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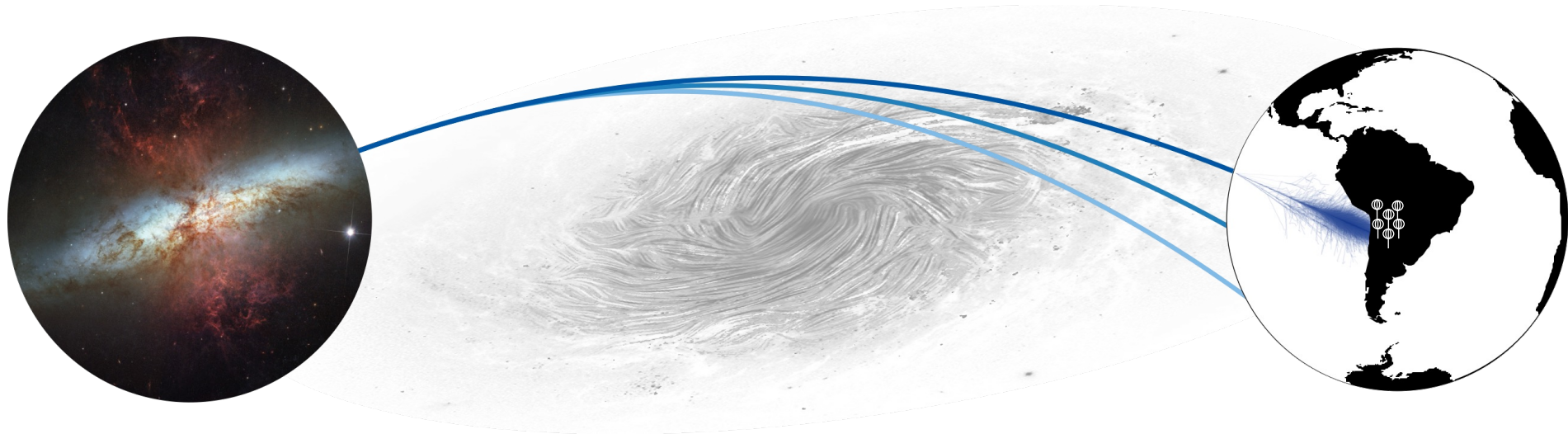
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Applications of Information Field Theory in Astroparticle Physics

MODE Workshop 2024



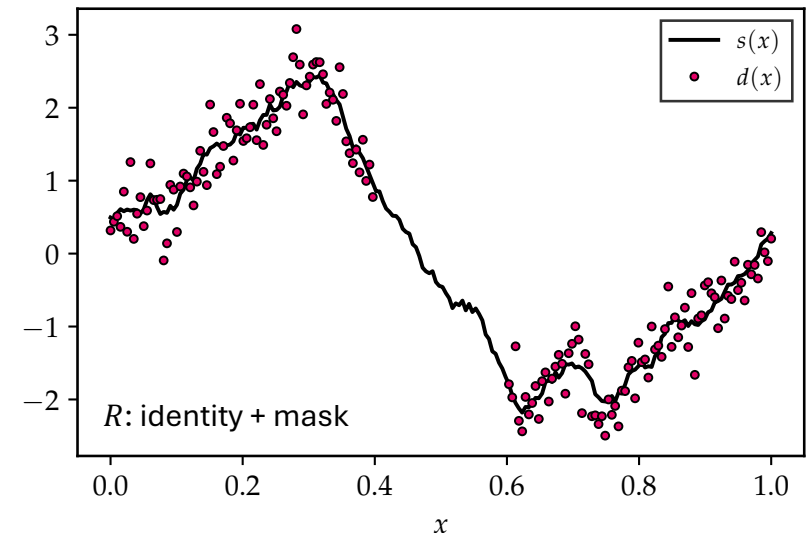
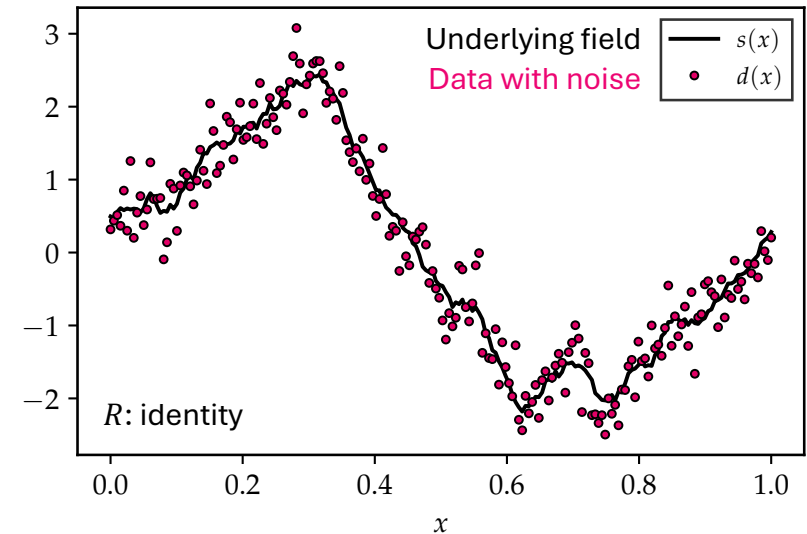
Frederik Krieger, Martin Erdmann, Alex Reuzki, Josina Schulte, Michael Smolka, Maximilian Straub

Information Field Theory (IFT)

- Information theory for fields
- Based on Bayesian statistics
- Use physical assumptions via priors
- Learn **correlation structure** of fields
- Invert measurement: $d = R[s] + n$
 - d : Data
 - R : Response
 - s : Field
 - n : noise

Goal: Infer **posterior distribution** of continuous field with **sparse and noisy data** using physical assumptions via priors

Example for measurement



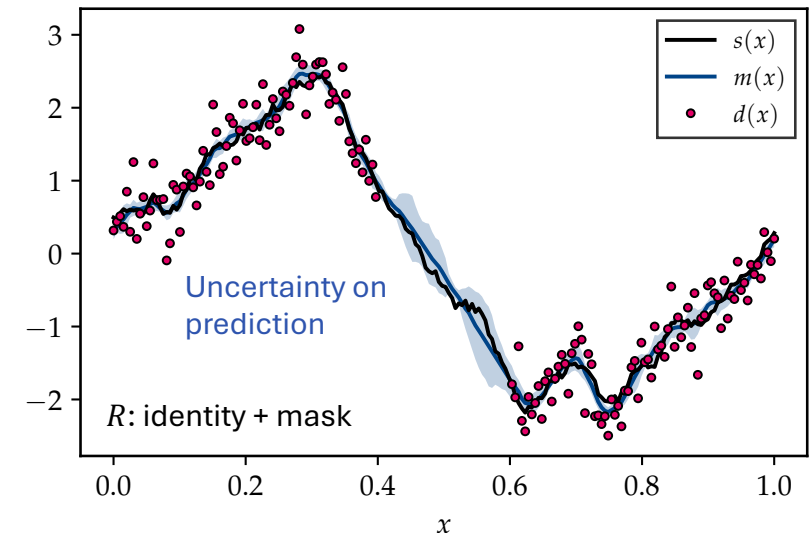
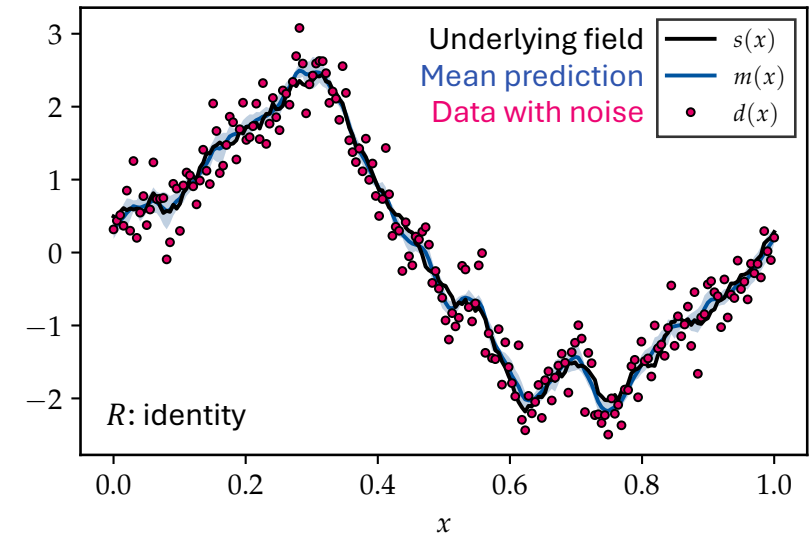
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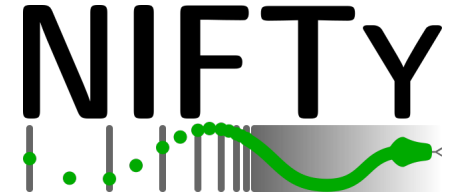
Example for posterior inference



Information Field Theory: Posterior inference

NIFTy

- Numerical Information Field Theory
- Software package written in python
- Multiple inference algorithms, e.g.
 - Metric Gaussian Variational Inference (MGVI)
 - Geometric Variational Inference (geoVI)

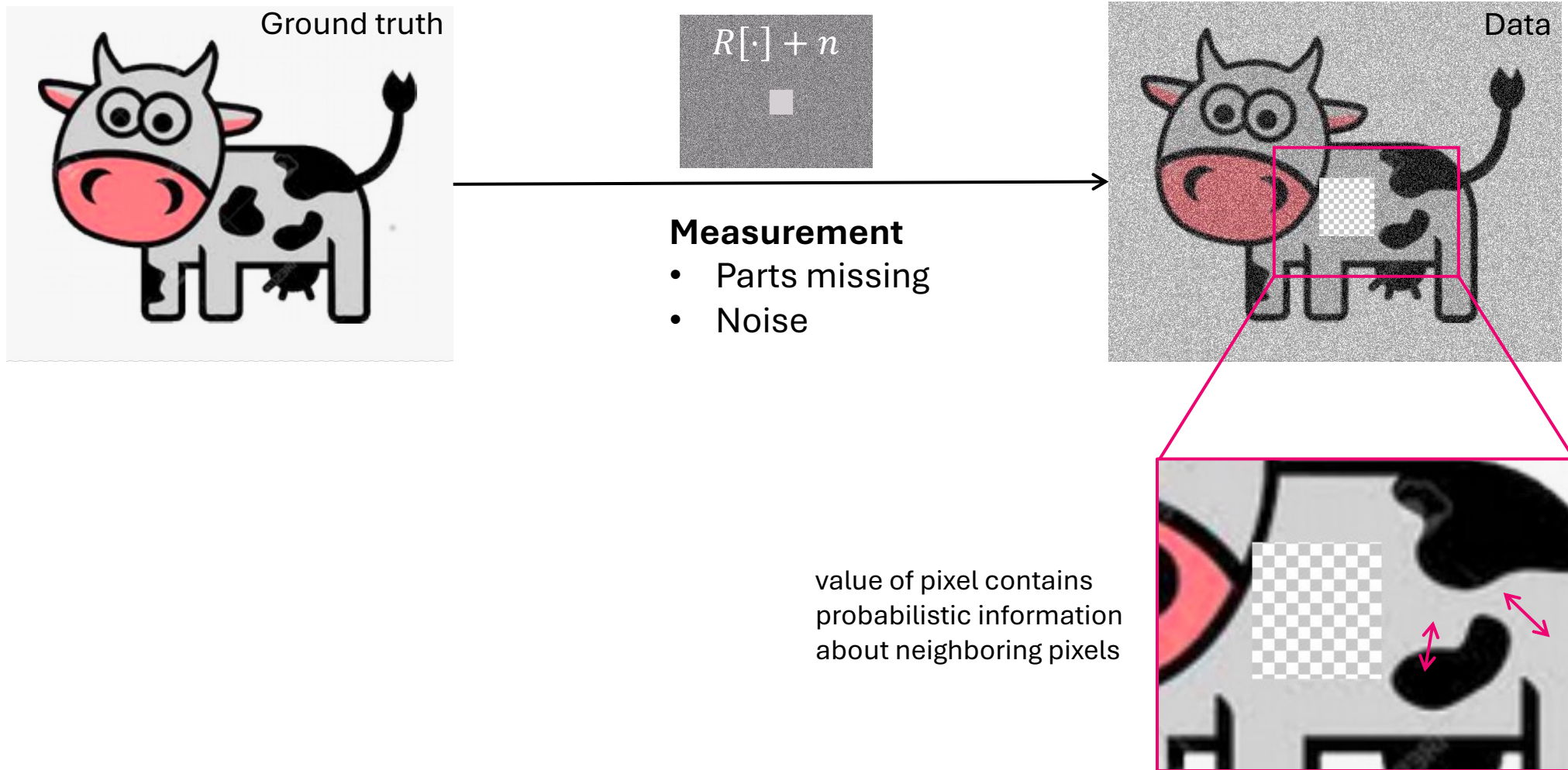


JAX

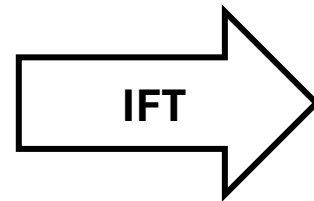
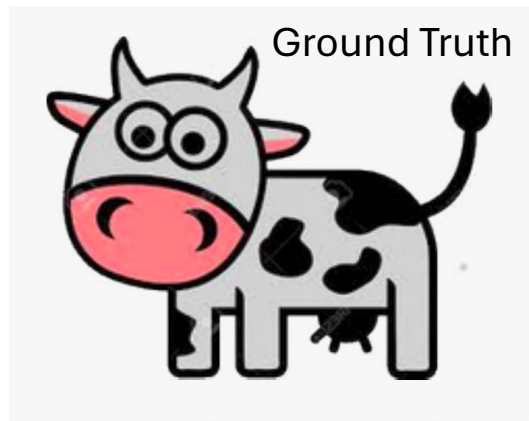
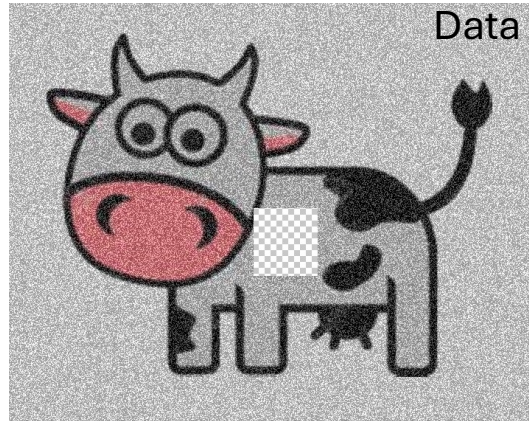
- Open-source machine learning framework from Google
 - NumPy-like syntax
 - Just-In-Time compilation (JIT)
 - Execution on GPUs and TPUs
 - Automatic differentiation
- Forward model “**field** → **data**” can be written as computational graph with NumPy-like syntax and **differentiable**



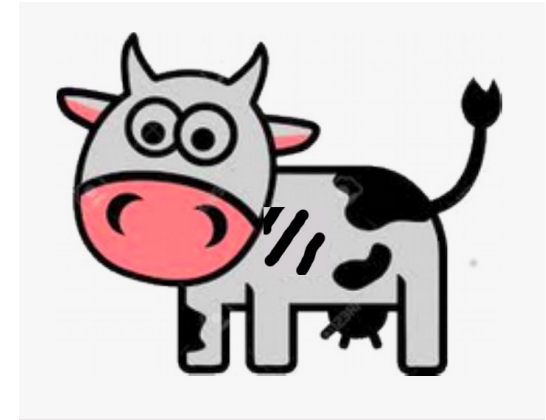
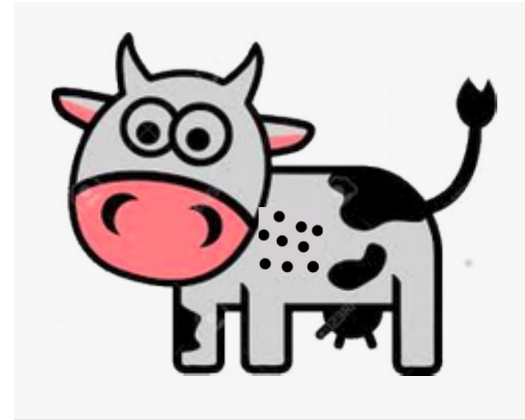
Information Field Theory: Correlation structures



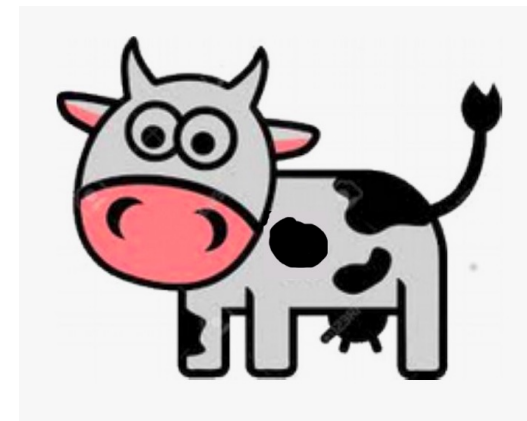
Information Field Theory: Correlation structures



Infer posterior distribution



Low probability

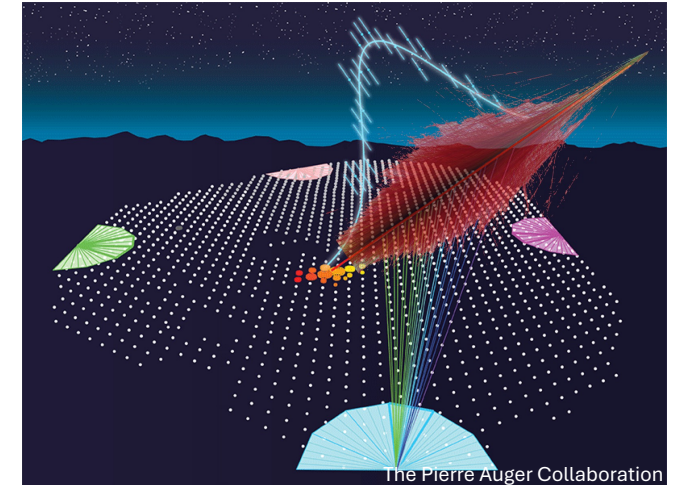


High probability

Applications in astroparticle physics

Ultra-high-energy cosmic rays (UHECRs)

- Charged nuclei moving through the Universe with energies $\geq 10^{18}$ eV that can be detected at Earth
- Origin still uncertain
- Initiate particle shower in atmosphere



1.

Calibration of radio antennas at the Pierre Auger Observatory

- Extensive air showers emit radio signals
- Measurement with radio antennas at the Pierre Auger Observatory in Argentina
- Relative calibration of radio antennas

2.

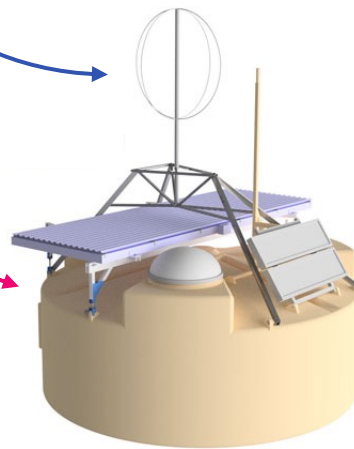
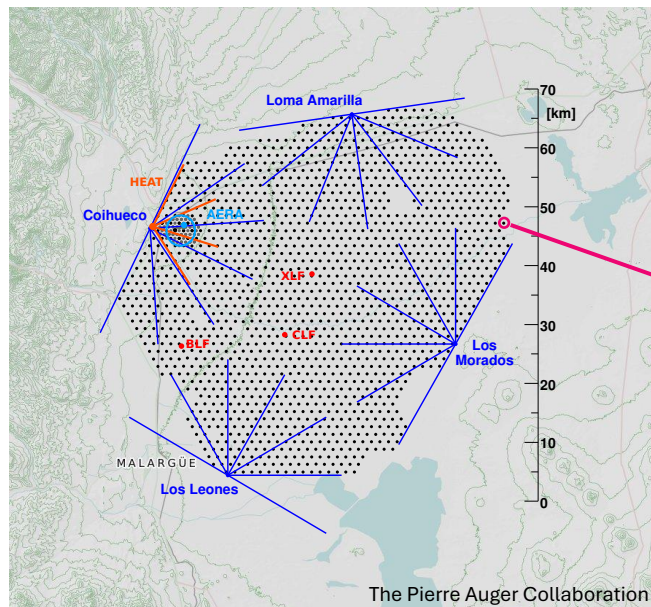
UHECR deflections in the Galactic magnetic field

- Reconstruction of UHECR flux before entering the Galaxy using sparse data of arrival directions measured at Earth
- Pinpoint to possible source directions

Radio detection of air showers at the Pierre Auger Observatory

The Pierre Auger Observatory

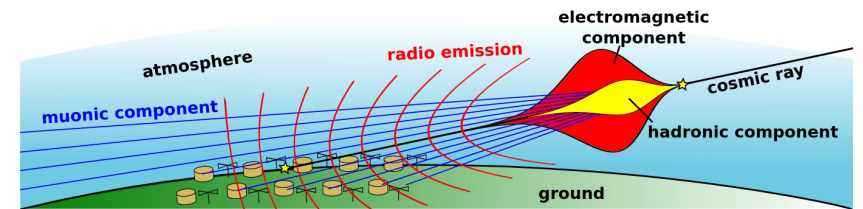
- Located in Argentina
- Hybrid detection of air showers
 - Fluorescence detector
 - Surface detector
 - **Radio detector**



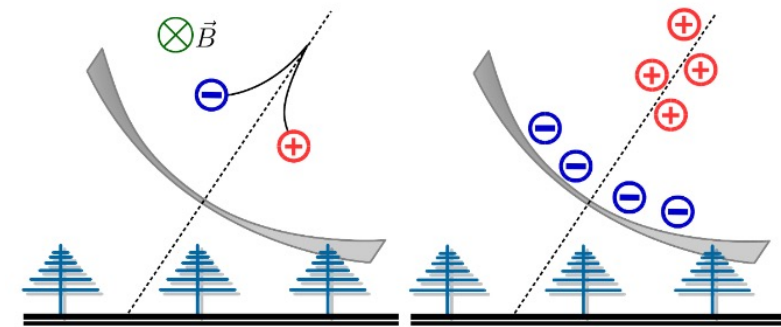
× 1660

Radio emission in air showers

- Geomagnetic effect
 - Deflection of e^- & e^+ in Earth's magnetic field
- Askaryan effect
 - Production of ionized air molecules



E. Holt (2018)

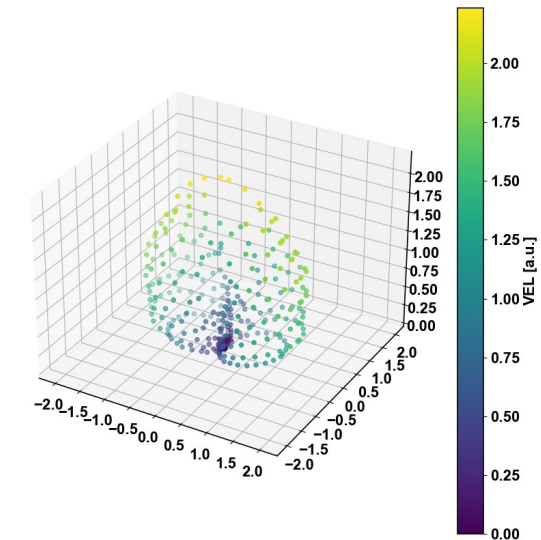


C. Glaser (2017)

Calibration of radio antennas with a drone

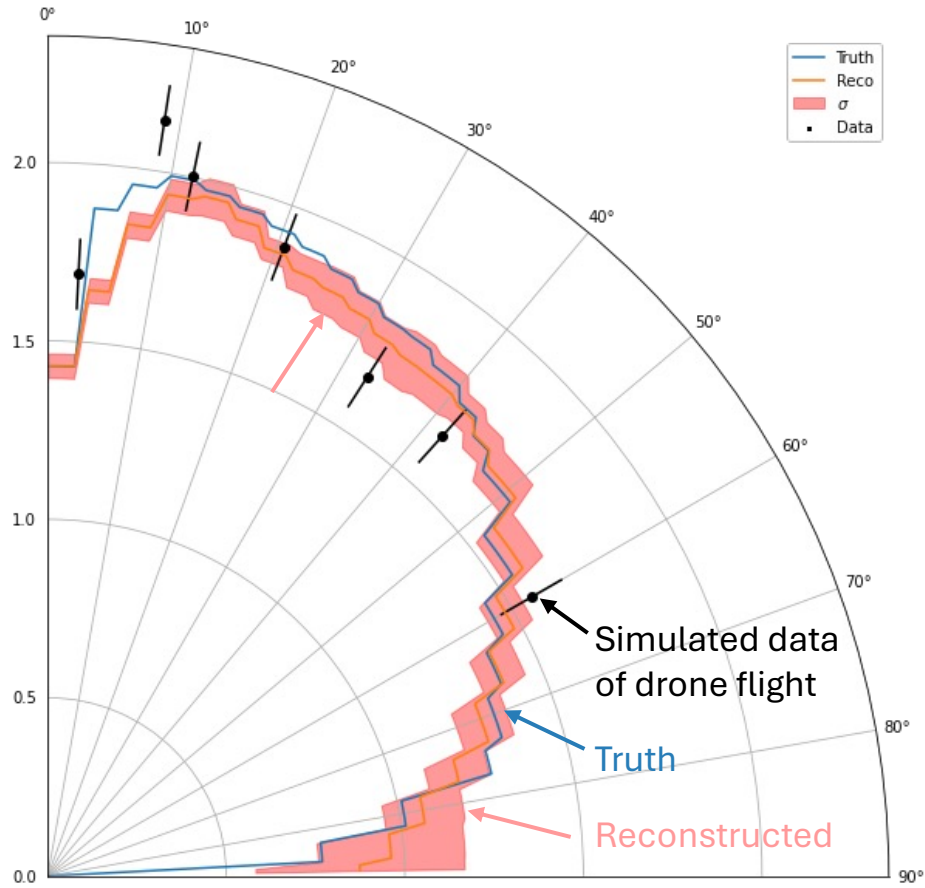
- Direction-dependent sensitivity of antennas (*antenna pattern*)
 - Relates incoming electric field to measured voltage in antenna as *Vector Effective Length* (VEL)
- Emit calibration signal from **drone**
- Measure signal amplitude
 - at various points in the sky
 - for multiple frequencies
 - with different polarizations

→ **Sparse data** on sphere

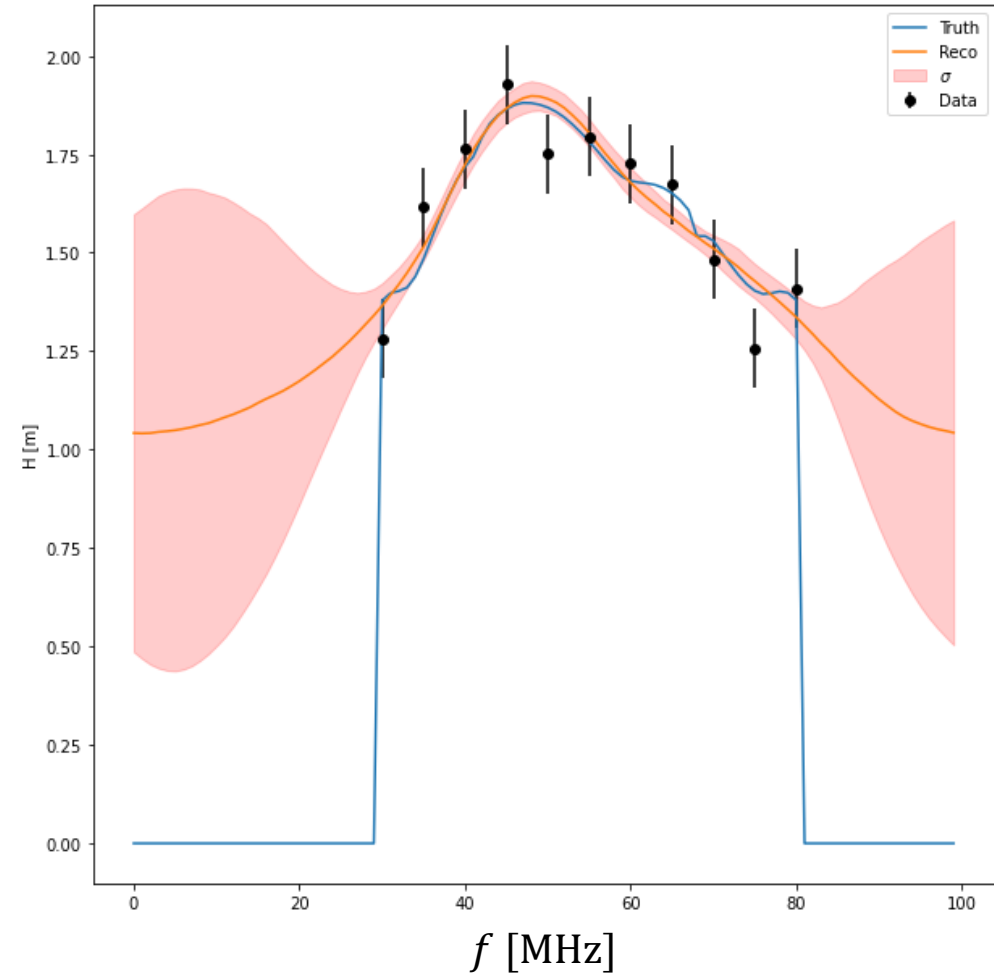


Radio antenna calibration with IFT

Antenna directional dependence



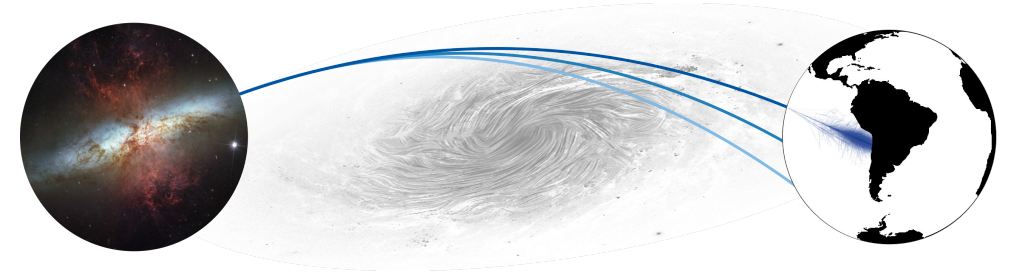
Antenna frequency dependence



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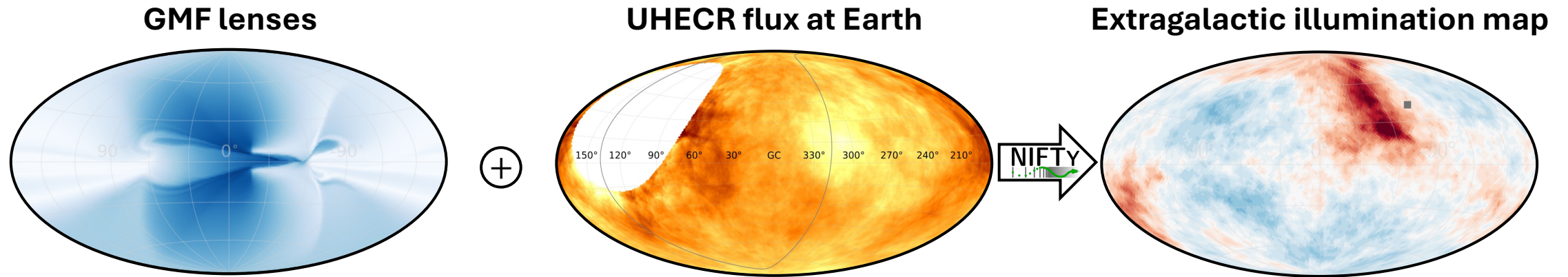
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UHECR deflections in the Galactic magnetic field

- Reconstruction of UHECR flux before entering the Galaxy using sparse data of arrival directions measured at Earth
- Pinpoint to possible source directions

UHECR deflections in the Galactic magnetic field (GMF)



By using GMF lenses and UHECR flux measured at Earth → reconstruct extragalactic illumination map

describe deflections in GMF

UHECR flux arriving at the edge of the Galaxy

UHECR deflections in the Galactic magnetic field: Toy model

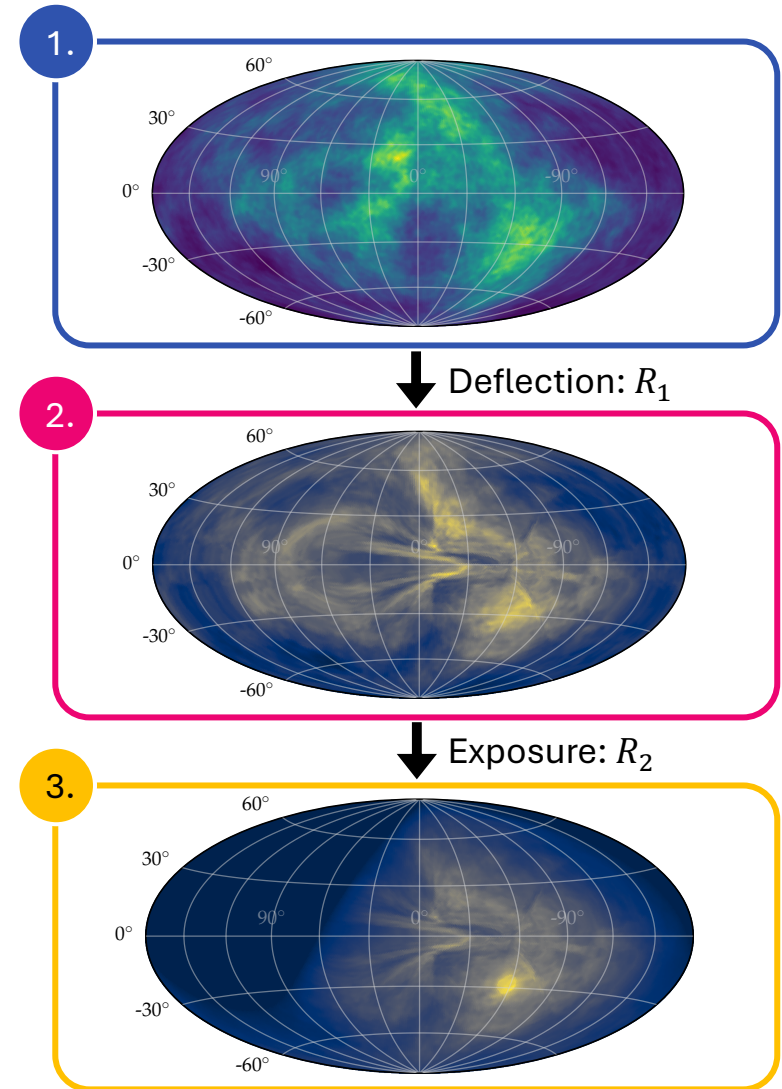
Simplifications

- Composition: only protons
- Fixed energy of $10^{19.2}$ eV
- No turbulent GMF component

Forward model

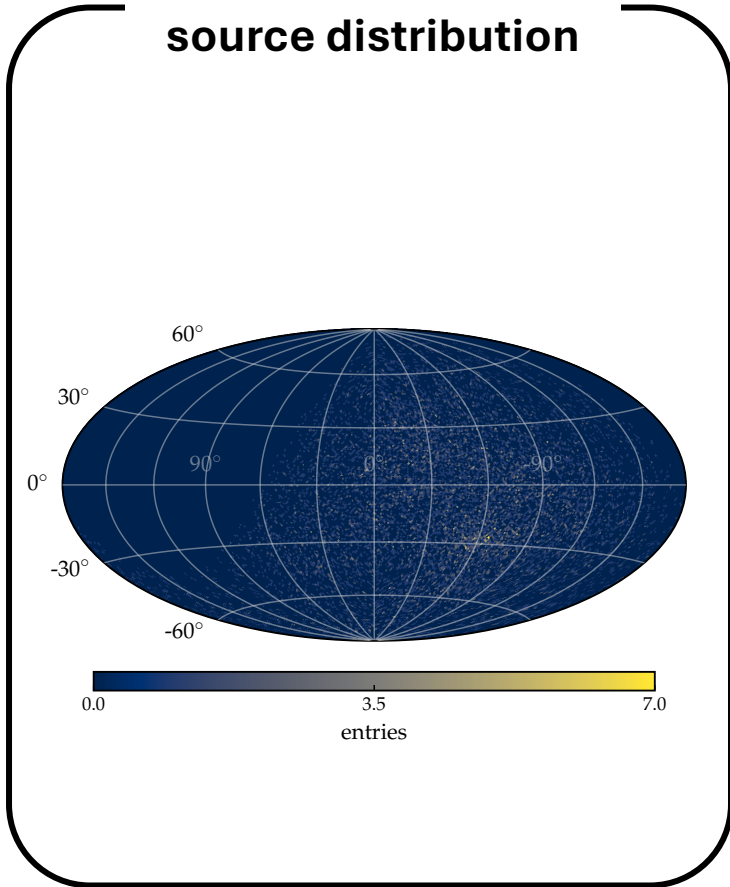
- 1) Sample extragalactic UHECR flux map
- 2) Deflect with GMF lens
- 3) Apply exposure and norm to number of measured events

$$d = R[s] + n \quad R[\cdot] = R_2[R_1[\cdot]]$$

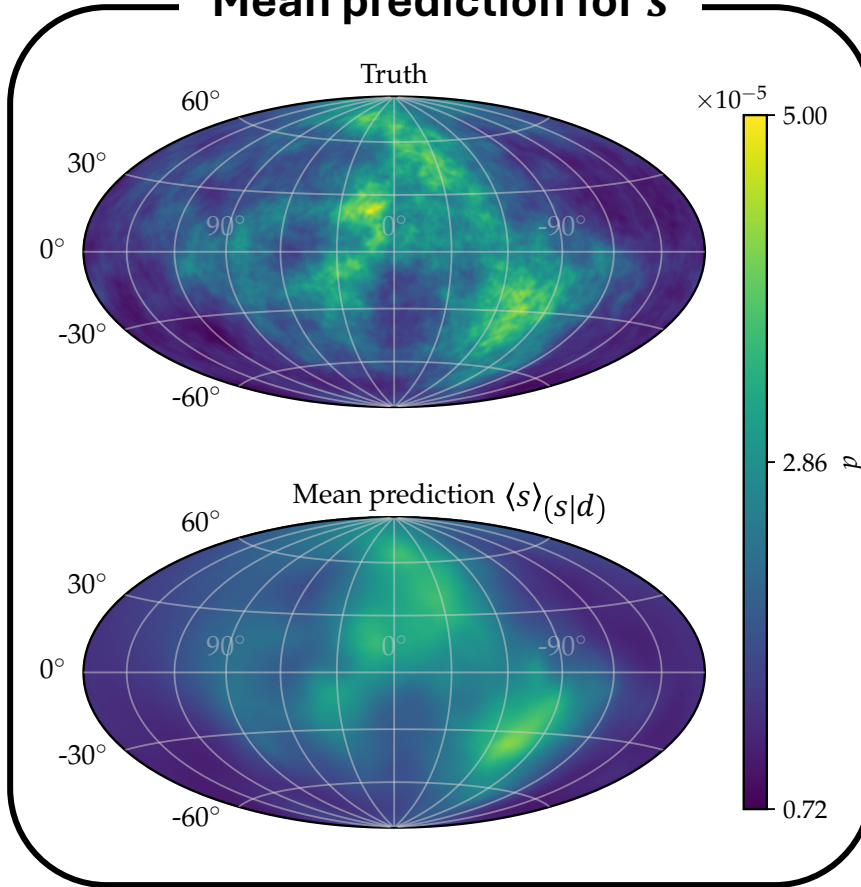


UHECR deflections in the Galactic magnetic field: Exemplary fit

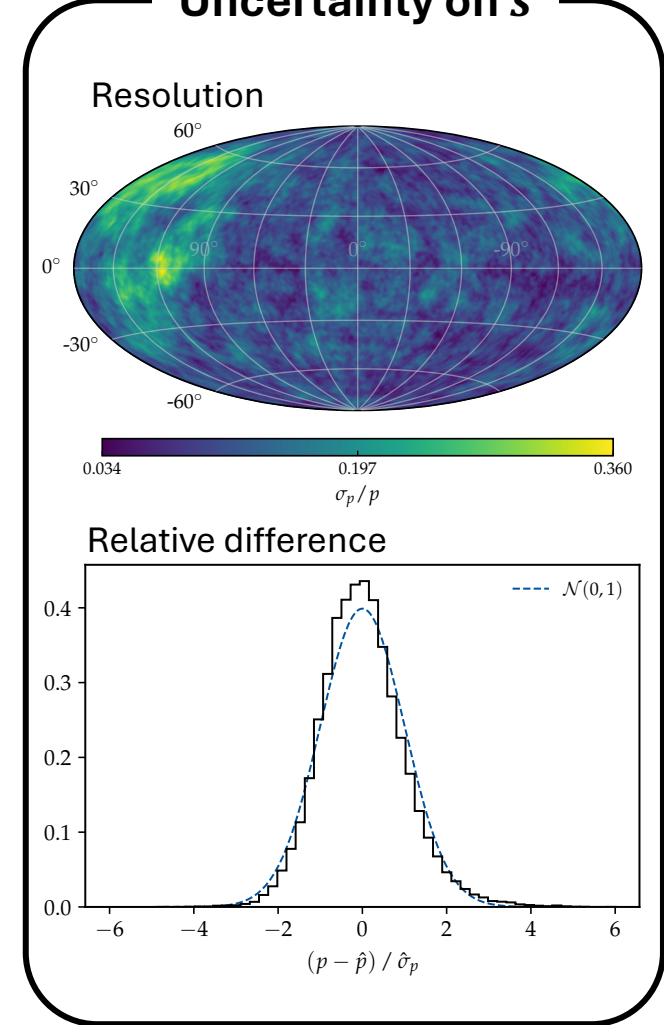
Simulated data d for source distribution



Mean prediction for s



Uncertainty on s



$$d = R[s] + n$$

Summary

- **Information Field Theory** as Bayesian framework to infer posterior distributions of continuous fields using sparse and noisy data
- **NIFTy** with underlying machine-learning framework **JAX** used for Just-In-Time compilation, computation on GPUs and **automatic differentiation**
- Promising applications in both **detector calibration** and **analysis** for astroparticle physics
 - Radio detector at the Pierre Auger Observatory
 - Deflections of ultra-high-energy cosmic rays in the Galactic magnetic field



Backup slides

Calibration strategy

Gain Calibration

Read-out Voltage

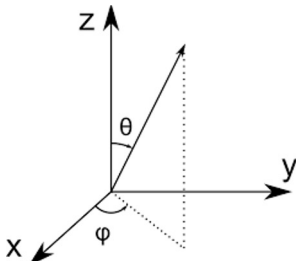
Incoming electric field

$$\mathcal{U}(\Phi, \Theta, f) = \left| \vec{H}_k(\Phi, \Theta, f) \right| \cdot \left| \vec{\mathcal{E}}_k(f) \right|$$

Vector Effective Length (VEL)

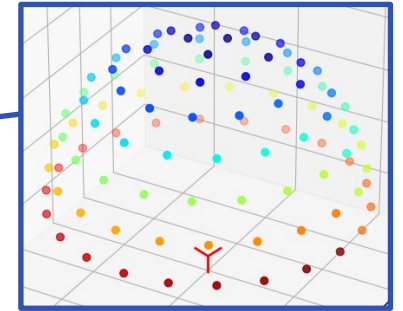
VEL for transmission measurements:

$$\left| H(\Phi, \Theta, f) \right| \propto R \cdot \sqrt{P(\Phi, \Theta, f)}$$



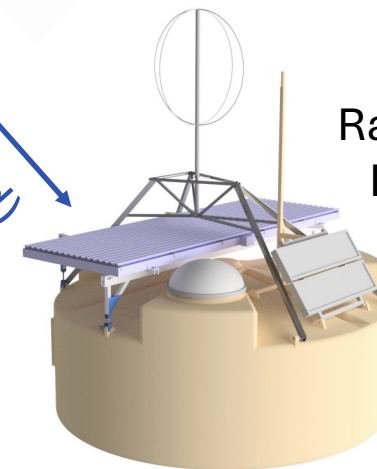
Position (ϕ, θ)

Automated flight



- Fly to “Waypoints”
- **Stop** for 6s
- **Automatically aim** at antenna

Distance R (far-field)



Rapid triggering
DAQ (≈ 1 Hz)



UHECR deflections in the Galactic magnetic field

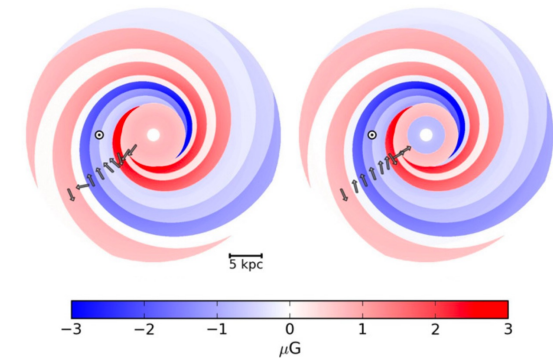
Galactic magnetic field (GMF)

- Order of magnitude: μG
- **Coherent** and **turbulent** component
- Multiple models with large uncertainties (JF12, UF23, ...)

Deflections of UHECRs

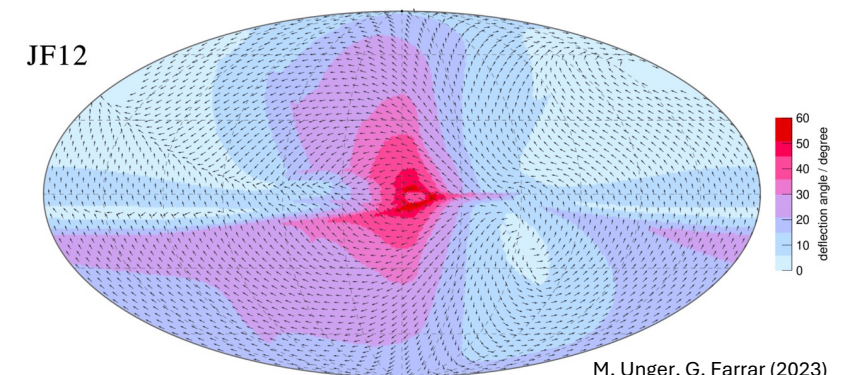
- Lorentz force: $\frac{d\vec{v}}{dt} = q \vec{v} \times \vec{B}(\vec{x})$
- Dependent on initial direction and rigidity
 $R = \frac{E}{Ze}$ of particle

coherent GMF component of JF12 model



R. Jansson, G. Farrar (2012)

deflection direction & strength of 20 EeV proton



M. Unger, G. Farrar (2023)

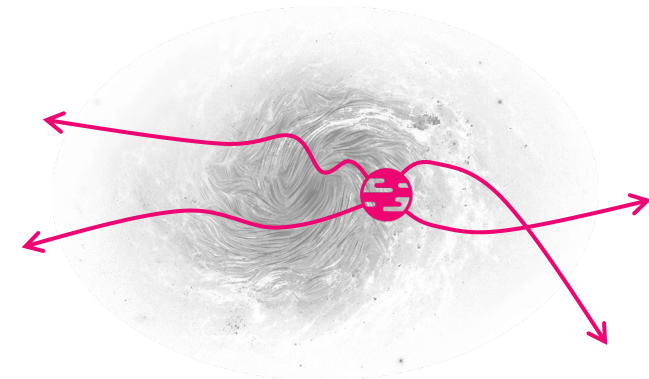
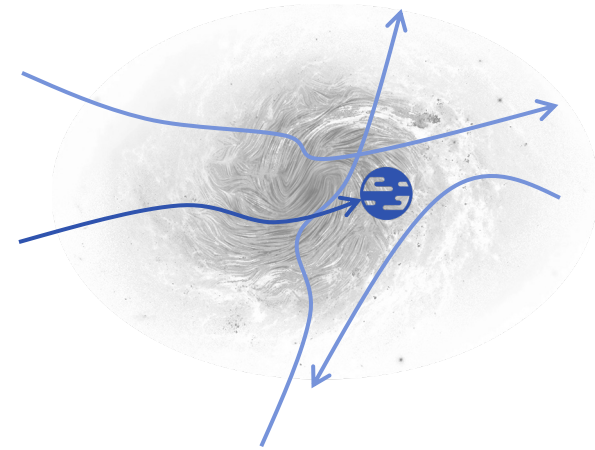
GMF lenses for UHECR deflections

- only tiny fraction of UHECRs that enter the Galaxy arrive at Earth
→ forward simulation is expensive
- **idea:** backpropagation of UHECRs with opposite charge

GMF lenses

- backtracking of N particles originating at Earth in direction of pixel i and count how many CRs are deflected in direction j
- lens is matrix L with entries $L_{ij} = \frac{n(i \rightarrow j)}{N}$
- relates detected sky map at Earth \vec{a} to extragalactic illumination map \vec{b} via

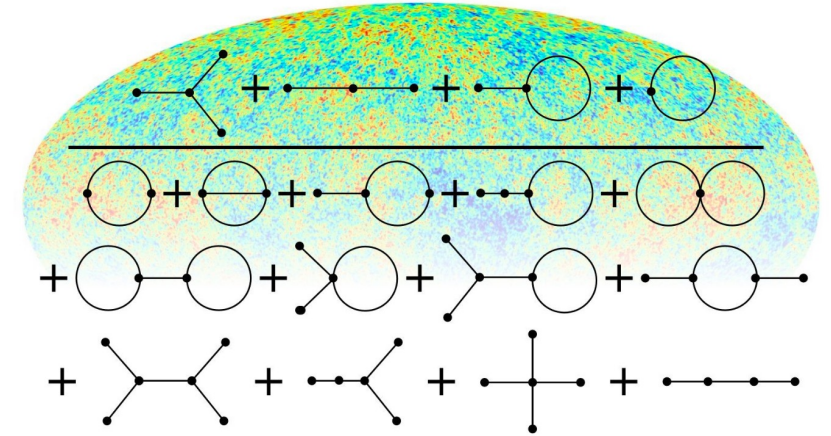
$$\vec{a} = L \vec{b}$$



Information Field Theory: Connection to field theory and statistical mechanics

Bayesian probabilities

- $\mathcal{P}(s|d) = \frac{\mathcal{P}(d,s)}{\mathcal{P}(d)} \equiv \frac{e^{\mathcal{H}(d,s)}}{\mathcal{Z}(d)}$
 - $\mathcal{P}(s|d)$: Posterior distribution
 - $\mathcal{P}(d, s)$: Joint probability
 - $\mathcal{P}(d) = \mathcal{Z}(d)$: evidence / partition function
 - $\mathcal{H}(d, s)$: Information Hamiltonian
- Rewriting allows to use techniques from statistical & quantum field theories
- Information is additive: $\mathcal{H}(d, s) = \mathcal{H}(d|s) + \mathcal{H}(s)$



[arXiv:1804.03350](https://arxiv.org/abs/1804.03350)