



# FORMOSA

24.07.05

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on behalf of the FORMOSA collaboration

LLP2024, Tokyo

# New physics and dark matter → Hidden Valley

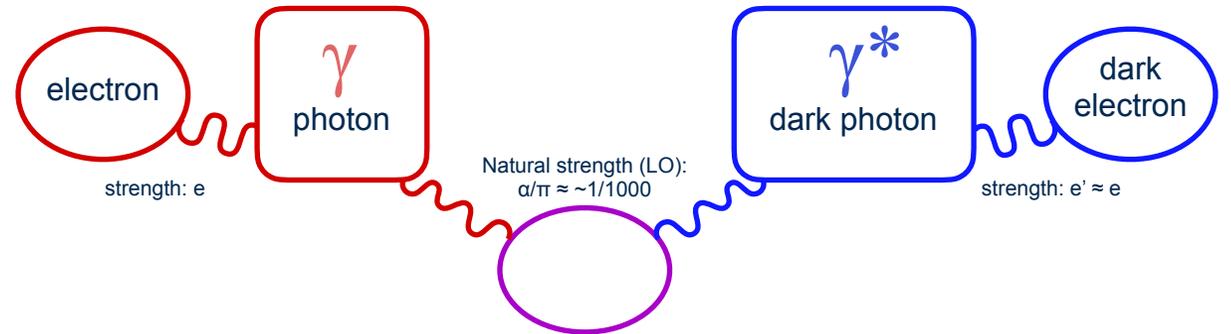
No signs of new physics  
seen at the LHC (yet)

SM extensions that  
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# New physics and dark matter → Hidden Valley

No signs of new physics seen at the LHC (yet)

SM extensions that include dark (or hidden) sectors give the most plausible hint



Photons and dark photons are both U(1), they can interact via **kinetic mixing**

Interaction with dark electrons is around **1/1000 as strong as the standard model**

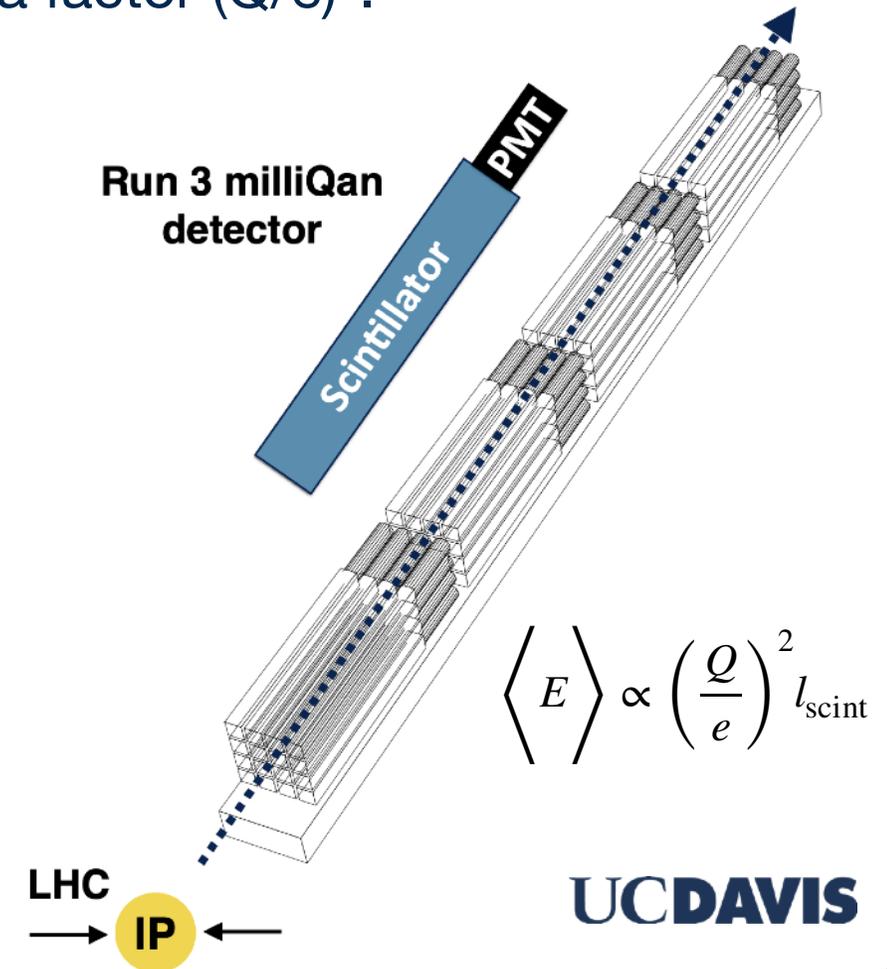
# Direct millicharged particle searches at the LHC

Millicharged particles (**MCPs**) are well motivated in dark sector theories, but difficult to detect because the interaction strength is reduced by a factor  $(Q/e)^2$ .

**Core concept:** Use array of efficient long scintillator bars + PMTs to detect ionisation from MCPs.

## Challenges:

- Expect few scintillation photons to be produced  
→ must be able to detect single scintillation photons
- Well controlled backgrounds → “point” at the interaction point, triggering on sets of signals within small time windows ( $\sim 15$  ns)



# The milliQan experiment

Located in P5, looking at the central region of the CMS interaction point (see [Neha's talk](#)).

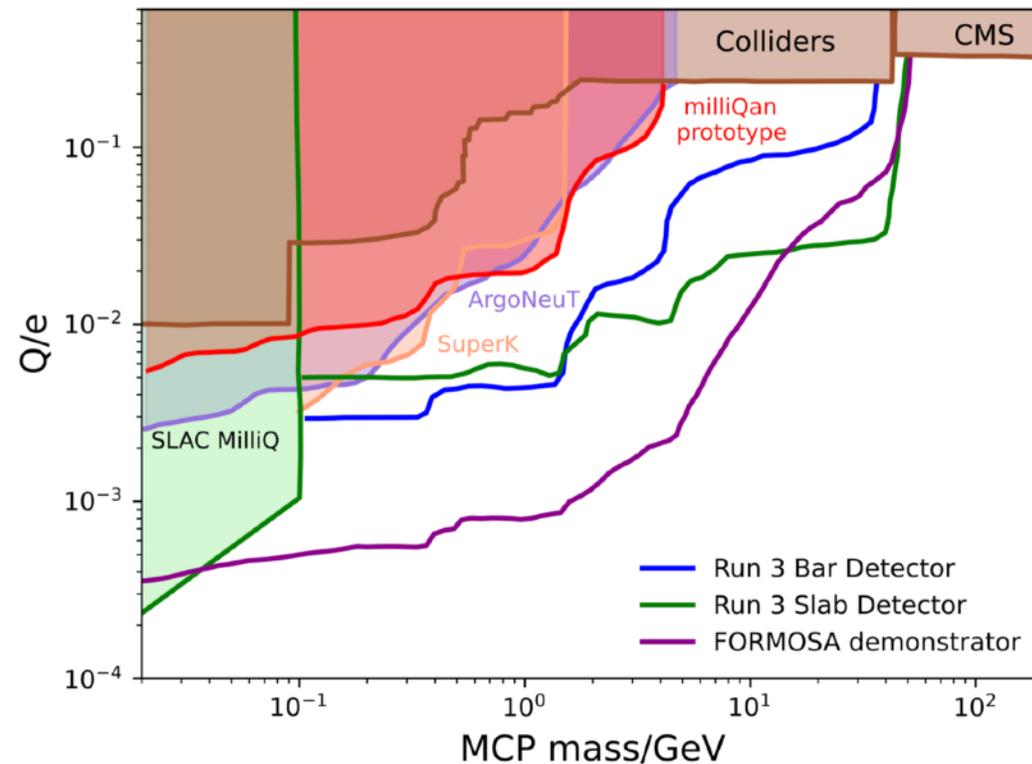
**First search** for MCPs at a hadron collider with new sensitivity carried out with the milliQan demonstrator.

Run 3 milliQan experiment ongoing

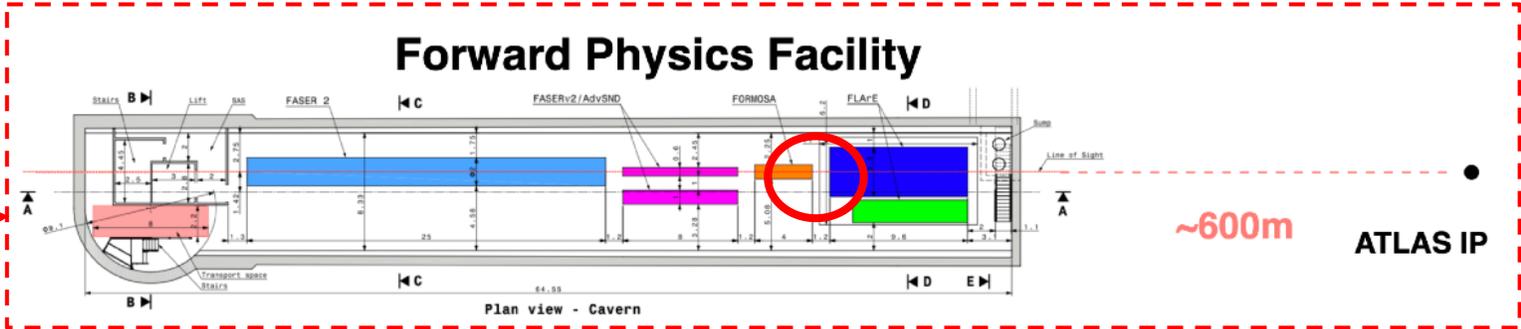
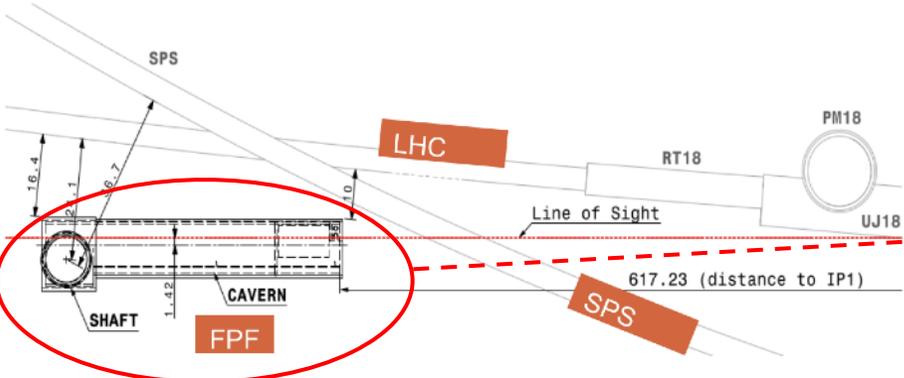
Overall, this means that **FORMOSA is not starting from zero!**

Tons of expertise acquired on the milliQan experiment: R&D, manufacturing, installation, calibration, commissioning, backgrounds, operation, analysis

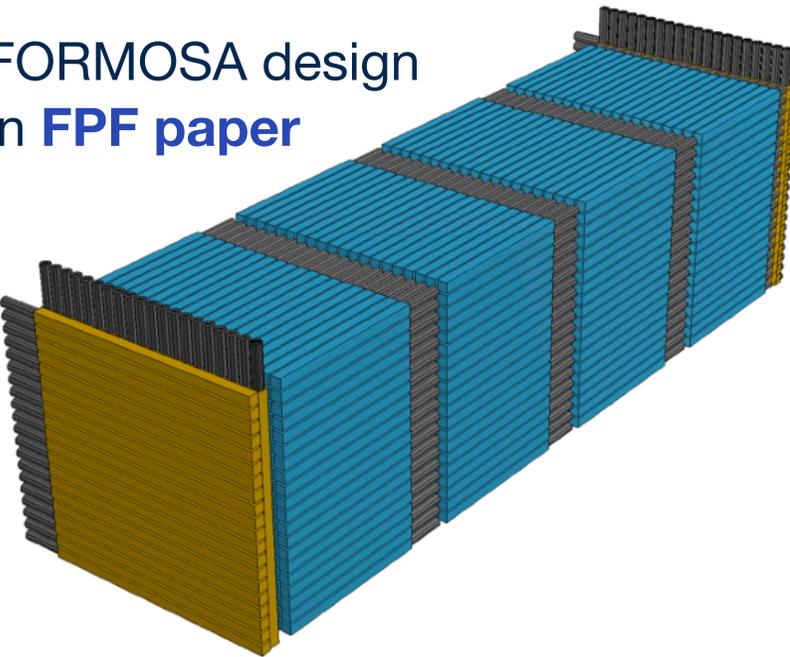
→ **FORMOSA is the natural next step**



# FORMOSA



FORMOSA design in **FPF** paper



Aimed at the ATLAS IP, located in the very-forward region ( $\eta > 9$ )

Expect to see **~250x rate** of millicharged particle detection in the forward region compared to the central one (i.e. milliQan,  $\eta \sim 0$ )

20rows x 20cols x 4layers of bars for detection

**Shielded** from most SM particles produced at the IP by **~ 600m of rock**

**Shielded** from LHC beam-gas radiation by **>10m of rock**

**Main background: beam muons** → segmented beam-muon veto panels



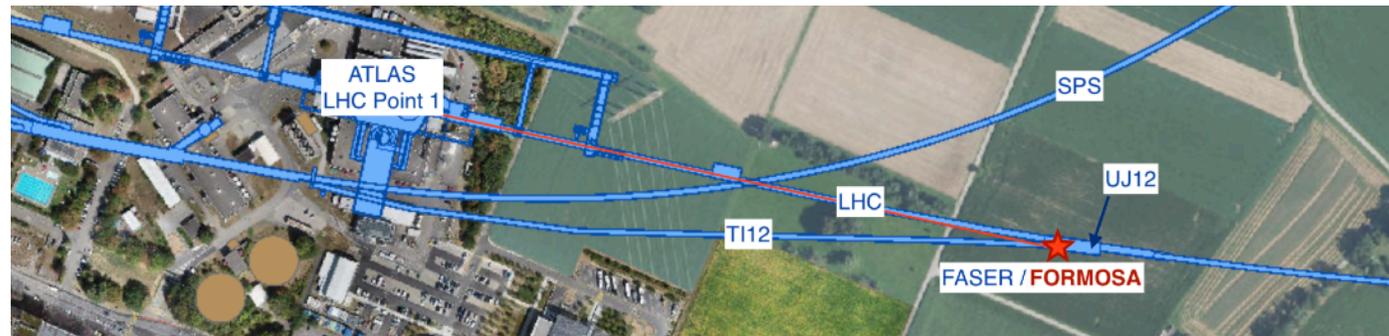
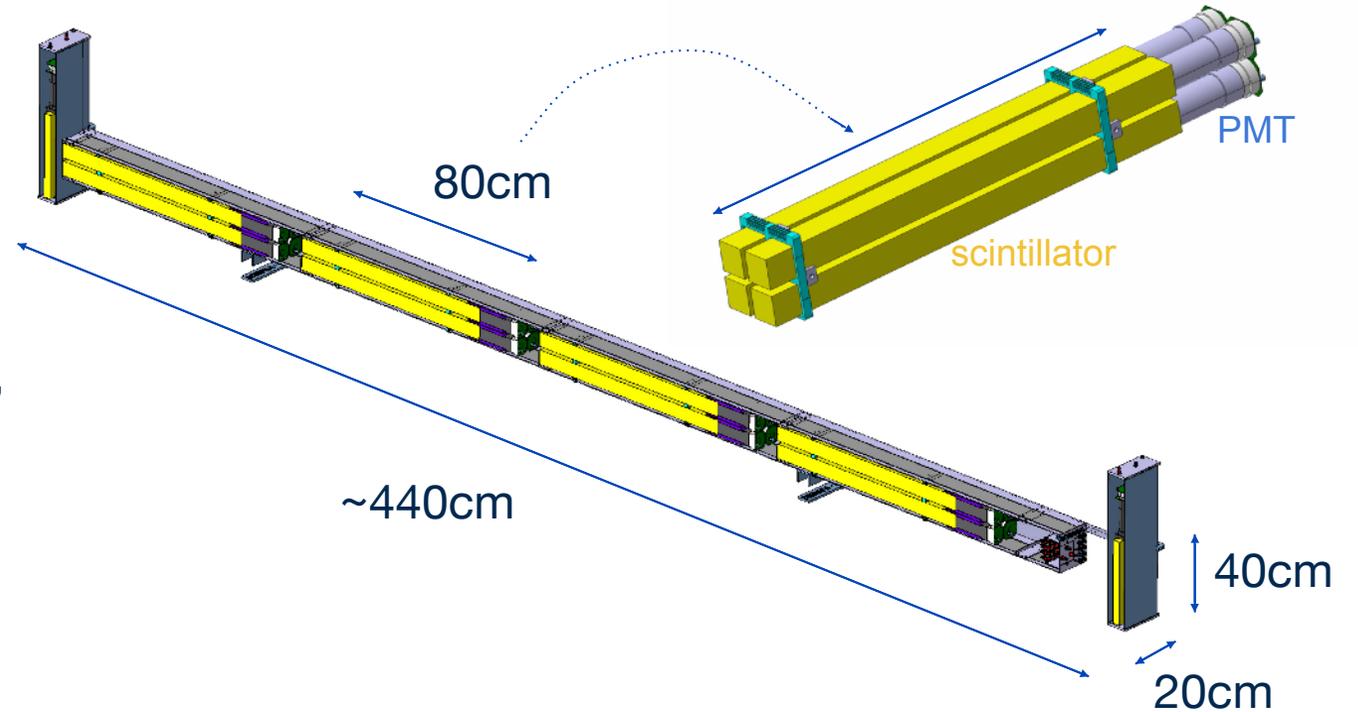
# The FORMOSA demonstrator

A small-scale version of the full FORMOSA was installed during YETS 2023

→ 2 rows x 2 columns x 4 layers

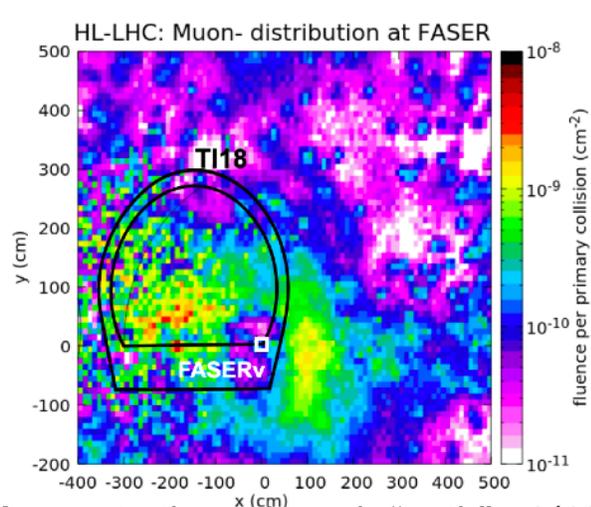
Located in the UJ12 cavern (behind FASER), in the opposite side of the to-be FPF

The demonstrator allows to **prove concept** and **target new phase space**



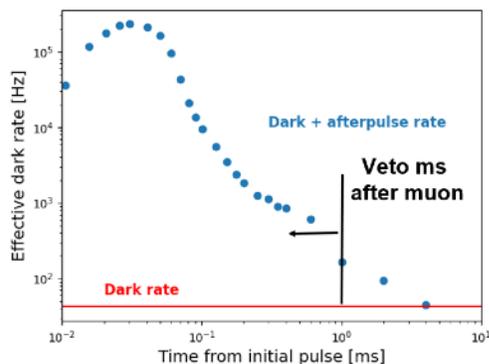
# Main backgrounds at the demonstrator

Dominant expected background in the forward region:  
**afterpulses initiated by through-going muons**



Muon rate through each "path"  $\sim 1/40\text{ms}$

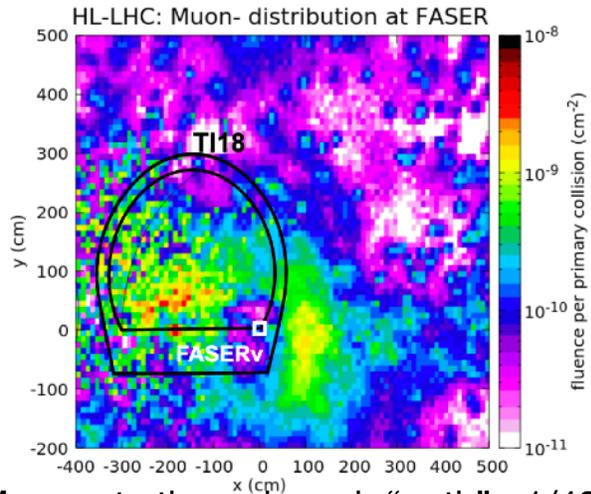
Measure afterpulsing induced by LED pulses



Can devote a ms of  
deadtime to cool down  
following a beam muon.  
But muons are frequent  
( $\sim 100\text{Hz}$  for demonstrator),  
so can't wait too long

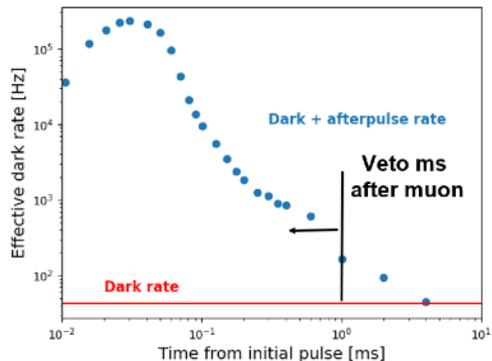
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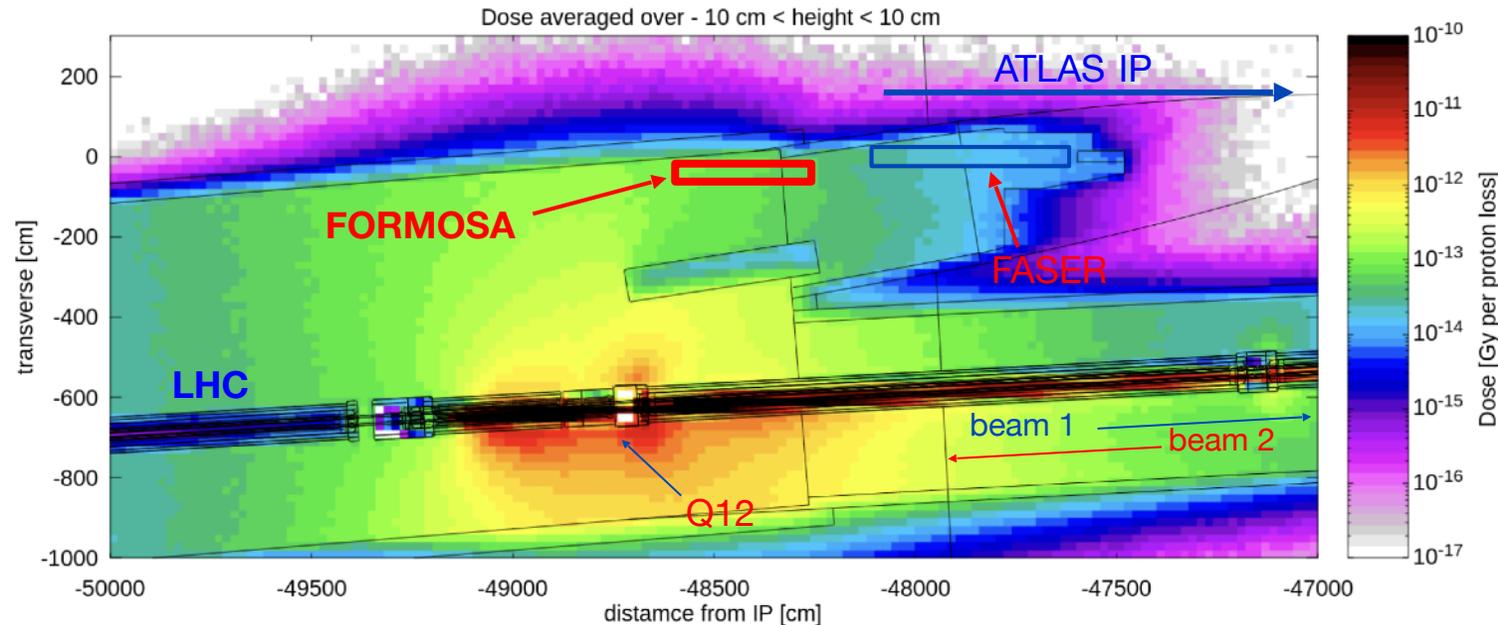
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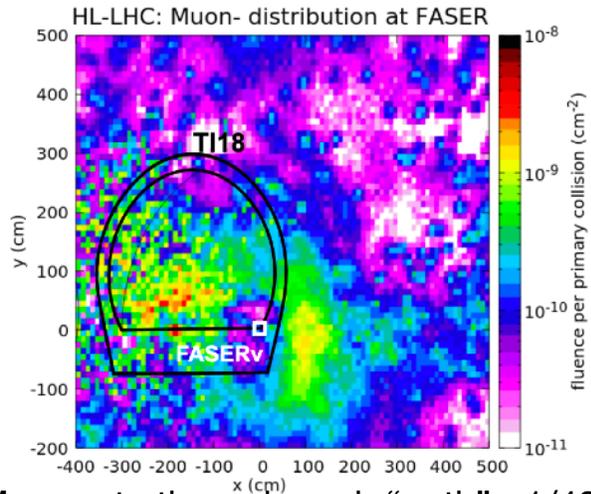
An additional challenge: **beam1-gas background**  
**(N.B. this background will NOT be present in the FPF)**



*Absolute average dose due to proton losses at Q12 (simulation provided by FASER)*

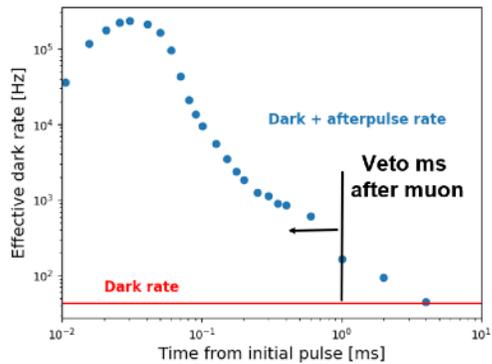
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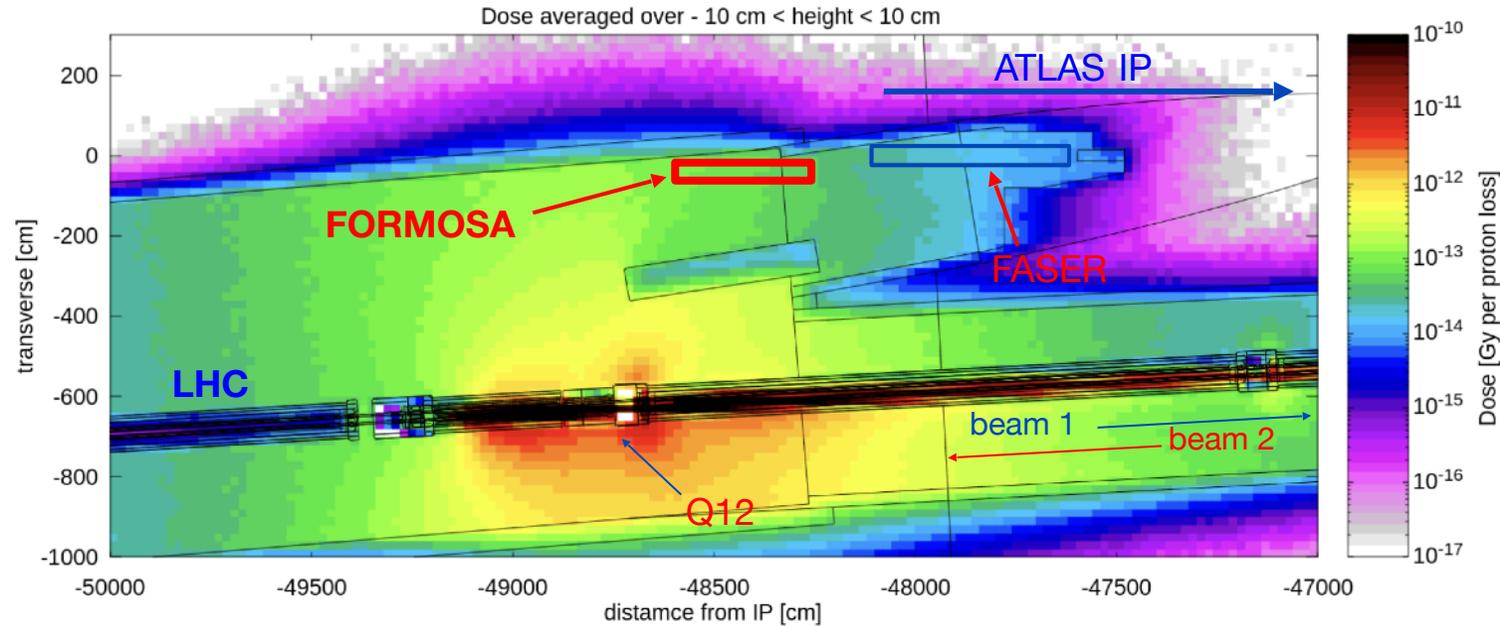
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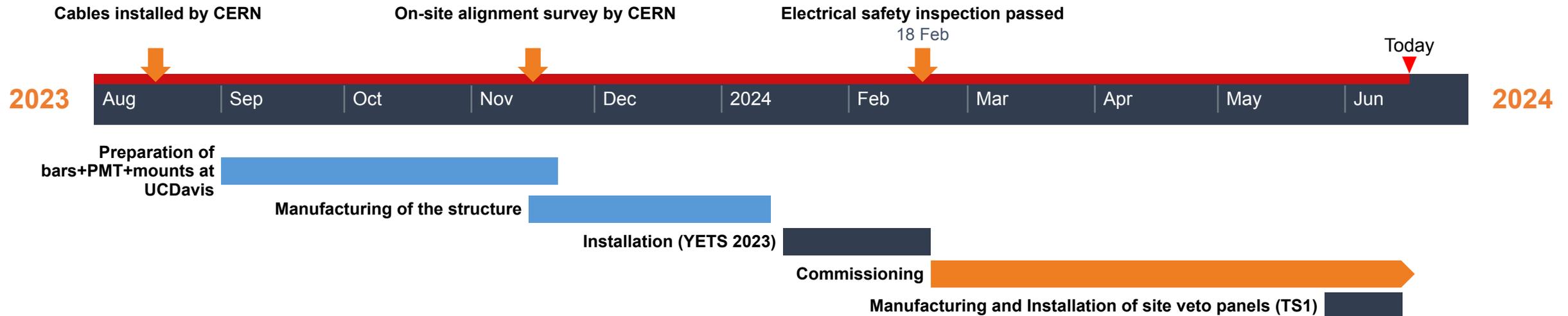
Absolute average dose due to proton losses at Q12 (simulation provided by FASER)

Side veto panels

Large front+back muon veto panels

# Manufacturing and installation

*The demonstrator was built primarily by UC Davis hands!*



The entire preparation and installation of the main body took about 7 months (+2 weeks for the side veto panels)

# Preliminary preparations



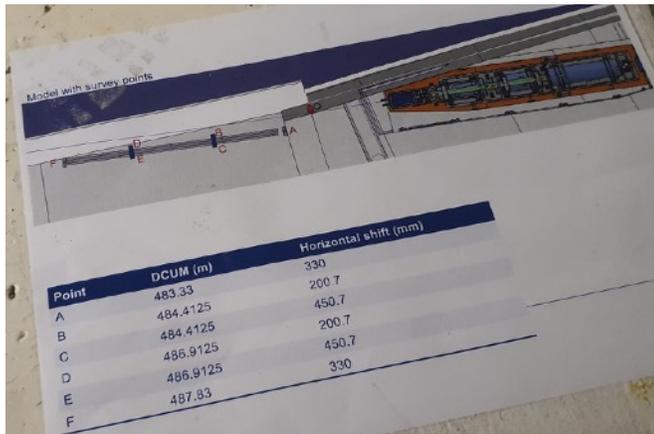
**August:**  
Cabling installed by cern



REU student: N. Gonzalez  
(Holding a scintillator bar and PMT)



Grad students: J. Steenis, S. Kelly



**November:**  
On-site alignment survey by CERN



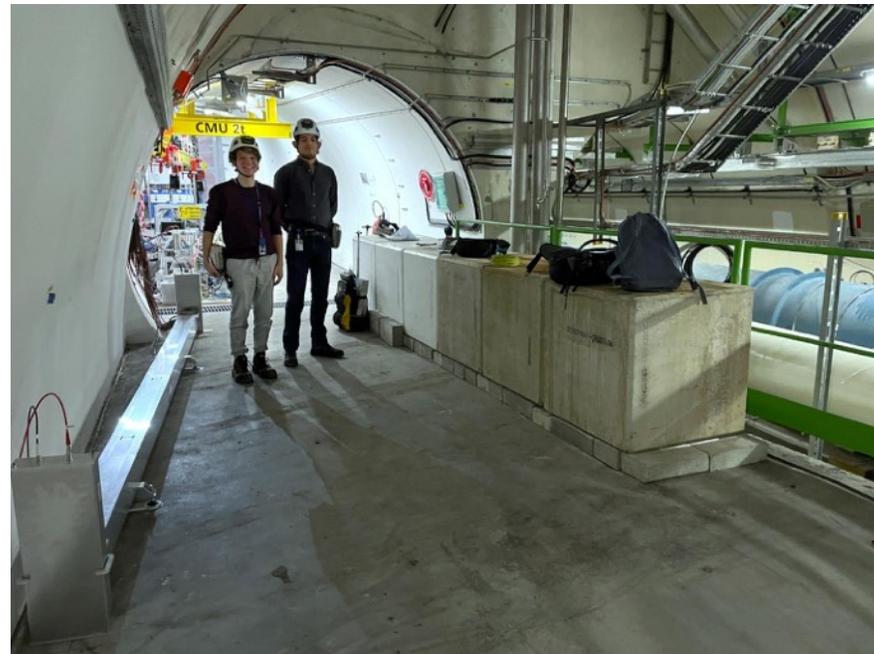
**September - November:**  
Bars+PMT+mount prepared by  
under/grad students at UC Davis

# Machining, assembly and installation



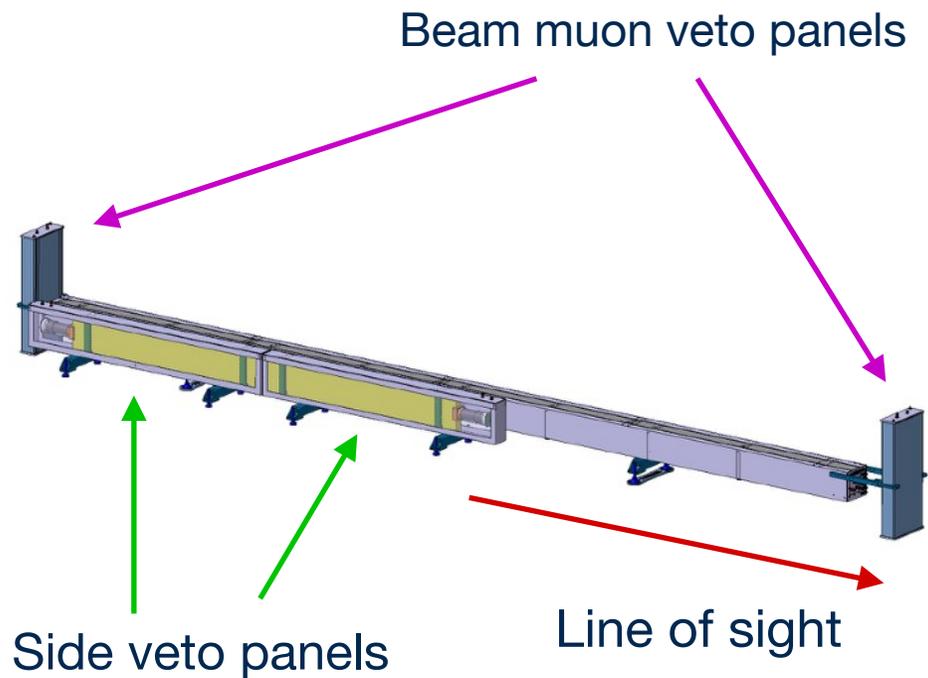
← **November - Mid-January:**  
Machining the structure

**Mid-January -  
Mid-February:**  
Installation →



Grad student: J. Steenis  
Postdoc: J.S. Tafoya V.

# The demonstrator, as of the 14th of June 2024



Grad student: J. Steenis  
Postdoc: J.S. Tafoya V.

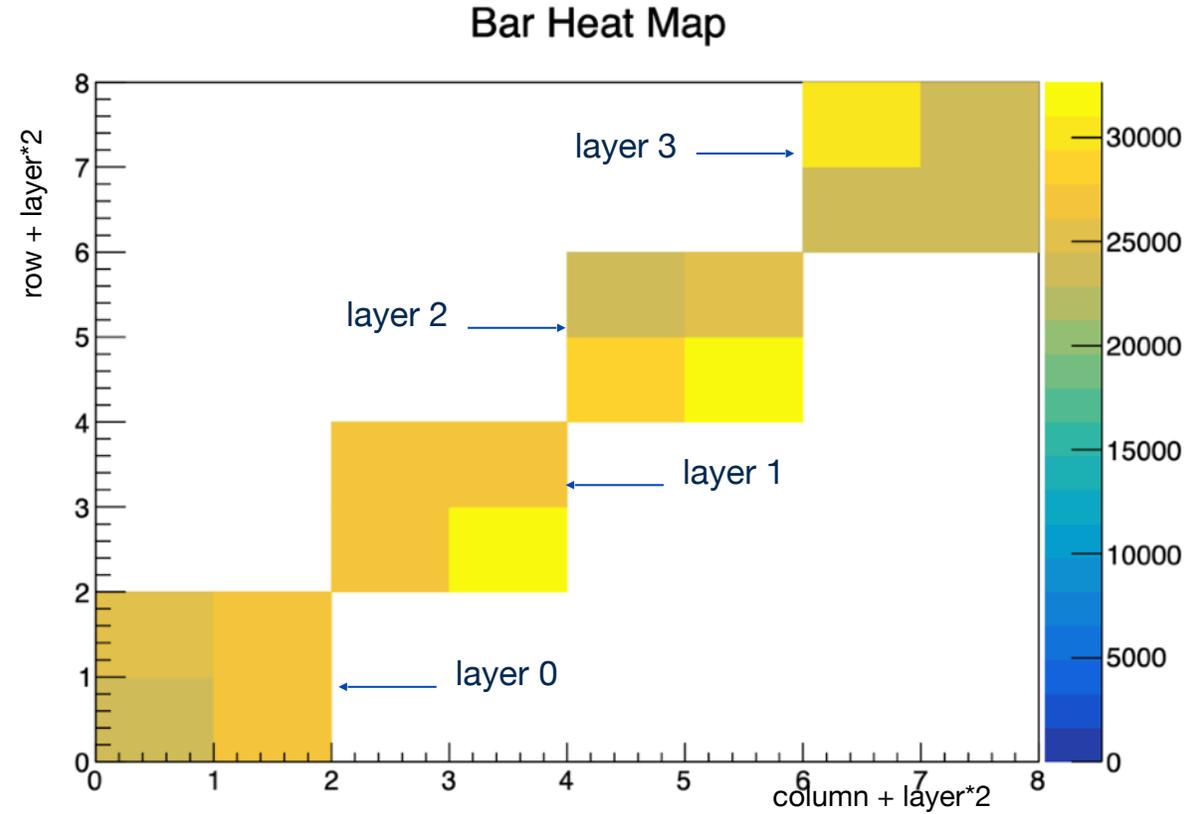
# No-beam data

Heat maps suggest correct activation on all the bars

Recorded data is consistent with that measured before the installation

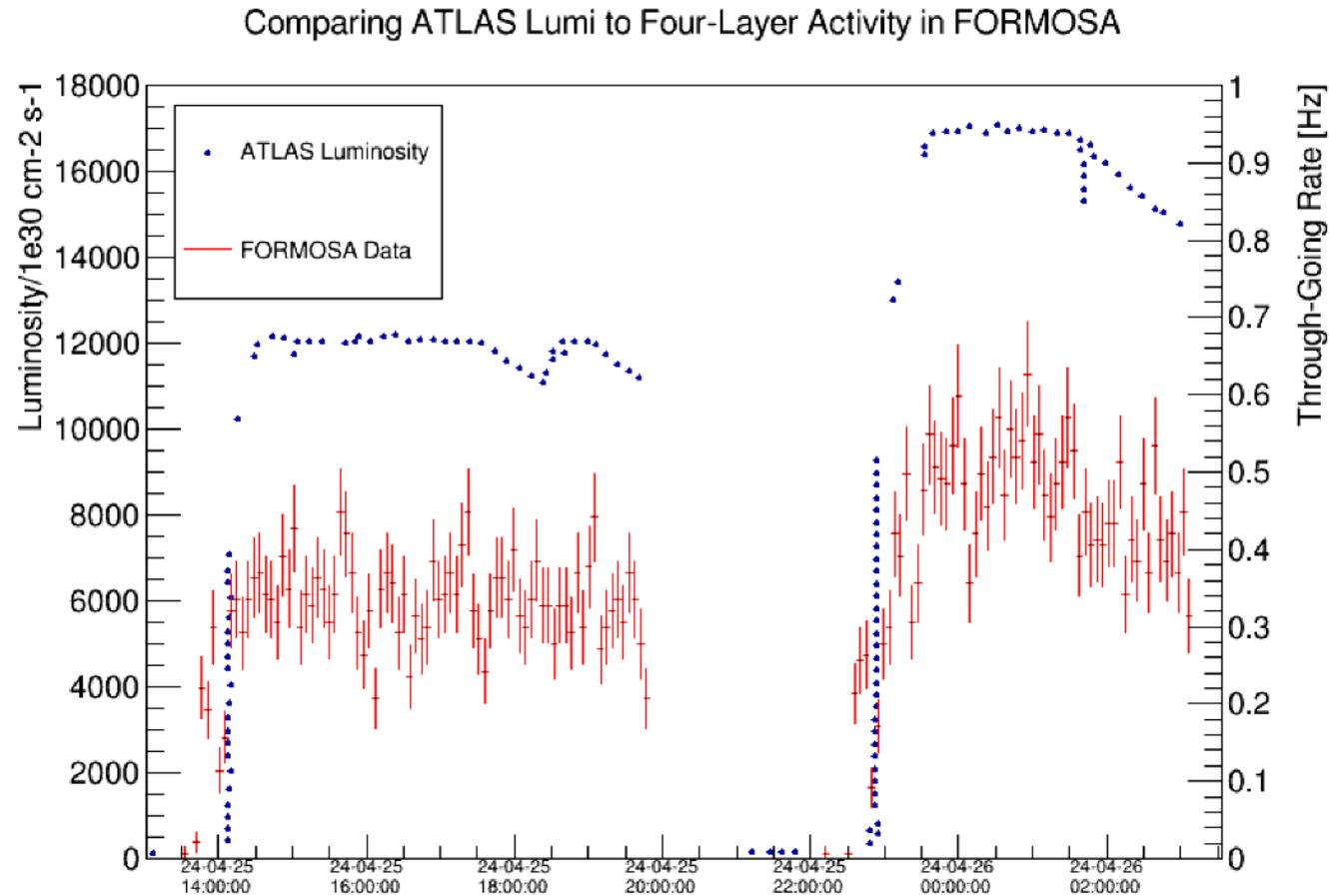
In no-beam conditions, we've measured adequate recording rates:

- bars + veto panels: ~5 Hz



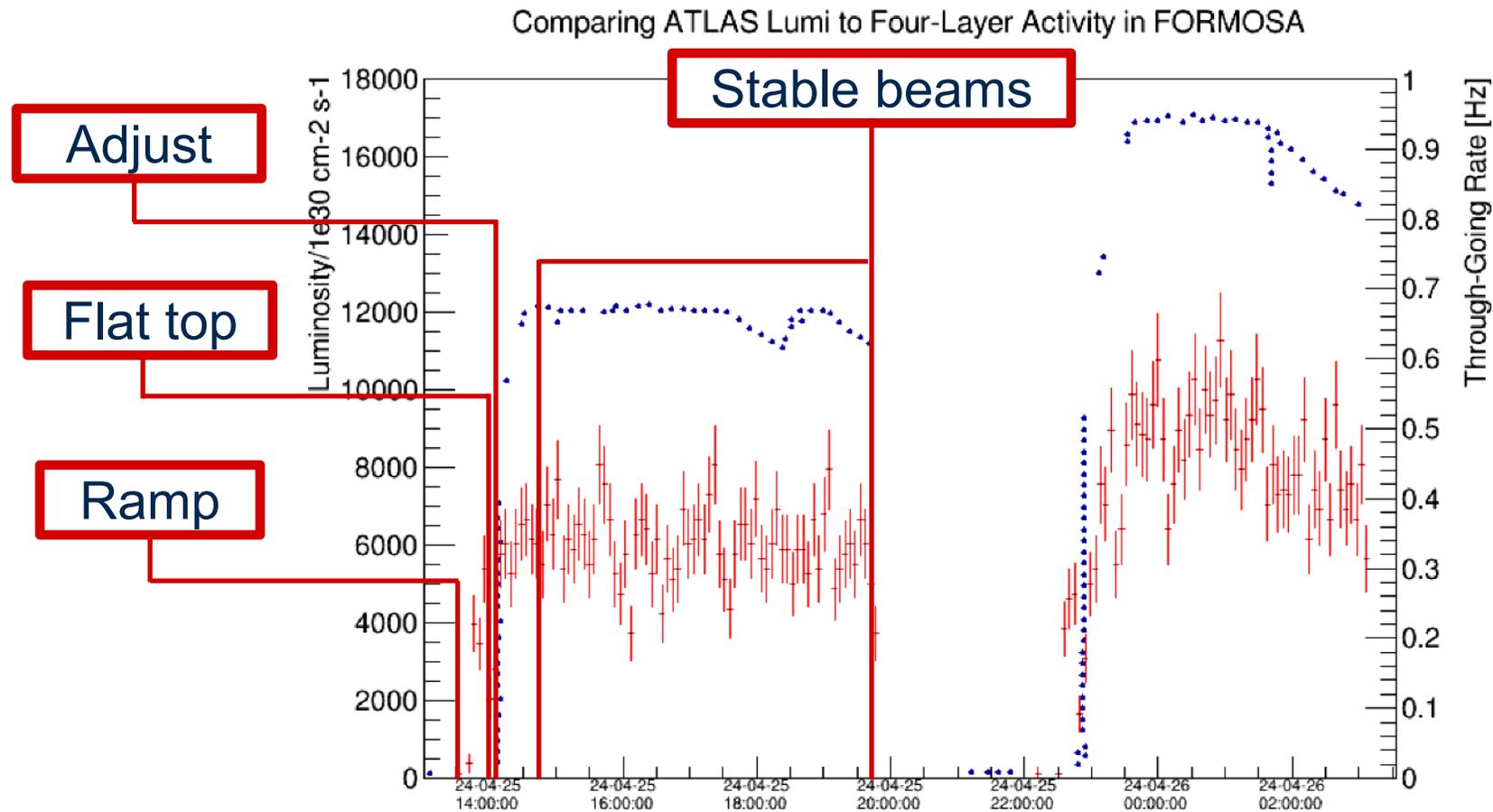
*Measured from a day-long run taken on 2024/02/27*

# First look at data (prior side veto panels)



- We see the increased activity in FORMOSA when beam circulates regardless of type
- During stable beams, there is a strong correlation between ATLAS luminosity and and 4-layer trigger rate

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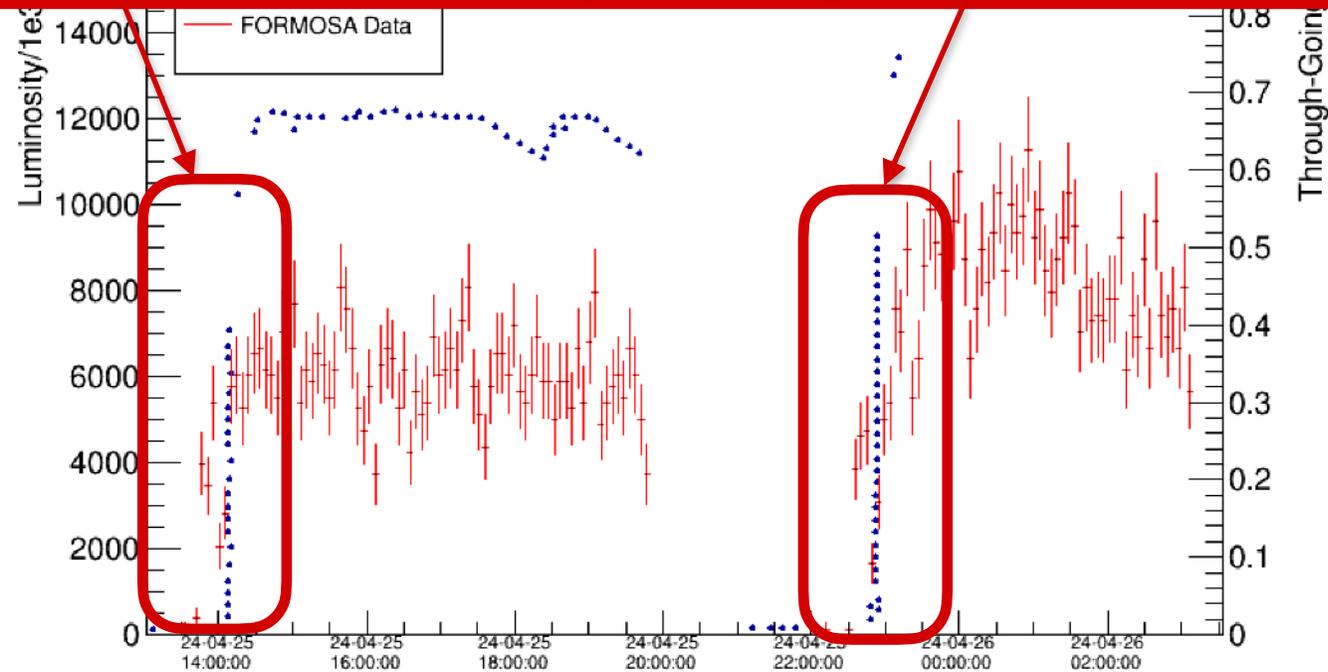


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# First look at data (prior side veto panels)

Expecting new side veto panels to improve the overall triggering rates, particularly pre-stable-beam detections (which certainly should be vetoed on).

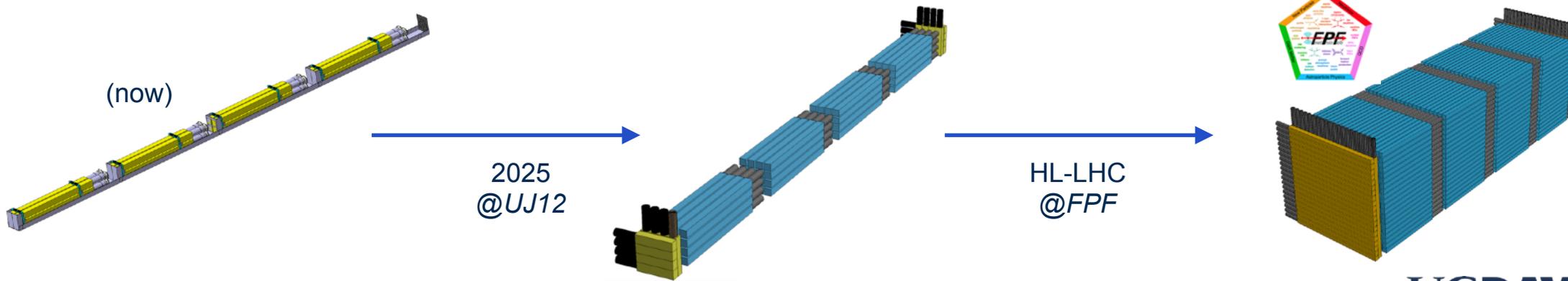
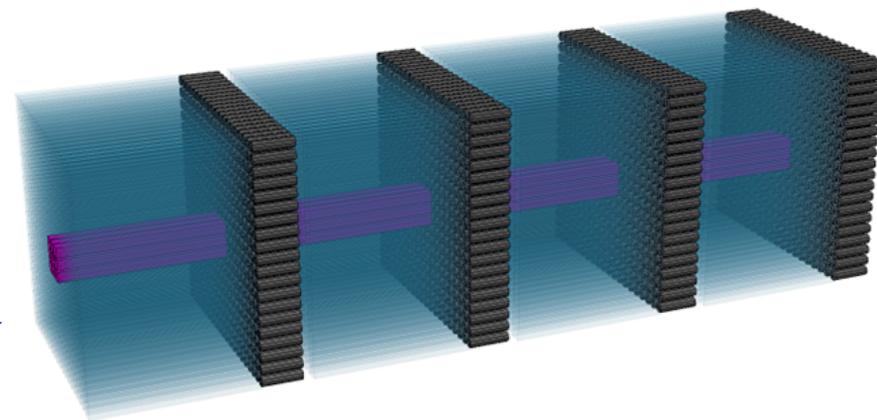
**Currently working on it!**



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# The future of FORMOSA

- Plan to expand the cross section of the detector:  
e.g. 4bars x 4bars x 4layers with segmented veto panels  
→ Could allow initial search with new sensitivity to fully prove feasibility
- Consider smaller subdetector using CeBr3 in the full FORMOSA  
→ Provides up to factor of  $\sim 4$  improvement in low charge sensitivity (below  $Q/e = 10^{-4}$ !)



# P5 outcomes and timeline

- FORMOSA fits P5 recommendation for new “agile” project portfolio (ASTAE)
  - From the P5 report: “Experiments at the proposed Forward Physics Facility at CERN like FASER2 and **FORMOSA** would be sensitive to the hidden sectors through the Vector and Heavy Neutral Lepton portals.”
- Timeline for experiment depends on when funding from ASTAE is realised (likely > 2026)
- We can build and commission FORMOSA in **~2-3 years** after construction funding is received

# 5

Explore  
New  
Paradigms  
in Physics



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# Outlook

- A small-scale version of FORMOSA was successfully installed at UJ12
- Beam data is already providing valuable information for the understanding of beam-muon backgrounds
- An additional background (beam-gas, not present at the FPF) affects data taking at UJ12. Already applying hardware- and software-based counter-measures.
- Opportunity to expand the demonstrator is being actively studied as we analyse current demonstrator data

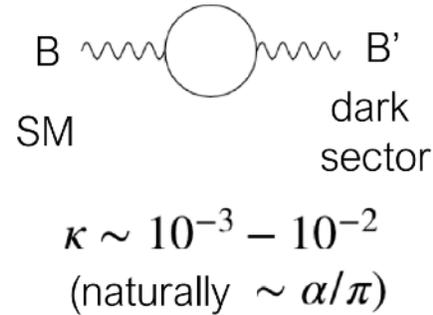
# Backup

# Why millicharged particles?

Standard motivation: Introduce new, hidden U(1) with a massless field  $A'$ , a “dark photon” that couples to a massive “dark fermion”  $\psi'$

$$\mathcal{L}_{\text{dark-sector}} = -\frac{1}{4}A'_{\mu\nu}A'^{\mu\nu} + i\bar{\psi}'(\underbrace{\gamma^\mu\partial_\mu + ie'\gamma^\mu A'_\mu + iM_{\text{mCP}}}_{\text{“dark fermion” with mass } M_{\text{mCP}}, \text{ charge } e'})\psi' - \frac{\kappa}{2}A'_{\mu\nu}B^{\mu\nu}$$

massless “dark photon”
mixing term



- $\Psi'$  has mass  $M_{\text{mCP}}$  and charge under the new U(1) of  $e'$
- Gauge transformation of  $A'_\mu \rightarrow A'_\mu + \kappa B_\mu$  introduces coupling  $\bar{\psi}'\kappa e'\gamma^\mu B_\mu\psi'$
- Conclusion: Coupling arises between dark fermion and SM photon of charge  $\kappa e' \cos \theta_W$ . **mCP parameters are entirely defined by their mass and charge**

see e.g. [arXiv:2104.07151v2](https://arxiv.org/abs/2104.07151v2) for more details

# Collecting data

Measure MCPs, which produce few scintillation photons per bar

- Expect an MCP to come from the L.O.S. and interact with a bar in each layer pointing back to the IP

Cosmic background:

- Activation of multiple bars within the same layer

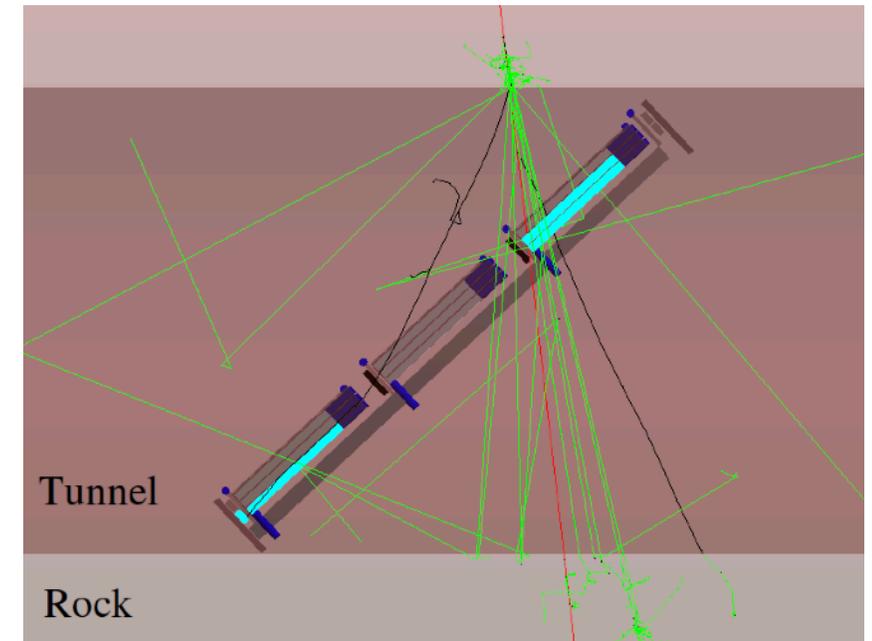
Beam muons:

- Apply dead time to the bars (i.e. to veto measurements) when the panels get activated
- Collect and labelling this data would also allow to better understand the effects of afterpulsing

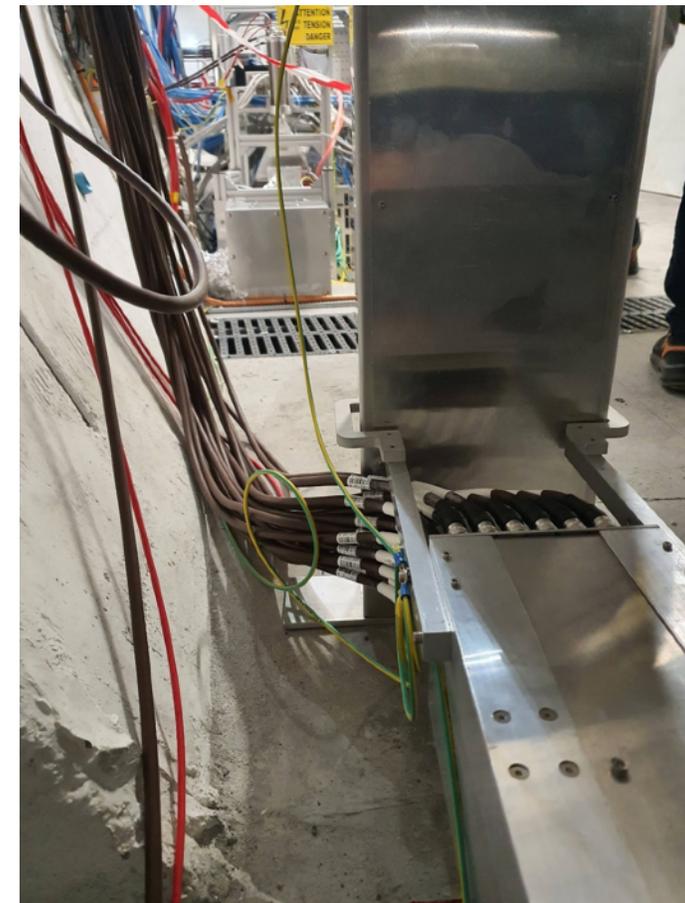
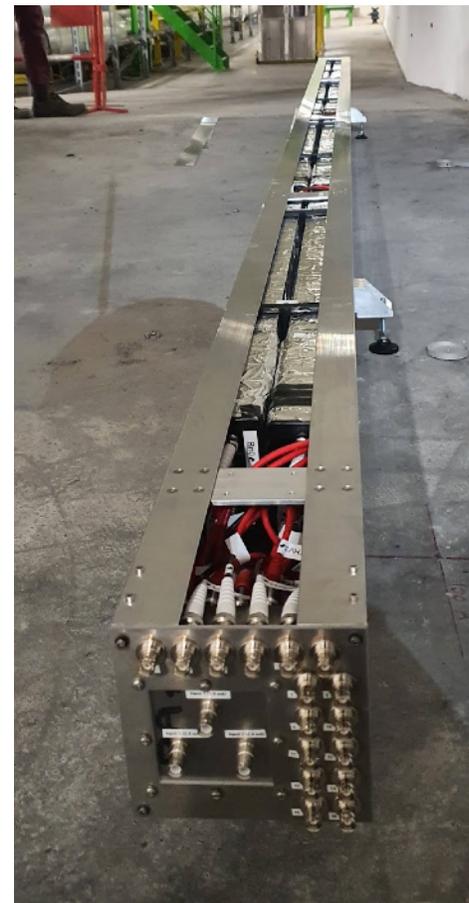
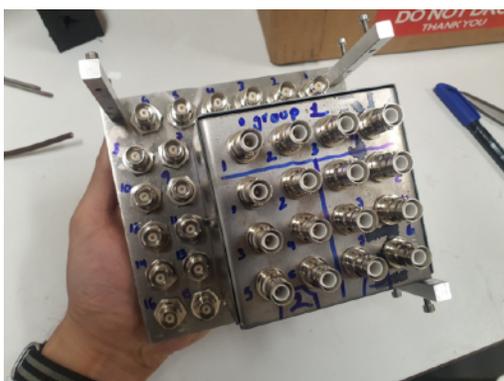
**Predicted rate of signal triggers ~ 1Hz**

**Predicted rate including all triggers ~4Hz**

Cosmic shower background, simulated for the milliQan demonstrator

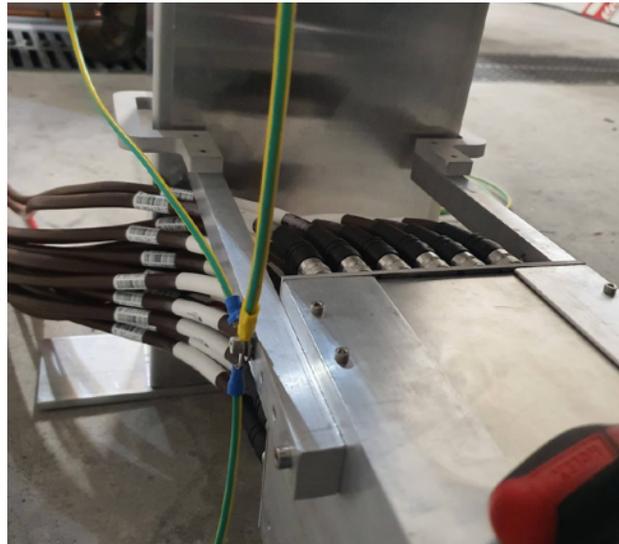
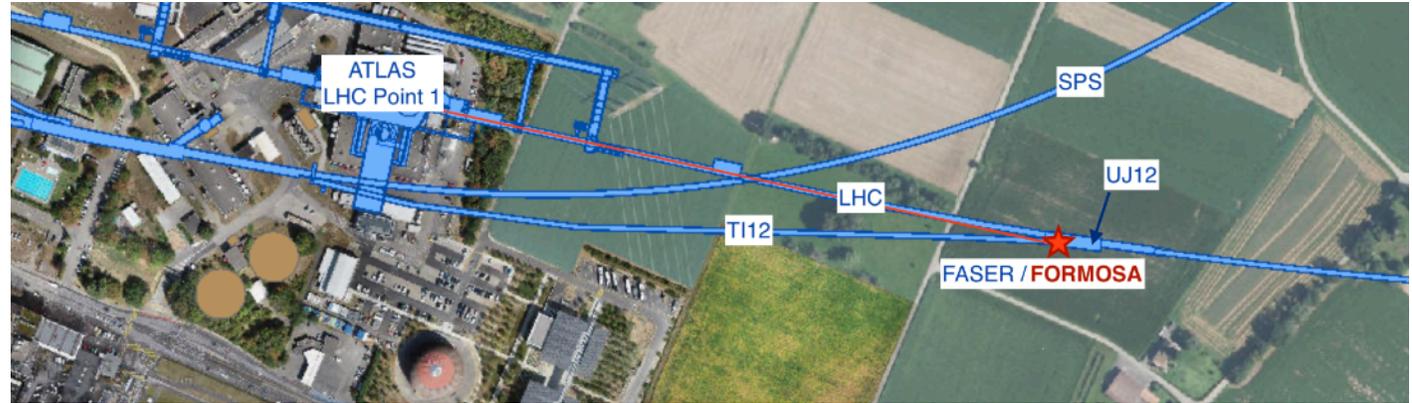


# Internal cabling



We use a patch panel (which also doubles as an HV splitter) to completely run all the cables on the inside → The final structure is fully closed

# Installation completed in February!

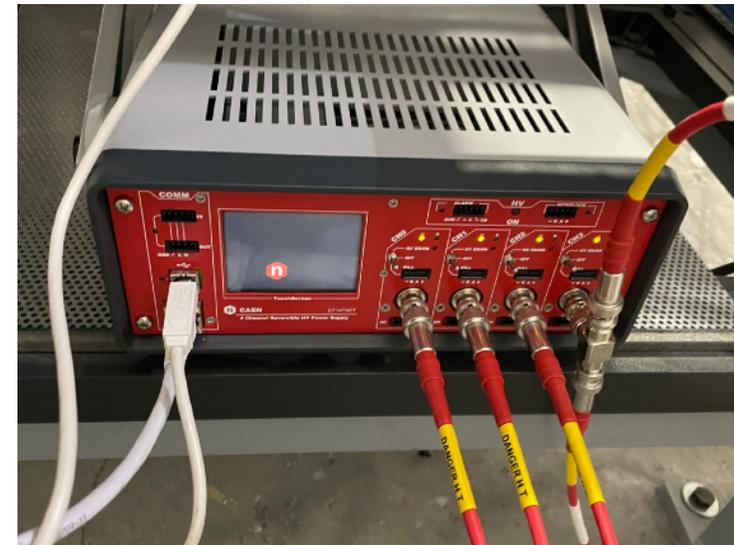
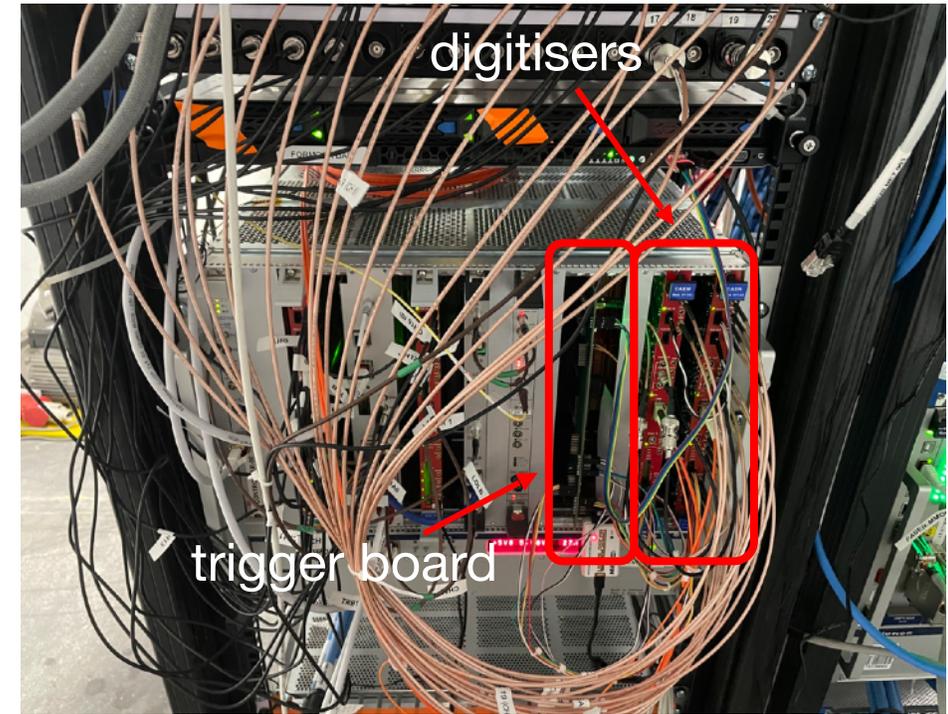


**Early February:**  
Electrical safety inspection passed.  
We are good to run remotely!

**Now:** LHC tunnel closed for Run 3 operations. **Commissioning FORMOSA!**

# Readout and HV

- Similar readout to that of milliQan:  
CAEN digitisers + custom trigger board
  - **New:** muon veto implemented at trigger board level
- Using our own HV power supply, which powers up the entire demonstrator through our patch panel.
- Everything is installed on the FASER rack at TI12



# Calibration

**SPE = Single PhotoElectron** i.e. physical electron emitted from the PMT's photocathode

MCPs produce just a **few scintillation photons**

→ we must be able to measure single photons

→ requires excellent calibration and identification of each PMT's SPE peak

PMT's signal grows linearly with the number of PE  
reaches saturation at 100ths or 1000ths of PE

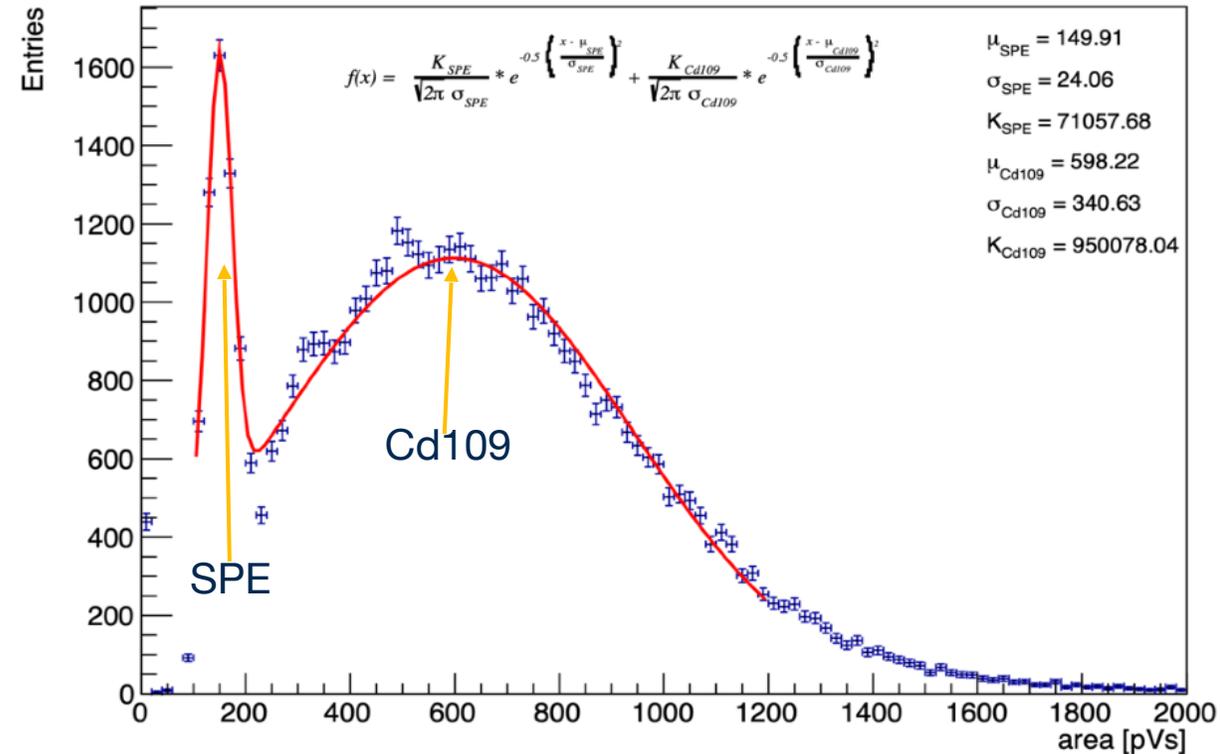
Calibration runs:

**Dark rate** → measure SPE peak, can be done at any time

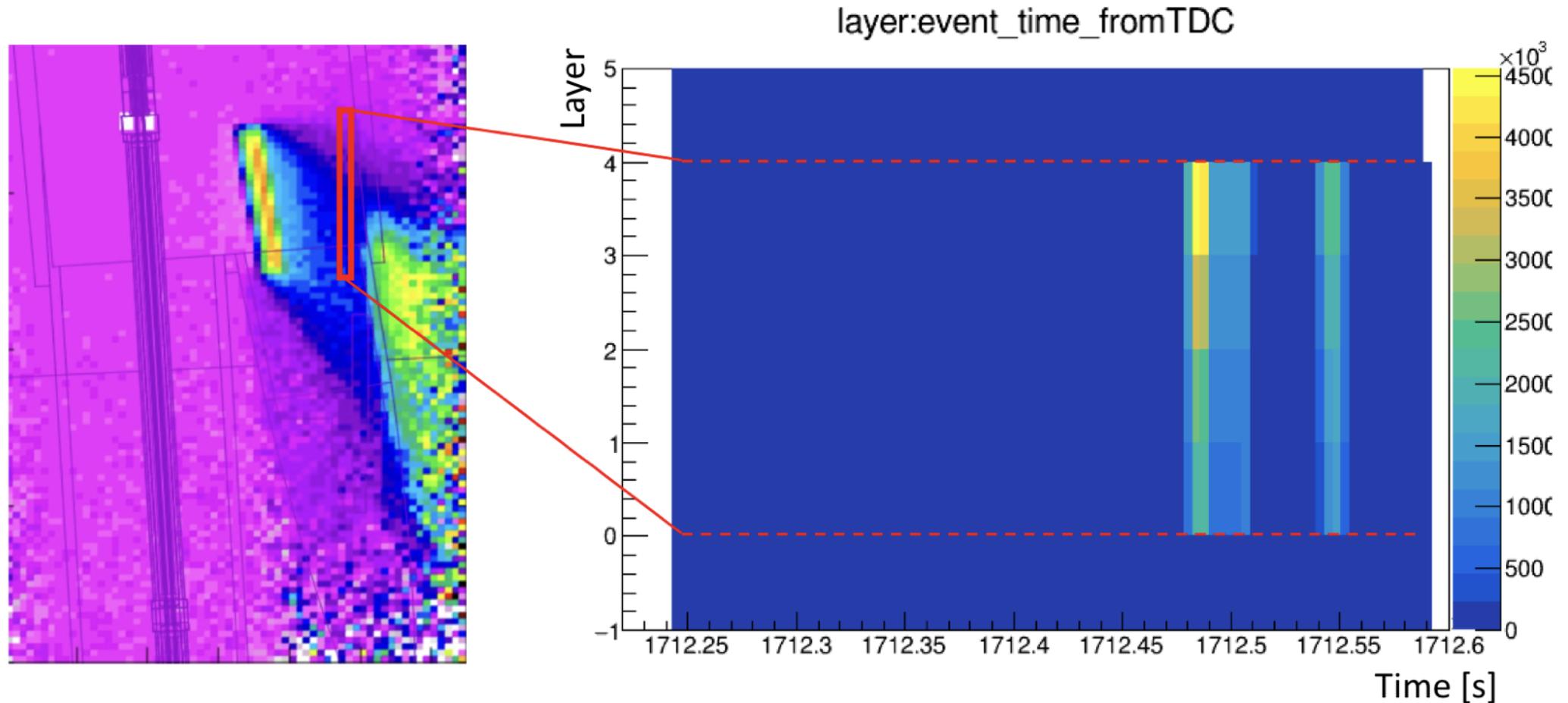
**Source data** → induce signal with a Cd109 radioactive source, done for each scintillator before installation.

**Clear separation** between SPE and source-induced signals → particularly true when looking at **pulse area**

## FORMOSA bar 0



# Detecting beam1-gas background



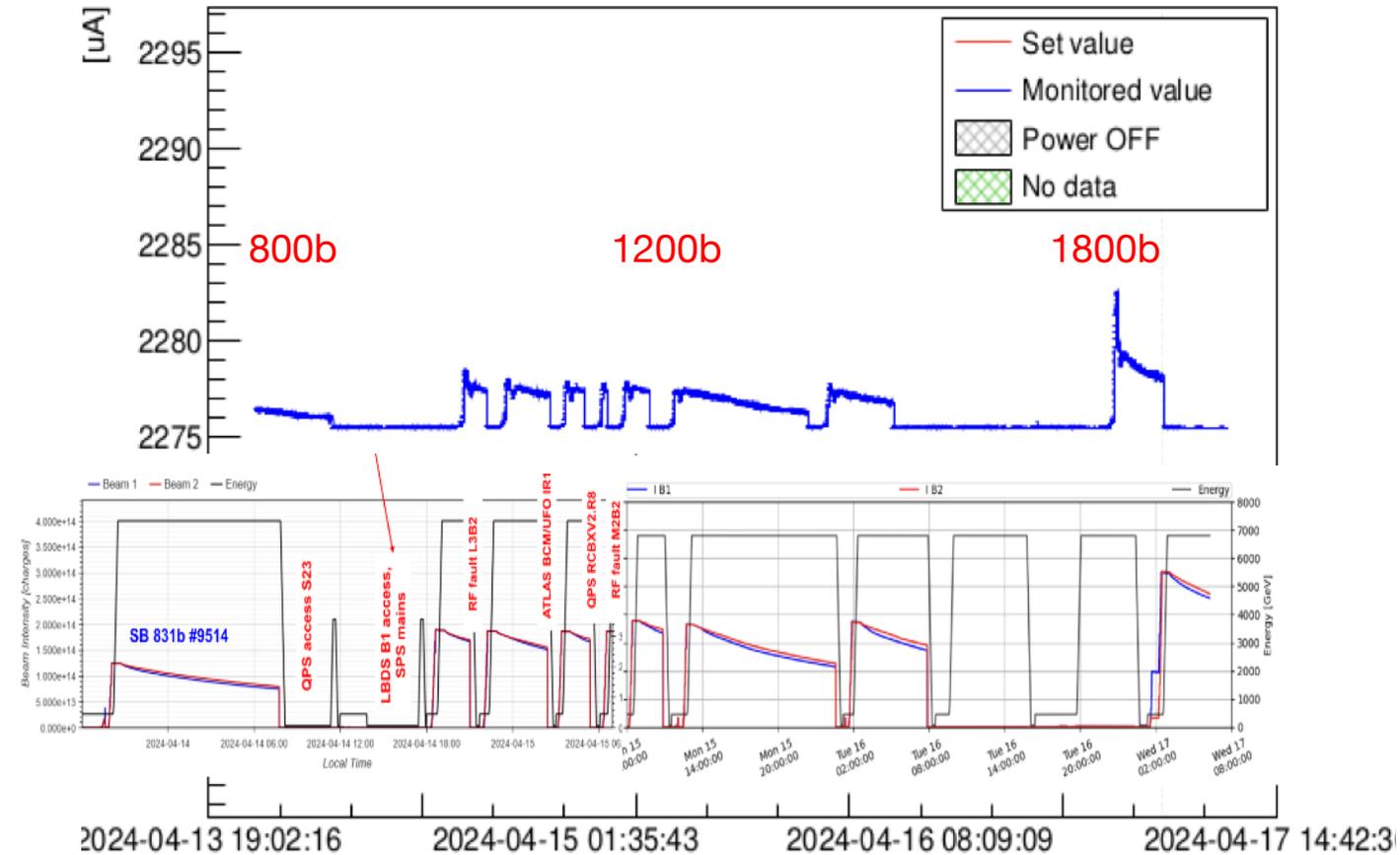
**Study done by J. Steenis.** It clearly shows activation that follows a gradient as in the beam background simulation!

# Beam-induced activation

We have measured activity at FORMOSA when the LHC circulates beam regardless of the type (scrubbing, VdM scans, stable beams).

- An easy way to sense beam activity is by looking at the PMT's current draw: the more activation they get, the more current they consume.
- This is heavily correlated to the circulation of the beam itself.
- We are detecting **beam background**

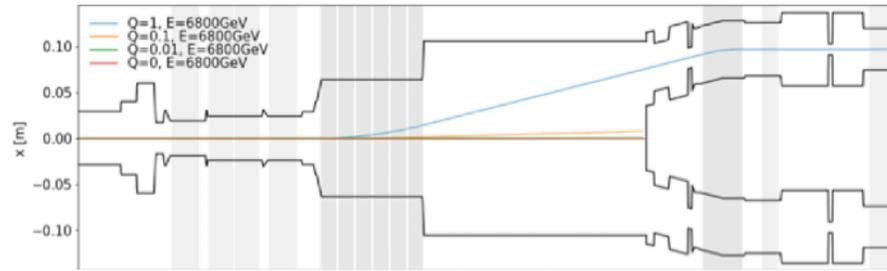
HVPS ch1



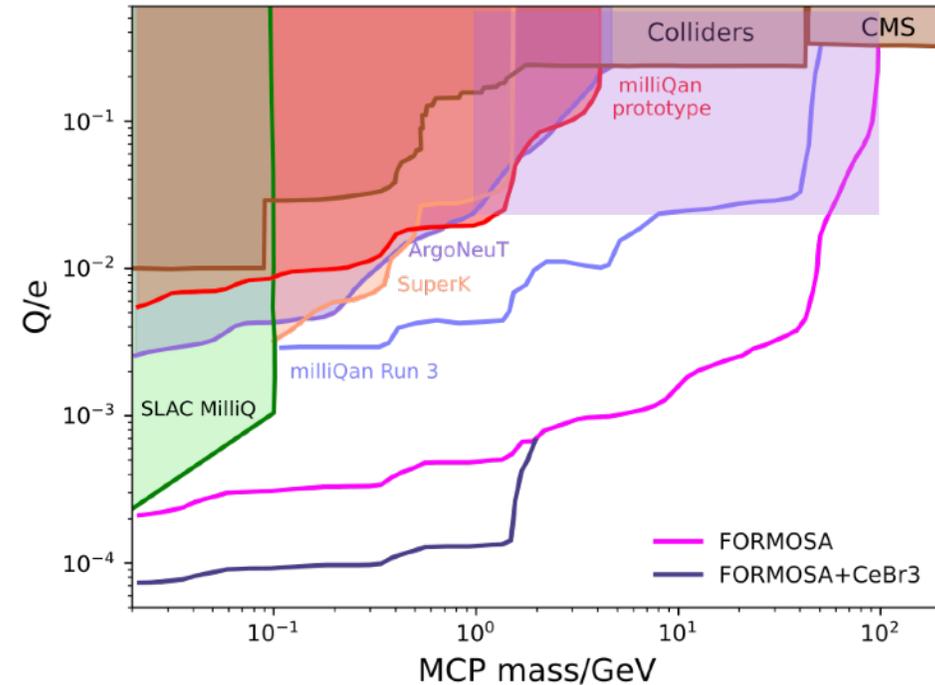
Current draw of ch1 over several days, compared to LHC activity in the same period (LHC status recovered from the [LHC morning meetings](#))

# FORMOSA limit uncertain here

## MCPs propagation



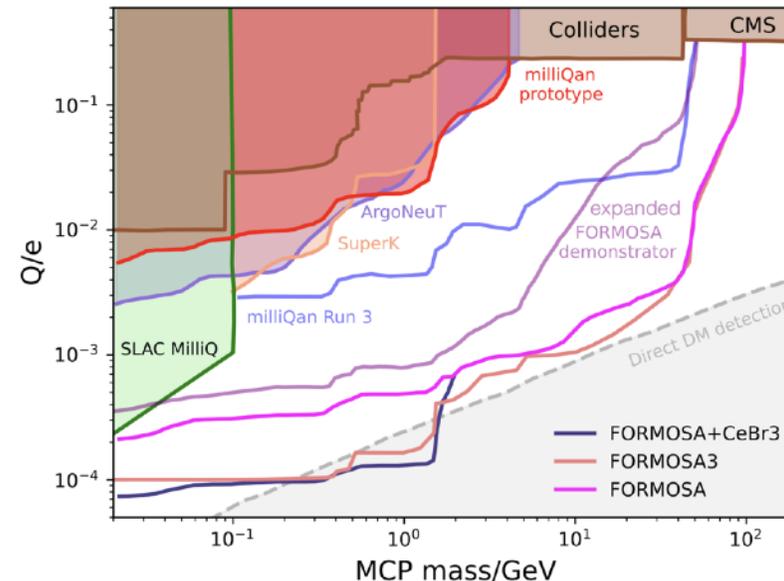
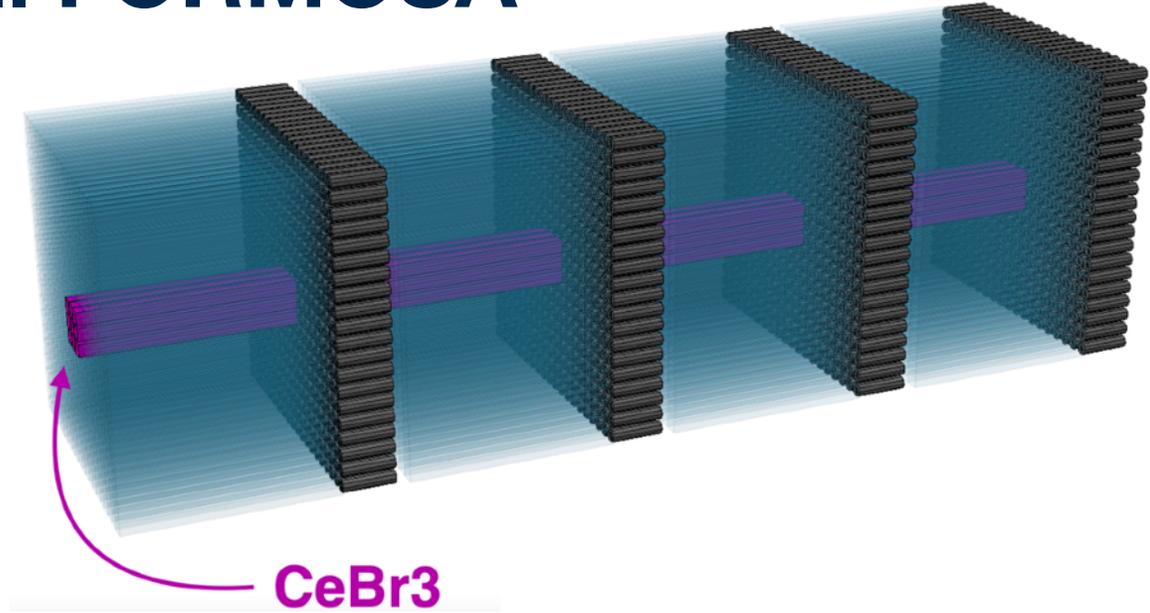
Field impact on different charges (F. Kling)



- Current exclusions assume **no impact** on MCPs from rock/LHC material/magnetic field
- Very reasonable for  $Q < \sim 0.1$  but what about higher charges? → need to evaluate probability for MCPs to reach detector!
- Ongoing work: use FORESEE together with propagation tools developed for milliQan (updated with LHC BDSIM model) to evaluate reach for higher charges

# Potential changes in the full FORMOSA

- Potential sub-detector made of CeBr3:
  - ~35x more photons/cm compared to plastic scintillators, fast with low internal radioactivity
  - studying in lab
- Considering whether a 3 layer-FORMOSA would provide enough background rejection
- Explore the versatility of the veto panels



# Collaboration



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UNIVERSITY OF  
**Nebraska**  
Lincoln

The FORMOSA collaboration is comprised largely from the milliQan collaboration.

Lots (and in fact, most) of our experience is easily transferable to FORMOSA.

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