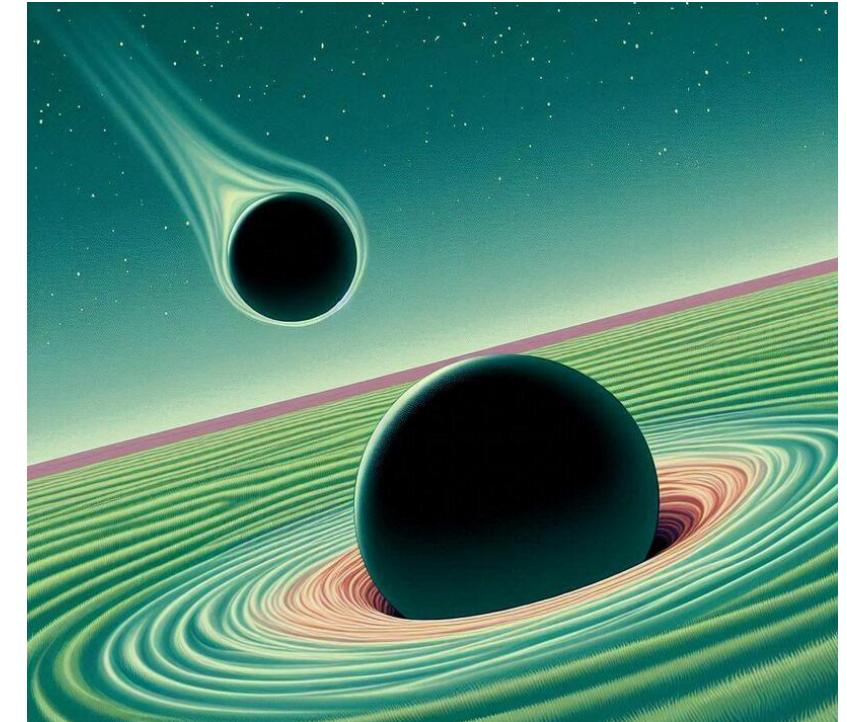
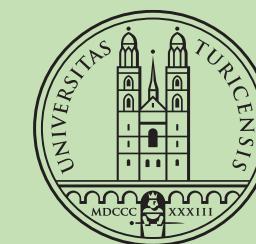


# Waveforms in the Post-Minkowskian Expansion



Based on 2312.14859 with Harald Ita, Manfred Kraus and Johannes Schlenk

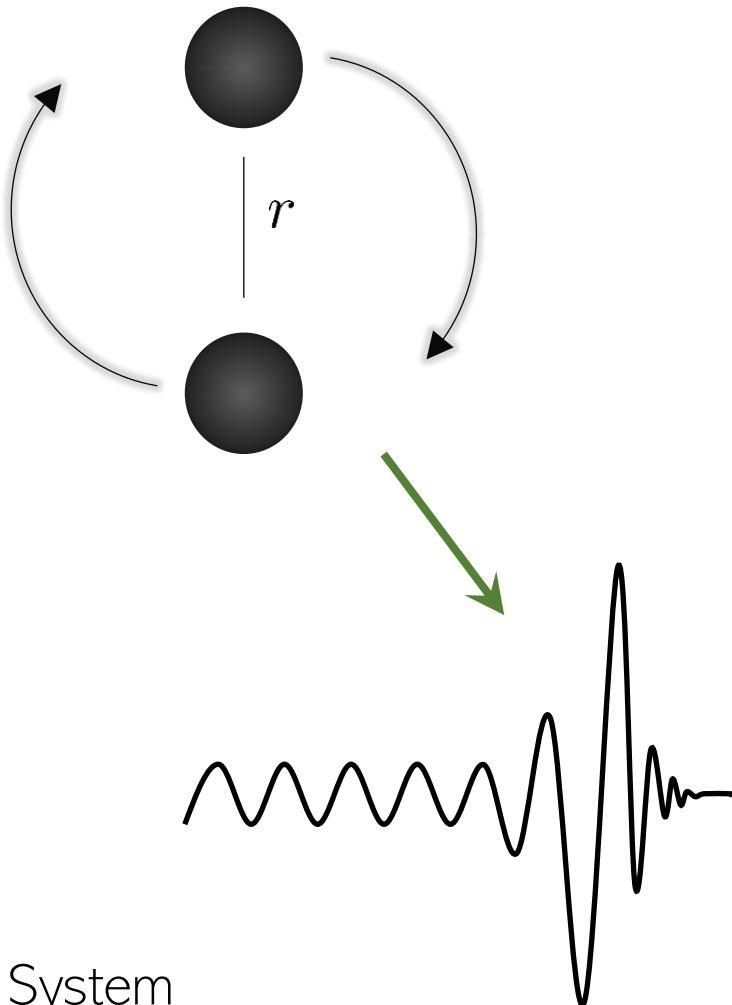
SPS Annual Meeting 2024  
Lara Bohnenblust  
11. September 2024



**University of  
Zurich**<sup>UZH</sup>

# Gravitational Scattering

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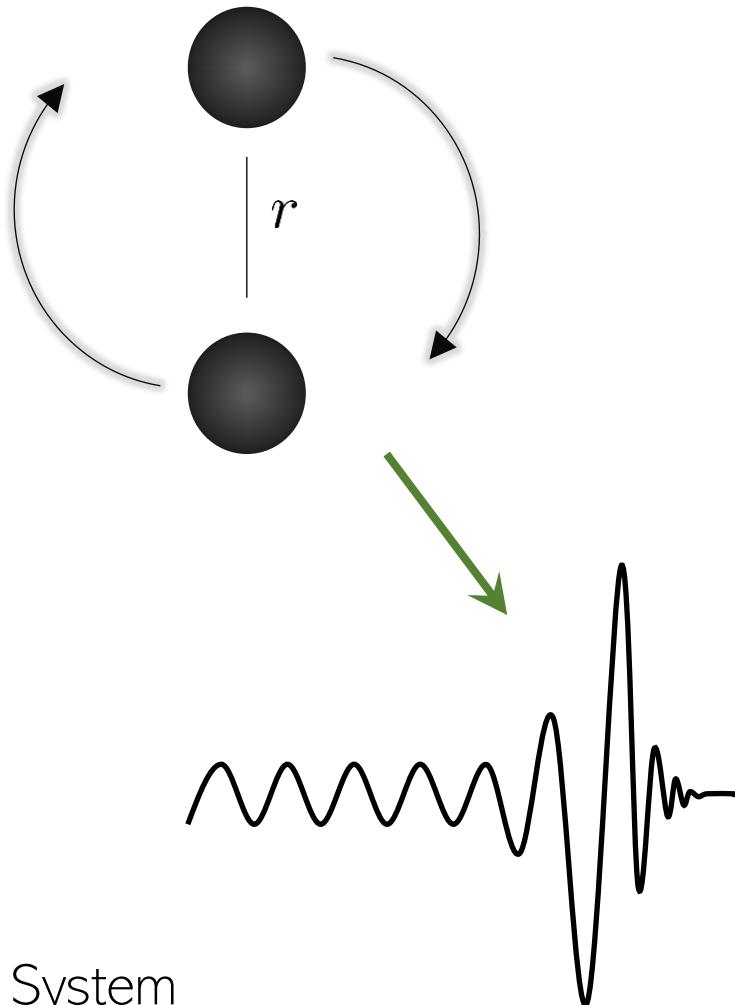


Bound System

$$v^2 \sim \frac{Gm}{r} \ll 1$$

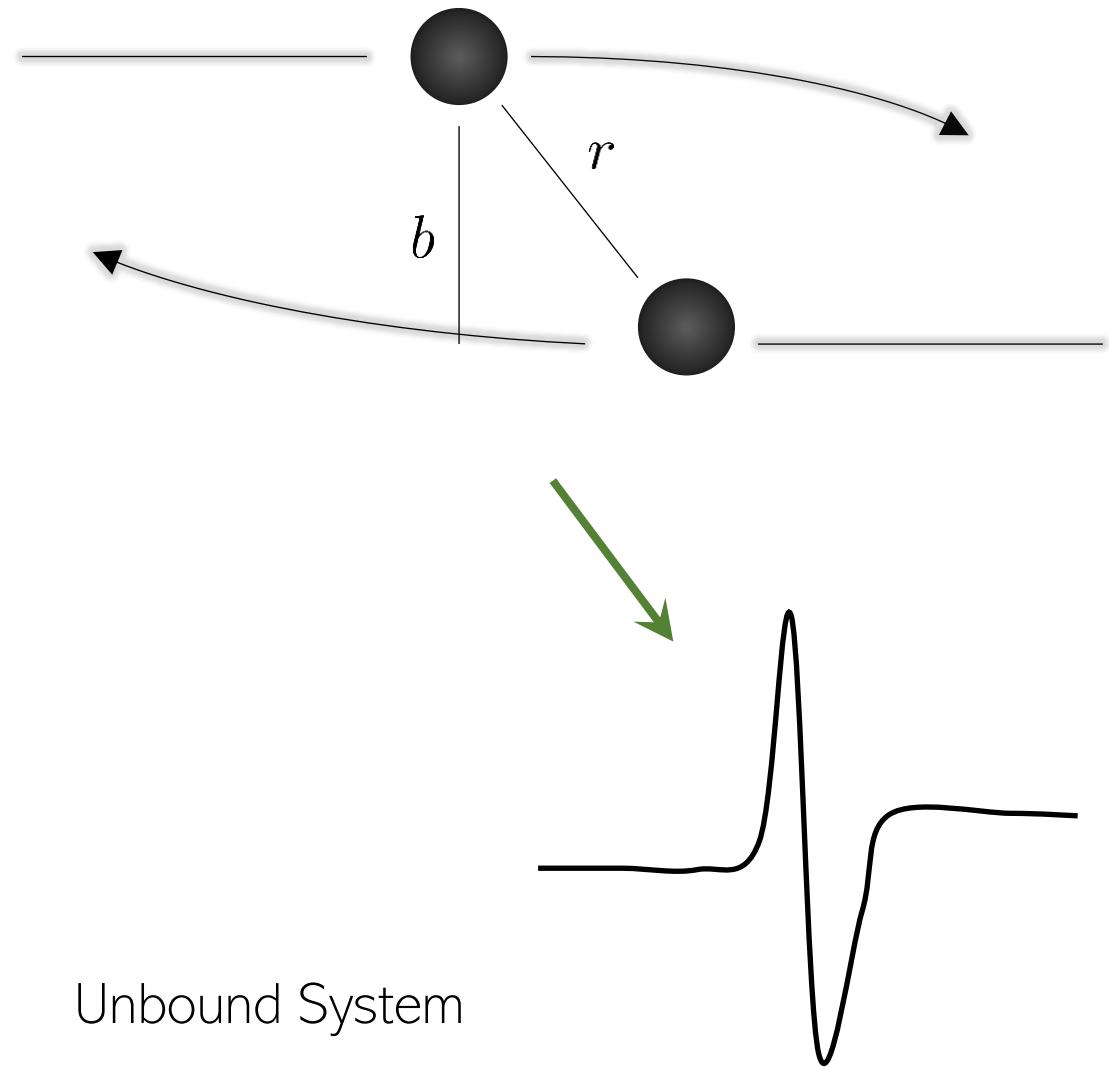
# Gravitational Scattering

---



Bound System

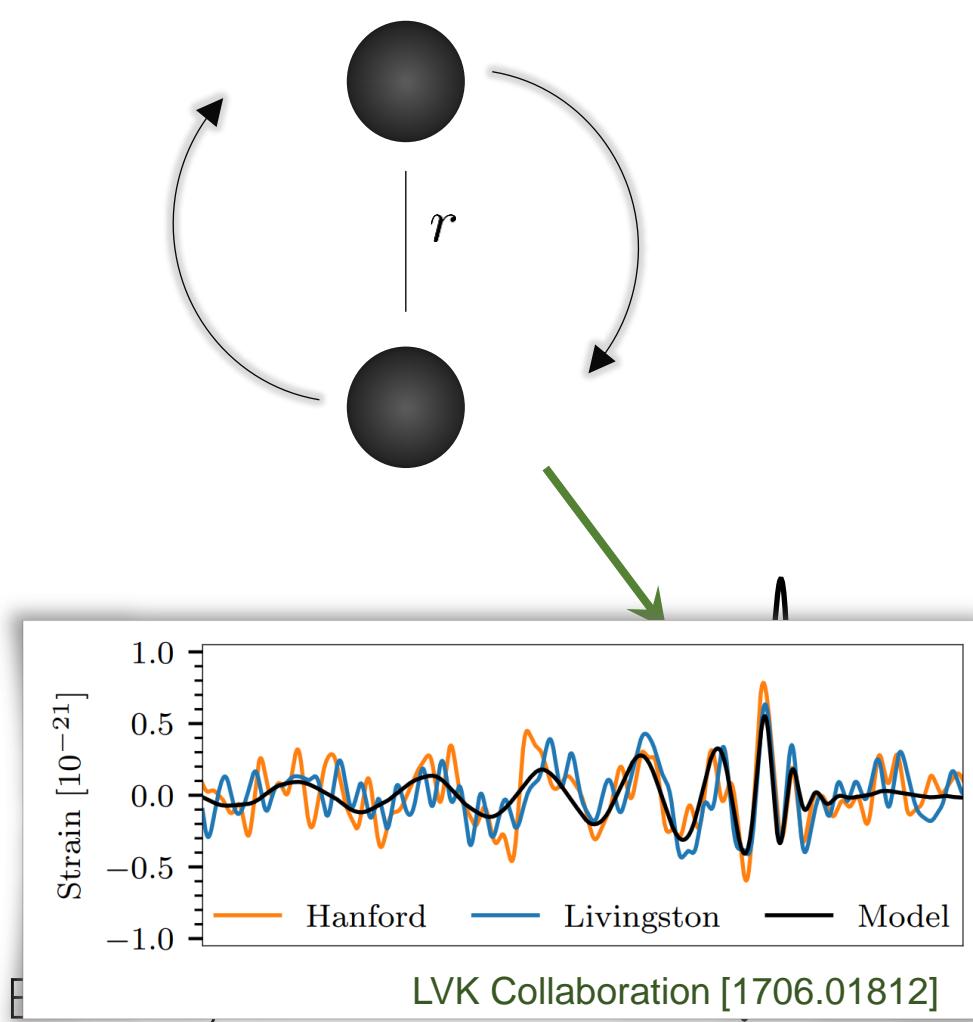
$$v^2 \sim \frac{Gm}{r} \ll 1$$



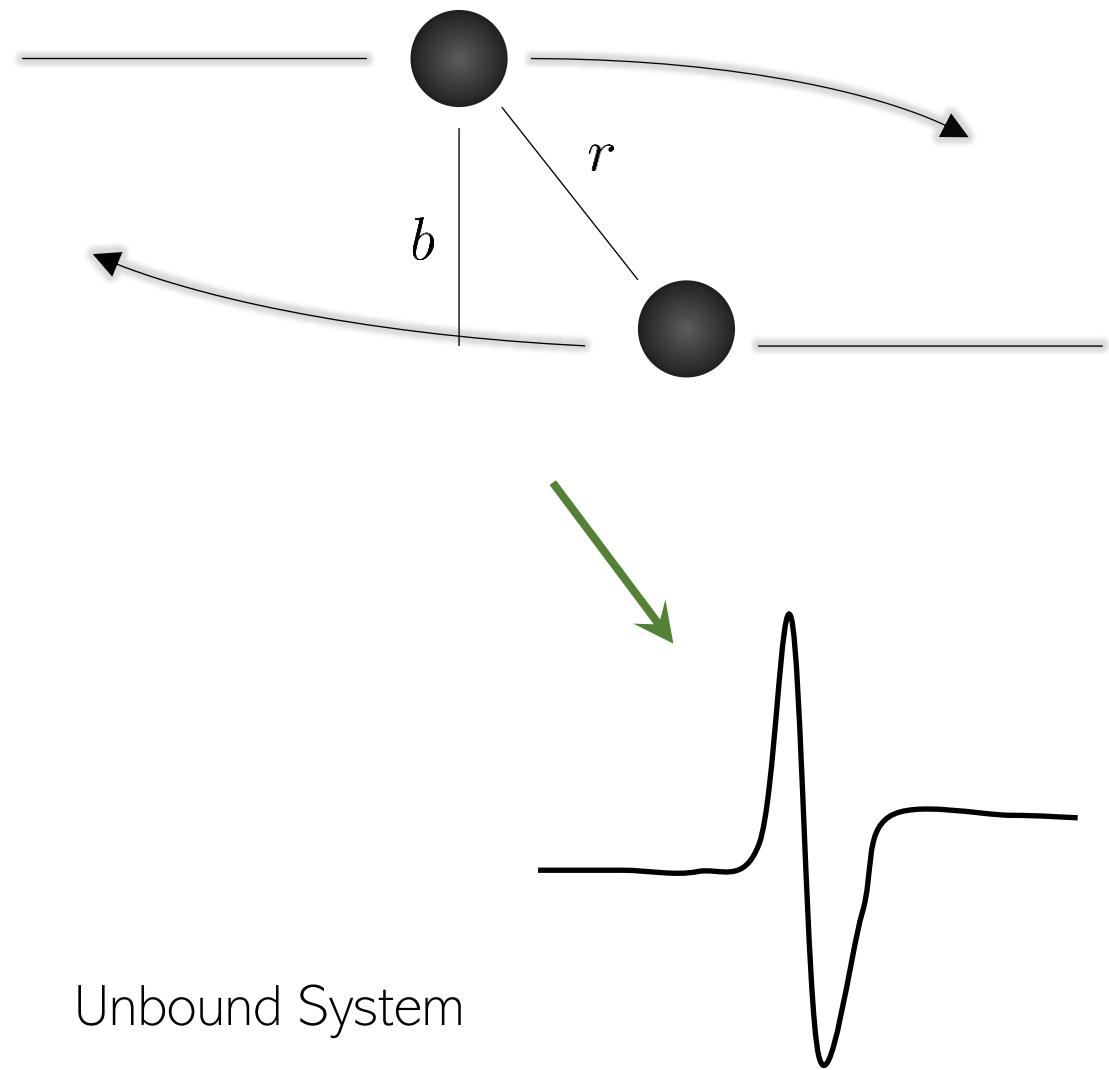
Unbound System

$$\frac{Gm}{r} \ll 1$$

# Gravitational Scattering

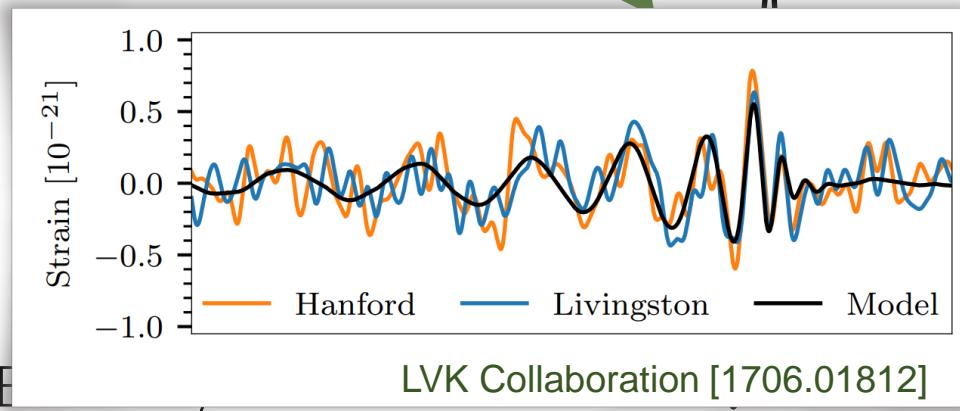
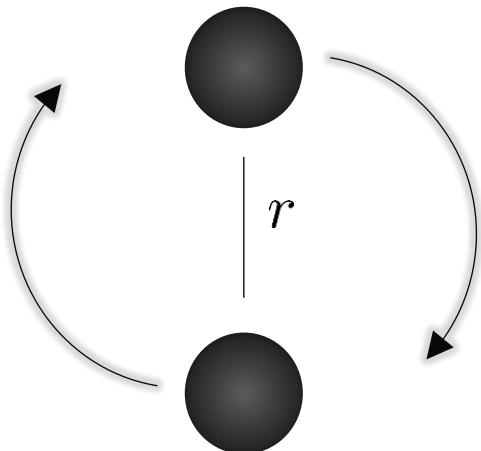


$$v^2 \sim \frac{Gm}{r} \ll 1$$

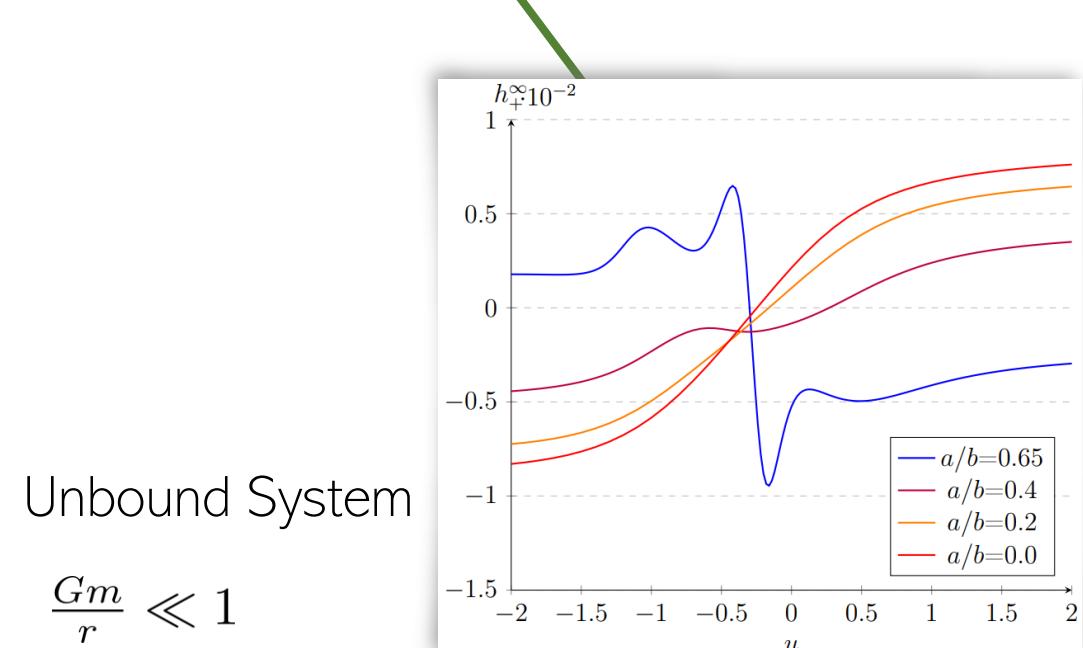
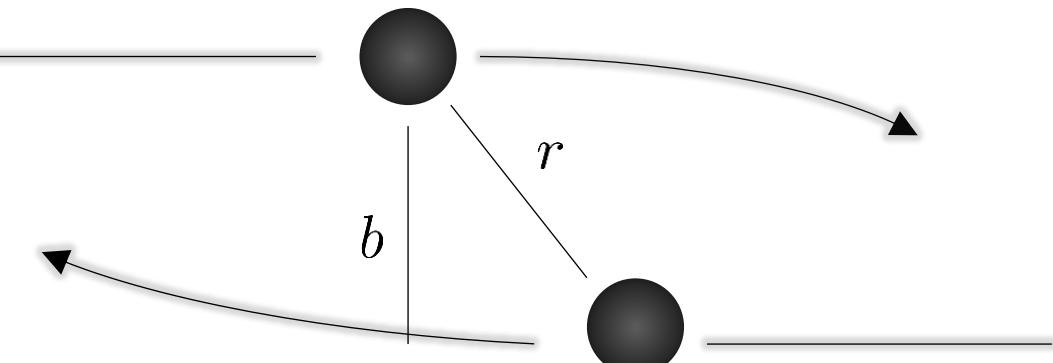


$$\frac{Gm}{r} \ll 1$$

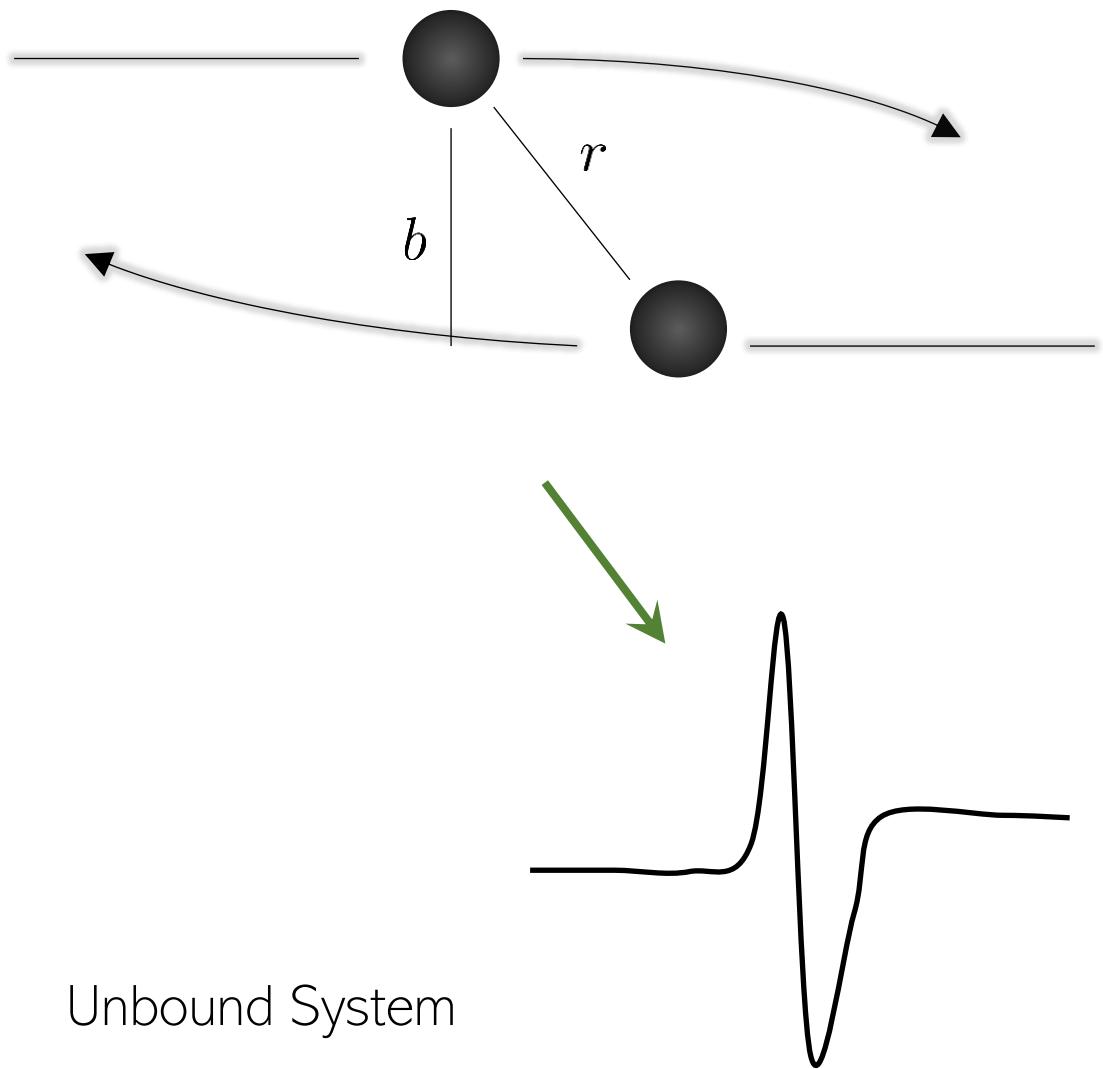
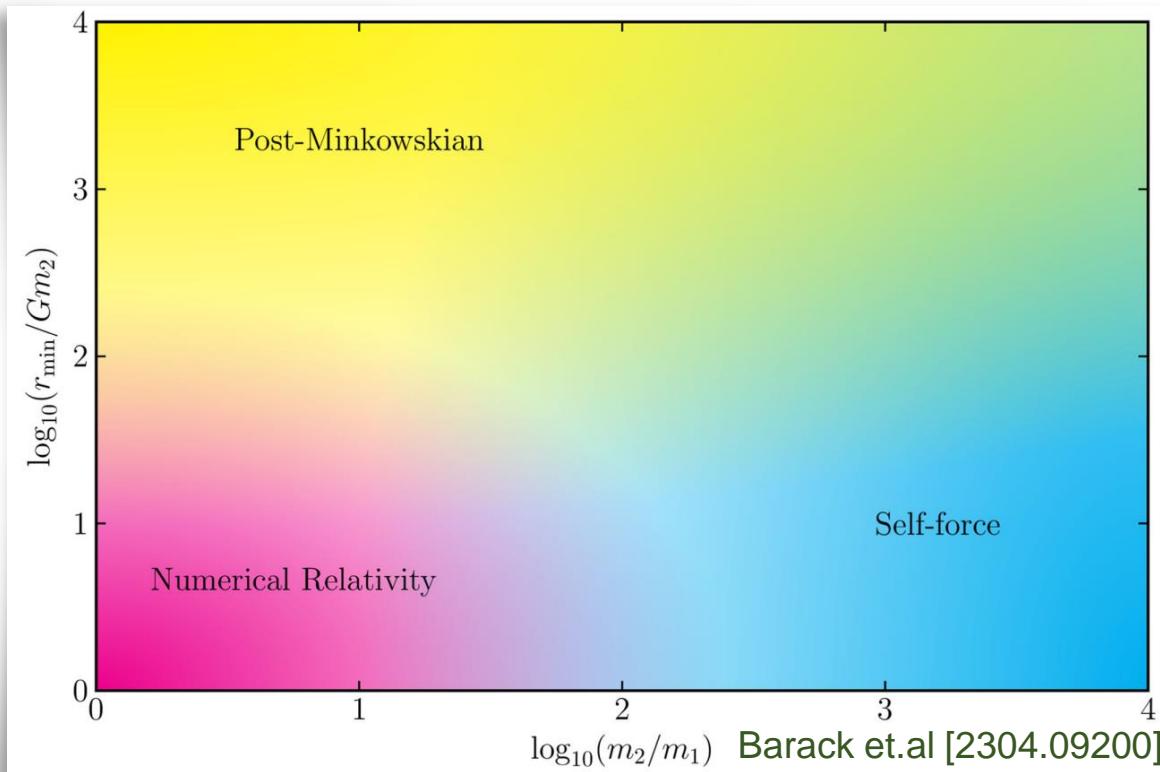
# Gravitational Scattering



$$v^2 \sim \frac{Gm}{r} \ll 1$$



# Gravitational Scattering



Unbound System

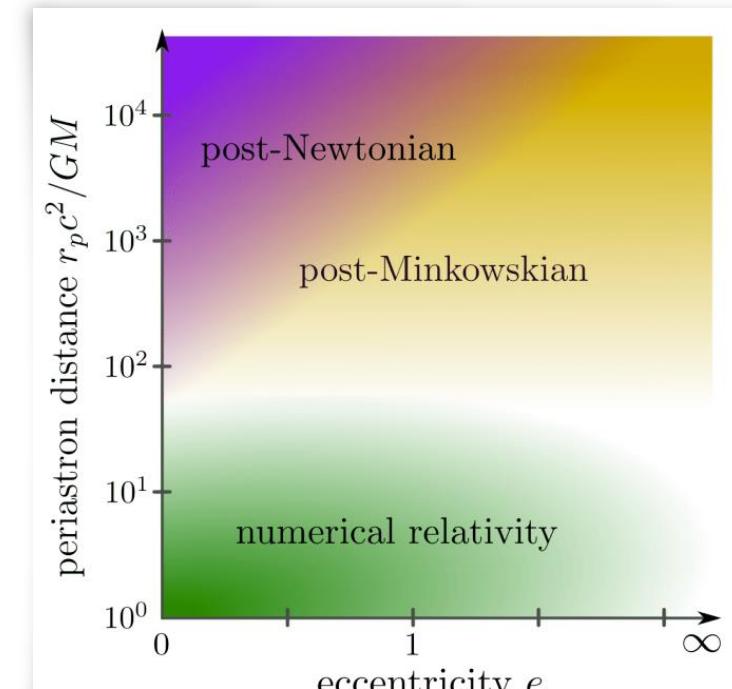
$$\frac{Gm}{r} \ll 1$$

# PN vs PM

---

	OPN	1PN	2PN	3PN	4PN	5PN	6PN
0PM		$(1 + v^2 + v^4 + v^6 + v^8 + v^{10} + v^{12} + v^{14} + \dots)$					
1PM	$G$		$(1 + v^2 + v^4 + v^6 + v^8 + v^{10} + v^{12} + \dots)$				
2PM	$+ G^2$			$(1 + v^2 + v^4 + v^6 + v^8 + v^{10} + \dots)$			
3PM	$+ G^3$				$(1 + v^2 + v^4 + v^6 + v^8 + \dots)$		
4PM	$+ G^4$					$(1 + v^2 + v^4 + v^6 + \dots)$	
5PM	$+ G^5$						$(1 + v^2 + v^4 + \dots)$
6PM	$+ G^6$						$(1 + v^2 + \dots)$
7PM	$+ G^7$						$(1 + \dots)$
$+ \dots$							

Buonanno, Khalil, O'Connell, Roiban, Solon, Zeng [2204.05194]

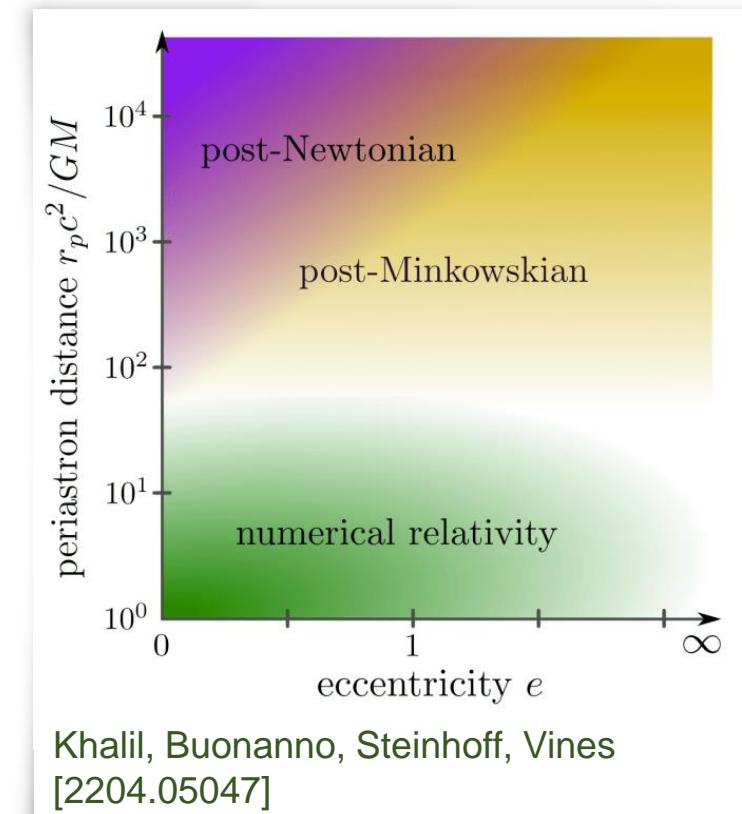


Khalil, Buonanno, Steinhoff, Vines  
[2204.05047]

# PN vs PM

	OPN	1PN	2PN	3PN	4PN	5PN	6PN
OPM	$(1 + v^2 + v^4 + v^6 + v^8 + v^{10} + v^{12} + v^{14} + \dots)$						
1PM	$G$	$(1 - v^2 - v^4 - v^6 - v^8 - v^{10} - v^{12} - \dots)$					
2PM	$+ G^2$		$(1 - v^2 - v^4 - v^6 - v^8 - v^{10} - \dots)$				
3PM	$+ G^3$			$(1 + v^2 + v^4 + v^6 + v^8 + \dots)$			
4PM	$+ G^4$				$(1 - v^2 + v^4 + v^6 + \dots)$		
5PM	$+ G^5$					$(1 + v^2 + v^4 + \dots)$	
6PM	$+ G^6$					$(1 + v^2 + \dots)$	
7PM	$+ G^7$						$(1 + \dots)$
$+\dots$							

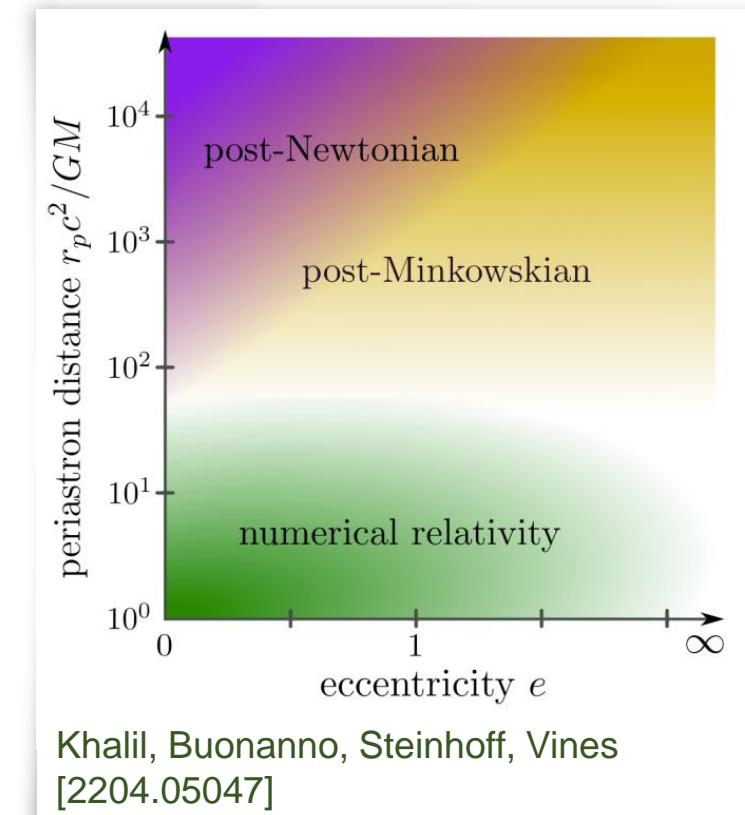
Buonanno, Khalil, O'Connell, Roiban, Solon, Zeng [2204.05194]



# PN vs PM

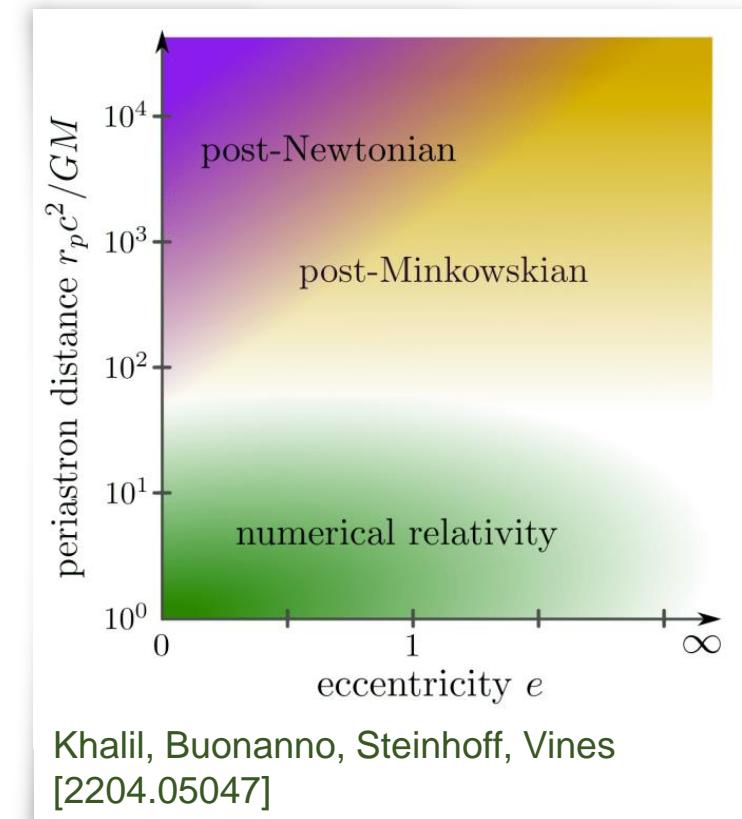
	OPN	1PN	2PN	3PN	4PN	5PN	6PN
OPM	$(1 + v^2 + v^4 + v^6 + v^8 + v^{10} + v^{12} + v^{14} + \dots)$						
1PM	$G$	$(1 - v^2 - v^4 - v^6 - v^8 - v^{10} - v^{12} - \dots)$					
2PM	$+ G^2$		$(1 - v^2 - v^4 - v^6 - v^8 - v^{10} - \dots)$				
3PM	$+ G^3$			$(1 + v^2 + v^4 + v^6 + v^8 + \dots)$			
4PM	$+ G^4$				$(1 - v^2 + v^4 + v^6 + \dots)$		
5PM	$+ G^5$					$(1 + v^2 + v^4 + \dots)$	
6PM	$+ G^6$						$(1 + v^2 + \dots)$
7PM	$+ G^7$						$(1 + \dots)$
$\vdots$	$\vdots$						

Buonanno, Khalil, O'Connell, Roiban, Solon, Zeng [2204.05194]



# PN vs PM

	OPN	1PN	2PN	3PN	4PN	5PN	6PN
OPM		$(1 + v^2 + v^4 + v^6 + v^8 + v^{10} + v^{12} + v^{14} + \dots)$					
1PM	$G$	$(1 - v^2 + v^4 + v^6 + v^8 + v^{10} + v^{12} + \dots)$					
2PM	$+ G^2$	$(1 - v^2 + v^4 + v^6 + v^8 + v^{10} + \dots)$					
3PM	$+ G^3$	$(1 + v^2 + v^4 + v^6 + v^8 + \dots)$					
4PM	$+ G^4$	$(1 - v^2 + v^4 + v^6 + \dots)$					
5PM	$+ G^5$	$(1 + v^2 + v^4 + \dots)$					
6PM	$+ G^6$	$(1 + v^2 + \dots)$					
7PM	$+ G^7$	$(1 + \dots)$					
	$+ \dots$						



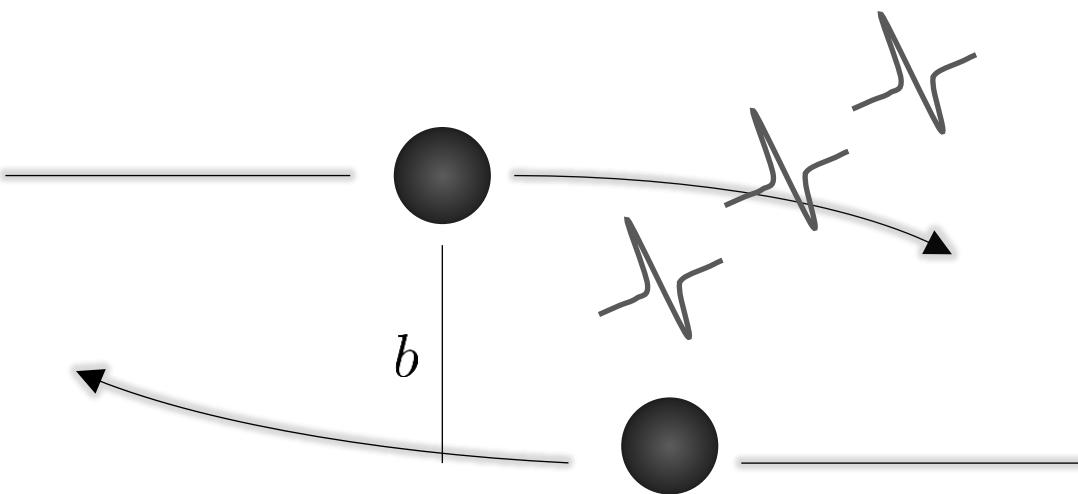
# The Post-Minkowskian (PM) Expansion

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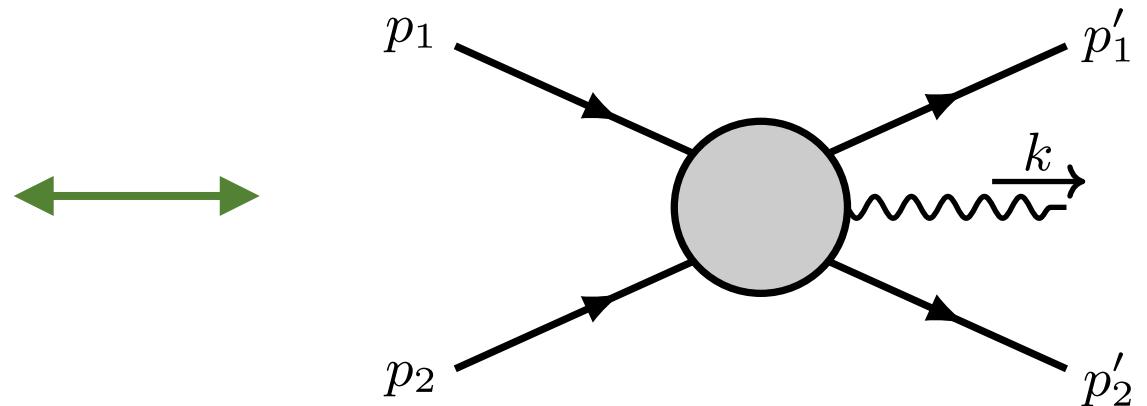
$$\mathcal{O} \sim \sum_i c_i(p_1, p_2, k, b) \cdot \left( \frac{Gm}{b} \right)^i$$

# The Post-Minkowskian (PM) Expansion

$$\mathcal{O} \sim \sum_i c_i(p_1, p_2, k, b) \cdot \left( \frac{Gm}{b} \right)^i$$



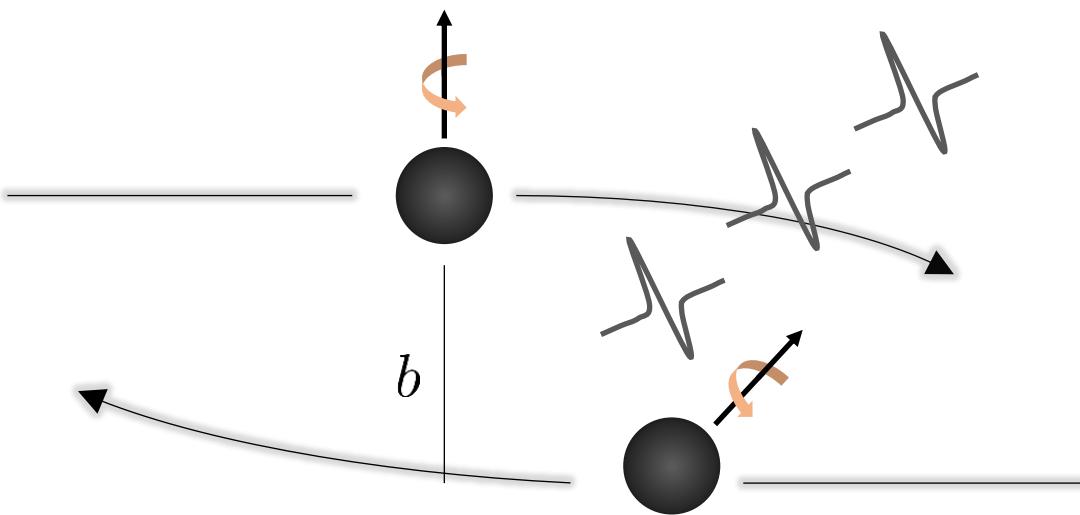
Scattering encounter



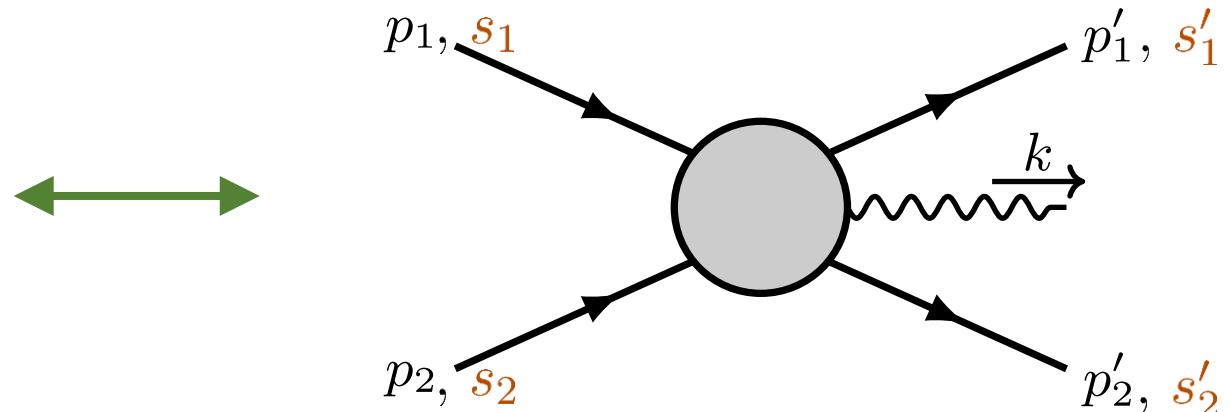
Scattering amplitude

# The Post-Minkowskian (PM) Expansion

$$\mathcal{O} \sim \sum_{ijk} c_{ij}(p_1, p_2, k, b) \cdot \left(\frac{Gm}{b}\right)^i \left(\frac{S_1}{m_1 b}\right)^j \left(\frac{S_2}{m_2 b}\right)^k$$



Scattering encounter



Scattering amplitude

Minimally-coupled scalar field

$$\mathcal{L}_{(\phi,m)} = \frac{1}{2} \sqrt{|g|} \left[ g^{\mu\nu} (\partial_\mu \phi)(\partial_\nu \phi) - m^2 \phi^2 \right]$$

Minimally-coupled Proca field

$$\mathcal{L}_{(V,m)} = -\frac{1}{4} \sqrt{|g|} \left[ g^{\mu\rho} g^{\nu\sigma} F_{\mu\nu} F_{\rho\sigma} - 2m^2 g^{\mu\nu} V_\mu V_\nu \right], \quad F_{\mu\nu} = \partial_\mu V_\nu - \partial_\nu V_\mu$$

Black-hole scattering

$$\mathcal{L}_{\text{scalar}} = \mathcal{L}_{\text{EH}} + \mathcal{L}_{(\phi_1,m_1)} + \mathcal{L}_{(\phi_2,m_2)}$$

$$\mathcal{L}_{s-1} = \mathcal{L}_{\text{EH}} + \mathcal{L}_{(V_1,m_1)} + \mathcal{L}_{(\phi_2,m_2)}$$

# PM waveforms

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$$\mathcal{O} \sim \sum_{ijk} c_{ij}(p_1, p_2, k, b) \cdot \left(\frac{Gm}{b}\right)^i \left(\frac{S_1}{m_1 b}\right)^j \left(\frac{S_2}{m_2 b}\right)^k$$

	$S^0$	$S^1$	$S^2$	$S^3$	$S^4$	$S^\infty$
Tree ( $G^2$ )	Kovacs & Thorne '78 Jakobsen, Mogull, Plefka, Steinhoff '21		Jakobsen, Mogull, Plefka, Steinhoff '21		De Angelis, Gonzo, Novichkov '23	Brandhuber, Brown, Chen, Gowdy, Travaglini '23 Aoude, Haddad, Heissenberg, Helset '23
1-loop ( $G^3$ )	Brandhuber, Brown, Chen, De Angelis, Gowdy '23 Herderschee, Roiban, Teng '23 Georgoudis, Heissenberg, Vazquez-Holm '23	LB, Ita, Kraus, Schlenk '23				

Comparison to the well-established MPM waveforms:

Bini, Damour, De Angelis, Geralico, Herderschee, Roiban, Teng [2402.06604]  
Georgoudis, Heissenberg, Russo [2402.06361]

# Waveform from Amplitudes

Kosower, Maybee, O'Connell [1811.10950]  
Cristofoli, Gonzo, Kosower, O'Connell [2107.10193]

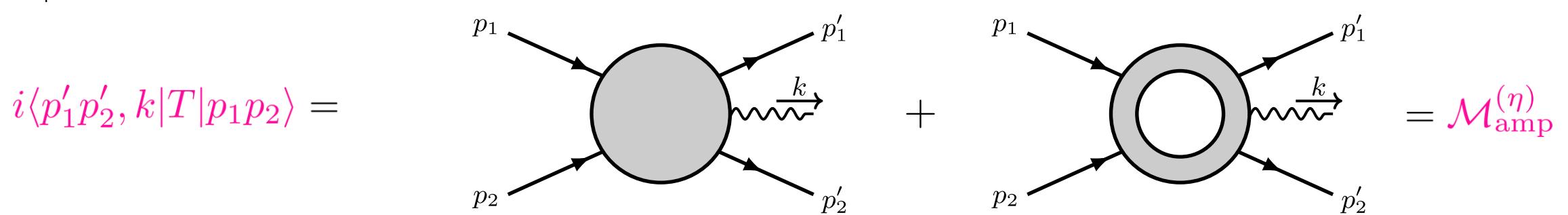
$$h^{(\eta)} = \lim_{r \rightarrow \infty} r \text{FT} \left[ i \langle p'_1, p'_2, k | T | p_1, p_2 \rangle + \langle p'_1, p'_2 | T^\dagger | r_1 r_2 \rangle \langle r_1, r_2, k | T | p_1, p_2 \rangle + \text{c.c.} \right]$$

# Waveform from Amplitudes

Kosower, Maybee, O'Connell [1811.10950]  
Cristofoli, Gonzo, Kosower, O'Connell [2107.10193]

$$h^{(\eta)} = \lim_{r \rightarrow \infty} r \text{FT} \left[ i \langle p'_1, p'_2, k | T | p_1, p_2 \rangle + \langle p'_1, p'_2 | T^\dagger | r_1 r_2 \rangle \langle r_1, r_2, k | T | p_1, p_2 \rangle + \text{c.c.} \right]$$

Up to order  $G^{5/2}$ :



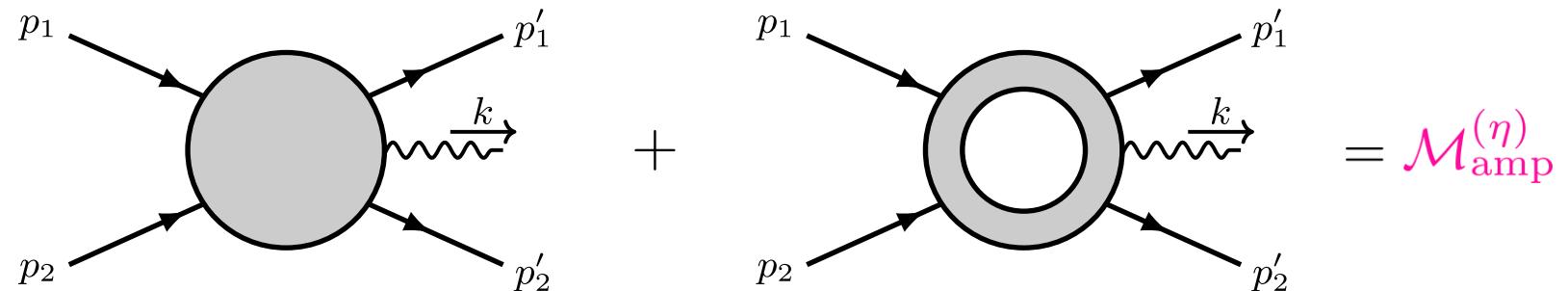
# Waveform from Amplitudes

Kosower, Maybee, O'Connell [1811.10950]  
 Cristofoli, Gonzo, Kosower, O'Connell [2107.10193]

$$h^{(\eta)} = \lim_{r \rightarrow \infty} r \text{FT} \left[ i \langle p'_1, p'_2, k | T | p_1, p_2 \rangle + \langle p'_1, p'_2 | T^\dagger | r_1 r_2 \rangle \langle r_1, r_2, k | T | p_1, p_2 \rangle + \text{c.c.} \right]$$

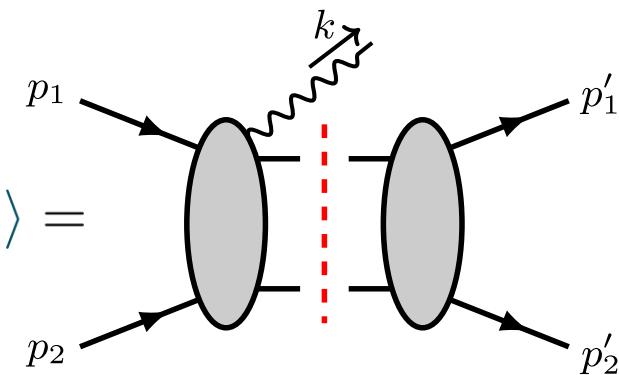
Up to order  $G^{5/2}$ :

$$i \langle p'_1 p'_2, k | T | p_1 p_2 \rangle =$$



$$+ \quad \quad \quad = \mathcal{M}_{\text{amp}}^{(\eta)}$$

$$\langle p'_1, p'_2 | T^\dagger | r_1 r_2 \rangle \langle r_1, r_2, k | T | p_1, p_2 \rangle =$$



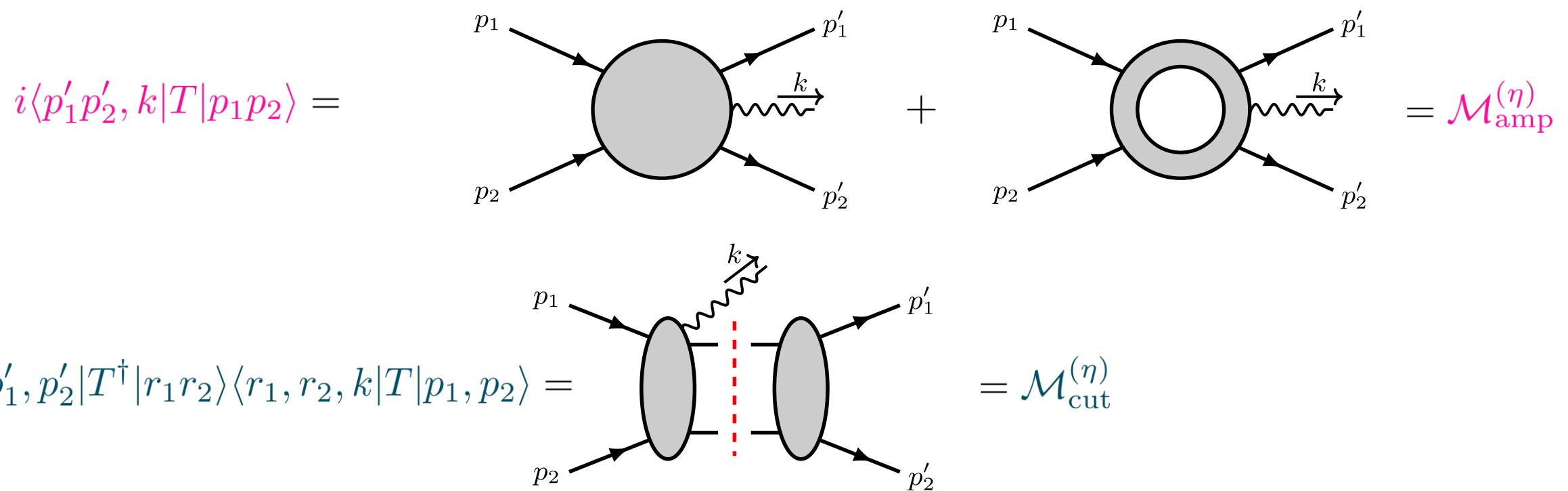
$$= \mathcal{M}_{\text{cut}}^{(\eta)}$$

# Waveform from Amplitudes

Kosower, Maybee, O'Connell [1811.10950]  
 Cristofoli, Gonzo, Kosower, O'Connell [2107.10193]

$$h^{(\eta)} = \lim_{r \rightarrow \infty} r \text{FT} \left[ i \langle p'_1, p'_2, k | T | p_1, p_2 \rangle + \langle p'_1, p'_2 | T^\dagger | r_1 r_2 \rangle \langle r_1, r_2, k | T | p_1, p_2 \rangle + \text{c.c.} \right]$$

Up to order  $G^{5/2}$ :



$$h^{(\eta)}(\tau) \sim \int_0^\infty d\omega e^{-i\omega\tau} \text{FT} \left[ \left( \mathcal{M}_{\text{amp}}^{(\eta)} + \mathcal{M}_{\text{cut}}^{(\eta)} + \text{c.c.} \right) \delta^{(D)}(q_1 + q_2 + k) \right]$$

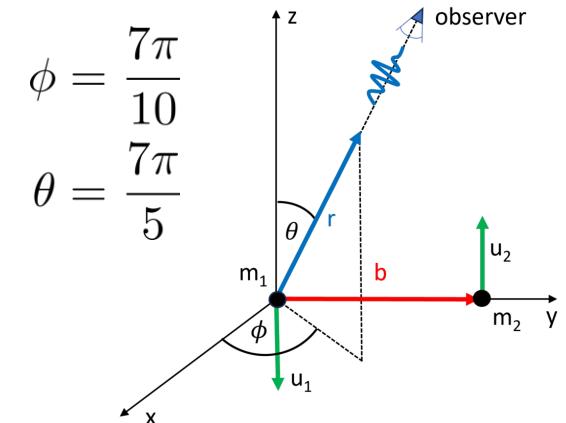
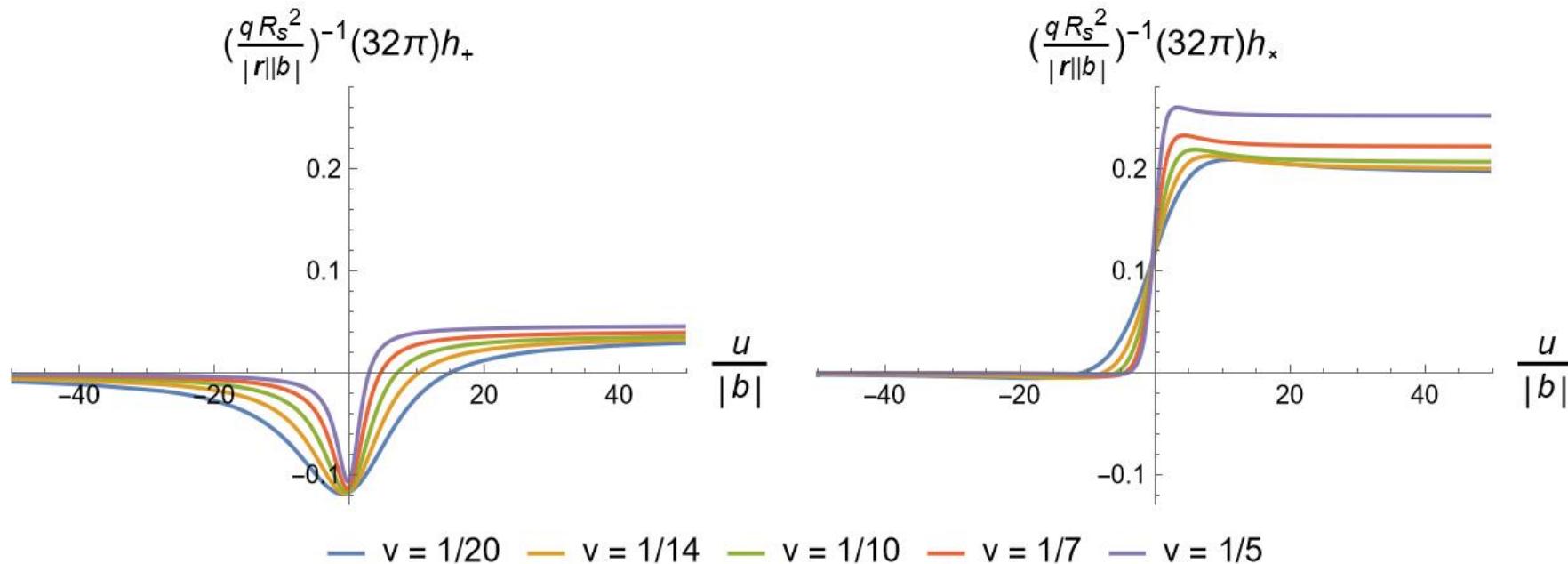
$$\mathcal{M} = \mathcal{M}^{\text{tree}} + \mathcal{M}^{\text{IR}} + \mathcal{M}^{\text{UV}} + \mathcal{M}^{\text{tail}} + \mathcal{M}^{\text{finite}}$$

Scalar case checked against

Brandhuber, Brown, Chen, De Angelis, Gowdy [2303.06111]  
Herderschee, Roiban, Teng [2303.06112]  
Georgoudis, Heissenberg, Vazquez-Holm [2312.14710]

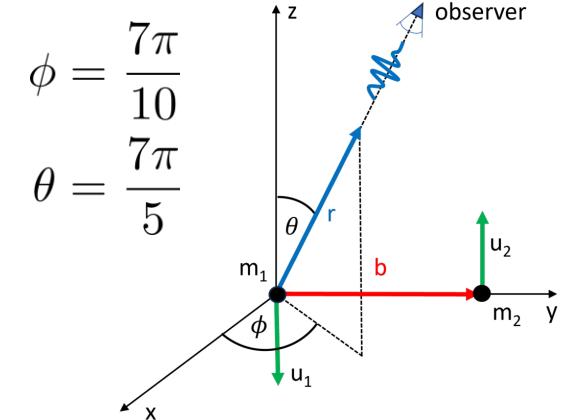
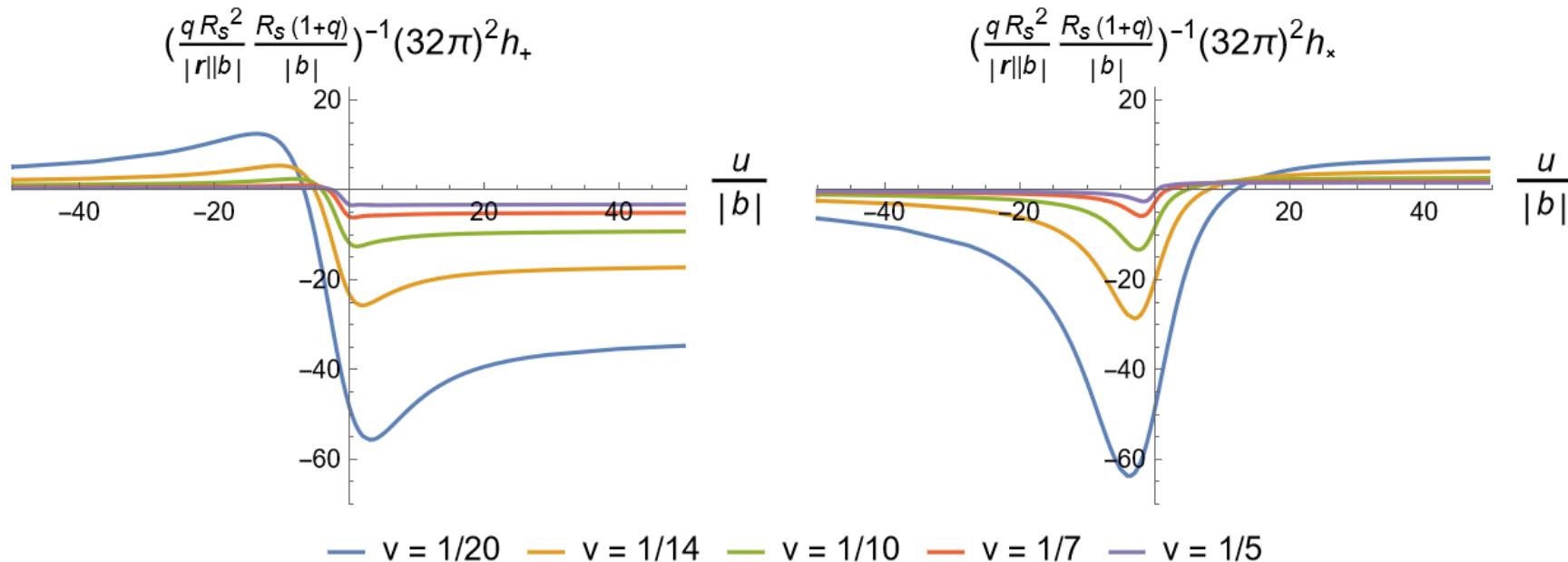
# Waveform: Time-domain

$\mathcal{M}_{\text{tree}}$  (LO) contribution:



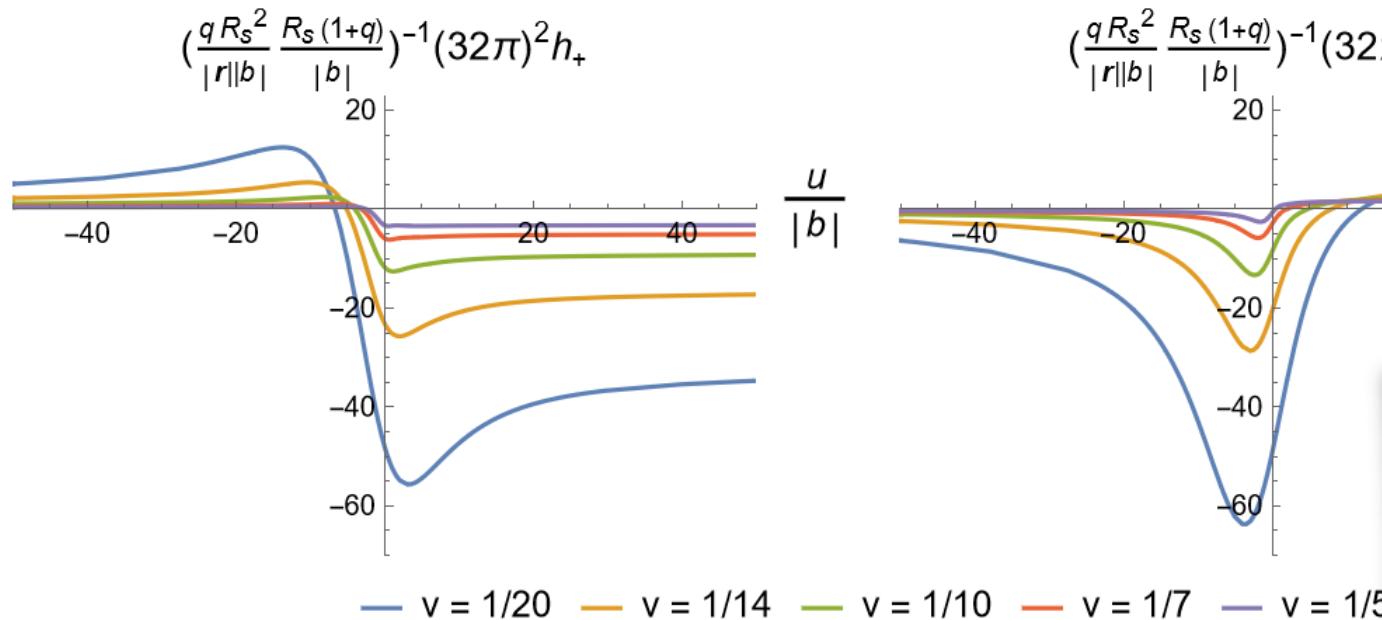
# Waveform: Time-domain

$\mathcal{M}_{\text{finite}}(\text{NLO})$  contribution:



# Waveform: Time-domain

$\mathcal{M}_{\text{finite}}(\text{NLO})$  contribution:



LISA sensitivity curve

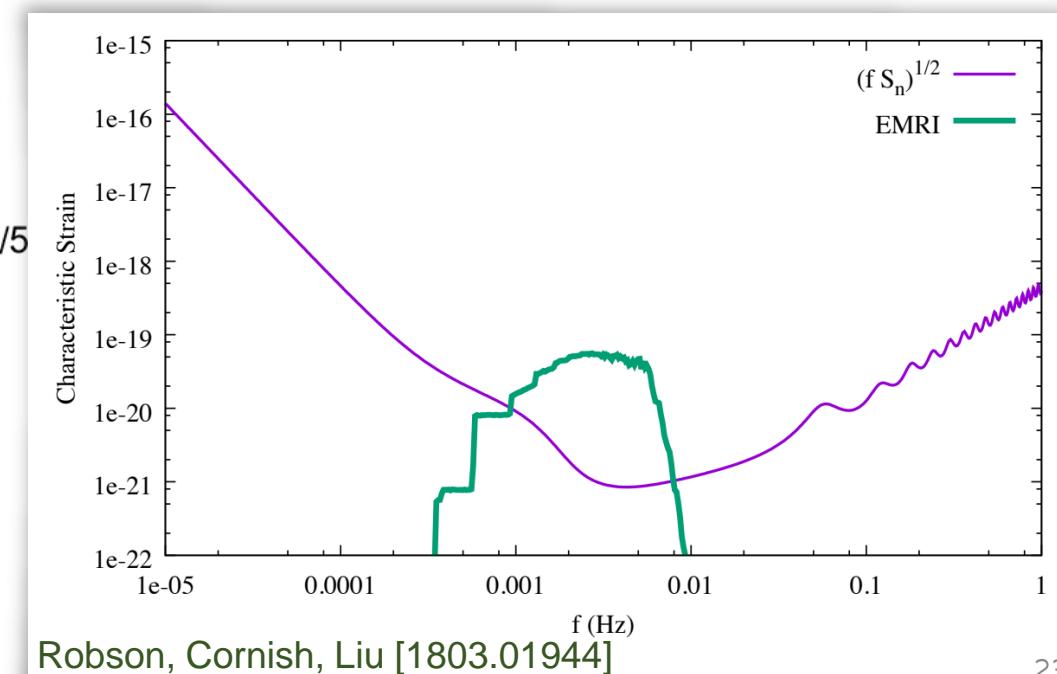
Detectability example

$$|b| = 1000 r_S^1,$$

$$m_1 = 5 \cdot 10^4 M_\odot, m_2 = 50 M_\odot,$$

$|\mathbf{r}| \approx 10^3 \text{ kpc}$ , distance to Andromeda

$$\begin{aligned} T &\sim 10^4 \text{ s}, \\ f &\sim 10^{-4} \text{ Hz}, \\ h_{\text{LO}} &\sim 10^{-13}, \\ h_{\text{NLO}} &\sim 10^{-16} \end{aligned}$$



# Outlook

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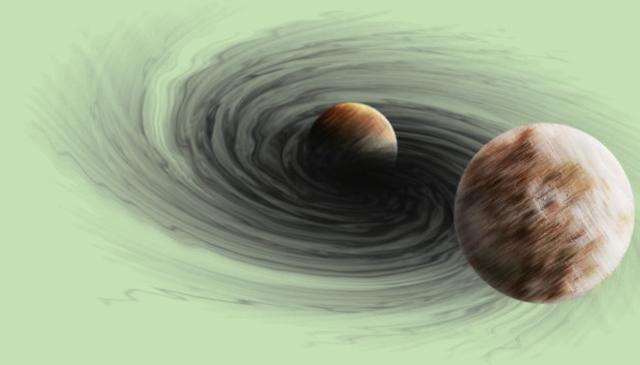
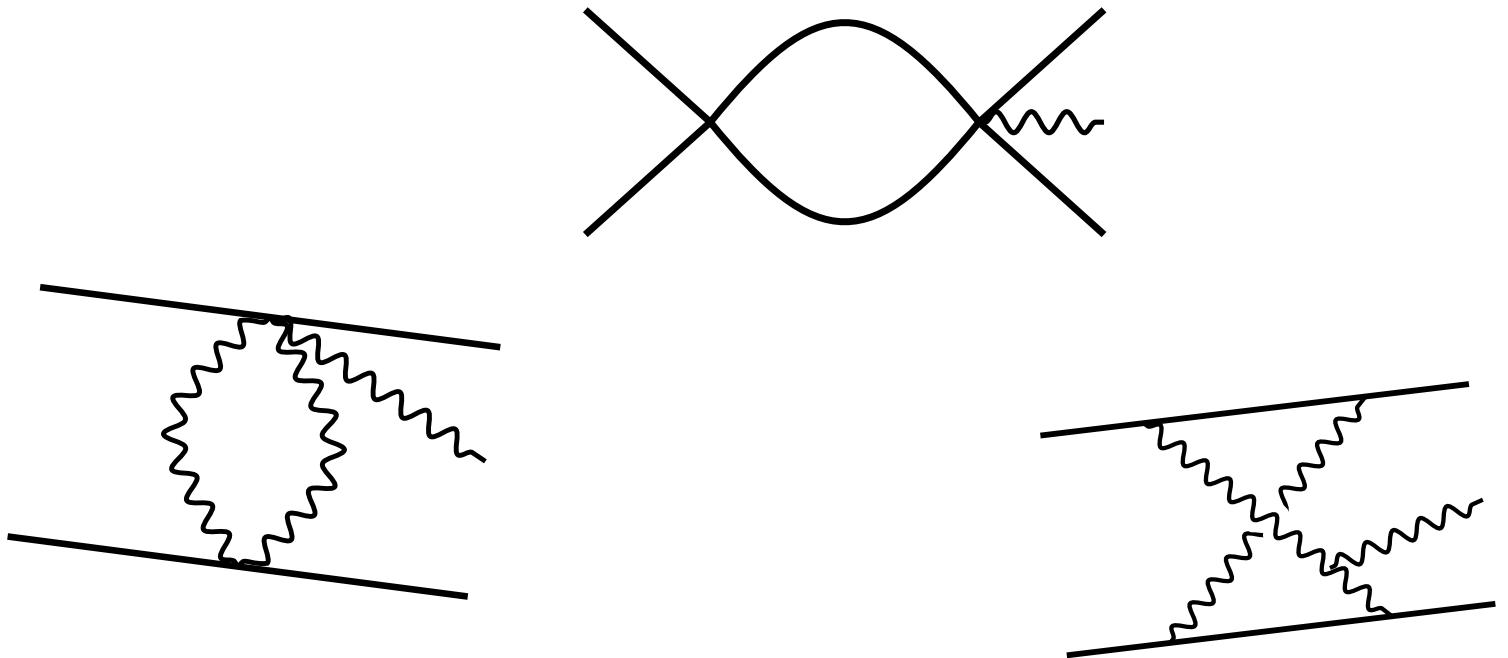
## Outlook

- ▶  $G^4$  and  $S^2$  spin corrections
- ▶ Rate and parameter space of scattering?
- ▶ Bound system **Adamo, Gonzo, Ilderton [2402.00124]**
- ▶ Fast and stable Fourier transformation required **Brunello, De Angelis [2403.08009]**



University of  
Zurich<sup>UZH</sup>

Thanks!



# Backup Slides

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# PM waveforms

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$$\begin{aligned} & G^2 (1 + S + S^2 + S^3 + S^4 + S^5 + S^6 + S^7 + \dots) \\ & + G^3 (1 + S + S^2 + S^3 + S^4 + S^5 + S^6 + \dots) \\ & + G^4 (1 + S + S^2 + S^3 + S^4 + S^5 + \dots) \\ & + G^5 (1 + S + S^2 + S^3 + S^4 + \dots) \\ & + G^6 (1 + S + S^2 + S^3 + \dots) \\ & + G^7 (1 + S + S^2 + \dots) \\ & + G^8 (1 + S + \dots) \\ & + \dots \end{aligned}$$

$$\mathcal{M} = \mathcal{M}^{\text{tree}} + \mathcal{M}^{\text{IR}} + \mathcal{M}^{\text{UV}} + \mathcal{M}^{\text{tail}} + \mathcal{M}^{\text{finite}}$$

**Tree** (including  $\epsilon$ - corrections)

$$\begin{aligned}\mathcal{M}_D^{\text{tree}} &= -\frac{\kappa^3 \bar{m}_1^2 \bar{m}_2^2}{4} \varepsilon_{\mu\nu}^{(\eta)} \left[ \frac{4P^\mu P^\nu}{q_1^2 q_2^2} + \frac{2y}{q_1^2 q_2^2} (Q^\mu P^\nu + P^\mu Q^\nu) + \left( y^2 - \frac{1}{D_s - 2} \right) \left( \frac{Q^\mu Q^\nu}{q_1^2 q_2^2} - \frac{P^\mu P^\nu}{\omega_1^2 \omega_2^2} \right) \right] \\ &= \mathcal{M}_{D=4}^{\text{tree}} + \epsilon \mathcal{M}_\epsilon^{\text{tree}} + \mathcal{O}(\epsilon),\end{aligned}$$

$$P^\mu = -\omega_1 u_2^\mu + \omega_2 u_1^\mu, \quad Q^\mu = (q_1 - q_2)^\mu + \frac{q_1^2}{\omega_1} u_1^\mu - \frac{q_2^2}{\omega_2} u_2^\mu,$$

$$D_s = 4 - 2\epsilon$$



# Waveform: Finite Contribution

LB, Ita, Kraus, Schlenk [2312.14859]

$$\mathcal{M} = \mathcal{M}^{\text{tree}} + \mathcal{M}^{\text{IR}} + \mathcal{M}^{\text{UV}} + \mathcal{M}^{\text{tail}} + \mathcal{M}^{\text{finite}}$$

Finite amplitude contribution

$$\begin{aligned}\mathcal{M}_{\text{finite}}^{\text{amp}} = & i \left[ r_1 + \frac{r_2}{\sqrt{y^2 - 1}} + \frac{r_3}{\sqrt{\omega_2^2 - q_1^2}} + \frac{r_4}{\sqrt{\omega_1^2 - q_2^2}} + \frac{r_5}{\sqrt{-q_1^2}} + \frac{r_6}{\sqrt{-q_2^2}} \right] \\ & + r_7 + \frac{r_8}{\sqrt{\omega_1^2 - q_2^2}} \log \left( \frac{\sqrt{\omega_1^2 - q_2^2} - \omega_1}{\sqrt{-q_2^2}} \right) + \frac{r_9}{\sqrt{\omega_2^2 - q_1^2}} \log \left( \frac{\sqrt{\omega_2^2 - q_1^2} - \omega_2}{\sqrt{-q_1^2}} \right) \\ & + \frac{r_{10}}{\sqrt{y^2 - 1}} \log \left( y + \sqrt{y^2 - 1} \right) + r_{11} \log \left( \frac{\omega_2^2}{\omega_1^2} \right) + r_{12} \log \left( \frac{q_1^2}{q_2^2} \right)\end{aligned}$$

# Waveform: Finite Contribution

LB, Ita, Kraus, Schlenk [2312.14859]

$$\mathcal{M} = \mathcal{M}^{\text{tree}} + \mathcal{M}^{\text{IR}} + \mathcal{M}^{\text{UV}} + \mathcal{M}^{\text{tail}} + \mathcal{M}^{\text{finite}}$$

Finite cut contribution

$$\begin{aligned}\mathcal{M}_{\text{finite}}^{\text{cut}} = & \frac{1}{\sqrt{y^2 - 1}} \left[ r_{13} + r_{14} \log(y^2 - 1) + r_{15} \log\left(\frac{\omega_1}{\omega_2}\right) + r_{16} \log\left(\frac{-q_1^2}{\omega_1 \omega_2}\right) + r_{17} \log\left(\frac{-q_2^2}{\omega_1 \omega_2}\right) \right] \\ & + r_{18} \log\left(y + \sqrt{y^2 - 1}\right)\end{aligned}$$

# Waveform: Tail Contribution

LB, Ita, Kraus, Schlenk [2312.14859]

$$\mathcal{M} = \mathcal{M}^{\text{tree}} + \mathcal{M}^{\text{IR}} + \mathcal{M}^{\text{UV}} + \mathcal{M}^{\text{tail}} + \mathcal{M}^{\text{finite}}$$

Tail contribution (scale-dependent)

$$\mathcal{M}^{\text{tail}} = -\log\left(\frac{\omega_1\omega_2}{\mu_{\text{IR}}^2}\right) \mathcal{W}_S \mathcal{M}_{D=4}^{\text{tree}} - \log\left(\frac{\omega_1\omega_2}{\mu_{\text{UV}}^2}\right) \overline{\mathcal{M}}^{\text{UV}}$$

# Waveform: IR Divergence

LB, Ita, Kraus, Schlenk [2312.14859]

$$\mathcal{M} = \mathcal{M}^{\text{tree}} + \mathcal{M}^{\text{IR}} + \mathcal{M}^{\text{UV}} + \mathcal{M}^{\text{tail}} + \mathcal{M}^{\text{finite}}$$

IR divergence

$$\mathcal{M}^{\text{IR}} = \left[ \frac{1}{\epsilon} - \log \left( \frac{\mu_{\text{IR}}^2}{\mu^2} \right) \right] \mathcal{W}_S \mathcal{M}_{D=4}^{\text{tree}}$$

Weinberg soft factor

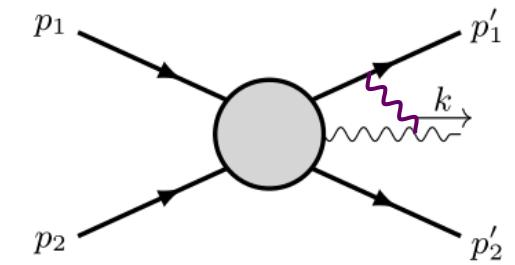
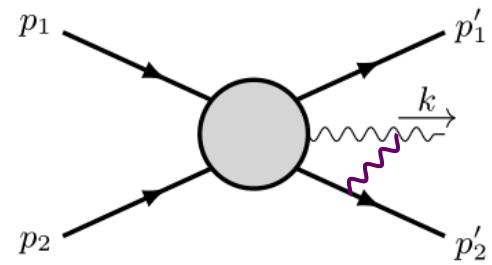
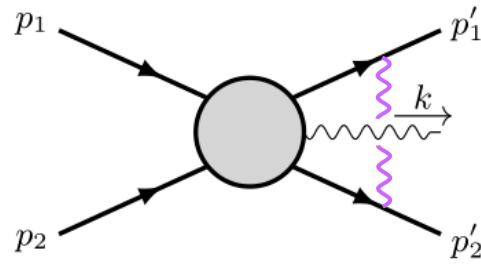
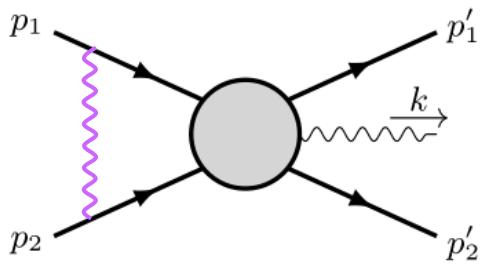
$$\mathcal{W}_S = iG(\bar{m}_1\omega_1 + \bar{m}_2\omega_2) \left( 1 + \frac{y(2y^2 - 3)}{2(y^2 - 1)^{3/2}} \right)$$

Cut contribution

# Waveform: IR Divergence

Weinberg soft theorem

Weinberg, 1965



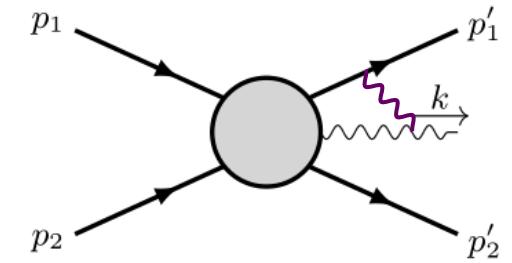
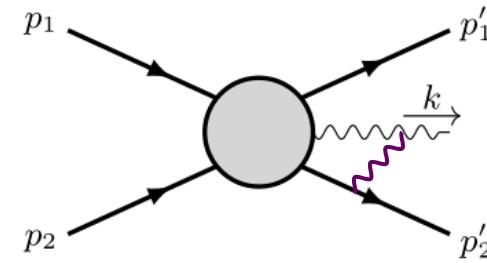
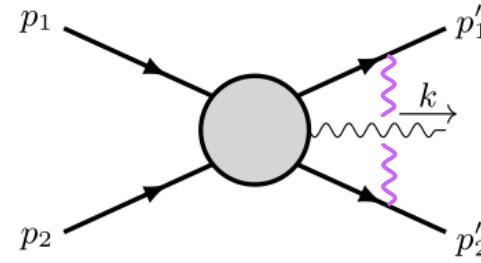
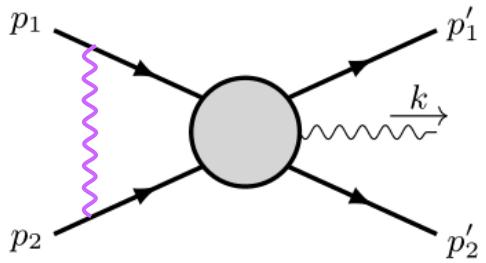
$$\text{Re}[W_S] = 0$$

$$\text{Im}[W_S] = \frac{1}{4} \sum_{\substack{i,j=1 \\ i \neq j}}^n c_{ij} \text{Im}[f_{ij}] = -\frac{\pi}{2} (2c_{12} + 2c_{1'2'} + c_{1'k} + c_{2'k}), \quad \text{Im}[f_{ij}] = \begin{cases} -\pi \Theta[(p_i \cdot p_j)], & m_{i/j} = 0, \\ -2\pi \Theta[(p_i \cdot p_j)], & \text{else}. \end{cases}$$

# Waveform: IR Divergence

Weinberg soft theorem

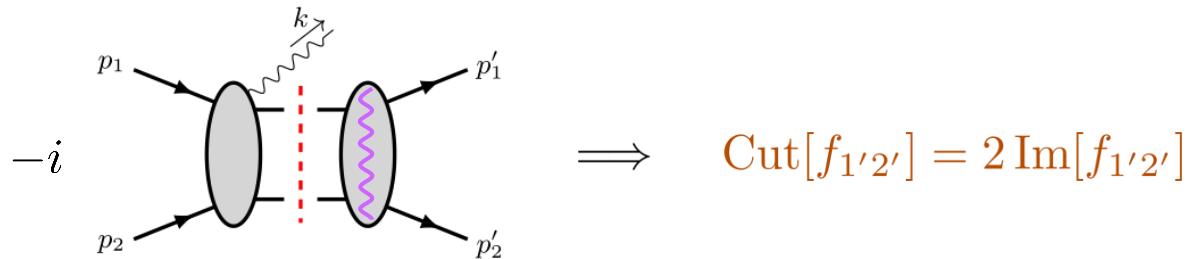
Weinberg, 1965



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Subtract cut:



$$\text{Im}[\mathcal{W}_S] = \text{Im}[W_S] - \text{Cut}_{1'2'}(W_S) = -\frac{\pi}{2} (2c_{12} - 2c_{1'2'} + c_{1'k} + c_{2'k}).$$

# Waveform: IR Divergence

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Exponentiation

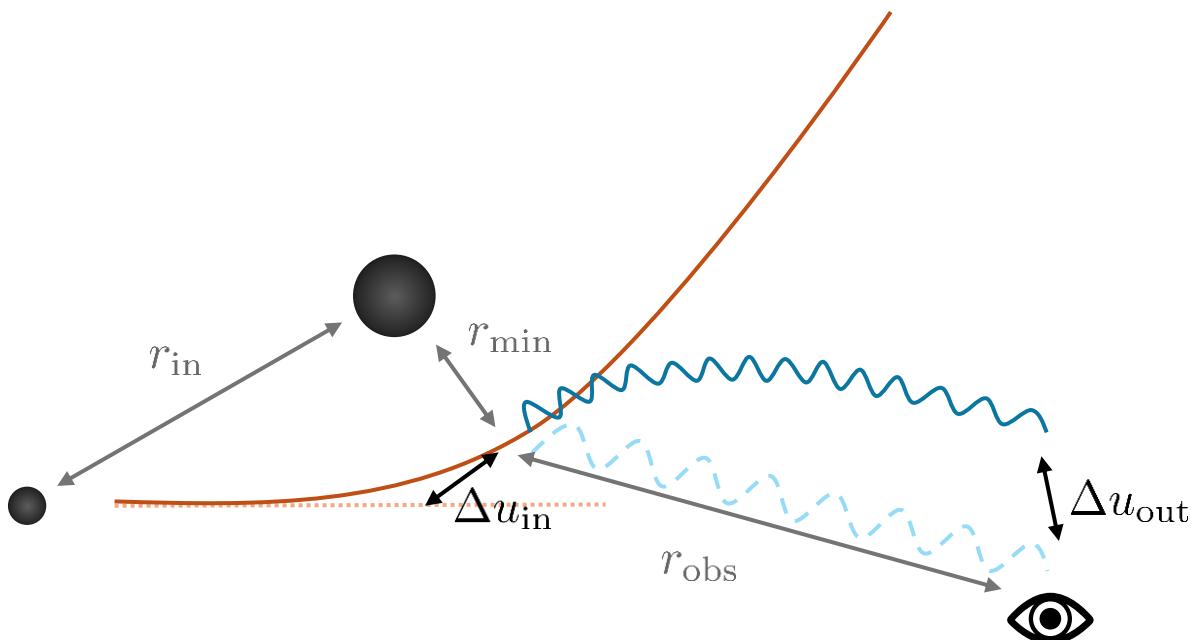
$$e^{-i\omega u} \mathcal{M} = e^{-i\omega \left[ u - \left( \frac{1}{\epsilon} - \log \frac{\mu_{\text{IR}}^2}{\mu^2} \right) \frac{\mathcal{W}_S}{i\omega} \right]} (\mathcal{M}^{\text{tree}} - \mathcal{W}_S \mathcal{M}_\epsilon^{\text{tree}} + \mathcal{M}^{\text{UV}} + \mathcal{M}^{\text{tail}} + \mathcal{M}^{\text{finite}}) + \mathcal{O}(G^3)$$

# Waveform: IR Divergence

Exponentiation

$$e^{-i\omega u} \mathcal{M} = e^{-i\omega \left[ u - \left( \frac{1}{\epsilon} - \log \frac{\mu_{\text{IR}}^2}{\mu^2} \right) \frac{\mathcal{W}_S}{i\omega} \right]} (\mathcal{M}^{\text{tree}} - \mathcal{W}_S \mathcal{M}_\epsilon^{\text{tree}} + \mathcal{M}^{\text{UV}} + \mathcal{M}^{\text{tail}} + \mathcal{M}^{\text{finite}}) + \mathcal{O}(G^3)$$

Classical interpretation



Shapiro delay

$$\Delta u_{\text{out}} = \frac{2G}{\omega} (\bar{m}_1 \omega_1 + \bar{m}_2 \omega_2) \log \left( \frac{r_{\text{obs}}}{r_{\min}} \right)$$

Deflection of incoming trajectory

$$\Delta u_{\text{in}} = \frac{2G}{\omega} (\bar{m}_1 \omega_1 + \bar{m}_2 \omega_2) \frac{y(2y^2 - 3)}{2(y^2 - 1)^{3/2}} \log \left( \frac{r_{\text{in}}}{r_{\min}} \right)$$

Cut contribution

Caron-Huot, Giroux, Hannesdottir, Mizera [2308.02125]

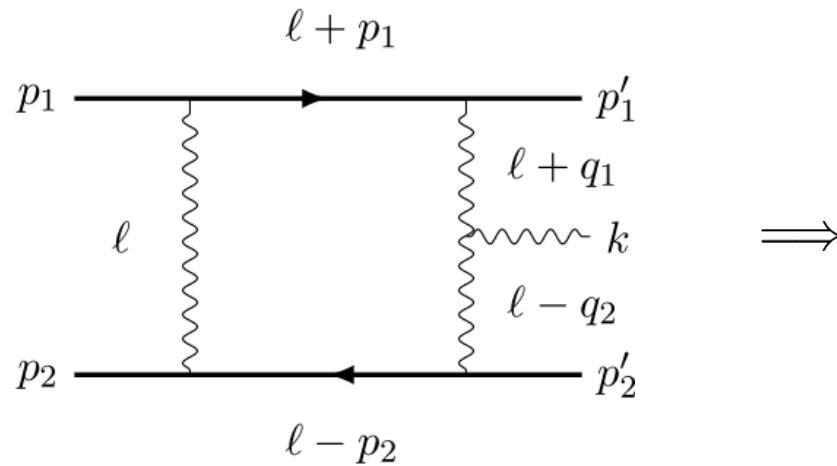
# Waveform: UV Divergence

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$$\mathcal{M}^{\text{UV}} = - \left( \frac{\kappa}{2} \right)^5 \frac{i}{8\pi\epsilon} (\bar{m}_1\omega_1 + \bar{m}_2\omega_2) \frac{\bar{m}_1^2 \bar{m}_2^2 (\omega_1^2 + \omega_2^2 + y\omega_1\omega_2)(1 - 2y^2)^2}{\omega_1\omega_2^3 (y^2 - 1)^{3/2}} F_1^{(2h)}$$

Contact term that integrates to  $\delta(|b|)$  in the FT and does not contribute to the far-field waveform!

Where does the UV pole come from?

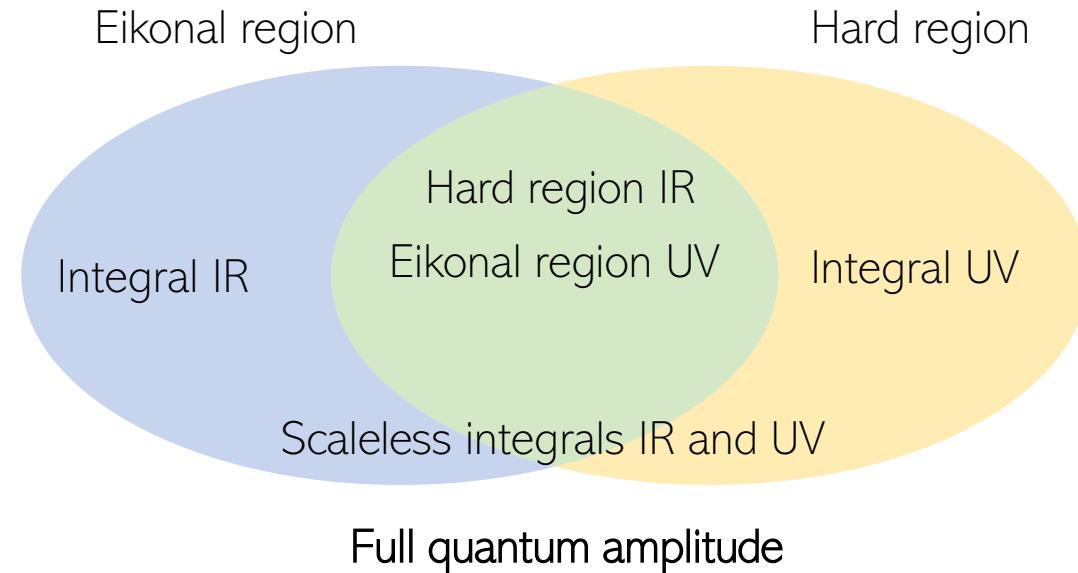


$$G^{\text{eik}} \sim e^{\epsilon\gamma_E} \int \frac{d^D \ell}{i\pi^{D/2}} \frac{1}{\ell^2 [2\ell \cdot p_1] (\ell + q_1)^2 (\ell - q_2)^2 [-2\ell \cdot p_2]} , \quad \ell^\mu \ll \bar{m}_i$$

$$G^{\text{hard}} \sim e^{\epsilon\gamma_E} \int \frac{d^D \ell}{i\pi^{D/2}} \frac{1}{[\ell^2]^3 [(\ell + p_1)^2 - m_1^2] [(\ell - p_2)^2 + m_2^2]} , \quad \ell^\mu \sim \bar{m}_i$$

Contributions polynomial in  $q_i$

# Waveform: UV Divergence



$$\mathcal{I}_{\text{box}} = \int \frac{d^D \ell}{(2\pi)^D} \frac{1}{\ell^2 [(\ell + p_1 - k)^2 - m_1^2] (\ell - q_2)^2 [(\ell - p_2)^2 - m_2^2]}$$

Eikonal expansion of second massive propagator

$$\frac{1}{(\ell - p_2)^2 - m_2^2 + i\epsilon} = \frac{1}{-2\bar{m}_2 u_2 \cdot \ell + i\epsilon} - \frac{\ell \cdot (\ell - q_2)}{[-2\bar{m}_2 u_2 \cdot \ell + i\epsilon]^2} + \dots$$

$\sim \ell^{-1}$                      $\sim \ell^0$

$\ell \rightarrow \infty$

