# Stress testing the Standard Model at colliders: from Amplitudes to Events

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Oxford Theoretical Studies of Particles and Strings retreat, March 2024



#### The QCD&collider group

Gavin





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Jack

#### Jasmine







Xiao

Federica





Giulio

#### Radek





Peter

#### The LHC: stress-testing the Standard Model

#### What have we learned so far Overall, the Standard Model works very well

#### Standard Model Total Production Cross Section Measurements



# What have we learned so far

Many "vanilla" BSM scenarios excluded at the EW scale **10 TeV** 

CMS String resonance # Arresonace # Mage resonance # Mage reso	Overview of CMS EXO results	15-73 Hel 1 15-54 Hel 12-2014 (Det 4) 15-14 Hel 12-2014 (Det 4) 15-14 Hel 1 15-14 Hel 1 15-73 Hel 10-2014 (Det 4) 15-73 Hel 10-2014 (Det 4) 15-74 Hel	647 (2) 247 (2+ 24) 259 (2+ 24) 259 (2+ 24)	March 2023 35 fb <sup>-1</sup> 35 fb <sup>-1</sup> 35 fb <sup>-1</sup> 35 fb <sup>-1</sup> 35 fb <sup>-1</sup> 35 fb <sup>-1</sup>			SUSY		
H = 4, predestation (risk), $(\frac{1}{2}, + 80, -10) = 0.210, 0.01$ p = 27(y + 3) p = 27	ABIT-R. STORY (10) LINKO UK, # 40         BERNET (10) LINKO UK, # 40	۲۰۰۰         ۲۰۰۰۰         ۲۰۰۰۰         ۲۰۰۰۰         ۲۰۰۰۰         ۲۰۰۰۰         ۲۰۰۰۰         ۲۰۰۰۰         ۲۰۰۰۰         ۲۰۰۰۰         ۲۰۰۰۰        ۲۰۰۰۰         ۲۰۰۰۰         ۲۰۰۰۰         ۲۰۰۰۰         ۲۰۰۰۰         ۲۰۰۰۰         ۲۰۰۰۰۰         ۲۰۰۰۰۰         ۲۰۰۰۰         ۲۰۰۰۰ <td< th=""><th>4494 (20) 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2</th><th><math display="block">\begin{array}{c} \begin{array}{c} \mu_{1}\mu_{2}\mu_{2}\\ \hline \textbf{ATLAS SUSY Sea}\\ \hline \textbf{March 2023}\\ \hline \textbf{Model}\\ \hline \\ \hline</math></th><th>rches* - 95% CLSignature<math>\int \mathcal{L} d</math><math>0 e, \mu</math>2-6 jets<math>E_T^{miss}</math><math>0 e, \mu</math>2-6 jets<math>E_T^{miss}</math><math>0 e, \mu</math>2-6 jets<math>E_T^{miss}</math><math>1 e, \mu</math>2-6 jets<math>E_T^{miss}</math><math>1 e, \mu</math>2-6 jets<math>E_T^{miss}</math><math>0 e, \mu</math>7-11 jets<math>E_T^{miss}</math><math>0 e, \mu</math>6 jets<math>E_T^{miss}</math><math>0 e, \mu</math>3 b<math>E_T^{miss}</math><math>0 e, \mu</math>2 b<math>E_T^{miss}</math></th><th>Lower <math>f_{t} [f_{b}^{-1}]</math> 139 <math>\frac{\bar{q}}{\bar{q}}</math> [13 139 <math>\frac{\bar{q}}{\bar{g}}</math> [13 139 <math>\frac{\bar{g}}{\bar{g}}</math> 139 <math>\frac{\bar{g}}{\bar{g}}</math> 139 <math>\frac{\bar{g}}{\bar{g}}</math> 139 <math>\frac{\bar{g}}{\bar{g}}</math> 139 <math>\frac{\bar{g}}{\bar{g}}</math> 139 <math>\frac{\bar{g}}{\bar{g}}</math> 139 <math>\frac{\bar{g}}{\bar{g}}</math> 139 <math>\frac{\bar{g}}{\bar{g}}</math> 139 <math>\frac{\bar{g}}{\bar{g}}</math> 139 <math>\frac{\bar{g}}{\bar{g}}</math></th><th>Limits Mass limit * Bx Degen.] 0.9 Forbidde</th><th>1.8 1.15-1. 15 1.25 2.255</th><th>5 <math>m(\tilde{x}_1^0) = 400 \text{ GeV}</math> <math>m(\tilde{q}) = m(\tilde{x}_1^0) = 5 \text{ GeV}</math> 95 <math>m(\tilde{x}_1) = 5 \text{ GeV}</math> 95 <math>m(\tilde{x}_1) = 1000 \text{ GeV}</math> 2.2 <math>m(\tilde{x}_1^0) = 000 \text{ GeV}</math> 2.2 <math>m(\tilde{x}_1^0) = 700 \text{ GeV}</math> 97 <math>m(\tilde{x}_1^0) = 200 \text{ GeV}</math> <math>m(\tilde{x}_1) = 100 \text{ GeV}</math> 97 <math>m(\tilde{x}_1^0) = 200 \text{ GeV}</math> <math>m(\tilde{x}_1^0) = 500 \text{ GeV}</math> <math>m(\tilde{x}_1^0) = 300 \text{ GeV}</math></th></td<>	4494 (20) 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2	$\begin{array}{c} \begin{array}{c} \mu_{1}\mu_{2}\mu_{2}\\ \hline \textbf{ATLAS SUSY Sea}\\ \hline \textbf{March 2023}\\ \hline \textbf{Model}\\ \hline \\ \hline$	rches* - 95% CLSignature $\int \mathcal{L} d$ $0 e, \mu$ 2-6 jets $E_T^{miss}$ $0 e, \mu$ 2-6 jets $E_T^{miss}$ $0 e, \mu$ 2-6 jets $E_T^{miss}$ $1 e, \mu$ 2-6 jets $E_T^{miss}$ $1 e, \mu$ 2-6 jets $E_T^{miss}$ $0 e, \mu$ 7-11 jets $E_T^{miss}$ $0 e, \mu$ 6 jets $E_T^{miss}$ $0 e, \mu$ 3 b $E_T^{miss}$ $0 e, \mu$ 2 b $E_T^{miss}$	Lower $f_{t} [f_{b}^{-1}]$ 139 $\frac{\bar{q}}{\bar{q}}$ [13 139 $\frac{\bar{q}}{\bar{g}}$ [13 139 $\frac{\bar{g}}{\bar{g}}$ 139 $\frac{\bar{g}}{\bar{g}}$ 139 $\frac{\bar{g}}{\bar{g}}$ 139 $\frac{\bar{g}}{\bar{g}}$ 139 $\frac{\bar{g}}{\bar{g}}$ 139 $\frac{\bar{g}}{\bar{g}}$ 139 $\frac{\bar{g}}{\bar{g}}$ 139 $\frac{\bar{g}}{\bar{g}}$ 139 $\frac{\bar{g}}{\bar{g}}$ 139 $\frac{\bar{g}}{\bar{g}}$	Limits Mass limit * Bx Degen.] 0.9 Forbidde	1.8 1.15-1. 15 1.25 2.255	5 $m(\tilde{x}_1^0) = 400 \text{ GeV}$ $m(\tilde{q}) = m(\tilde{x}_1^0) = 5 \text{ GeV}$ 95 $m(\tilde{x}_1) = 5 \text{ GeV}$ 95 $m(\tilde{x}_1) = 1000 \text{ GeV}$ 2.2 $m(\tilde{x}_1^0) = 000 \text{ GeV}$ 2.2 $m(\tilde{x}_1^0) = 700 \text{ GeV}$ 97 $m(\tilde{x}_1^0) = 200 \text{ GeV}$ $m(\tilde{x}_1) = 100 \text{ GeV}$ 97 $m(\tilde{x}_1^0) = 200 \text{ GeV}$ $m(\tilde{x}_1^0) = 500 \text{ GeV}$ $m(\tilde{x}_1^0) = 300 \text{ GeV}$
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utble         [Link]         1.0         #           utble         [Link]         1.0         #           Utble         [Link]         1.0         #           Utble         [Link]         1.0         #           Utble         [Link]         [Link]         #<	881-8878 (1993) 811-8878 (1993) 811-88		2 	$\begin{array}{c} \chi_{1}\chi_{1} \; \text{va}\; \ell_{L}/\tilde{r} \\ \chi_{1}\chi_{1} \; \text{va}\; \ell_{L}/\tilde{r} \\ \tilde{t}, \tilde{\tau}, \tilde{\tau}, \tilde{\tau}, \tilde{\tau}, \tilde{t}_{1}^{0} \\ \tilde{\ell}_{LR}\tilde{\ell}_{LR}, \tilde{\ell} \rightarrow \tilde{\ell}_{1}^{0} \\ \tilde{H}\tilde{H}, \tilde{H} \rightarrow h\tilde{G}/Z\tilde{G} \end{array}$ Direct $\tilde{\chi}_{1}^{\dagger}\tilde{\chi}_{1}^{-} \text{ prod., long-lived } \tilde{\chi}_{1}^{\dagger} \\ Stable \tilde{g} \text{ R-hadron} \\ \text{Metastable } \tilde{g} \text{ R-hadron}, \tilde{g} \rightarrow qq\tilde{\chi}_{1}^{0} \\ \tilde{\ell}, \tilde{\ell} \rightarrow \ell\tilde{G} \end{array}$	$\begin{array}{cccc} 2 e, \mu & E_T^{miss} \\ 2 r & E_T^{miss} \\ e, \mu & 0 \text{ jets } E_T^{miss} \\ e, \mu & 2 \text{ l jet } E_T^{miss} \\ 0 e, \mu & 2 \text{ l jet } E_T^{miss} \\ 0 e, \mu & 2 \text{ l set } E_T^{miss} \\ 0 e, \mu & 2 \text{ large jets } E_T^{miss} \\ 2 e, \mu & 2 \text{ large jets } E_T^{miss} \\ \end{array}$ Disapp. trk 1 jet $E_T^{miss}$ pixel dE/dx $E_T^{miss}$ pixel dE/dx $E_T^{miss}$ pixel dE/dx $E_T^{miss}$ pixel dE/dx $E_T^{miss}$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	L, <sup>7</sup> R,L] 0.16-0.3 0.12-0.39 0.256 0.13-0.23 0.29-0.88 0.55 0.45-0.93 0.77 0.66 0.21 (ĝ) =10 ns] 0.34 0.36		$\begin{array}{c} \mathfrak{m}(\ell,\tilde{r})=0.5(\mathfrak{m}(\ell_1)+\mathfrak{m}(\ell_1))\\ \mathfrak{m}(\tilde{r})=0\\ \mathfrak{m}(\tilde{r})=0\\ \mathfrak{m}(\tilde{r})=0\\ \mathfrak{m}(\tilde{r})=0\\ \mathfrak{m}(\tilde{r})-\mathfrak{m}(\tilde{r})=10 \   \mathrm{Gev}\\ BR(\tilde{r}_1^0\to Z\tilde{G})=1\\ BR(\tilde{r}_1^0\to Z\tilde{G})=1\\ BR(\tilde{r}_1^0\to Z\tilde{G})=BR(\tilde{r}_1^0\to h\tilde{G})=0.5\\ \hline\\ \hline\\ Pure Wino\\ Pure Wino\\ Pure Wino\\ Pure Wino\\ Pure Gym\\ Signion\\ 2.05\\ 2.2 \\ \mathfrak{m}(\tilde{r}_1^0)=100 \   \mathrm{GeV}\\ \mathfrak{r}(\tilde{r})=0.1 \   \mathrm{ns}\\ \mathfrak{r}(\tilde{r})=10 \   \mathrm{ns}\\ \mathfrak{r}(\tilde{r})=10 \   \mathrm{ns} \end{array}$
0.001 Selection of observed exclusion limits at 95% C.L. (theory uncertainties are not includes)	n uuu 100 Mas Scale [FeV] <b>1 Te</b>	V		$\begin{array}{c} \tilde{\chi}_{1}^{\dagger} \tilde{\chi}_{1}^{T} / \tilde{\chi}_{1}^{0}, \tilde{\chi}_{1}^{\dagger} \rightarrow Z\ell \rightarrow \ell\ell\ell \\ \tilde{\chi}_{1}^{\dagger} \tilde{\chi}_{1}^{0} \tilde{\chi}_{2}^{0} \rightarrow WWZ\ell\ell\ell\ellrv \\ \tilde{\chi}_{1}^{\dagger} \tilde{\chi}_{1}^{0} \tilde{\chi}_{1}^{0} \rightarrow qqq \\ \tilde{\chi}_{1}^{\dagger} \tilde{\chi}_{1}^{0} \ell \delta \\ \tilde{\chi}_{1}^{\dagger} \tilde{\chi}_{1}^{0} - \hbar\delta \\ \tilde{\chi}_{1}^{\dagger} \tilde{\chi}_{1}^{0} - \hbar\delta \\ \tilde{\chi}_{1}^{\dagger} \tilde{\chi}_{1}^{0} - \hbar\delta \\ \tilde{\chi}_{1}^{\dagger} \tilde{\chi}_{2}^{0} / \tilde{\chi}_{1}^{0} \chi_{1}^{0} \rightarrow \delta \\ \tilde{\chi}_{1}^{\dagger} / \tilde{\chi}_{2}^{0} / \tilde{\chi}_{1}^{0} \chi_{1}^{0} \rightarrow \delta \\ \tilde{\chi}_{1}^{\dagger} / \tilde{\chi}_{2}^{0} / \tilde{\chi}_{1}^{0} \chi_{1}^{0} \rightarrow \delta \\ \tilde{\chi}_{1}^{\dagger} / \tilde{\chi}_{2}^{0} / \tilde{\chi}_{1}^{0} \chi_{1}^{0} \rightarrow \delta \\ \tilde{\chi}_{1}^{\dagger} / \tilde{\chi}_{2}^{0} / \tilde{\chi}_{1}^{0} \chi_{1}^{0} \rightarrow \delta \\ \tilde{\chi}_{1}^{\dagger} / \tilde{\chi}_{2}^{0} / \tilde{\chi}_{1}^{0} \chi_{1}^{0} \rightarrow \delta \\ \tilde{\chi}_{1}^{\dagger} / \tilde{\chi}_{2}^{0} / \tilde{\chi}_{1}^{0} \chi_{1}^{0} \rightarrow \delta \\ \tilde{\chi}_{1}^{\dagger} / \tilde{\chi}_{2}^{0} / \tilde{\chi}_{1}^{0} \chi_{1}^{0} \rightarrow \delta \\ \tilde{\chi}_{1}^{\dagger} / \tilde{\chi}_{2}^{0} / \tilde{\chi}_{1}^{0} \chi_{1}^{0} \rightarrow \delta \\ \tilde{\chi}_{1}^{\dagger} / \tilde{\chi}_{2}^{0} / \tilde{\chi}_{1}^{0} \chi_{1}^{0} \rightarrow \delta \\ \tilde{\chi}_{1}^{0} / \tilde{\chi}_{1}^{0} / \tilde{\chi}_{1}^{0} \chi_{1}^{0} \rightarrow \delta \\ \tilde{\chi}_{1}^{0} / \tilde{\chi}_{1}^{0} / \tilde{\chi}_{1}^{0} \chi_{1}^{0} \rightarrow \delta \\ \tilde{\chi}_{1}^{0} / \tilde{\chi}_{1}^{0} / \tilde{\chi}_{1}^{0} \chi_{1}^{0} \rightarrow \delta \\ \tilde{\chi}_{1}^{0} / \tilde{\chi}_{1}^{0} / \tilde{\chi}_{1}^{0} \chi_{1}^{0} \chi_$	$\begin{array}{cccc} 3 \ e, \mu & & \\ 4 \ e, \mu & 0 \ \text{jets} & E_T^{\text{miss}} & \\ 4 \ -5 \ \text{large jets} & & \\ & 4 \ -5 \ \text{large jets} & \\ & & \\ $	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$ \begin{array}{    l  l  l  l  l  l  l  l  l  l  l  l $	1.55 1.3 1 0.4-1.45 1.6	Pure Wino $m(\tilde{x}_{1}^{0})=200 \text{ GeV}$ Large $\lambda_{112}^{\prime}$ $m(\tilde{x}_{1}^{0})=200 \text{ GeV}$ bino-like $m(\tilde{x}_{1}^{0})=200 \text{ GeV}$ bino-like $m(\tilde{x}_{1}^{0})=500 \text{ GeV}$ $BR(\tilde{t}_{1} \rightarrow be/b\mu)>20\%$ $BR(\tilde{t}_{1} \rightarrow de/b\mu)>20\%$ $BR(\tilde{t}_{1} \rightarrow de/b\mu)=100\%$ , cos $d_{t}=1$ Pure higgsino Mass scale [TeV]

#### 1 TeV

## Is the SM established? Can we go home?

Ultra-precise data / theory comparison: some tensions

Example: W mass

Input in the SM. But enters radiative corrections and the set of t

Overview of m<sub>w</sub> Measurements LEP Combination PR. 532, p119-244, (2013) **ATLAS** Preliminary  $\sqrt{s} = 7 \text{ TeV}, 4.6 \text{ fb}^{-1}$ D0 (Run 2) PRL 108, p151804 (2012) CDF (Run 2) Science 376, 6589, p170 (2022) LHCb 2022 JHEP 01, 036 (2022) **ATLAS 2017** EPJ-C 78-2, p110 (2018) Measurement Stat. Unc. Total Unc. **ATLAS 2023** this work SM Prediction 80200 80300 80400 m<sub>w</sub> [MeV]

Require exquisite theoretical control over Z/W bosons predictions

Testing ground for "precision" programme

similar situation:  $(g\mathchar`-2)_\mu$ 

#### Is the SM established? Can we go home?

 $\begin{aligned} \mathcal{Z} &= -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} \\ &+ i F \mathcal{B} \mathcal{F} \\ &+ \mathcal{F} \mathcal{B} \mathcal{F} \\ &+ \mathcal{F} \mathcal{B} \mathcal{F}_{j} \mathcal{F}_{j} \mathcal{P} + h.c. \end{aligned}$ Yukawa sector: exploration  $+\left|\mathcal{D}_{\mathcal{M}}\varphi\right|^{2}-\bigvee(\phi)$ only just beginning Gauge sector: well-studied, Higgs potential: basically but now with a scalar field unexplored



### Moving forward

We are only at the beginning of the LHC programme



- So far: only about 5% of data
- Big jump in statistics, no big jump in energy → precision data/theory comparison more and more important

#### Collider predictions: how to get there



proton, Q~GeV, non perturbative High-energy scattering, perturbative

From high-energy to hadrons:  $PS \rightarrow Silvia$ 

#### Collider predictions: how to get there





#### Collider predictions: how to get there $d\sigma = \int dx_1 dx_2 f(x_1) f(x_2) d\sigma_{\text{part}}(x_1, x_2) F_J(1 + \mathcal{O}(\Lambda_{\text{QCD}}^n / Q^n))$ $+\alpha_s^3\Delta\sigma_{N^3LO}+\ldots$ $+\alpha_s^2 \Delta \sigma_{\rm NNLO}$ $(1 + \alpha_s \Delta \sigma_{\rm NLO})$ $\sigma_{\rm part} = \sigma_{\rm LO}$ NLO: N<sup>3</sup>LO: **NNLO:** $\sim 10/20\%$ ~few % ≲ % 3L 00 00 $\alpha_s \sim 0.1$

#### From "nice" theories to real-life QCD

#### "Nice" $\equiv$ N=4 sYM, in the planar limit

<u>Issue</u>: QCD is non (dual) (super) conformal  $\rightarrow$  most cool N=4 tricks don't work





No "N=4"-like tricks (nice integrands...)  $\rightarrow$  "textbook" way  $\otimes$ 

$$\mathcal{A} = \sum_{\substack{j \in \mathcal{A} \\ j \in \mathcal{I}}} \left[ \int_{\mu} d^{d}k \ \partial_{\mu} [v^{\mu} \mathcal{F}(k)] = 0 \\ \text{Integral reduction, IBPs} \right] \qquad \text{Major bottleneck}$$
$$\mathcal{A} = \sum_{\substack{j \in \mathcal{I} \\ j \in \mathcal{I}}} c_{j} \times \text{MI}_{j}$$

#### Some recent results from us & friends





<u>Numerical methods for</u> complex Electroweak amplitudes





	1L	2L	3L
Number of diagrams	6	138	3299
Number of integral topologies	1	2	3
Number of integrals before IBPs and symmetries	209	20935	4370070
Number of master integrals	6	39	486
Size of the Feynman diagrams list [kB]	4	90	2820
Size of the result before integral reduction [kB]	276	54364	19734644
Size of the result in terms of MIs [kB]	12	562	304409
Size of the result in terms of HPLs [kB]	136	380	1195



#### Interesting way forward: intersection theory

$$\mathscr{A} = \sum c_i \times \mathrm{MI}_i$$

Can we "project" over MIs? 
$$c_i = \langle \widetilde{\mathrm{MI}}_i | \mathscr{A} \rangle$$

$$I = \int \prod d^d k_i \frac{\mathcal{N}}{D_1 \dots D_n} \longrightarrow \int \frac{dz_1 \dots dz_n f(\mathbf{z})}{z_1^{n_1} \dots z_n^{n_n}} \times u(\mathbf{z}) = \int_{\mathbf{C}_{\mathbf{R}}} \mathbf{u}(\mathbf{z}) \varphi_{\mathbf{L}}(\mathbf{z})$$

 $\varphi \sim \varphi_L + (d + \omega \wedge)\xi, \quad \omega = d \ln(u) \longrightarrow \varphi_L \in H^n_{\omega}$ 

 $I = \langle \varphi_L C_R] = \sum c_i \langle e_i C_R] \longrightarrow c_i = \langle \varphi_L e_j \rangle [\langle e_j e_i \rangle]^{-1}, \quad |e_i\rangle \in H^n_{-\omega}$ 

# From high-energy scattering to observable final states: parton showers

[Graphics: Gavin Salam]

#### Parton showers in a nutshell



- Hard process: high-scale, perturbative, few legs...
- ... triggering a cascade of QCD radiation, down to hadronic scale
- Up to the very last step, computable from pQCD
- ~ Markov chain semi (but not entirely) classical process

 $d\mathcal{P}_{\text{split}} \sim \alpha_s(k_t) \frac{dk_t^2}{k_t^2} \frac{dz}{z}$ 

#### The two faces of parton showers

- PS 1: a <u>highly tuneable device</u>, able to accurately reproduces datas over a multitude of configurations (*"with 4 parameters I can fit an elephant"*)
  - huge amount of intuition, ingenuity; ab initio control not required

- PS 2: a <u>highly predictive</u> tool, till the hadronic scale. Exquisite control *ab initio* mandatory
  - more and more crucial with the rise of AI/ML: PS predictions are "the truth" that we feed the machines...
  - highly non-trivial, both technically and conceptually, to go beyond the standard "(N)LL" paradigm

#### Recent results from the group & friends

<u>A robust theoretical definition for "PS</u>

<u>accuracy" + the first NLL shower</u>



Gavin P. Salam

DESY Theory Seminar, June 2022

<u>Steps towards higher-order PS</u>

$$\begin{aligned} \mathbb{K}_{q}^{\mathrm{R}}[G_{q},G_{g}] &= \sum_{(A)} \frac{1}{S_{2}} \int d\Phi_{3}^{(A)} P_{1\to3}^{(A)} \left\{ G_{f_{1}}(x \, z_{p} \, (1-z), t_{1,2}) \, G_{f_{2}}(x \, (1-z_{p}) \, (1-z), t_{1,2}) \right. \\ & \times G_{q}(x \, z, t_{12,3}) - G_{f_{12}}(x \, (1-z), t_{12,3}) \, G_{q}(x \, z, t_{12,3}) \right\} \frac{\Delta_{q}(t)}{\Delta_{q}(t_{1,2})} \\ & + \int d\Phi_{3}^{(B)} P_{1\to3}^{(B)} \left\{ G_{g}(x \, (1-z), t_{1,23}) \, G_{g}(x \, z \, (1-z_{p}), t_{2,3}) \right. \\ & \times G_{q}(x \, z \, z_{p}, t_{2,3}) - G_{g}(x \, (1-z), t_{1,23}) \, G_{q}(x \, z, t_{1,23}) \right\} \frac{\Delta_{q}(t)}{\Delta_{q}(t_{2,3})} \, \Theta(t_{2,3} - t_{1,3}) \,, \quad (C.2) \end{aligned}$$



Accurate PS: beyond large-N

#### Recent results from the group & friends



## Making use of all of this: pheno studies

## An example: jet flavour

- <u>What we want</u>
  - back to the start: we want to measure Higgs Yukawa couplings  $\rightarrow$  H  $\rightarrow$  bb(cc...) decay  $\rightarrow$  need to measure "b-quarks"
  - many (B)SM resonances decay to quarks/gluons  $\rightarrow$  want to measure them
- <u>The obvious problem</u>
  - quark and gluons are not asymptotic states... "what is a b-quark?"
- <u>The "in principle" solution...</u>
  - find observables which are strongly correlated to "I had a b-quark in the hard process", yet insensitive to IR physics. *"Jet flavour"*
- <u>The "good" solution</u>
  - something you can compute, to high precision
  - something that experimentalists can measure, to high precision
  - something that is flexible enough for a wide array of studies



## A final disclaimer: only a subset of what the group is doing, more stuff going on

## (heavy-ions, IR subtractions, nonperturbative effects, Higgs physics...)

<u>If you want to hear more: ask us/</u> <u>come to our Friday 1pm-2pm</u> <u>Journal Club (also on zoom)</u>



Thank you very much for your attention

#### From "nice" theories to real-life QCD: issues

#### "Nice" $\equiv$ N=4 sYM, in the planar limit

<u>Issue</u>: QCD is non (dual) (super) conformal → most cool N=4 tricks don't work

 $\underline{\textbf{Issue}}: EW \ particles \twoheadrightarrow leading-N_c \neq planar$ 



#### From "nice" theories to real-life QCD: issues

#### "Nice" $\equiv$ N=4 sYM, in the planar limit

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 $\underline{\textbf{Issue}}: EW \ particles \twoheadrightarrow leading-N_c \neq planar$ 



No parametric enhancement of the "simple" contributions

#### Beyond leading colour: 2



[Chen, Gehrmann, Glover, Huss, Mo (2022)]

#### Beyond leading colour: 3

In some cases:  $1/N_c^2 \ge [i\pi]^2 \rightarrow$  no longer suppressed



 $\delta$ NNLO/NLO: 3%

#### Beyond leading colour: 3

In some cases:  $1/N_c^2 \ge [i\pi]^2 \rightarrow$  no longer suppressed



**δNNLO/NLO:** up to 1.5%

δNNLO/NLO: 3%

[Liu, Melnikov, Penin (2019)]

#### High scale: top quark effects



similar problems for EW corrections, massive W/Z

No "N=4"-like tricks (nice integrands...)  $\rightarrow$  "textbook" way  $\otimes$ 

$$\mathcal{A} = \sum_{\substack{j \in \mathcal{A} \\ j \in \mathcal{I}}} \int_{\mathbf{d}^{d}k} \partial_{\mu} [v^{\mu} \mathcal{F}(k)] = 0}$$
Integral reduction, IBPs
$$\mathcal{A} = \sum_{\substack{j \in \mathcal{L} \\ i \in \mathcal{I}}} c_{i} \times \mathbf{MI}_{i}$$

$$\mathscr{A} = \sum c_i \times \mathrm{MI}_i$$

Computing MIs: binary situation

- 1. They are GPLs  $\rightarrow \bigoplus$
- 2. They are not (elliptic, CY...)  $\rightarrow$  hic sunt leones  $\diamondsuit$

$$\mathscr{A} = \sum c_i \times \mathrm{MI}_i$$

Computing MIs: binary situation

1. They are GPLs  $\rightarrow \bigoplus$ 

2. They are not (elliptic, CY...)  $\rightarrow$  hic sunt mansuefacti leones

Lot of progress in understanding space of required functions, and in providing <u>fast and reliable numerical evaluations</u>

$$\mathbf{E}_4\left(\begin{smallmatrix}n_1 & \dots & n_m \\ c_1 & \dots & c_m \end{smallmatrix}; z, \vec{q}\right) = \int_0^z dt \,\psi_{n_1}(c_1, t, \vec{q}) \,\mathbf{E}_4\left(\begin{smallmatrix}n_2 & \dots & n_m \\ c_2 & \dots & c_m \end{smallmatrix}; t, \vec{q}\right)$$

...ask the experts in the audience!





Lee, von Manteuffel, Schabinger, Smirnov, Smirnov, Steinhauser (2022)

2→2 scattering:



[Bargiela, Chakraborty, FC, von Manteuffel, Tancredi]

2→3:



 $2 \rightarrow n \ge 4$ : 1-loop, numerical

[Abreu, Badger, Brønnum-Hansen, Bargiela, Borowka, Buccioni, Chawdhry, Chen, Chicherin, Czakon, de Laurentis, Dormans, Duhr, Dunbar, Febres-Cordero, Frellesvig, Gambuti, Gehrmann, Gluza, Hartanto, Heinrich, Henn, Ita, Jones, Jehu, Kajda, Kosower, Liu, Lo Presti, Manteuffel, Ma, Maître, Mitev, Mitov, Page, Peraro, Perkins, Poncelet, Schabinger, Sotnikov, Tancredi, Usovitch, Wasser, Weinzierl, Zhang...]



#### $2 \rightarrow n \ge 4$ : 1-loop, numerical





Lee, von Manteuffel, Schabinger, Smirnov, Smirnov, Steinhauser (2022)





2→3:



 $2 \rightarrow n \ge 4$ : 1-loop, numerical





#### $2 \rightarrow n \ge 4$ : 1-loop, numerical



#### From amplitudes to phenomenology

A glimpse of recent results: exploring the  $2\rightarrow 3$  frontier



A rich phenomenology, to be explored

## QCD amplitudes: <u>some</u> open questions

- Some obvious questions:
- better ways of approaching QCD amplitudes
- •how to get more complicated ones (e.g. 5pt, 2 legs off-shell  $\rightarrow$  di-boson)

#### QCD amplitudes: <u>some</u> open questions

Perhaps less obvious questions:• are we really minimising non-physical information?



#### QCD amplitudes: <u>some</u> open questions

Perhaps less obvious questions:• are we choosing the right MIs?

One-loop amplitudes for 
$$e^+e^-$$
 to four partons

Zvi Bern<sup>a,1</sup>, Lance Dixon<sup>b,2</sup>, David A. Kosower<sup>c,3</sup> <sup>a</sup> Department of Physics, University of California, Los Angeles, Los Angeles, CA 90024, USA <sup>b</sup> Stanford Linear Accelerator Center, Stanford University, Stanford, CA 94309, USA <sup>c</sup> Service de Physique Théorique, Centre d'Etudes de Saclay, F-91191 Gif-sur-Yvette cedex, France<sup>4</sup>

Received 13 August 1997; accepted 20 October 1997

1L: "optimal" QCD MIs are not UT...

$$L_{0}(r) = \frac{\ln(r)}{1 - r}, \qquad L_{1}(r) = \frac{L_{0}(r) + 1}{1 - r},$$

$$L_{s-1}(r_{1}, r_{2}) = Li_{2}(1 - r_{1}) + Li_{2}(1 - r_{2}) + \ln r_{1} \ln r_{2} - \frac{\pi^{2}}{6},$$

$$L_{s_{0}}(r_{1}, r_{2}) = \frac{1}{(1 - r_{1} - r_{2})} Ls_{-1}(r_{1}, r_{2}),$$

$$L_{s_{1}}(r_{1}, r_{2}) = \frac{1}{(1 - r_{1} - r_{2})} [Ls_{0}(r_{1}, r_{2}) + L_{0}(r_{1}) + L_{0}(r_{2})],$$

Guiding principle: singularity structure of the QCD amplitude (which is not uniform transcendental...) **Beyond** amplitudes

#### Amplitude techniques: more general

A "revolution" in precision phenomenology: reverse unitarity

## Higgs boson production at hadron colliders in NNLO QCD

Charalampos Anastasiou, Kirill Melnikov

Stanford Linear Accelerator Center, Stanford University, Stanford, CA 94309, USA Received 9 July 2002; received in revised form 4 September 2002; accepted 19 September 2002

$$\begin{vmatrix} p_{1} & \cos \varphi & q_{H} \\ g & g & q_{1} \\ p_{2} & \cos \varphi & q_{2} \end{vmatrix} ^{2} \sim \int \frac{\mathrm{d}^{d} q_{1} \,\mathrm{d}^{d} q_{2} \,\delta(q_{1}^{2}) \delta(q_{2}^{2}) \delta(q_{H}^{2} - m_{H}^{2}) [\cdots]}{[(q_{H} - p_{1})^{2}]^{2} [(q_{2} - p_{2})^{2}]^{2}}$$
$$2i\pi \delta(p^{2} - m^{2}) \rightarrow \frac{1}{p^{2} - m^{2} + i0} - \frac{1}{p^{2} - m^{2} - i0}.$$

Maps phase-space integration into loop integrals  $\rightarrow$  amplitude technology

#### Amplitude techniques: more general

Dilepton Rapidity Distribution in the Drell-Yan Process at Next-to-Next-to-Leading Order in QCD

Charalampos Anastasiou,<sup>1</sup> Lance Dixon,<sup>1</sup> Kirill Melnikov,<sup>2</sup> and Frank Petriello<sup>1</sup> <sup>1</sup>Stanford Linear Accelerator Center, Stanford University, Stanford, California 94309, USA <sup>2</sup>Department of Physics & Astronomy, University of Hawaii, Honolulu, Hawaii 96822, USA (Received 25 June 2003; published 31 October 2003)

$$\frac{d\sigma_{ij}}{2e^{2Y}dY} = \int d\Pi_f |\mathcal{M}_{ij}|^2 \delta\left(e^{2Y} - \frac{E+p_z}{E-p_z}\right).$$

"Observable": constraint in the integration domain

$$\delta(x) \rightarrow \frac{1}{2\pi i} \left[ \frac{1}{x - i0} - \frac{1}{x + i0} \right]$$

Reverse unitarity trick still works, but generalised propagators (also:  $\theta$  functions...)

Can we understand the properties of these "generalised" Feynman integrals (e.g. regularity conditions to fix b.c. of diff. eq. etc)

## Amplitude techniques: more general

Understanding the structure of QCD amplitudes: crucial beyond amplitudes themselves

For example

- Soft/collinear regions: insight into large  $\alpha_s \ln^k(v \ll 1)$  contributions to the cross section. Starting from 2L: non-trivial limits in the non-planar sector [see G. Gambuti's talk]
- Soft/collinear limits beyond leading power: insight into NP regime of QCD
- (Multi)-Regge limit: insight into physics of high-density QCD / NNLL important information for parton distribution functions at N<sup>3</sup>LO [<u>see E.Gardi's talk]</u>
- Non-planar soft/collinear interplay at 2 loops and beyond: expose "Glauber" modes, responsible for potential breaking of collinear factorisation (foundation of perturbative QCD at hadron colliders!) [see G. Gambuti's talk]

and many more!



... also few percent:  $\sim \alpha \rightarrow SM$  at the quantum level (e.g. W-mass)



## The Higgs potential: a long journey $V = \lambda (H^{\dagger}H - v^2)^2 \rightarrow 2\lambda v^2 h^2 + \sqrt{2\lambda} v h^3 + \lambda h^4/4$



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... also few percent: ~  $\alpha \rightarrow SM$  at the quantum level (e.g. W-mass)

$$\varphi_L(\mathbf{z}) = \frac{f(\mathbf{z})}{z_1^{n_1} \dots z_N^{n_N}} dz_1 \dots dz_N$$

$$0 = \int_{C_R} d(u\xi_L) = \int_{C_R} du \wedge \xi_L + ud\xi_L \qquad 0 = \int_{C_R} u\left(\frac{du}{u} \wedge \xi_L + d\right)\xi_L = \int_{C_R} u\nabla_\omega\xi_L$$

$$\int_{C_R} u(\mathbf{z})\varphi_L = \int_{C_R} u(\mathbf{z})(\varphi_L + \nabla_\omega \xi_L) \qquad I = \int_{C_R} u(\mathbf{z})\varphi_L(\mathbf{z}) = \langle \varphi_L | C_R ]$$

$$\left\langle \varphi_{L} | \varphi_{R} \right\rangle = \frac{1}{(2\pi i)^{n}} \int_{X} \varphi_{L} \wedge \varphi_{R} \qquad \left\langle \varphi_{L}^{(1)} | \varphi_{R}^{(1)} \right\rangle = \sum_{p \in \mathcal{P}_{1}} \operatorname{Res}_{z_{1}=p} \left( \psi^{(p)} \varphi_{R}^{(1)} \right) \tag{2.21}$$

where  $\mathcal{P}_1$  is the set of poles of  $\omega_1 = \frac{\partial \log u(z_1)}{\partial z_1}$  (including infinity), and  $\psi_p^{(1)}$  is the solution to the differential equation

$$\boldsymbol{\nabla}_{\omega}\psi^{(p)} = d\psi^{(p)} + \omega_1 \wedge \psi^{(p)} = \varphi_L^{(1)} \tag{2.22}$$