

ASTROCENT



Quality Control of PEN wavelength shifters for DarkSide-20k veto

Sarthak Choudhary¹

with André Cortez¹, Marcin Kuźniak¹, Maciej Kuźwa¹, Grzegorz Nieradka¹, T. Sworobowicz¹, Tomasz
Szczęśniak², Łukasz Świdorski²

On behalf of the DarkSide-20k Collaboration

1. AstroCeNT, Nicolaus Copernicus Astronomical Center, PAN

2. National Center for Nuclear Research



European Union



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POLAND

Outline

I. Introduction

II. ArGSet

- ✓ SiPM characterization

III. Data Analysis

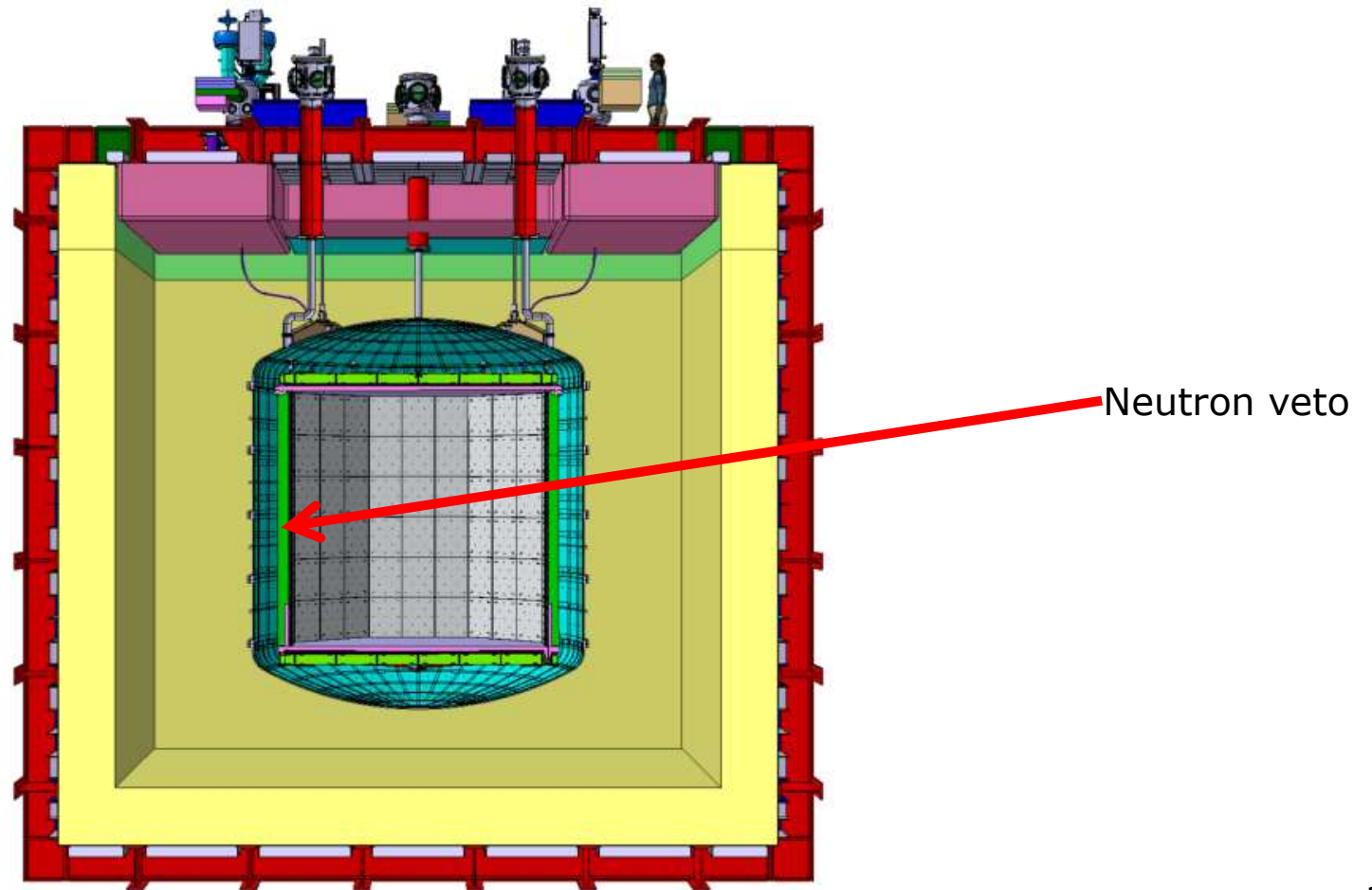
- ✓ Event selection cuts

- ✓ Argon triplet lifetime verification

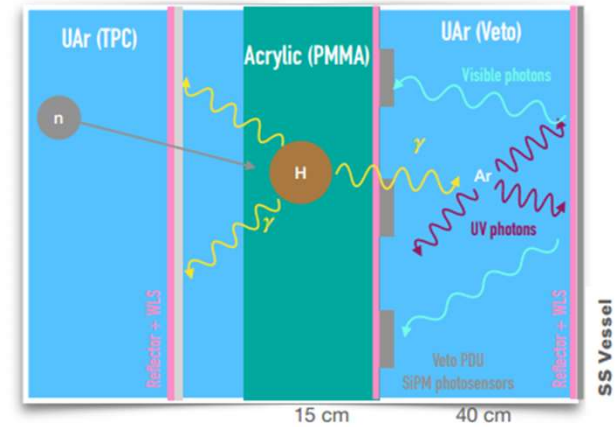
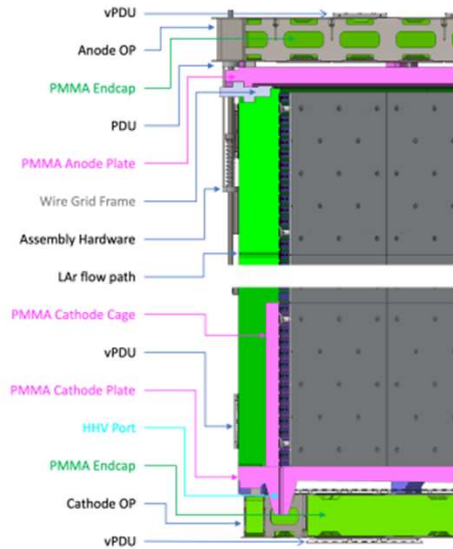
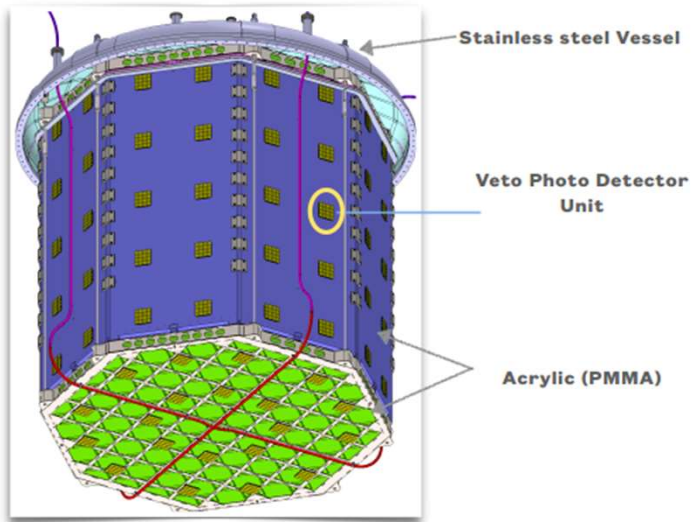
- ✓ Full WF charge distribution

IV. Summary and Conclusion

DarkSide-20k neutron veto



DarkSide-20k neutron veto



- Neutrons can mimic WIMP signal. **PSD** is useless against neutron events.
- The UAr volume between the SS vessel and **PMMA** serves as a veto volume with **~40 cm thickness**.

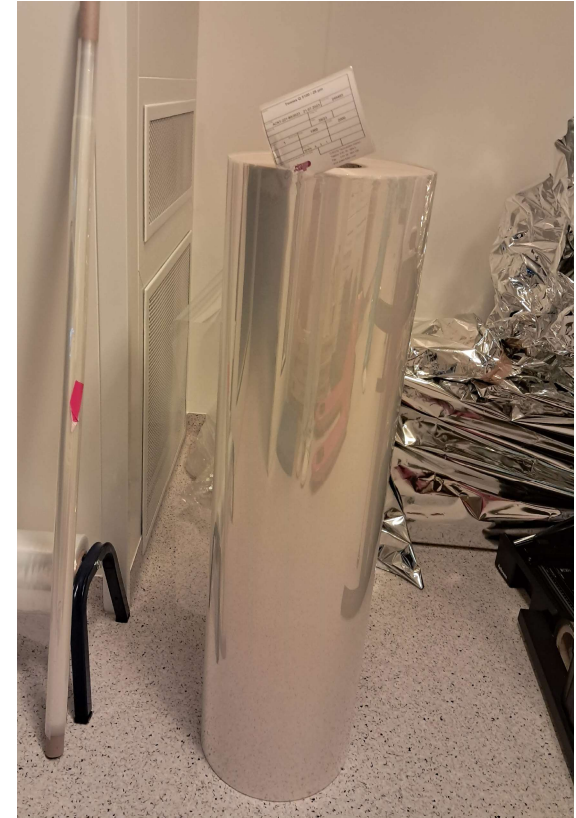
Veto Working Principle

- Neutrons are moderated in the acrylic shell and then captured by **Hydrogen**.
- H emits γ -rays **2.2 MeV**.
- γ -rays interact in the liquid argon buffers.
- LAr scintillation light is wavelength shifted and detected by **~1920 SiPM-based photosensors**.

Slide courtesy: I. Ahmad (IDM 2024)

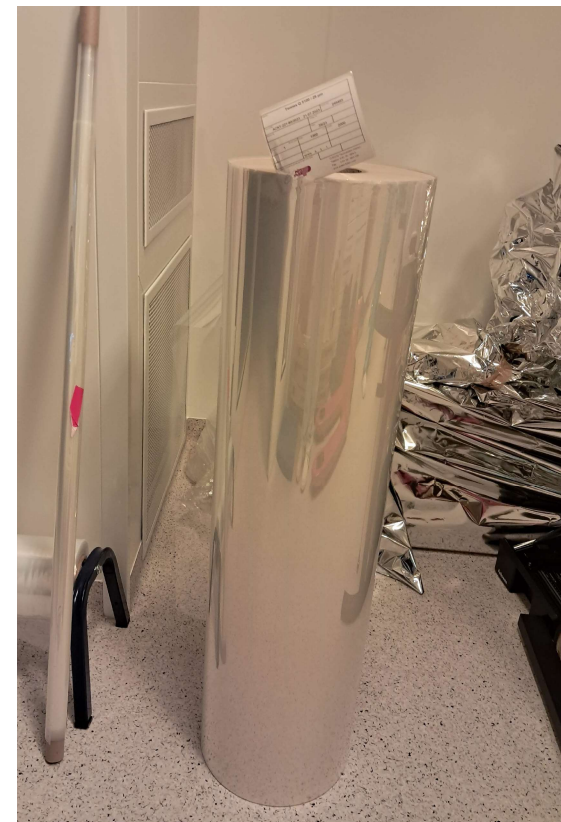
PEN rolls

- ❑ Polyethylene naphthalate (PEN) is proposed alternate material to tetraphenyl butadiene (TPB).
- ❑ PEN rolls:
More than 4000 m² of PEN procured and available at AstroCeNT in form of rolls. Sufficient excess material available for quality control tests.
- ❑ Roll:
Length = 2 km
Width = 1.2 m
- ❑ The rolls undergo uniformity test.



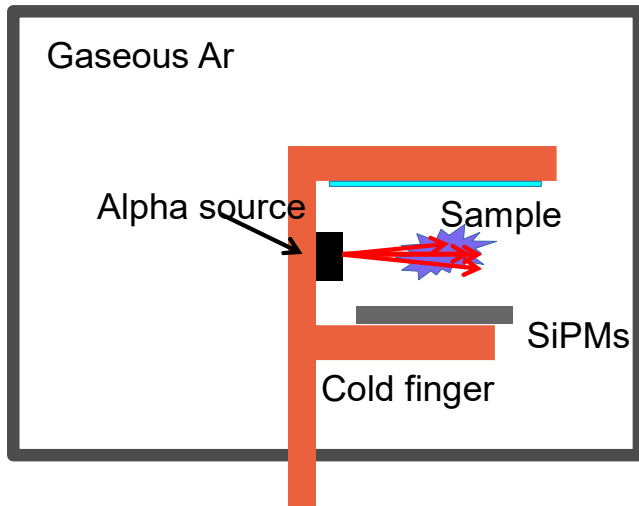
PEN quality control strategy

- ❑ Keep witness samples from each sheet of PEN to be installed in the detector
- ❑ Before sheet production test the uniformity of the roll with ~50 samples:
 - Each sample to be measured at 190 nm excitation with a spectrophotometer equipped with an integrating sphere
 - Subset of samples to be tested with 128 nm excitation at room temperature
 - Several samples to be tested at 128 nm and in cryogenic conditions (currently the main bottleneck)
- ❑ Correlate cryogenic measurement result with other methods



ArGSet

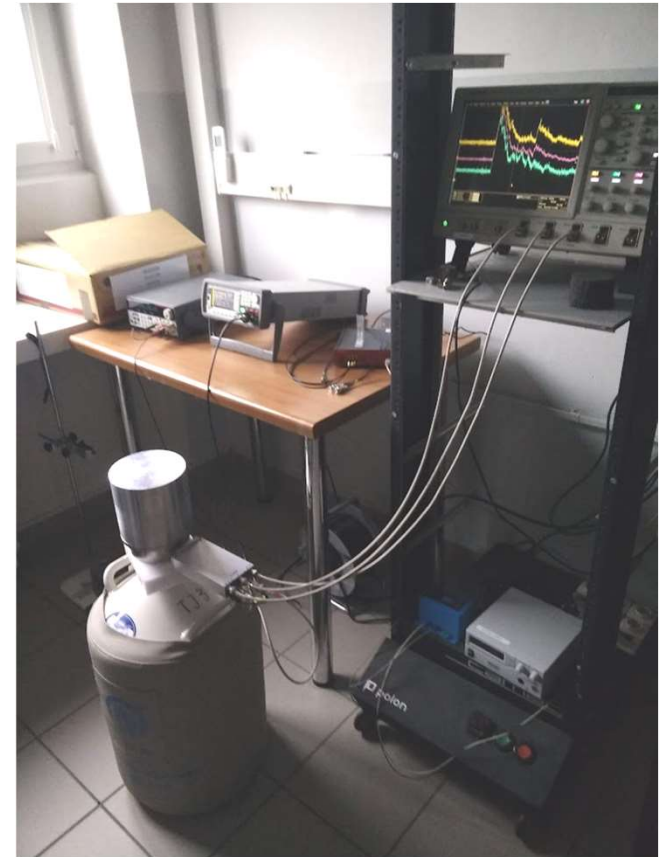
- An ArGon Gas Setup for measurement of Wavelength Shifting Efficiency (WSLE)
- WSLE depends on the excitation wavelength AND temperature
- ArGSet is an alpha-excited gaseous Ar cell with a cryogenic stage



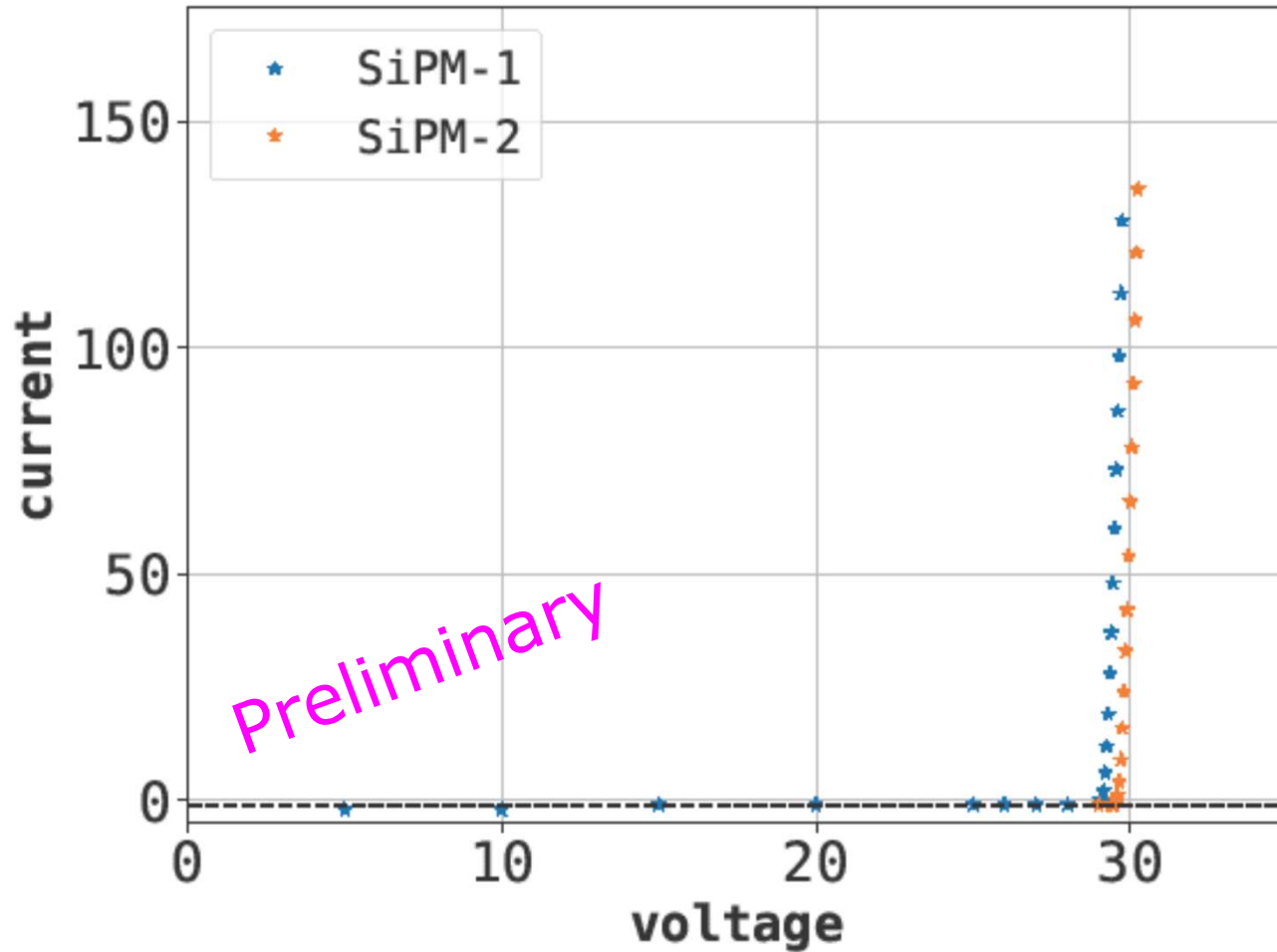
Two SiPMs:
Hamamatsu S14160



Ref: Kuzniak, LIDINE-2023



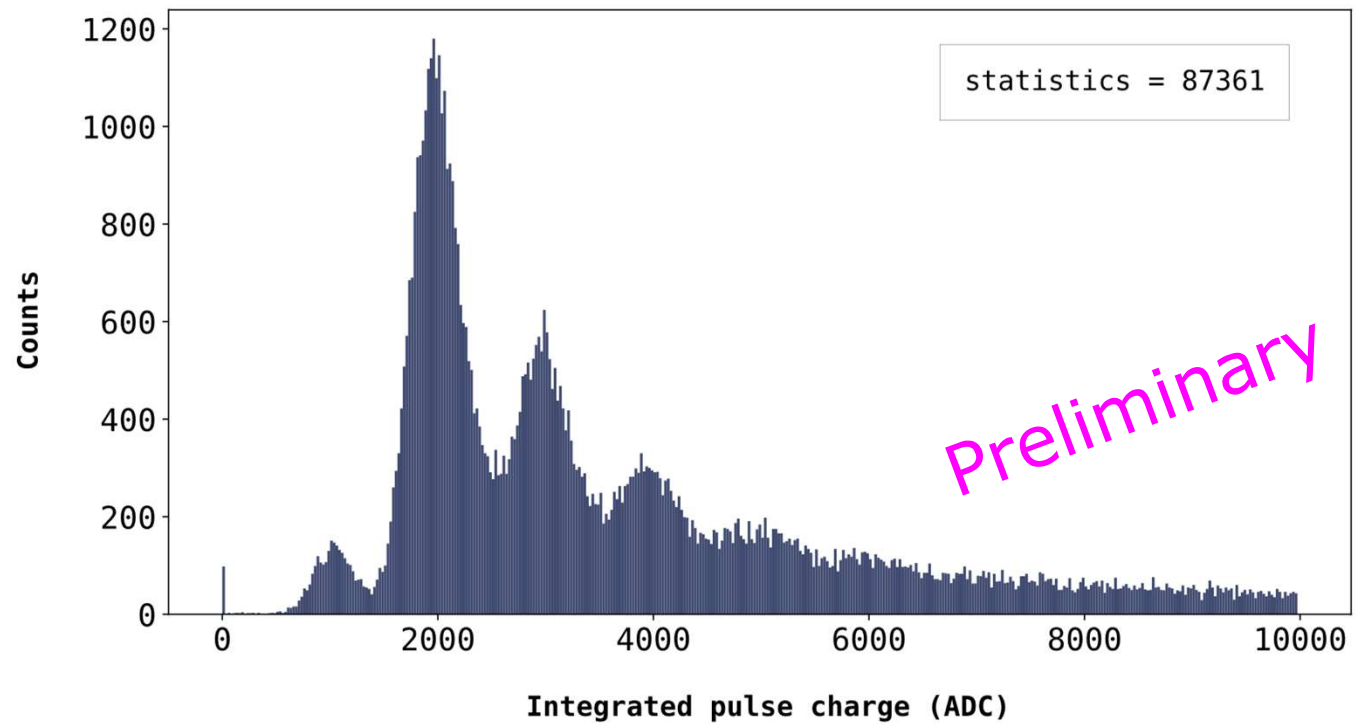
VI curve



- Found the breakdown voltage for both SiPM's
- The SiPM are operated at the same Over-Voltage, i.e., gains balanced.

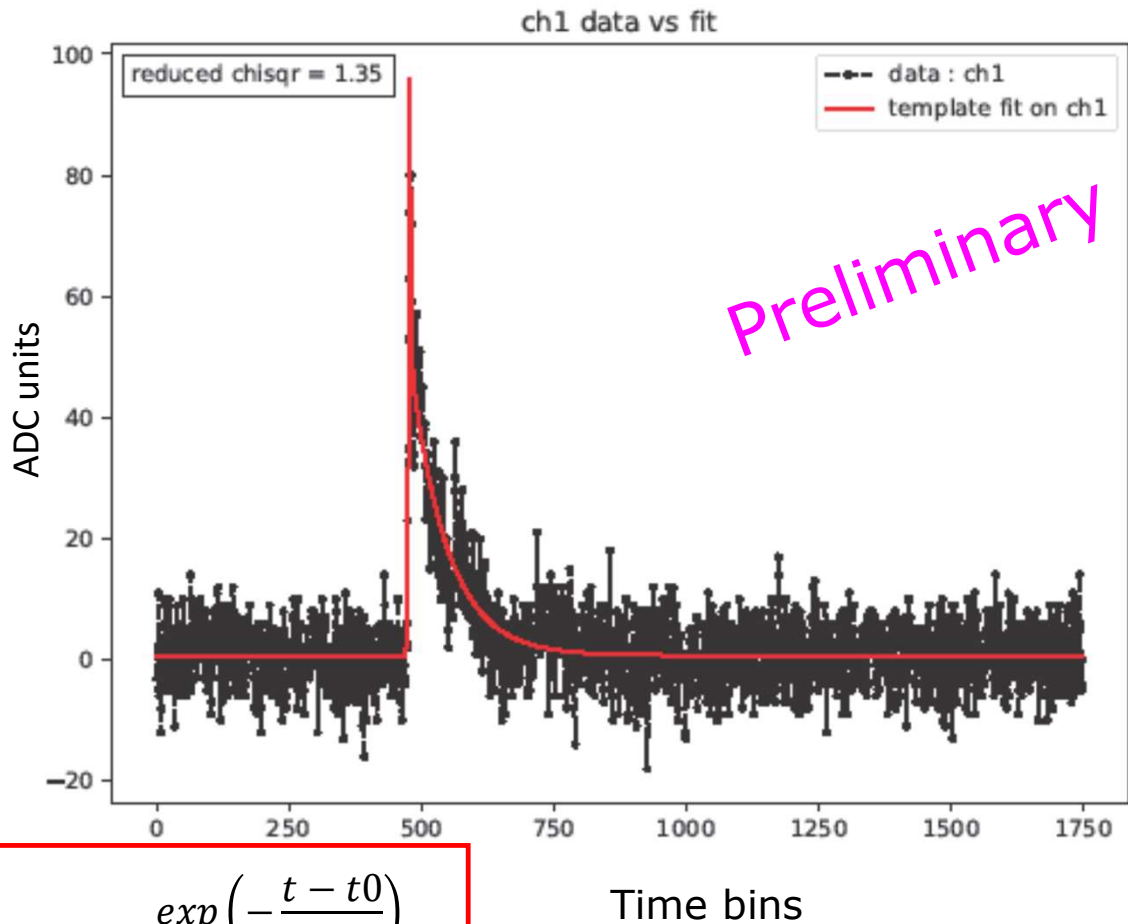
PhotoElectron Charge Spectrum (fingerplot)

- ❑ Cell is in vacuum (no Ar)
- ❑ Data taken with LED pulser.
- ❑ SiPM at Liquid Nitrogen temperature.
- ❑ Filtered waveform from summation of two SiPM's
- ❑ AR filter is tuned to waveform for these SiPM's (see next slide).
- ❑ SPE charge: 1564 +/- 14.57 (ADC)
- ❑ SNR: 4.47



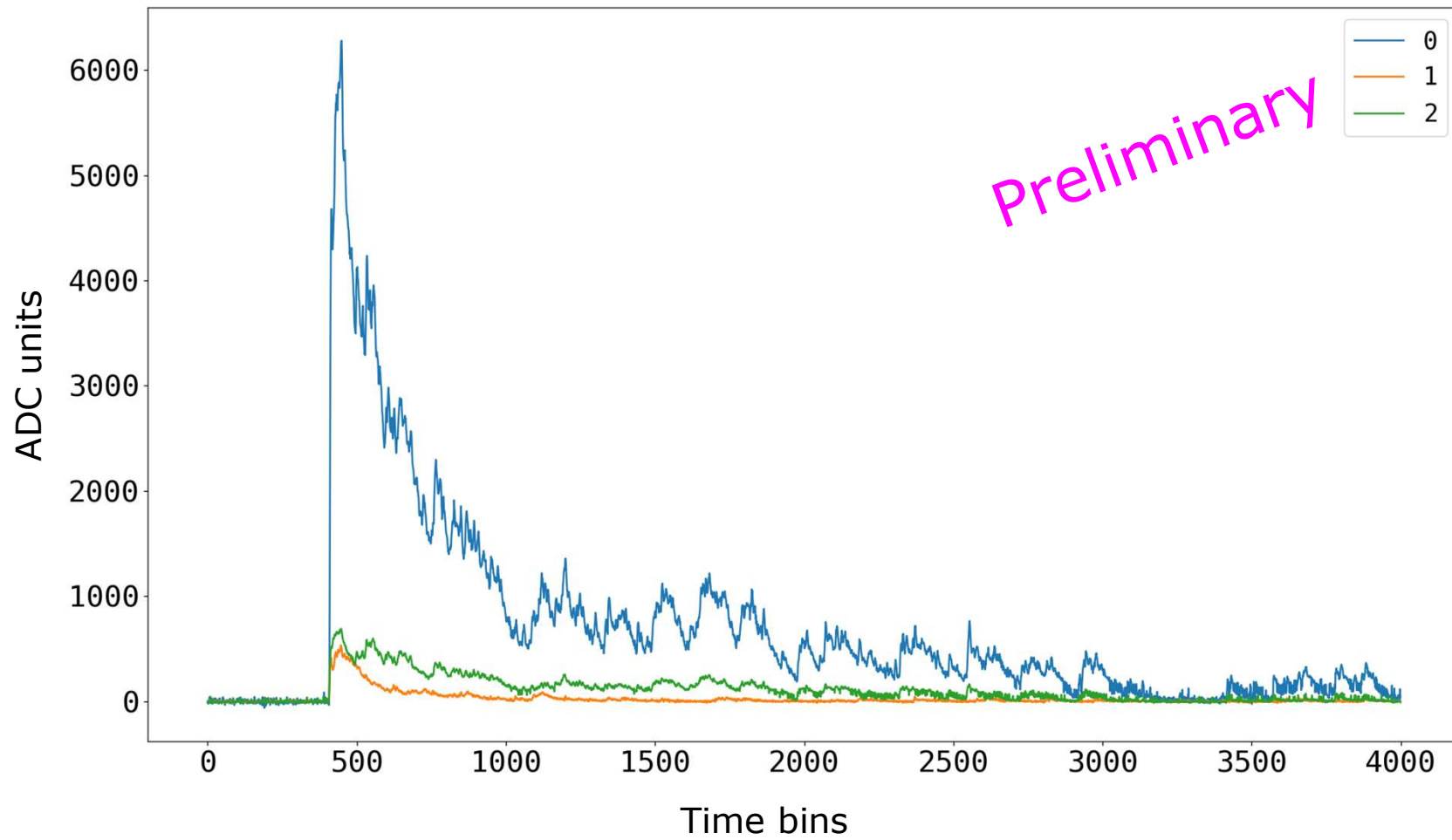
AR filter optimization

- ❑ Auto-regressive (AR) Filter used to reduce noise
- ❑ AR filter parameters were tuned accordingly to match the pulse parameters



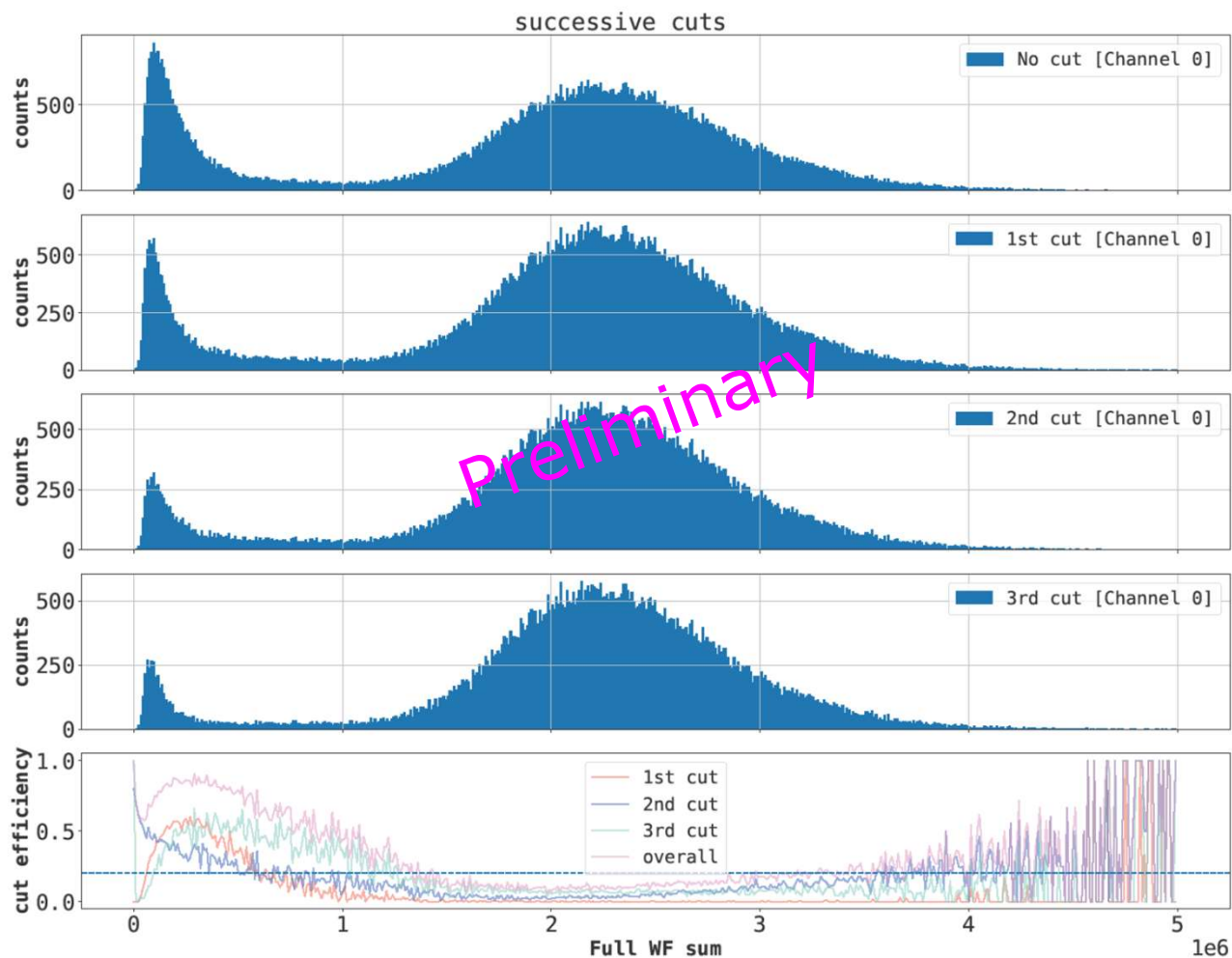
$$(1 - \alpha) * gauss(t - t_0) + \alpha * heaviside(t - t_0) * \frac{\exp\left(-\frac{t - t_0}{\tau}\right)}{\tau}$$

Raw waveform

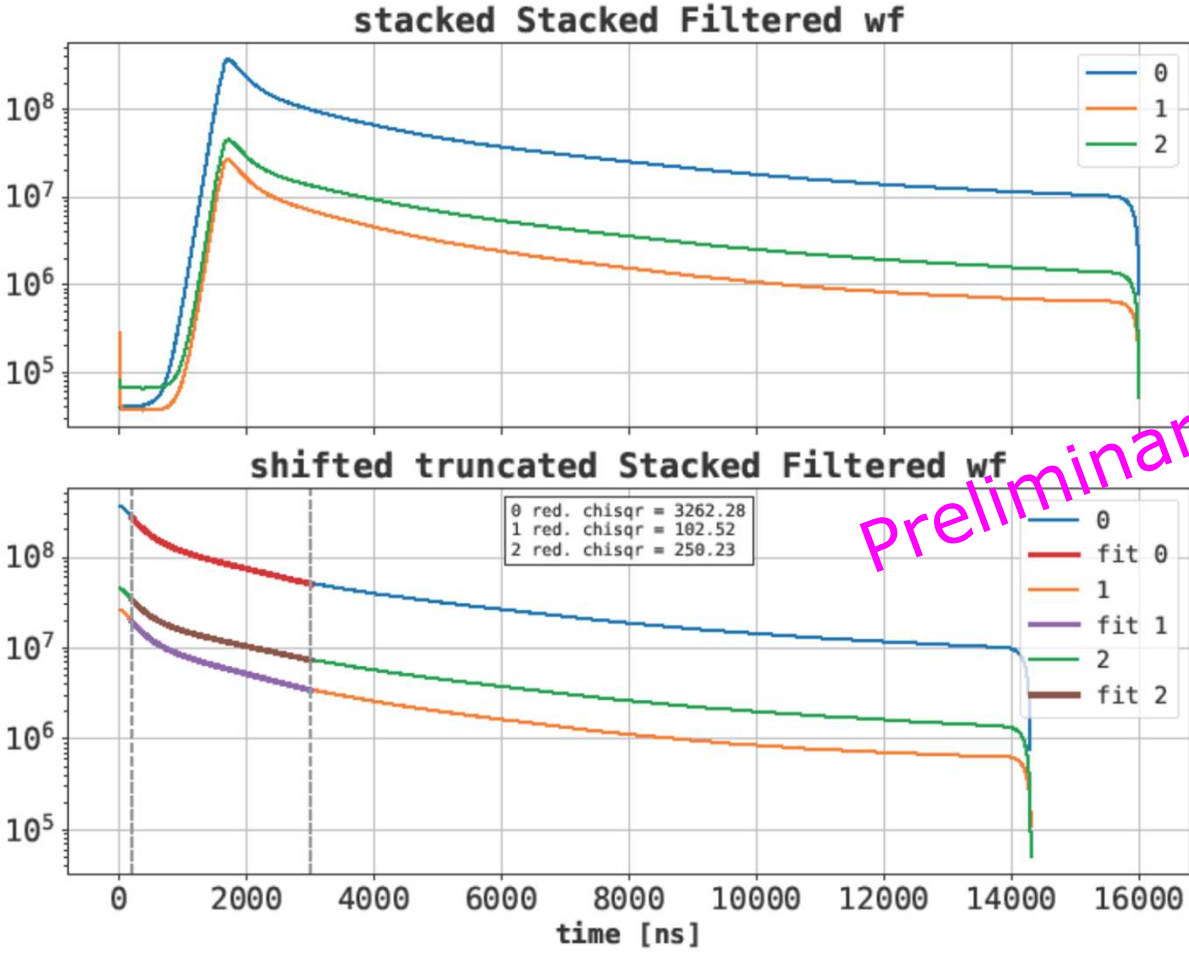


Event selection cuts

- Events selection cuts are based on following quantities:
- 1st cut: Sum of PreTrigger window
- 2nd cut: Centre of Mass for waveform
- 3rd cut: Time separation between pulses in SiPM-1 and SiPM-2

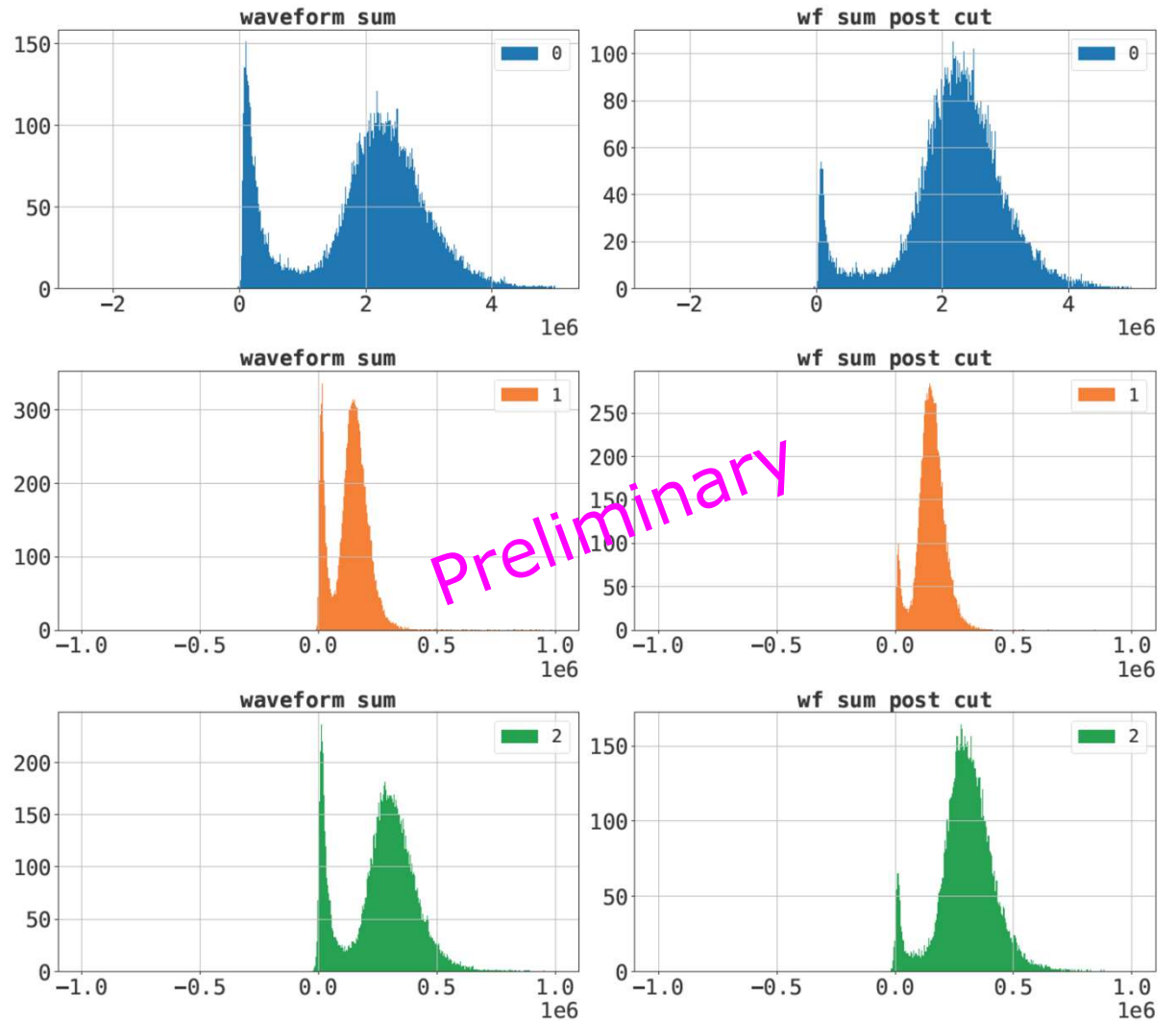


Cross check of gaseous Ar triplet time constant



Mean lifetime =
2450 +/- 2.722 ns

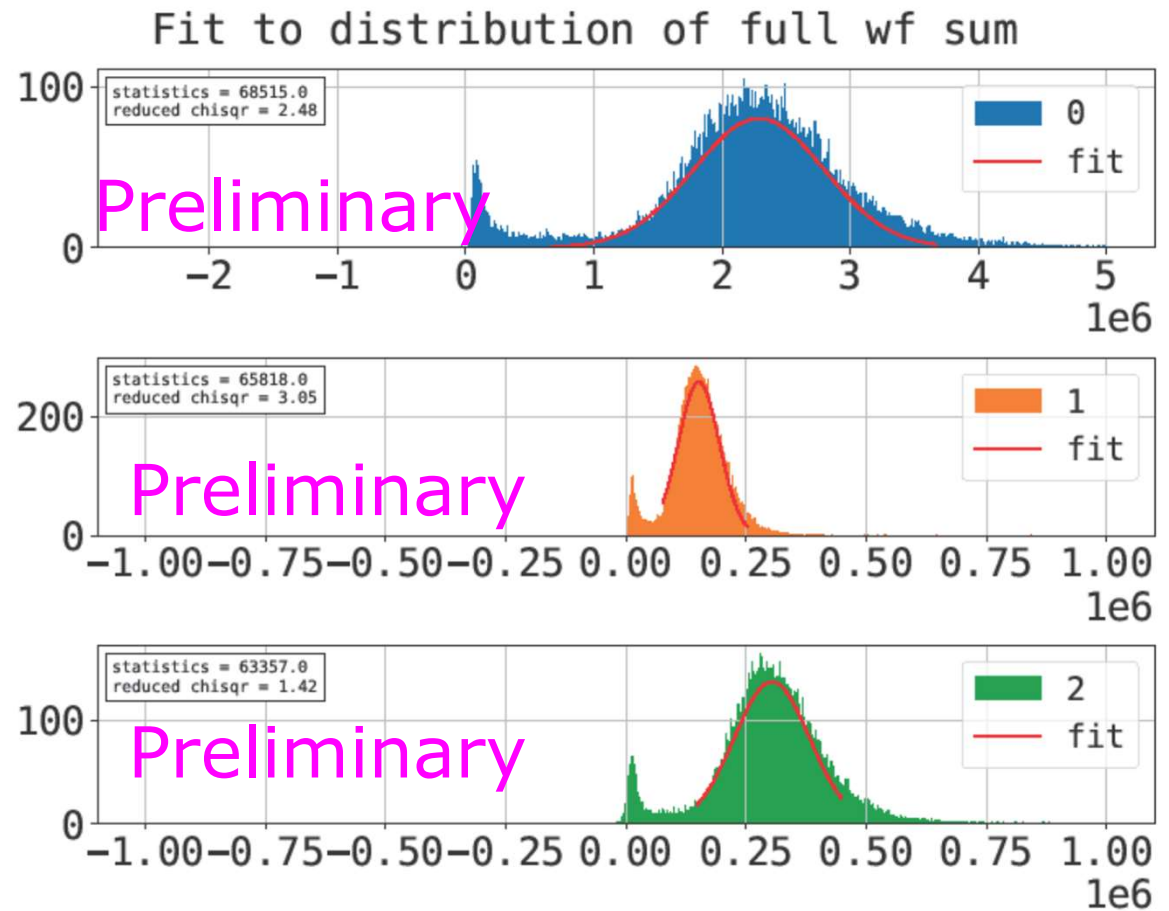
Comparison of charge distribution before and after applying the cuts



Charge distribution post cut

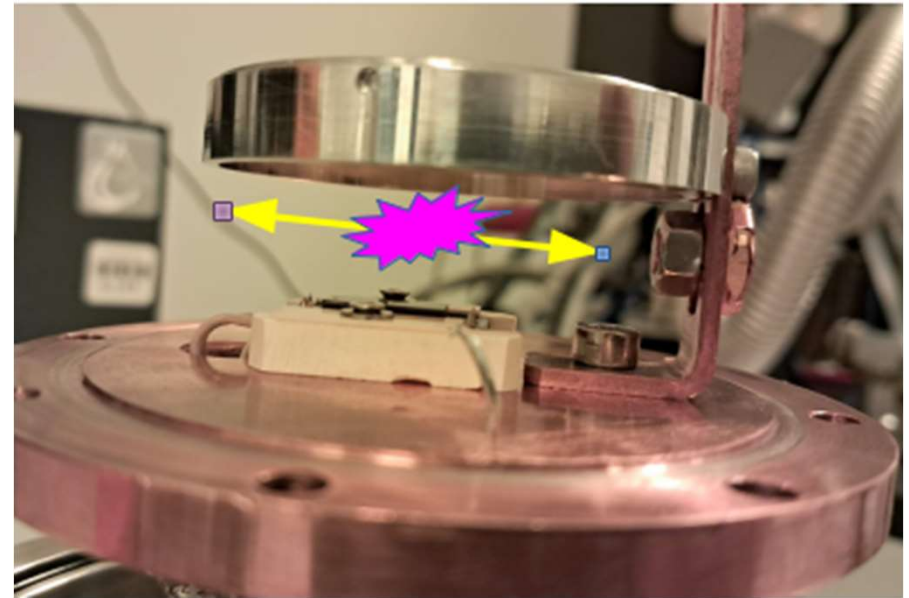
Ch 0	
Integral (ADC) M_{eff}	2.286e+06
Photoelectrons N_{pe}	1462

- Number of detected photoelectrons from an alpha scintillation event is consistent with expectations



Status

- Following the SiPM calibration, collected cryogenic data for 4 PEN samples and TPB reference samples (twice)
- Revealed instrumental issues:
 - Intermittent noise – efficiently removed with cuts in most of the runs – likely a grounding issue to be investigated and fixed
 - Asymmetry of signal collected in both SiPMs varies from run to run
 - Could be a hint of Ar pressure different between runs
 - Dangerous, as it can move the position of the Bragg maximum/primary VUV light source
 - Under investigation, pressure transducer to be added
 - Working out a correction for the collected data
 - Temperature stability acceptable for several hours of data taking
- Comparison of samples as soon as the varying asymmetry issue is understood
- For now planning to proceed with the room temperature and near UV measurements



Summary

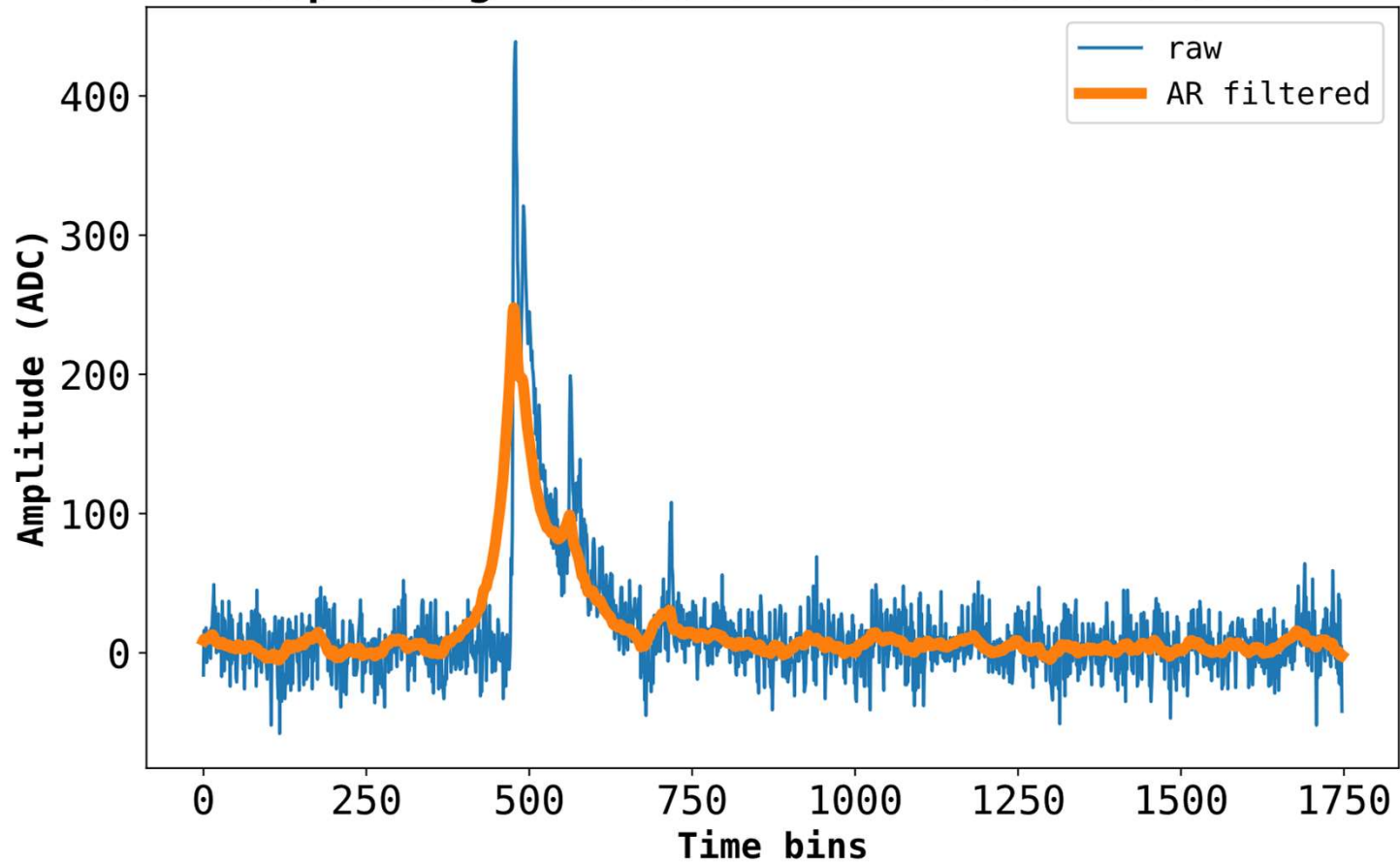
- A low temperature setup for measurement of wavelength shifting efficiency at VUV wavelength is presented
 - Overall the light yield and performance consistent with expectations
- SiPM calibration and analysis flow is in now place
- The application for quality control of PEN wavelength shifters is discussed:
 - Several samples already measured, but more work on systematic needed for a robust comparison
 - Better control/monitoring of gaseous Ar pressure is likely needed
- Starting to characterize larger quantities of PEN samples:
 - With near UV excitation at room temperature
 - With VUV at room temperature

Thank You

back up slides

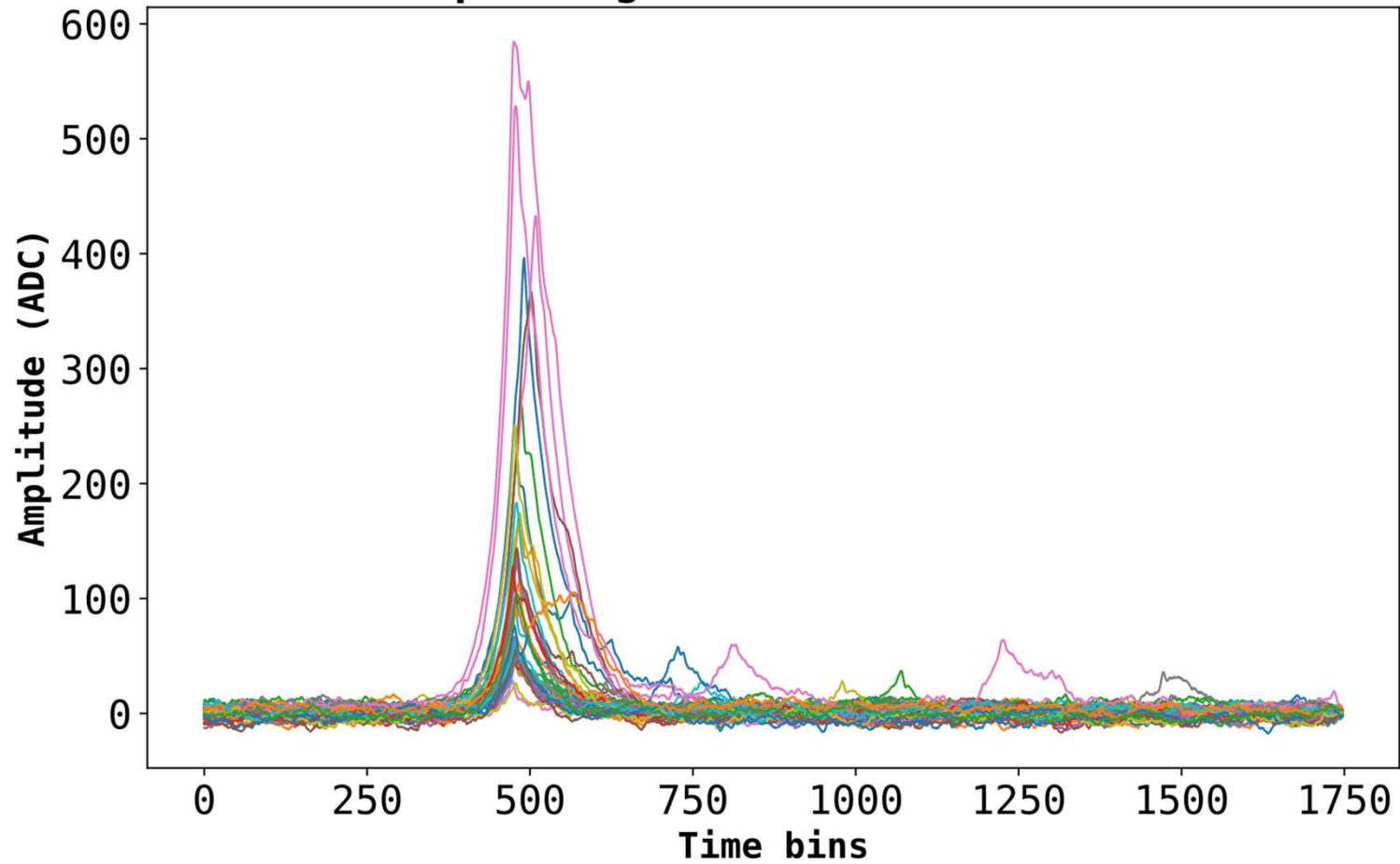
Raw vs AR filtered WF

coplotting raw and AR filtered waveforms



AR filtered

Over-plotting AR filtered waveforms



Charge distribution post cut

	Ch 0
Mean	2.286e+06
N _{pe}	1462

