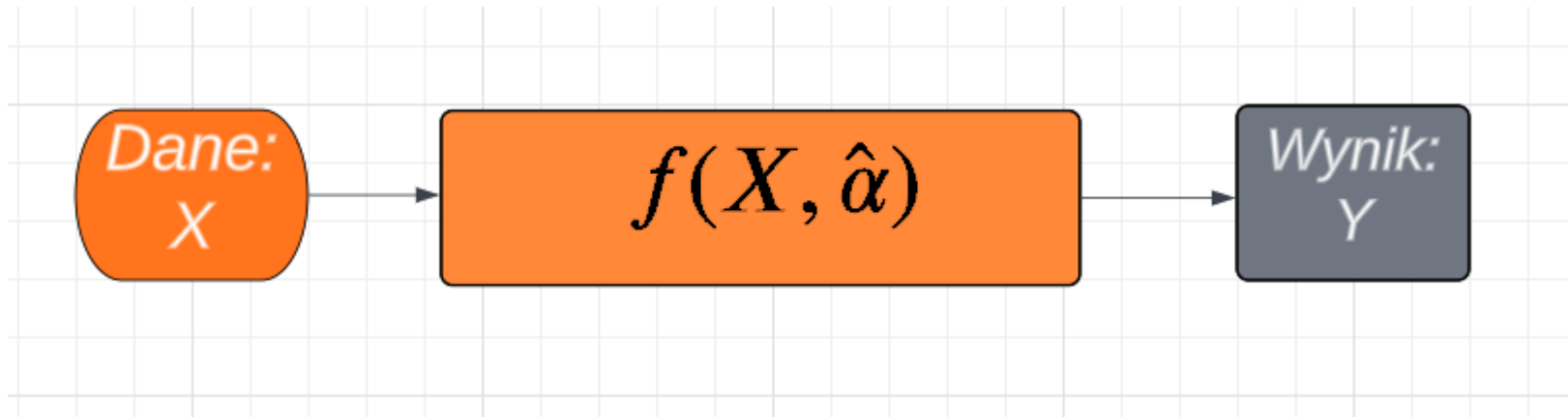


# **Machine learning in physics**

Artur Kalinowski  
University of Warsaw

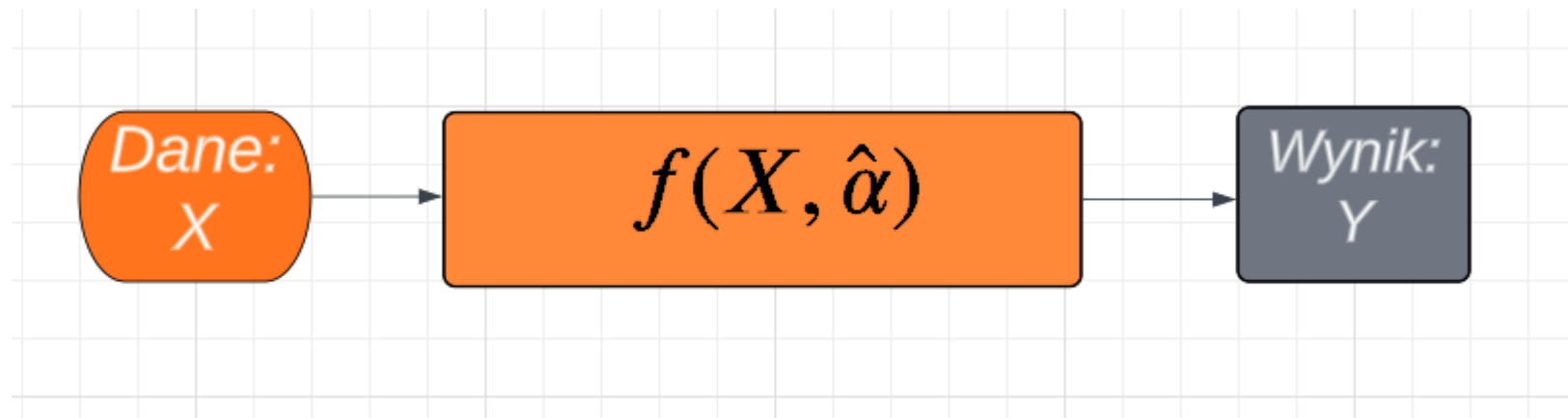
**X:** n dimensional input data

**Y:** k dimensional output data



**X**: n dimensional input data

**Y**: k dimensional output data



Function fitting:

- **$n \sim 2$ ,  $k \sim 1$**
- **basic function given *explicite***
- **expansion coefficients *could* be interpretable**

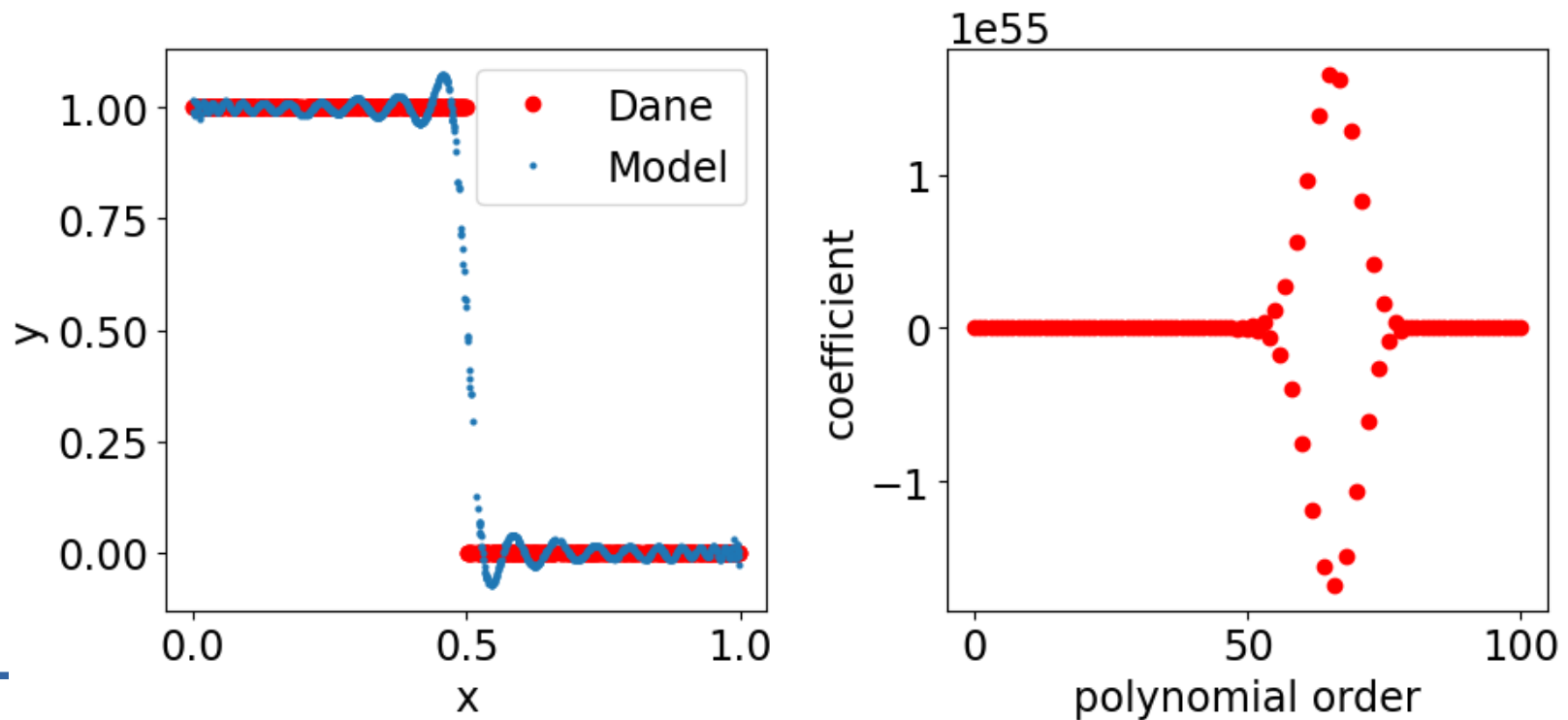
*Taylor theorem (J. Gregory, 1671):*

Every, continuous, differentiable, function  $f(x): \mathbb{R} \rightarrow \mathbb{R}$  can be approximated by a polynomial:

$$f(x, \theta) = \sum_{n=0}^{\infty} \theta_n x^n \simeq \theta_0 + \theta_1 x + \theta_2 x^2 + \dots + \theta_n x^n$$

Coefficients  $\theta_i$  are derivatives of  $f(x)$ .

In the case of unknown function (“data”) coefficients can be found by a numerical procedure. Usually...



## Fourier theorem (1807) :

Every continuous, differentiable, and periodic function  $f(x): \mathbb{R} \rightarrow \mathbb{R}$  can be approximated in a basis of sines i cosines:

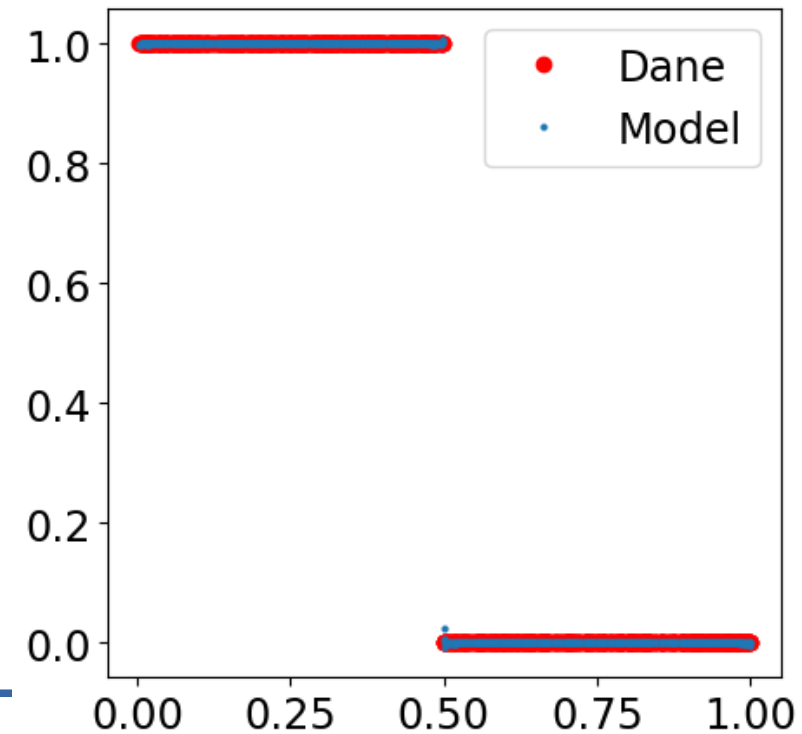
$$f(x, \theta) = \frac{\theta_{0,0}}{2} + \sum_{n=1}^{\infty} \theta_{0,n} \cos(n\omega x) + \theta_{1,n} \sin(n\omega x), \quad \omega = \frac{2\pi}{T}$$

Coefficients  $\theta_i$  can be found analytically.  
In the case of unknown function ("data") coefficients can be found by a numerical procedure.

```

1 nMax = 25000
2 omega = 2*np.pi/1.0
3
4 Y_model = np.full_like(X,0.5)
5 for n in range(1,nMax+1, 2):
6     Y_model += 2.0/np.pi*1/n*np.sin(omega*n*X)

```

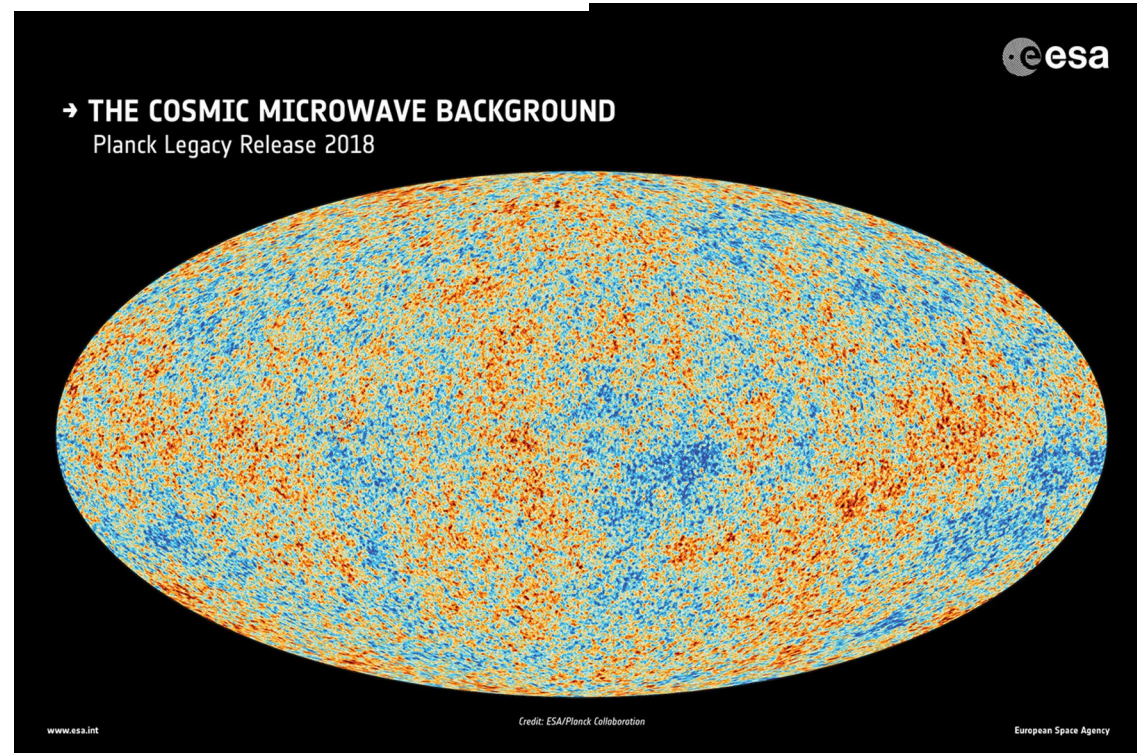


## Spherical harmonics (Laplace, 1782) :

Every, continuous, differentiable, function on sphere  $f(x): \mathbb{R}^2 \rightarrow \mathbb{R}$  can be approximated in a basis of function solving a Laplace's equation - spherical harmonics:

$$f(\theta, \varphi, w) = \sum_{n=1}^{\infty} w_{l,m} Y_m^l(\theta, \varphi)$$

Coefficients  $w$  can be found analytically. In the case of unknown function ("data") coefficients can be found by a numerical procedure.



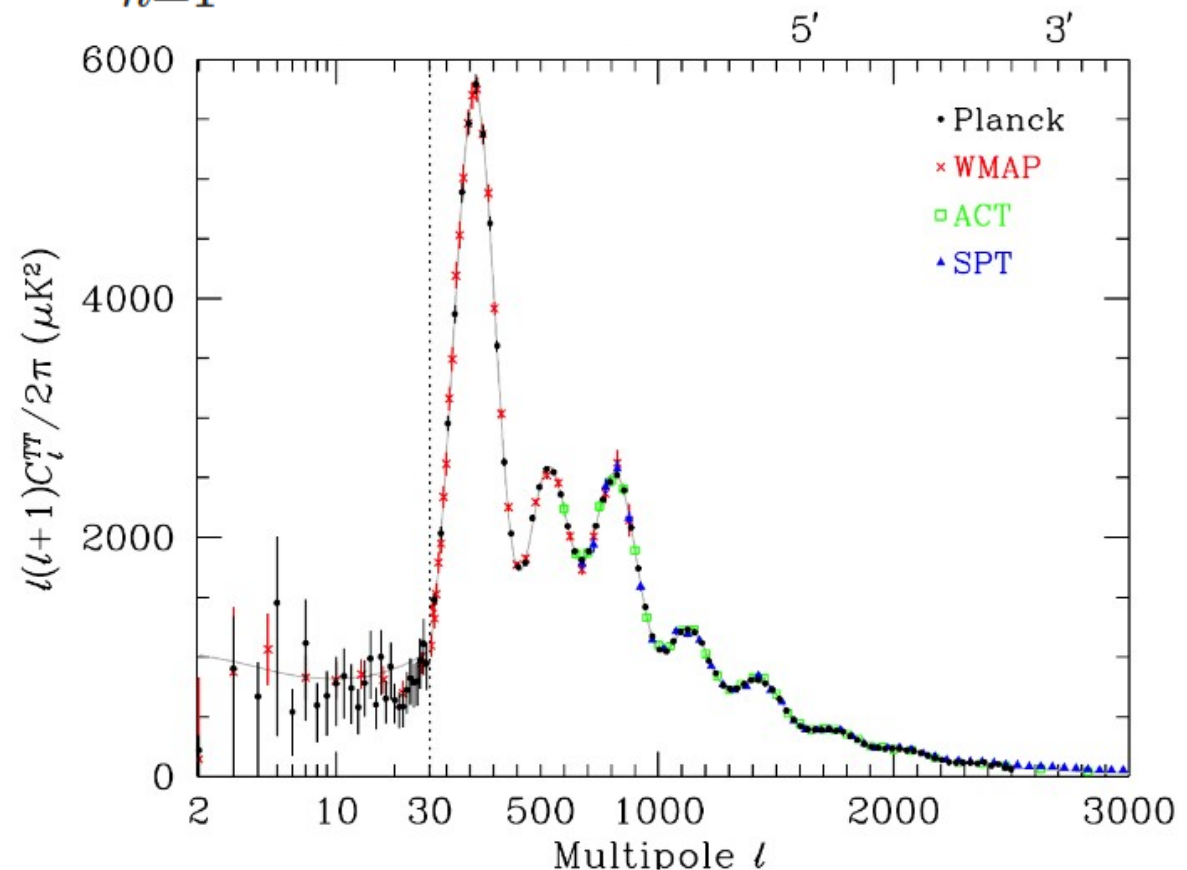
ESA/Planck Collaboration

## Spherical harmonics (Laplace, 1782) :

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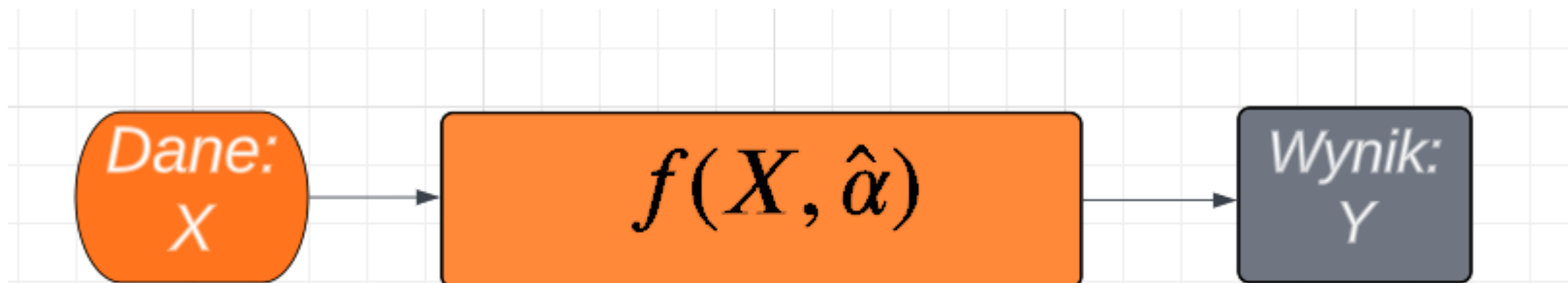
$$f(\theta, \varphi, w) = \sum_{n=1}^{\infty} w_{l,m} Y_m^l(\theta, \varphi)$$

Coefficients  $w$  can be found analytically. In the case of unknown function ("data") coefficients can be found by a numerical procedure.



**X**:  $n$  dimensional input space

**Y**:  $k$  dimensional output space



- $n \sim 10^l$ ,  $k \sim 10^m$ ,  $l, m \sim 6$
- basis functions defined *implicite* - through data flow architecture
- expansion coefficients are uninterpretable



A sigmoid function: any non polynomial function fulfilling conditions:

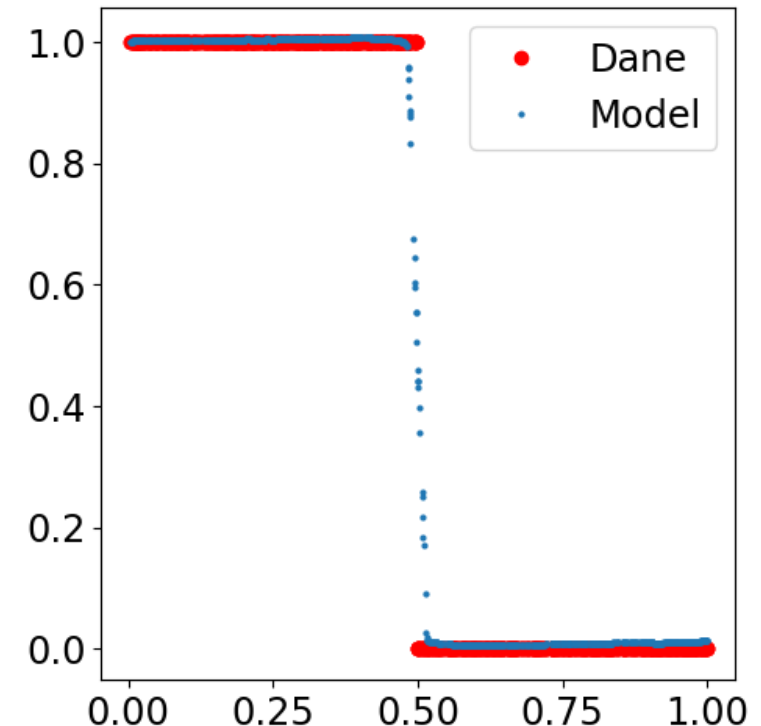
$$A(\theta, x) = A\left(\sum_{i=1}^n \theta_i x_i + b\right) \quad \lim_{x_i \rightarrow -\infty} A(x) \rightarrow 0 \quad \lim_{x_i \rightarrow +\infty} A(x) \rightarrow 1$$

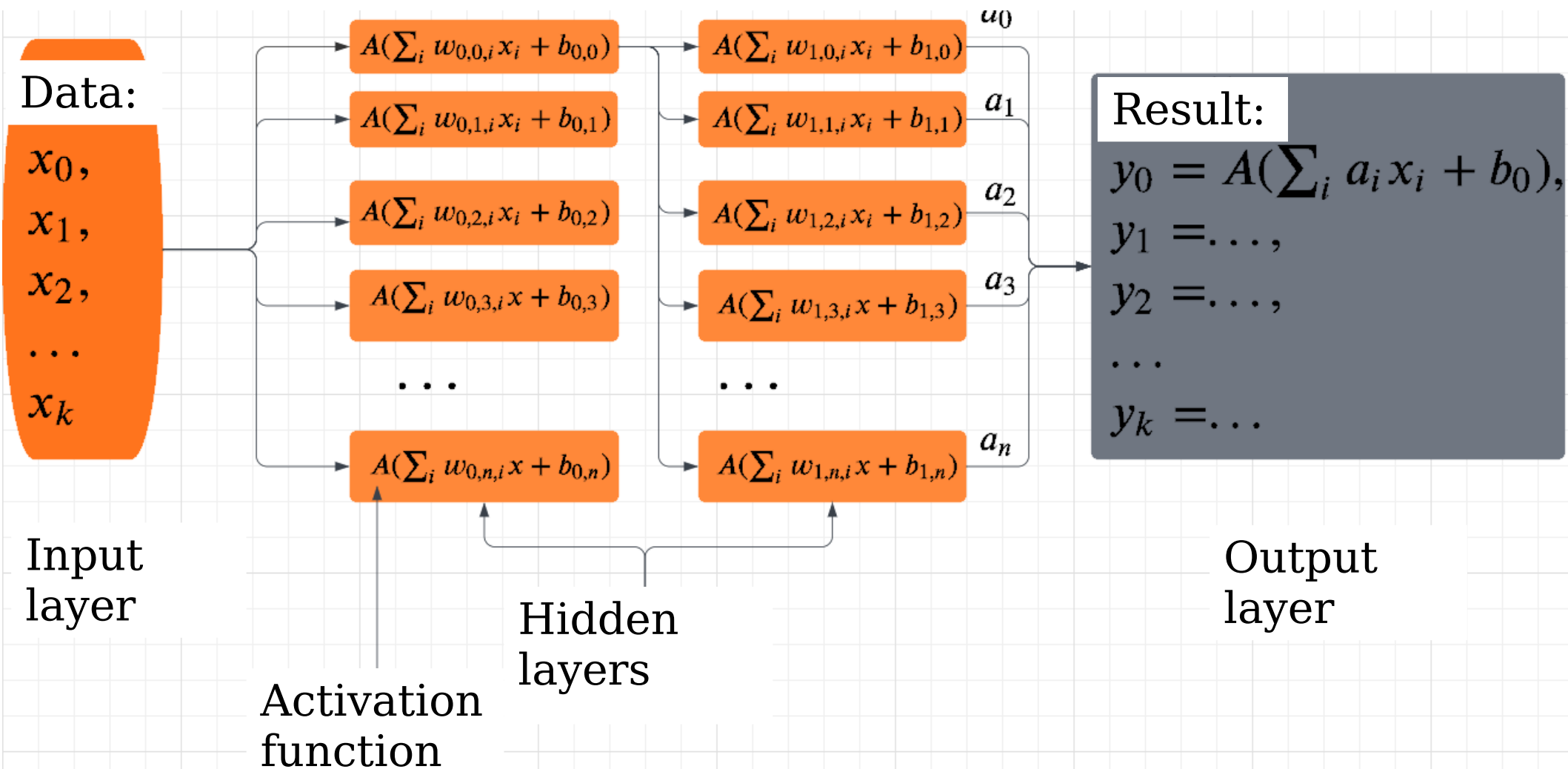
*Universal approximator theorem (Cybenko, 1989):*

Every continuous function  $f(x) \mathbb{R}^n \rightarrow \mathbb{R}$  can be approximated in basis of sigmoidal functions:

$$f(x, \theta) \simeq \sum_n w_n A(\theta_n, x)$$

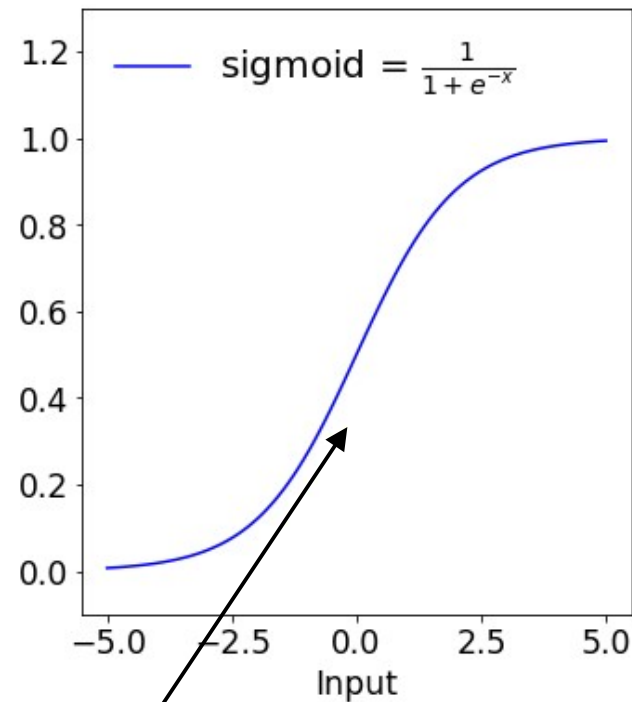
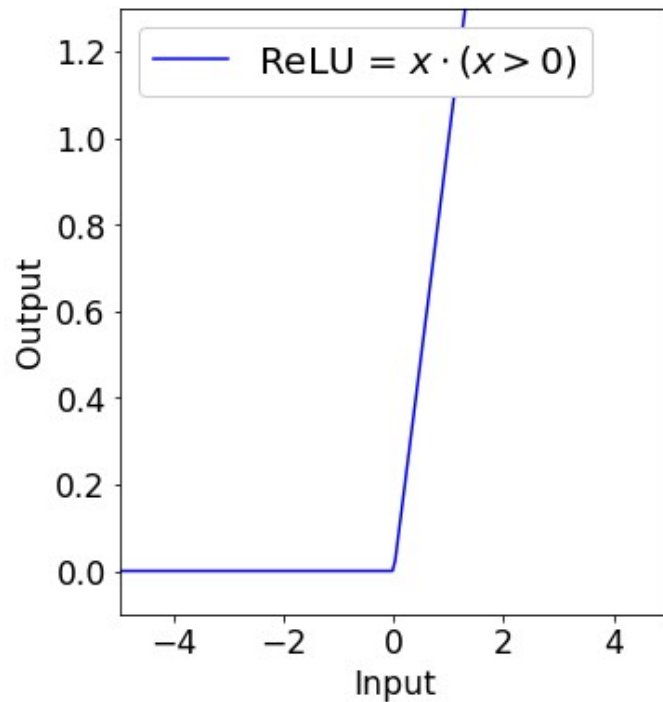
Coefficients  $\theta_i, w_i$  do not have in general an analytic form, but can be found using a numerical procedure



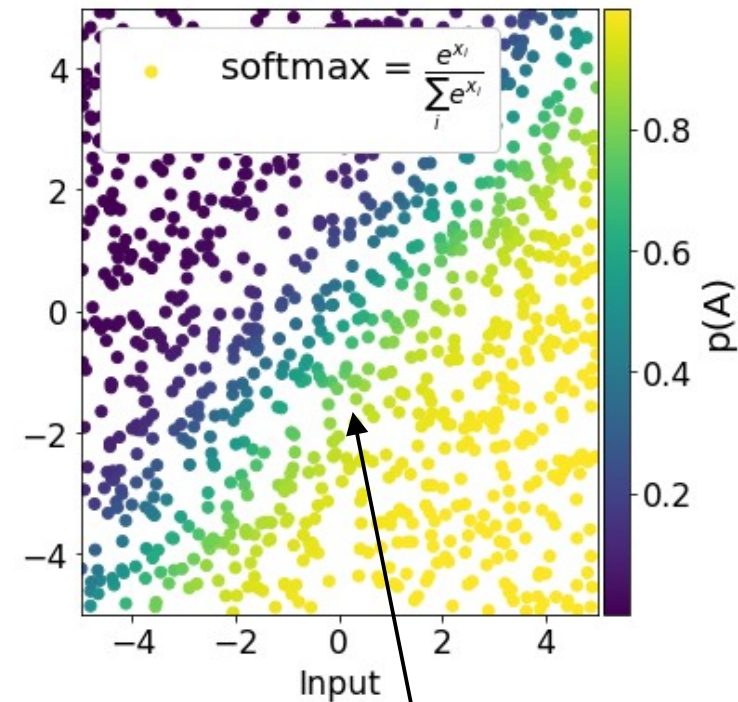


standard for hidden layers

standard for output layers  
(when probability is the target)



two classes probability



many classes probability

**The task:** superconductivity compound classification:  
Is superconductive? **YES/NO**

## Input data:

- set of compound elements,
- 22 features per element + stoichiometric coefficients

## Data sets:

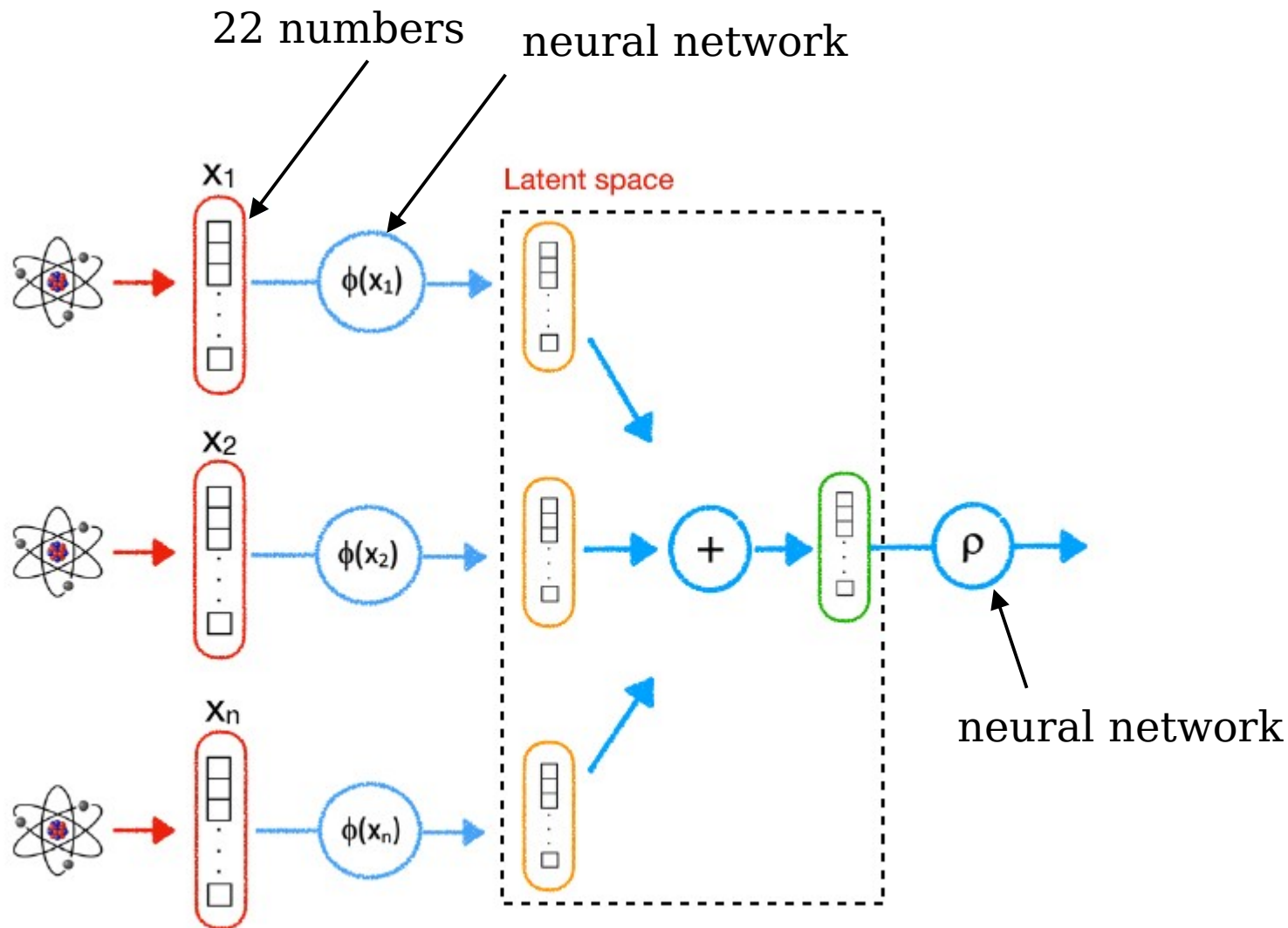
### Training and validation:

- 16 395 superconductive compounds,
- 50 000 normal compounds

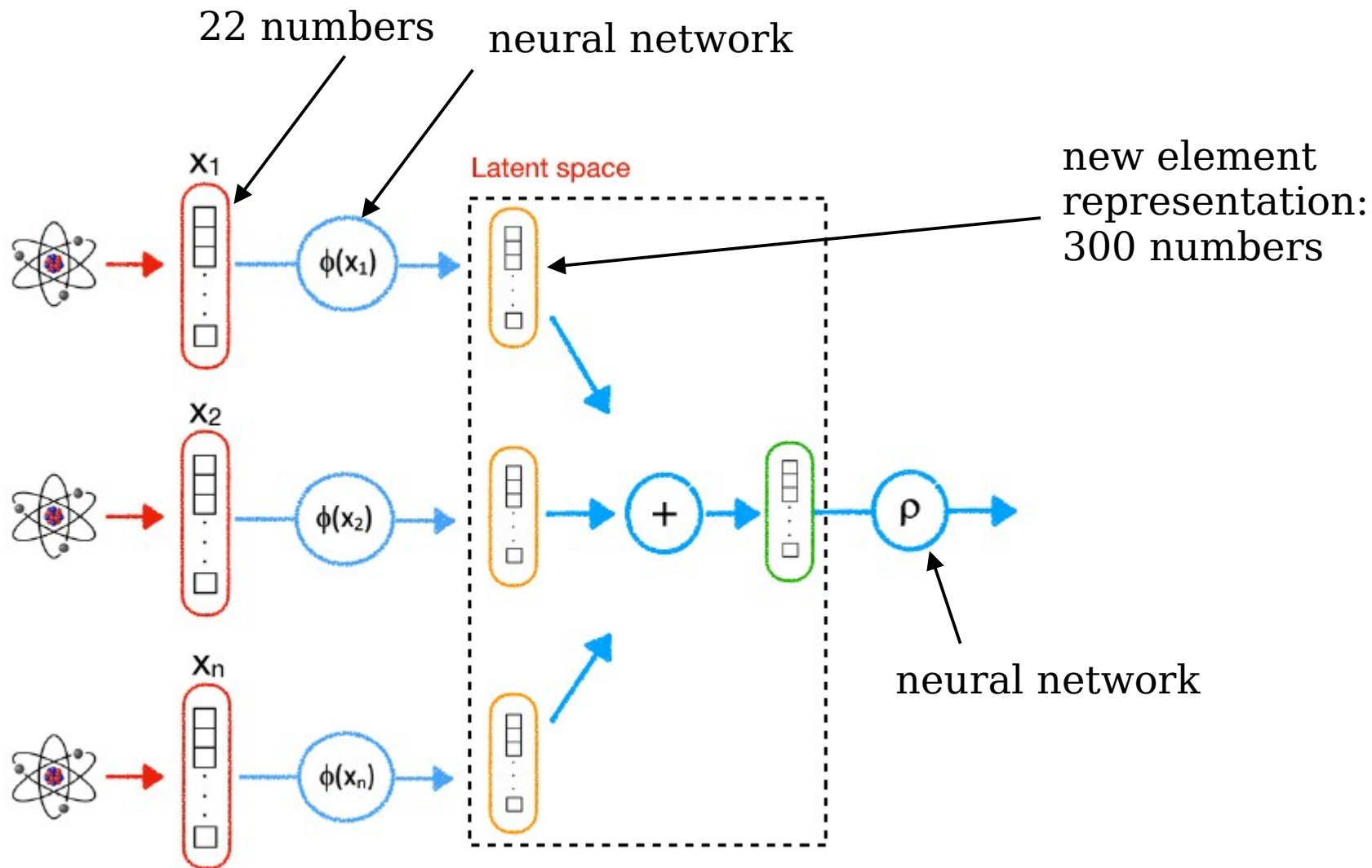
### Test:

- 207 compounds including 39 superconductors

Pereti, C., Bernot, K., Guizouarn, T. et al. From individual elements to macroscopic materials: in search of new superconductors via machine learning. npj Comput Mater 9, 71 (2023).  
<https://doi.org/10.1038/s41524-023-01023-6>

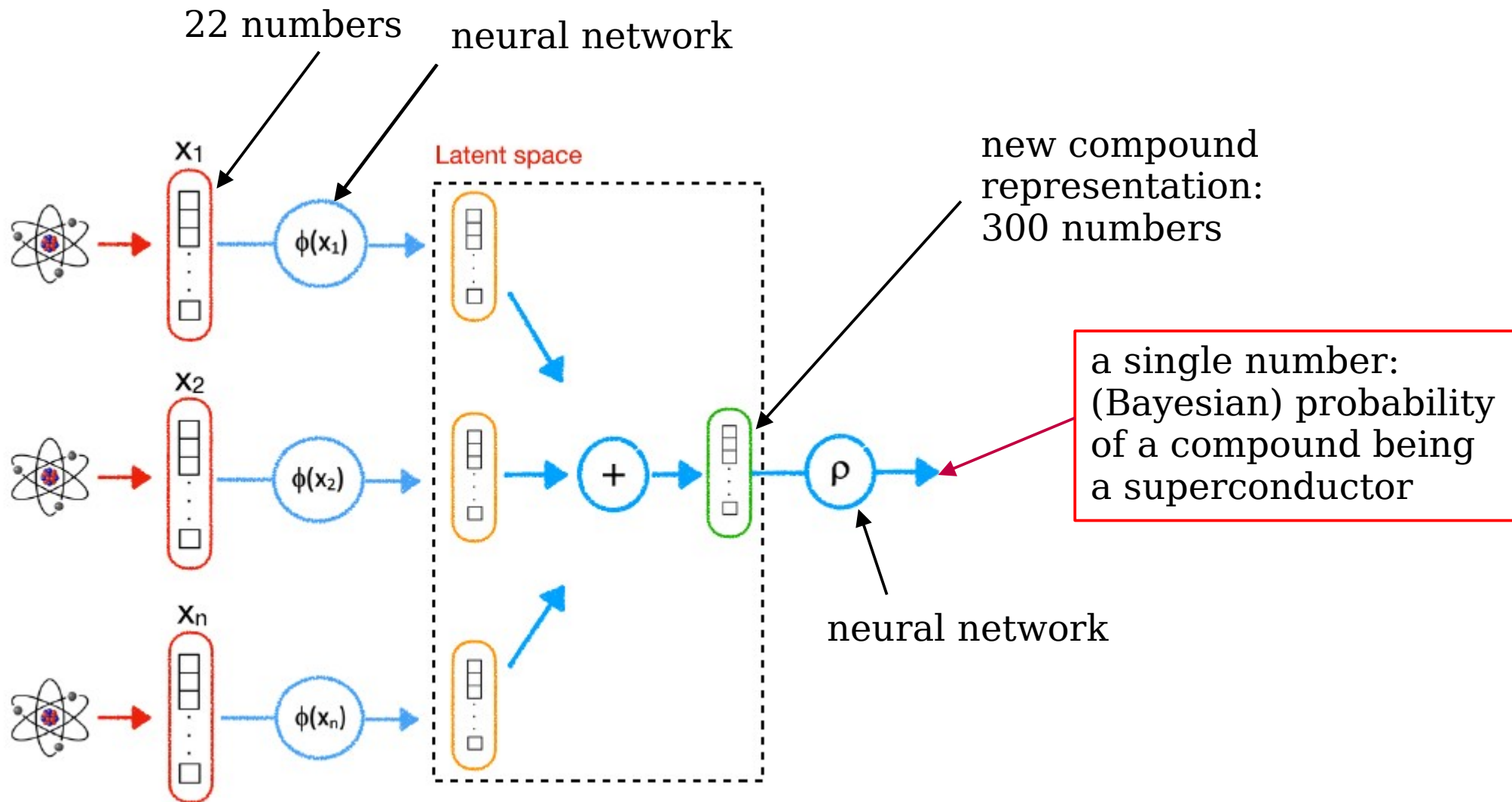


Pereti, C., Bernot, K., Guizouarn, T. et al. From individual elements to macroscopic materials: in search of new superconductors via machine learning. npj Comput Mater 9, 71 (2023).  
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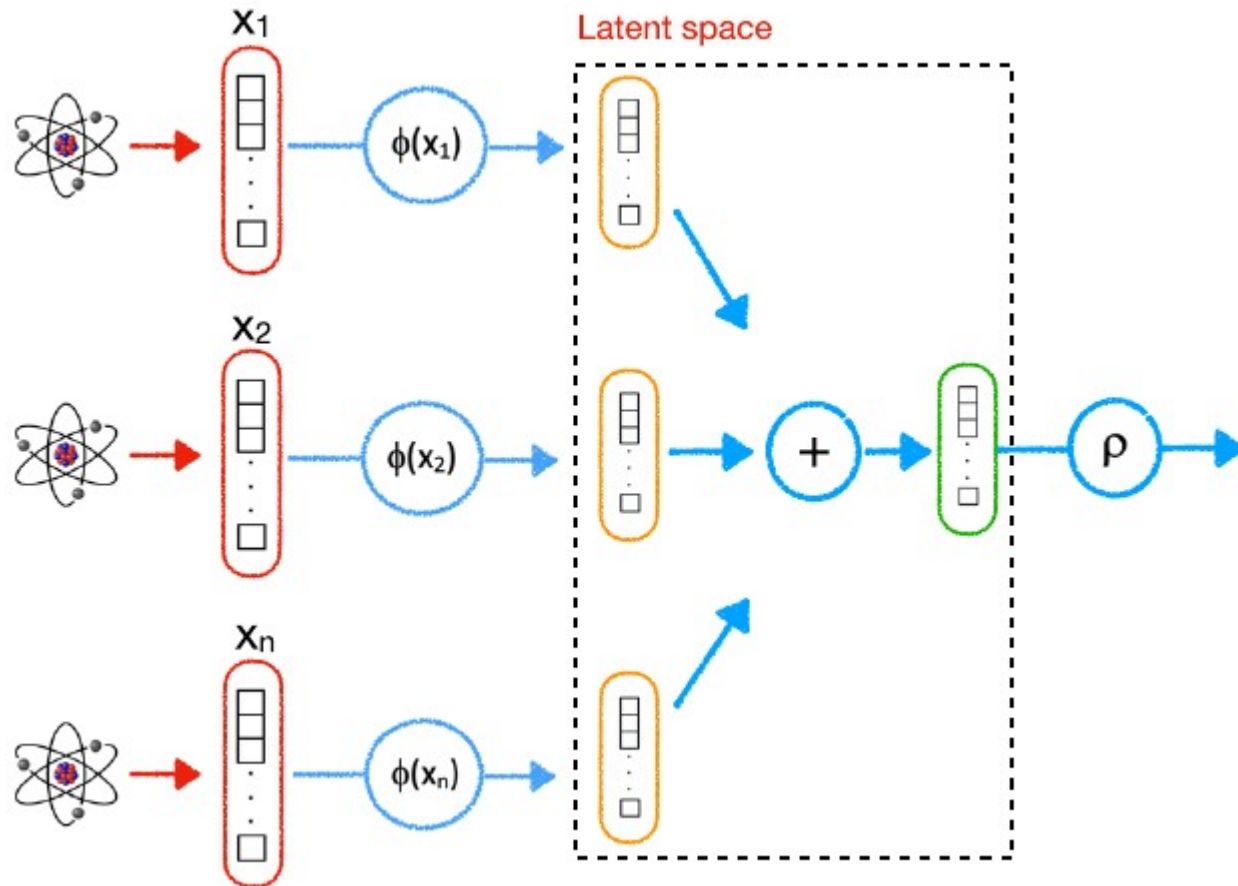


Pereti, C., Bernot, K., Guizouarn, T. et al. From individual elements to macroscopic materials: in search of new superconductors via machine learning. npj Comput Mater 9, 71 (2023).  
<https://doi.org/10.1038/s41524-023-01023-6>





Pereti, C., Bernot, K., Guizouarn, T. et al. From individual elements to macroscopic materials: in search of new superconductors via machine learning. npj Comput Mater 9, 71 (2023).  
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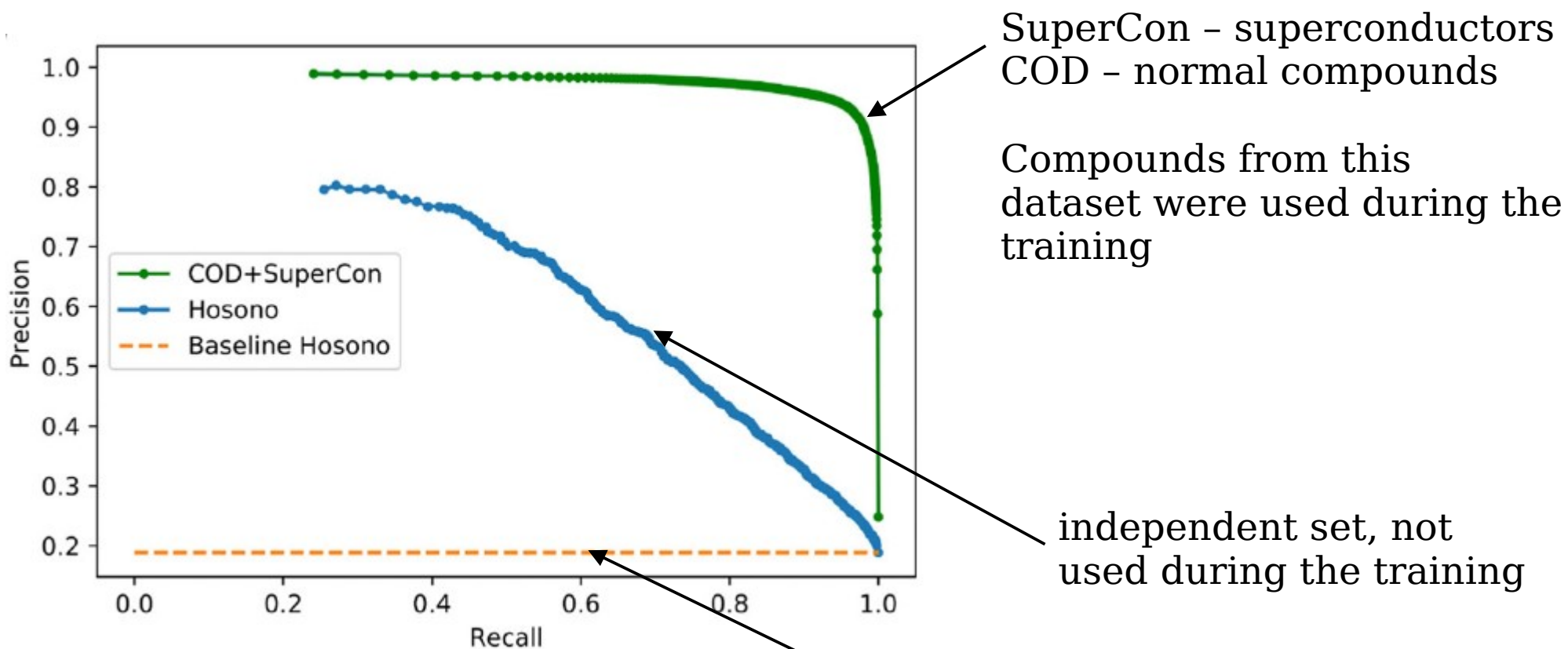


## Training:

60 sessions with random split into training and validation datasets in 80:20 proportions

Pereti, C., Bernot, K., Guizouarn, T. et al. From individual elements to macroscopic materials: in search of new superconductors via machine learning. *npj Comput Mater* 9, 71 (2023).  
<https://doi.org/10.1038/s41524-023-01023-6>





decision - **average** from 60 models responses

Pereti, C., Bernot, K., Guizouarn, T. et al. From individual elements to macroscopic materials: in search of new superconductors via machine learning. npj Comput Mater 9, 71 (2023).  
<https://doi.org/10.1038/s41524-023-01023-6>

decision -  
**average** from 60  
models responses

$$\text{avg} = \frac{1}{60} \left[ \sum p \right] > 0.5$$

$$\text{avg} = \frac{1}{3} (0.1) +$$

$$\frac{1}{3} (0.51) +$$

$$\frac{1}{3} (0.51) =$$

$$\frac{1.12}{3}$$

decision -  
**„voting”** of 60  
models

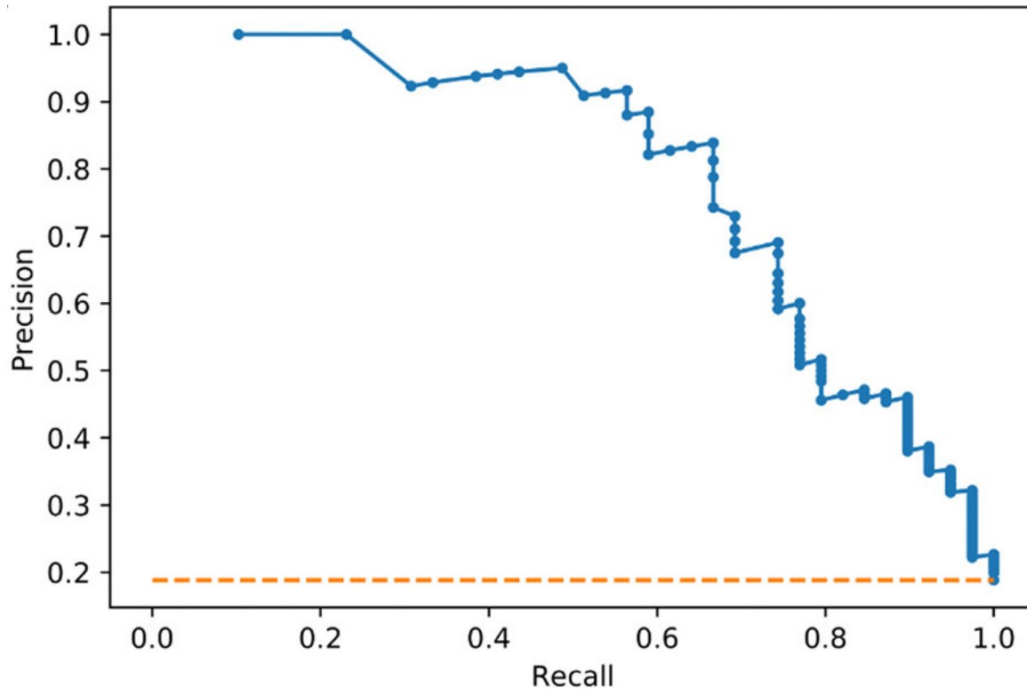
$$\text{vote} = \frac{1}{60} \left[ \sum p > 0.5 \right] > 0.5$$

$$\text{vote} = \frac{1}{3} (0.1 > 0.5) +$$

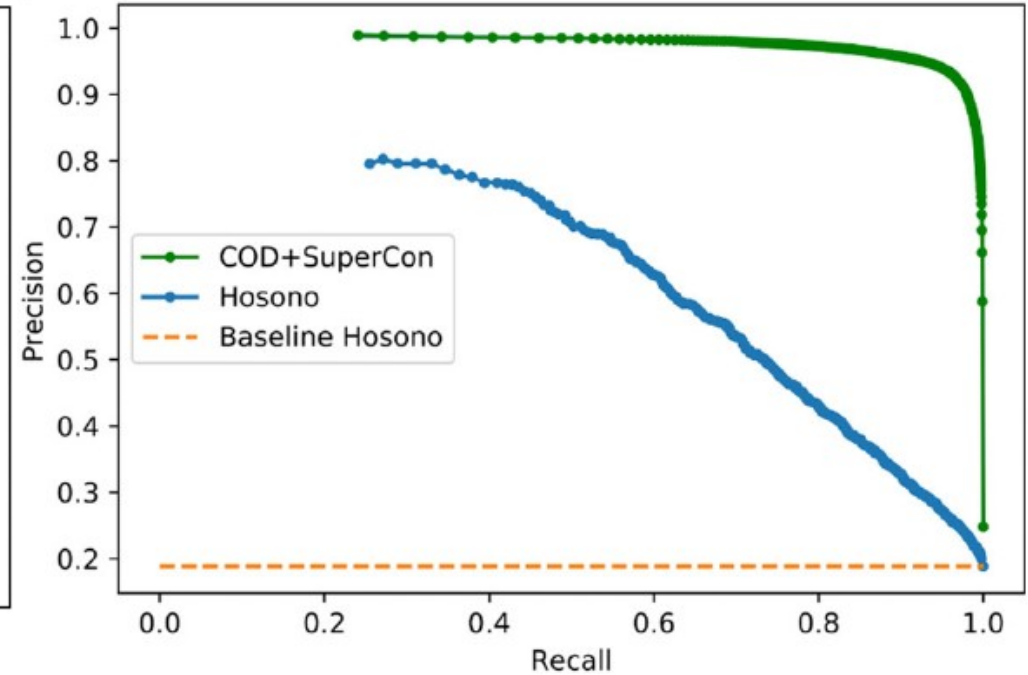
$$\frac{1}{3} (0.51 > 0.5) +$$

$$\frac{1}{3} (0.51 > 0.5) =$$

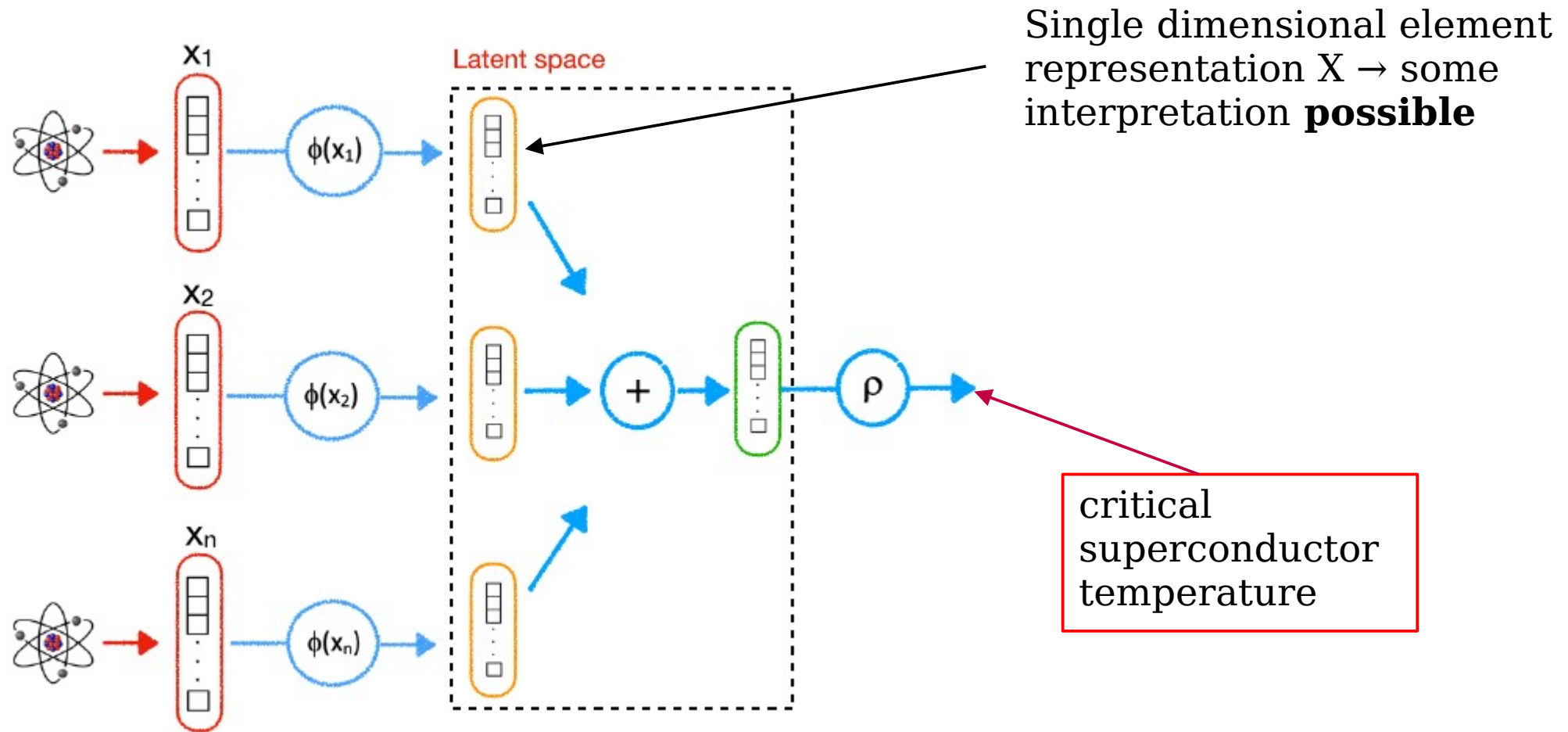
$$\frac{2}{3}$$



decision -  
**„voting”** of 60  
models

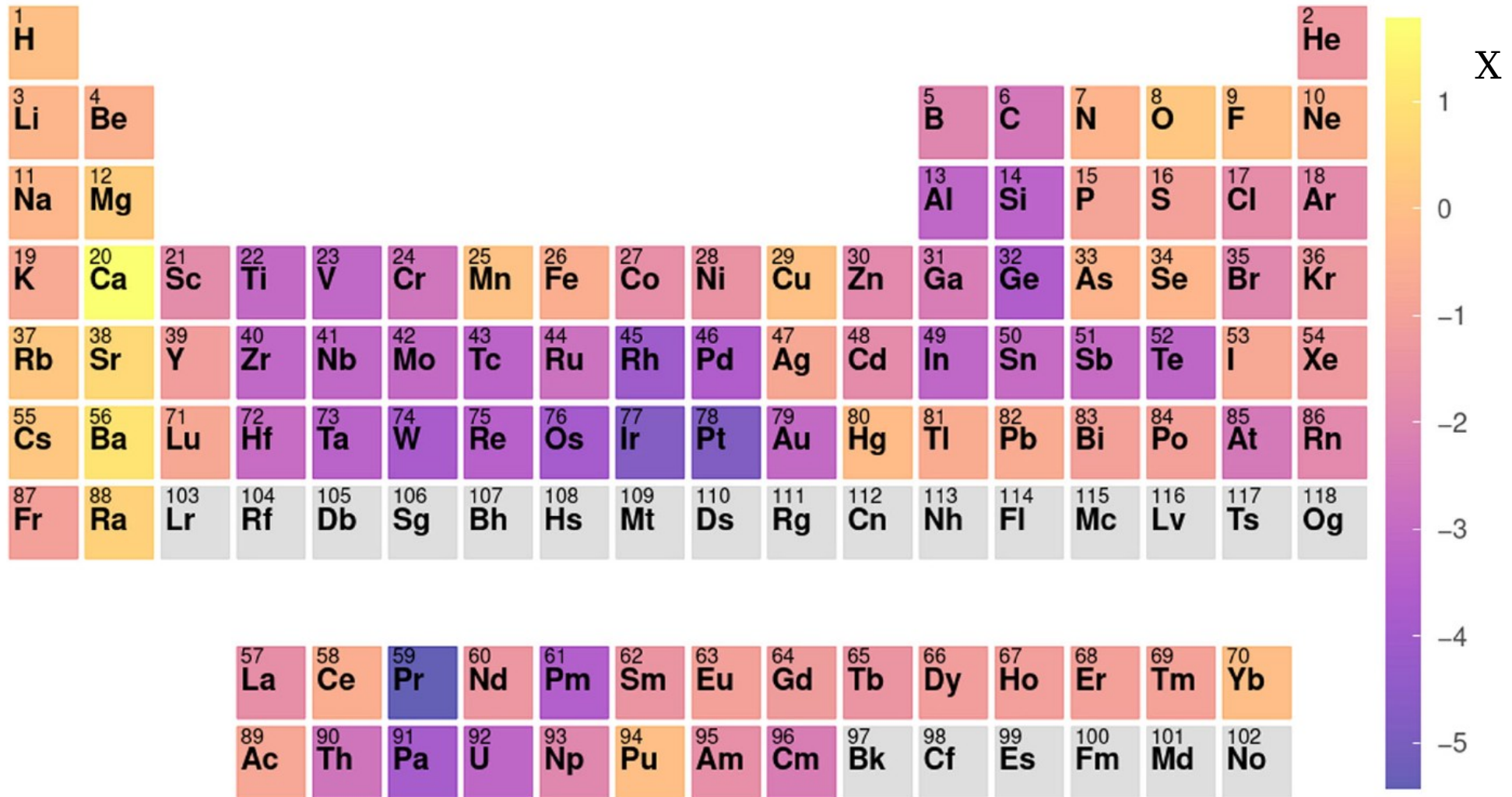


decision -  
**average** from 60  
models responses



Pereti, C., Bernot, K., Guizouarn, T. et al. From individual elements to macroscopic materials: in search of new superconductors via machine learning. npj Comput Mater 9, 71 (2023).  
<https://doi.org/10.1038/s41524-023-01023-6>

Contribution of an element on critical temperature value:  
large  $X \rightarrow$  increased  $T_c$



Pereti, C., Bernot, K., Guizouarn, T. et al. From individual elements to macroscopic materials: in search of new superconductors via machine learning. npj Comput Mater 9, 71 (2023).  
<https://doi.org/10.1038/s41524-023-01023-6>

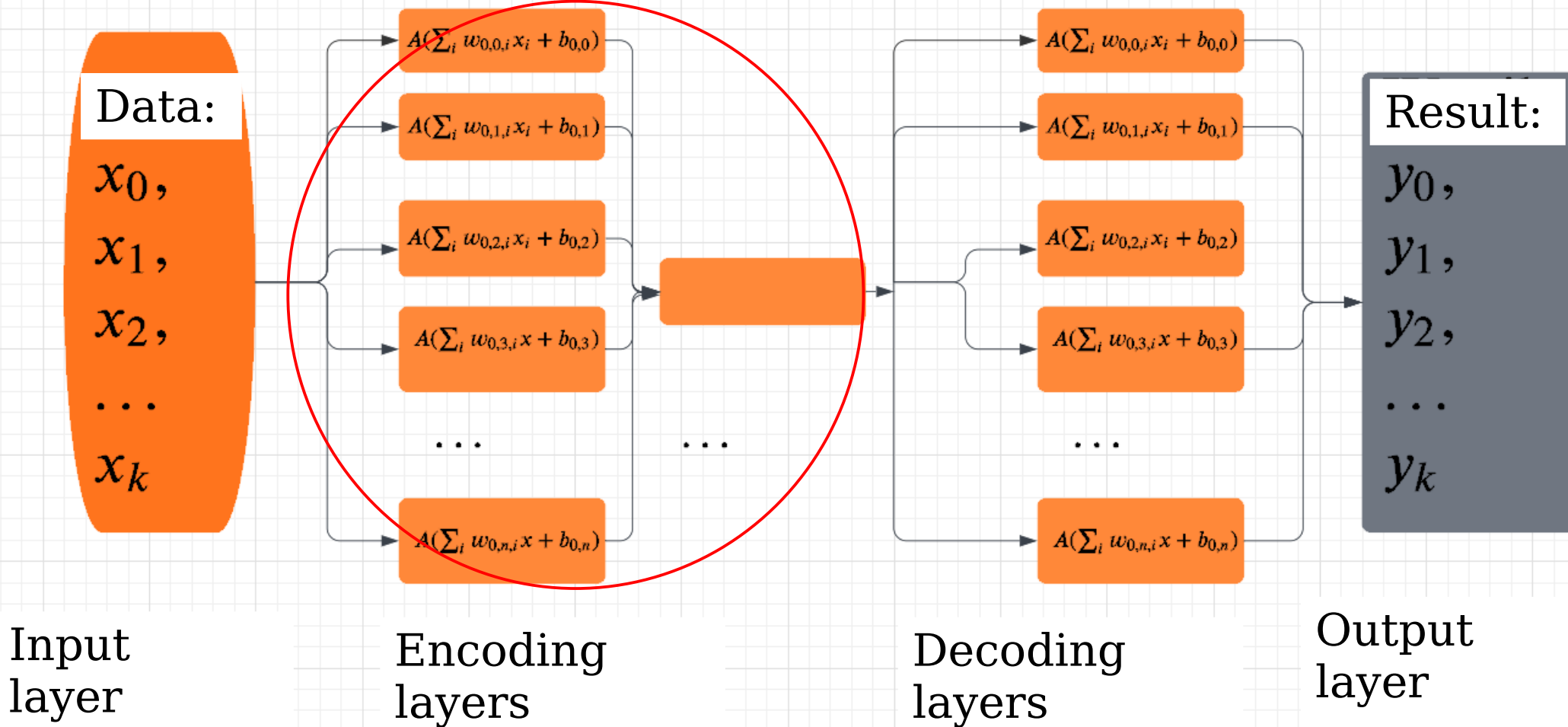
**Task:** seek a function:  $\mathbf{E}(\mathbf{x}): \mathbf{R}^n \rightarrow \mathbf{R}^m$ ,  $m \neq n$ , such that keeps maximum amount of information from the original input data

**Solution:** creation of two functions:  $E(x): x \rightarrow z$ ,  $D(z) = E^{-1}(z): z \rightarrow x$

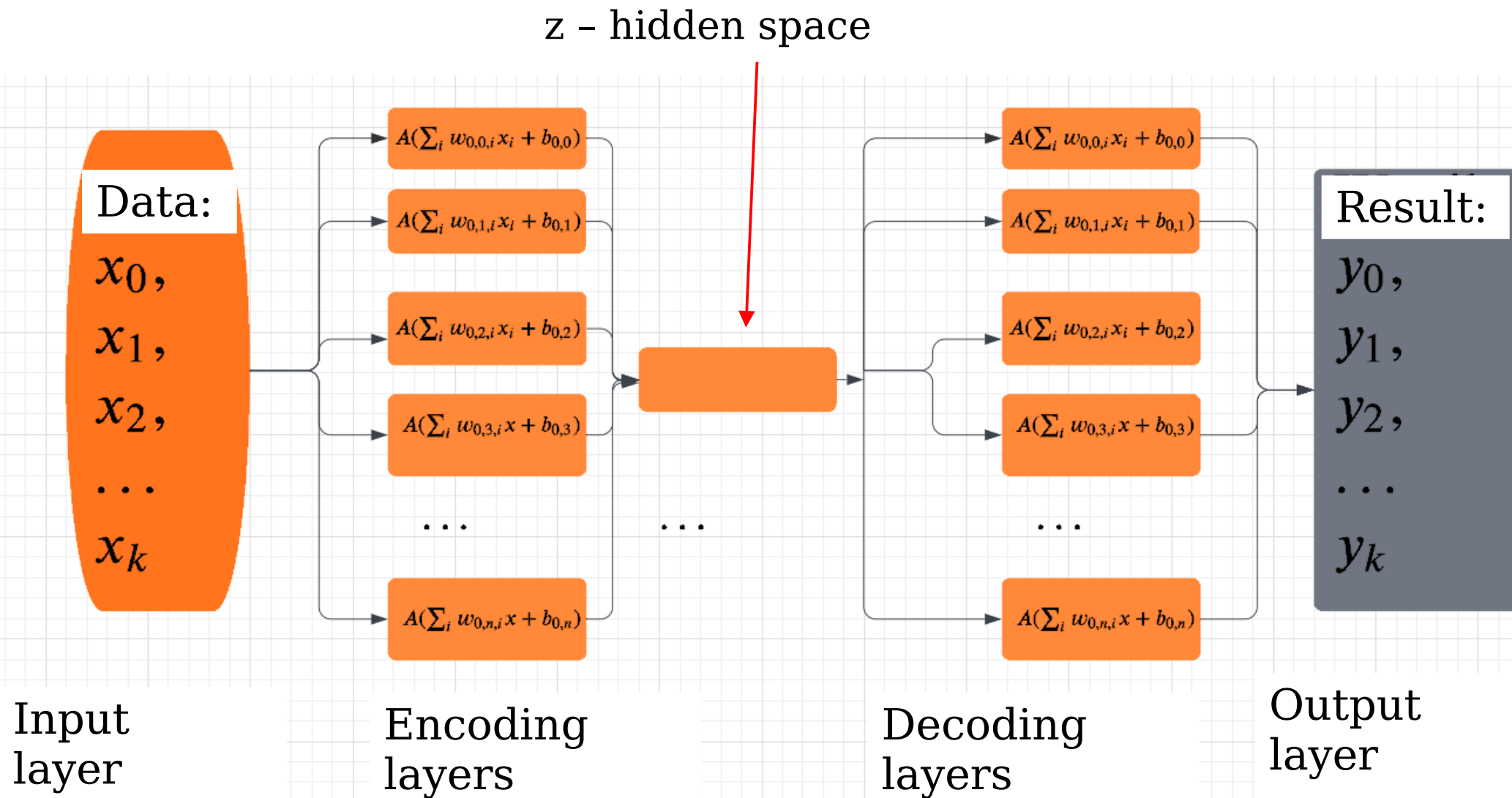
$$\mathbf{x} = \mathbf{D}(\mathbf{E}(\mathbf{x}))$$

**$\mathbf{E}(\mathbf{x})$  and  $\mathbf{D}(\mathbf{z})$  parametrization:** given *implicite* by the neural network connections

$E(x)$

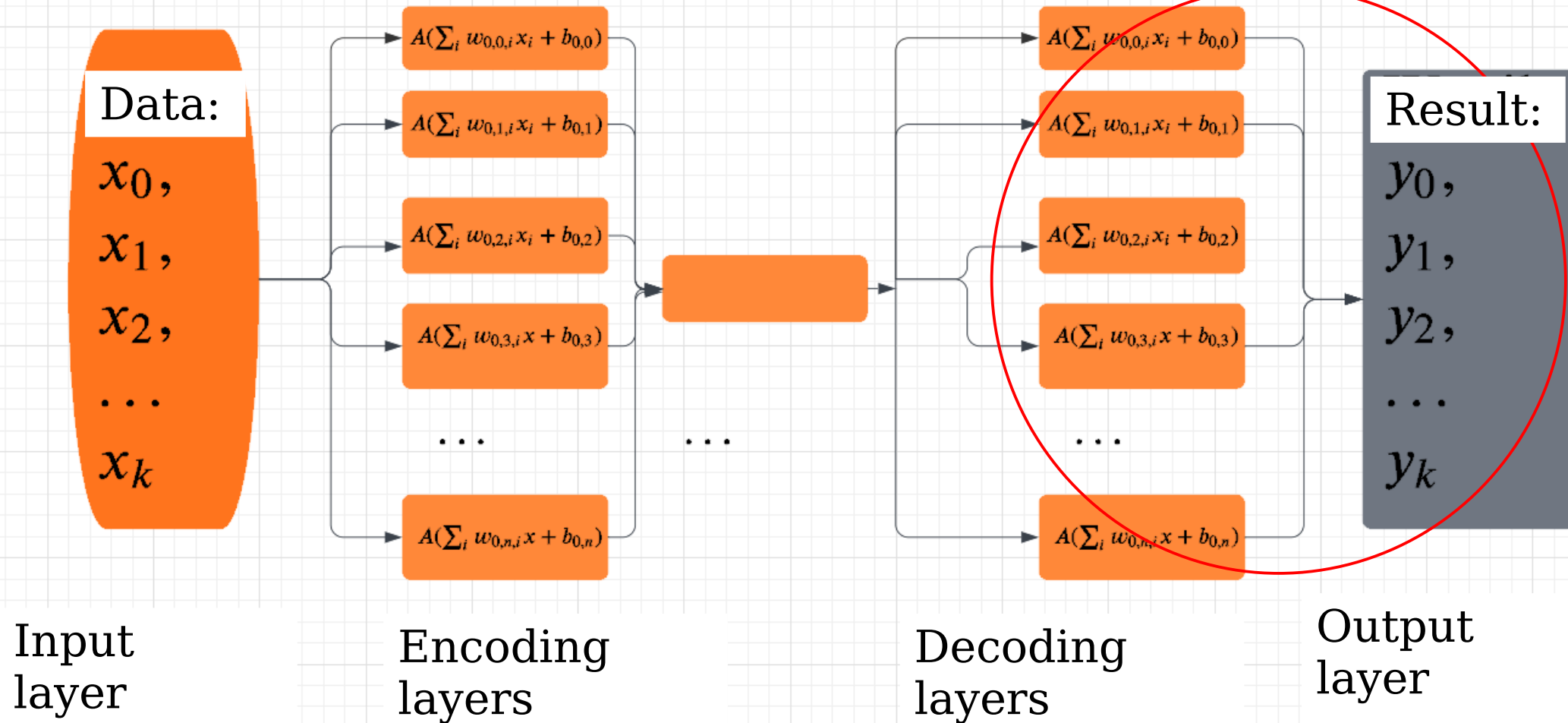








$D(z)$



## The task:

identification of active galaxy nuclei (AGN) with high redshift  $z \sim 3$

## Input data:

- light spectra in form of intensity in 914 wave length bins

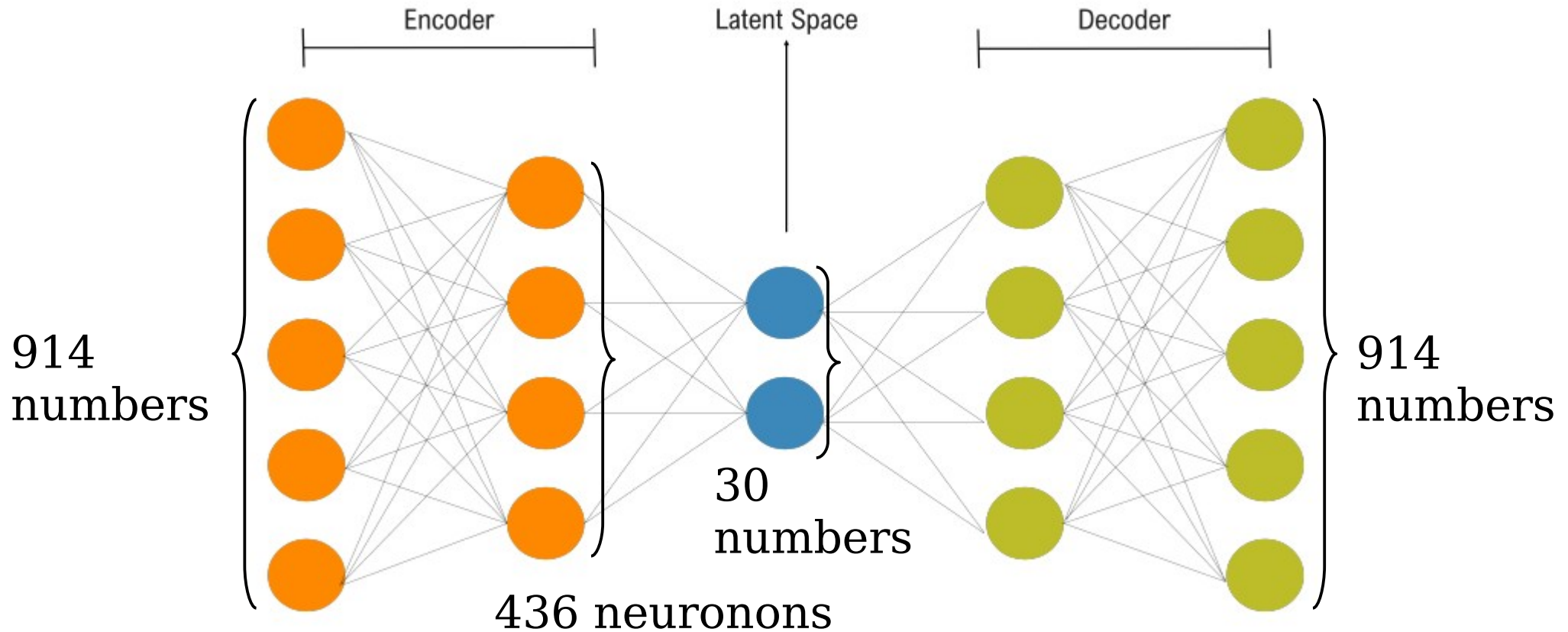
## Datasets:

### Training and validation:

- 2458 spectra: 23% AGN, 22% stars, 55% normal galaxies with (High- $z$ ) or low (Low- $z$ )  $z$
- Division: 4:1

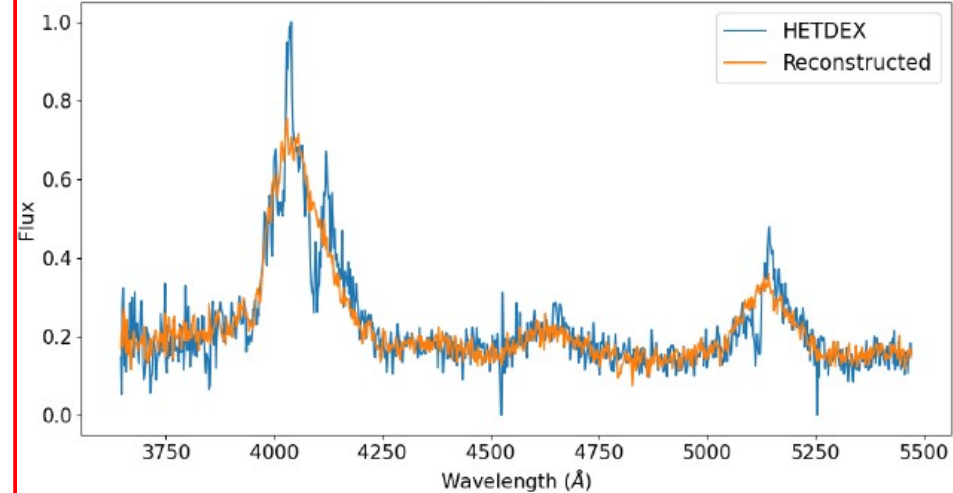
### Untagged set:

- 716 objects with  $z \sim 3$  (Photo- $z$ )

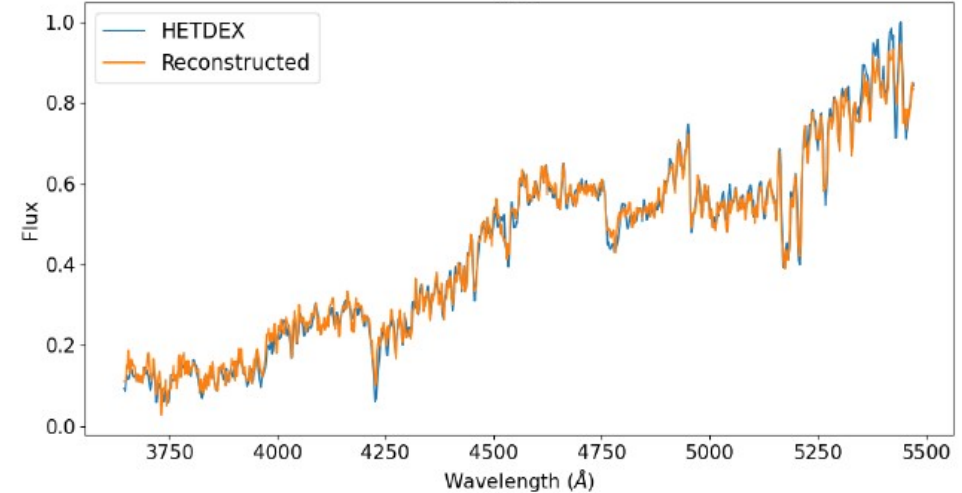


Identifying Active Galactic Nuclei at  $z \sim 3$  from the HETDEX Survey Using Machine Learning  
[arXiv:2302.11092](https://arxiv.org/abs/2302.11092) [[astro-ph.GA](https://arxiv.org/abs/2302.11092)]

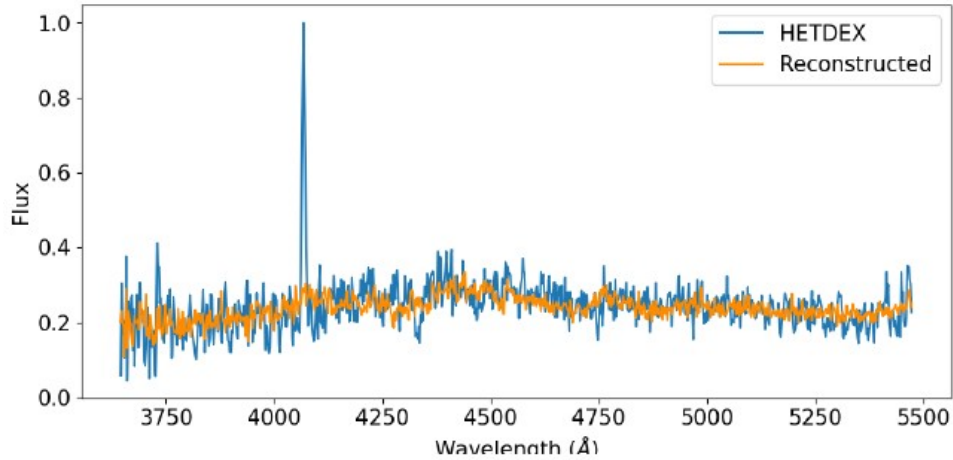
AGN



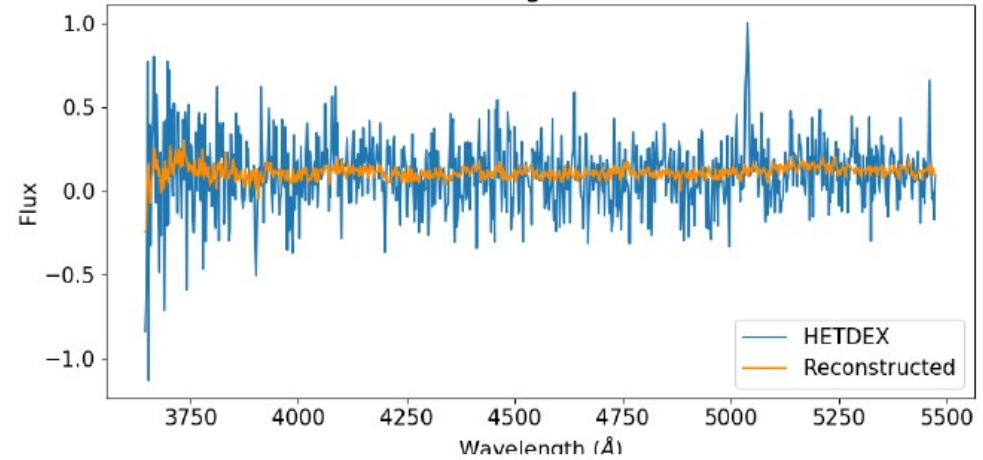
Star



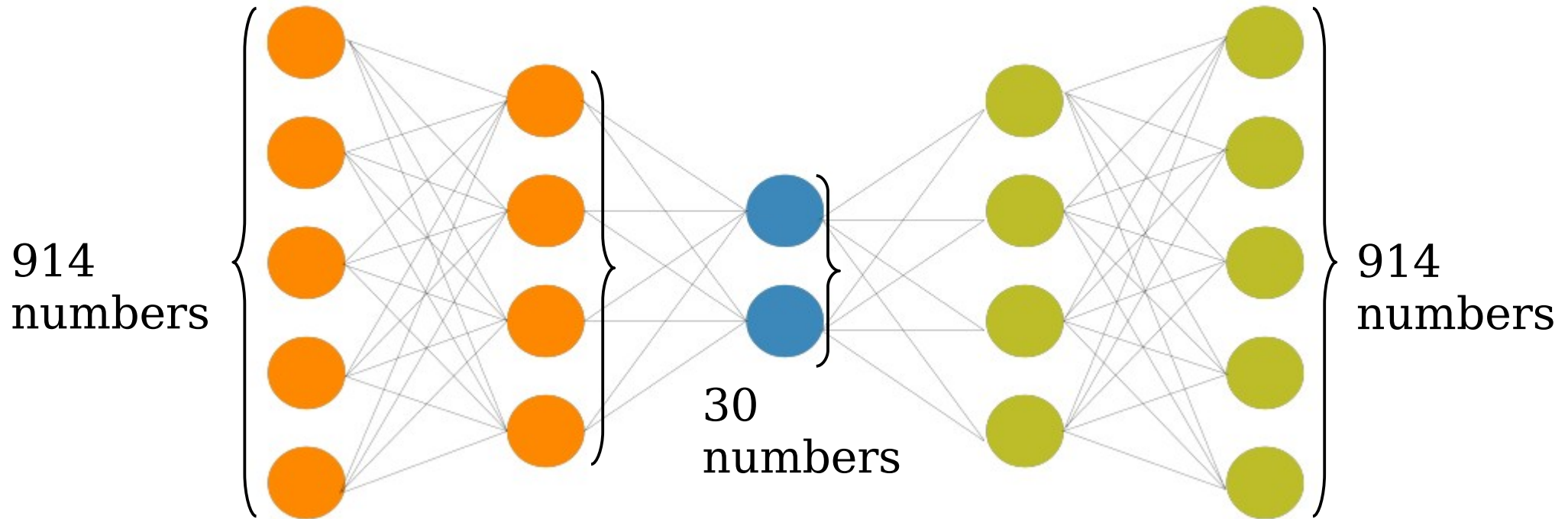
Low-z



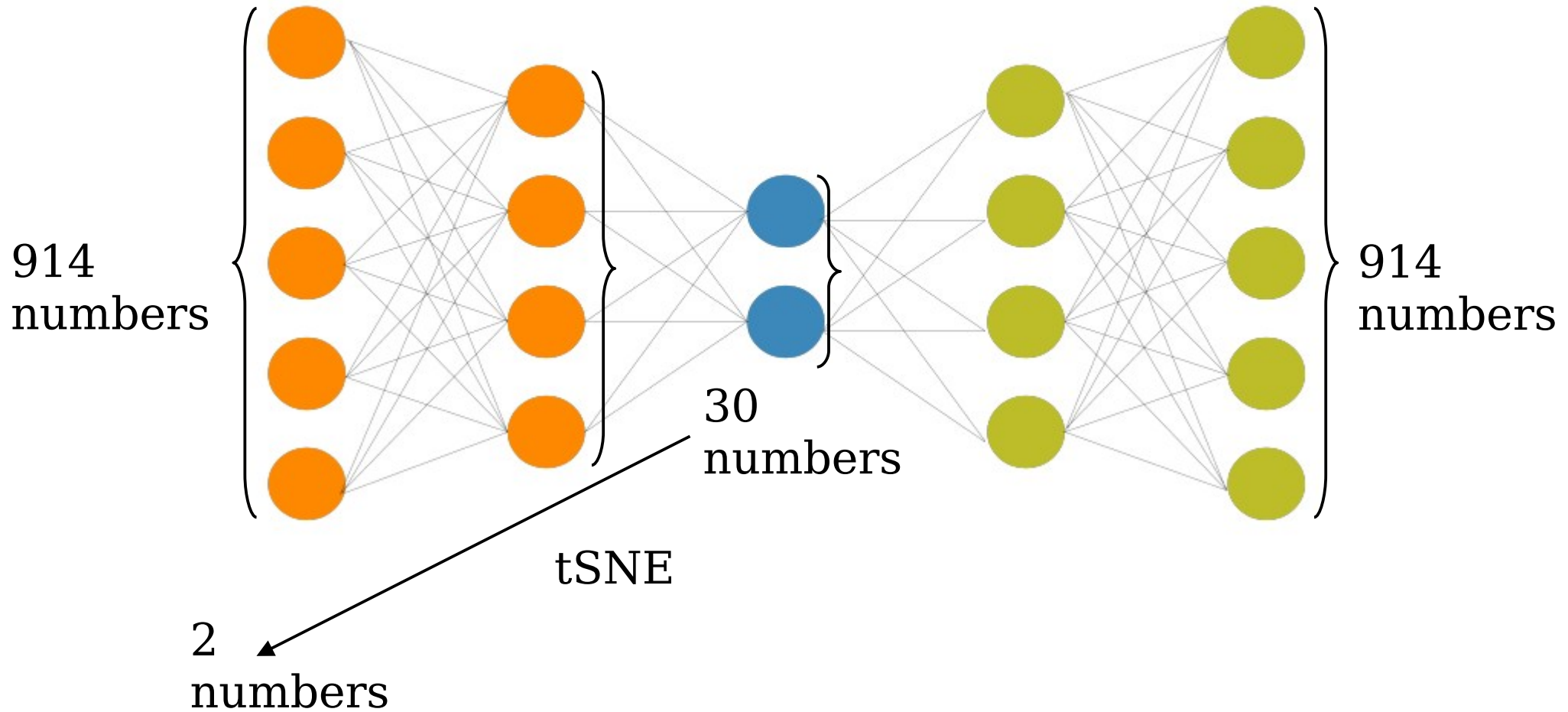
High-z



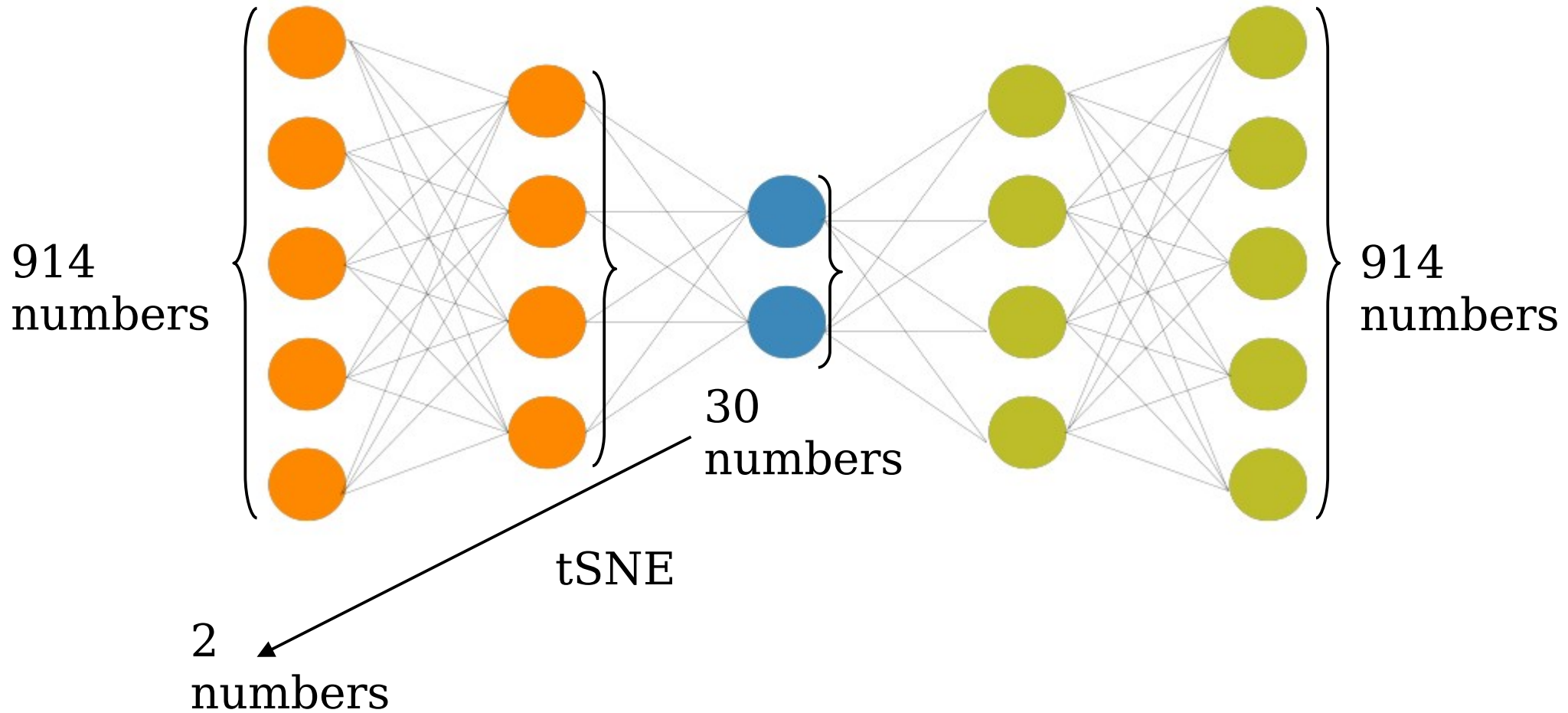
Identifying Active Galactic Nuclei at  $z \sim 3$  from the HETDEX Survey Using Machine Learning  
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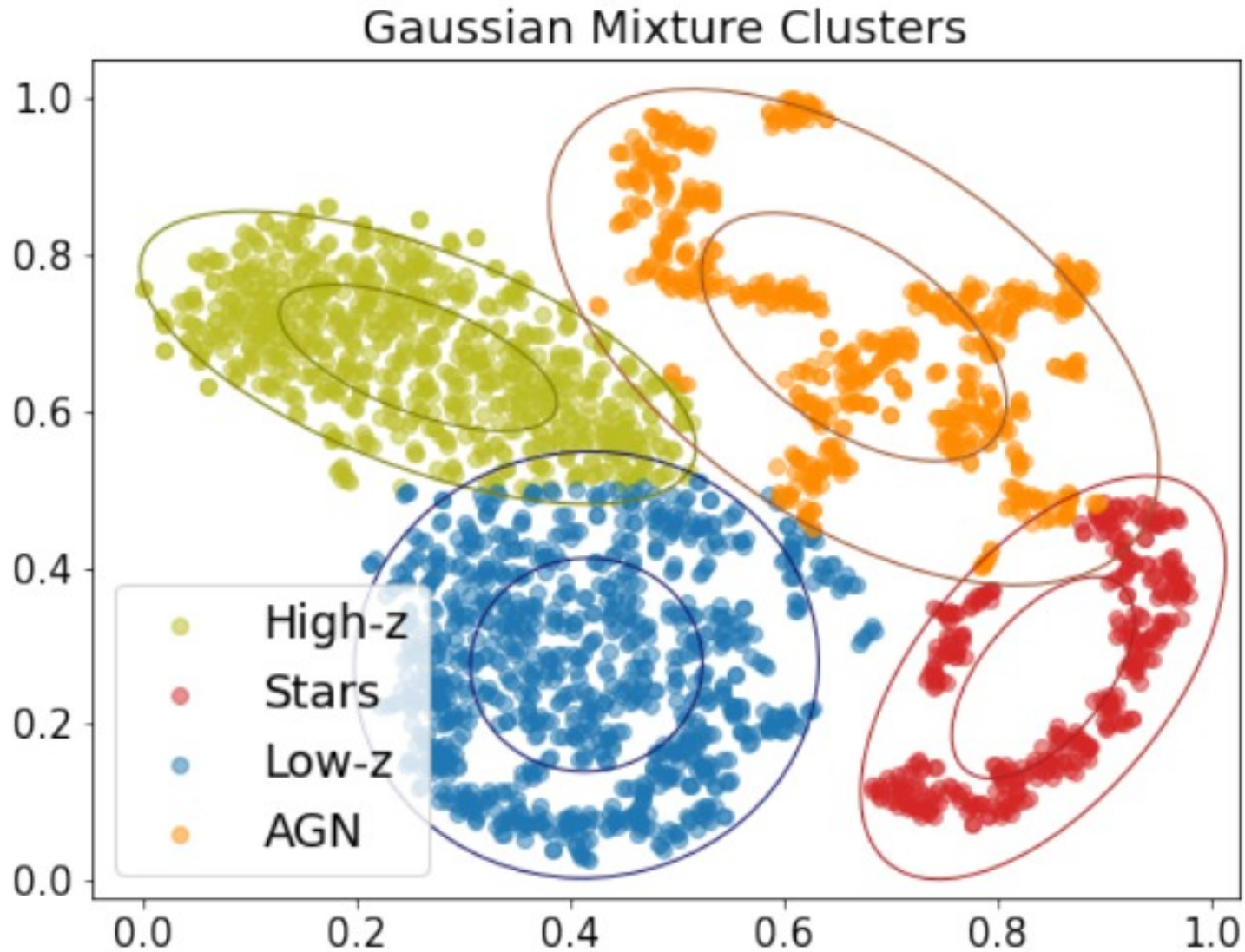




Gaussian mixture - one Gaussian for each category

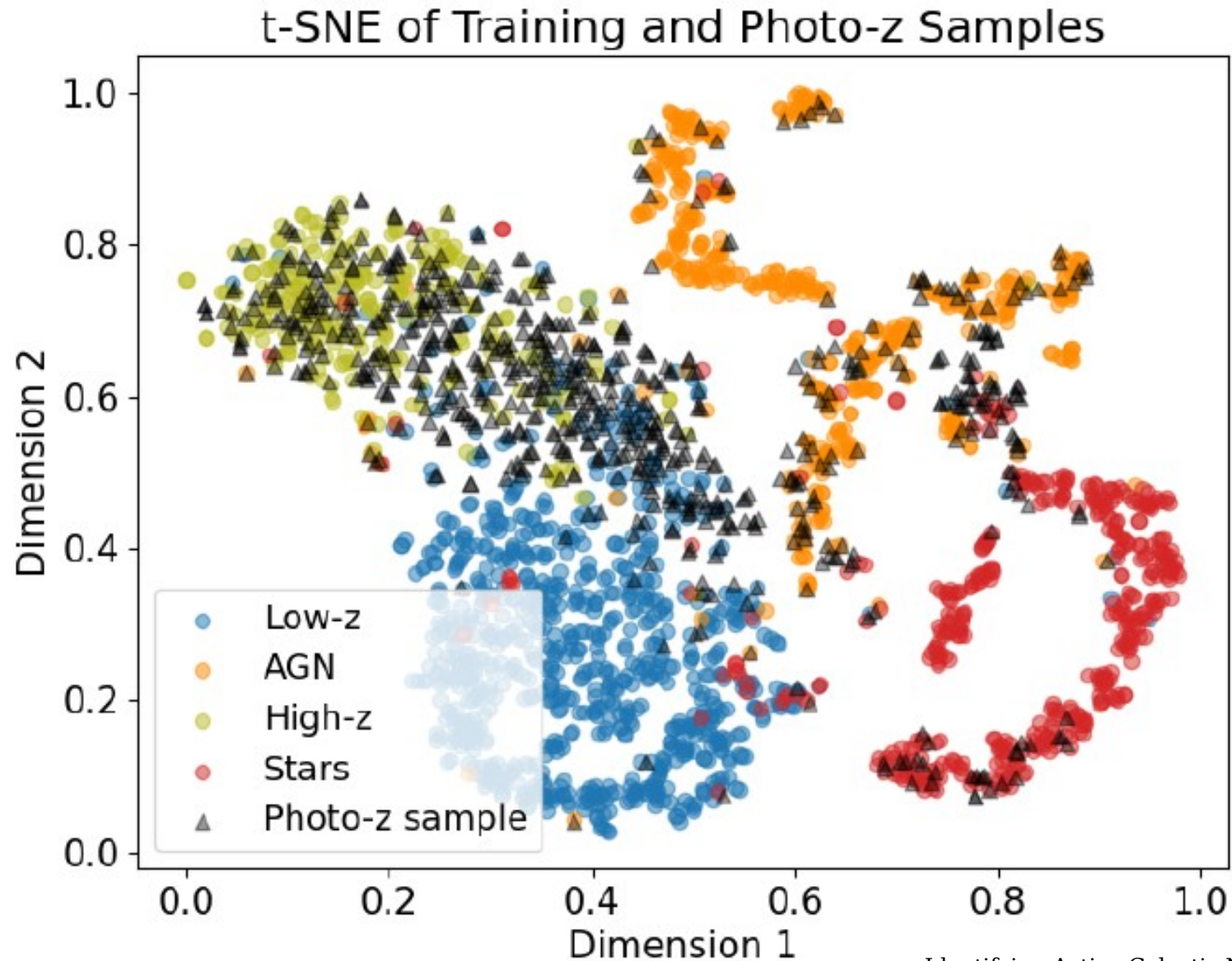
$$p(x_1, x_2) = \sum_{i=0}^3 \alpha_i \text{Gauss}(\vec{\mu}_i, \text{Cov}_i)$$

Identifying Active Galactic Nuclei at  $z \sim 3$  from the HETDEX Survey Using Machine Learning  
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the HETDEX Survey Using Machine Learning  
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[arXiv:2302.11092](https://arxiv.org/abs/2302.11092) [[astro-ph.GA](https://arxiv.org/abs/2302.11092)]

## The task:

muon transverse momentum calculation ( $p_T$ )

Warsaw  
CMS group  
work

## Input data:

- “hit pattern” - muon position at four points, momentum direction in two points

## Datasets:

### Training:

- 5.5 M muons with various  $p_T$

### Testing:

- 0.5M muons with various  $p_T$

## Naive Bayes:

- calculate hit configuration likelihood assuming hit positions are independent between layers:

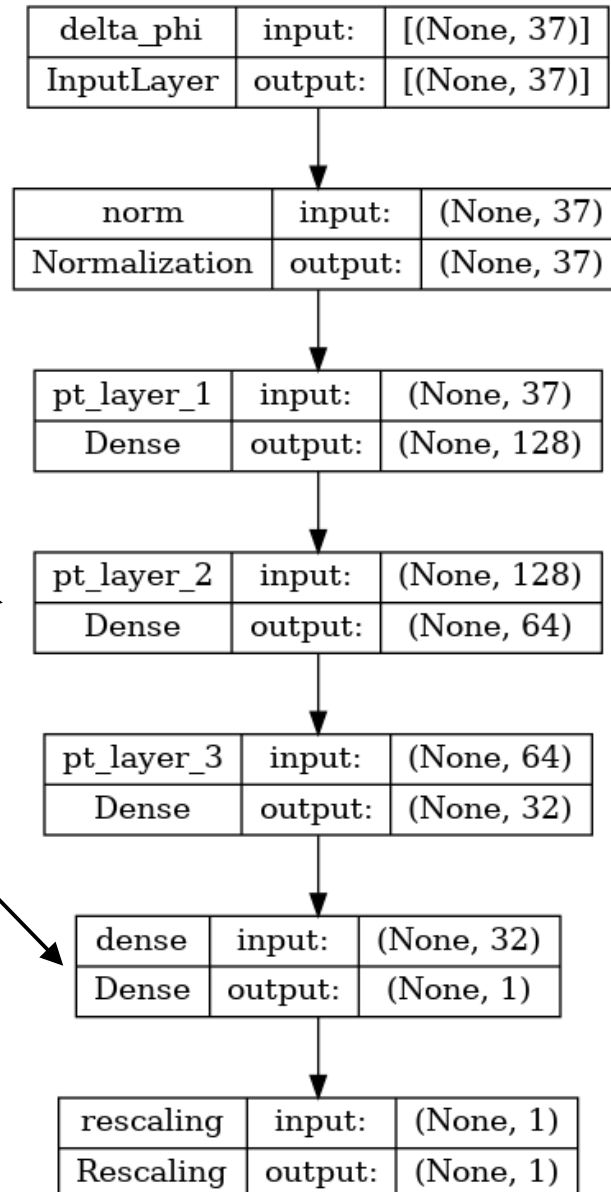
$$L(\text{hit configuration} | p_T) = \prod_{\text{layers}} p(\text{hit position in layer } i | p_T)$$

- select  $p_T$  giving the largest likelihood value

## Machine Learning:

- very standard fully connected network:

4 layers of neurons



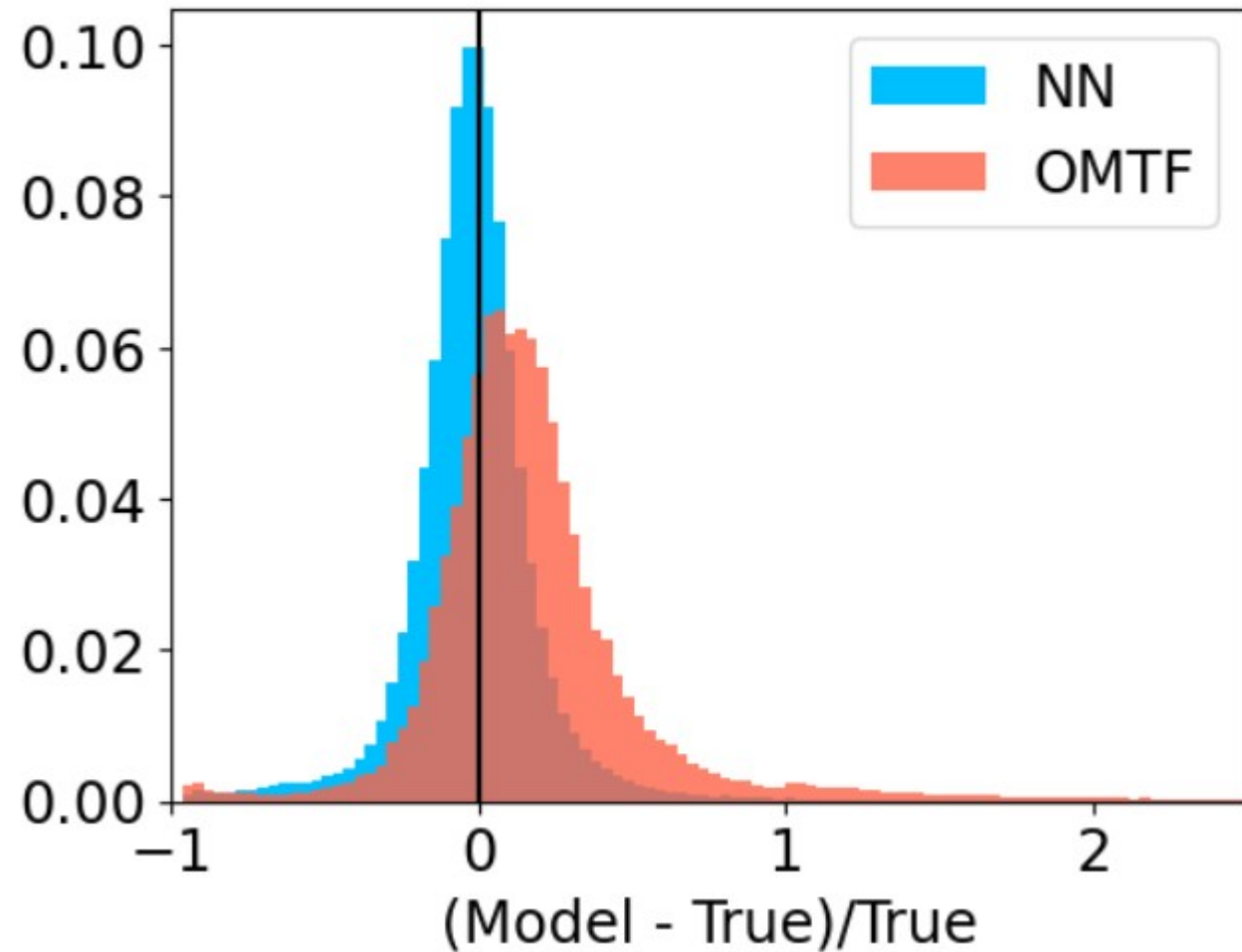
NN value centered at true value:

- good for measurement

$$\langle p_T^{measured} \rangle = p_T^{true}$$

- bad for event selection

$$P(p_T^{measured} \geq p_T^{true}) = 0.5$$



NN value centered at true value →  
“bug/feature” can be easily fixed with manual scaling:

$$p_T = 1.15 \cdot p_T^{NN}$$

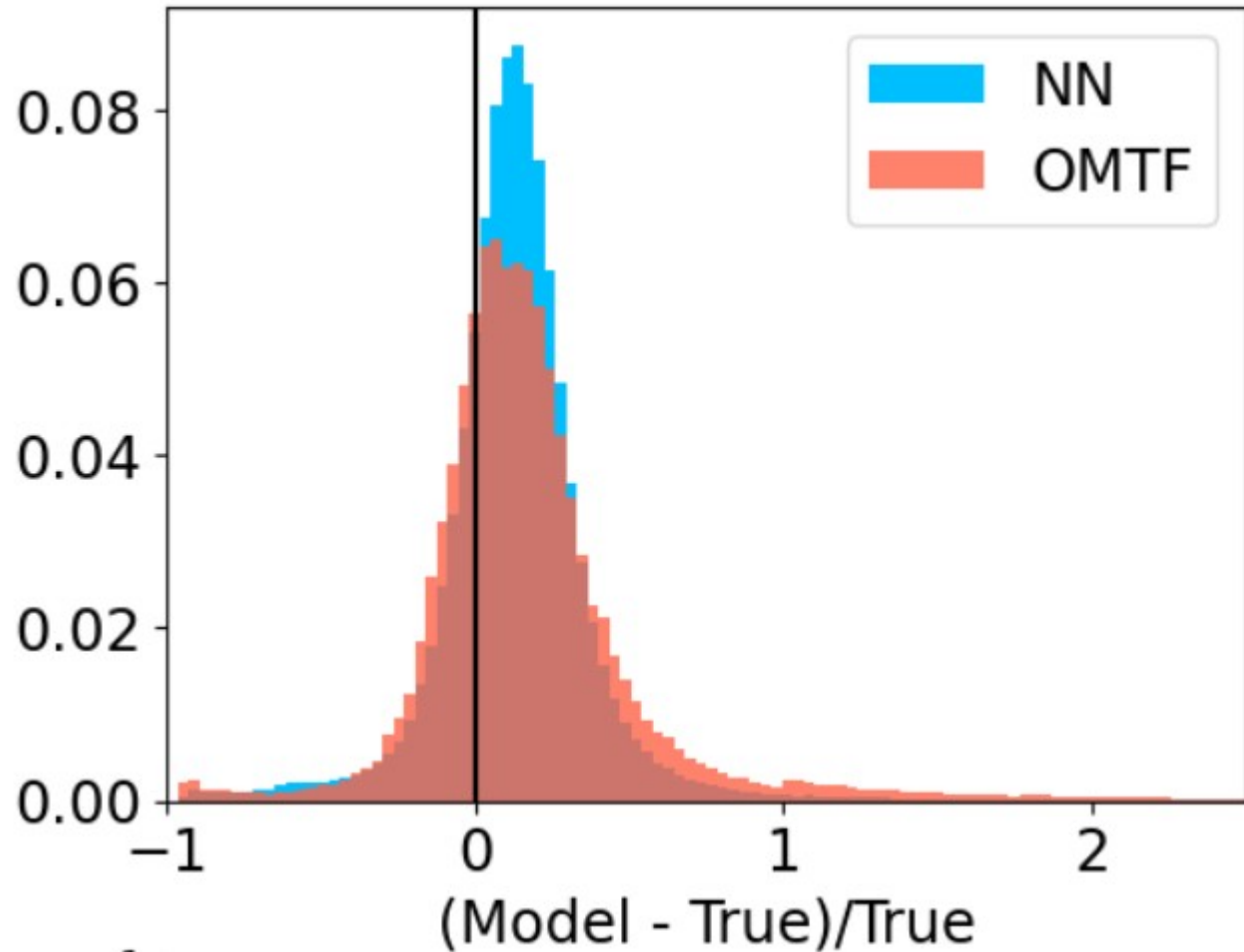
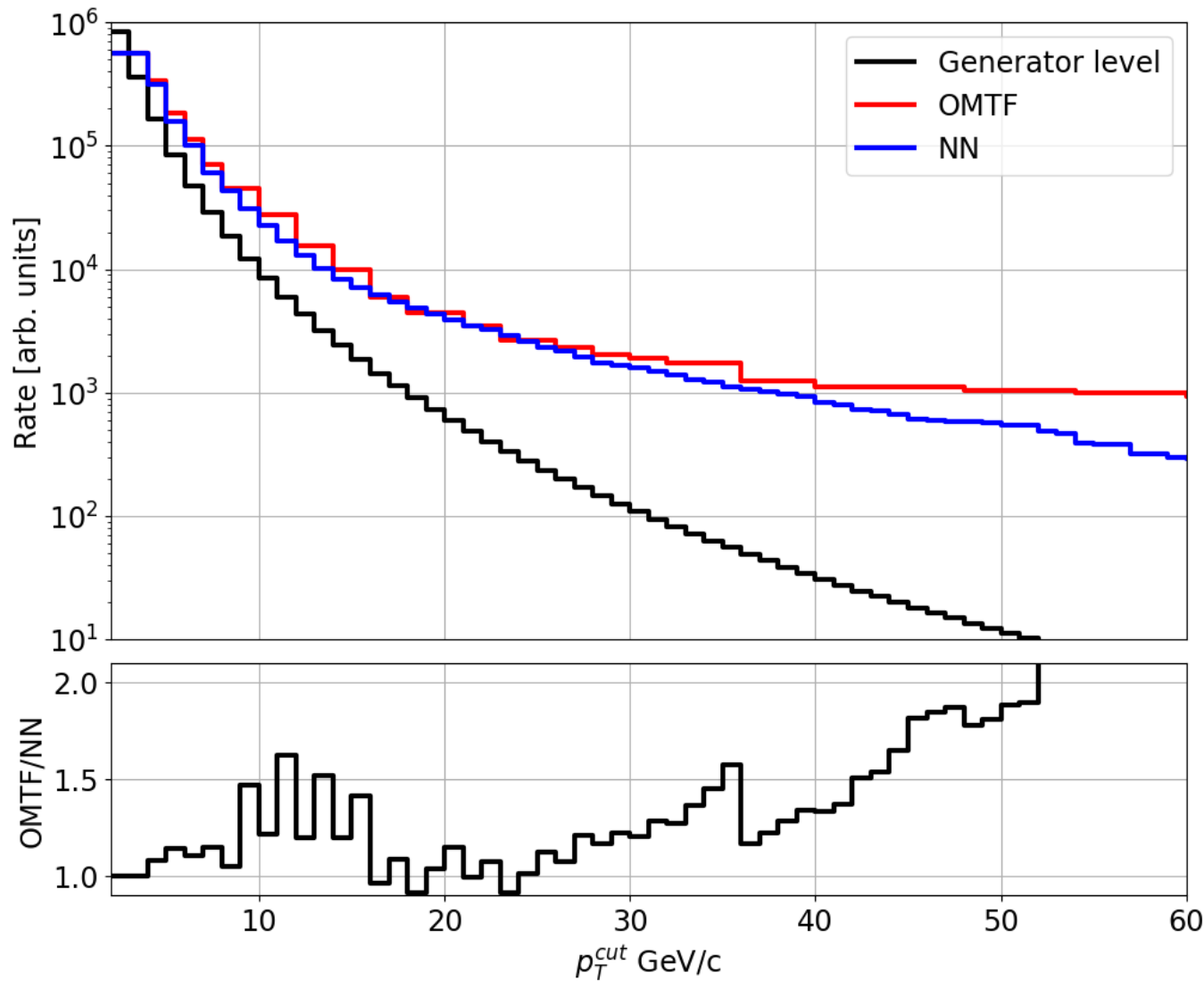
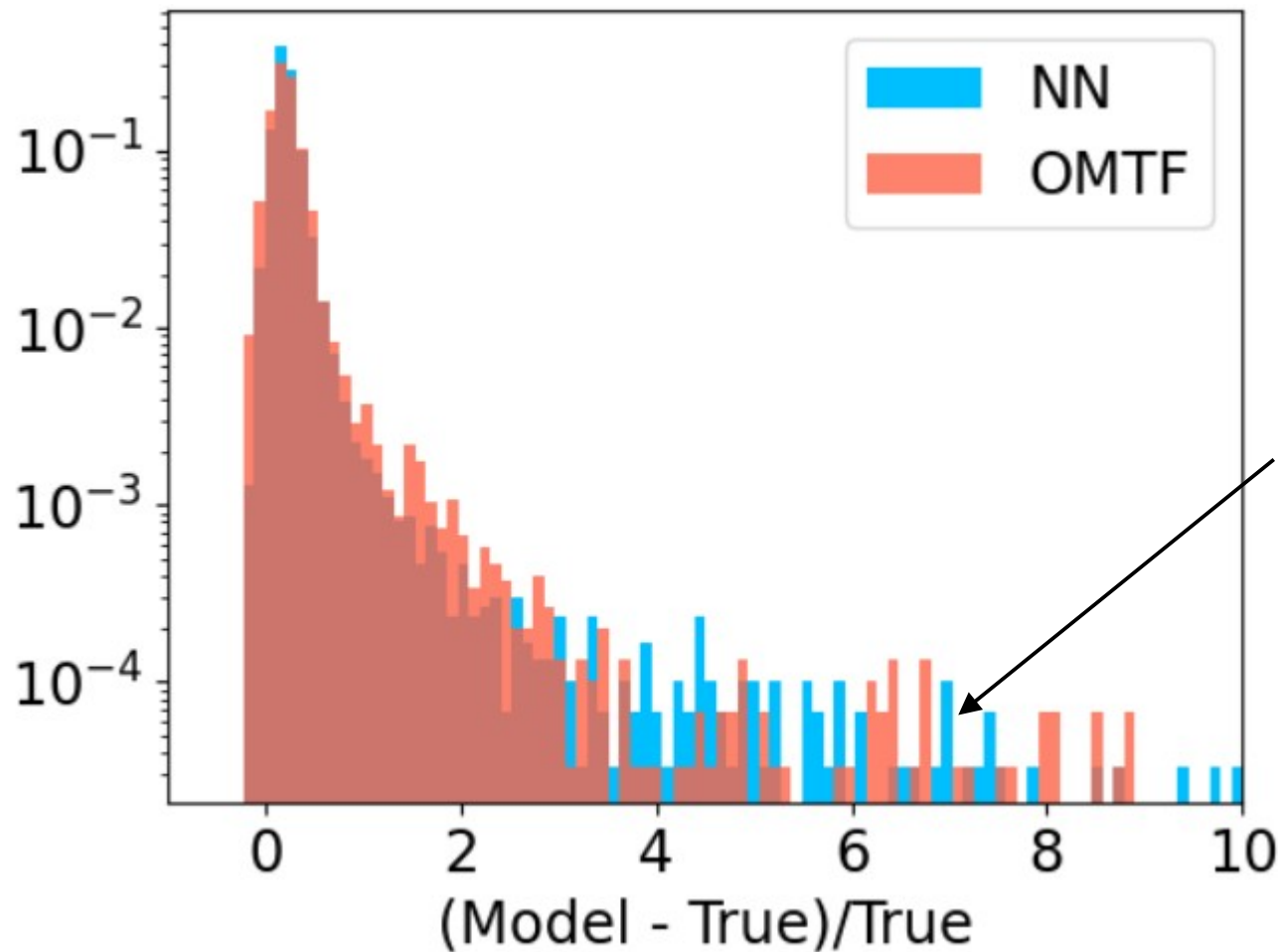


Figure of merit:  
accepted event rate

$$R[\text{ev/s}] = N[\text{ev/s}] : p_T^{\text{measured}} > p_T^{\text{cut}}$$



$$p_T^{\text{generated}} < 5 \text{ GeV}/c$$



The rate is dominated by highly overestimated low  $p_T$  muons



**Solution:**

$$p_T^{measured} = \min(\text{Naive Bayes, NN})$$

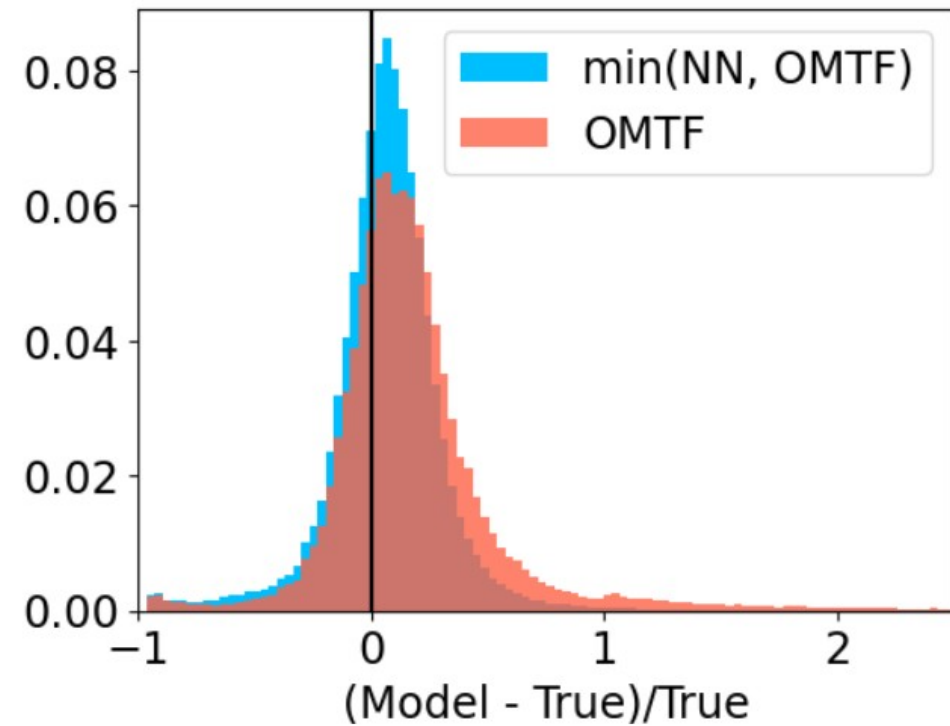
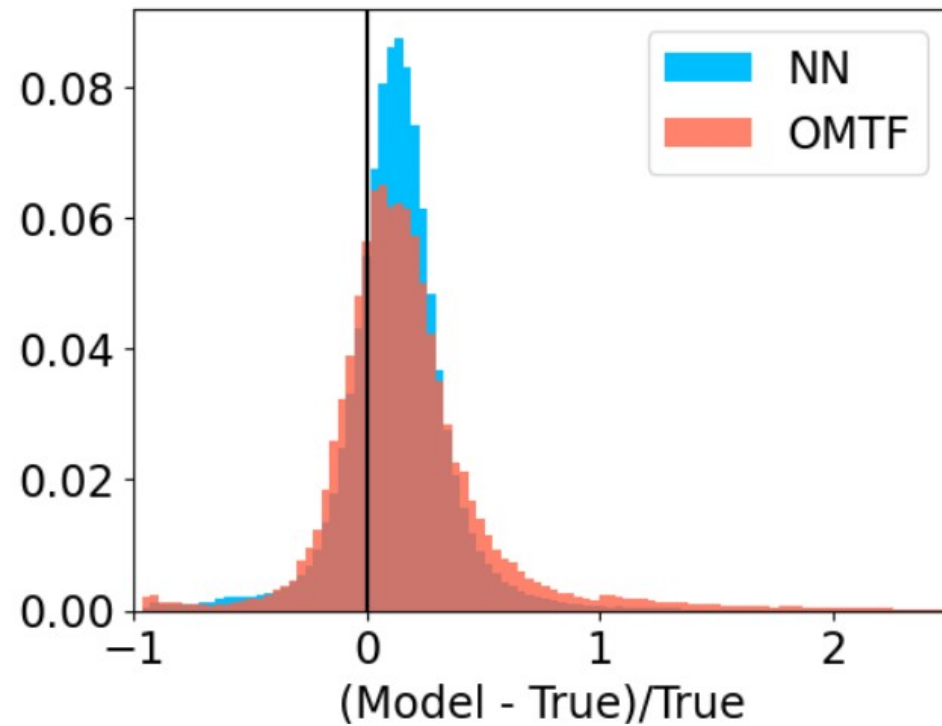
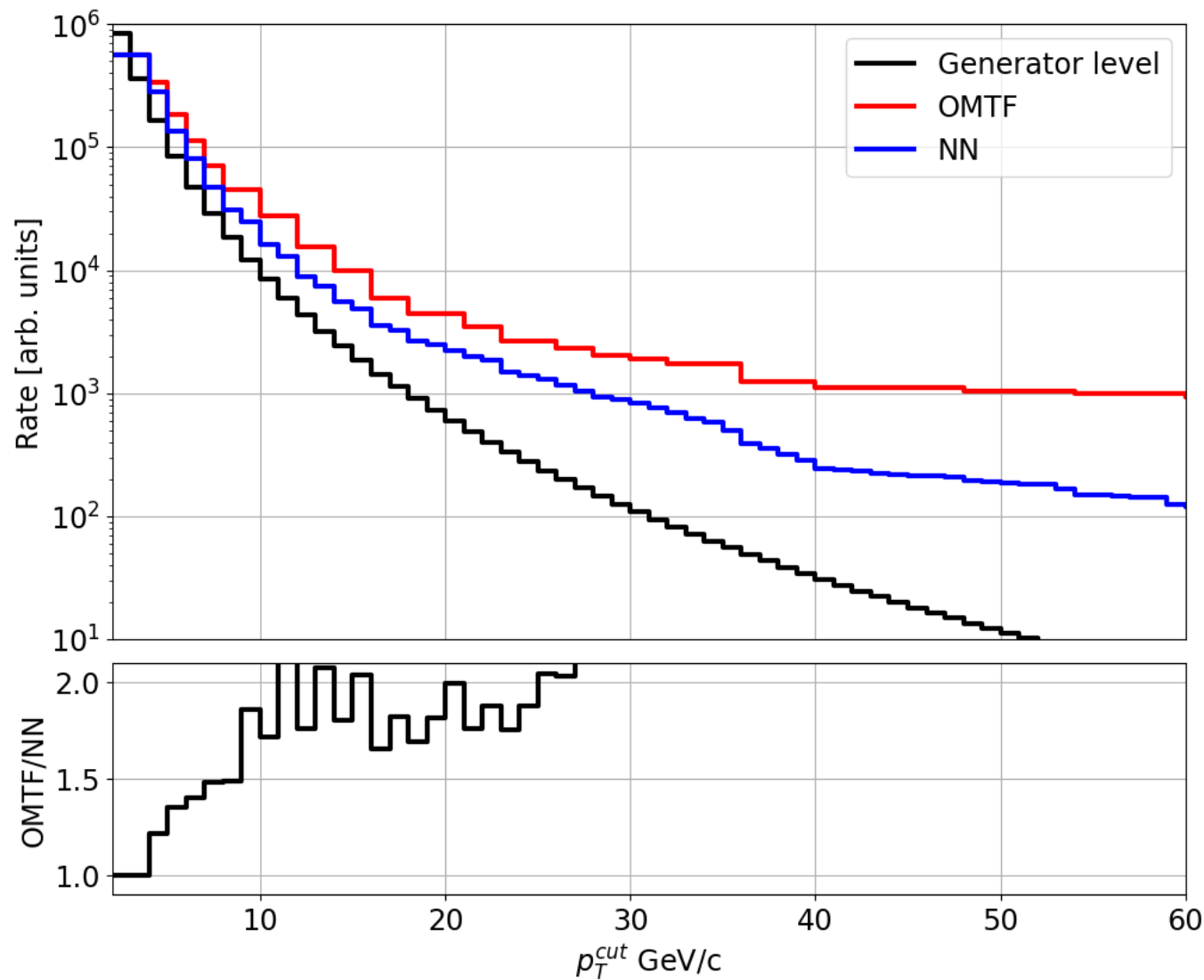


Figure of merit:  
accepted event rate

$$R[\text{ev/s}] = N[\text{ev/s}] : p_T^{\text{measured}} > p_T^{\text{cut}}$$



- Machine learning is not a magic ward – this is yet another technology

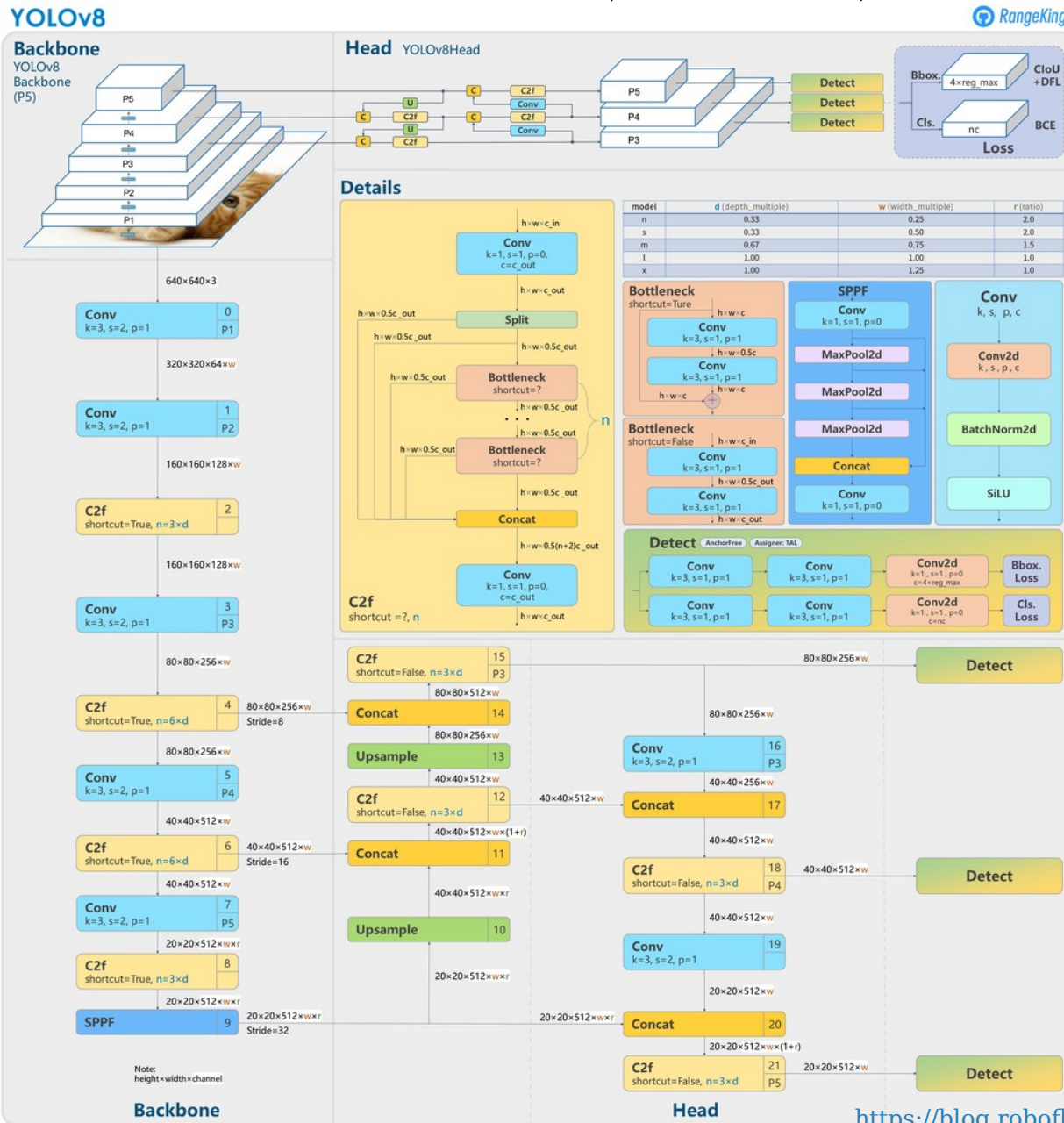
but keep in mind what Arthur C. Clare said: “Any sufficiently advanced technology is indistinguishable from magic.”

- „ordinary” ML users should concentrate on creative problem formulation instead of attempting to invent a new, complicated, architecture

- for some time ML algorithms will require a human assist in result post processing

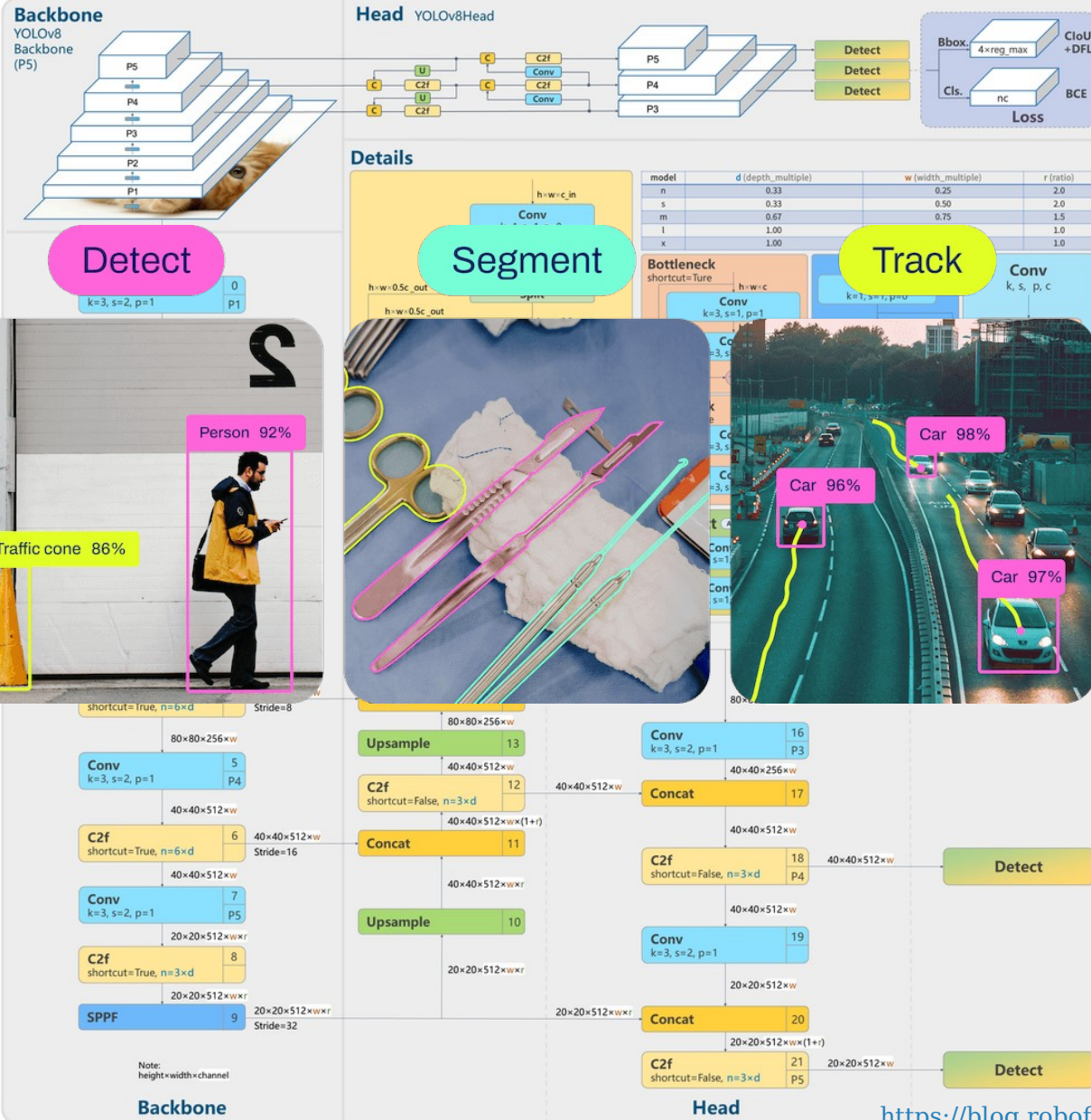
but I do not think this will be longer than LHC Run 5 time scale

Backup slides





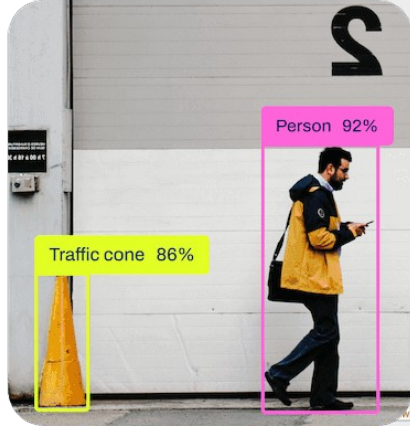
## YOLOv8



Classify



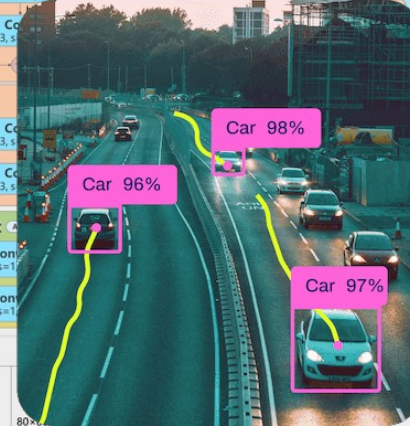
Detect



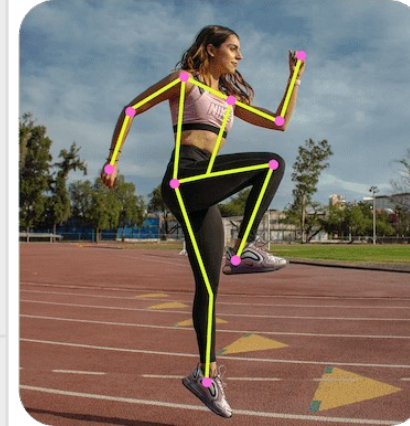
Segment



Track



Pose



<https://blog.roboflow.com/whats-new-in-yolov8/>

## Classification

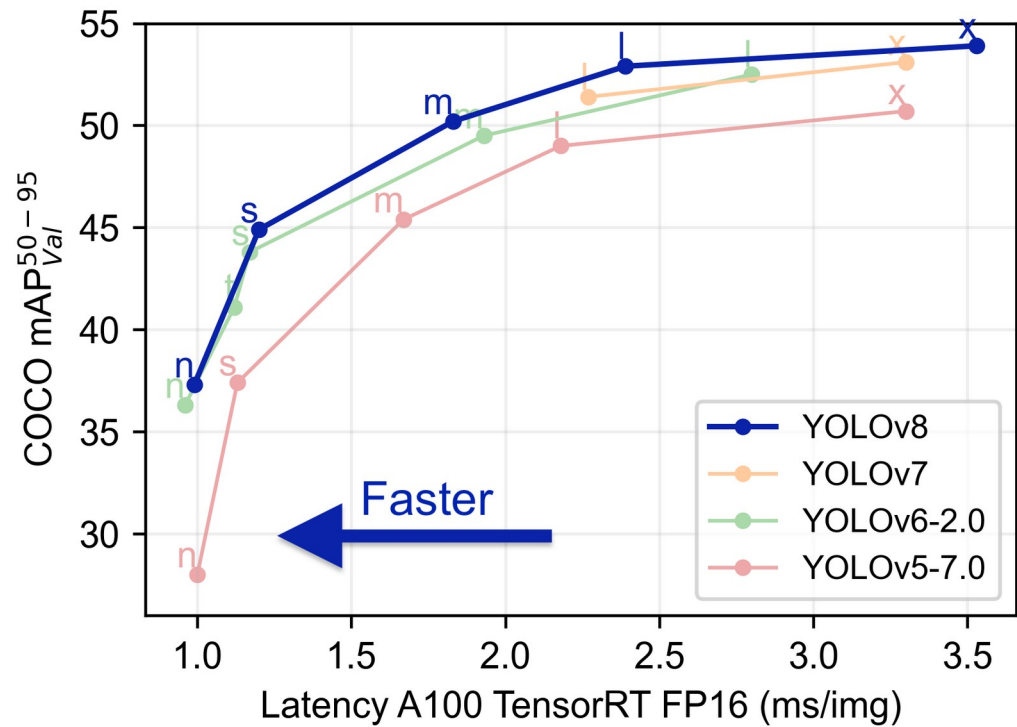
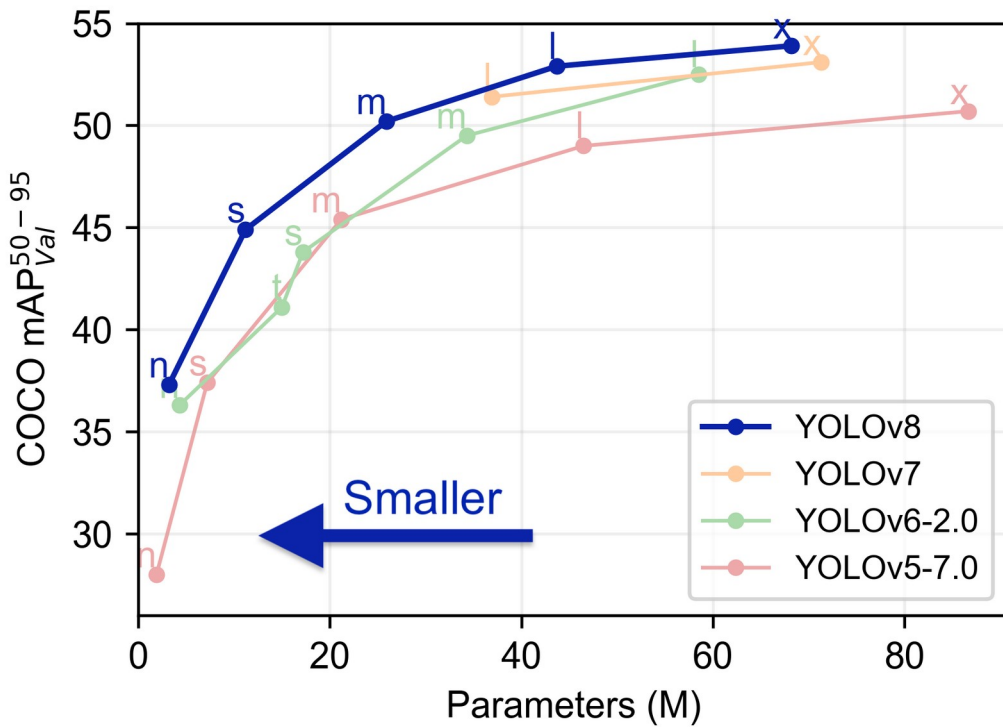
Model	size (pixels)	acc top1	acc top5	Speed CPU ONNX (ms)	Speed A100 TensorRT (ms)	params (M)	FLOPs (B) at 640
<a href="#">YOLOv8n-cls</a>	224	66.6	87.0	12.9	0.31	2.7	4.3
<a href="#">YOLOv8s-cls</a>	224	72.3	91.1	23.4	0.35	6.4	13.5
<a href="#">YOLOv8m-cls</a>	224	76.4	93.2	85.4	0.62	17.0	42.7
<a href="#">YOLOv8l-cls</a>	224	78.0	94.1	163.0	0.87	37.5	99.7
<a href="#">YOLOv8x-cls</a>	224	78.4	94.3	232.0	1.01	57.4	154.8

Analysis of 1M events takes 64 CPU h



## Detection

Model	size (pixels)	mAP <sup>val</sup> 50-95	Speed CPU ONNX (ms)	Speed A100 TensorRT (ms)	params (M)	FLOPs (B)
<a href="#">YOLOv8n</a>	640	37.3	80.4	0.99	3.2	8.7
<a href="#">YOLOv8s</a>	640	44.9	128.4	1.20	11.2	28.6
<a href="#">YOLOv8m</a>	640	50.2	234.7	1.83	25.9	78.9
<a href="#">YOLOv8l</a>	640	52.9	375.2	2.39	43.7	165.2
<a href="#">YOLOv8x</a>	640	53.9	479.1	3.53	68.2	257.8



**Task:** cell colony counting

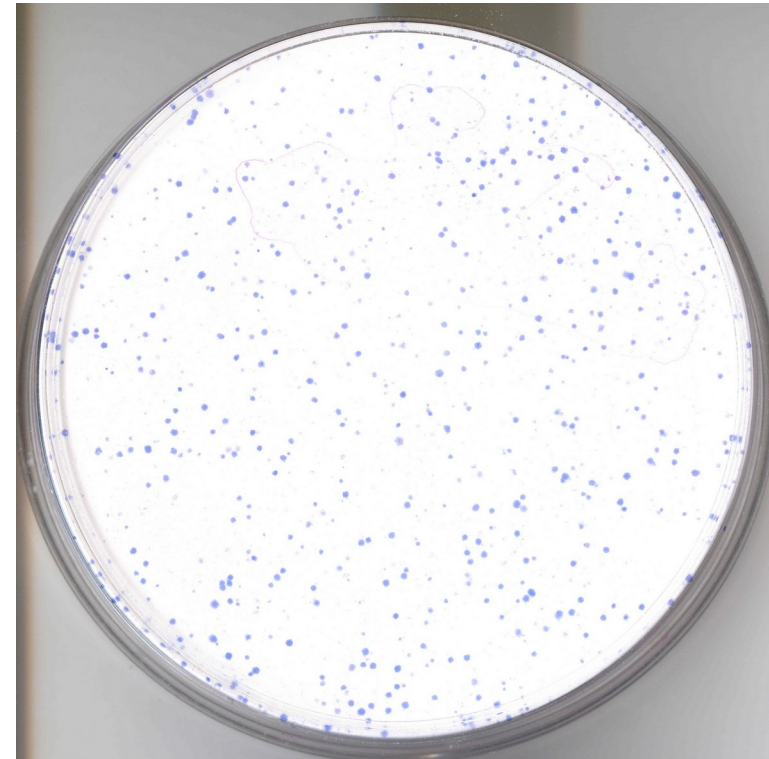
## Input data:

- 57 Petri dish high resolution (4390x5059) photos

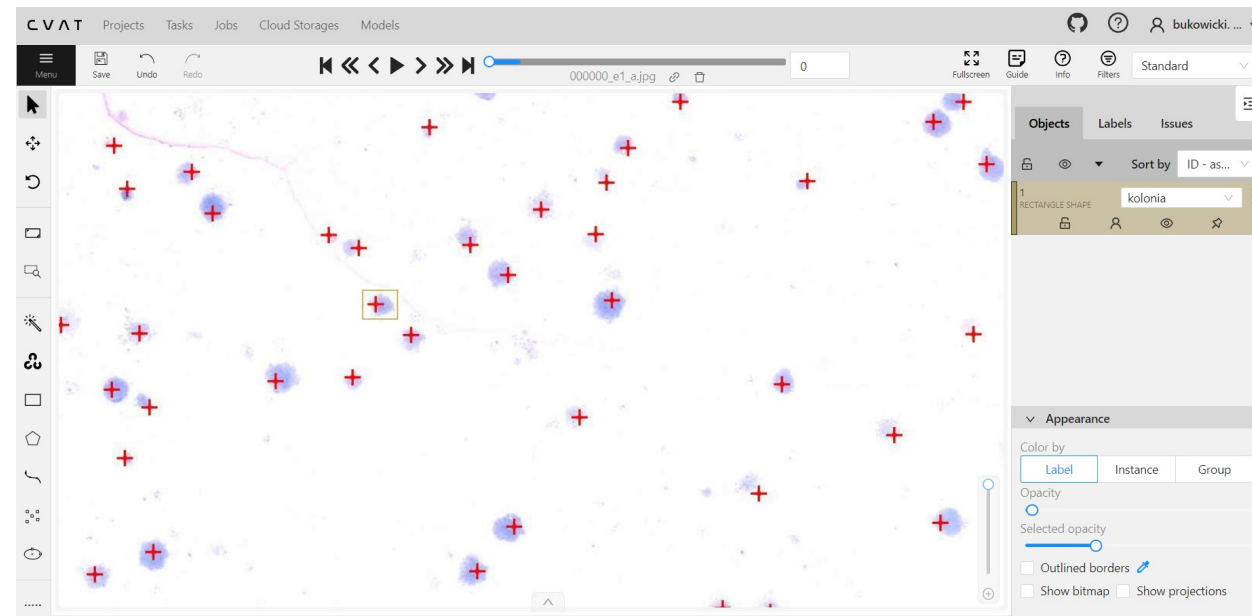
## Datasets:

**Training and validation:** 243 + 73 slices  
of 640x640

**Testing:** 32 slices



- labeling



- training (single GPU, Google Colaboratory)

```
# Build a new model from YAML, transfer pretrained weights to it and start training  
yolo detect train data=coco128.yaml model=yolov8n.yaml pretrained=yolov8n.pt epochs=100 imgsz=640
```



- prediction with a 640x640 window



Zastosowanie uczenia maszynowego do automatycznej analizy testu klonogenego przeprowadzonego na komórkach ssaków

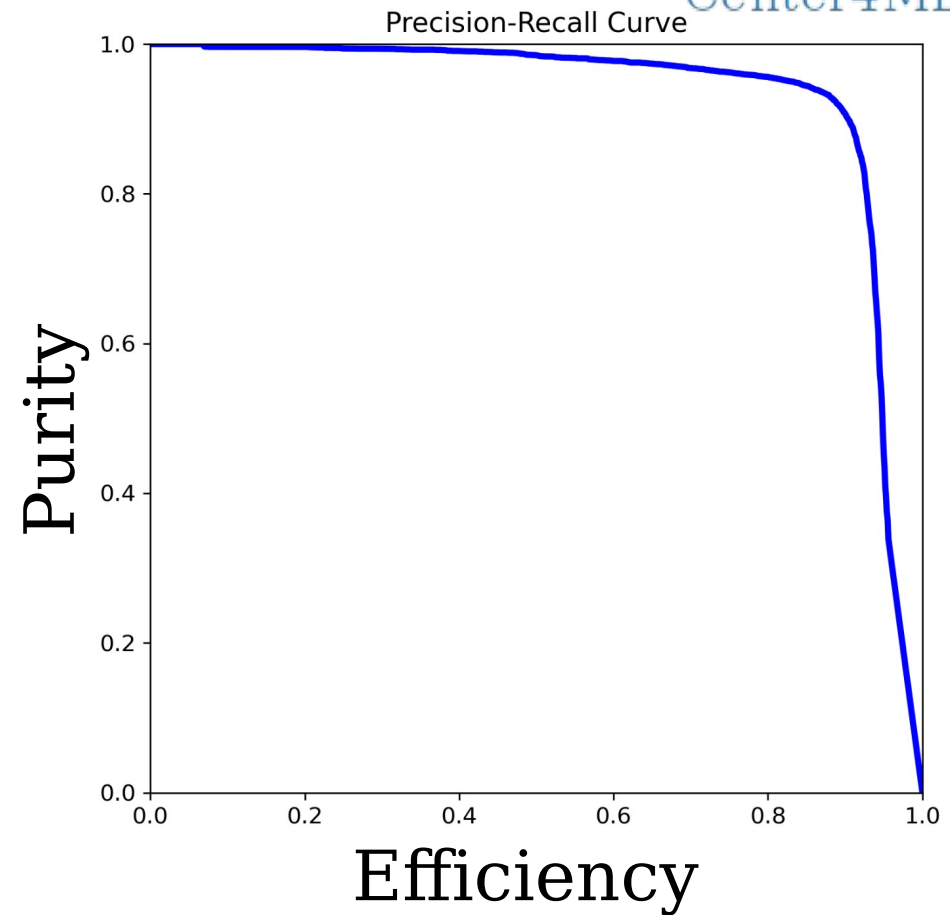


- merged image



- the ROC is close to ideal

YOLO provides a convenient user interface easing the training and prediction steps



**Center4ML is a IDUB funded service that provide a ready to use ML models for any UW based group.**

**The service is free of charge.**



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