3. Dark Maffer

3.1 Evidence for DM

Galaxy Rotation Curves
 (Jour Oorth 1932, Vera Rubin 1970)
 Ourge Click (radius R)
 Id

$$I_{1} \text{ Bulge } (approx. spherical) \\ \overline{T} = \frac{G \text{ in } H(r)}{r^{2}} \stackrel{!}{=} \text{ in } \frac{v^{2}(r)}{r} \\ v(r) - \sqrt{\frac{G \cdot M(r)}{r}} = \sqrt{\frac{4}{3}\pi Gr^{2}} \frac{g}{balge}$$

2) Dick: a little more complicated

3) Outside file disk
is galaxy can be beated as a pointlike

$$\overline{F} = \frac{Gun H(R)}{r^2} \stackrel{!}{=} un \frac{u^2(r)}{r}$$

is $v(r) = \sqrt{\frac{G \cdot H(R)}{r}}$
 \sqrt{r}
bulg disk outside r

Tiny fluctuations in Tacross the sky from primordial quantum fluctuations can be observed on the CTCB.

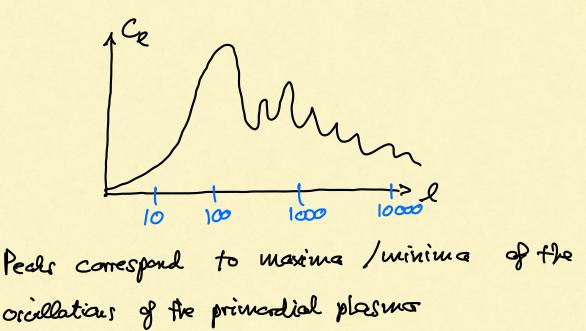
The principality plasma ascillates as matter is drown towards overdense regions, falls in, get pushed back by radiction pressure, etc.

L7 CHB power spectrum $T(\theta, \phi) = T \alpha_{lm} Y_{lm}(\theta, \phi)$ temperature map in galactic coordinates

$$a_{lm} = \int dar \theta \, d\phi \, Y_{lm}^{*}(\theta, \phi) \cdot T(\theta, \phi)$$

$$define \quad \left[C_{l} \equiv \frac{1}{2l+1} \sum_{m} |a_{lm}|^{2} \right]$$

$$CMB \quad Prover Steatrum$$



For instance: First peak around R ~ 100 corresponds to an oscillation made whose oscillations were interrupted by recombination when it had reacod its first maximum

Hore DN tends to suppress the peaks because photons have to travel out of deeper potential wells is more gravitational realshift.

· Structure Formation 17 see Songan lecture from the CMB Primardial density flectuation calculate subsequent explicition of functions until today, where they have evolved into galaxies Compare predicted distribution of galaries to observations to constrain properties of DO1 Conclusion: DM should be non-relativistic early cn.

3.2 DM condidates - Overwiew

- · Weakly Interacting Massive Particles (WIMPs)
- Axions
- · Sterile Neutrinas
- · Primordial black holes