





Status and prospects of the PTOLEMY project

M Messina on behalf of the PTOLEMY collaboration

Marcello Messina INFN-LNGS, LHC Days in Split 2024, Hvar

OUTLINE

• Short physics introduction

• PTOLEMY detector concept

Conclusion and Outlook

The Gold-mine of Cosmologist

CMB: The oldest electromagnetic radiation in the universe



- Universe is **expanding**: Hubble's law: v = H₀D (~70 km/s/Mpc), 1919.
- Cosmic microwave background, Penzias & Wilson, 1964
- Abundance of **primordial elements**: ⁴He, ²H, ⁷Li (?)
- Galaxies morphology and stars populations in time
- Primordial gas cloud (without heavy elements), 2011

The Big Bang



V decoupling The present Universe

emerges from an Ultra-dense and high temperature initial state

> Time of decoupling: 1 second neutron/proton ratio (a) start of nucleosynthesis Temperature: $T_{\nu}=1.95$ K Number density: $n_{\nu}=112/cm^{3}$ Velocity distribution: $<v_{\nu}> ~T_{\nu}/m_{\nu}$

NEUTRINO FEATURES

• What we do know about neutrinos:

they are massive well measured Δm_i^2 cosmic neutrino background should be out there

• What we <u>don't know</u> about neutrinos: absolute mass scale (50n $(m_v < 0.45 eV)$ [KATRIN – recent publication arXiv:2406.13516v1

 $\begin{array}{ll} \text{mass ordering} \\ \textbf{(50meV < m_{light} \simeq m_e orm_{\tau})} \\ \text{From Cosmology several} \\ \text{limits at 95 \% CL on } \boldsymbol{\Sigma} \text{m}_{\boldsymbol{V}} \\ \text{from 0.56 to 0.11 eV} \end{array}$

cosmic neutrino background yet to be seen

PTOLEMY - RELIC NEUTRINO DETECTION STRATEGY



Energy E [eV]

Signature of CvB

0



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INDUCED BETA DECAY

HYSICAL REVIEW

VOLUME 128, NUMBER 3

NOVEMBER 1, 1962

Universal Neutrino Degeneracy

STEVEN WEINBERG* Imperial College of Science and Technology, London, England (Received March 22, 1962)

Modern cosmological theories imply that the universe is filled with a shallow degenerate Fermi sea of neutrinos. In the steady state and oscillating models (and perhaps also the "big bang" theories) it can be shown rigorously that the proportion of filled neutrino levels (plus the proportion of filled antineutrino levels) is precisely one up to a finite Fermi energy E_F . The proof takes into account both absorption and the repressive effects of already filled levels on neutrino emission. Experiment shows that $E_F \leq 200$ eV for antineutrinos and $E_F \leq 1000$ eV for neutrinos. The degenerate neutrinos could be observed (if $E_F > 10$ eV) by looking for apparent violations of energy conservation in β^- decay. In the steady state and evolutionary cosmologies E_F is much too low to ever be observed, but in the oscillating cosmologies $E_F \simeq 5R_c$ MeV, where R_c is the minimum radius of the universe in units of its present radius; thus experiment already shows that the universe will contract by a factor over 10³, if at all. Astronomical evidence plus Einstein's field equation (without cosmological constant) require in an oscillating cosmology that $E_F < 2 \times 10^{-3}$ eV (so $R_c < 10^{-9}$) and suggest that higher energy neutrinos may represent the bulk of the energy of the universe. A model universe incorporating this idea is constructed.

Cocco, Mangano, Messina calculatedin the casde $m_{ m v} eq 0$ case in 2007



 $m_{\nu}=0$



Detailed evaluation on 2007 $JCAP \ 06 \ (2007) \ 015$ of $\sigma \times \tau$ renewed the dormant discussion on relic neutrino detection and paved the view to a possible experiment. Several authors confirmed the cross section evaluation and added informations:

> J. Phys. G: Nucl. Part. Phys. 35 025001 JCAP 08 (2014) 038

PHYSICS OUTCOME

- Direct neutrino mass
- CNB
- Sterile Neutrinos
- Dirac/Majorana

DETECTOR CONCEPT

PonTecorvo / Observatory for Light Early-universe Massive-neutrino Yield





- PTOLEMY aims at using TES detectors with an envisaged resolution of $\Delta E \simeq 0.05 eV$
- However:

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- 1. TES' perform best with energies O(10 eV) need to slow the electrons down
- 2. TES' are slow response detectors need to reduce the number of electrons coming from β -decay

PTOLEMY FILTER

Exponential decaying magnetic field



$$B_x = B_0 \cos\left(\frac{x}{\lambda}\right) e^{-z/\lambda} ,$$

$$B_y = 0 ,$$

$$B_z = -B_0 \sin\left(\frac{x}{\lambda}\right) e^{-z/\lambda} .$$

$$E_x = 0,$$

$$E_y = E_0 \cos\left(\frac{y}{\lambda}\right) e^{-z/\lambda},$$

$$E_z = -E_0 \sin\left(\frac{y}{\lambda}\right) e^{-z/\lambda}.$$

$$V_{E\times B}^{y}(z)|_{x,y=0} = \frac{E\times B}{B_x^2} = \frac{E_z B_x \hat{y}}{B_x^2} = \frac{E_z}{B_x} \hat{y}$$

$$\boldsymbol{V}_{\nabla B-C} = \frac{1}{2}m(\boldsymbol{v}_{\perp}^{2} + 2\boldsymbol{v}_{\parallel}^{2})\frac{\boldsymbol{B} \times \boldsymbol{\nabla}_{\perp} \boldsymbol{B}}{q\boldsymbol{B}^{3}} = (T_{\perp} + 2T_{\parallel})\frac{\boldsymbol{B} \times \boldsymbol{\nabla}_{\perp} \boldsymbol{B}}{q\boldsymbol{B}^{3}}$$

$$\frac{dT_{\perp}}{dt} = \frac{\mu}{B^2} \boldsymbol{E} \cdot (\boldsymbol{\nabla} B \times \boldsymbol{B})$$

• Exponential decaying electric field





Prog.Part.Nucl.Phys. 106 (2019) 120-131

JINST 17 (2022) 05, P05021

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 A new electromagnetic filter idea based on RF detection and dynamic E setting



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 A new electromagnetic filter idea based on RF detection and dynamic E setting

∆V known to 1 ppm precision



PTOLEMY: ATOMICT TARGET

PTOLEMY has to deal with large instrumented mass.
 Distribute atomic tritium on a solid state substrate (e.g.







Loading chamber Tritium loading chamber



T loading chamber designed and built in Rome in the framework of the PTOLEMY R&D. Loading technique tested up to 90 % loading efficiency measured with H and D.

Quadrupole Mass Spectrometer: SRS RGA 100

how to estimate the H (or D):C upload → directly from a quantification of the sp3 bond spectroscopic signal from the XPS C 1s core level:





Abdelnabi et alii, Nanomat. 11, 130 (2021)

1 MG Betti et al, Dielectric response and excitations of hydrogenated free-standing graphene, Carbon Trends, September 2023, 100274;

2 MG Betti et al, Homogeneous Spatial Distribution of Deuterium Chemisorbed on Free-Standing Graphene, Nanomaterials 2022, 12(15), 2613;

A SPECIAL MAGNET

BEING REBUILT IN A LARGER SIZE AND WILL BE INSTALLED AT THE LNGS KEY ELEMENTS TO REALIZE THE TRANSVERSE DRIFT AND THE DEMONSTRATOR OF THE PTOLEMY PROJECT



Measured B field shape as expected CST 1.0 Data 0.2 0.8 E 8 0.1 Ξ0.6 ^م_{0.4} 0.2 0.0 0.0 0.0 0.2 -0.4 -0.20.40.6-0.4 -0.20.0 0.2 0.40.6 *Z* [m] Z [m]

Construction ASG/Suprasys consortium of a SC dipole with special attention to the fringe fleld

Saddle point key features of the field map

nate system 🔕 🞅 🍞 😭 🐼 🛷 🚰 🛍 🐔 🕼 🤭 🕐 🛃 🔹

Detailed simulation of the filter



More than a single geometry under study

easy to accommodate micrograms

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Preliminary estimation of sensitivity curves



CALORIMETER



Now: 0.11 eV @ 0.8 eV and 106 mK and 10x10 μm² TiAuTi 90nm [Ti(45nm) Au(45nm)] ^(τ ~137 ns)

Design Goal (PTOLEMY): $\Delta E_{FWHM} = 0.05 \text{ eV} @ 10 \text{ eV}$

translates to $\Delta E \propto E^{\alpha}$ ($\alpha \leq 1/3$) $\Delta E_{FWHM} = 0.022 \text{ eV} @ 0.8 \text{ eV}$

Based on the expertise of the INRiM an important results have been achieved on electron measurement with TES. Key elements of the measurements: performing TES and new e-source based on nanostructures





The PTOLEMY Collaboration Institution Board recently constituted



CONCLUSION

- PTOLEMY aims at eventually detect cosmic neutrino background on a long term time scale
- The detector prototype will be ready at LNGS by the next year
- Prototype baseline option is: T embedded on graphene; New concept EM filter; electron energy resolution measured in several steps (MCP/SDD).
 Ultimately operating TES with sub-eV energy resolution.
- Ultimate goals of the Demonstrator: instrumented mass ~ hundreds of μg, energy resolution 50-100 meV, T storage solution will come from optimization of atomic T support structure. Time scale 5 years.
- "Intermediate" physics results of the Demonstrator: neutrino mass measurements (or limit) at level or better than present limits.