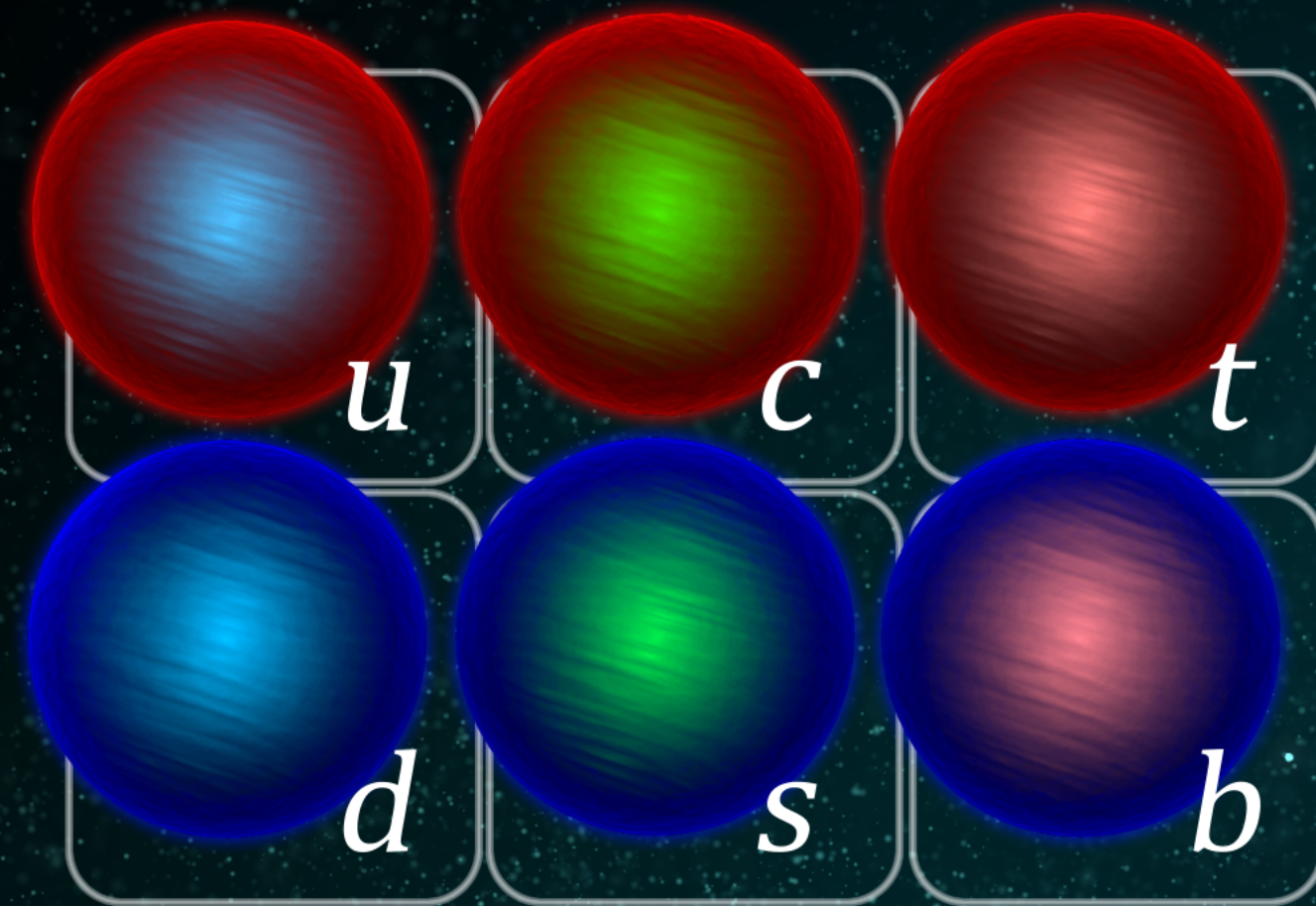




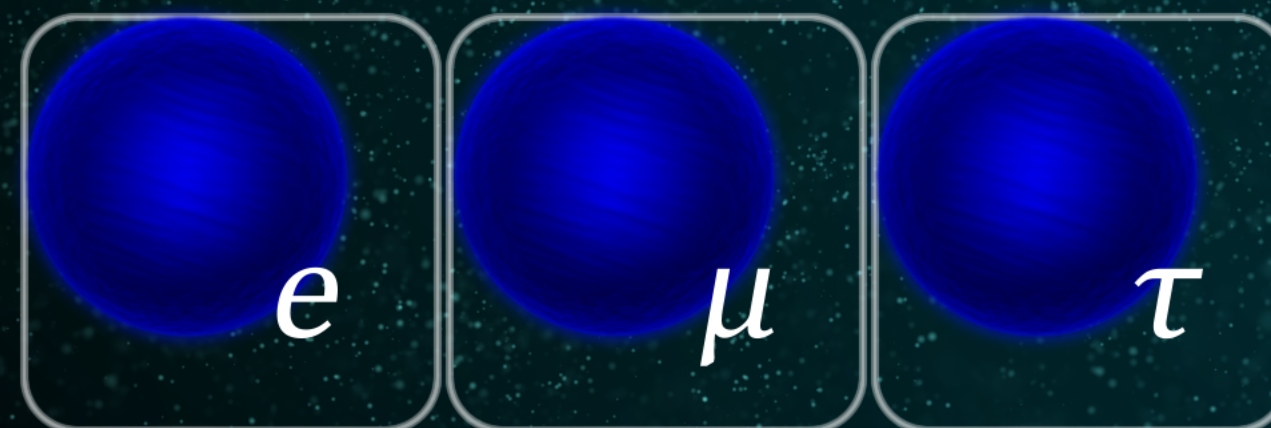
# Model-Independent Measurements

Monica Dunford - KIP Heidelberg

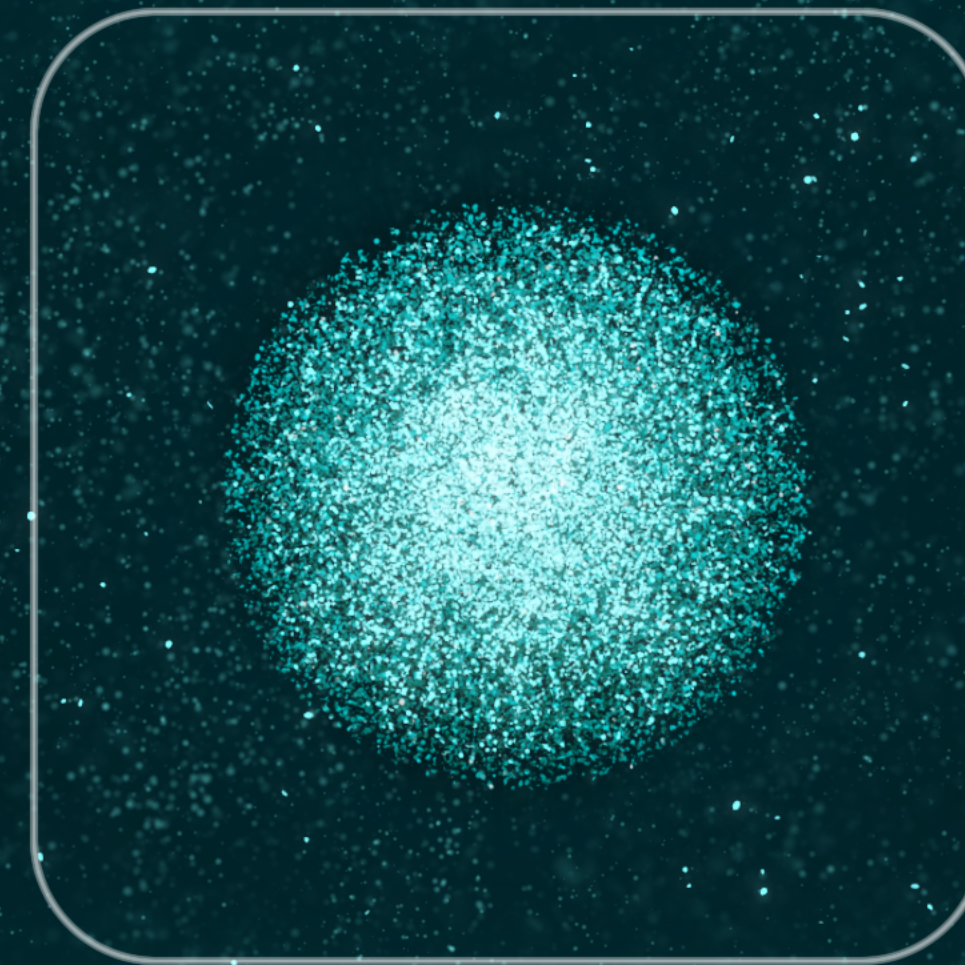
# The genesis to any measurement is the question itself



Quarks



Leptons

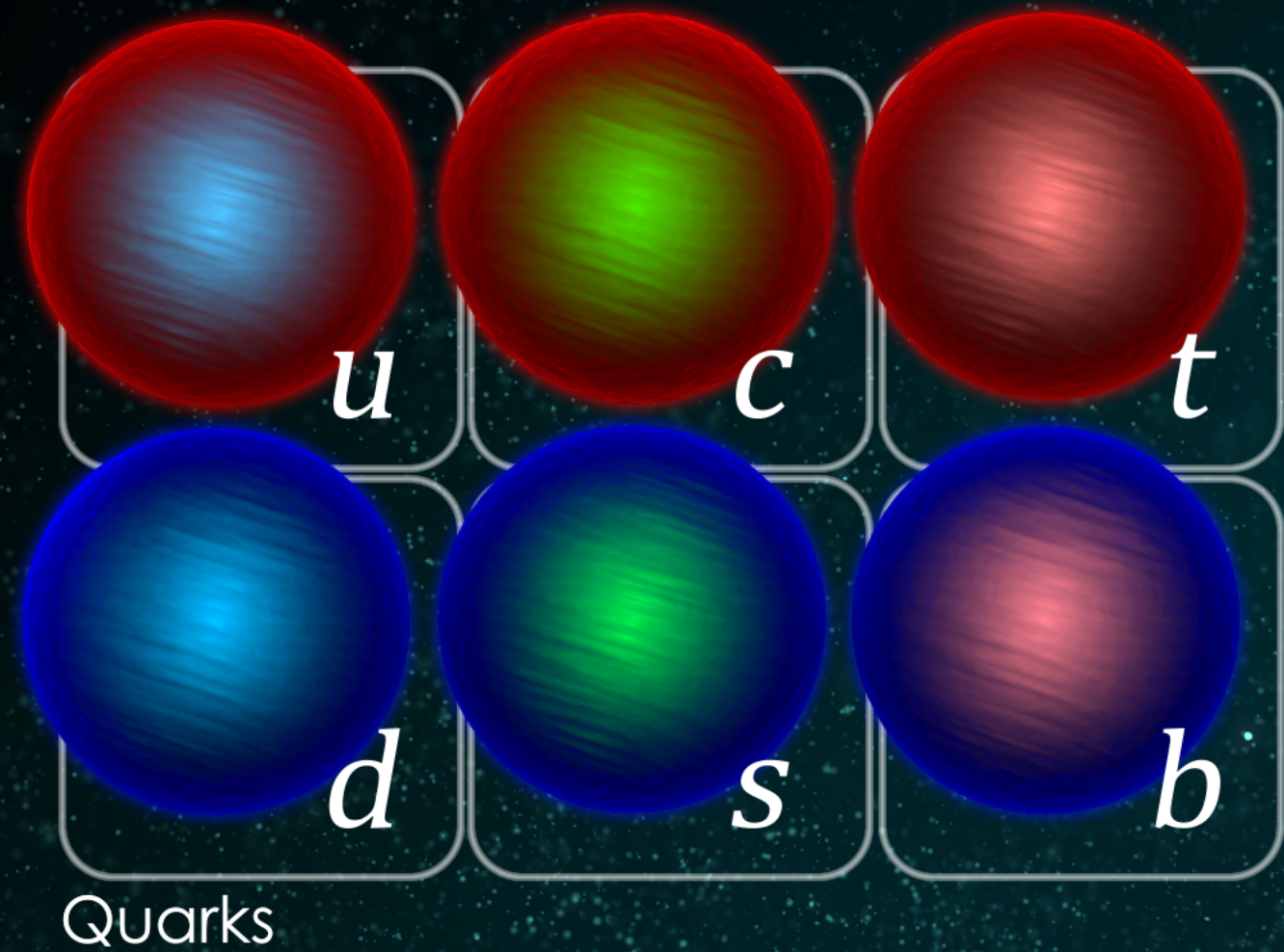


Higgs boson



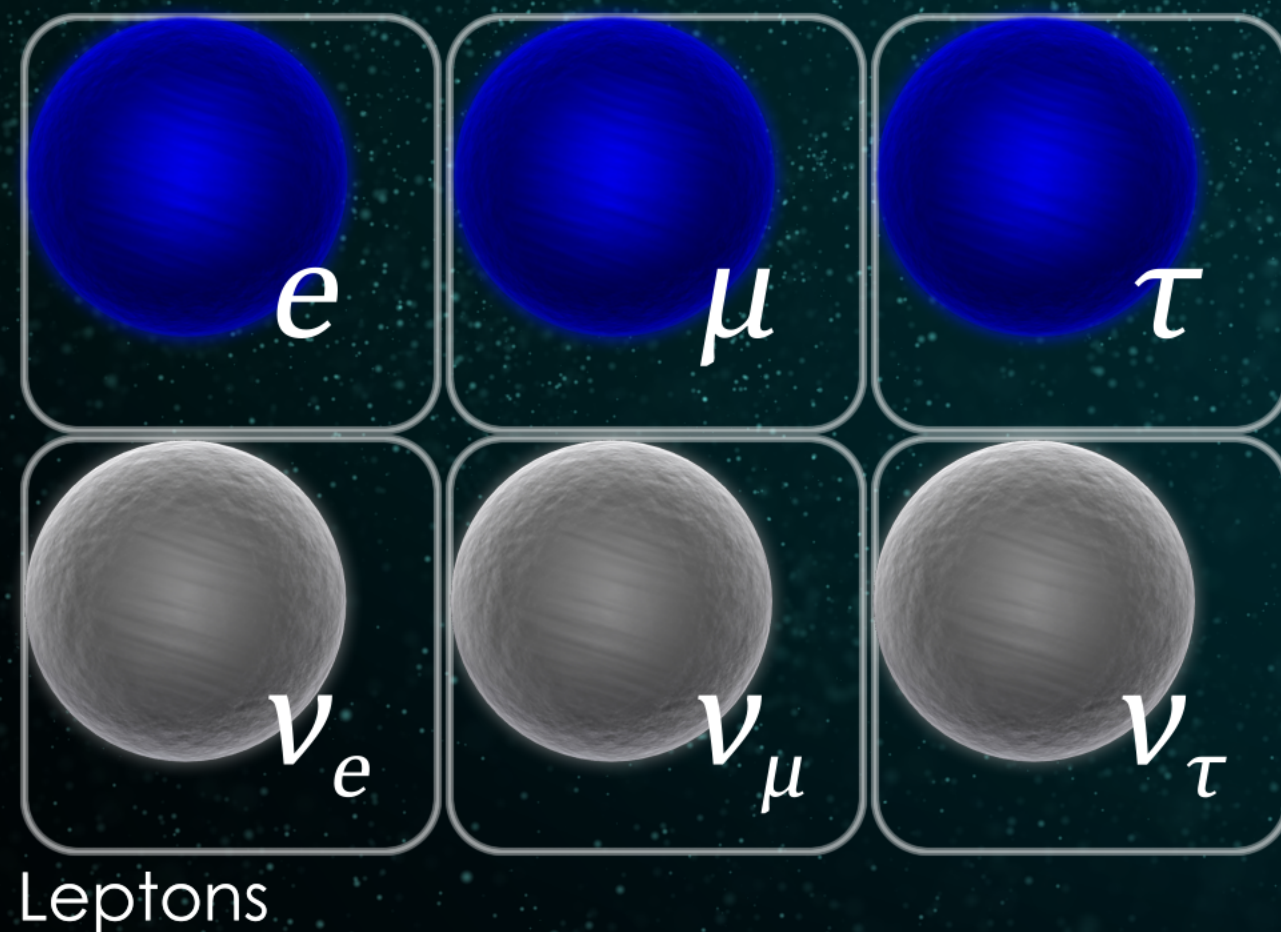
Forces

Mass hierarchy

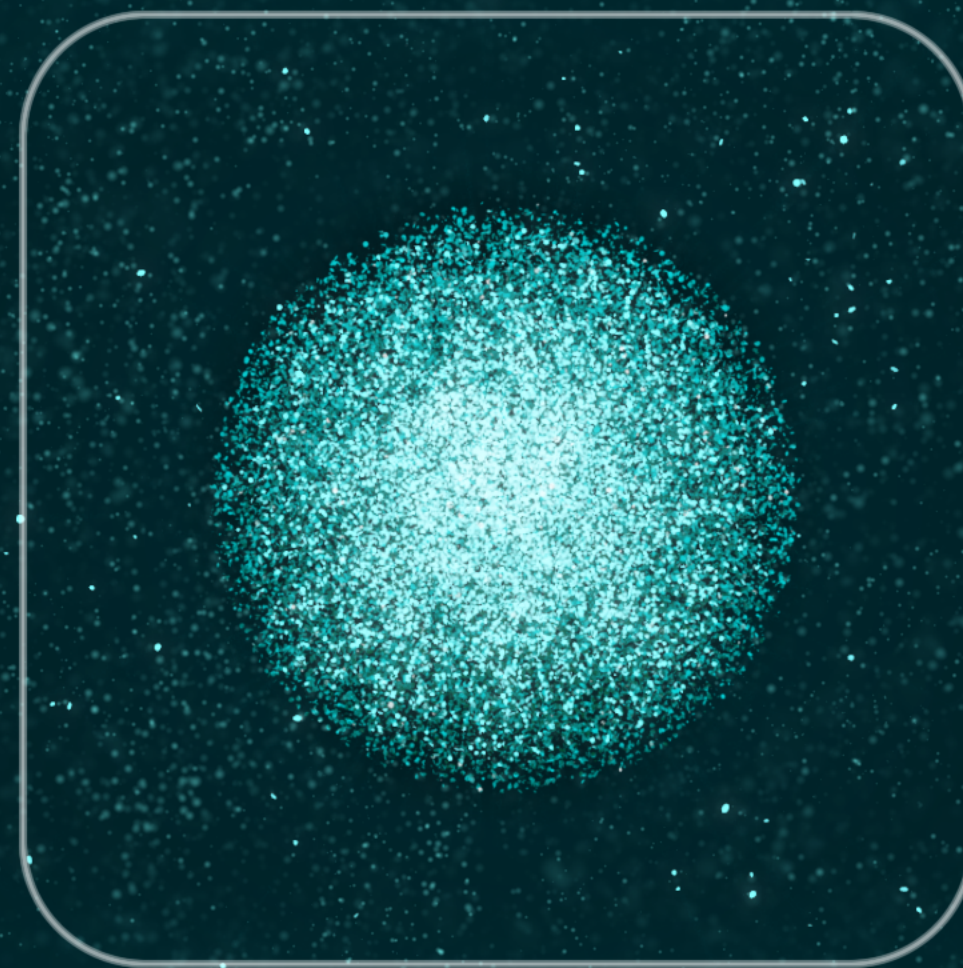


CP Violation

Additional particles?



# The genesis to any measurement is the question itself



Higgs boson

Composite Higgs?

Who breaks electroweak symmetry?

Why is  $\theta_{13}$  so small?

Dark energy?

Dark matter?



Forces

Gravity?

Unified forces?

Majorana vs. Dirac

Why is neutrino mixing so large?

Theories beyond the Standard Model?

What is dark matter?

Is it a particle?

Is it heavy and fast or light and slow?

Dark matter is massive. How does it get its mass?

Is dark matter just modified gravity?

Or something else like primordial blackholes?

How is it distributed?

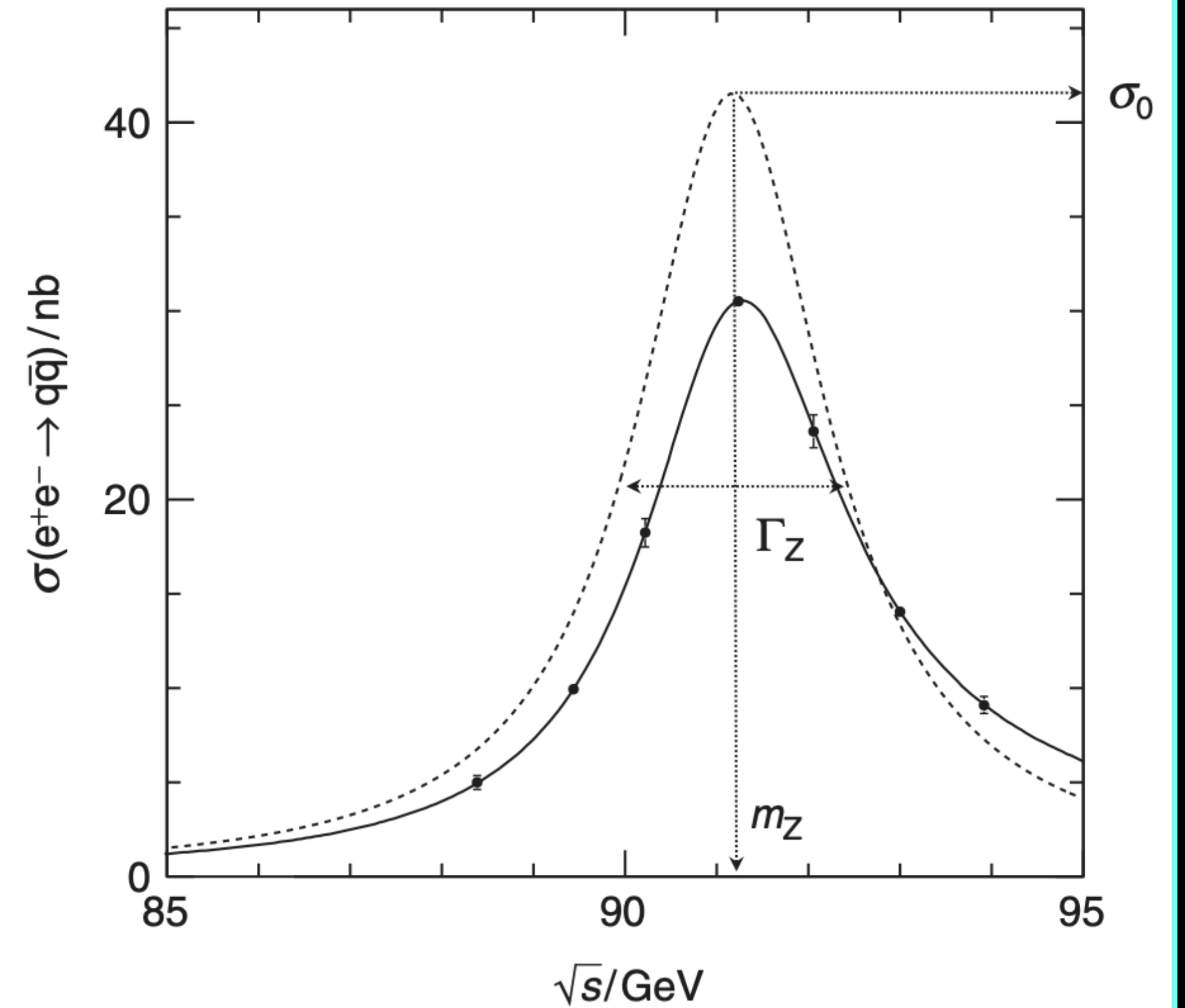
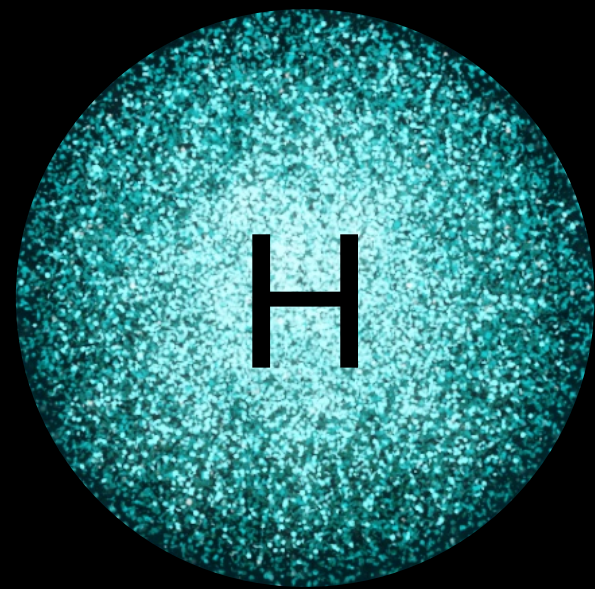
Is the Earth in a dark matter deficit?

What role does dark matter play in galaxy formation?

Each of these questions are important and the measurements to answer them are completely different

Building those measurements, both experimentally and theoretically, is where all the fun is

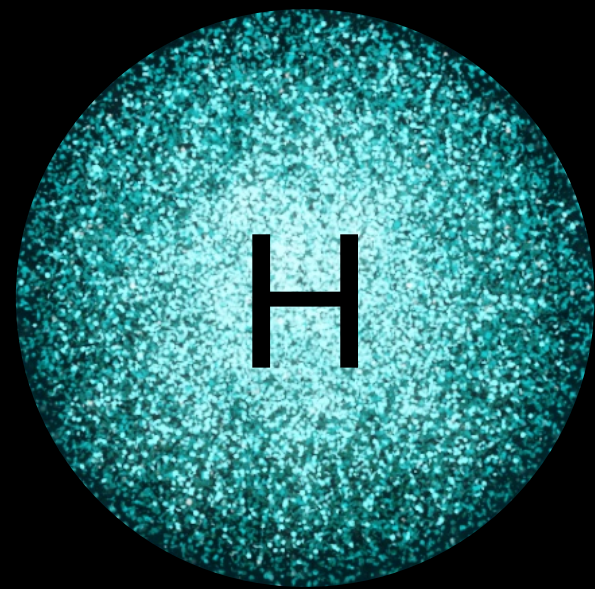
Dark matter has mass. How?



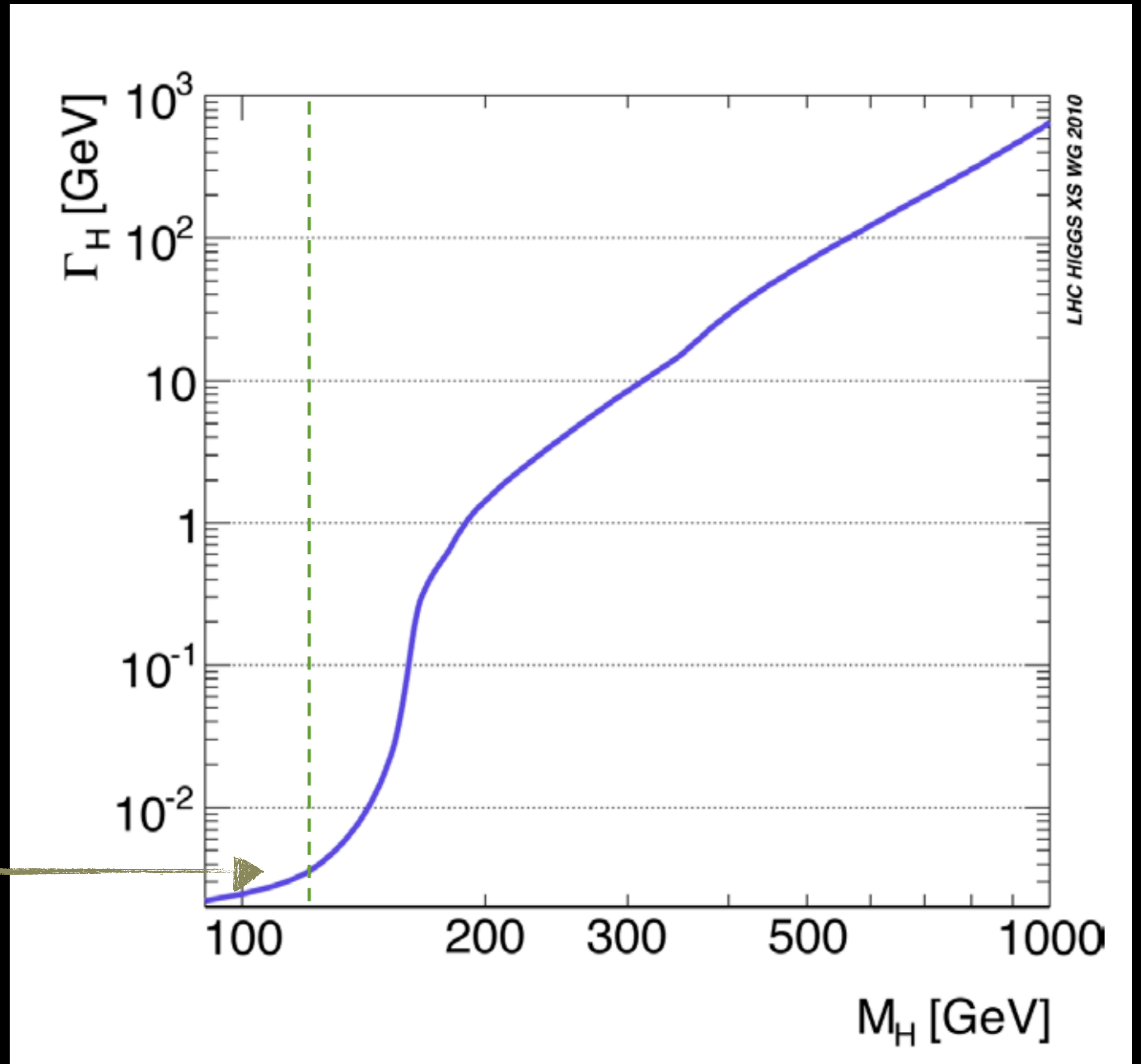
$$\sigma^0(e^+e^- \rightarrow Z \rightarrow f\bar{f}) = \frac{12\pi}{m_Z^2} \frac{\Gamma_{ee}\Gamma_{ff}}{\Gamma_Z^2}$$

$$\Gamma_Z = 3\Gamma_{\ell\ell} + \Gamma_{\text{hadrons}} + N_\nu\Gamma_{\nu\nu}$$

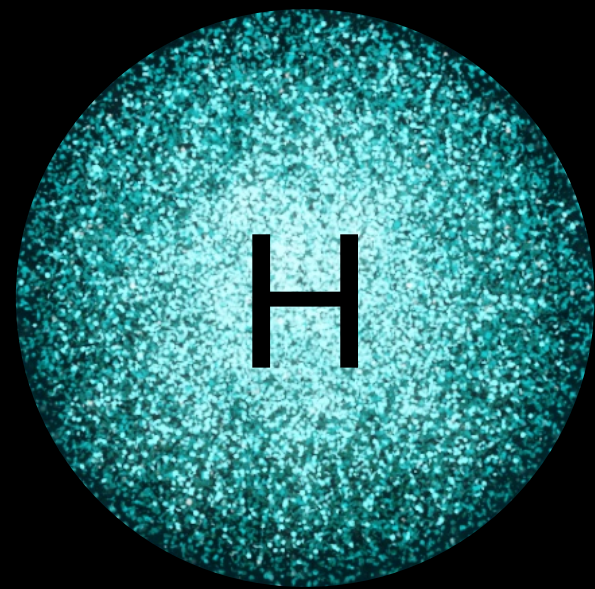
# Dark matter has mass. How?



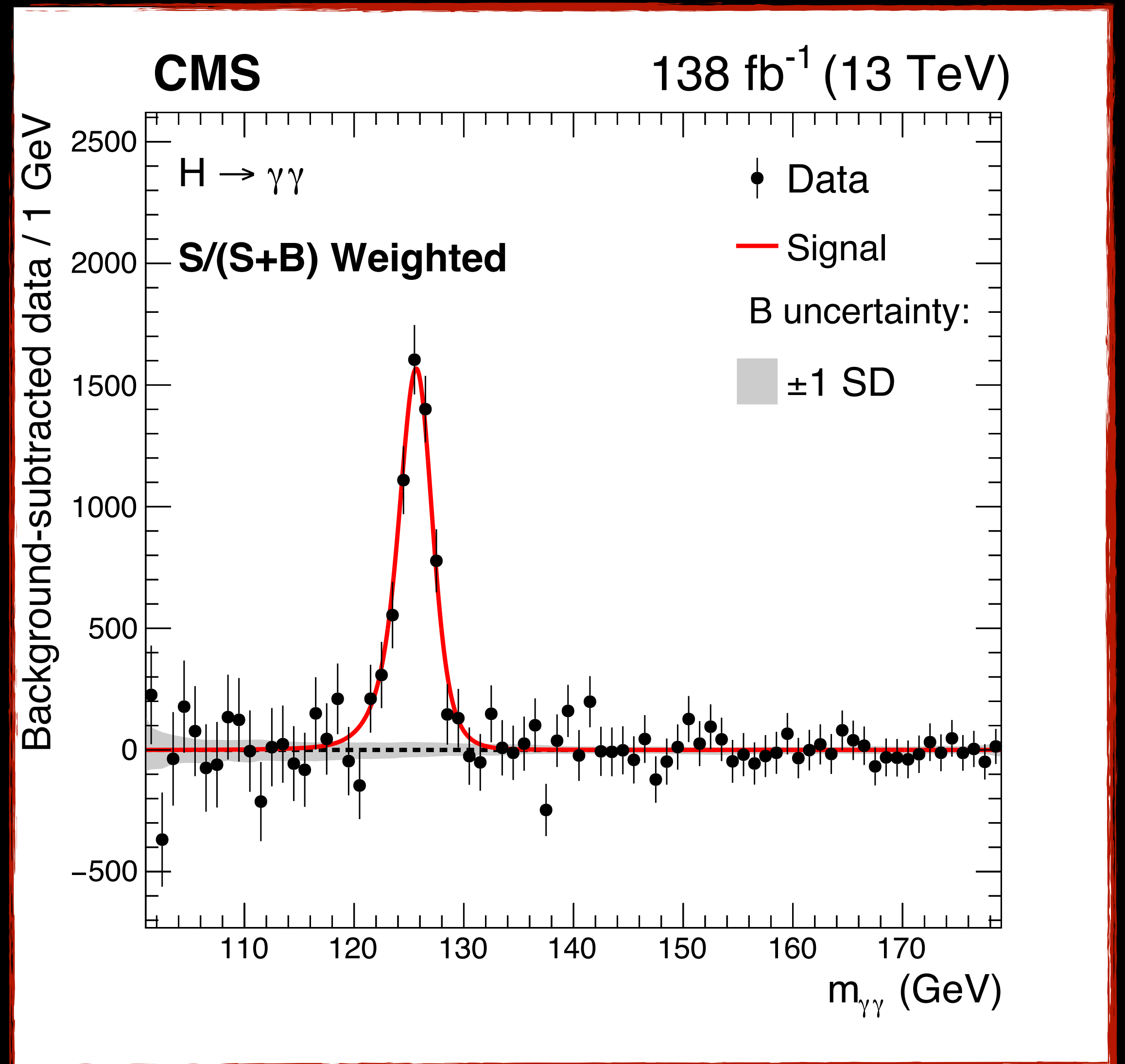
- Prediction 4.07 MeV



# Dark matter has mass. How?

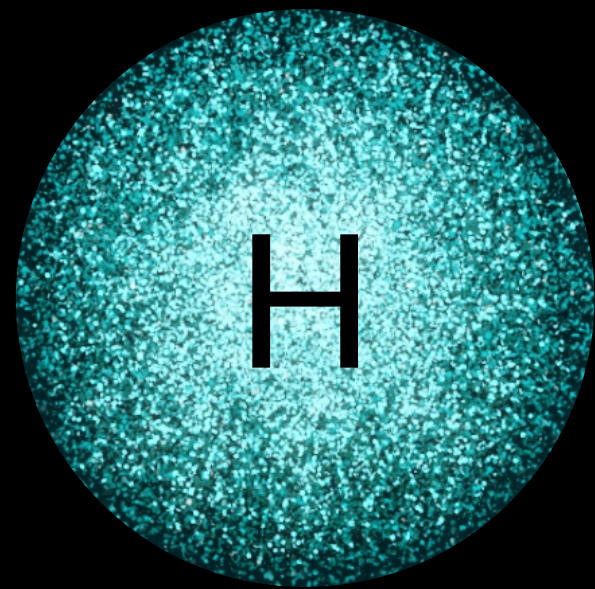


- Prediction 4.07 MeV

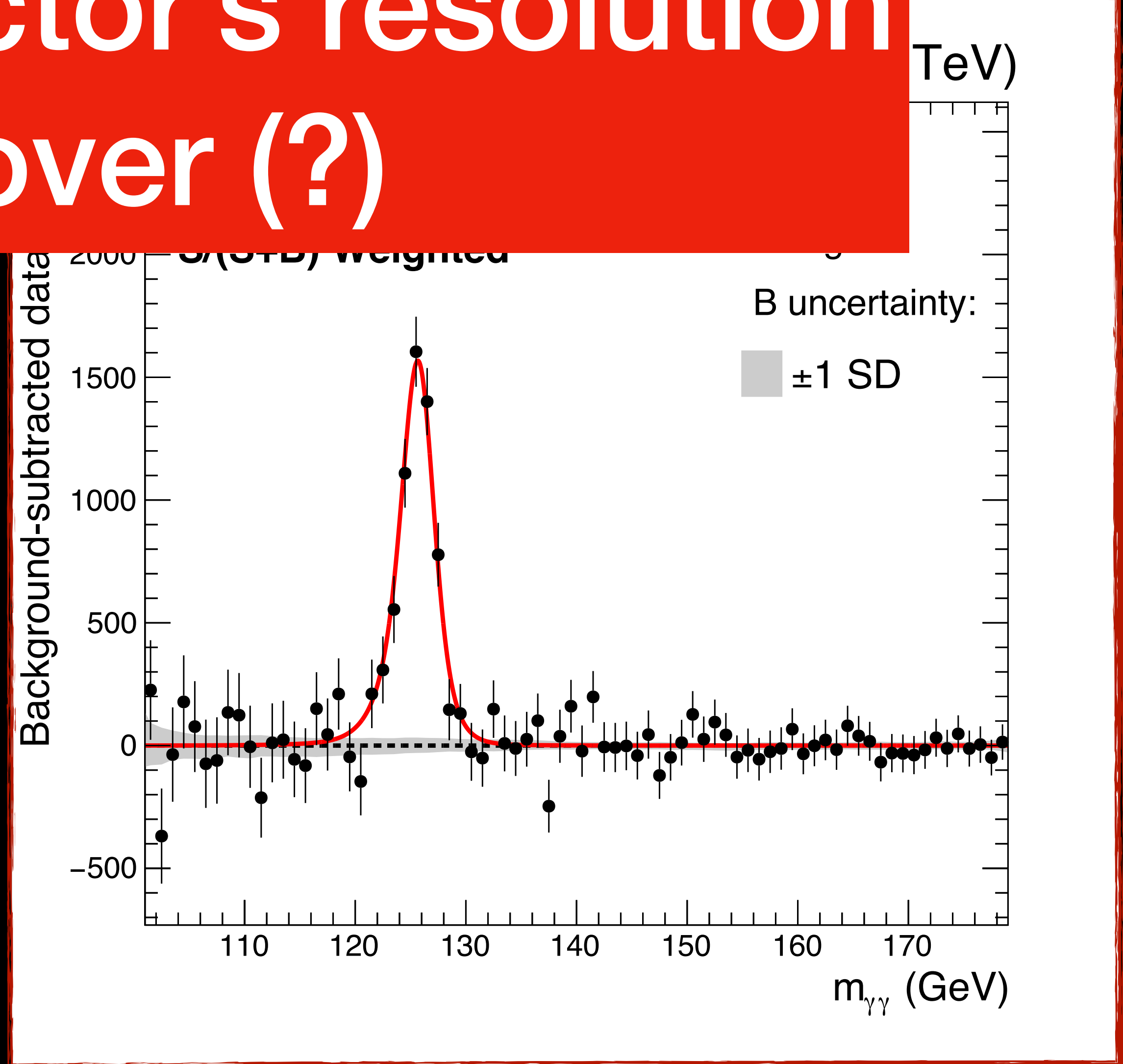


# This is the detector's resolution Game over (?)

mass. How?



- Prediction 4.07 MeV

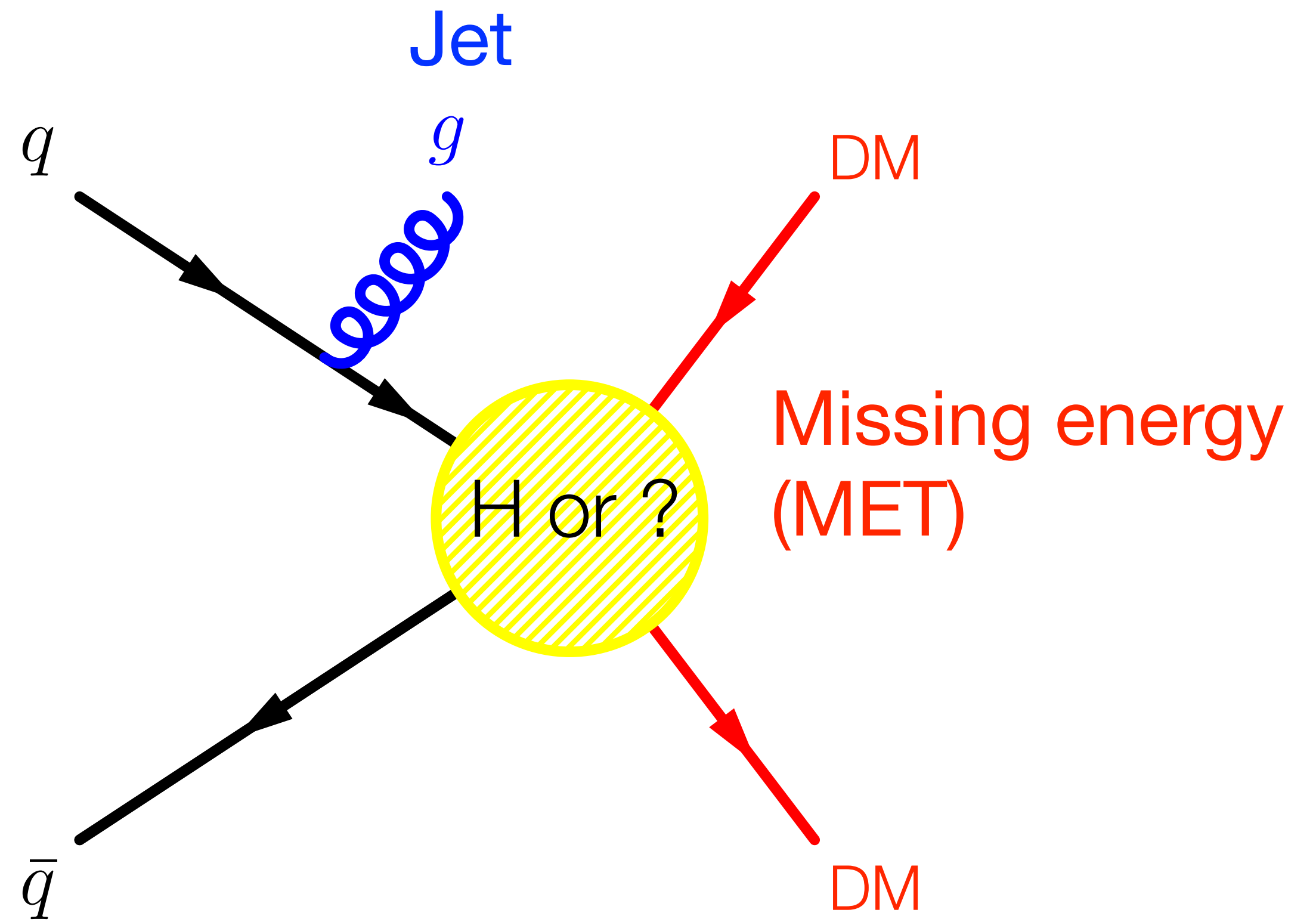




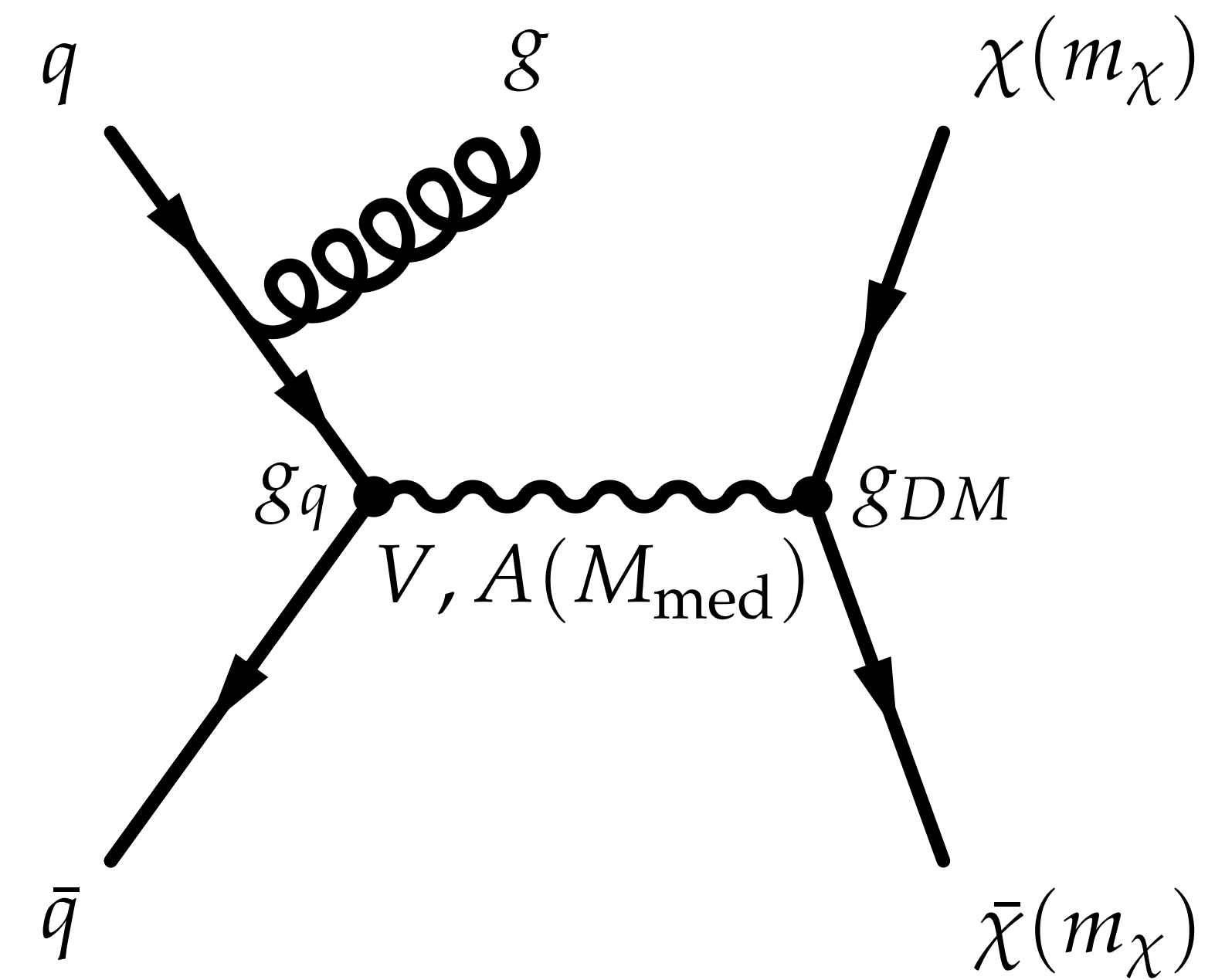
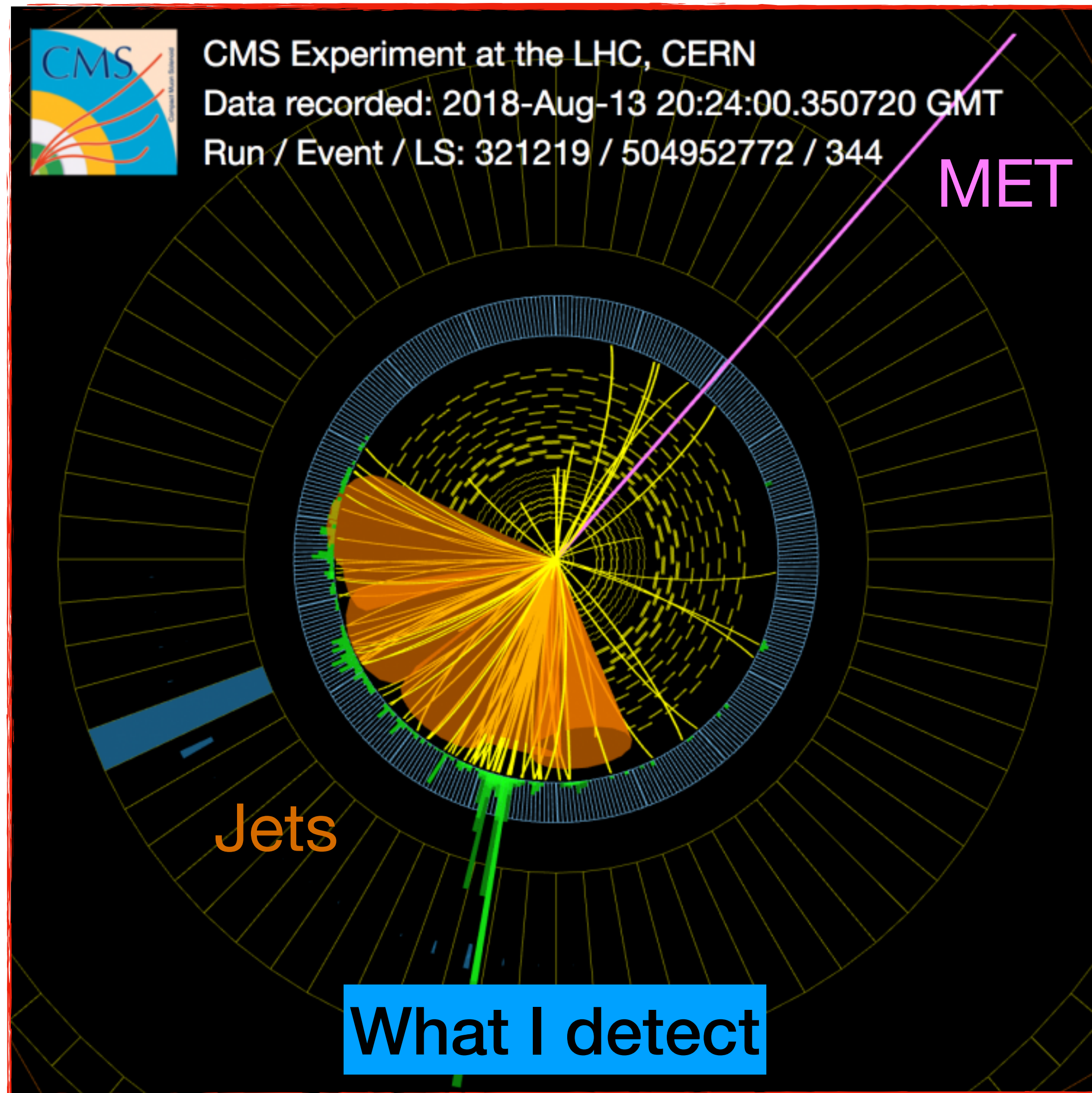
# Back to the drawing board

---

- Search for a Higgs boson (or anything else) decaying to dark matter
- Dark matter doesn't interact (i.e. don't detect it)
- See only its 'shadow'. Detect it via missing transverse energy

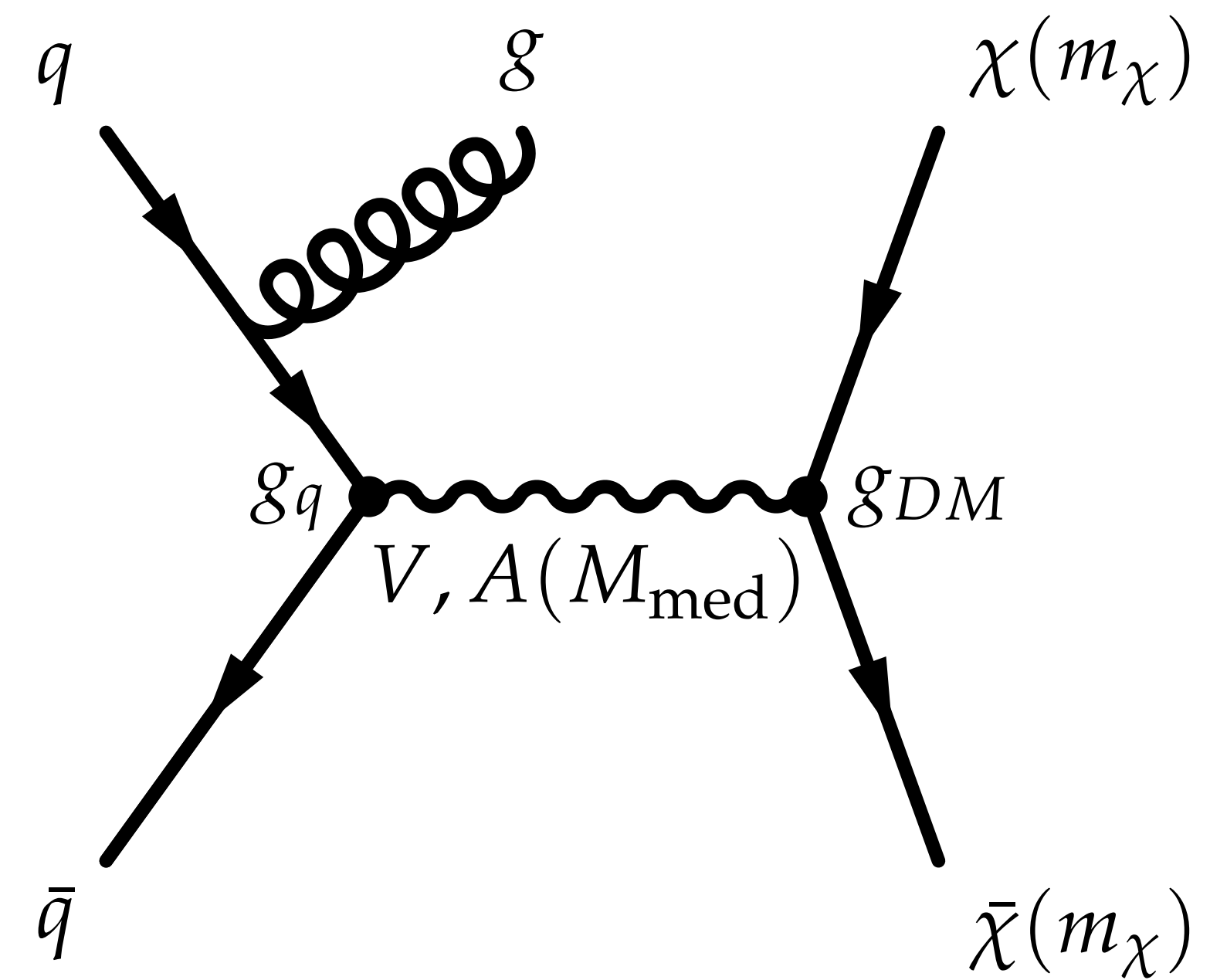
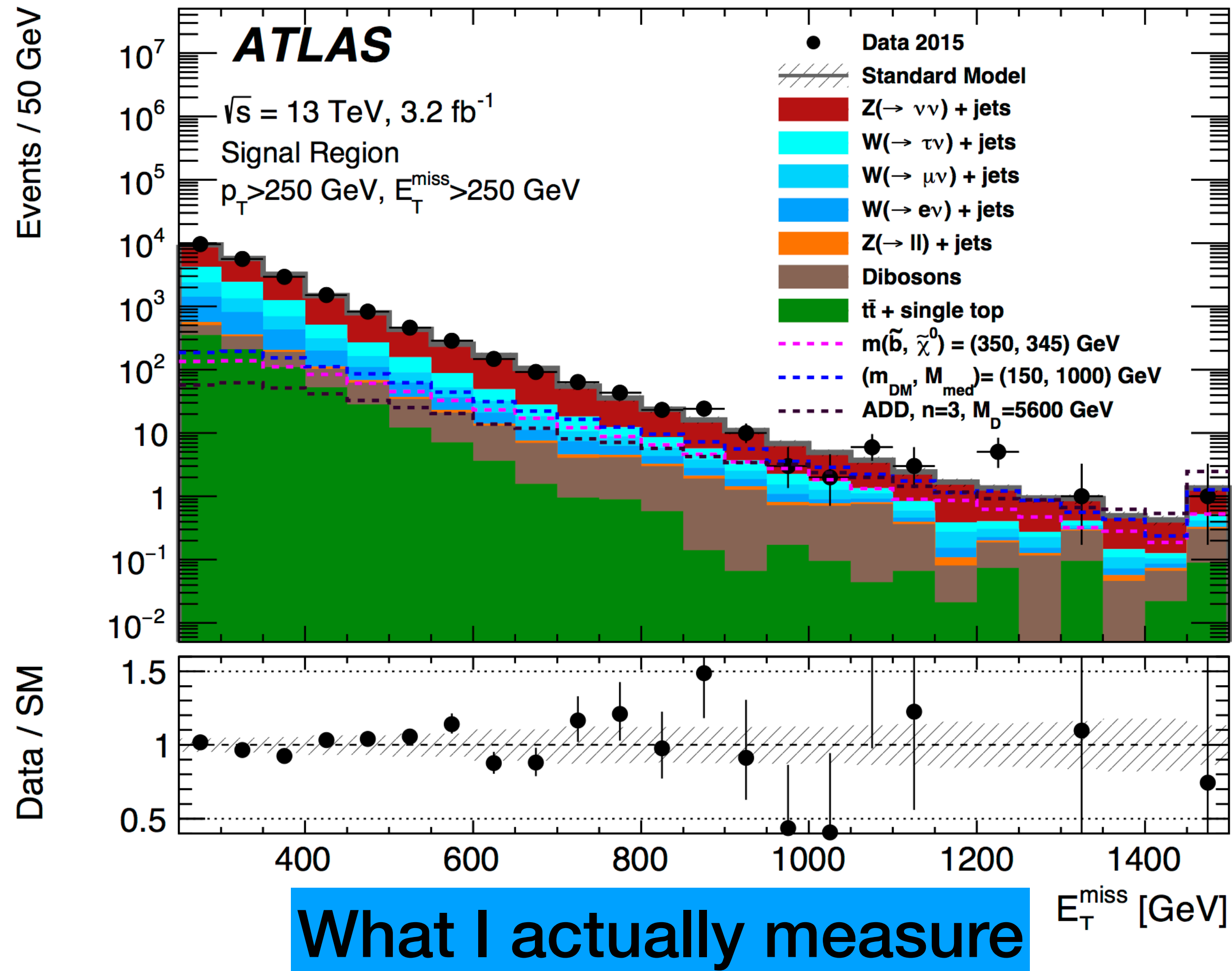


# Detecting dark matter



What I want to measure

# Dark matter and the Higgs

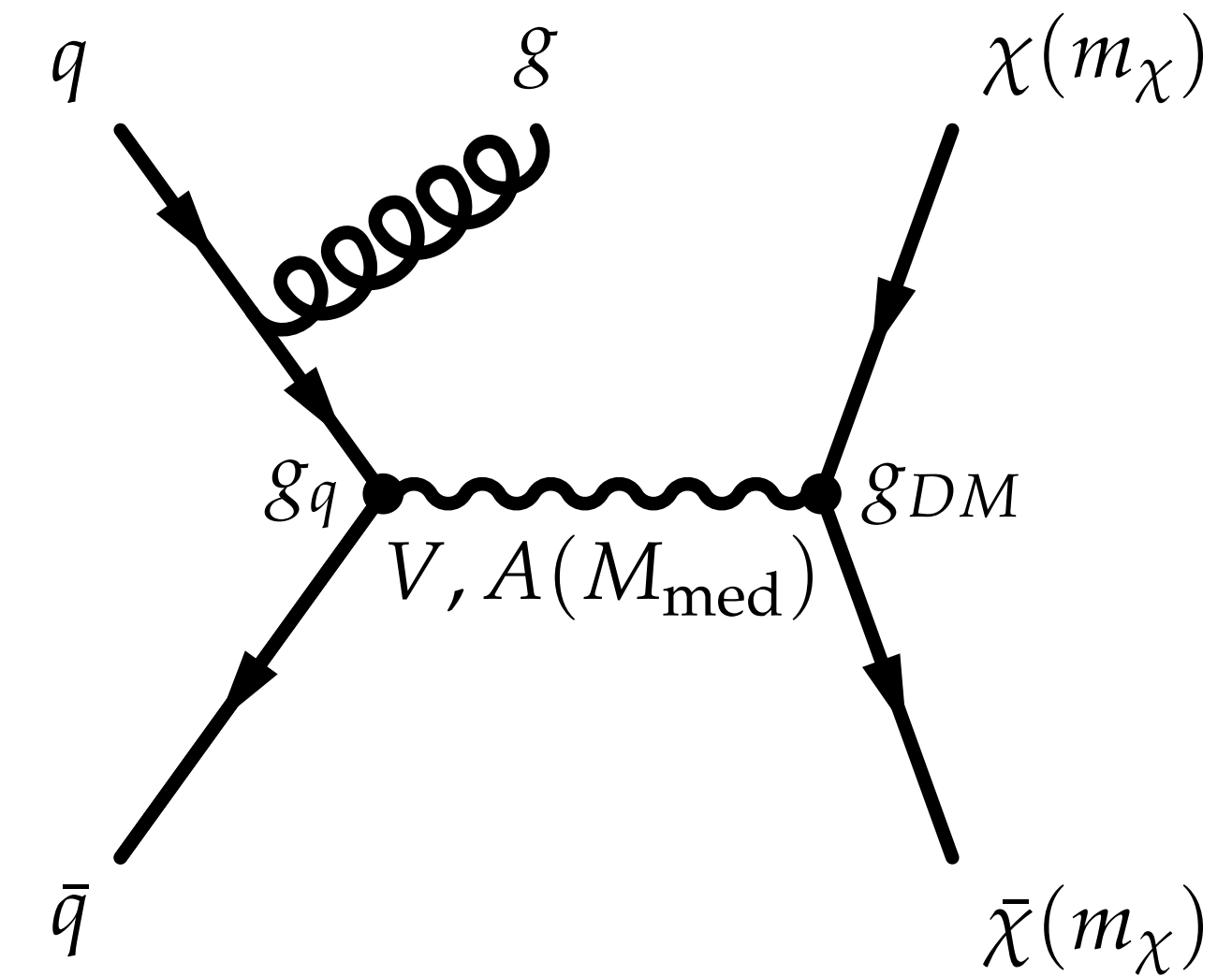
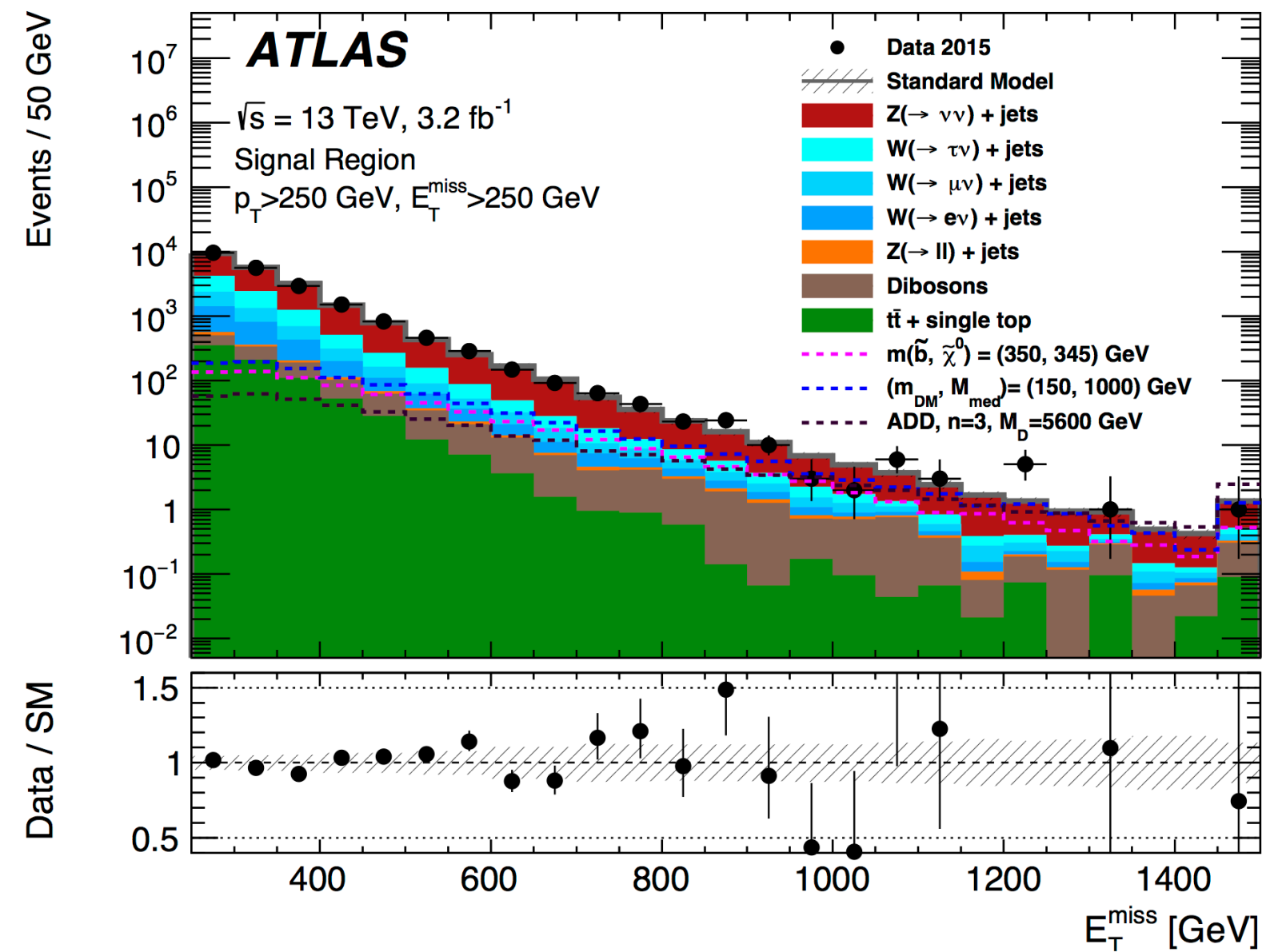


**What I want to measure**

# How do these actually connect?

Number of events

Cross section or couplings

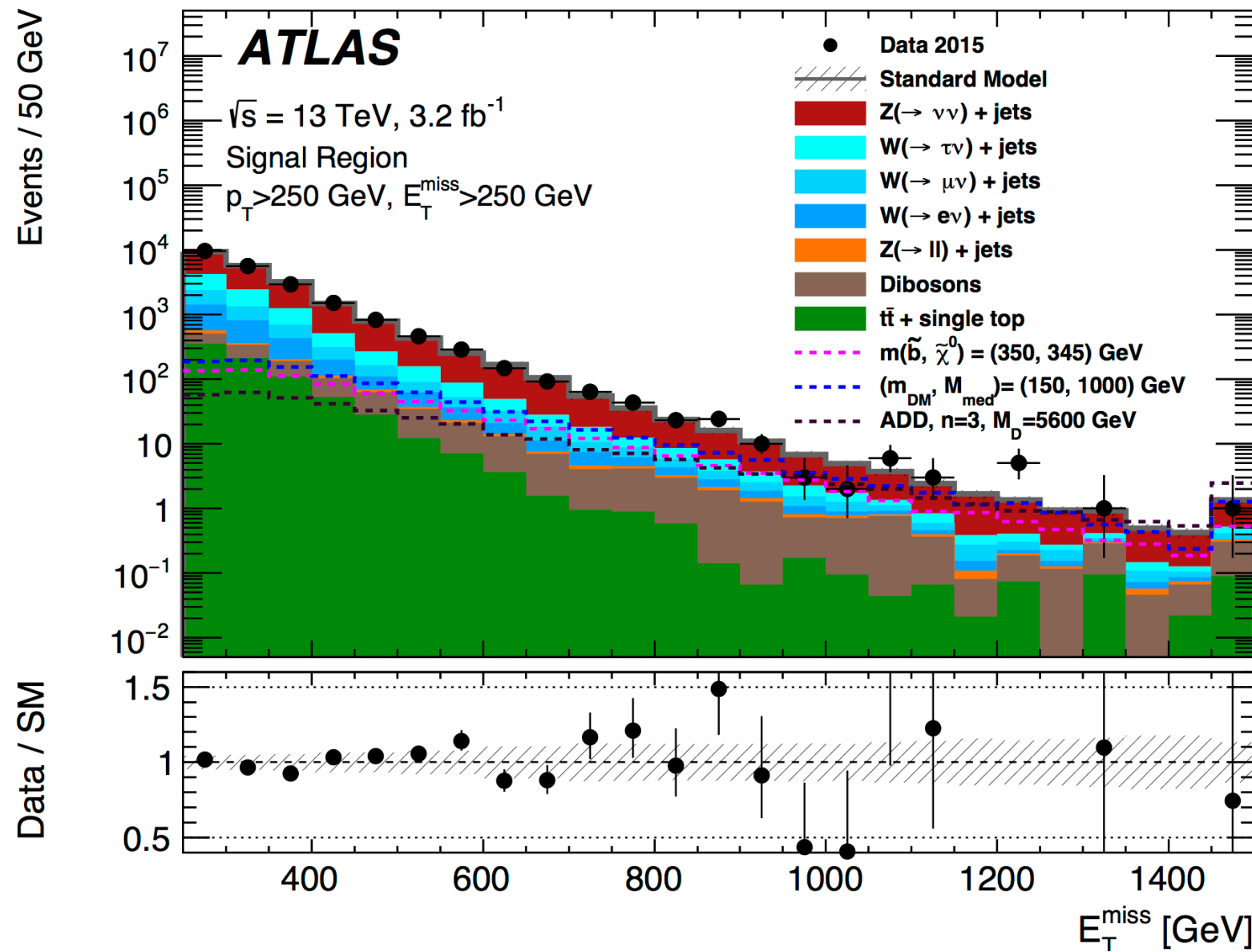


What I actually measure

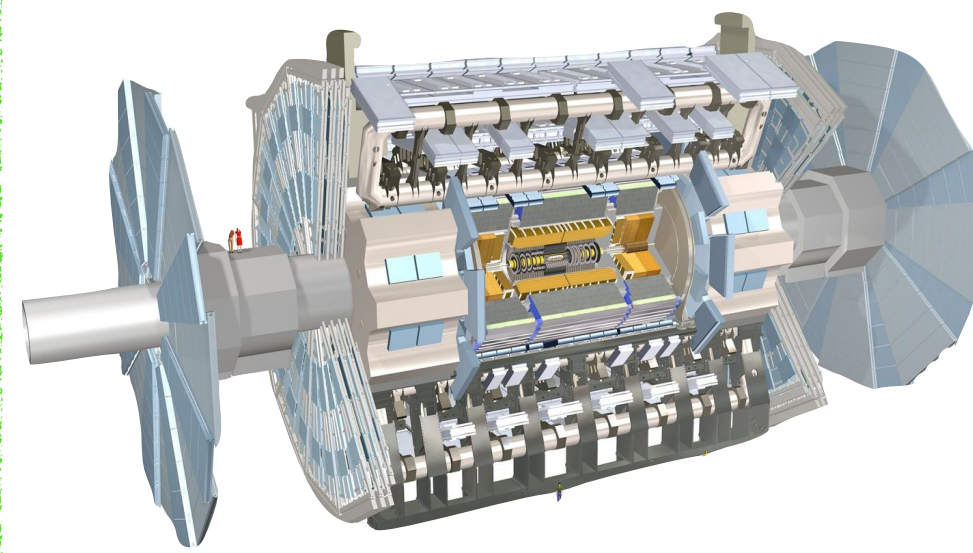
What I want to measure

# How do these actually connect?

Number of events



Detector effects

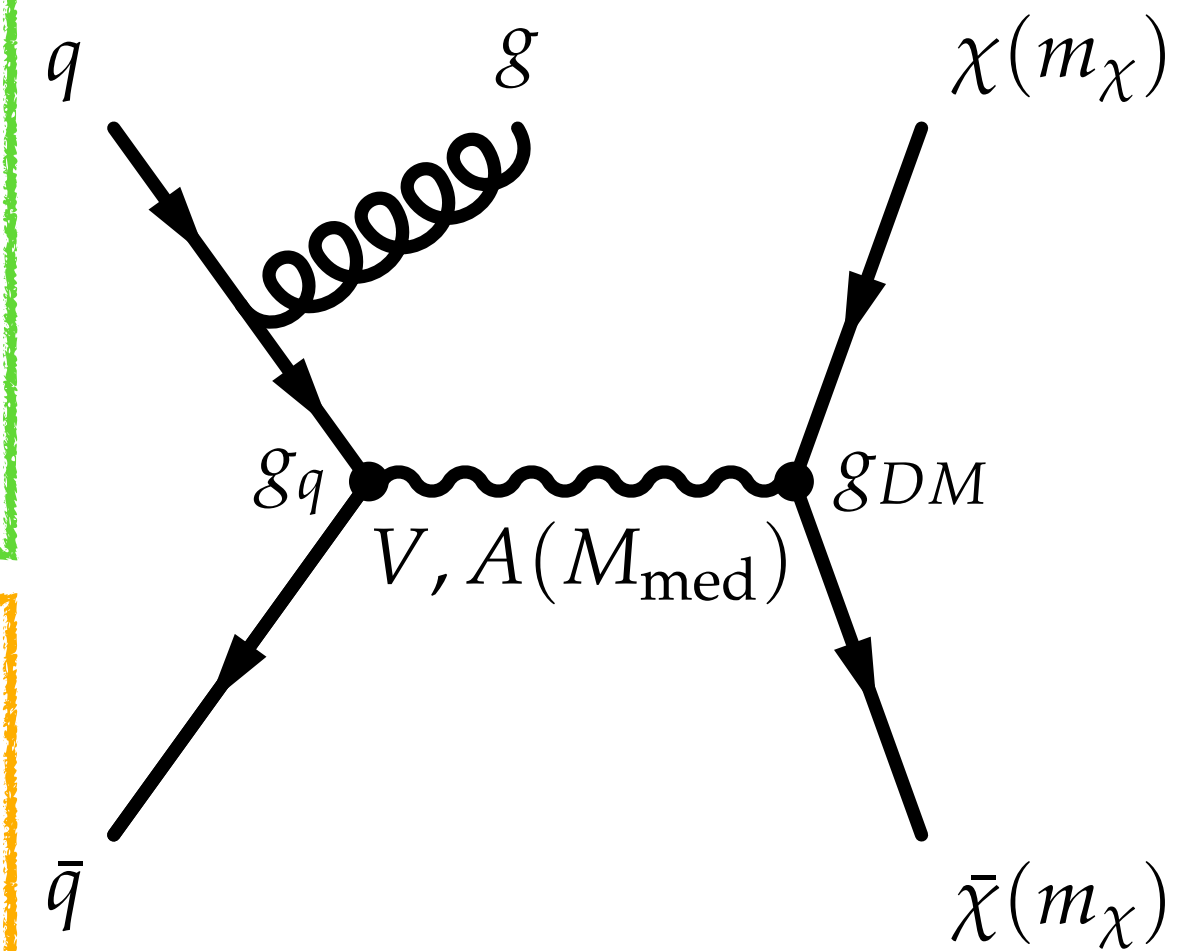


Calibration  
 Finite resolution  
 Finite coverage  
 Non-perfect efficiencies  
 Measurement uncertainties

Final state particle definition  
 Like lifetime cut (10ps)  
 Pre/post hadronization  
 Fiducial selection

Phase space/particle definition

Cross section or couplings



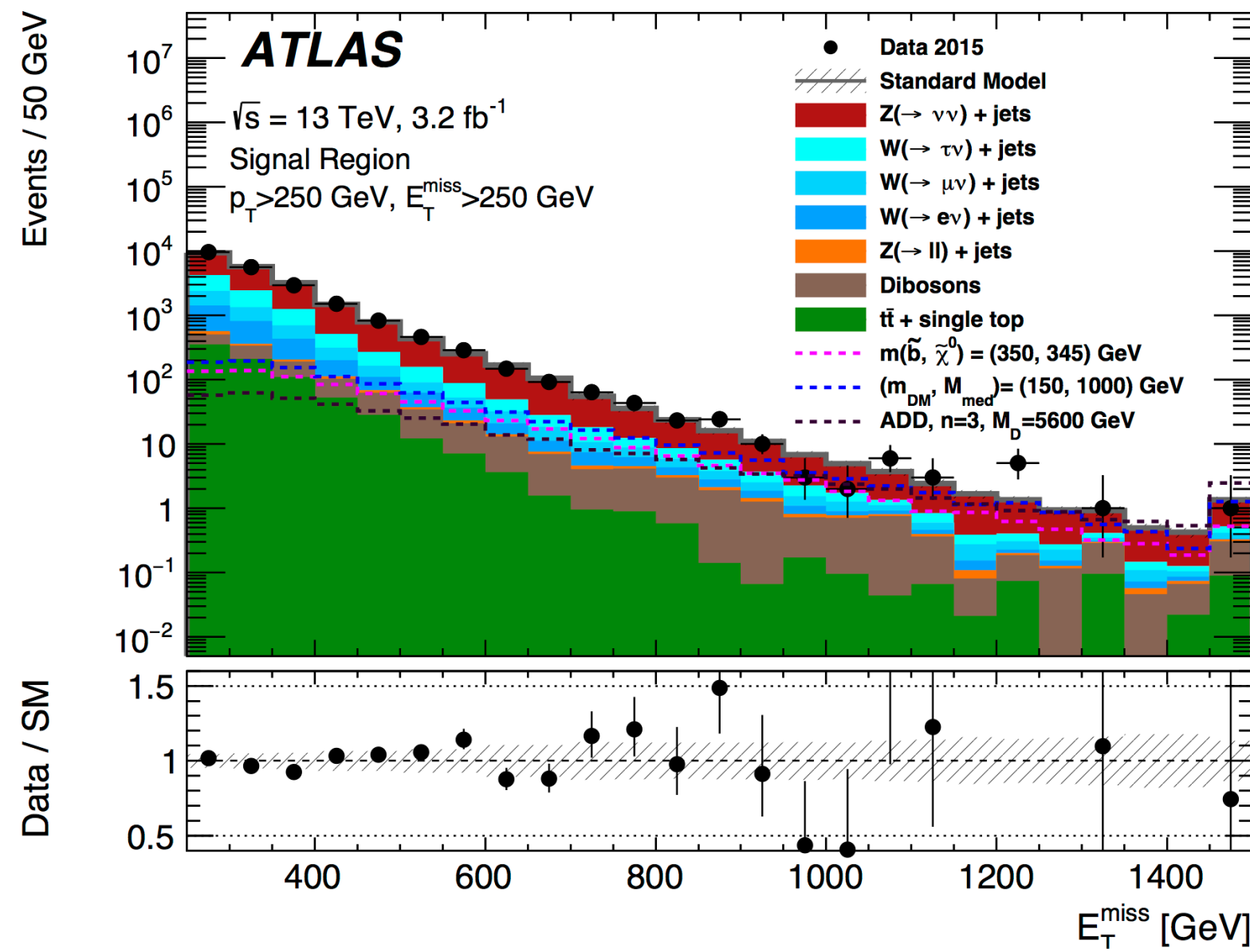
What I actually measure

What I want to measure

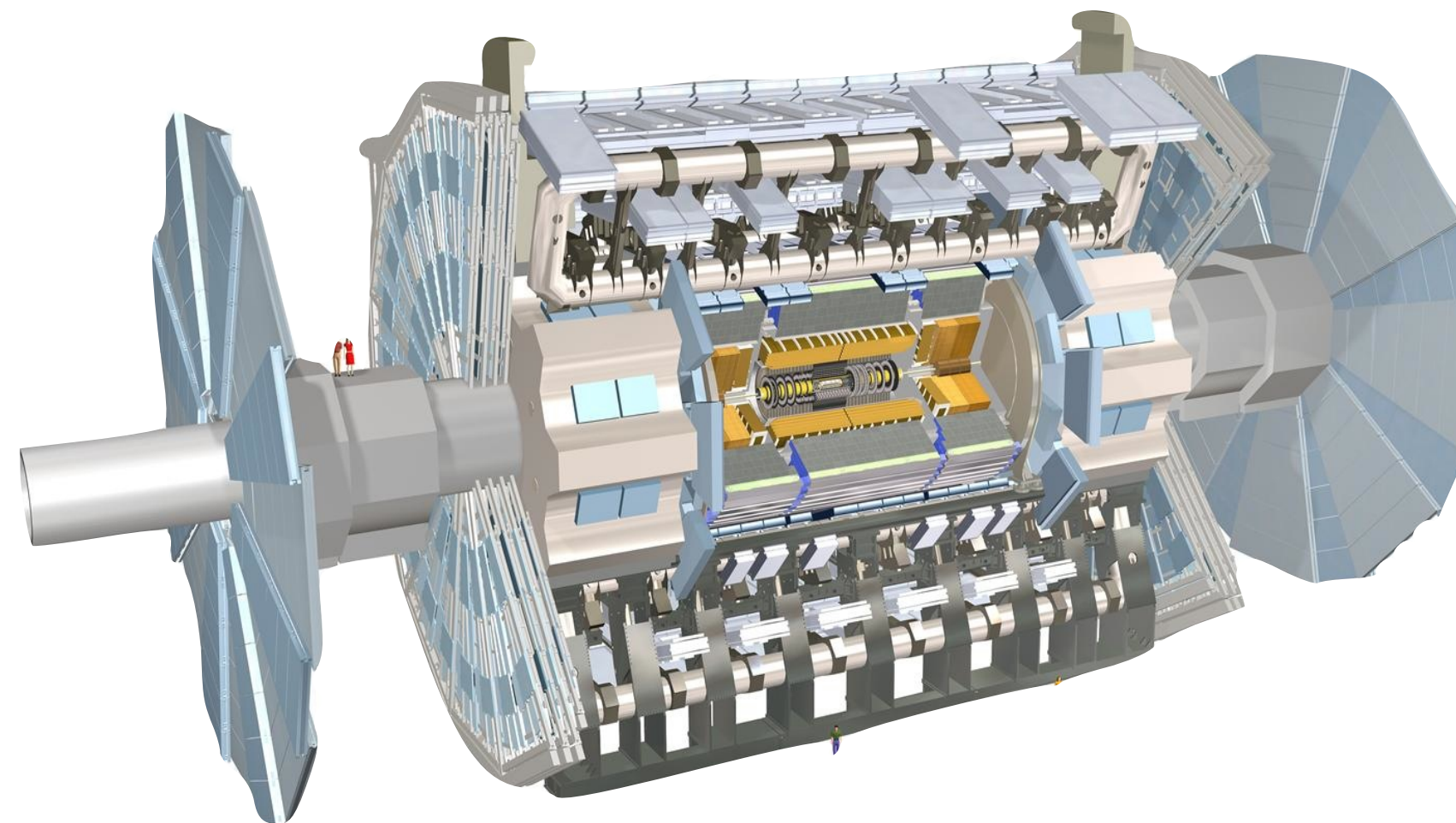
# How do these actually connect?

Use a MC + Simulation ...

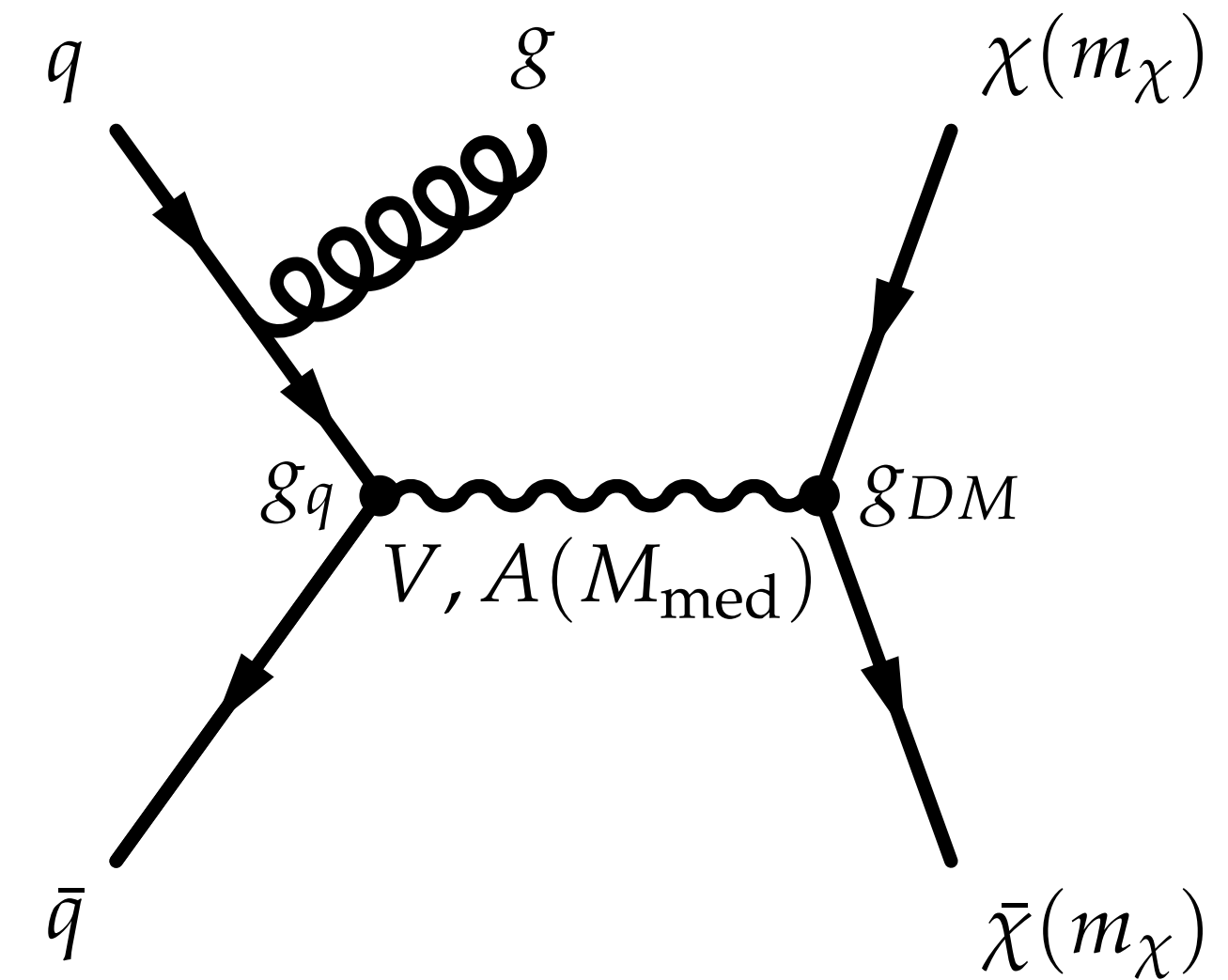
Number of events



To correct for all detector effect (i.e unfolding)



Cross section or couplings



To predict the data

# How do these actually connect?

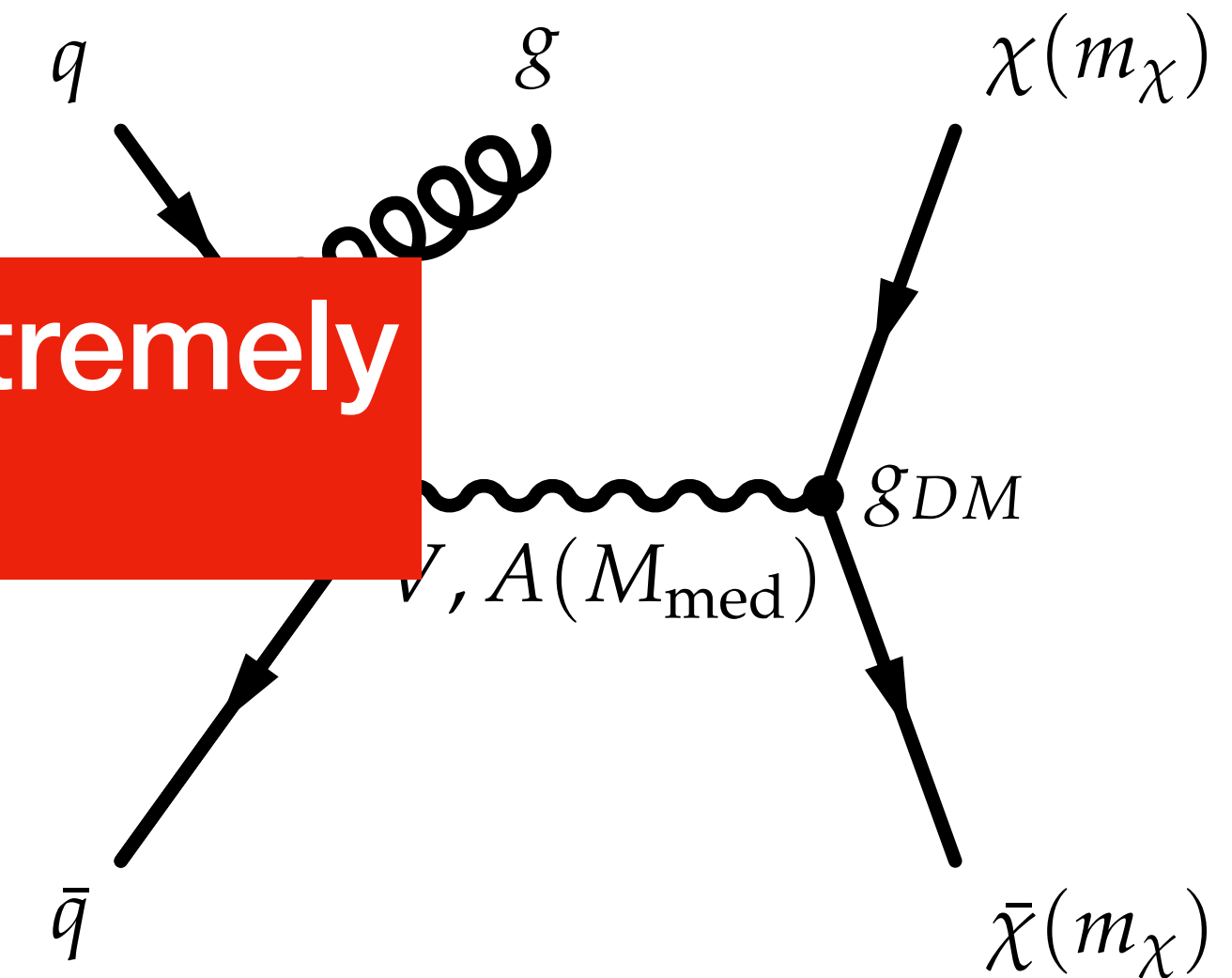
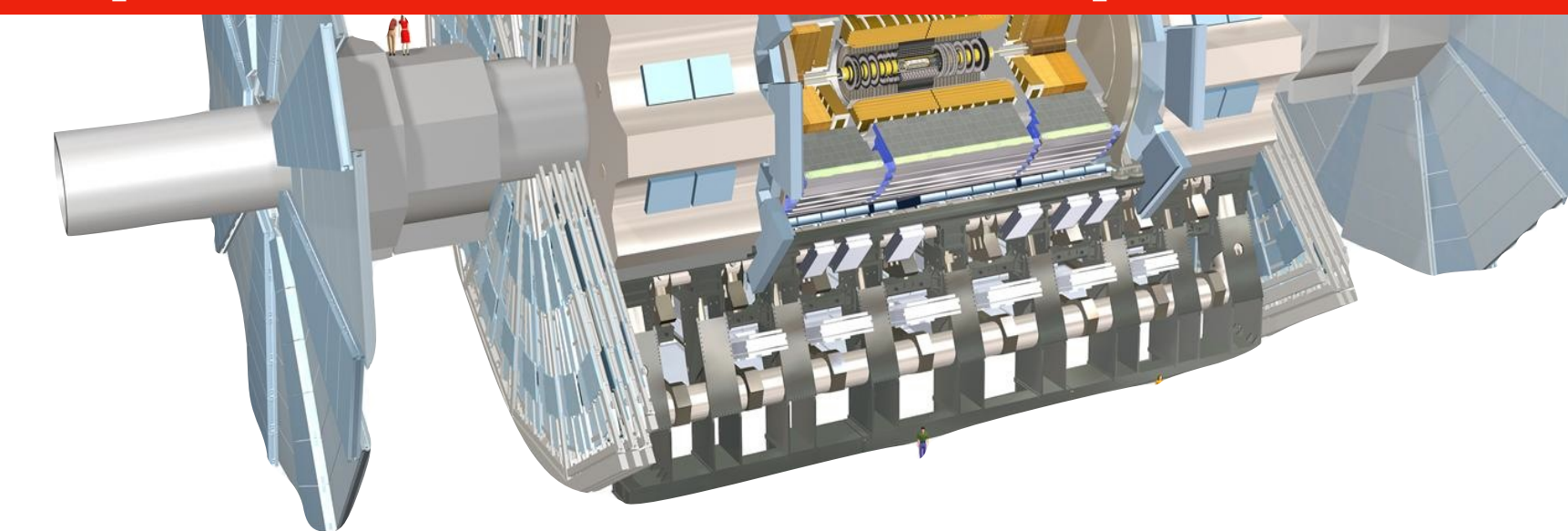
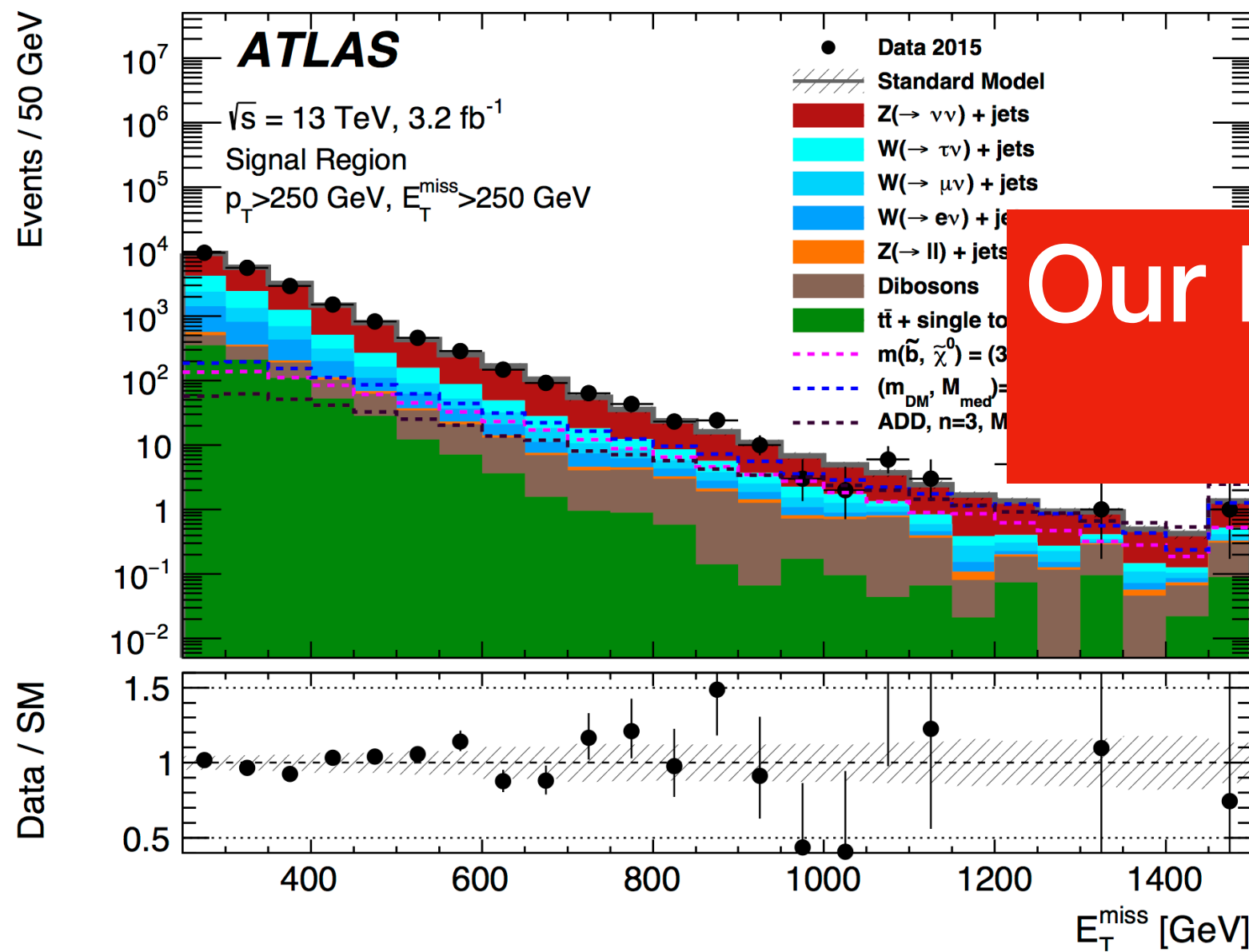
Use a MC + Simulation ...

Cross section or couplings

Number of events

To correct for all detector effect (i.e unfolding)

Our Monte Carlo and simulation are extremely powerful and unique tools!



To predict the data

Designing a measurement



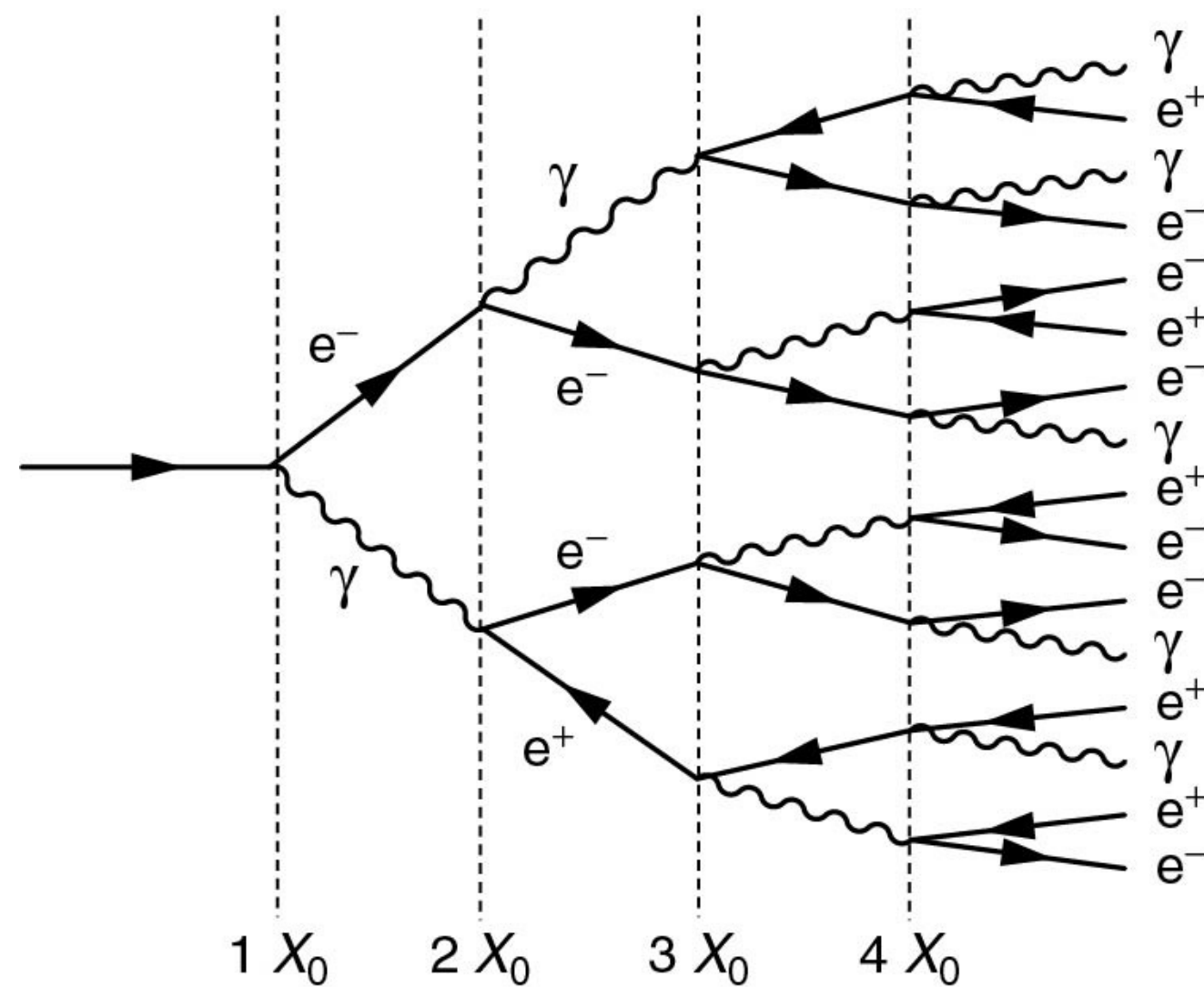
# Designing a measurement - some key ingredients

---

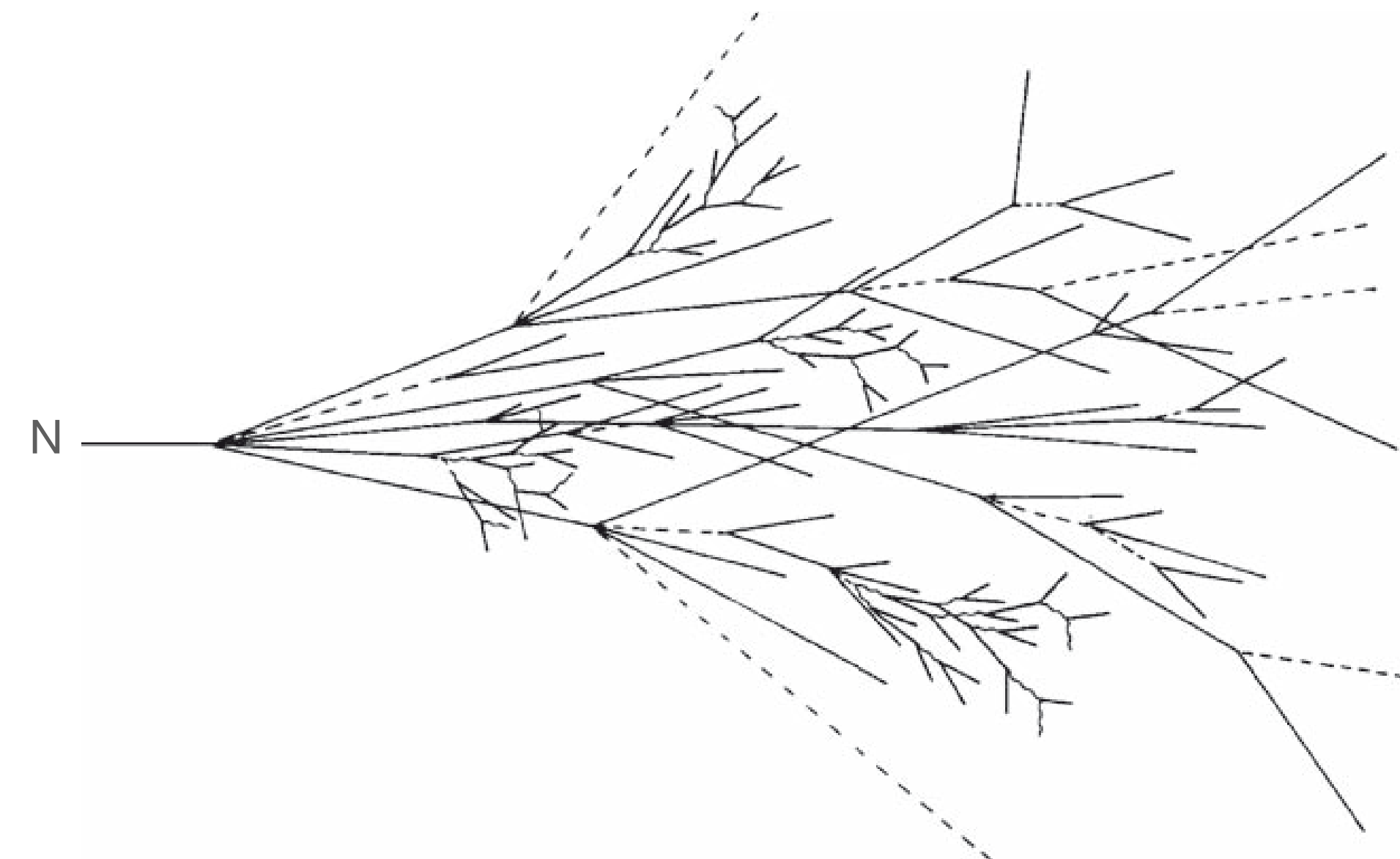
- What is the final state and how precisely can you reconstruct it? Is it precise enough for your question?
- How large is the signal (what you want to measure/find) compared to backgrounds with the same final state (other processes that are in your way)?
- What will be the dominate uncertainty? Can it be controlled or constrained?

# A brief aside... Experimentally not all particles are equal

- Muons, electrons, photons are most precise, hadrons less so
- Comes down to how particles lose energy



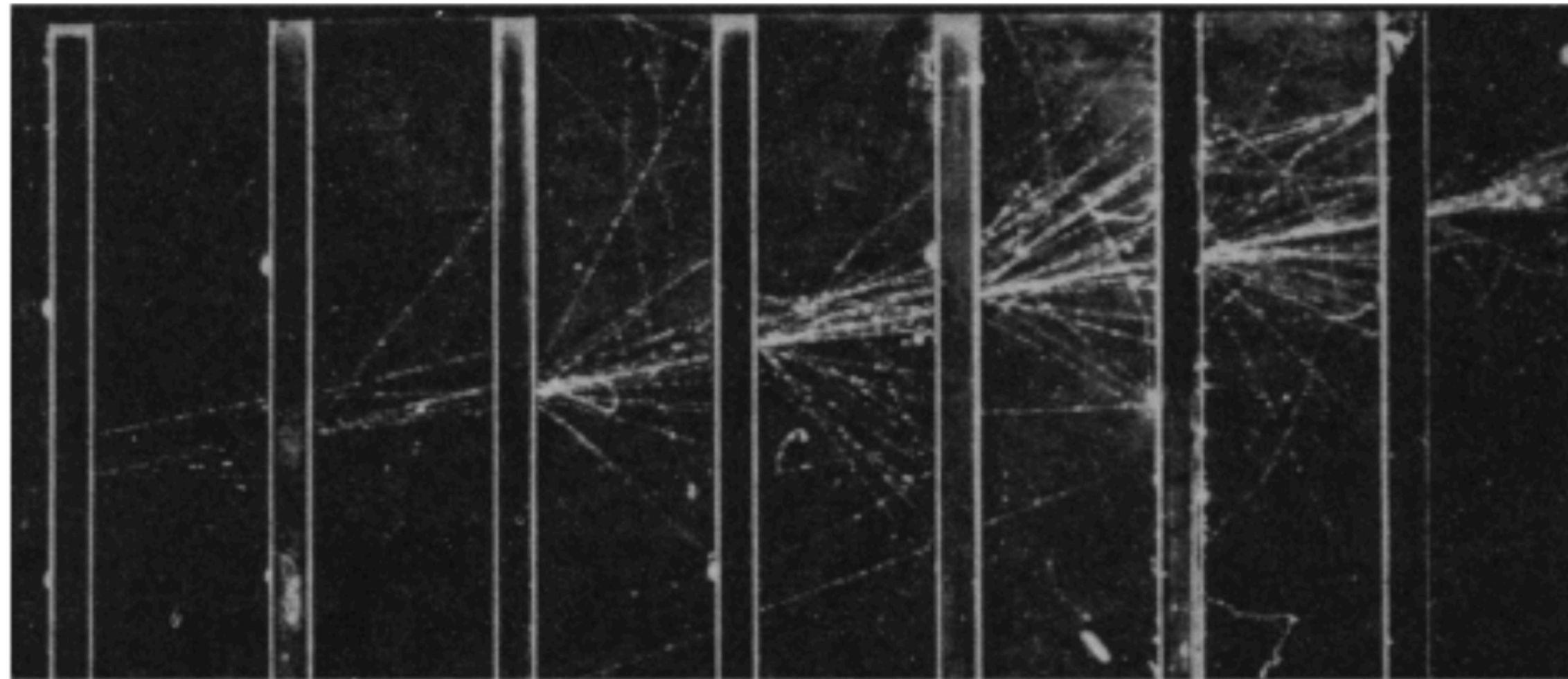
**EM shower - many interactions, fluctuations are small (i.e. EM showers look similar)**



**Hadronic shower - energy loss via strong interaction and ionisation. Final state varies, large shower-to-shower variations**

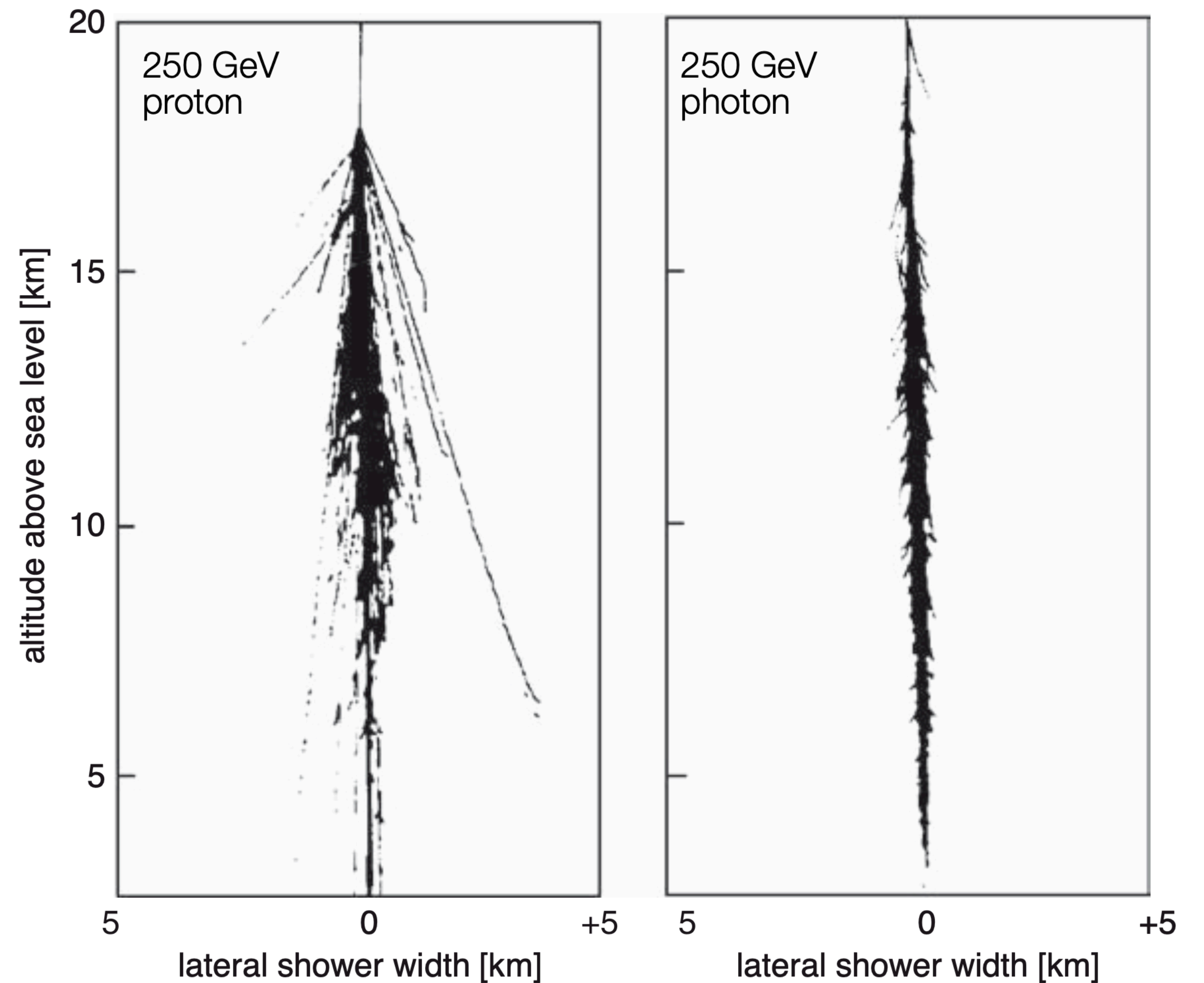
# A brief aside... Experimentally not all particles are equal

- Some examples



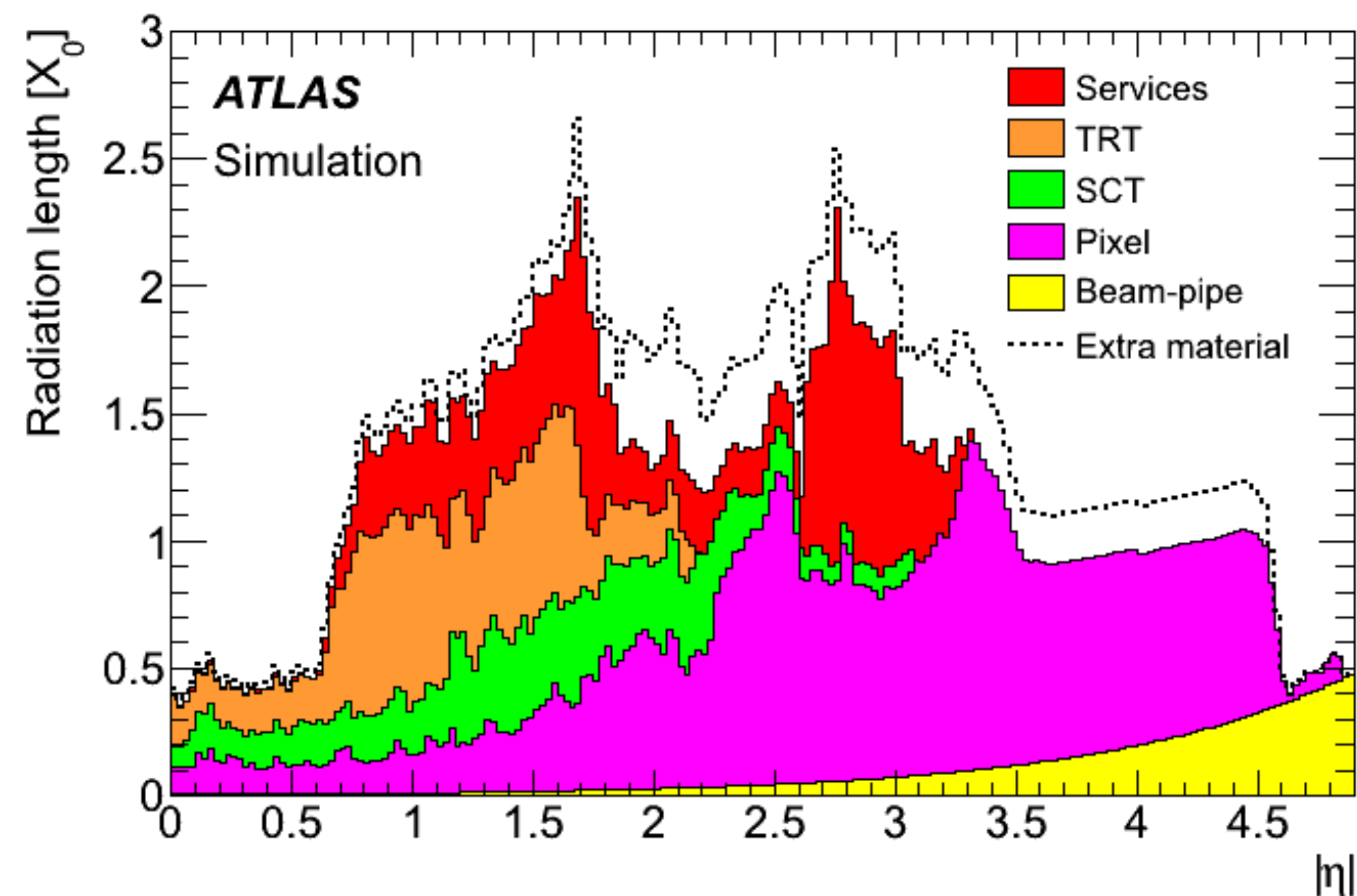
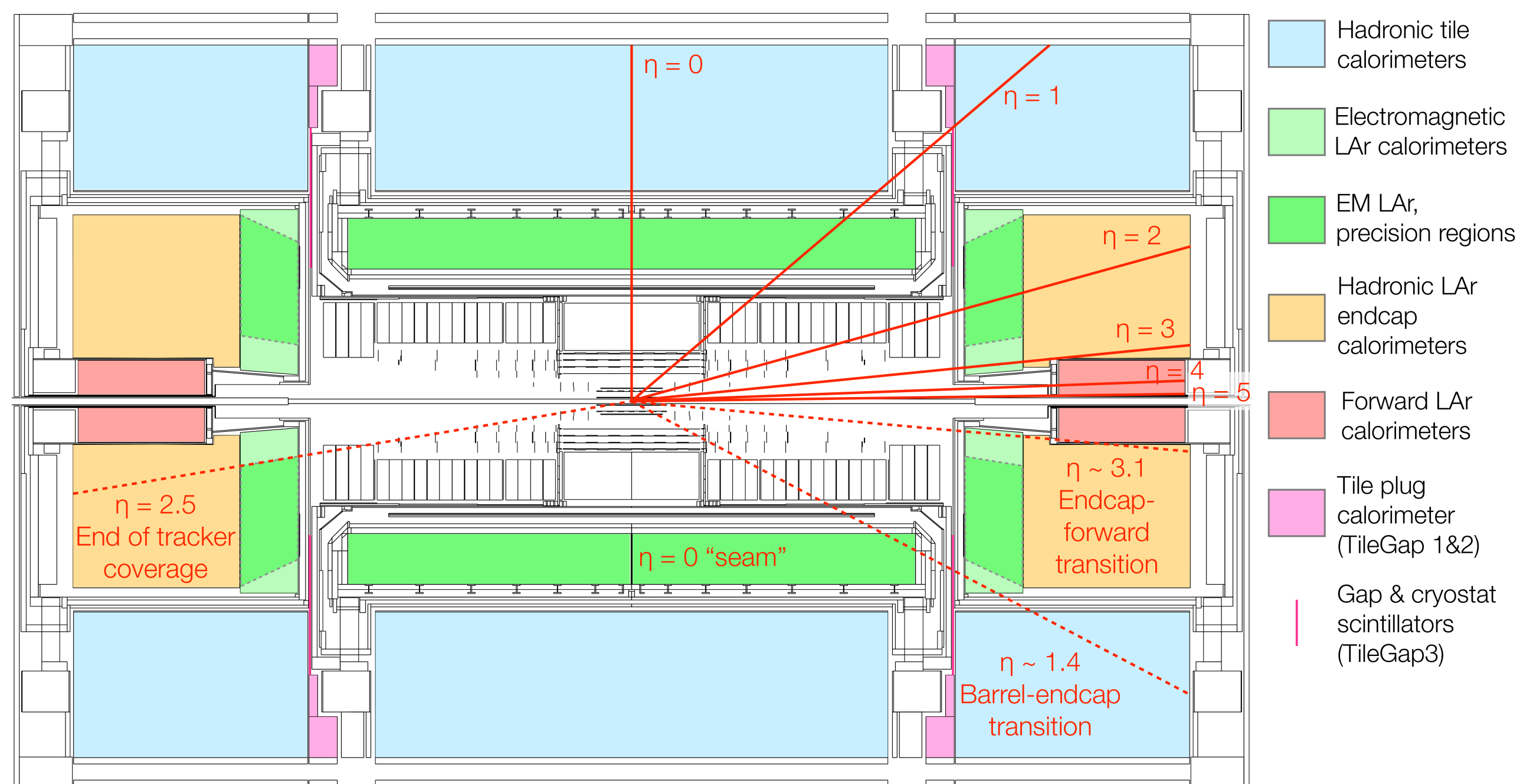
EM shower in a sampling calorimeter

## Air shower simulation



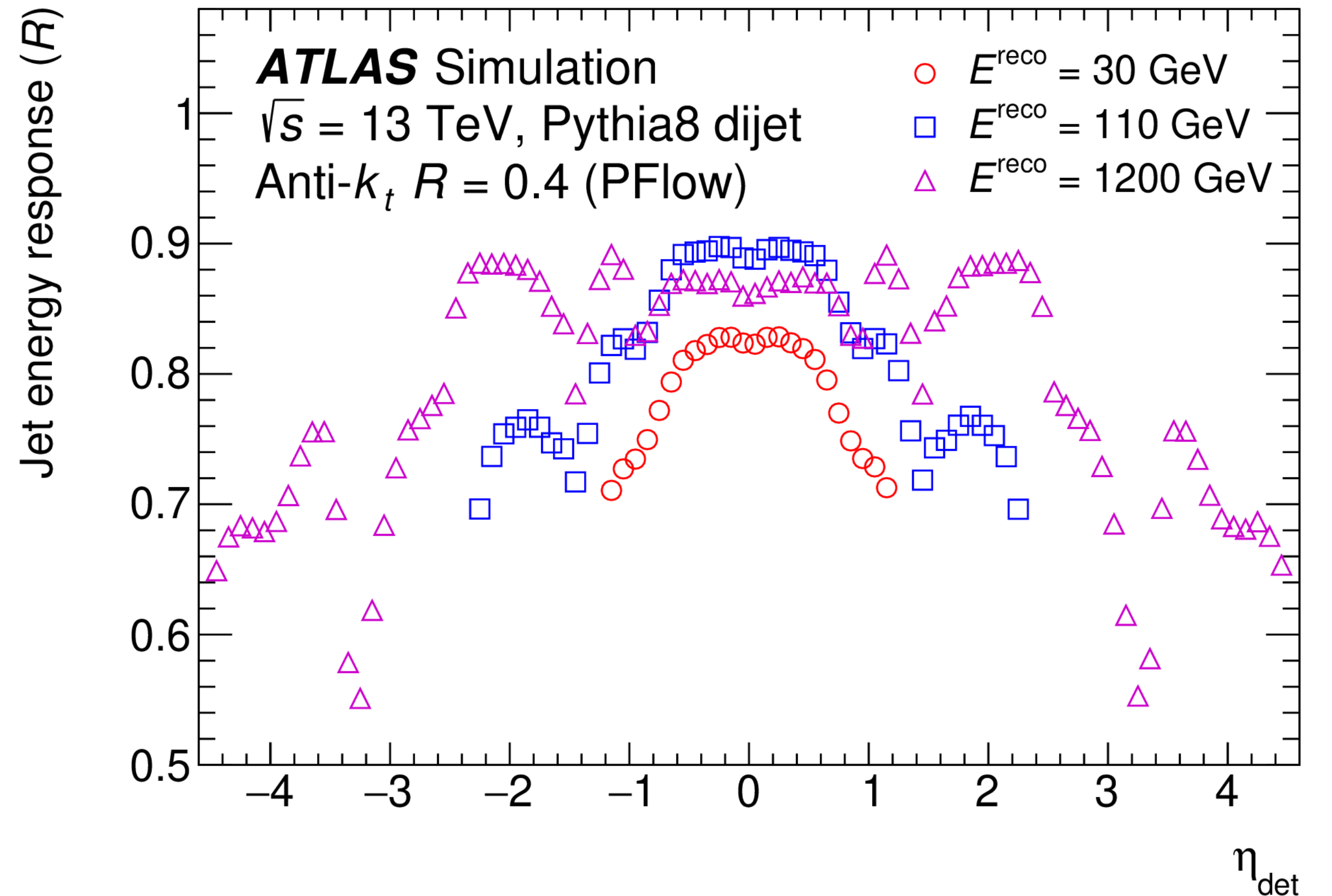
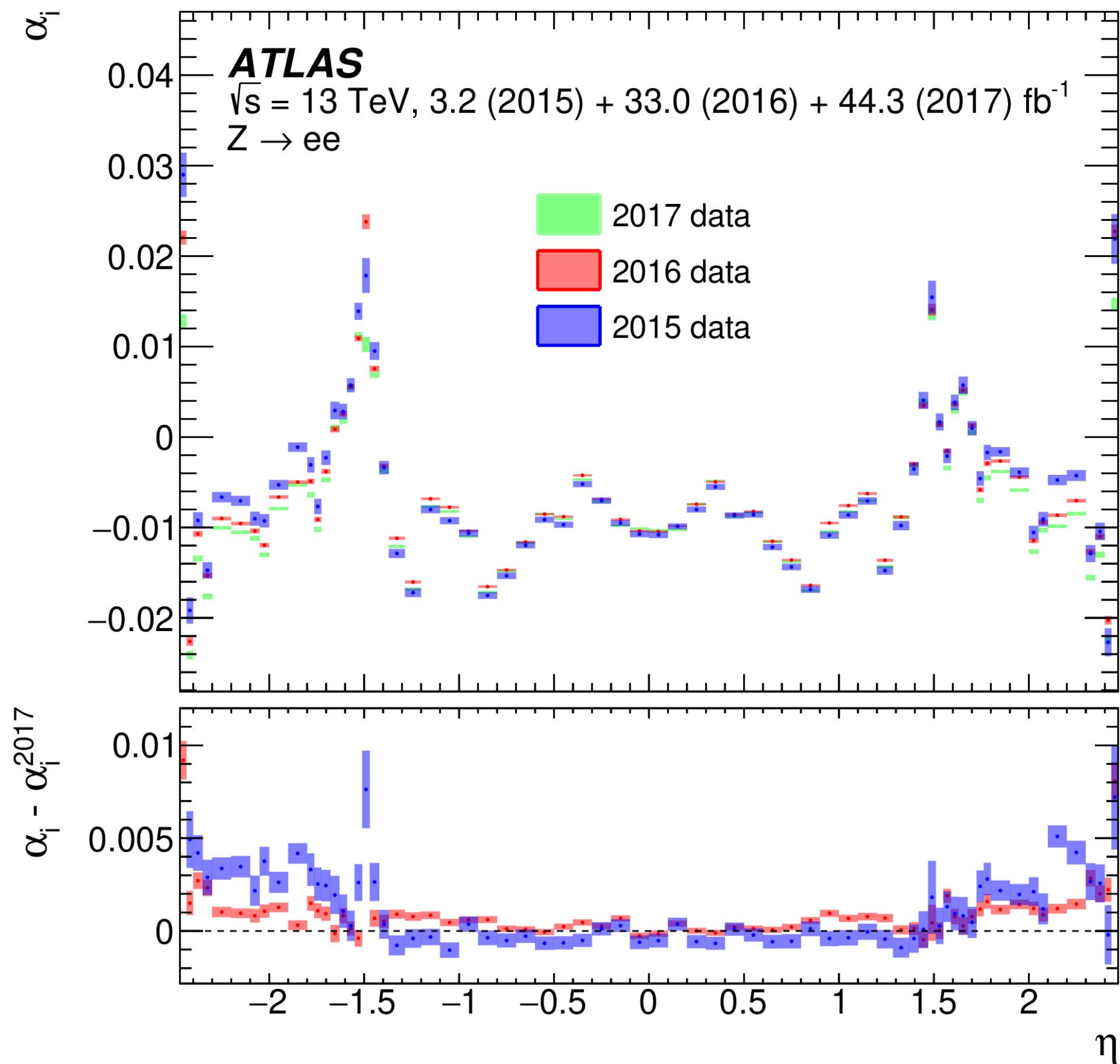
# Detector corrections

- Must account and correct for many effects - like pile-up, different detector technologies, material budget



# Detector corrections

- Often have strong  $\eta$  dependence, some corrections can be large

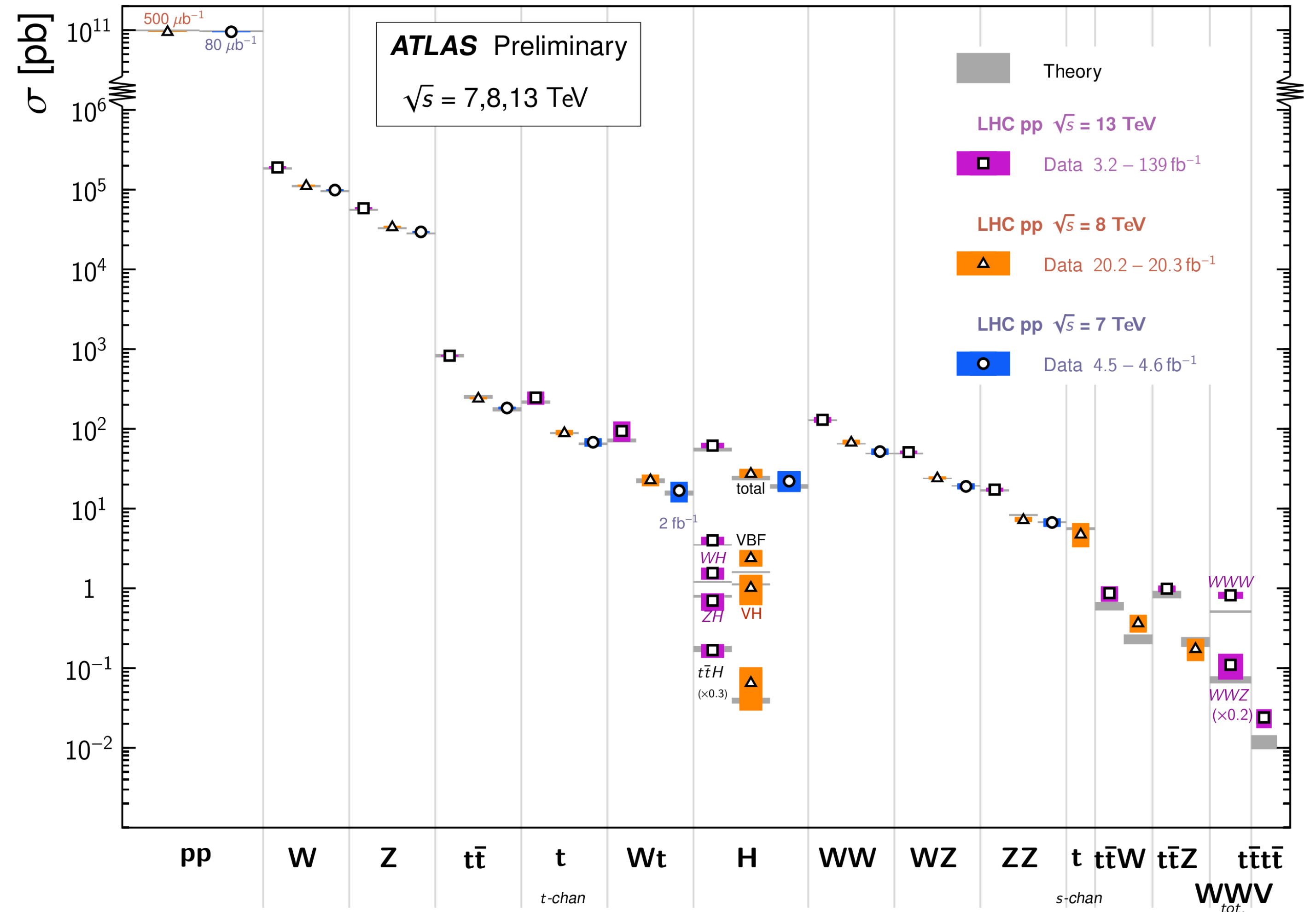


# A brief aside... Theoretically not all particles are equal

- QCD dominates (i.e. we produce a lot of jets)
- For the W/Z decays to hadrons dominate
- For Higgs, decays to b-quarks and taus dominate

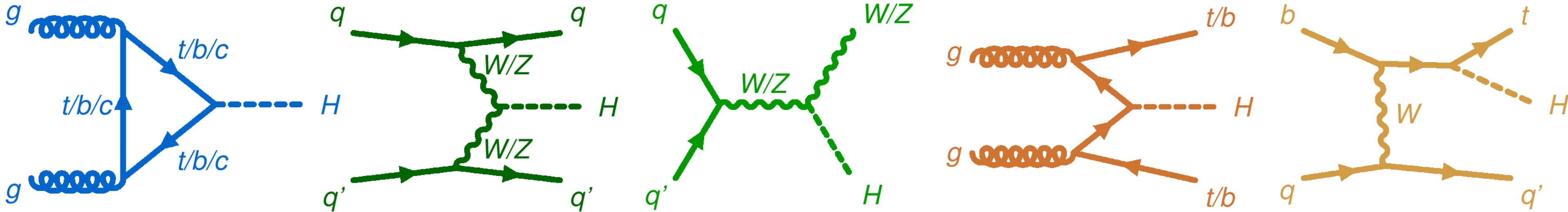
Standard Model Total Production Cross Section Measurements

Status: February 2022



# As a general rule - A Higgs example

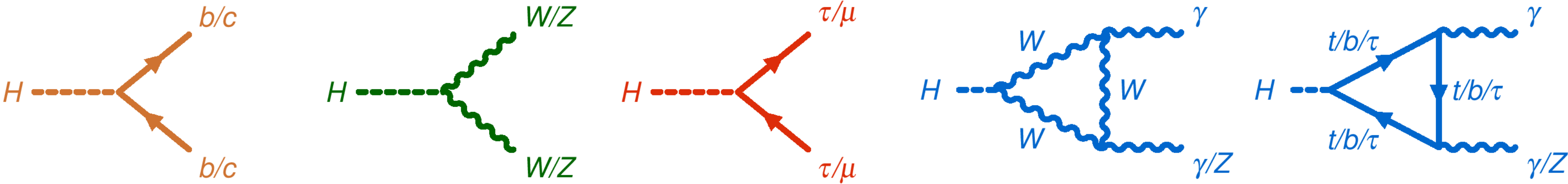
Production



Probability

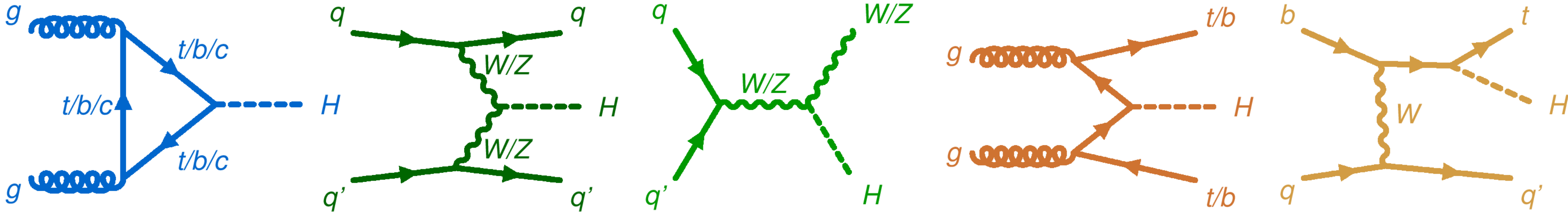


Decay



# As a general rule - A Higgs example

Production

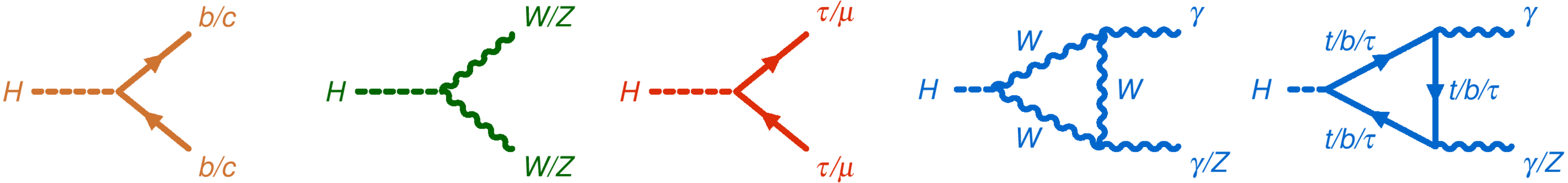


Probability



Experimentally

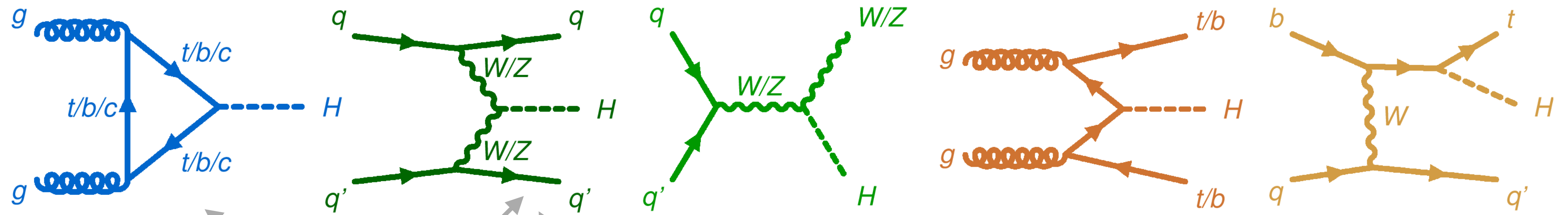
Decay



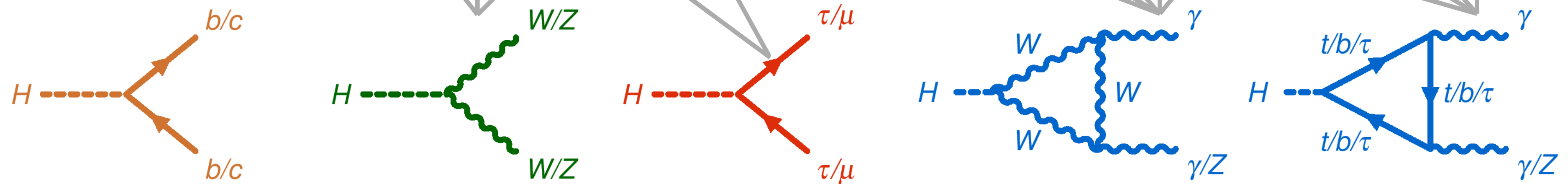


In practice - measure them all

Production

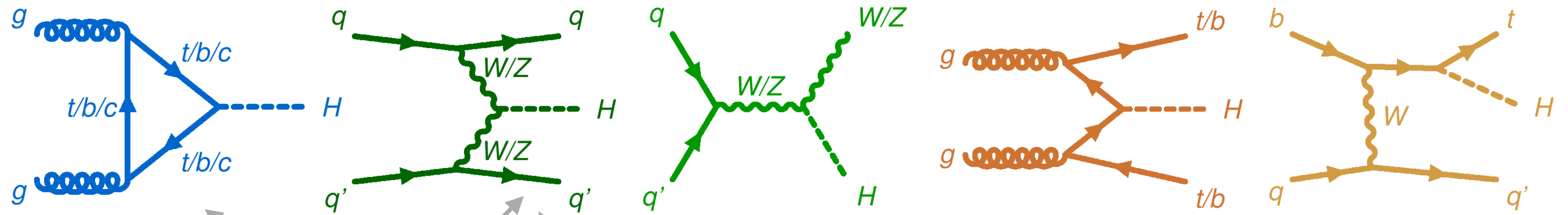


Decay



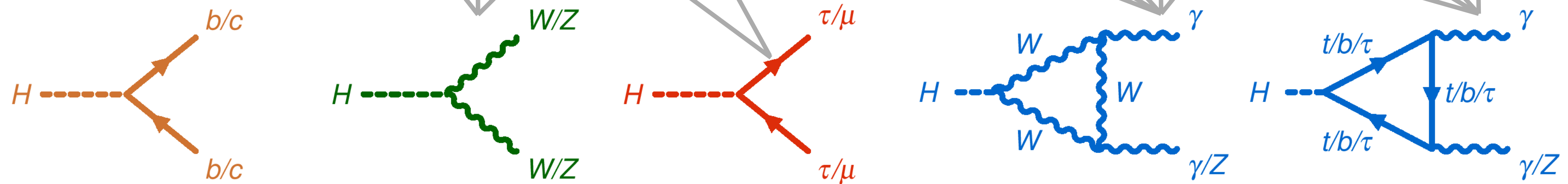
In practice - measure them all

Production

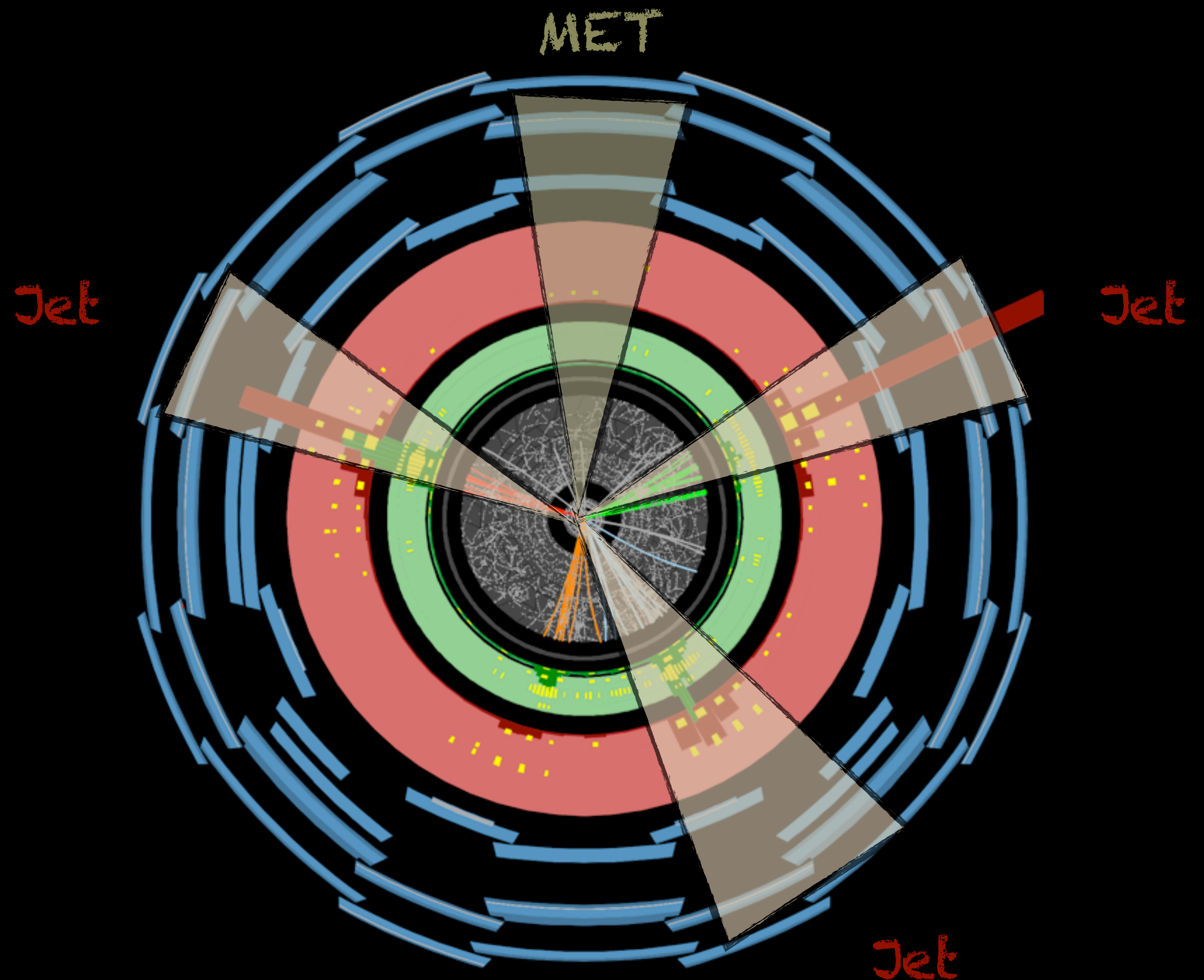


Each measurement brings different experimental and theoretical strengths and weaknesses

Decay



- How precisely can you reconstruct the final state?  
Is it precise enough?



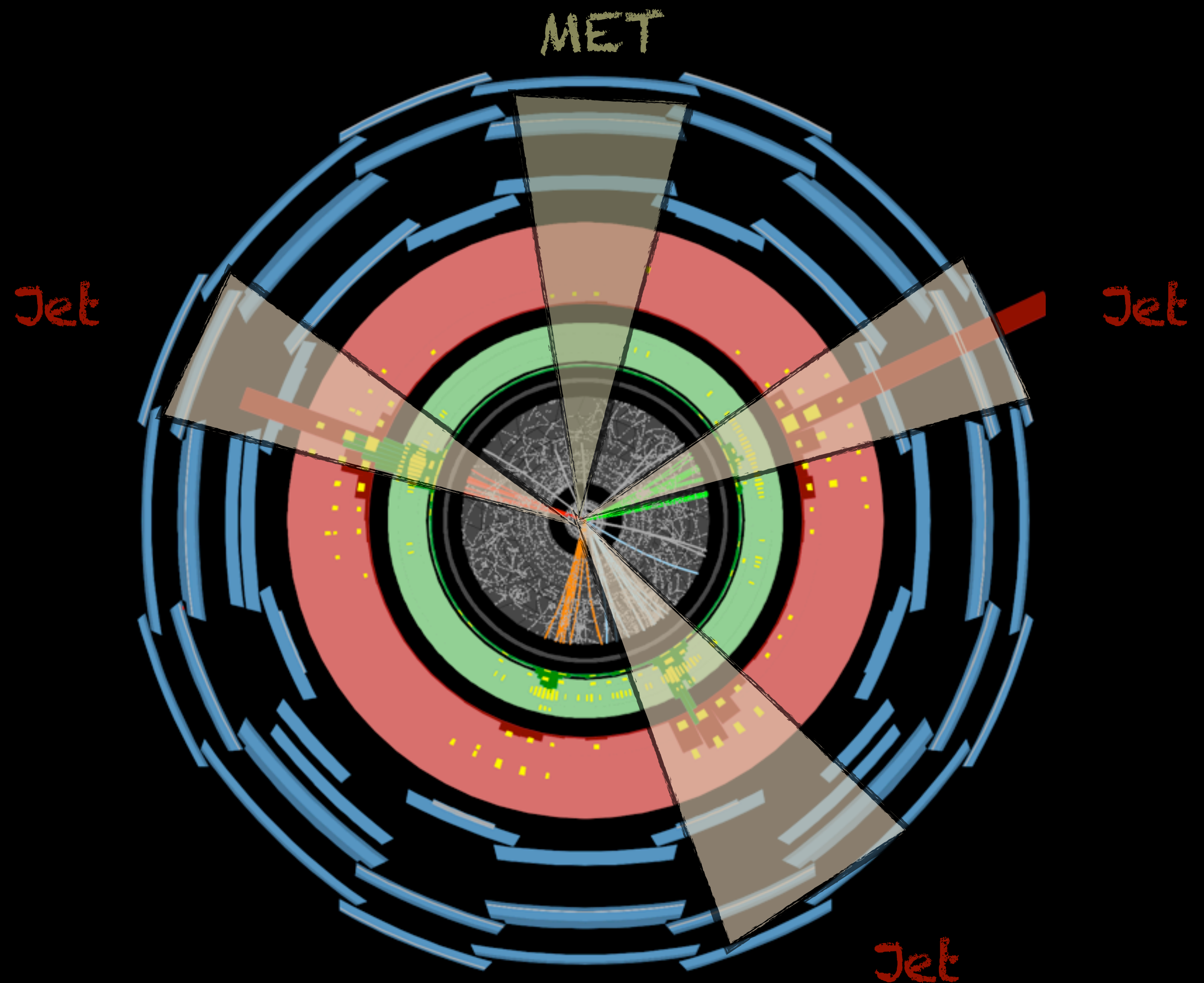
$$\text{Energy}_{xy} \text{ DM} = \text{Energy}_{xy} \text{ before} - \text{Energy}_{xy} \text{ Jets}$$

- How precisely can you reconstruct the final state?  
Is it precise enough?

Missing energy is the weakest link

Relies on measuring everything else

The **detector resolution** of each other object and **the uncertainty** on the measurement of each other object is propagated to the determination of missing energy



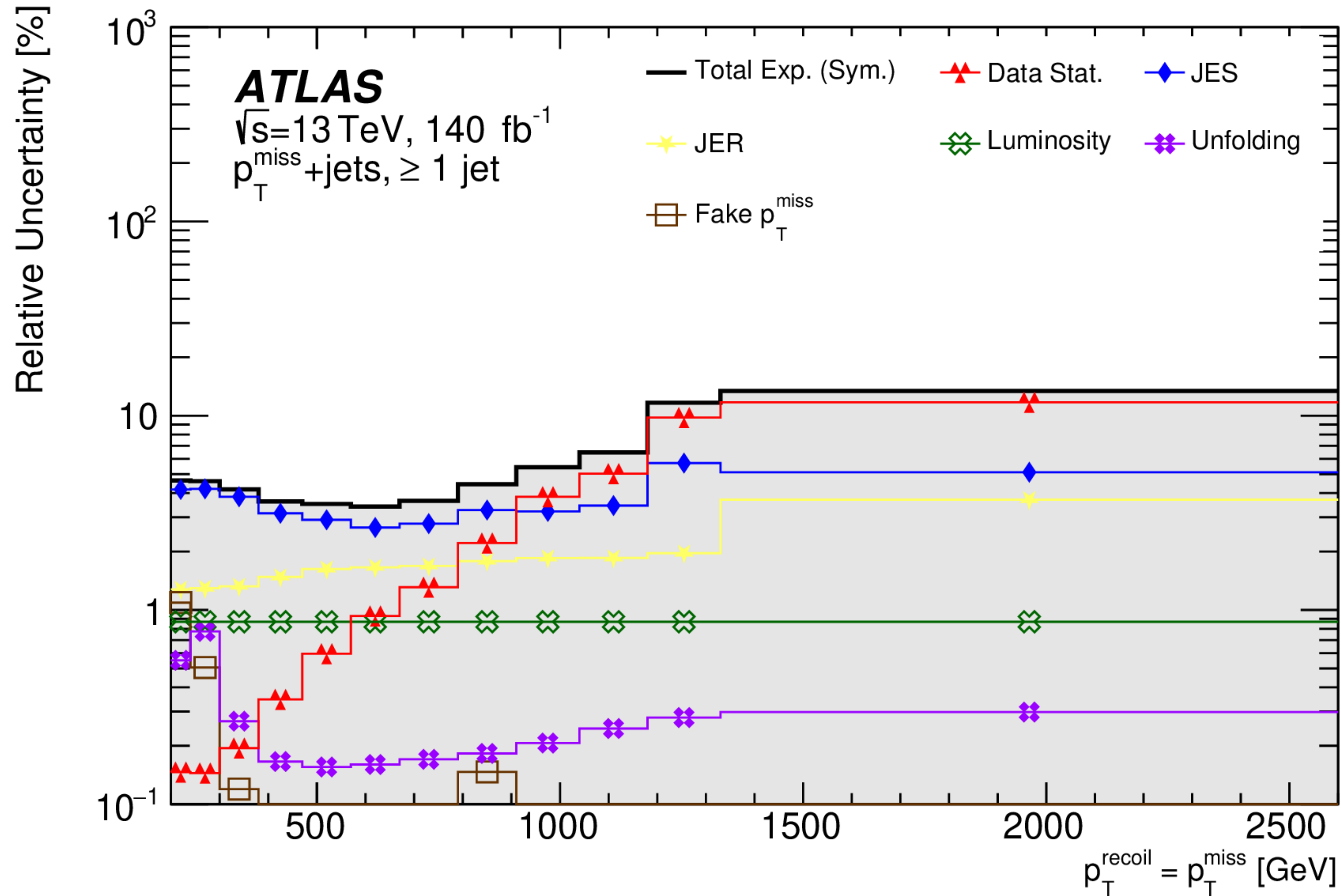
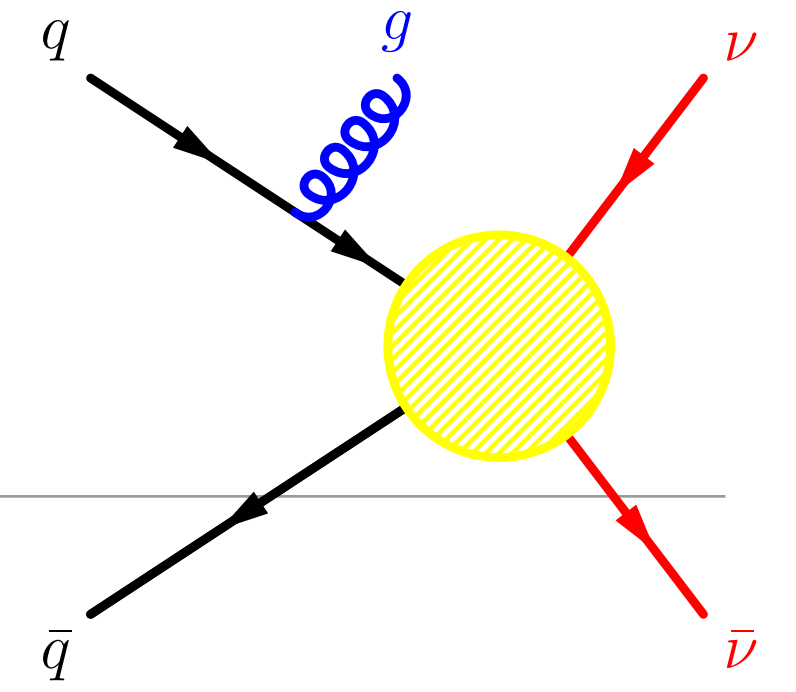
Energy<sub>xy</sub> DM

=

-

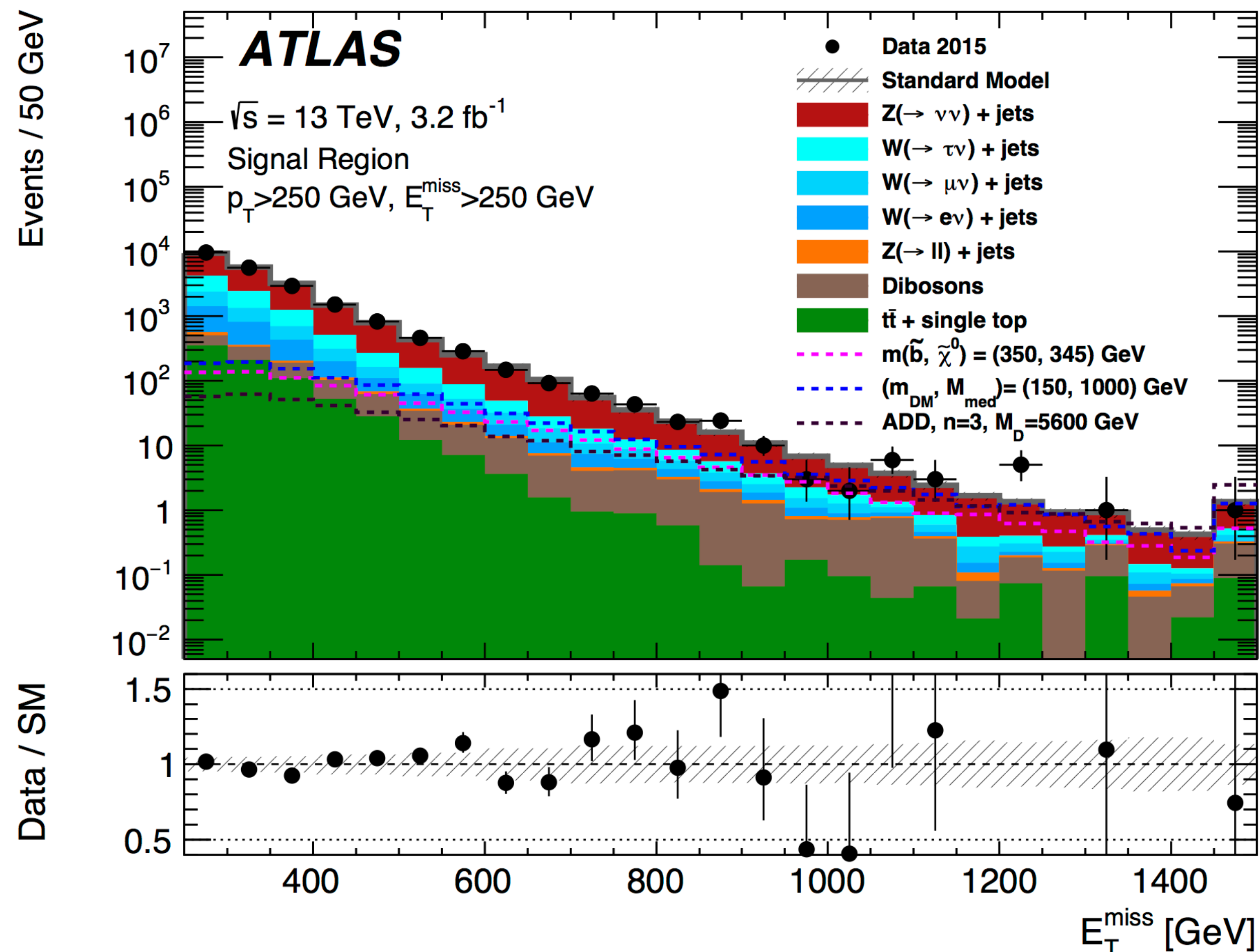
Energy<sub>xy</sub> Jets

# Reconstructing the final state



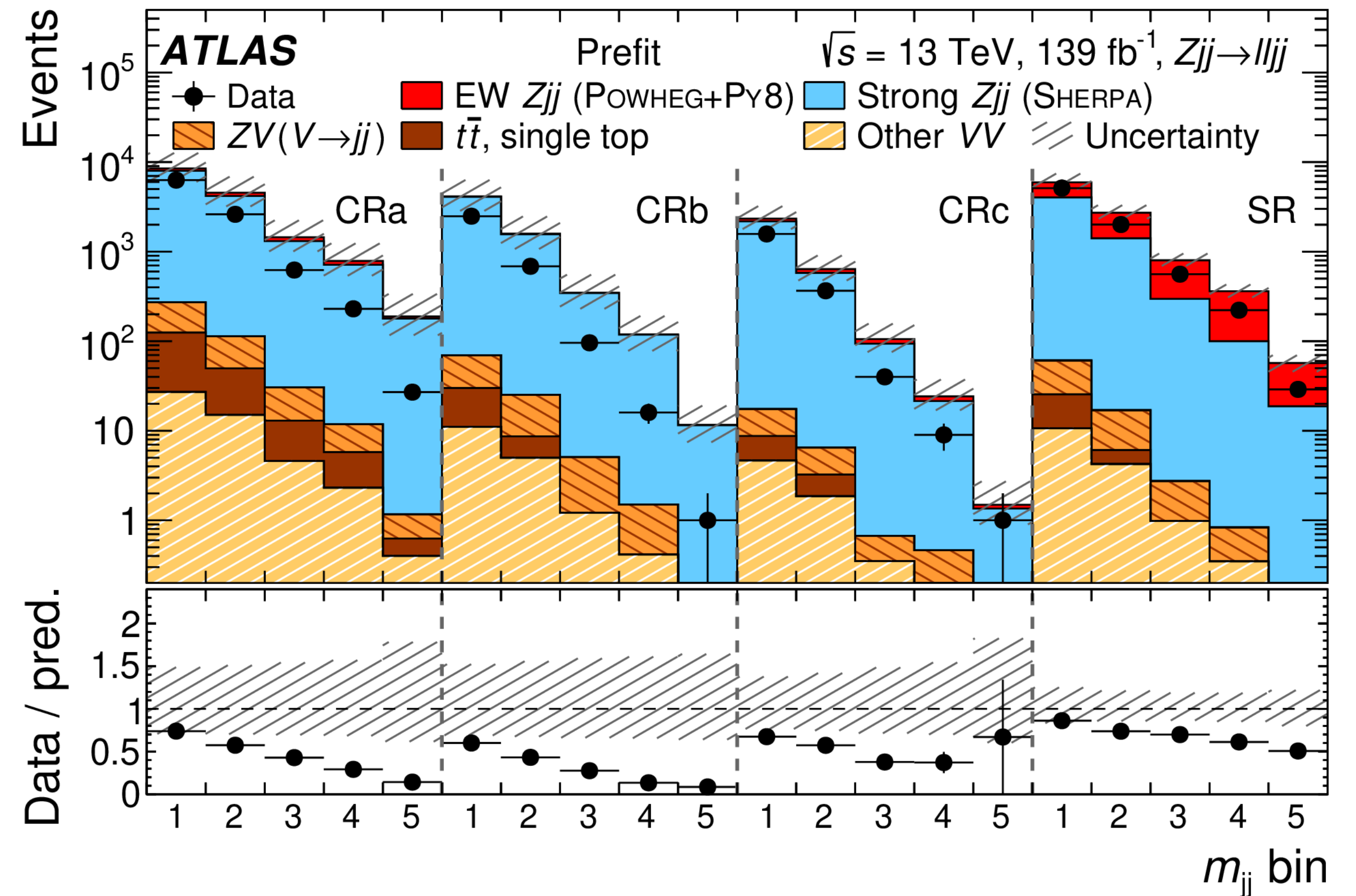
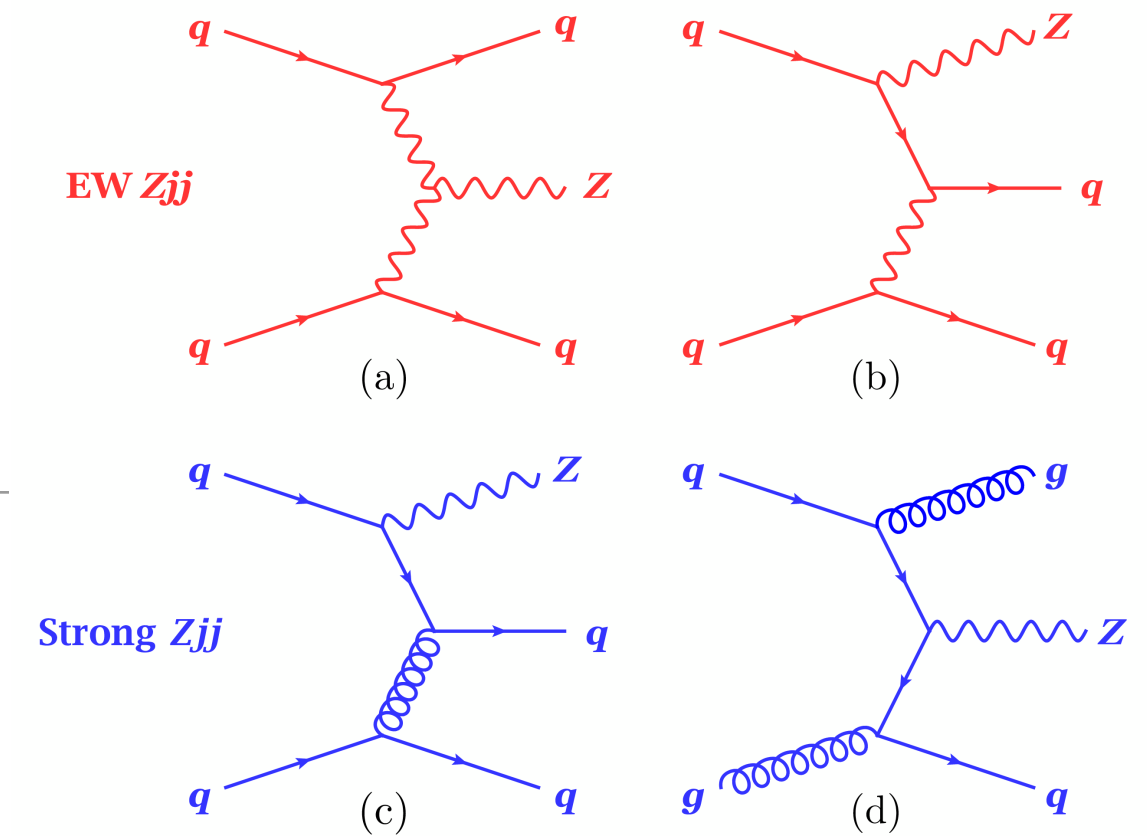
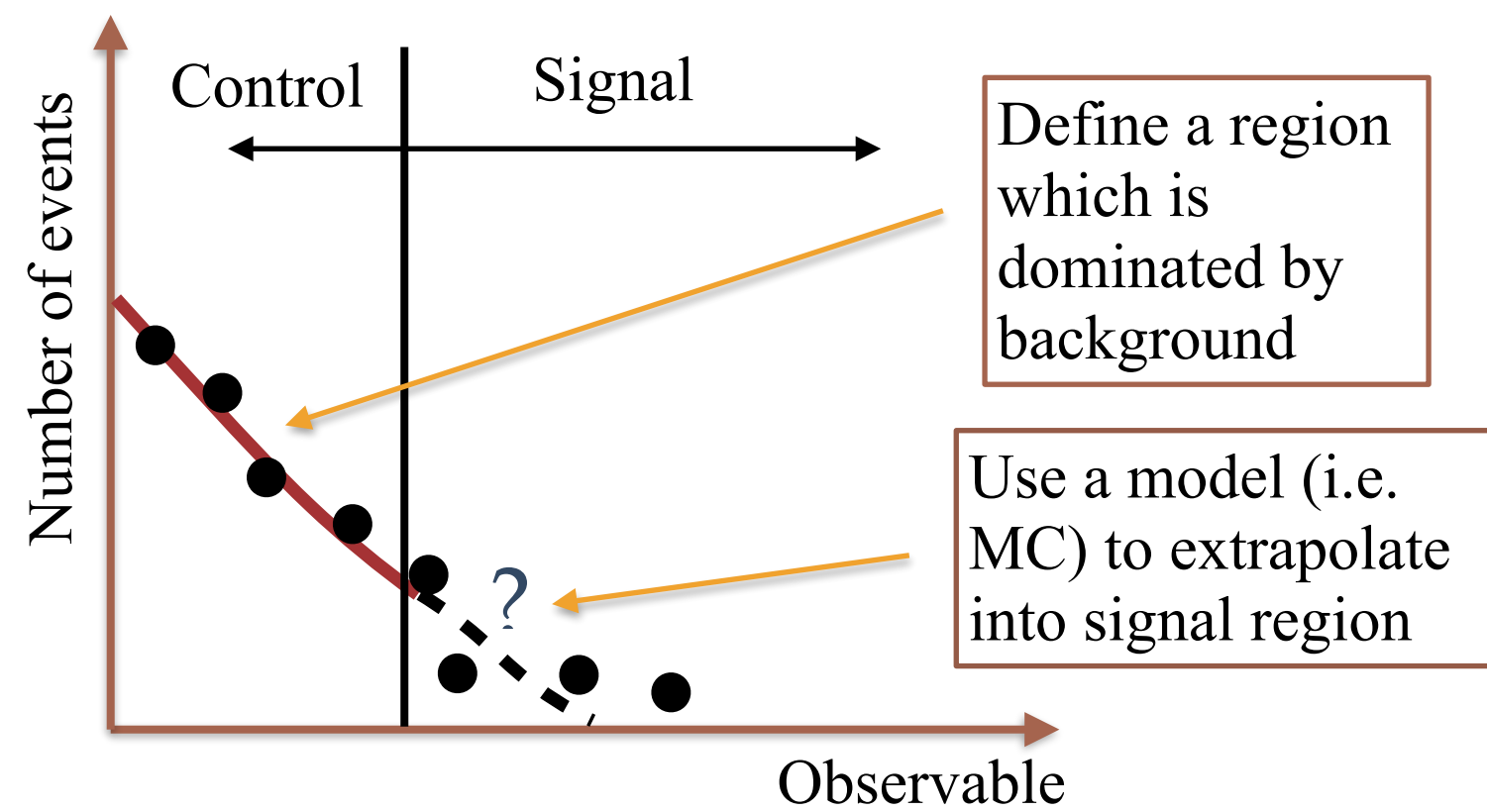
# Backgrounds

- How large is the signal compared to backgrounds with the same final state?
- Three main types of backgrounds
  - Backgrounds from the same final state (i.e.  $Z$  to neutrinos + jets)
  - Backgrounds where an object was outside the acceptance or not reconstructed (i.e.  $W$  decays to leptons)
  - Backgrounds due to misconstruction (i.e. multijet production)



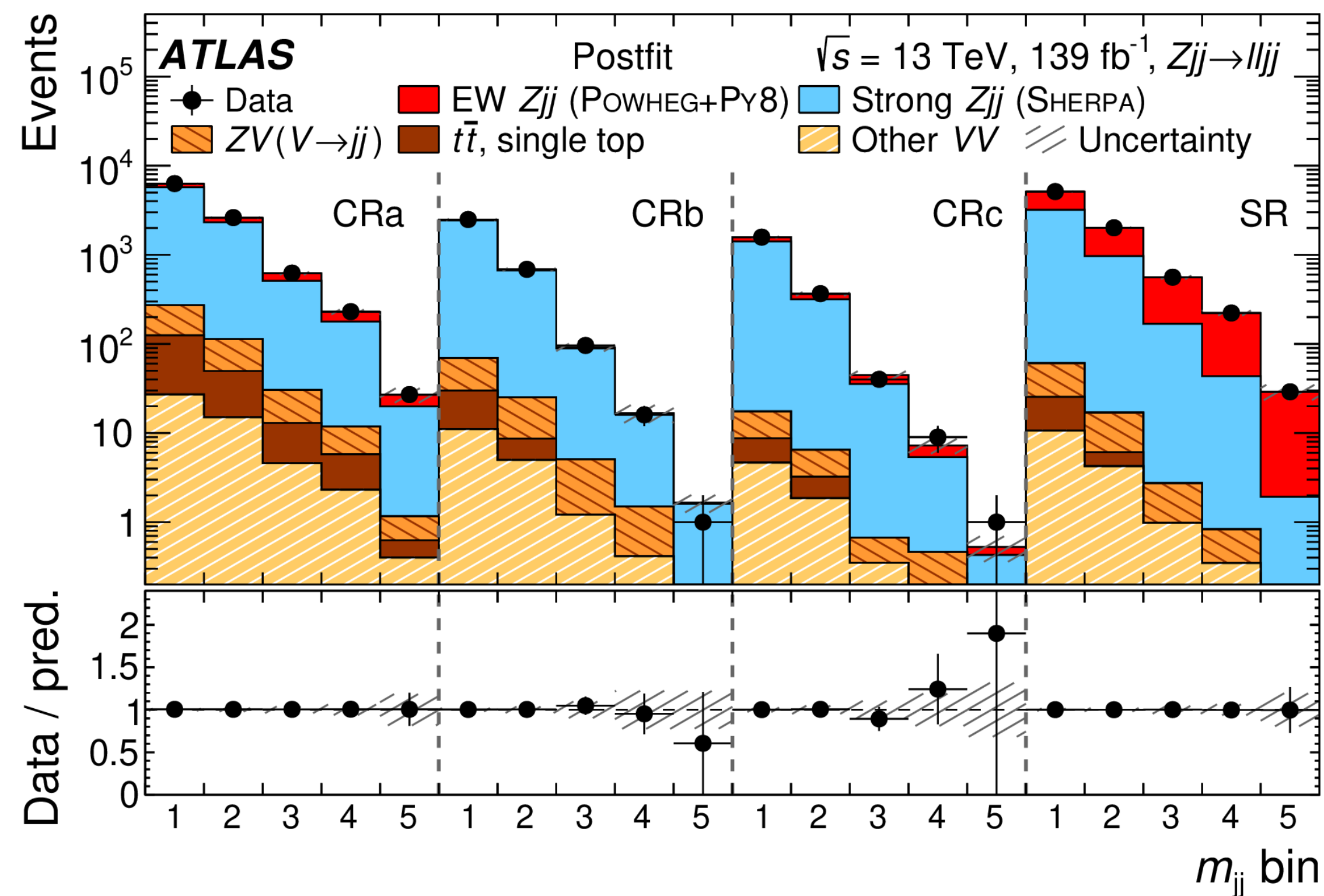
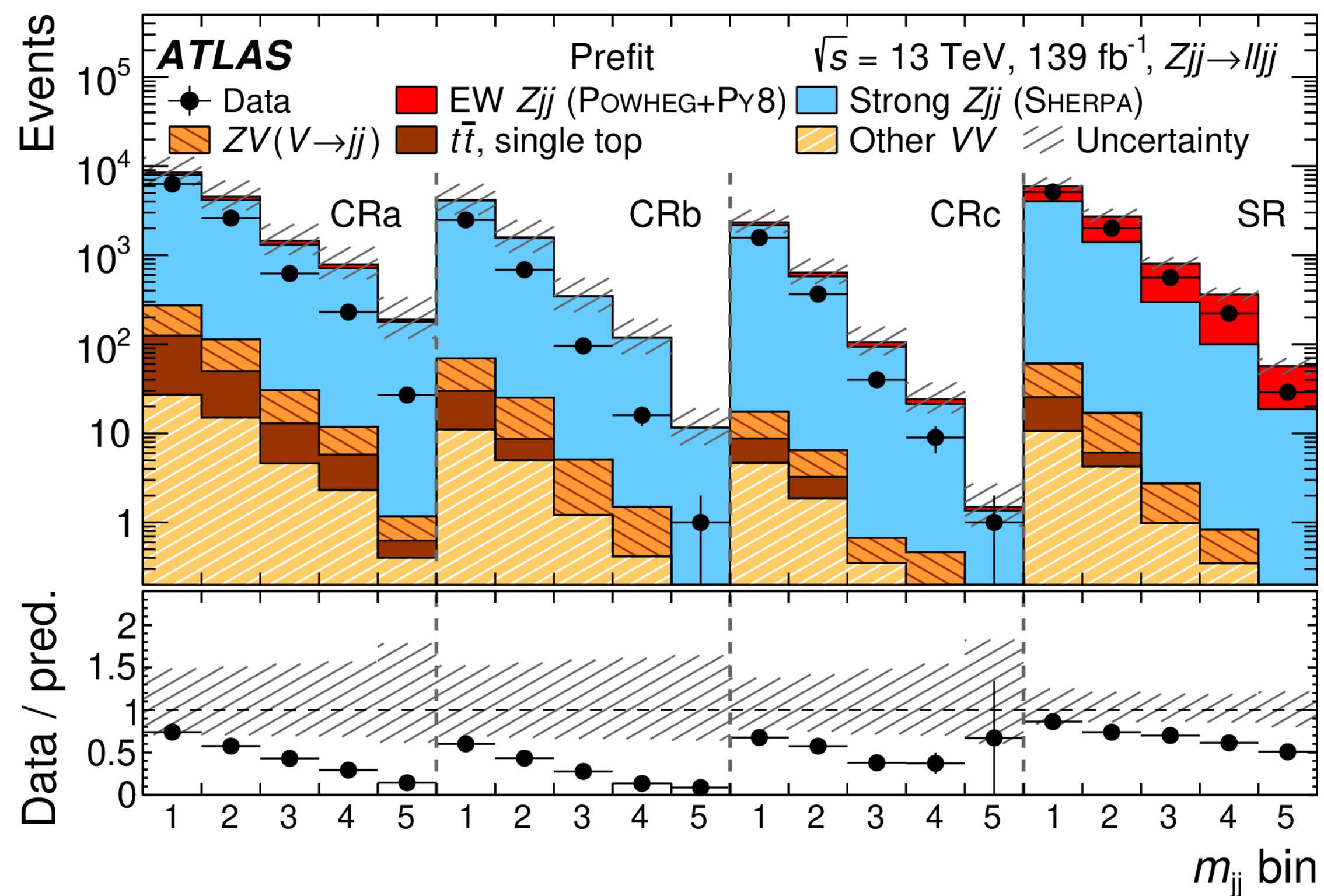
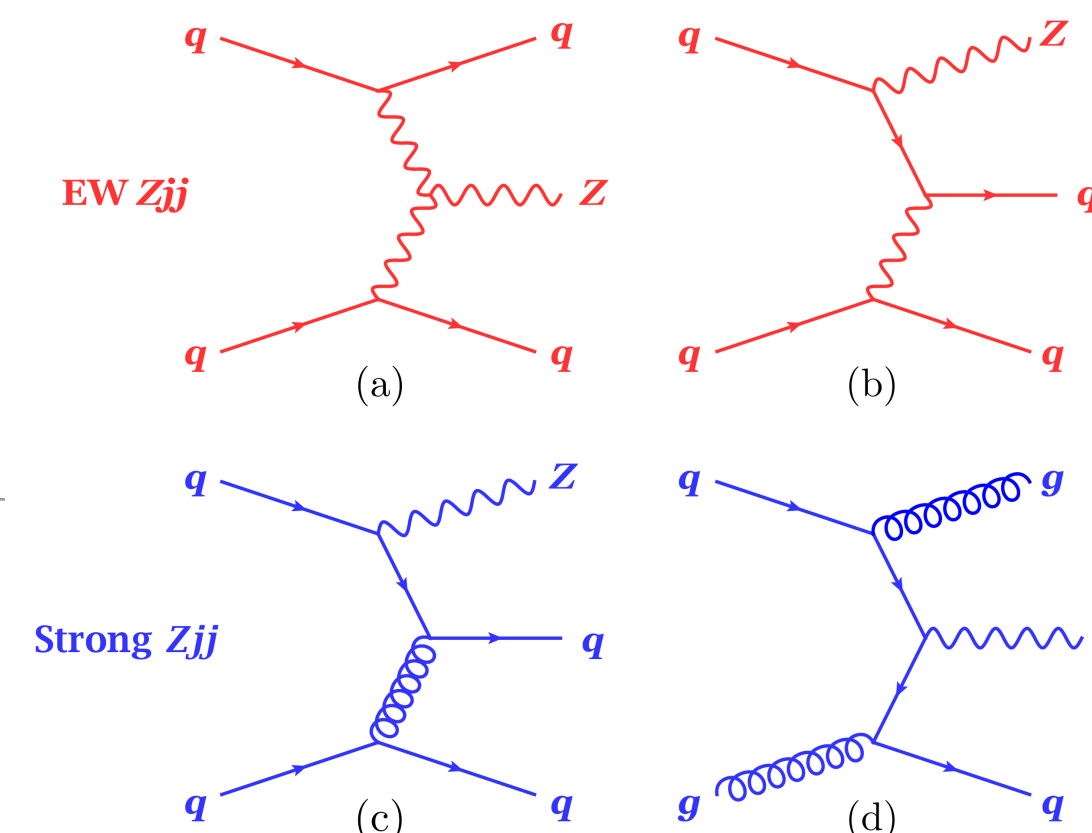
# Uncertainties

- What will be the dominate uncertainty? Can it be controlled or constrained?



# Uncertainties

- What will be the dominate uncertainty? Can it be controlled or constrained?

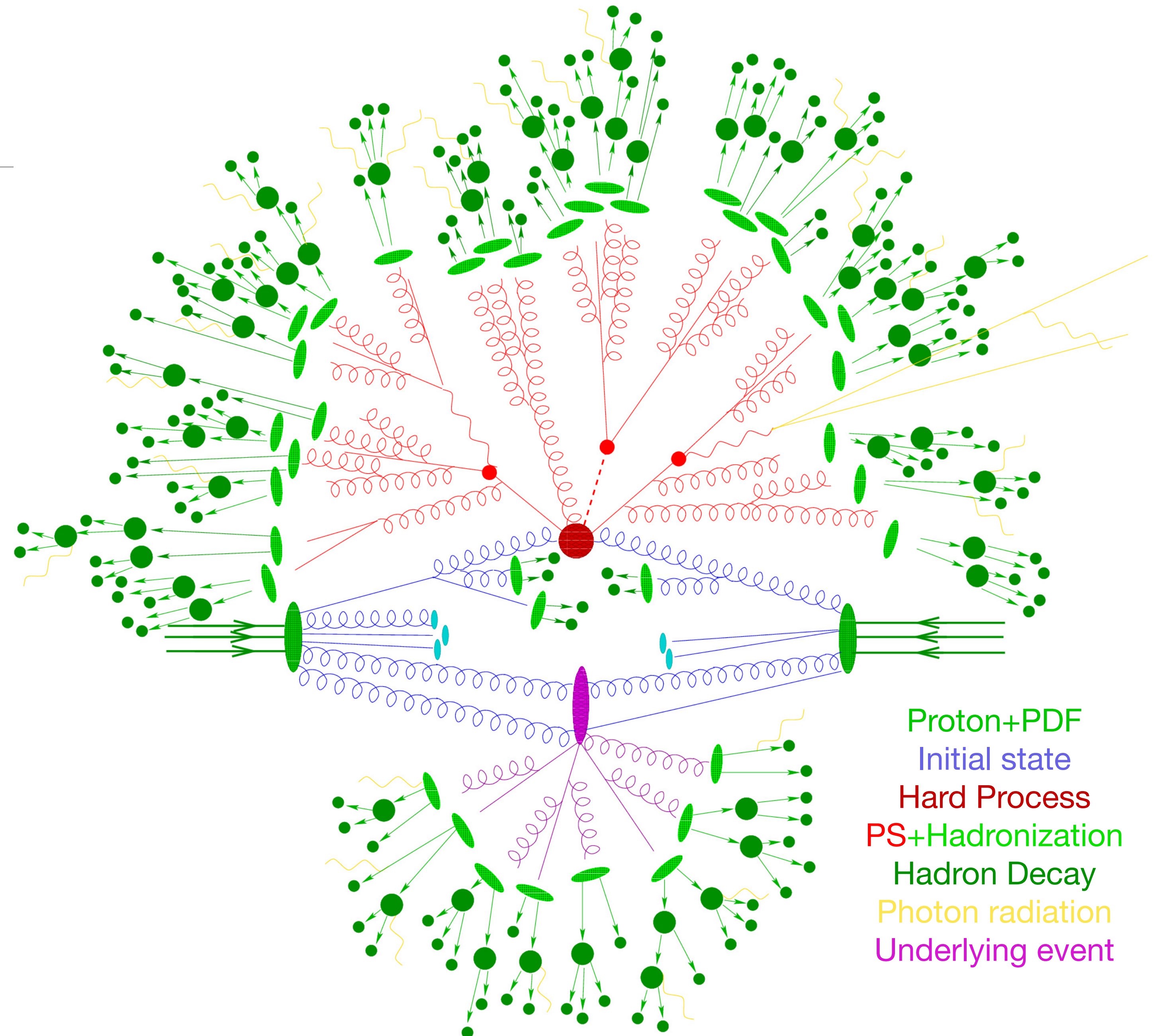




Defining the phase space

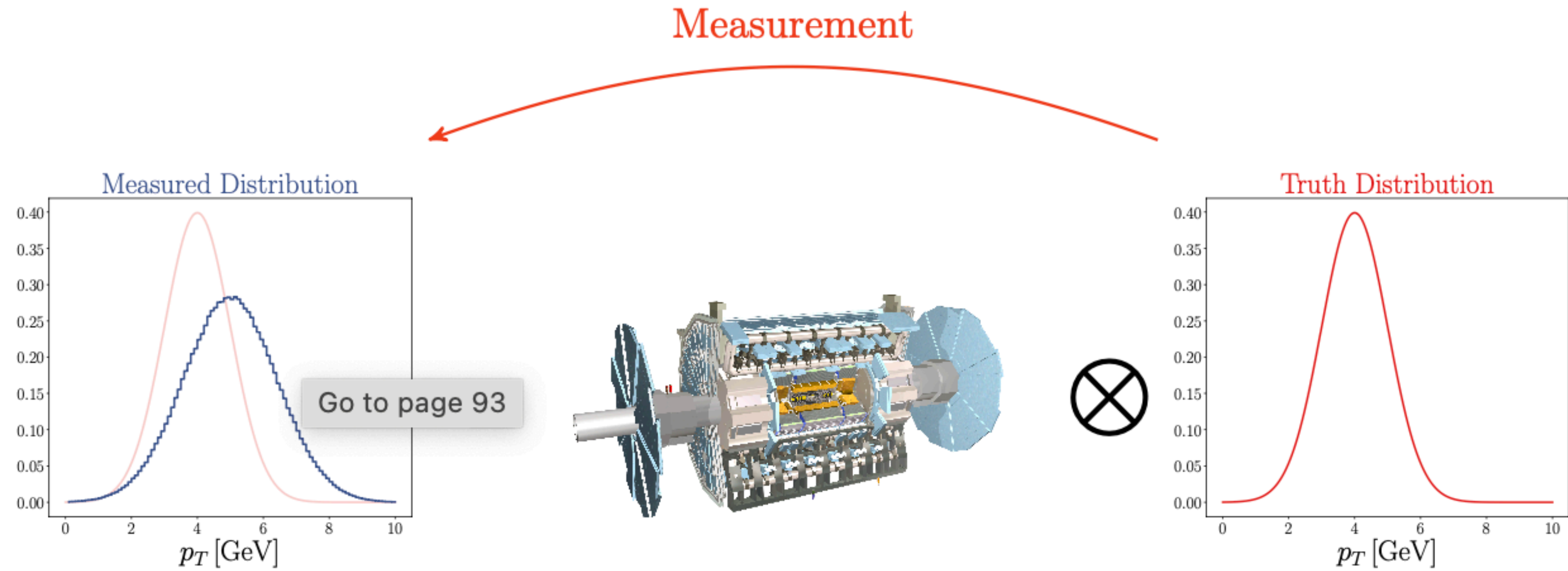
# Defining the phase space

- The detector has a certain kinematic region (i.e. it's not a  $4\pi$  detector and it can't measure  $p_T = 0$ )
- Fiducial phase space is a criteria applied to final state 'truth' particles



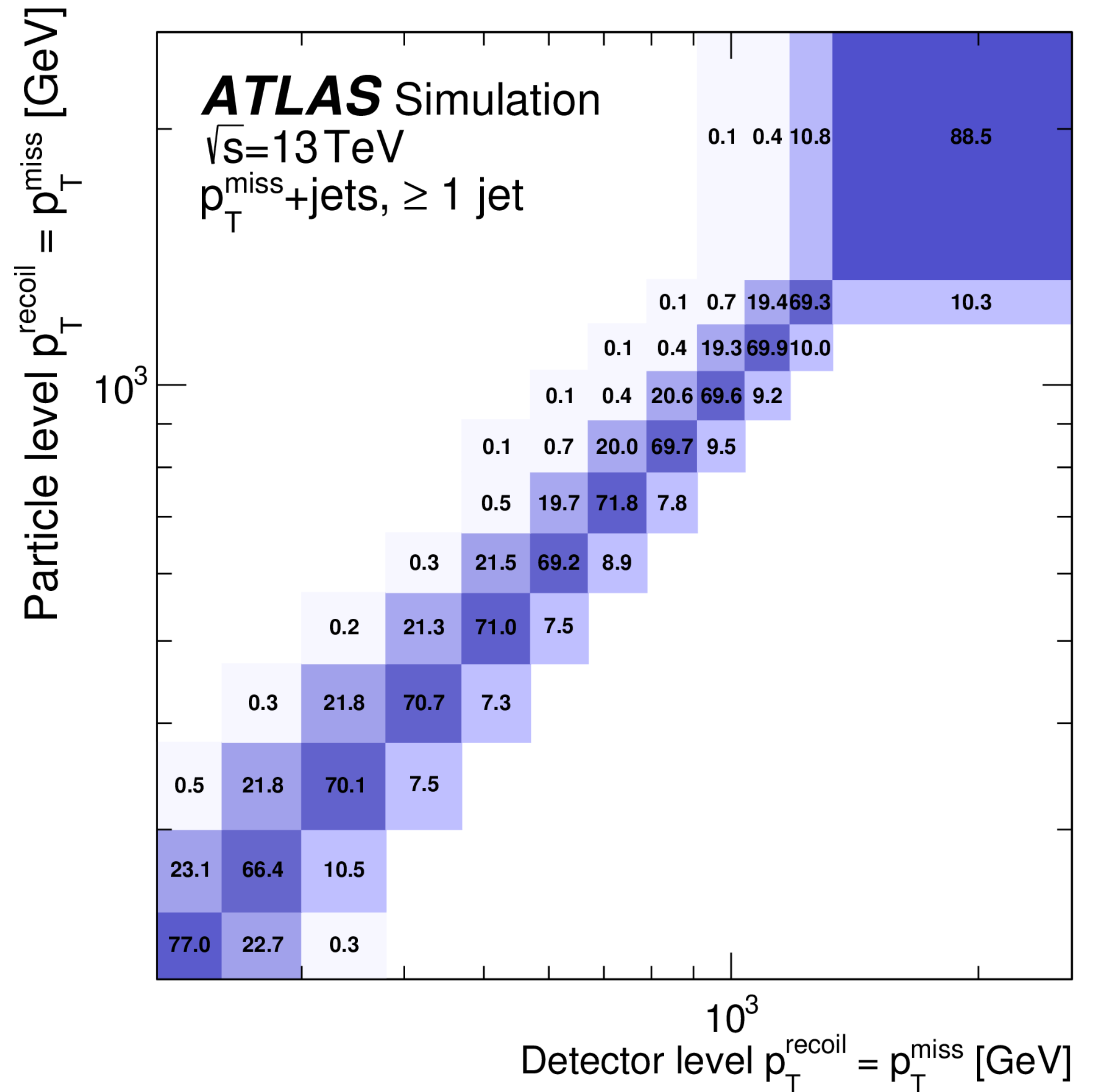
Unfolding - Correcting for detector effects

# Unfolding - the basics



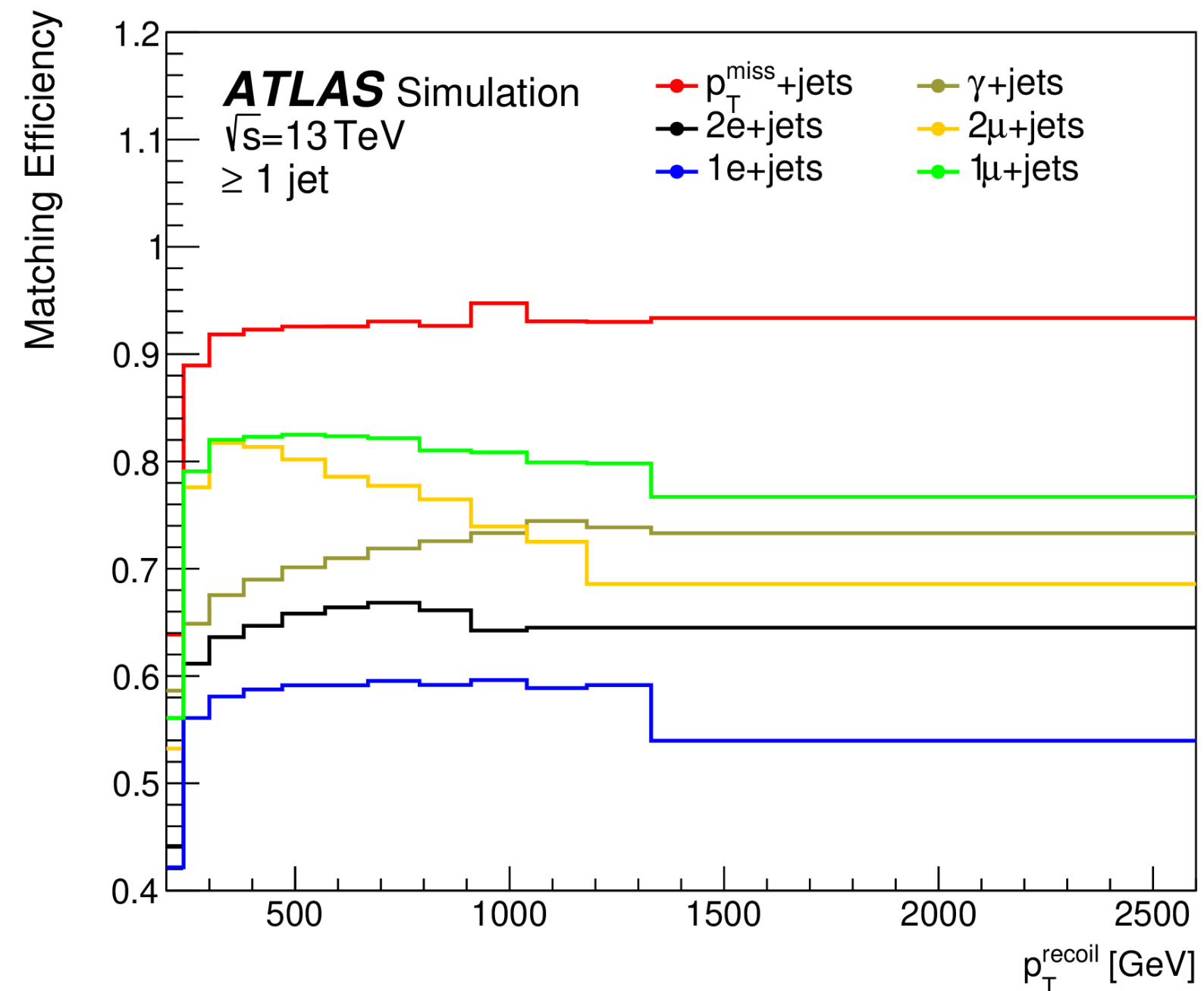
# Unfolding - the basics

- Migration matrix - maps truth to reconstructed object
- Fewer migrations is better



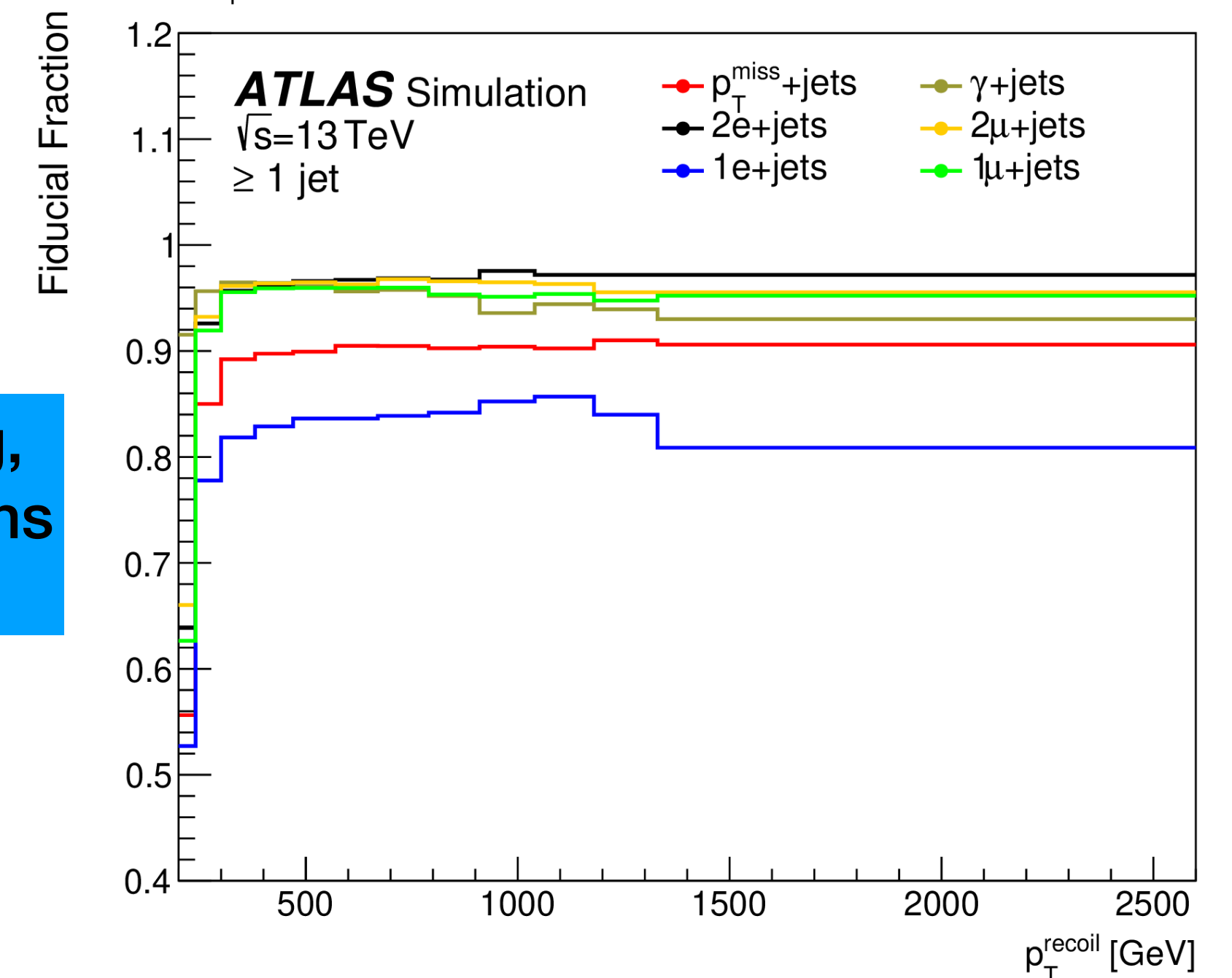
# Unfolding - the basics

- Additional corrections for
  - Events in truth phase that are not at detector level (matching efficiency)
  - Events at detector-level but in truth the phase space (shows as the fiducial fraction, fraction where an event is in both)
- Closer to 1.0 is better

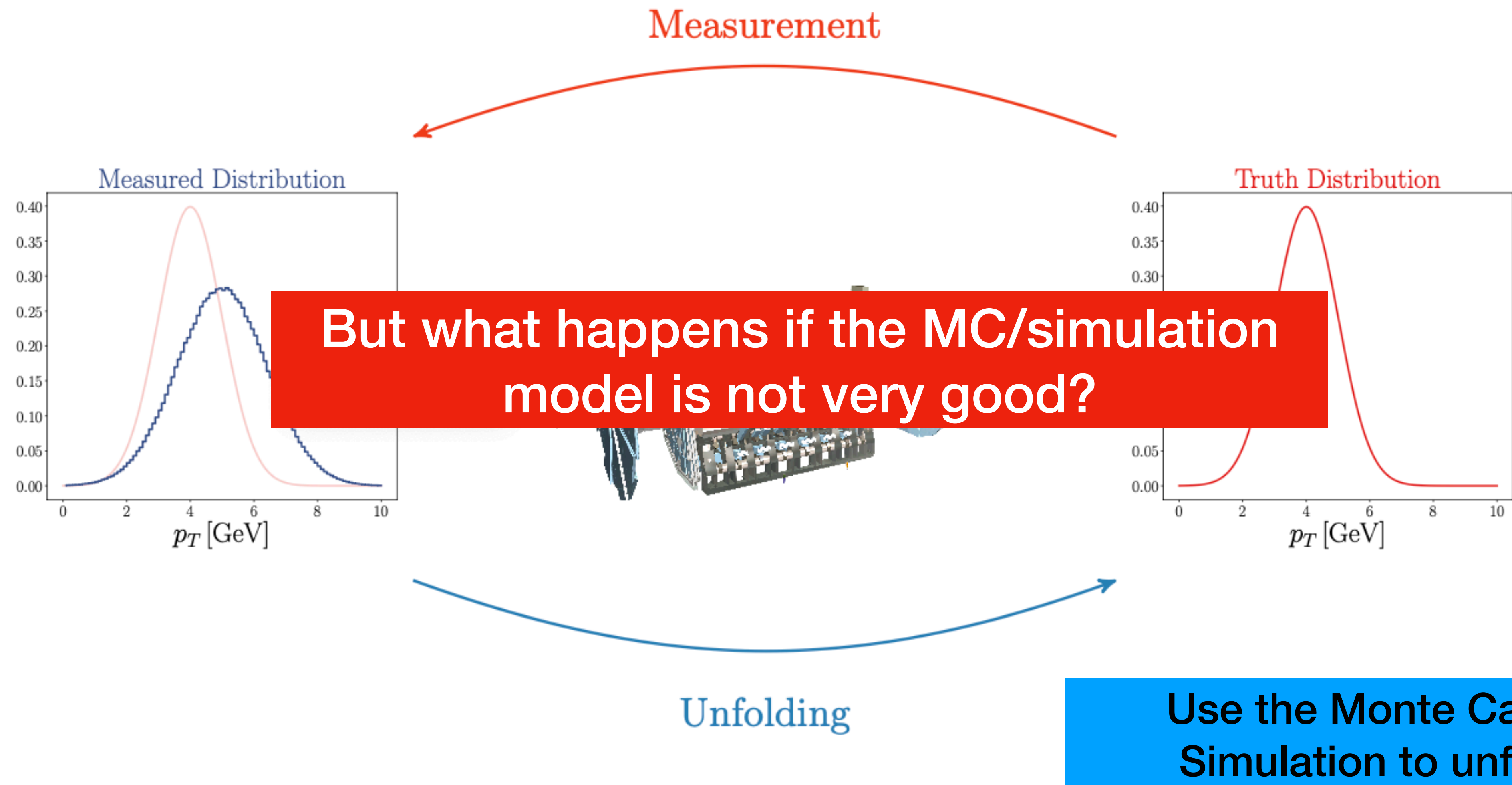


A poor choice of phase space definition can affect this

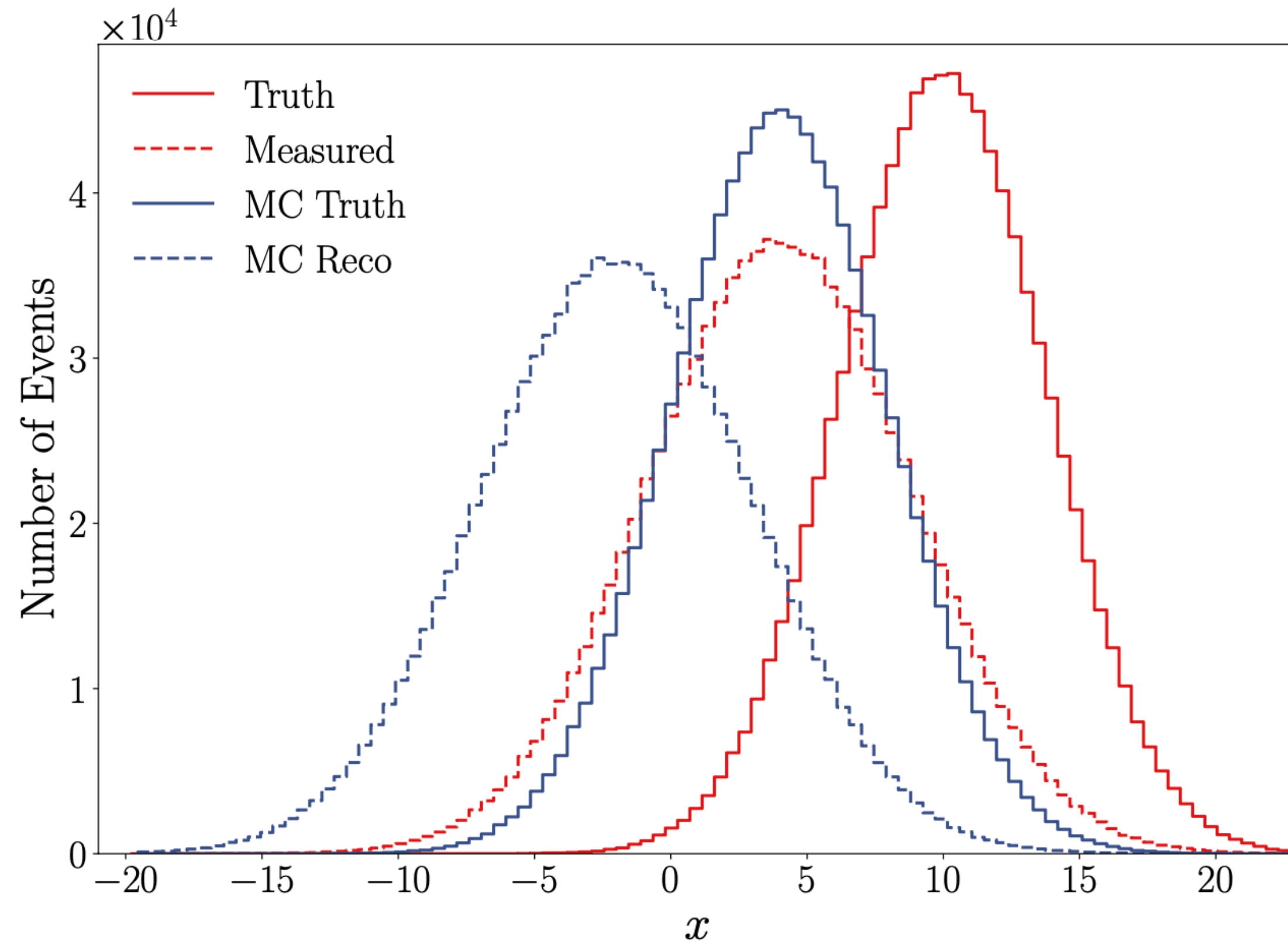
Poor choice of binning, poor detector resolutions can affect this



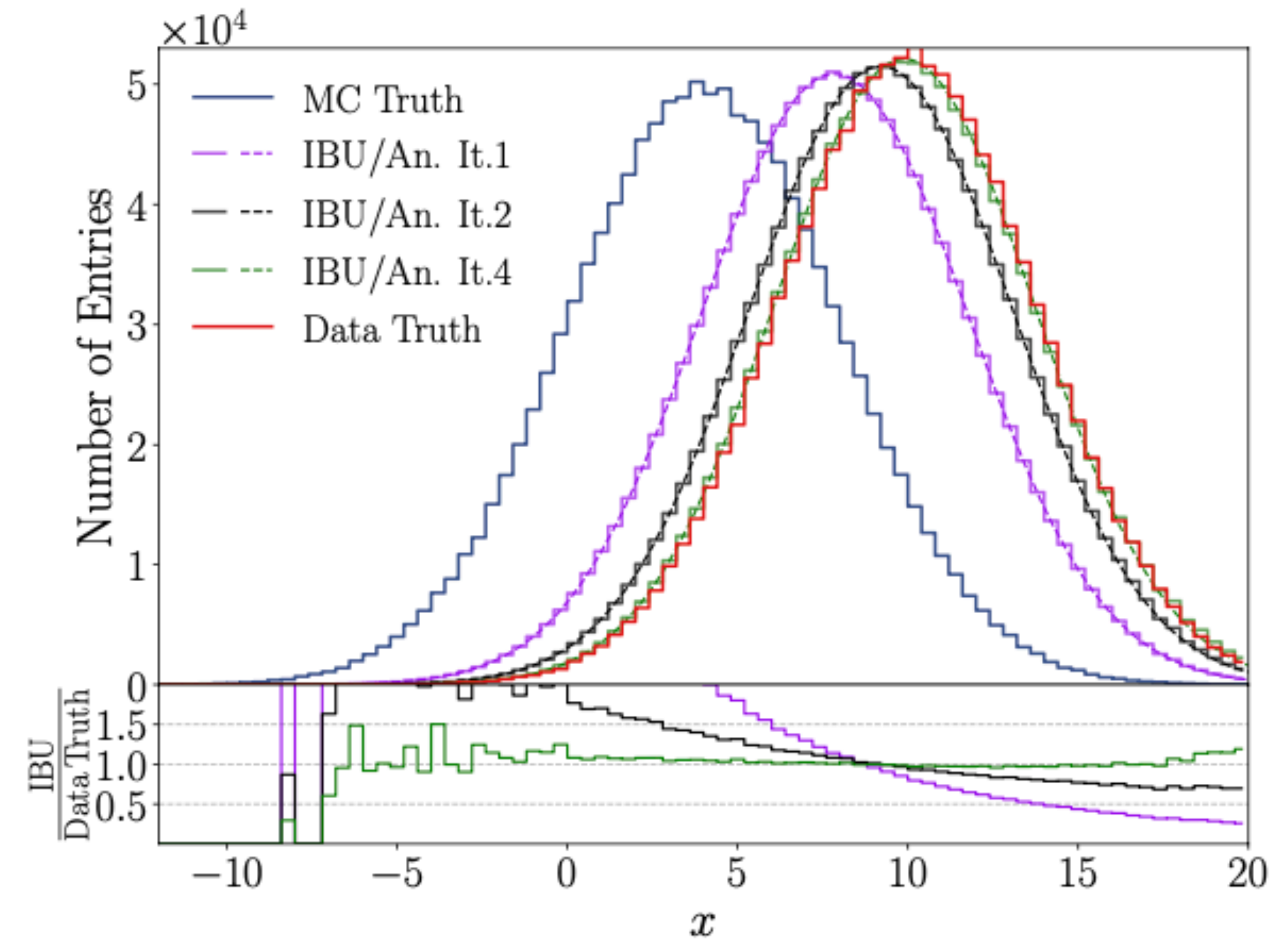
# Unfolding - the basics



# A poor MC/simulation model example



**A toy example of a poor model**

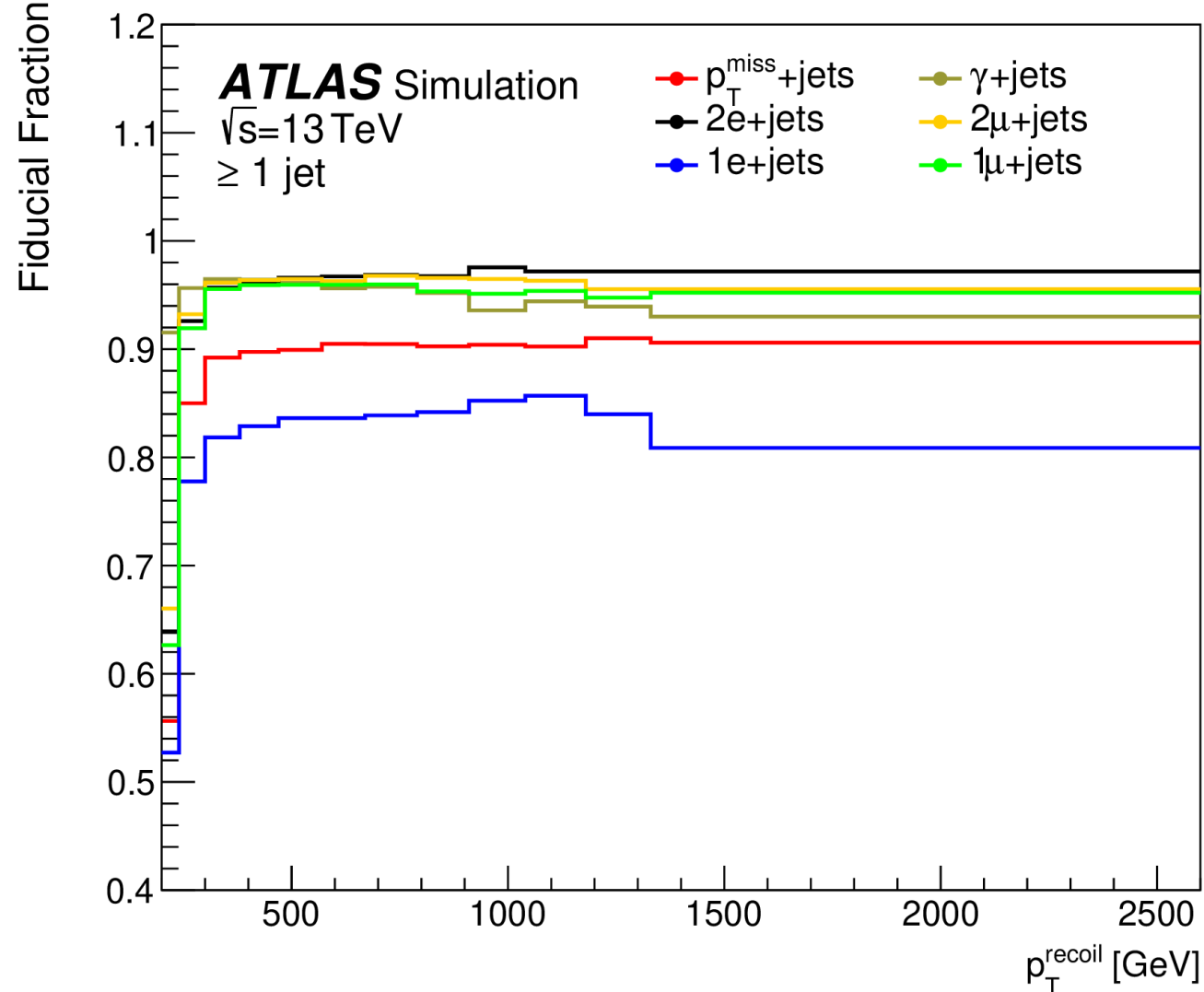
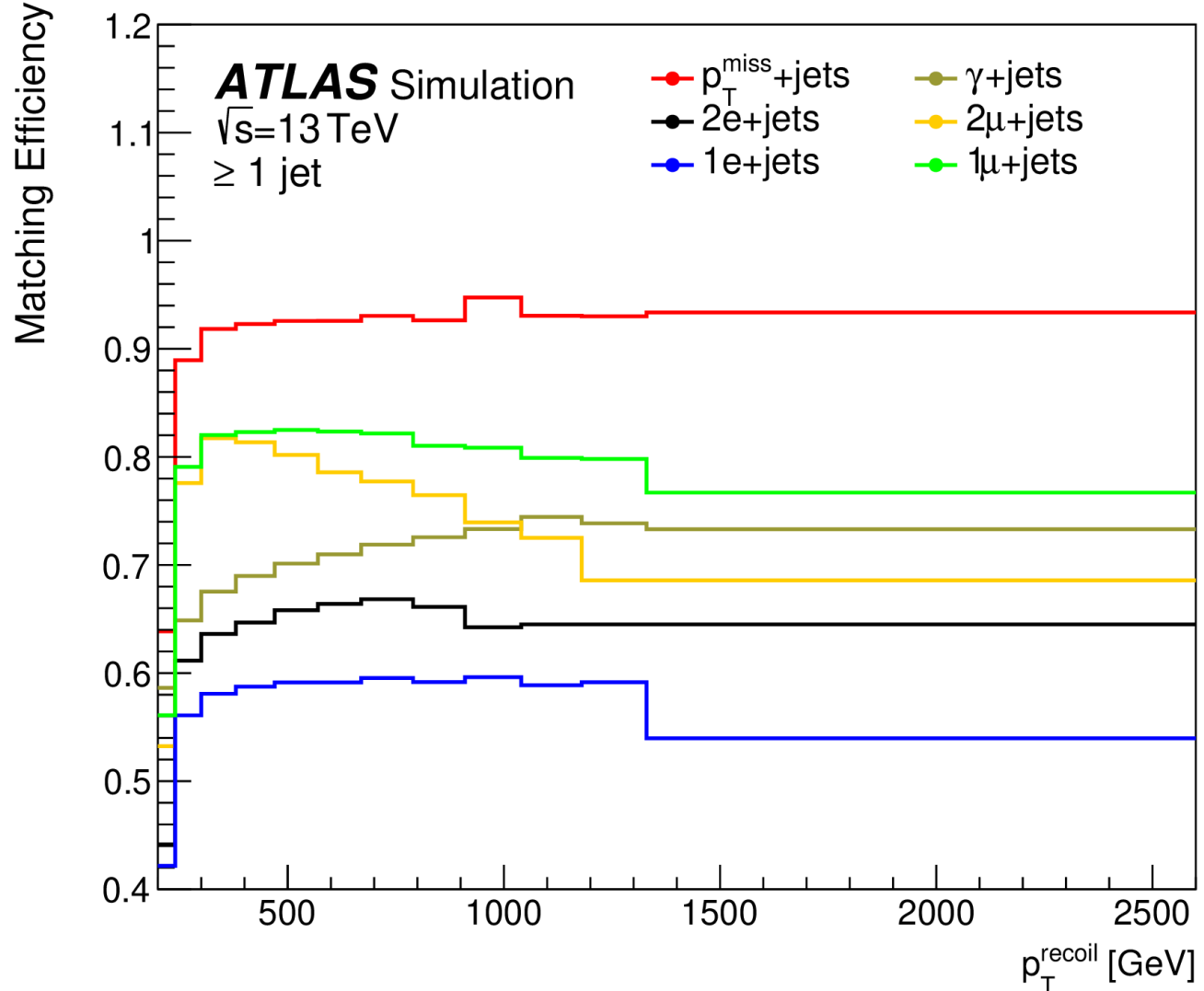
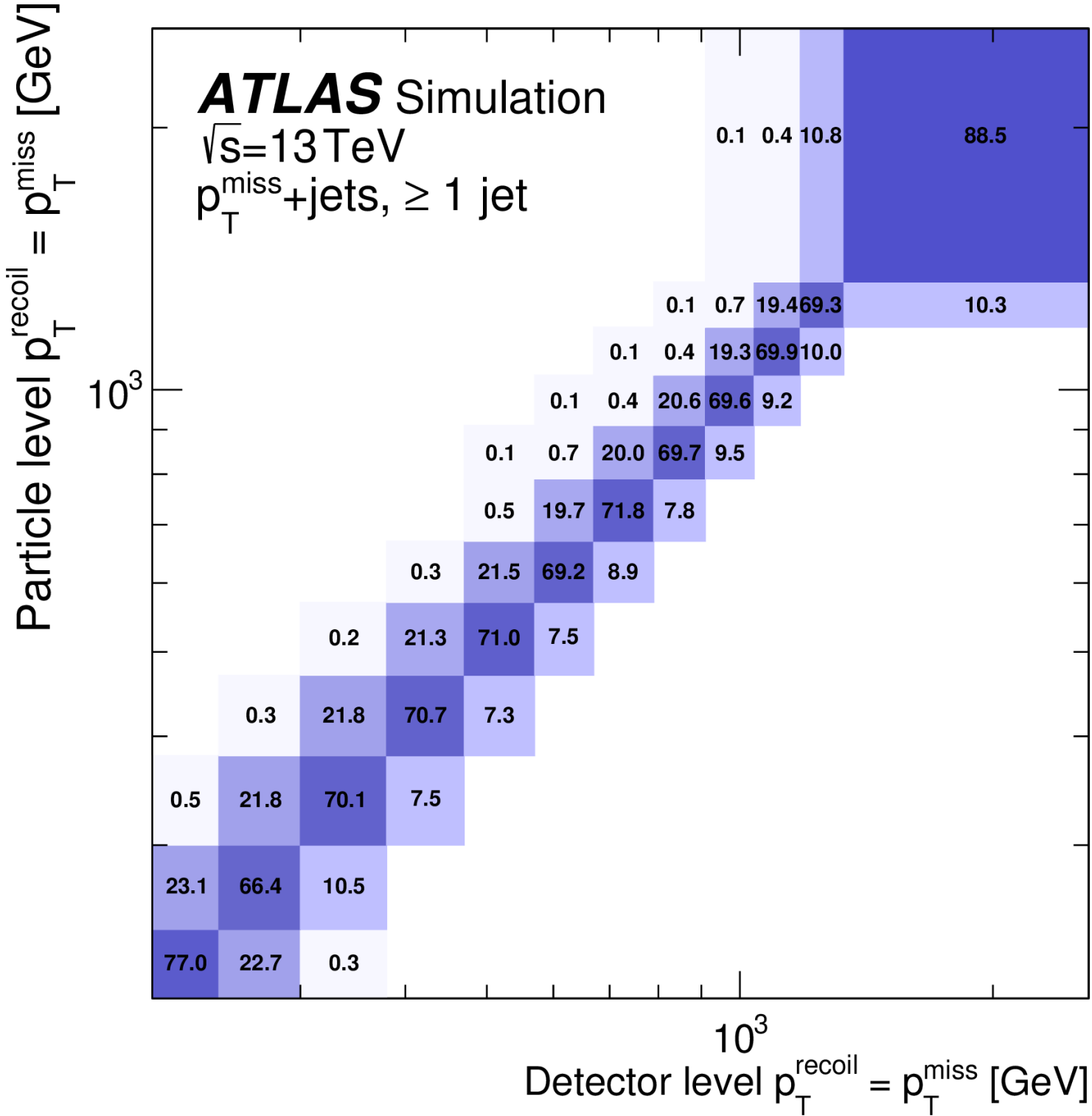


**To avoid model dependences on the measure, do an iterative approach**



# A word of caution

- We use iterative unfolding to reduce dependencies on the underlying model



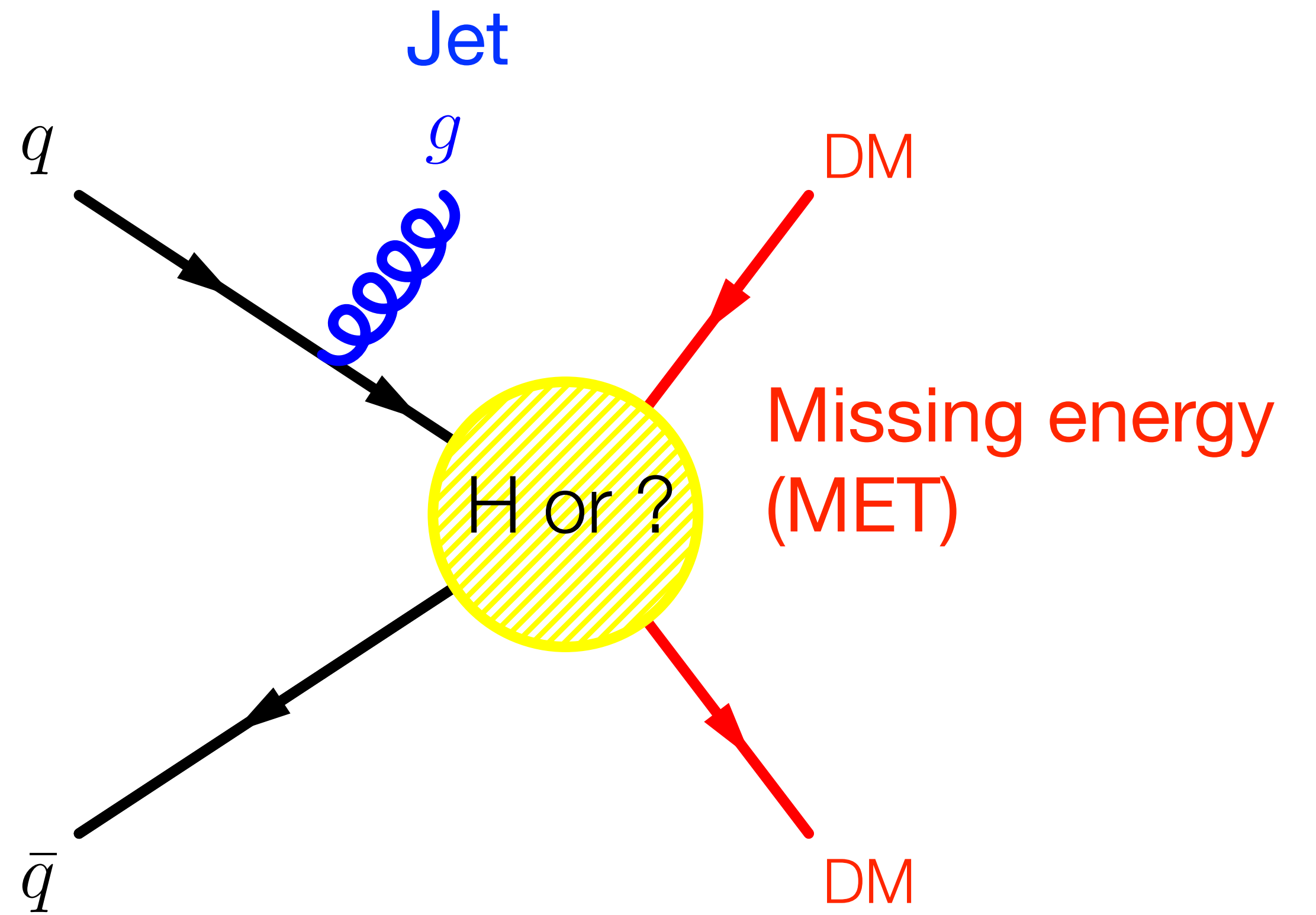
Iterations largely improve the migration matrix

These are more susceptible to model dependencies and therefore the optimal phase space and measurement definition for the question at hand is critical

# A real example with all the pieces in place

---

My question - Does DM couple to the Higgs?



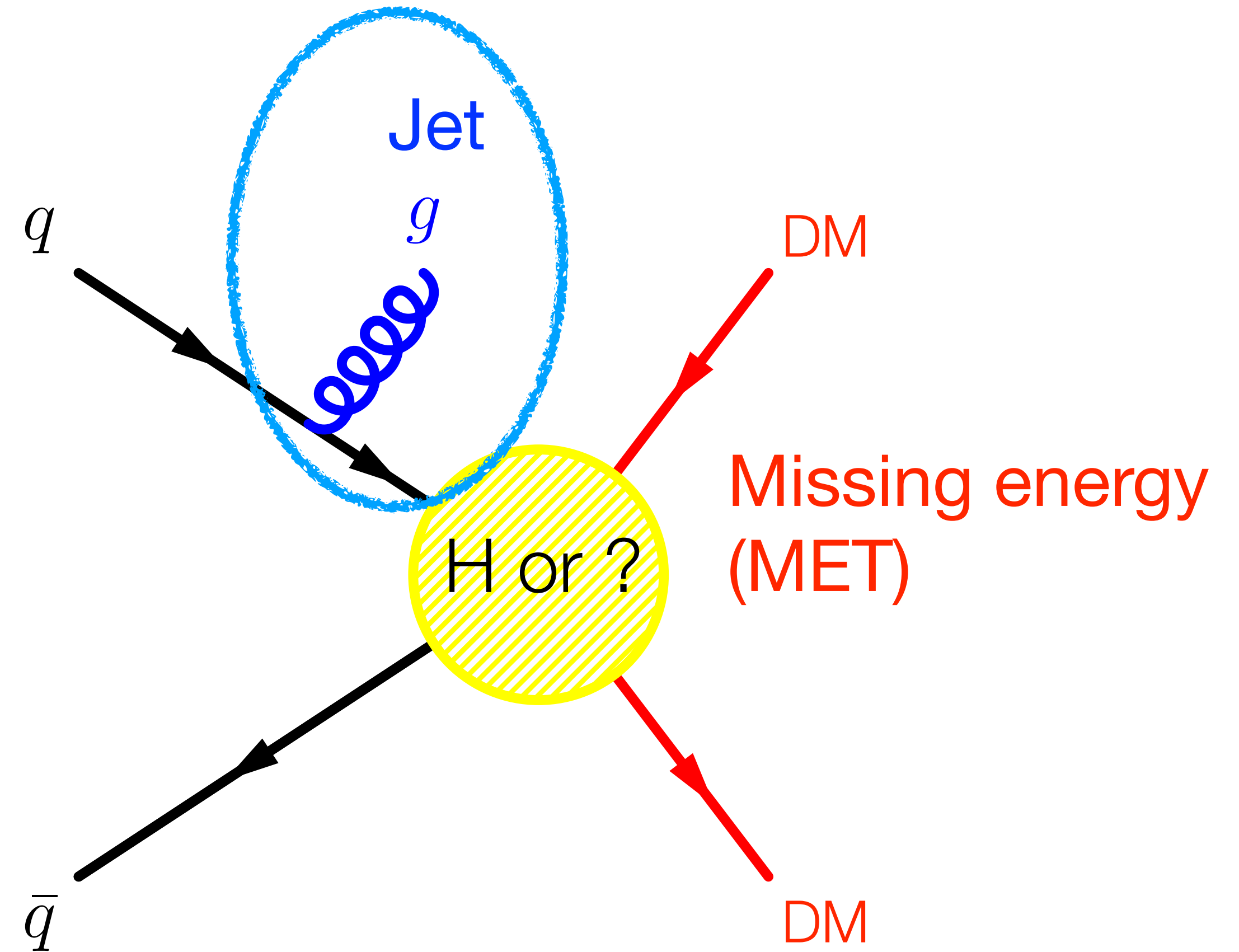
# A real example with all the pieces in place

My question - Does DM couple to the Higgs?

My measurement -  
**the process will be rare —**  
**Focus on jets (high cross section)**

Backgrounds will be large — Will measure them via control regions

Jets are imprecise — Will reduce the uncertainties via a ratio



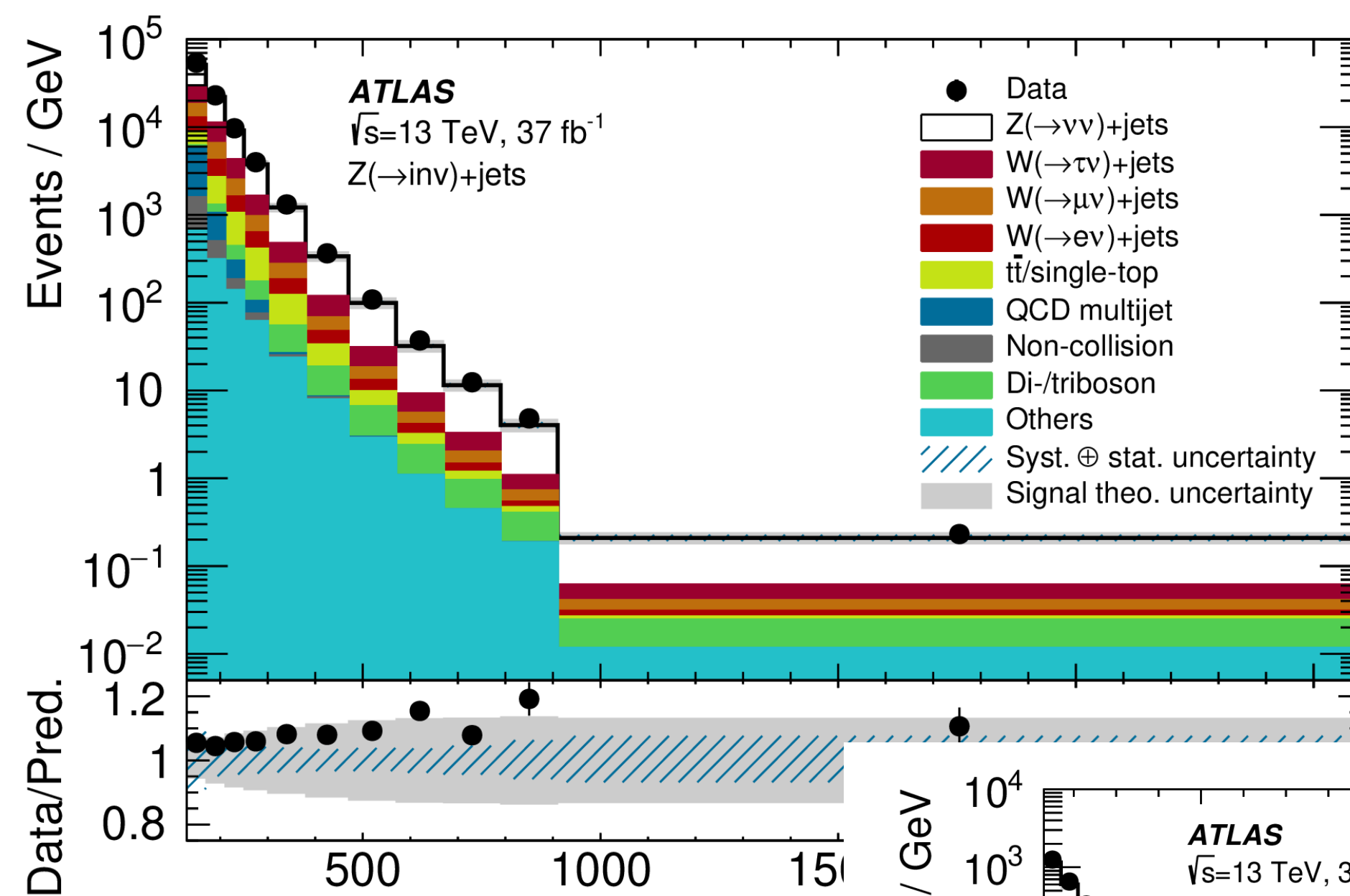
# A real example with all the pieces in place

My question - Does DM couple to the Higgs?

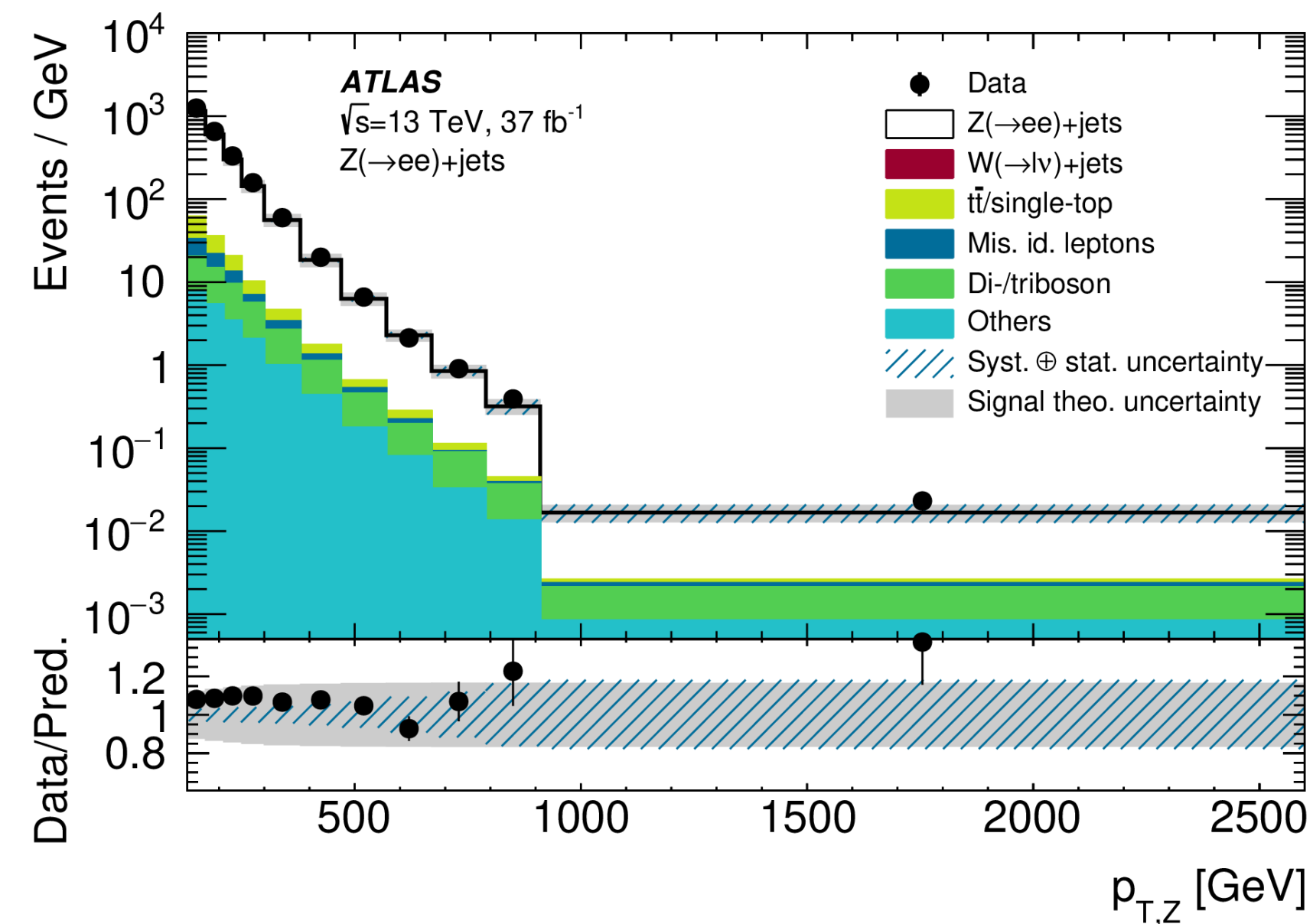
My measurement - the process will be rare — Focus on jets (high cross section)

Backgrounds will be large — Will measure them via control regions

Jets are imprecise — Will reduce the uncertainties via a ratio



Background enhanced (control region), rely on theory to extrapolate to my signal region



What I want to measure

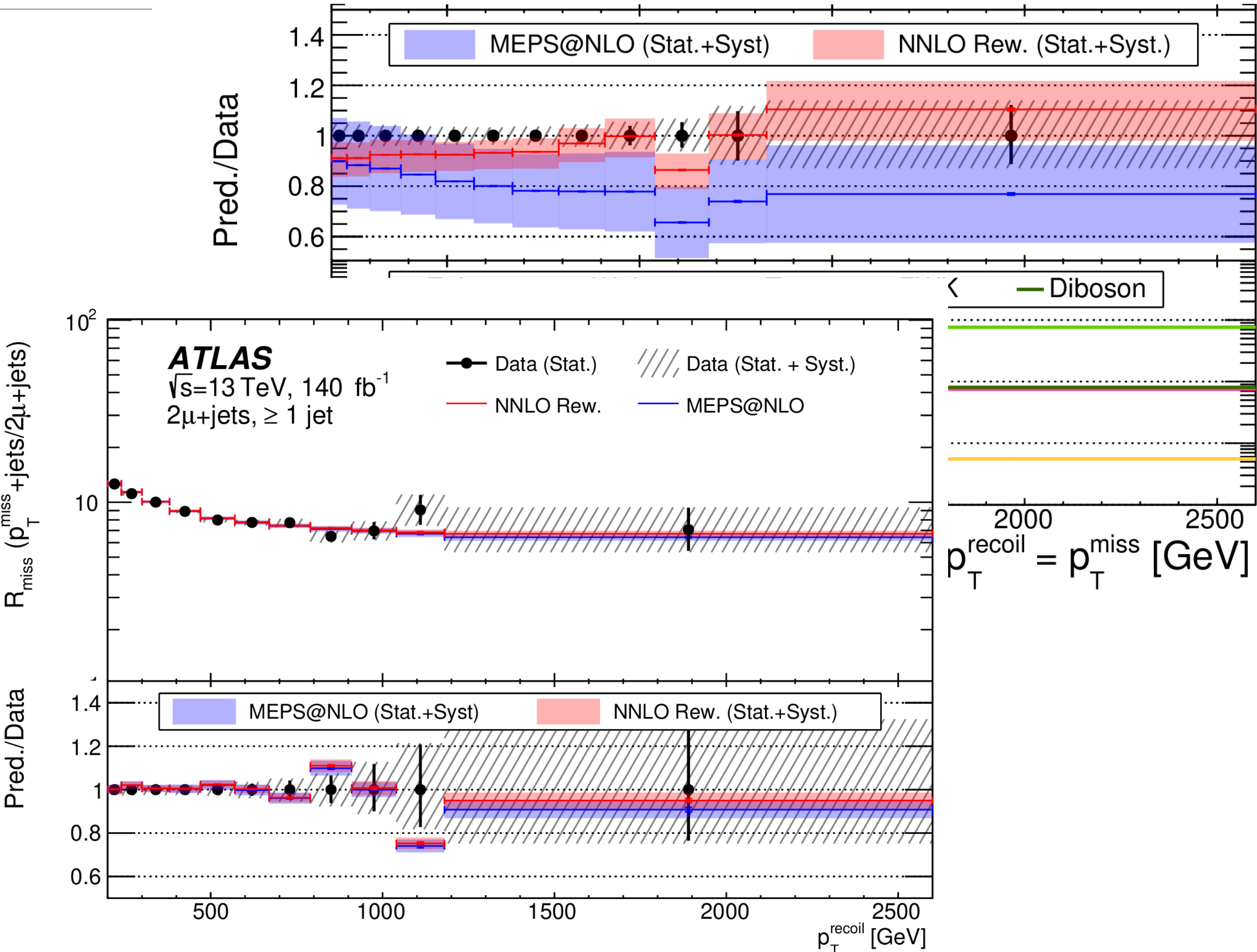
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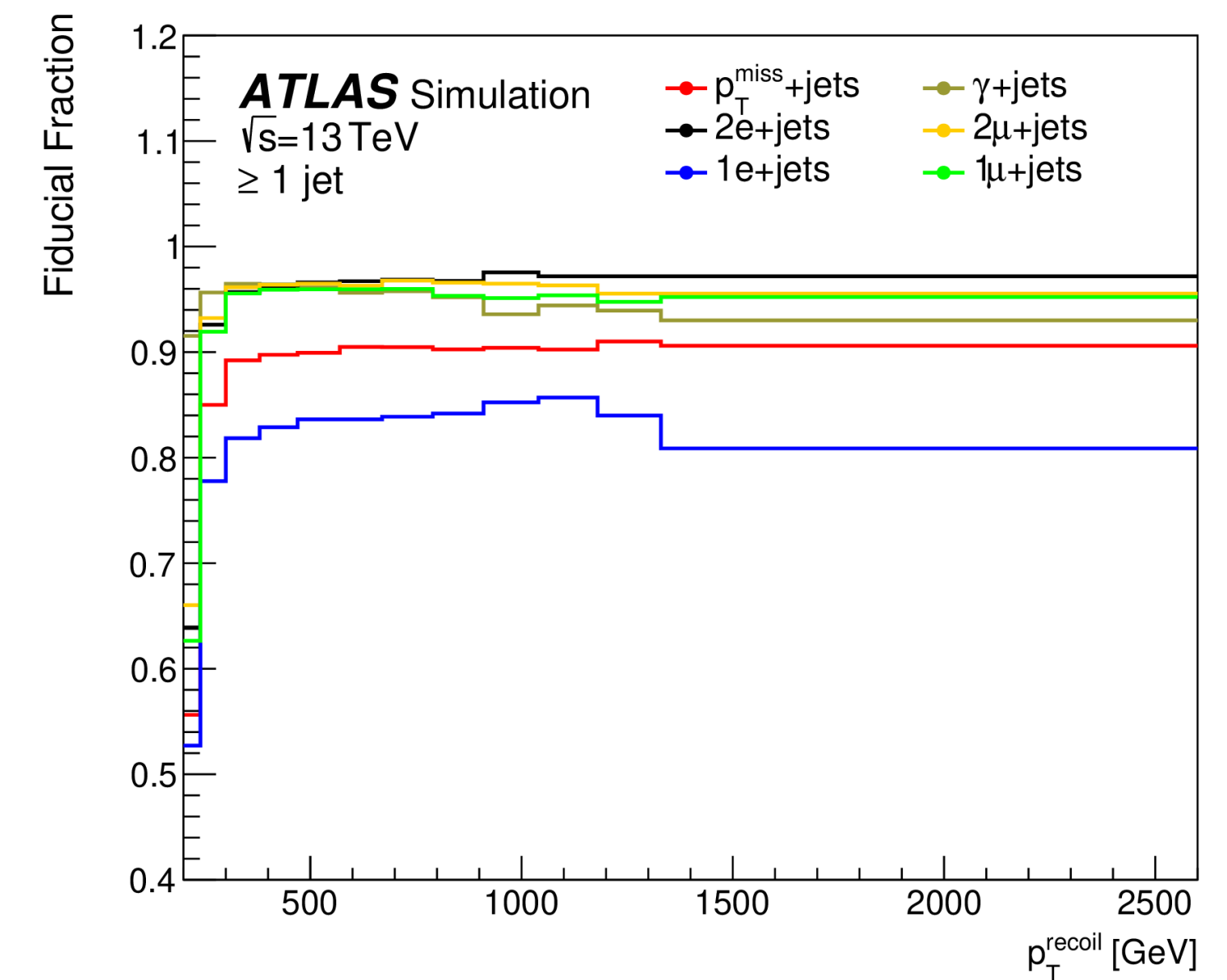
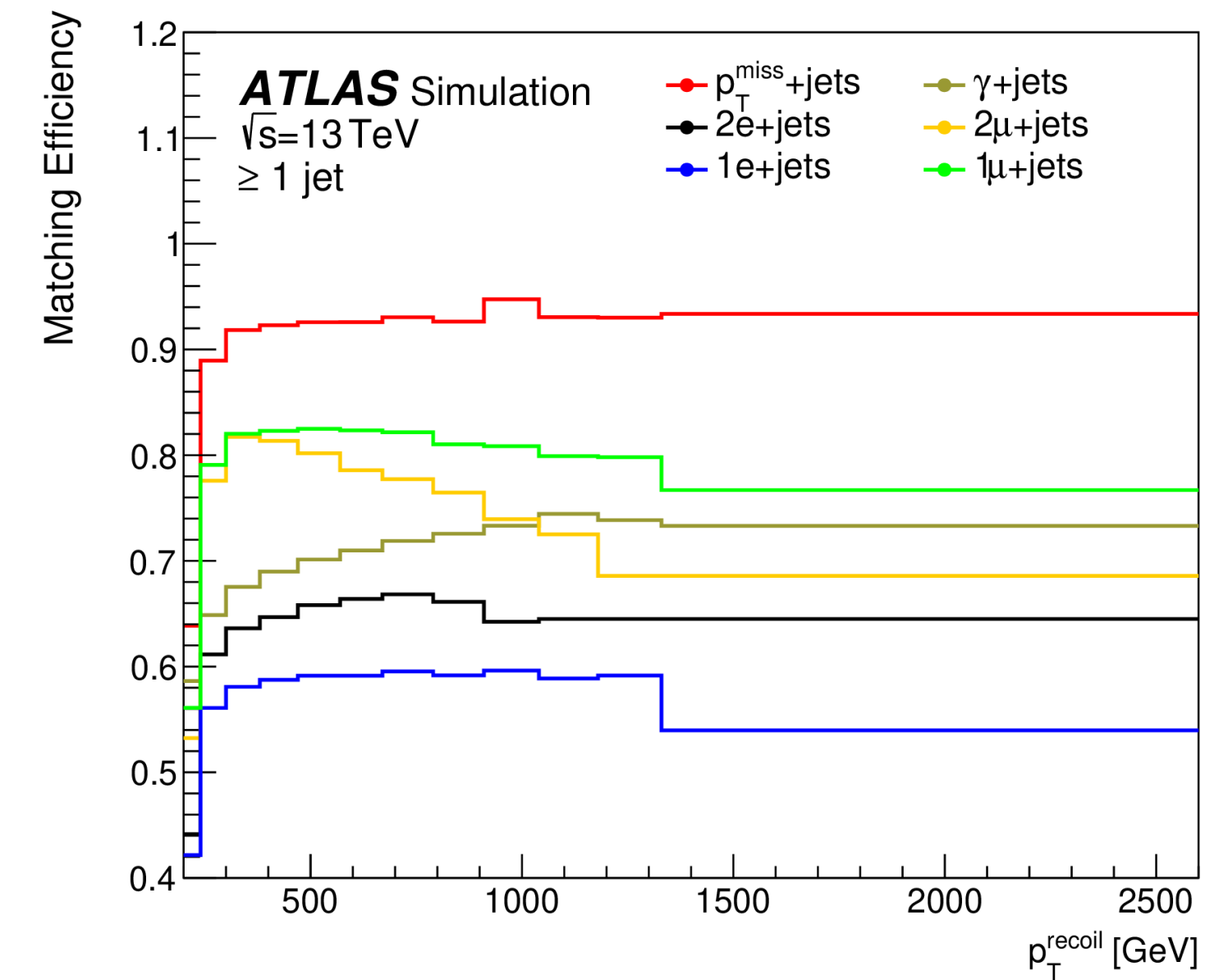
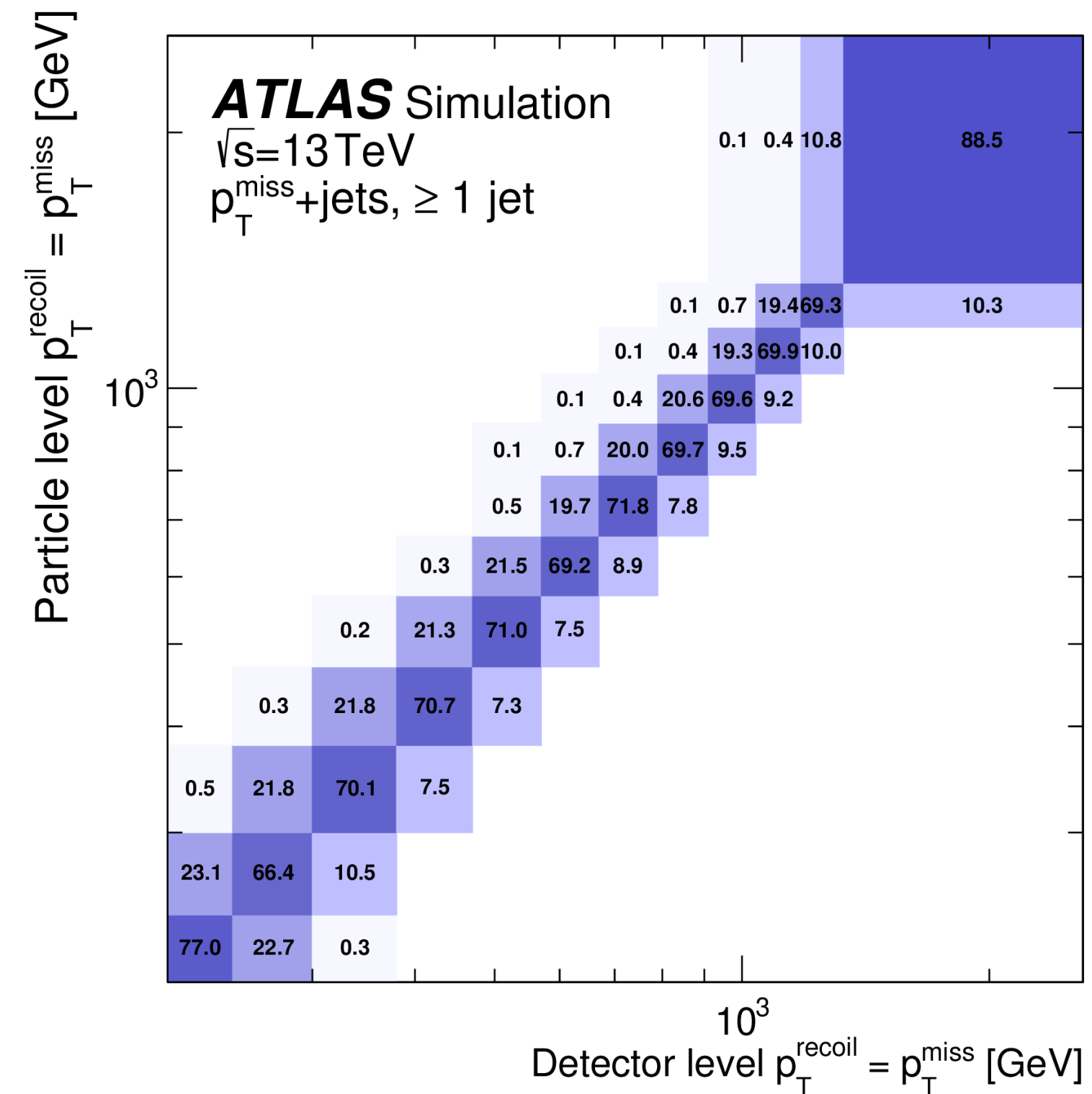
Backgrounds will be large — Will measure them via control regions

**Jets are imprecise — Will reduce the uncertainties via a ratio**



# A real example with all the pieces in place

My phase space

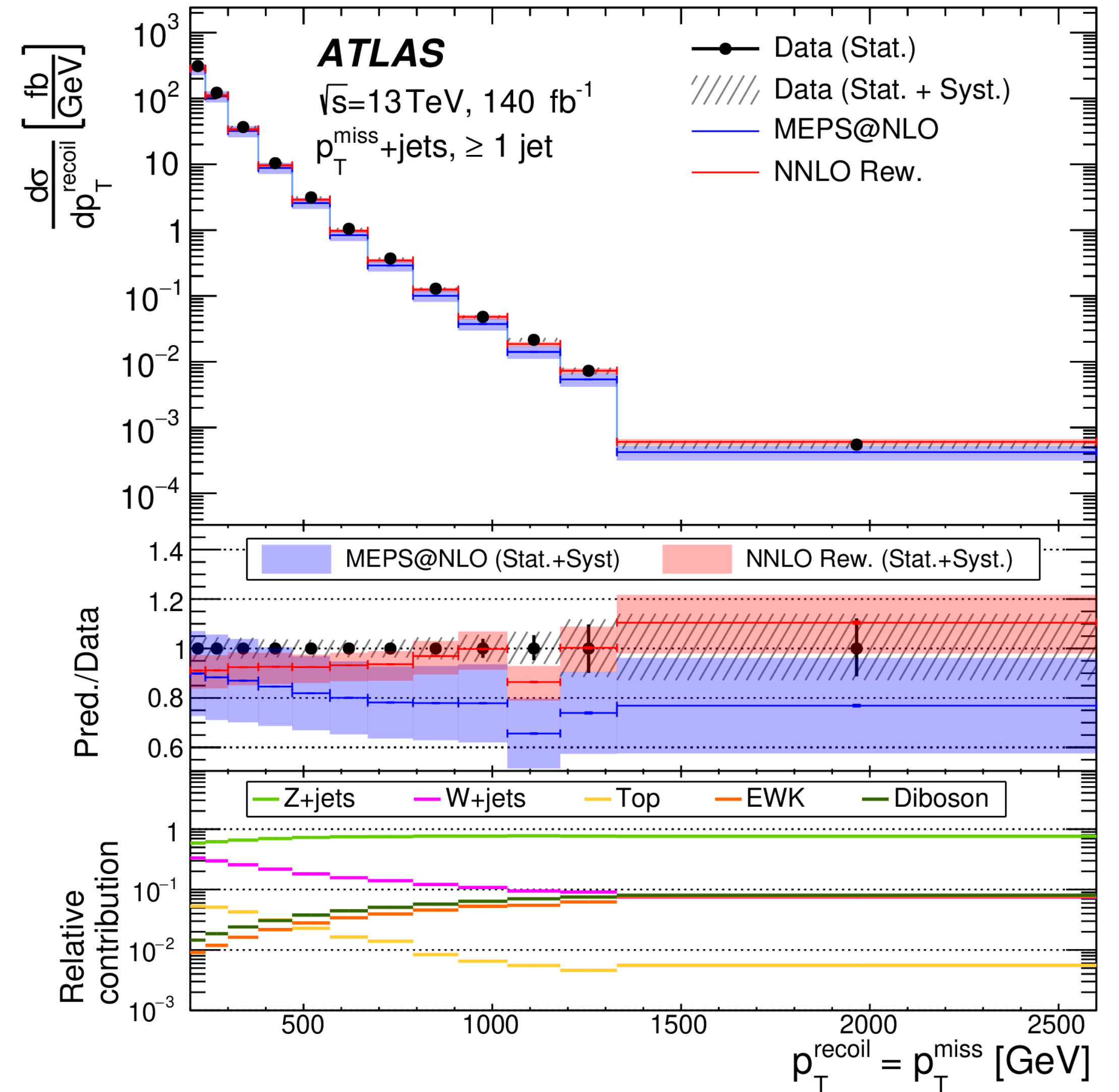


# A real example with all the pieces in place

My phase space

My results

**Fully corrected! Usable by anyone with a theory**

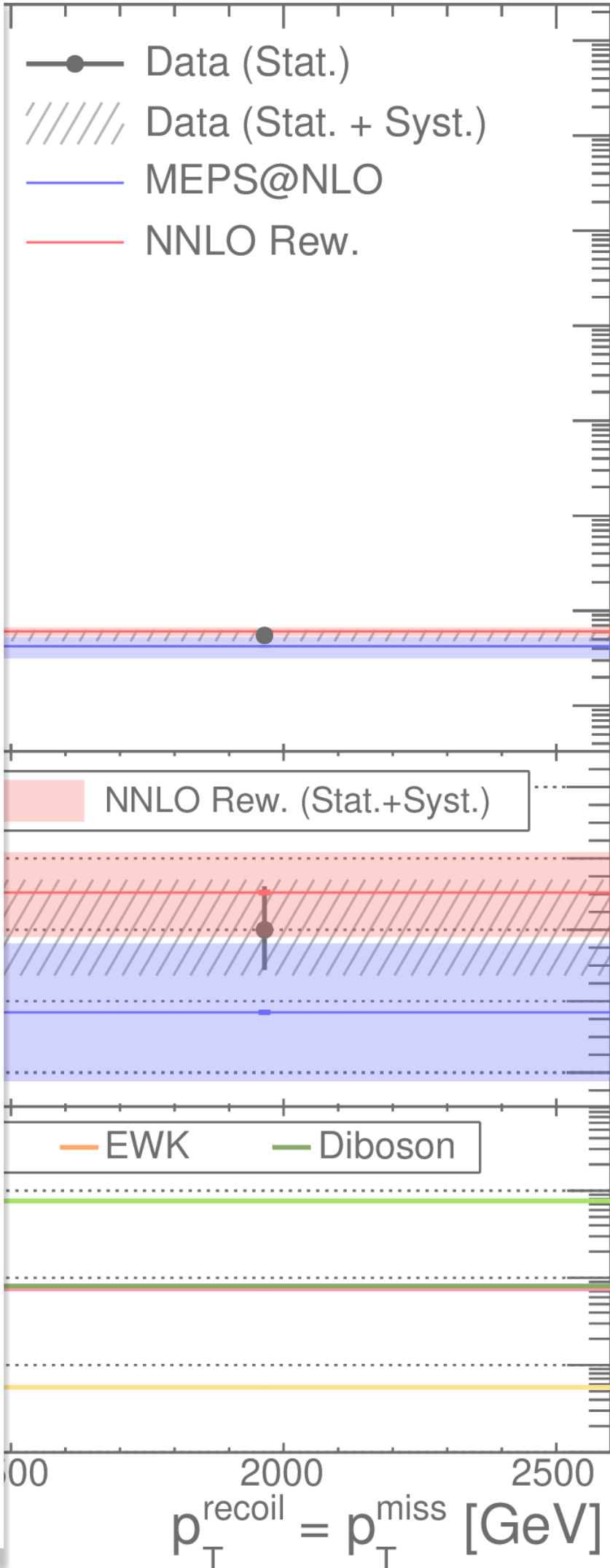
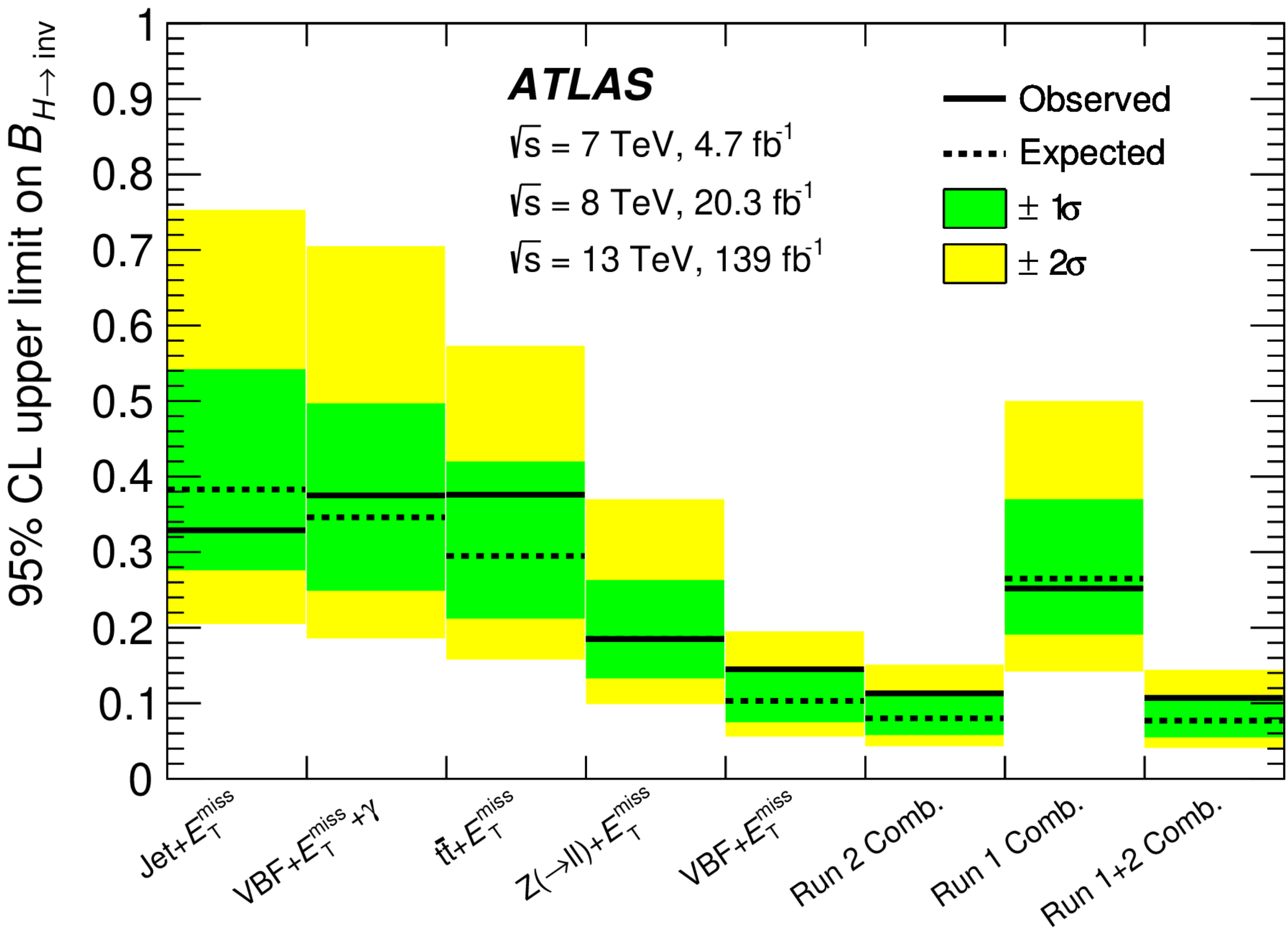


# A real example with all the pieces in place

My pha

My

Fully correc  
anyone w



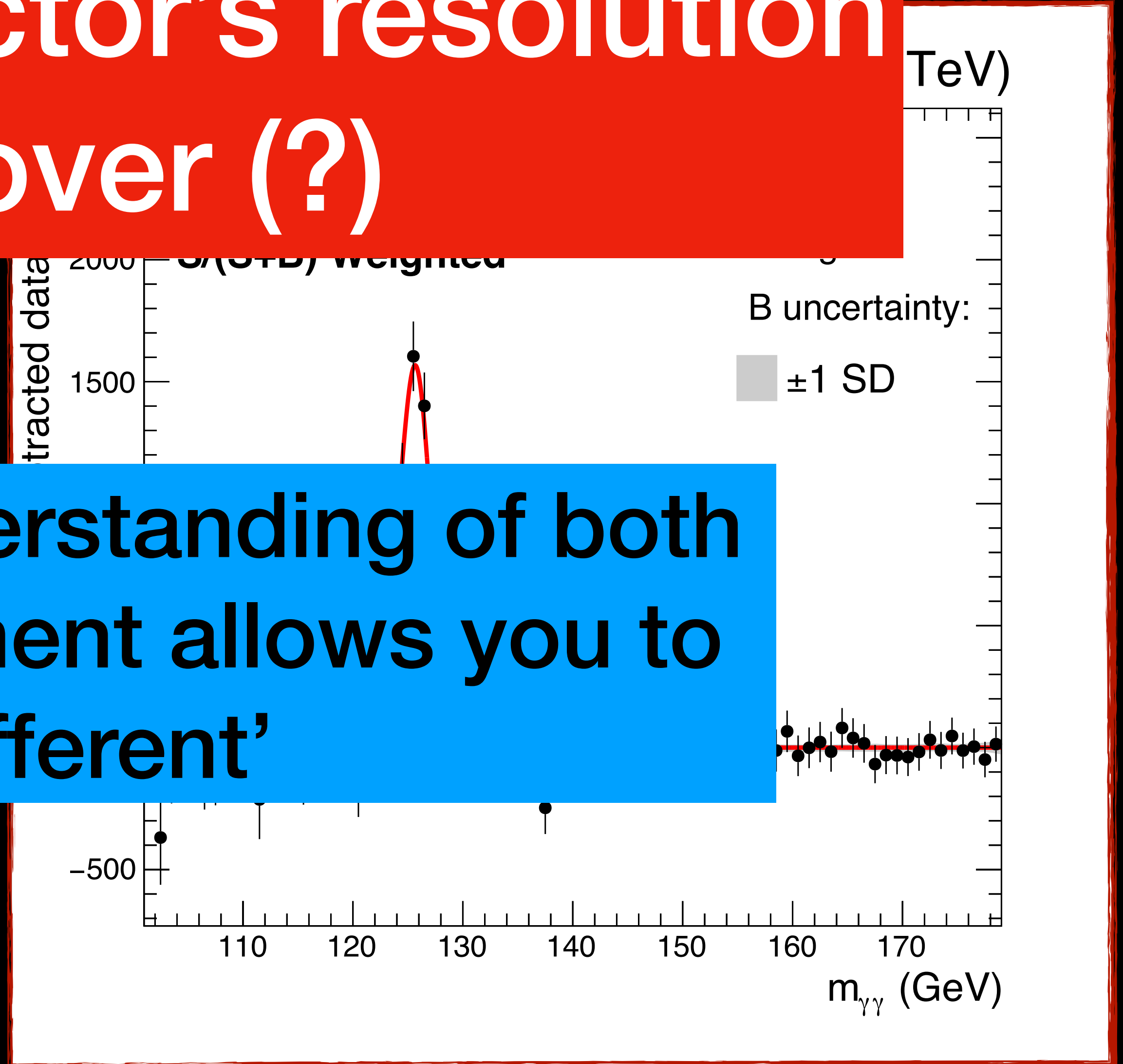


# This is the detector's resolution Game over (?)

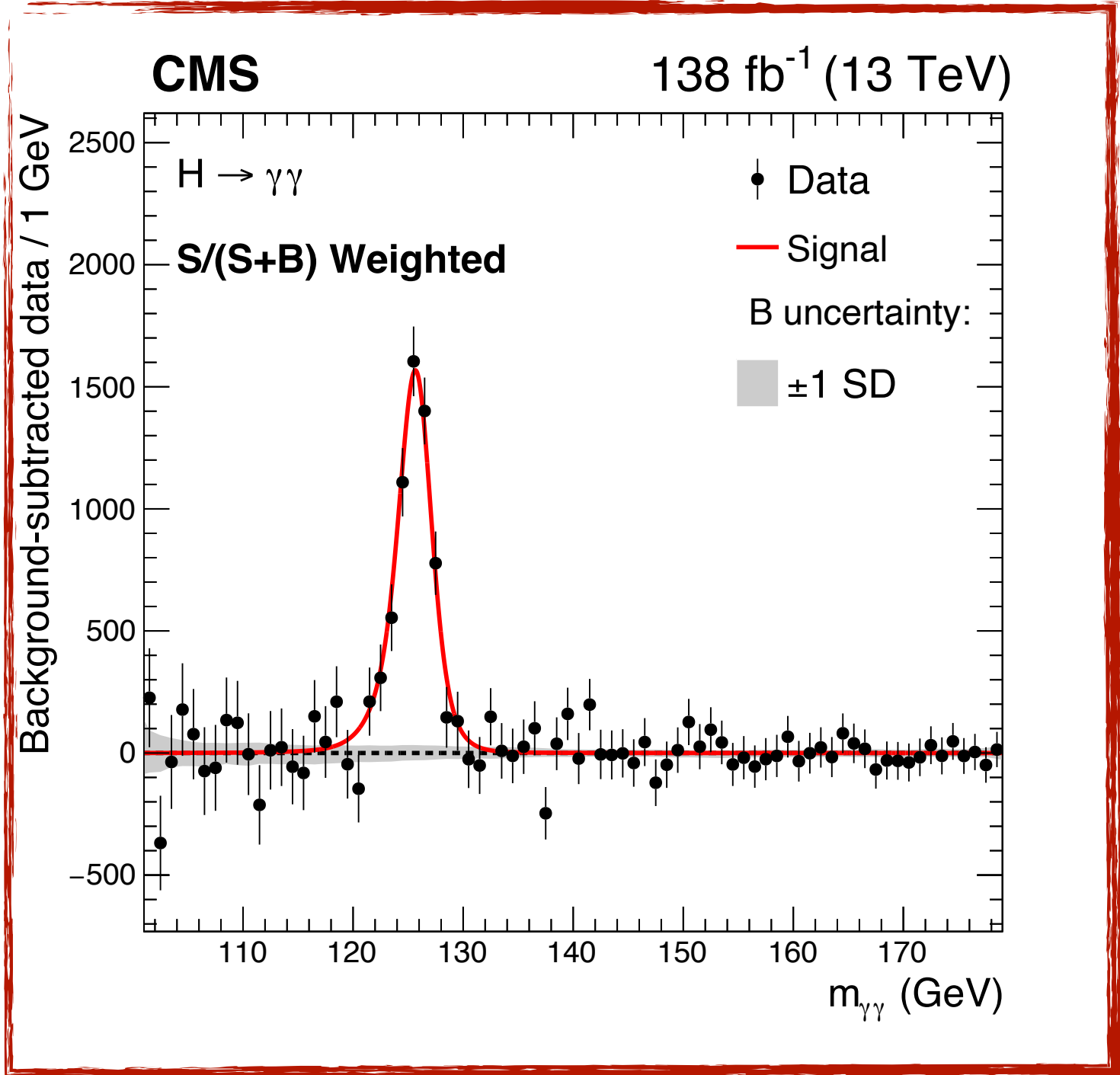
Mass. How?

Having a good understanding of both  
theory and experiment allows you to  
'think different'

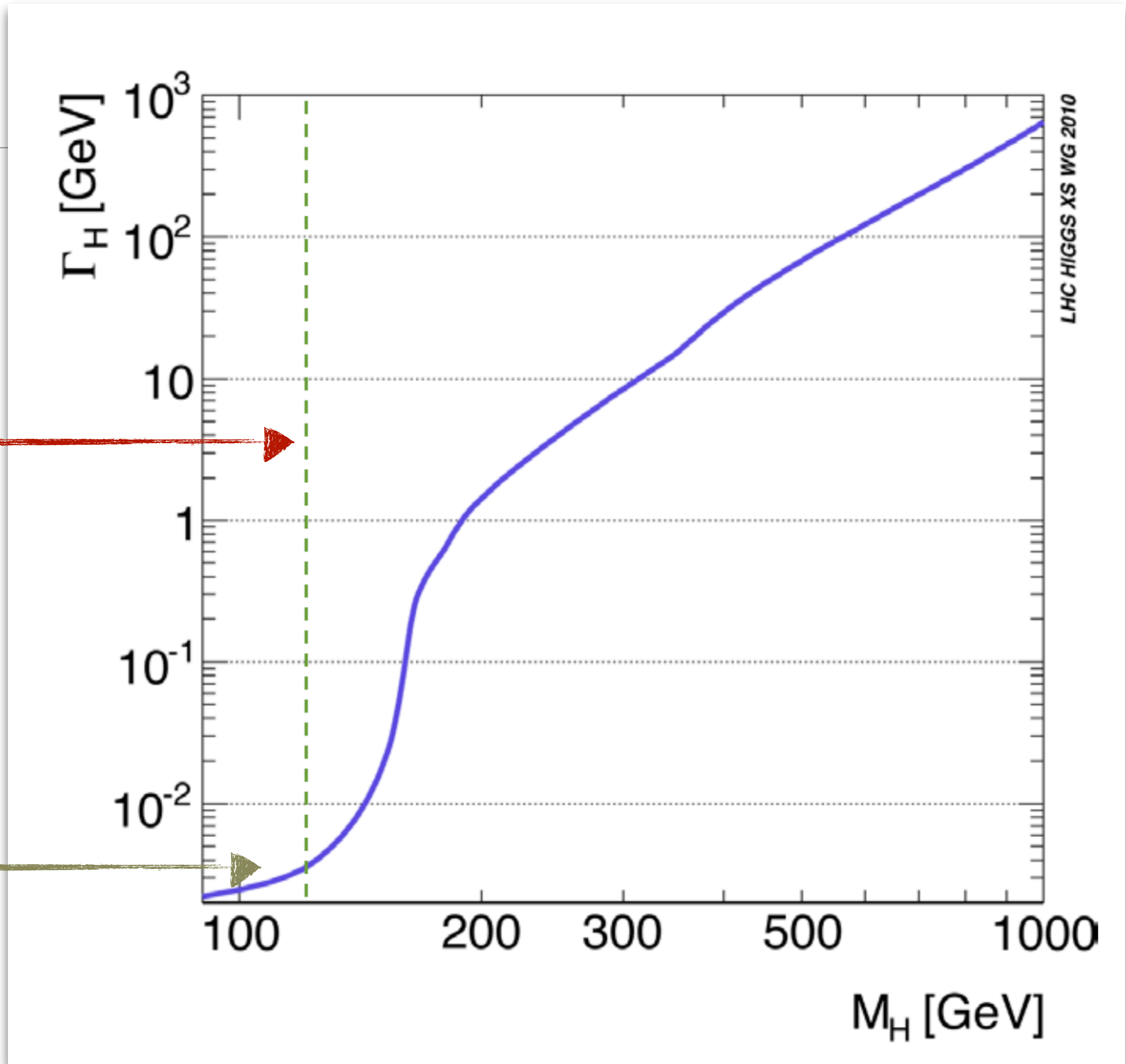
- Prediction 4.07 MeV



# An impossible width



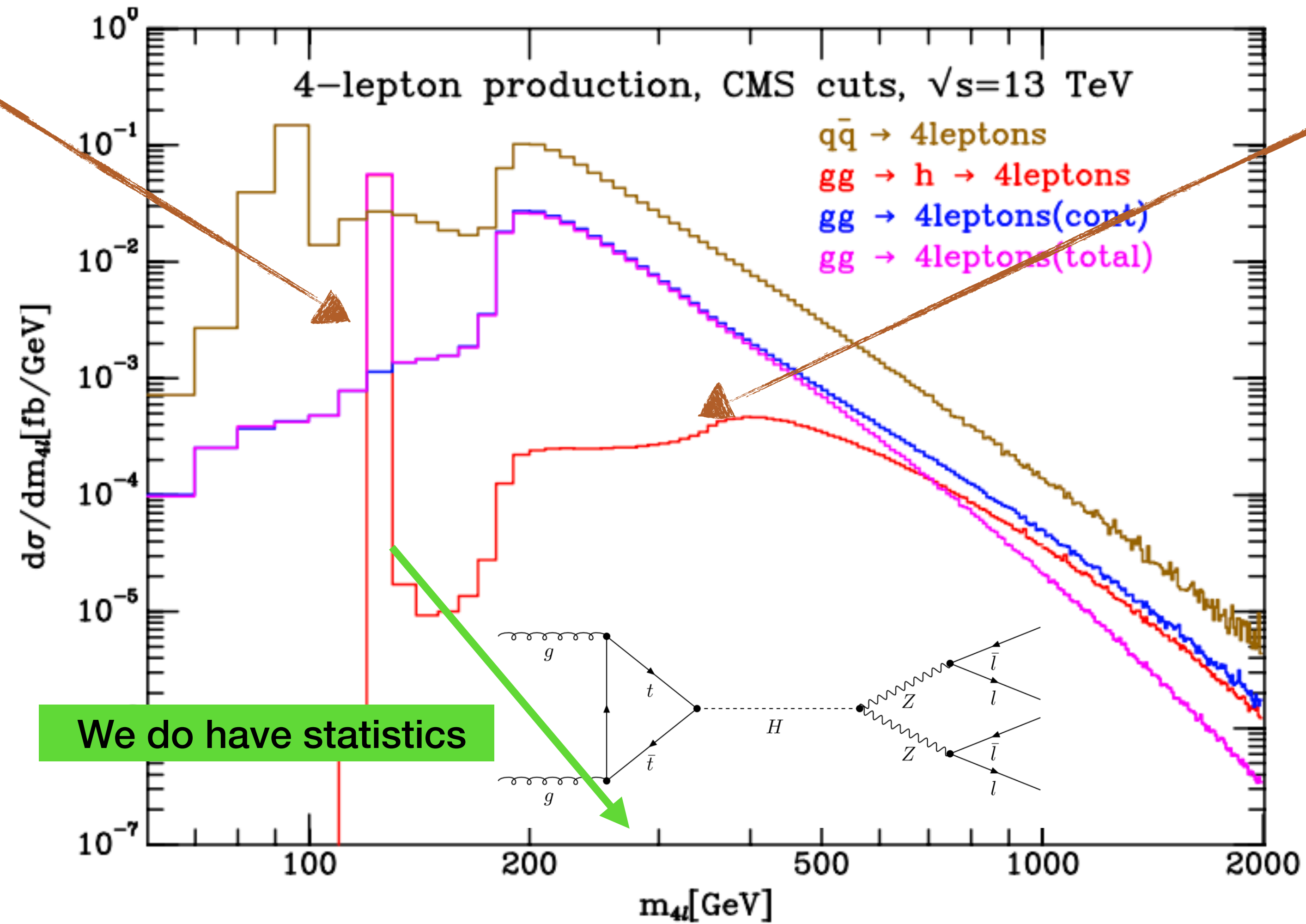
- Prediction 4.07 MeV



# Particle vs. Propagator

On-shell

Off-shell

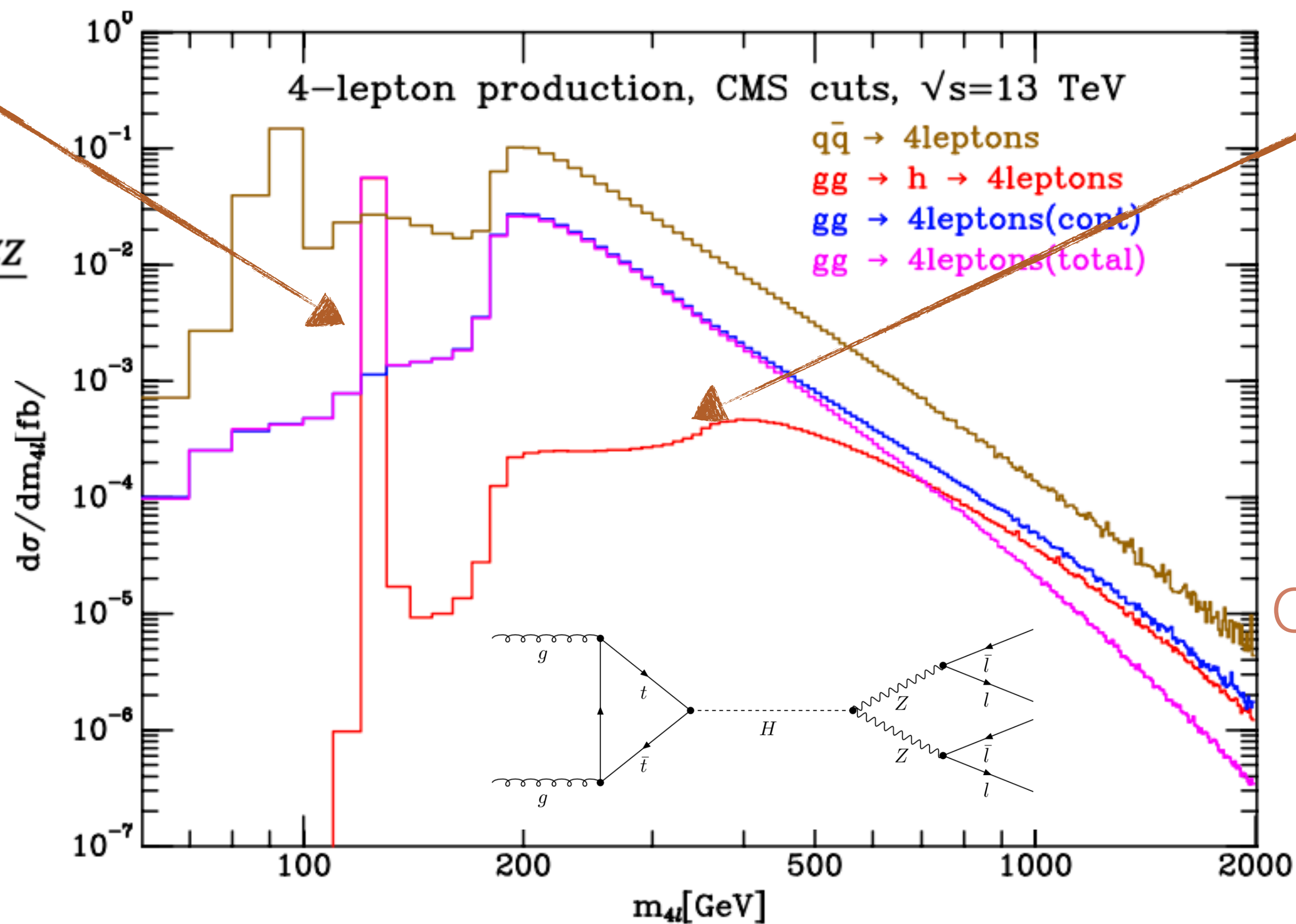


# Particle vs. Propagator

On-shell

$$\sigma_{\text{on-shell}}^{pp \rightarrow H \rightarrow ZZ} \propto \frac{g_{Hgg}^2 g_{HZZ}^2}{m_H \Gamma_H}$$

Width dependence



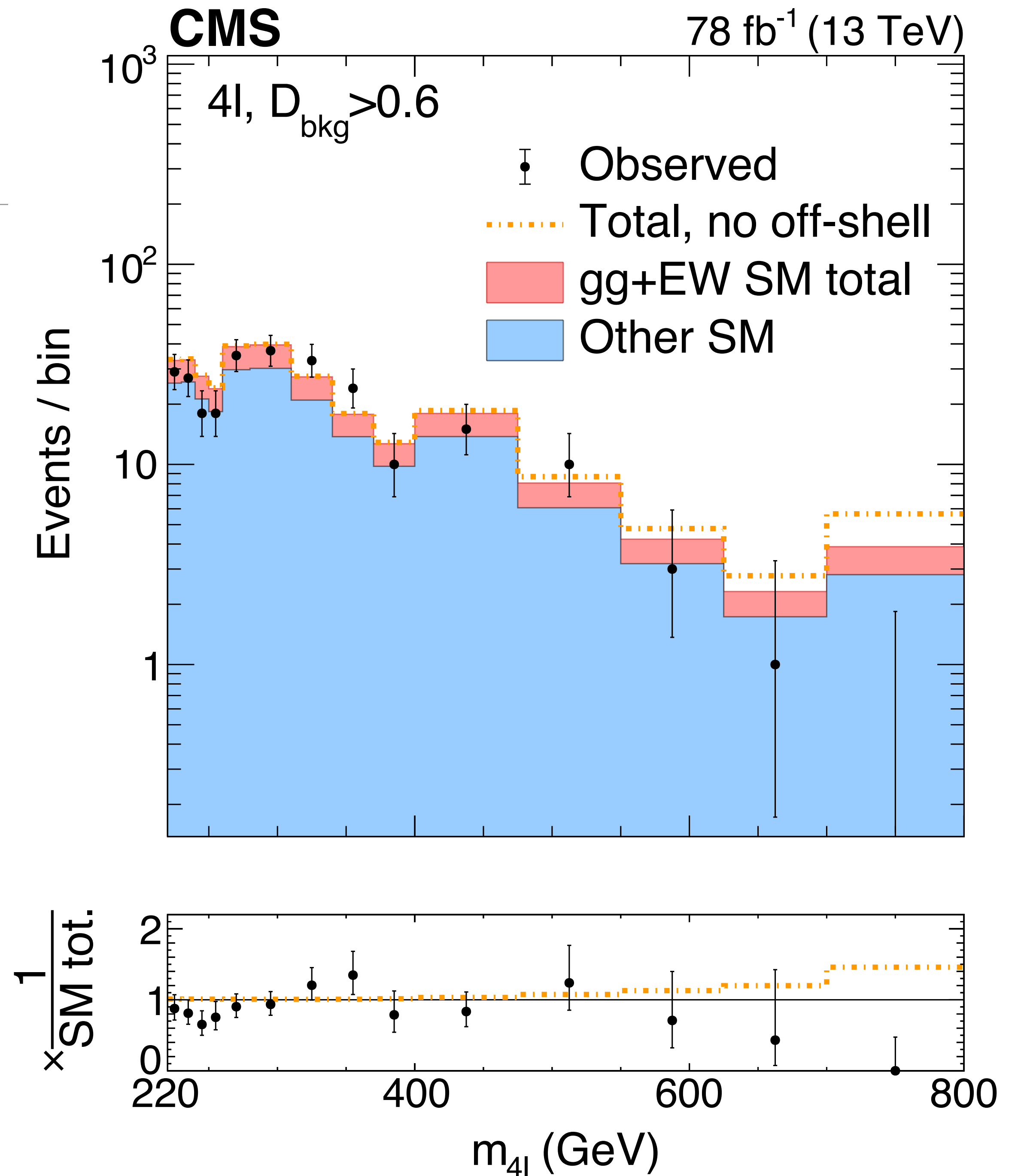
Off-shell

$$\frac{d\sigma_{\text{off-shell}}^{pp \rightarrow H \rightarrow ZZ}}{dm_{ZZ}} \propto \frac{g_{Hgg}^2 g_{HZZ}^2}{(m_{ZZ}^2 - m_H^2)^2}$$

Coupling, mass dependence

# Fitting it all together

- [enter here... a huge amount of measurement fanciness]
- Fit 117 categories together in both on-shell and off-shell to extract the width



Obtaining the width

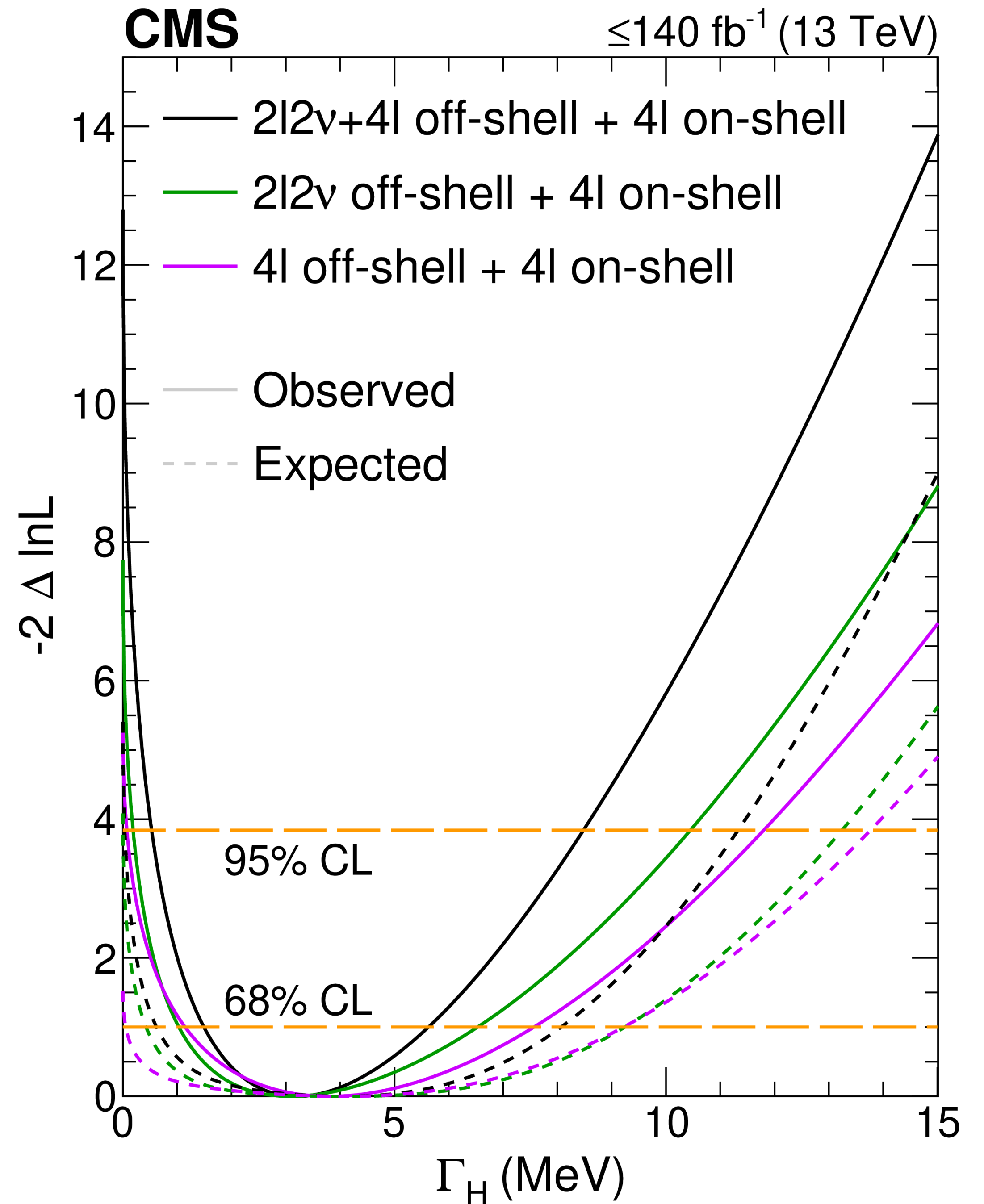
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**CMS**

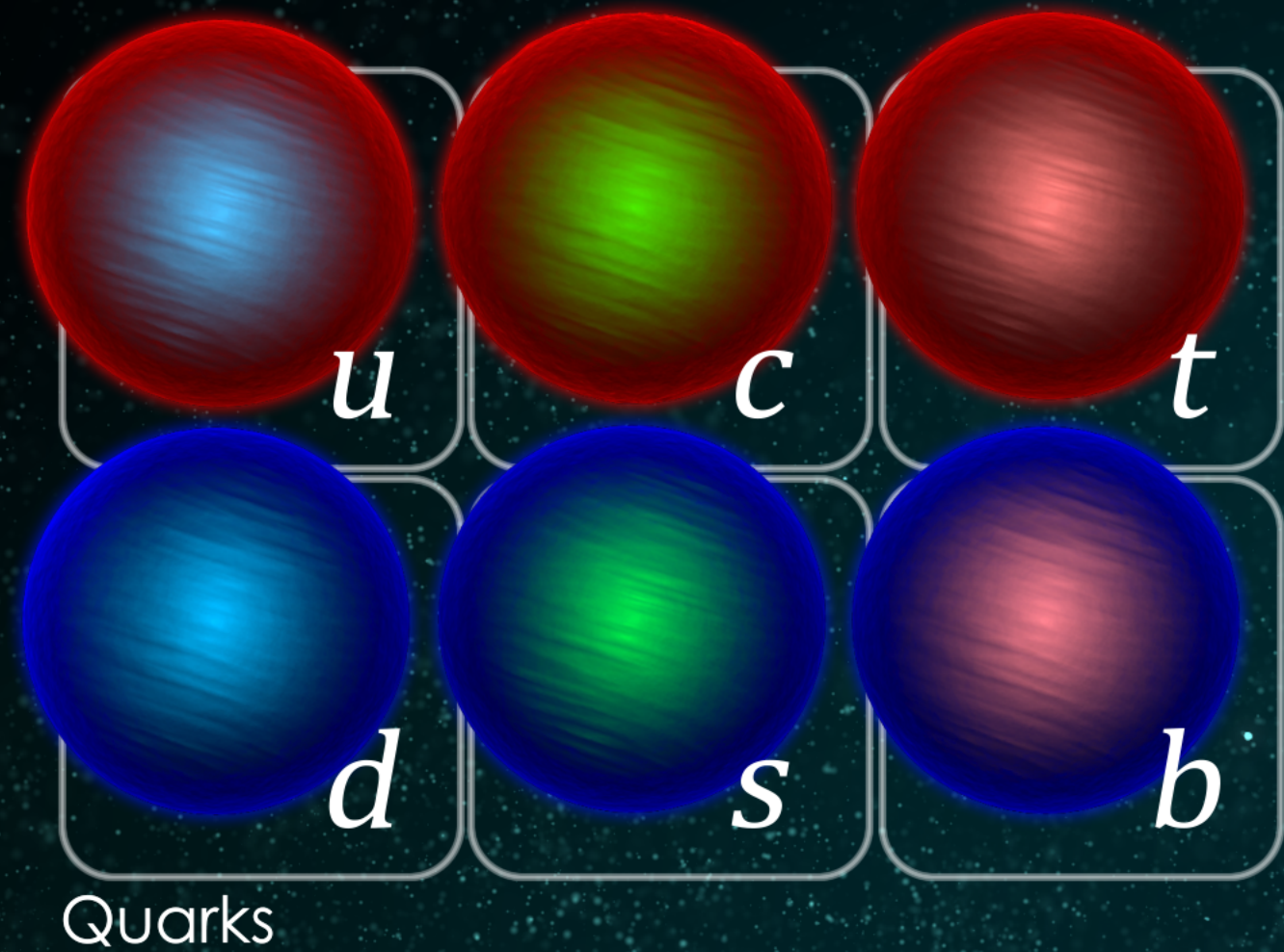
$$\Gamma = 3.2^{+2.4}_{-1.7} \text{ MeV}$$

**ATLAS**

$$\Gamma = 4.6^{+2.6}_{-2.5} \text{ MeV}$$

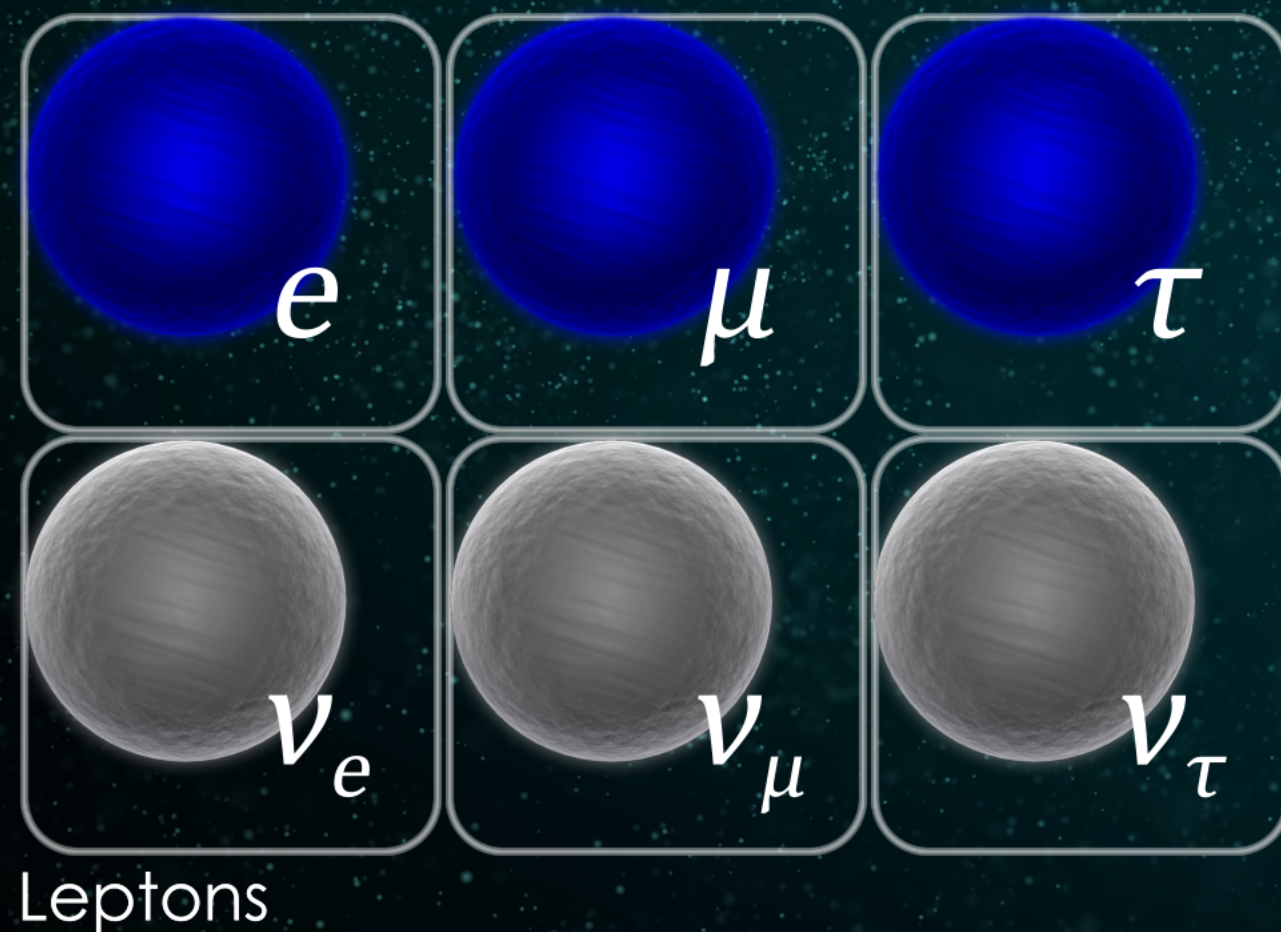


Mass hierarchy

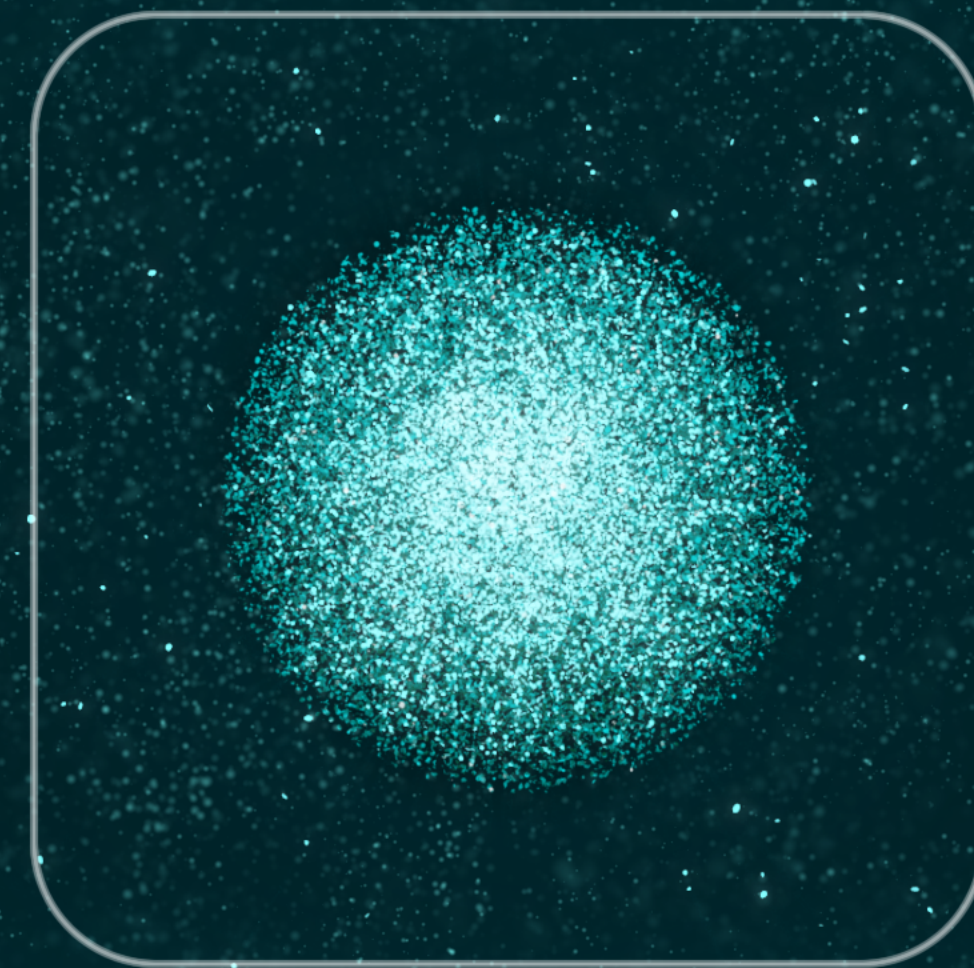


CP Violation

Additional particles?



# The genesis to any measurement is the question itself



Higgs boson

Composite Higgs?

Who breaks electroweak symmetry?

Why is  $\theta_{13}$  so small?

Dark energy?

Dark matter?



Forces

Gravity?

Unified forces?

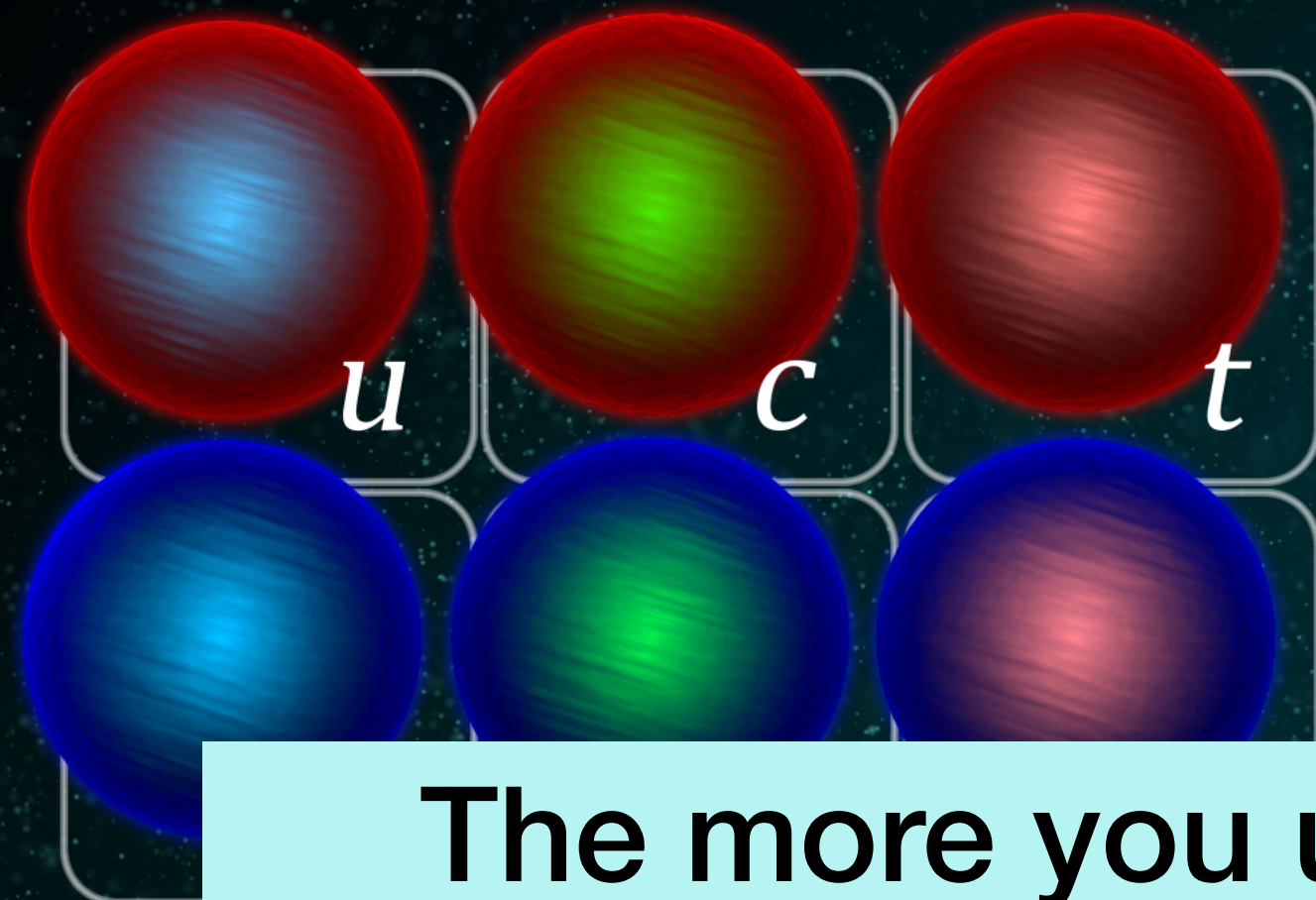
Majorana vs. Dirac

Why is neutrino mixing so large?

Theories beyond the Standard Model?

Mass hierarchy

# The genesis to any measurement is the question itself



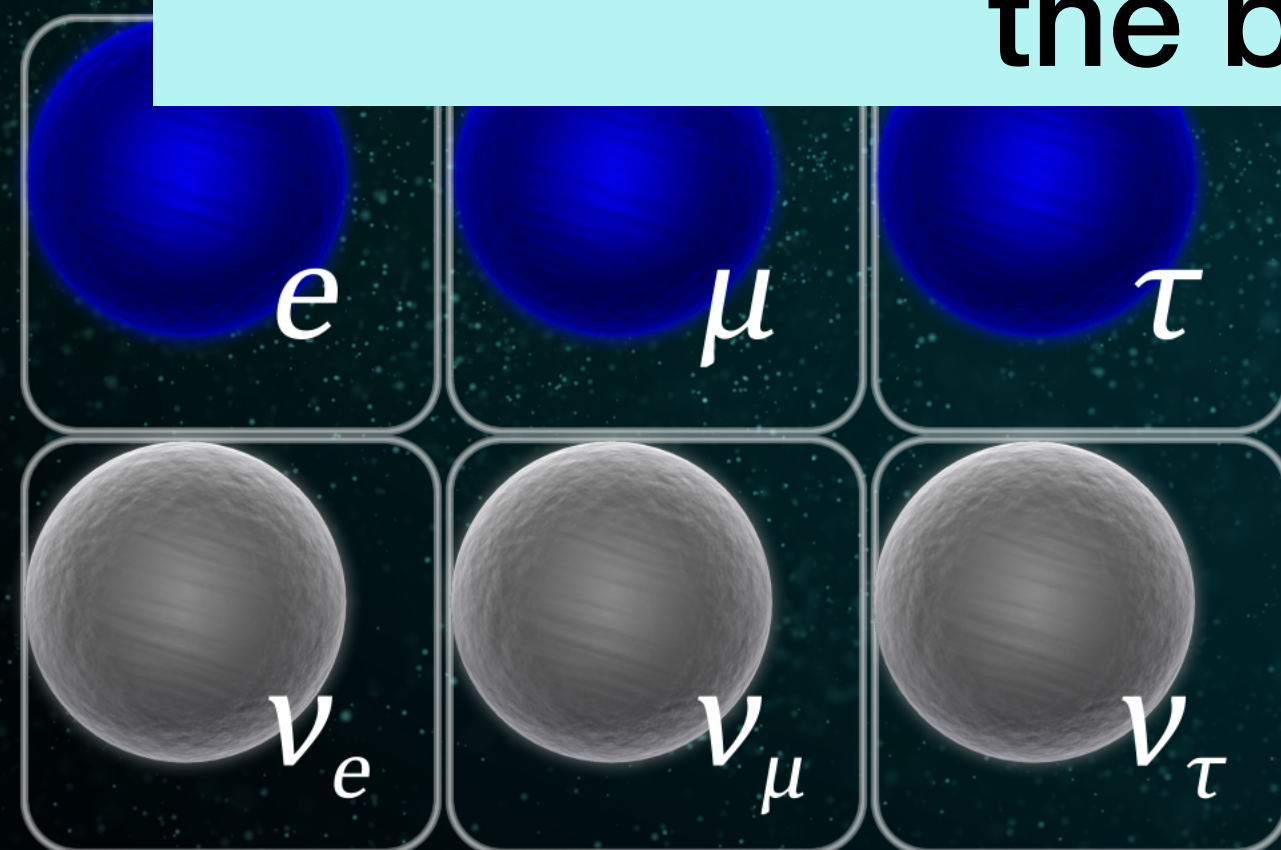
Dark energy?

Dark matter?

The more you understand about what goes into the experimental measurement and the theory estimations —

the better to unlock the answer

Additional particles



Higgs boson

Composite Higgs?



Forces

Gravity?

Who breaks electroweak symmetry?

Unified forces?

Why is  $\theta_{13}$  so small?

Majorana vs. Dirac

Why is neutrino mixing so large?

Theories beyond the Standard Model?



**Extras**