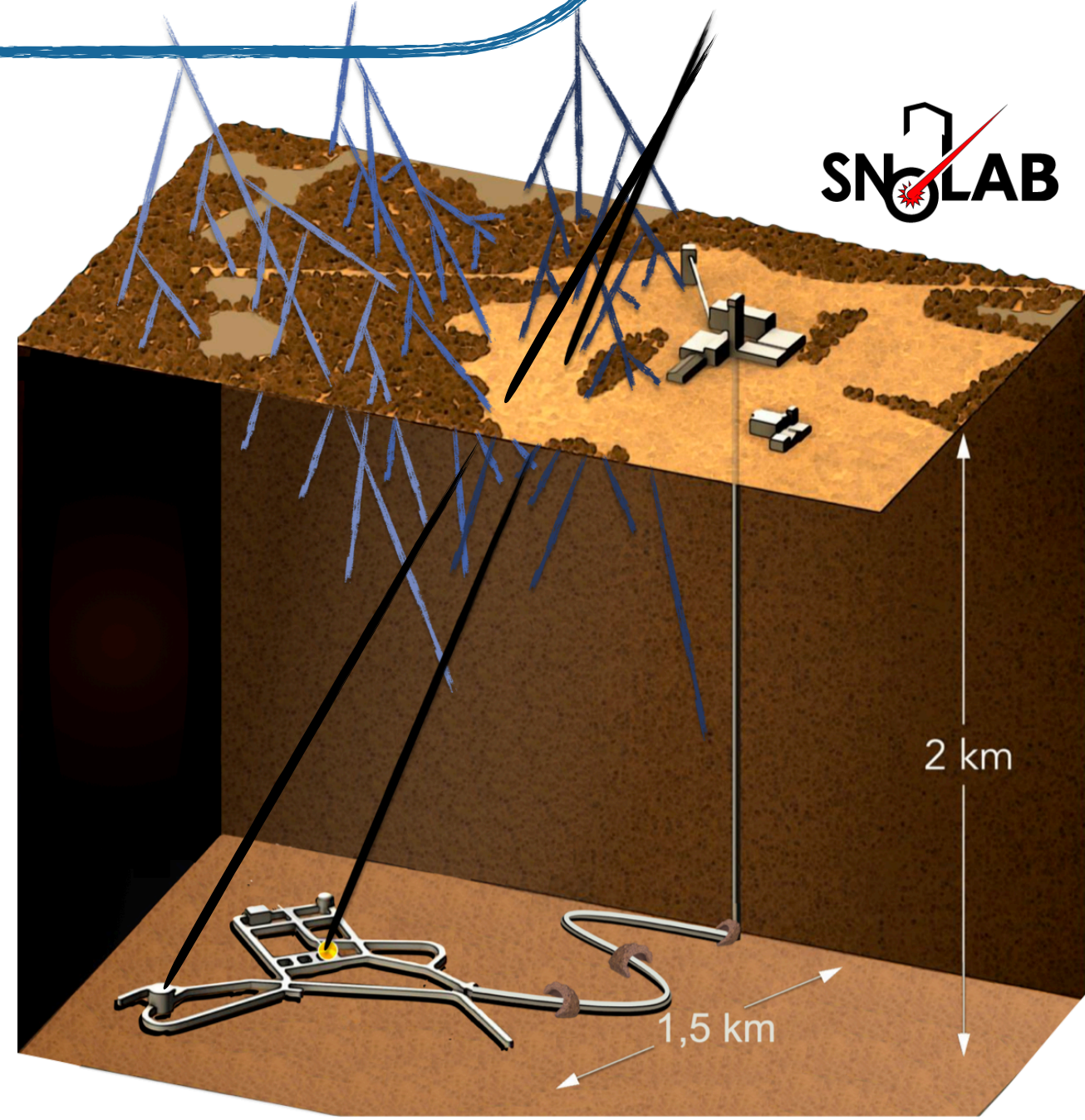
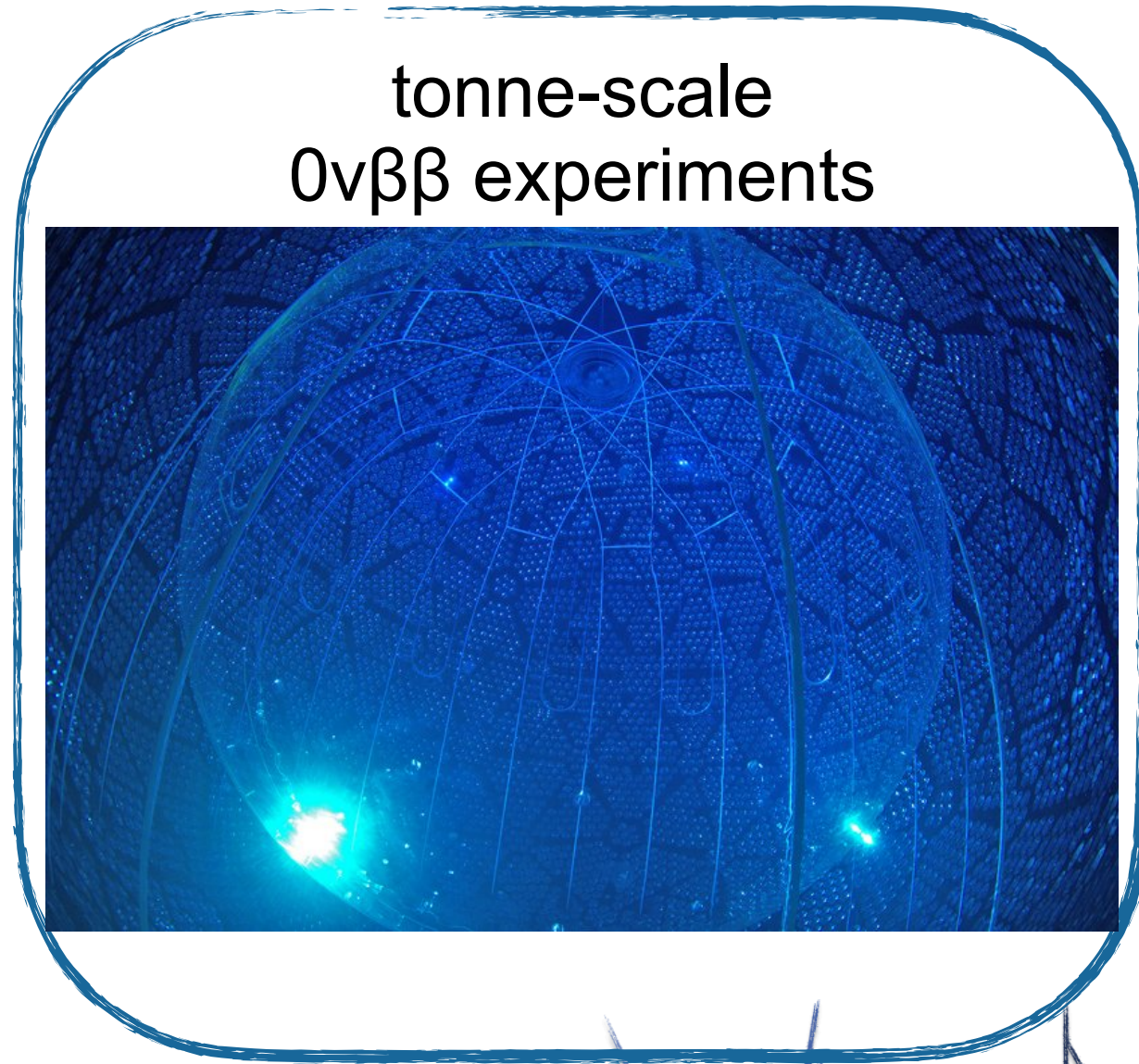
L. Baudis, arXiv:2404.19524



Scintillation light emitted by LAr and LXe: 128 nm and 178 nm, respectively

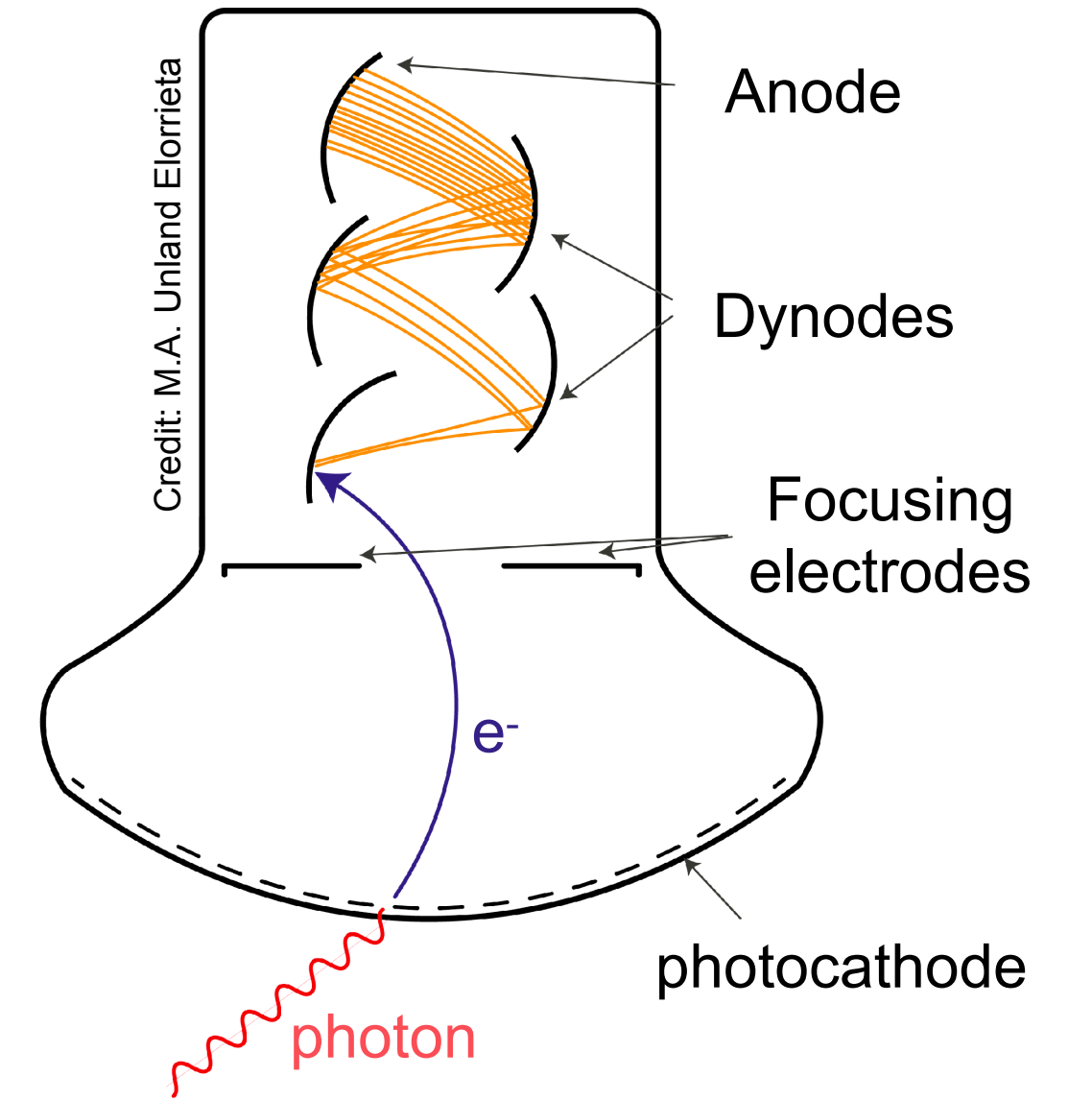


# Ton scale $0\nu\beta\beta$ detectors



SNO: 1 kt of  $D_2O$   
SNO+: 0.78 kt of liquid scintillator

7 kt  $H_2O$

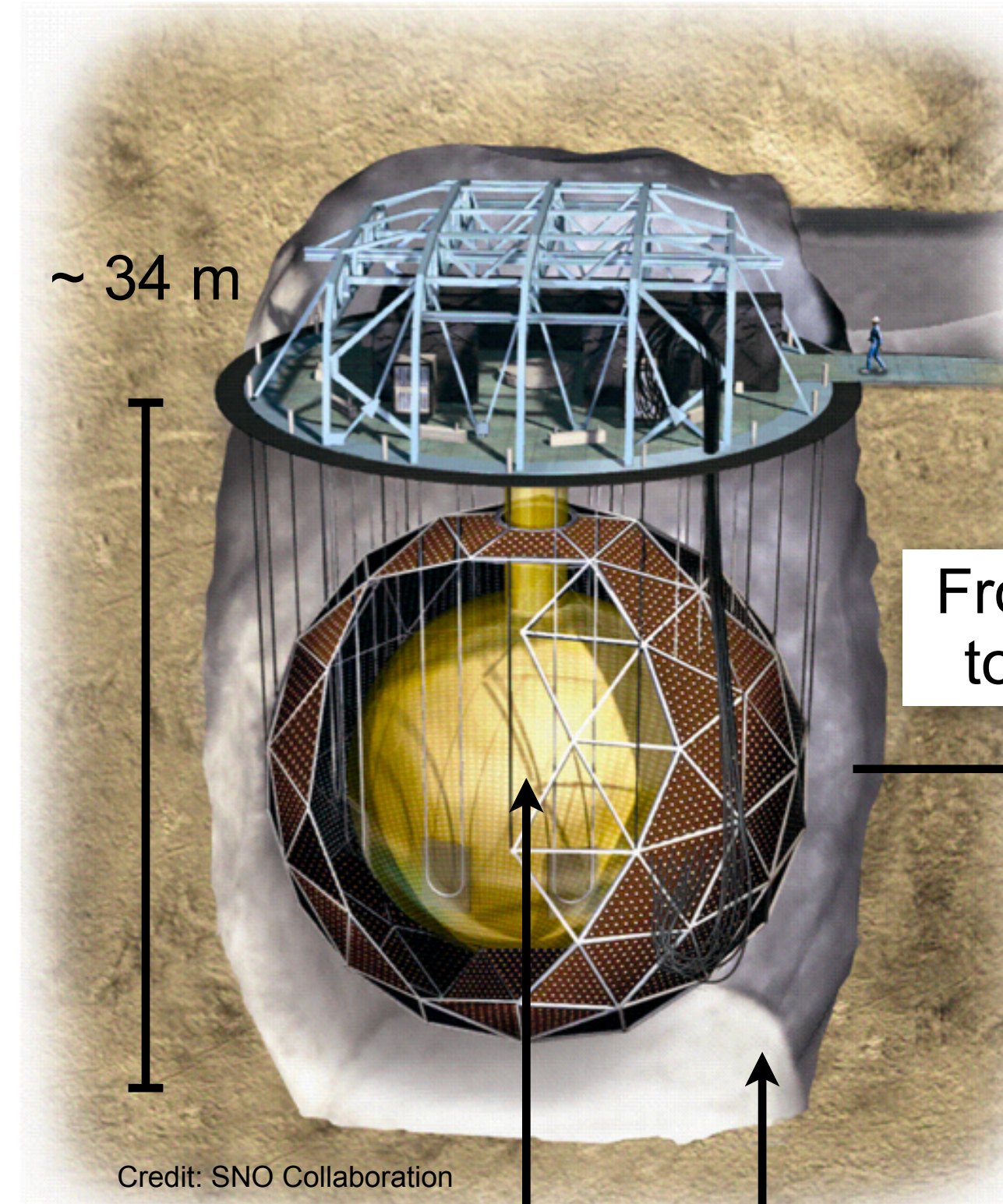
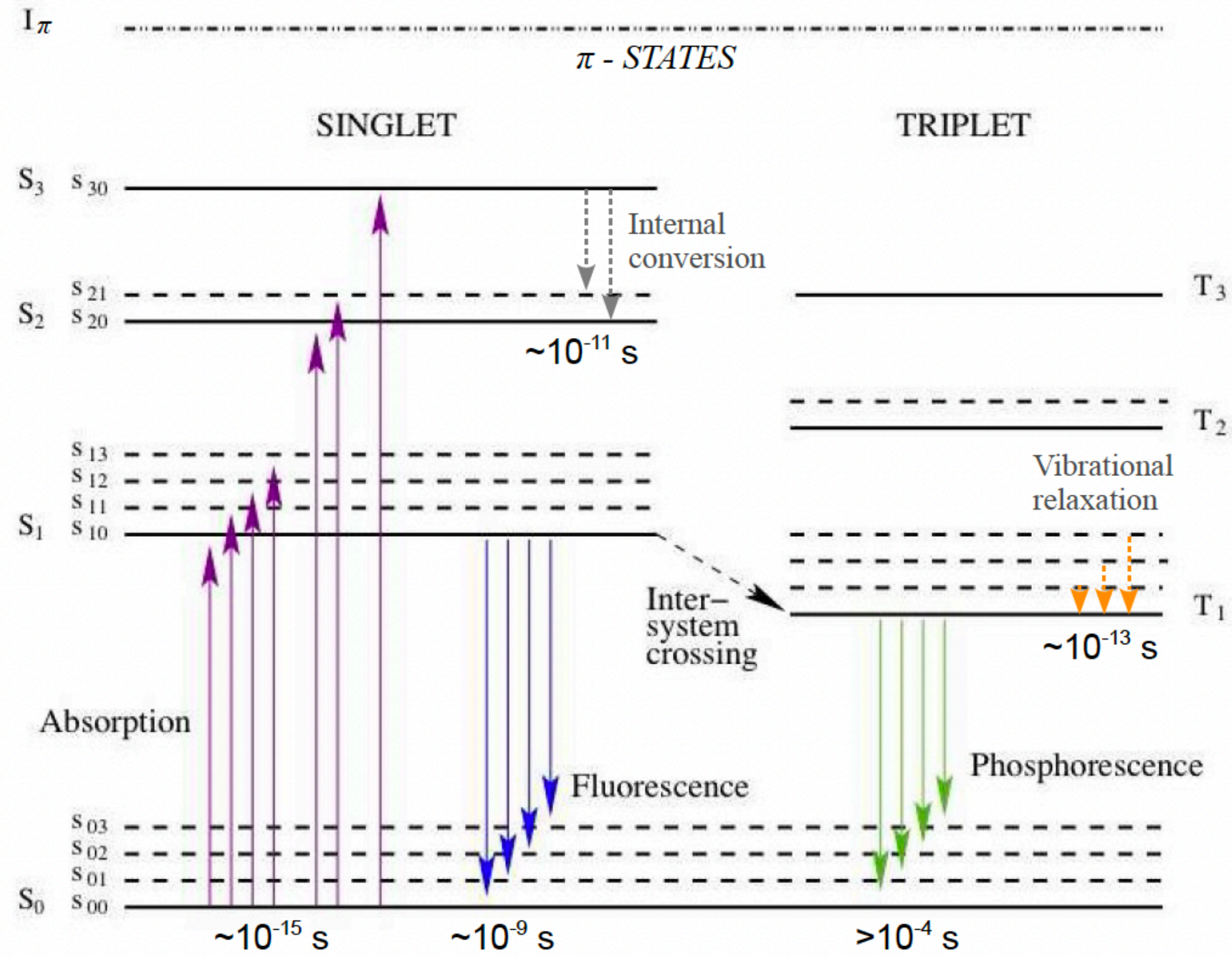


> 9000 PMTs with  $\varnothing$  20cm

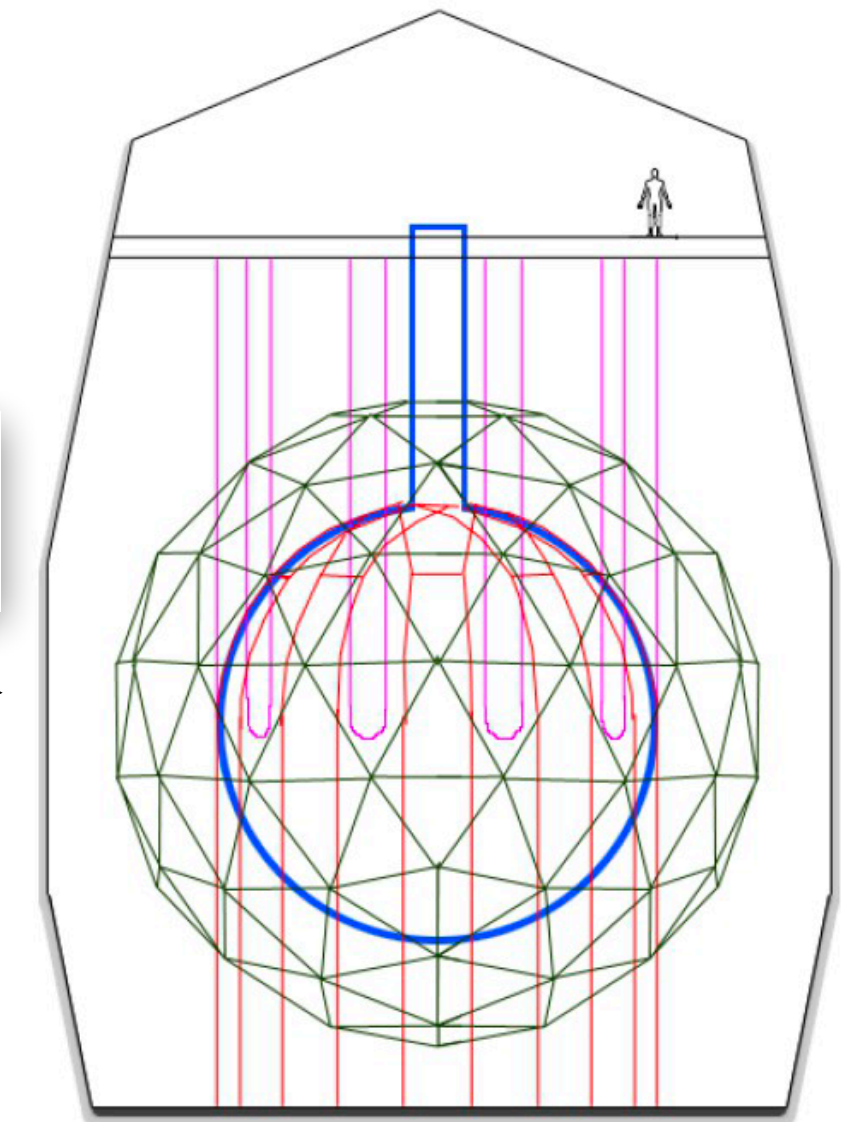


# Ton scale $0\nu\beta\beta$ detectors

## Scintillation light from an organic molecule



From SNO  
to SNO+



SNO: 1 kt of  $D_2O$

SNO+: 0.78 kt of liquid scintillator

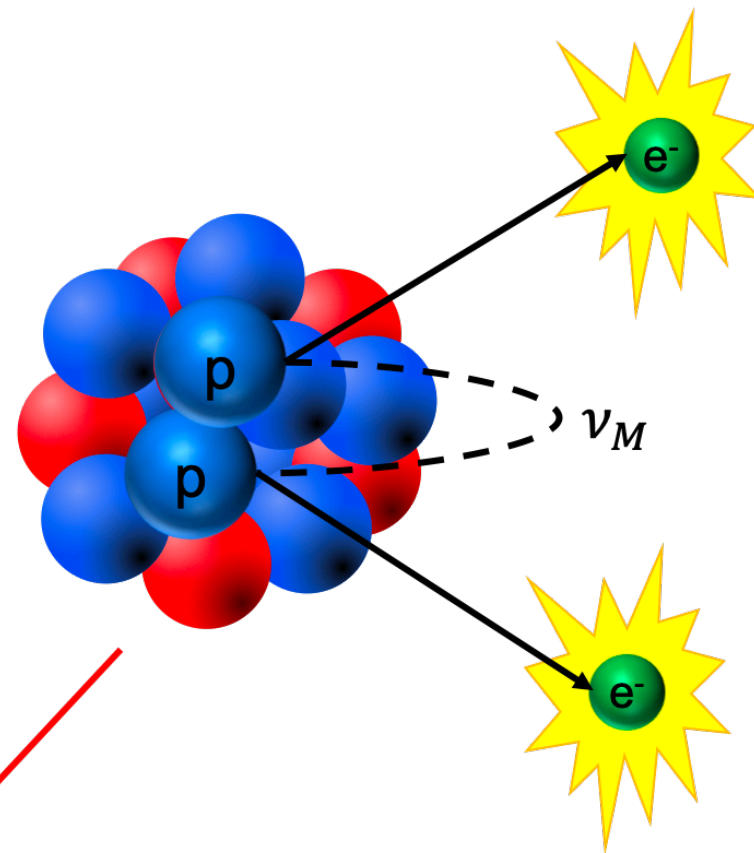
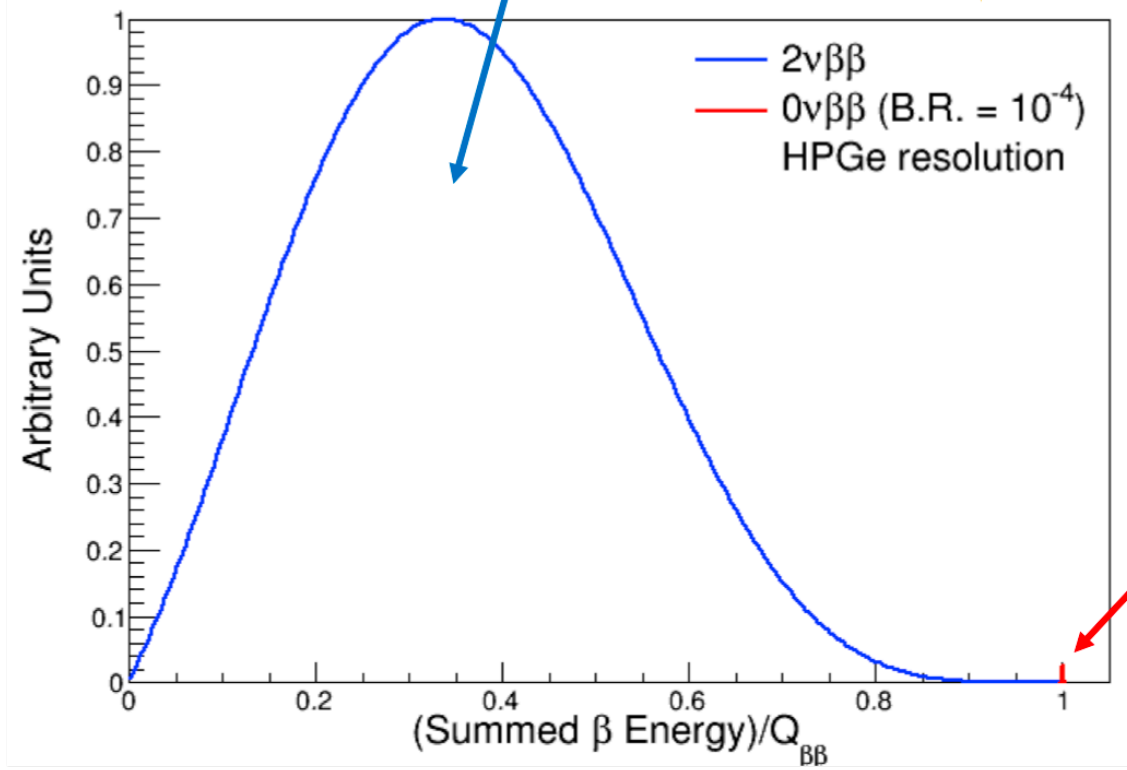
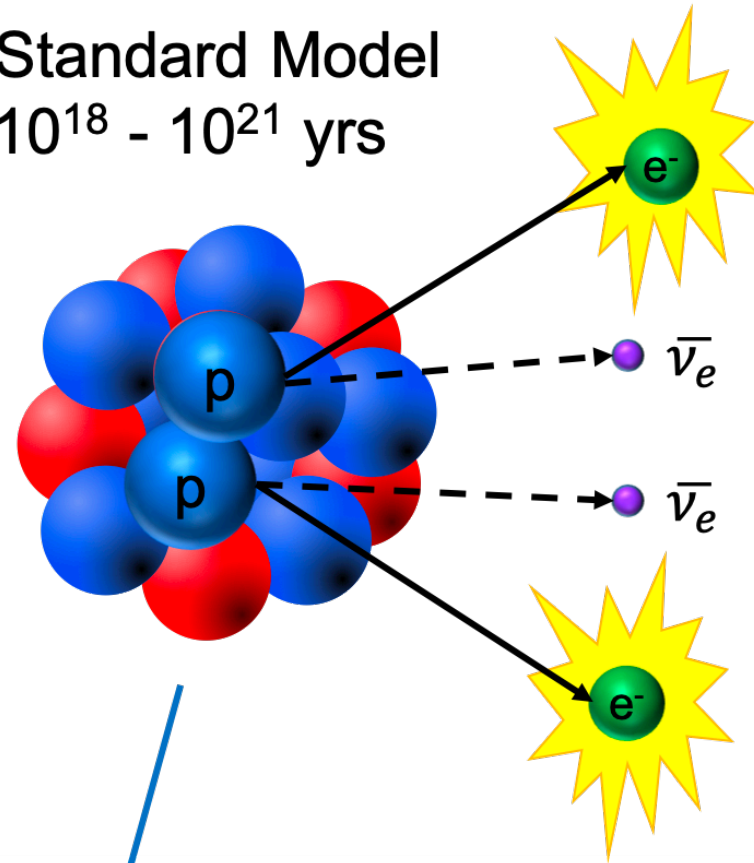
7 kt  $H_2O$



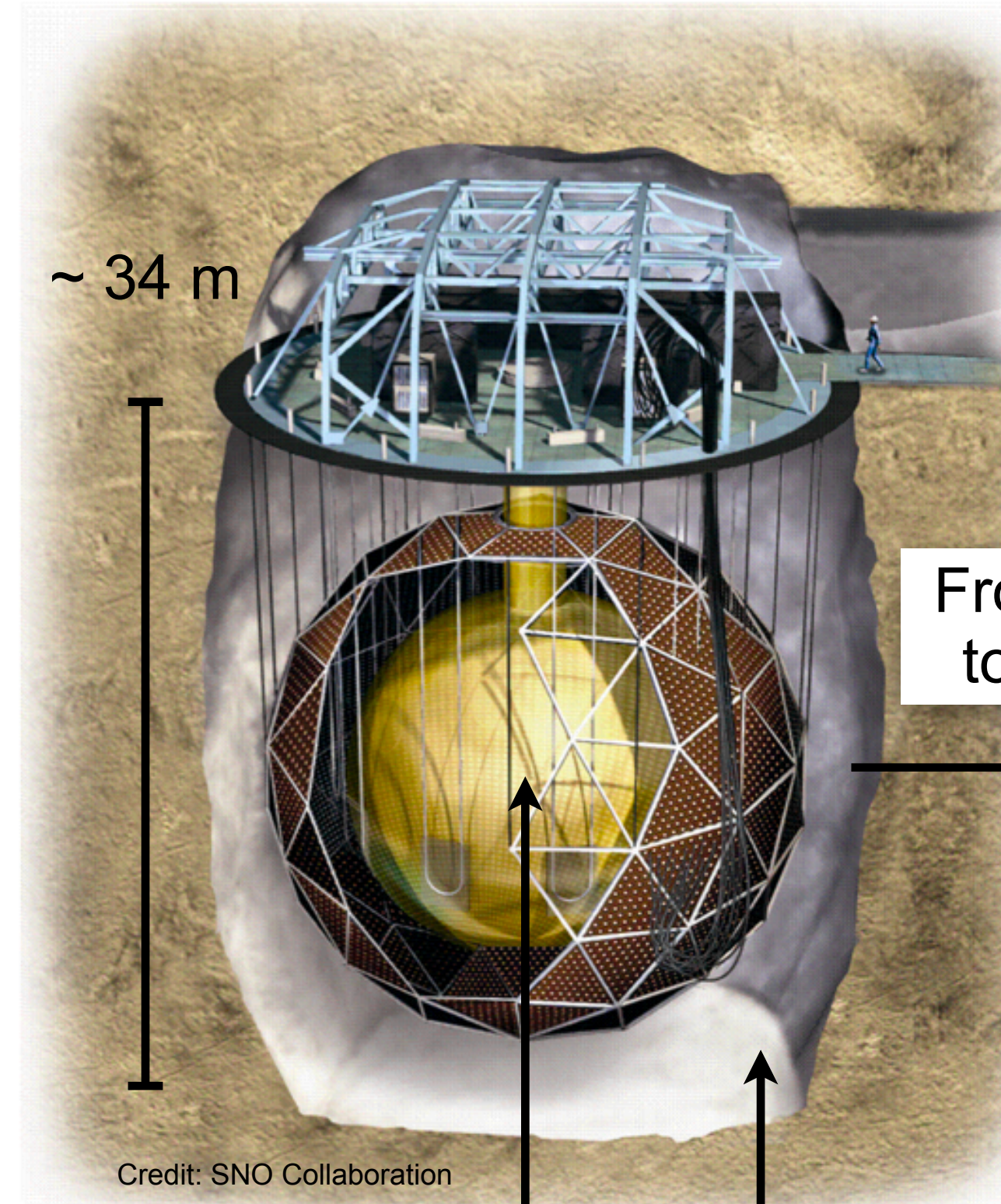
# Ton scale $0\nu\beta\beta$ detectors

## Searching for the $0\nu\beta\beta$ decay of $^{130}\text{Te}$

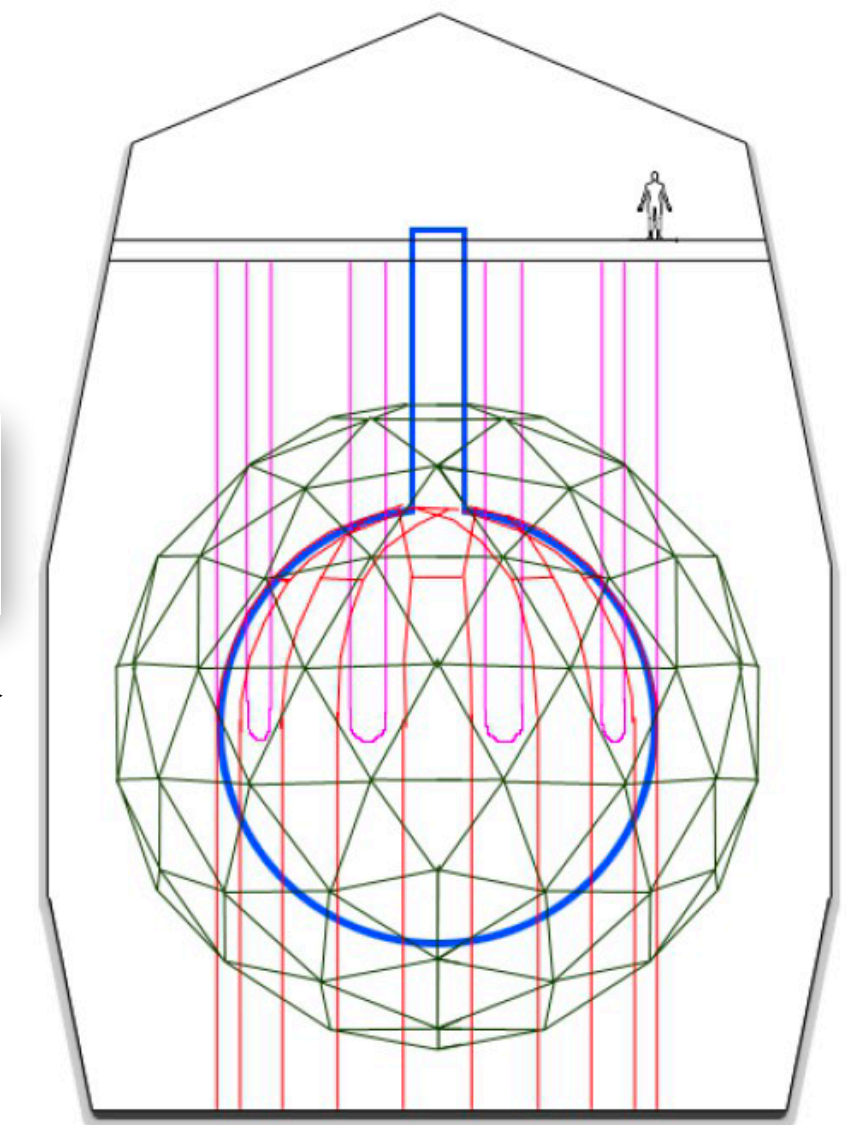
$2\nu\beta\beta$  – Standard Model  
 $\tau_{1/2} = 10^{18} - 10^{21}$  yrs



$0\nu\beta\beta$  – New Physics!  
 $\tau_{1/2} > 10^{26}$  yrs



From SNO  
to SNO+



SNO: 1 kt of  $\text{D}_2\text{O}$

SNO+: 0.78 kt of liquid scintillator

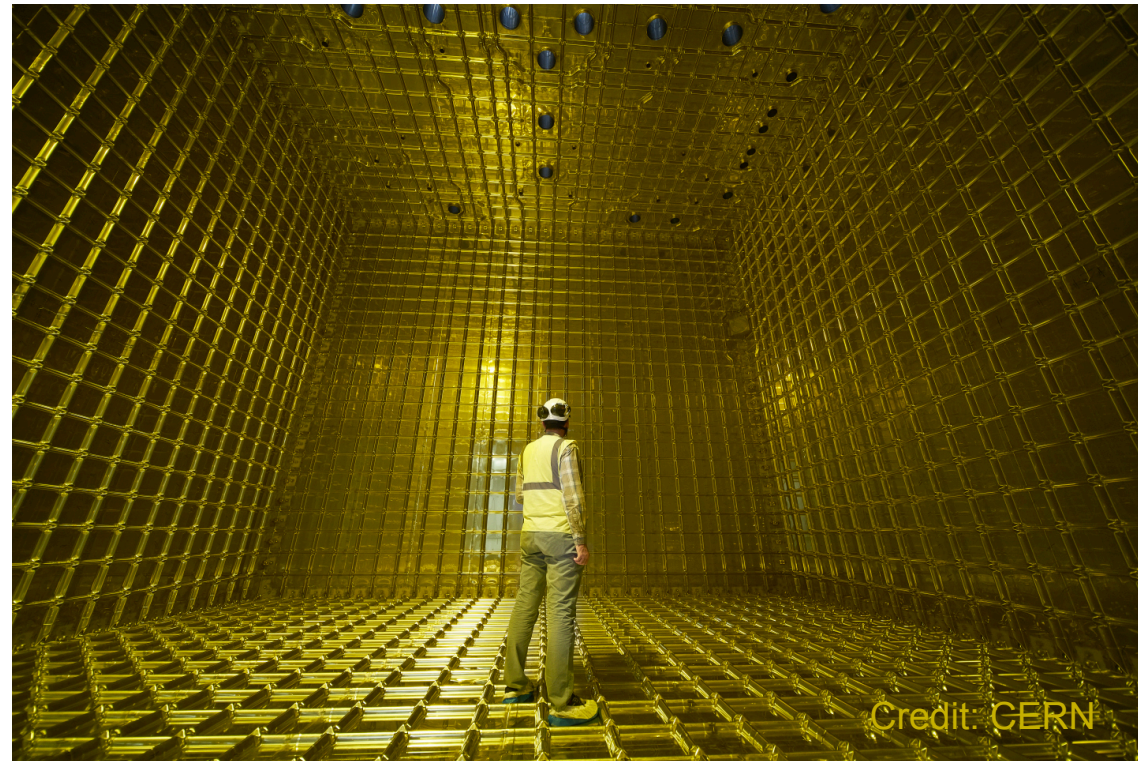
+  $^{130}\text{Te}$

7 kt  $\text{H}_2\text{O}$

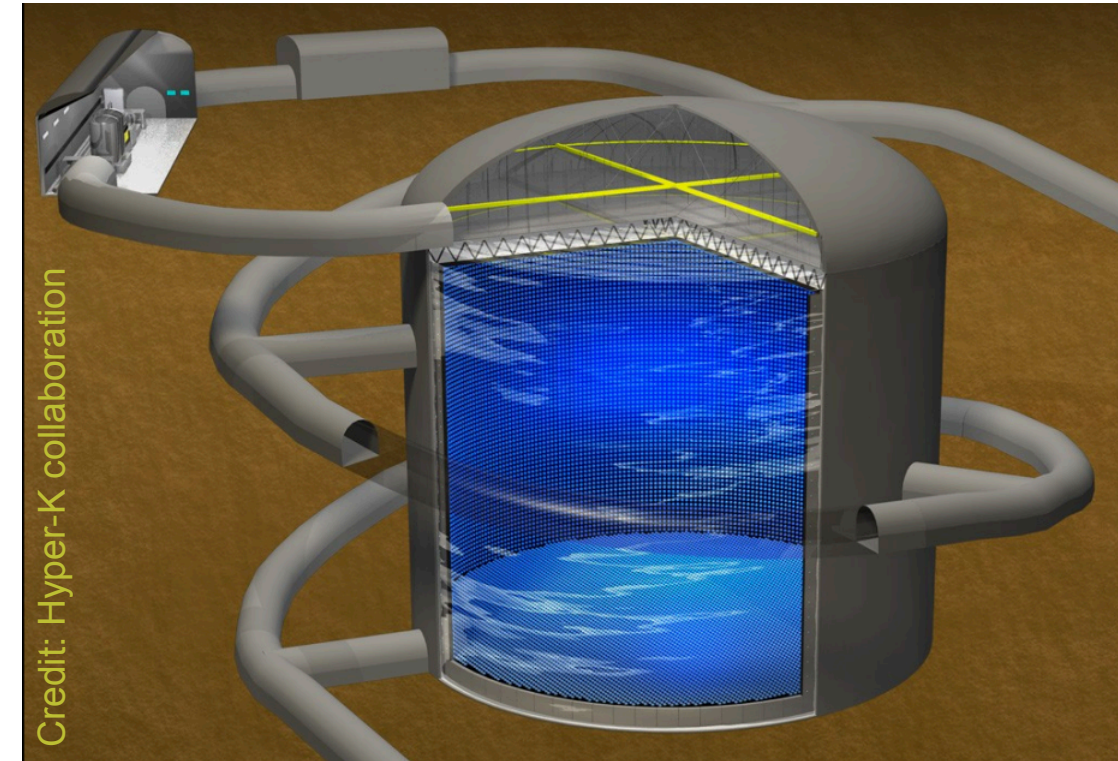


# Long baseline accelerator-neutrino programmes

DUNE programme

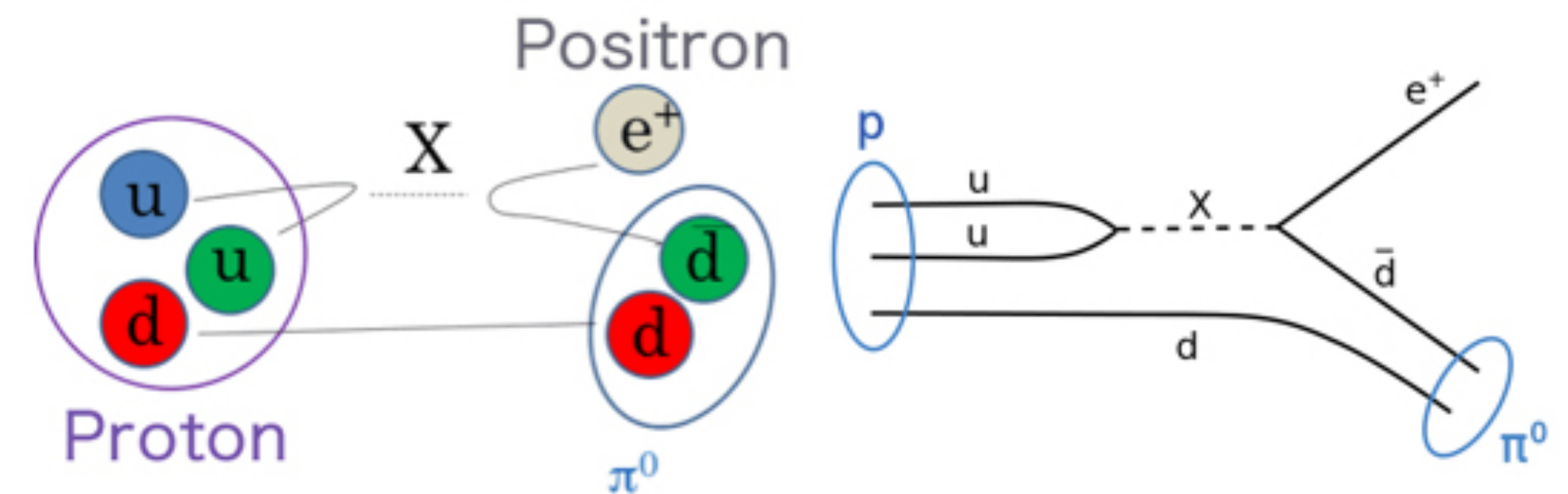
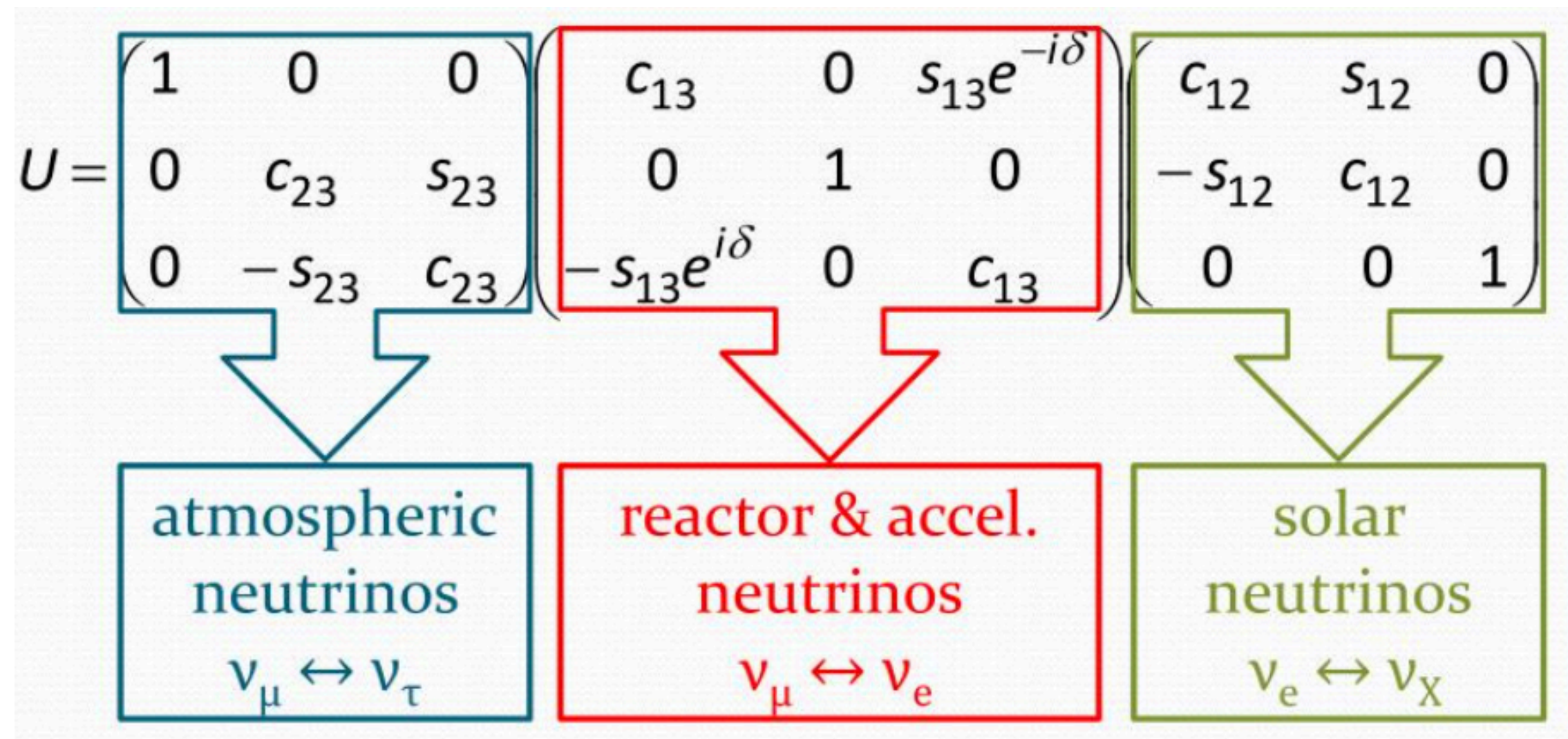


Hyper-K programme



## Science drivers:

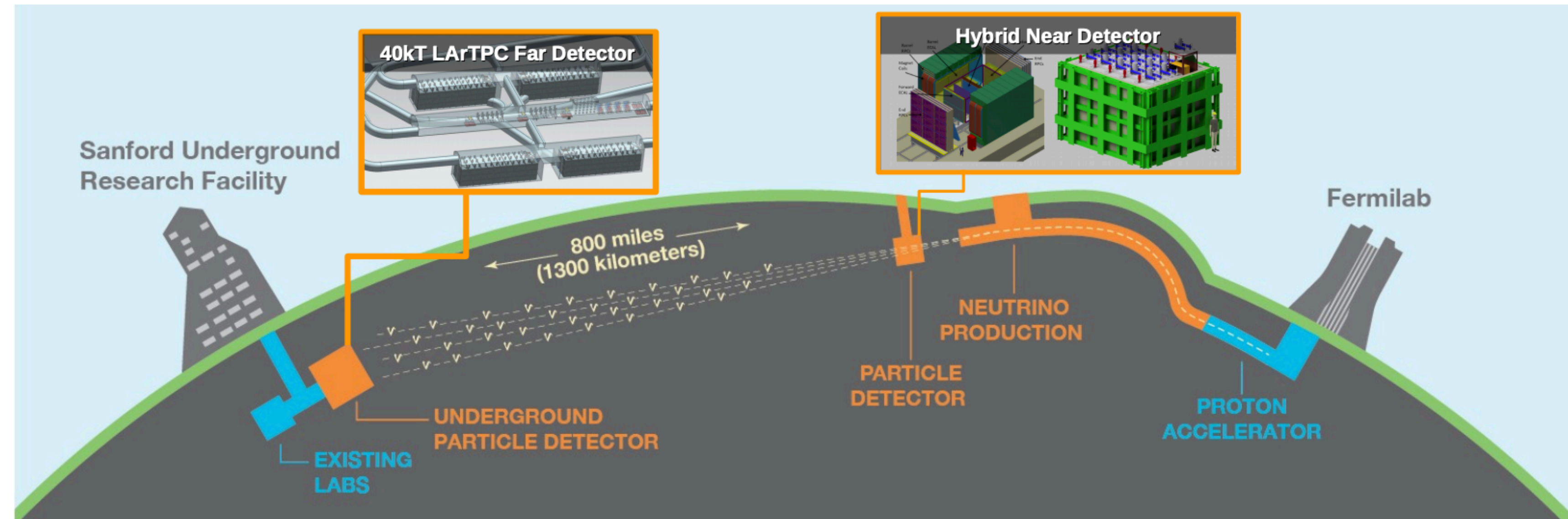
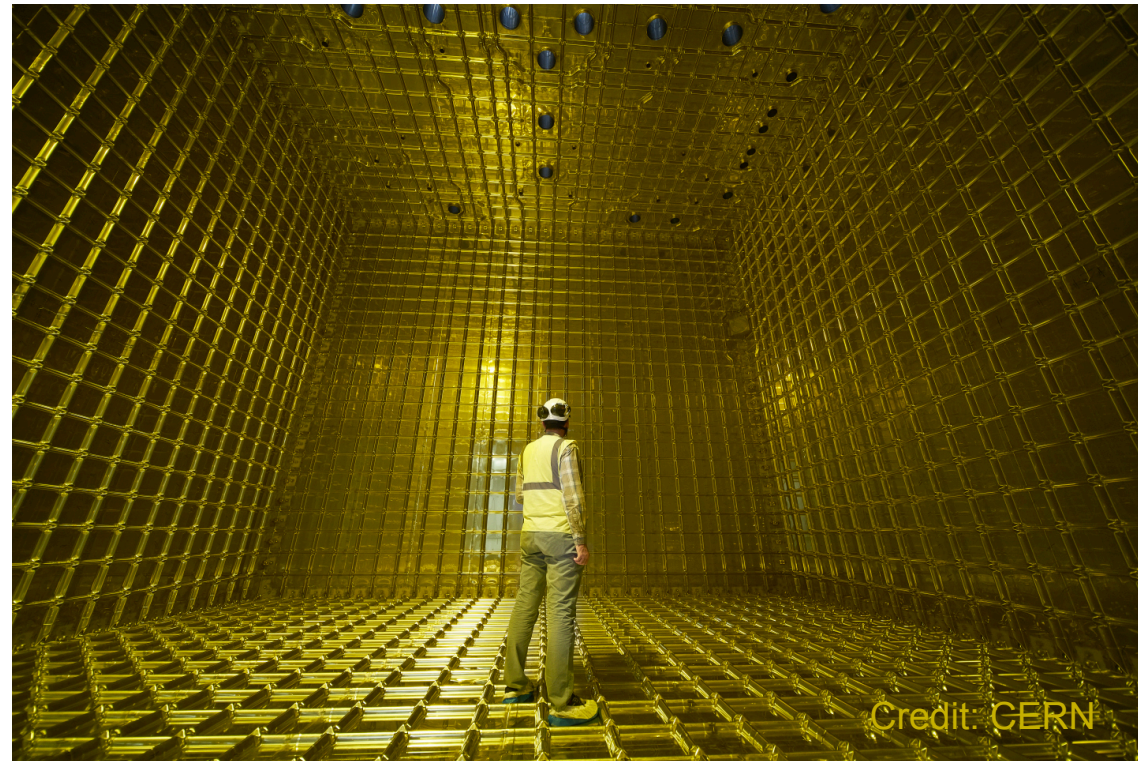
- discovery of CP violation in neutrino oscillations
- precision measurement of the PMNS matrix parameters
- study of solar and supernova neutrinos
- proton decay





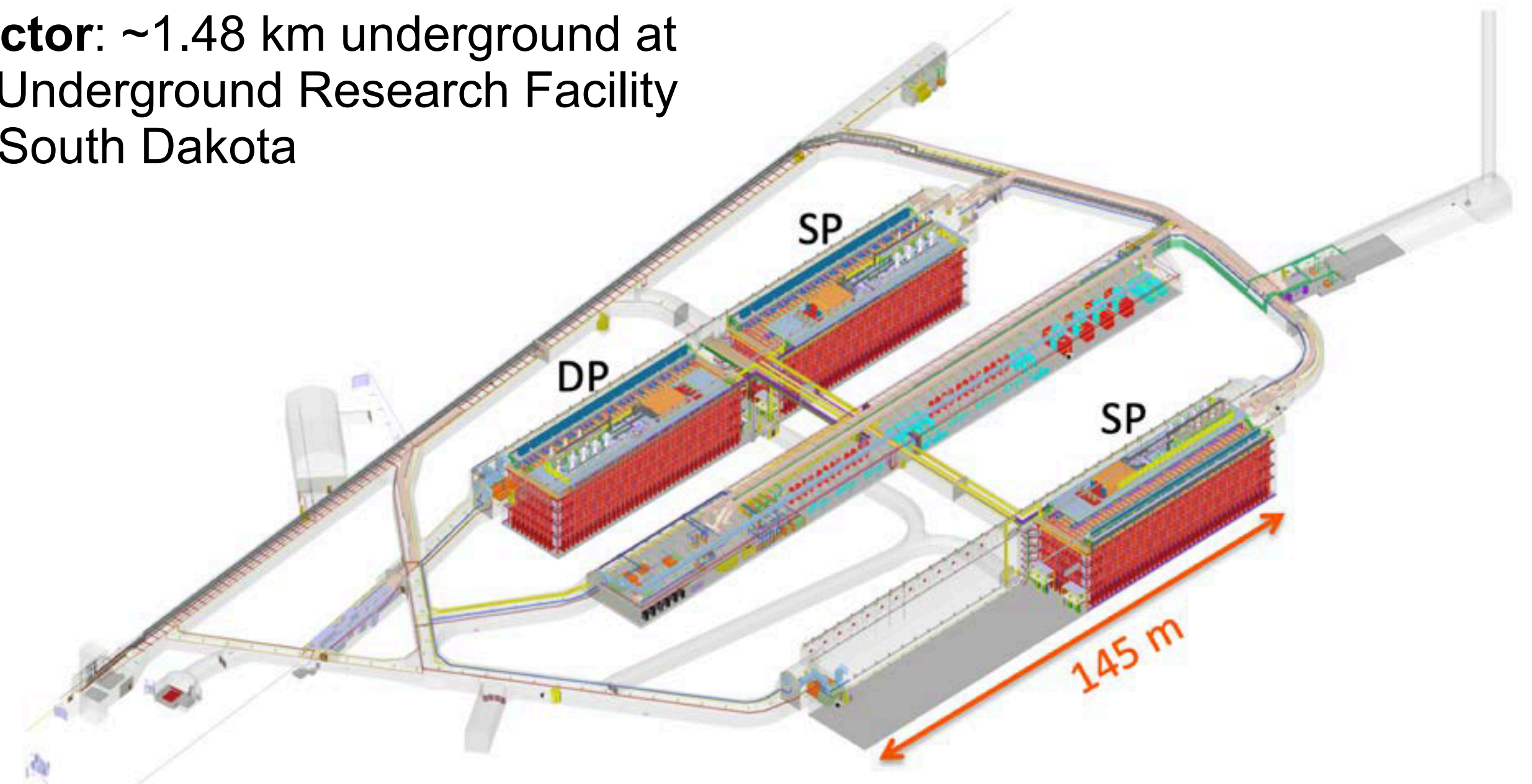
# Deep Underground Neutrino Experiment (DUNE)

## DUNE programme



**Far detector:** ~1.48 km underground at Sanford Underground Research Facility in Lead, South Dakota

- 4 x 10-kt fiducial LAr TPC modules
  - 2 modules horizontal drift
  - 1 vertical drift module
  - 1 “opportunity” module

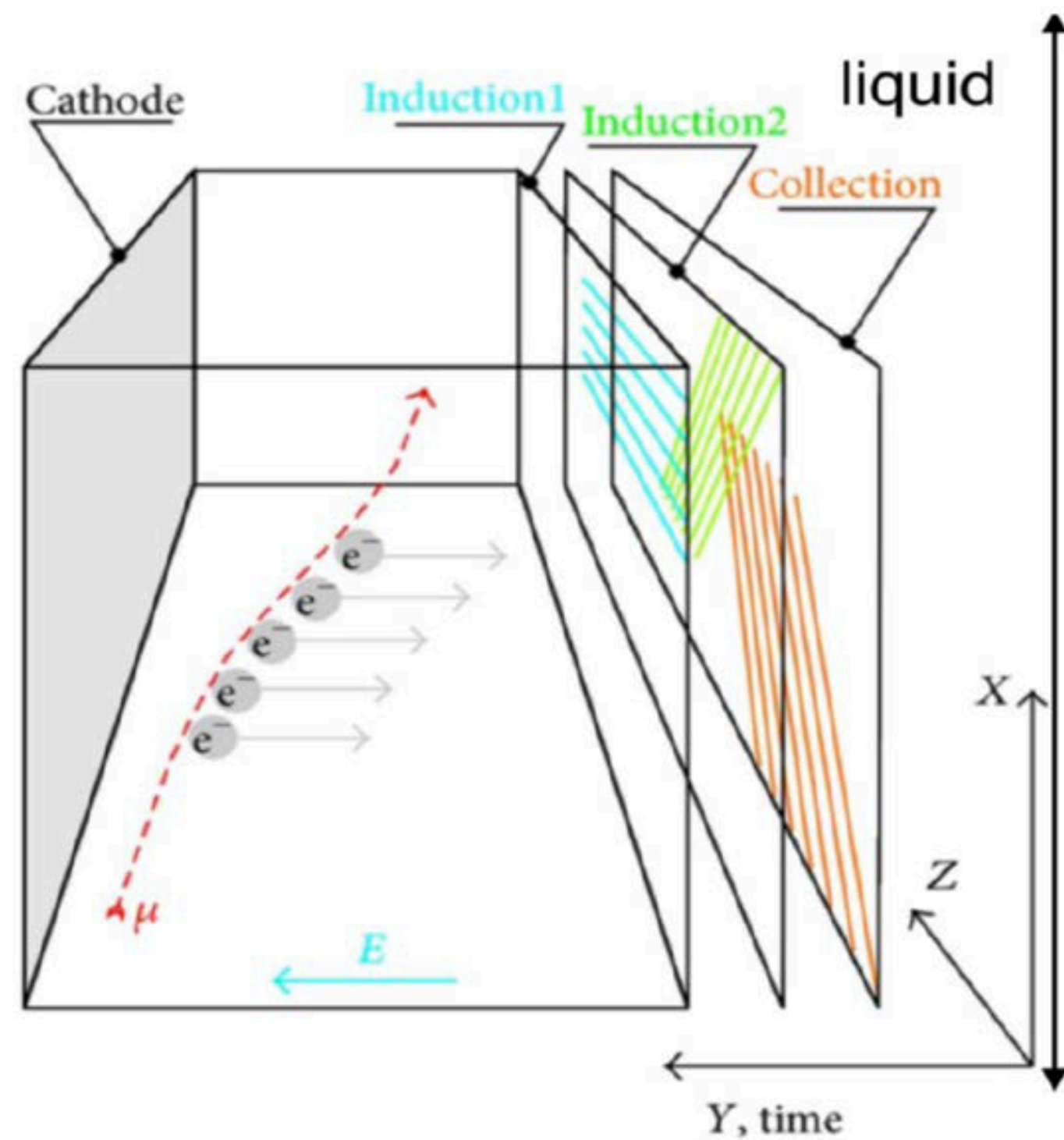




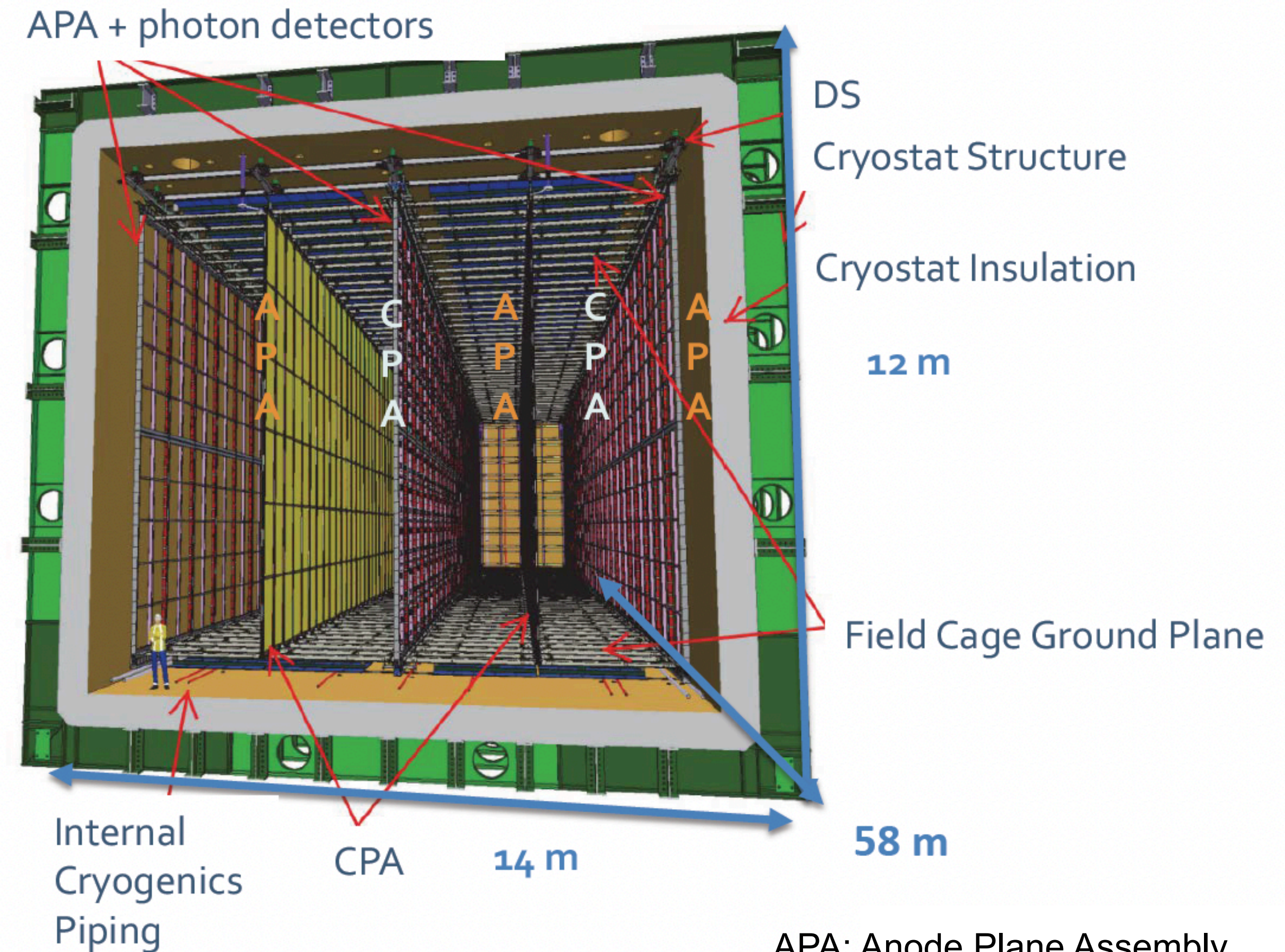
# Deep Underground Neutrino Experiment (DUNE)

## ■ LAr TPC technology

- Excellent 3D imaging capabilities – ( $O(\text{mm})$  over large volume)
- Excellent energy measurement capability – fully active calorimeter
- Particle ID by  $dE/dx$ , range, event topology



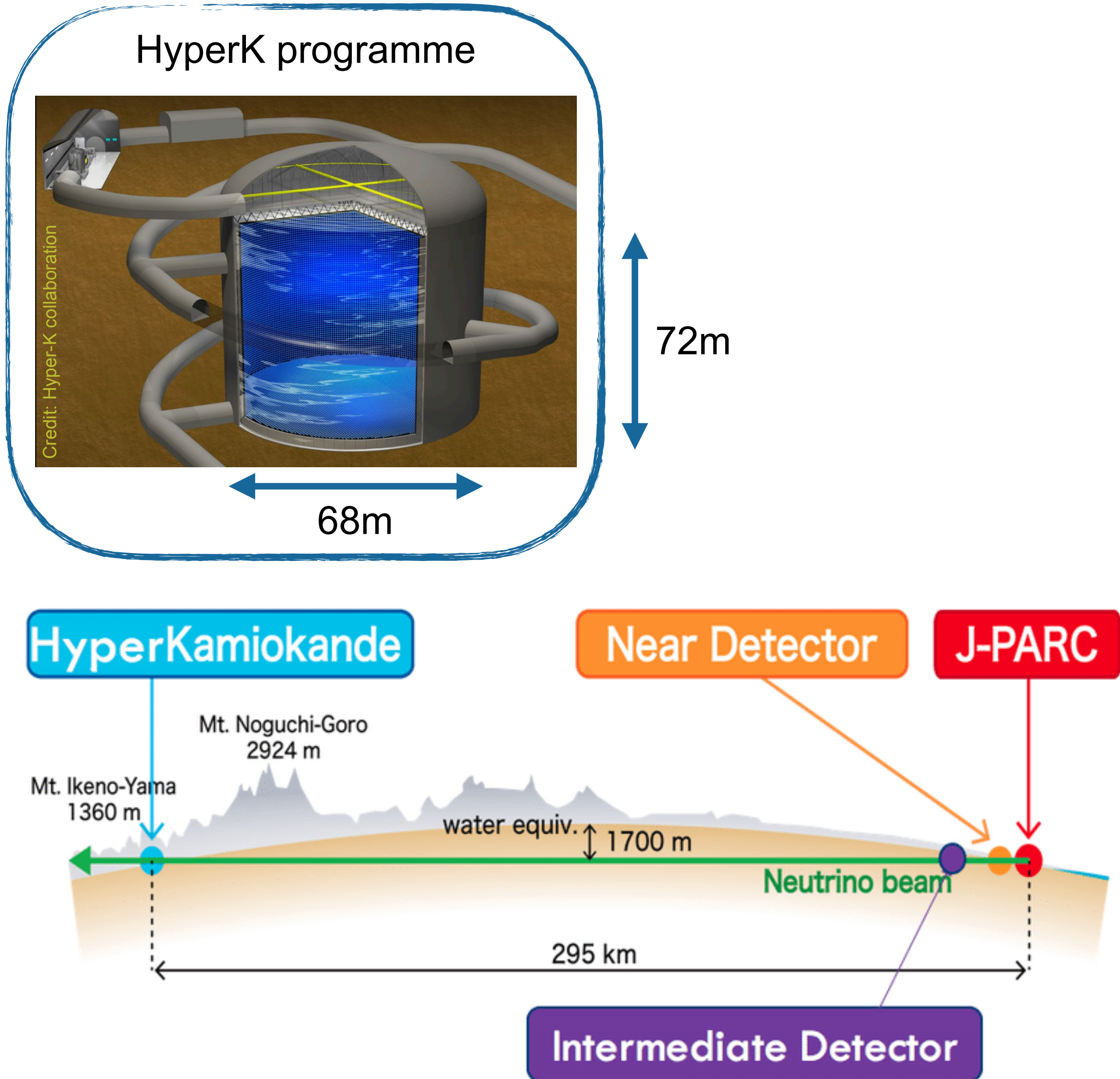
Far detector 1 (horizontal drift)



APA: Anode Plane Assembly  
CPA: Cathode Plane Assembly

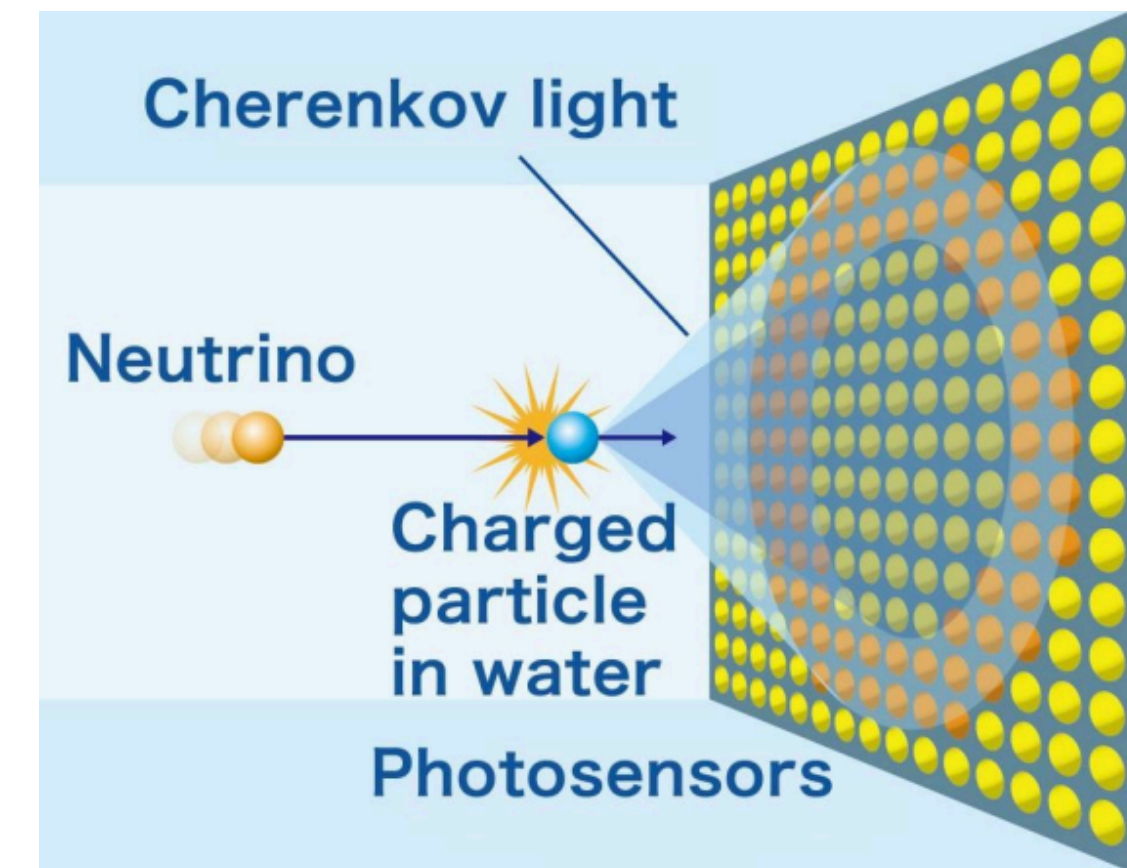
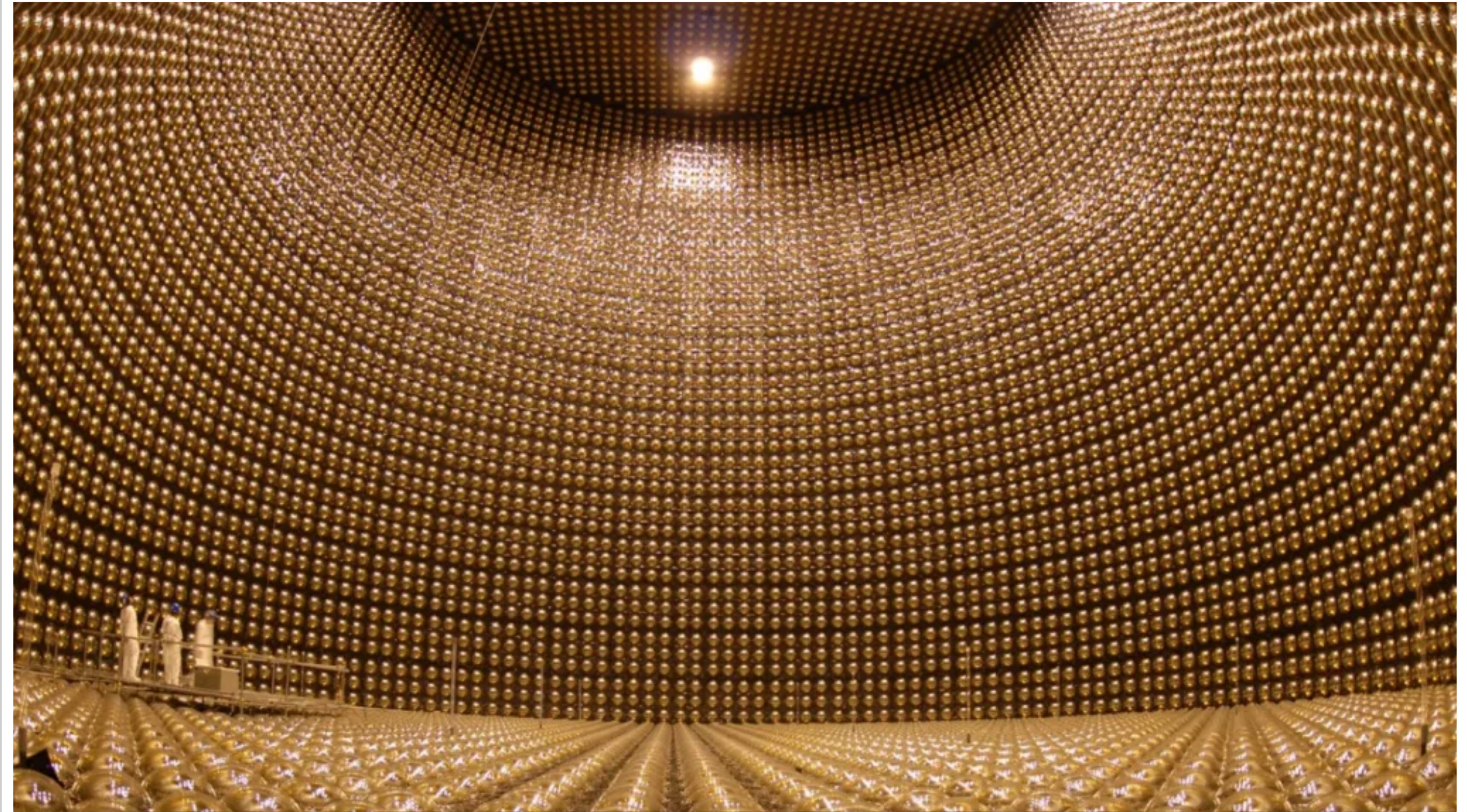


# Hyper-Kamiokande



## Super-Kamiokande

Credit: Kamioka Observatory, ICRR, University of Tokyo

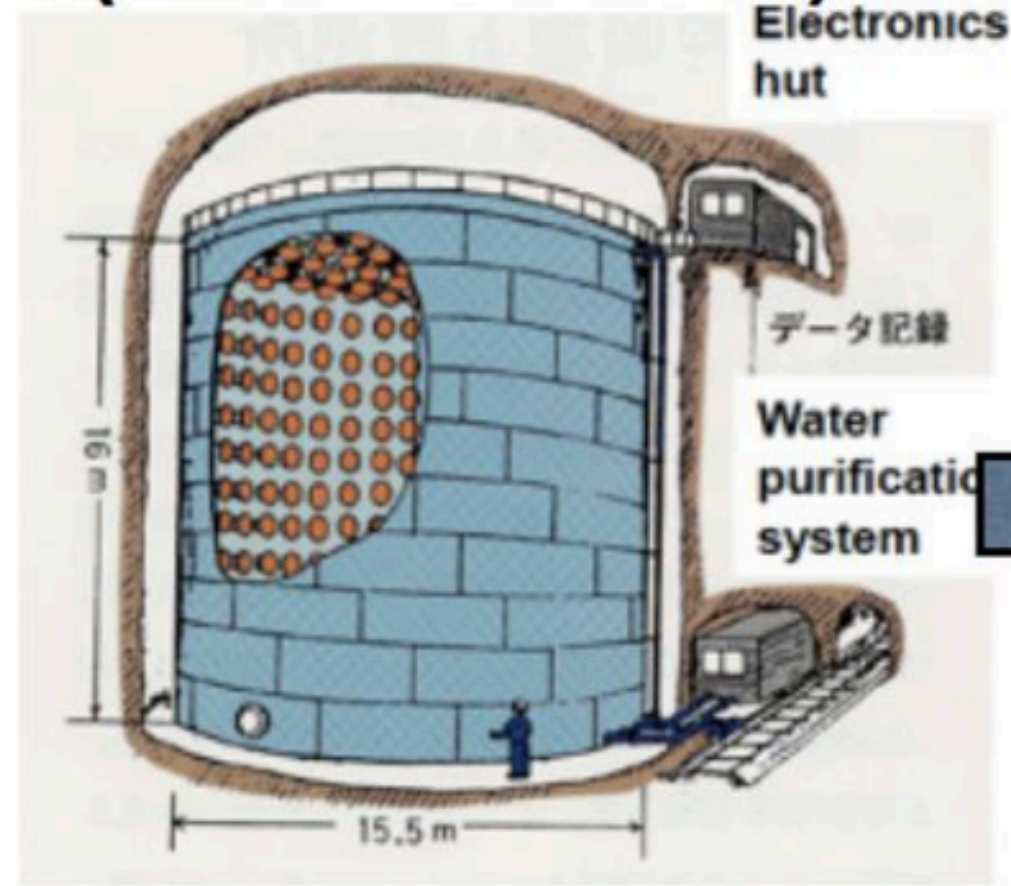






# Hyper-Kamiokande

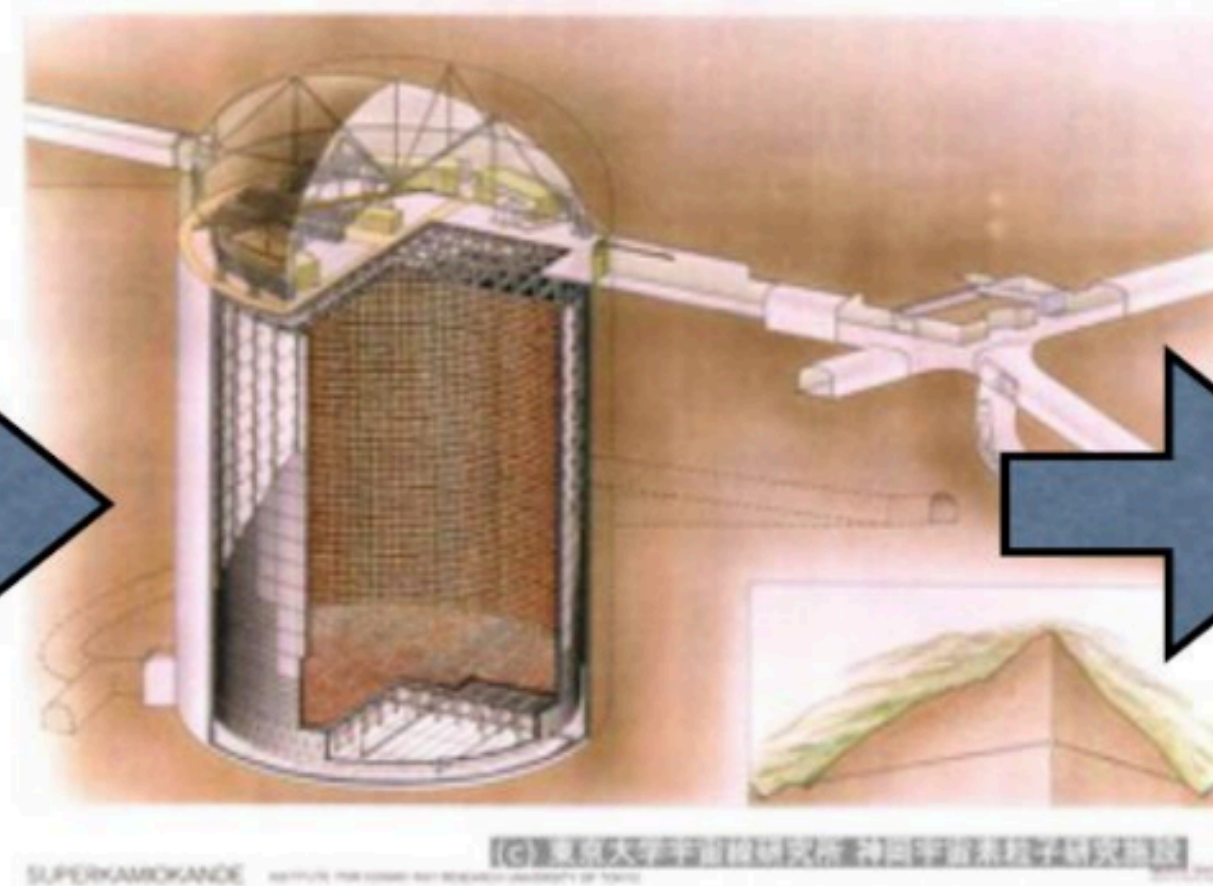
Kamiokande  
(1983-1996)



3kton

20% coverage  
with 50cm PMT

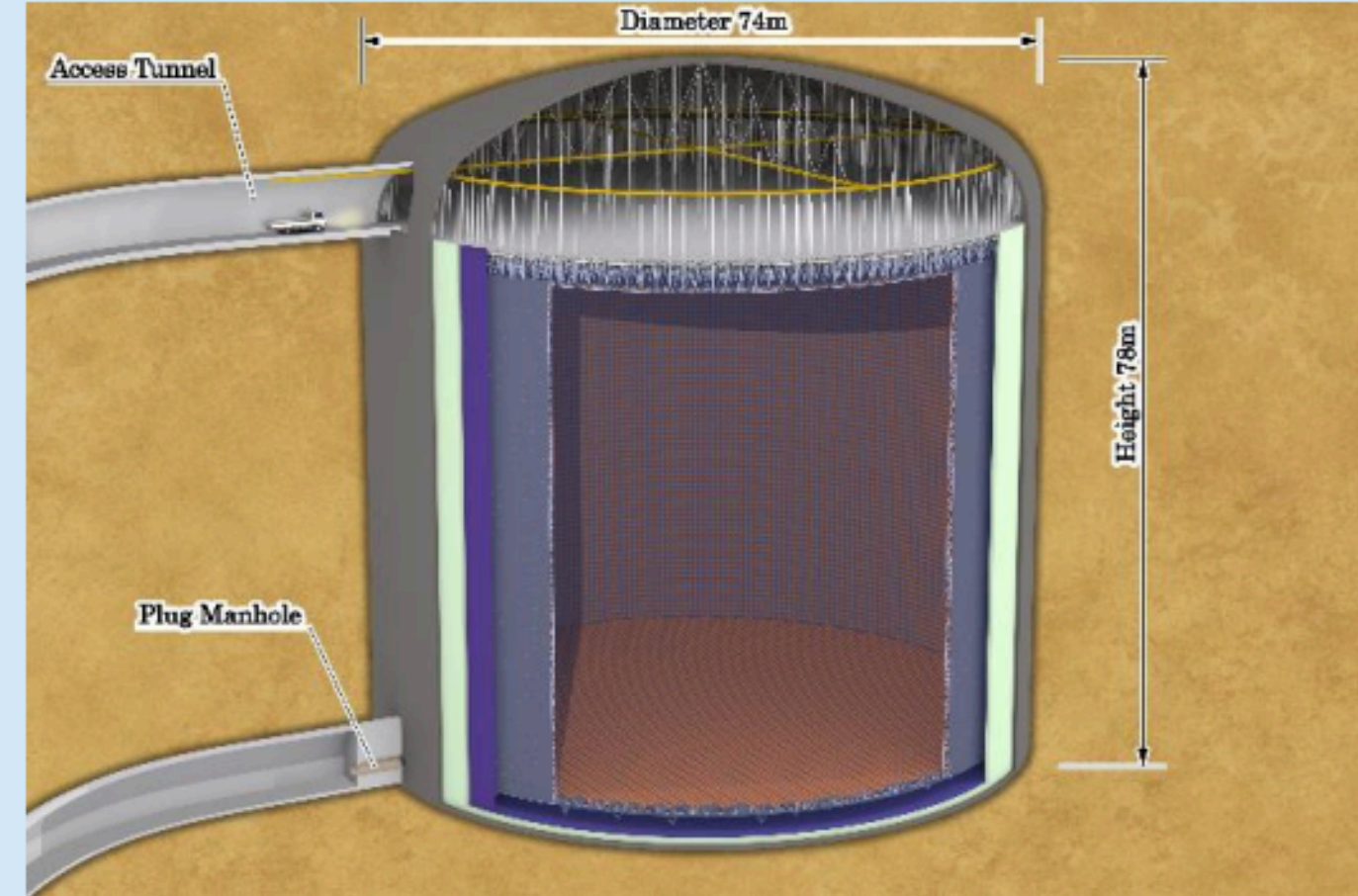
Super-Kamiokande  
(1996-)



50kton

40% coverage  
with 50cm PMT

Hyper-Kamiokande  
(~2026-)



260kton

40% coverage  
with high-QE 50cm PMT

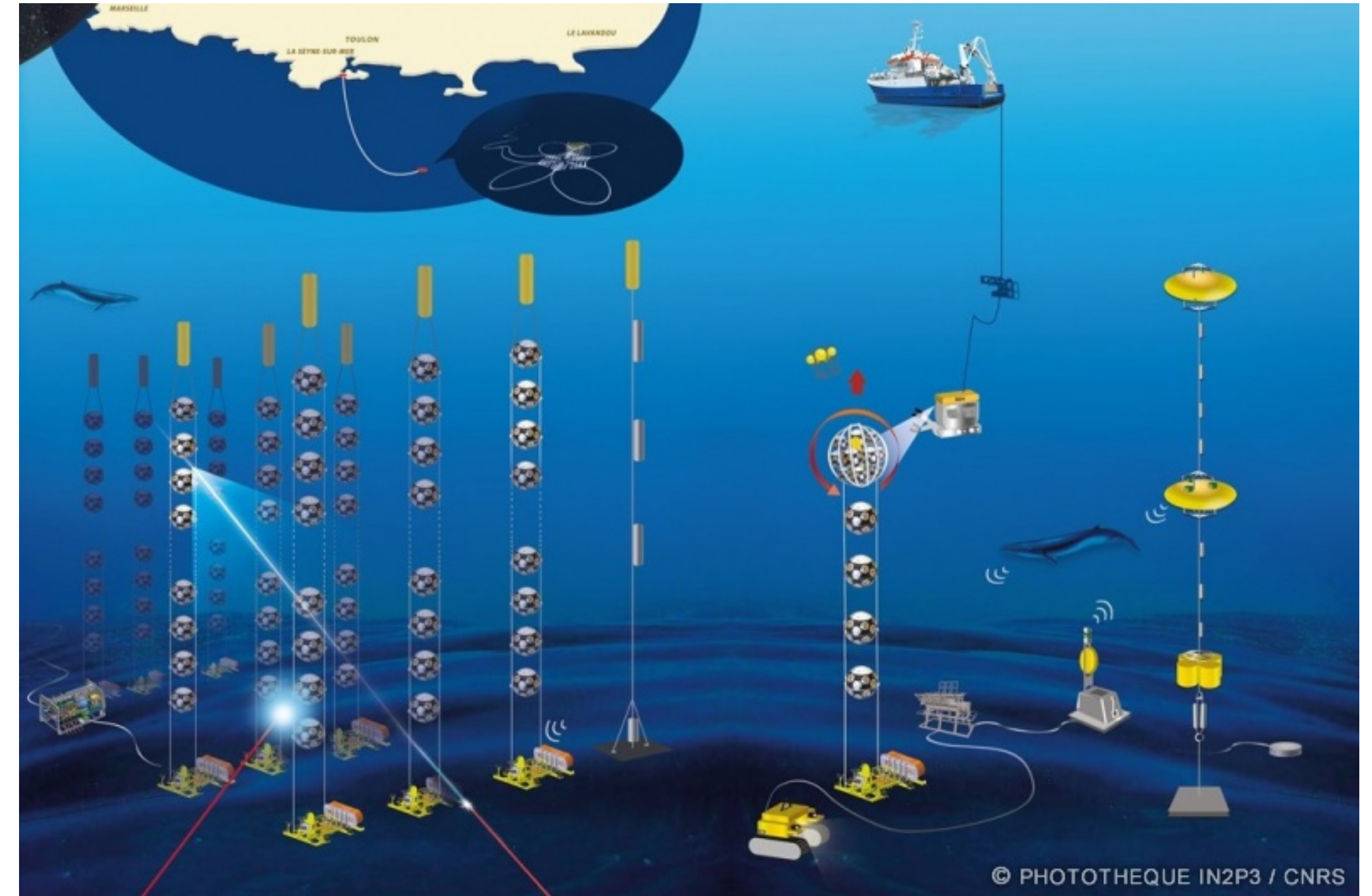
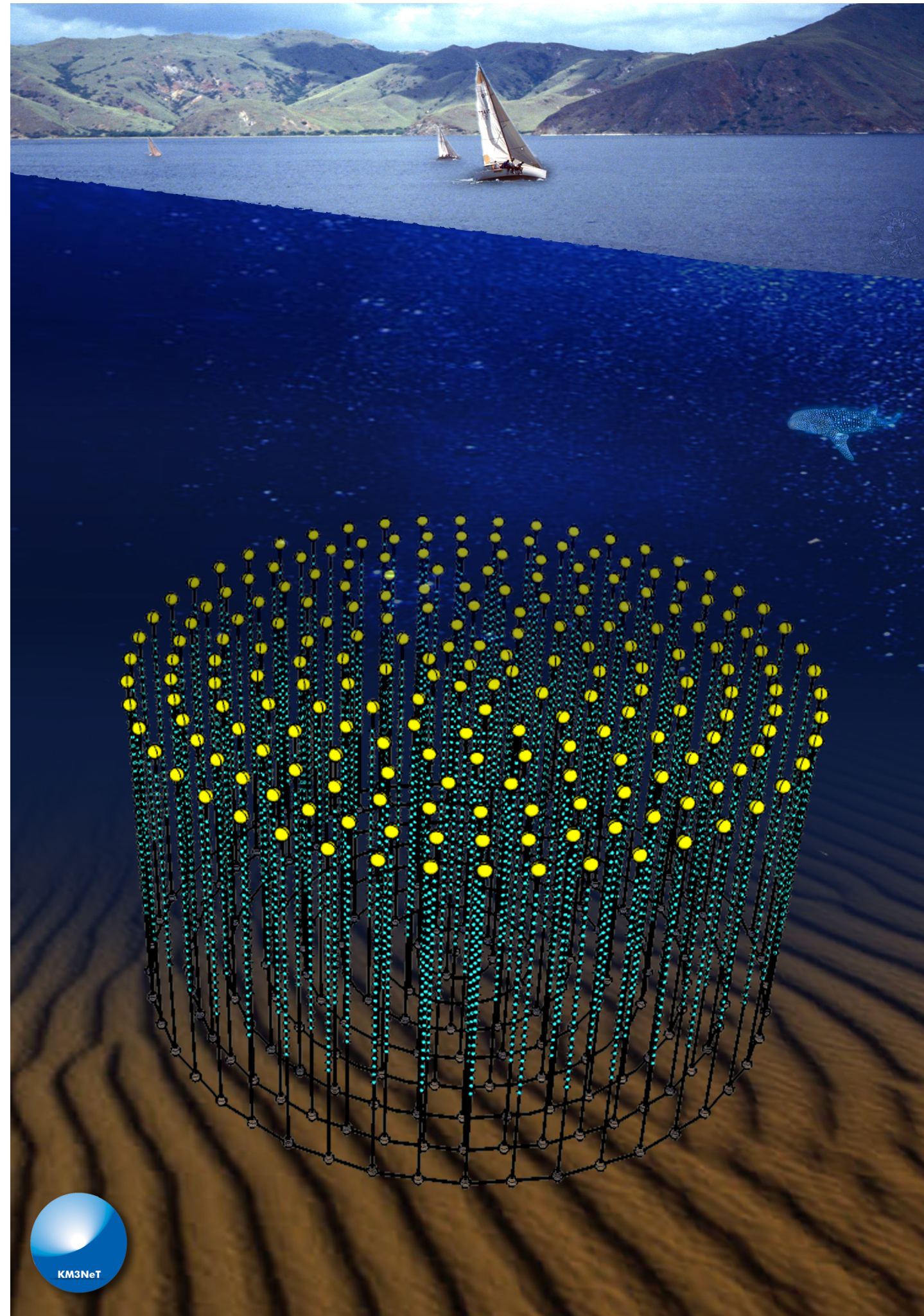
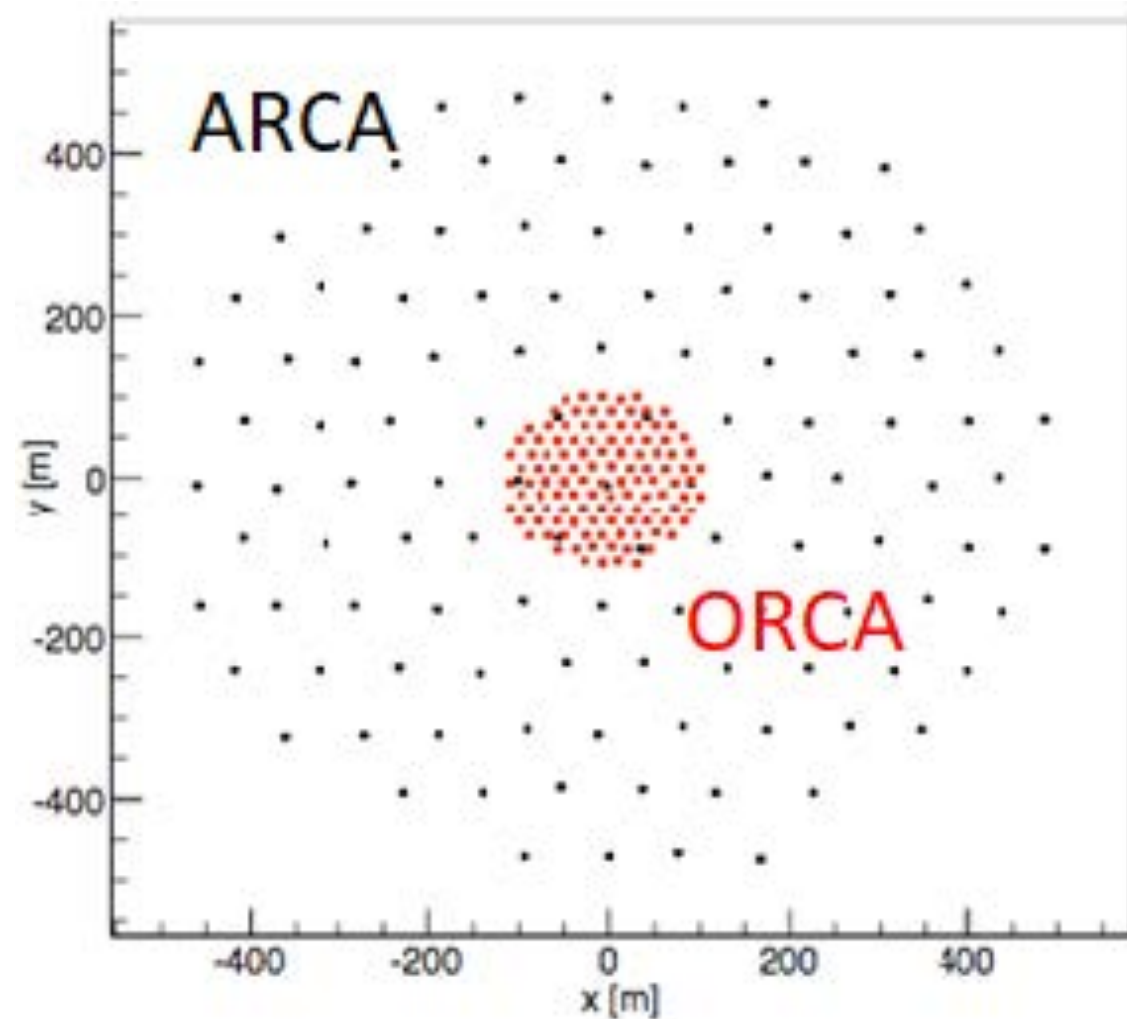
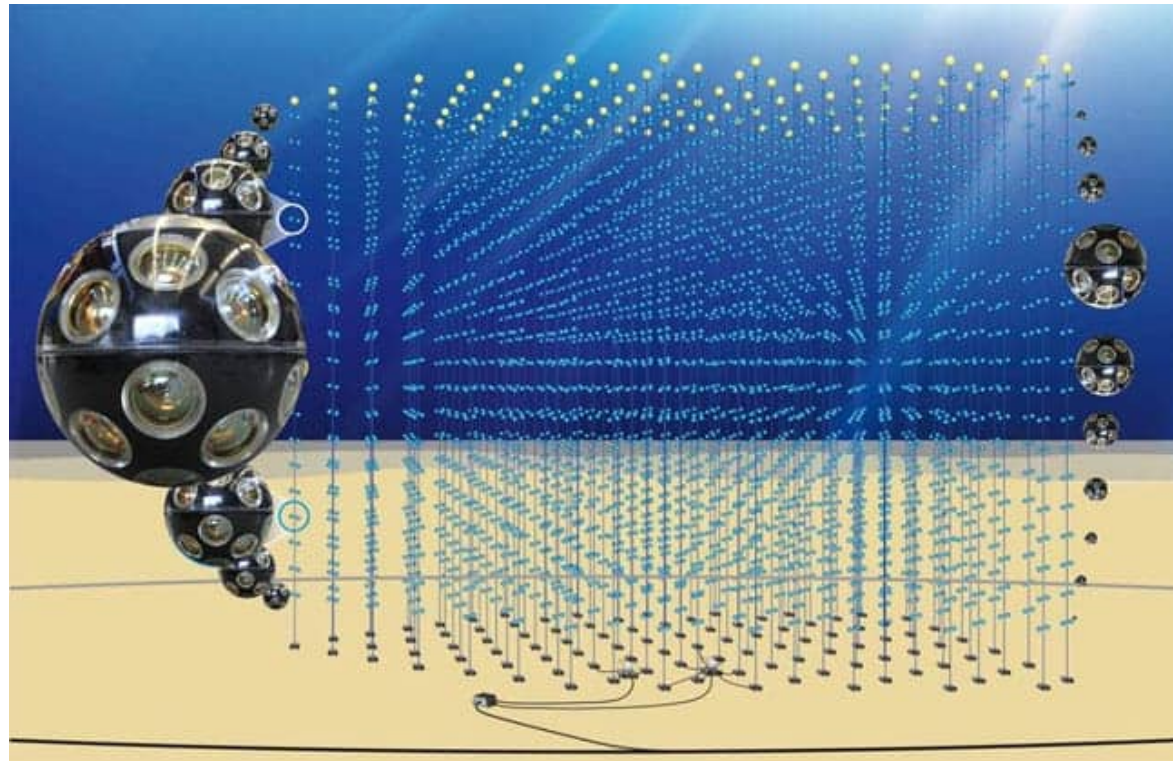
x17

x10



# KM3NeT

neutrino telescopes

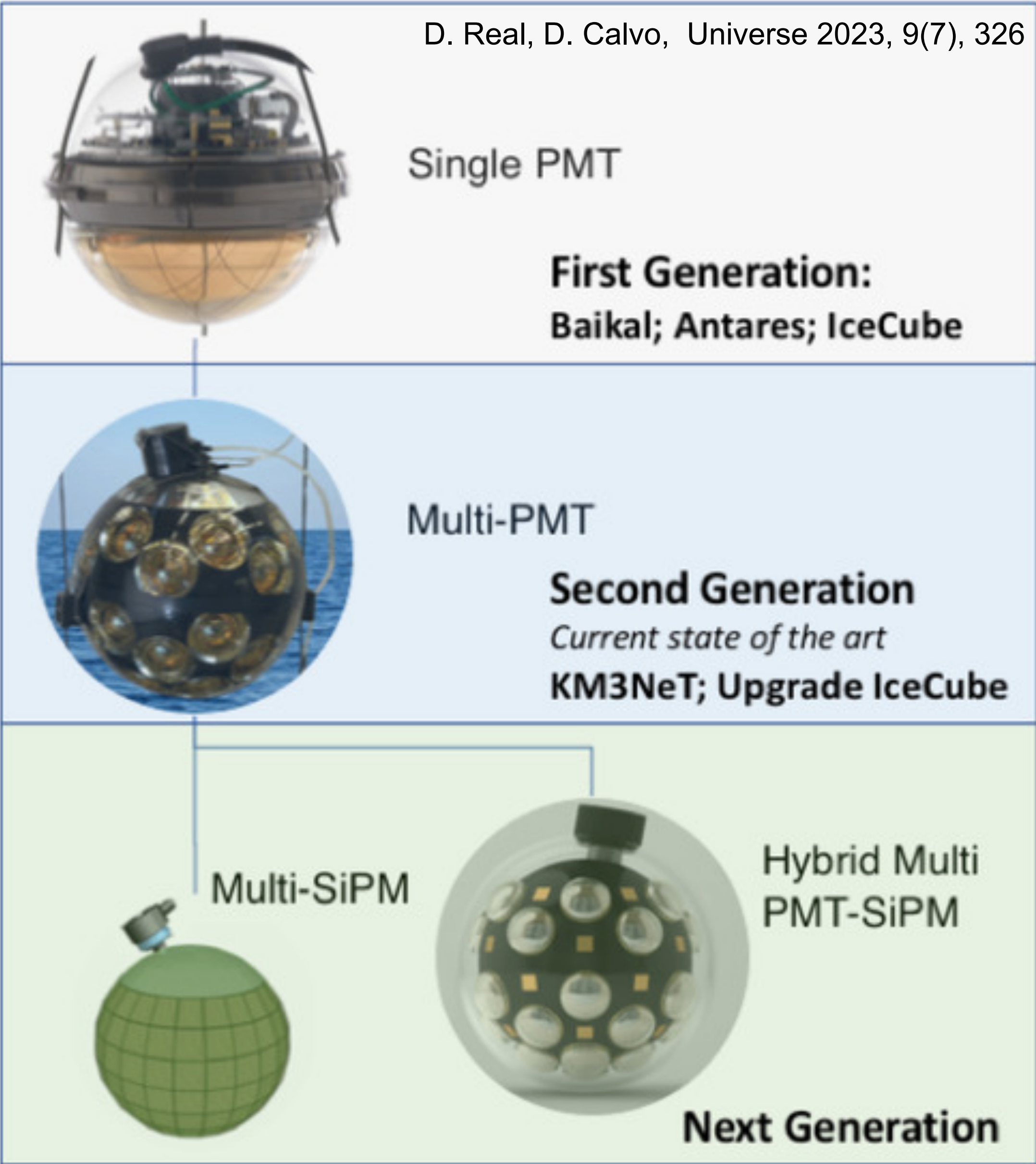


- ARCA at 3.5 km depth in front of Sicily
  - neutrinos from supernovae, GRBs or colliding neutron stars
- ORCA at 2.5 km depth in front of Toulon
  - optimized for energies around 10 GeV (atmospheric neutrinos)





# KM3NeT





**Enabling technologies**





# Enabling technologies (near-term, next 5-10 years)

1

Properties of  
**noble liquids**

2

**Charge collection**  
in noble liquids

3

**Purification**, cryogenics,  
infrastructure, and  
integration for noble  
liquids

4

**Light collection** in noble  
liquids and other liquids

5

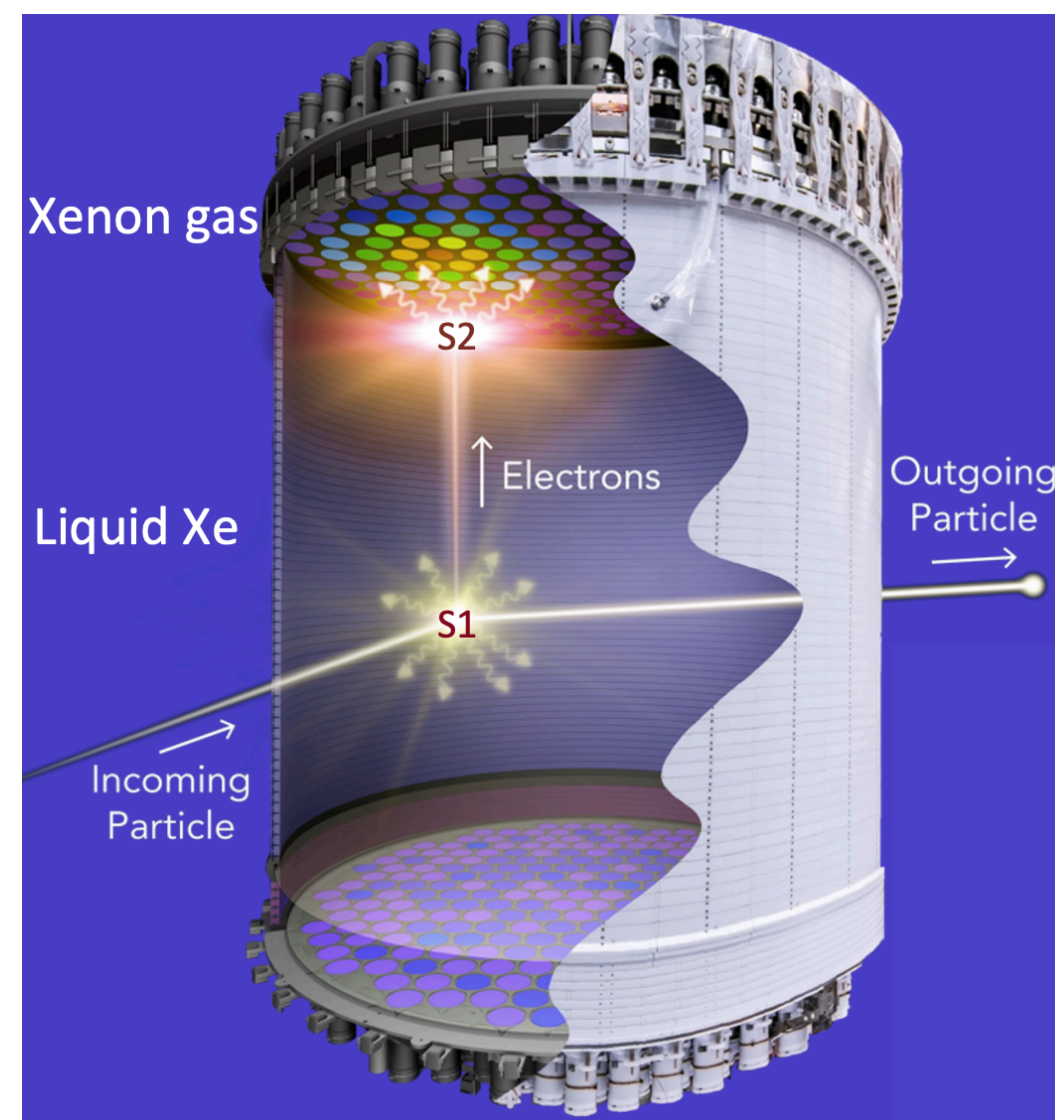
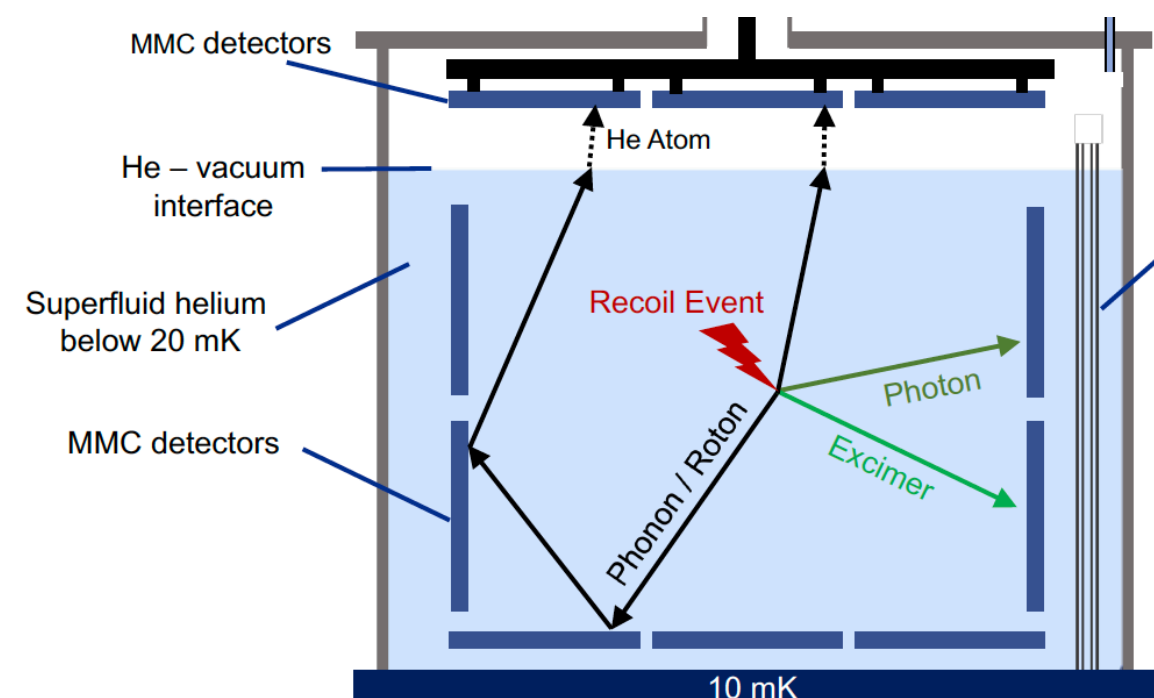
Liquid **scintillator**  
and **water** detectors



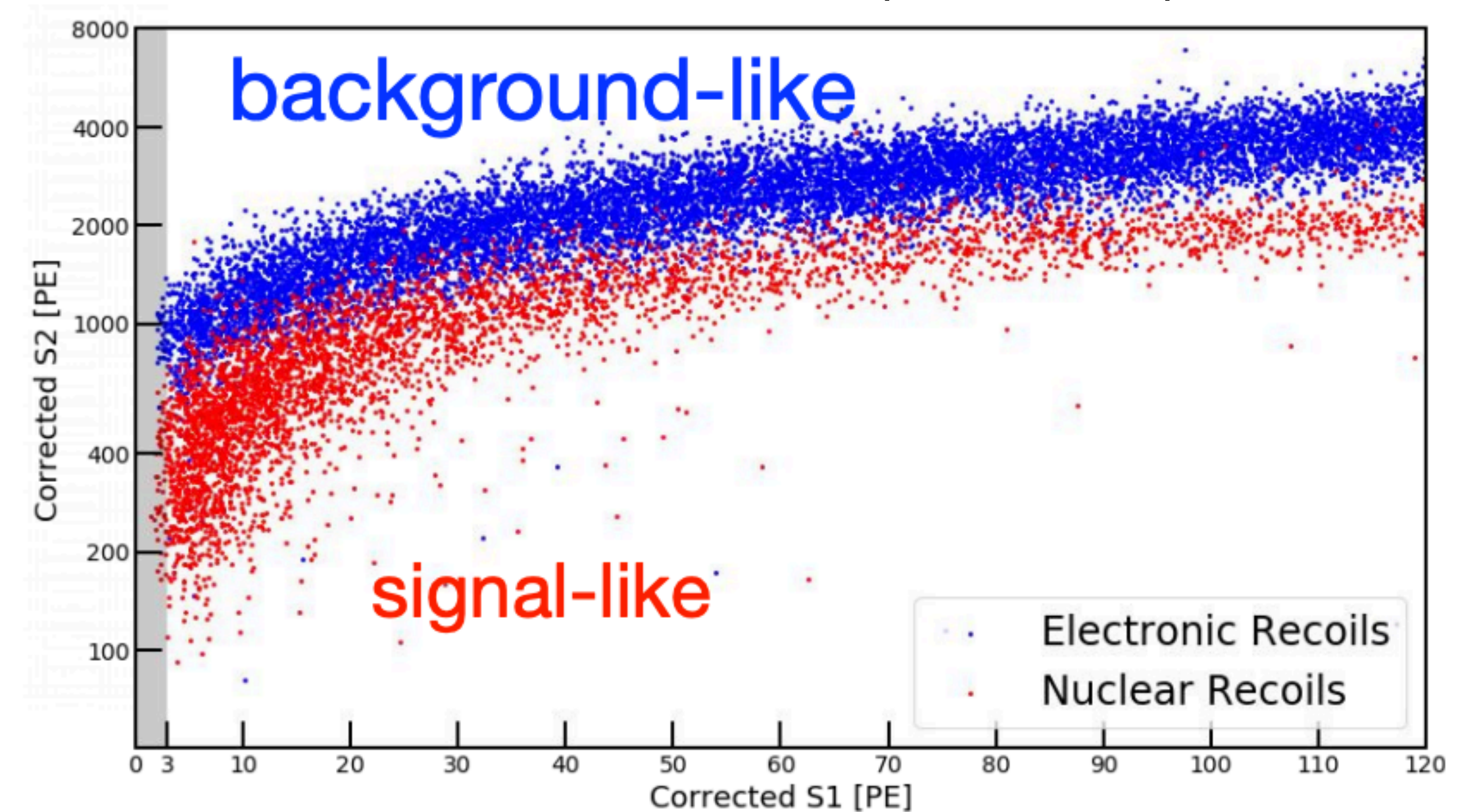
# 1 Properties of noble liquids

## ■ intrinsic key properties

- multiple signal channels (light and potentially charge or quasiparticles)
- high scintillation yield
- high, interaction-type-dependent charge or quasiparticle yield
- long electron lifetime



PSD in xenon (LXe TPC)



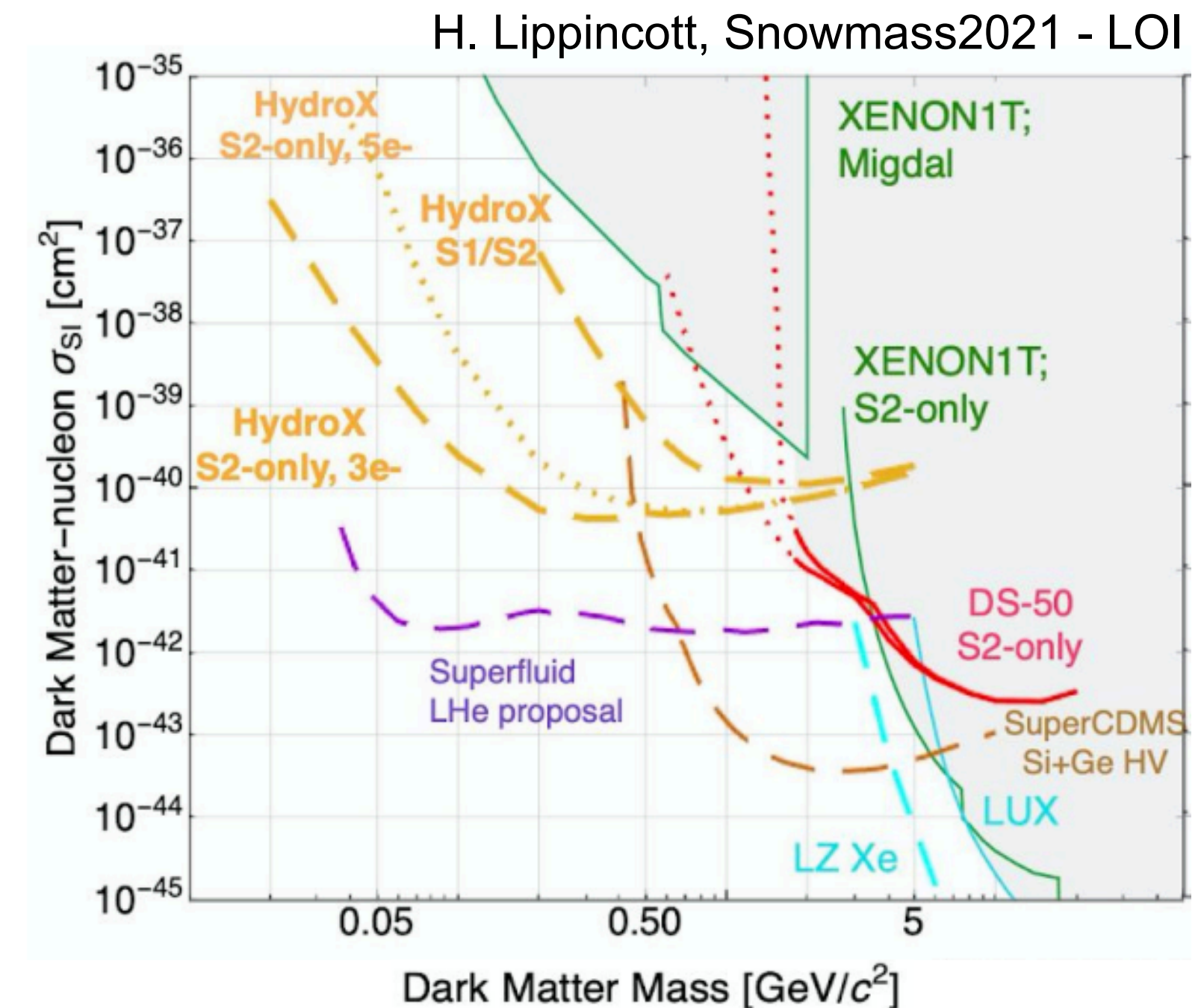
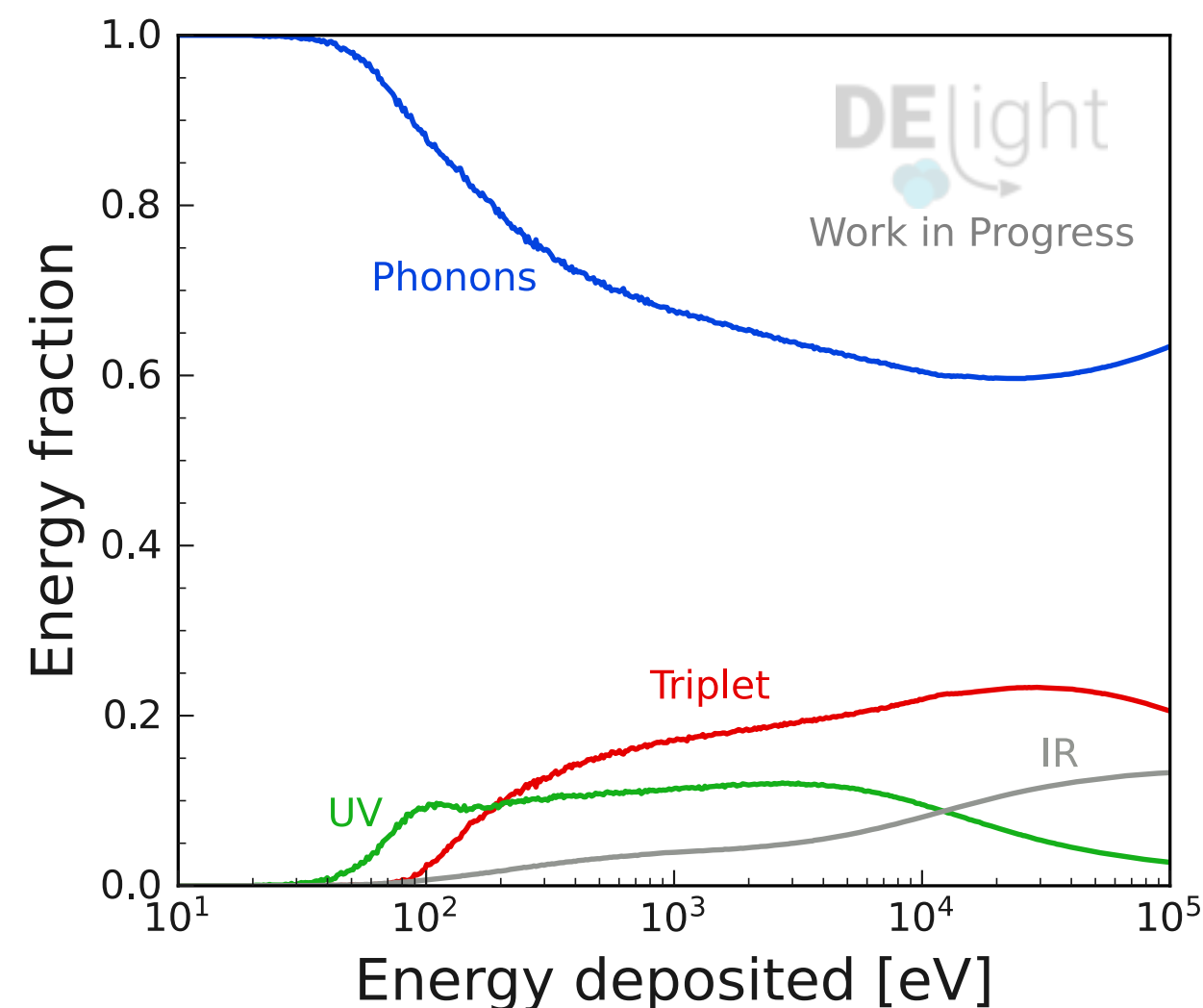


# 1 Properties of noble liquids - R&D

## ■ Microphysics of liquids

- Better understanding of energy **partitioning** to improve energy resolution and threshold
- **Doping techniques**
  - **Xe in Ar**: to shift light from VUV to UV for higher quantum efficiency (avoid WLS coatings)
  - **H<sub>2</sub> in Xe**: low-mass target to increase sensitivity at low DM masses (e.g. HydroX as upgrade to multi-ton scale Xe detectors)

Signal partitioning in LHe  
(nuclear recoil)





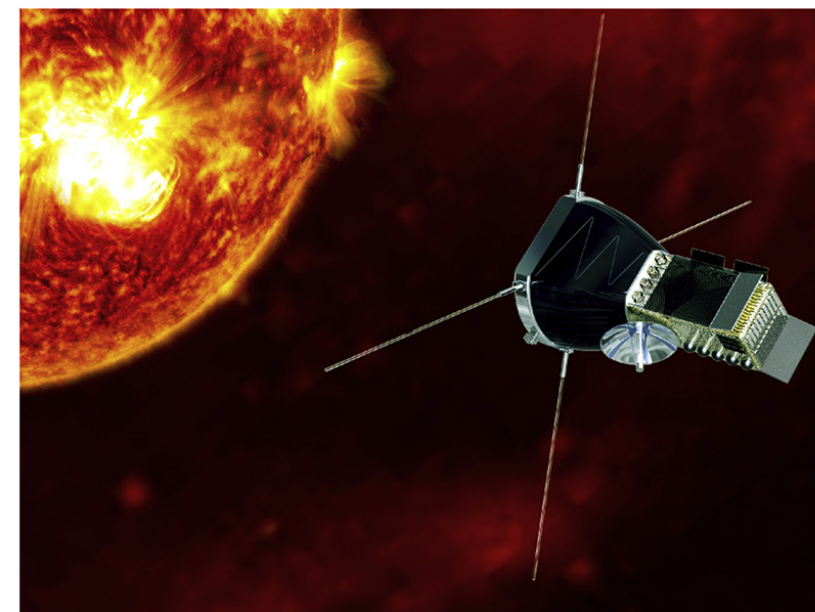
# 1 Properties of noble liquids - R&D

## Light emission / detection

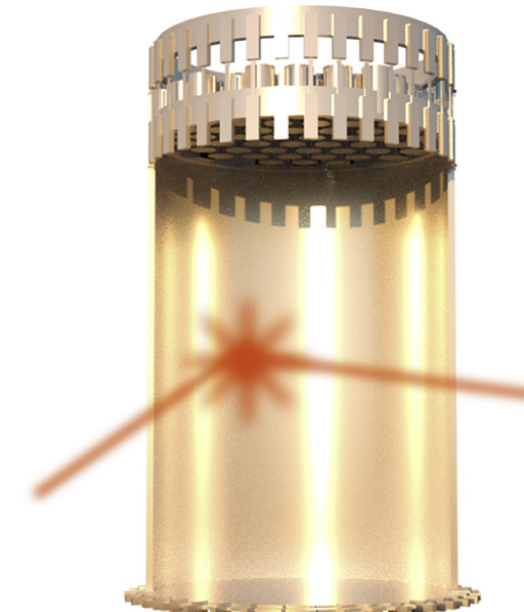
- developing photon detectors with **VUV-sensitivity**
- exploiting **sub-dominant emissions** (e.g. near-IR light component in LAr)

W. Zheng, Volume 23, Issue 6, 26 June 2020, 101145

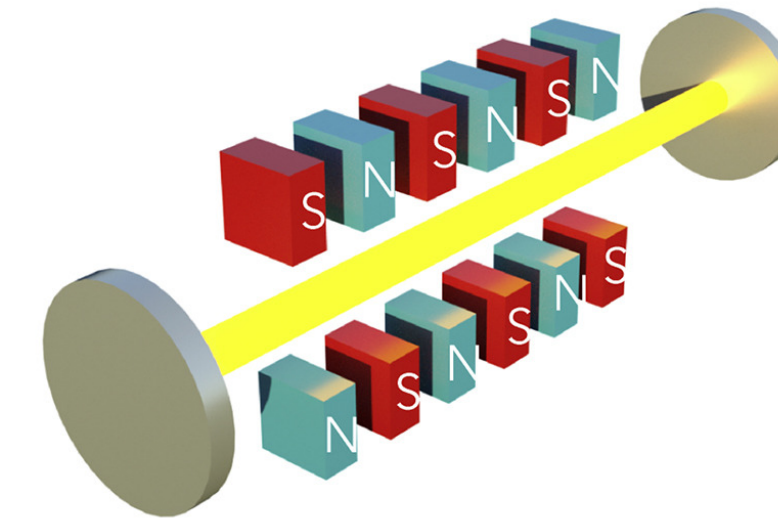
**A** Space science:  
VUV solar radiation monitoring



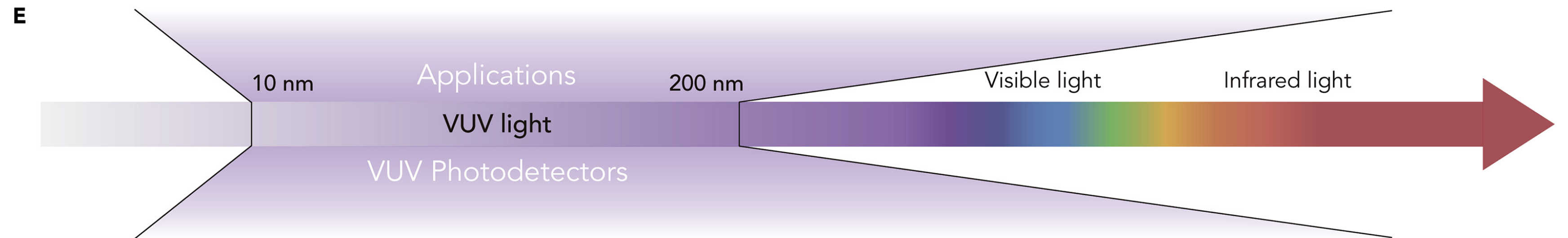
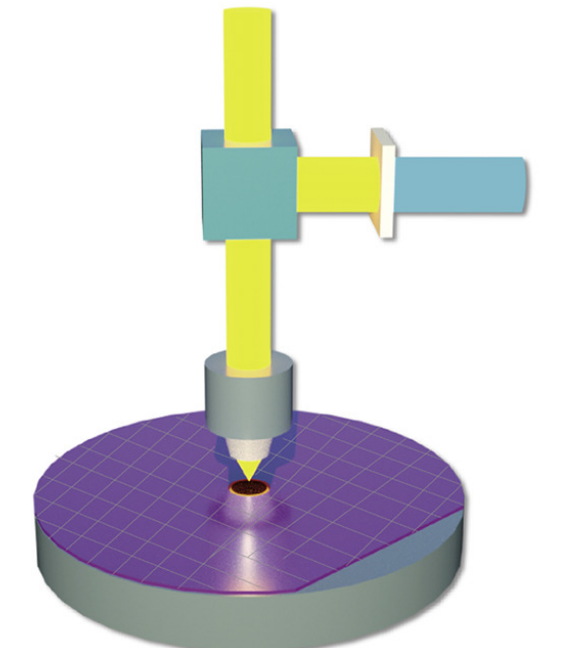
**B** High-energy physics:  
Detecting dark matter



**C** Large-scale scientific facility:  
VUV FEL radiation diagnosis



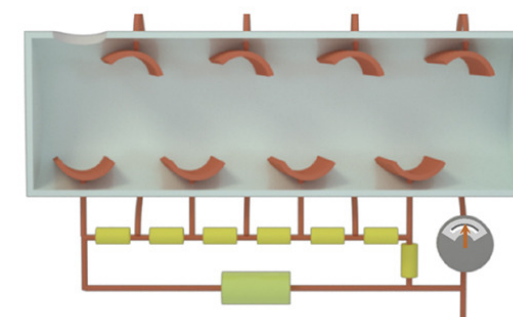
**D** Electronic industry:  
High-resolution lithography



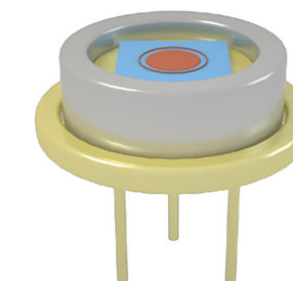
Scintillator detector



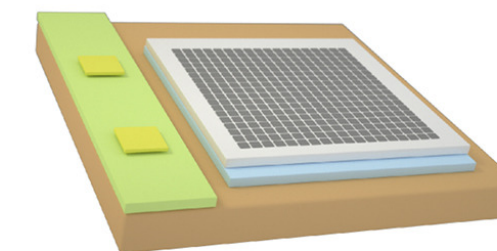
Photomultiplier tube



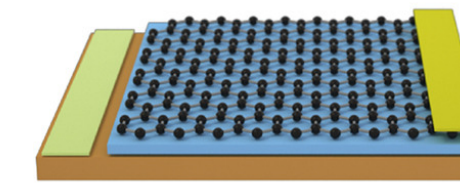
Silicon diode



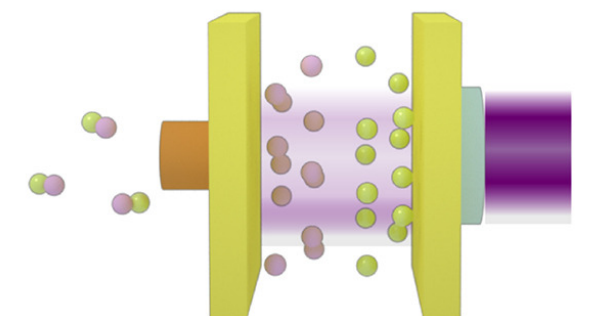
Silicon photomultiplier



Ultra-wide bandgap semiconductor photodetector



Gaseous detector

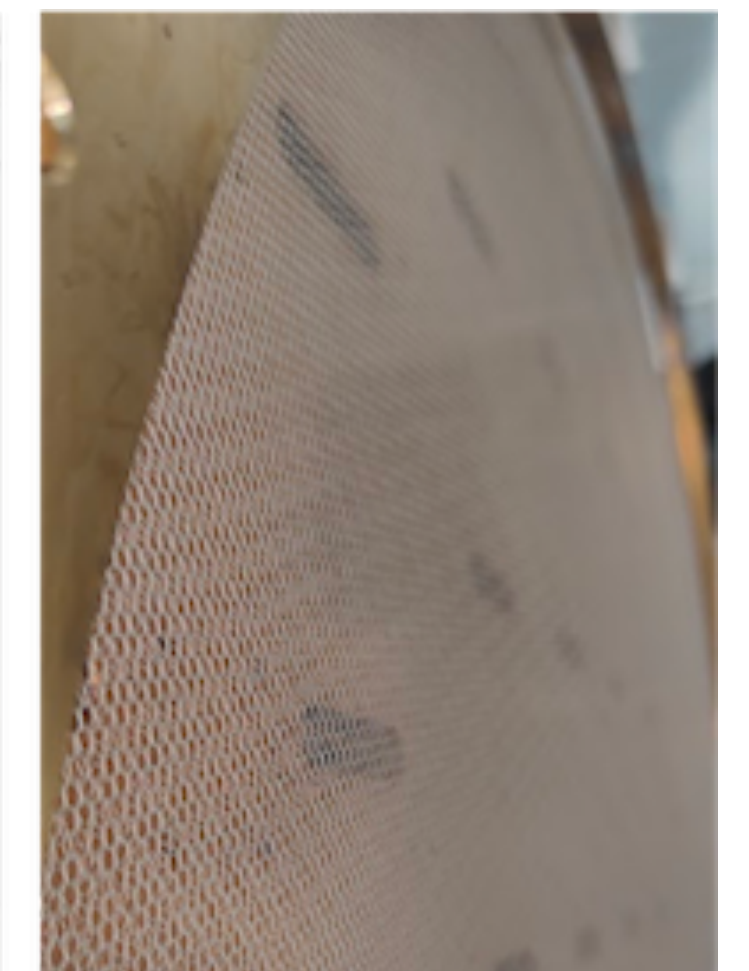
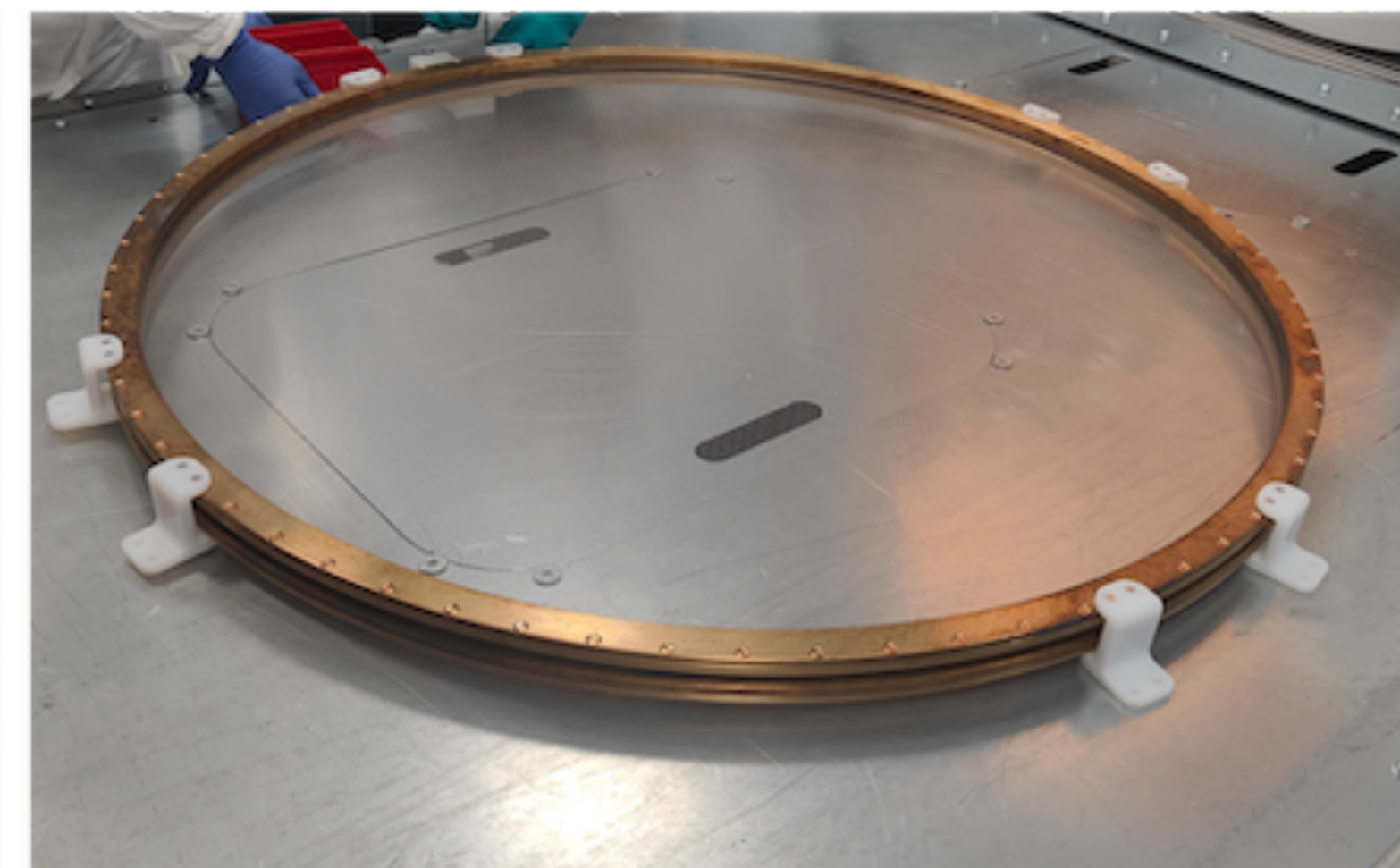
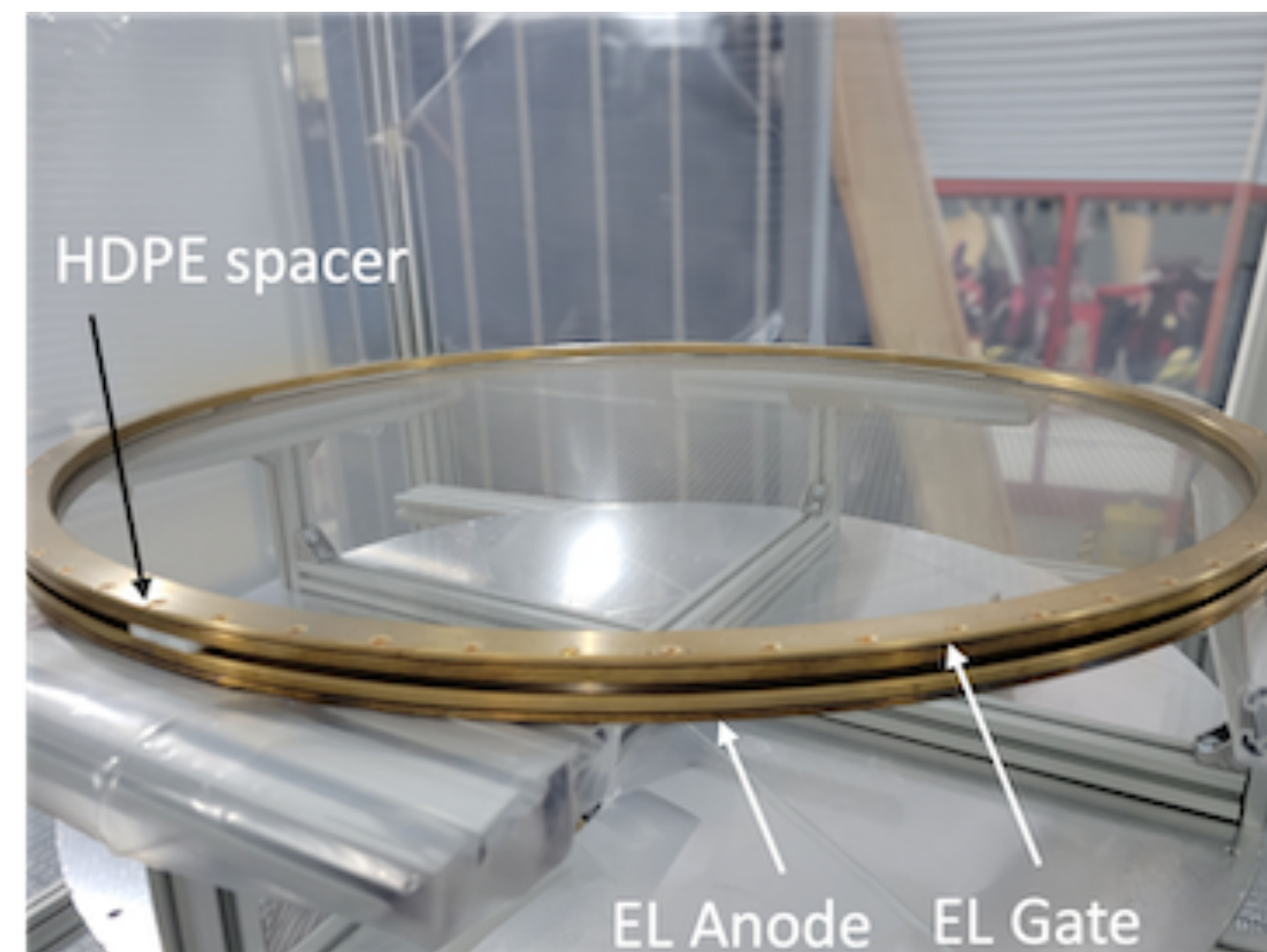
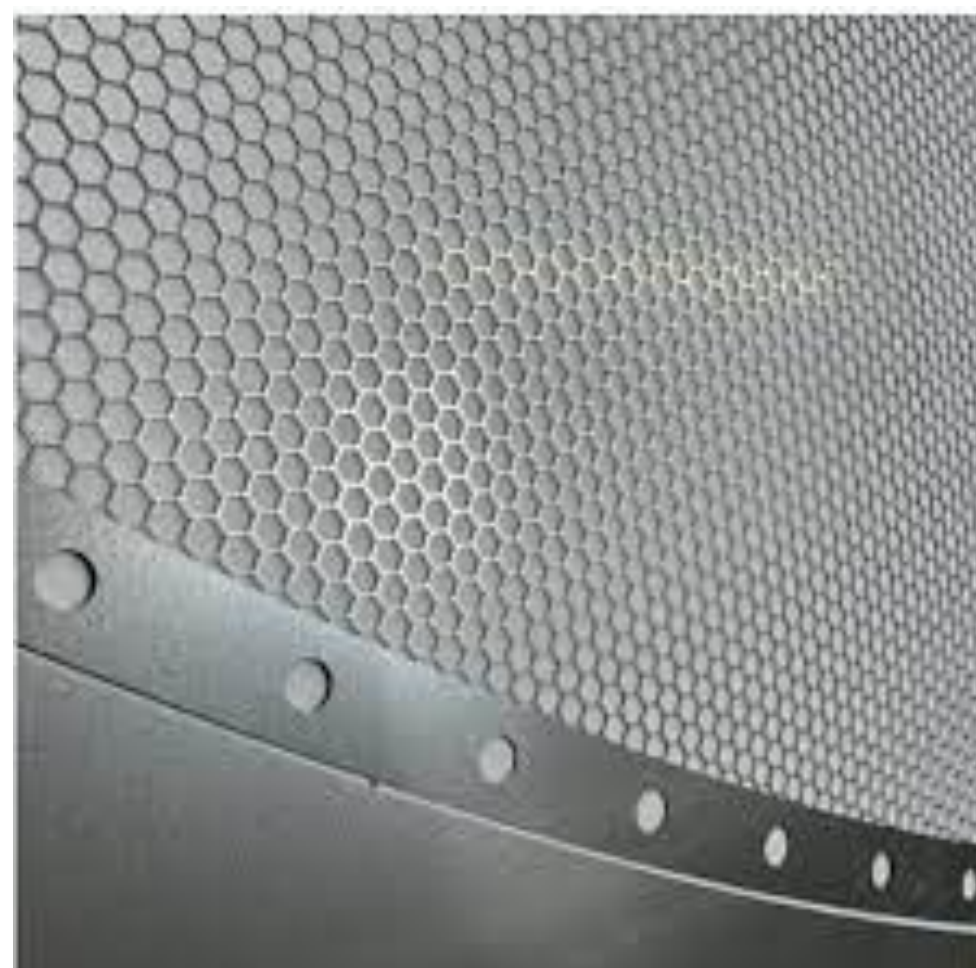




# 1 Properties of noble liquids - R&D

- Higher voltages to drift electrons over up to 10 m scale
  - **HV feed-throughs** must deliver 50 kV or more to the cathode
  - **Dielectric properties** of liquid must be better understood
  - Detailed models of light emission from **HV breakdown** required
  - **Electrodes** with large ( $>2.5$  m) diameters: wire, mesh/ woven, micro-pattern

NEXT-100, arXiv:2311.03528v2

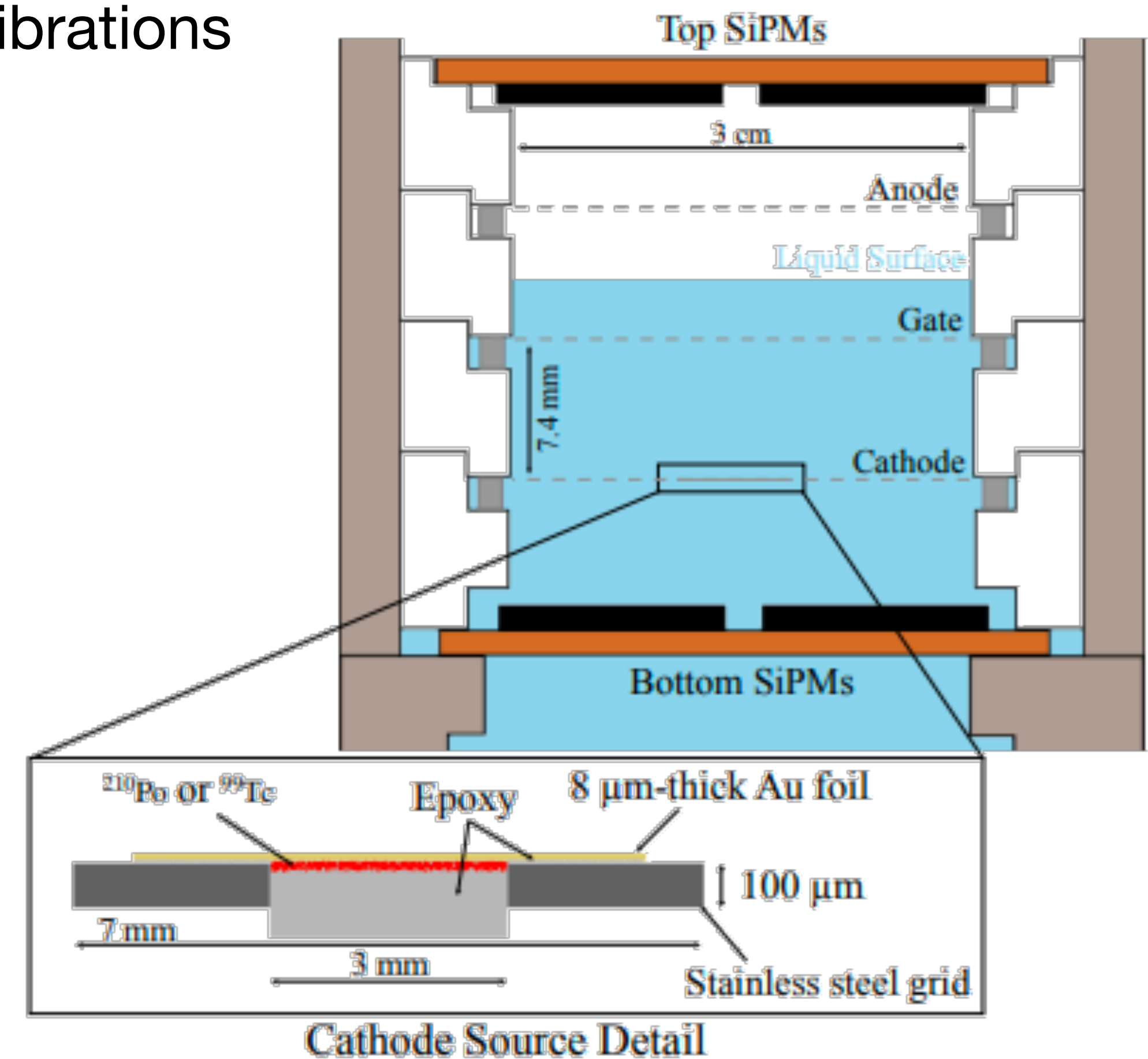
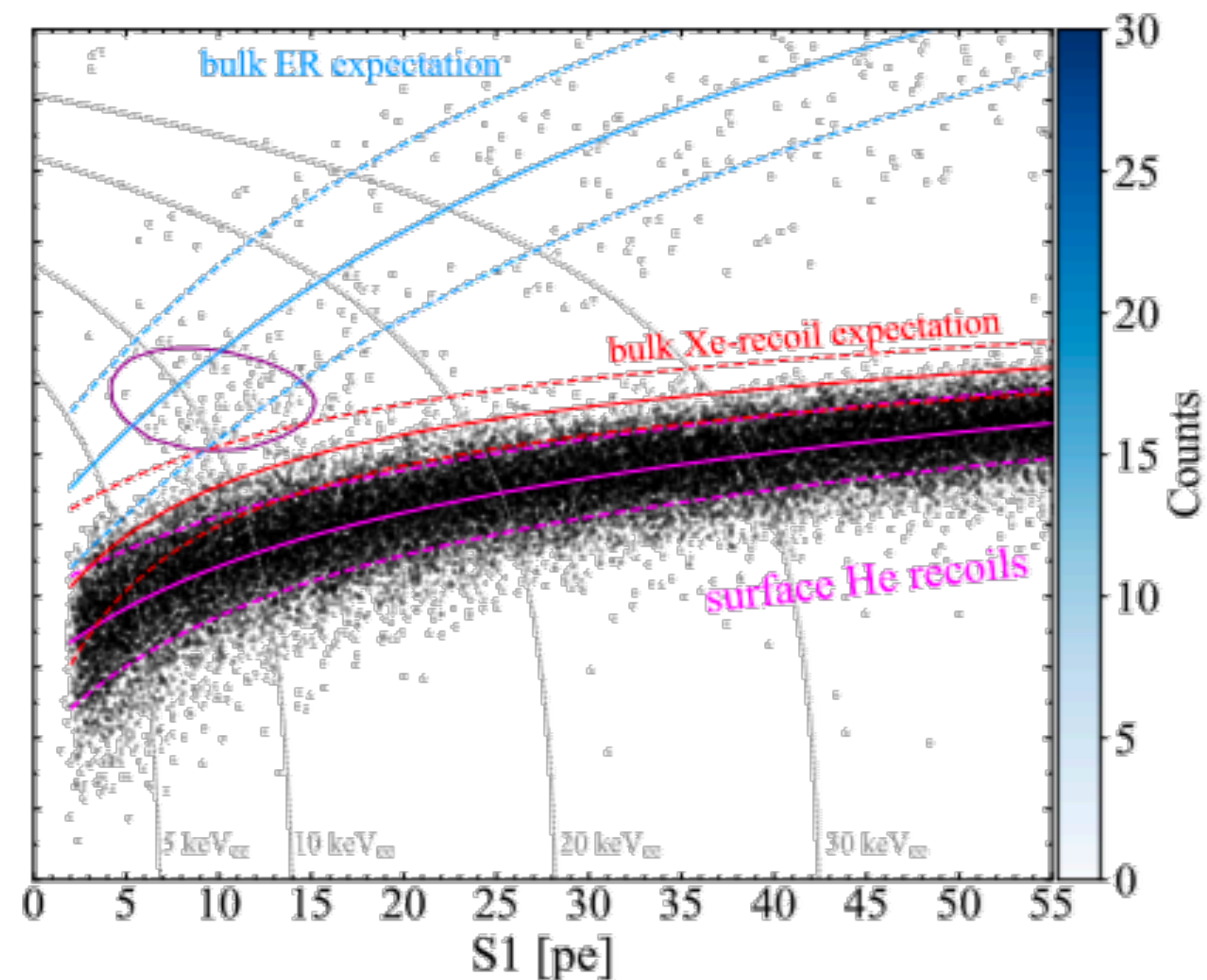




# 1 Properties of noble liquids - R&D

## ■ Calibration

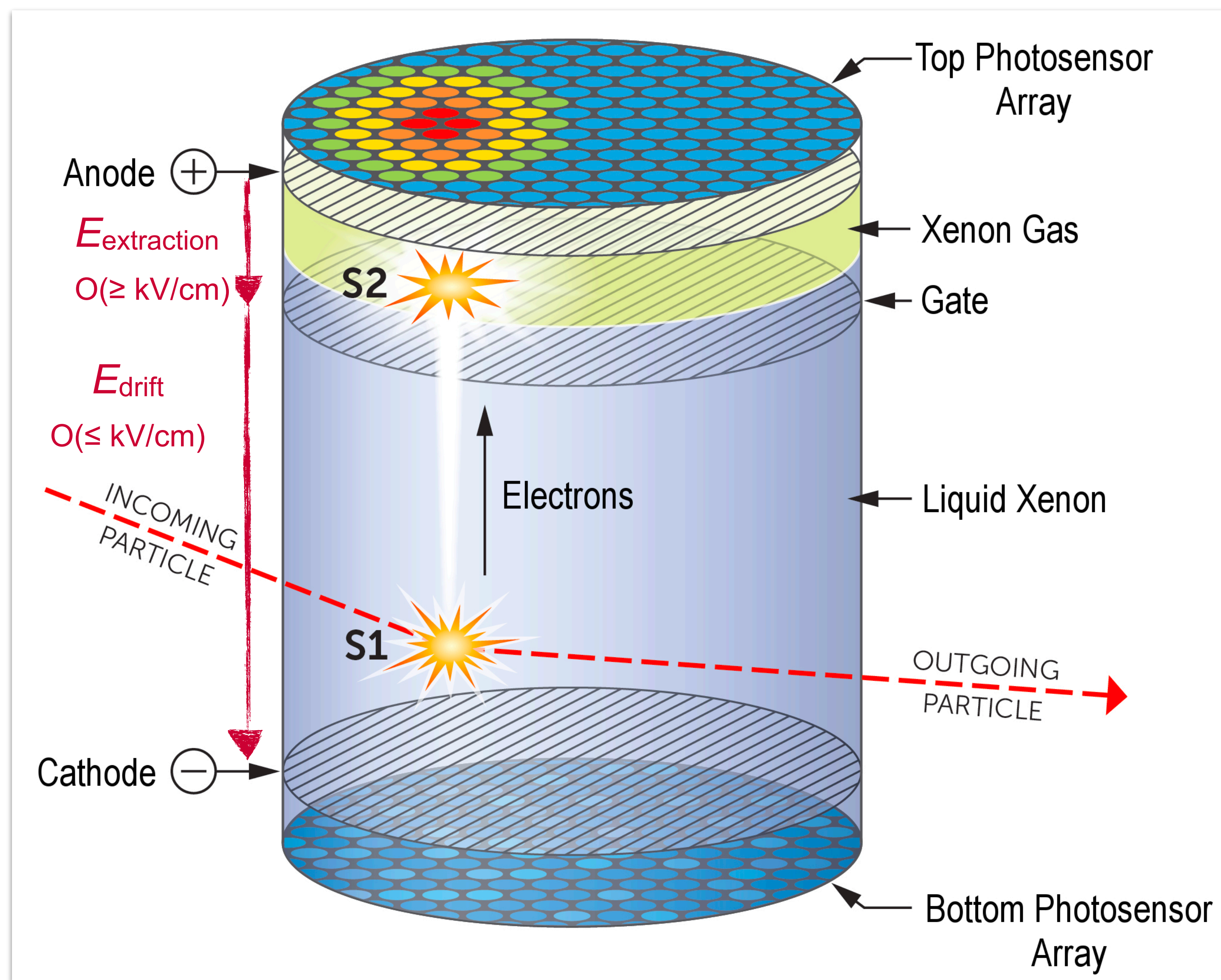
- New methods needed to ensure calibration capabilities throughout large volumes
- New methods needed for low energy nuclear recoil calibrations





## 2 Charge collection in noble liquids

- Charges from ionization with density of
  - 26,000 e<sup>-</sup>/cm for 4.0 MeV/cm in LXe
  - 9,000 e<sup>-</sup>/cm for 2.1 MeV/cm in LAr
- Overall acquired charge
  - **≥ 50,000 e<sup>-</sup> for LXe**
  - **~ 20,000 e<sup>-</sup> for LAr**
- O(1 mm/μs) drift velocity for 1 kV/cm E-field
- Charge diffusion during drift informs design granularity of the collecting anode

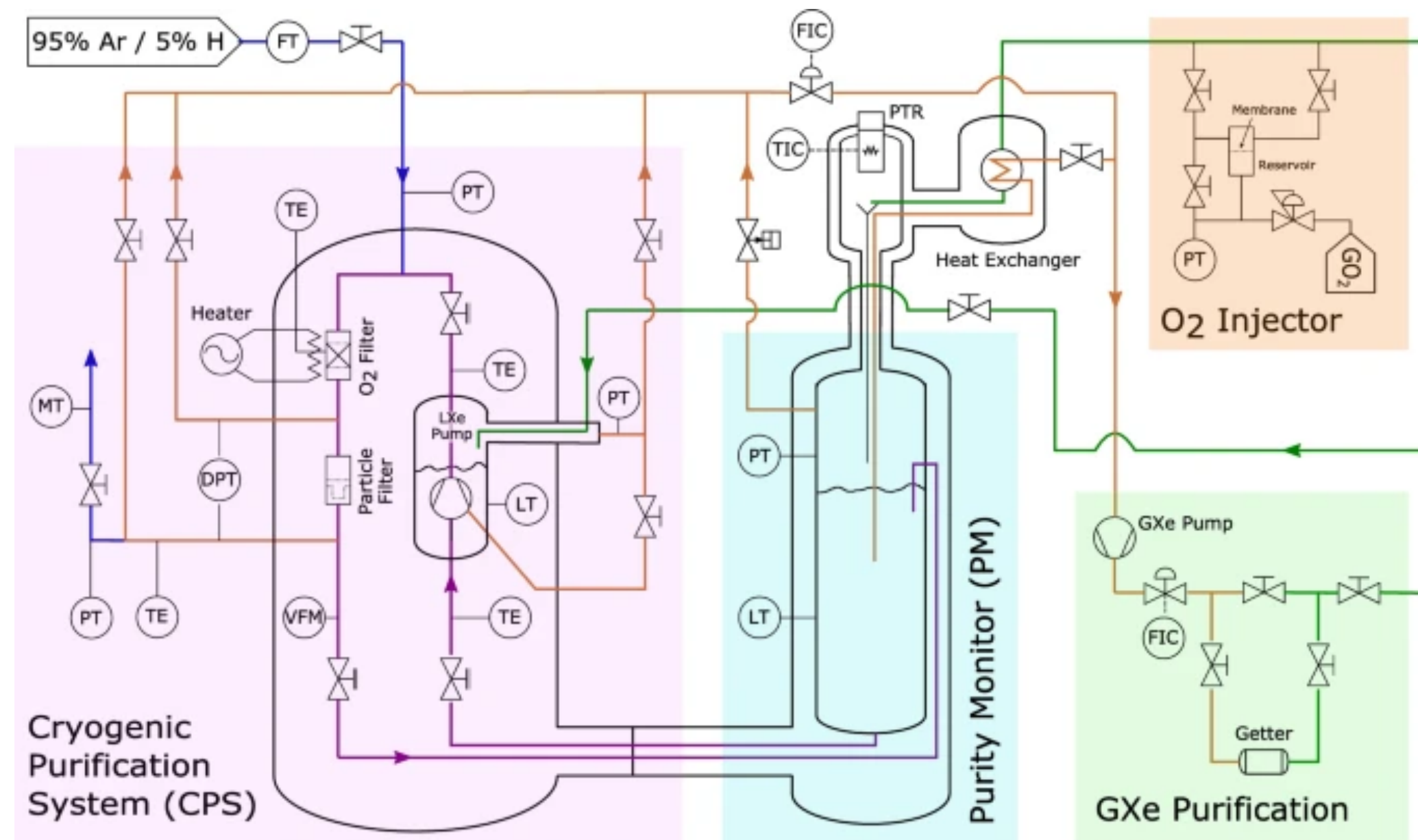




## 2 Charge collection in noble liquids - R&D

### ■ Liquid impurities

- LXe and LAr must be purified ( $H_2O$ , electronegative impurities) for high light and charge yield
- Mean  $e^-$  lifetimes for LXe / LAr around 1-10 ms in a  $\sim 9$  t total mass for impurity concentration around 10 ppt
- Modern technologies allow notably lower impurity concentrations



Xeclipse apparatus at Columbia University, E. Aprile et al.



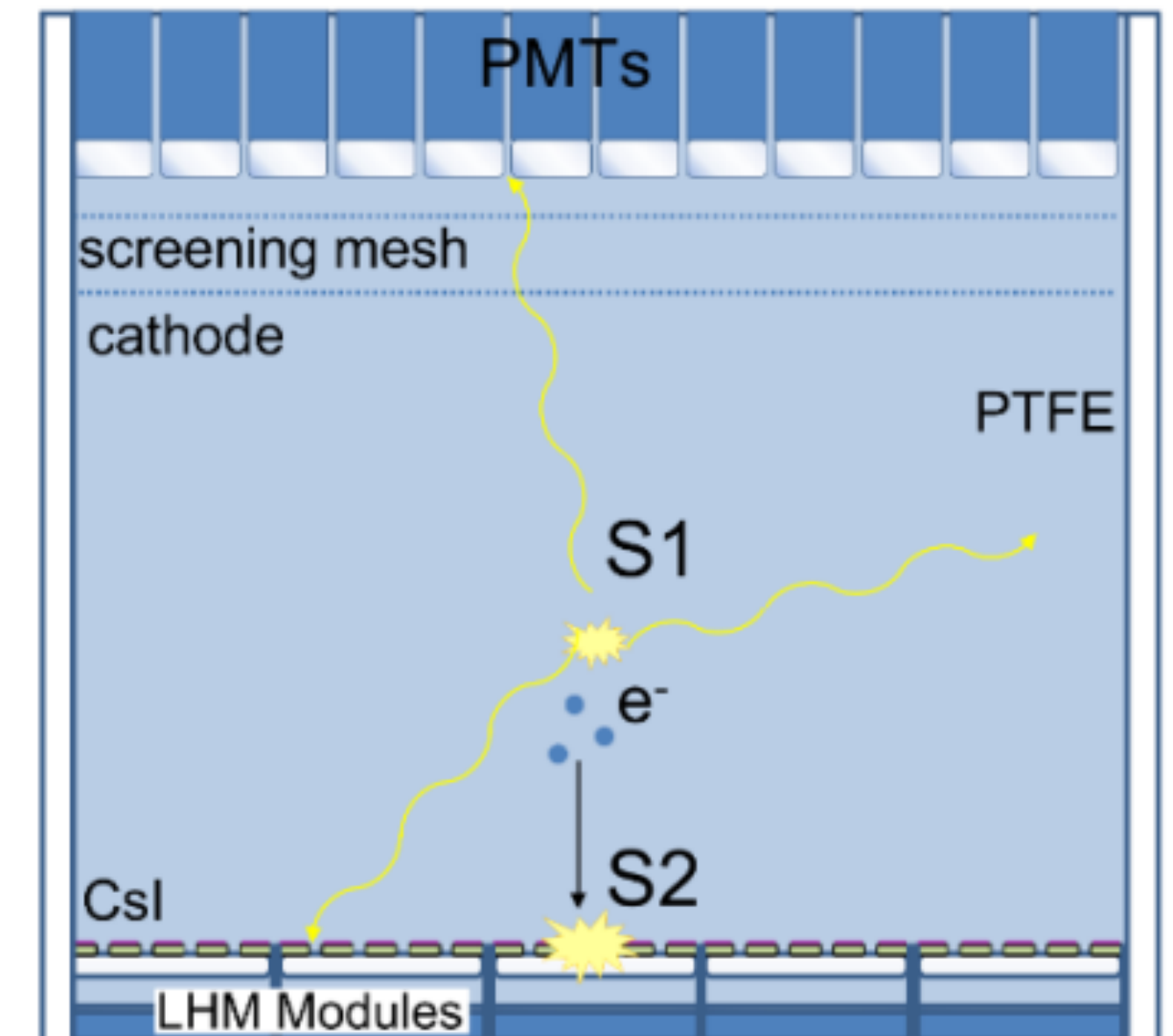
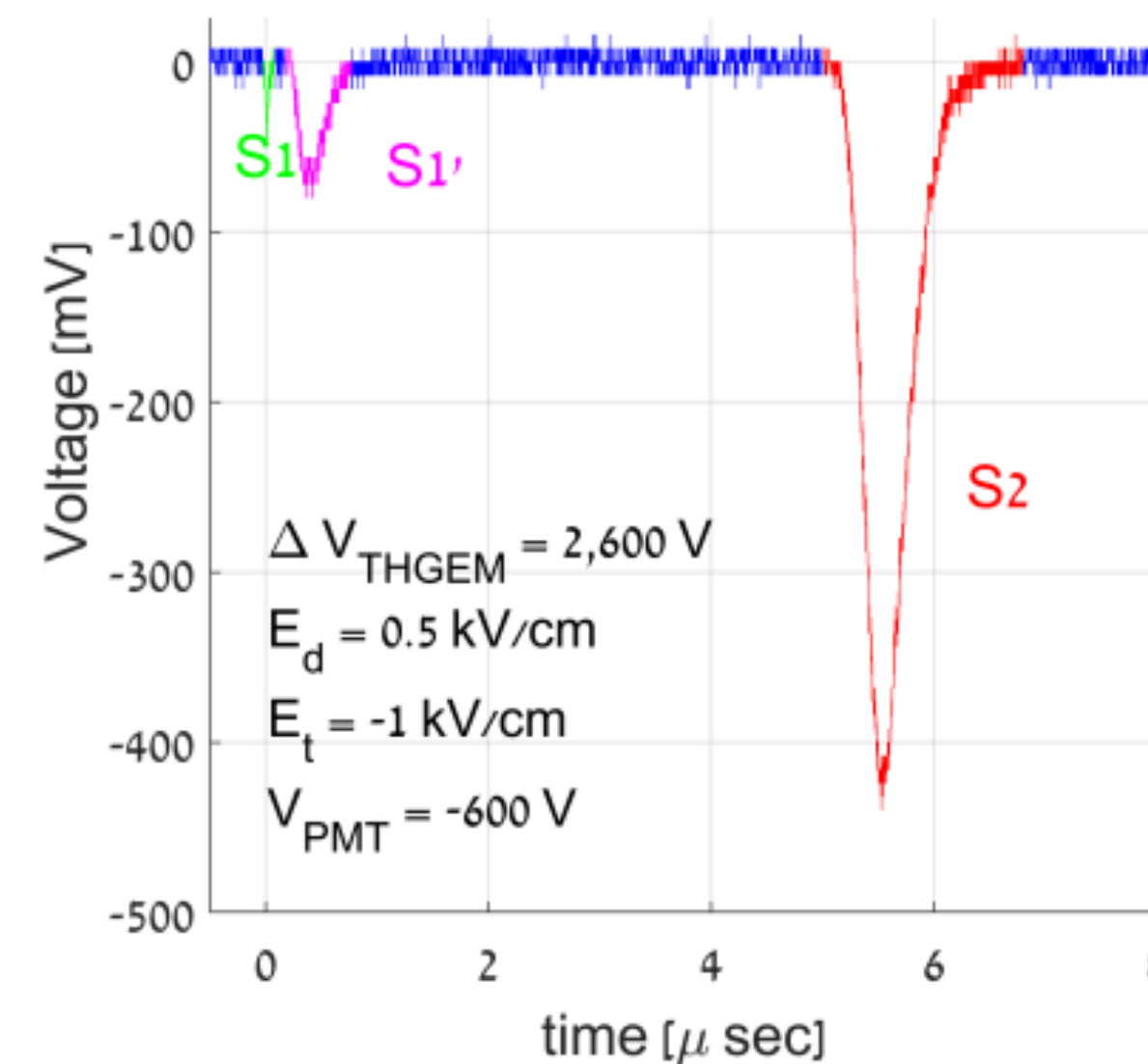
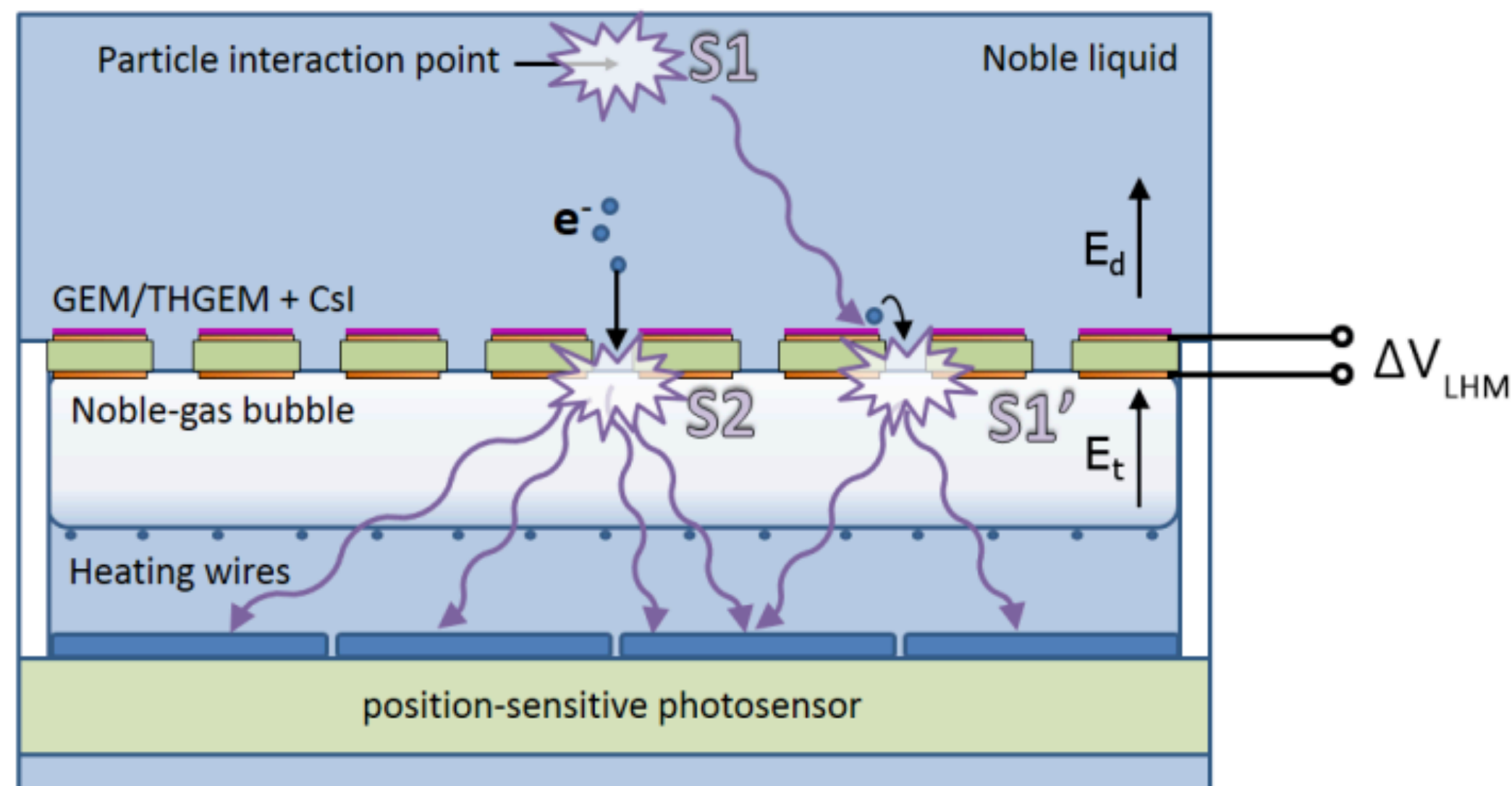


## 2 Charge collection in noble liquids - R&D

A. Breskin, 2022 JINST 17 P08002

### ■ Charge amplification

- several orders of magnitude **amplification in gas phase** via electroluminescence
- **amplification in liquid** requires R&D
  - doping LAr with small percentages of LXe suggests amplification by factor  $\geq 100$  in LAr
  - exploitation of localized electroluminescence suggests potential amplification in LXe





# 2 Charge collection in noble liquids - R&D

Images from L. Baudis and J. Asaadi

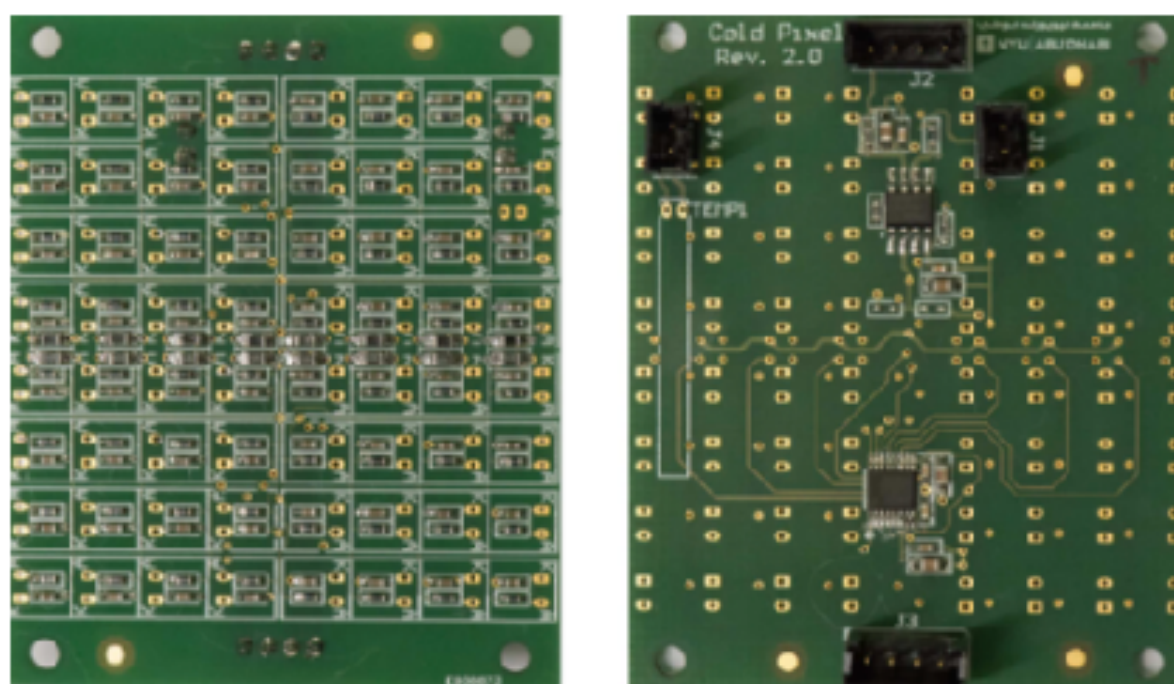
## Charge readout structures

- High lumi applications (e.g. LBN near-detectors): **Pixel electrodes on PCBs** instead of “standard” wire planes for X-Y measurement, to eliminate risk of reconstruction ambiguities

- Amplification stability**

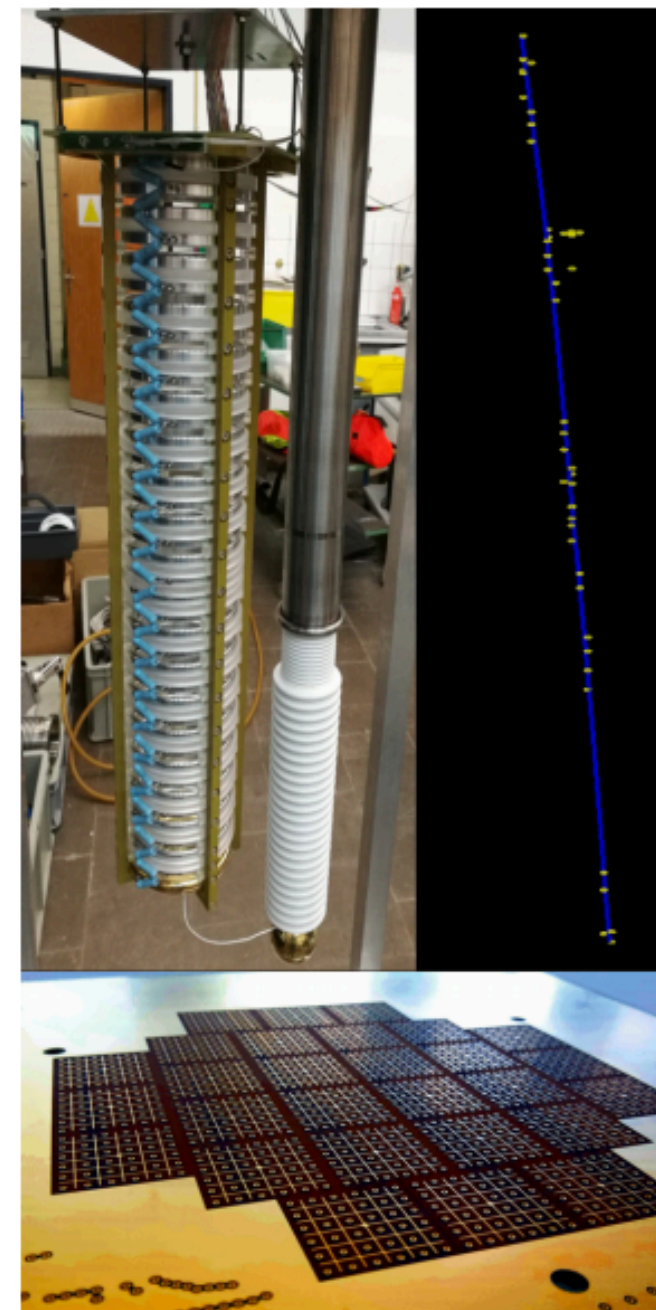
- Cryogenic low-noise, low-radioactivity, low-heat dissipation readout**

- Integrated, simultaneous readout of light and charge**



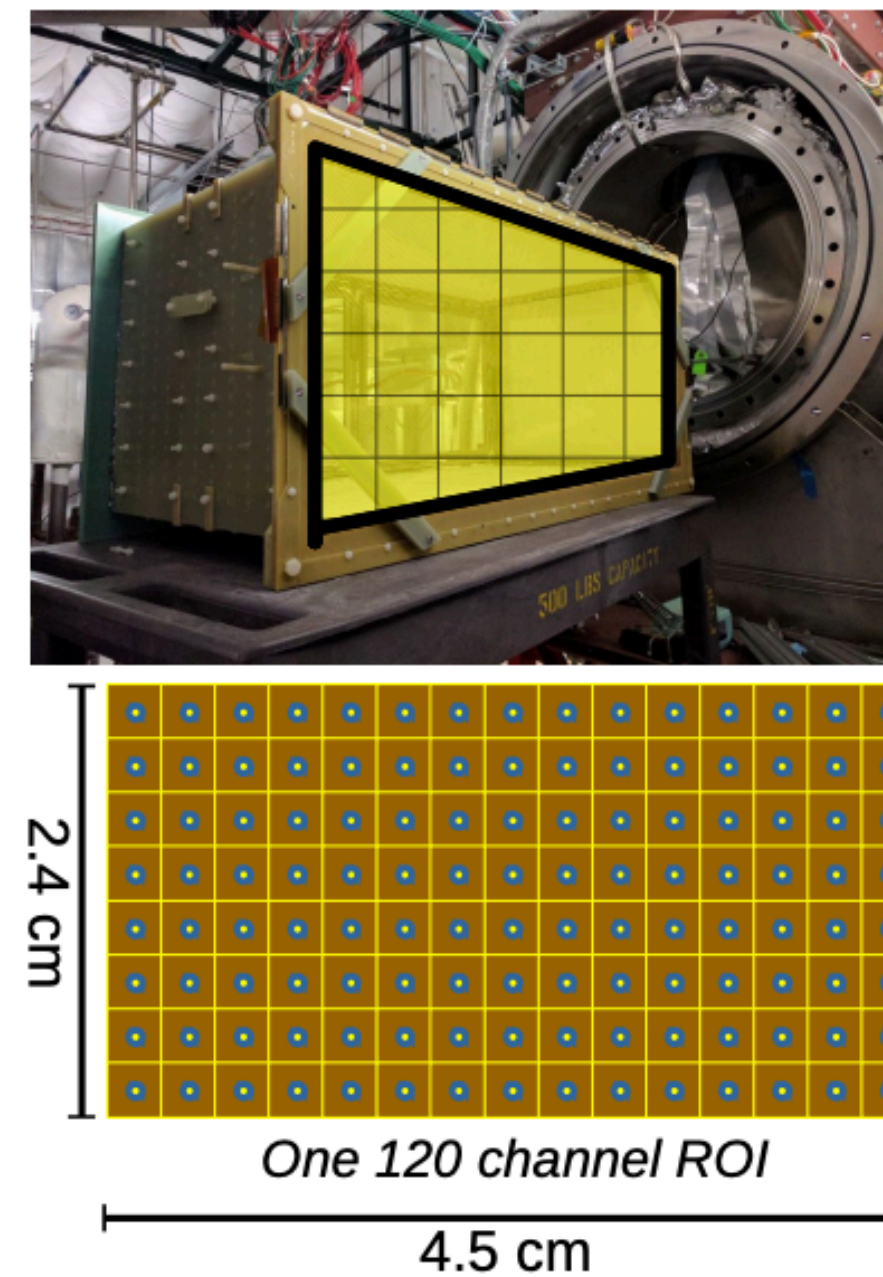
Cryogenic preamp for SiPMs, NIM 936, 2019

**Demonstrated  
Summer 2016**



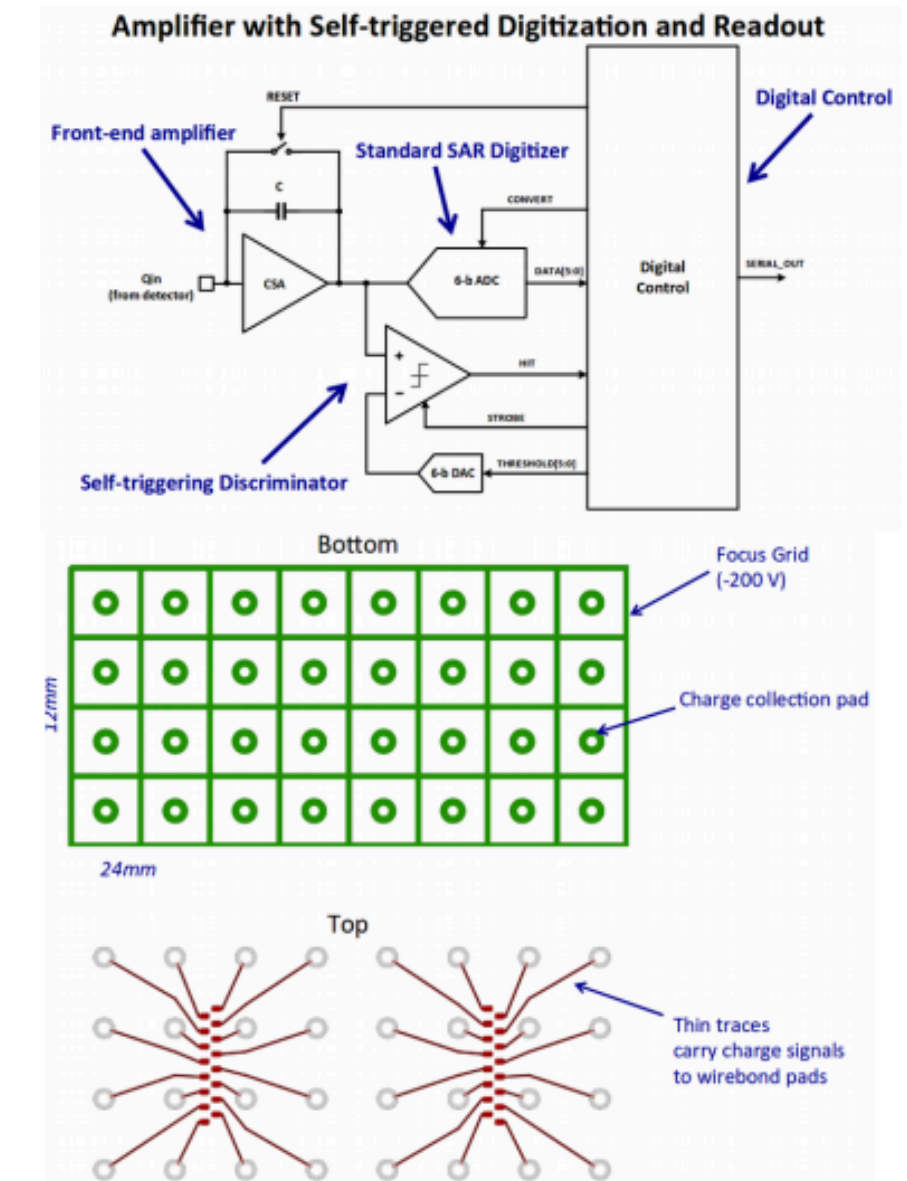
**Multiplexed Pixel  
LArTPC Demonstrator**  
(1008 pixels readout with 32 channels)

**Taking data Nov-Dec 2017**



**Highly-multiplexed Pixel  
LArTPC in a Testbeam**  
(28,800 pixels readout with 480 channels)

**Benchtop chip tests  
December 2017**



**Truely 3d front end ASIC  
for LArTPC digital LArTPC  
Pixel Readout (LARPIX)**

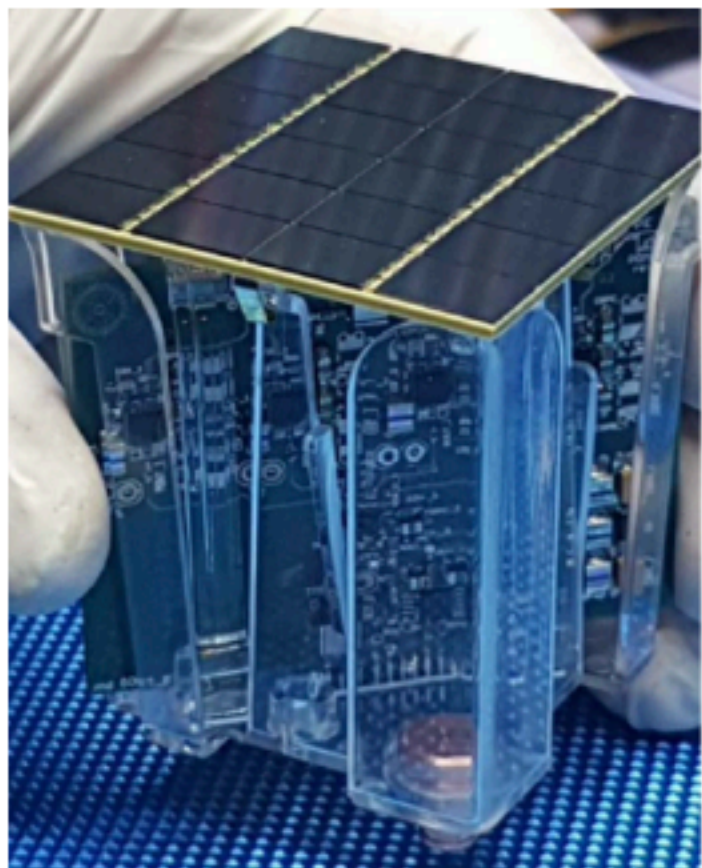


# 2 Charge collection in noble liquids - R&D

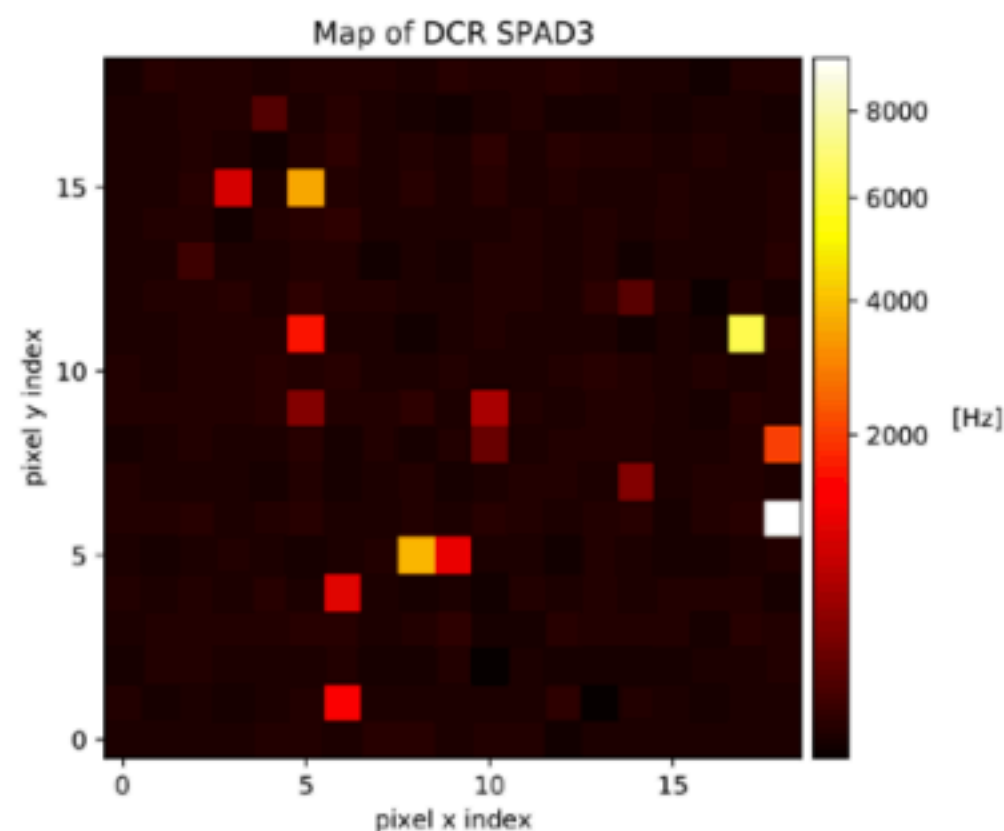
## ■ Light and charge

- **Photomultipliers**: established technology, low DCR ( $\sim 0.02$  Hz/mm<sup>2</sup>), high QE (mean around 34%, up to  $> 40\%$  at 175 nm)
  - *issues*: lower radioactivity required, after-pulsing due to vacuum leaks and light emission
- **SiPM arrays**: lower radioactivity / area, lower voltage
  - *issues*: dark count rate (too high by  $\sim$  factor 50 at least)
  - $\Rightarrow$  low-field SiPMs (reduce band-to-band tunneling), digital SiPMs

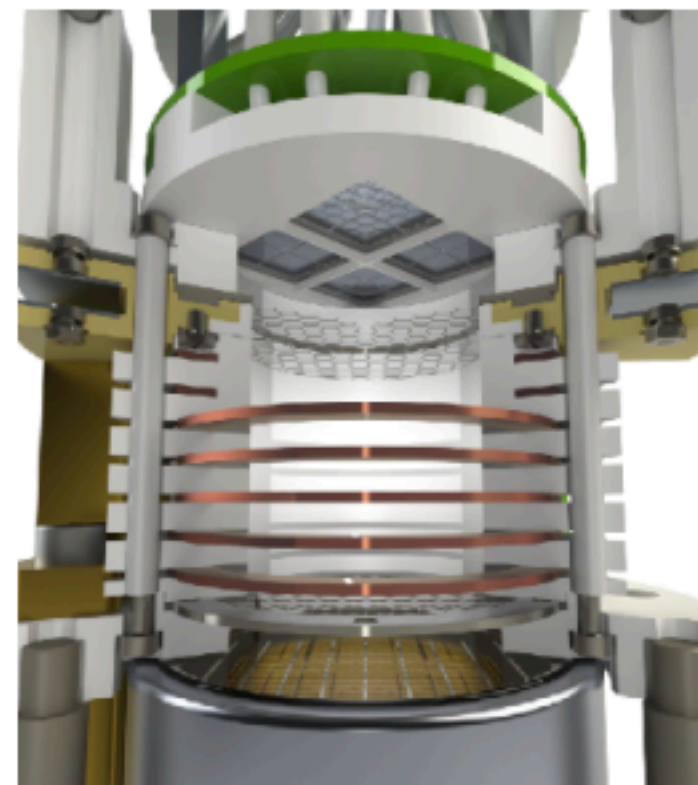
SiPM arrays for Darkside-20k



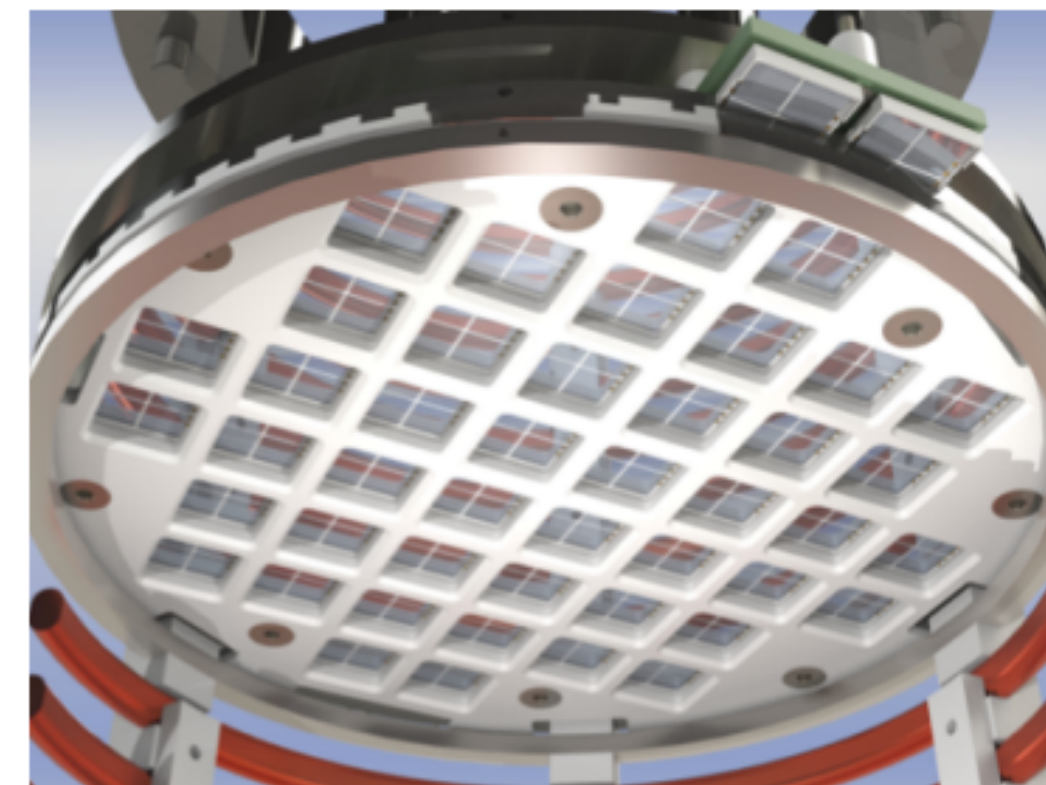
Digital SiPM



Two-phase TPC with SiPM array

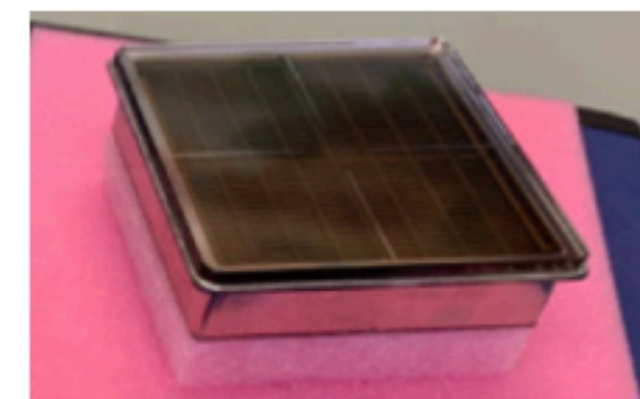


EPJ-C 80, 2020



SiPM array, DARWIN demo

PMT array for XENONnT



2''x 2'' flat panel PMT (R12699) R&D for DARWIN



3'' (R1311 low-rad PMT by XMASS), JINST 15, 2020

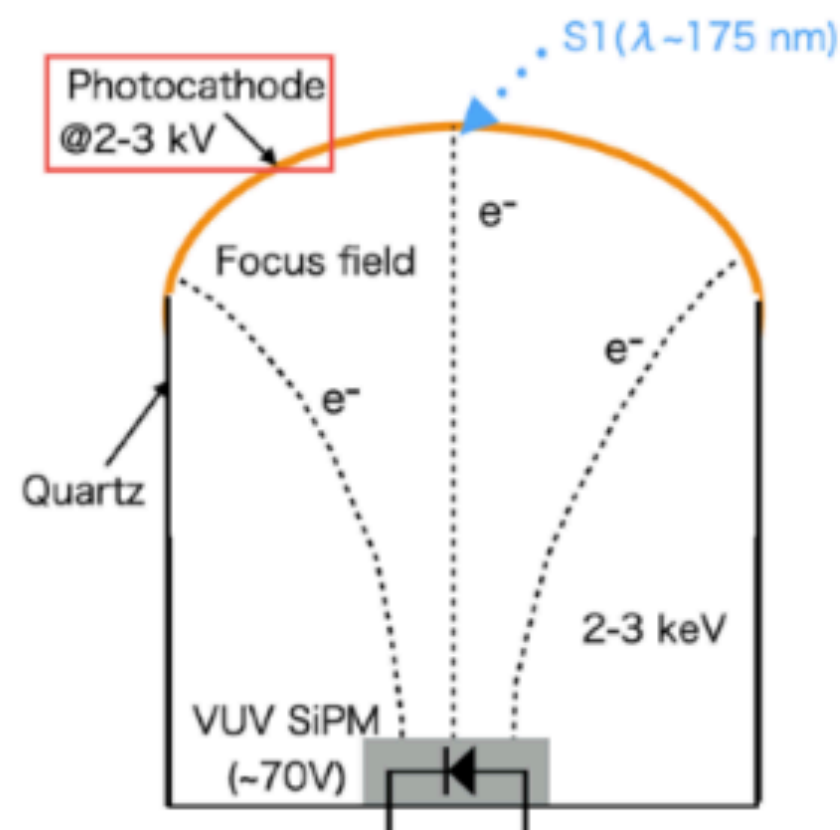
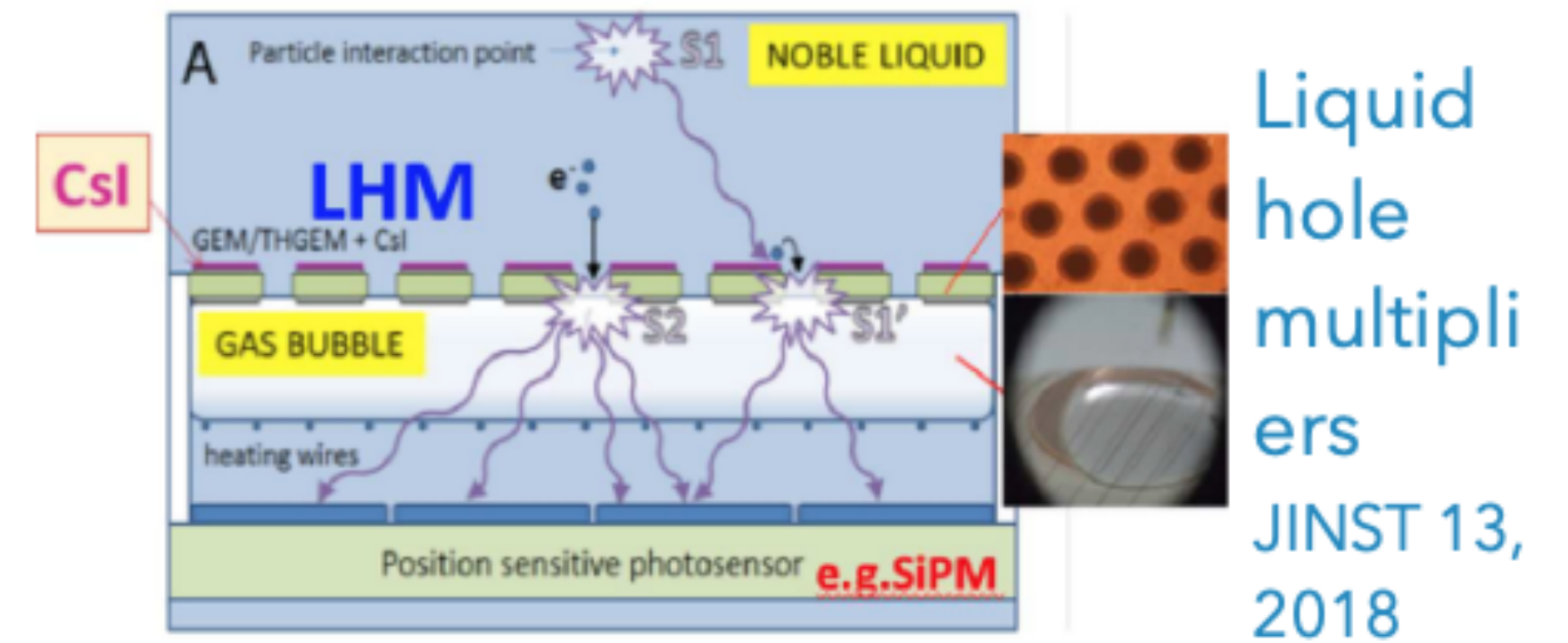


# 2 Charge collection in noble liquids - R&D

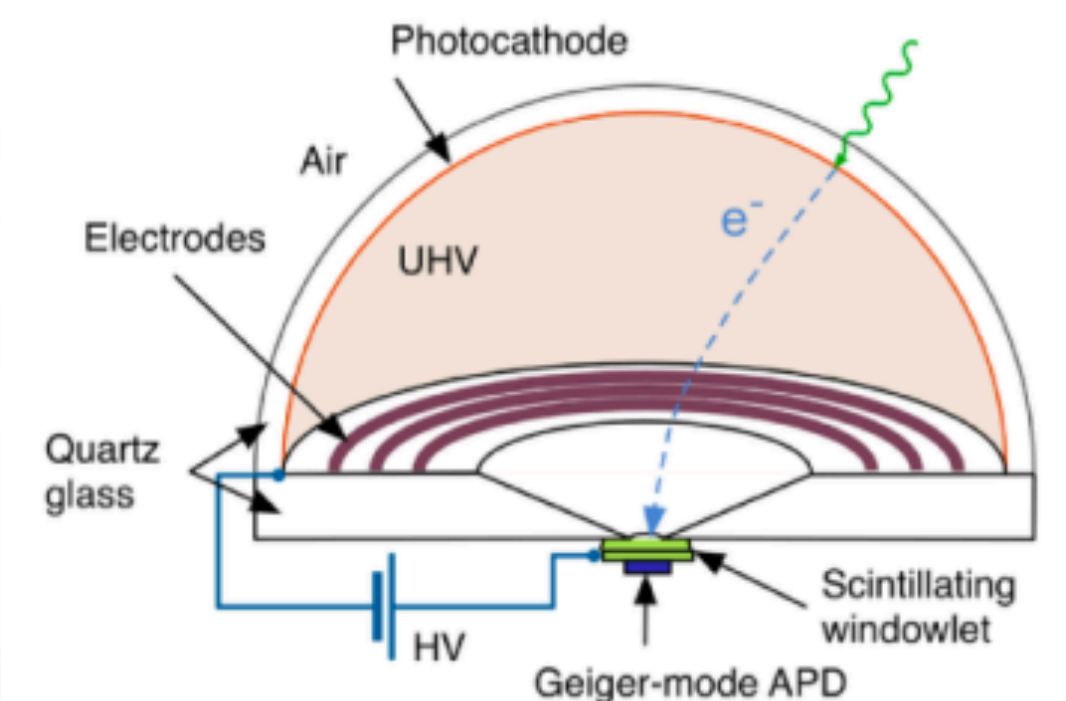
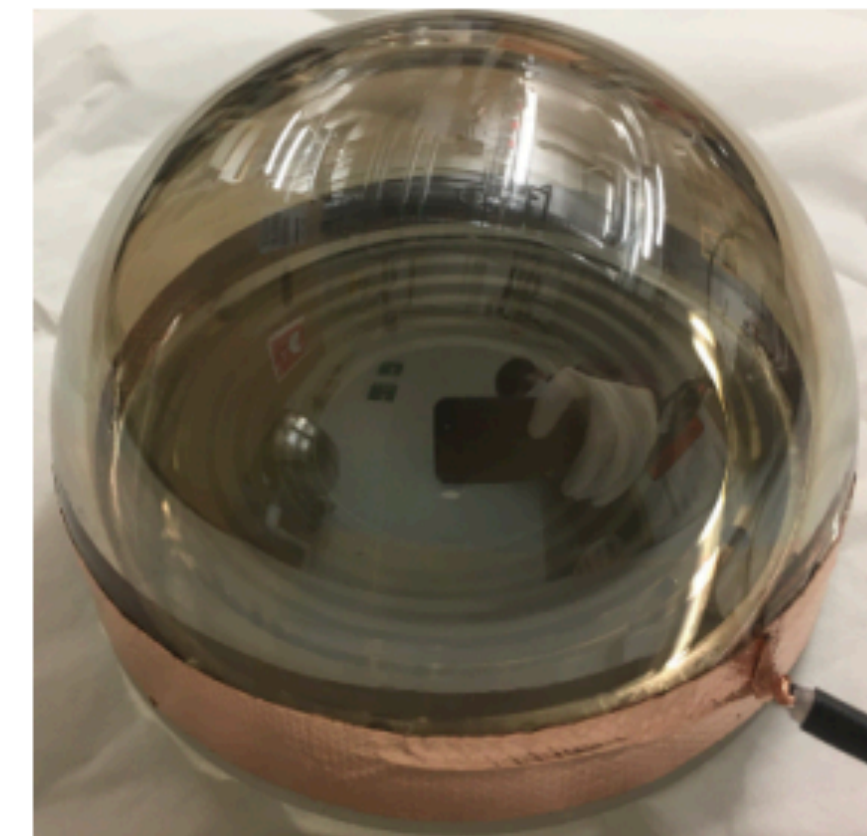
Images from L. Baudis

## Light and charge

- **Hybrid sensors:** e.g., ABALONE, VSiPM, SIGHT
  - SiPM + Quartz + photocathode: reduced radioactivity compared to PMTs
  - lower DCR compared to SiPM arrays (photosensitive area difference)
- **Bubble-assisted Liquid Hole Multipliers:** local vapor bubble underneath GEM-like perforated electrode in LXe



Hybrid photosensor: Hamamatsu XE5859



Hybrid photosensor: ABALONE; left (DARWIN R&D with SiPM); right: NIM 954, 2020





## 2 Charge collection in noble liquids - R&D

- Cryogenic front-end electronics
  - Number of readout channels ~ number of sensors and/or meters-long TPC wires
  - Challenges for cryogenic front-end electronics include heat dissipation
  - Cryogenic optical transmission of signals (e.g. in DUNE and DarkSide-20k) requires transmission driver and receiver development
  - Pixel electronics require independent front-end channels for each pixel, likely operating in cold (even immersed in the cryogenic liquid) with a potentially destructive dissipation heat
  - The handling of high bandwidths, the improvement of the S/N ratio, and the reduction of cost/channel will be crucial for large-scale pixel readout

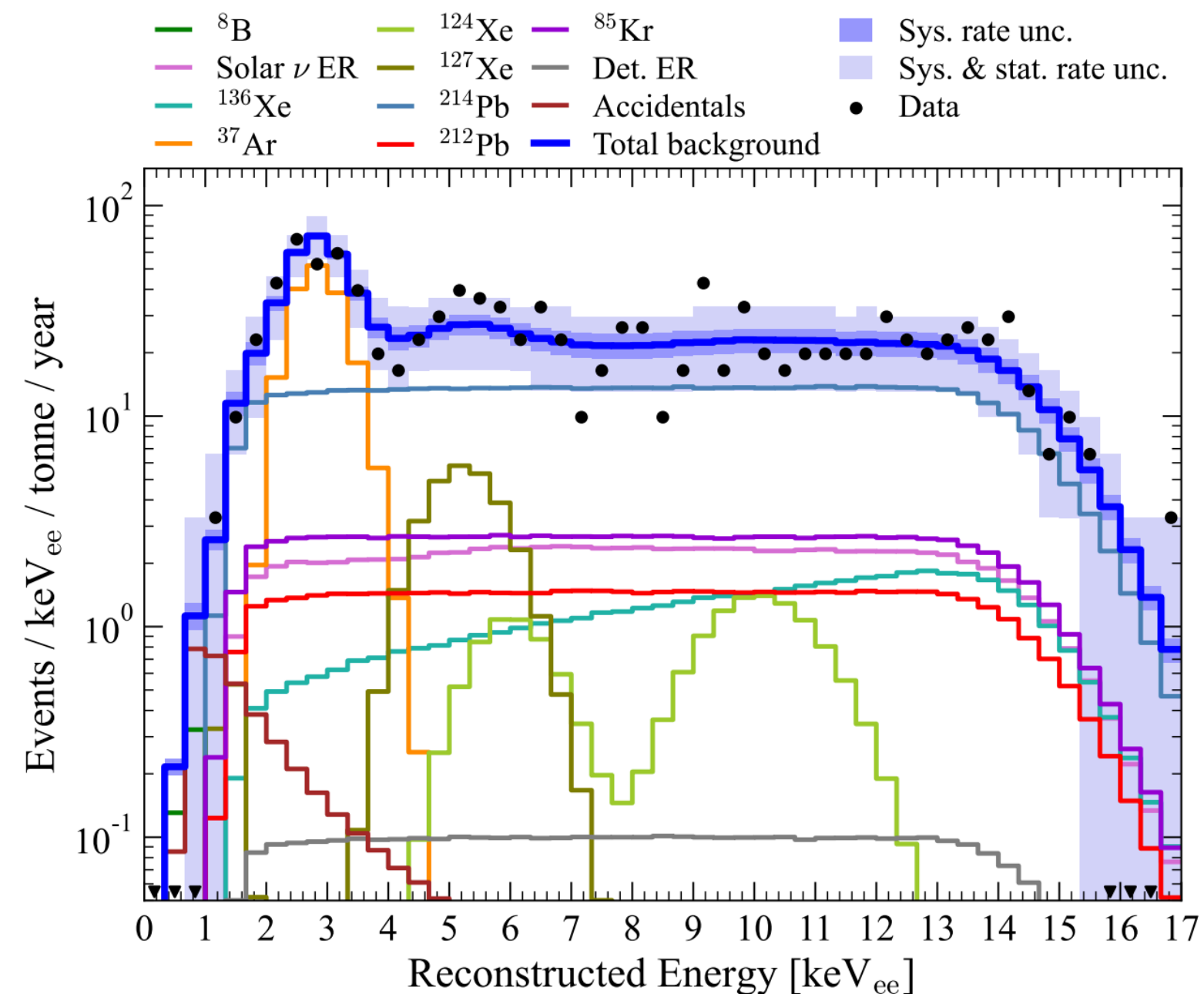




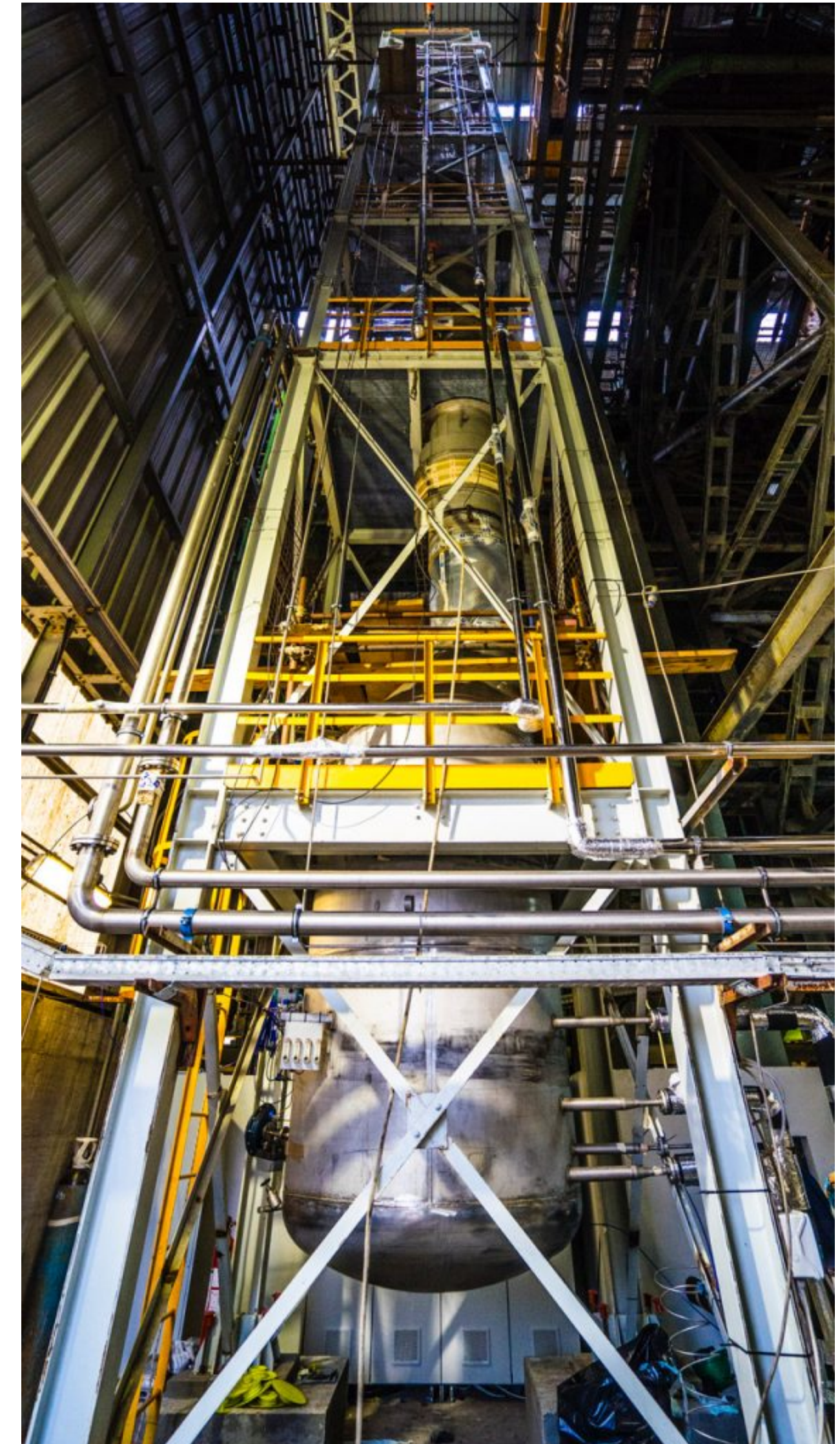
# 3 Purification, cryogenics, infrastructure for noble liquids

## ■ Target radiopurity

- Primarily of concern for the low-background experiments
- Near absence of impurities like  $^{39}\text{Ar}$ ,  $^{222}\text{Rn}$ ,  $^{85}\text{Kr}$  is crucial
- LAr: large-scale cryogenic  $^{40}\text{Ar}$  distillation efforts currently under development (ARIA project)
- LXe: focus on  $^{222}\text{Rn}$ : continuous cryogenic distillation needed



Distillation column, ARIA project



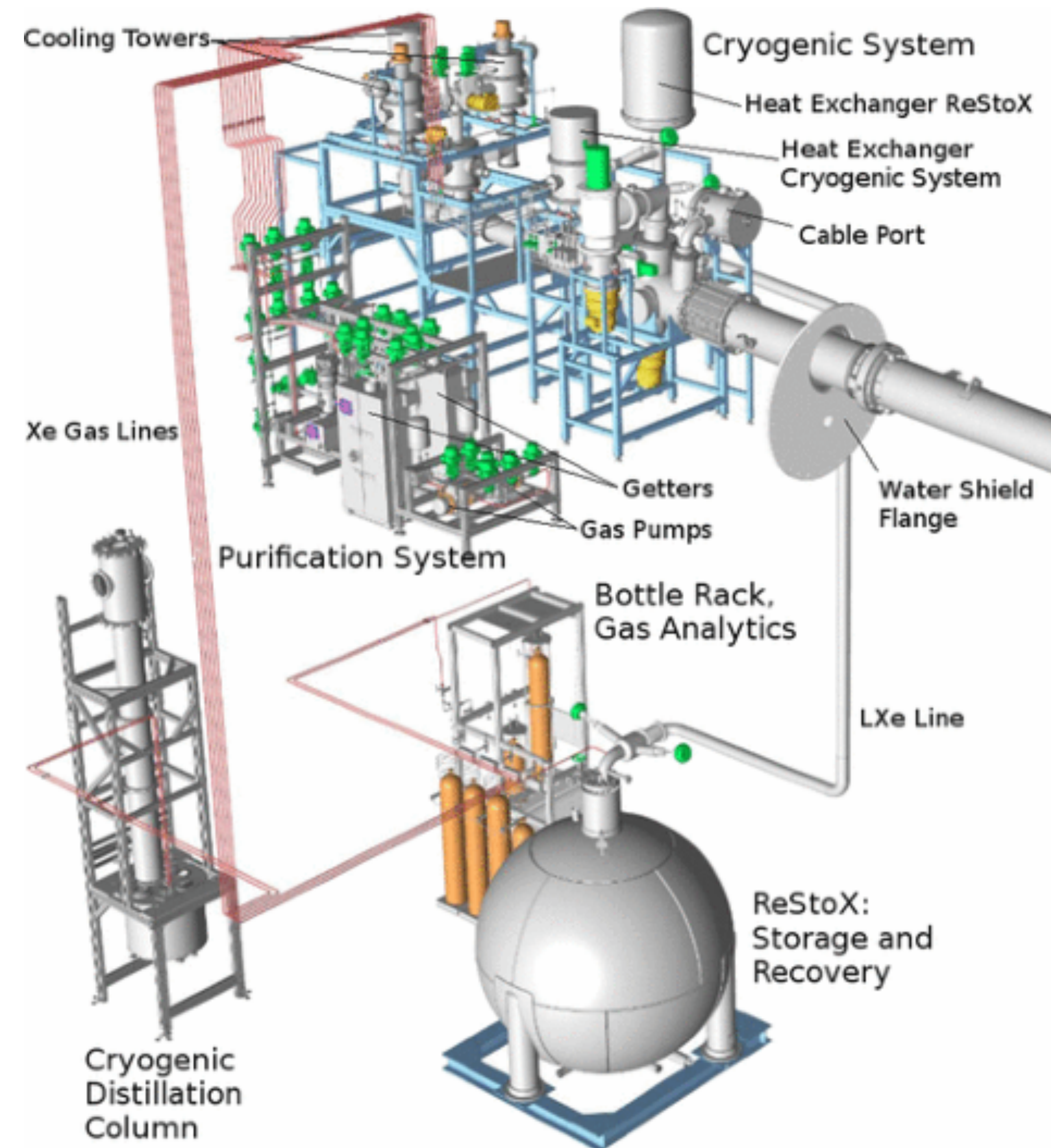




# 3 Purification, cryogenics, infrastructure for noble liquids

## ■ Cryogenics

- increasingly complex cryogenic systems
- need to facilitate circulation of LAr/LXe to maintain cleanliness
- thermodynamic solutions necessitates ultra-clean compressors and heat exchangers
- technologies needs similar to those of particle accelerators, and detector cooling solutions in HEP experiments
  - for different reasons, both communities need clean, non-polluting components in their respective cryogenic and cooling systems
- multi-ton storage components with rapid noble liquid recovery systems in case of emergency



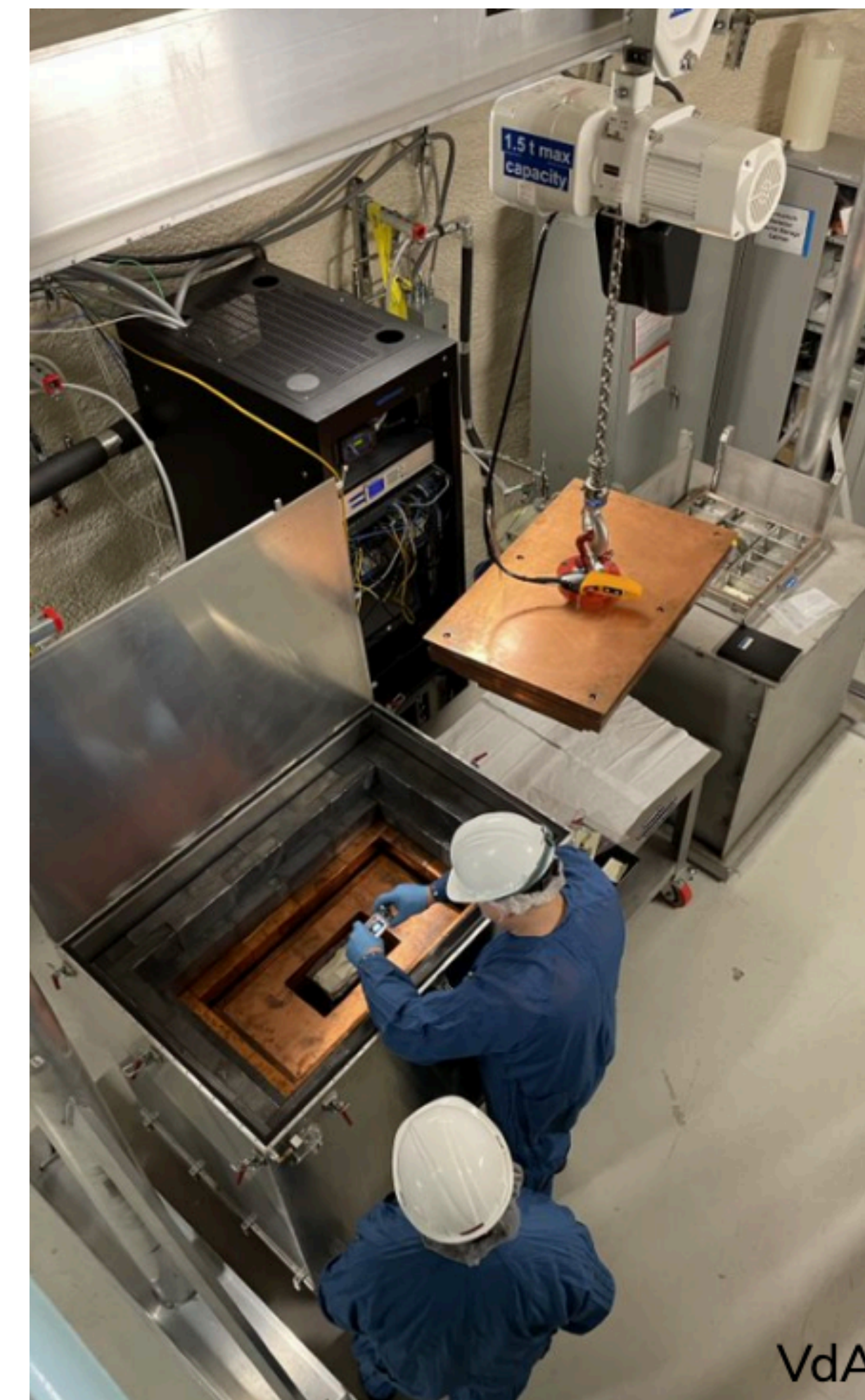
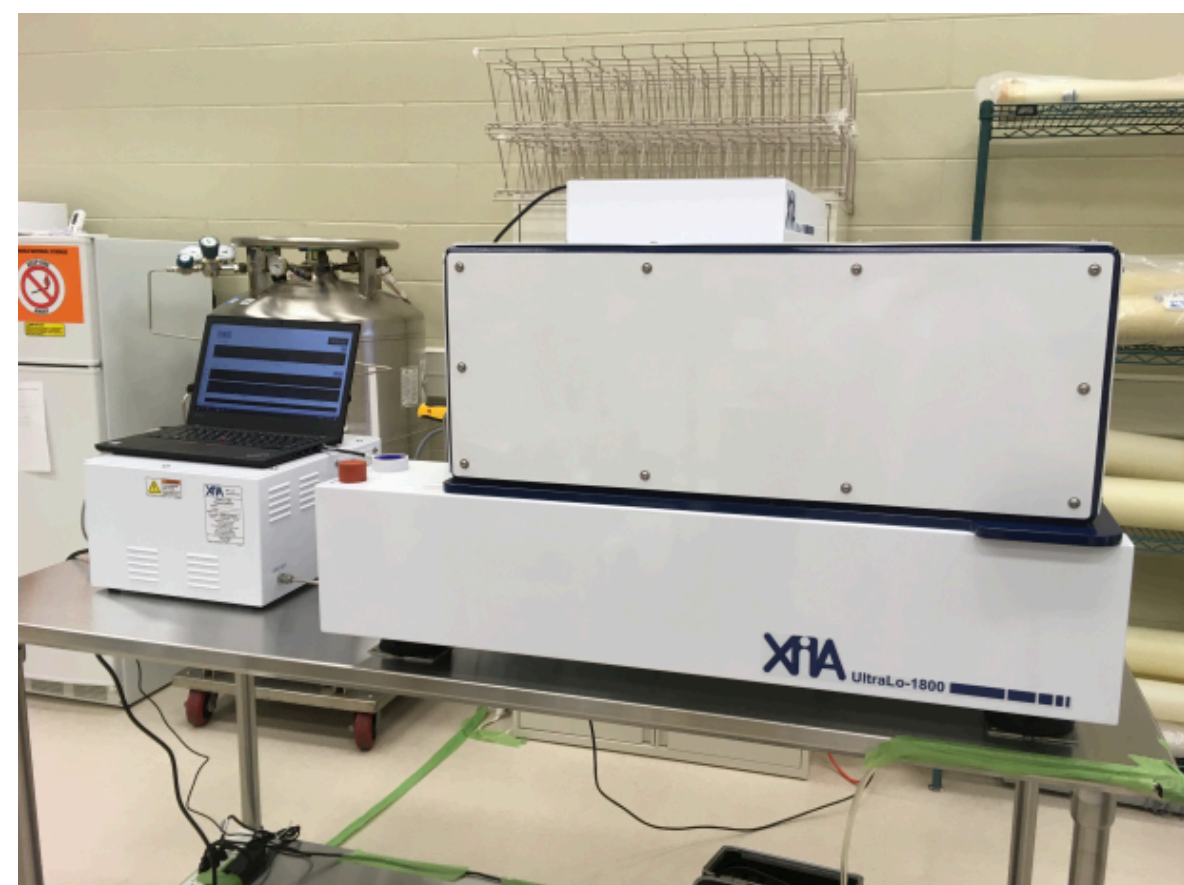
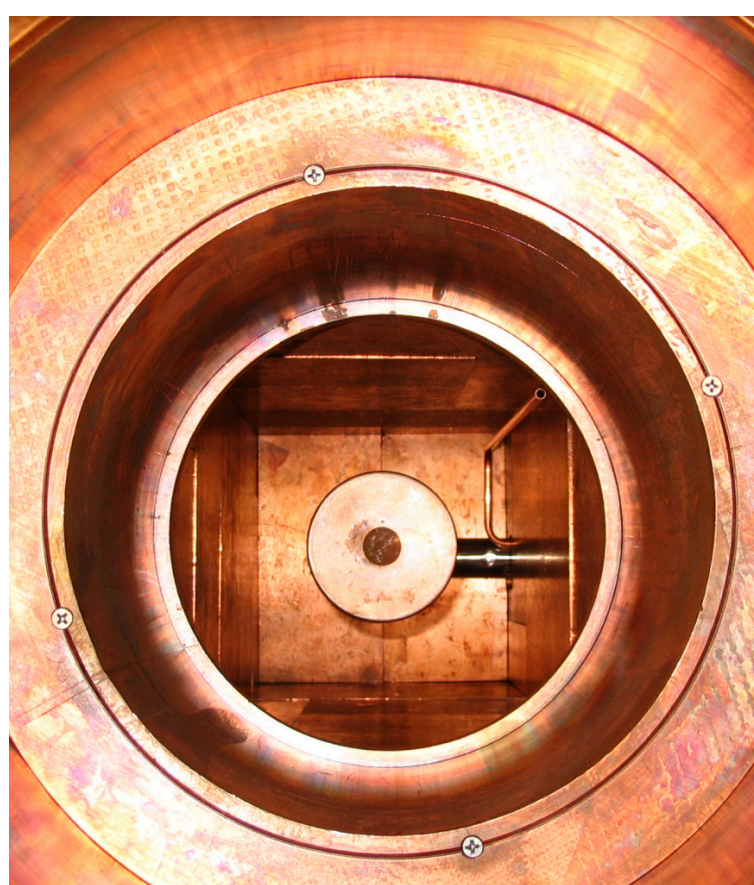




# 3 Purification, cryogenics, infrastructure for noble liquids

## ■ Radioassay

- extensive screening campaigns in UG labs
- highly specialized radon emanation measurements
- R&D for mitigation of emanation backgrounds by investigating material selection procedures and options for surface treatment in systematic way
- ultra-sensitive trace analysis capabilities need to developed for a much larger scale

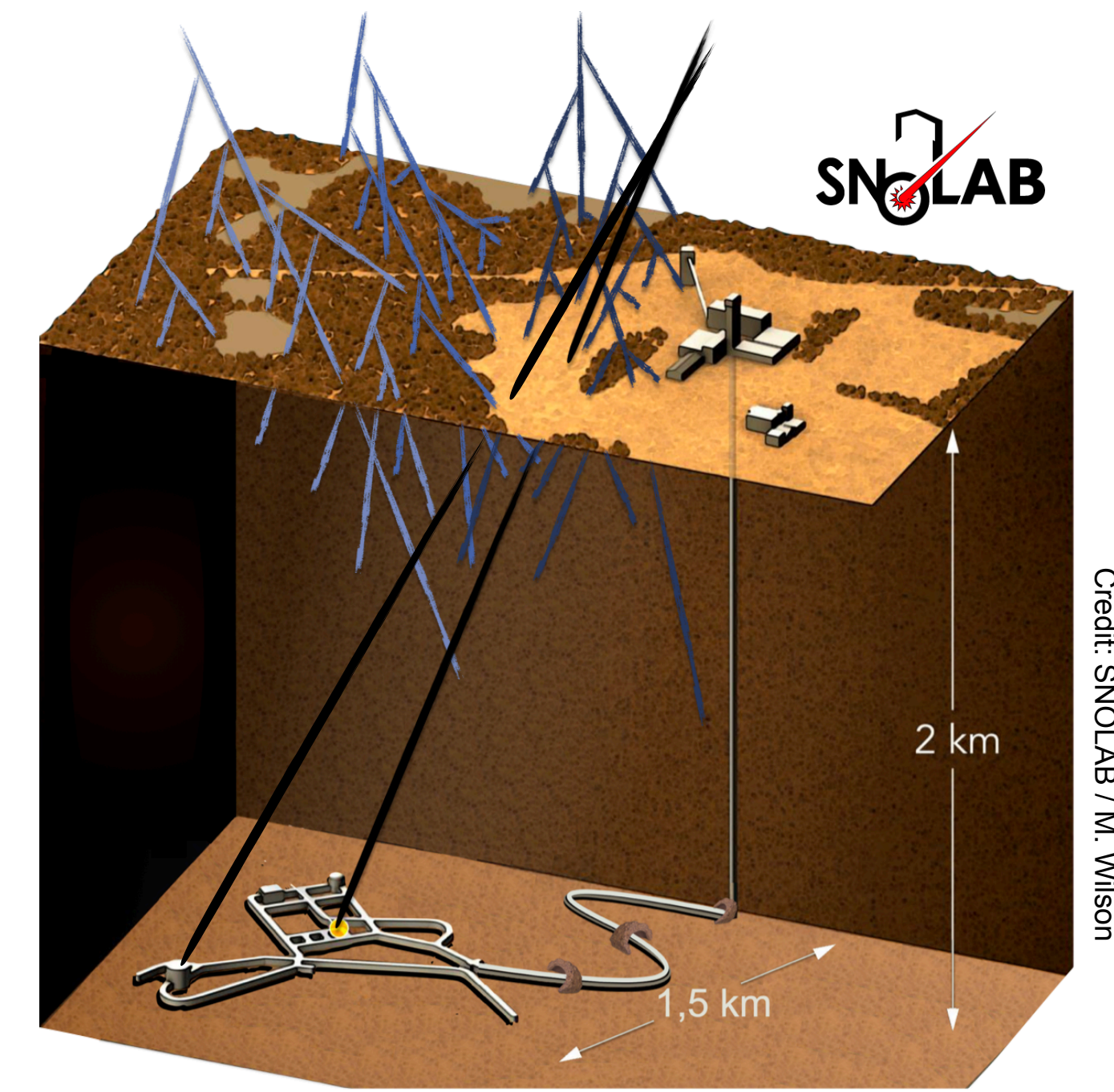
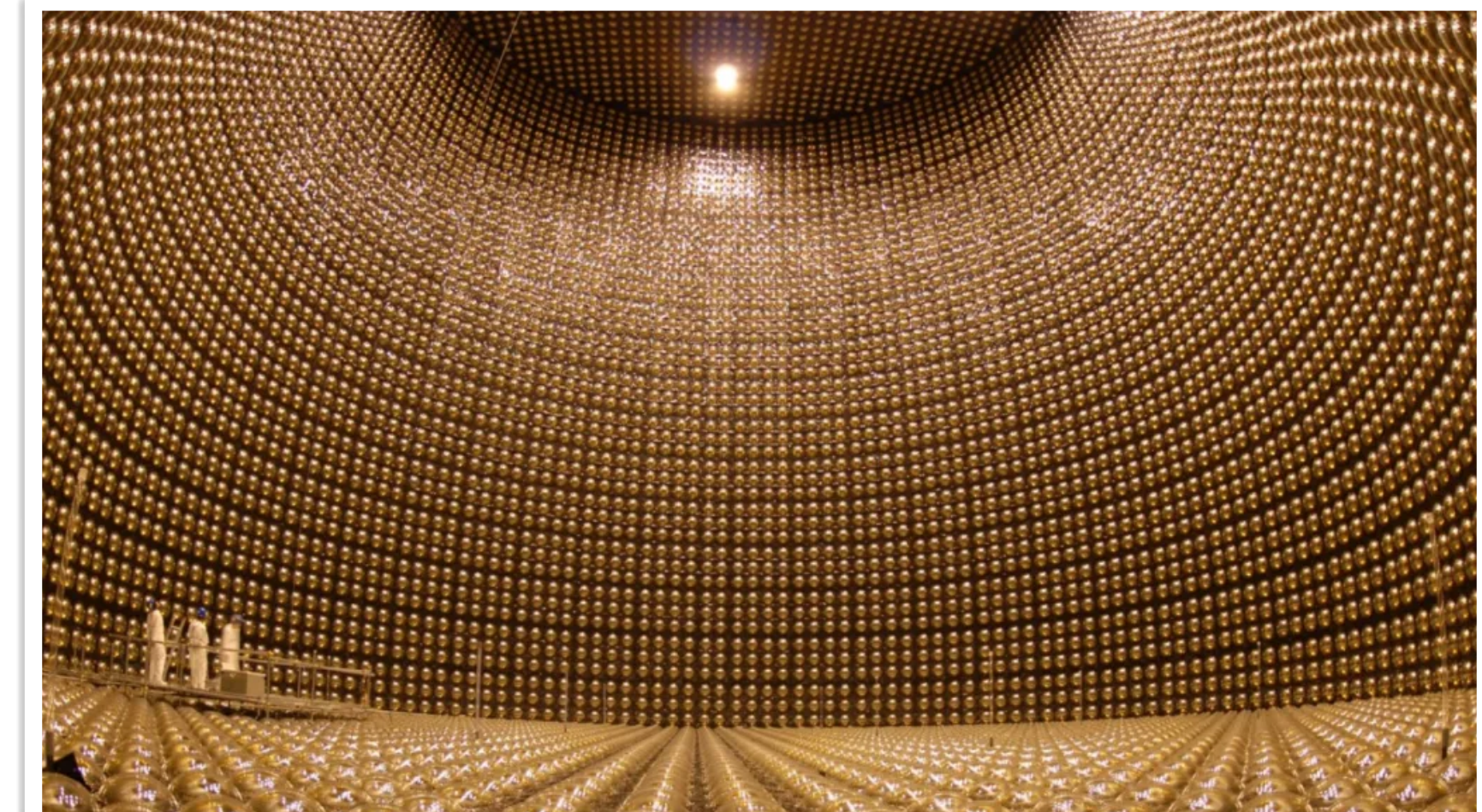




# 4 Light collection in noble liquids and other liquids

## ■ Main challenges:

- photosensor coverage over huge surfaces
- extreme environments (e.g. underground, undersea, and / or at cryogenic temperatures)
- single photon sensitivity
- long term operation and stability ( $\geq 10$  years between detector accesses)
- numbers of readout channels as well as data volume
- very different time scales (both  $\sim$ ns and  $\sim$ ms at same time)
- radiopurity of light collection, sensing, front-end electronics and signal transmission technologies







## 4 Light collection in noble liquids and other liquids

- Huge photosensitive areas
  - reducing dark current rate (DCR)
  - reducing channel count through summing SiPM signals in arrays
  - increasing S/N through front-end electronics design (mitigate large output capacitance of arrays)
  - improving timing resolution

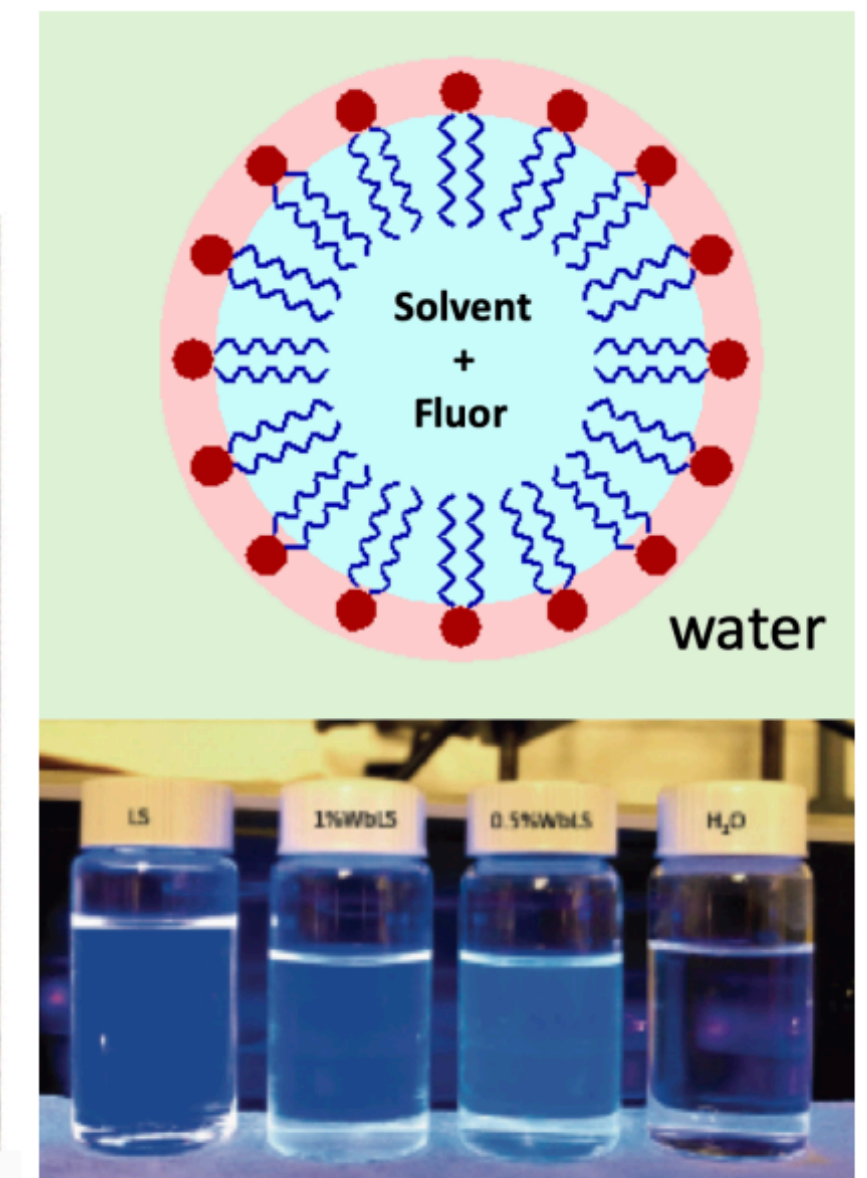
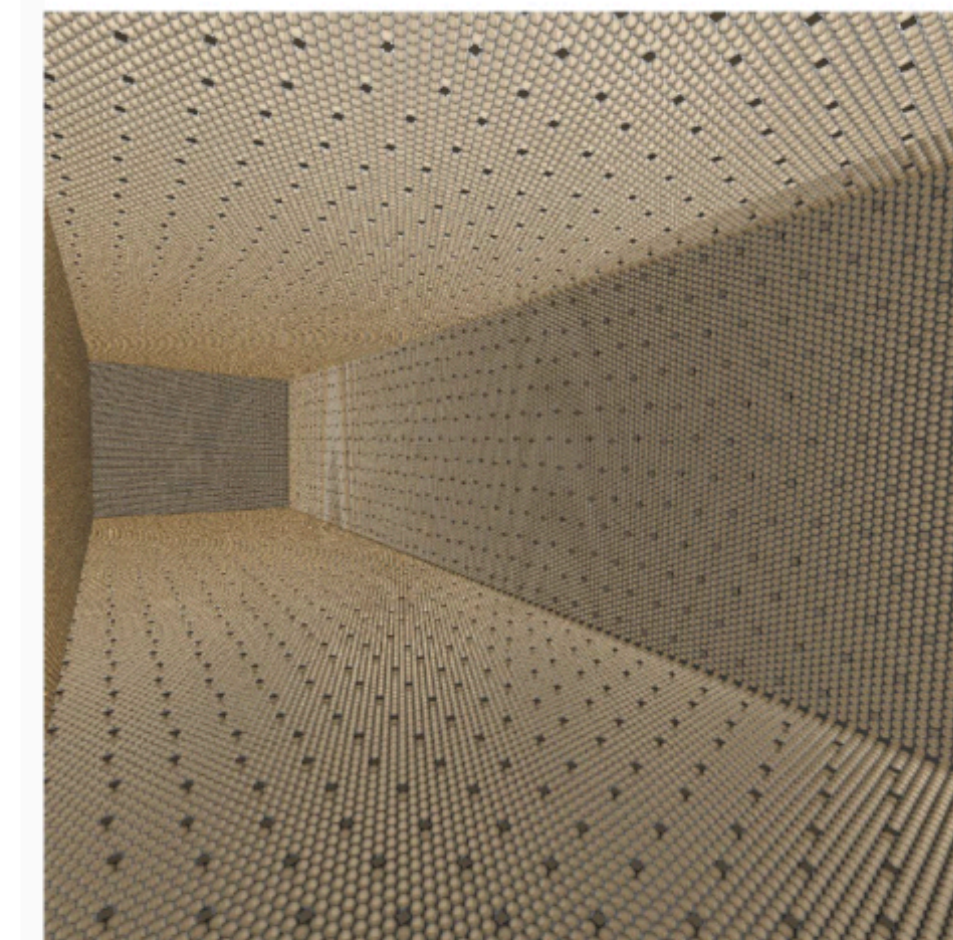
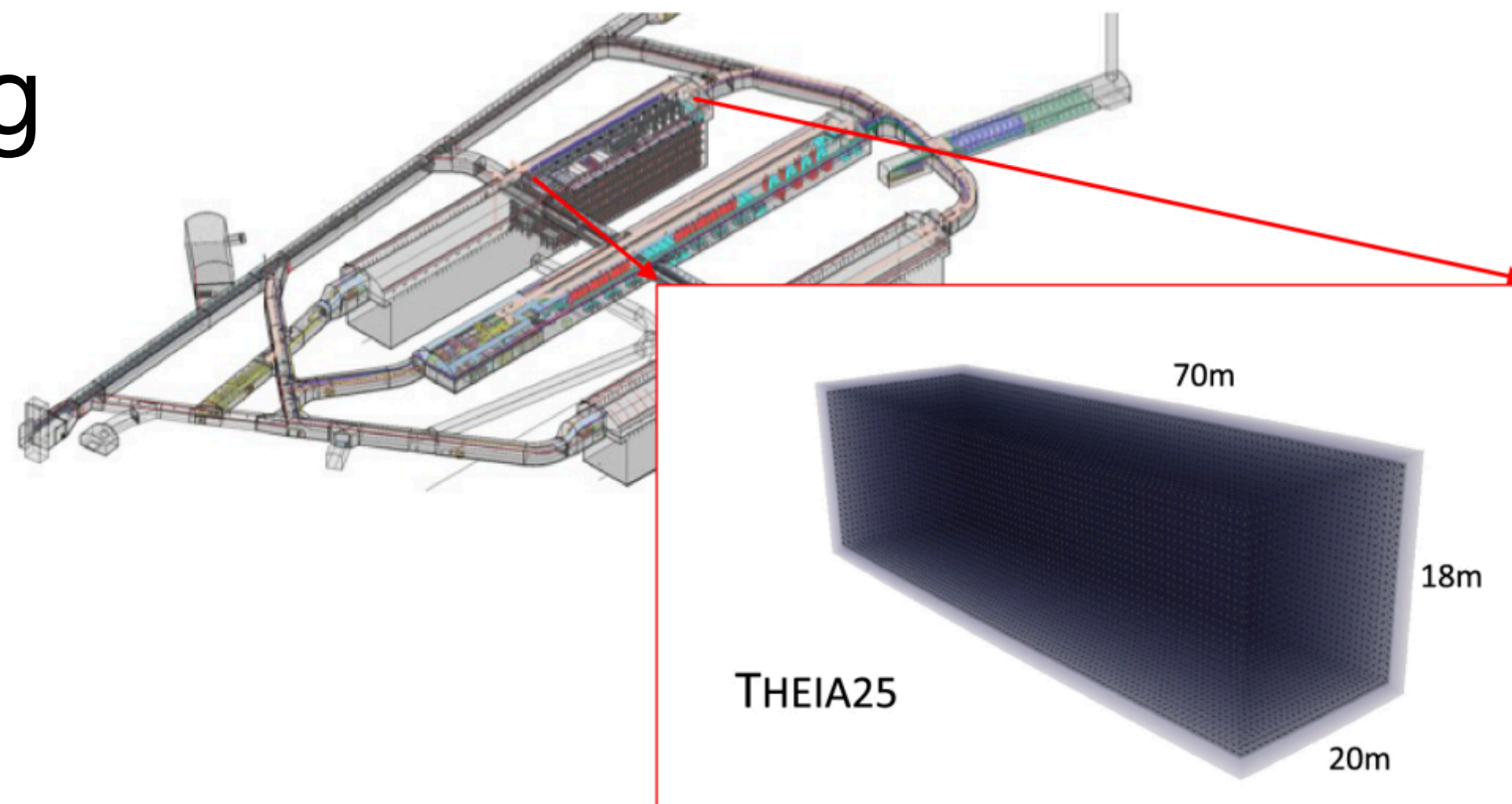
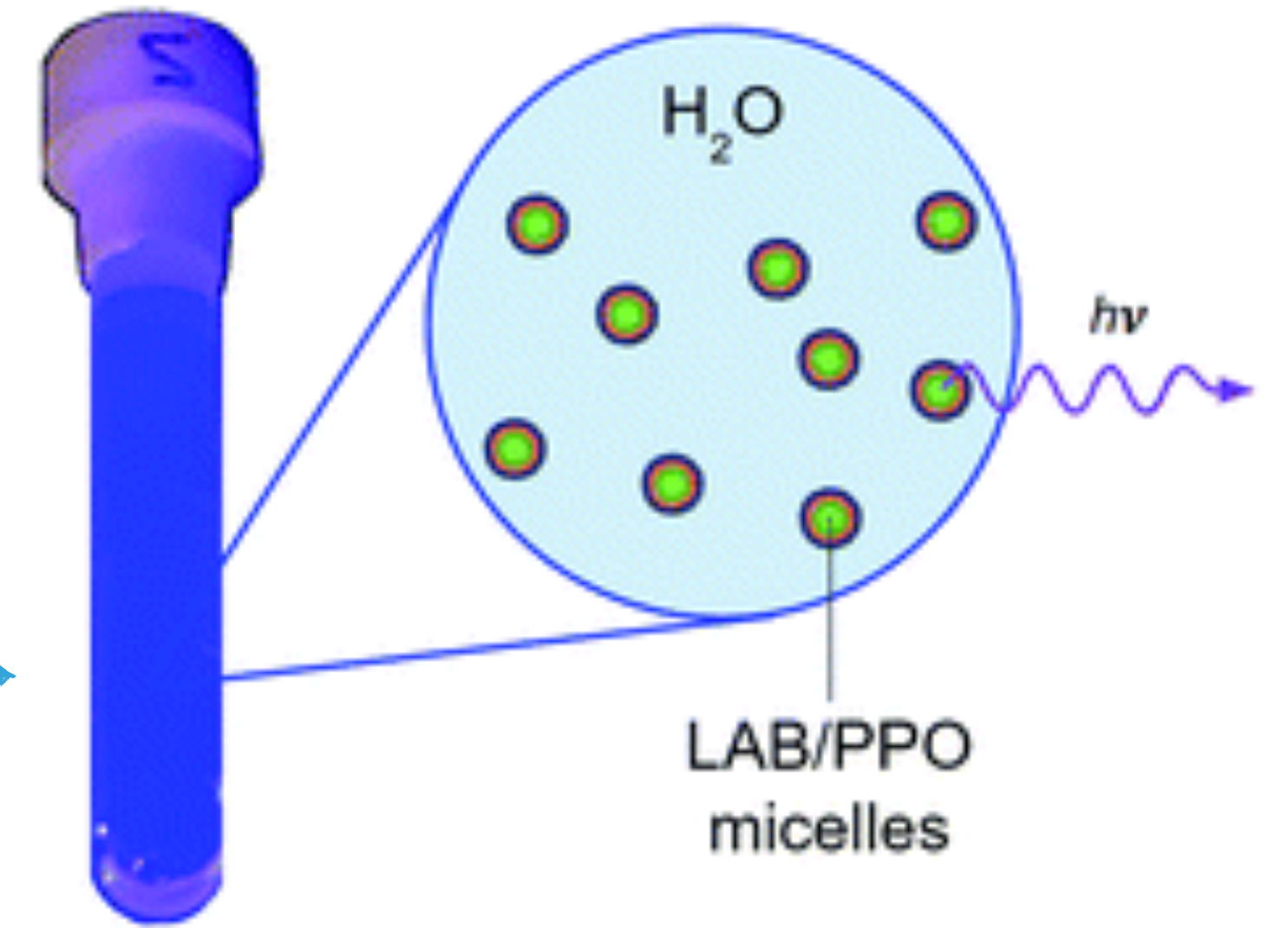
Experiment	Type	Photon detector	Area (m <sup>2</sup> )
nEXO	LXe	SiPMs (FBK [Ch2-18], Hamamatsu [Ch2-19]), digital 3D-SiPM	5
DARWIN	LXe	PMTs, SiPMs or Hybrids (SIGHT, ABALONE)	8
TAO	LSci	FBK SiPMs	10
DarkSide-20k	LAr	SiPMs (FBK NUV-HD triple-dopant)	30
ARGO	LAr	SiPM is baseline option	200
DUNE	LAr	Light guide or trap + SiPM	10-1000



# 5 Liquid scintillator and water detectors

- Radiopurity, radiopurity, radiopurity...
- Liquid composition
- Low temperatures
- Opacity
- Hybrid detectors for event reconstruction
- Liquid doping

Credit: THEIA collaboration

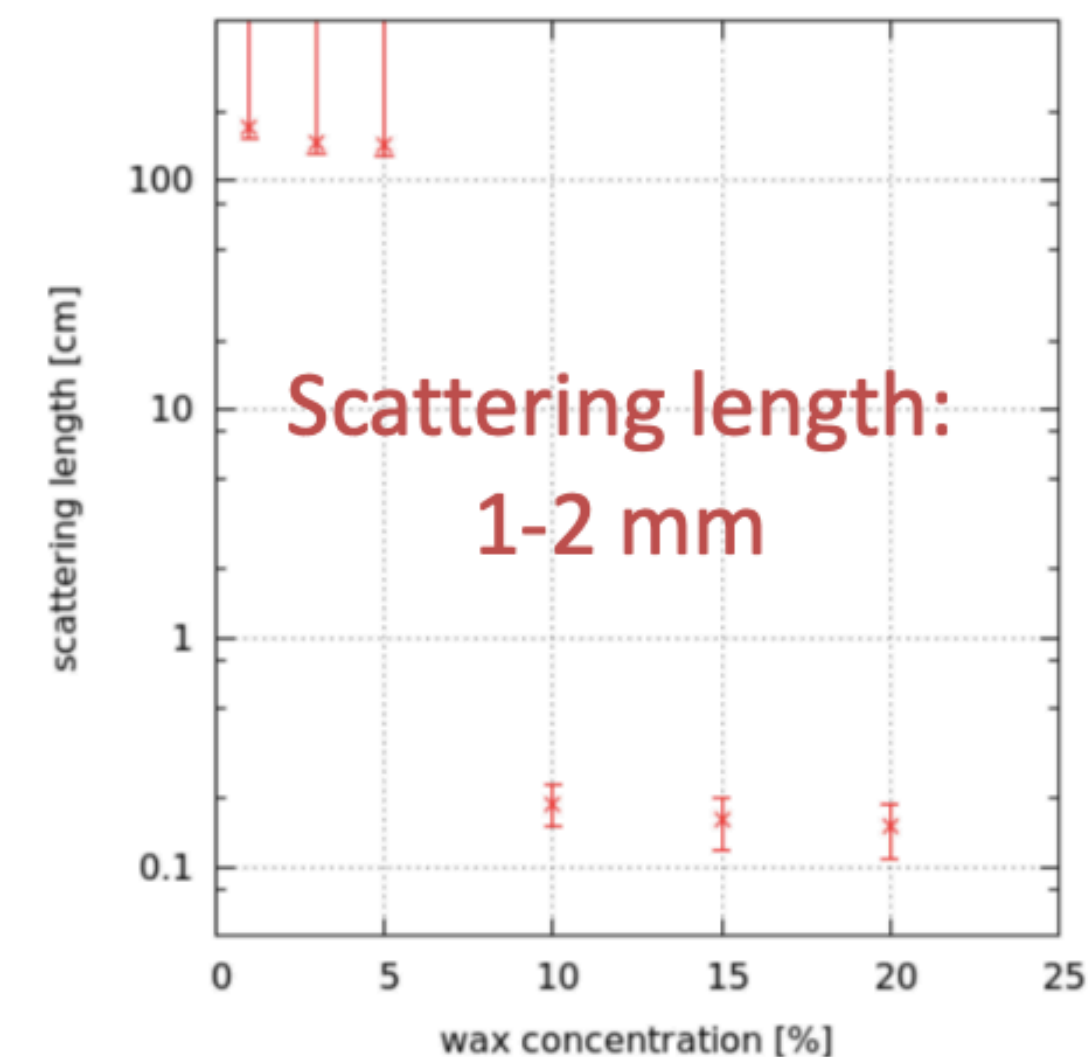
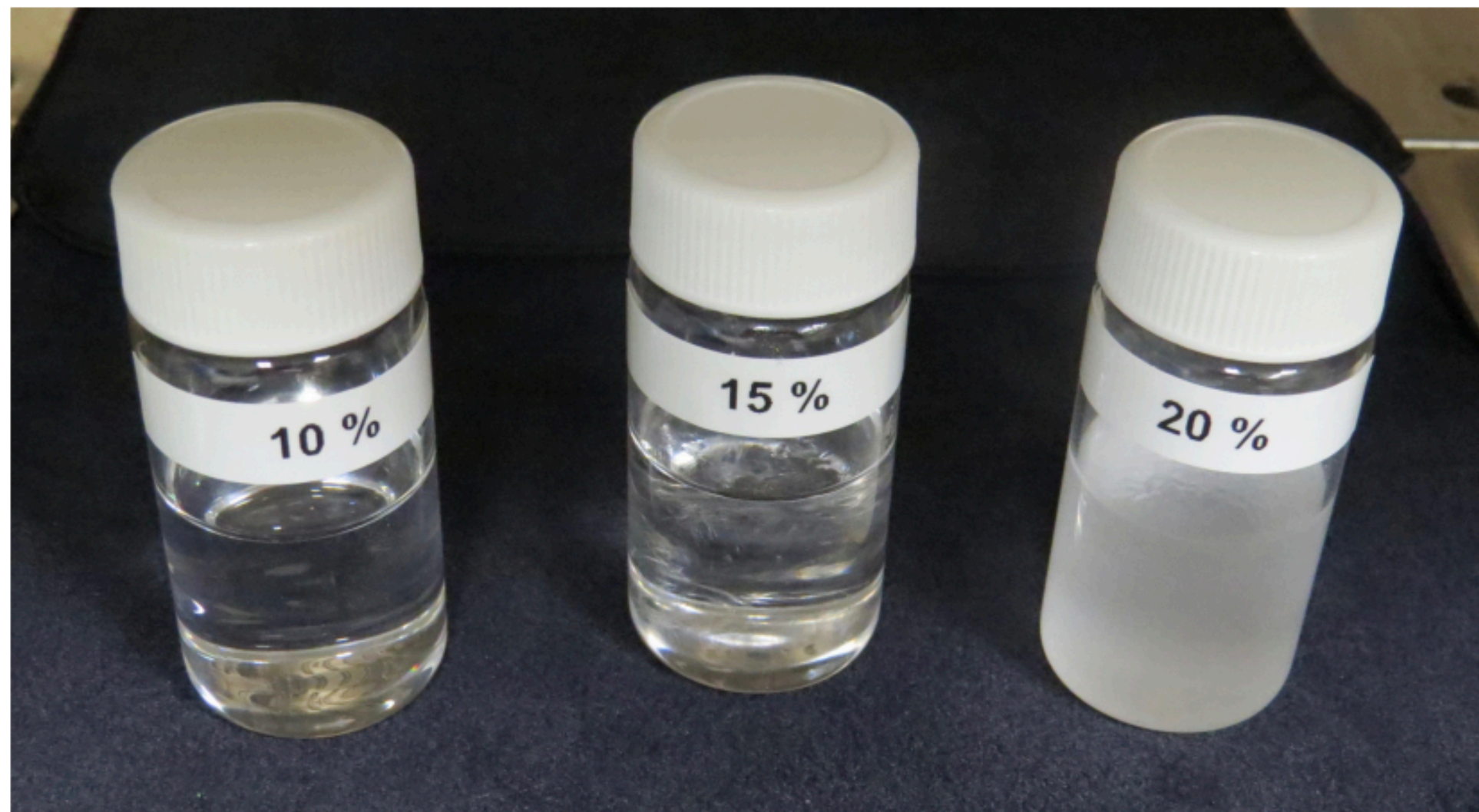




# 5 Liquid scintillator and water detectors

## Opaque scintillators

- Light collection at origin using fibers + fast readout
- Highly improved vertex resolution by light confinement
- Wax based scintillator (NoWaSH)
- High metal loading possible (up to 10%), suspensions
- Potential applications: reactor, geo,  $0\nu\beta\beta$ , solar,...







# Required R&D readiness (high-level, non-exhaustive overview)



Earliest feasible start date, such that R&D is not the limiting factor.