

SPONTANEOUS CP VIOLATION FROM *S*EVERAL ANGLES

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^(*) Who is applying for postdocs

Spontaneous CP violation

PHYSICAL REVIEW D

VOLUME 8, NUMBER 4

15 AUGUST 1973

A Theory of Spontaneous T Violation*

T. D. Lee

Department of Physics, Columbia University, New York, New York 10027

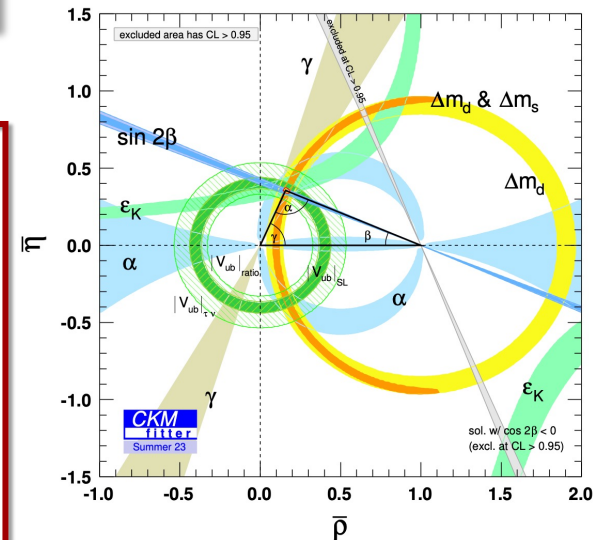
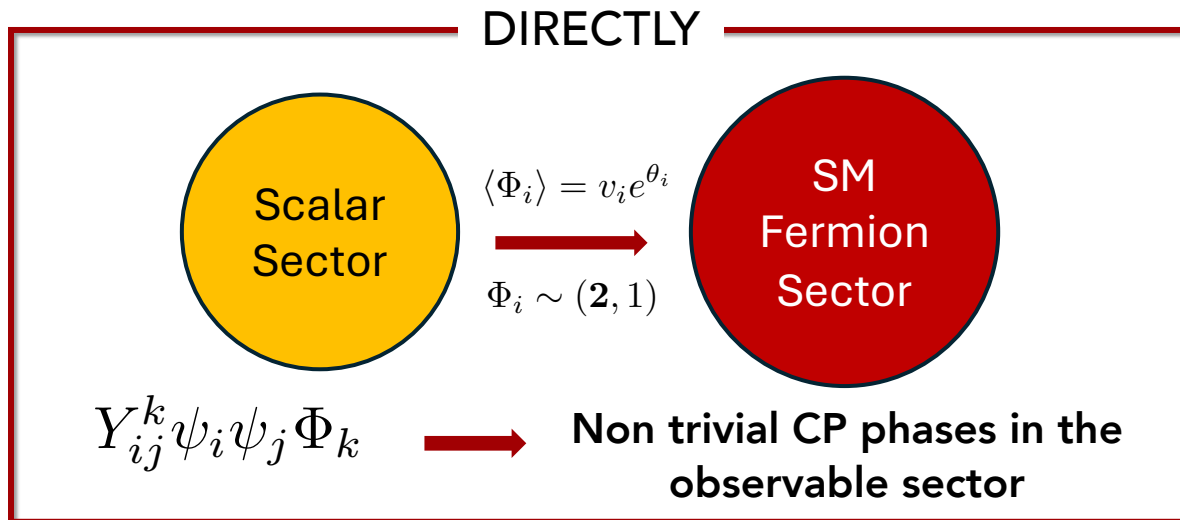
(Received 11 April 1973)

A theory of spontaneous T violation is presented. The total Lagrangian is assumed to be invariant under the time reversal T and a gauge transformation (e.g., the hypercharge gauge), but the physical solutions are not. In addition to the spin-1 gauge field and the known matter fields, in its simplest form the theory consists of two complex spin-0 fields. Through the spontaneous symmetry-breaking mechanism of Goldstone and Higgs, the vacuum expectation values of these two spin-0 fields can be characterized by the shape of a triangle and their quantum fluctuations by its vibrational modes, just like a triangular molecule. T violations can be produced among the known particles through virtual excitations of the vibrational modes of the triangle which has a built-in T -violating phase angle. Examples of both Abelian and non-Abelian gauge groups are discussed. For renormalizable theories, all spontaneously T -violating effects are finite. It is found that at low energy, below the threshold of producing these vibrational quanta, T violation is always quite small.

First model of Spontaneous CP Violation (SCPV)

T. D. Lee (1973)

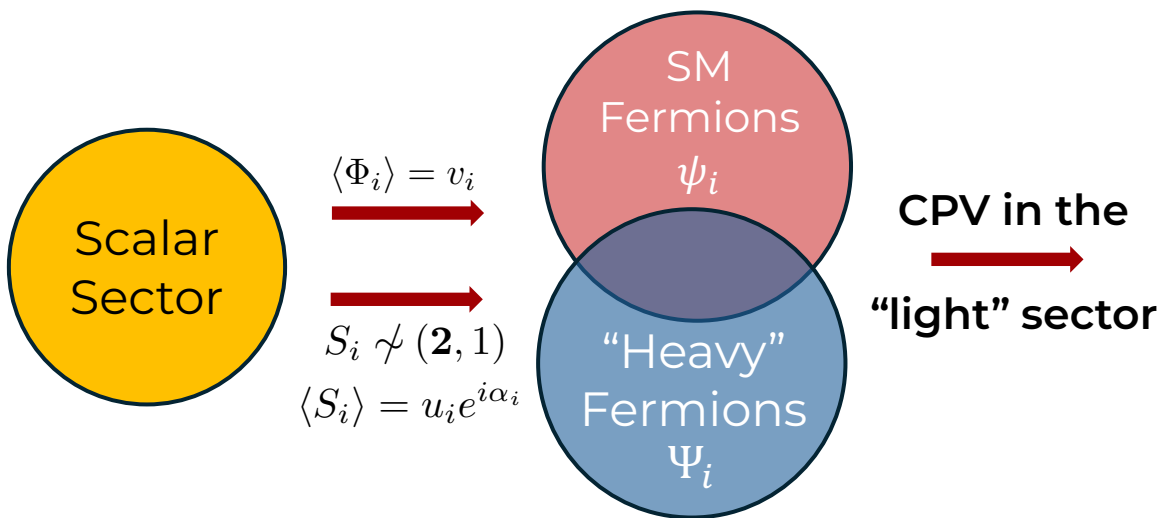
SCPV is only appealing if it can lead to **non-trivial CPV effects** in the fermion sector (CKM, PMNS,...)



HOW?

Spontaneous CP violation

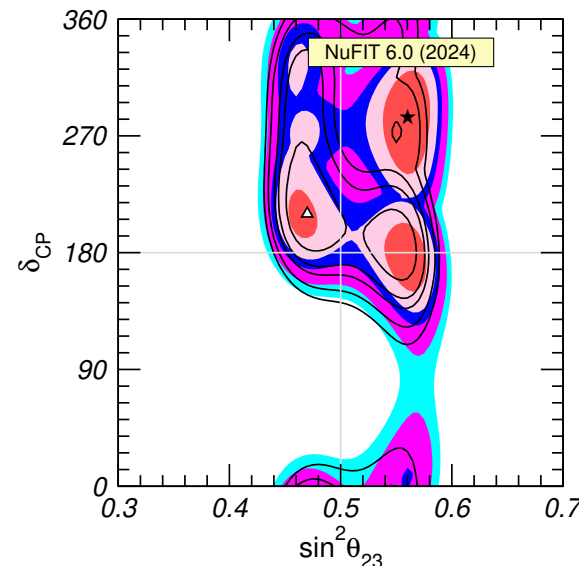
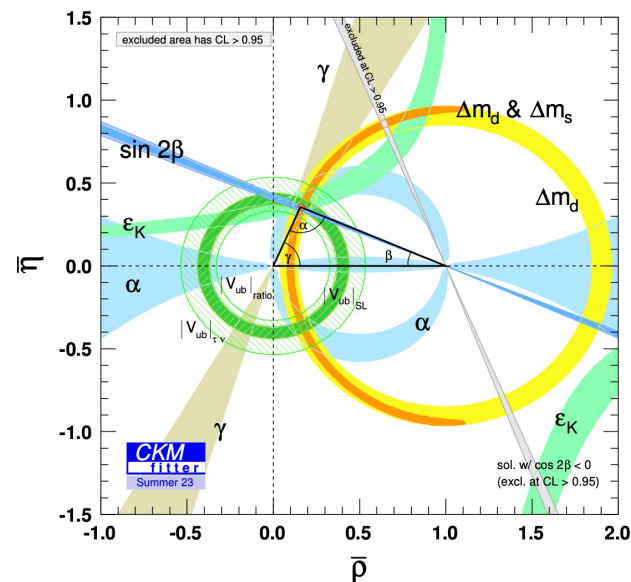
CPV in the **"light SM sector"** via a **heavy-fermion** portal



Heavy, light, heavy-light Yukawas

$$Y_{ij}^{(k)} \psi_i \psi_j \Phi_k + \tilde{Y}_{ij}^{(k)} \Psi_i \psi_j \Phi_k + \Gamma_{ij}^{(k)} \Psi_i \psi_j S_k + \dots$$

CPV + some extra nice things...



Spontaneous CP violation



Physics Letters B

Volume 267, Issue 1, 5 September 1991, Pages 95-99



A minimal model with natural suppression of strong CP violation

Luis Bento ^a*, Gustavo C. Branco ^b, Paulo A. Parada ^b

A Common Origin for all CP Violations

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(Dated: March 22, 2022)

We put forward the conjecture that all CP violating phenomena may have a common origin. In order to illustrate our idea, we present a minimal model where CP is spontaneously broken at a high energy scale, through the phase in the vacuum expectation value of a complex scalar singlet. This single phase is the origin of both low energy CP violation in the quark and leptonic sectors, as well as leptogenesis. We also show that in this framework the strong CP problem may be solved in a simple way through the introduction of a Z_4 symmetry which allows for the implementation of the Nelson-Barr mechanism.

(BPR Model) Z_4 SYMMETRY:

$$D^0 \rightarrow -D^0, S \rightarrow -S, \nu_R^0 \rightarrow i\nu_R^0, \psi_l^0 \rightarrow i\psi_l^0, e_R^0 \rightarrow ie_R^0$$

$$\mathcal{L}_q = \overline{\psi}_q^0 G_u \phi u_R^0 + \overline{\psi}_q^0 G_d \tilde{\phi} d_R^0 + (f_q S + f_q' S^*) \overline{D}_L^0 d_R^0 + \tilde{M} \overline{D}_L^0 D_R^0 + \text{h.c.}$$

$$\mathcal{L}_l = \overline{\psi}_l^0 G_l \phi e_R^0 + \overline{\psi}_l^0 G_\nu \tilde{\phi} \nu_R^0 + \frac{1}{2} \nu_R^{0T} C (f_\nu S + f_\nu' S^*) \nu_R^0 + \text{h.c.}$$

$$V = V_{SM} + (\mu^2 + \lambda_1 S^* S + \lambda_2 \phi^\dagger \phi) (S^2 + S^{*2}) + \lambda_3 (S^4 + S^{*4})$$

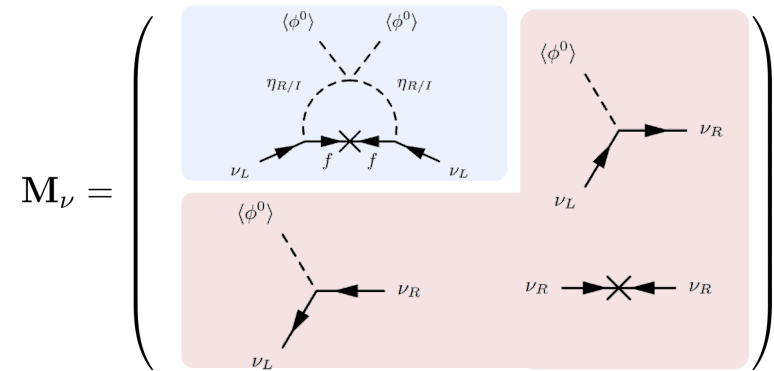
$$\langle \phi^0 \rangle = \frac{v}{\sqrt{2}}, \quad \langle S \rangle = \frac{V \exp(i\alpha)}{\sqrt{2}}$$



CPV in the "light" quark and neutrino sectors

Minimal Scoto-Seesaw I (S-STI) model

	ν_R	η	f
Rojas, Srivastava, Valle'19			
$SU(2)_L$	1	2	1
$U(1)_Y$	0	1/2	0
Z_2	+	-	-
Multiplicity	1	1	1



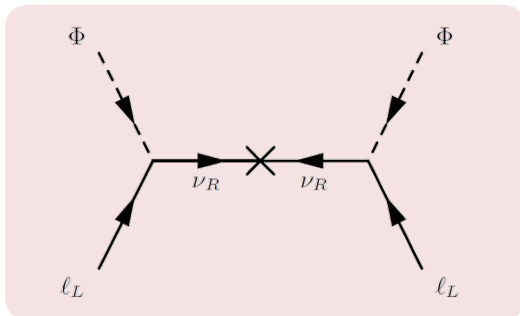
$$-\mathcal{L} = \underbrace{\bar{\ell}_L \mathbf{Y}_\nu^* \tilde{\Phi} \nu_R + \frac{1}{2} M_R \bar{\nu}_R \nu_R^c}_{\text{atmospheric}} + \underbrace{\bar{\ell}_L \mathbf{Y}_f^* \tilde{\eta} f + \frac{1}{2} M_f \bar{f} f^c}_{\text{solar}} + \text{H.c.}$$

Generates the **atmospheric** neutrino mass scale

Generates the **solar** neutrino mass scale

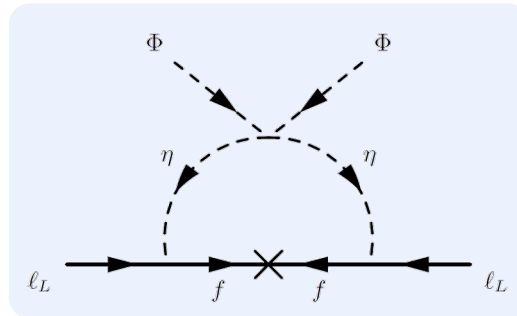
At the effective level:

$$= -\frac{v^2}{2} \frac{\mathbf{Y}_\nu \mathbf{Y}_\nu^T}{M_R} + \mathcal{F}(M_f, m_{\eta_R^0}, m_{\eta_I^0}) M_f \mathbf{Y}_f \mathbf{Y}_f^T$$



Type-I seesaw contribution

Minkowski'77; M. Gell-Mann, P. Ramond and R. Slansky'79; T. Yanagida'79; S.L. Glashow'80; Mohapatra and G. Senjanovic'80; Schechter & Valle'80.



Scotogenic contribution

Tao, PRD 54 (1996) 5693; Ma, PRD 73 (2006) 077301

- **Economical framework** for neutrino masses and mixing (one massless neutrino)
- **Scalar or fermion DM**
- **New LFV contributions** mediated by dark fermions/scalars

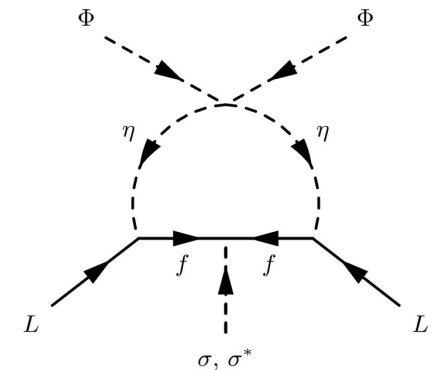
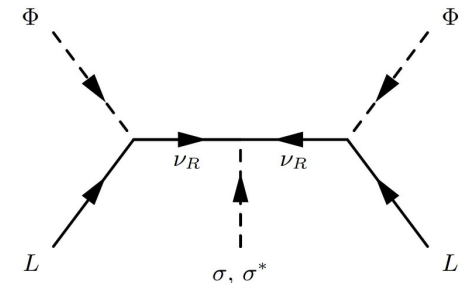
Scoto-seesaw + SCPV from a scalar singlet $\langle \sigma \rangle = ue^{i\theta}$

$$\frac{1}{2}(y_R\sigma + \tilde{y}_R\sigma^*)\overline{\nu}_R\nu_R^c + \frac{1}{2}(y_f\sigma + \tilde{y}_f\sigma^*)\bar{f}f^c + \text{H.c.}$$

$$\mathbf{M}_\nu = -v^2 e^{i(\theta_f - \theta_R)} \frac{\mathbf{Y}_\nu \mathbf{Y}_\nu^T}{|M_R|} + \mathcal{F}(|M_f|, m_{\eta_R}, m_{\eta}) |M_f| \mathbf{Y}_f \mathbf{Y}_f^T$$

$$|M_{R,f}|^2 = [y_{R,f}^2 + \tilde{y}_{R,f}^2 + 2y_{R,f}\tilde{y}_{R,f} \cos(2\theta_{R,f})] u^2$$

$$\tan(\theta_f - \theta_R) = \frac{(y_f\tilde{y}_R - y_R\tilde{y}_f) \sin(2\theta)}{y_R y_f + \tilde{y}_R \tilde{y}_f + (y_R \tilde{y}_f + y_f \tilde{y}_R) \cos(2\theta)}$$



Minimal scoto-seesaw

$$1 \nu_R + 1 f$$

(too many parameters, no predictions in the)

+

Abelian Flavour symmetries



Two massless neutrinos and/or
 vanishing mixing angles

Minimal model that leads to
 predictions in the neutrino sector:

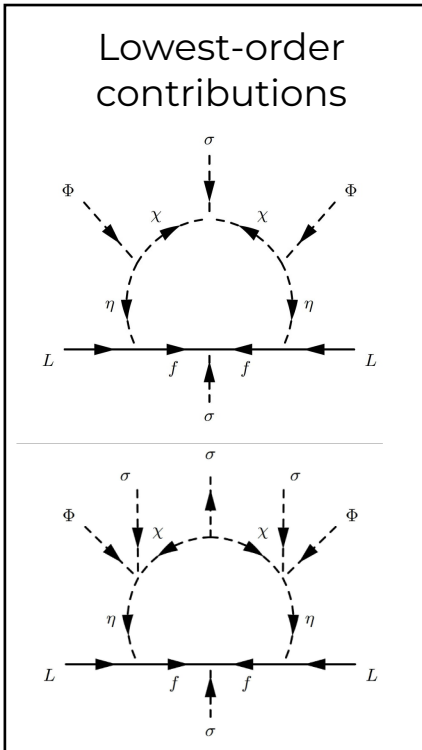
$$\mathbf{Z}_8 \text{ flavour symmetry} + 2 \nu_R + 1 f$$

$$\downarrow \langle \sigma \rangle = ue^{i\theta}$$

\mathbf{Z}_2 DM symmetry

Flavour Structures:

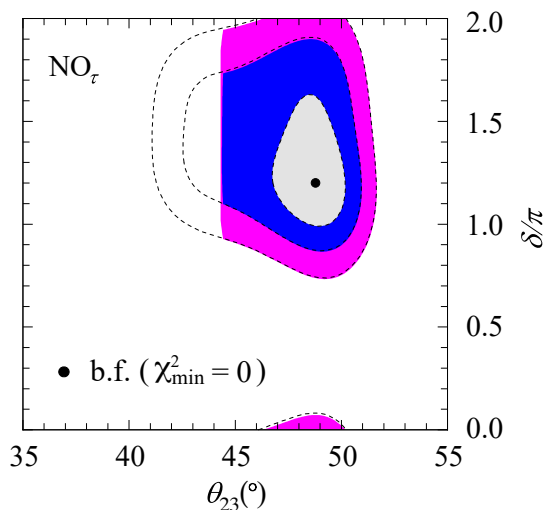
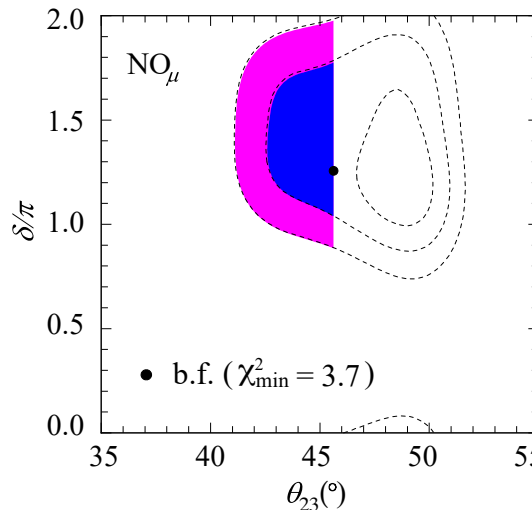
$$\mathbf{Y}_\nu = \begin{pmatrix} \times & 0 \\ 0 & \times \\ \times & 0 \end{pmatrix}, \quad \mathbf{Y}_f = \begin{pmatrix} \times \\ 0 \\ \times \end{pmatrix}, \quad \mathbf{Y}_\ell = \begin{pmatrix} \times & 0 & \times \\ 0 & \times & 0 \\ \times & 0 & \times \end{pmatrix}, \quad \mathbf{M}_R = \begin{pmatrix} 0 & \times \\ \cdot & \times \end{pmatrix}$$



	Fields	SU(2) _L × U(1) _Y	$\mathcal{Z}_8^{e-\mu} \rightarrow \mathcal{Z}_2^D$	$\mathcal{Z}_8^{\mu-\tau} \rightarrow \mathcal{Z}_2^D$	$\mathcal{Z}_8^{e-\tau} \rightarrow \mathcal{Z}_2^D$
Fermions	L_e, e_R	(2, -1/2), (1, 0)	$\omega^6 \equiv -i \rightarrow +1$	$\omega^0 \equiv 1 \rightarrow +1$	$\omega^6 \equiv -i \rightarrow +1$
	L_μ, μ_R	(2, -1/2), (1, 0)	$\omega^6 \equiv -i \rightarrow +1$	$\omega^6 \equiv -i \rightarrow +1$	$\omega^0 \equiv 1 \rightarrow +1$
	L_τ, τ_R	(2, -1/2), (1, 0)	$\omega^0 \equiv 1 \rightarrow +1$	$\omega^6 \equiv -i \rightarrow +1$	$\omega^6 \equiv -i \rightarrow +1$
	ν_R^1	(1, 0)	$\omega^6 \equiv -i \rightarrow +1$	$\omega^6 \equiv -i \rightarrow +1$	$\omega^6 \equiv -i \rightarrow +1$
	ν_R^2	(1, 0)	$\omega^0 \equiv 1 \rightarrow +1$	$\omega^0 \equiv 1 \rightarrow +1$	$\omega^0 \equiv 1 \rightarrow +1$
	f	(1, 0)	$\omega^3 \rightarrow -1$	$\omega^3 \rightarrow -1$	$\omega^3 \rightarrow -1$
Scalars	Φ	(2, 1/2)	$\omega^0 \equiv 1 \rightarrow +1$	$\omega^0 \equiv 1 \rightarrow +1$	$\omega^0 \equiv 1 \rightarrow +1$
	σ	(1, 0)	$\omega^2 \equiv i \rightarrow +1$	$\omega^2 \equiv i \rightarrow +1$	$\omega^2 \equiv i \rightarrow +1$
	η	(2, 1/2)	$\omega^5 \rightarrow -1$	$\omega^5 \rightarrow -1$	$\omega^5 \rightarrow -1$
	χ	(1, 0)	$\omega^3 \rightarrow -1$	$\omega^3 \rightarrow -1$	$\omega^3 \rightarrow -1$

In cases $\text{NO}_{\mu,\tau}$ the model predicts the **octant** of the **atmospheric mixing angle**

In the remaining cases the model does not constrain the mixing parameters.





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Flavour and dark matter in a scoto/type-II seesaw model

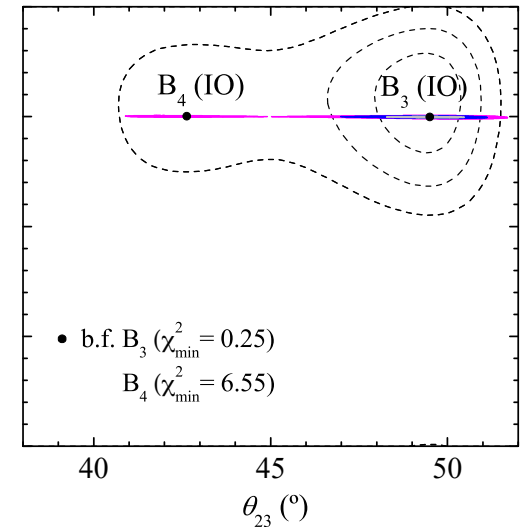
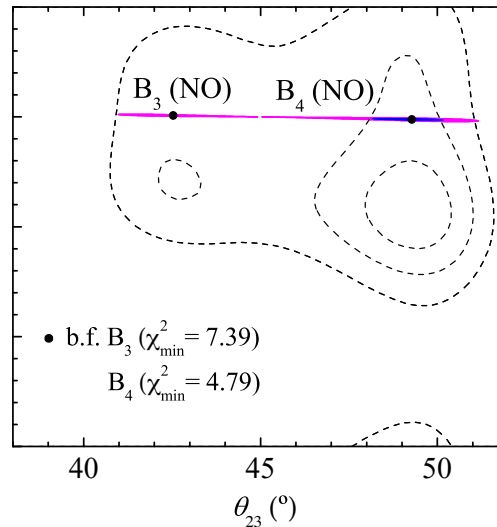
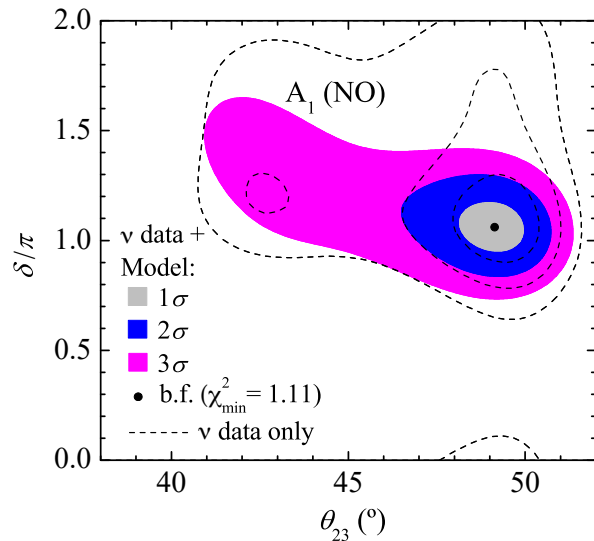
JHEP08(2022)030

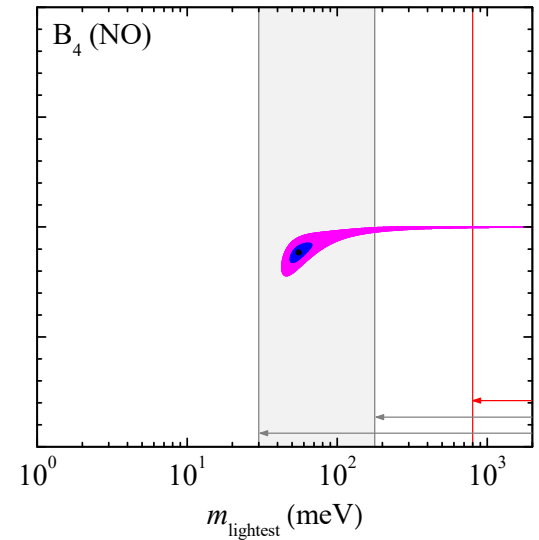
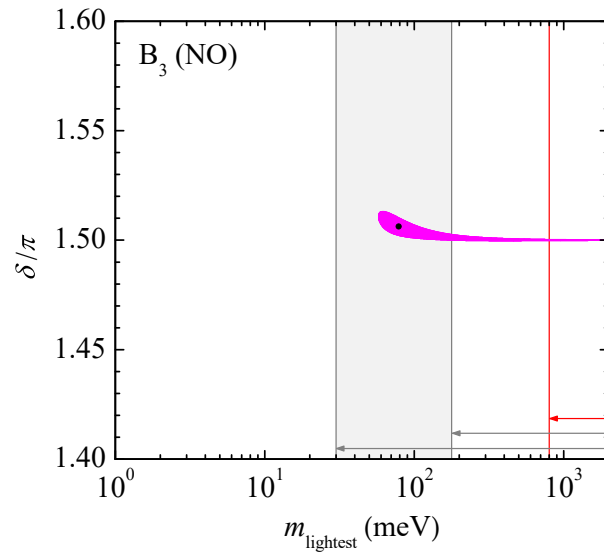
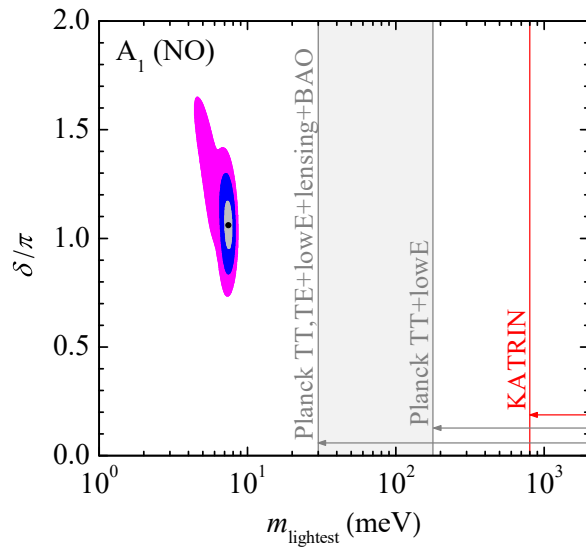
D.M. Barreiros, H.B. Câmara and F.R. Joaquim

	Fields	$SU(2)_L \otimes U(1)_Y$	$Z_8^{e-\mu} \rightarrow Z_2$	$Z_8^{e-\tau} \rightarrow Z_2$	$Z_8^{\mu-\tau} \rightarrow Z_2$
Fermions	ℓ_{eL}, e_R	$(2, -1/2), (1, -1)$	$1 \rightarrow +$	$1 \rightarrow +$	$\omega^2 \rightarrow +$
	$\ell_{\mu L}, \mu_R$	$(2, -1/2), (1, -1)$	$\omega^6 \rightarrow +$	$\omega^2 \rightarrow +$	$1 \rightarrow +$
	$\ell_{\tau L}, \tau_R$	$(2, -1/2), (1, -1)$	$\omega^2 \rightarrow +$	$\omega^6 \rightarrow +$	$\omega^6 \rightarrow +$
	f	$(1, 0)$	$\omega^3 \rightarrow -$	$\omega^3 \rightarrow -$	$\omega^3 \rightarrow -$
Scalars	Φ	$(2, 1/2)$		$1 \rightarrow +$	
	Δ	$(3, 1)$		$1 \rightarrow +$	
	σ	$(1, 0)$		$\omega^2 \rightarrow +$	
	η_1	$(2, 1/2)$		$\omega^3 \rightarrow -$	
	η_2	$(2, 1/2)$		$\omega^5 \rightarrow -$	

Very restricted flavour structure for the effective neutrino mass texture

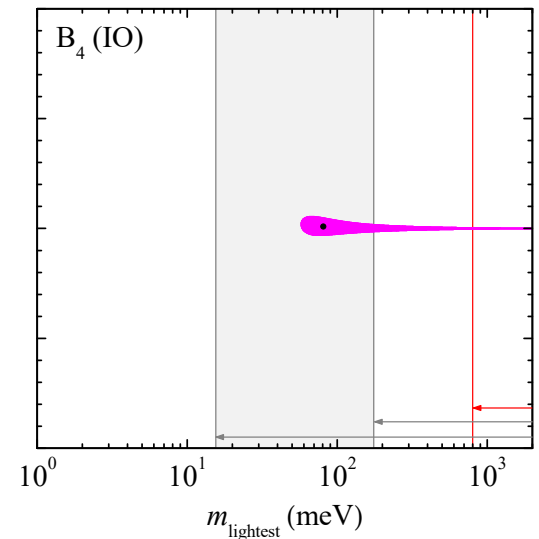
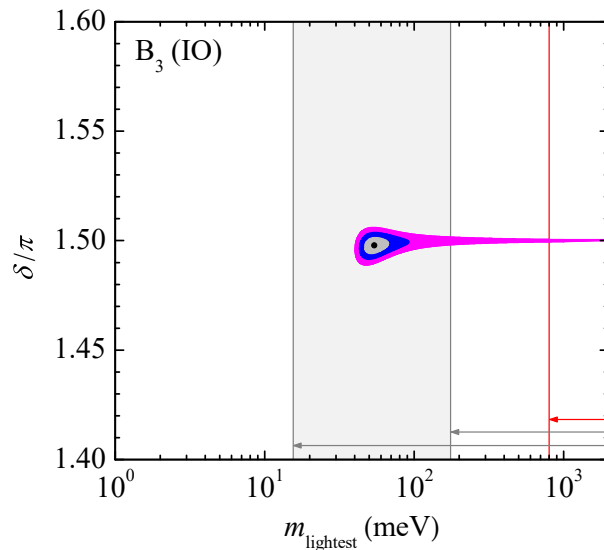
$$Z_8^{e-\mu} \rightarrow B_4 : \begin{pmatrix} \times & \times & 0 \\ \cdot & \times & \times \\ \cdot & \cdot & 0 \end{pmatrix}, Z_8^{e-\tau} \rightarrow B_3 : \begin{pmatrix} \times & 0 & \times \\ \cdot & 0 & \times \\ \cdot & \cdot & \times \end{pmatrix}, Z_8^{\mu-\tau} \rightarrow A_1 : \begin{pmatrix} 0 & 0 & \times \\ \cdot & \times & \times \\ \cdot & \cdot & \times \end{pmatrix}$$



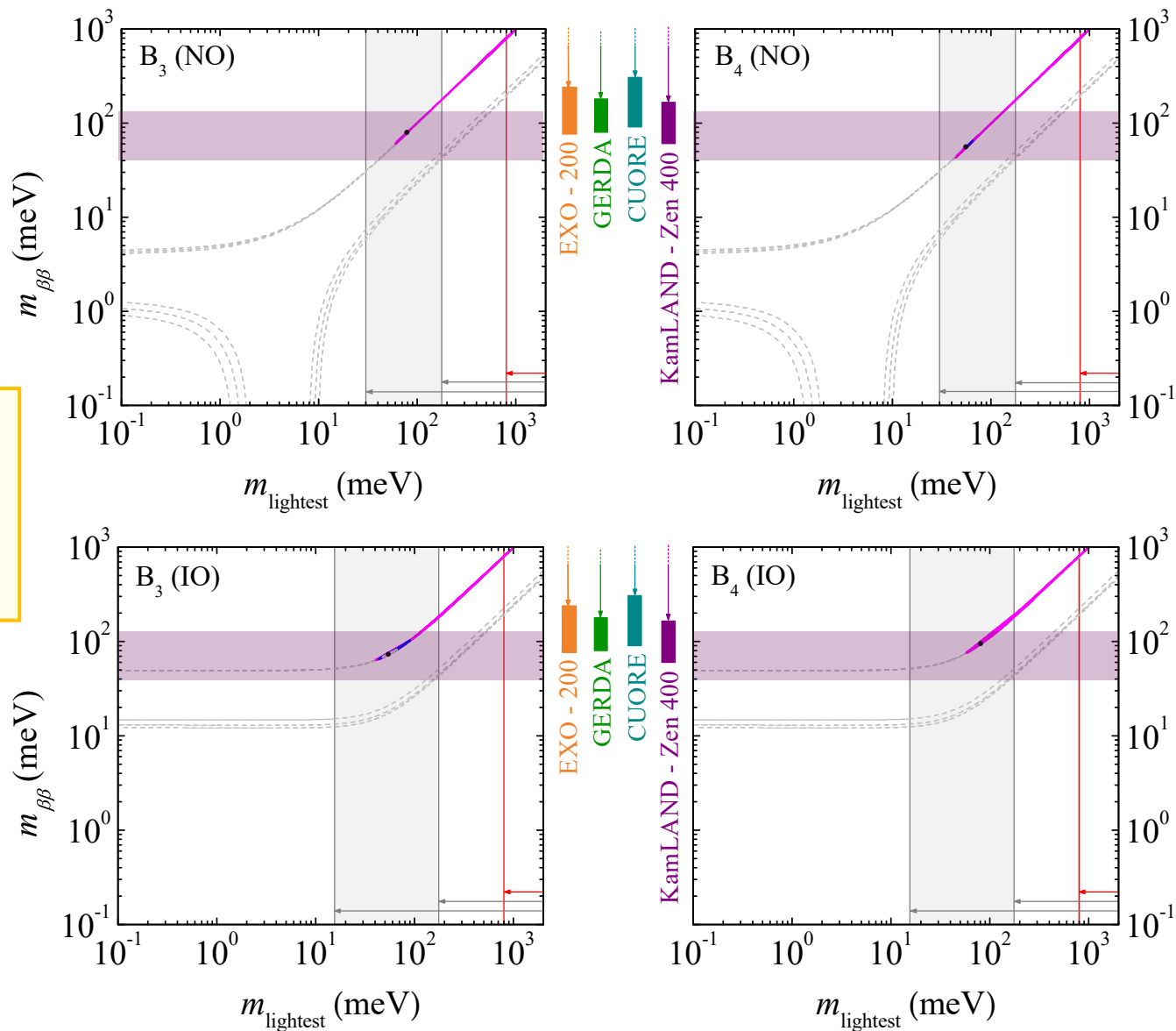


➤ Case A1 **is compatible** with the current bounds coming from cosmology. A **direct measurement** of neutrino mass by KATRIN **would exclude this case.**

➤ For the remaining cases, the lower bound on the **lightest neutrino mass** in tension with bounds from cosmology



Flavour and DM in the S-ST2 model



Best present limit from

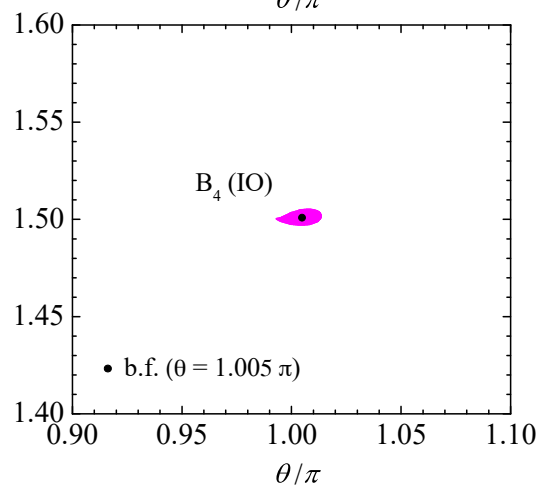
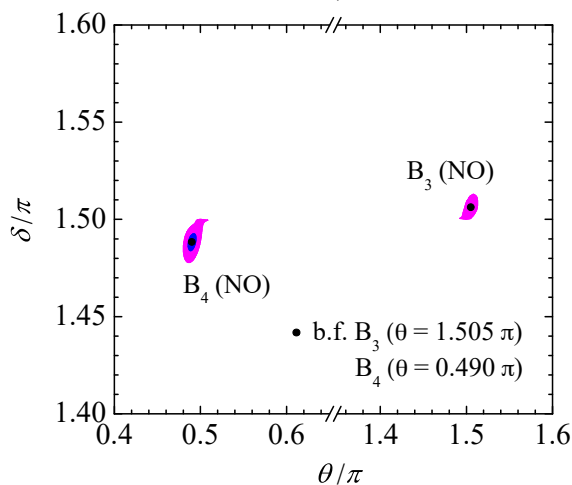
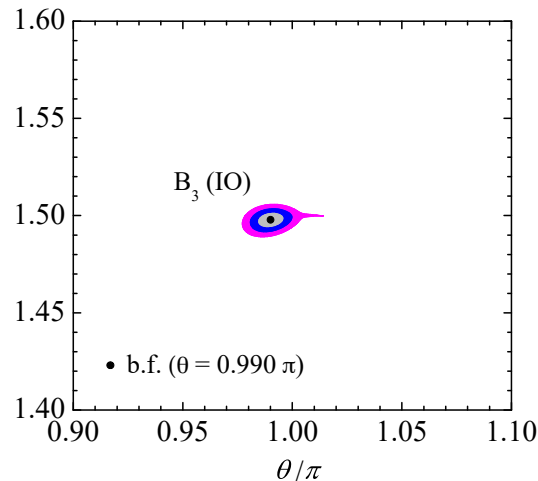
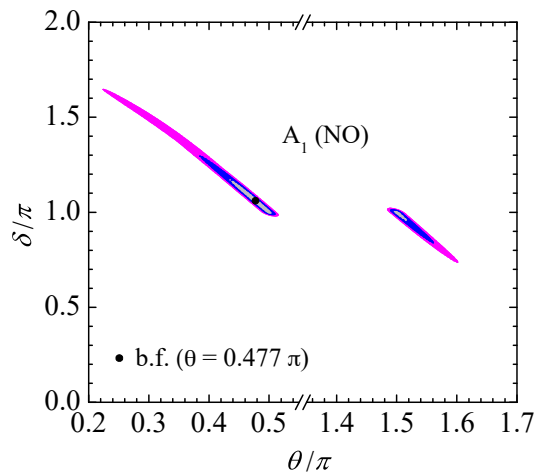
KamLAND-Zen

$$m_{\beta\beta} < 0.036 - 0.156$$

[Abe et al.; PRL 130 \(2023\) 051801](#)

The σ VEV phase is determined by **neutrino data !!!**

$$\langle \sigma \rangle = u e^{i\theta} \rightarrow \cotan \theta = -\frac{1}{\sin \xi} \left[\cos \xi - \frac{\left| \left(\widehat{\mathbf{M}}_\nu \right)_{12} \right|^2}{\left| \left(\widehat{\mathbf{M}}_\nu \right)_{11} \right| \left| \left(\widehat{\mathbf{M}}_\nu \right)_{22} \right|} \mathcal{R} \right]$$

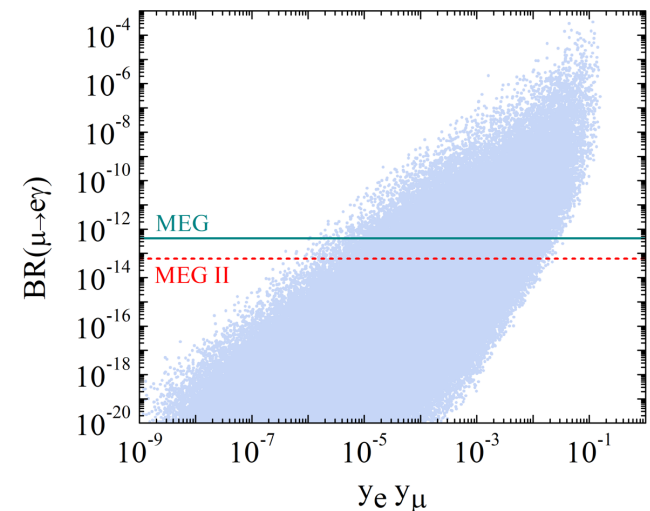


LFV radiative and 3-body decays

$$\mathbf{Y}_f^1 = \begin{pmatrix} y_e \\ 0 \\ 0 \end{pmatrix}, \mathbf{Y}_f^2 = \begin{pmatrix} 0 \\ y_\mu \\ 0 \end{pmatrix}$$

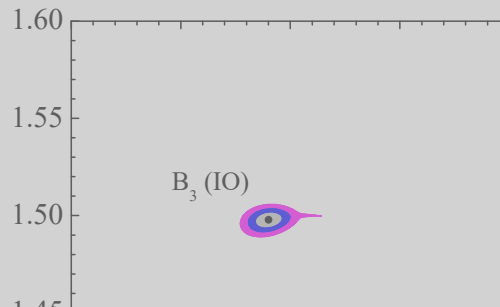
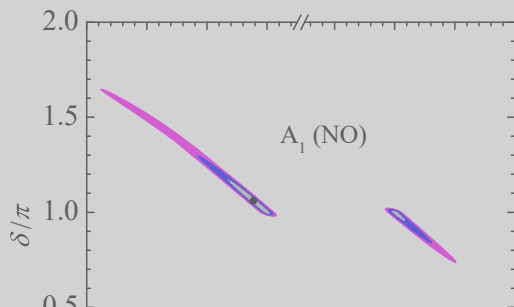
$$\mathbf{Y}_\Delta = \begin{pmatrix} y_1 & 0 & 0 \\ 0 & 0 & y_2 \\ 0 & y_2 & 0 \end{pmatrix} e^{-i\theta}$$

Contributions to LFV processes induced by dark fields and scalars from the triplet



The σ VEV phase is determined by **neutrino data !!!**

$$\langle \sigma \rangle = u e^{i\theta} \rightarrow \cotan \theta = -\frac{1}{\sin \xi} \left[\cos \xi - \frac{\left| \left(\widehat{M}_\nu \right)_{12} \right|^2}{\left| \left(\widehat{M}_\nu \right)_{11} \right| \left| \left(\widehat{M}_\nu \right)_{22} \right|} \mathcal{R} \right]$$

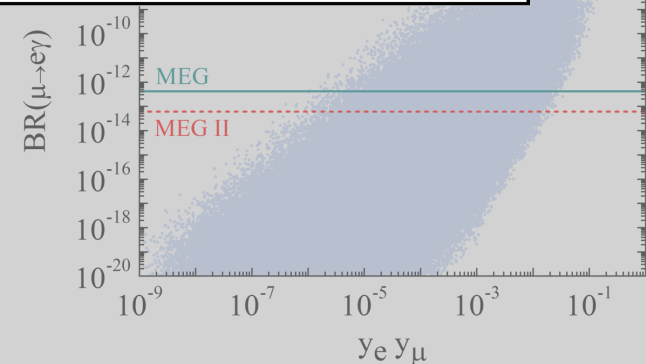
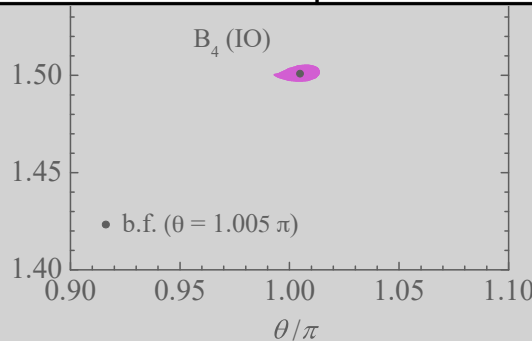
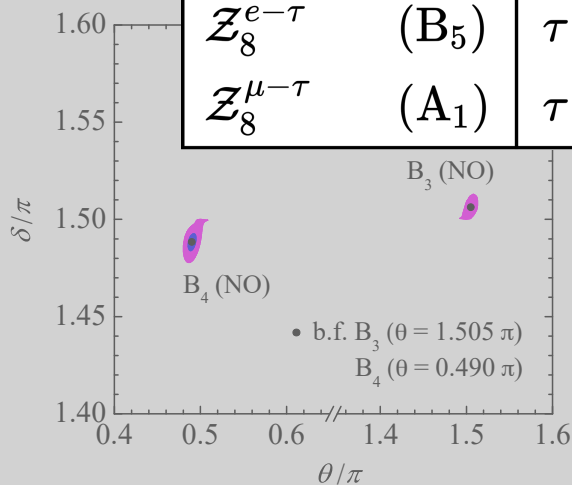


LFV radiative and 3-body decays

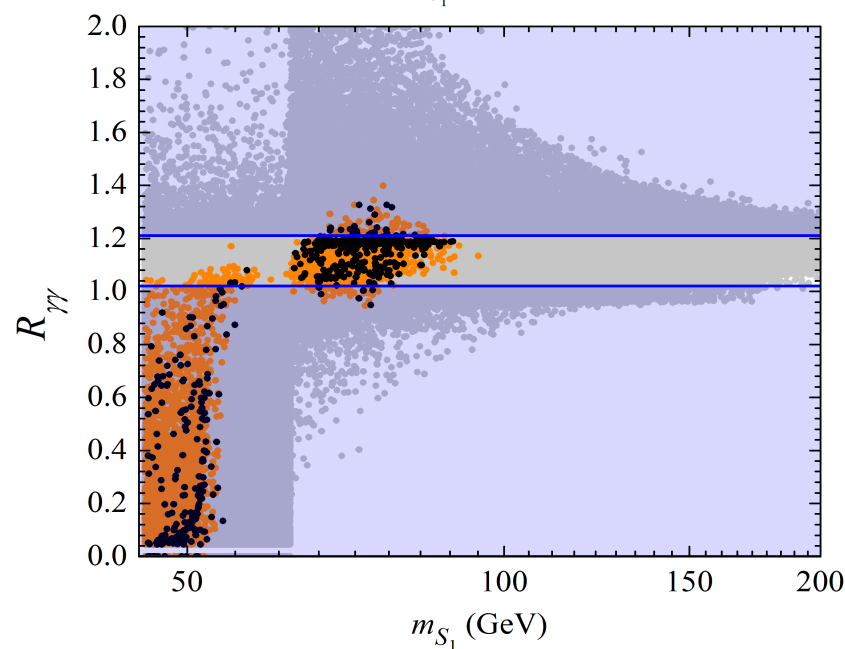
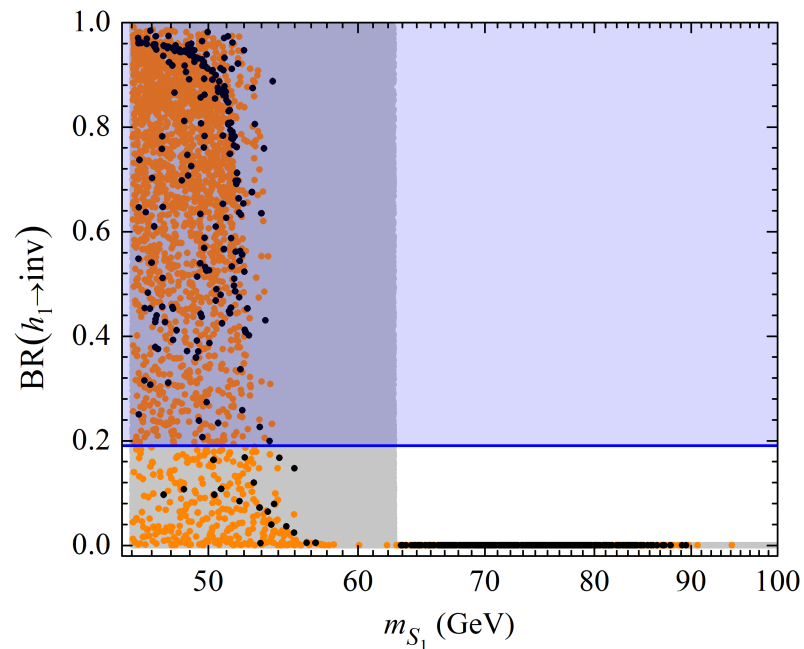
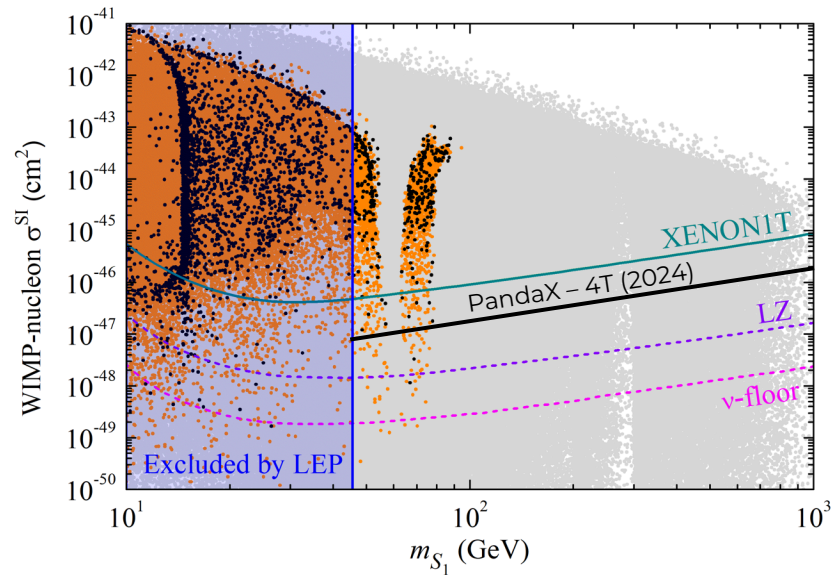
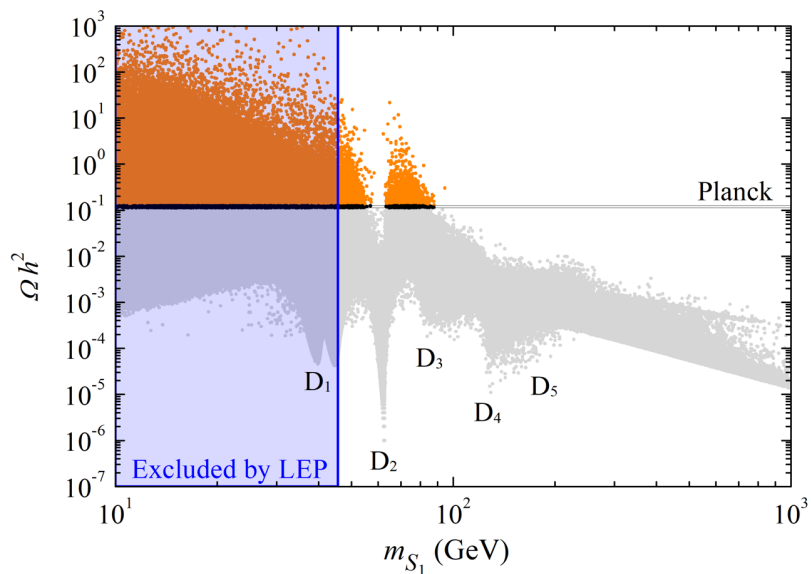
$$\mathbf{Y}_f^1 = \begin{pmatrix} y_e \\ 0 \\ 0 \end{pmatrix}, \mathbf{Y}_f^2 = \begin{pmatrix} 0 \\ y_\mu \\ 0 \end{pmatrix}$$

$$\mathbf{Y}_\Delta = \begin{pmatrix} y_1 & 0 & 0 \\ 0 & 0 & y_2 \\ 0 & y_2 & 0 \end{pmatrix} e^{-i\theta}$$

Cases	Type-II seesaw	Scotogenic
$\mathcal{Z}_8^{e-\mu}$ (B ₄)	$\tau^- \rightarrow \mu^+ e^- e^-$	$\mu \rightarrow e\gamma, \mu \rightarrow 3e, \mu - e$ conversion
$\mathcal{Z}_8^{e-\tau}$ (B ₅)	$\tau^- \rightarrow \mu^+ e^- e^-$	$\tau \rightarrow e\gamma, \tau \rightarrow 3e$
$\mathcal{Z}_8^{\mu-\tau}$ (A ₁)	$\tau^- \rightarrow e^+ \mu^- \mu^-$	$\tau \rightarrow \mu\gamma, \tau \rightarrow 3\mu$



Flavour and DM in the S-ST2 model



PHYSICAL REVIEW D **108**, 095003 (2023)

Dark-sector seeded solution to the strong CP problem

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²AHEP Group, Institut de Física Corpuscular–CSIC/Universitat de València,
Parc Científic de Paterna. C/ Catedrático José Beltrán, 2 E-46980 Paterna (Valencia), Spain

☉ (Received 13 March 2023; accepted 27 September 2023; published 2 November 2023)

$$-\mathcal{L}_{\text{Yuk}} \supset \mathbf{Y}_u \bar{q}_L \tilde{\Phi} u_R + \mathbf{Y}_d \bar{q}_L \Phi d_R \\ + \mathbf{Y}_\xi \overline{D_{2L}} d_R \xi + \mathbf{Y}_\chi \overline{D_{1L}} d_R \chi^* + \text{H.c.}$$

$$-\mathcal{L}_{\text{Yuk}} \supset y_\chi \overline{B_L} D_{2R} \chi + y_\xi \overline{B_L} D_{1R} \xi^* \\ + y'_\chi \overline{D_{2L}} B_R \chi^* + y'_\xi \overline{D_{1L}} B_R \xi + \text{H.c.}$$

$$-\mathcal{L}_{\text{mass}} = m_B \overline{B_L} B_R + m_{D_{1,2}} \overline{D_{1,2L}} D_{1,2R} + \text{H.c.}$$

Tree-level quark mass matrix

$$\mathcal{M}_d^{(0)} = \begin{pmatrix} \mathbf{M}_d & 0 \\ 0 & m_B \end{pmatrix}$$

$$\bar{\theta} = \arg[\det(\mathbf{M}_u)] + \arg[\det(\mathcal{M}_d)] = 0$$

and CKM is real

	Fields	G_{SM}	$Z_8 \rightarrow Z_2$
Fermions	q_L	$(\mathbf{3}, \mathbf{2}, 1/6)$	$\omega^2 \rightarrow +$
	u_R	$(\mathbf{3}, \mathbf{1}, 2/3)$	$\omega^2 \rightarrow +$
	d_R	$(\mathbf{3}, \mathbf{1}, -1/3)$	$\omega^2 \rightarrow +$
	$B_{L,R}$	$(\mathbf{3}, \mathbf{1}, -1/3)$	$\omega^6 \rightarrow +$
	$D_{1L,1R}$	$(\mathbf{3}, \mathbf{1}, -1/3)$	$\omega^7 \rightarrow -$
	$D_{2L,2R}$	$(\mathbf{3}, \mathbf{1}, -1/3)$	$\omega^3 \rightarrow -$
Scalars	Φ	$(\mathbf{1}, \mathbf{2}, 1/2)$	$1 \rightarrow +$
	σ	$(1, 1, 0)$	$\omega^2 \rightarrow +$
	χ	$(1, 1, 0)$	$\omega^3 \rightarrow -$
	ξ	$(1, 1, 0)$	$\omega \rightarrow -$

@ one loop: $\mathcal{M}_d = \mathcal{M}_d^{(0)} + \Delta \mathcal{M}_d$

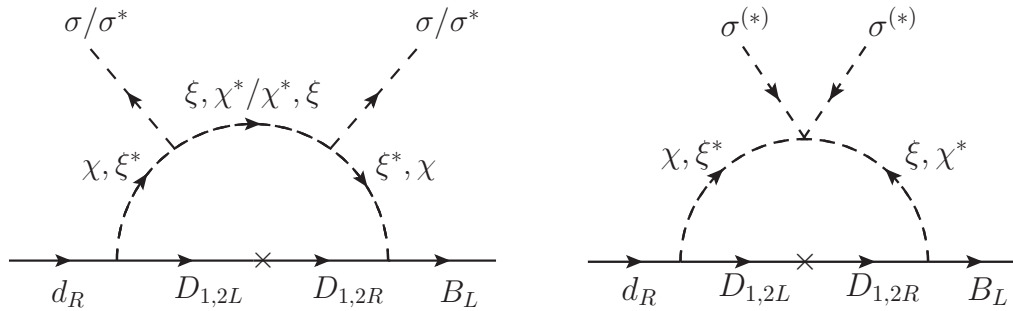
$$\Delta \mathcal{M}_d^{(1)} = \begin{pmatrix} 0 & 0 \\ \Delta \mathbf{M}_{Bd} & \Delta m_B \end{pmatrix}$$

In terms of operators:

$$\Delta \mathcal{M}_d^{(1)} = \begin{pmatrix} 0 & 0 \\ \overline{B_L} d_R \sigma^{(*)2} & \overline{B_L} B_R (\Phi^\dagger \Phi), \overline{B_L} B_R |\sigma|^2 \end{pmatrix}$$

COMPLEX $\Delta \mathbf{M}_{Bd}$ \longrightarrow COMPLEX CKM!

$$\bar{\theta} = \arg[\det(\mathbf{M}_u)] + \arg[\det(\mathcal{M}_d)] = 0 \quad \text{@ one loop}$$

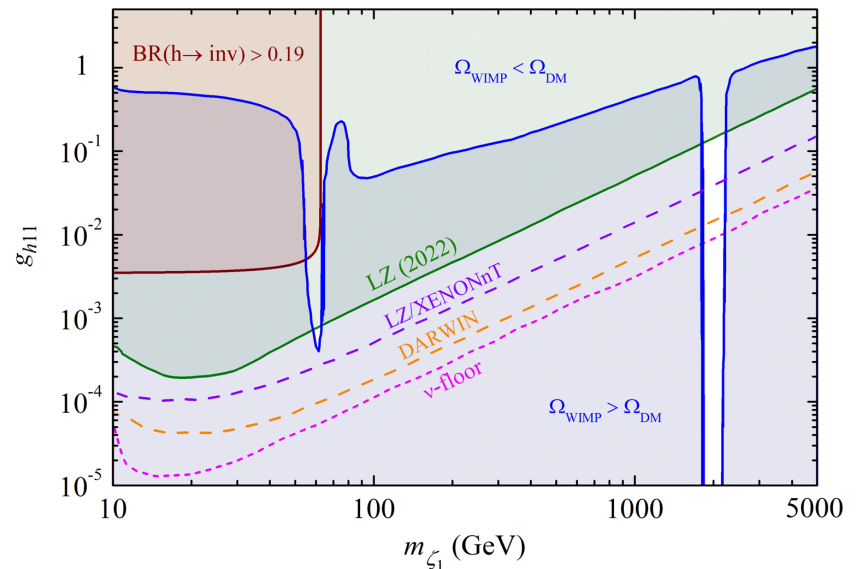
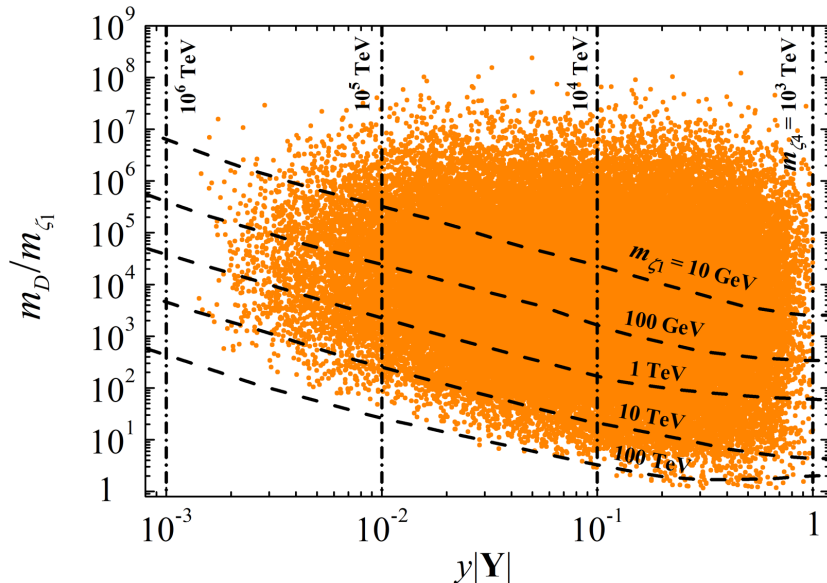


ONE-LOOP CONTRIBUTIONS TO $\Delta\mathbf{M}_{Bd}$:

$$|\Delta\mathbf{M}_{Bd}| \sim \frac{1}{16\pi^2} \lambda_{\sigma\zeta\zeta} |\mathbf{Y}_\zeta| y_\zeta \frac{v_\sigma^2}{m_\zeta^2} m_D$$

$$|\Delta\mathbf{M}_{Bd}| \sim \frac{1}{16\pi^2} |\mathbf{Y}_\zeta| y_\zeta \frac{\mu_\zeta^2}{m_\zeta^2} \frac{v_\sigma^2}{m_\zeta^2} m_D$$

LIGHT-QUARK MASSES: $\mathbf{M}_{\text{light}}^2 \simeq \mathbf{M}_d \mathbf{M}_d^T - \frac{\mathbf{M}_d \Delta\mathbf{M}_{Bd}^\dagger \Delta\mathbf{M}_{Bd} \mathbf{M}_d^T}{\tilde{m}_B^2}$





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Scalar-singlet assisted leptogenesis with CP violation from the vacuum

D. M. Barreiros,^a H. B. Câmara,^a R. G. Felipe^{b,a} and F. R. Joaquim^a

CPV needed for Leptogenesis to work is communicated to the RH neutrino sector at **high scales** through the VEVs of scalar singlets:

$$u_i > \Lambda_{\text{Leptog.}} \gg v$$

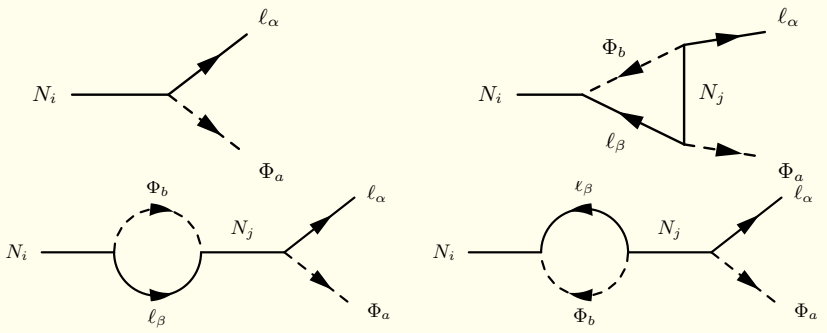
HEAVY SCALAR SINGLETS: $S_k = \frac{1}{\sqrt{2}} (u_k e^{i\theta_k} + S_{Rk} + iS_{Ik})$

LIGHT SCALAR DOUBLETS: $\Phi_a = \begin{pmatrix} \phi_a^+ \\ \phi_a^0 \end{pmatrix} = \frac{1}{\sqrt{2}} \begin{pmatrix} \sqrt{2}\phi_a^+ \\ v_a e^{i\varphi_a} + \phi_{Ra}^0 + i\phi_{Ia}^0 \end{pmatrix}$

θ_k

$$-\mathcal{L}_{\text{Yuk.}} = \bar{\ell}_L \mathbf{Y}_\ell^a \Phi_a e_R + \bar{\ell}_L \mathbf{Y}_D^{a*} \tilde{\Phi}_a \nu_R + \frac{1}{2} \bar{\nu}_R (\mathbf{M}_R^0 + \mathbf{Y}_R^k S_k + \mathbf{Y}'_R S_k^*) \nu_R^c + \text{H.c.}$$

$$\mathbf{M}_R = \mathbf{M}_R^0 + \frac{u_k}{\sqrt{2}} (\mathbf{Y}_R^k e^{i\theta_k} + \mathbf{Y}'_R e^{-i\theta_k}) \xrightarrow[\text{mass basis}]{\text{In the N}} \mathbf{Y}^{a*} \bar{\ell}_L N \Phi_a : \mathbf{Y}^{a*} = \mathbf{V}_L^\dagger \mathbf{Y}_D^{a*} \mathbf{U}_R$$



$$\epsilon_{i\alpha}^a = \frac{\Gamma(N_i \rightarrow \Phi_a l_\alpha) - \Gamma(N_i \rightarrow \Phi_a^\dagger \bar{l}_\alpha)}{\sum_{\beta=e,\mu,\tau} \sum_{b=1}^{n_H} [\Gamma(N_i \rightarrow \Phi_b l_\beta) + \Gamma(N_i \rightarrow \Phi_b^\dagger \bar{l}_\beta)]}$$

Fukujita & Yanagida'86; Covi, Roulet Vissani'96;

A simple realisation with **high-energy SCPV**

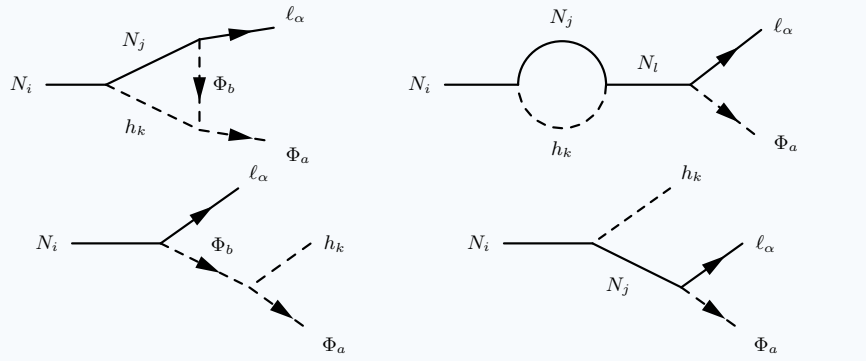
	Fields	SU(2) _L × U(1) _Y	Z ₈ ^e	Z ₈ ^μ	Z ₈ ^τ
Fermions	l _{eL}	(2, -1/2)	ω ⁵	ω ⁷	ω ⁶
	l _{μL}	(2, -1/2)	ω ⁷	ω ⁵	ω ⁵
	l _{τL}	(2, -1/2)	ω ⁶	ω ⁶	ω ⁷
	e _R	(1, -1)	ω ⁴	ω ⁷	ω ⁶
	μ _R	(1, -1)	ω ⁷	ω ⁴	ω ⁴
	τ _R	(1, -1)	ω ⁶	ω ⁶	ω ⁷
	ν _{R1}	(1, 0)	ω ⁶	ω ⁶	ω ⁶
	ν _{R2}	(1, 0)	1	1	1
Scalars	Φ ₁	(2, 1/2)		1	
	Φ ₂	(2, 1/2)		ω	
	S	(1, 0)		ω ²	

Only one case compatible with neutrino data: Z₈^μ

$$\mathbf{Y}_\ell^1 = \begin{pmatrix} y_1 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & y_4 \end{pmatrix}, \quad \mathbf{Y}_\ell^2 = \begin{pmatrix} 0 & 0 & y_2 \\ 0 & y_3 & 0 \\ 0 & 0 & 0 \end{pmatrix}, \quad \mathbf{Y}_D^1 = \begin{pmatrix} 0 & 0 \\ 0 & 0 \\ y_{D_3} & 0 \end{pmatrix},$$

$$\mathbf{M}_R^0 = \begin{pmatrix} 0 & 0 \\ \cdot & m_R \end{pmatrix}, \quad \mathbf{Y}'_R = \begin{pmatrix} 0 & y_{R_S} \\ \cdot & 0 \end{pmatrix}$$

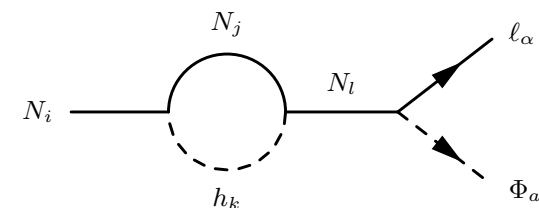
In this minimal setup, only the **new wavefunction diagrams** are relevant



$$\epsilon_{i\alpha}^a \simeq \frac{\sum_{k=1}^{2n_S} [\Gamma(N_i \rightarrow \Phi_a l_\alpha h_k) - \Gamma(N_i \rightarrow \Phi_a^\dagger \bar{l}_\alpha h_k)]}{\sum_{\beta=e,\mu,\tau} \sum_{b=1}^{n_H} [\Gamma(N_i \rightarrow \Phi_b l_\beta) + \Gamma(N_i \rightarrow \Phi_b^\dagger \bar{l}_\beta)]}$$

M. Le Dall and A. Ritz, PRD 90 (2014)

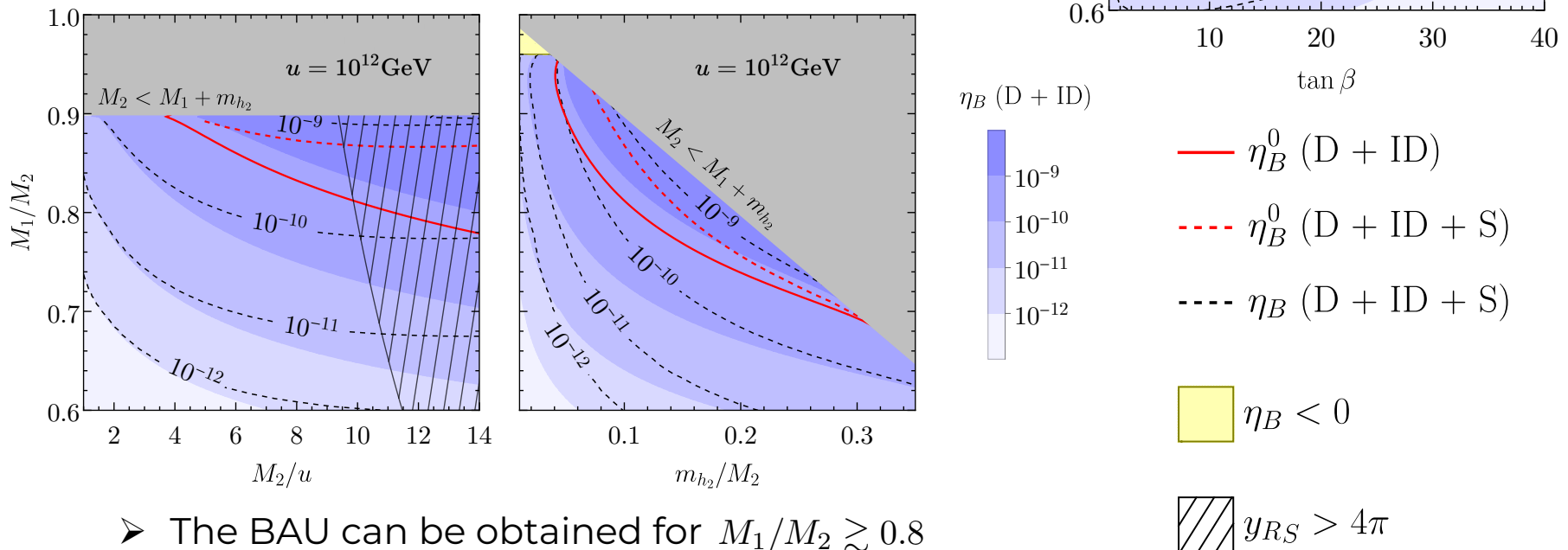
H I G G S
P O R T A L
L G



➤ Best-fit neutrino parameters for IO

Case	θ_{12} ($^\circ$)	θ_{13} ($^\circ$)	θ_{23} ($^\circ$)	δ/π	α/π	$m_{\beta\beta}$ (meV)	m_β (meV)	$\sum_i m_i$ (meV)
$\mathcal{Z}_8^\mu(\text{IO})$	35.48	8.60	49.62	1.88	0.92	16.6	49.2	99.7

Case	θ/π	θ_L/π	(x, y, z) (meV)
$\mathcal{Z}_8^\mu(\text{IO})$	1.89	7.29×10^{-2}	(0.325, 32.8, 0.426)



➤ The BAU can be obtained for $M_1/M_2 \gtrsim 0.8$

Conclusions

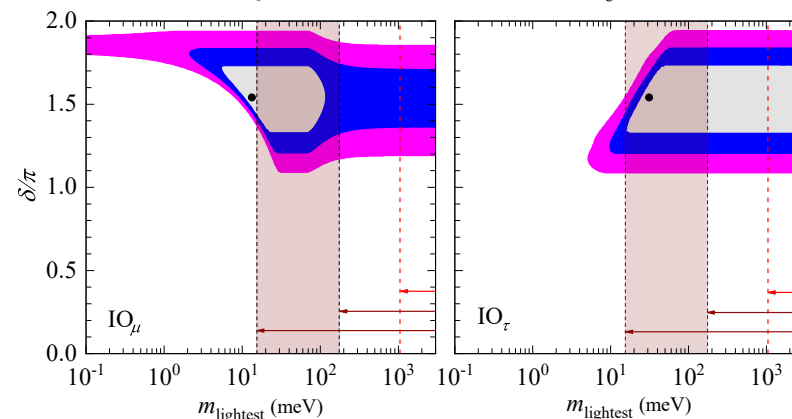
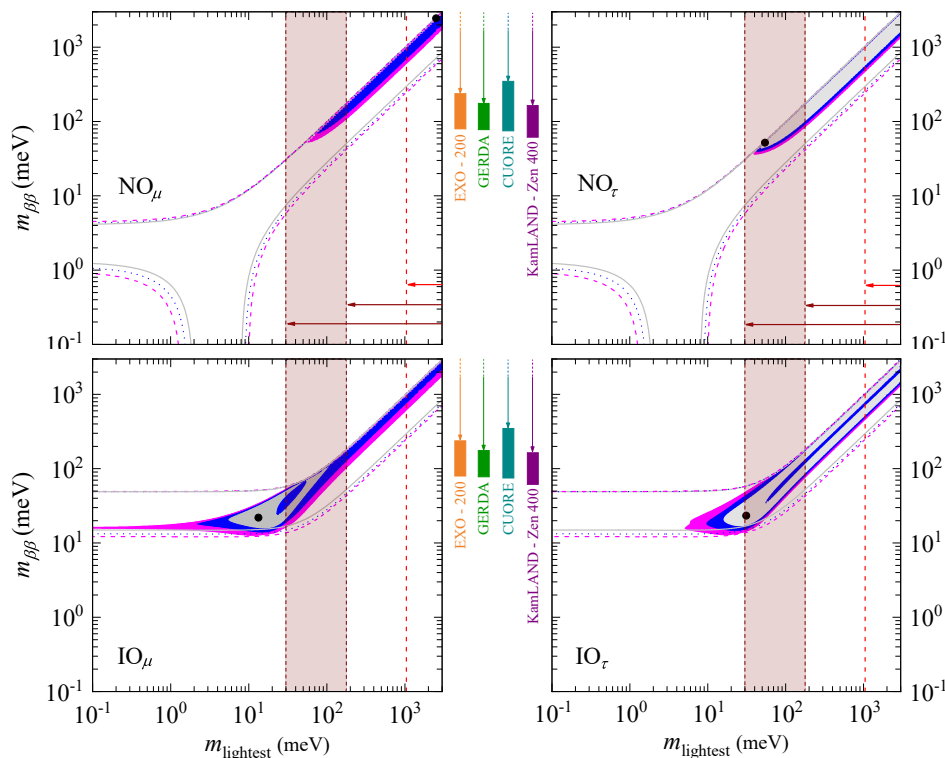
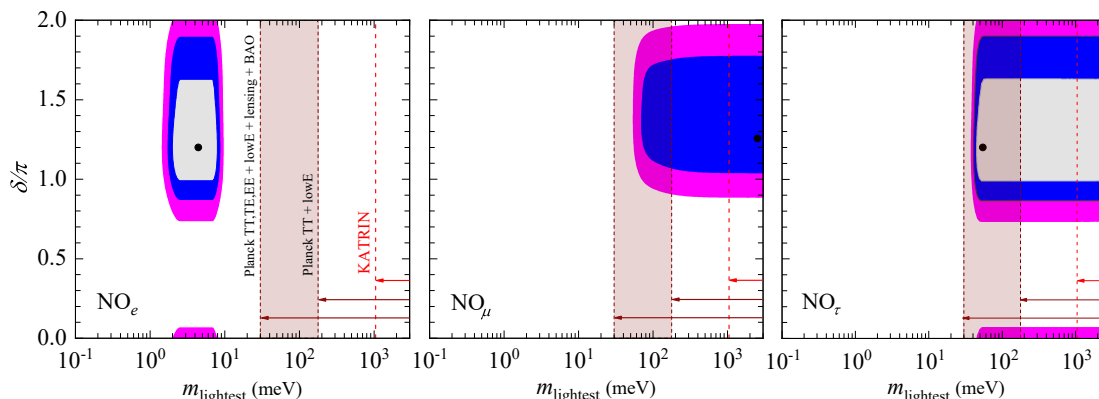
- **SCPV** is very appealing from the theory point of view.
- **In the standard paradigm** SCPV comes from doublets.
- **SCPV from the Scalar Singlet Portal** coupled to a heavy fermion sector has some nice features:
 - **Seesaw induced** low-energy CPV (e.g. scoto-seesaw)
 - Connections with Dark matter and **SCP problems**
 - **Singlet-assisted** leptogenesis
- Implementation in the framework of **low-energy** neutrino mass generation mechanisms.

Câmara and FRJ, [JHEP 05 \(2021\) 021](#)

Thank you!

Neutrinoless double beta decay

$$\beta\beta_{0\nu}$$



Lower bound on the **lightest neutrino mass** in tension with bounds from cosmology