

About cross sections relevant for UHECRs

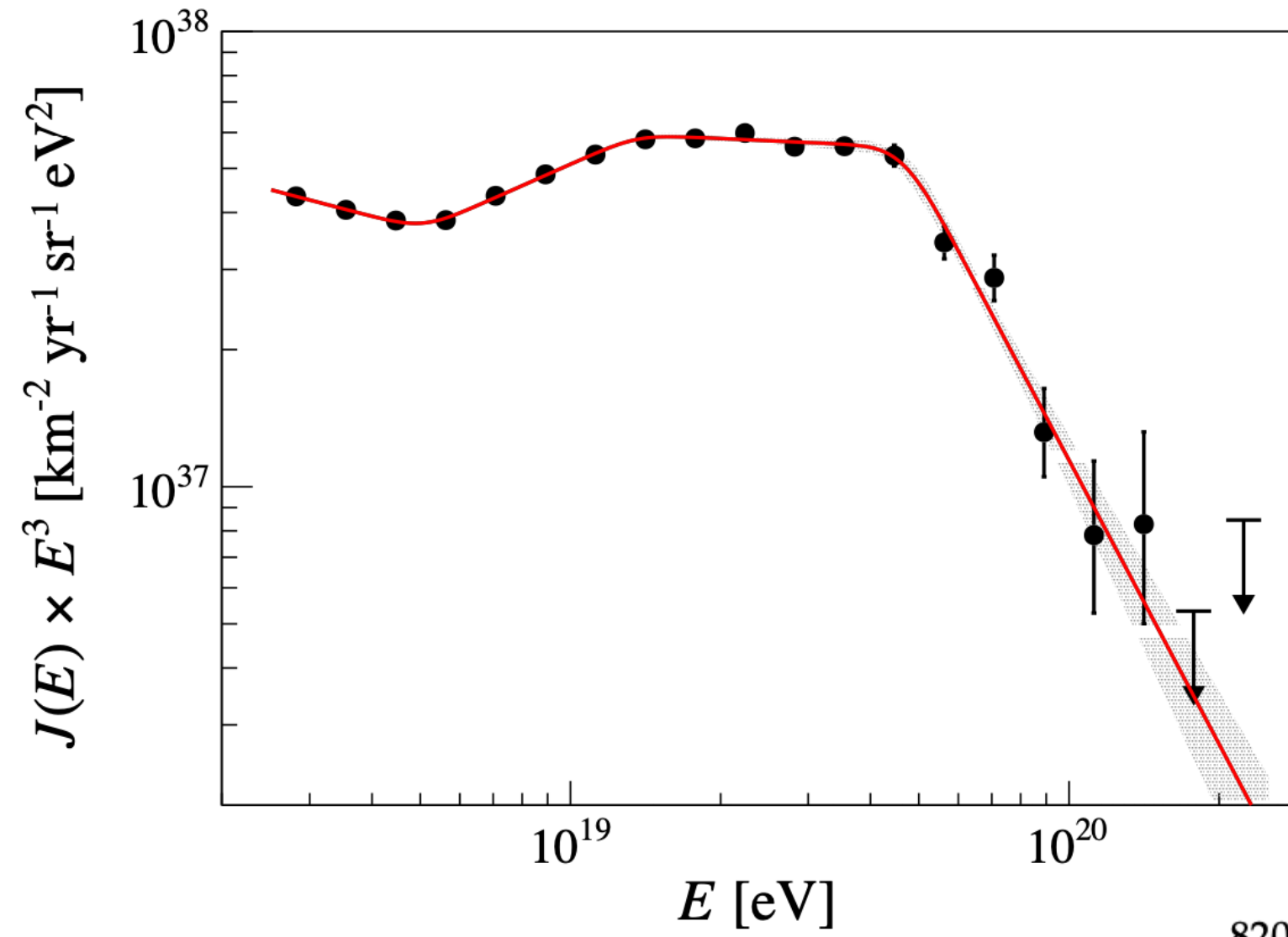
Denise Boncioli

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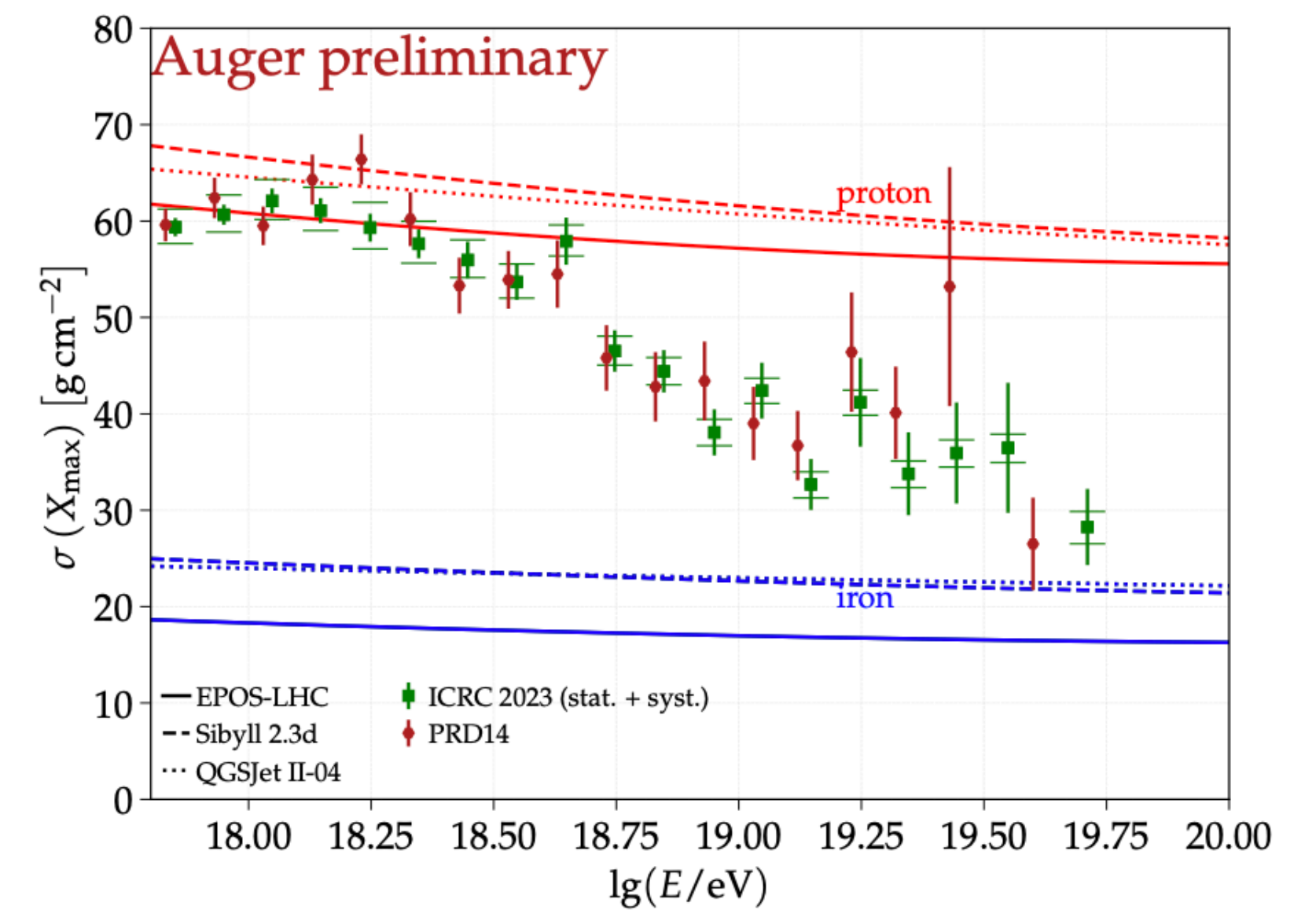
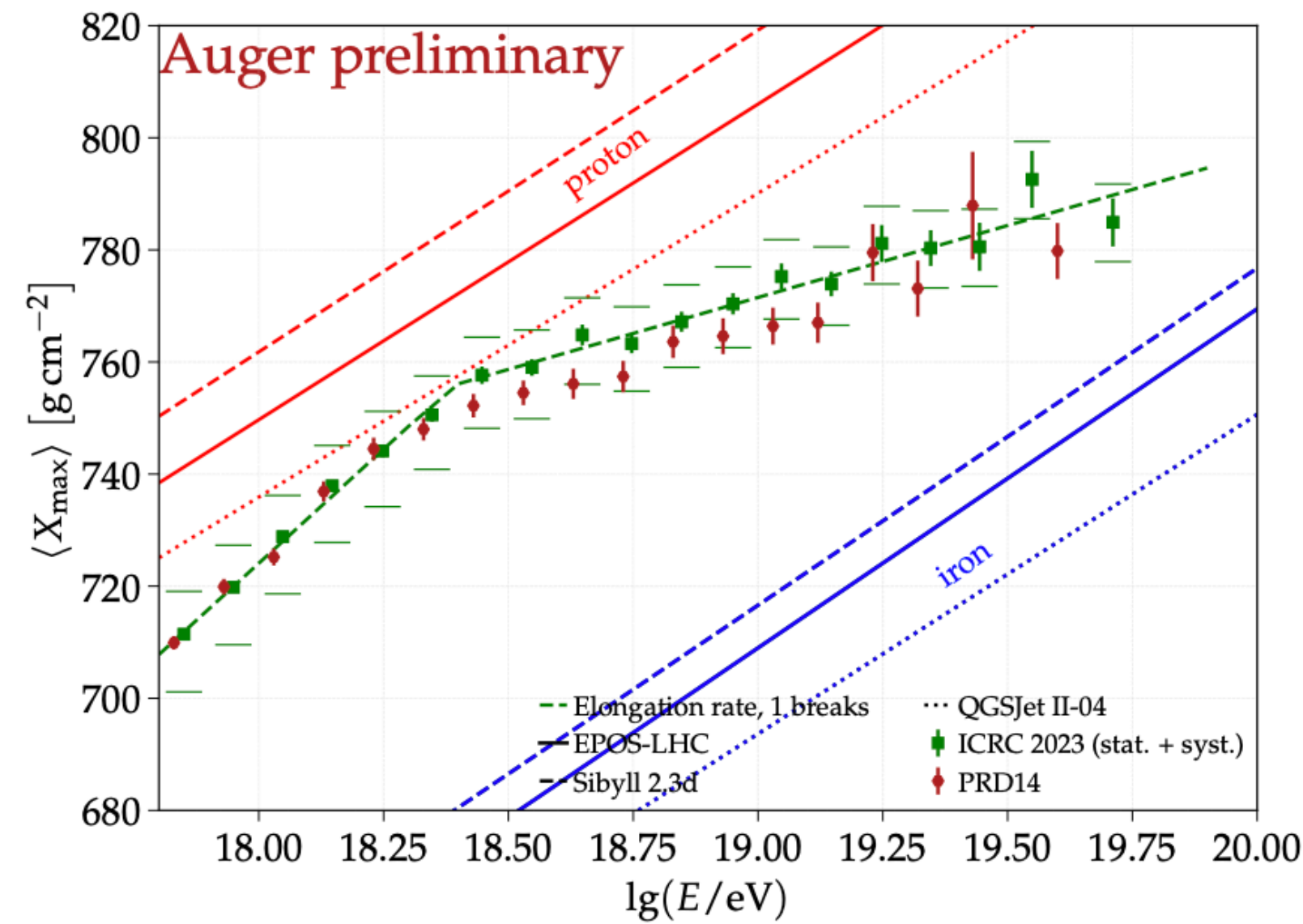
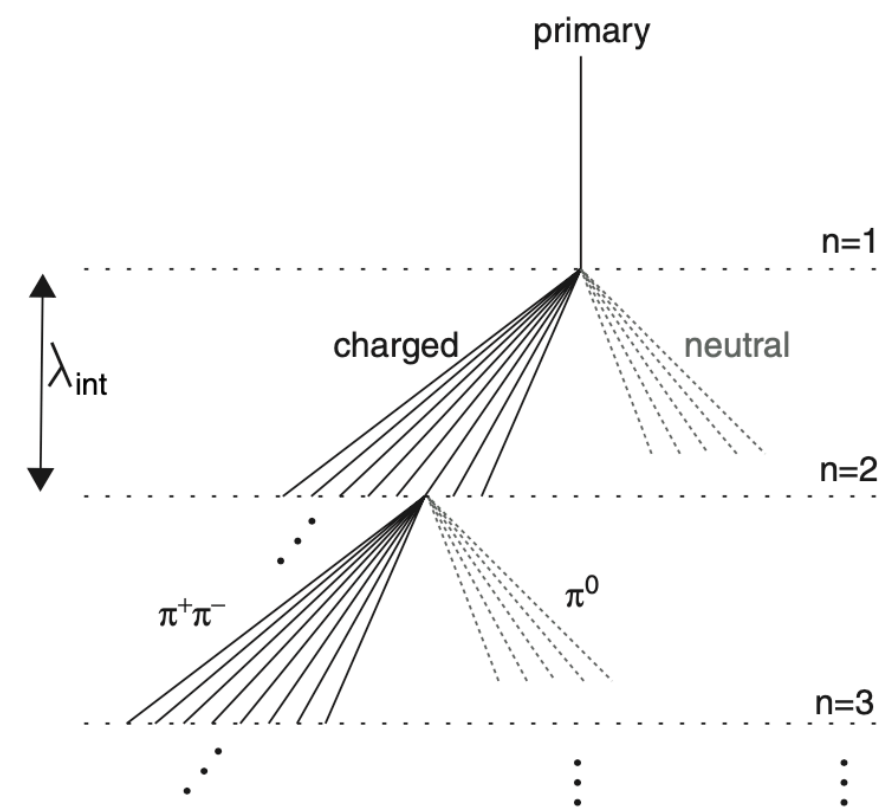
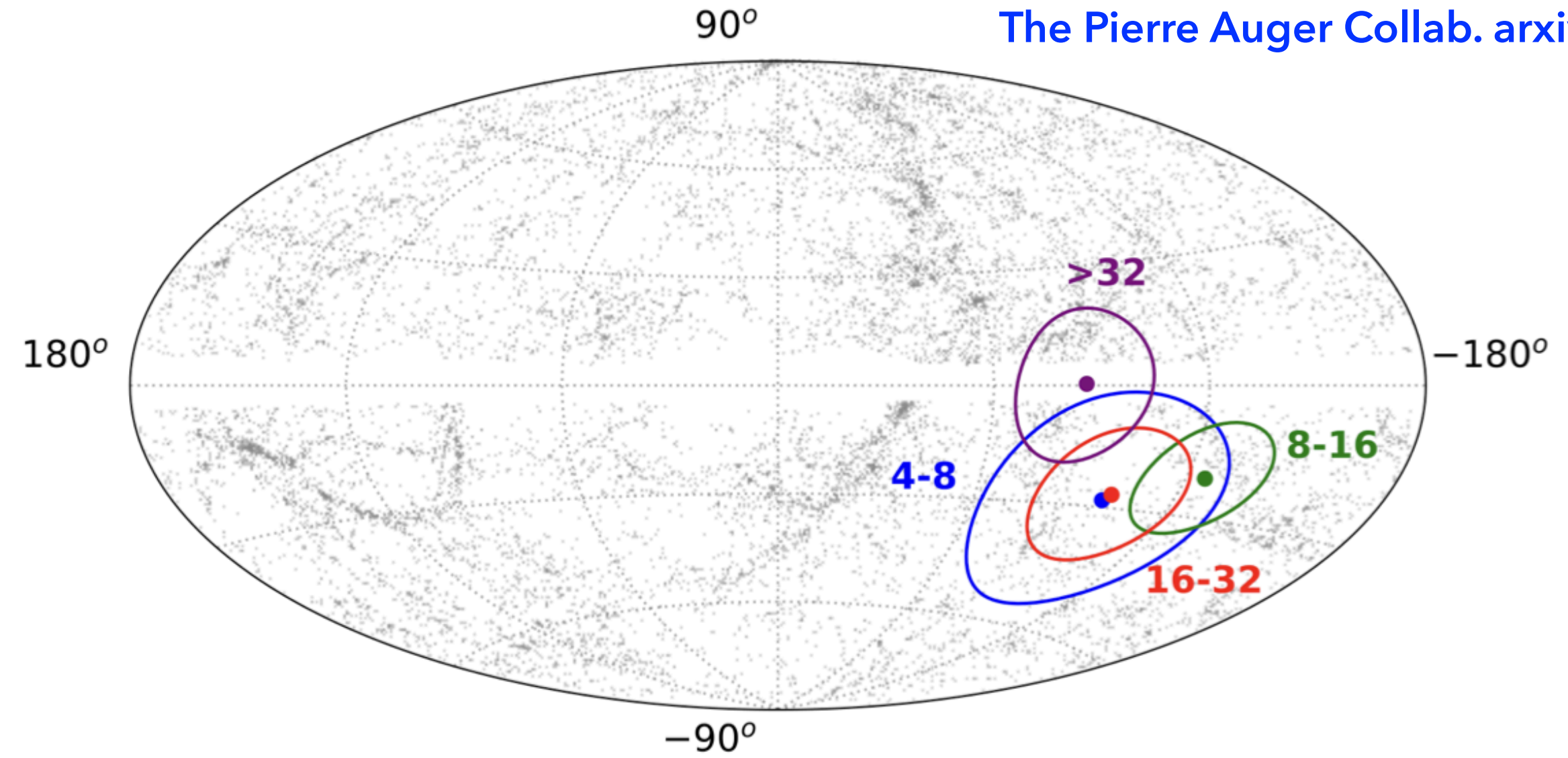
XSCRC24, 16-18 October 2024

WHY ARE WE INTERESTED IN NUCLEAR REACTIONS IN THE CONTEXT OF UHECRS?

The Pierre Auger Collab. PRD&PRL 2020

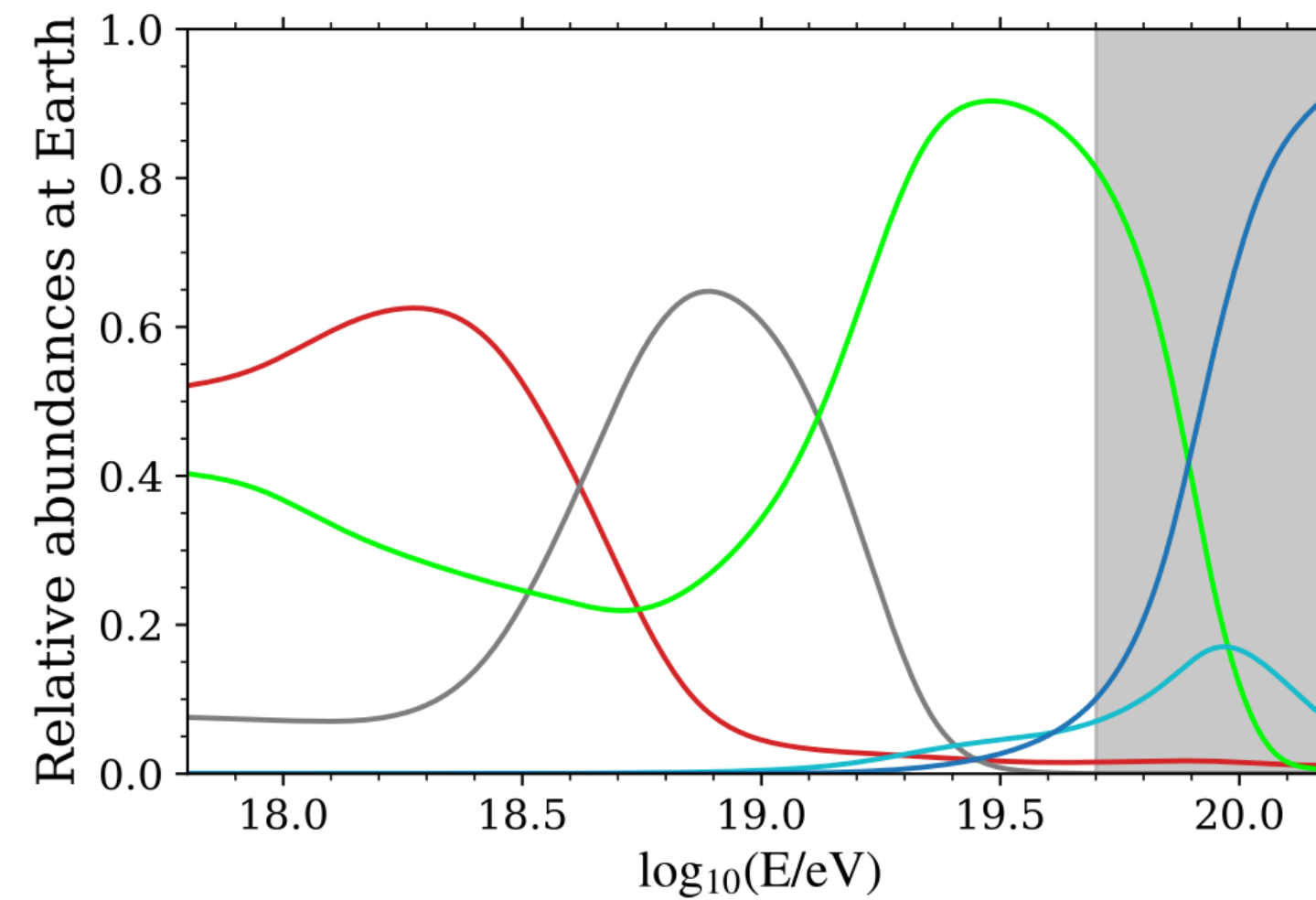
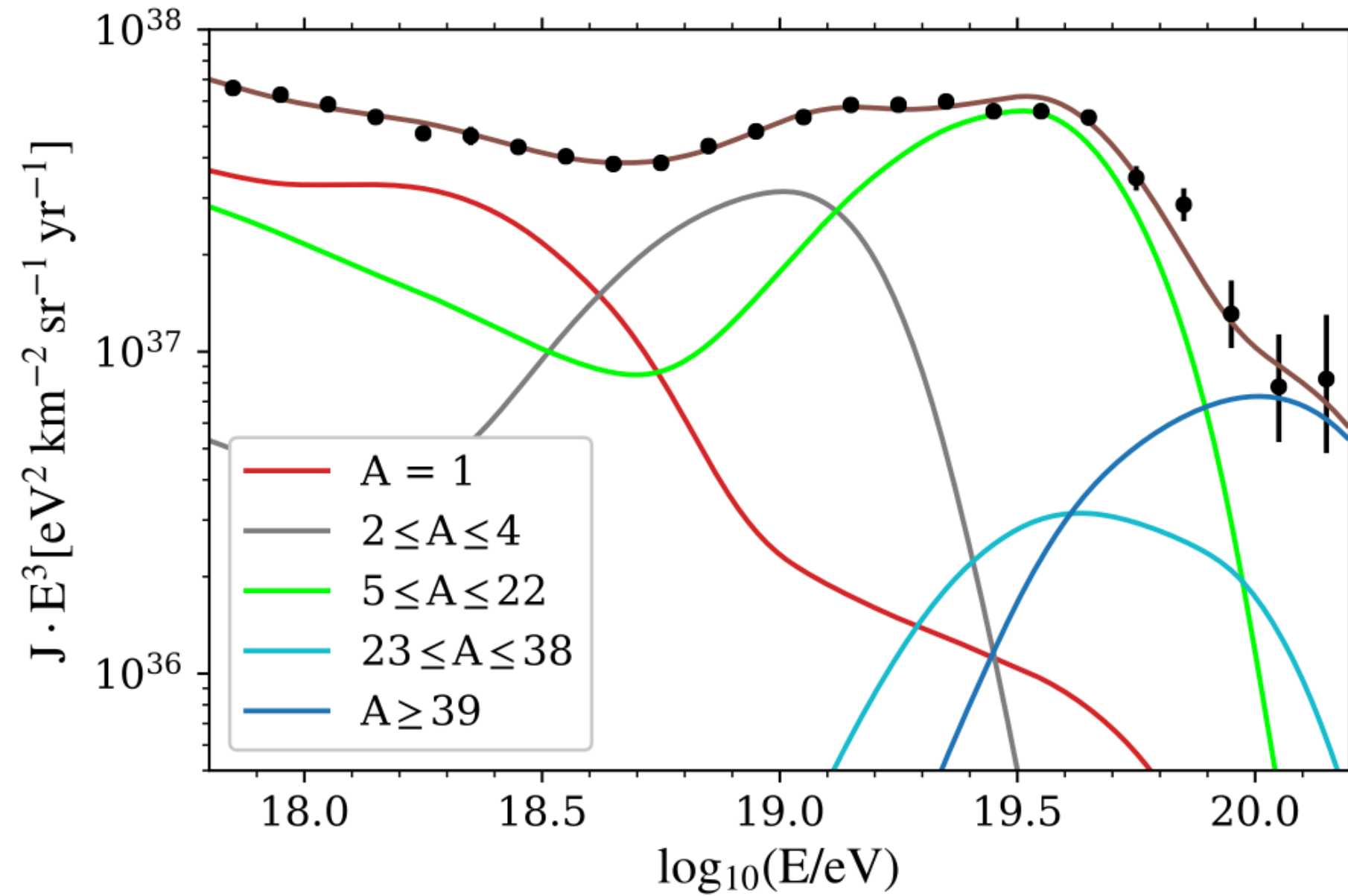
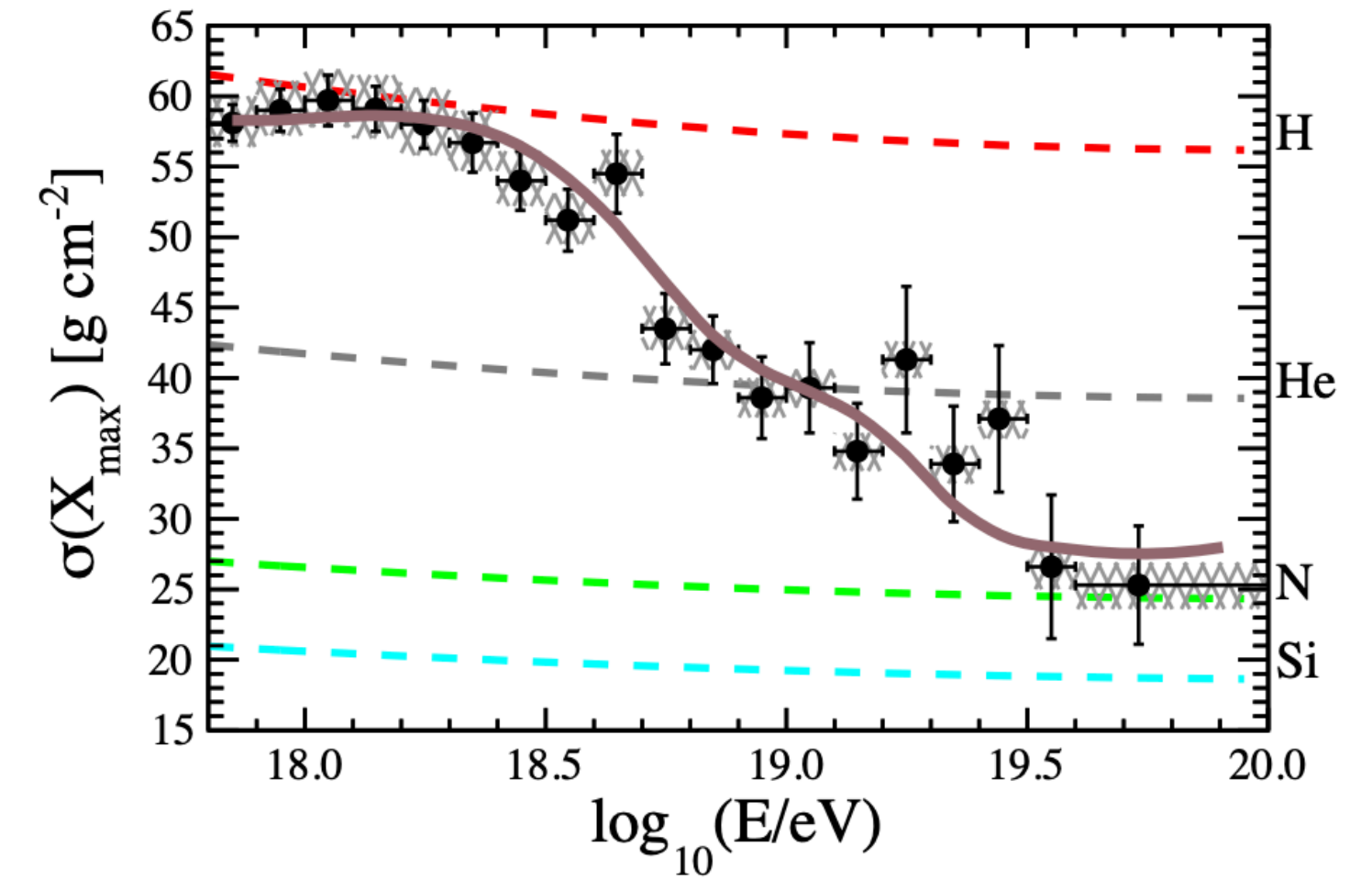
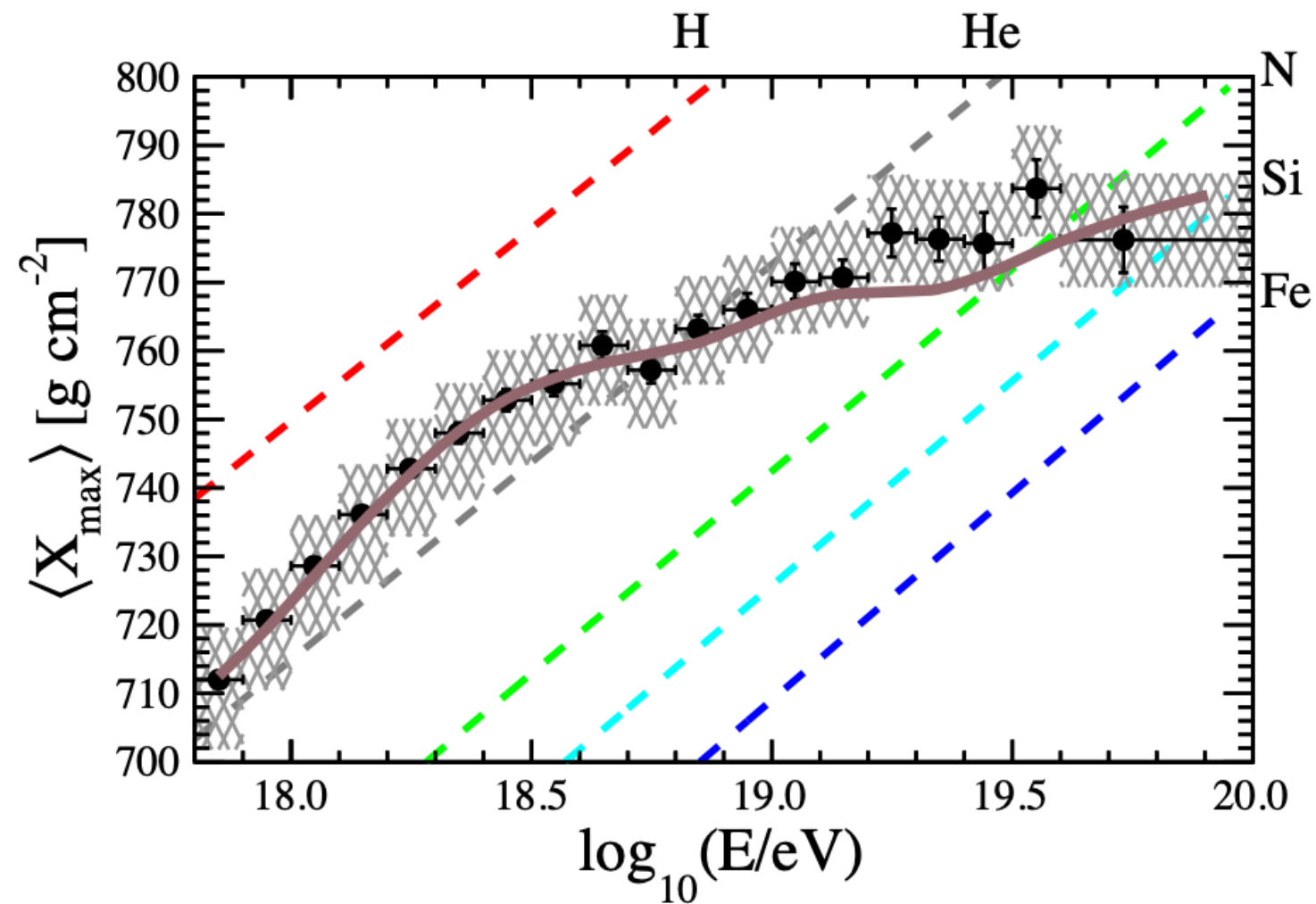
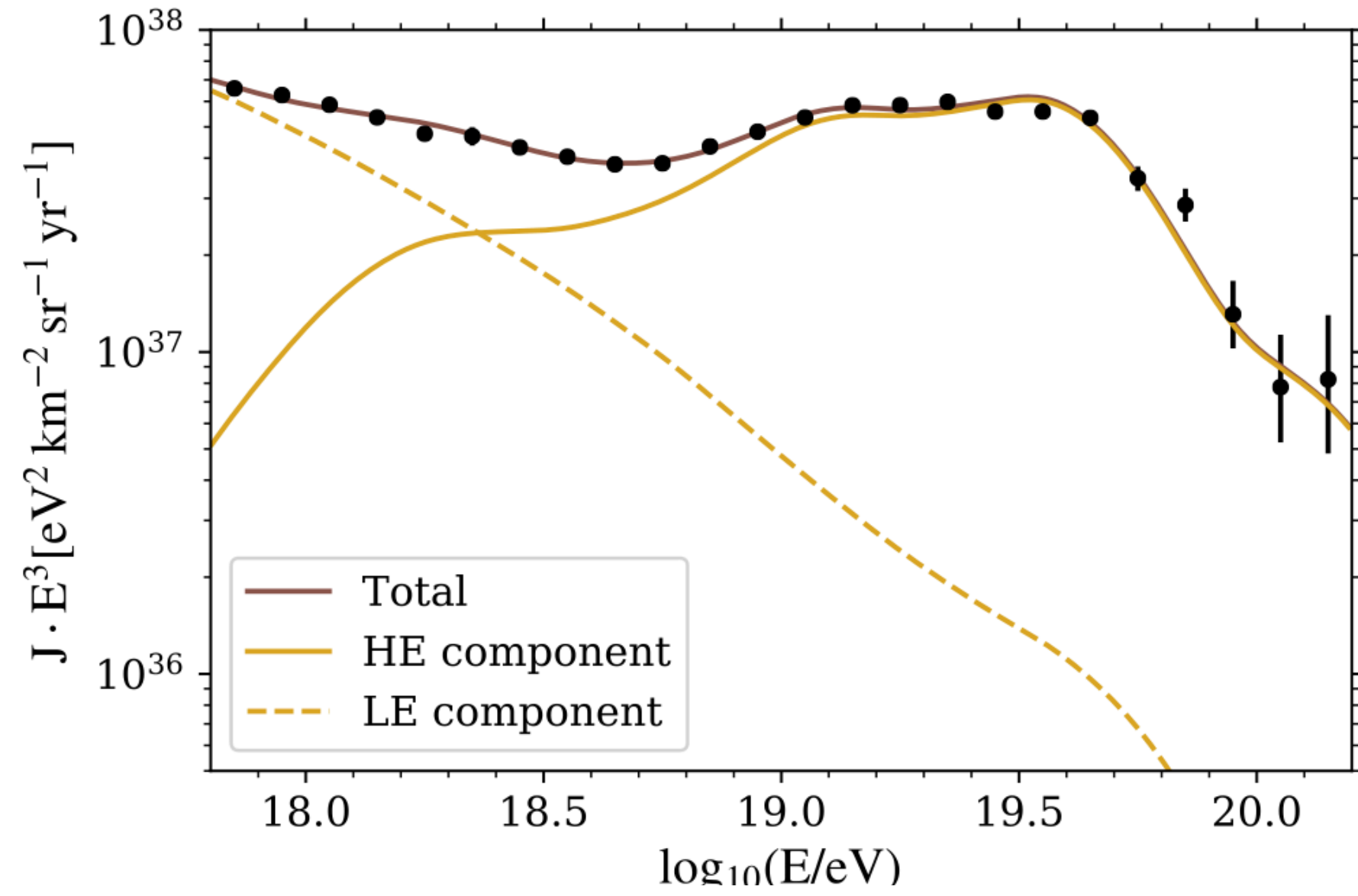


The Pierre Auger Collab. arxiv:2408.05292



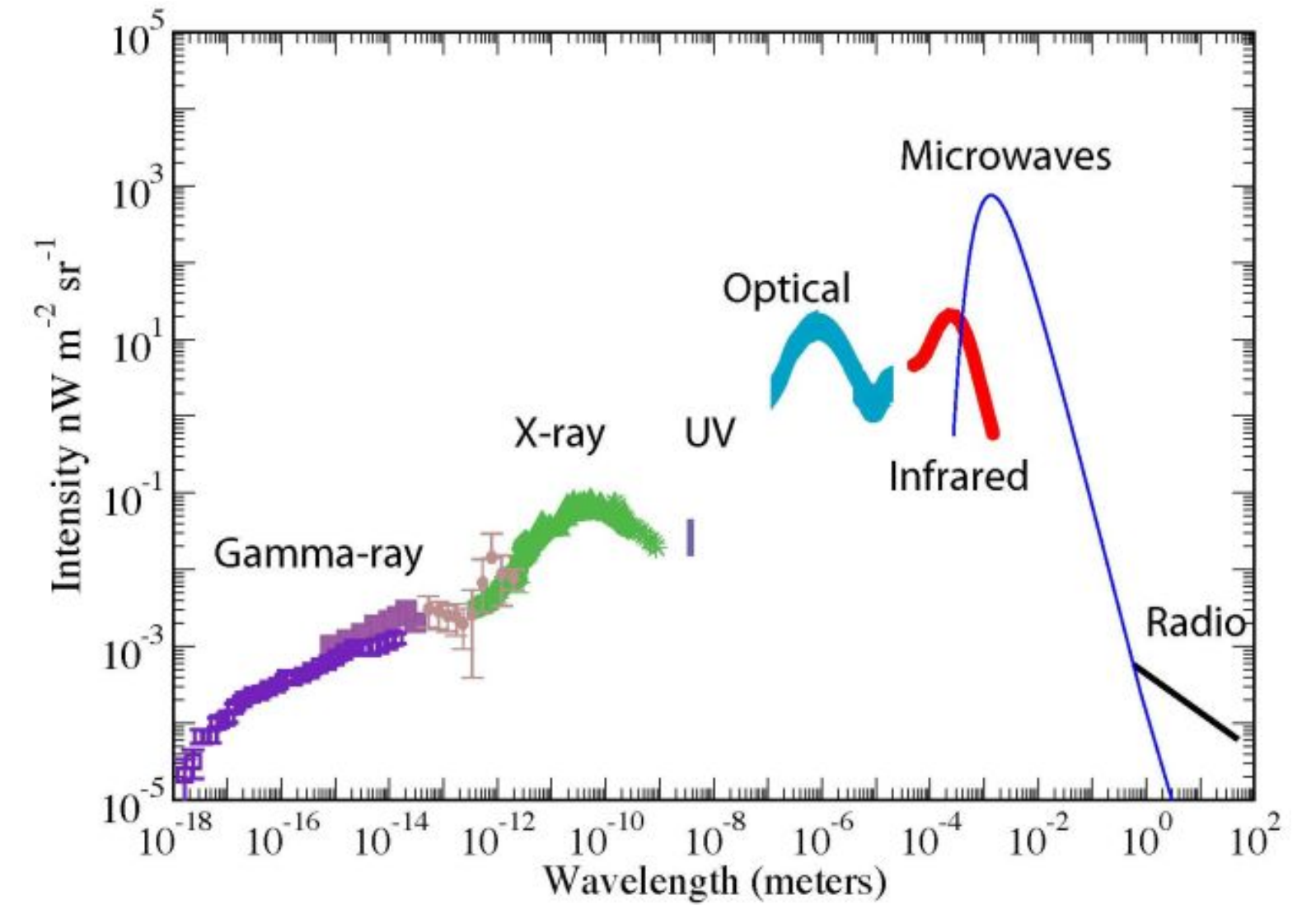
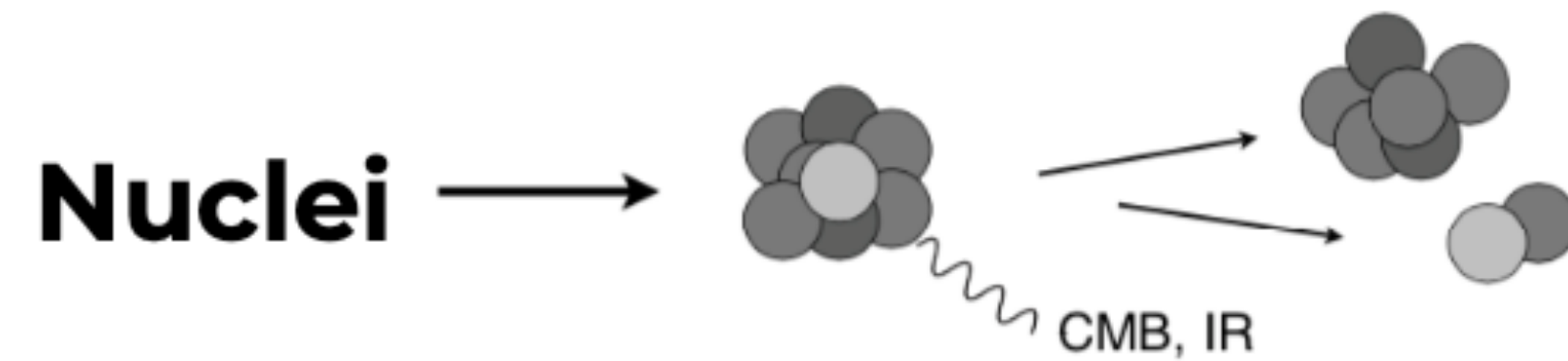
The Pierre Auger Collab. ICRC23

ASTROPHYSICAL SCENARIOS EXPLAINING UHECR DATA

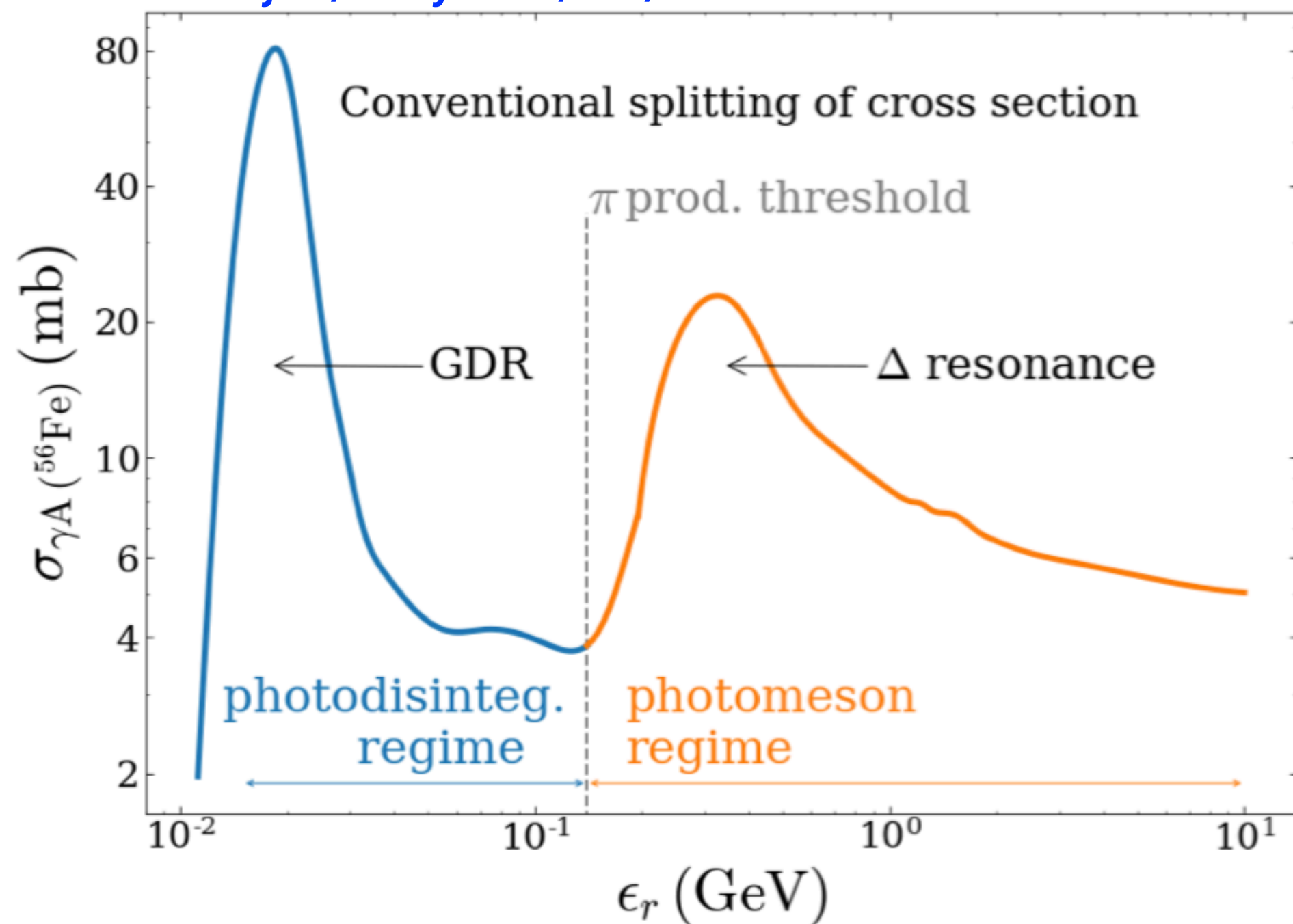


Aloisio, Berezhinsky & Blasi JCAP 2014;
Mollerach & Roulet PRD 2020; Das et al, Eur.Phys.J. 2021; Luce et al, ApJ 2022; The Auger Collab. JCAP 2023

PHOTO-DISINTEGRATION OF NUCLEI IN EXTRAGALACTIC SPACE



Morejon, Fedynitch, DB, Biehl & Winter JCAP 2019



- Regimes of photo-disintegration reactions
 - Giant Dipole Resonance (GDR): protons and neutrons can be considered as penetrating fluids; absorption of photons determines vibrations; ejection of one/two nucleons is dominant
 - Quasi-Deuteron (QD), 20-150 MeV: the photon wavelength becomes comparable with the nuclear dimensions; photon interacts with nucleon pair; ejection of pair + possibly other nucleons
- Conservation of Lorentz factor

Photodisintegration implemented in several UHECR propagation codes, such as

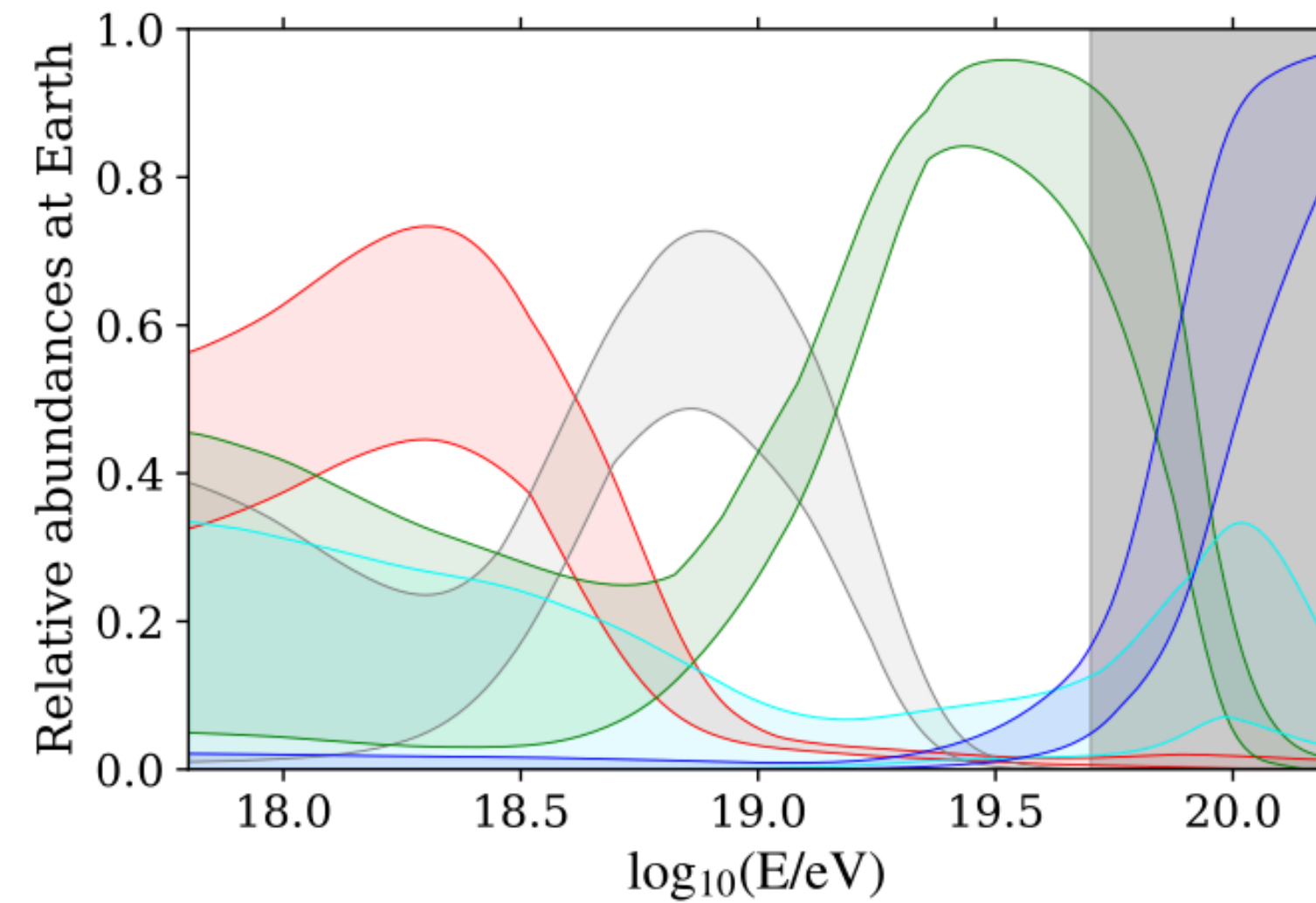
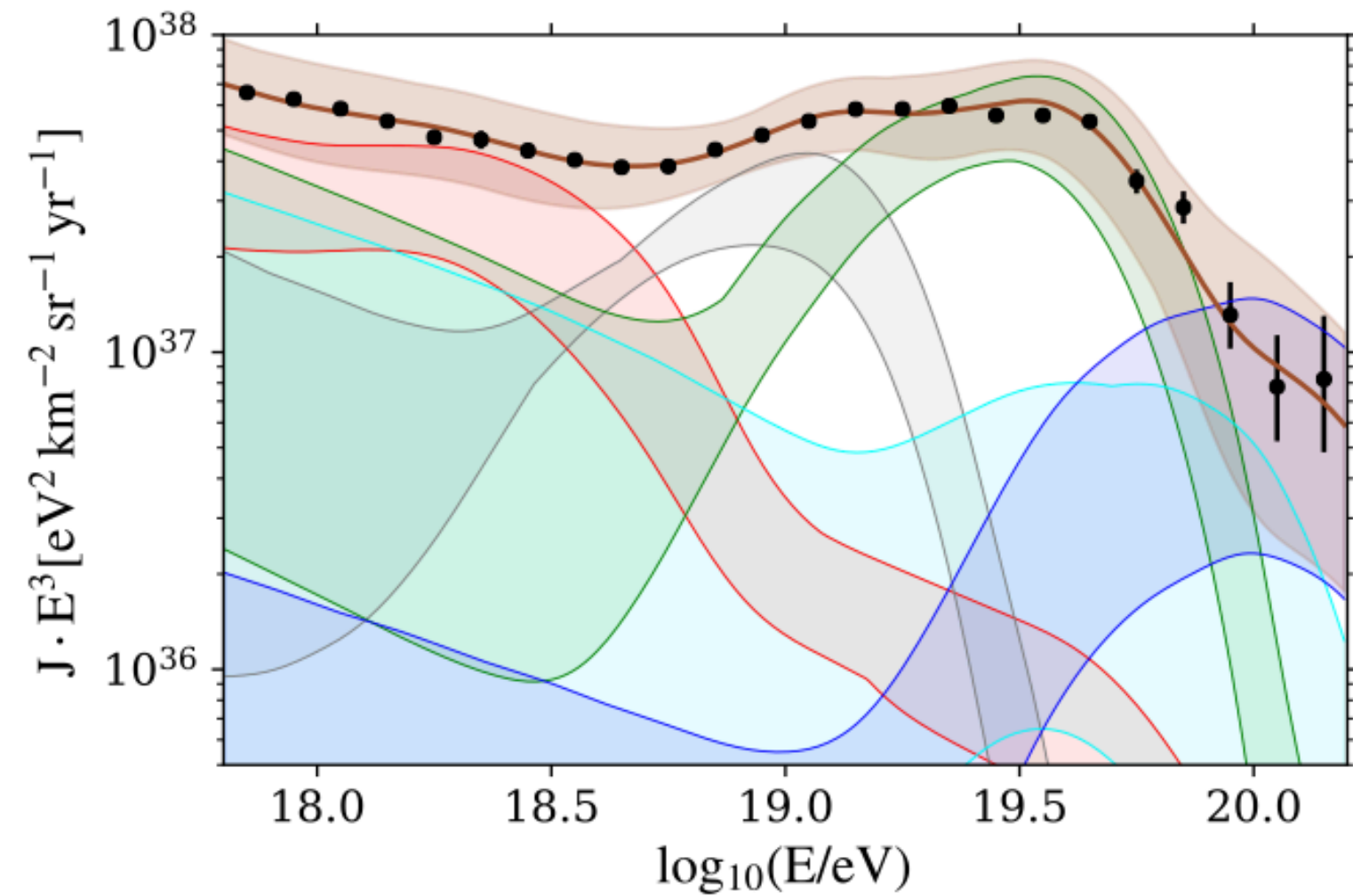
CRPropa (Alves Batista et al, JCAP 2022), SimProp (Aloisio, DB, di Matteo, Grillo, Petrera, Salamida JCAP 2017), Prince (Fedynitch & Heinze)...

ASTROPHYSICAL SCENARIOS EXPLAINING UHECR DATA

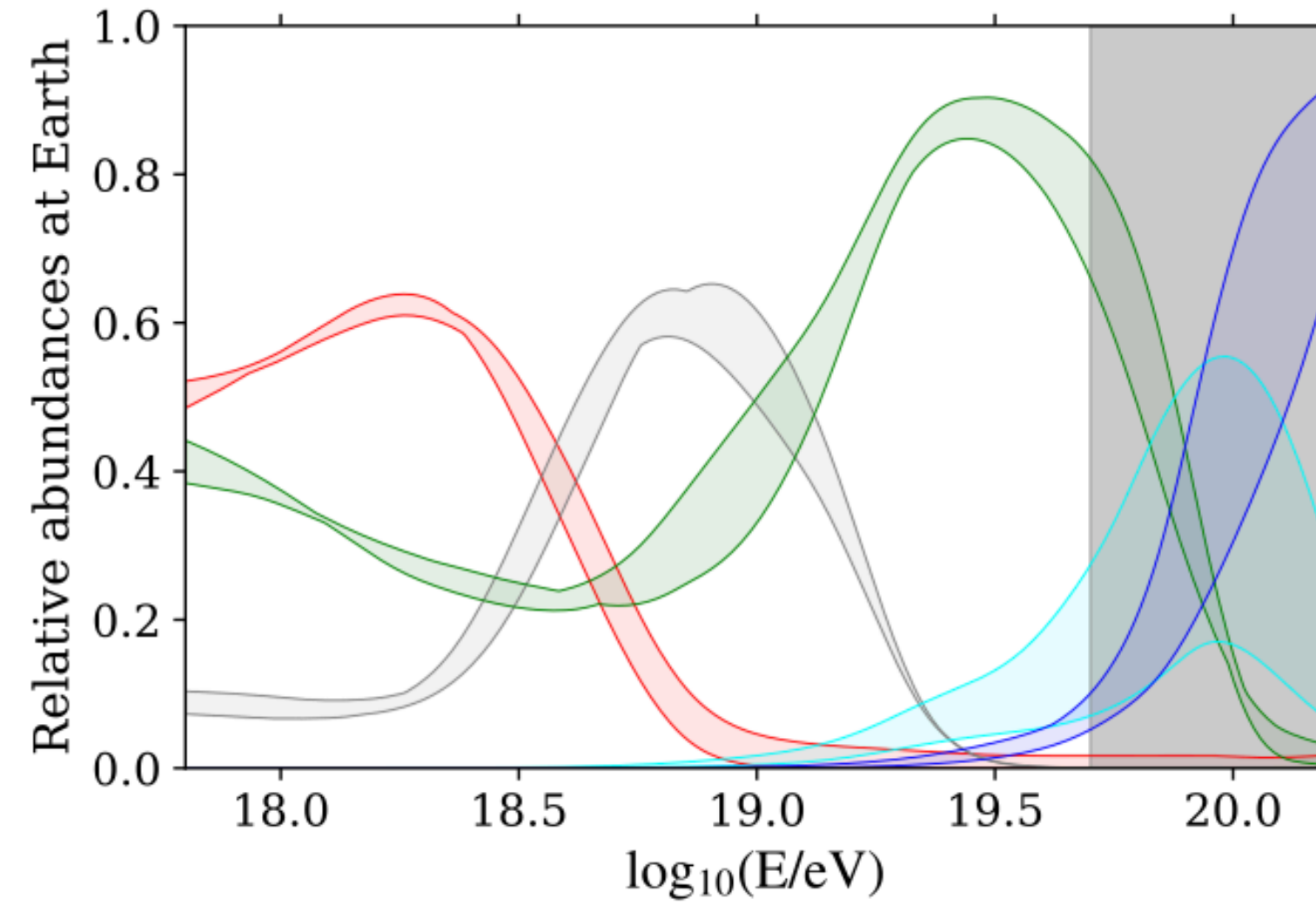
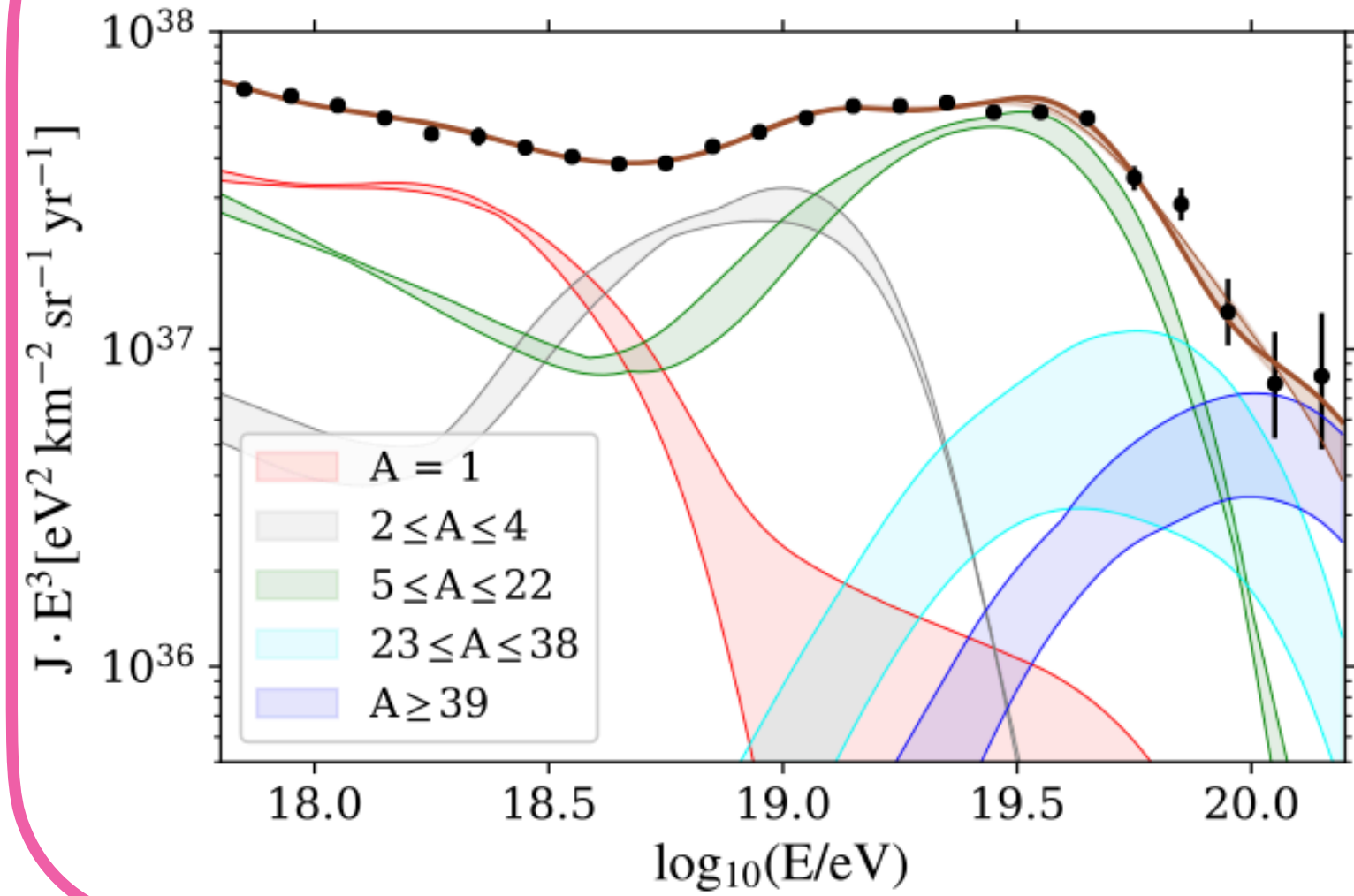
- Uncertainties to be taken into account
 - UHECR measurements:
 - Energy scale: 14%
 - X_{\max} : 6-9 g/cm²
 - Shower models
 - Propagation models (cross section models and extragalactic background light models)

ASTROPHYSICAL SCENARIOS EXPLAINING UHECR DATA

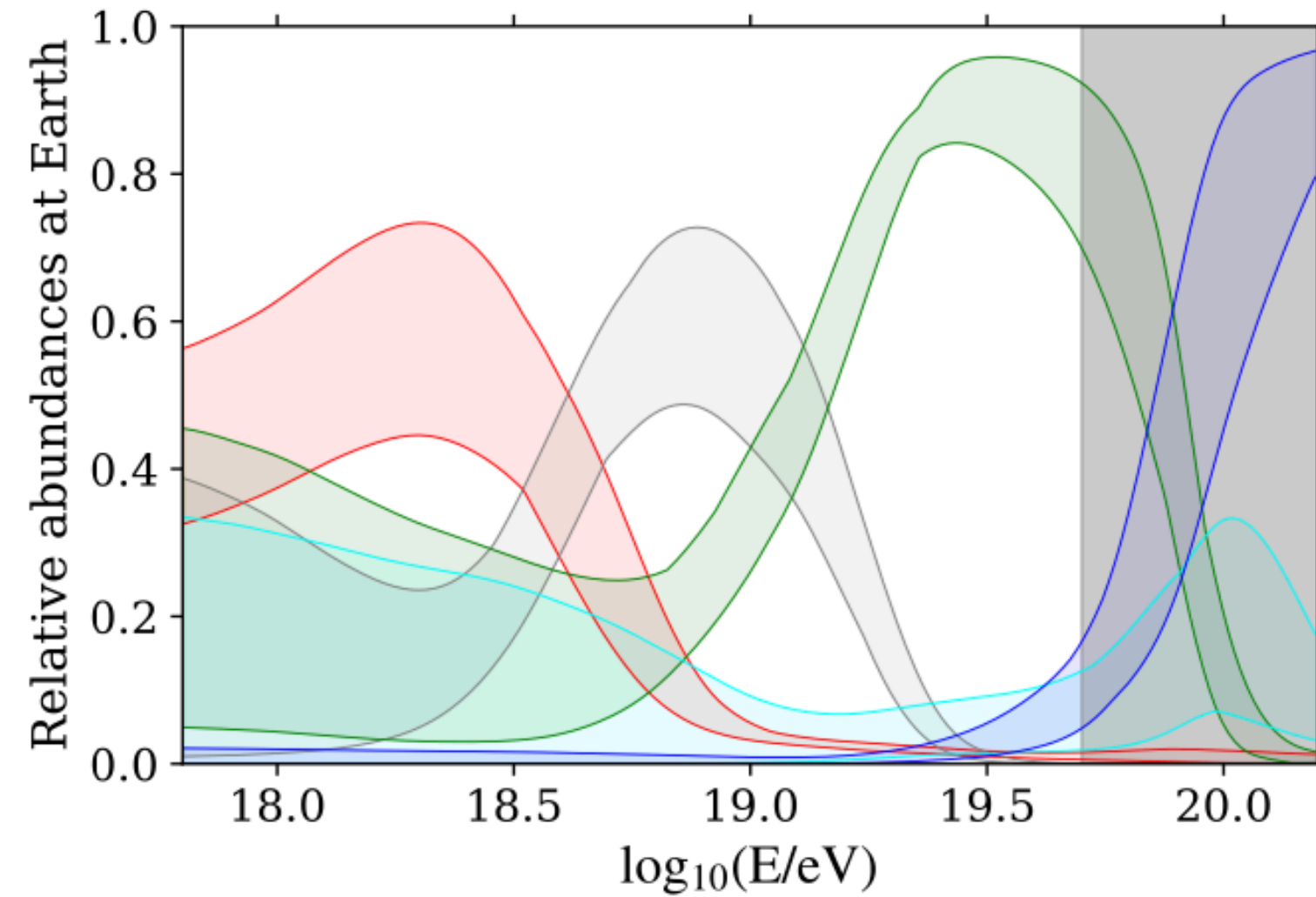
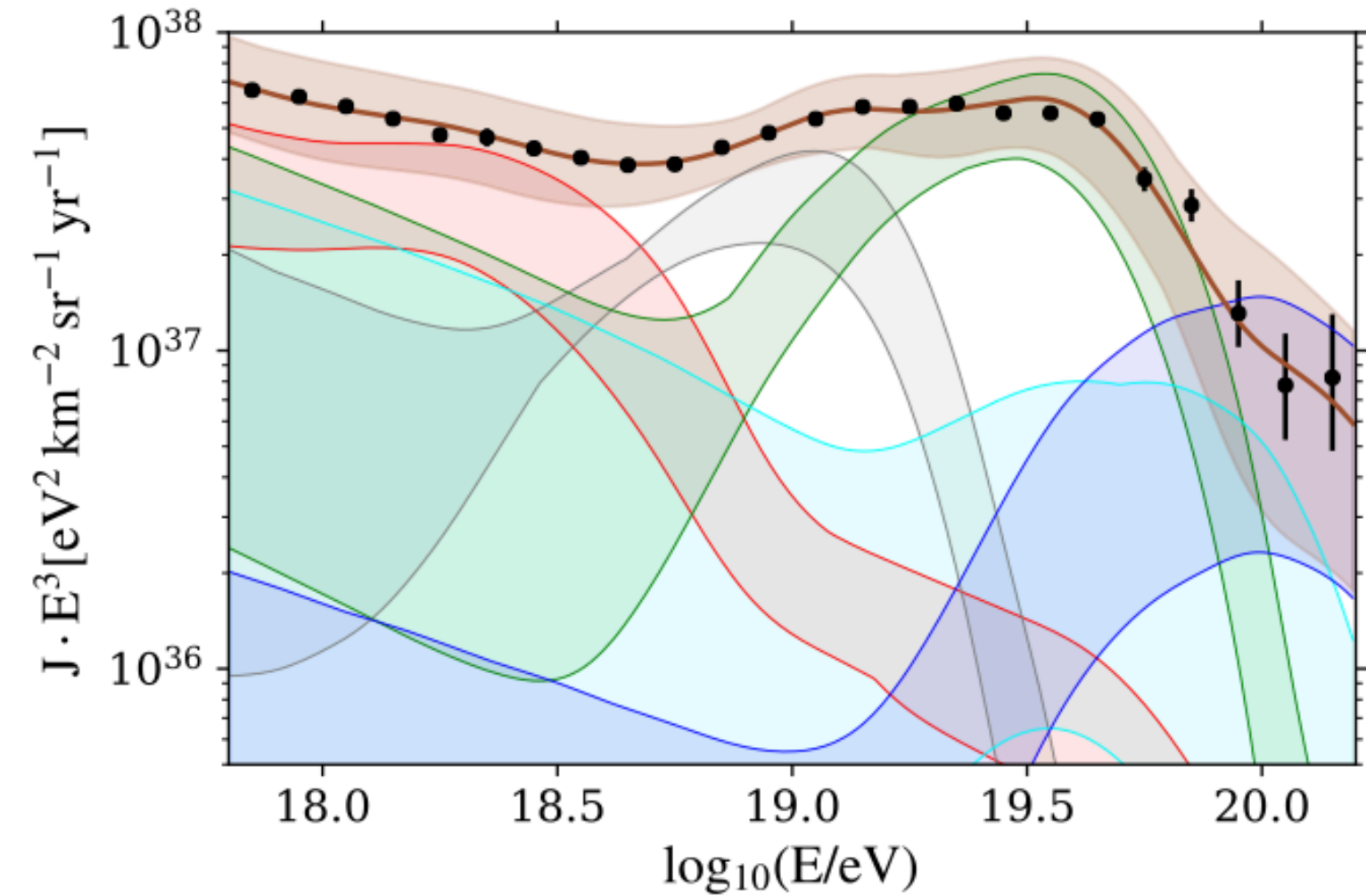
- Uncertainties to be taken into account
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 - Energy scale: 14%
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ASTROPHYSICAL SCENARIOS EXPLAINING UHECR DATA



	TALYS		PSB	
	LE	HE	LE	HE
Gilmore EBL				
$\mathcal{L}_0 / (10^{44} \text{ erg Mpc}^{-3} \text{ yr}^{-1})^*$	11.4	5.1	11.1	4.9
γ	3.52 ± 0.03	-1.99 ± 0.11	3.51 ± 0.03	-1.89 ± 0.18
$\log_{10}(R_{\text{cut}}/V)$	>19.4	18.15 ± 0.01	>19.5	18.16 ± 0.02
$I_{\text{H}} (\%)$	48.7	0.0	49.1	0.2
$I_{\text{He}} (\%)$	7.3	23.6	11.1	48.3
$I_{\text{N}} (\%)$	44.0	72.1	39.8	41.5
$I_{\text{Si}} (\%)$	0.0	1.3	0.0	8.5
$I_{\text{Fe}} (\%)$	0.0	3.1	0.0	1.5
δ_{HIM}	1.0 (limit)		$0.96^{+0.04}_{-0.12}$	
$D_J (N_J)$	56.6 (24)		50.7 (24)	
$D_{X_{\text{max}}} (N_{X_{\text{max}}})$	516.5 (329)		529.0 (329)	
$D (N)$	573.1 (353)		579.7 (353)	

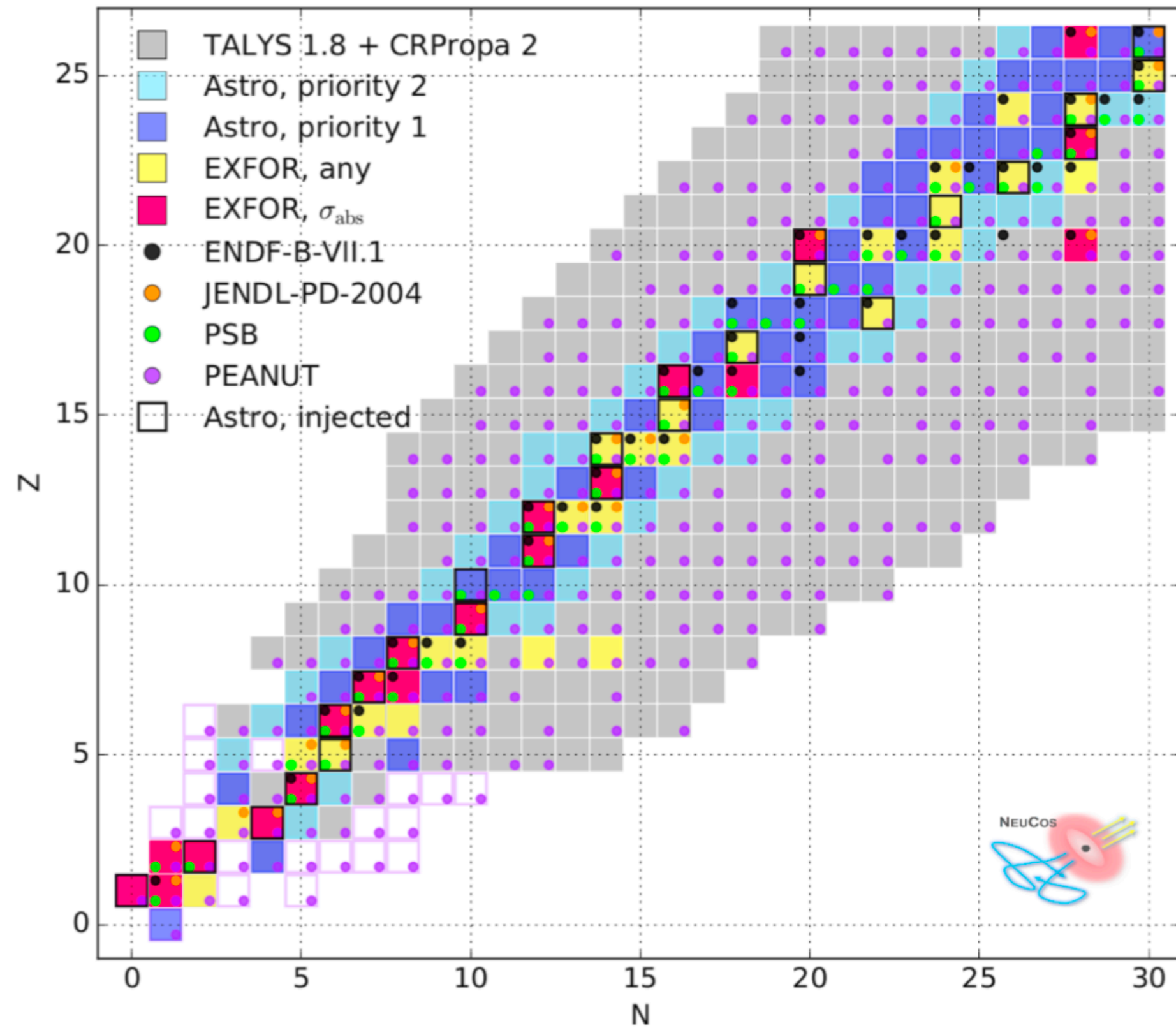


TALYS	EPOS-LHC		SIBYLL 2.3d	
	LE	HE	LE	HE
Gilmore EBL				
$\mathcal{L}_0 / (10^{44} \text{ erg Mpc}^{-3} \text{ yr}^{-1})^*$	11.4	5.1	10.8	4.9
γ	3.52 ± 0.03	-1.99 ± 0.11	3.40 ± 0.02	-1.30 ± 0.19
$\log_{10}(R_{\text{cut}}/V)$	>19.4	18.15 ± 0.01	18.26 ± 0.05	18.19 ± 0.02
$I_{\text{H}} (\%)$	48.7	0.0	15.6	0.0
$I_{\text{He}} (\%)$	7.3	23.6	46.2	20.9
$I_{\text{N}} (\%)$	44.0	72.1	38.2	70.7
$I_{\text{Si}} (\%)$	0.0	1.3	0.0	5.4
$I_{\text{Fe}} (\%)$	0.0	3.1	0.0	3.0
$D_J (N_J)$	56.6 (24)		42.7 (24)	
$D_{X_{\text{max}}} (N_{X_{\text{max}}})$	516.5 (329)		592.2 (329)	
$D (N)$	573.1 (353)		634.9 (353)	

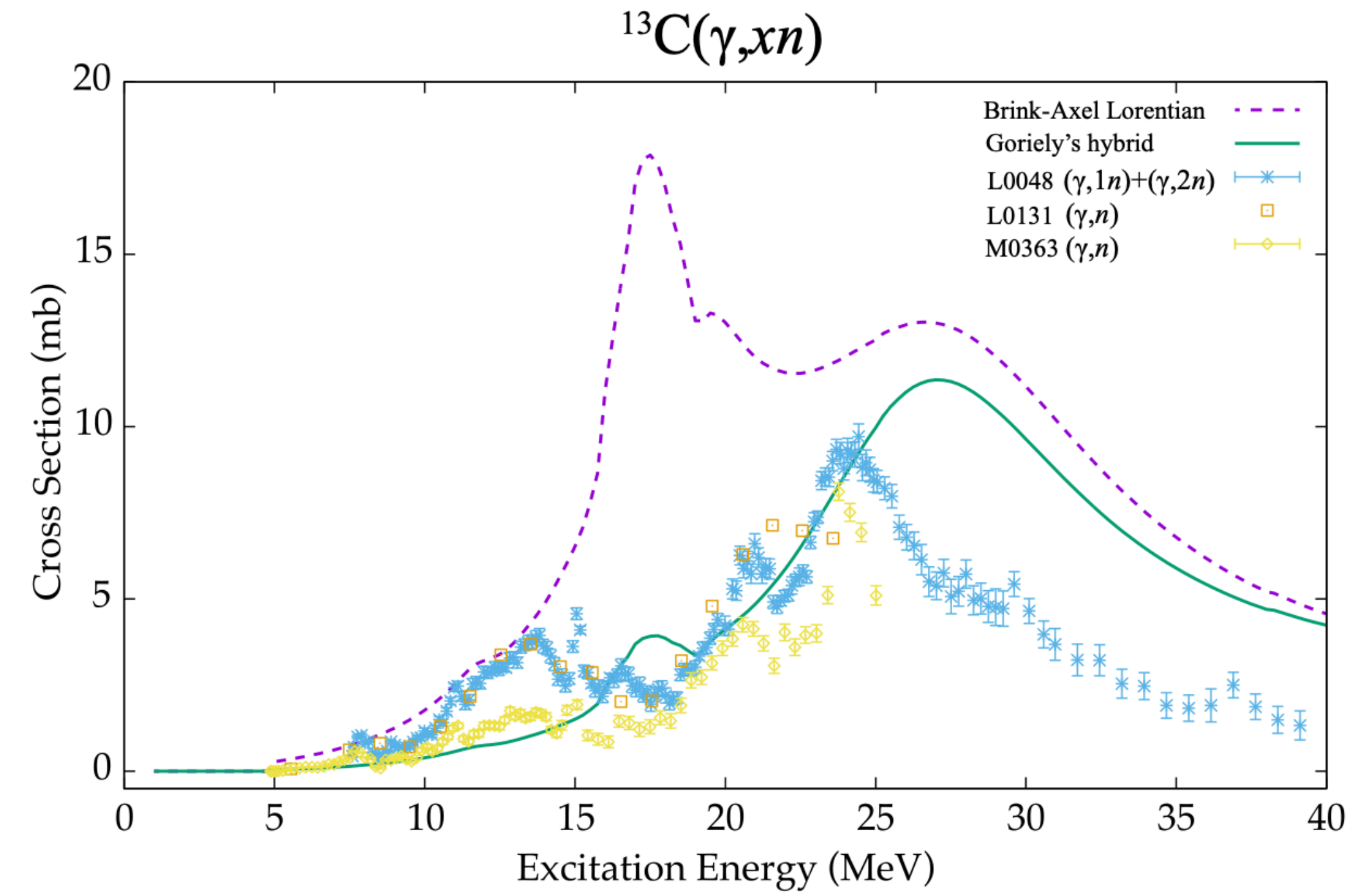
UNCERTAINTIES IN PHOTO-DISINTEGRATION

- Available measurements are sparse

DB, Fedynitch & Winter, Sci. Reports 2017



Tamii et al, PANDORA project, Eur.Phys.J. 2023



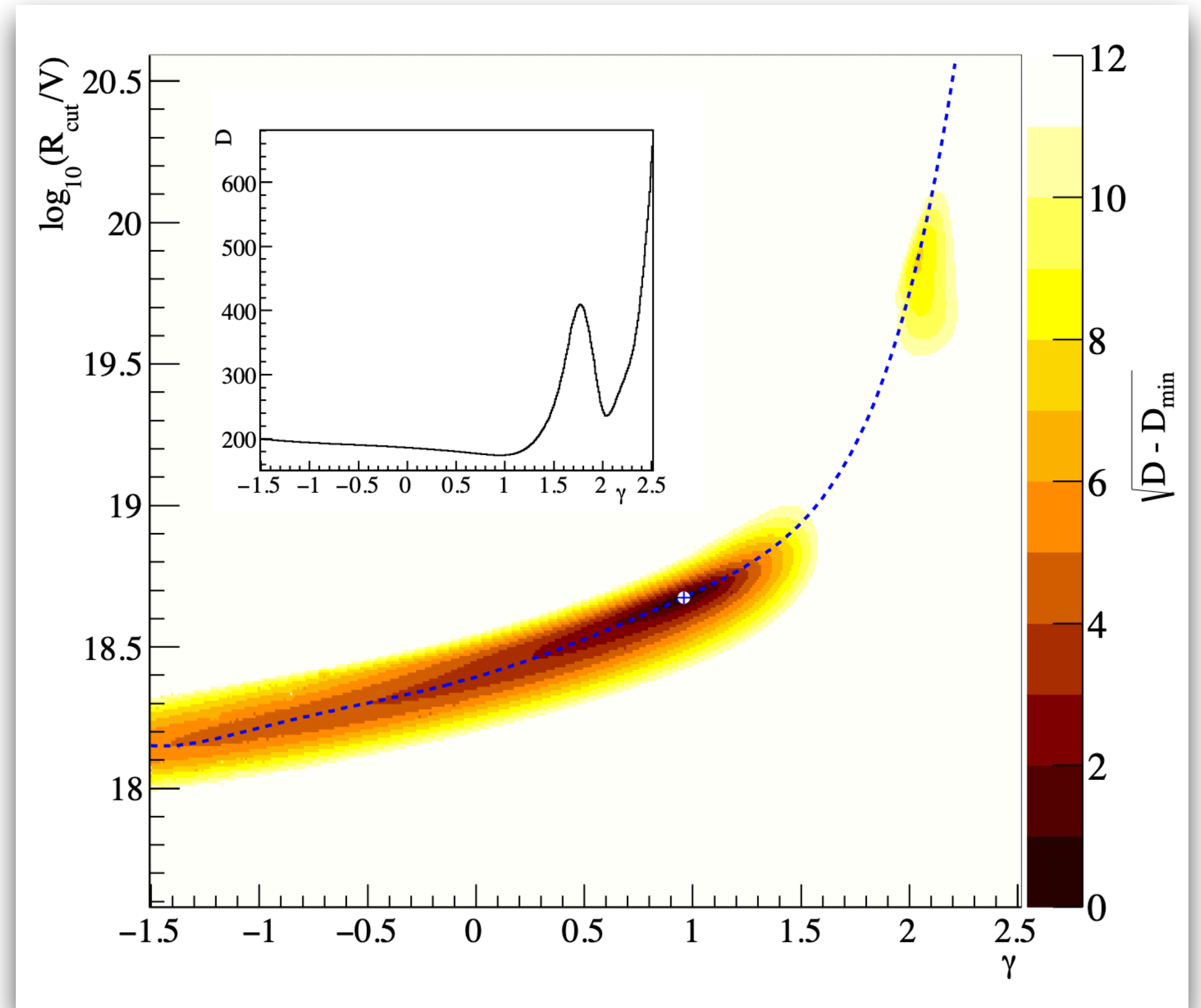
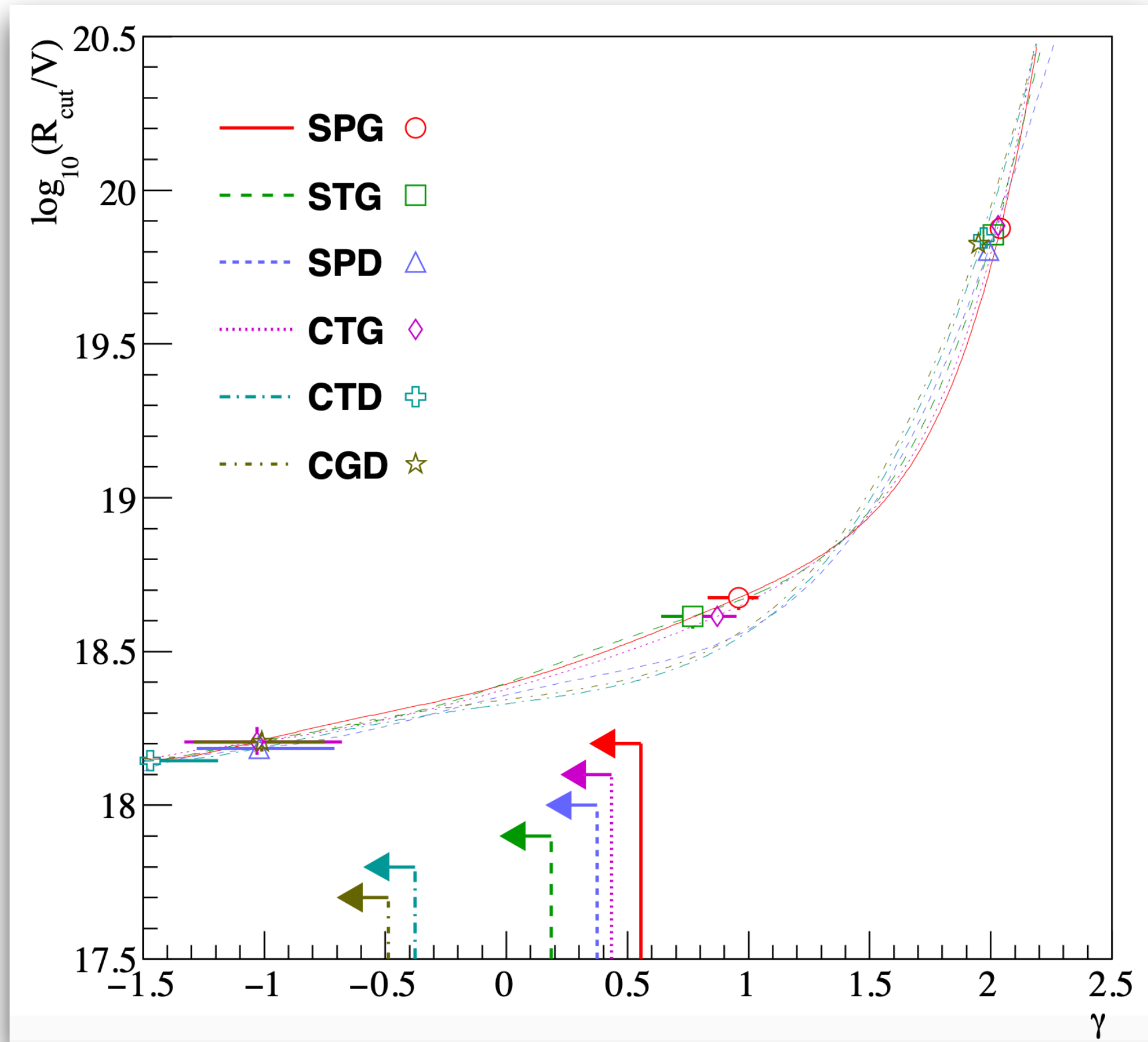
- Theoretical models do not always reproduce (available) data

- Studies of sensitivity of UHECR data to uncertainties in photo-disintegration: [Khan et al Astropart.Phys. 2005](#); [Allard et al JCAP 2006](#); [Alves Batista, DB, di Matteo, van Viet & Walz JCAP2015](#); [The Pierre Auger Collab. JCAP 2017](#); [DB, Fedynitch & Winter, Sci.Report 2017](#); [Alves Batista, DB, di Matteo & van Vliet JCAP2019](#); [Heinze, Fedynitch, DB & Winter ApJ 2019](#); [The Pierre Auger Collab. JCAP 2023](#); [Kido et al. Astropart.Phys. 2023](#)

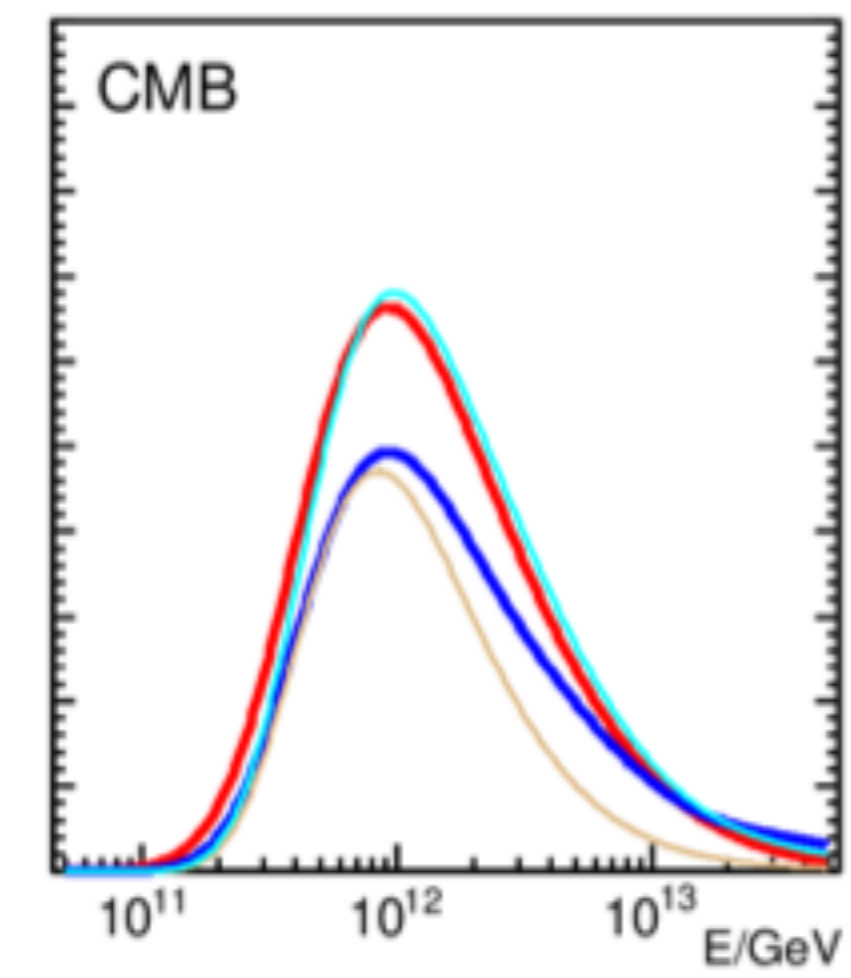
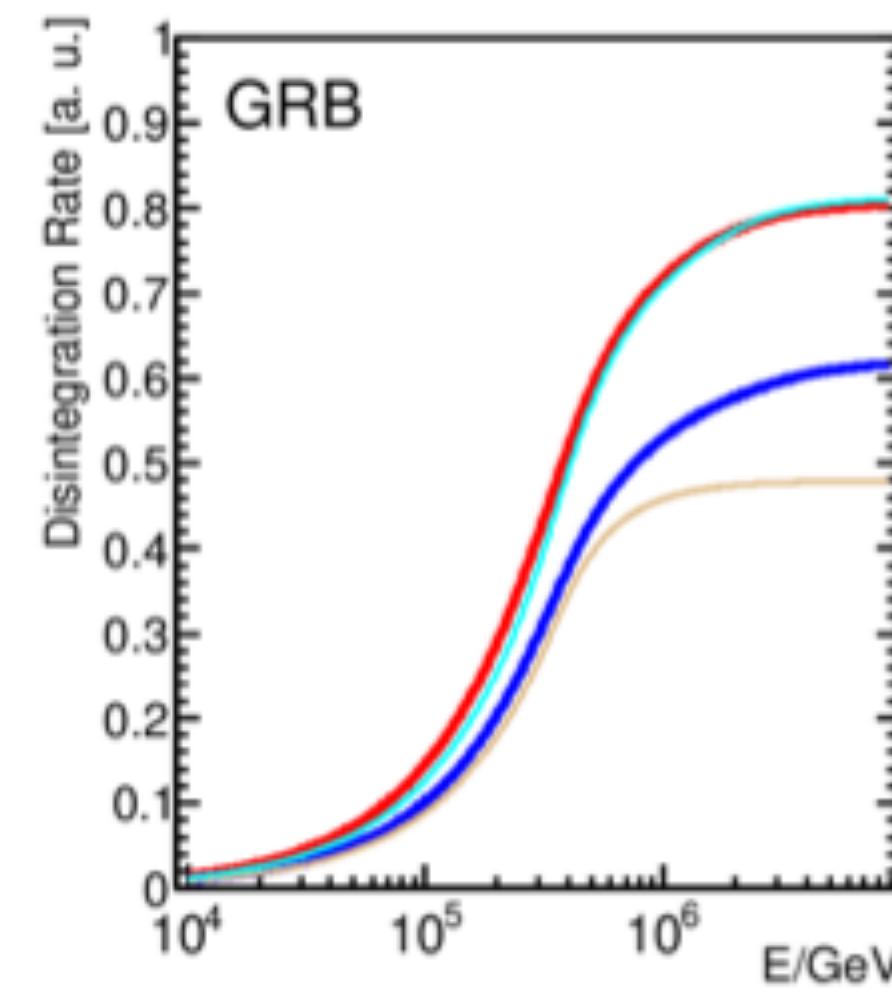
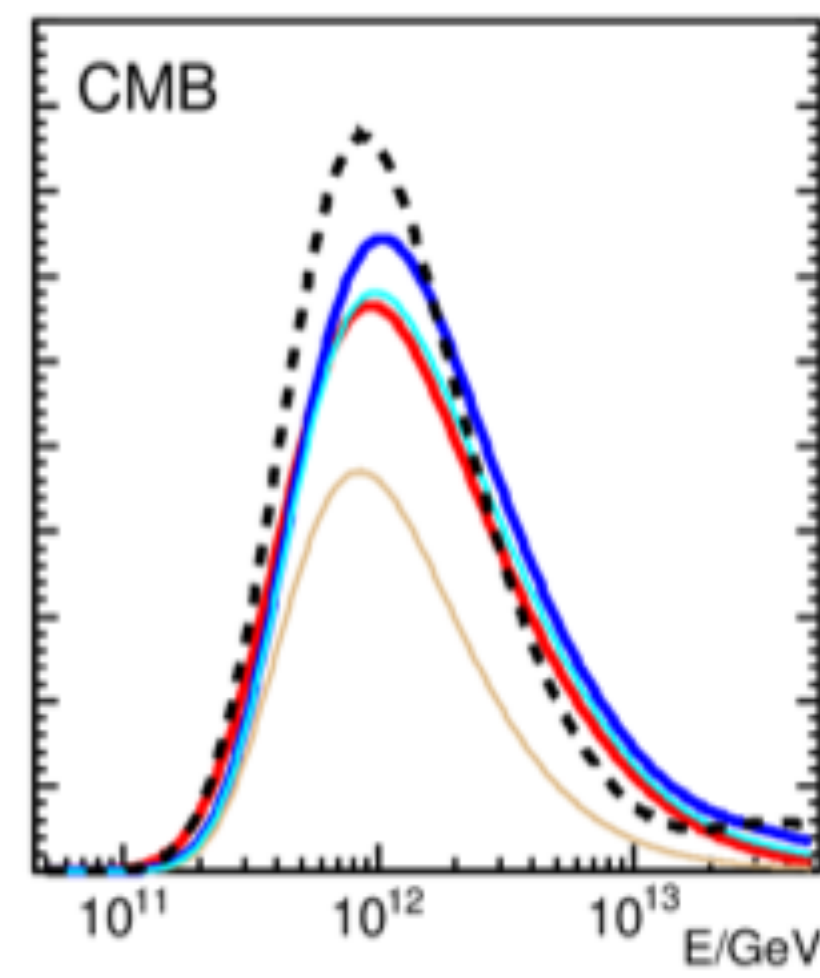
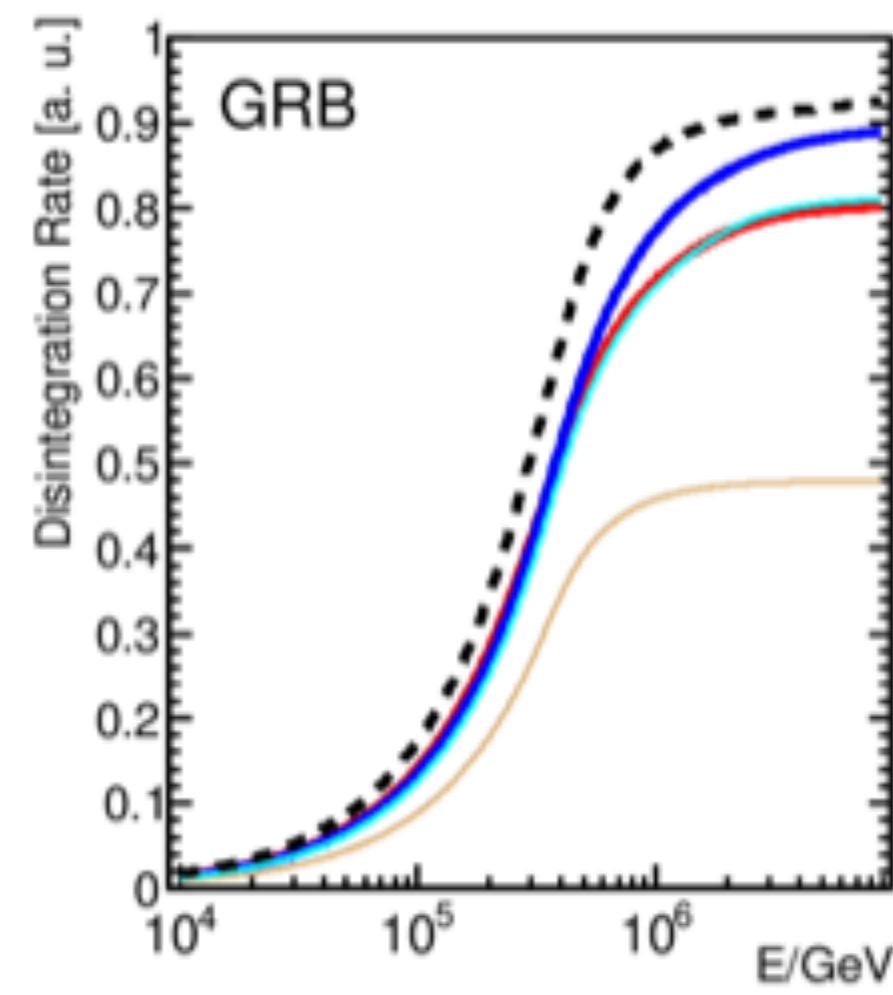
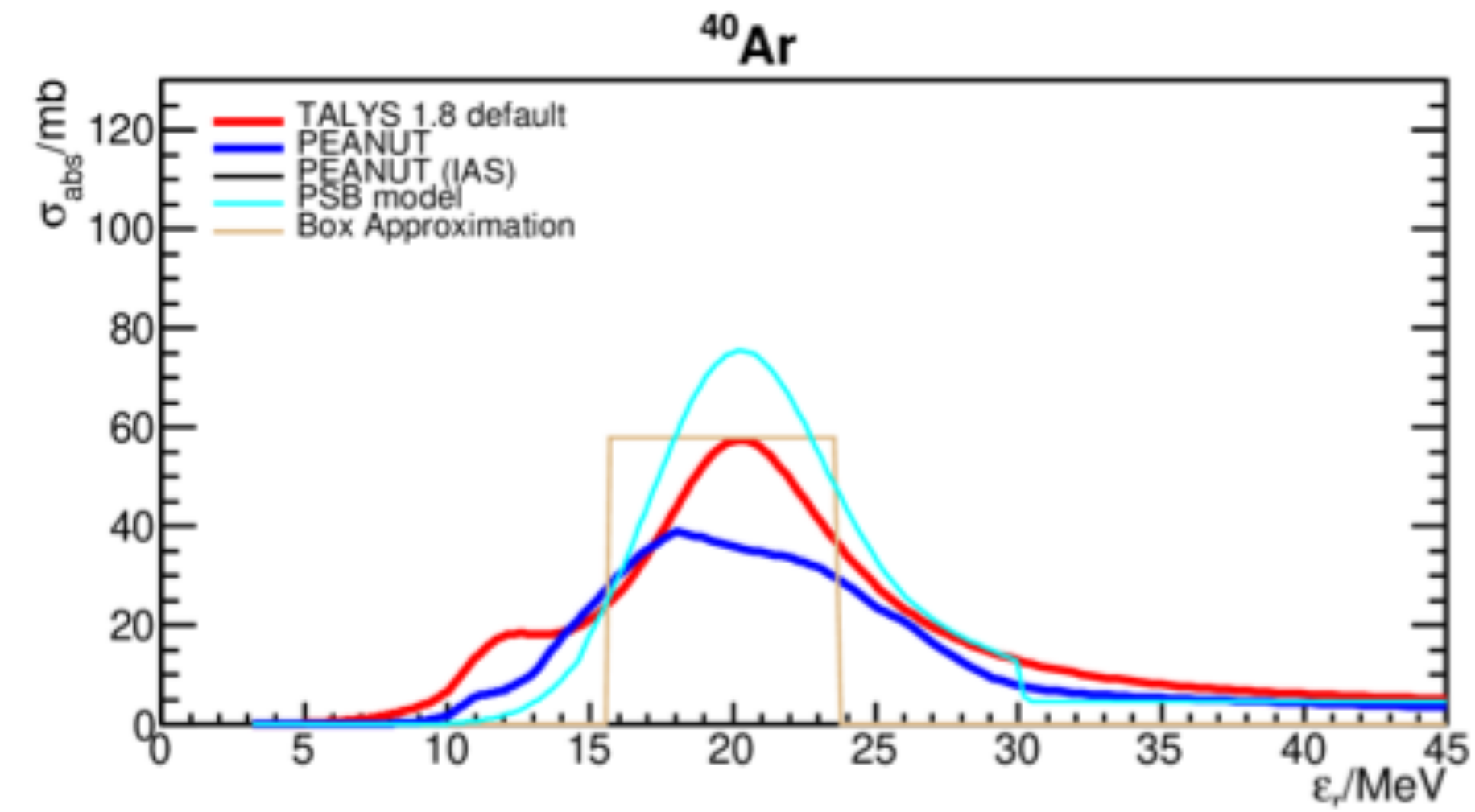
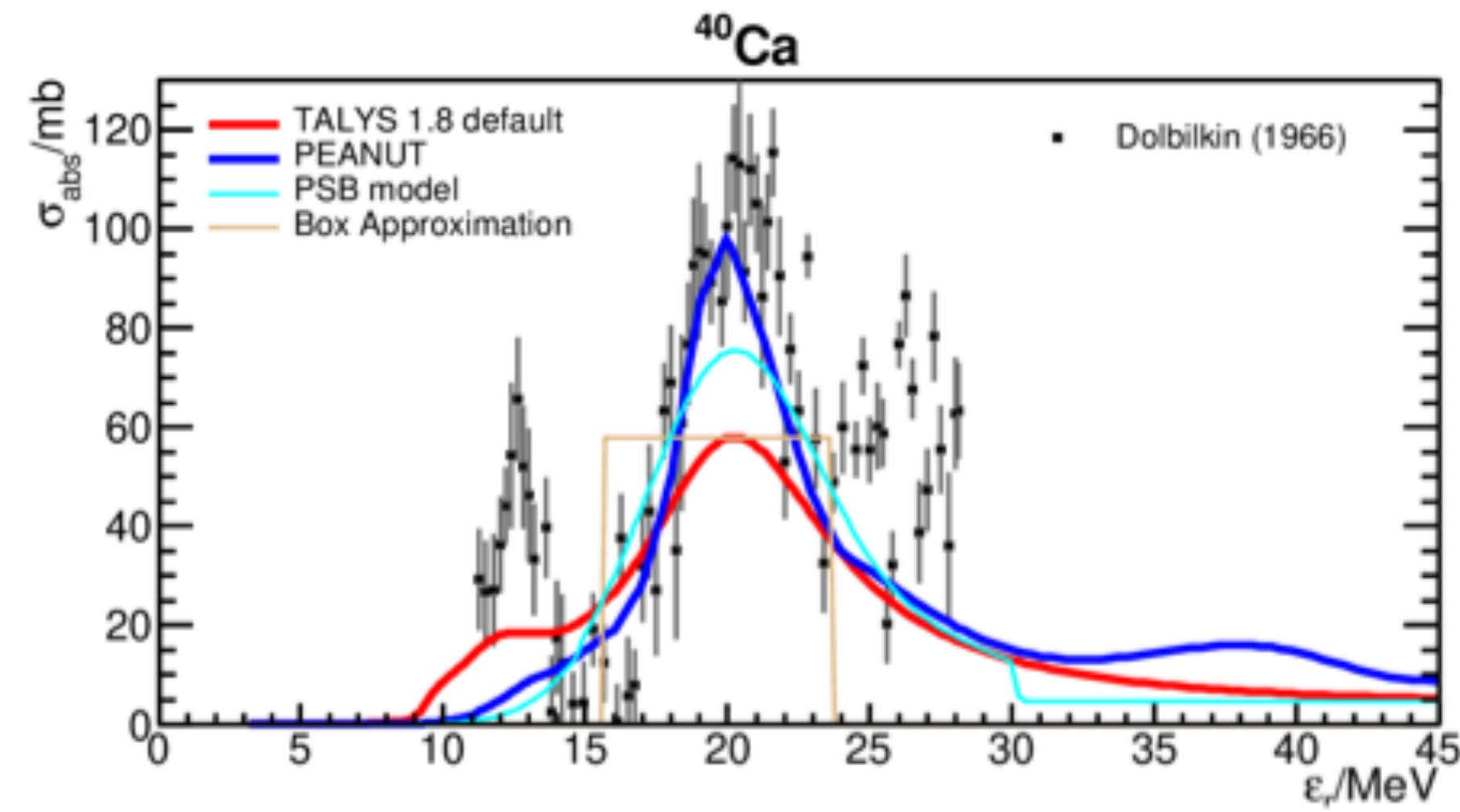
- New measurements: [The PANDORA project, Eur.Phys.J. 2023](#)

BACKUP SLIDES

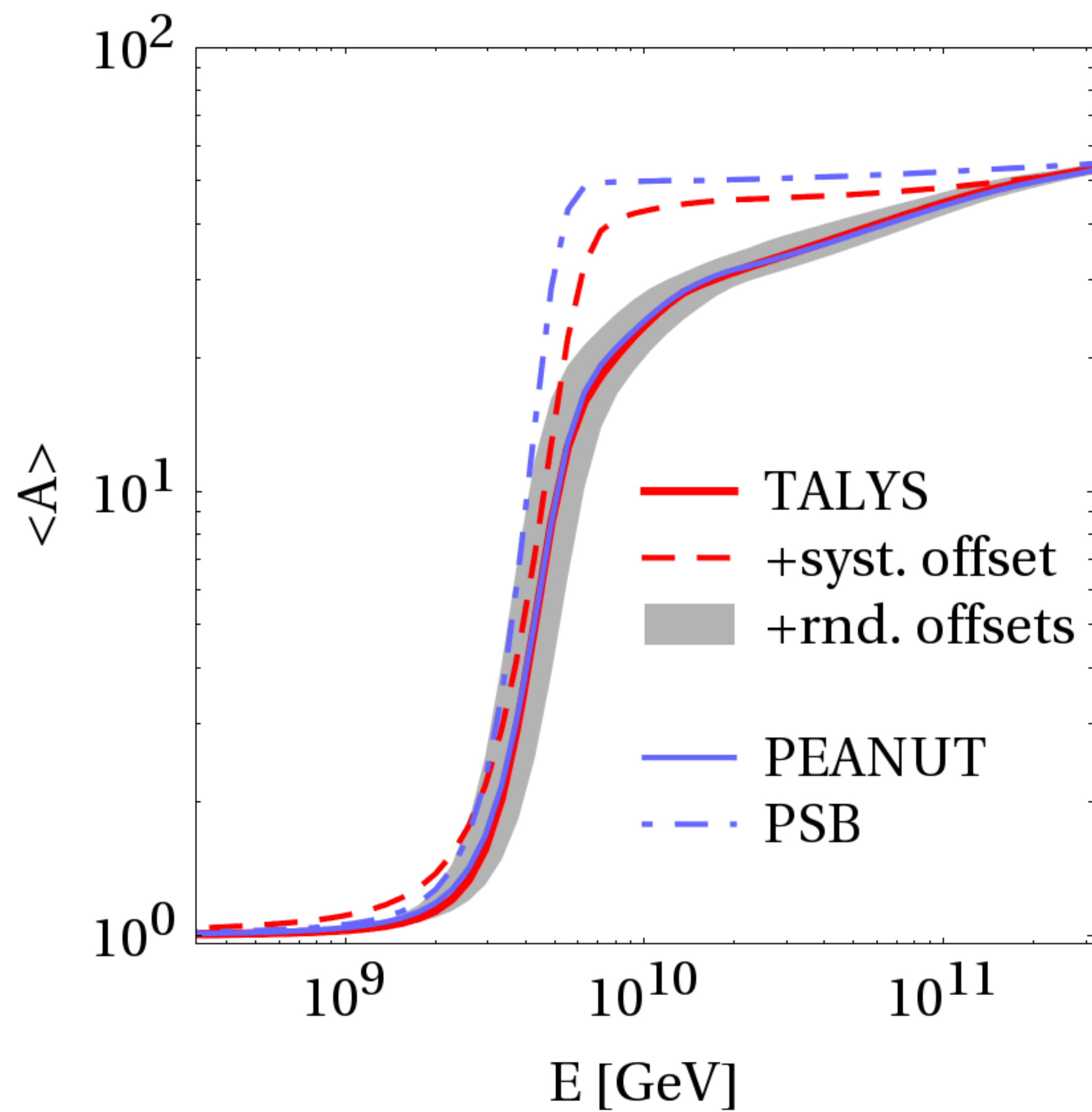
ASTROPHYSICAL SCENARIOS EXPLAINING UHECR DATA



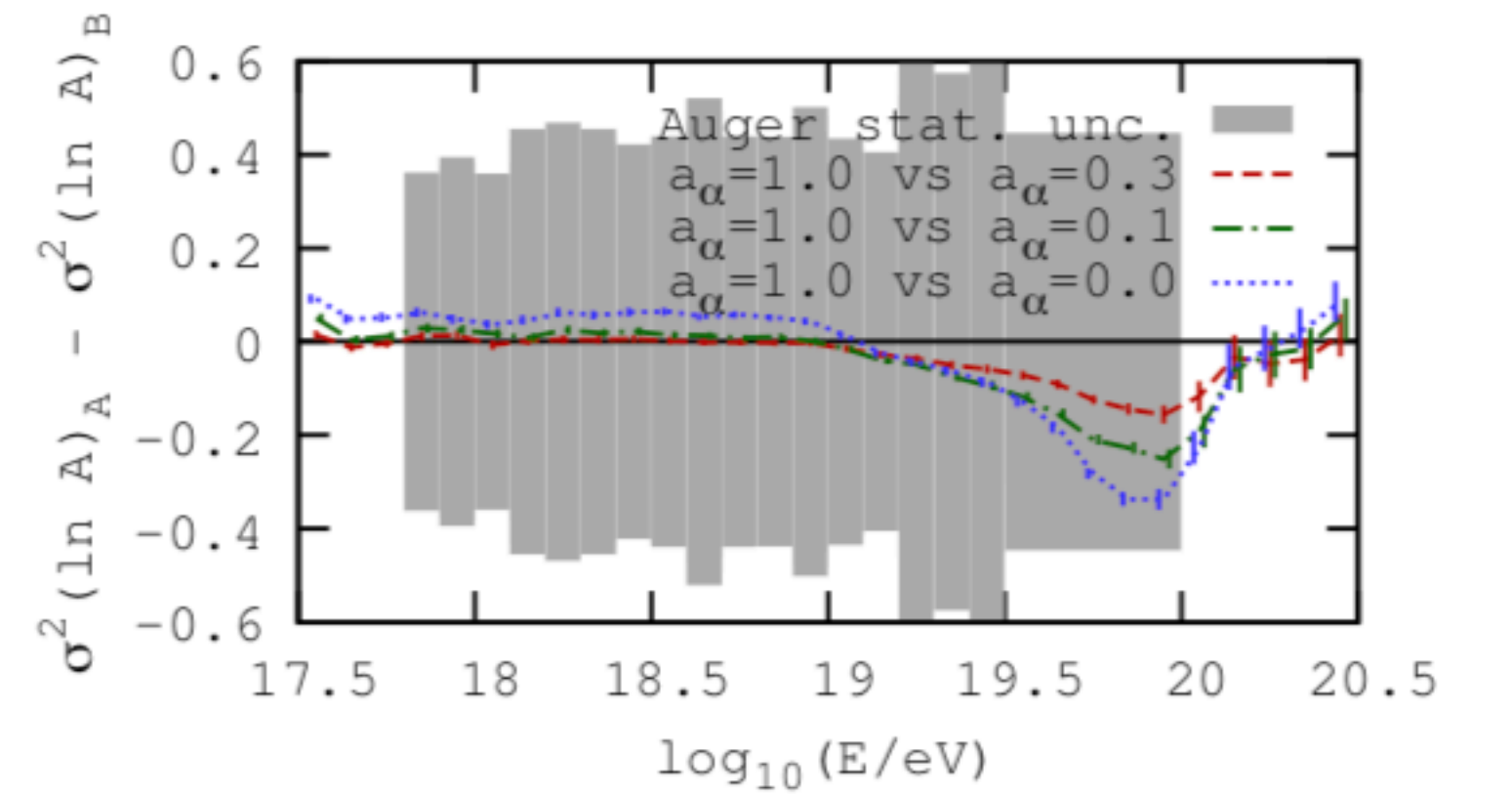
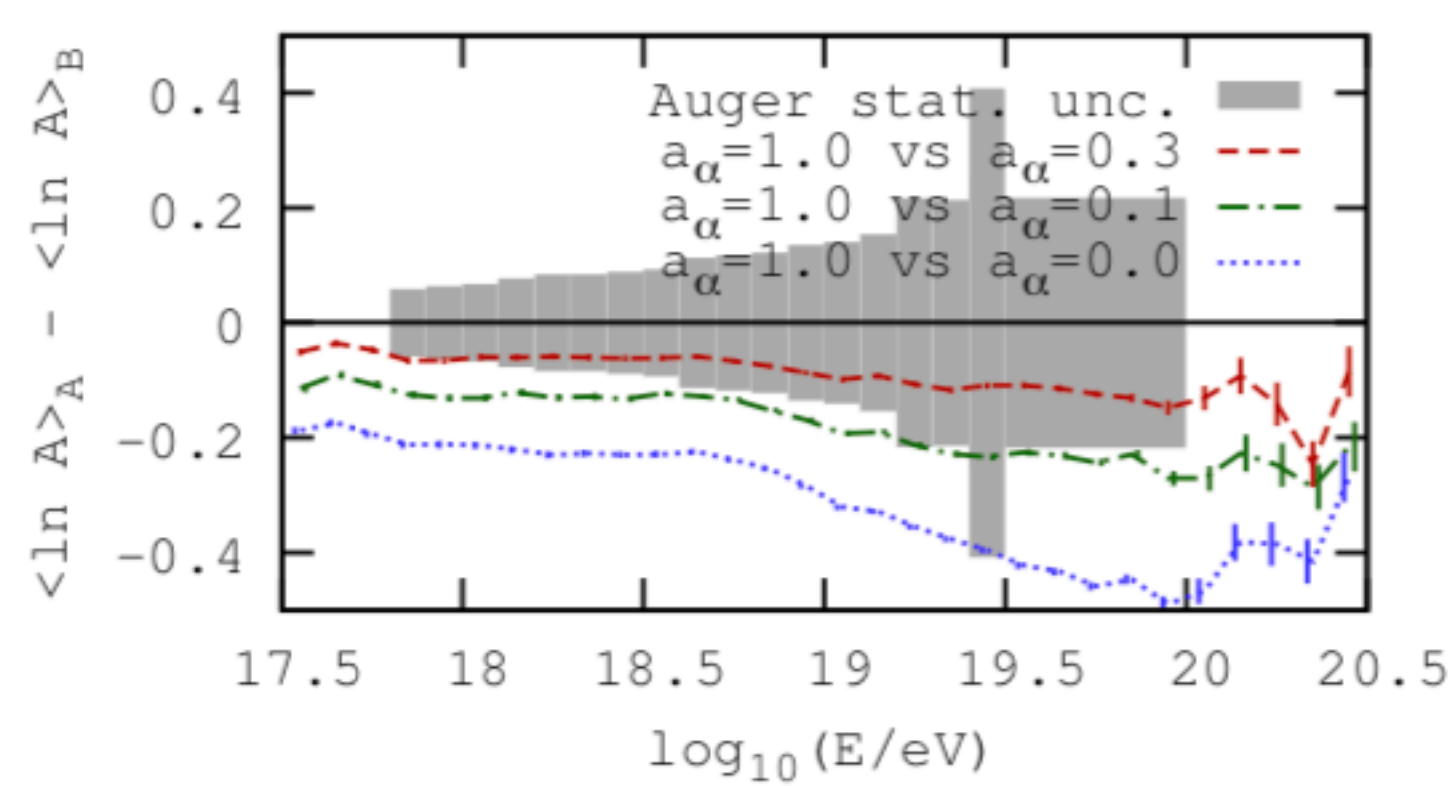
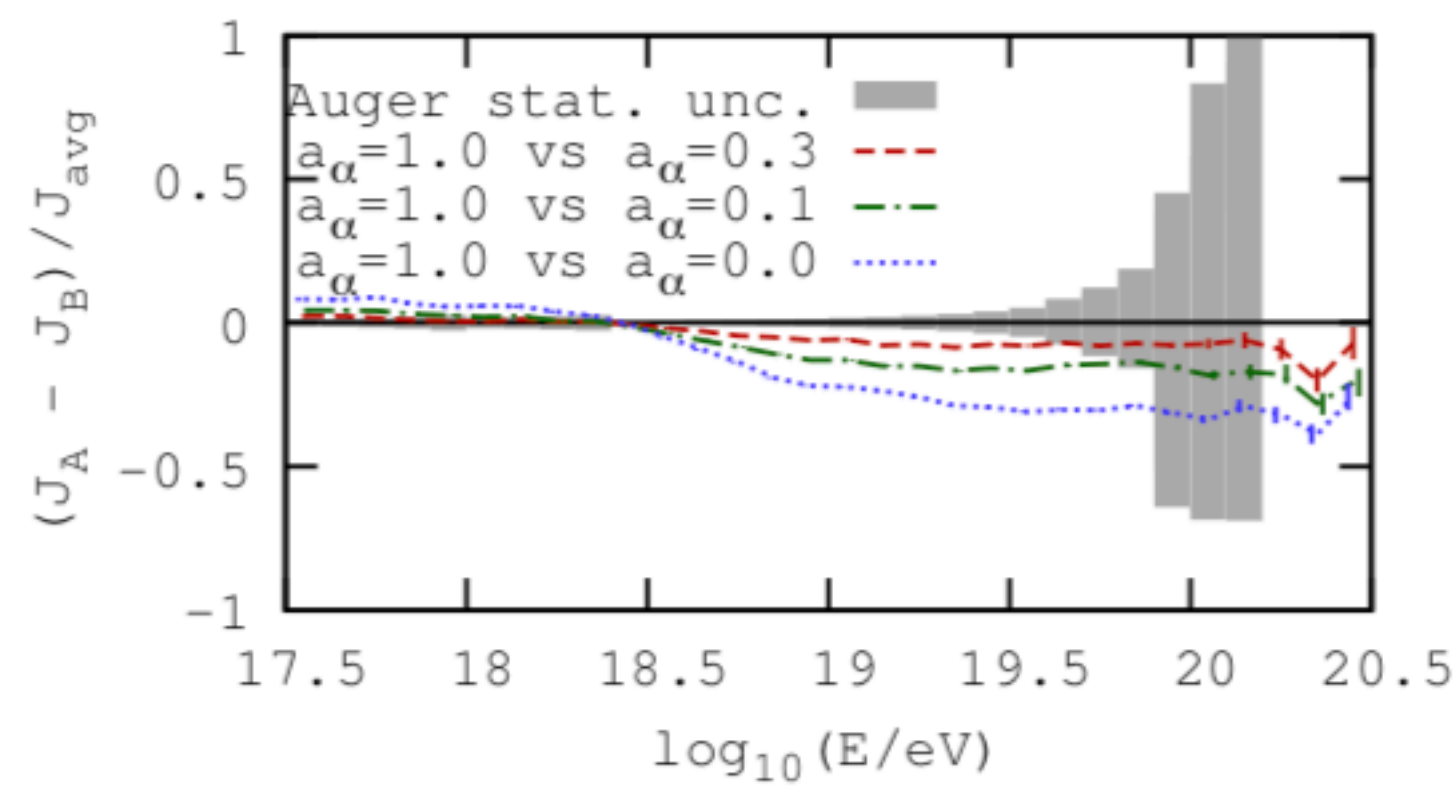
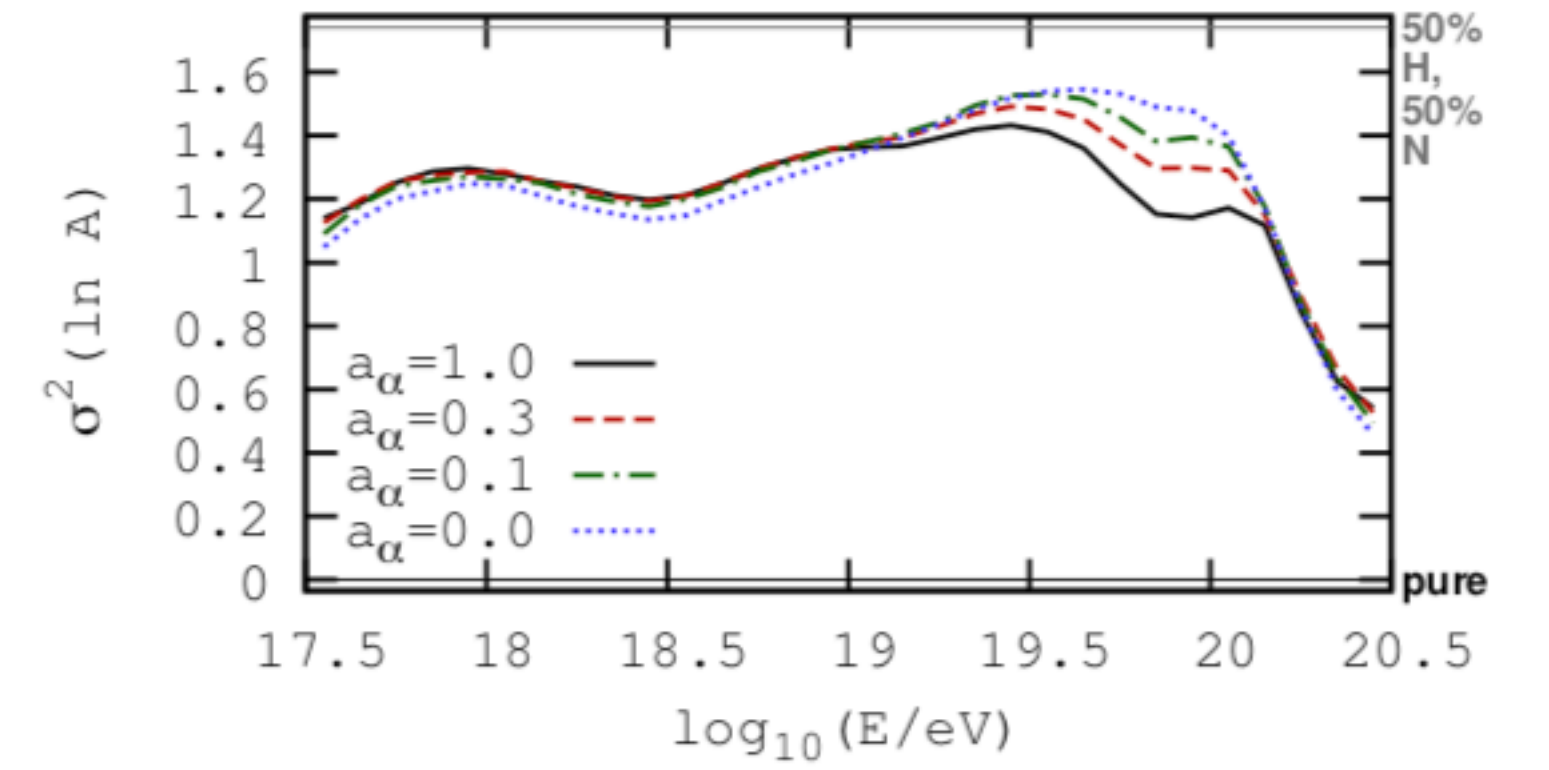
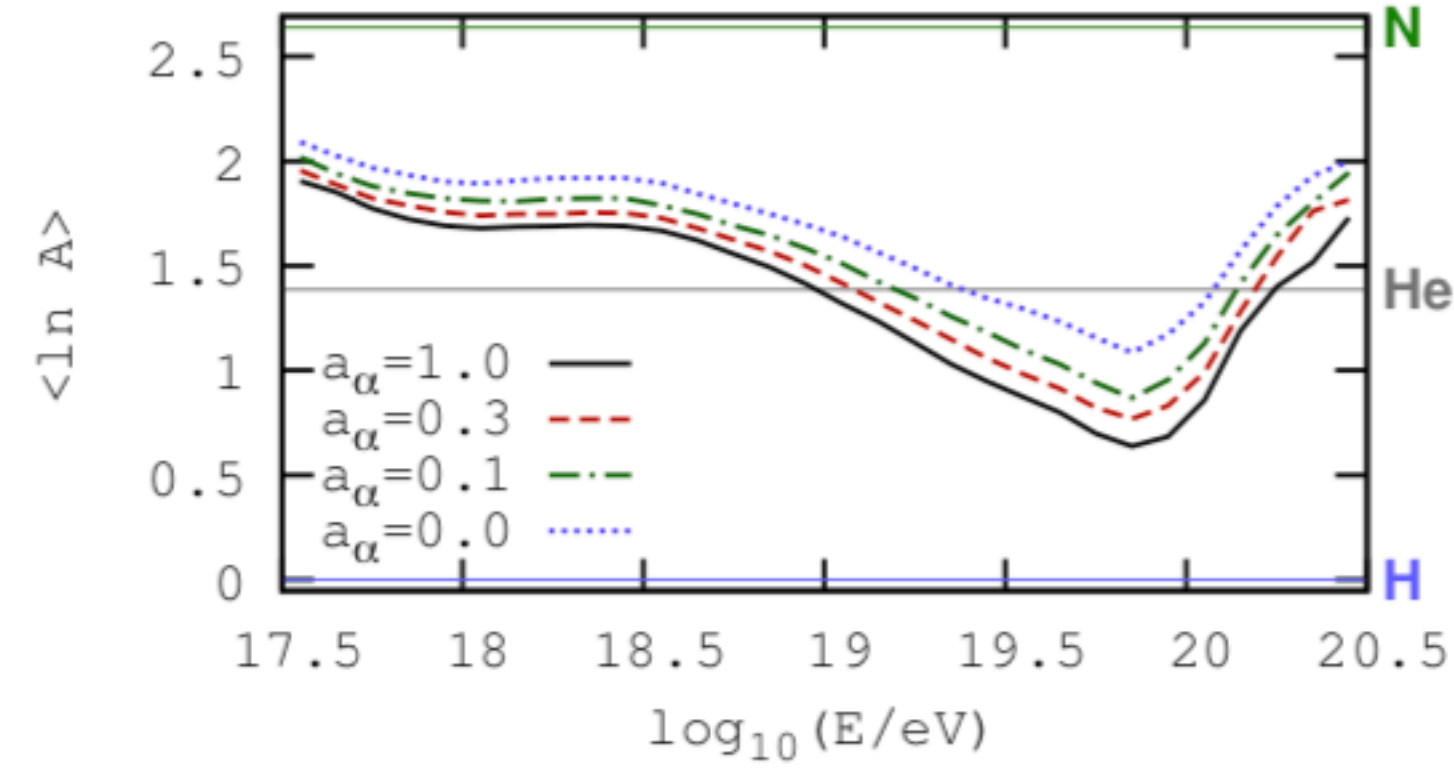
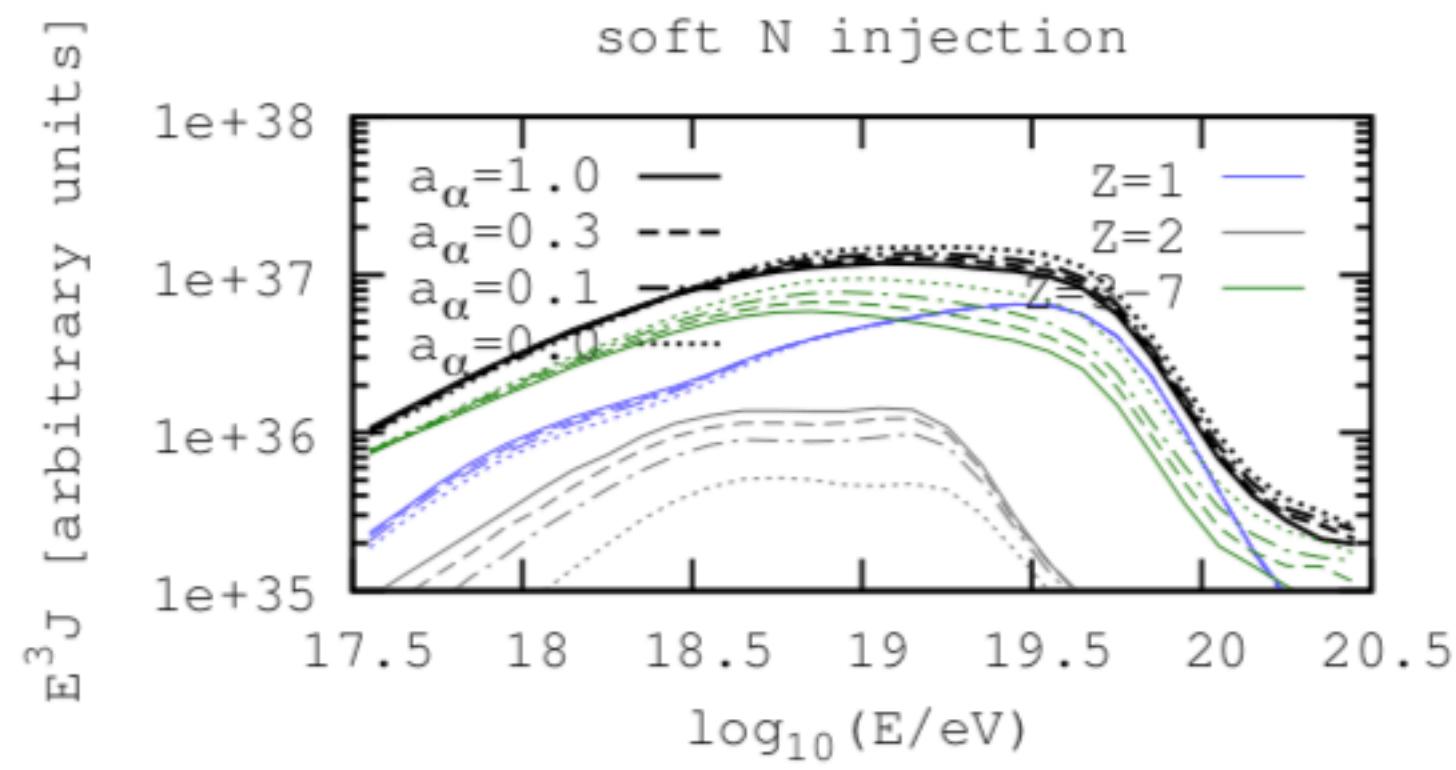
CROSS SECTION UNCERTAINTIES AND DISINTEGRATION RATES



CROSS SECTION UNCERTAINTIES AND AVERAGE MASS



CROSS SECTION UNCERTAINTIES AND EXPECTED SPECTRA



CROSS SECTION UNCERTAINTIES AND EXPECTED SPECTRA

