Antinucleon-nucleus interactions with the INCL code



XSCRC2024: Cross sections for Cosmic Rays @ CERN

16-18 October 2024 - Cern



What is INCL?

INCL and CR

Antiprotons

Antineutrons

(light) Antinuclei

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What is INCL? Generalities

light particle + nucleus (~20 MeV < E_p < ~20 GeV)

INCL: Intra Nuclear Cascade of Liège



- After the cascade follows a deexcitation (usually we use ABLA)
- INCL (and ABLA) implemented in Geant4

Antinucleon-nucleus interactions with the INCL code



• INCL energy range matches CR's spectra



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Spallation, cosmic rays, meteorites, and planetology J.-C. David and I. Leya / Prog. Part. Nucl. Phys. 109 (2019) 103711

<u>conclusions</u>: models more useful/accurate than ever but era of precision...

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📥 A review paper

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conclusions: models more useful/accurate than ever but era of precision...

Recently have been asked to put pbar (AD, FAIR)

So if pbar, why not nbar?

Then antideuteron, anti-³He

for DM experiments...



nucleus interactions ne INCL code

Antiprotons

In-flight E_{at rest}=200 MeV < E < 10 GeV



Main ingredients

- Cross sections
 - Elastic
 - Annihilation
 - Production
 - Charge exchange
- Final products (types; momenta)
- Potential (p)



Main ingredients

- Annihilation nucleon (p or n)
- Position of the Annihilation
- Final products (types; momenta)

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10

12

OK

10

Antinucleon-nucleus interactions with the INCL code



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Antiprotons Results

Multiplicities ³He & ⁴He



At rest

• ³He: underestimation (< x1.5)

• ⁴He: rather good

W. Markiel et al. Nuclear Physics A 485.3 (1988), 445–460.

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Antiprotons Results

Residue production



Here, cumulative production (progenitors accounted for)

At rest

Not bad at all, is it?

(same reliabilty as in p + A)

E. F. Moser et al., Z. Phys.A – AtomicNuclei 333, 89-105 (1989)

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Mass distributions $\overline{p} + {}^{98}Mo \rightarrow Z$

Antineutrons

In-flight E_{at rest}=14 MeV < E < 10 GeV



Main ingredients

- Cross sections
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 - Annihilation
 - Production
 - Charge exchange
- Final products (types; momenta)
- Potential (n)



• Annihilation nucleon (p or n)

n

- Position of the Annihilation
- Final products (types; momenta)

« At rest »

 $E < E_{at rest} = 14 MeV$



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XSCRC2024: Cross sections for Cosmic Rays @ CERN

with the INCL code

\overline{d} , \overline{t} , $\overline{{}^{3}He}$, $\overline{{}^{4}He}$

- INCL treats d, t, ³He, ⁴He-induced reactions (and more)
- Now \bar{n} and \bar{p} -induced reactions available
- So, why not \overline{d}_{1} , \overline{t}_{1} , $\overline{{}^{3}He}_{1}$, $\overline{{}^{4}He}$ -induced reactions?

It's in progress... but at an (very) early stage with antideuteron!

First results are encouranging.



S.P. Denisov et al., Nuclear Physics B 31.2 (1971), pp. 253–260.



<u>Conclusions</u>

pbarin INCL and Geant4 (since Geant4-11.2)nbaralmost in INCL (some checks) and planned for next year in Geant4light antinucleiin progress...

• Rather good results,

- but place for improvements
 - π high multiplicities (refinement)
 - ~underestiamted
 - d overestimated; t and ³He underestimated
 - К⁰ ~ОК
 - K^{+/-} underestimated
 - Λ underestimated
- Some not-so-well-known ingredients ...

р

K

(potential, position of the annihilation, on which nucleon (n? p?) the annihilation...)



A project (NuRBS: Nuclear Reaction model improvement with Bayesian Statistics) has been funded (2024->2027) – CEA & Bern U. (and IAEA+Coruña U.)

Goals: Building tools for biasing and parameter optimisation Applying them to INCL and ABLA for several cases

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And thanks to the students

D. Zharenov O. Lourgo

(antiprotons + antineutrons) (antineutrons + antideuterons)

J. Hirtz who gave advices

and



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backup



Antiprotons Results



1400 **Reaction Cross section** 1200 Solid lines: INCL calculations 1000 dm Dashed curves: at rest normalization Cross 800 $\sigma_{reac} = \pi R^2 \left(1 + \frac{Z e^2 (m_{\bar{p}} + M_{target})}{4\pi\epsilon_0 E_{kin} R M_{target}} \right)$ reaction o 600 C12 Ne20 Al27 400 Ca40 Cu63 200 600 800 1000 400 antiProton Momentum (MeV/c)

> Antinucleon-nucleus interactions sinteractions

1200

VCCDC2024. Crace costions for Cosmic Dave @ CEDN VSCPC2021. Cross sections for Cosmic Pays @ CEPN

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preliminary results d(6.1 GeV/c per nucleon) + Ta multiplicity			
		π+/π-	р
	Exp.	5.08 ± 0.08	7.26 ± 0.16
	INCL	4.98	3.43
	bias	0.04%	51.7%

V. F. Andreyev et al., Il Nuovo Cimento A 103.8 (1990), pp. 1163–1176

S.P. Denisov et al., Nuclear Physics B 31.2 (1971), pp. 253–260.

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