

# Characterization of the Hamamatsu R12699-406-M4 2-inch Photomultipliers

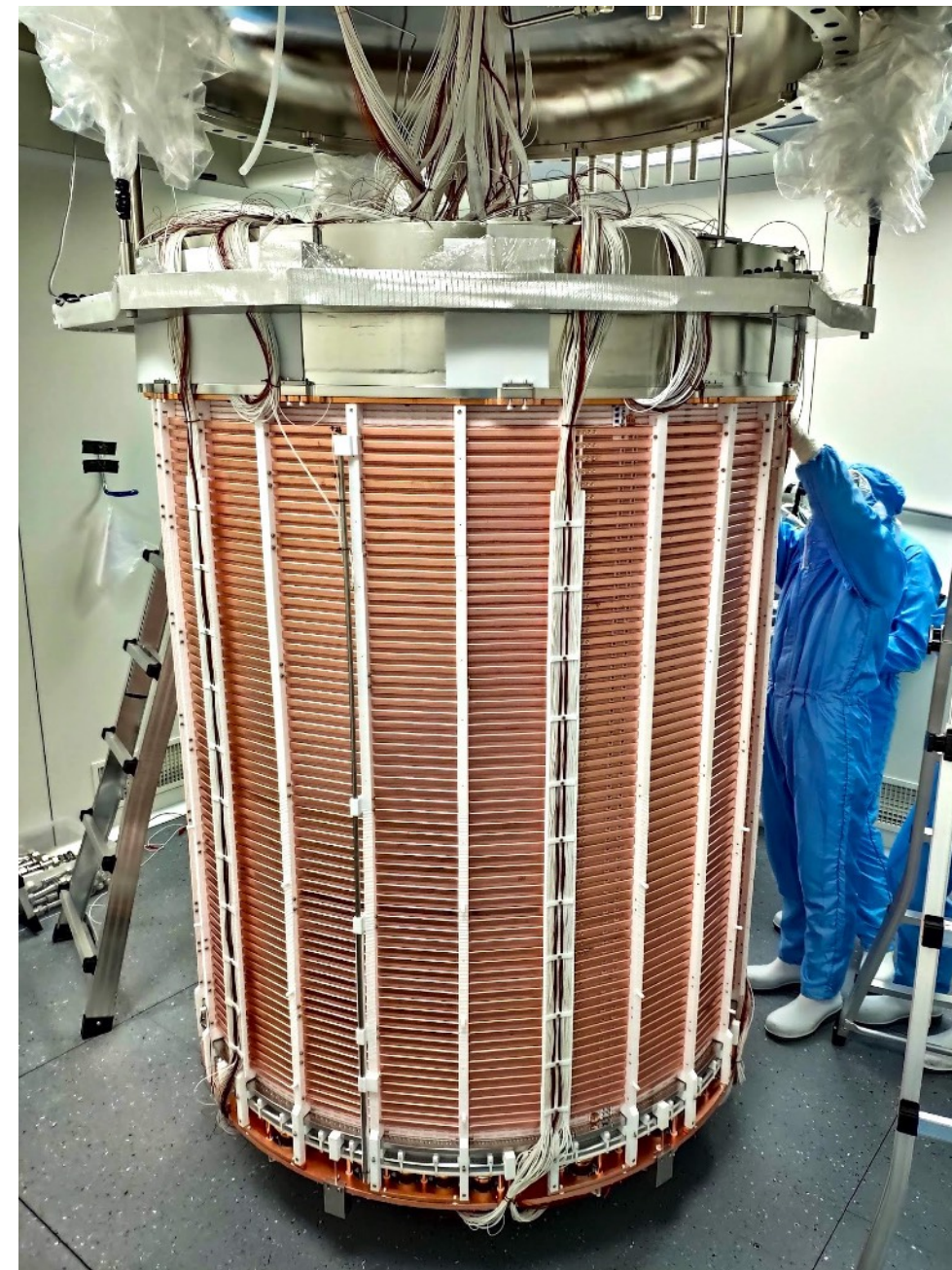
in MarmotX and XAMS

# Rare event searches

## Xenon time projection chambers

**WIMP DM**  
 Spin-dependent  
 Spin-independent  
 Light WIMPs  
*and more*

**Other DM candidates**  
 Dark photons  
 Axion-like particles  
*and more*



**Neutrinos**  
 Atmospheric & solar neutrinos  
 Neutrino Magnetic Moment  
 Super Nova (Early Warning System)  
 B8 CE $\nu$ NS

**Weak decays**  
 Double Electron Capture  
 Spectral shape measurements  
 $0\nu\beta\beta$  &  $2\nu\beta\beta$

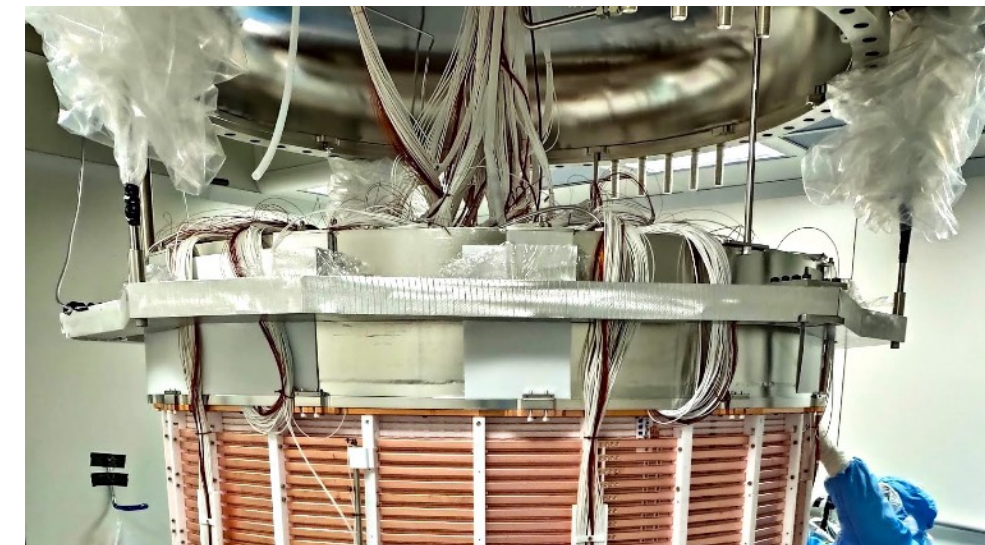
XENON10	XENON100	XENON1T	XENONnT (present)	DARWIN/ XLZD (future)
14kg	62kg	~2 t	~6 t	~37 t



# Rare event searches

## Xenon time projection chambers

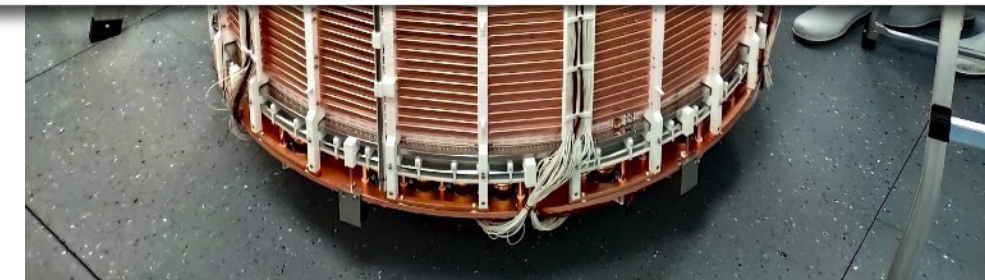
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**Neutrinos**  
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**Low energy  $\rightarrow$  few keV to tens of MeV**

**Other DM candidates**  
Dark photons  
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*and more*



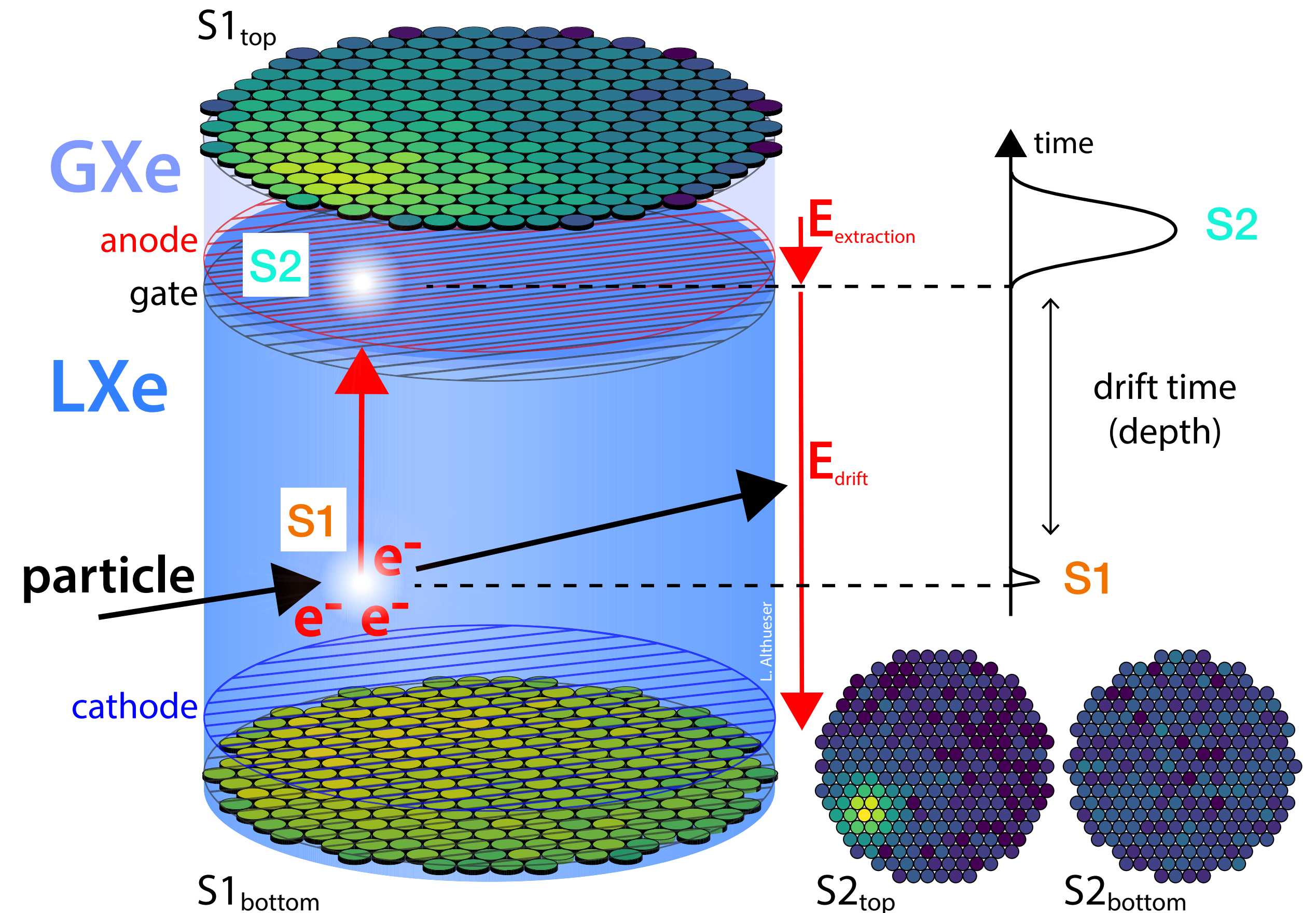
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# Dual-phase xenon TPC

## Working principle

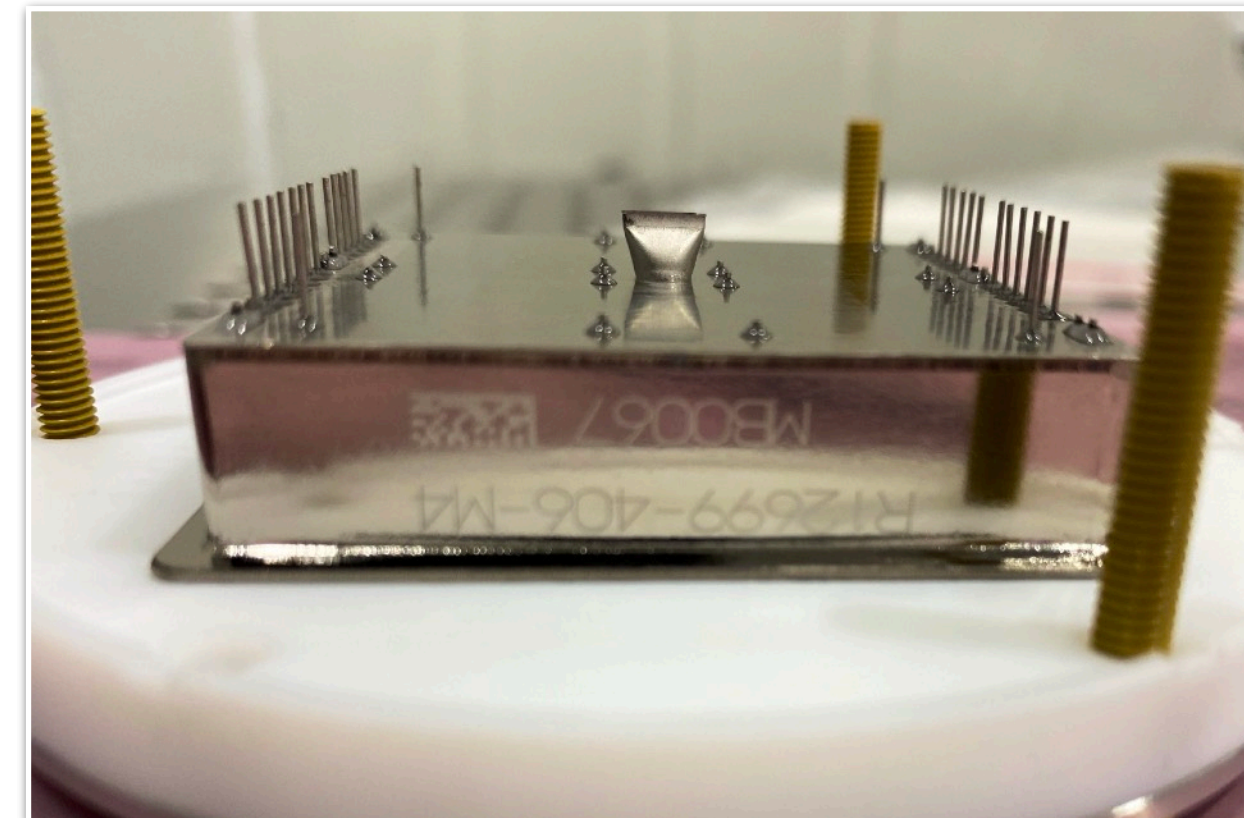
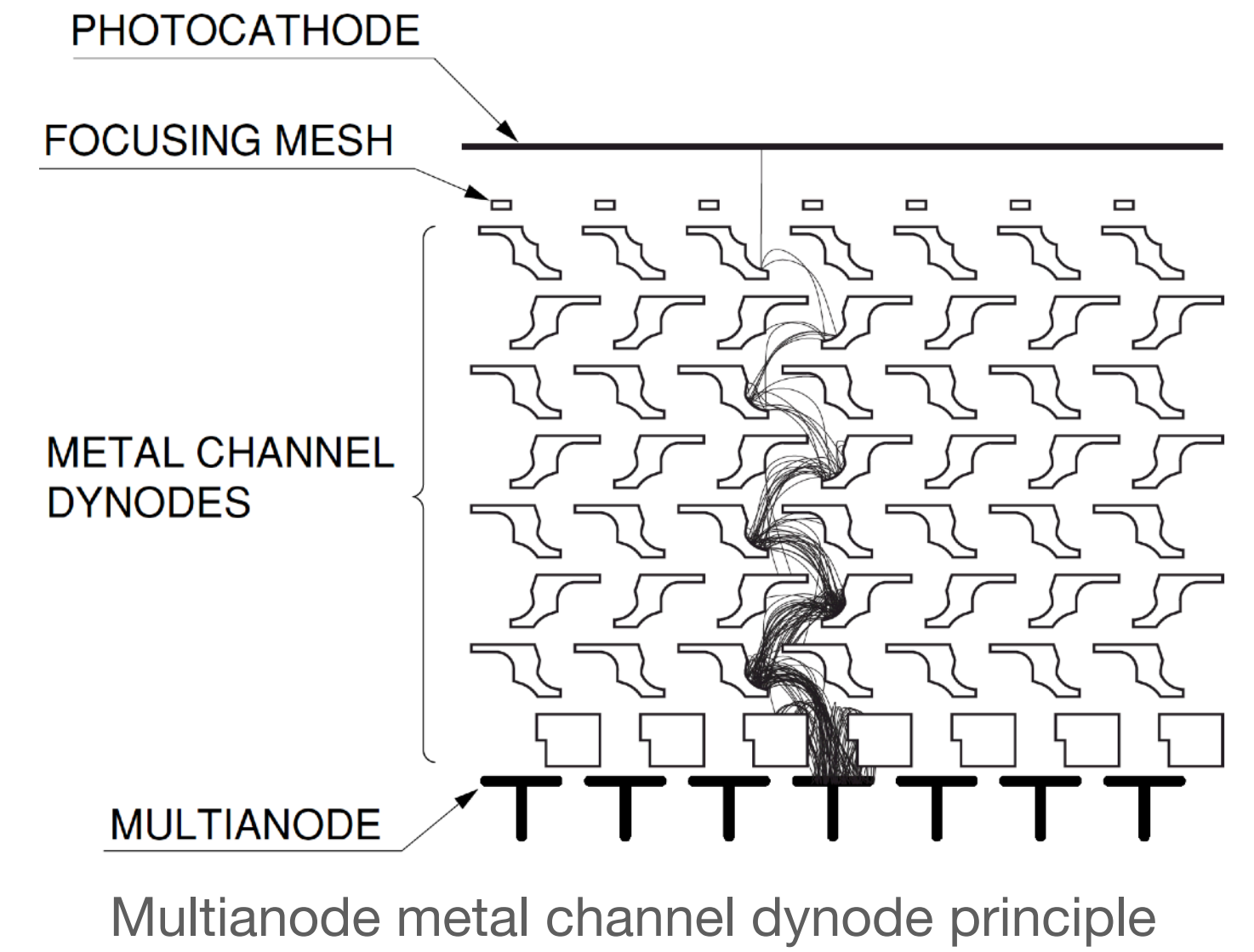
- ▶ Readout of both scintillation and ionization signals
- ▶ Prompt scintillation light: **S1**
- ▶ Secondary (proportional) scintillation light: **S2**
- ▶ Reconstruction of
  - ▶ 3D position ( $x, y, z$ )
  - ▶ Energy
  - ▶ Interaction type (ER/NR) through **S1/S2** ratio
- ▶ Self-shielding → fiducial volume



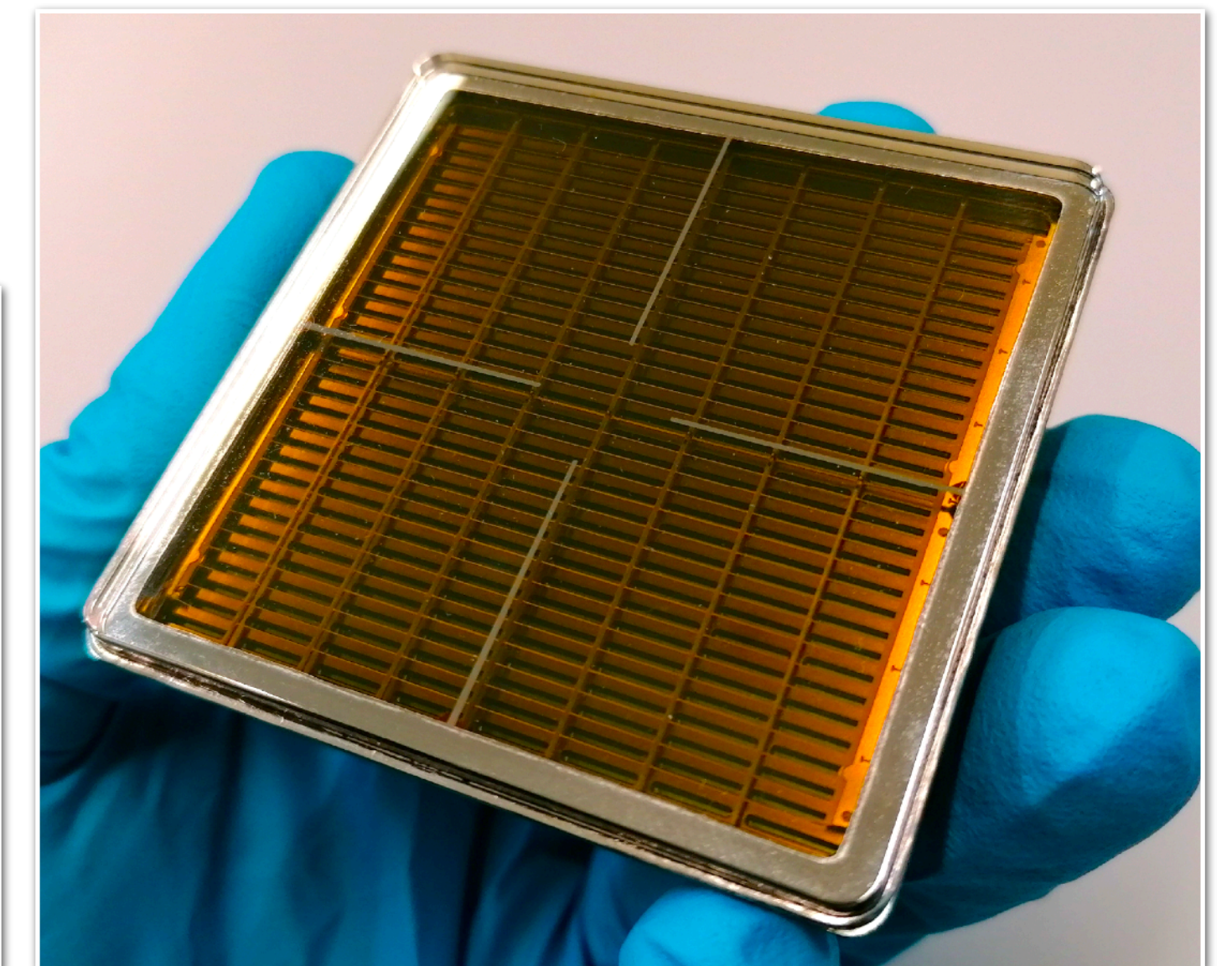


# R12699-406-M4 2-inch PMT

What makes this model interesting?



Low profile PMT



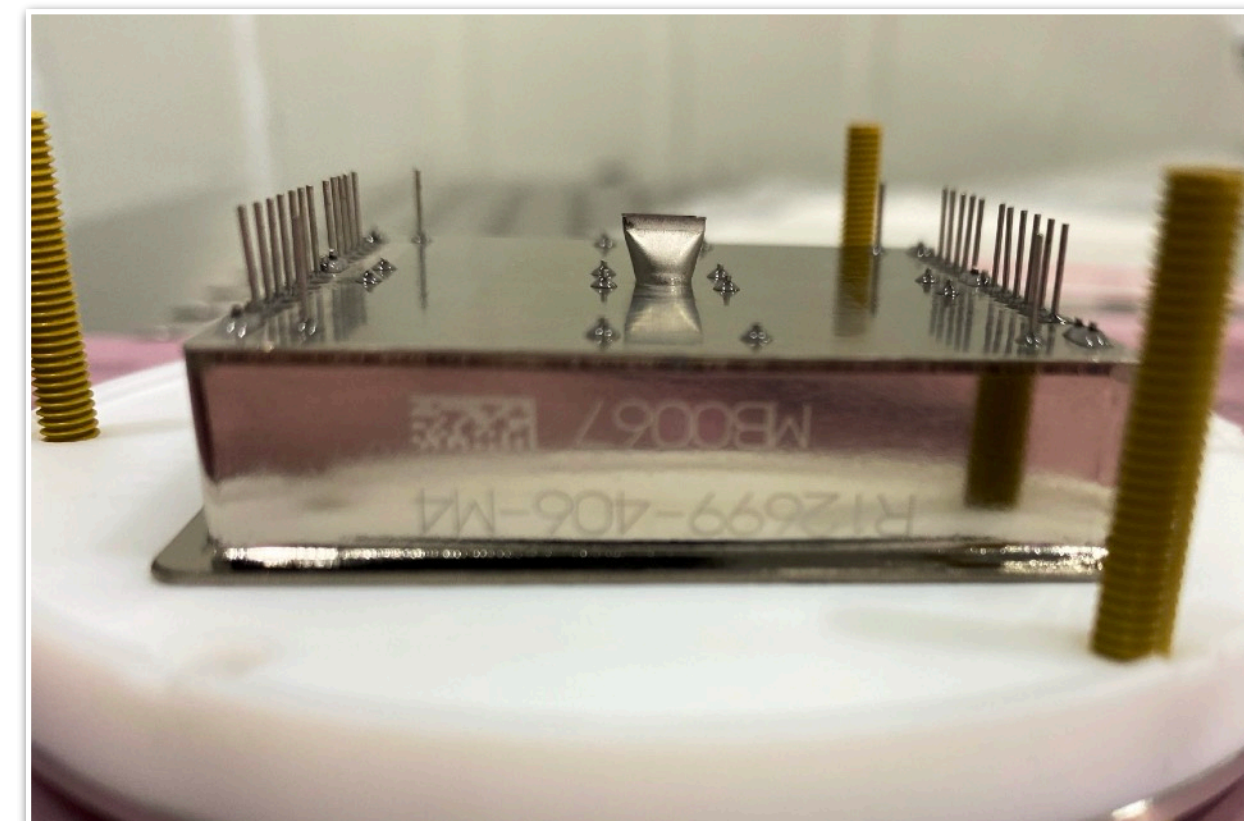
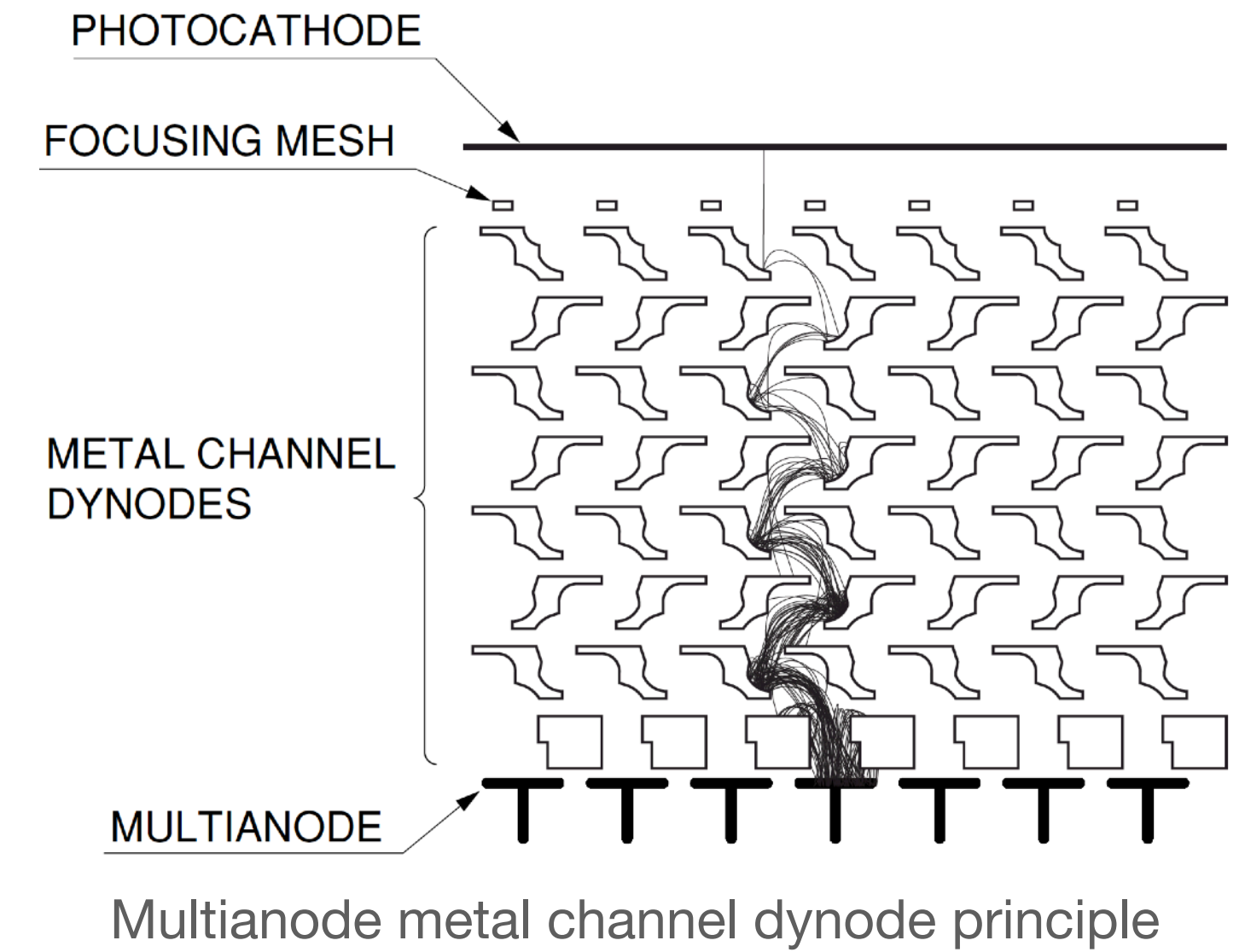
Hamamatsu R12699 M4 2-inch PMT 52 x 52 x 15 mm



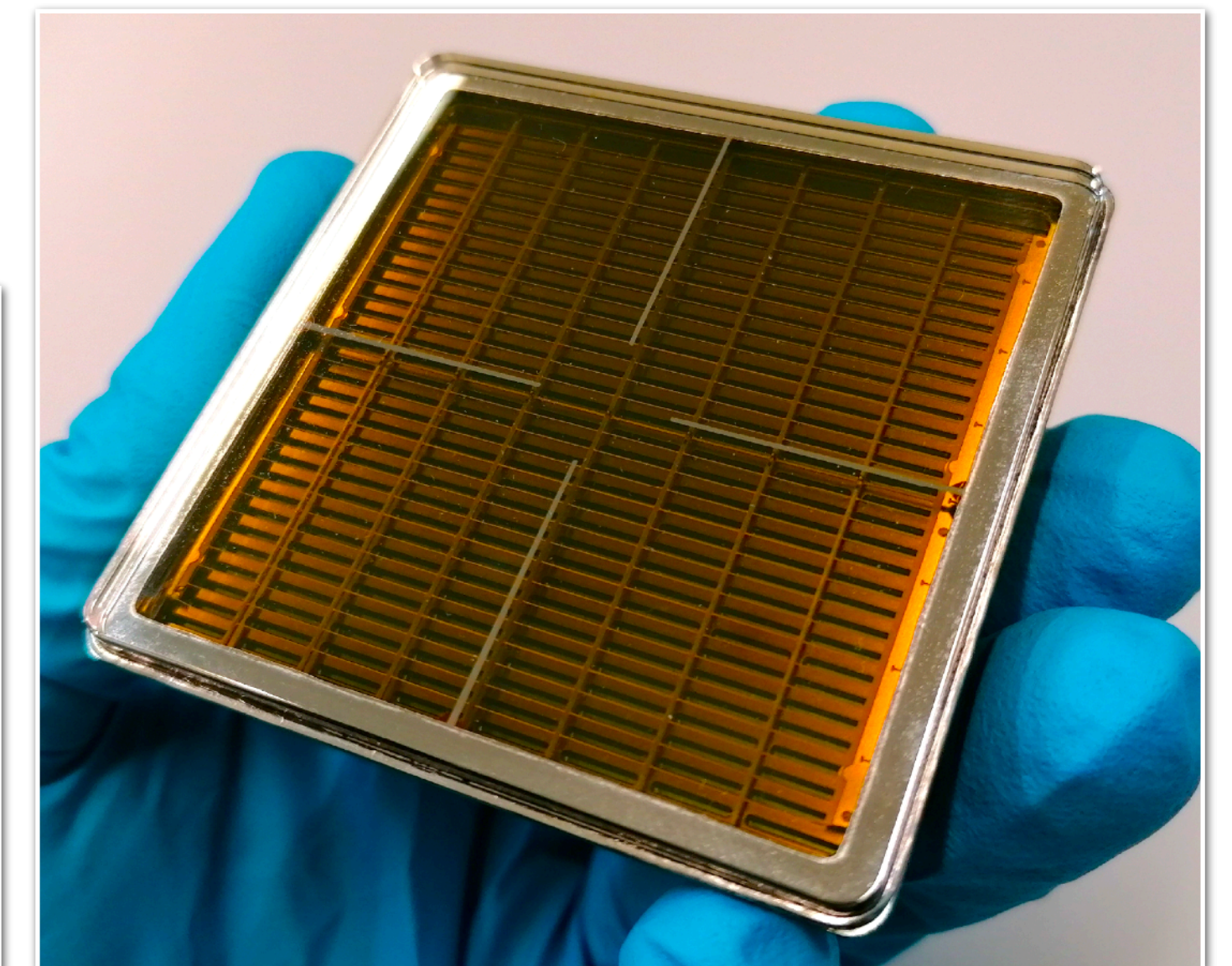
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## What makes this model interesting?

- ▶ Low profile
  - ▶ Low buoyancy
  - ▶ Sub-ns rise-time and transit-time spread (TTS) (i.e. very fast)



Low profile PMT



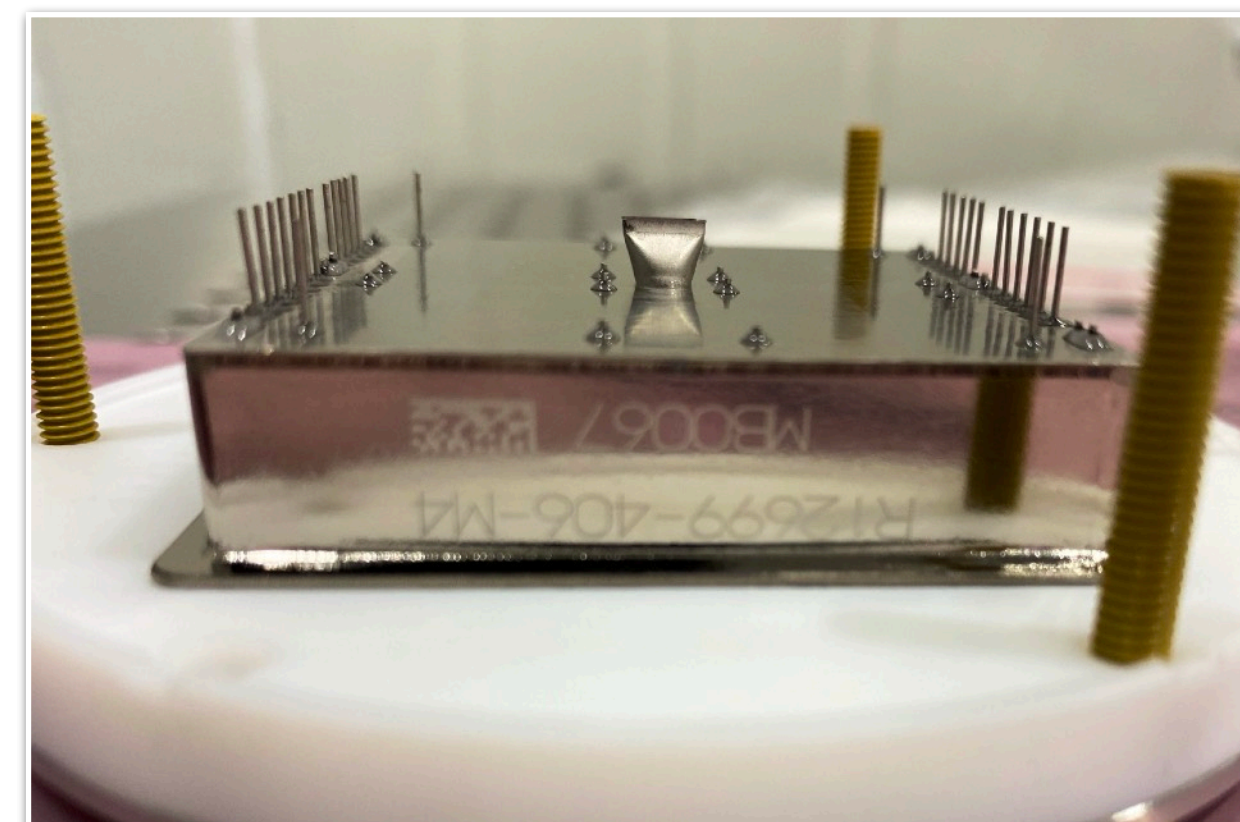
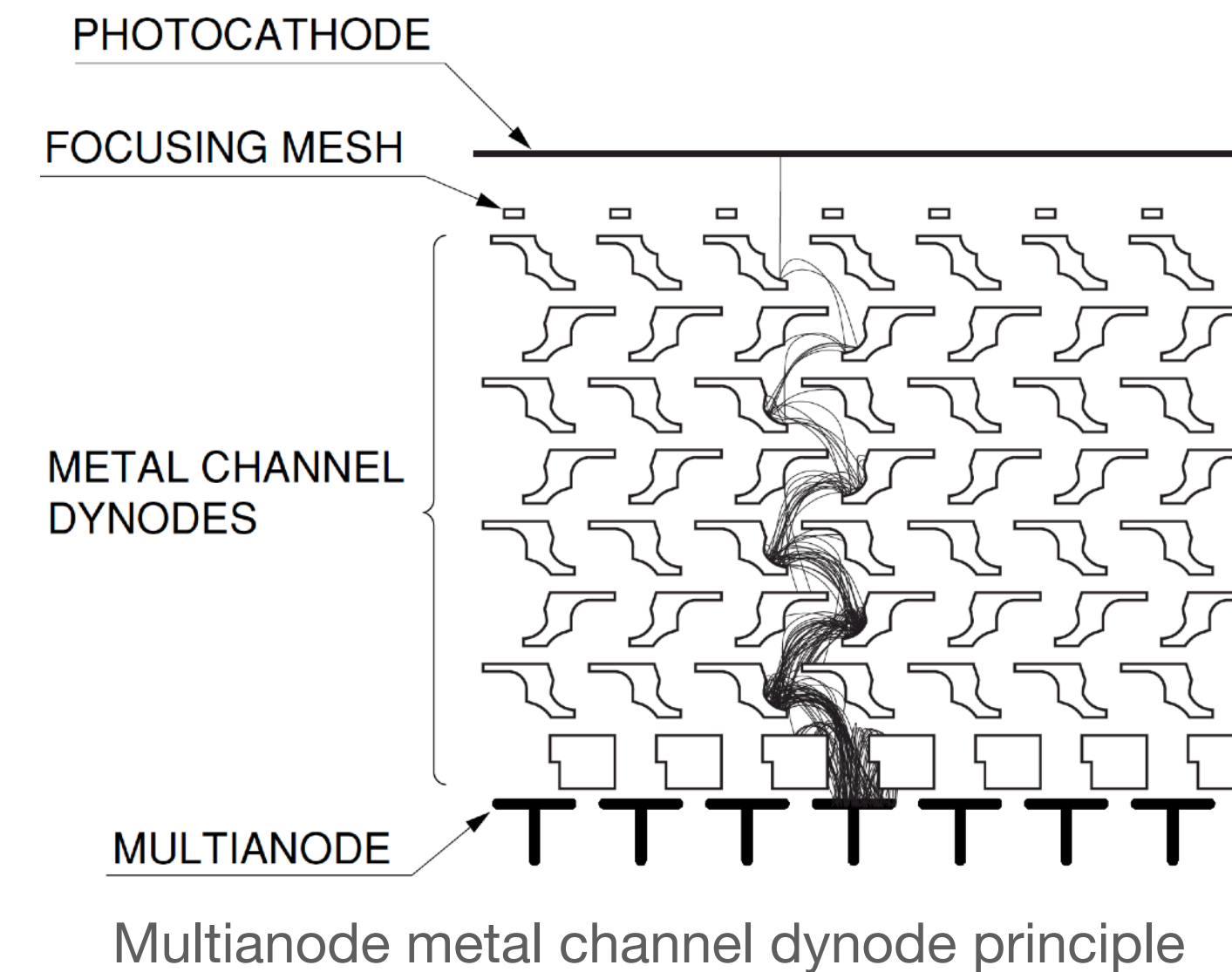
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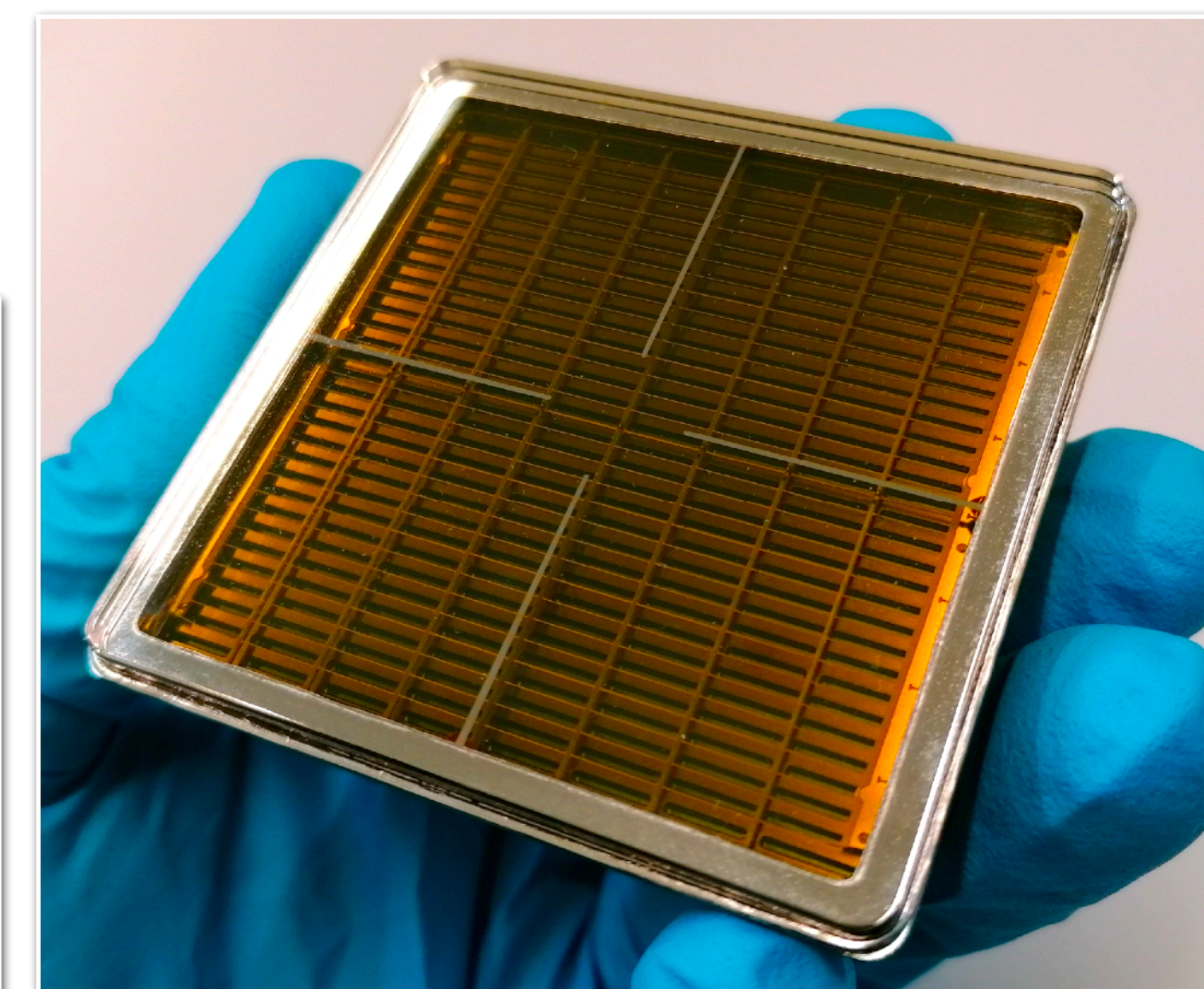
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- ▶ Multi-anode readout
  - ▶ Less HV cables per channel
  - ▶ Variable granularity



Low profile PMT



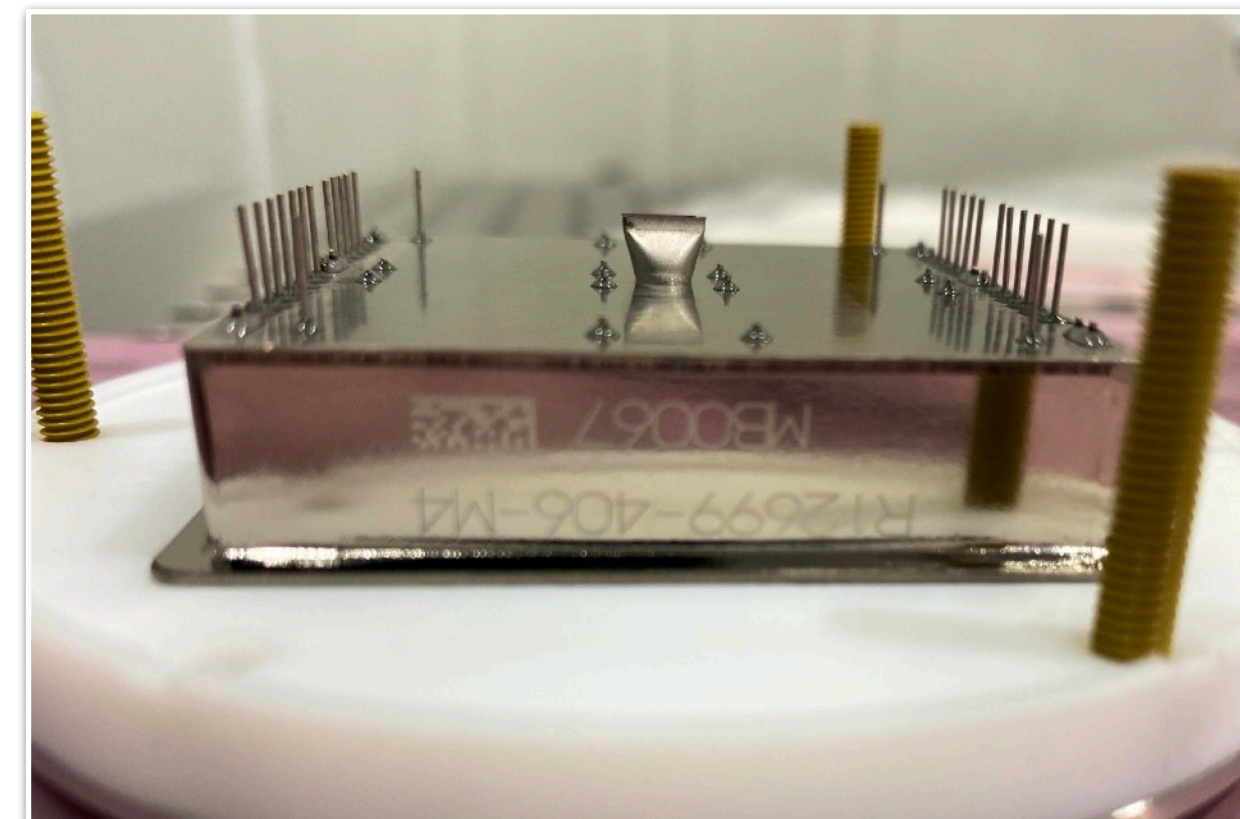
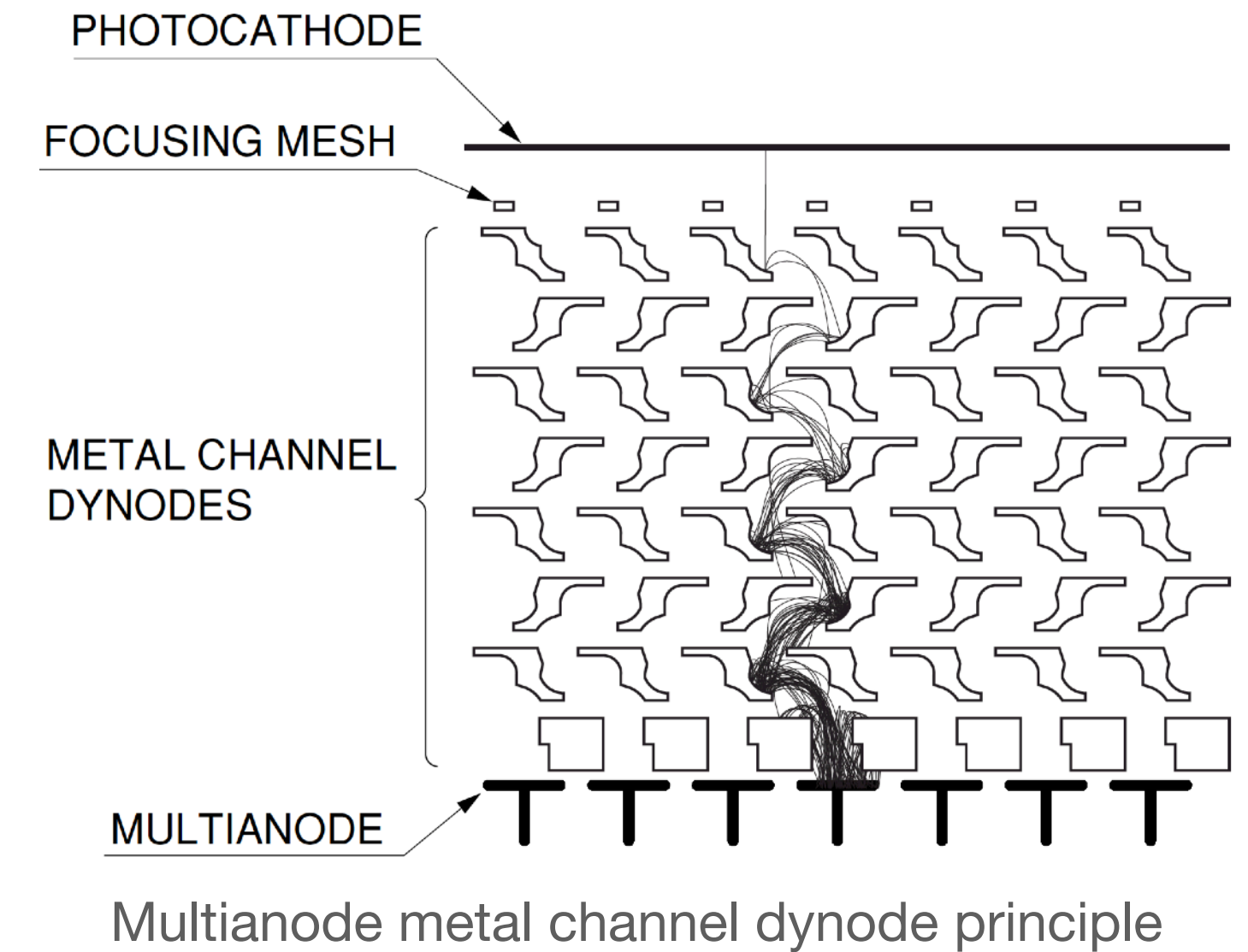
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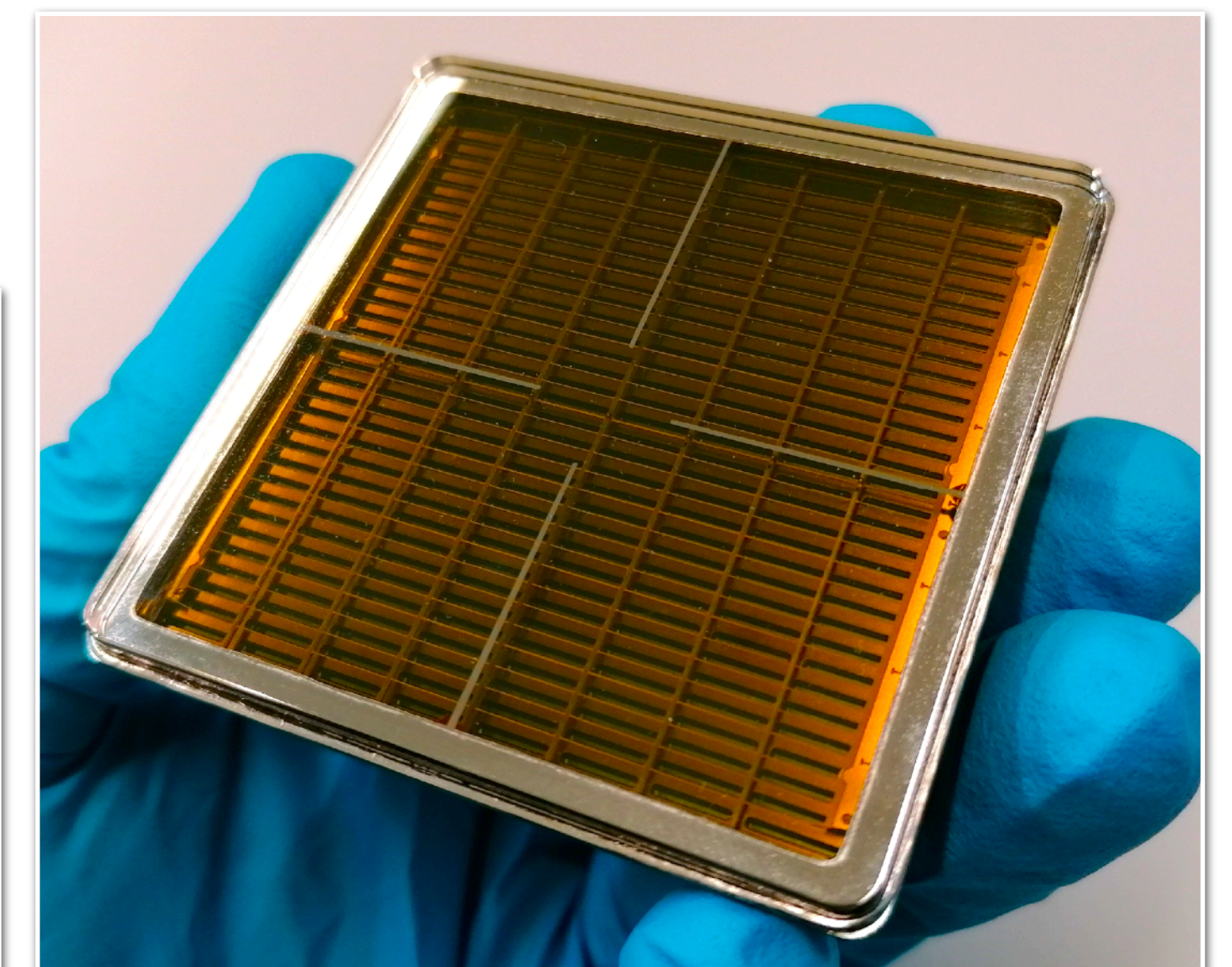
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  - ▶ Variable granularity
- ▶ 75% photocathode coverage
- ▶ QE of 33% (similar to 32.5% of R11410-21 XENONnT PMTs)
- ▶ Improved radioactivity\*



Low profile PMT



Hamamatsu R12699 M4 2-inch PMT 52 x 52 x 15 mm

\* See XeSAT2024 [presentation](#) by Y. Meng (PandaX)

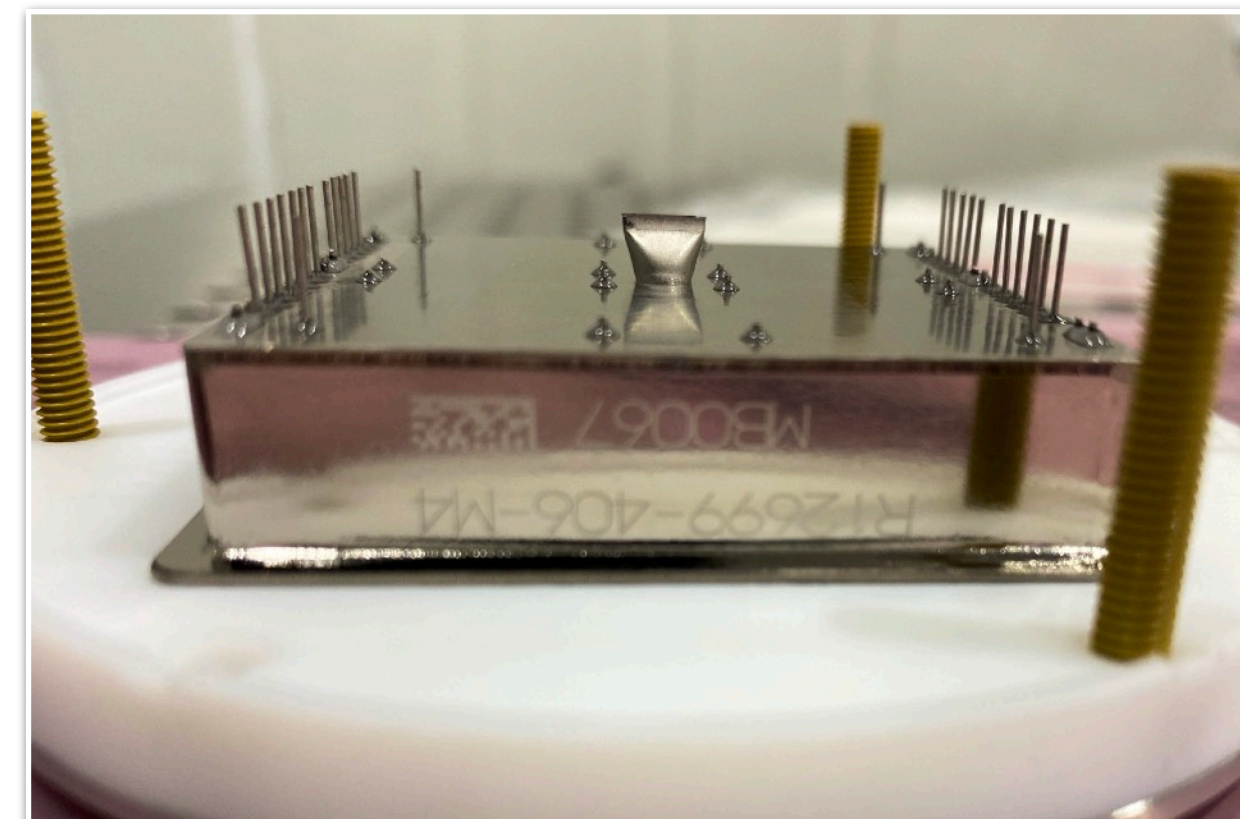
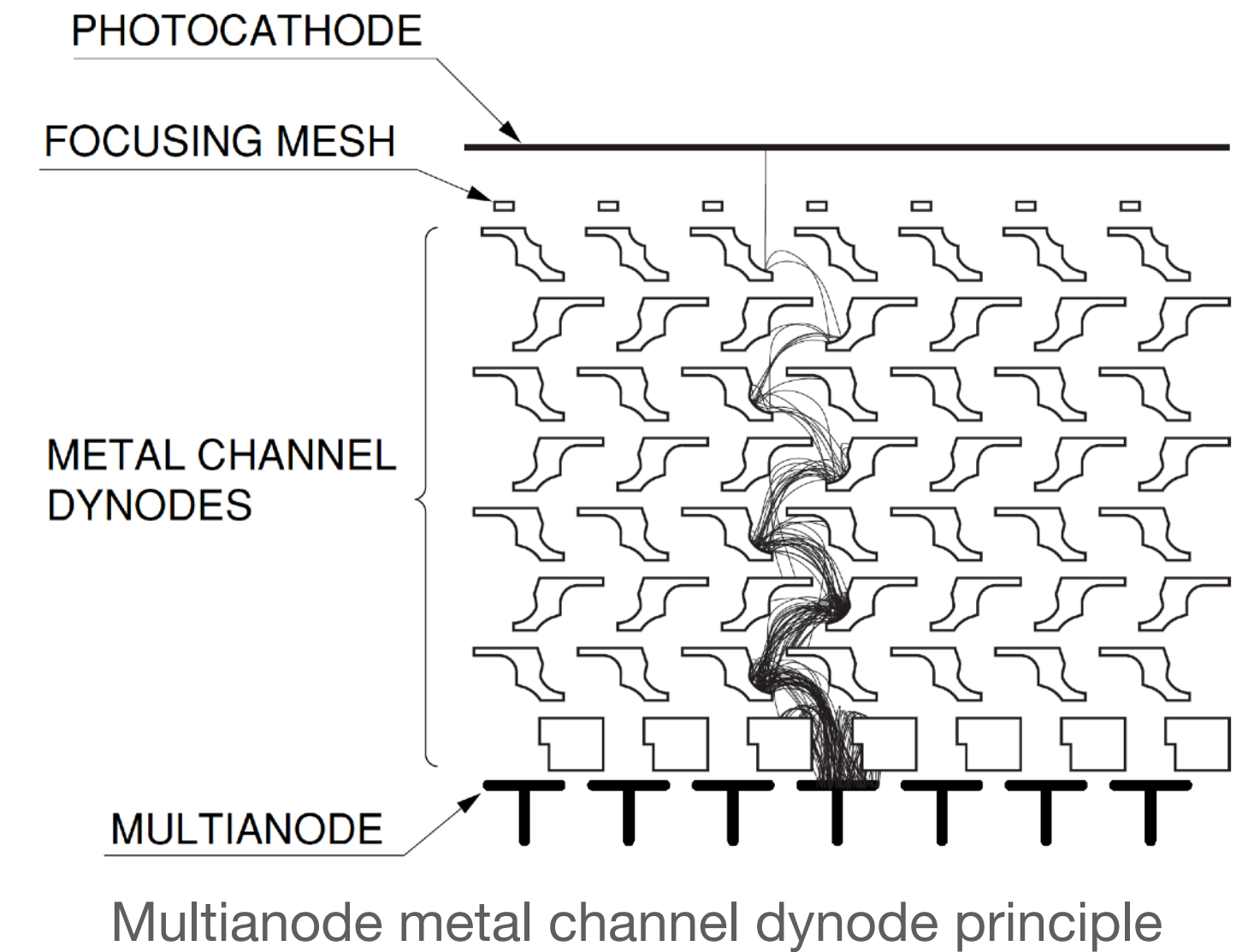


# R12699-406-M4 2-inch PMT

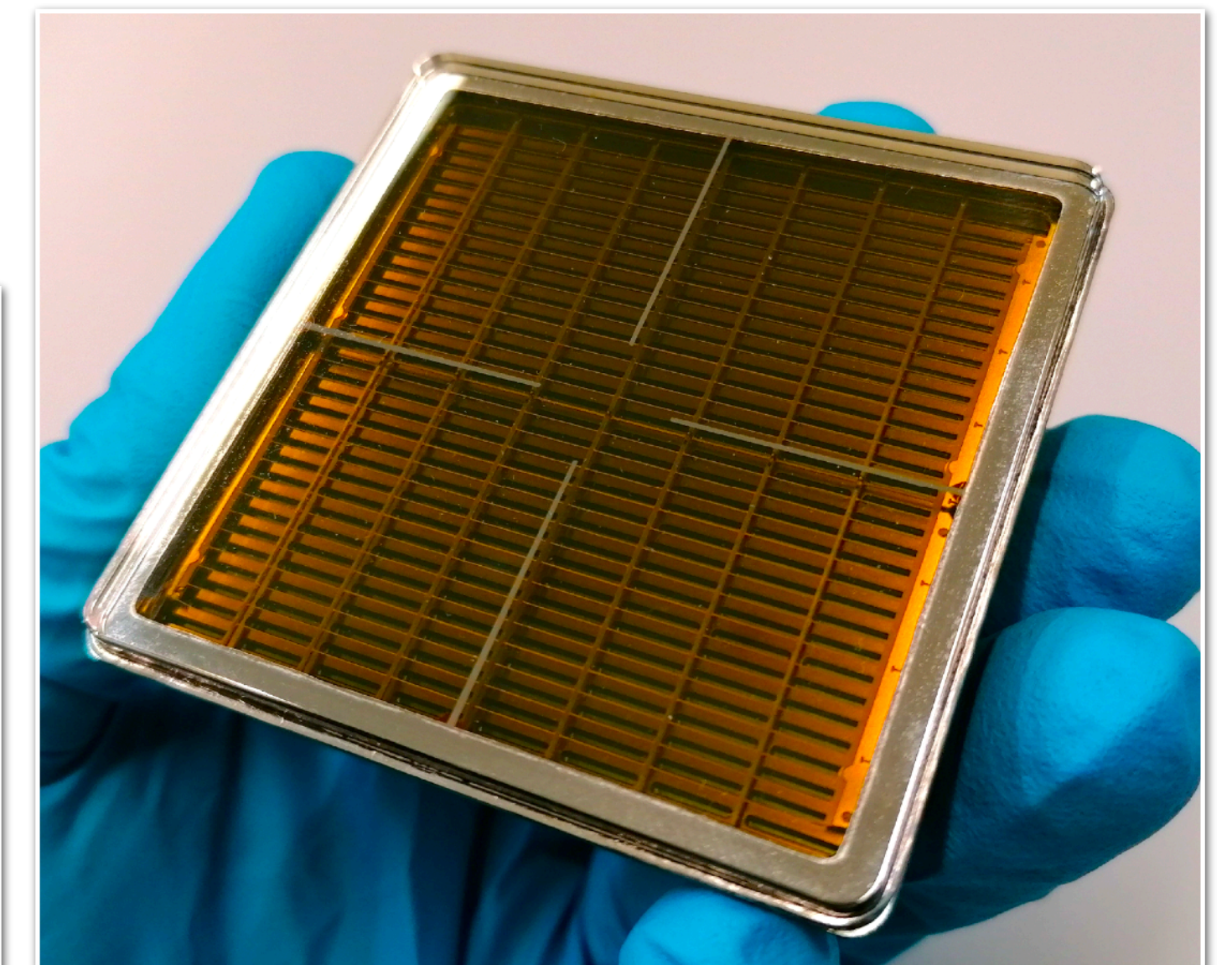
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From now on “2-inch PMTs”



Low profile PMT



Hamamatsu R12699 M4 2-inch PMT 52 x 52 x 15 mm

\* See XeSAT2024 [presentation](#) by Y. Meng (PandaX)



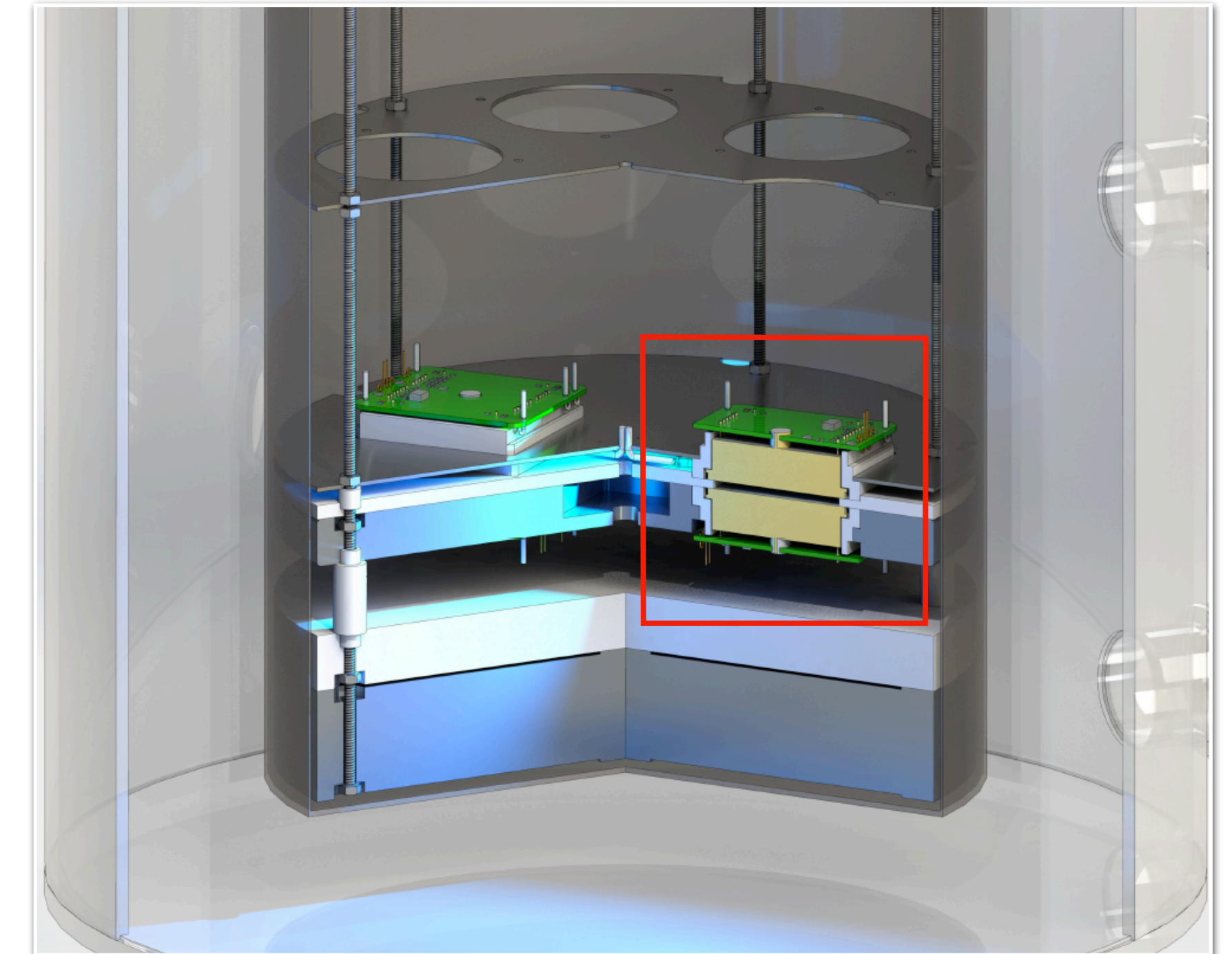
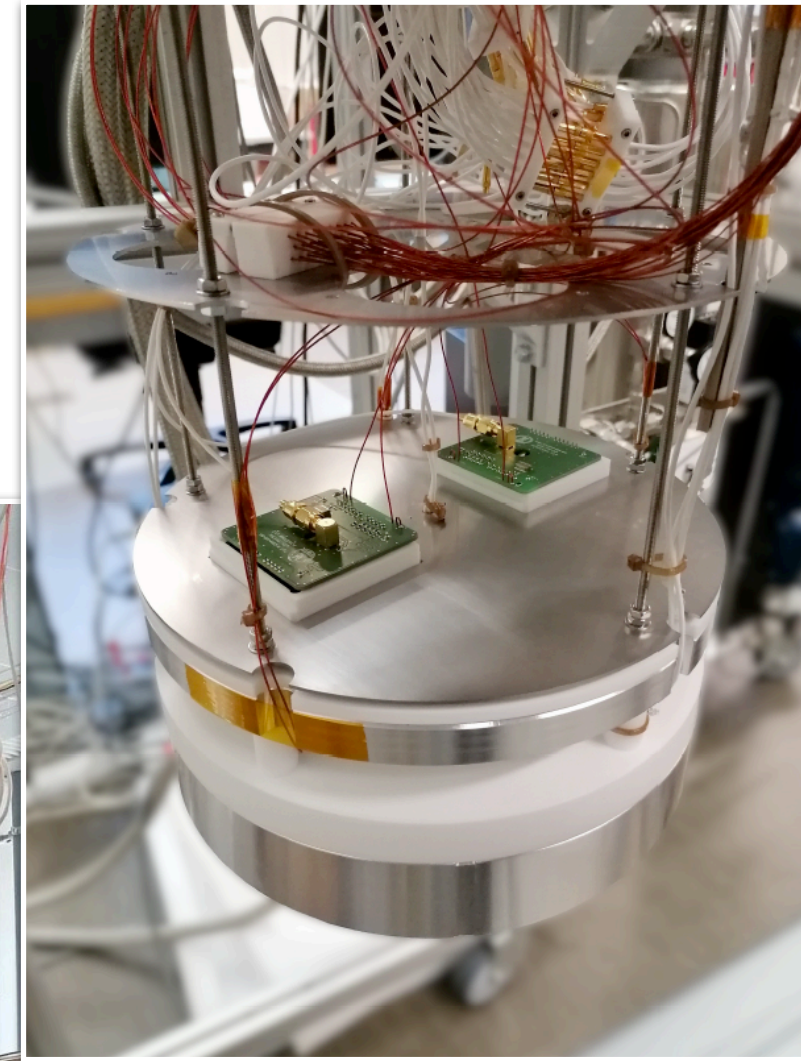
# MarmotX

Universität  
Zürich

## Photosensor R&D



Full MarmotX setup



MarmotX cryostat with two pairs of face-to-face 2-inch PMTs

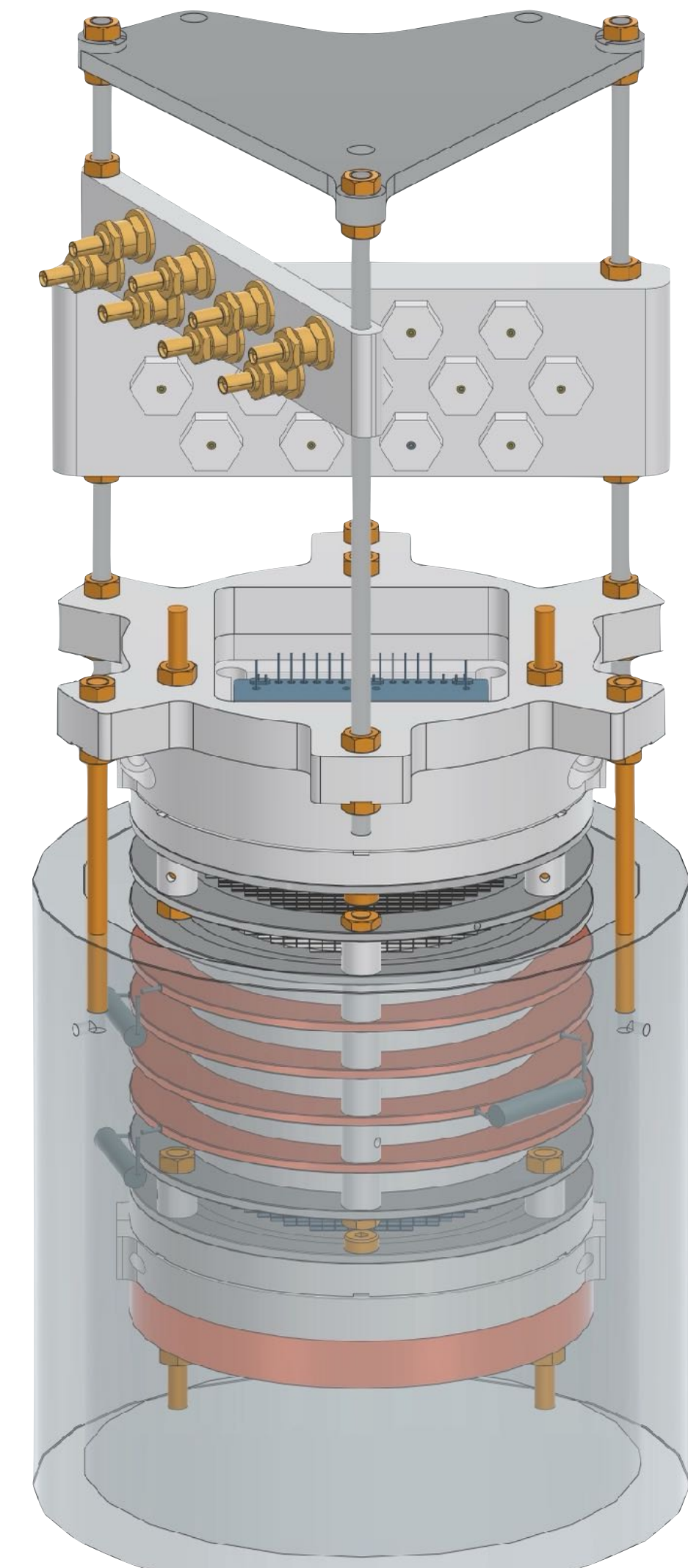
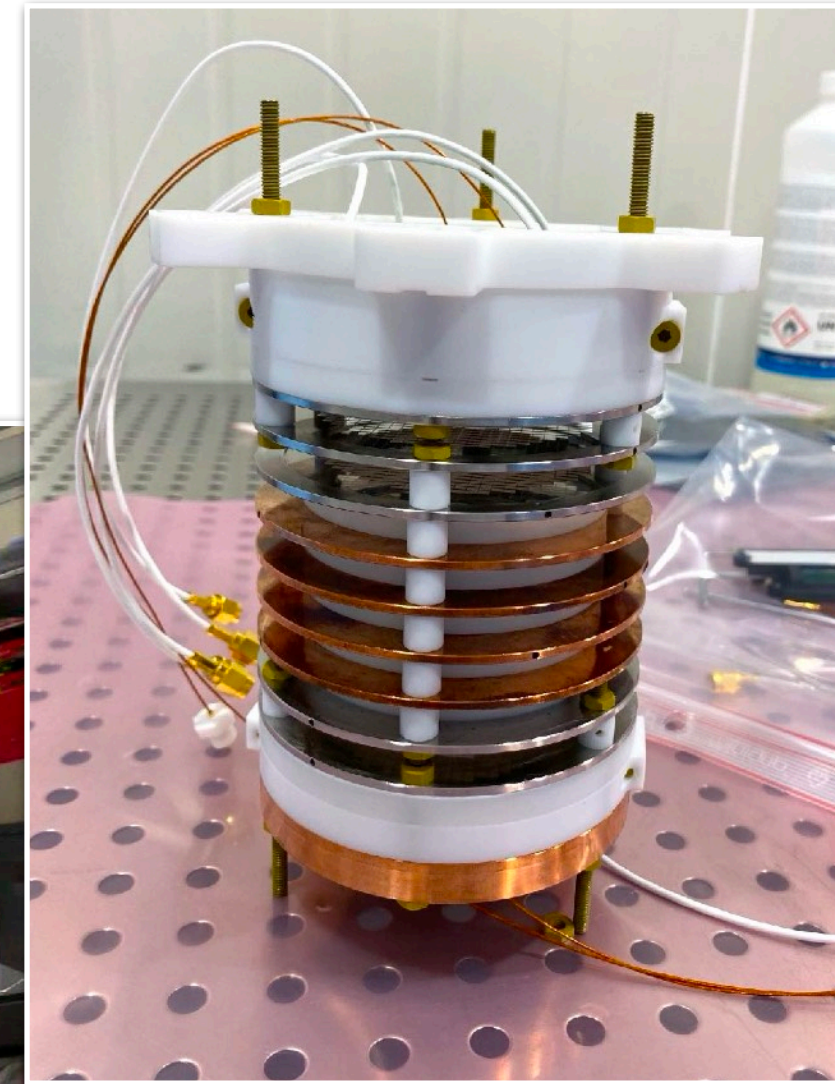


# XAMS

## Dual-phase xenon TPC



Full XAMS setup



XAMS dual-phase xenon TPC with two 2-inch PMTs installed



# Results

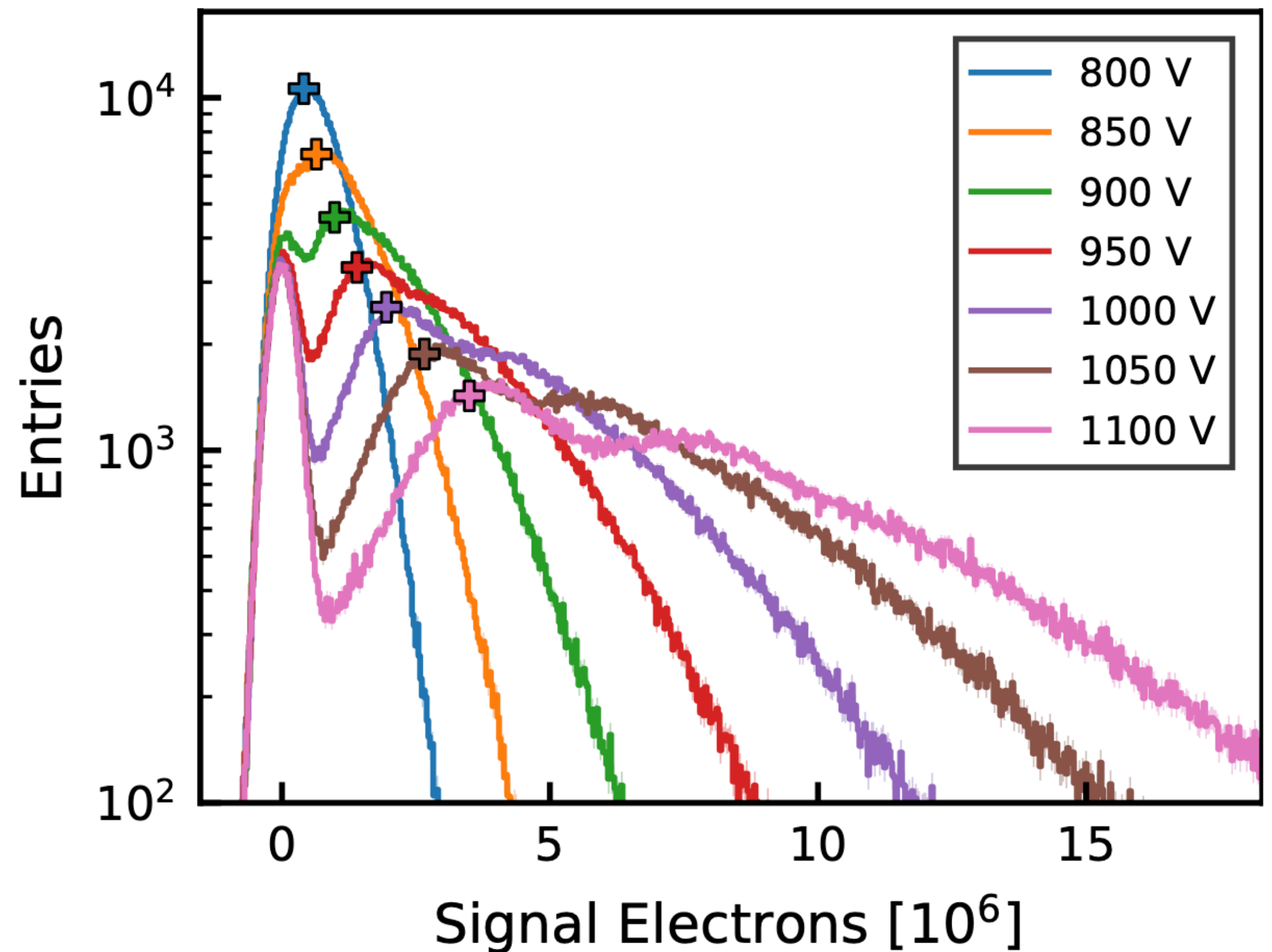
SPE response | Dark counts | Afterpulsing | Position reconstruction



# Characterization results

## Single photoelectron response

- ▶ Model independent approach as proposed by Saldanha et al. (2016)<sup>1</sup>
- ▶ SPE resolution  $\sigma_{1PE}/\mu_{1PE}$  factor  $1.23 \pm 0.14$  higher than R11410-21 PMTs
- ▶ Typical gain of  $2 \cdot 10^6$  exceeded for each PMT at nominal voltage ( $\sim 1000V$ )
- ▶ Long term gain stability tests ongoing



<sup>1</sup>Saldanha, R., Grandi, L., Guardincerri, Y., & Wester, T. (2017). Model independent approach to the single photoelectron calibration of photomultiplier tubes. *Nuclear Instruments and Methods in Physics Research Section A: Accelerators, Spectrometers, Detectors and Associated Equipment*, 863, 35-46.

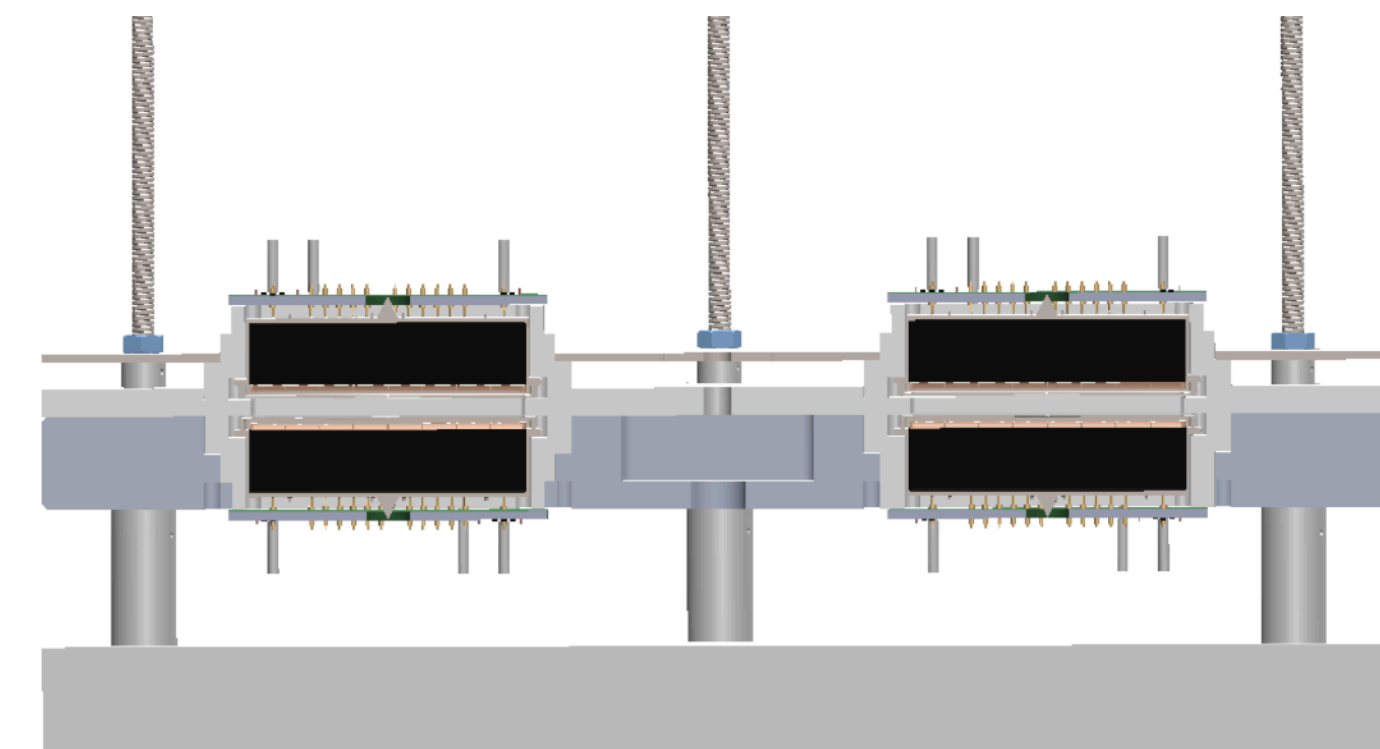
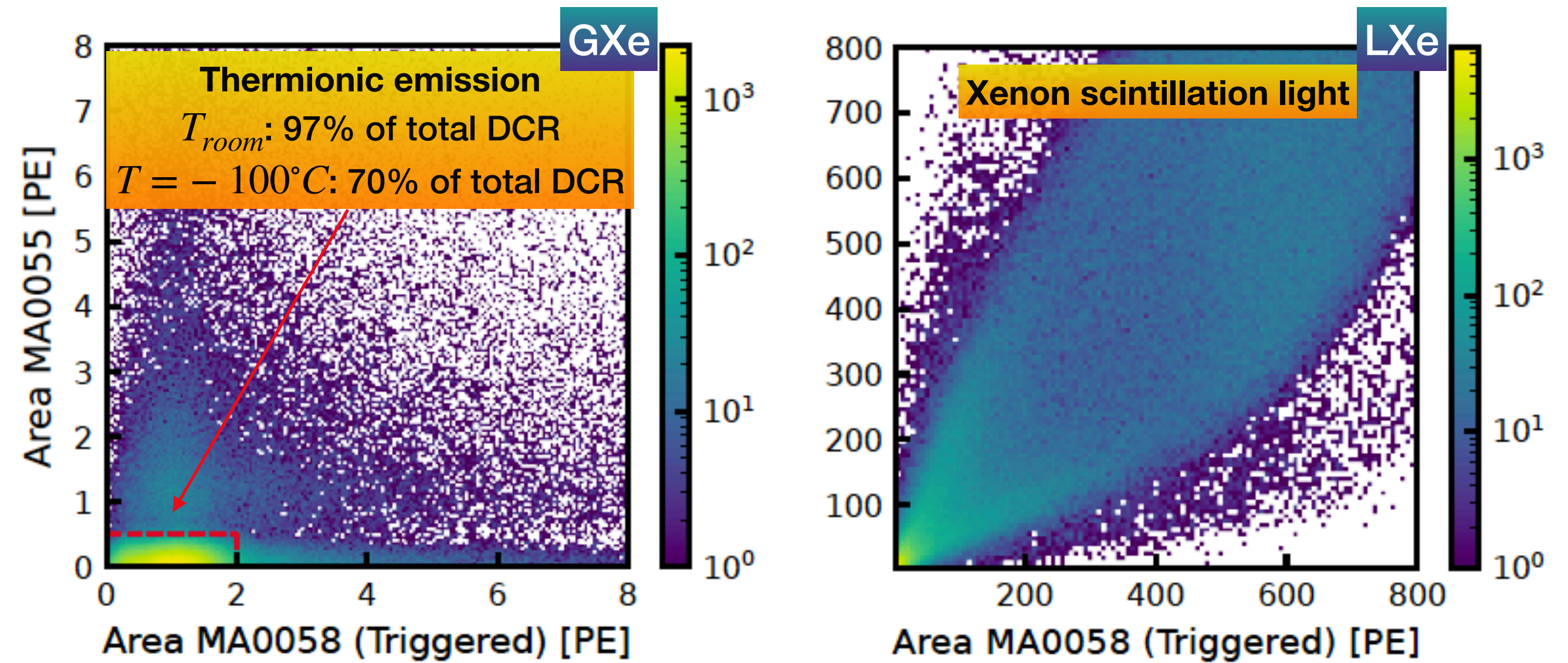


# Characterization results

## Dark counts

- Dark counts (DC) dangerous for accidental coincidences (ACs)
- Important background for WIMP and LowER searches in DARWIN

	DC rate at LXe temperature [Hz/cm <sup>2</sup> ]
<b>R12699-406-M4</b>	$0.4 \pm 0.2$
<b>R11410-21</b>	$1.4 \pm 0.7$



Two facing 2-inch PMTs in MarmotX



# Characterization results

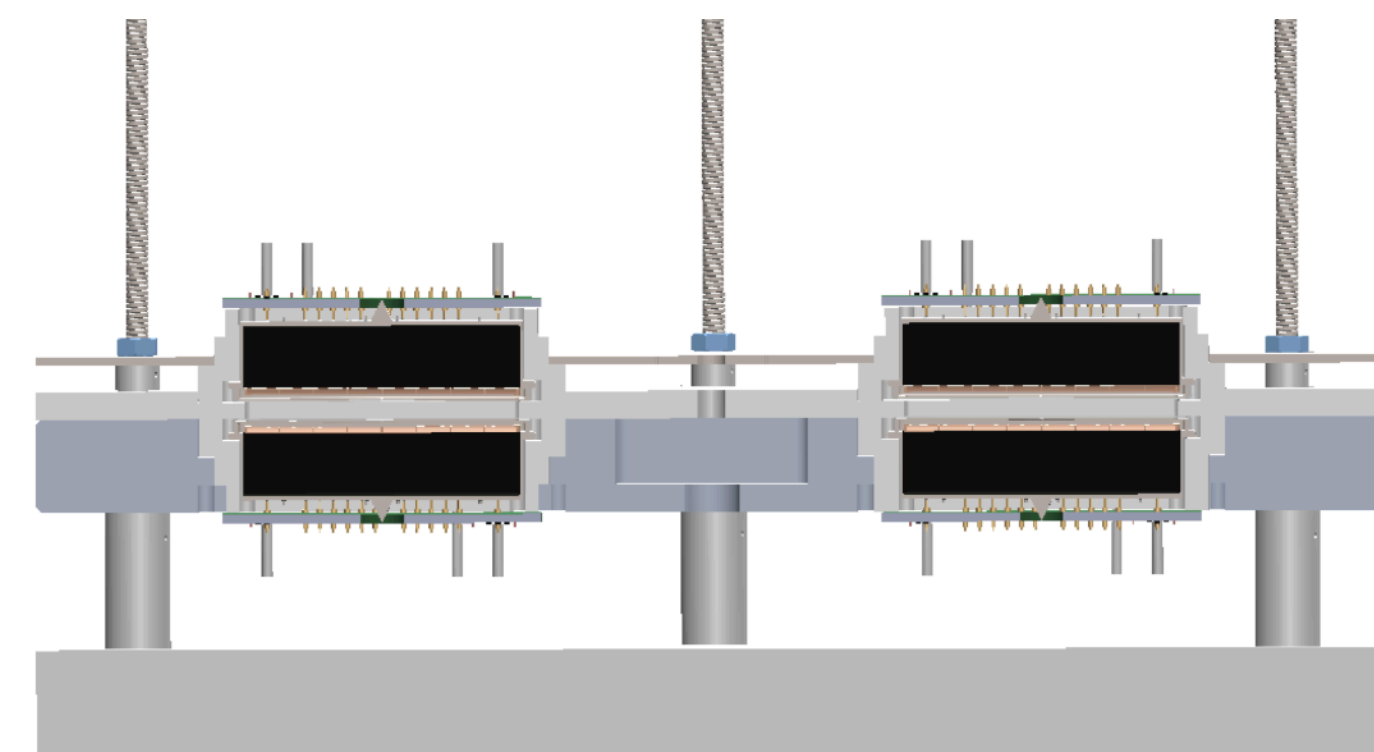
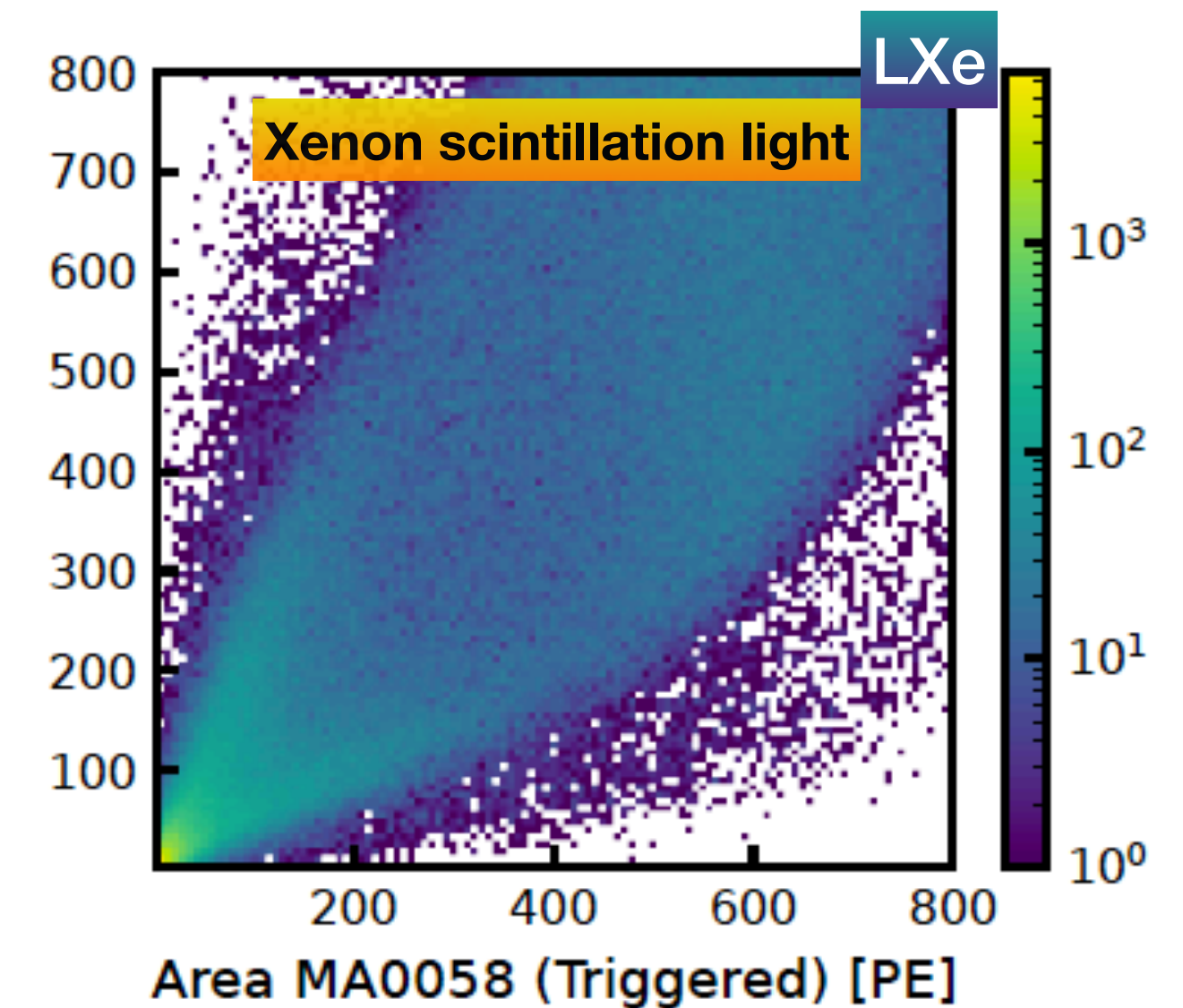
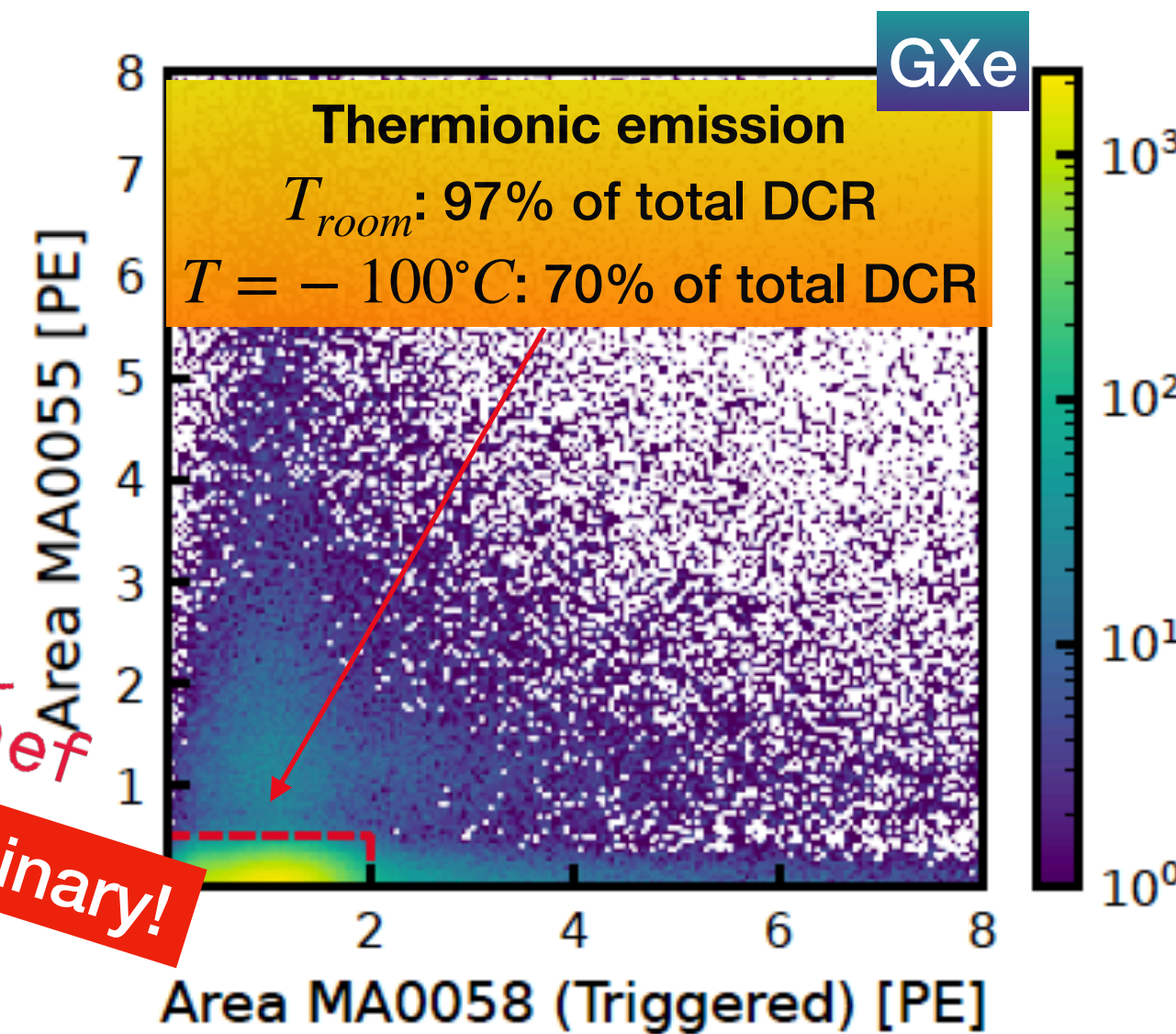
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	DC rate at LXe temperature [Hz/cm <sup>2</sup> ]	AC rate estimation for DARWIN in WIMP ROI* [events/year]
<b>R12699-406-M4</b>	$0.4 \pm 0.2$	$4 \cdot 10^4$
<b>R11410-21</b>	$1.4 \pm 0.7$	$3 \cdot 10^5$

\*4-fold coincidence, S1 ROI (4,20) PE and S2 ROI (100,1000) PE  
AC rate same order of magnitude for XENONnT sized detector

Nikhef  
Very preliminary!



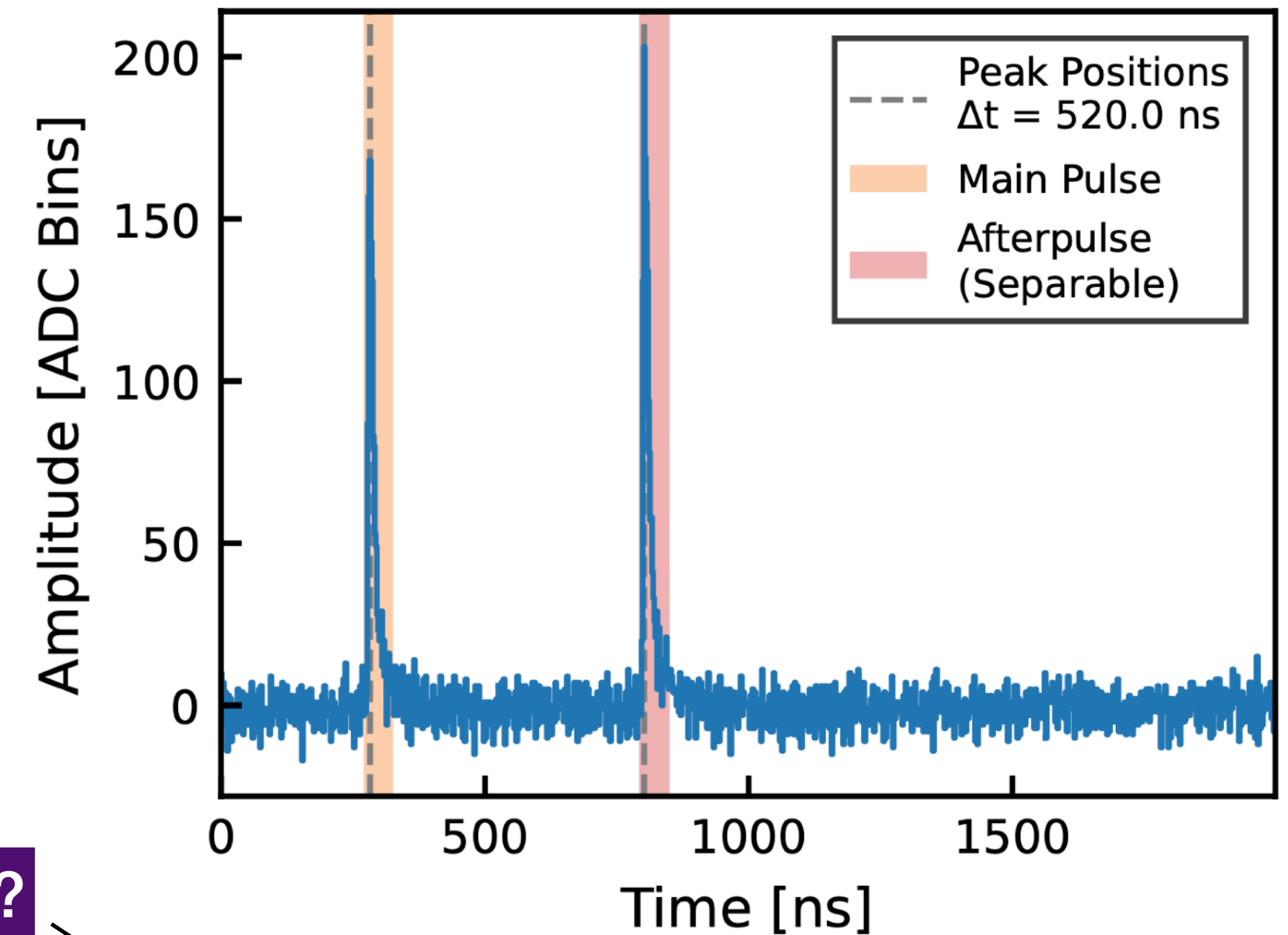
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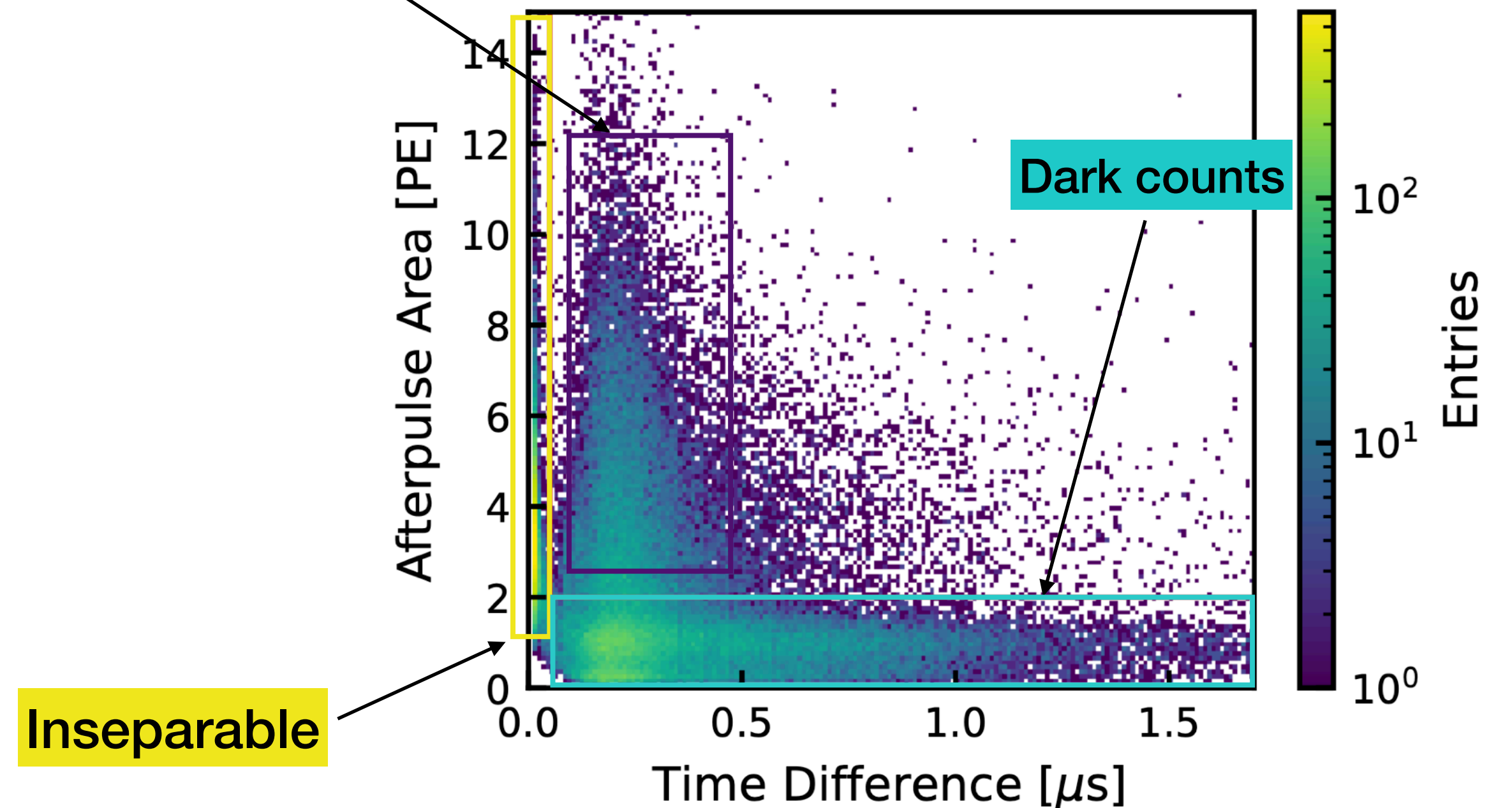
# Characterization results

## Afterpulsing

- ▶ Timing: ion drift path and mass-to-charge ratio
- ▶ Expected timing order of magnitude faster than R11410-21 PMTs
- ▶ Different AP treatment needed
- ▶ For 8PE/trigger occupancy, separable AP rate:  $(0.90 \pm 0.2)\%$ /PE  $\rightarrow$  hard to compare



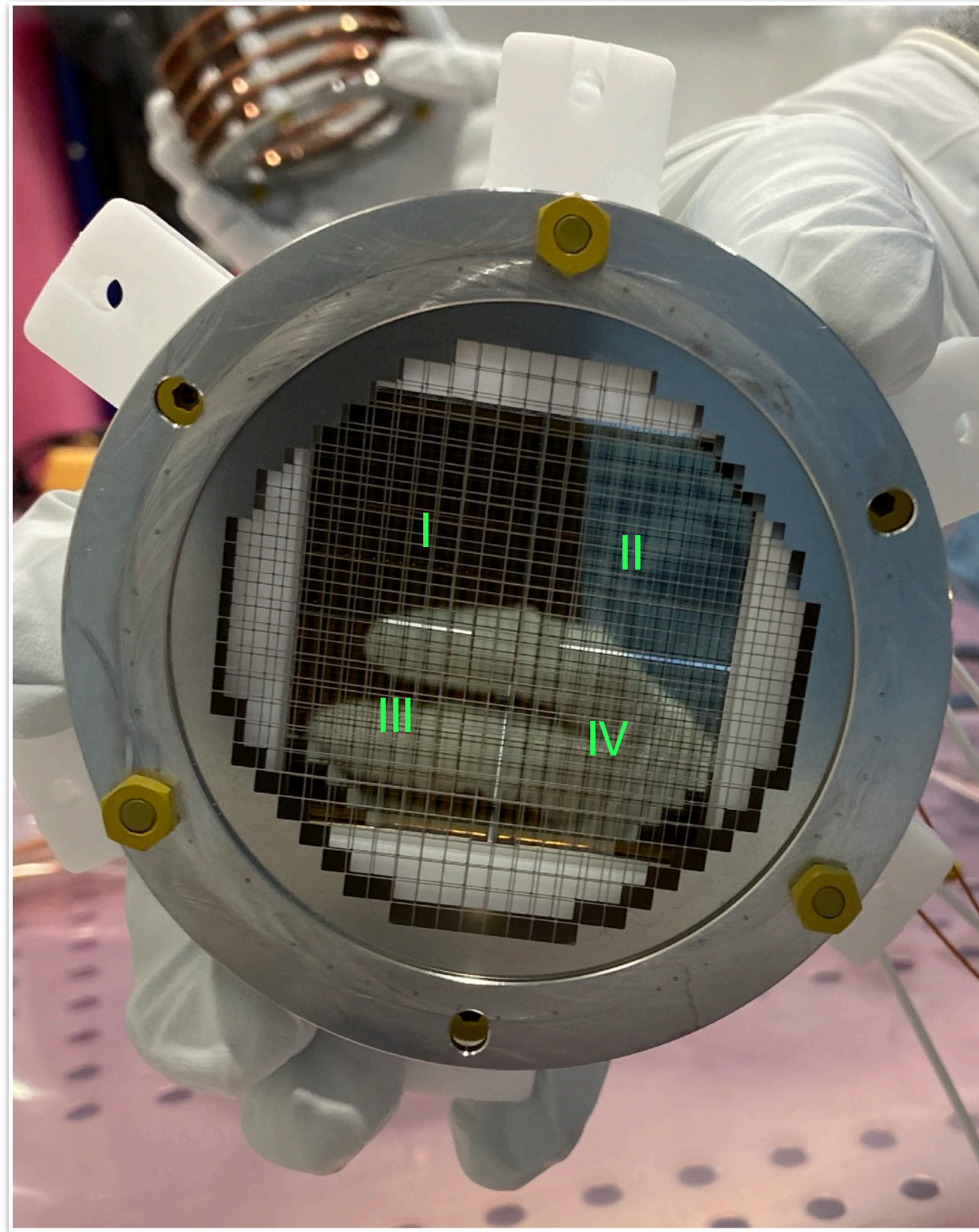
Heavy ions?





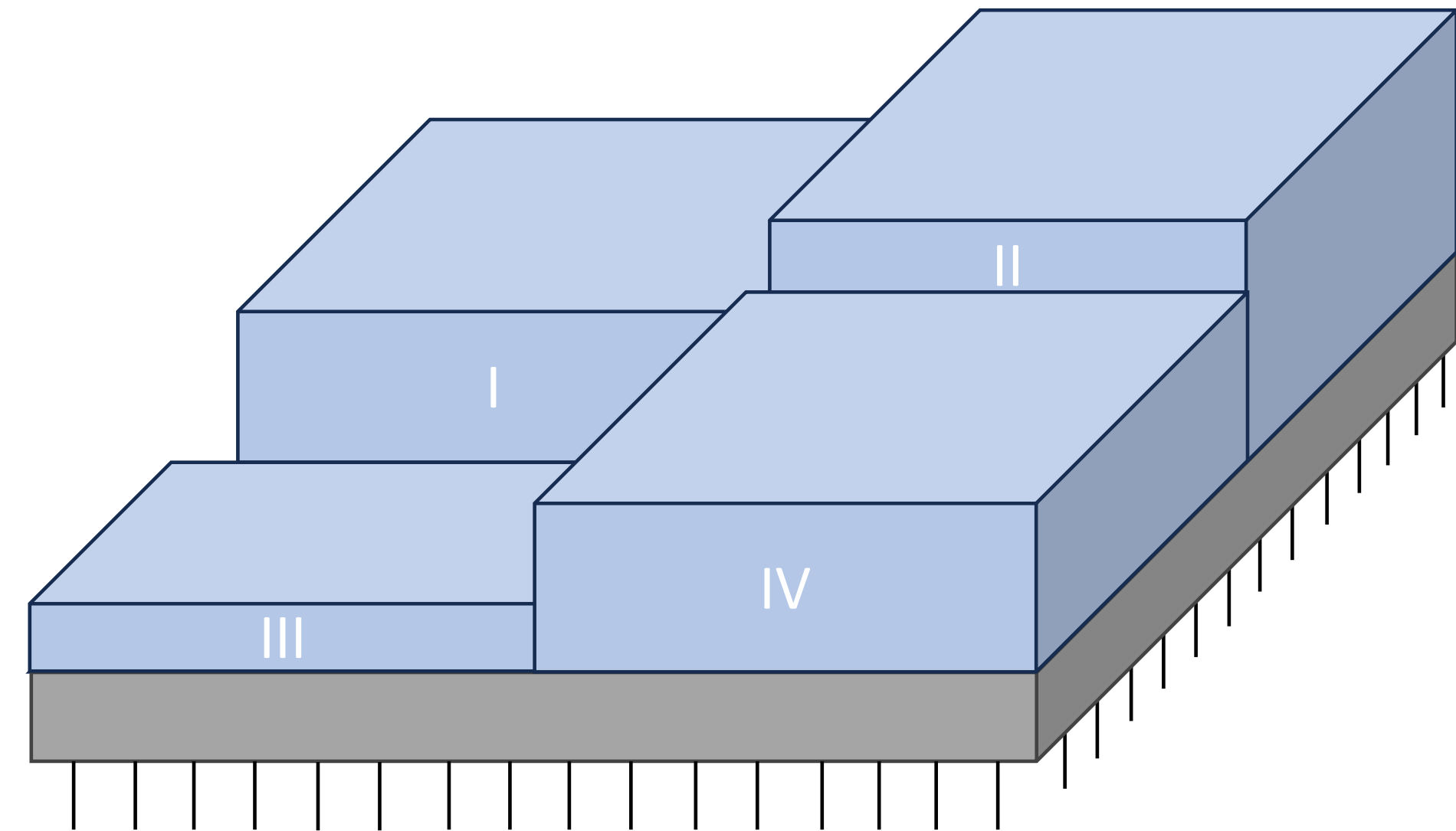
# Characterization results

## Position reconstruction



XAMS Top PMT holder including anode and top screening meshes

$$X = \sum_{i=1}^4 \frac{x_i I_i}{I_{total}}$$

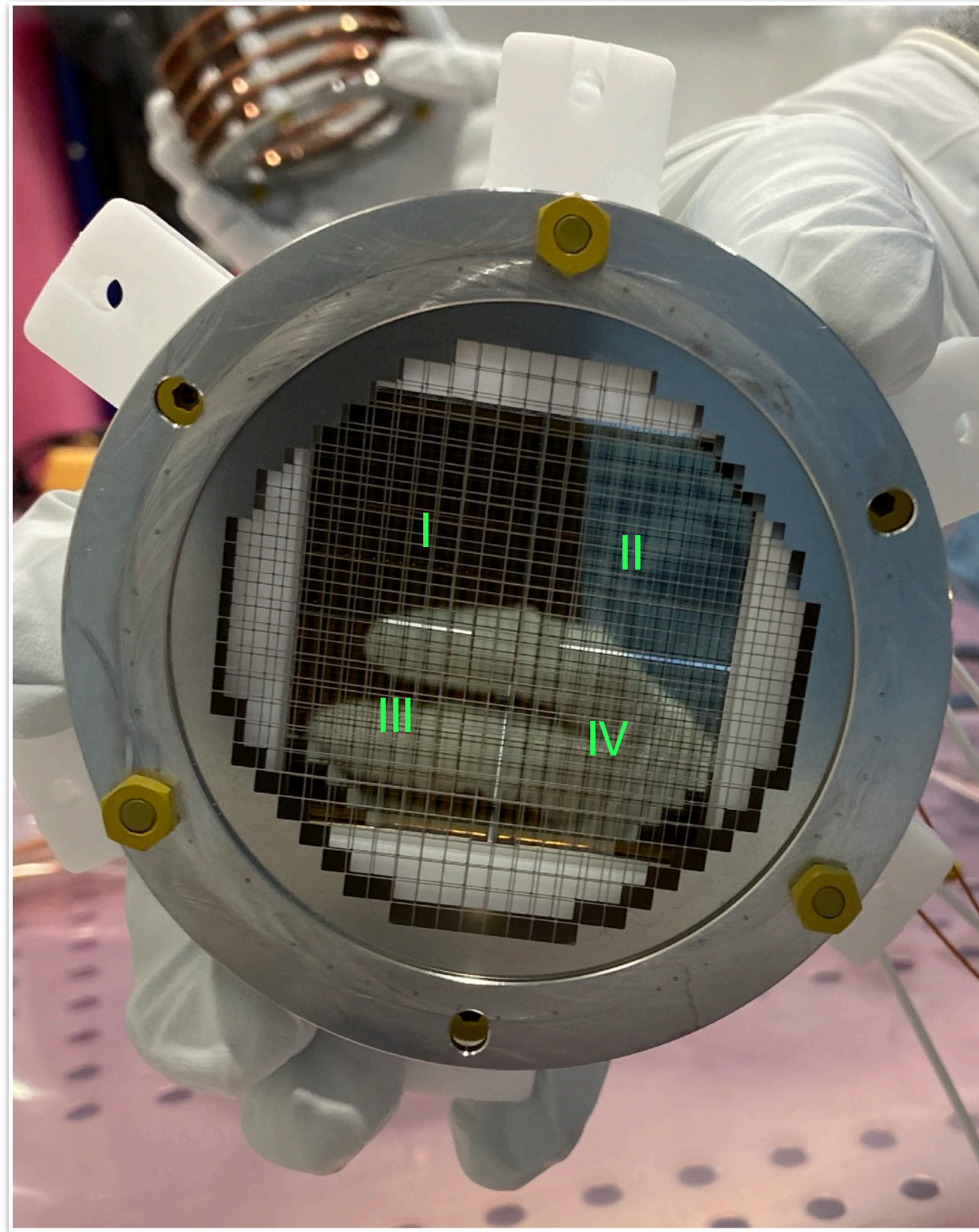


Center of gravity method

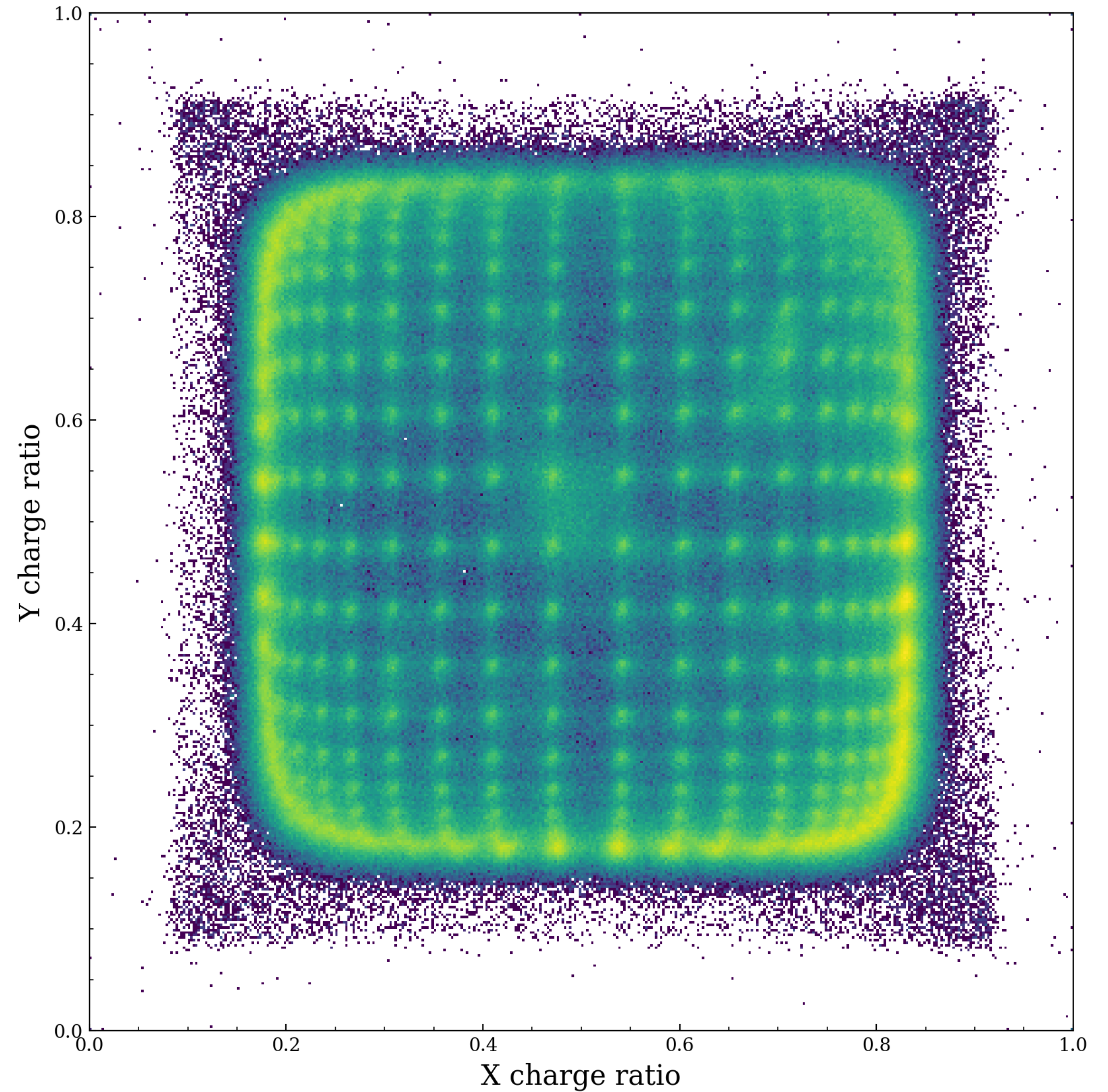


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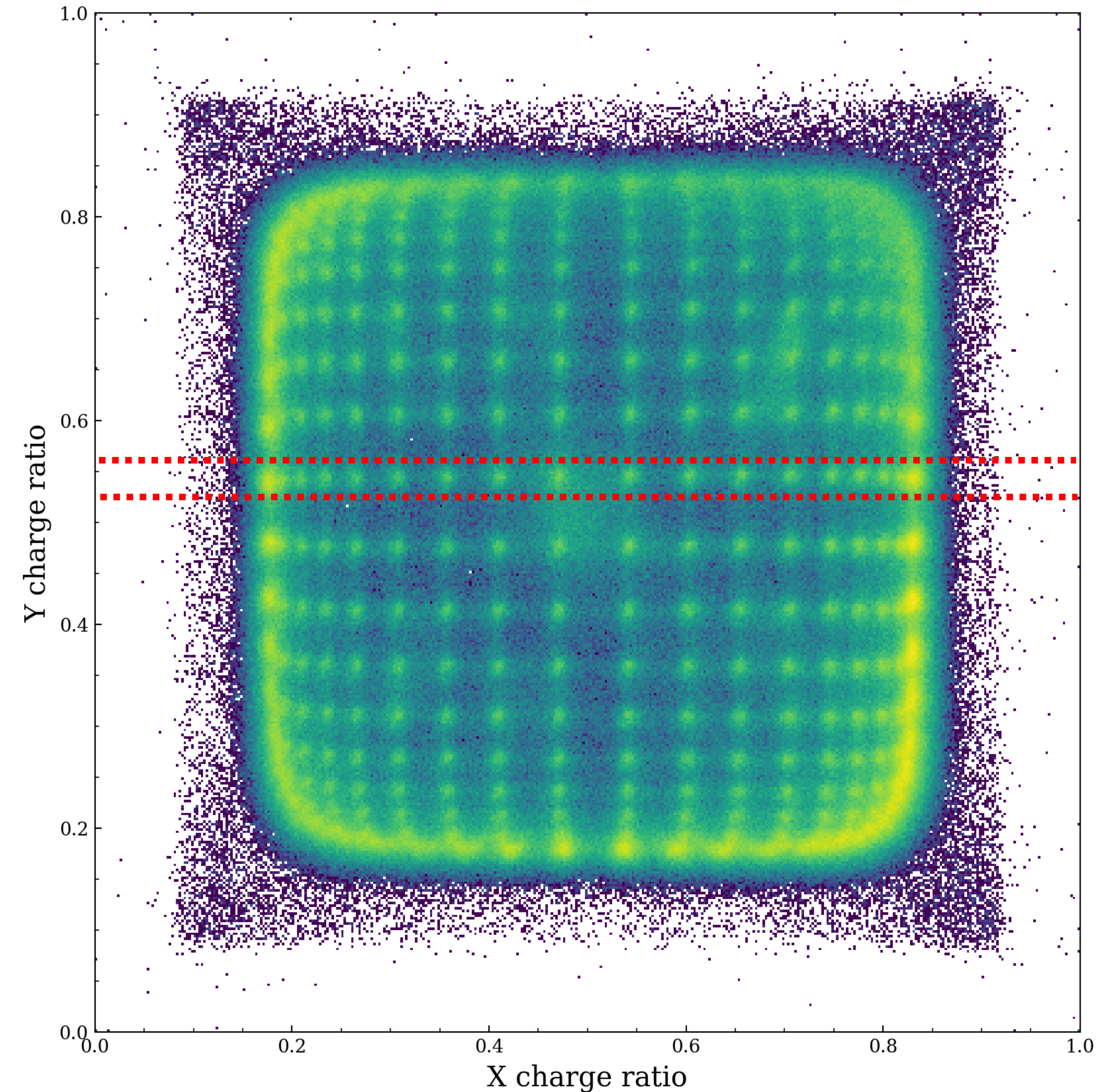
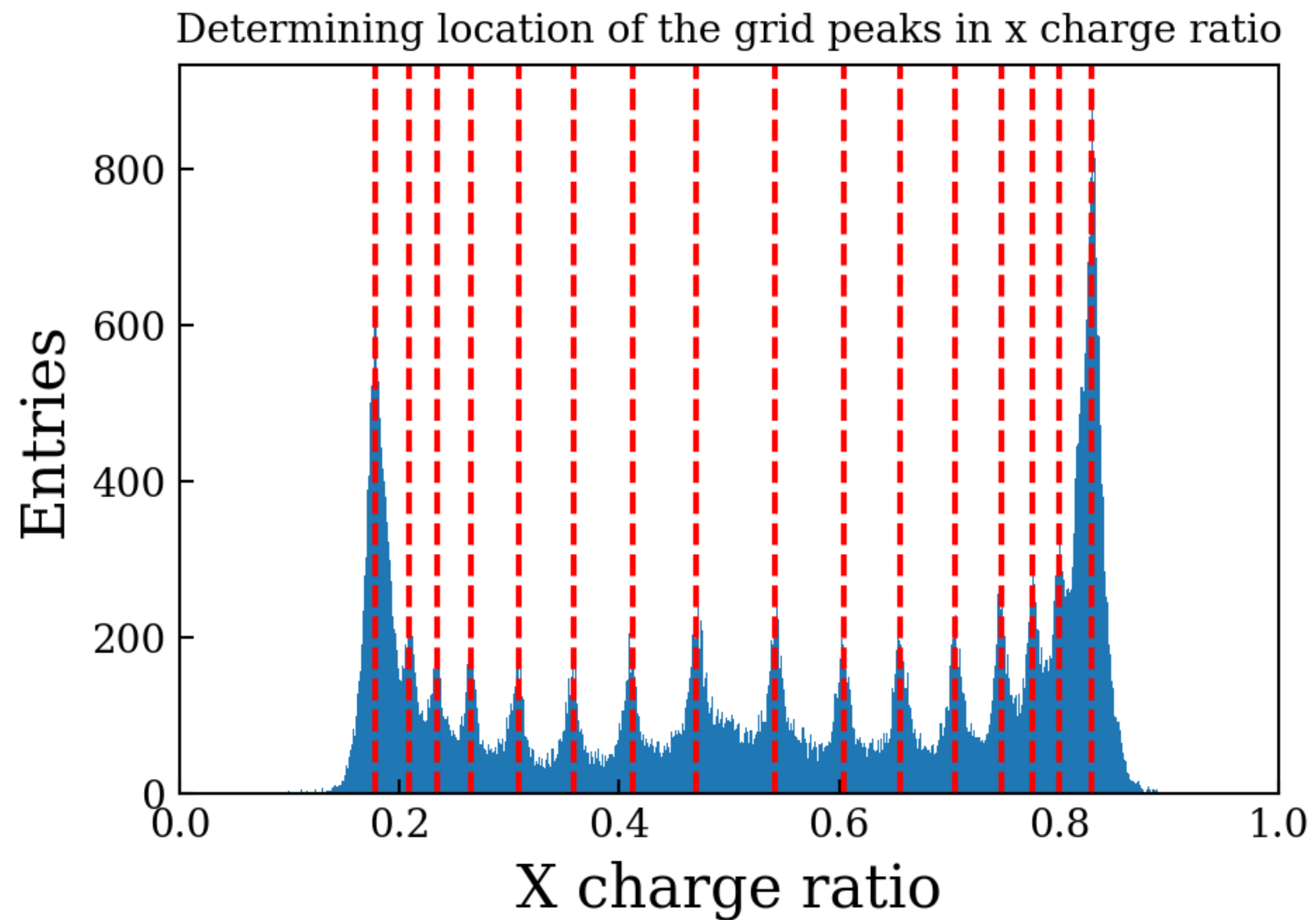
XAMS Top PMT holder including anode and top screening meshes





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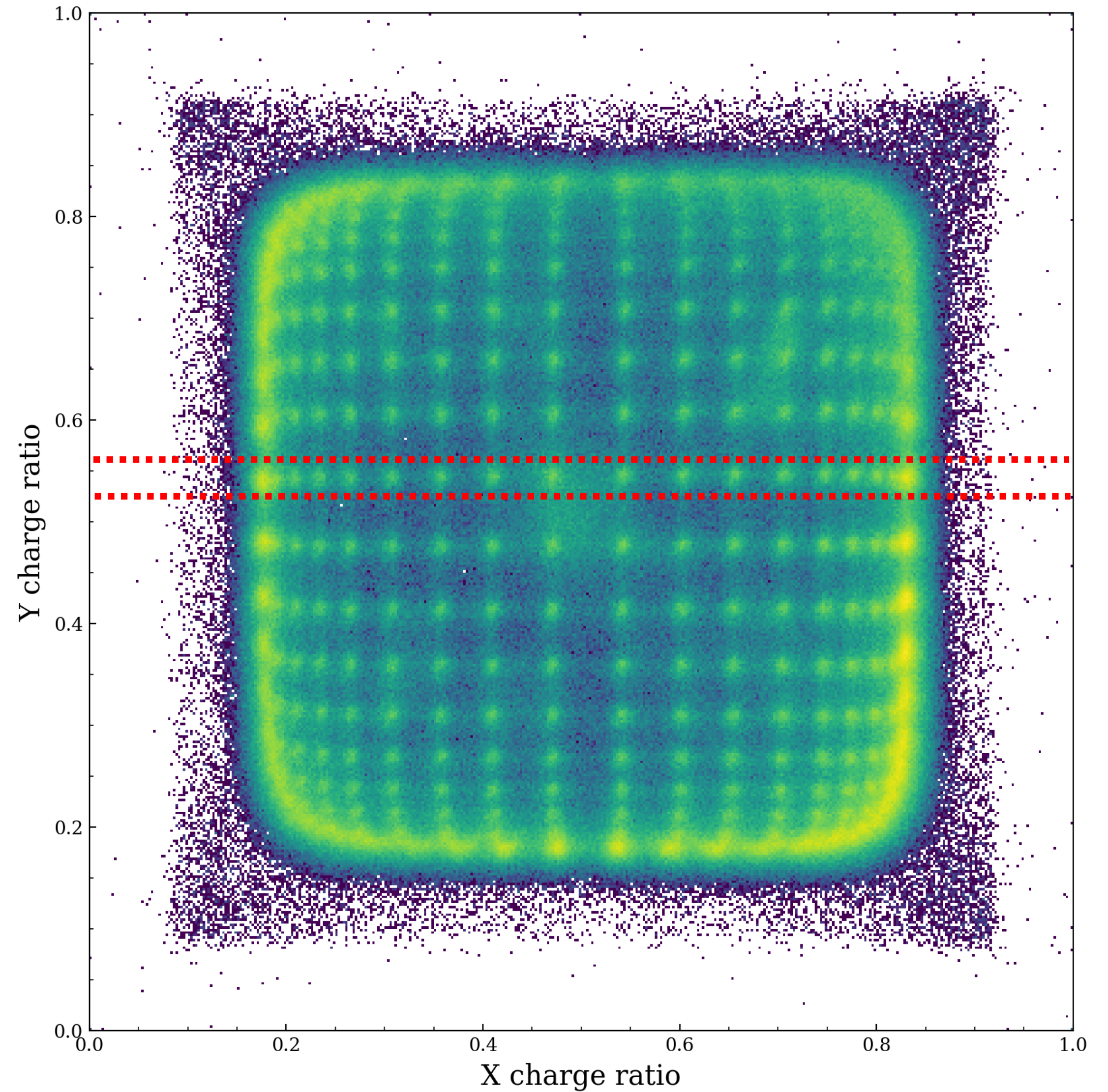
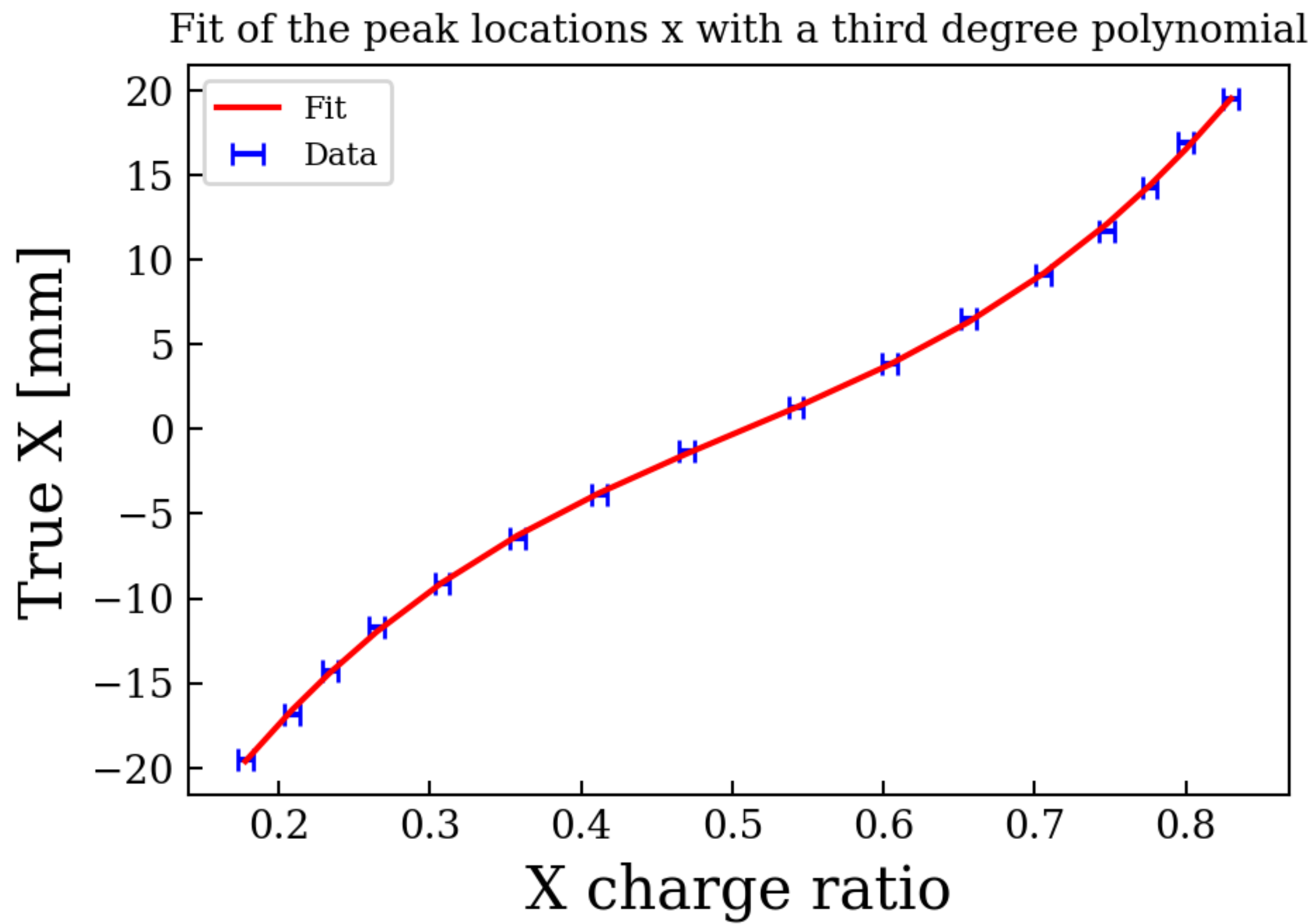
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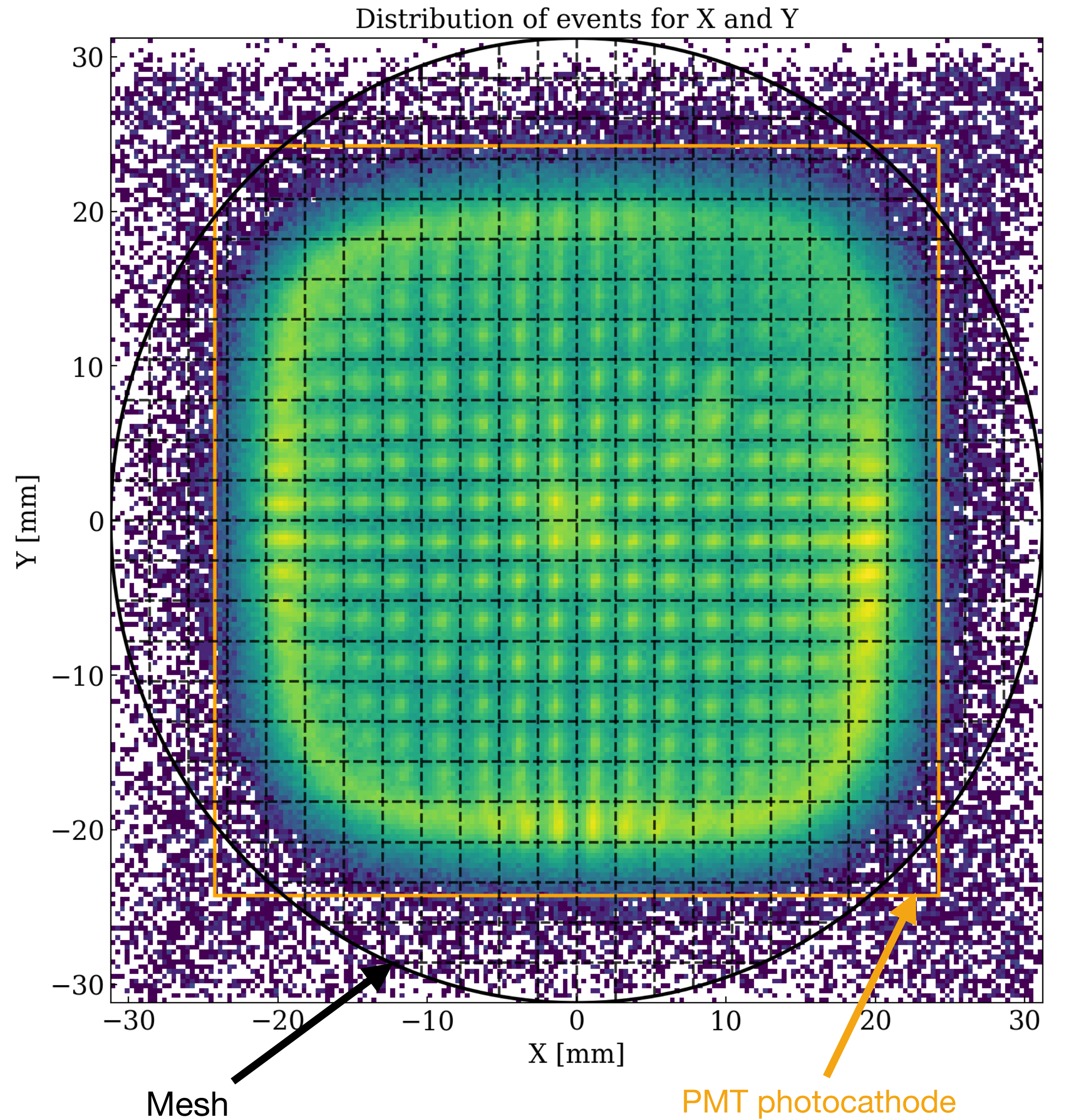
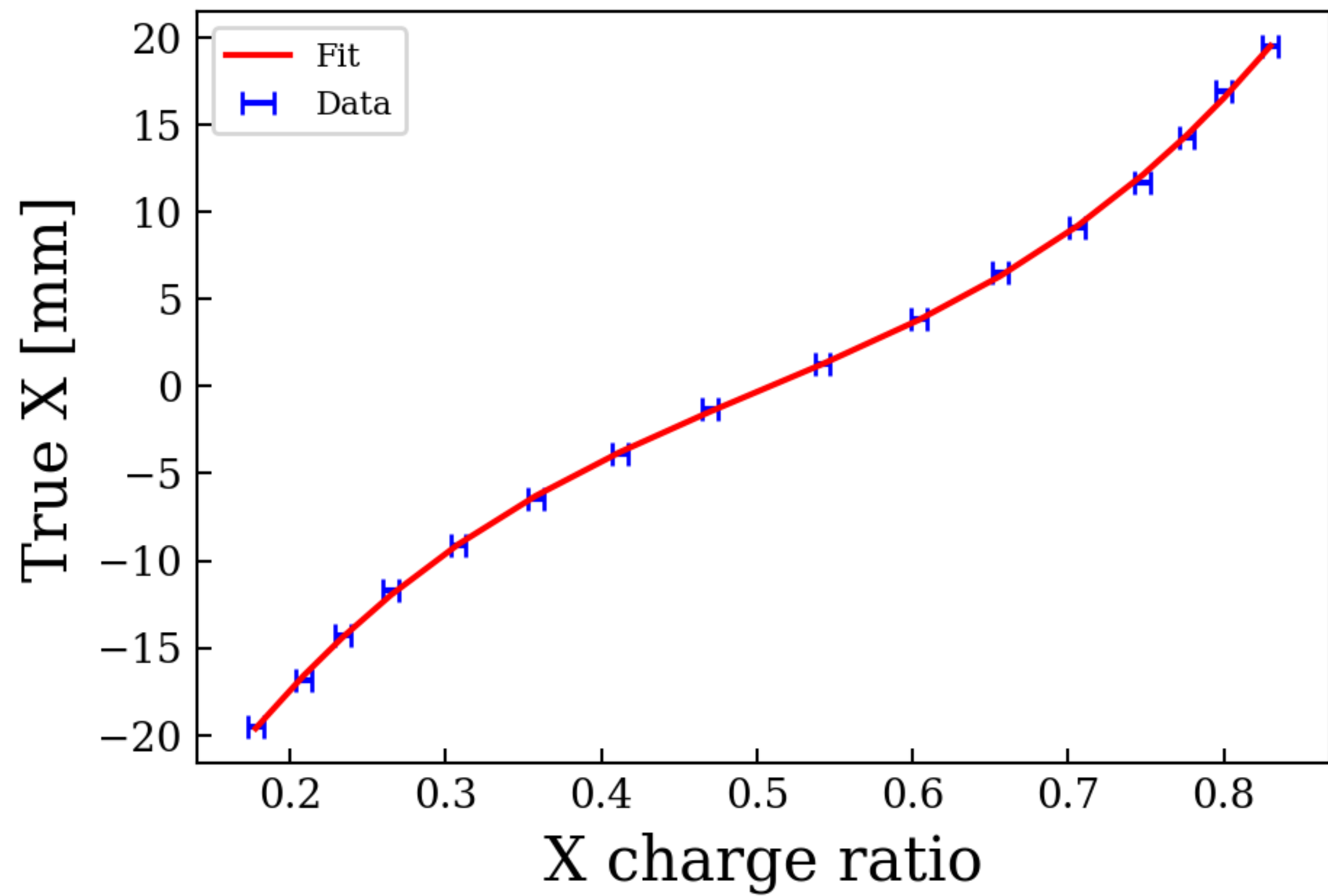




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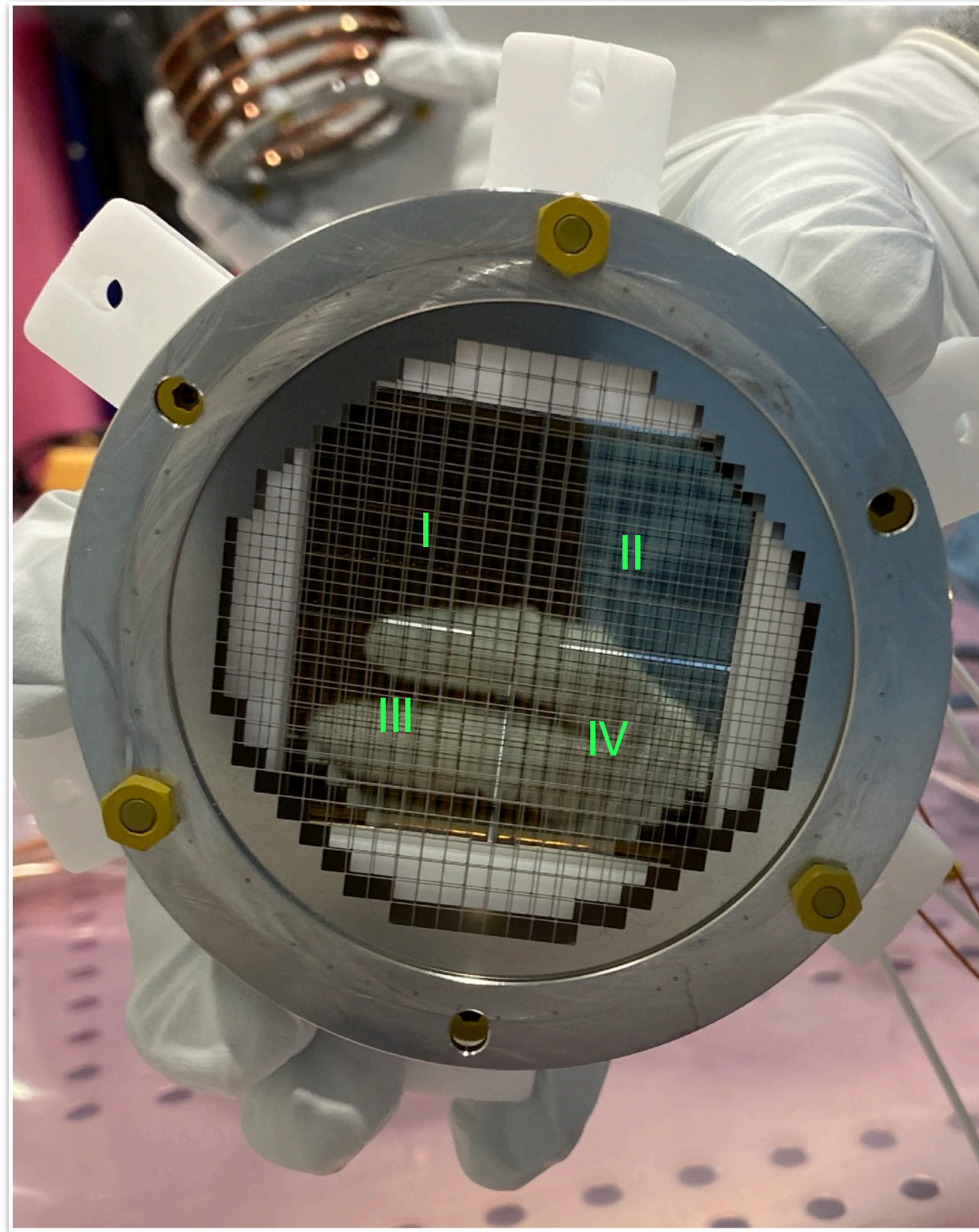
Fit of the peak locations x with a third degree polynomial



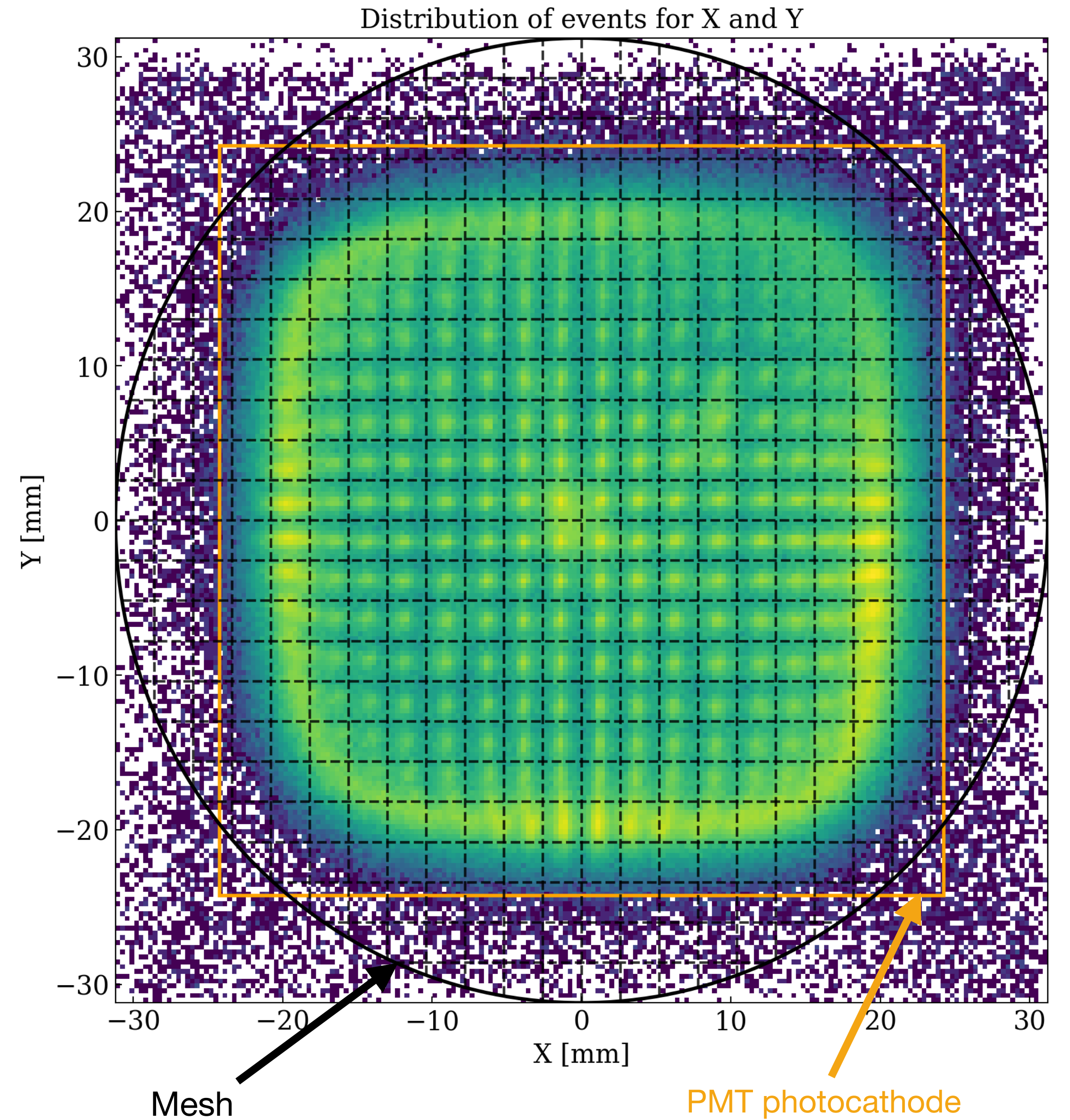


# Characterization results

## Position reconstruction



XAMS Top PMT holder including anode and top screening meshes

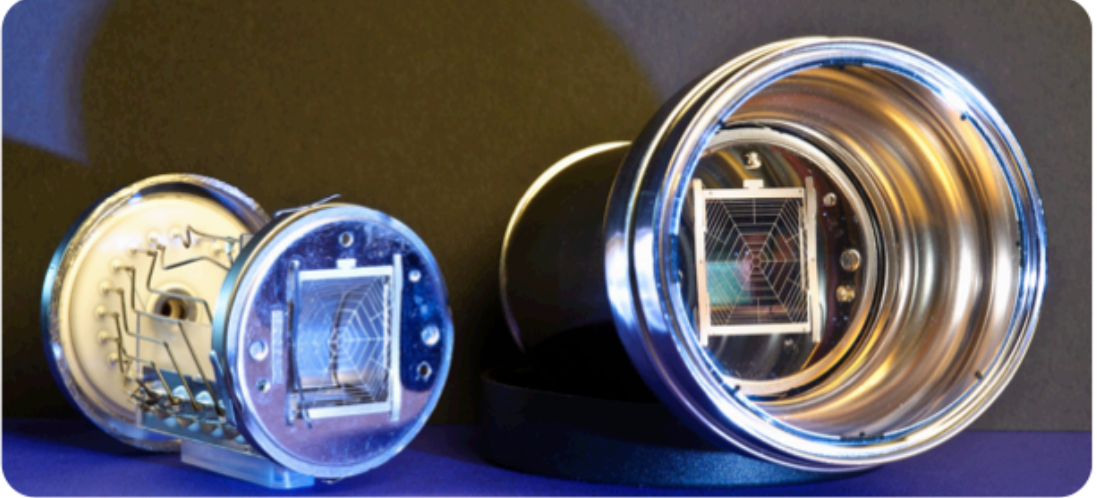





# Summary and comparison

## R12699 & R11410-21

+ position reconstruction!

	↑ 114 mm    ∅ 77.5 mm
	61.8 %
	Box & linear-focused / 12
	1500 V / 1750 V
	32.5 % / (34.5±2.8) %
	$3.5 \cdot 10^6$ / $(5.4 \pm 2.1) \cdot 10^6$
	5.5 / 46 / 9 ns
	0.5 μs
	$(1.4 \pm 0.7)$ Hz/cm <sup>2</sup>



	↑ 14.8 mm    ↔ 56 mm	
	75.0 %	
	Metal channel / 10	
	1000 V / 1100 V	
	33 % / <i>tbd</i>	
	$1.5 \cdot 10^6$ / $(3.0 \pm 0.8) \cdot 10^6$	
	1.2 / 5.9 / 0.41 ns	
	0.03-0.06 μs	
	$(0.5 \pm 0.2)$ Hz/cm <sup>2</sup>	

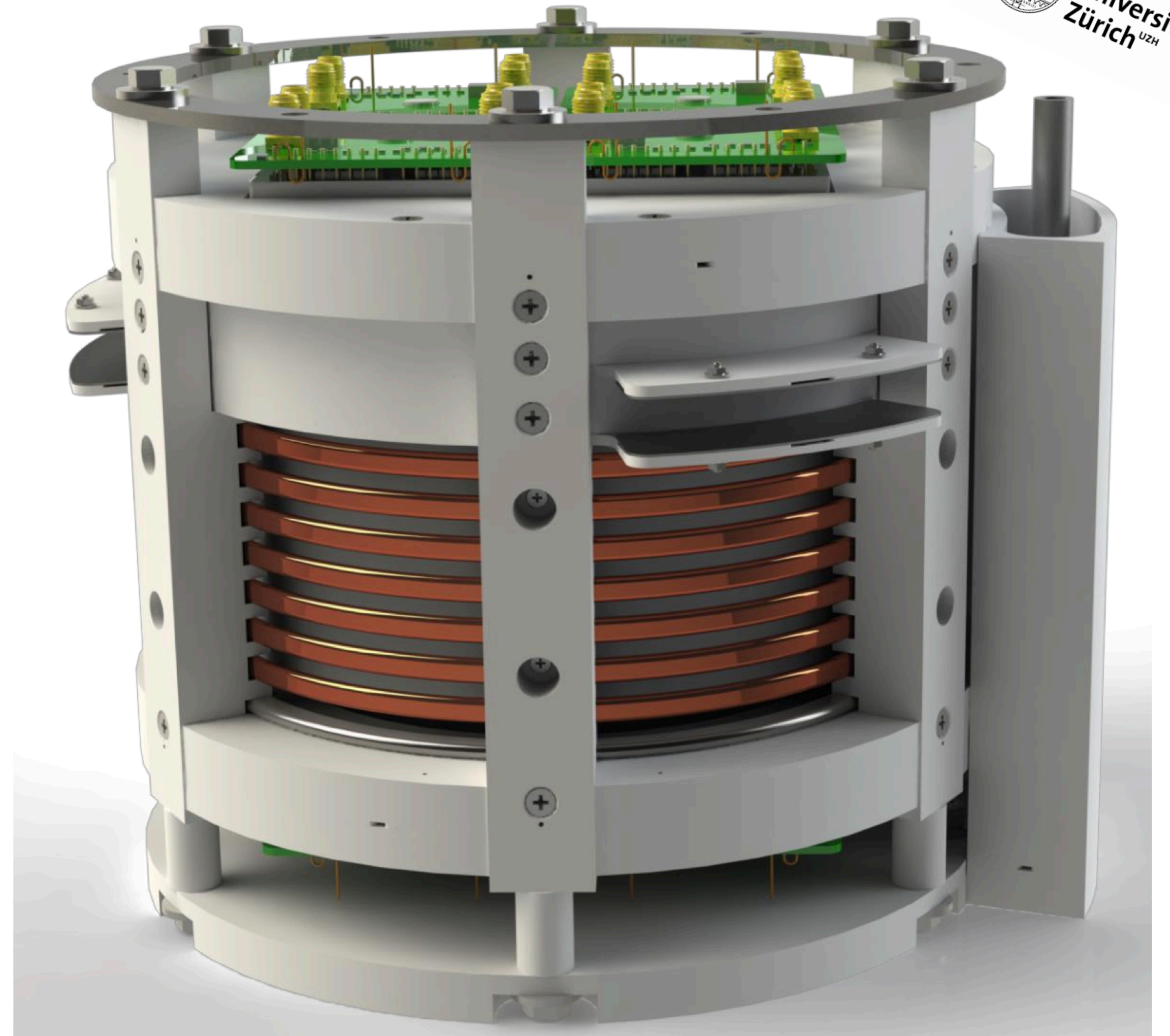
- Dimensions
- Packing density
- Dynodes (structure / number of stages)
- Operation voltage (nominal / maximum)
- Quantum efficiency at 175 nm (data sheet / effective)
- Gain (data sheet / effective)
- Time response (Rise time / Transit time / Transit time spread)
- Expected He<sup>+</sup> afterpulse delay
- Dark count rate (-100°C, >1/4 PE)



# Outlook

## Ongoing studies

- ▶ Long term stability tests
  - ▶ Gain/SPE response
  - ▶ Afterpulse rate
- ▶ Extend characterization to more PMTs
- ▶ Assembly of 2x4 TPC setup at UZH
- ▶ Improving position reconstruction resolution by building simulation of XAMS



Design of 2x4 setup at UZH

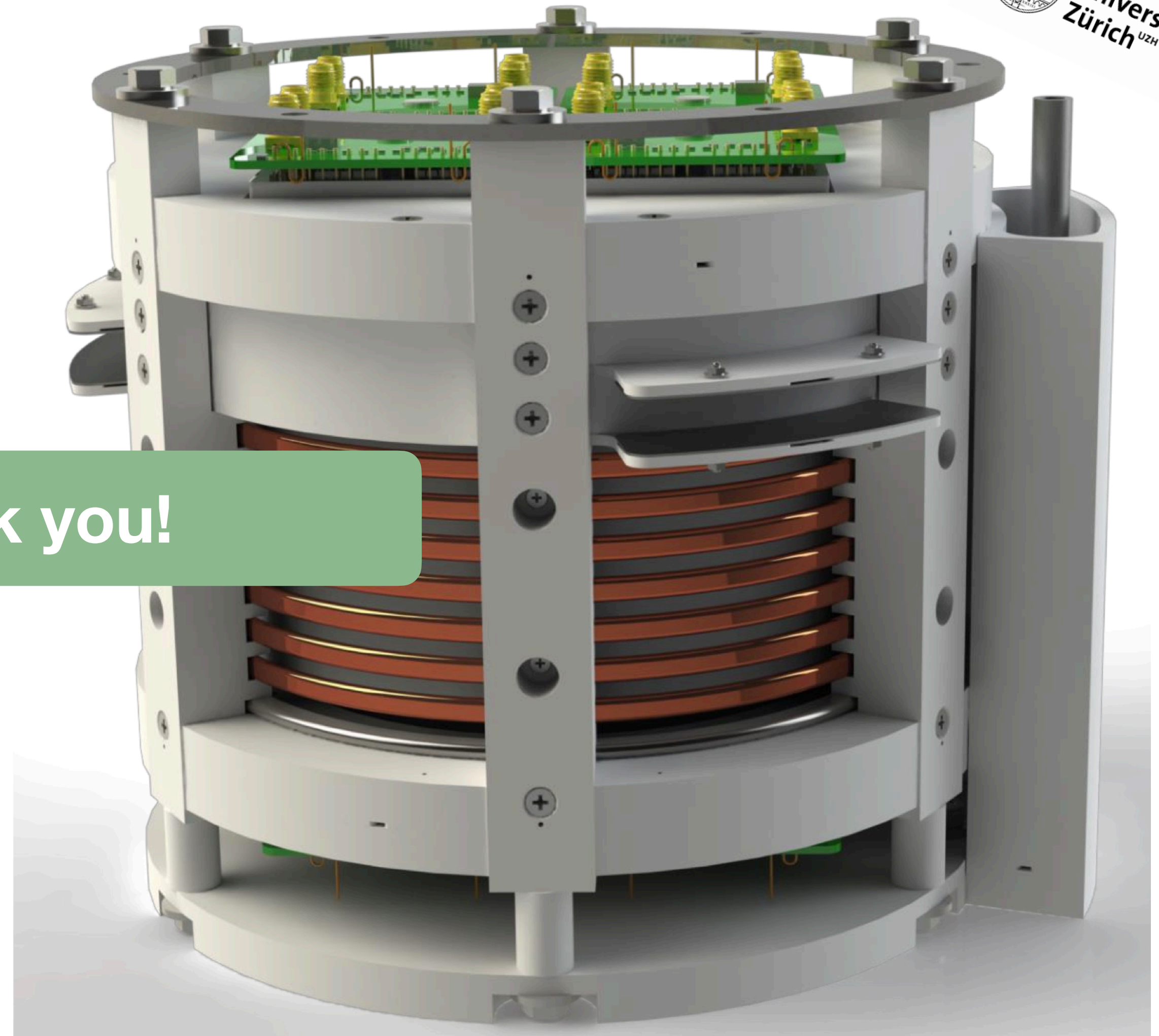


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Thank you!



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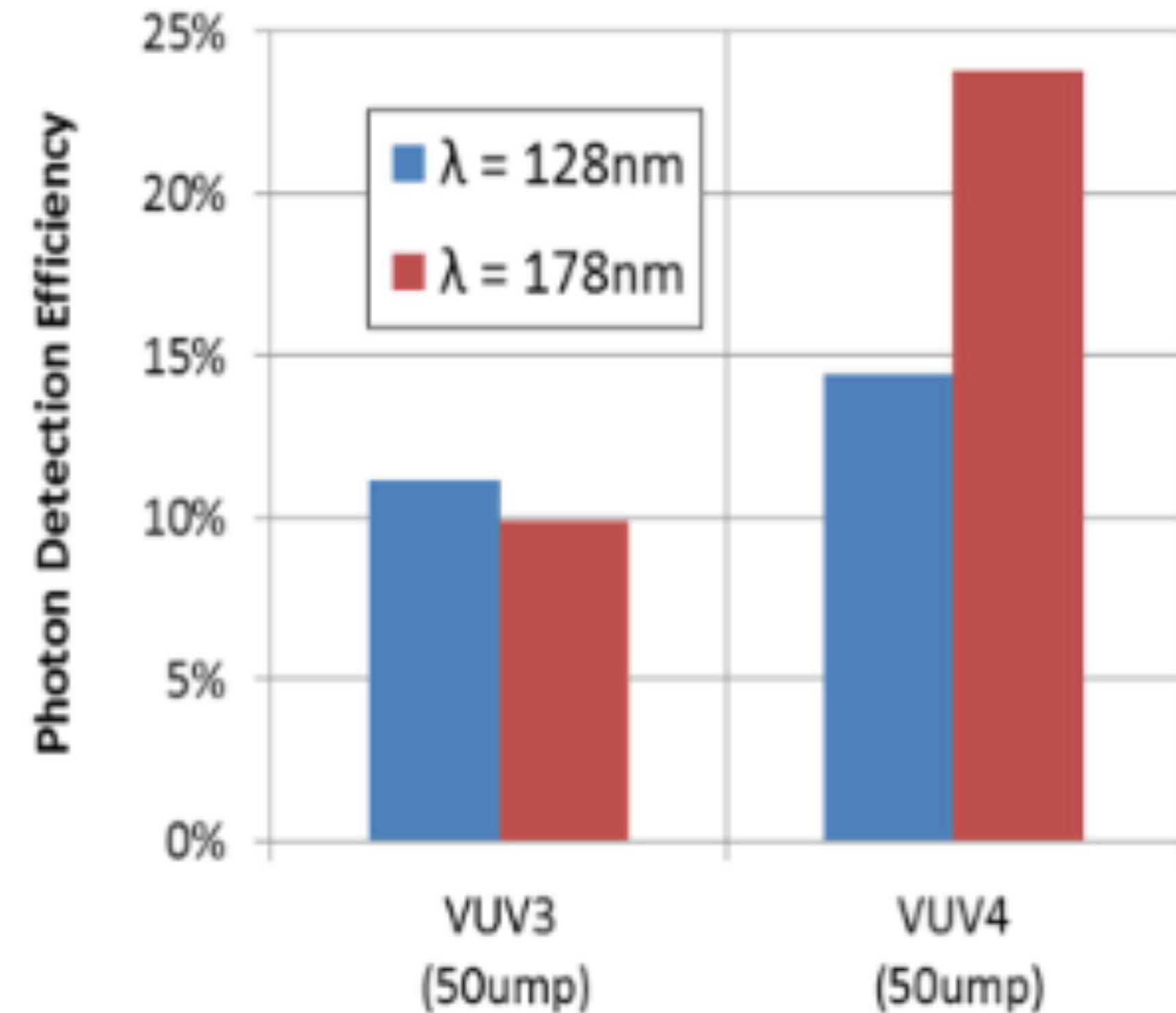
**Backup slides**



# SiPMs

## Pros and cons

- ▶ Pros:
  - ▶ No high voltage needed
  - ▶ Cheap
  - ▶ Low radioactivity
- ▶ Cons:
  - ▶ We need a large photosensitive surface  
→ high channel count (high DCR, lots of cables, high data stream, etc)
  - ▶ Lower QE ~20% for UV sensitive SiPMs



From Hamamatsu datasheet



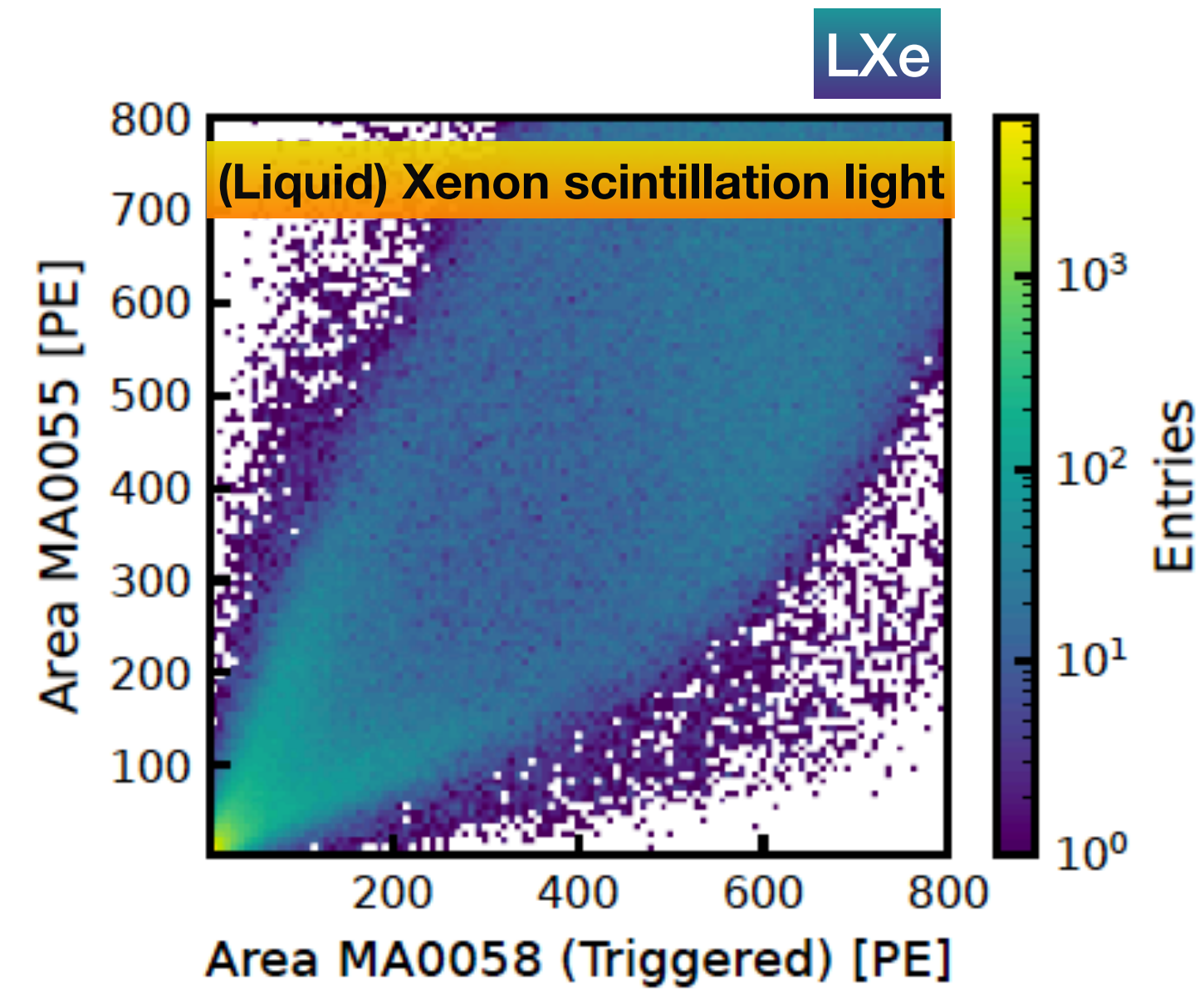
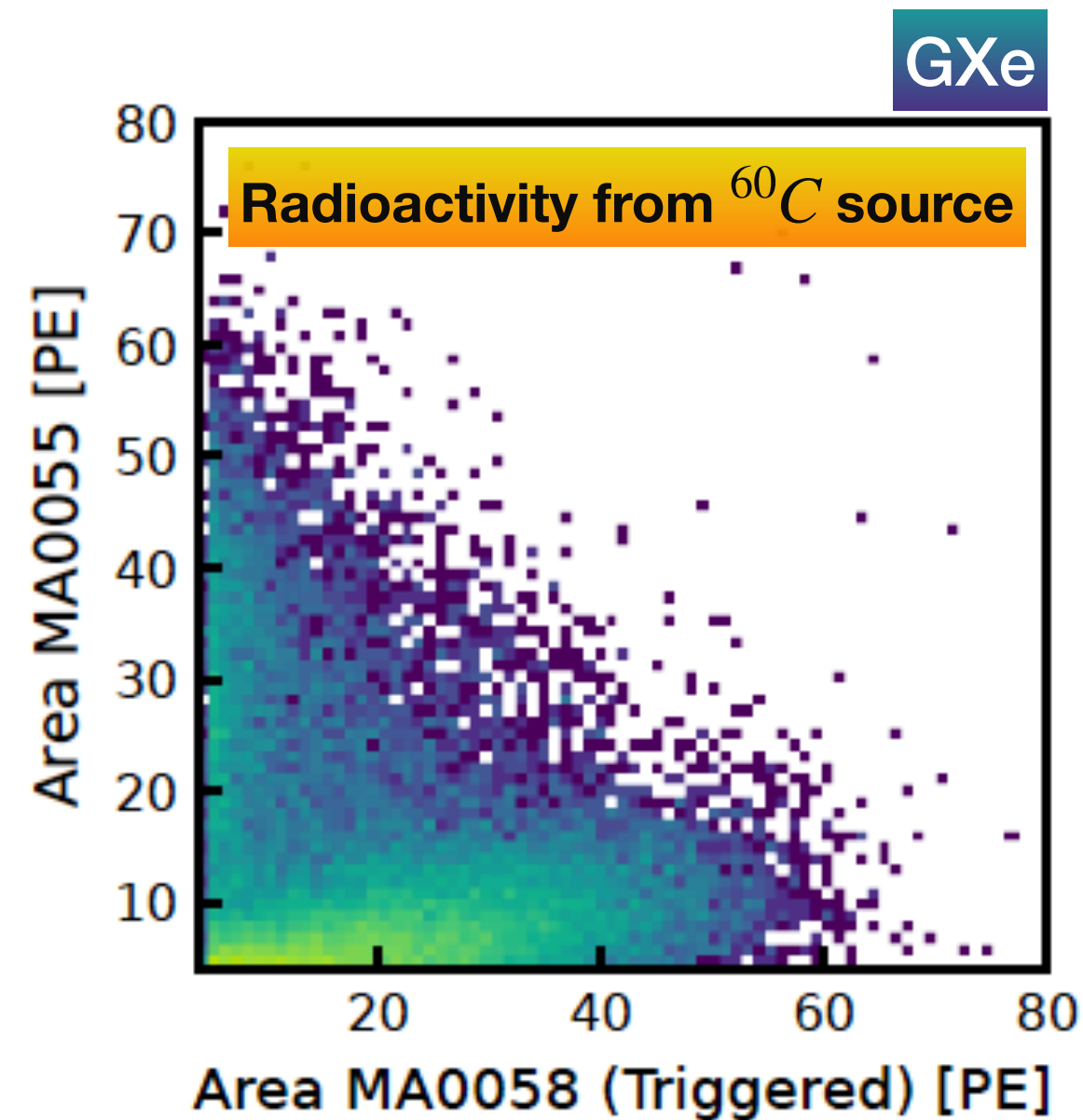
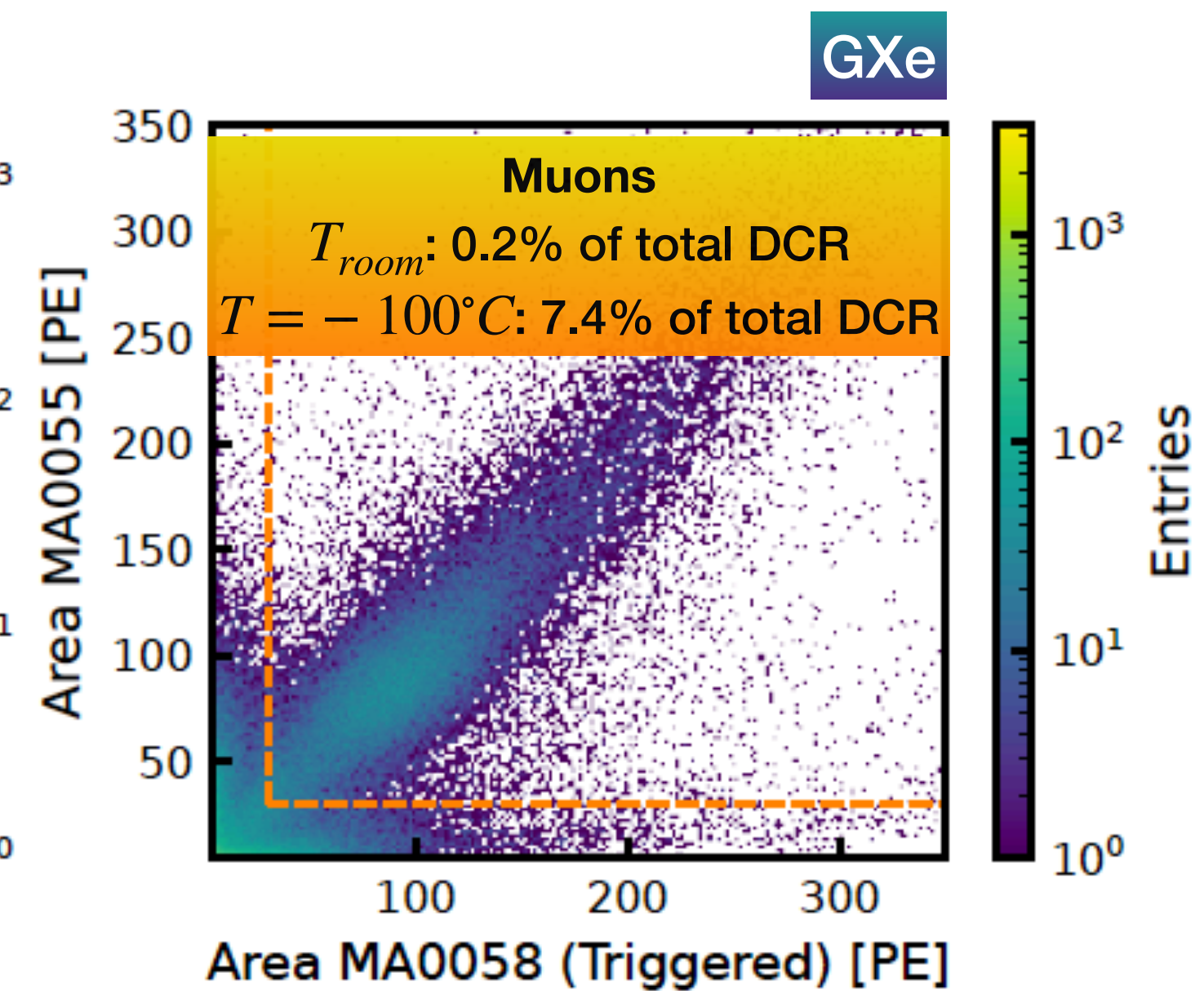
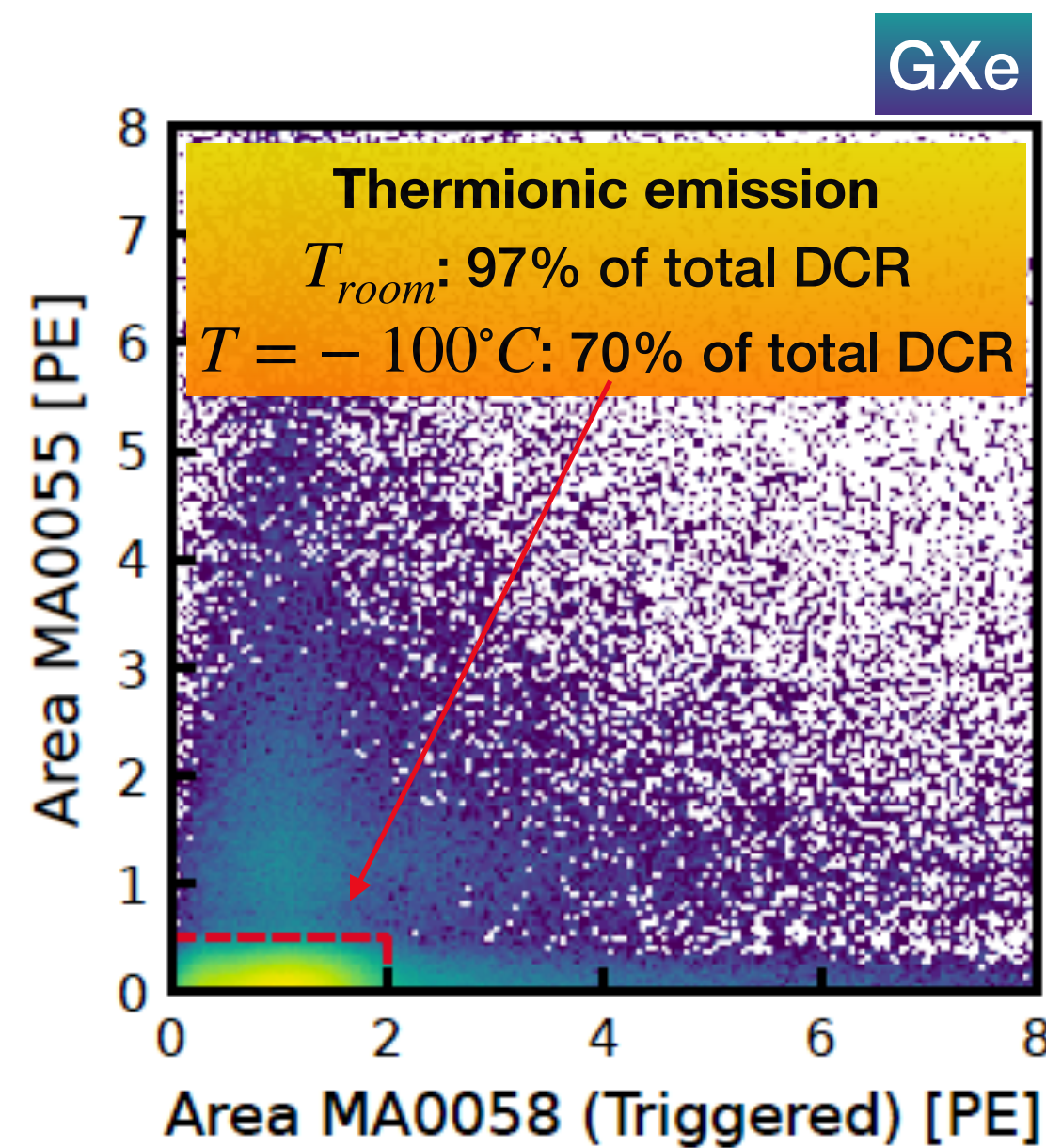
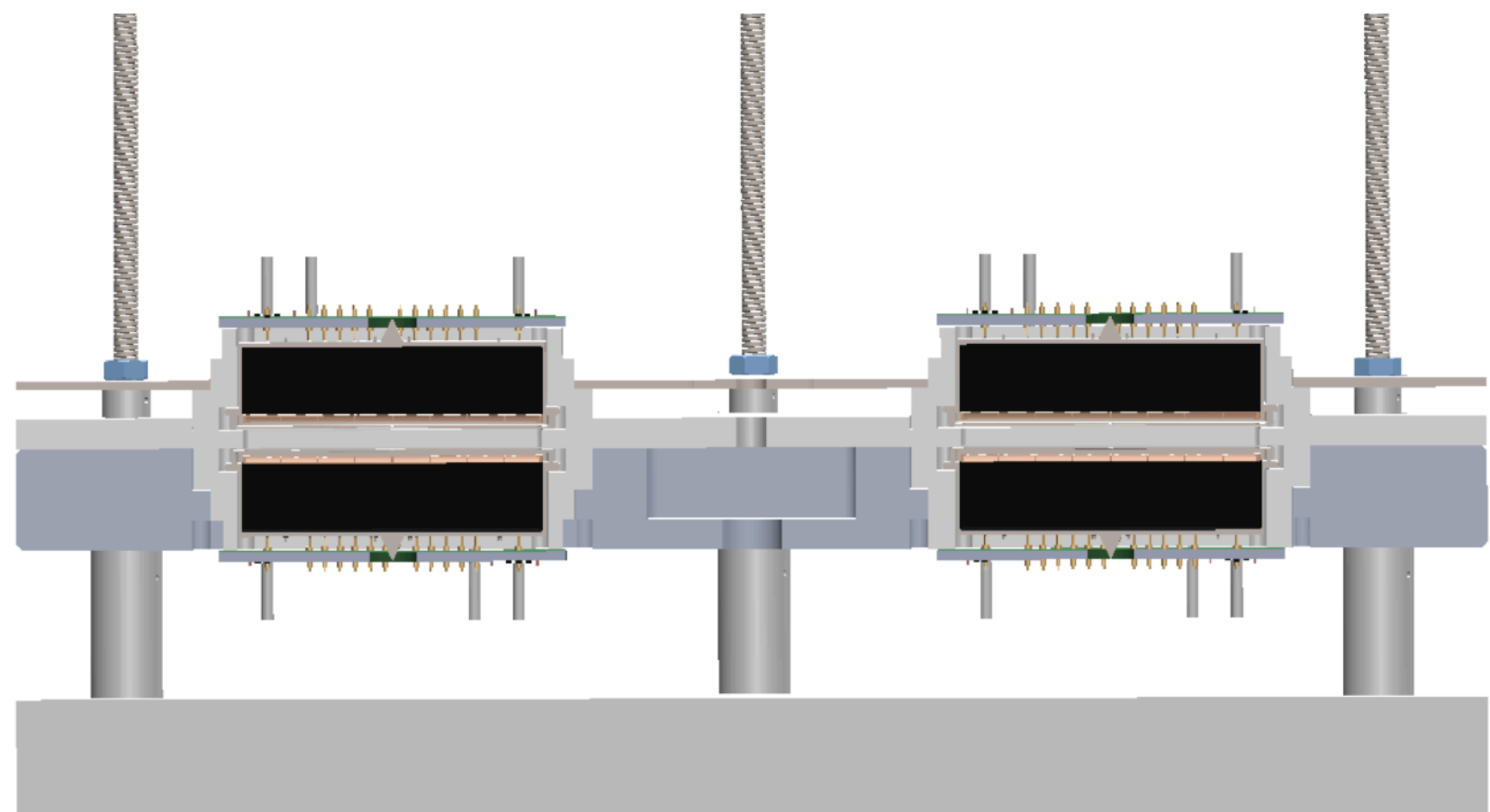
# DCR to AC estimation

## Very preliminary

- ▶ Semi-analytical (partially simulation, partially theory) model to predict AC rate as a function of:
  - ▶ Dark Count Rate
  - ▶ Detector parameters ( $g_1$ ,  $g_2$ , SEG, electric fields, etc)
  - ▶ Detector geometry (size, number of PMTs, etc)
- ▶ Matched to XENON1T data
- ▶ Extrapolated to DARWIN-sized detector



# DCR and LE



1/4 PE threshold



# Model independent approach<sup>1</sup>

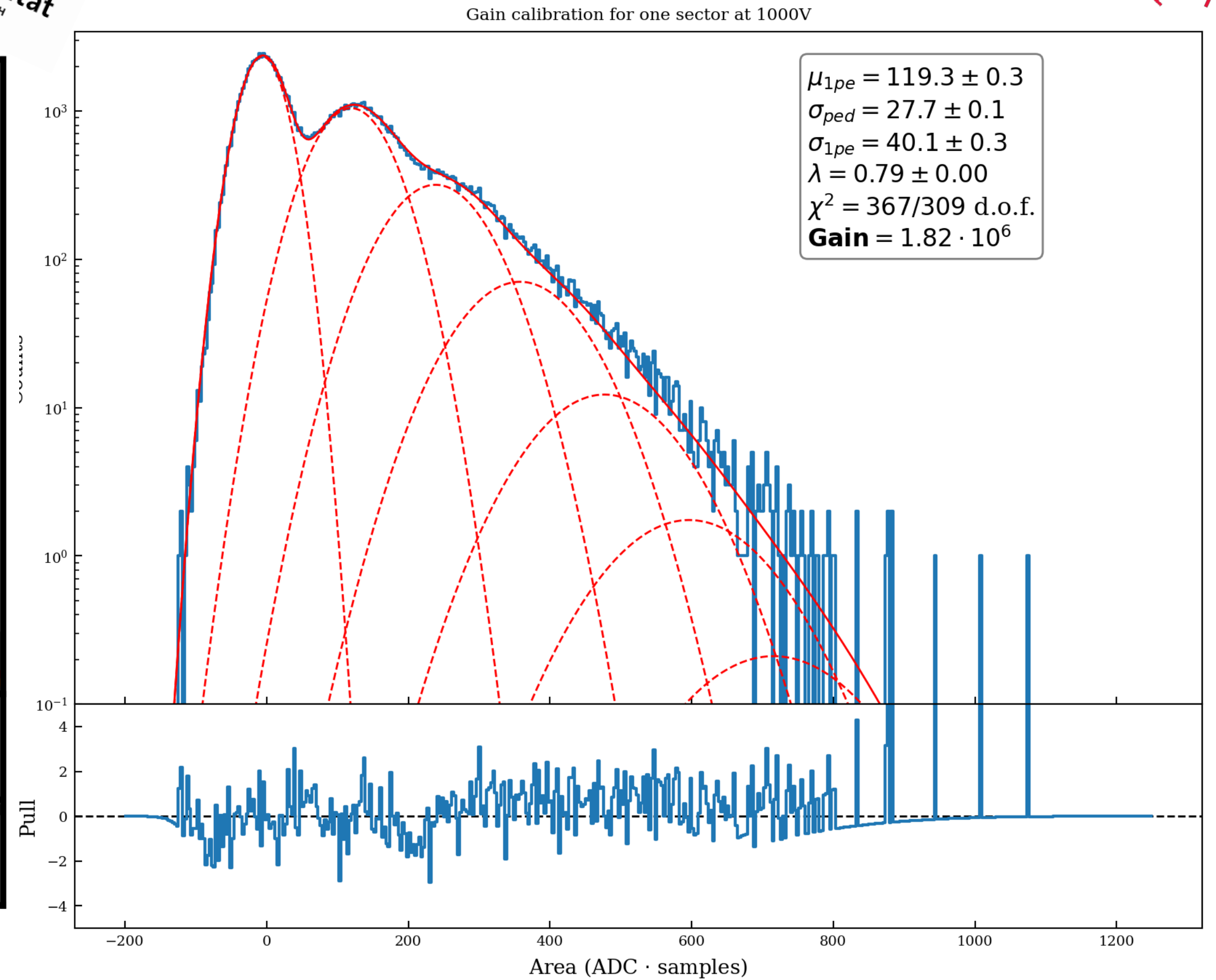
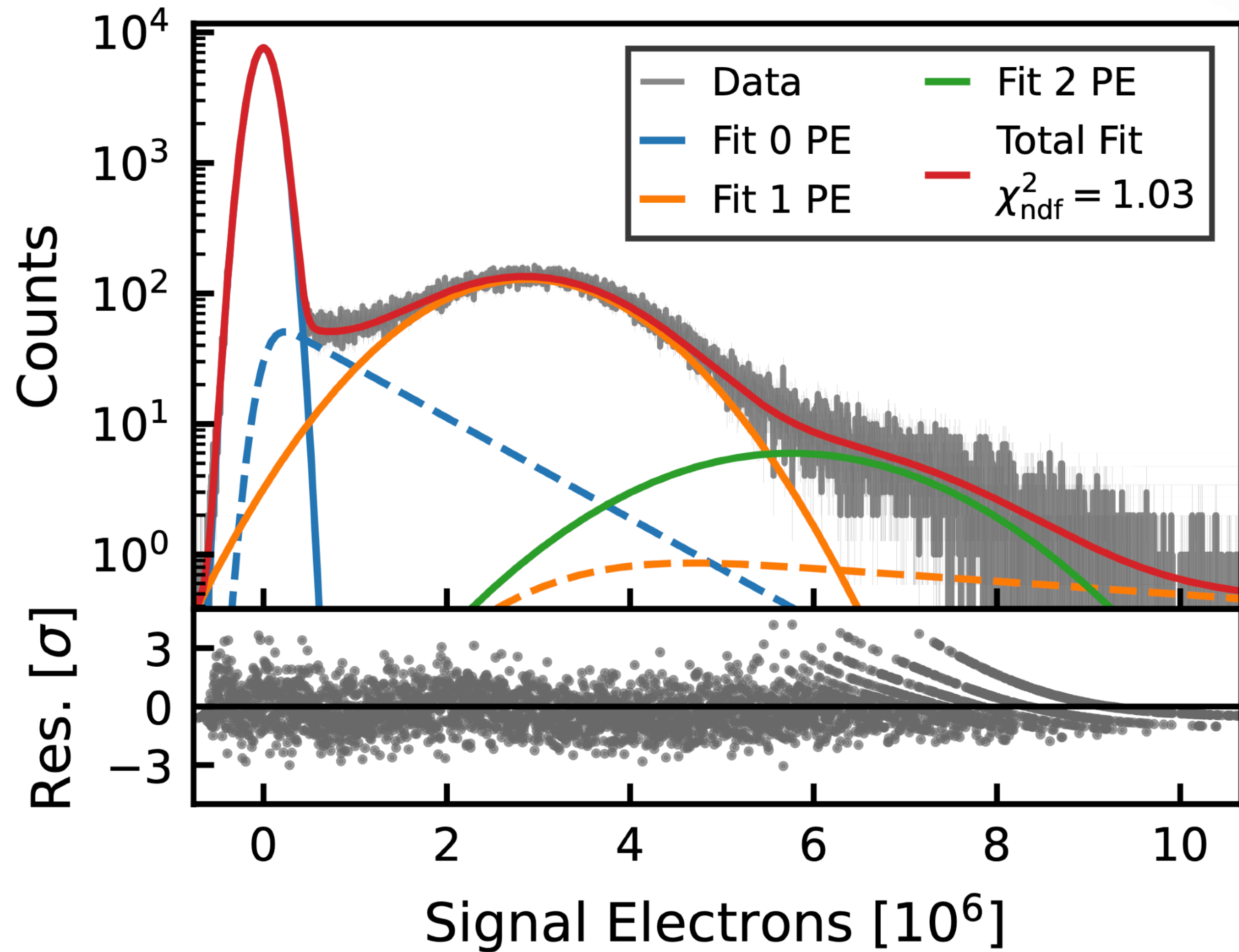
## Determination of the SPE response

- ▶ Fitting
  - ▶ Shape of PMT charge spectrum of **fully amplified** PE not known
  - ▶ SPE response under amplified due to sub-optimal trajectories through the dynode chain
  - ▶ Noise spectrum not known a priori
- ▶ Model independent approach
  - ▶ Full spectral shape of SPE response not needed
  - ▶ Only mean and variance of the SPE distribution + occupancy

<sup>1</sup>Saldanha, R., Grandi, L., Guardincerri, Y., & Wester, T. (2017). Model independent approach to the single photoelectron calibration of photomultiplier tubes. *Nuclear Instruments and Methods in Physics Research Section A: Accelerators, Spectrometers, Detectors and Associated Equipment*, 863, 35-46.

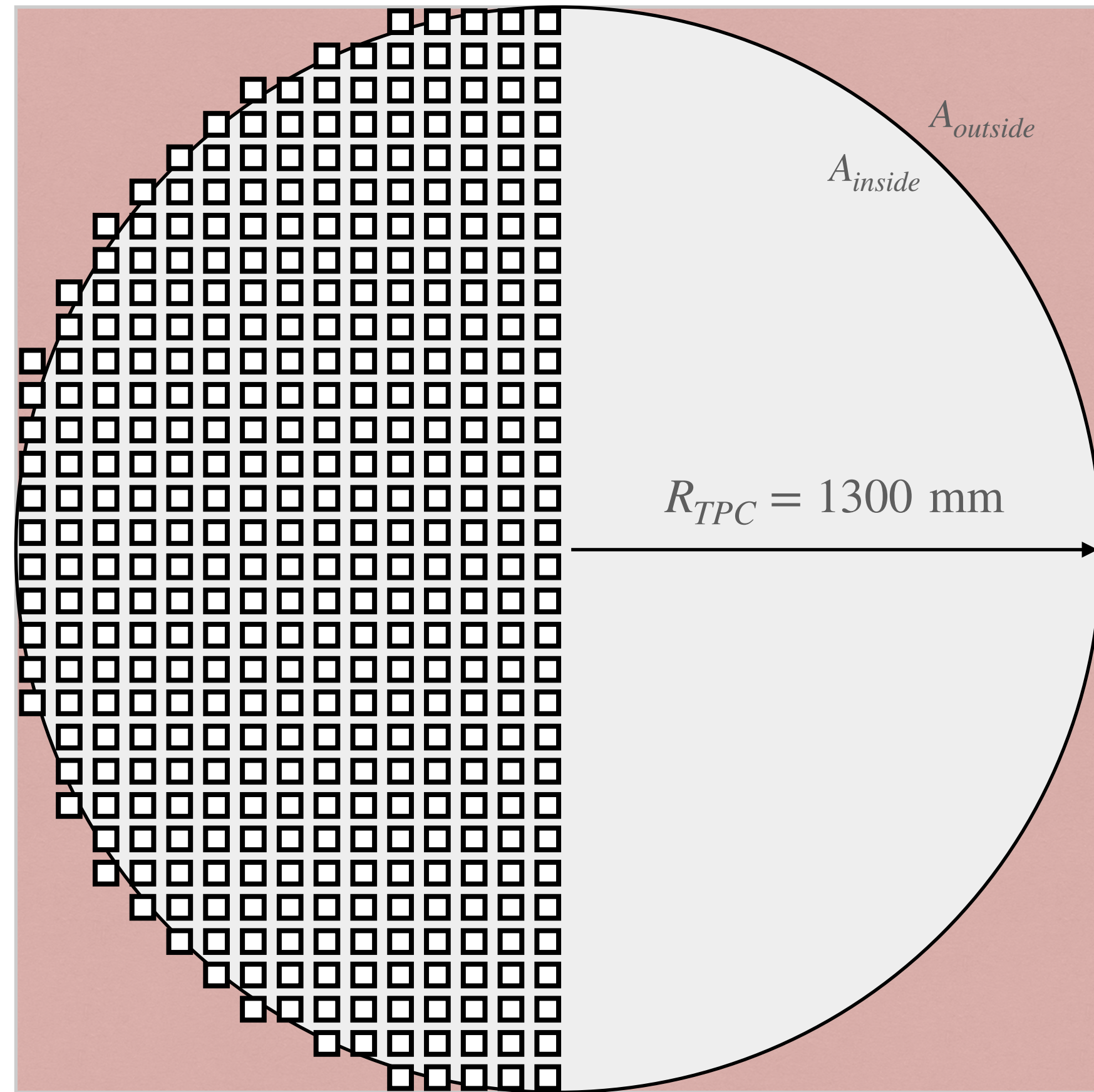


# SPE fitting





# Number of 2-inch PMTs in DARWIN



$$A_{outside} = (2 \cdot R_{TPC})^2 - A_{inside} = 1.45 \cdot 10^6 \text{ mm}^2$$

$$A_{inside} = \pi \cdot R_{TPC}^2 = 5.31 \cdot 10^6 \text{ mm}^2$$

$$A_{PMT} = (56 + 10)^2 \text{ mm}^2$$

*Margin*

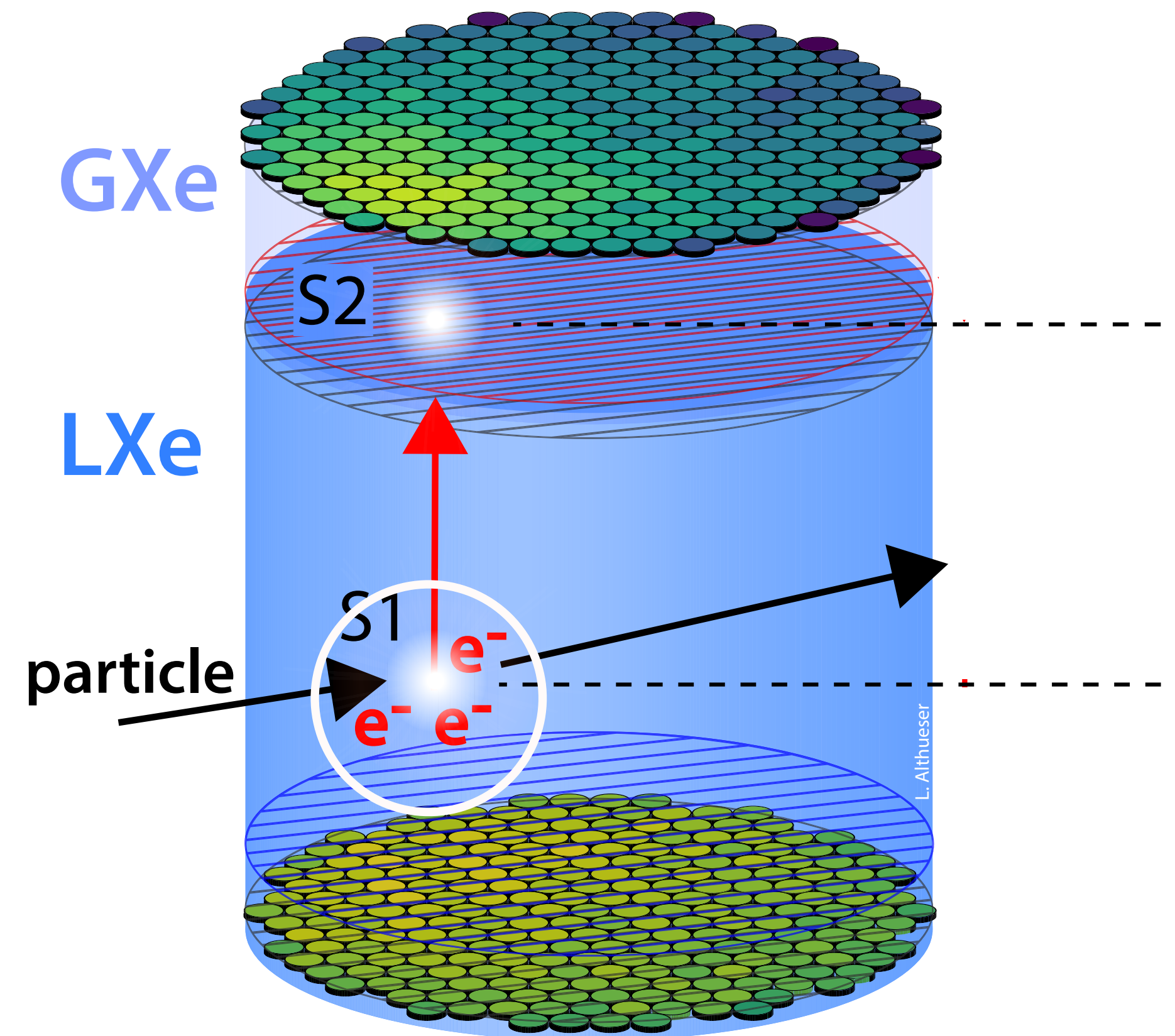
$$N_{PMT} = \frac{A_{inside} + \left\lfloor \frac{A_{outside} - A_{inside}}{A_{PMT}} \right\rfloor}{A_{PMT}} \cdot 2 \approx 2440$$



# Direct detection

## Dual-phase xenon TPC working principle

1. Particle interacts with the xenon atom, which ionizes and excites the xenon
2. Excited and ionized xenon forms dimer states and de-excites: S1 (mostly bottom PMT array)
3. Freed electrons drift up due to electric field
4. Between the gate and the anode, extraction field causes proportional scintillation of the xenon: S2 (mostly top PMT array).

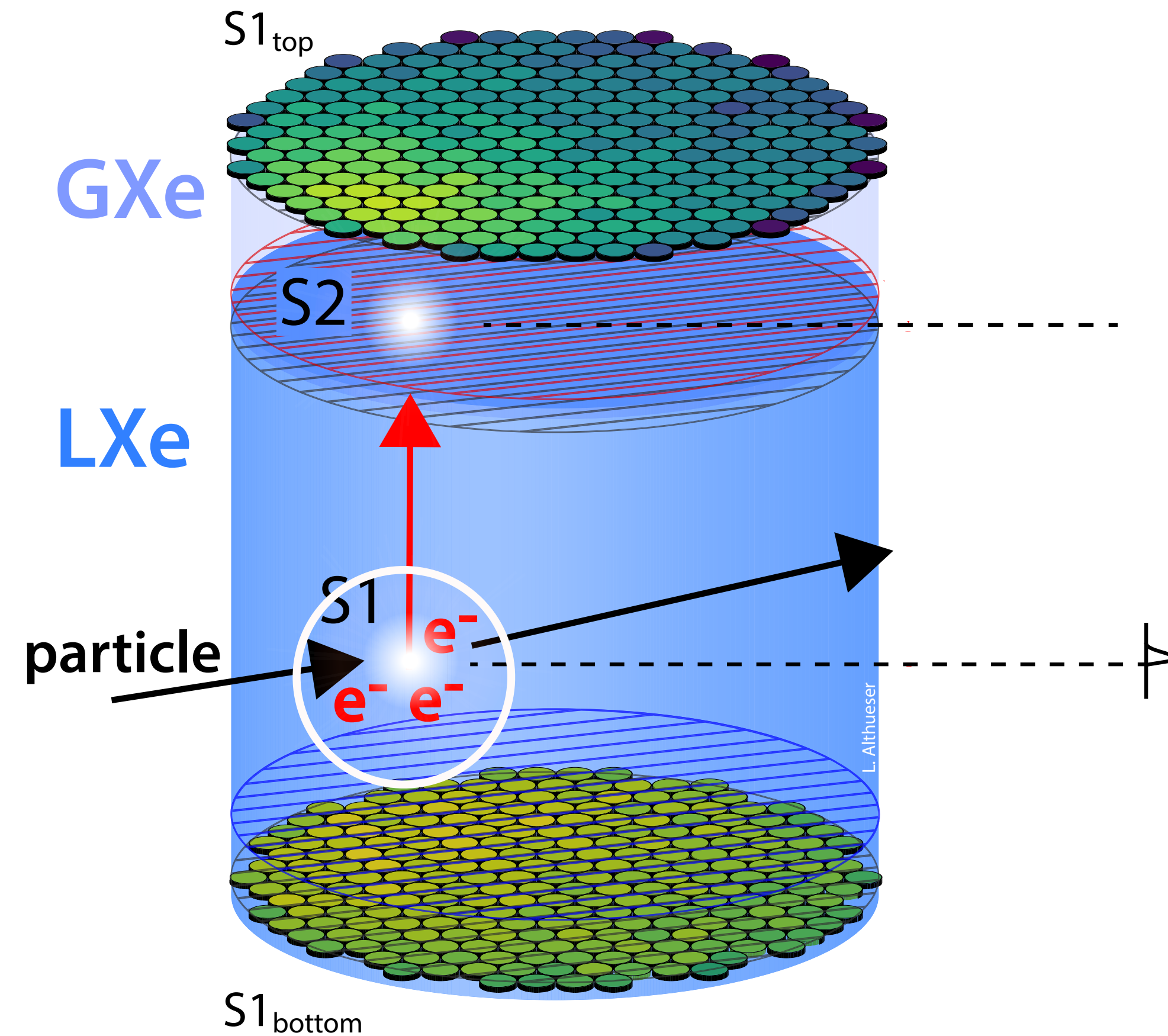




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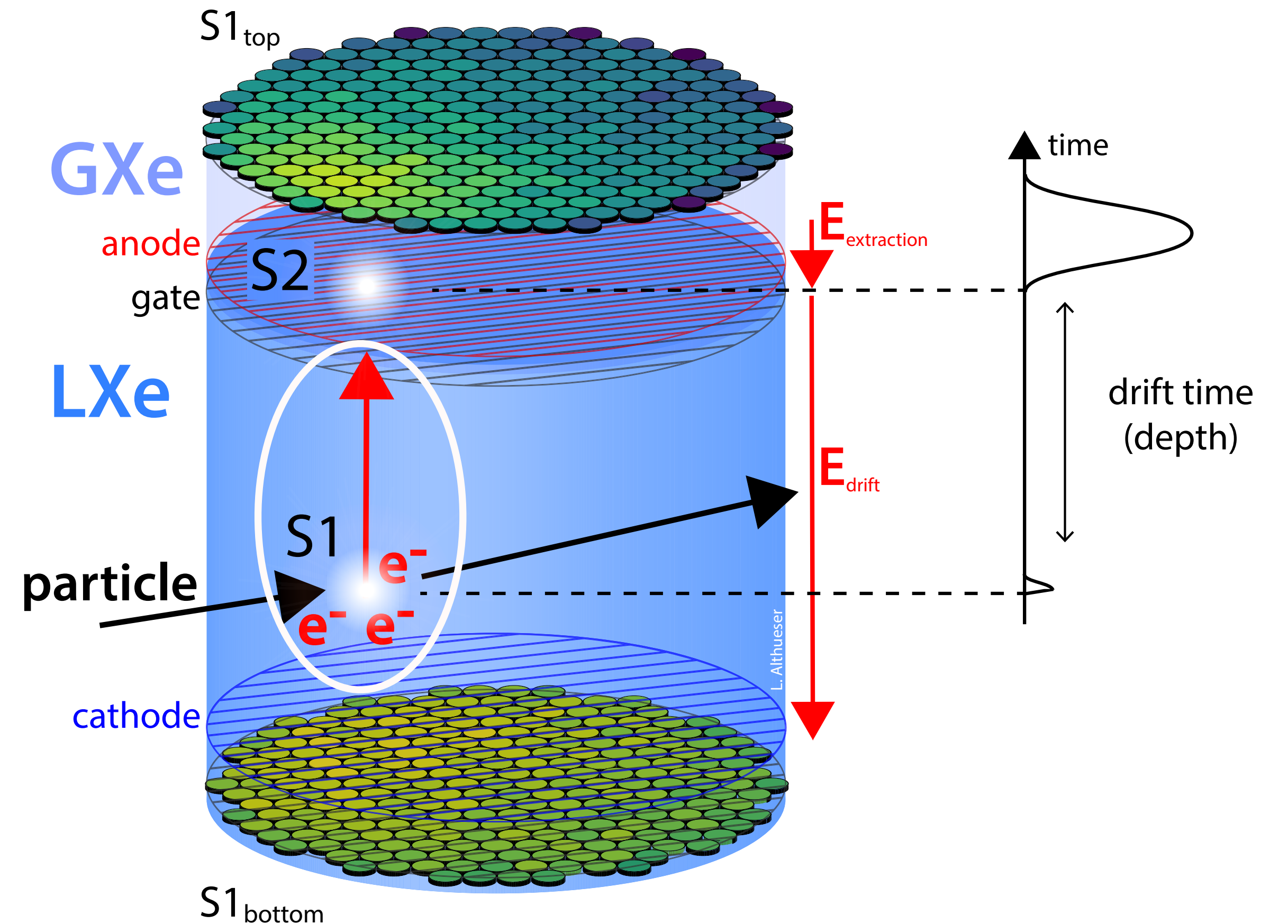




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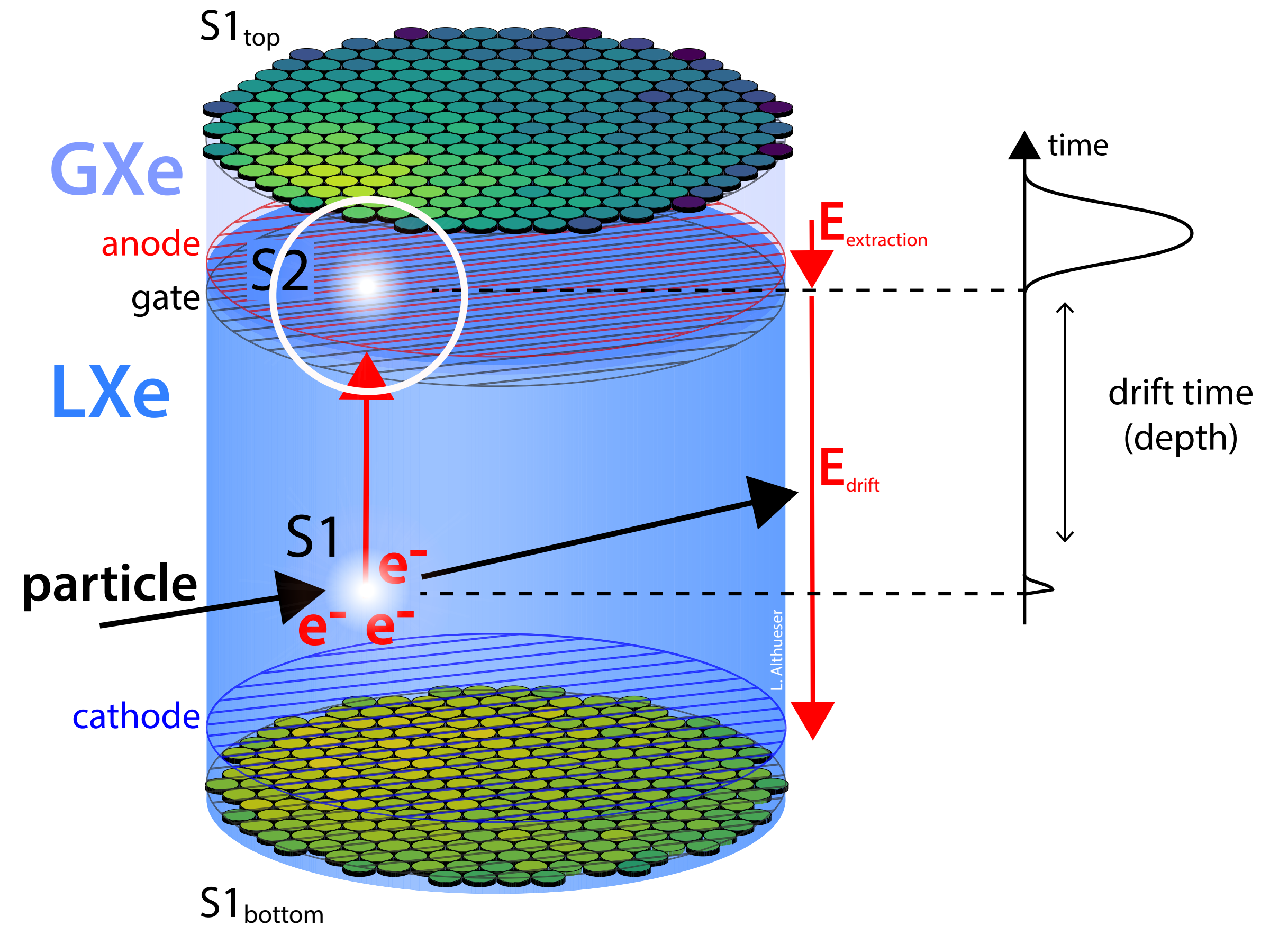




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