# Ground-state properties an techniques

#### lain Moore (iain.d.moore@jyu.fi)

Department of Physics, University of Jyväskylä, Finland

#### My home base...





From the NuPECC Long Range Plan Draft Executive Summary 2024

The northern most Accelerator Laboratory, some 250 km north of Helsinki.

#### A few remarks

- As this is a school, my goal is to enlighten those of you not already familiar with ground-state properties of nuclei and to present some of the techniques we use to measure them.
- Interrupt me any time for questions or clarifications. We will also have the Q&A session in the evening.
- Hopefully this will be relaxed...after all, I will sometimes use Comic Sans font in my slides instead of Times New Roman and I swap between black and blue!
- For more single-particle/shell structure, see Sean, for emergent collective structure, Magda. Theory will be tackled by Gianluca.







#### 2024 STFC Nuclear Physics Summer School

### (Some) open questions in nuclear physics research

#### Nuclear Astrophysics

- a. How can we better understand the synthesis of heavy elements and the chemical evolution of the visible universe
- b. What is the role of the strong interaction in stellar objects?
- с. ...

d.

#### Nuclear structure / nuclear matter

- a. What shapes can nuclei take, how do nuclear shells evolve, and what role do nuclear correlations play?
- b. What are the limits of the existence of nuclei, and what phenomena arise from open quantum systems?
- c. What are the mechanism bind nuclear reactions and nuclear fission?

#### Hadronic physics

- a. How does the majority of the visible mass of the universe emergye from the almost massless quarks?
- b. What are the properties of the quark-gluon plasma?
- c. How do nuclei and nuclear matter emerge from the underlying fundamental interactions?







#### Properties of a (radioactive) nucleus





### Outline of my two lectures



#### Lecture 1: Nuclear ground state properties

- The atomic mass, nuclear binding energy and nuclear structure
- Nuclear fingerprints on atomic lines the isotope shift
- Hyperfine structure

#### Lecture 2: Techniques of optical spectroscopy and selected results

- A short introduction to radioactive ion beam production
- Laser resonance ionization
- Doppler-free laser spectroscopy
- What can we learn from charge radii, nuclear moments and spins?



- A = mass number = neutron number + proton number
- Z = proton number (element)
- B the binding energy
  - Nuclear (~MeV/u), Atomic electrons (eV...keV)
  - Compare to absolute nucleon mass ~1 GeV
- Usually, atomic mass tables list the mass defect or excess rather than atomic mass:

 $\Delta(^{A}X) = [m(Z,N) - Au]c^{2}$ 

- "u" is the atomic mass unit (Definition: m(<sup>12</sup>C) = 12 u exactly)
- keV/c<sup>2</sup> (1 u = 931494.10242(28) keV/c<sup>2</sup> CODATA2018)

**Nuclear Binding Energy**  $M_{atom}(Z,N) = M_{nuc}(Z,N) + Zm_{\rho} - B_{\rho}(Z)$  $M_{nuc}(Z,N) = Zm_p + Nm_n + \frac{E(Z,N)}{c^2}$ 





## The King James bible for masses: Atomic Mass Table



#### Periodically published collection of atomic masses



W.J. Huang et al., Chinese Phys. C 45, 030002 (2021)





110

100

Nu

Page 3: Mass-excess uncertainties for all nuclei in ground state

NEUTRONS

## Nuclear binding energy





Figures from L.S. Paraschiv et al., Energy Reports 8 (2022) 342; P. Schwerdtfeger et al., The periodic table and the physics that drives it, Nature Reviews Chemistry (2020)



- When a nucleus is formed, the binding energy occurs due to a transformation of a mass quantity into energy.
- More bound = more binding energy
- Nuclear stability the nuclear landscape!
- Theory is needed to predict the limits of stability.
- Nuclear energy

Mass of each and every nuclear state is important!

• Some masses are needed with more precision, some less...



Figure from D. Lunney, Hyp. Int. 240 (2019) 48

~2500 masses experimentally determined (white boxes)

				·· 1 a
	<b>δ</b> m∕m	δm for m=100 u		$V_c = \frac{1}{2\pi} \frac{q}{m} B$
		(µu)	(keV)	
General physics & chemistry	≤ 10 <sup>-5</sup>	1000	1000	
Nuclear structure physics - separation of isobars	≤ 10 <sup>-6</sup>	100	100	
Astrophysics - separation of isomers	≤ 10 <sup>-7</sup>	10	10	MR-TC
Weak interaction studies	≤ 10 <sup>-8</sup>	1	1	
Metrology - fundamental constants Neutrino physics	≤ 10 <sup>-9</sup>	0.1	0.1	Penning traps
CPT tests	≤ 10 <sup>-10</sup>	0.01	0.01	
QED in highly-charged ions - separation of atomic states	≤ 10 <sup>-11</sup>	0.001	0.001	



#### Nuclear masses for astrophysics





"Masses of exotic nuclei", from the Handbook of Nuclear Physics (2022)

- Nuclear masses (binding energies) determine nucleosynthesis pathways
- Modeling the rp process:
  - X-ray bursts and light curves
    proton captures have exponential dependence on mass
- Theoretical models are critically important, e.g., for modeling the rprocess



See nuclear astrophysics lectures by Ann-Cecilie!

### Calculated and experimental S<sub>2n</sub> for erbium





- Stable Er (Z = 68) isotopes from N = 94 to N = 102.
- $S_{2n}$  = two-neutron separation energy
- $S_{2n} = 0$  is the neutron drip line
- $S_{2p}$  = two-proton separation energy

Figure from J. Erler et al., Nature 486 (2012) 509.

#### Mass extrapolations towards exotic nuclei





From "Ion traps in nuclear physics – Recent results and achievements", T. Eronen et al., PPNP 91 (2016) 259

#### Nuclear structure via mass probes



Courtesy of Lukas Nies, PLATAN 2024 conference.



### Highlighting the empirical shell gaps

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#### Properties of a (radioactive) nucleus





### Nuclear properties via the atomic fingerprint



Note: the heaviest of elements, e.g., Lr with Z=103, only theoretical estimates for atomic levels exist!

1 cm<sup>-1</sup> : ~30000 MHz (30 GHz) 1 eV: ~8000 cm<sup>-1</sup> UNIVERSITY OF JYVÄSKYLÄ

#### Isotope shifts of electronic transitions



Relative frequency (MHz)

The isotope shift is the frequency difference in an electronic transition between two isotopes of mass A and A'





#### What causes the isotope shift?

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The shift in the atomic transition frequency between different isotopes of the same element arises due to changes in nuclear mass and nuclear size.



### From the isotope shift to the nuclear size



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### From the isotope shift to the nuclear size





#### Properties of a (radioactive) nucleus





#### Let's return to the atomic structure of uranium...



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### Hyperfine interaction (in free atoms)

Hyperfine interaction = the interaction of nuclear magnetic and electric moments with electromagnetic fields (which are produced at the nucleus by the orbiting electrons)

Lets consider the effect on an atomic orbit of spin J (where J = L + S) – our fine structure orbital.





 $F=I+J, I+J-1, \ldots, \left|I-J\right|$ 

States of the same I and J but coupled to different angular momenta F have slightly different energies

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#### Nuclear magnetic dipole moment





The magnetic dipole moment of a state of spin I = expectation value of the z-component of the dipole operator  $\mu$ :

$$\mu(I) \equiv \langle I, m = I | \hat{\mu}_z | I, m = I \rangle = g I \mu_N$$

Measuring a magnetic moment is an excellent way to learn more about the wavefunction/configuration of the nucleus.

#### The magnetic dipole interaction



$$E(F) = \frac{A}{2}C$$

where C = F(F+1) - I(I+1) - J(J+1)

Access to nuclear spin I (number of hyperfine components) and  $\mu_I$ 

v=∆E/h

v=∆E/h

JYU. Since 1863.

### The electric quadrupole moment



$$eQ = \int_0^\infty \rho_n(\mathbf{r})(3z^2 - r^2) d\tau$$
 A spherical nucleus would have zero Q

Experiments measure the maximum "projection" of the intrinsic quadrupole moment along the quantization axis

Using angular momentum algebra, we get

$$Q_s = Q_0 \frac{3K^2 - I(I+1)}{(I+1)(2I+3)}$$

this assumes a well-defined deformation axis (not always a good approximation)

Note for nuclear spin I =0 and I =1/2 the spectroscopic quadrupole moment vanishes even if the intrinsic shape is deformed.





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The intrinsic Q moment can in turn be related to the quadrupole deformation parameter  $\beta_2$ 



Magda will discuss more on nuclear collectivity!

#### Side-tracking



#### How common is quadrupole deformation? Or, how uncommon are spherical nuclei?



#### The electric quadrupole interaction



#### Let's summarize...



- The hyperfine interaction manifests as a splitting and perturbation of the atomic fine structure lines.
- The energy of the F state is given by:

$$E(F) = \frac{A}{2}C + B \frac{\frac{3}{4}C(C+1) - I(I+1)J(J+1)}{2(2I-1)(2J-1)I \cdot J}$$
  
where  $C = F(F+1) - I(I+1) - J(J+1)$ 

• Experimentally we extract the A and B hyperfine coefficiencts from either the upper or lower fine structure state:

$$A = \frac{\mu_I B_J(0)}{I J} \qquad \qquad A \qquad \qquad B = e Q_s V_{JJ}(0)$$

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## In reality what we are measuring





#### Nuclear magnetic moment

$$\frac{A_1}{A_2} = \frac{\mu_1}{I_1} \frac{I_2}{\mu_2}$$

Reference nuclei (stable isotopes) with accurately known moments

$$\mu_1 = \frac{I_1}{I_2} \frac{A_1}{A_2} \mu_2$$

 $A_1, A_2$  from fit  $\mu_2$  from NMR

- To extract the hyperfine parameters A and B from experimental measurements, the magnetic field B<sub>J</sub> and electric field gradient V<sub>jj</sub> need to be extracted.
- This can be done in principle using atomic theory
- A more elegant way is to use reference nuclei with known moments. *B<sub>j</sub>* and *V<sub>jj</sub>* depend only on electronic structure.

Nuclear quadrupole moment

$$B = e Q V_{JJ}(0)$$



*B* from fit Electric field gradient from atomic physics

### Did anyone notice the little "+" sign 2 slides ago?



PHYSICAL REVIEW A 103, 032826 (2021)

Magnetic octupole moment of <sup>173</sup>Yb using collinear laser spectroscopy

R. P. de Groote <sup>(a)</sup>,<sup>1,\*</sup> S. Kujanpää<sup>(b)</sup>,<sup>1</sup> Á. Koszorús<sup>(a)</sup>,<sup>2</sup> J. G. Li<sup>(b)</sup>,<sup>3</sup> and I. D. Moore<sup>(a)</sup> <sup>1</sup>Department of Physics, University of Jyväskylä, PB 35(YFL) FIN-40351 Jyväskylä, Finland <sup>2</sup>Department of Physics, University of Liverpool, Liverpool L69 7ZE, United Kingdom <sup>3</sup>Institute of Applied Physics and Computational Mathematics, Beijing 100088, China



Defining K = F(F + 1) - I(I + 1) - J(J + 1), this can be written as (truncated at the octupole (k = 3) term):

$$E_F^{(1)} = \frac{AK}{2} + \frac{3B}{4} \frac{K(K+1) - I(I+1)J(J+1)}{(2I(2I-1)J(2J-1))} + \frac{5C}{4} \frac{K^3 + 4K^2 + \frac{4}{5}K(-3I(I+1)J(J+1) + I(I+1) + J(J+1) + 3) - 4I(I+1)J(J+1)}{I(I-1)(2I-1)J(J-1)(2J-1))}$$

- Measurements of the magnetic octupole constant C and moment  $\Omega$  are scarce!
- Currently only measured for about 18 elements, and all stable isotopes.
- Interpretation of such a moment is somewhat lacking and due to a lack of data, not much theoretical progress in over half a century.



Magnetic octupole

## And finally – discovering nuclear states





#### Is it excited or not?

- It might have been too long-lived for nuclear decay spectroscopy methods
- It could have been too low-lying to be separated by precision mass measurements
- Discovered by optical spectroscopy through the hyperfine pattern!
- However optical spectroscopy alone cannot say which state is the ground state and which is the isomer...



Cheal et al. Phys. Rev. C 82, 051302(R)

## Take home messages and outlook to lecture 2





- The nuclear mass and binding energy contain a wealth of structural information, but need supported by other techniques to understand the nature of structural changes.
- The nuclear spin can be determined unambiguously in a nuclear model-independent way e.g., counting the observed number of resonances.
  - assumes sufficient experimental resolution (L2)
- A measurement of hyperfine structure gives access to magnetic dipole and electric quadrupole moments (single particle and collective properties)
- The isotope shift is a gateway to the nuclear size (charge radii).
  - close connection with shell structure, deformation
  - complementary to information from binding energies (L2)





#### Back up material for lecture 1

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The interaction energy depends on angle  $\boldsymbol{\Theta}$ 

$$\mathbf{E} = -\boldsymbol{\mu} \cdot \mathbf{B}_{\mathbf{e}} = -\boldsymbol{\mu} \mathbf{B}_{\mathbf{e}} \cos \theta$$

Since 
$$\mu = g \mathbf{I} \mu_N$$
 and  $\mathbf{B}_e = -(\frac{B_e}{J}) \mathbf{J}$   
then the interaction Hamiltonian can  
be expressed as  
 $H_m = (\frac{g B_e \mu_N}{J}) \mathbf{I} \cdot \mathbf{J} = A \mathbf{I} \cdot \mathbf{J}$ 



The different energy shifts of the different F-states are then

$$\Delta E = \langle IJF \mid H_m \mid IJF \rangle = A \langle \mathbf{I}.\mathbf{J} \rangle$$

Wh

nere: 
$$\langle \mathbf{I}.\mathbf{J} \rangle = \frac{1}{2} \langle F^2 - I^2 - J^2 \rangle = \frac{1}{2} [F(F+1) - I(I+1) - J(J+1)]$$

#### The electric quadrupole interaction







Energy shifts of the F-states are then

$$\Delta E_Q = \frac{B}{4} \frac{\frac{3}{2}C(C+1) - 2I(I+1)J(J+1)}{I(2I-1)J(2J-1)}$$

Where: 
$$C = [F(F+1) - I(I+1) - J(J+1)]$$

The hyperfine factor "B" is measured by experiment:

$$B = eQ_s \langle \frac{\partial^2 V}{\partial Z^2} \rangle = eQ_s V_{JJ}$$

The electric field gradient  $V_{JJ}$  may be calibrated with an isotope with known  $Q_s$ 



Expanding the charge distribution in multipoles:

$$Q_q^n = eZ_{\sqrt{\frac{4\pi}{2n+1}}} \left\langle I \left| r_n^n Y_q^n(\theta_n, \varphi_n) \right| I \right\rangle$$

• Electric monopole =  $eZ\sqrt{4\pi} \langle I | Y_0^0 | I \rangle = eZ$ 

Electric dipole = 
$$eZ\sqrt{\frac{4\pi}{3}}\langle I | r Y_q^1 | I \rangle$$

• Electric quadrupole:

$$Q_q^2 = eZ\sqrt{\frac{4\pi}{5}}r^2Y_q^2$$

= 0



### Charge radii: back to basics

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The majority of modern day experiments apply electromagnetic probes to obtain (absolute) nuclear radius information – preferably point-like, e.g., electrons and muons.

0.90

0.1p

The mean-square charge radius can be defined as:

 $\langle r^2 \rangle = \frac{\int r^2 \rho_{ch}(r) dV}{\int \rho_{ch}(r) dV}$  where  $\rho_{ch}(r)$  is the nuclear charge density distribution.

Extensive studies of stable nuclei with electron scattering experiments revealed the charge density to be nearly constant in the nuclear volume.

The trend of the mean-square charge radii has the form:

$$angle = \frac{3}{5}(r_0A^{\frac{1}{3}})$$
 Thus the rms charge radius  $R = \sqrt{\langle r^2 \rangle}$  was seen to scale with  $A^{\frac{1}{3}}$ 

Great! So we can all go home.....however....



Figure from Hofstadter Nobel Prize lecture, 1961.

 $\langle r^2 \rangle$ 

#### Nuclear shell effects - ``kinky´´ trends!



- The effect of all neutron shell closures for N≥28 is visible as a kink in the charge radii
- Much theoretical effort employed, including ab initio and DFT approaches
- Odd-even staggering effects are probed via isotope shift measurements.

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Neutrons

12.

(126)

(82)

(50)

(28)

(20)

8

(2)

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