

# Gravitational waves from cosmic strings with friction

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# Plan

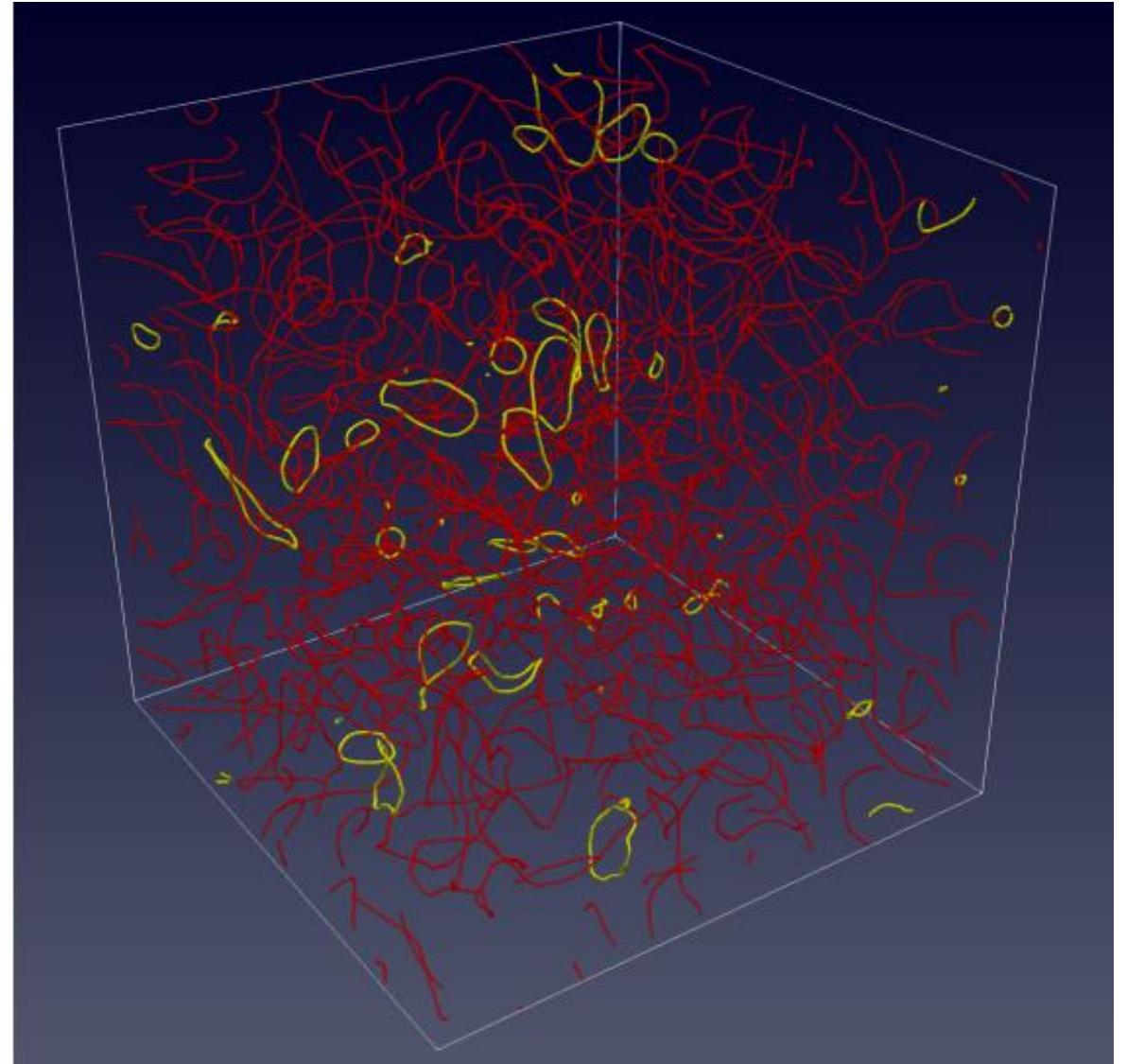
1. Production of cosmic strings
2. Evolution of a cosmic string network
3. Cosmic string loops
4. Stochastic Gravitational Wave Background (SGWB) calculation
5. SGWB with friction
6. Detectability

# Production of a cosmic string network

During phase transitions as a result of spontaneous symmetry breaking

- GUT strings
- Electroweak strings
- Axion strings
- ...

Cosmic superstrings

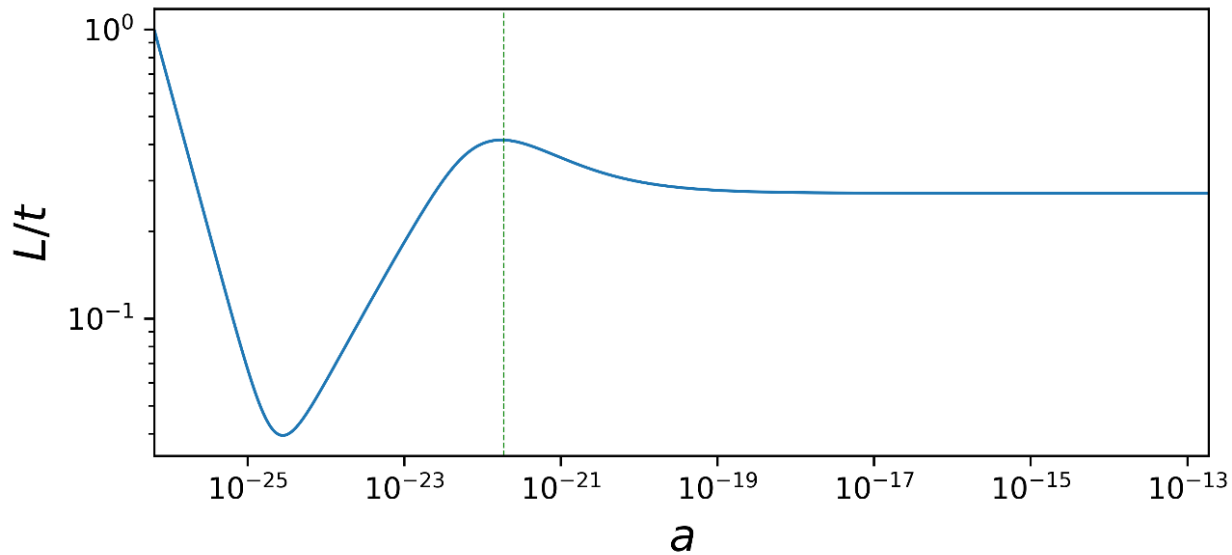


Naoya Kitajima, Kazunori Nakayama; 2212.13573

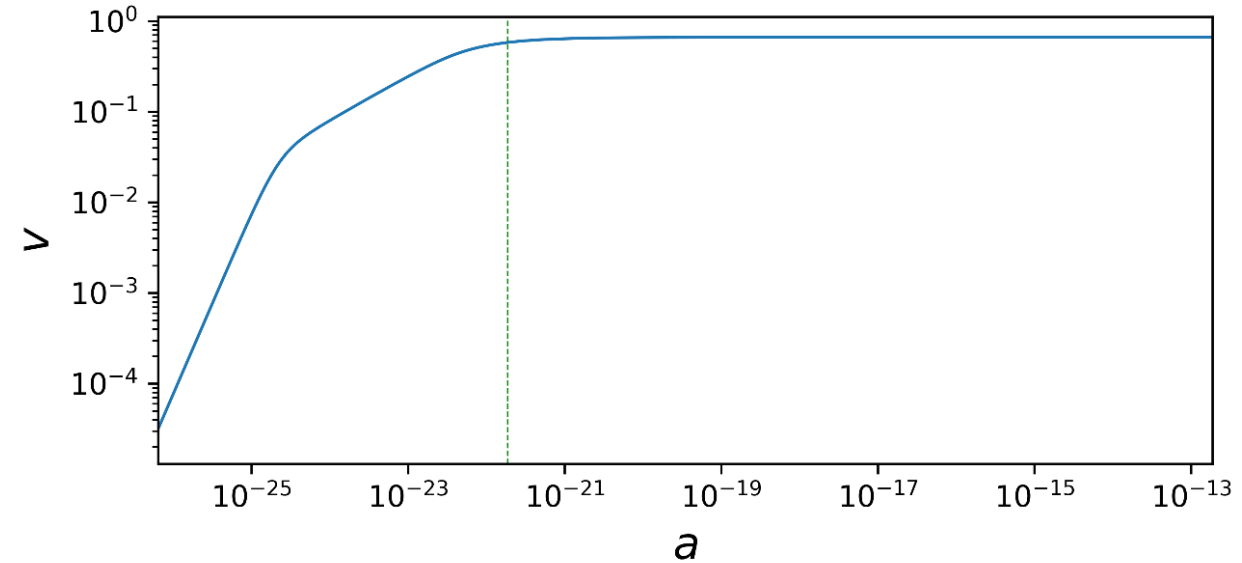
# The evolution of cosmic string networks

## Scaling regimes

$L$  - the characteristic lengthscale



$\bar{v}$  - the Root-Mean-Squared (RMS) velocity



Stretching

$$L = L_c \left( \frac{t}{t_c} \right)^{1/2}, \quad \bar{v} = \bar{v}_c \left( \frac{t}{t_c} \right)$$

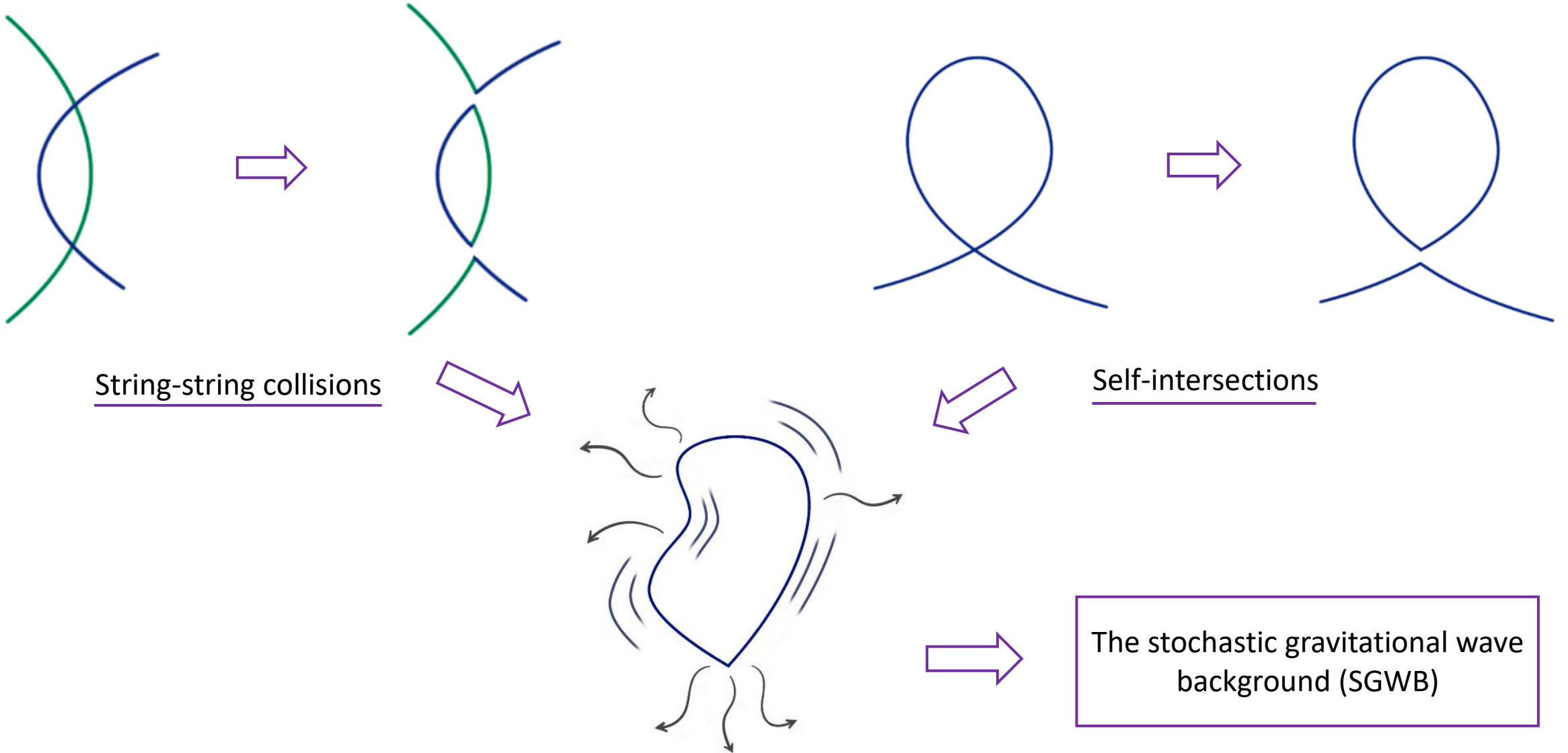
Kibble

$$\frac{L}{t_c} \sim \left( \frac{t}{t_c} \right)^{5/4}, \quad \bar{v} \sim \left( \frac{t}{t_c} \right)^{1/4}$$

Linear

$$L \sim t, \quad \bar{v} = \text{const}$$

# Cosmic string loops



# The SGWB. Calculation

1) Calculate the evolution of the network:  $L(t)$   $\bar{v}(t)$



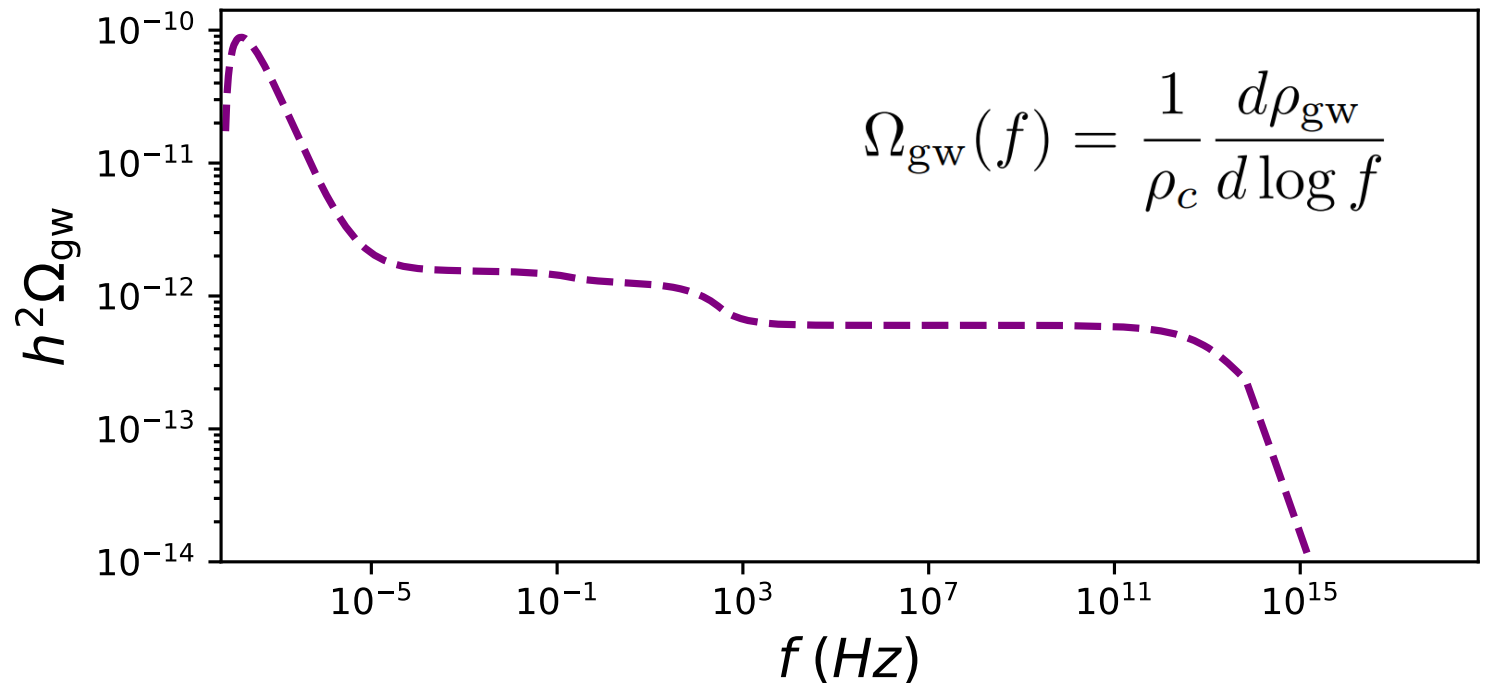
Find the loop production function:

$f(l, t)dl$  = the number density of loops with length between  $l$  and  $l + dl$  produced per unit time.

2) Evolution of the cosmic string loop:

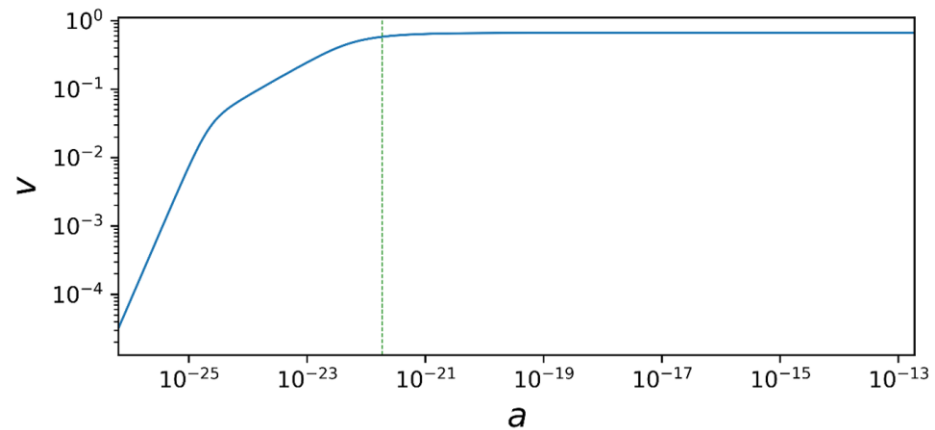
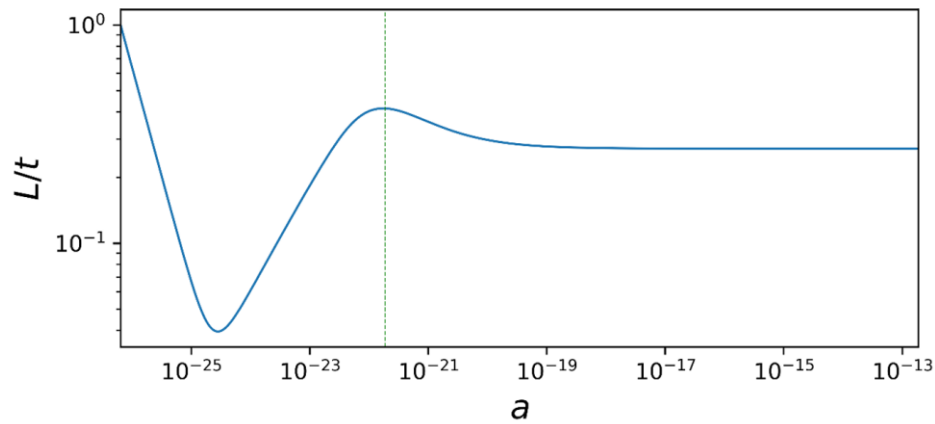
$$l = l_b - \Gamma G\mu (t - t_b)$$

$$\underline{l_b = \alpha L(t_b)}$$



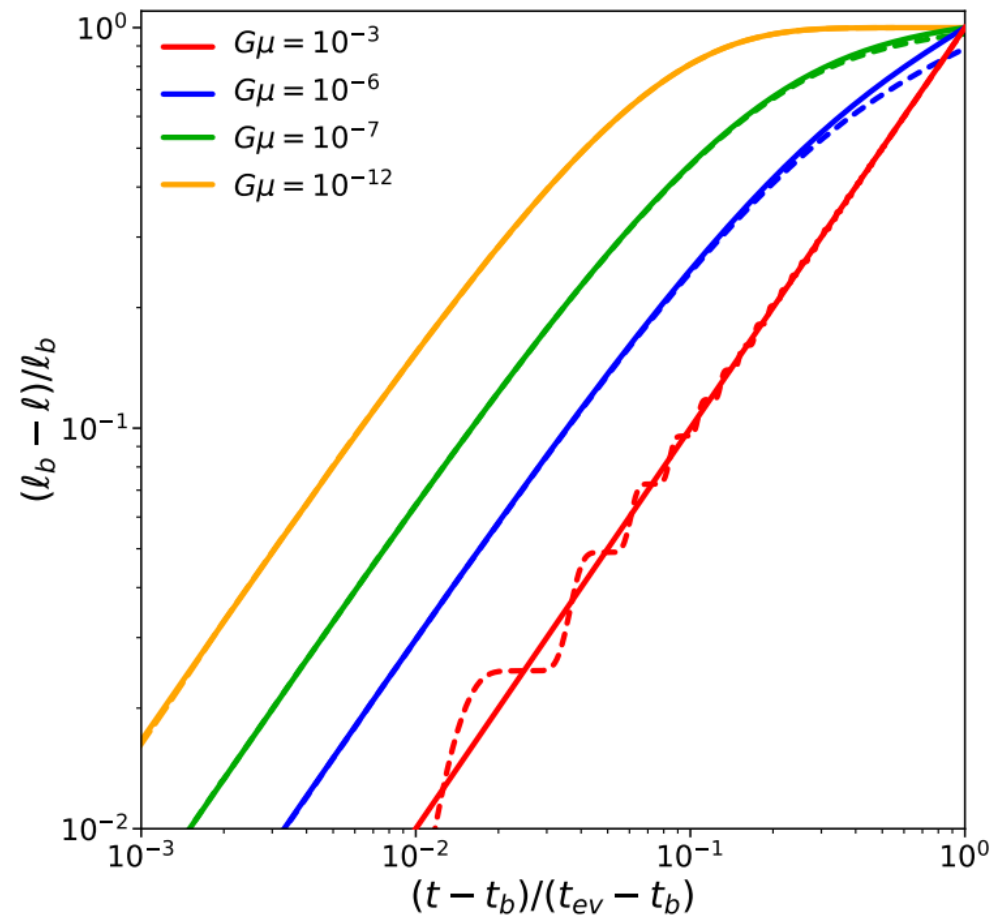
# The SGWB with friction. Calculation

Evolution of the cosmic string network:



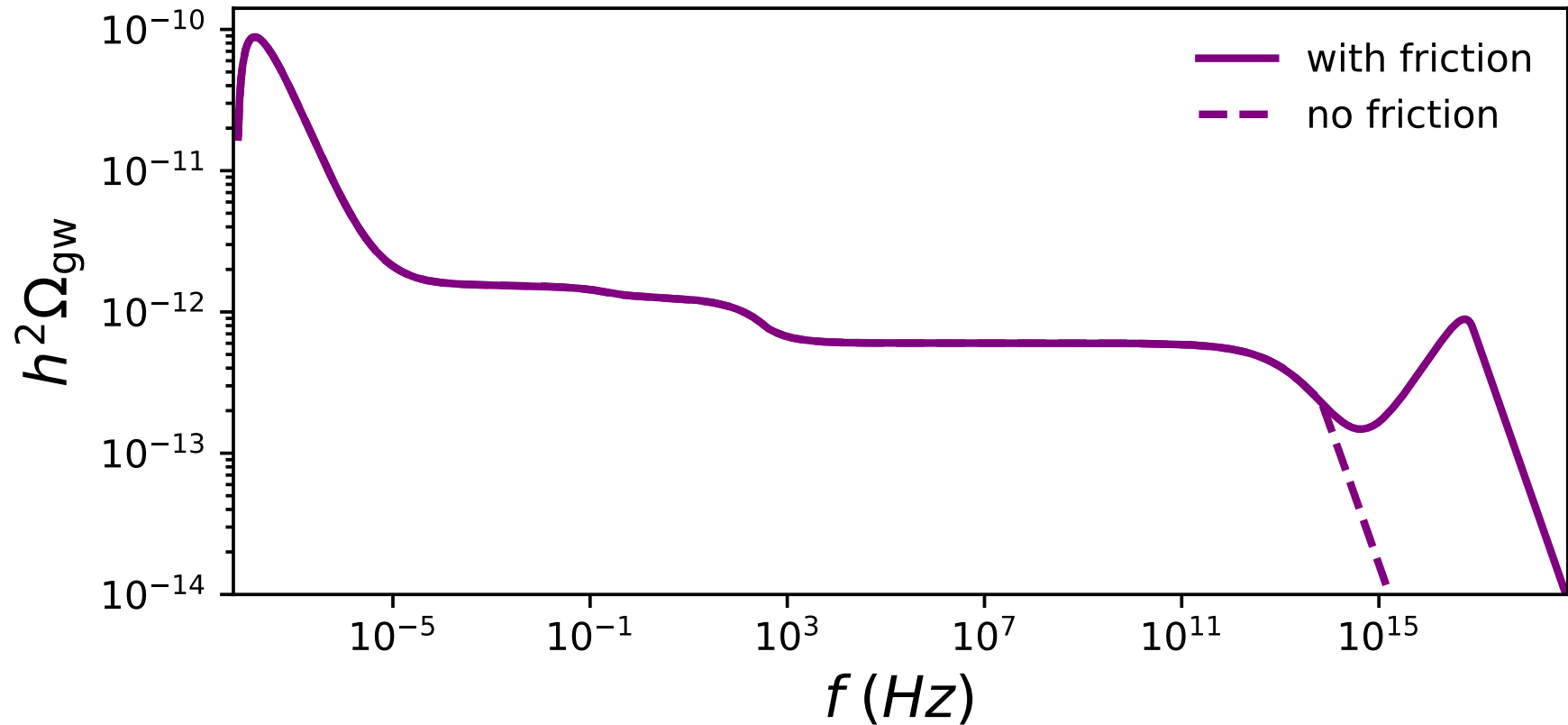
Evolution of the cosmic string loop:

$$\ell = \ell_b \exp \left[ t_f^{1/2} \left( t^{-1/2} - t_b^{-1/2} \right) \right] - \Gamma G\mu (t - t_b), \quad t_f = \frac{t_{pl}}{\chi^3} \left( \frac{\beta}{G\mu} \right)^2$$



# The SGWB with friction. Full spectrum.

$$G\mu = 10^{-10}, \alpha = 10^{-9}, \beta = 1, L_c = t_c$$



Cut-offs on the size scales

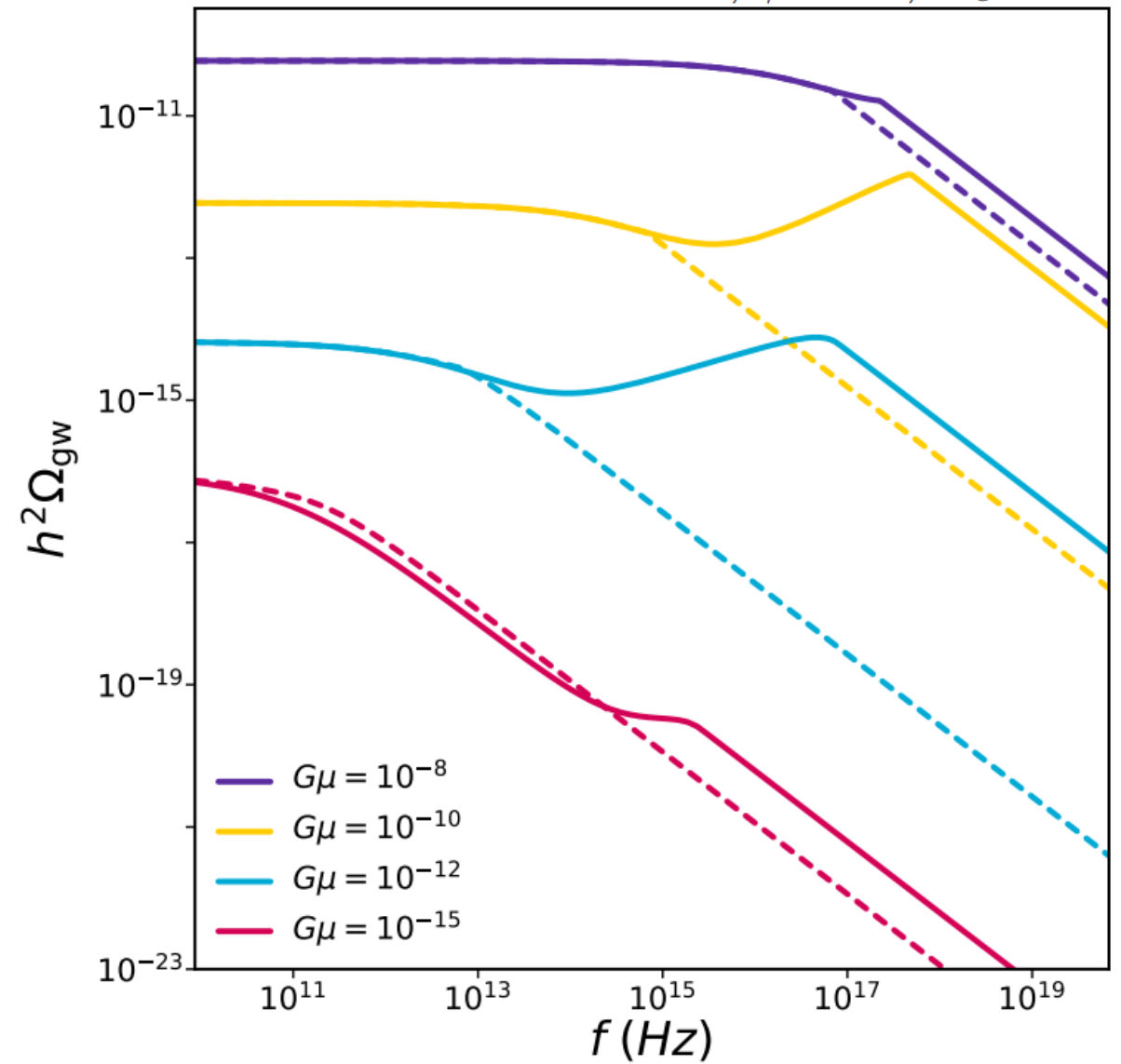
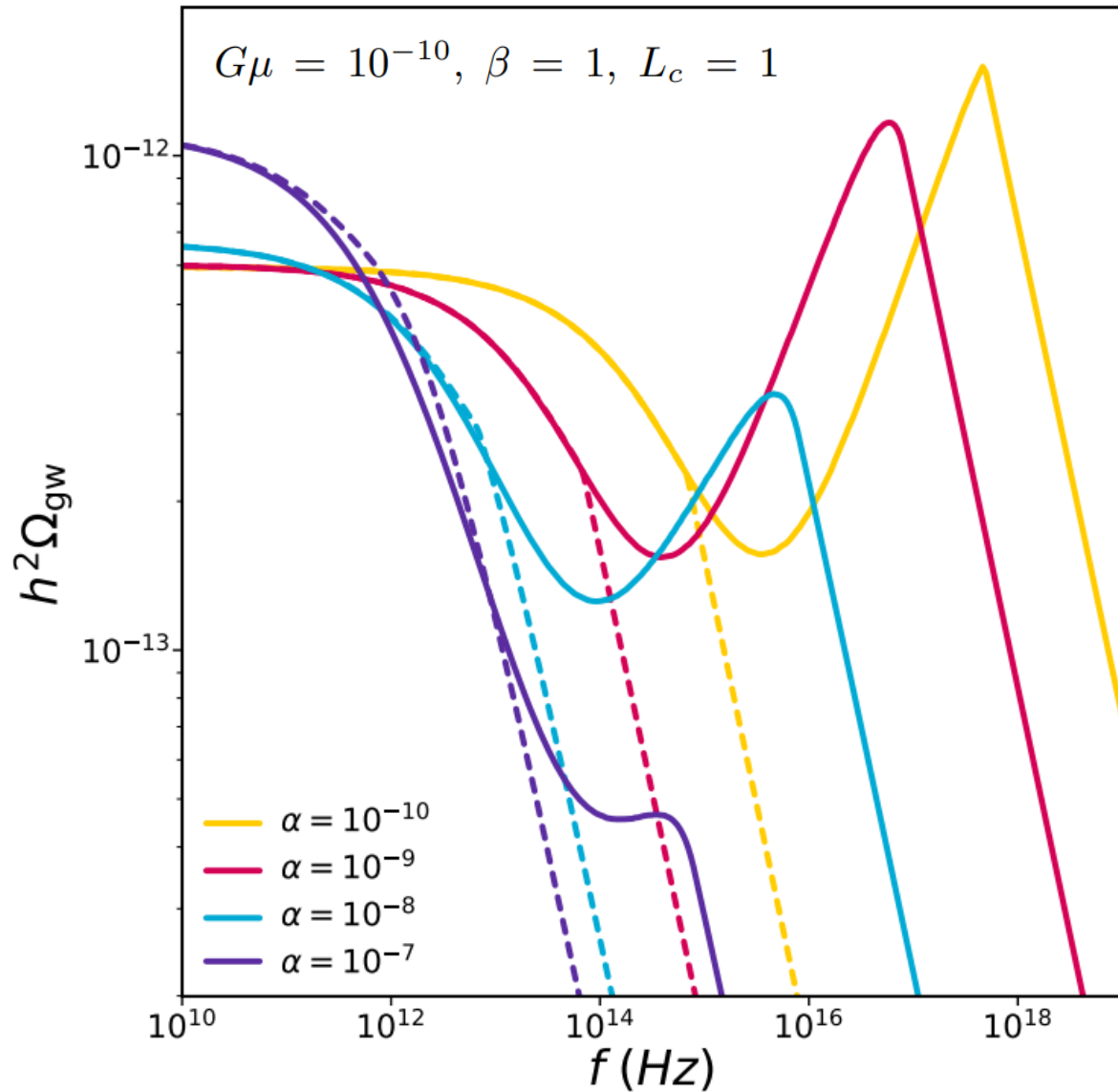
Length of loops  $> l_{pl}$

Gravitational backreaction scale  $> l_{pl}$

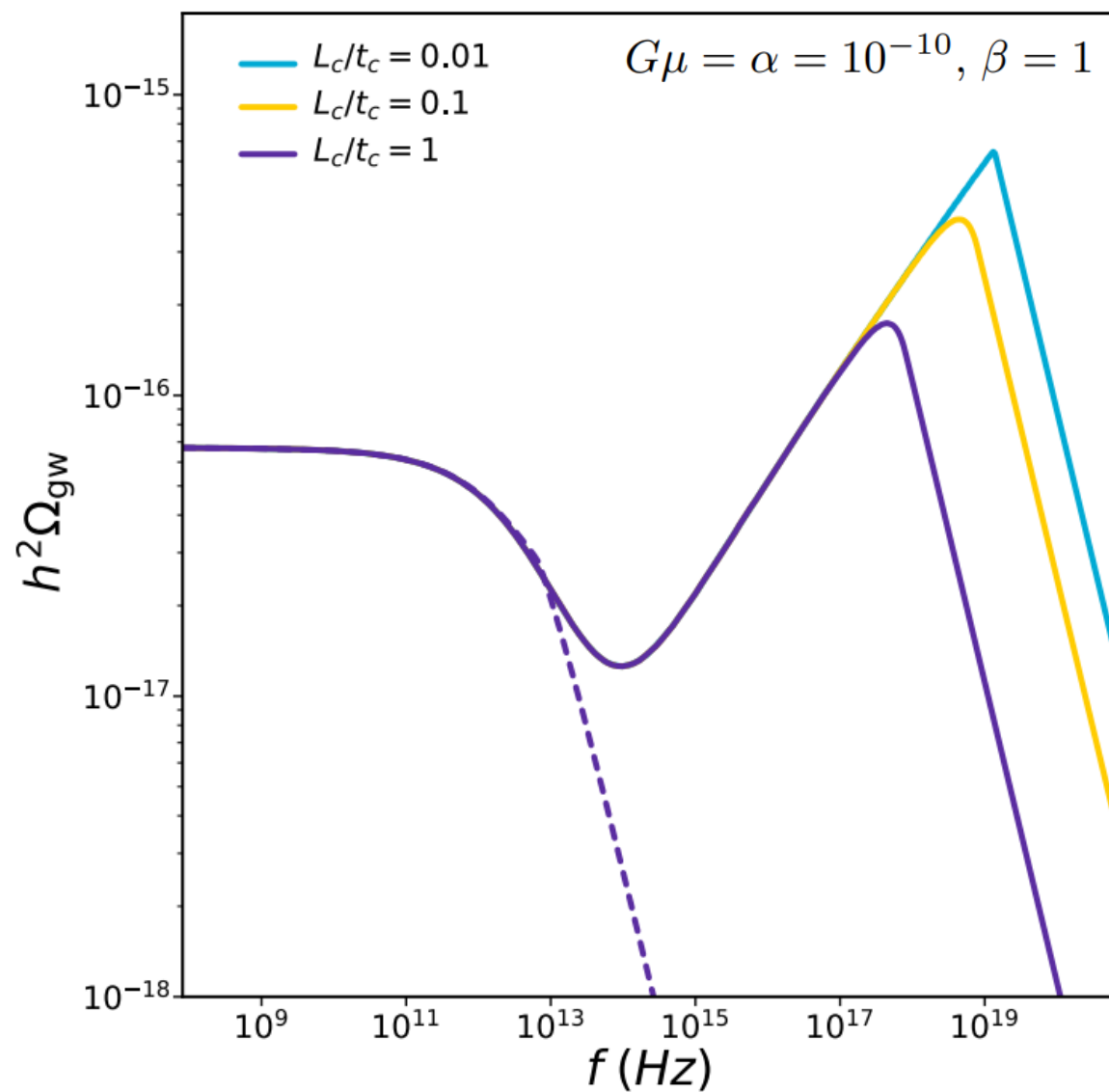
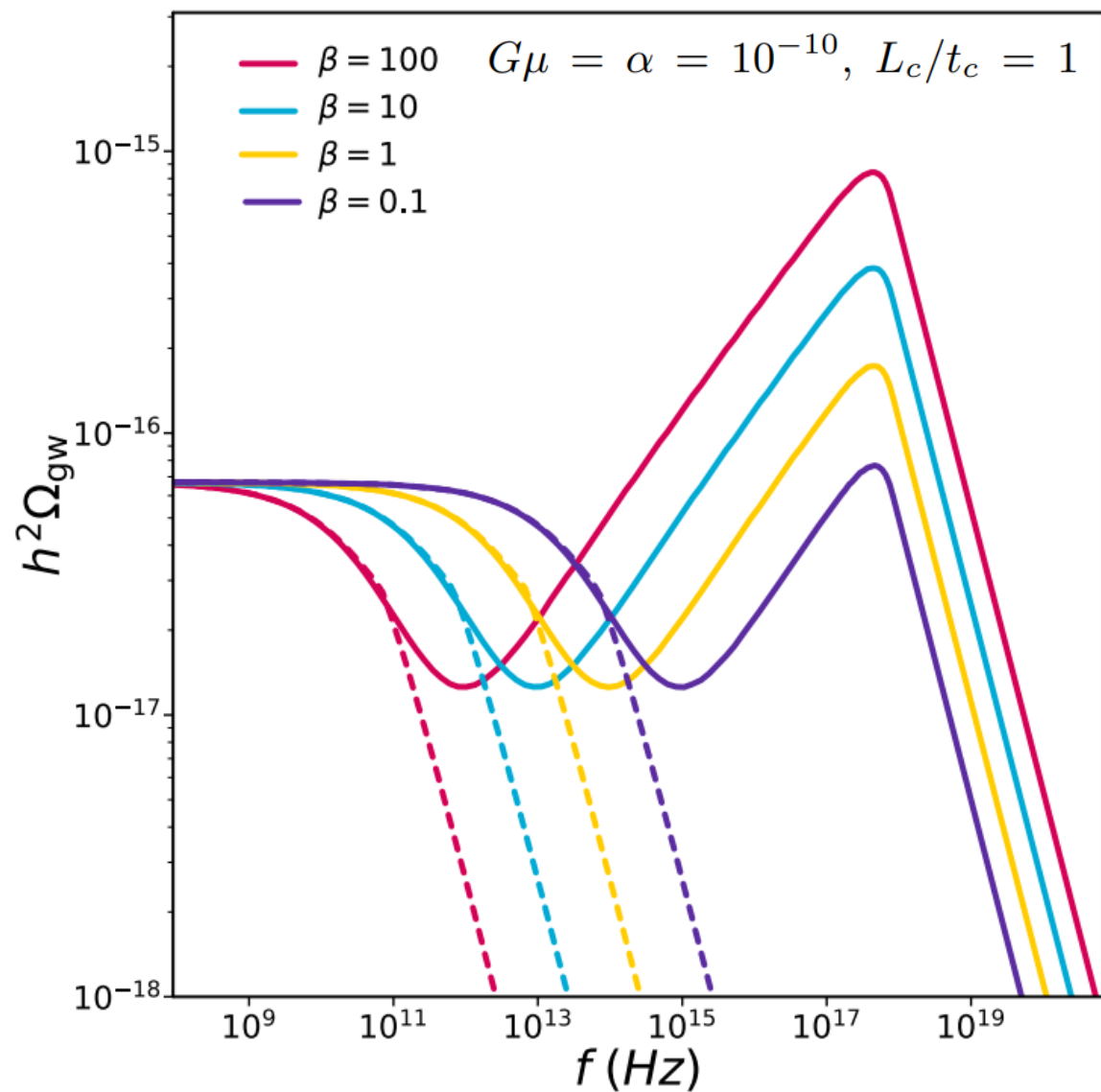


# The SGWB with friction. Parameters variation.

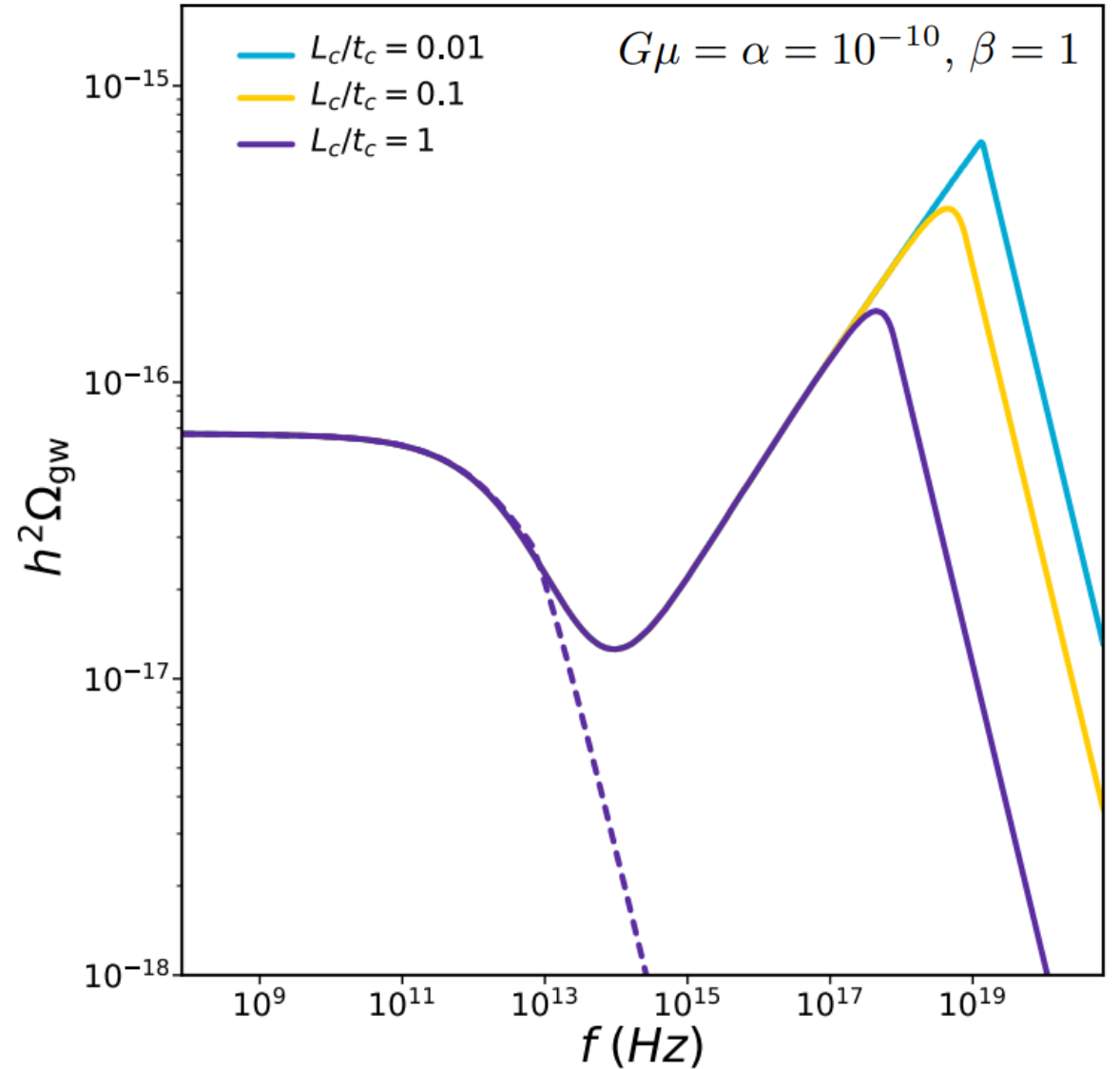
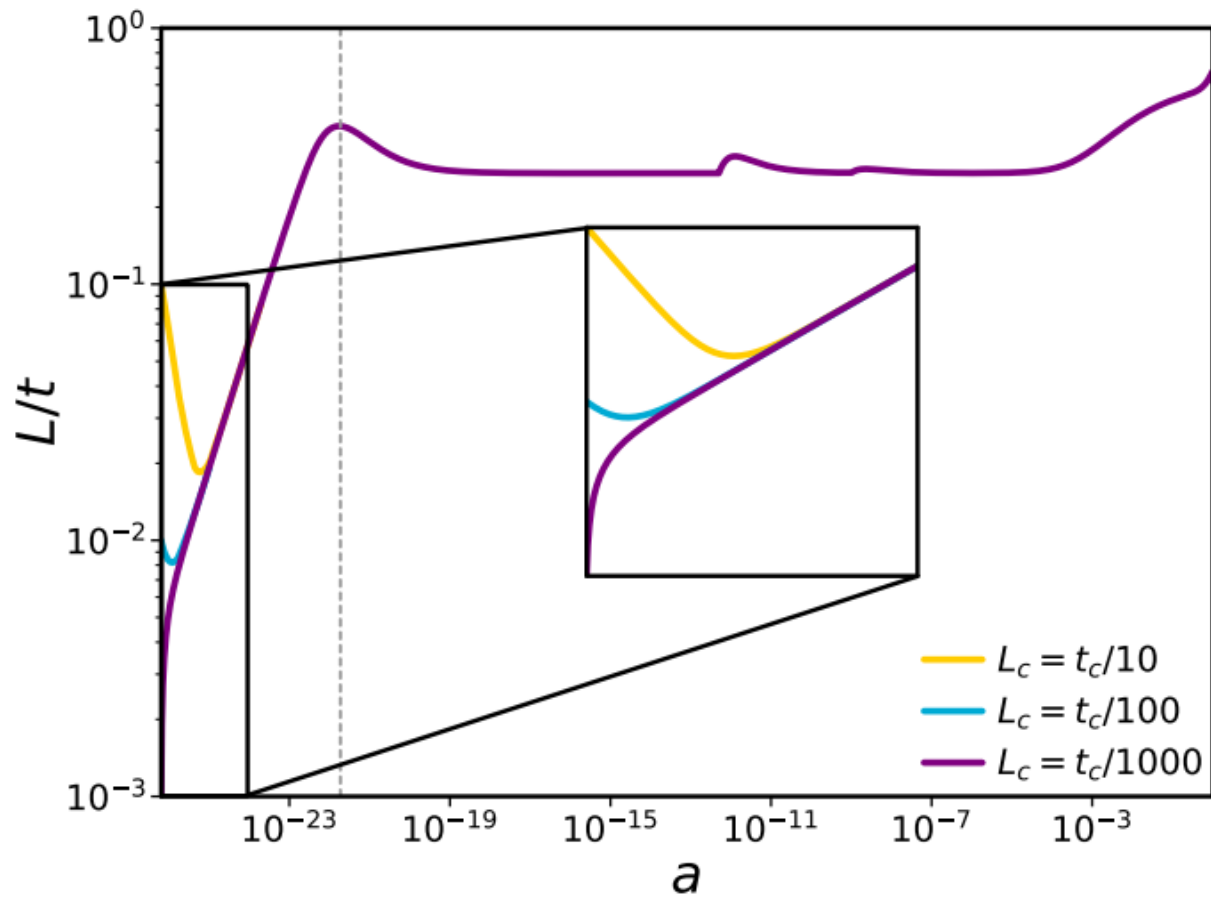
$$\alpha = 10^{-10}, \beta = 1, L_c = 1$$



# The SGWB with friction. Parameters variation.



# The SGWB with friction. Parameters variation.



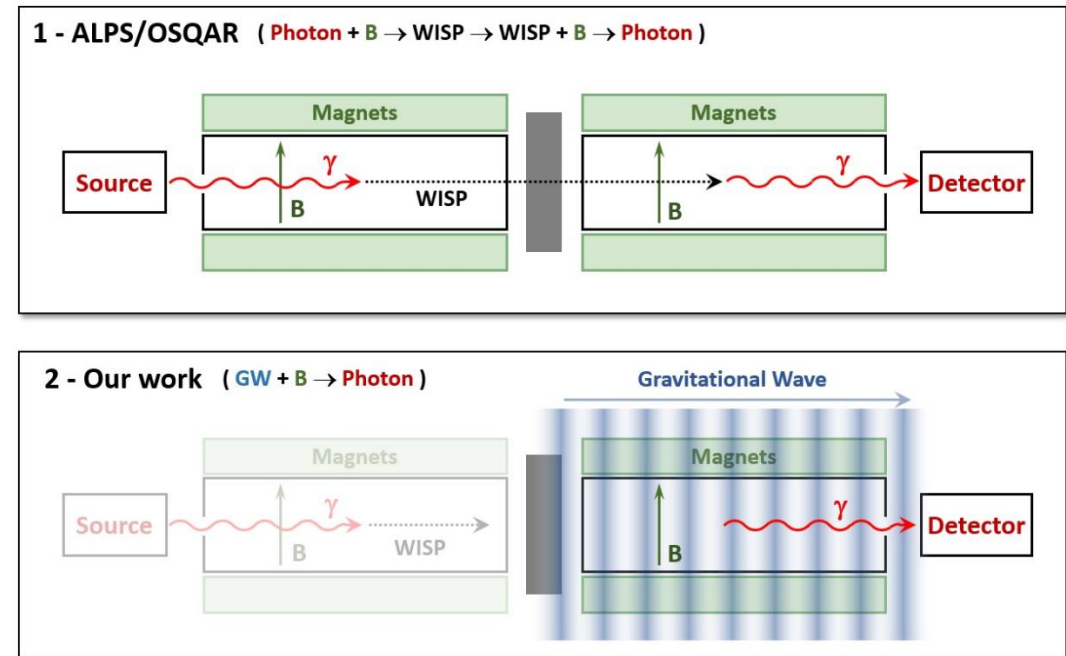
# Detectability

## HF GWs technical concepts:

- Laser interferometers
- Optically levitated sensors
- Magnetic conversion
- Resonant polarization rotation
- Microwave cavities
- Graviton-magnon resonance

Recent review arXiv:2011.12414

## Magnetic conversion



A. Ejlli et al.; 1908.00232

Magnetic fields of pulsars, planetary magnetospheres,  
galaxy clusters, artificial magnetic fields

# Conclusions

- There is a signature of friction in the UHF range of the spectrum
- It is strongly dependent on the particular particle physics theory of the early universe
- It's detection is challenging. There is a strong interest of the scientific community in this range of the spectrum.

The signature of friction opens a wider window for probing the physics of the very early universe, inaccessible by considering just frictionless period of evolution.