



Dark Matter: *recipe to search it at colliders*

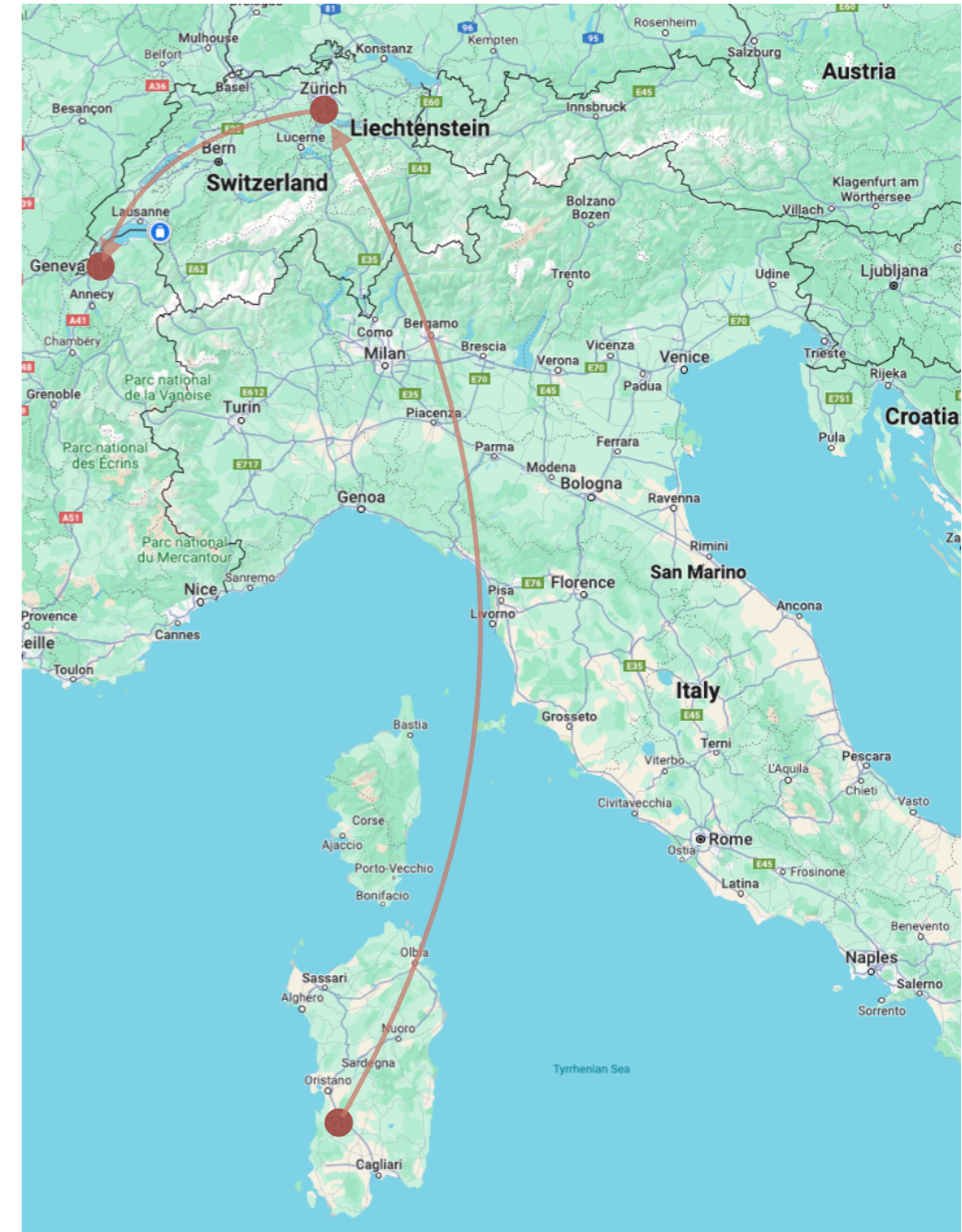
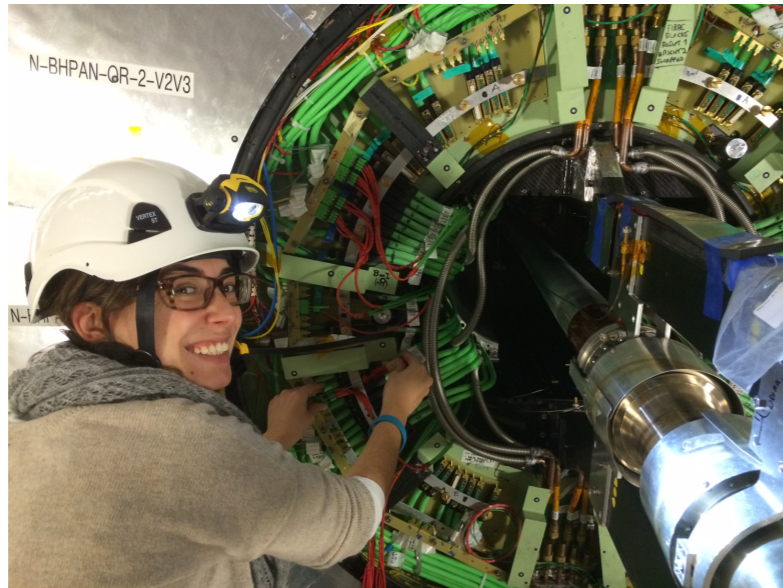
Deborah Pinna

(University of Wisconsin-Madison)

USCMS Undergraduate Summer Internship

10 July 2024

Few words about me ...



Deborah Pinna

- Scientist at University of Wisconsin-Madison (based at CERN)
- PhD in Zurich at UZH
- originally from Sardinia (Italy)

Research

- Searches for Dark Matter, Beyond the Standard Model particles and Standard Model measurements at the CMS Experiment

Hobbies

- climbing, painting, traveling

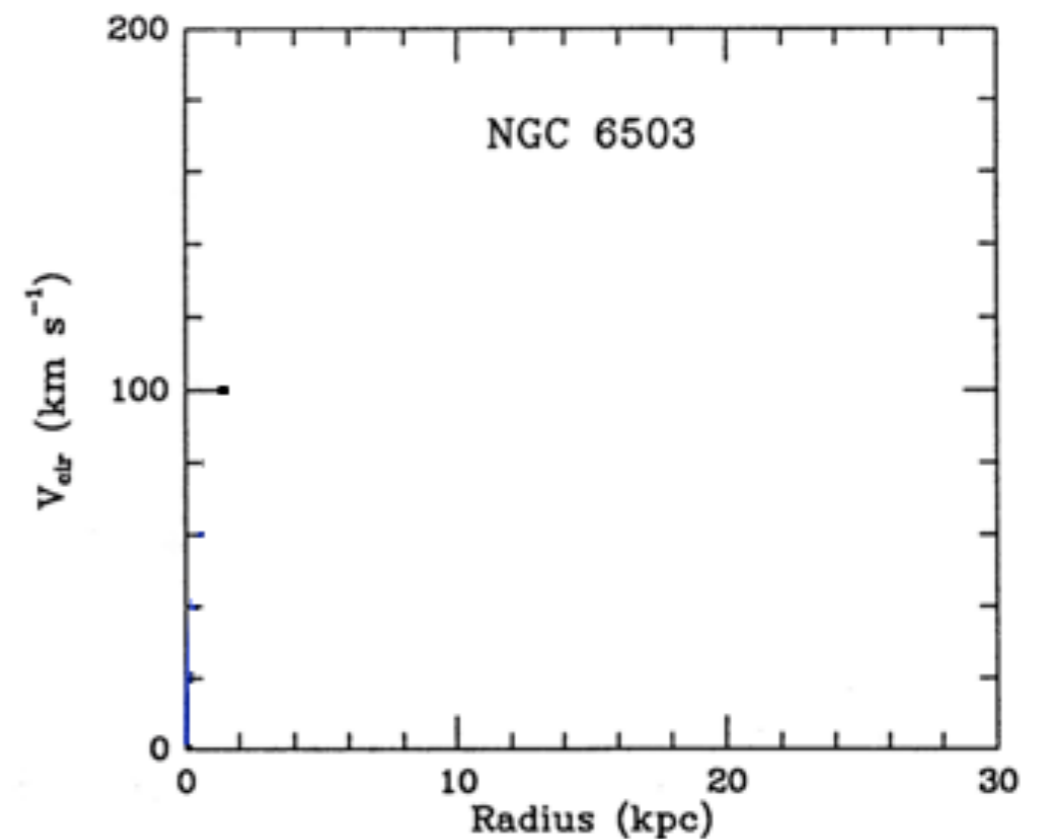
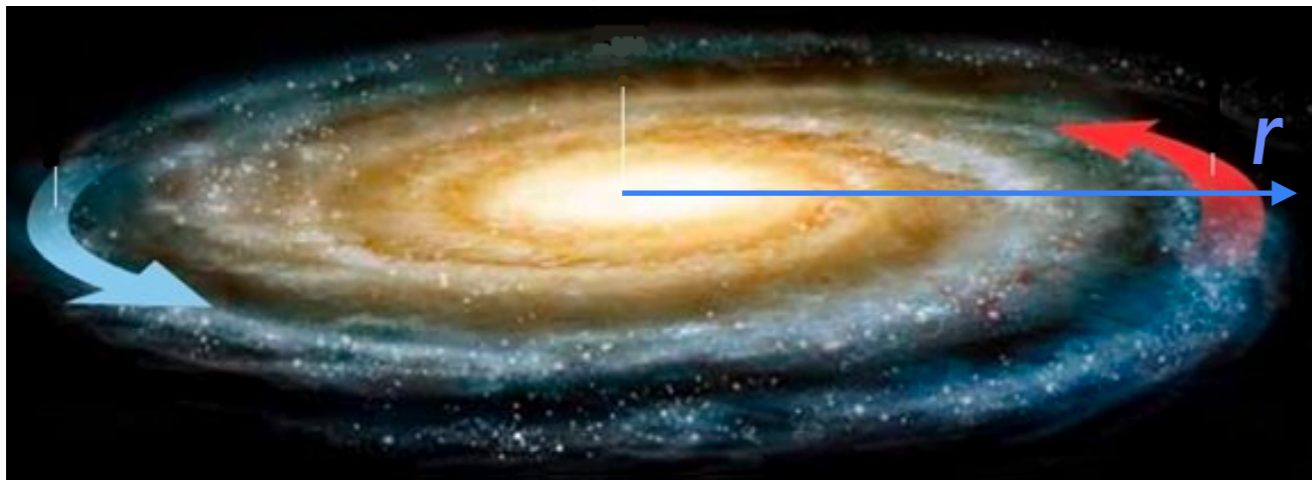
Dark matter, why?

► Different empirical evidence of DM from astrophysical observations at different scales

- first indication from Zwicky's dispersion velocity measurements of galaxies in Coma cluster
- existence of DM confirmed by measurements of stars and gas circular velocities within a galaxy by Ford and Rubin

- from Newtonian dynamics expected velocity $v(r)$ of these objects:

$$v(r) = \sqrt{\frac{GM(r)}{r}}$$



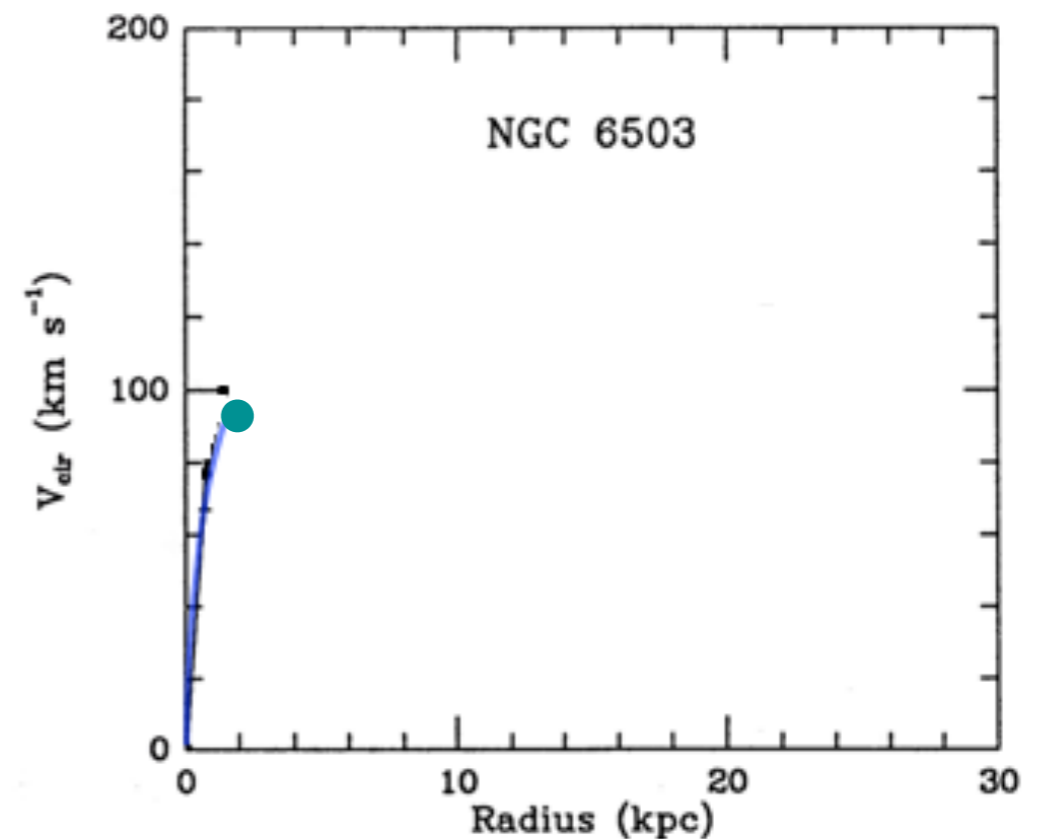
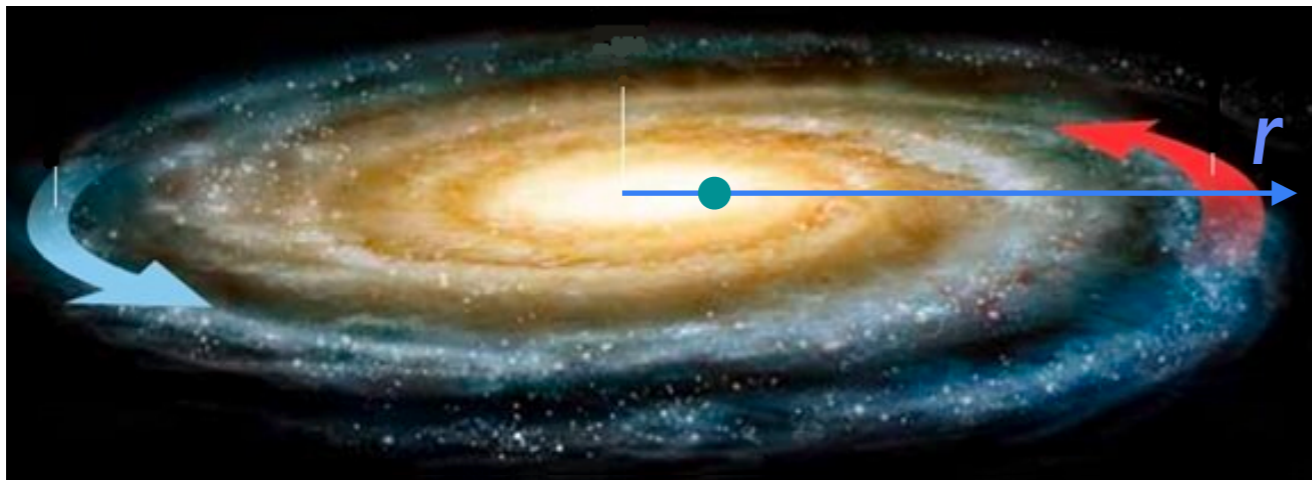
Dark matter, why?

► Different empirical evidence of DM from astrophysical observations at different scales

- first indication from Zwicky's dispersion velocity measurements of galaxies in Coma cluster
- existence of DM confirmed by measurements of stars and gas circular velocities within a galaxy by Ford and Rubin

- from Newtonian dynamics expected velocity $v(r)$ of these objects:

$$v(r) = \sqrt{\frac{GM(r)}{r}}$$



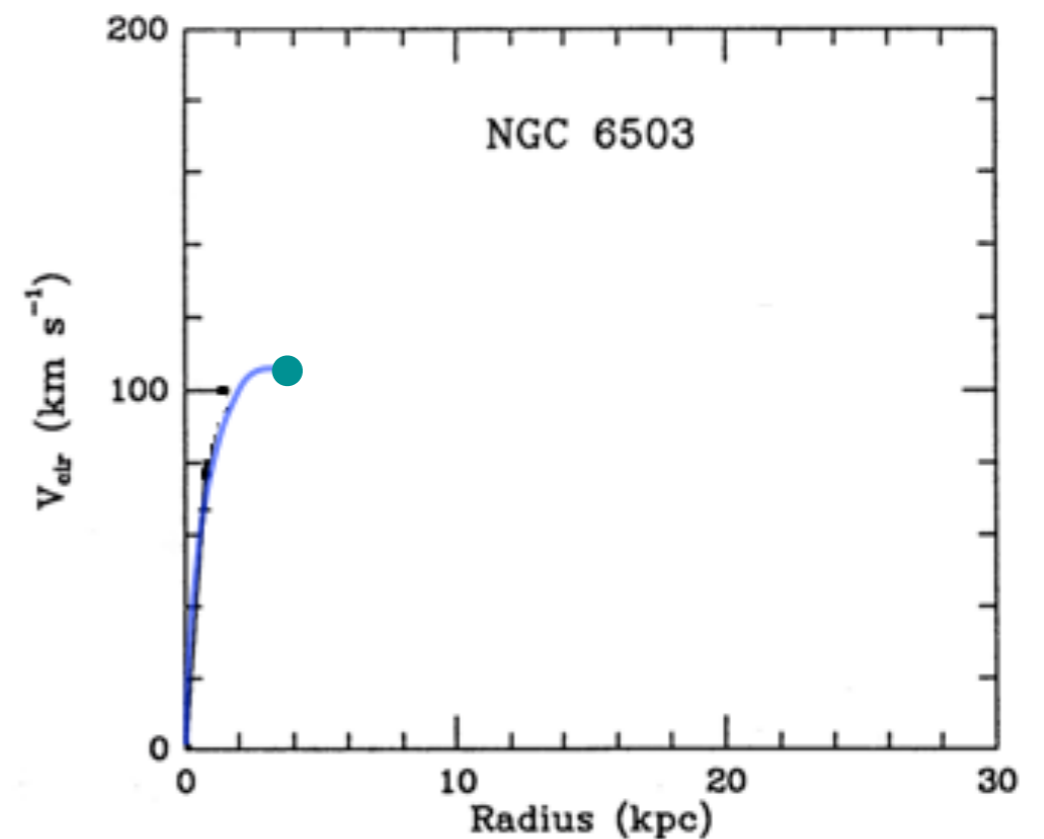
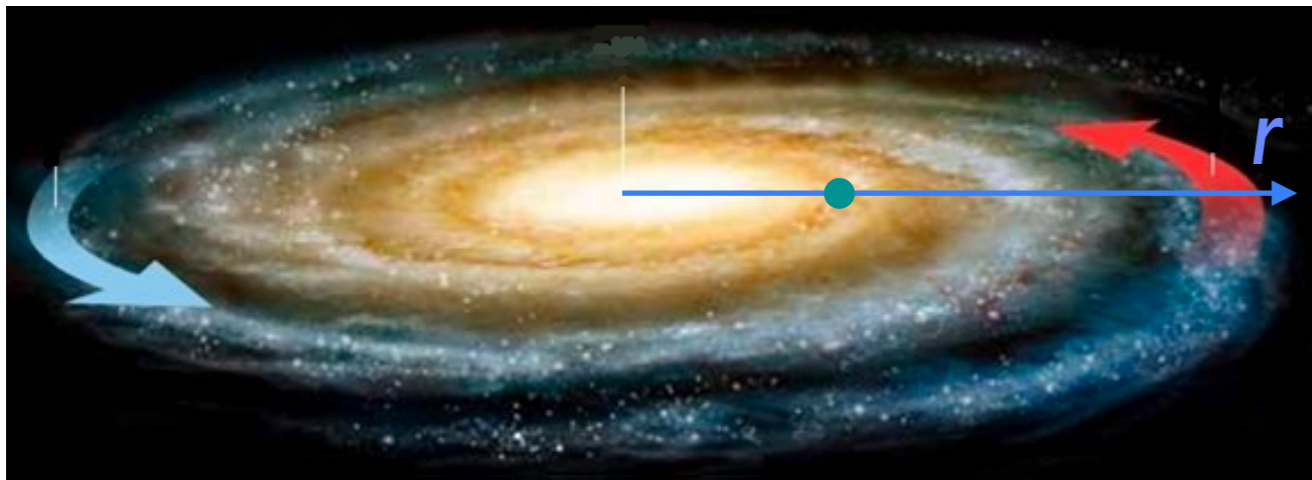
Dark matter, why?

► Different empirical evidence of DM from astrophysical observations at different scales

- first indication from Zwicky's dispersion velocity measurements of galaxies in Coma cluster
- existence of DM confirmed by measurements of stars and gas circular velocities within a galaxy by Ford and Rubin

- from Newtonian dynamics expected velocity $v(r)$ of these objects:

$$v(r) = \sqrt{\frac{GM(r)}{r}}$$



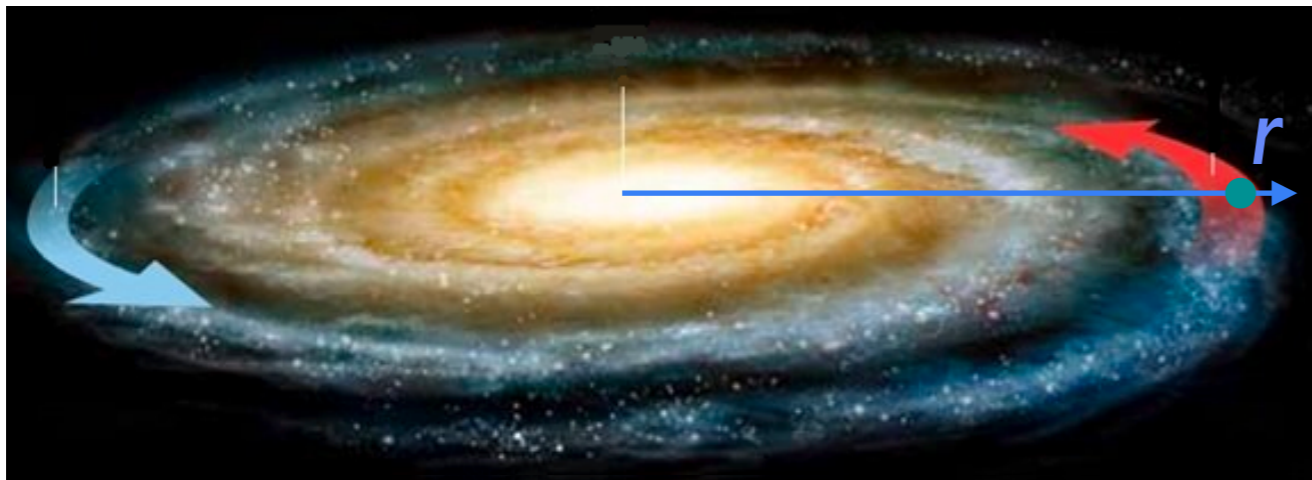
Dark matter, why?

► Different empirical evidence of DM from astrophysical observations at different scales

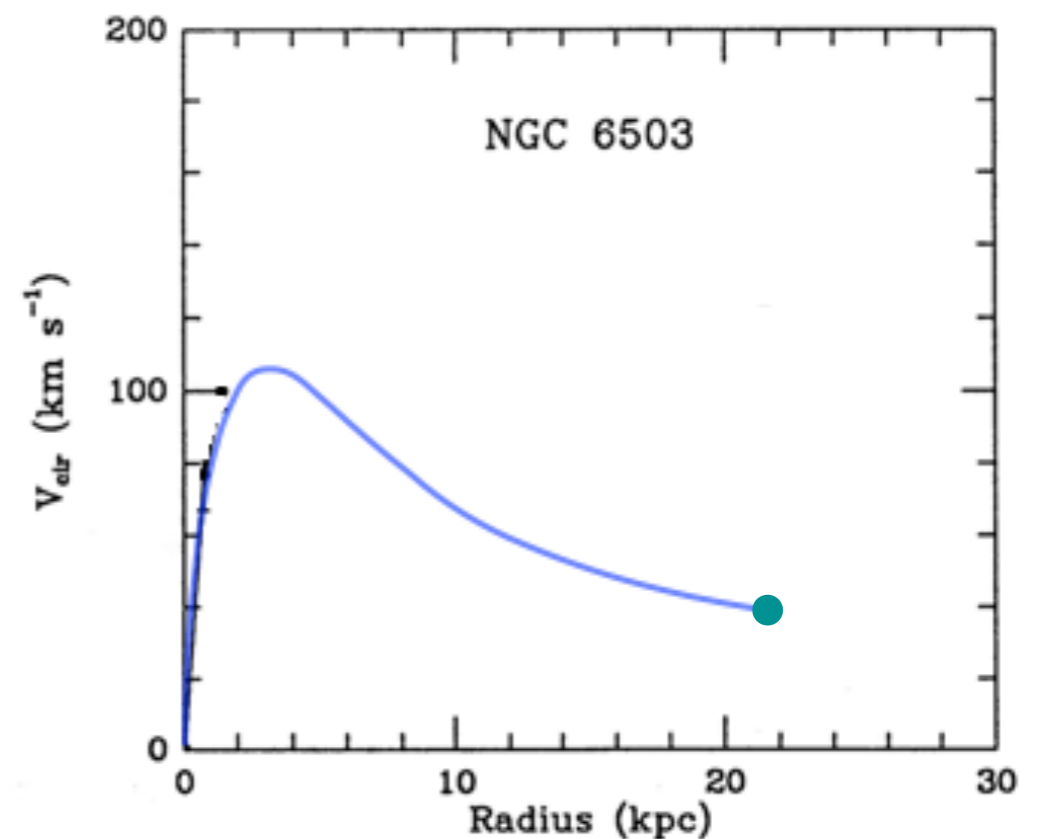
- first indication from Zwicky's dispersion velocity measurements of galaxies in Coma cluster
- existence of DM confirmed by measurements of stars and gas circular velocities within a galaxy by Ford and Rubin

- from Newtonian dynamics expected velocity $v(r)$ of these objects:

$$v(r) = \sqrt{\frac{GM(r)}{r}}$$



Is this what we measured?



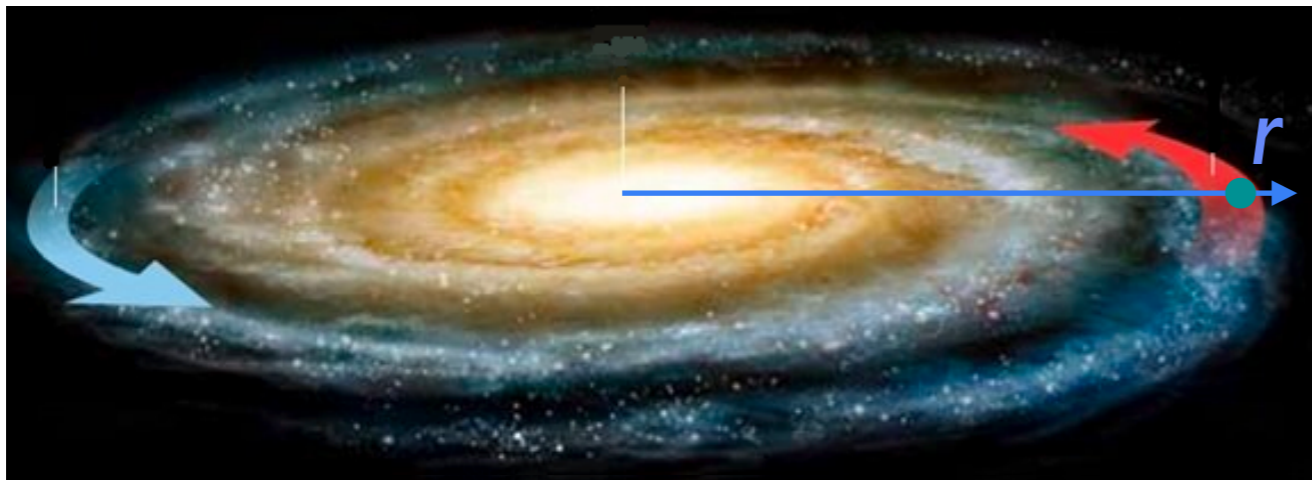
Dark matter, why?

► Different empirical evidence of DM from astrophysical observations at different scales

- first indication from Zwicky's dispersion velocity measurements of galaxies in Coma cluster
- existence of DM confirmed by measurements of stars and gas circular velocities within a galaxy by Ford and Rubin

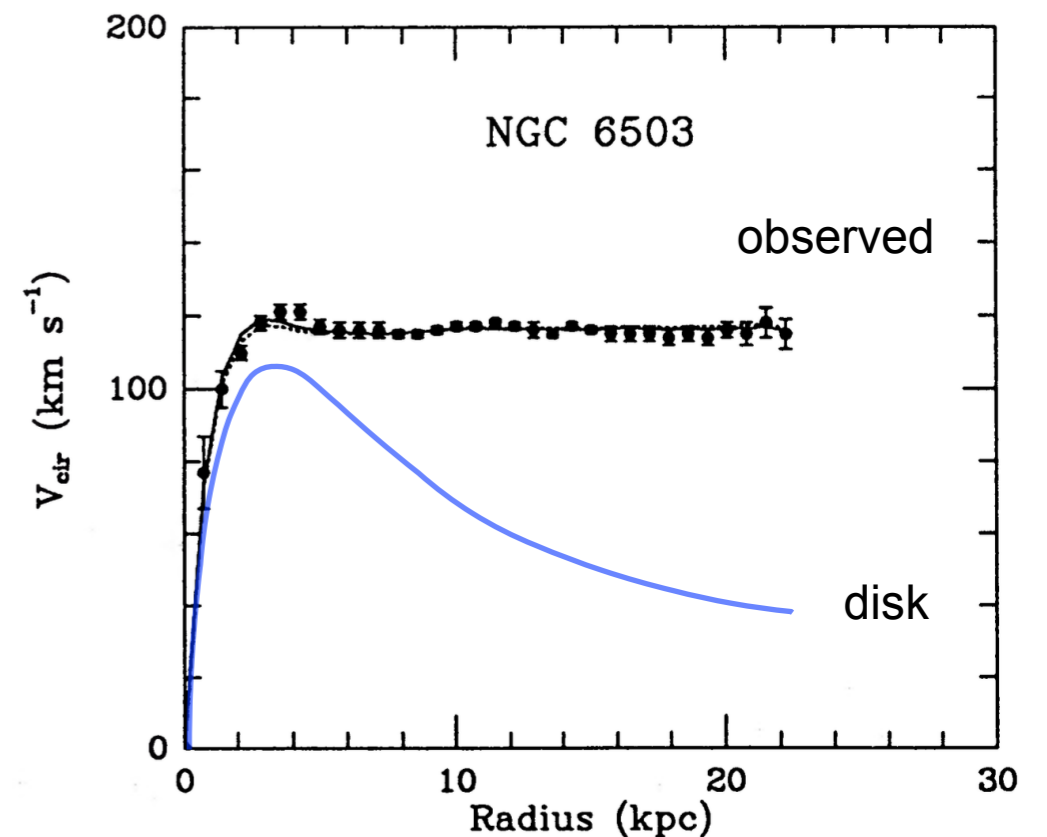
- from Newtonian dynamics expected velocity $v(r)$ of these objects:

$$v(r) = \sqrt{\frac{GM(r)}{r}}$$



Is this what we measured?

- $v \approx \text{const.} \rightarrow M \propto r$



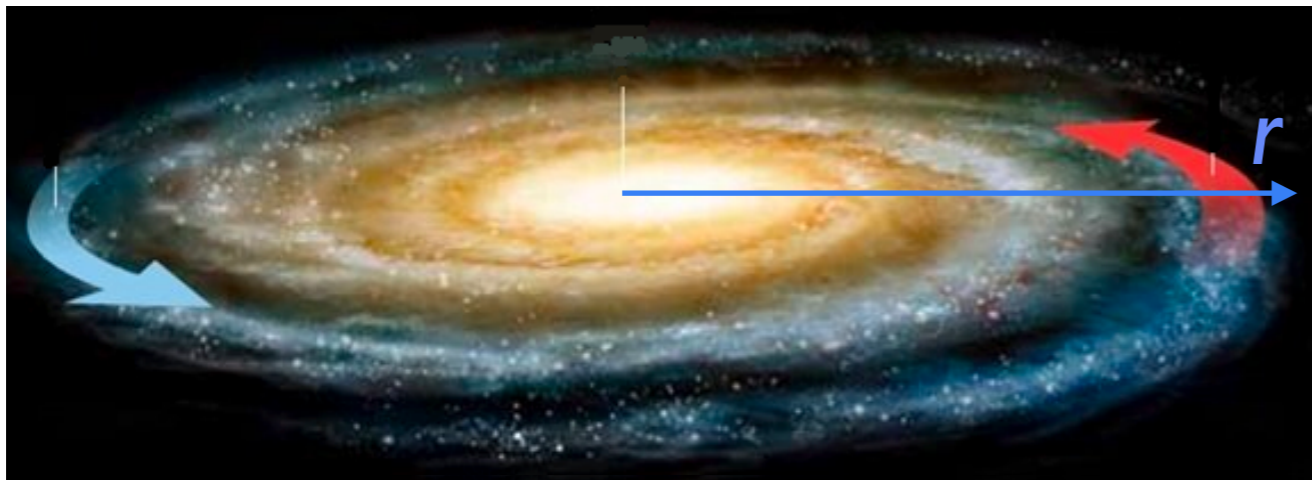
Dark matter, why?

► Different empirical evidence of DM from astrophysical observations at different scales

- first indication from Zwicky's dispersion velocity measurements of galaxies in Coma cluster
- existence of DM confirmed by measurements of stars and gas circular velocities within a galaxy by Ford and Rubin

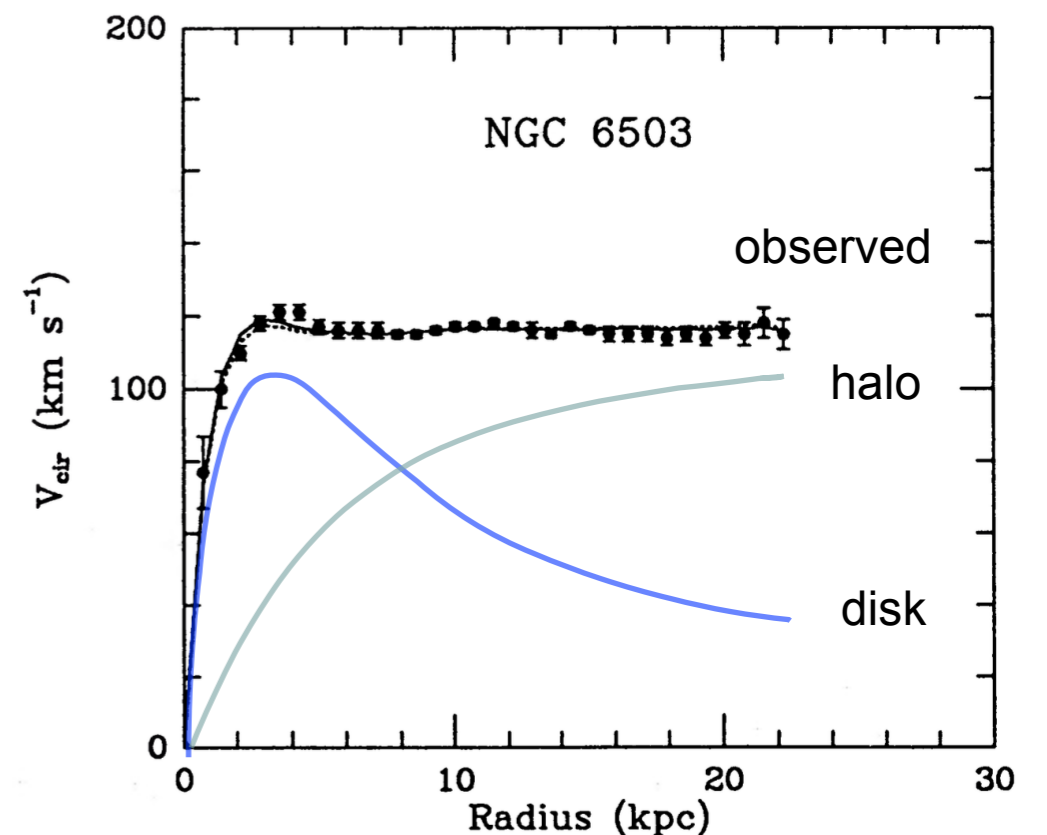
- from Newtonian dynamics expected velocity $v(r)$ of these objects:

$$v(r) = \sqrt{\frac{GM(r)}{r}}$$



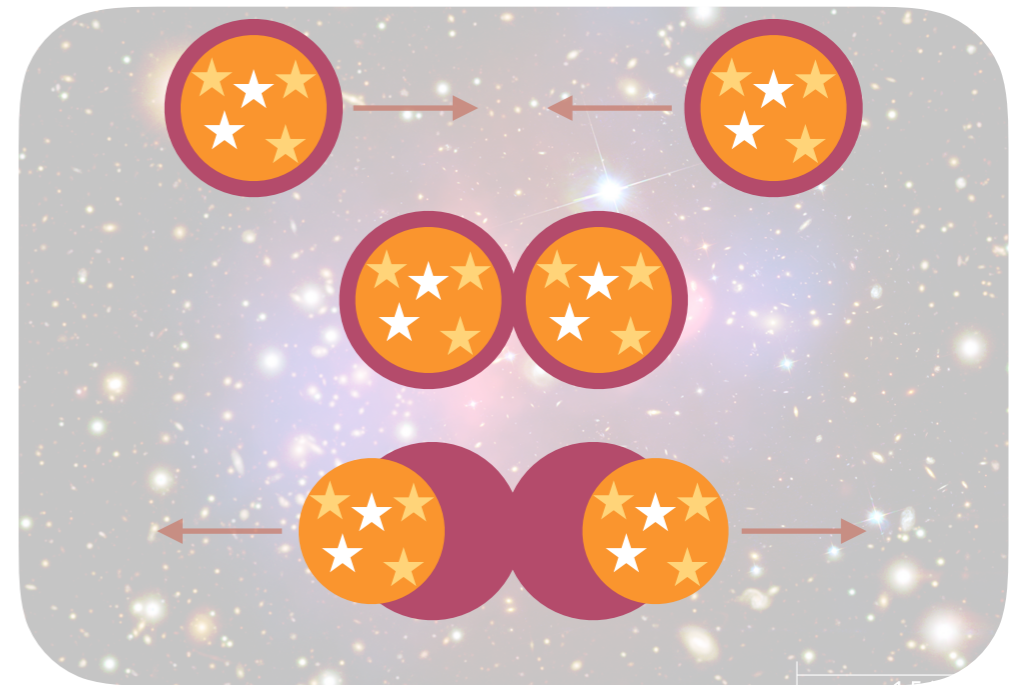
Is this what we measured?

- $v \approx \text{const.} \rightarrow M \propto r$
- non-luminous matter halo with spherical distribution in galaxy outer part



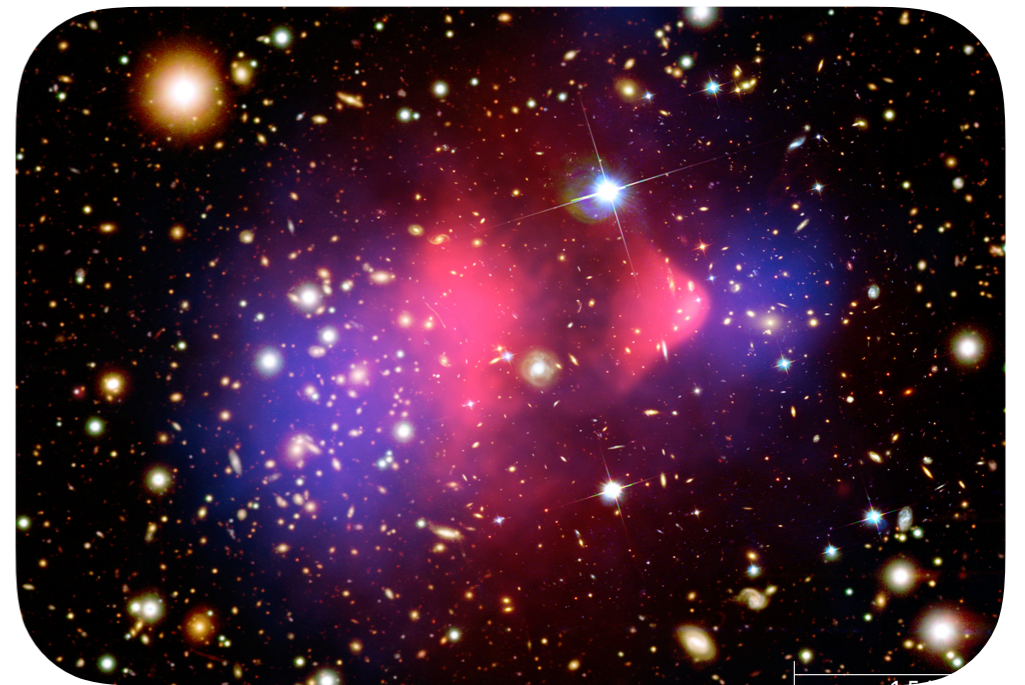
Dark matter, why?

- ▶ *Different empirical evidence of DM* from astrophysical observations at different scales
 - previous examples are based on gravity description, many attempt to explain by Modified Newtonian Dynamics (MOND)
 - from gravitational lensing confirmation of non-luminous matter presence in the universe
- ▶ *Merging of two clusters of galaxies*
 - *stars* behave as collisionless particles (orange and white)
 - *intracluster hot gas* experiences ram pressure, distributed toward the system centre after collision (pink clumps)



Dark matter, why?

- ▶ *Different empirical evidence of DM* from astrophysical observations at different scales
 - previous examples are based on gravity description, many attempt to explain by Modified Newtonian Dynamics (MOND)
 - from gravitational lensing confirmation of non-luminous matter presence in the universe
- ▶ *Merging of two clusters of galaxies*
 - *stars* behave as collisionless particles (orange and white)
 - *intracluster hot gas* experiences ram pressure, distributed toward the system centre after collision (pink clumps)
 - *From lensing highest mass density regions shown by blue regions*

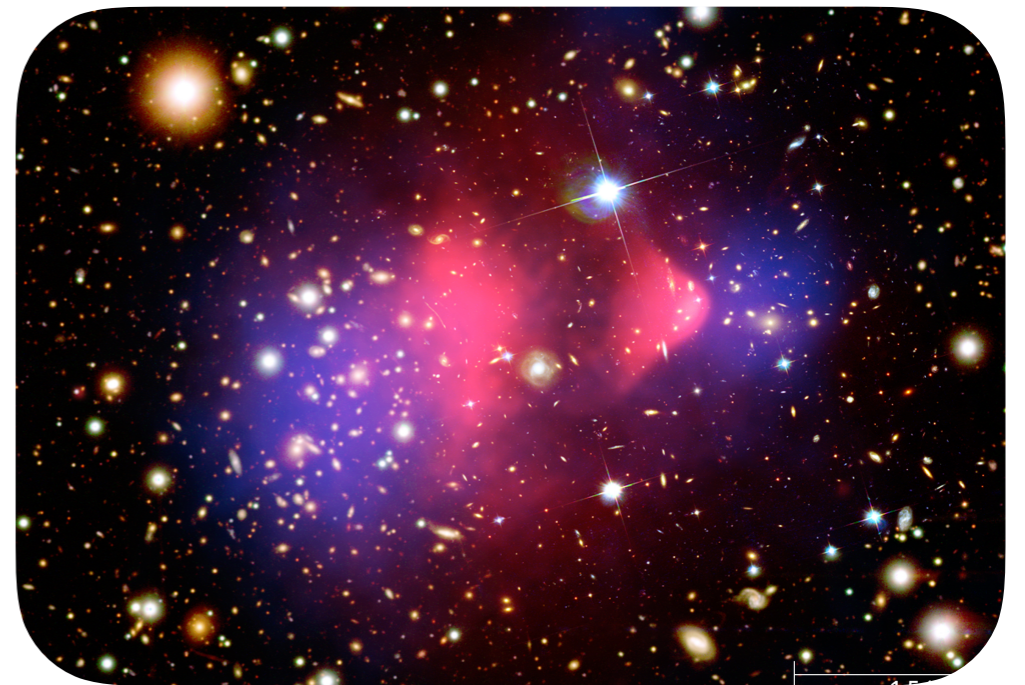


Dark matter, why?

- ▶ *Different empirical evidence of DM* from astrophysical observations at different scales
 - previous examples are based on gravity description, many attempt to explain by Modified Newtonian Dynamics (MOND)
 - from gravitational lensing confirmation of non-luminous matter presence in the universe

- ▶ *Merging of two clusters of galaxies*

- *stars* behave as collisionless particles (orange and white)
- *intracluster hot gas* experiences ram pressure, distributed toward the system centre after collision (pink clumps)
- *From lensing highest mass density regions shown by blue regions*



- If only visible matter in galaxies the highest mass concentration would coincide with hot gas distribution
- The observed separation points to presence of collisionless DM. This without assumptions on gravitational force law description

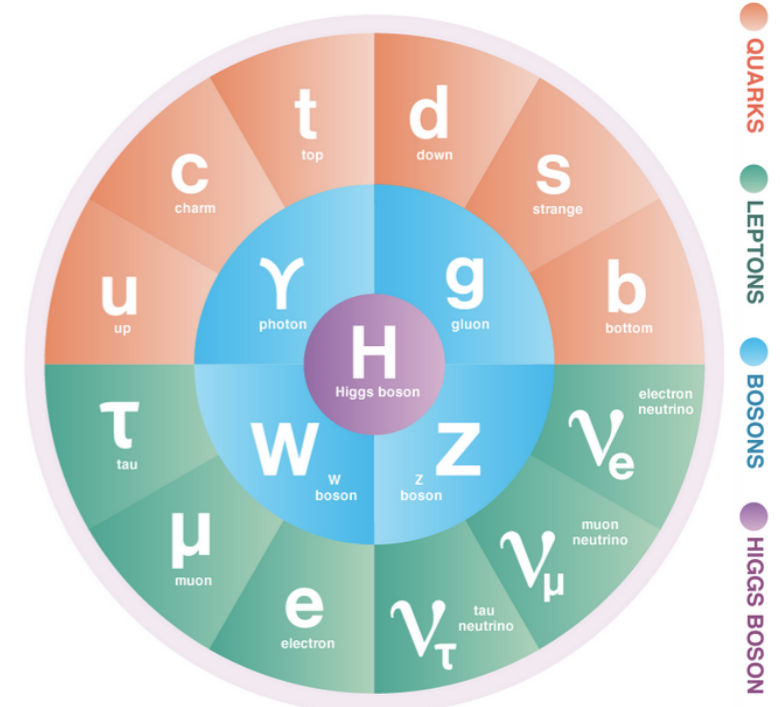
What do we know about DM?

► DM characteristics

- *stable* - on cosmological scale, relic density
- *electrically neutral* - does not significantly emit, reflect, or absorb light
- *massive* - interacts gravitationally
- *not made of baryons* (protons, neutrons) - 25% of our universe is made of DM from Cosmic Microwave



► A possible DM candidate in the Standard Model?



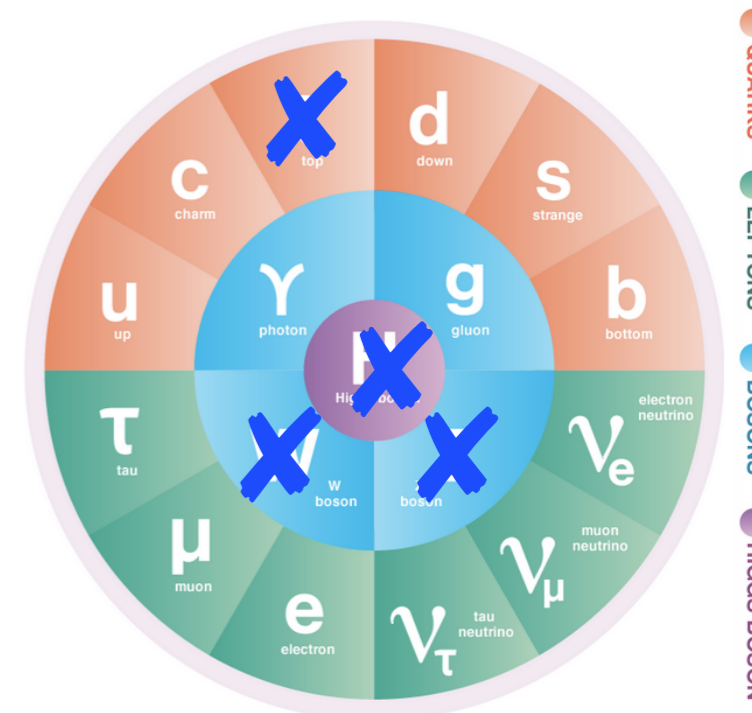
What do we know about DM?

► DM characteristics

- *stable* - on cosmological scale, relic density
- *electrically neutral* - does not significantly emit, reflect, or absorb light
- *massive* - interacts gravitationally
- *not made of baryons* (protons, neutrons) - 25% of our universe is made of DM from Cosmic Microwave



► A possible DM candidate in the Standard Model?



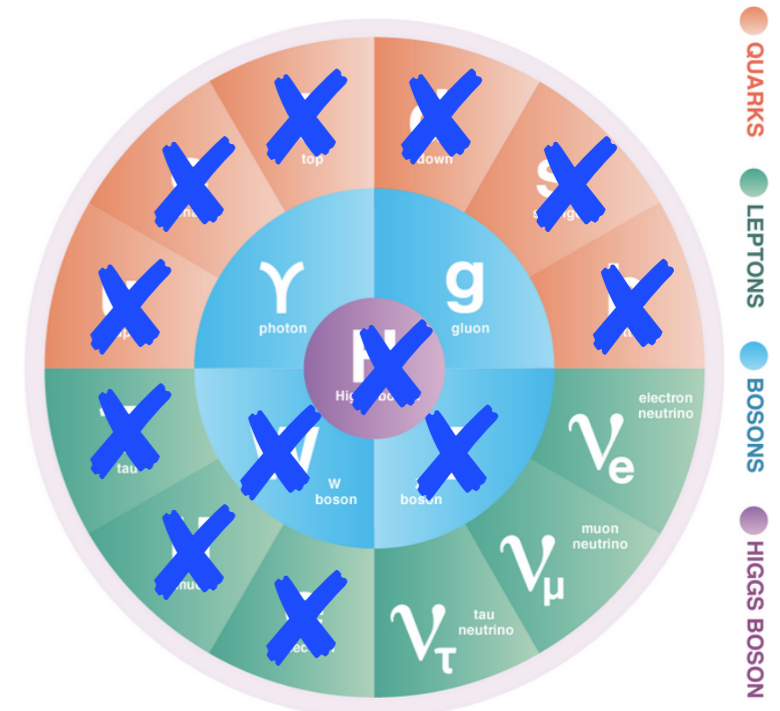
What do we know about DM?

► DM characteristics

- *stable* - on cosmological scale, relic density
- *electrically neutral* - does not significantly emit, reflect, or absorb light
- *massive* - interacts gravitationally
- *not made of baryons* (protons, neutrons) - 25% of our universe is made of DM from Cosmic Microwave



► A possible DM candidate in the Standard Model?



What do we know about DM?

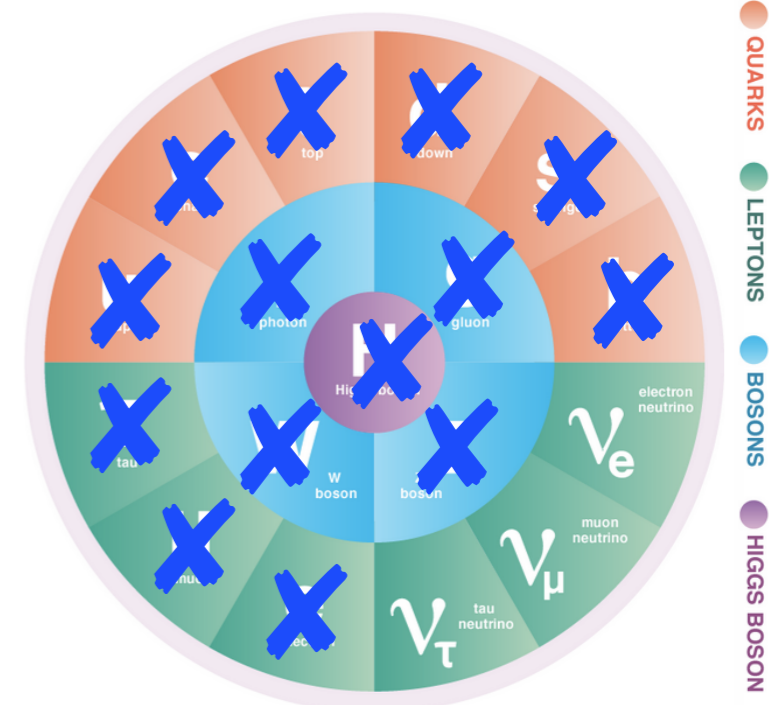
► DM characteristics

- *stable* - on cosmological scale, relic density
- *electrically neutral* - does not significantly emit, reflect, or absorb light
- *massive* - interacts gravitationally
- *not made of baryons* (protons, neutrons) - 25% of our universe is made of DM from Cosmic Microwave



► A possible DM candidate in the Standard Model?

- neutrinos?
 - neutrinos were relativistic when gravity began to bind large scale structure
 - if DM relativistic then larger structure would have formed earlier in the evolution of the universe
 - from observations, *dark matter non-relativistic* at the time of galaxies formation



What do we know about DM?

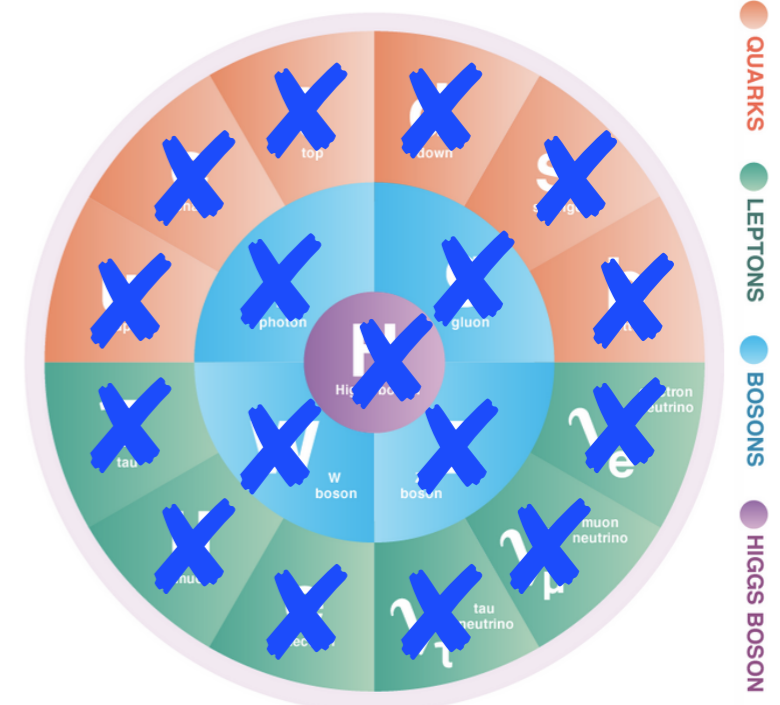
► DM characteristics

- *stable* - on cosmological scale, relic density
- *electrically neutral* - does not significantly emit, reflect, or absorb light
- *massive* - interacts gravitationally
- *not made of baryons* (protons, neutrons) - 25% of our universe is made of DM from Cosmic Microwave



► A possible DM candidate in the Standard Model?

- neutrinos?
 - neutrinos were relativistic when gravity began to bind large scale structure
 - if DM relativistic then larger structure would have formed earlier in the evolution of the universe
 - from observations, *dark matter non-relativistic* at the time of galaxies formation



What's beyond the Standard Model for DM?

Recap

- ▶ Empirical evidence of DM from astrophysical observations at different scales
 - interacts gravitationally, long lived and neutral
 - *no information about its nature*



We saw what DM cannot be, but what can be DM?

- most studied class of theories: let's assume DM is a weakly interacting massive particle

▶ Assuming DM-SM interactions enables different searches:

- indirect detection,
search for stable final SM products (neutrinos, gamma rays, positrons, antiprotons and their antiparticles) from annihilation of DM particles
- direct detection,
search for nuclear recoils produced in the elastic scattering of DM particles on nuclei
- colliders,
search for DM particles produced in high energy collisions



What's beyond the Standard Model for DM?

Recap

- ▶ *Empirical evidence of DM* from astrophysical observations at different scales
 - interacts gravitationally, long lived and neutral
 - *no information about its nature*

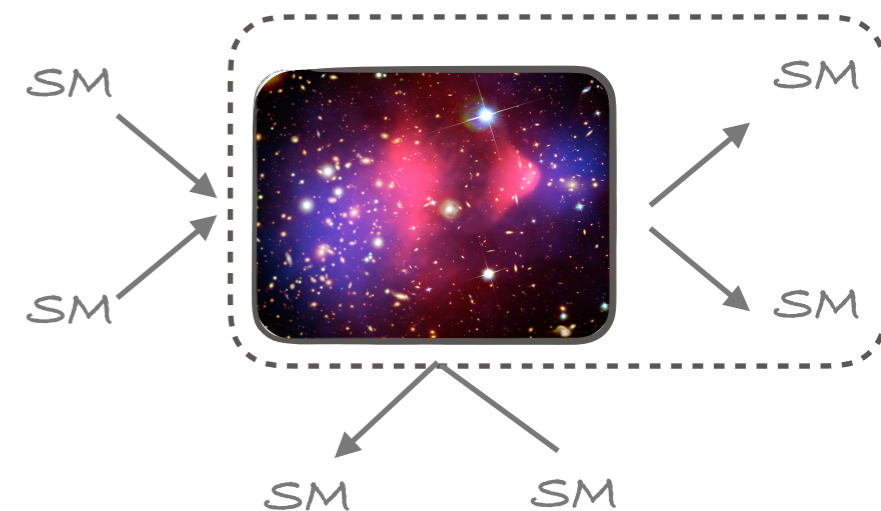


We saw what DM cannot be, but what can be DM?

- most studied class of theories: let's assume DM is a weakly interacting massive particle

▶ *Assuming DM-SM interactions enables different searches:*

- indirect detection,
search for stable final SM products (neutrinos, gamma rays, positrons, antiprotons and their antiparticles) from annihilation of DM particles
- direct detection,
search for nuclear recoils produced in the elastic scattering of DM particles on nuclei
- colliders,
search for DM particles produced in high energy collisions



What's beyond the Standard Model for DM?

Recap

- ▶ Empirical evidence of DM from astrophysical observations at different scales
 - interacts gravitationally, long lived and neutral
 - *no information about its nature*

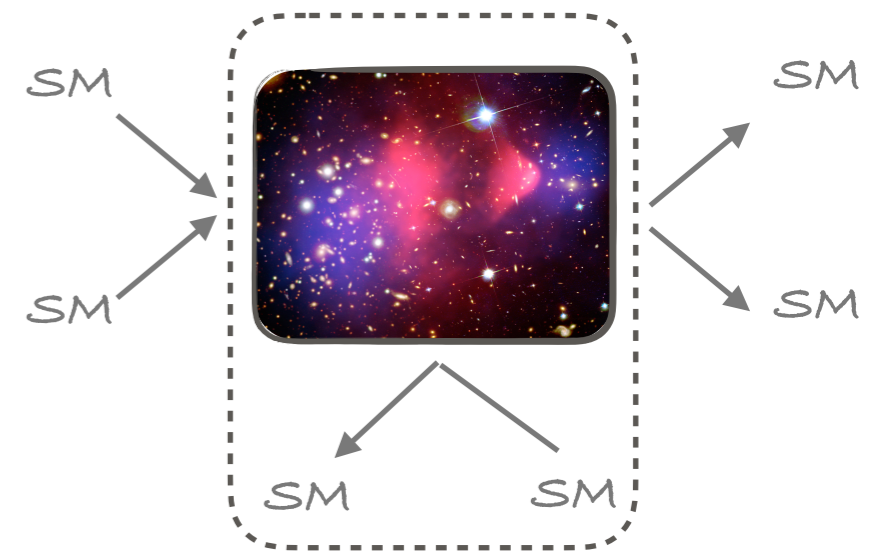


We saw what DM cannot be, but what can be DM?

- most studied class of theories: let's assume DM is a weakly interacting massive particle

▶ Assuming DM-SM interactions enables different searches:

- *indirect detection*,
search for stable final SM products (neutrinos, gamma rays, positrons, antiprotons and their antiparticles) from annihilation of DM particles
- direct detection,
search for nuclear recoils produced in the elastic scattering of DM particles on nuclei
- *colliders*,
search for DM particles produced in high energy collisions



What's beyond the Standard Model for DM?

Recap

- ▶ Empirical evidence of DM from astrophysical observations at different scales
 - interacts gravitationally, long lived and neutral
 - *no information about its nature*

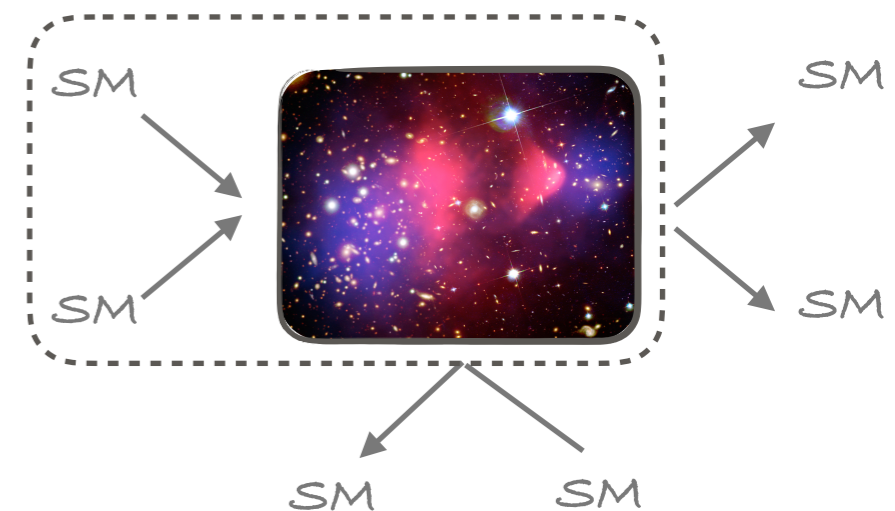


We saw what DM cannot be, but what can be DM?

- most studied class of theories: let's assume DM is a weakly interacting massive particle

▶ Assuming DM-SM interactions enables different searches:

- *indirect detection*,
search for stable final SM products (neutrinos, gamma rays, positrons, antiprotons and their antiparticles) from annihilation of DM particles
- *direct detection*,
search for nuclear recoils produced in the elastic scattering of DM particles on nuclei
- colliders,
search for DM particles produced in high energy collisions



What's beyond the Standard Model for DM?

Recap

- ▶ *Empirical evidence of DM* from astrophysical observations at different scales
 - interacts gravitationally, long lived and neutral
 - *no information about its nature*

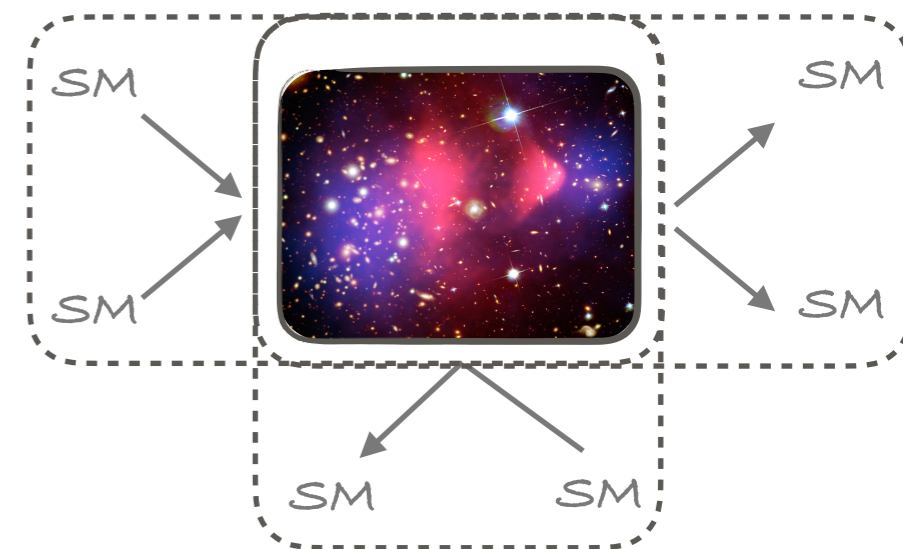


We saw what DM cannot be, but what can be DM?

- most studied class of theories: let's assume DM is a weakly interacting massive particle

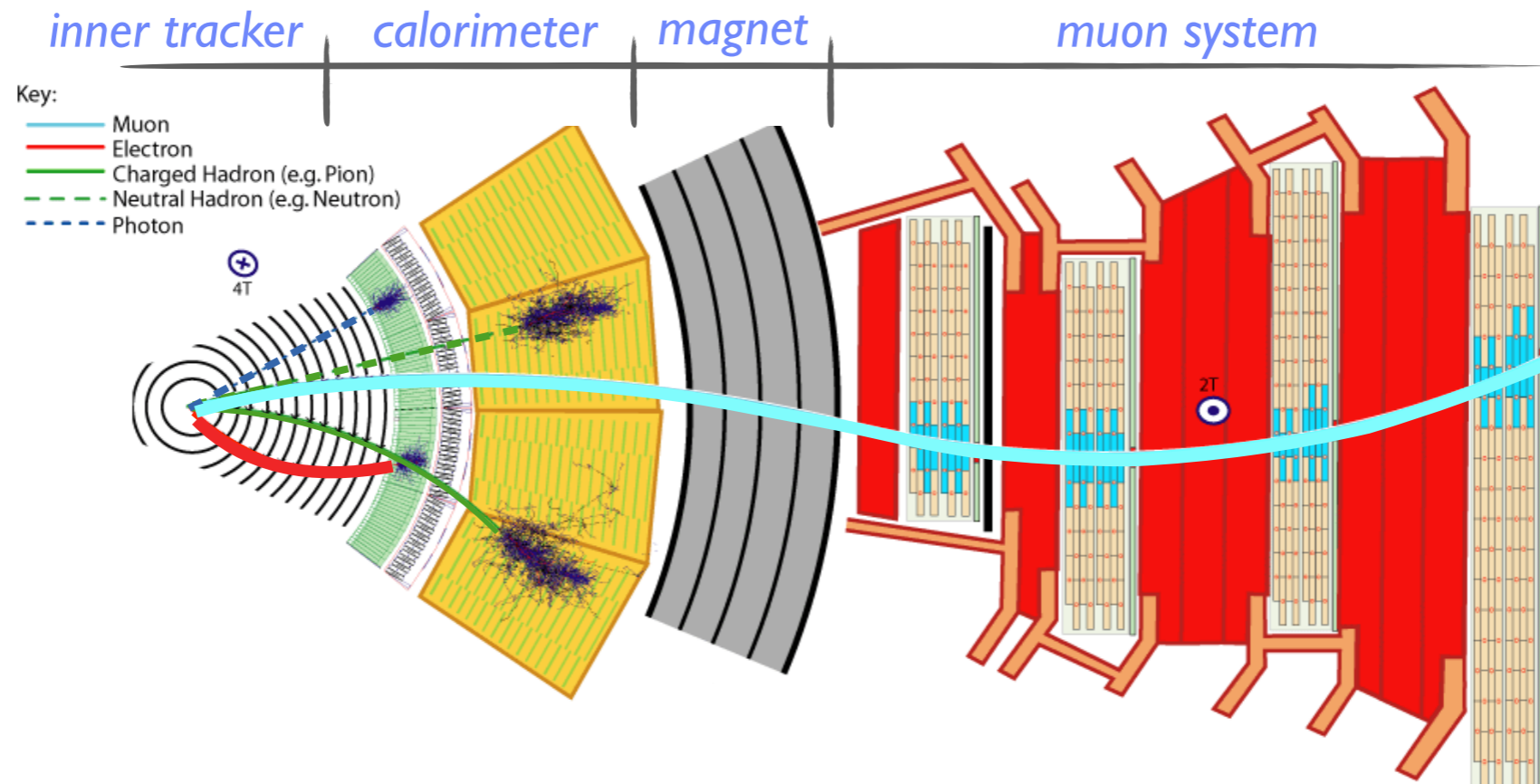
▶ *Assuming DM-SM interactions enables different searches:*

- *indirect detection*,
search for stable final SM products (neutrinos, gamma rays, positrons, antiprotons and their antiparticles) from annihilation of DM particles
- *direct detection*,
search for nuclear recoils produced in the elastic scattering of DM particles on nuclei
- *colliders*,
search for DM particles produced in high energy collisions



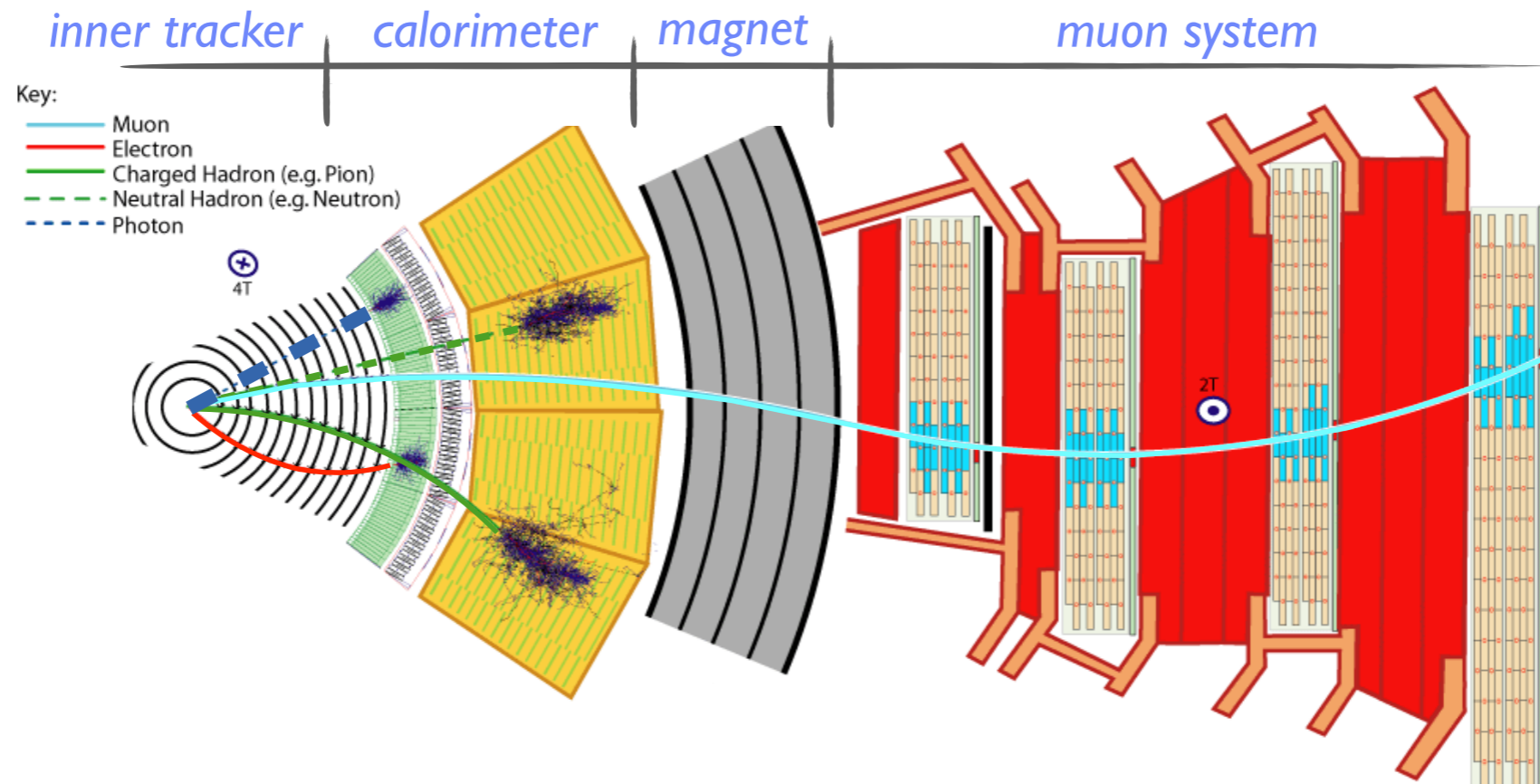
* *Complementarity essential*: eg. info about lifetime in case of DM discovery at colliders ($\sim 10^{-7}$ s), particle properties compared with cosmological constraints

How can we detect particles at colliders?



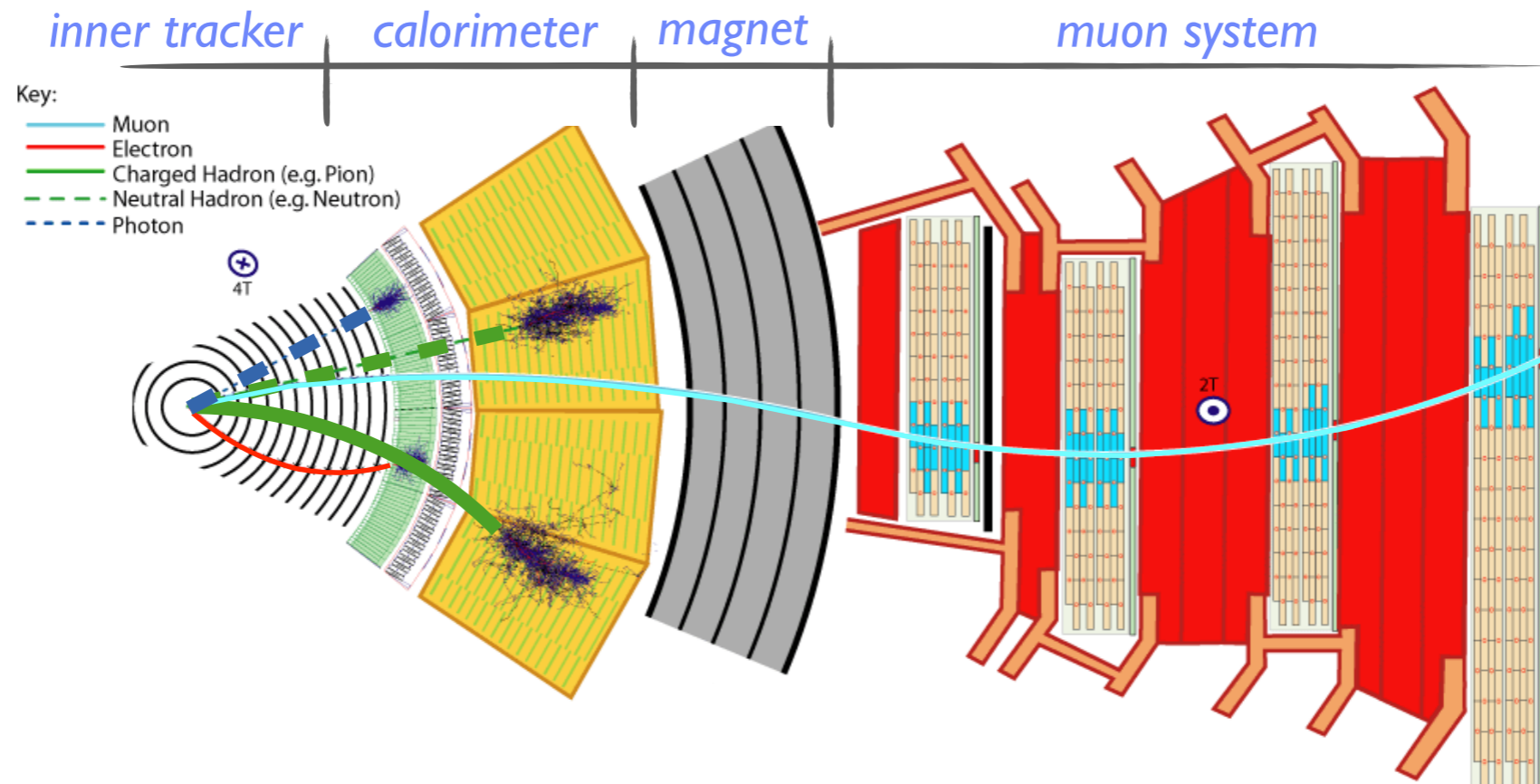
- **electron** (**muon**), from tracks in inner tracker and energy in calorimeter (track in muon spectrometer)
- *photon*, from energy deposits in electromagnetic calorimeter
- *jets from quarks and gluons*, produced partons hadronize in colour-neutral particles groups, so-called jet. Parton energy and momentum reconstructed clustering all particles from hadronization

How can we detect particles at colliders?



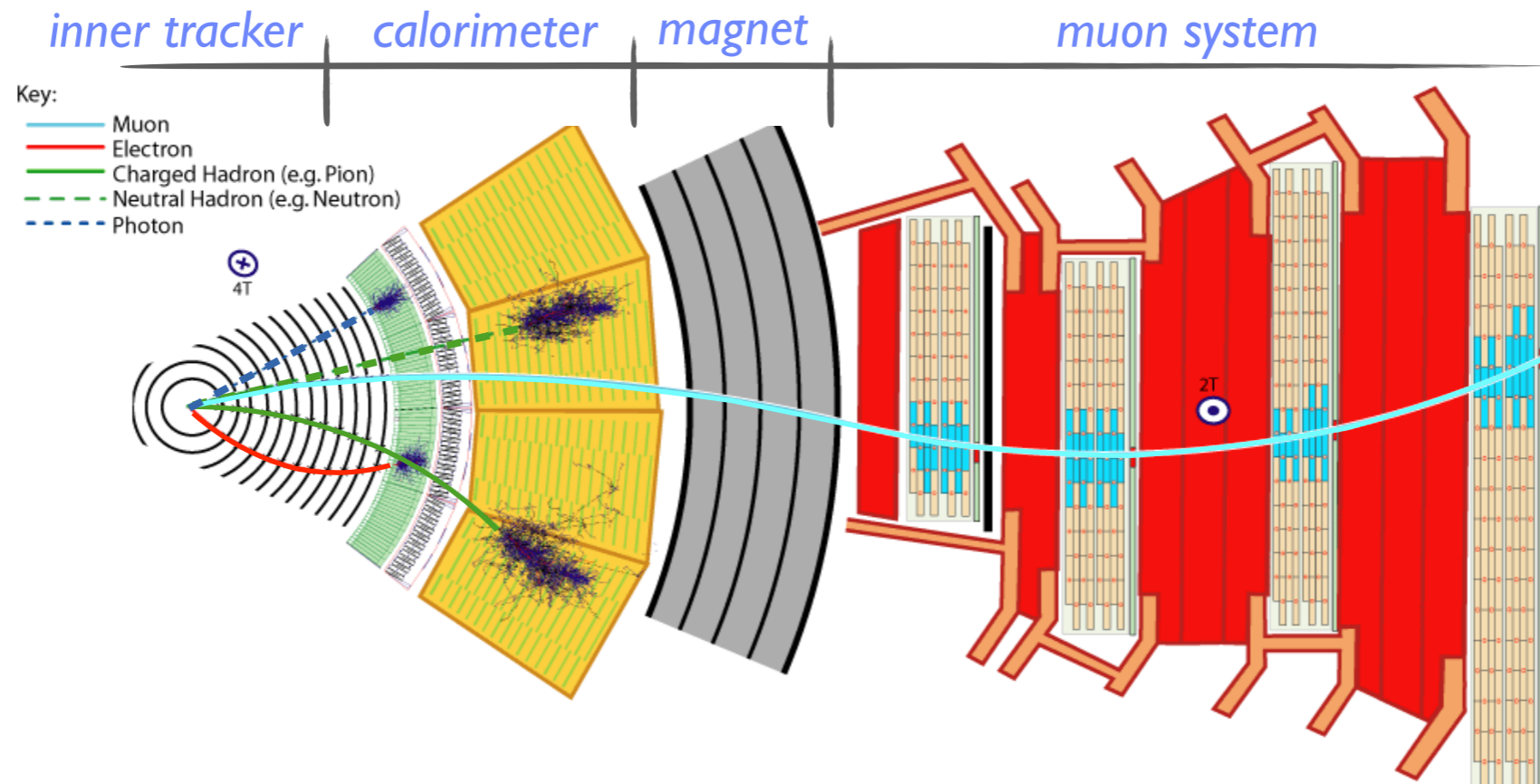
- *electron (muon)*, from tracks in inner tracker and energy in calorimeter (track in muon spectrometer)
- **photon**, from energy deposits in electromagnetic calorimeter
- *jets from quarks and gluons*, produced partons hadronize in colour-neutral particles groups, so-called jet. Parton energy and momentum reconstructed clustering all particles from hadronization

How can we detect particles at colliders?



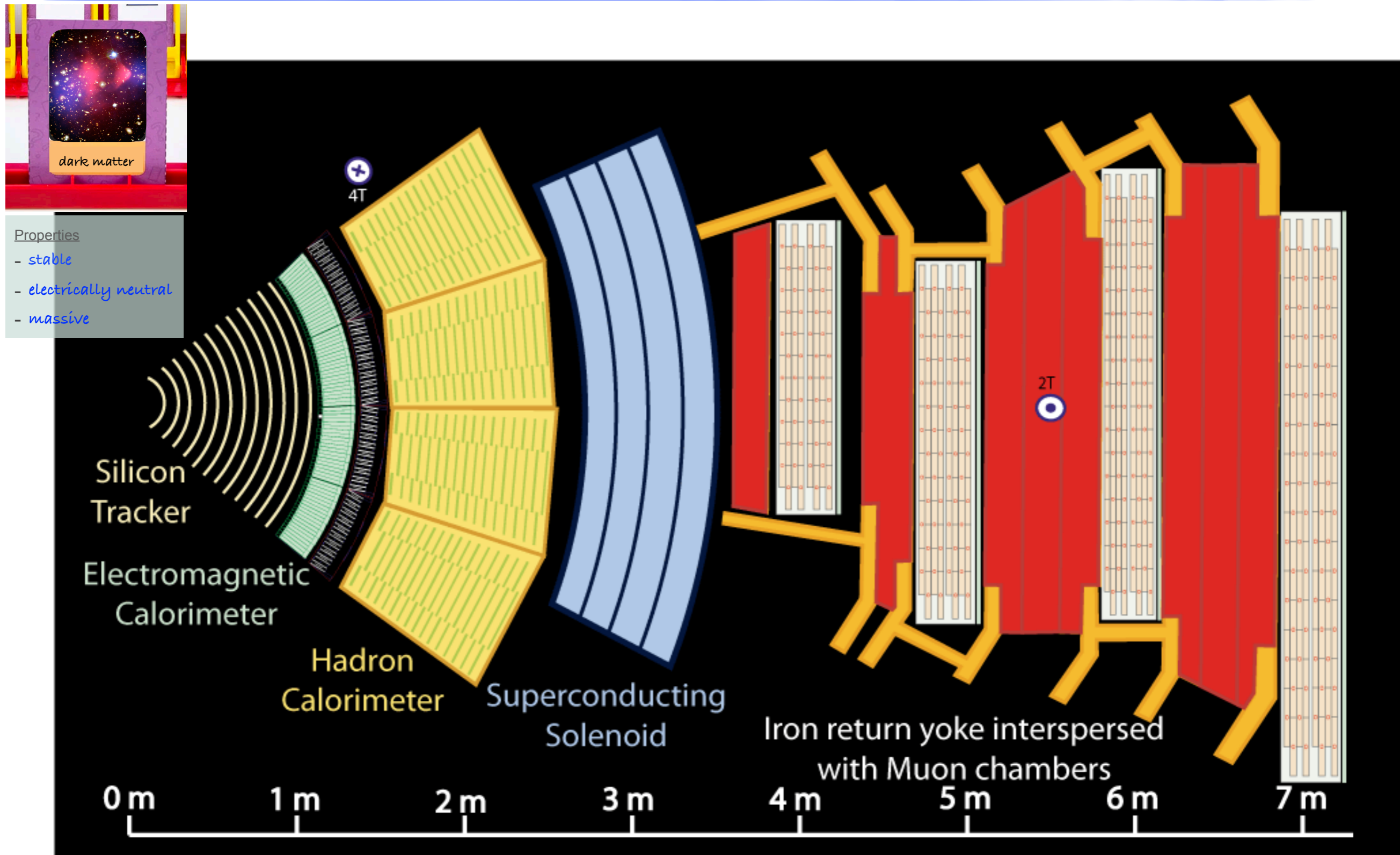
- *electron (muon)*, from tracks in inner tracker and energy in calorimeter (track in muon spectrometer)
- *photon*, from energy deposits in electromagnetic calorimeter
- **jets** from *quarks and gluons*, produced partons hadronize in colour-neutral particles groups, so-called jet. Parton energy and momentum reconstructed clustering all particles from hadronization

How can we detect particles at colliders?



- *electron (muon)*, from tracks in inner tracker and energy in calorimeter (track in muon spectrometer)
- *photon*, from energy deposits in electromagnetic calorimeter
- *jets from quarks and gluons*, produced partons hadronize in colour-neutral particles groups, so-called jet. Parton energy and momentum reconstructed clustering all particles from hadronization
- *Dark matter?*

How can we detect particles at colliders?



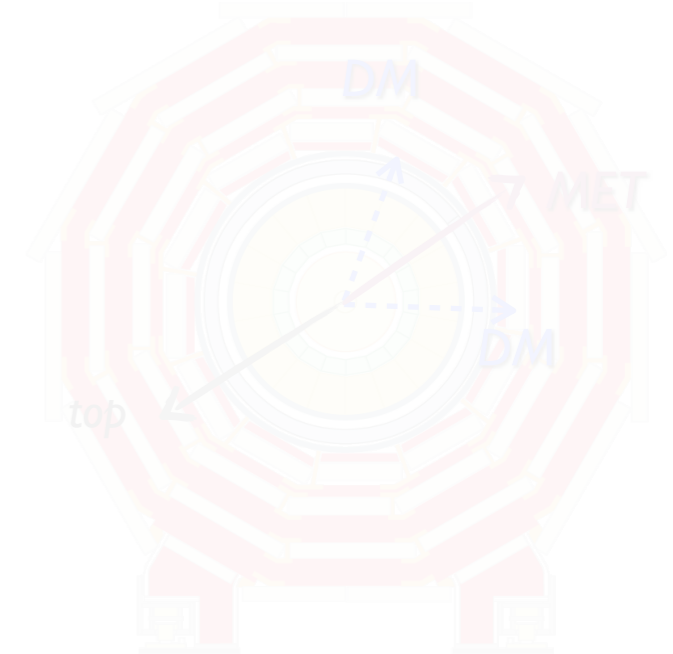
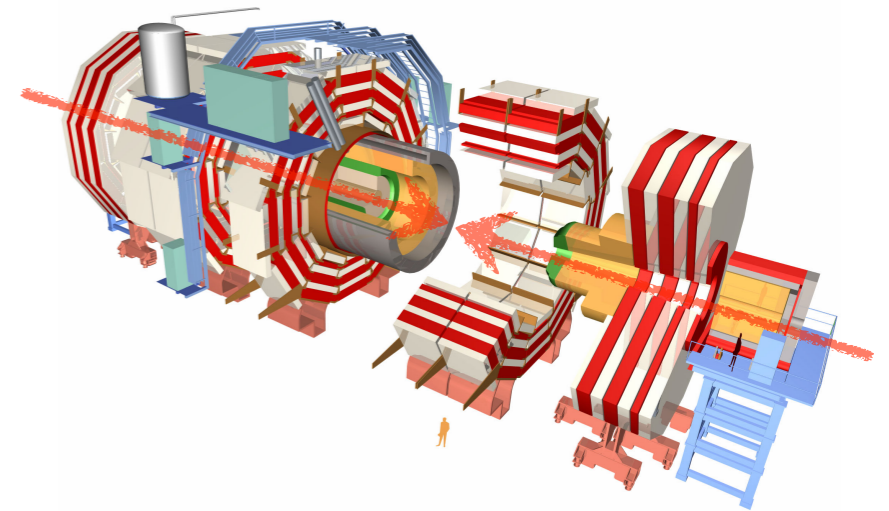
DM signature at colliders

- ▶ DM could be produced at colliders (rare process)
 - no direct trace in the detector, but could create a p_T imbalance (**MET**)

- conservation of momentum:

- no information about longitudinal momentum of colliding partons
- but total initial parton $p_T=0$
 - need to be conserved after the collision $\sum \vec{p}_T = 0$
 - if $\sum \vec{p}_T \neq 0$ some particles escaped the detector carrying $\vec{E}_T^{miss} = -\sum \vec{p}_T$
 - $|\vec{E}_T^{miss}| =$ missing transverse energy (MET)

- ▶ to see the invisible we need the visible ...
 - need visible particle to which DM particle recoils against
 - “mono-X searches”: X includes jets, vector bosons, top, ...



DM signature at colliders

► DM could be produced at colliders (rare process)

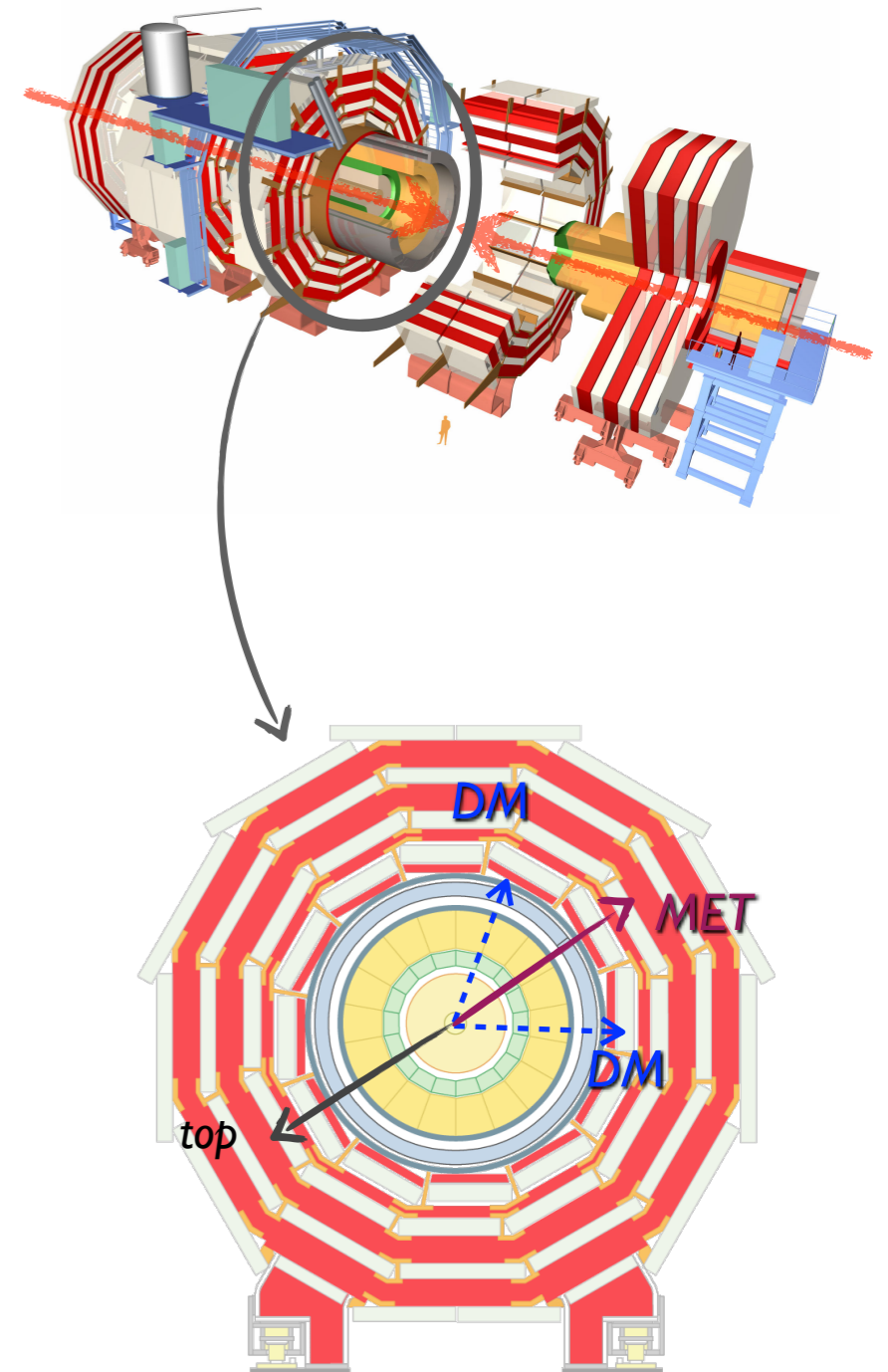
- no direct trace in the detector, but could create a p_T imbalance (**MET**)

- conservation of momentum:

- no information about longitudinal momentum of colliding partons
- but total initial parton $p_T=0$
 - need to be conserved after the collision $\sum \vec{p}_T = 0$
 - if $\sum \vec{p}_T \neq 0$ some particles escaped the detector carrying $\vec{E}_T^{miss} = -\sum \vec{p}_T$
 - $|\vec{E}_T^{miss}| =$ missing transverse energy (MET)

► to see the invisible we need the visible ...

- need visible particle to which DM particle recoils against
- "mono-X searches": X includes jets, vector bosons, top, ...



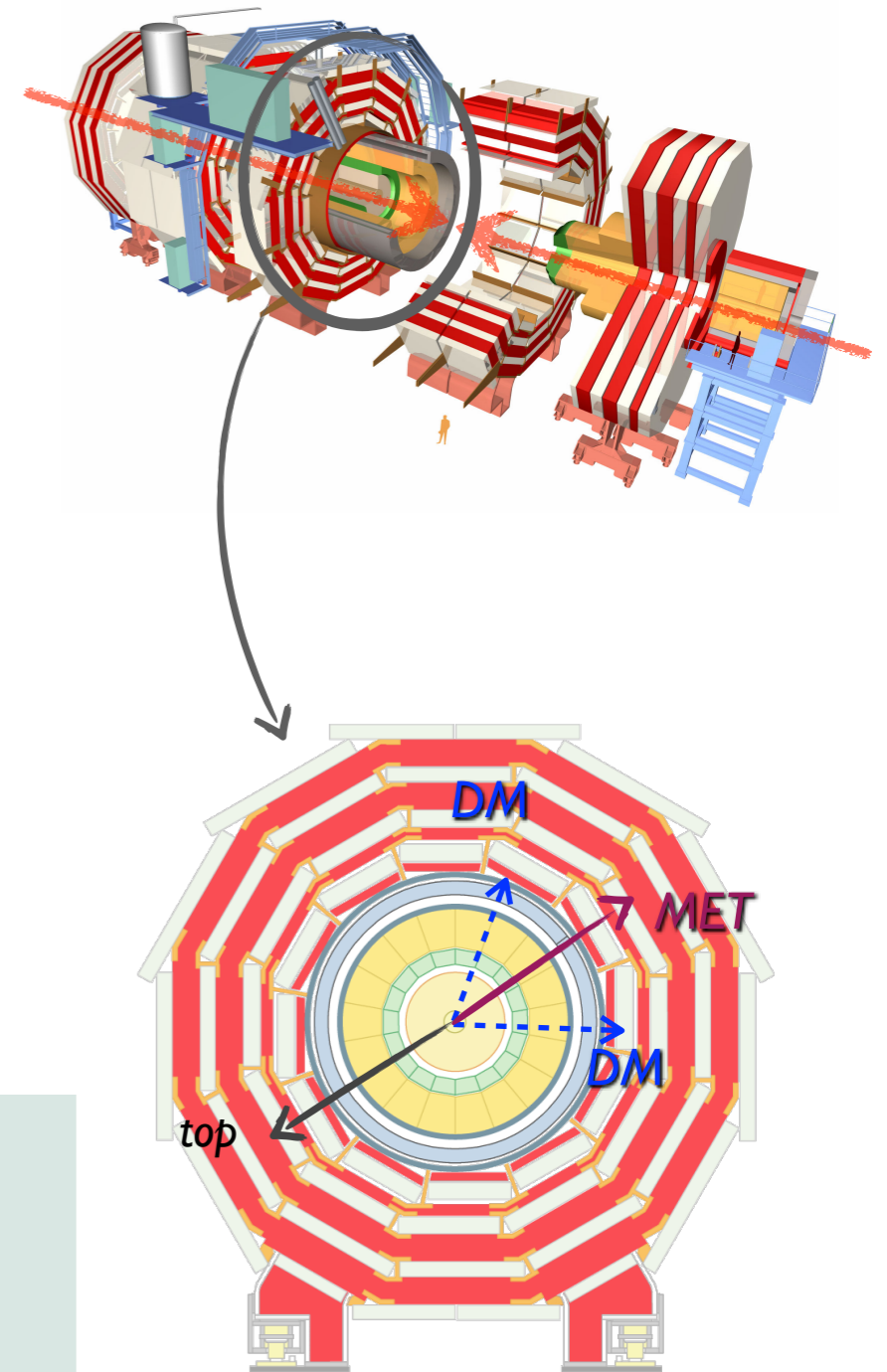
DM signature at colliders

► DM could be produced at colliders (rare process)

- no direct trace in the detector, but could create a p_T imbalance (**MET**)
- conservation of momentum:
 - no information about longitudinal momentum of colliding partons
 - but total initial parton $p_T=0$
 - need to be conserved after the collision $\sum \vec{p}_T = 0$
 - if $\sum \vec{p}_T \neq 0$ some particles escaped the detector carrying $\vec{E}_T^{miss} = -\sum \vec{p}_T$
 - $|\vec{E}_T^{miss}| =$ missing transverse energy (MET)

► to see the invisible we need the visible ...

- need *visible particle* to which DM particle recoils against
- “mono- X searches”: X includes jets, vector bosons, top, ...



Dark matter phenomenology: guess “who”



- ▶ We do not have information about the DM nature, how to discover DM?
 - we can remain very general and make very little assumptions
 - eg. for this board: “is it a 2D shape?”
 - we can make more assumptions and tests more specific models
 - eg. for this board: “is it a 2D shape, yellow color and with only 90° angles?”

Dark matter phenomenology: guess “who”



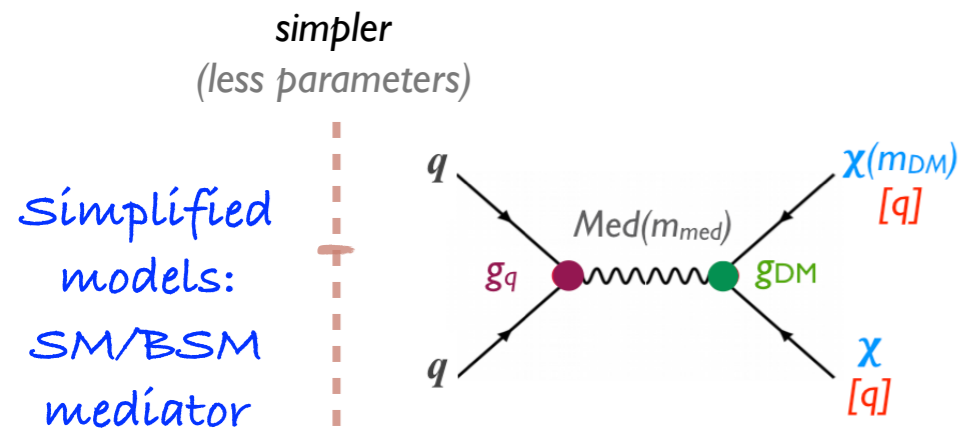
- ▶ We do not have information about the DM nature, how to discover DM?
 - we can remain very general and make very little assumptions
 - eg. for this board: “is it a 2D shape?”
 - we can make more assumptions and tests *more specific models*
 - eg. for this board: “is it a 2D shape, yellow color and with only 90° angles?”

Dark matter phenomenology: guess “who”



- ▶ We do not have information about the DM nature, how to discover DM?
 - we can remain very general and make very little assumptions
 - eg. for this board: “is it a 2D shape?”
 - we can make more assumptions and tests more specific models
 - eg. for this board: “is it a 2D shape, yellow color and with only 90° angles?”

Dark matter? phenomenology at colliders



Beyond SM mediator

- one new mediator (spin-1 or spin-0), one DM particle
- set of parameters (m_{DM} , m_{med} , g_q , g_{DM})

	vector	axial-vector	
spin-1	$g_q \sum_q V_\mu \bar{q} \gamma^\mu q$	$g_q \sum_q A_\mu \bar{q} \gamma^\mu \gamma^5 q$	💡 choose X to increase xsec or bkg rejection
spin-0	$g_q \frac{\phi}{\sqrt{2}} \sum_f y_f \bar{f} f$	$g_q \frac{iA}{\sqrt{2}} \sum_f y_f \bar{f} \gamma^5 f$	💡 choose X to exploit coupling \propto to quark mass (or increase xsec)



Higgs portal DM

- the SM-DM mediator could be the Higgs boson (spin-0)
- Higgs decay branching fractions not yet sufficiently constrained

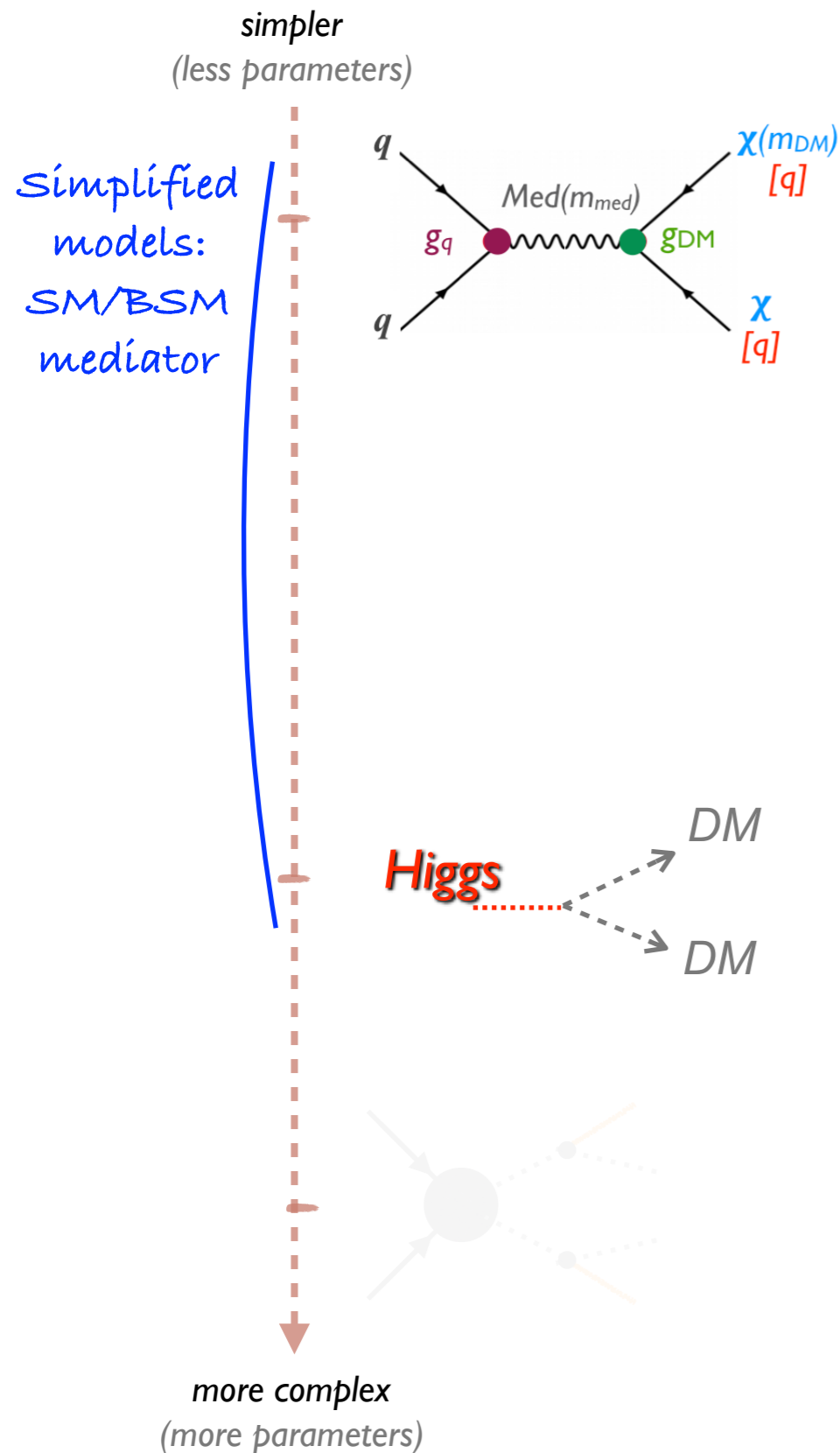
SUSY models

- see more on N. Strobbe's talk



more complex (more parameters)

Dark matter? phenomenology at colliders



Beyond SM mediator

- one new mediator (spin-1 or spin-0), one DM particle
- set of parameters (m_{DM} , m_{med} , g_q , g_{DM})

	vector	axial-vector	
spin-1	$g_q \sum_q V_\mu \bar{q} \gamma^\mu q$	$g_q \sum_q A_\mu \bar{q} \gamma^\mu \gamma^5 q$	💡 choose X to increase xsec or bkg rejection
spin-0	$g_q \frac{\phi}{\sqrt{2}} \sum_f y_f \bar{f} f$	$g_q \frac{iA}{\sqrt{2}} \sum_f y_f \bar{f} \gamma^5 f$	💡 choose X to exploit coupling \propto to quark mass (or increase xsec)

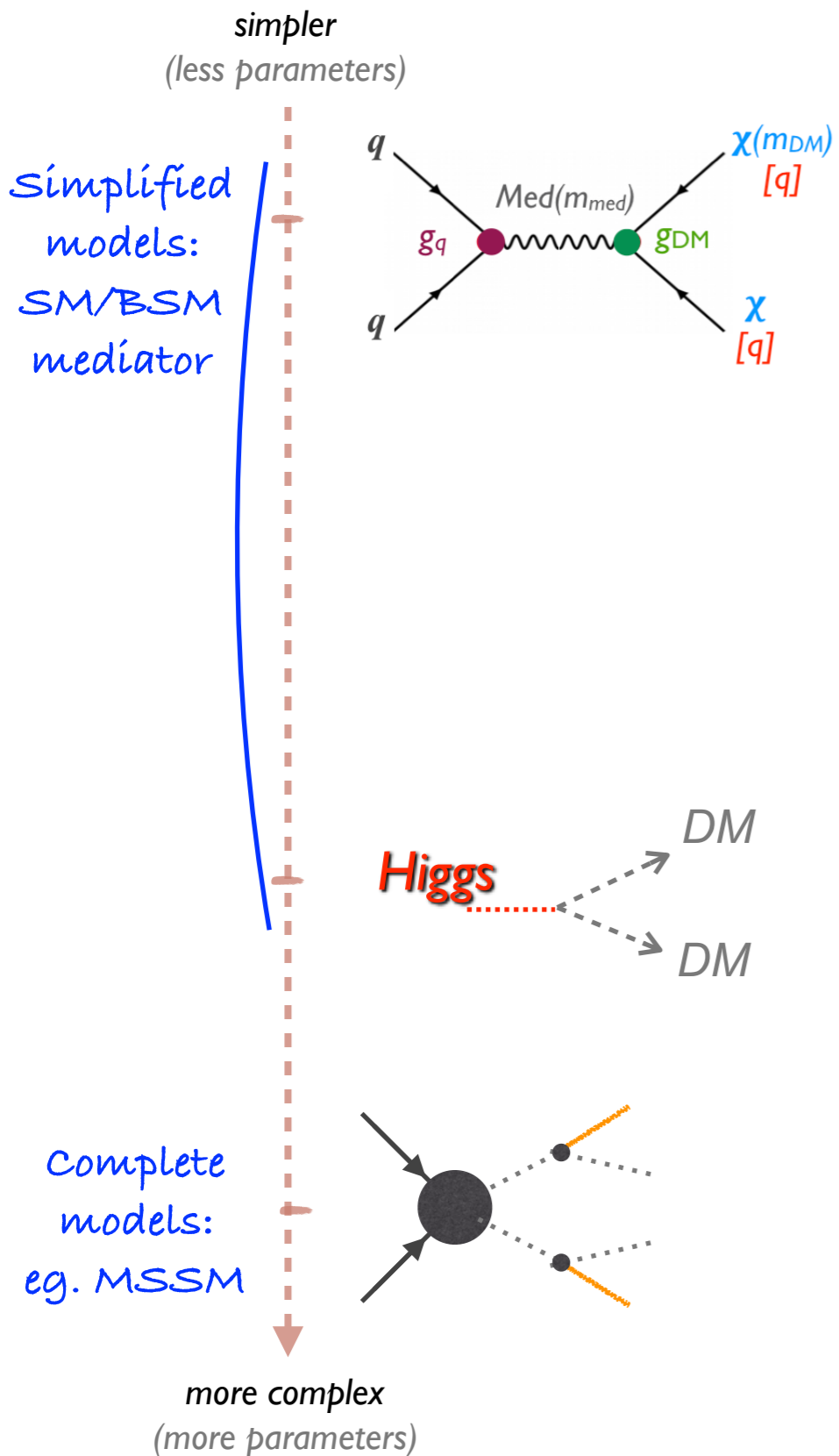
Higgs portal DM

- the SM-DM mediator could be the Higgs boson (spin-0)
- Higgs decay branching fractions not yet sufficiently constrained

SUSY models

- see more on N. Strobbe's talk

Dark matter? phenomenology at colliders



Beyond SM mediator

- one new mediator (spin-1 or spin-0), one DM particle
- set of parameters (m_{DM} , m_{med} , g_q , g_{DM})

	vector	axial-vector	
spin-1	$g_q \sum_q V_\mu \bar{q} \gamma^\mu q$	$g_q \sum_q A_\mu \bar{q} \gamma^\mu \gamma^5 q$	💡 choose X to increase xsec or bkg rejection
spin-0	$g_q \frac{\phi}{\sqrt{2}} \sum_f y_f \bar{f} f$	$g_q \frac{iA}{\sqrt{2}} \sum_f y_f \bar{f} \gamma^5 f$	💡 choose X to exploit coupling \propto to quark mass (or increase xsec)

Higgs portal DM

- the SM-DM mediator could be the Higgs boson (spin-0)
- Higgs decay branching fractions not yet sufficiently constrained

SUSY models

How do we search for DM at colliders?

Recap

- DM could be produced at colliders, rare process
- long lived and neutral, will appear as MET

► Signature: which DM process we want to study?

- phenomenology, eg. simplified model
- X visible particles, which decays?
- allow to identify main characteristics of process of interest (signal) 📖

Reminder:

<p>vector</p> $g_q \sum_q V_\mu \bar{q} \gamma^\mu q$	<p>axial-vector</p> $g_q \sum_q A_\mu \bar{q} \gamma^\mu \gamma^5 q$
<p>* choose X to increase xsec or bkg</p>	
<p>scalar</p> $g_q \frac{\phi}{\sqrt{2}} \sum_f y_f \bar{f} f$	<p>pseudoscalar</p> $g_q \frac{iA}{\sqrt{2}} \sum_f y_f \bar{f} \gamma^5 f$
<p>* choose X to exploit coupling \propto to quark mass (or increase xsec)</p>	



Pre-selection 📖

- in this case, large values of MET (from DM), 1 jet with high p_T (for energy conservation), no jets from b-quarks

How do we search for DM at colliders?

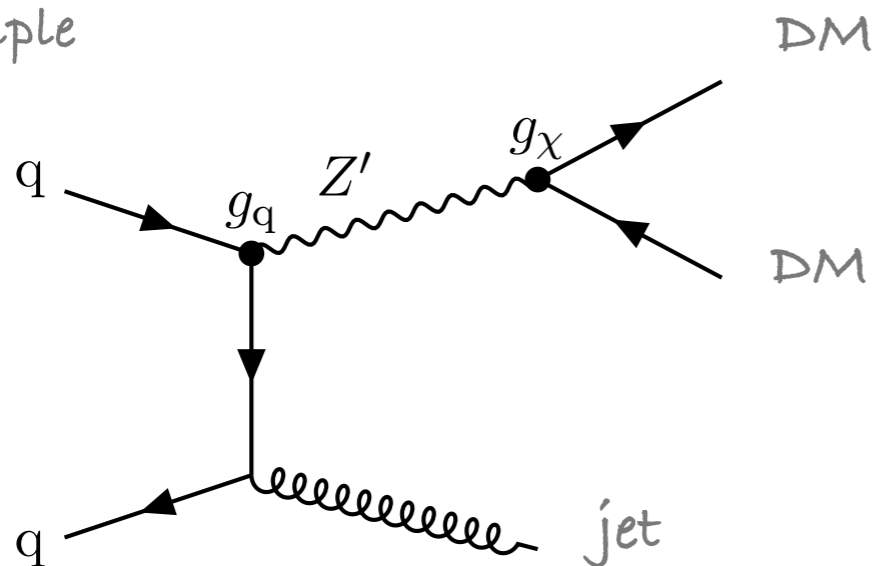
Recap

- DM could be produced at colliders, rare process
- long lived and neutral, will appear as MET

► Signature: which DM process we want to study?

- phenomenology, eg. simplified model
- \times visible particles, which decays?
- allow to identify main characteristics of process of interest
(signal) 📖

example



Reminder:

vector	axial-vector
$g_q \sum_q V_\mu \bar{q} \gamma^\mu q$	$g_q \sum_q A_\mu \bar{q} \gamma^\mu \gamma^5 q$
* choose χ to increase χ sec or $b\kappa_q$	
scalar	pseudoscalar
$g_q \frac{\phi}{\sqrt{2}} \sum_f y_f \bar{f} f$	$g_q \frac{iA}{\sqrt{2}} \sum_f y_f \bar{f} \gamma^5 f$
* choose χ to exploit coupling \propto to quark mass (or increase χ sec)	

Pre-selection 📖

- large values of MET (from DM), 1 jet with high p_T (for energy conservation), no jets from b-quarks

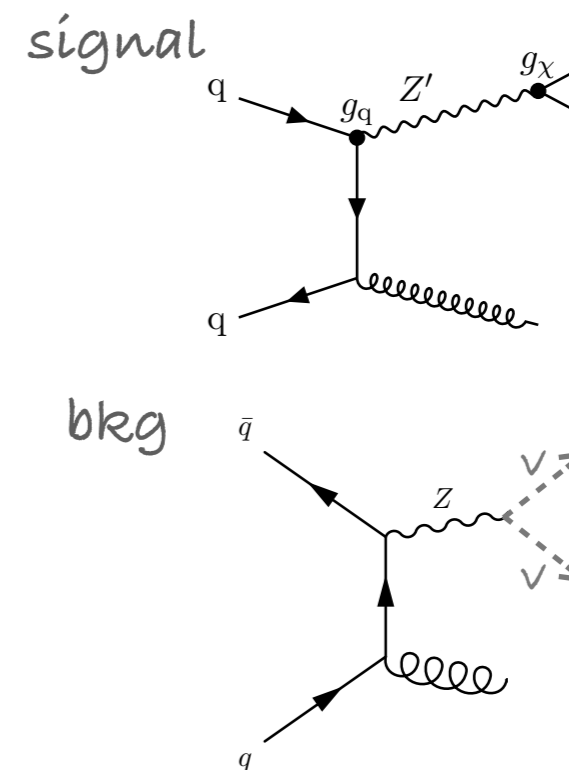
How do we search for DM at colliders?

Recap

- DM could be produced at colliders, rare process
- long lived and neutral, will appear as MET

► 1- Selection

- many SM processes can have similar characteristics (or fake them) as the signal - *SM background* 📄🔍
- these SM processes are much more probable than signal
- require additional criteria to enhance the signal vs background - signal region (SR) 📄



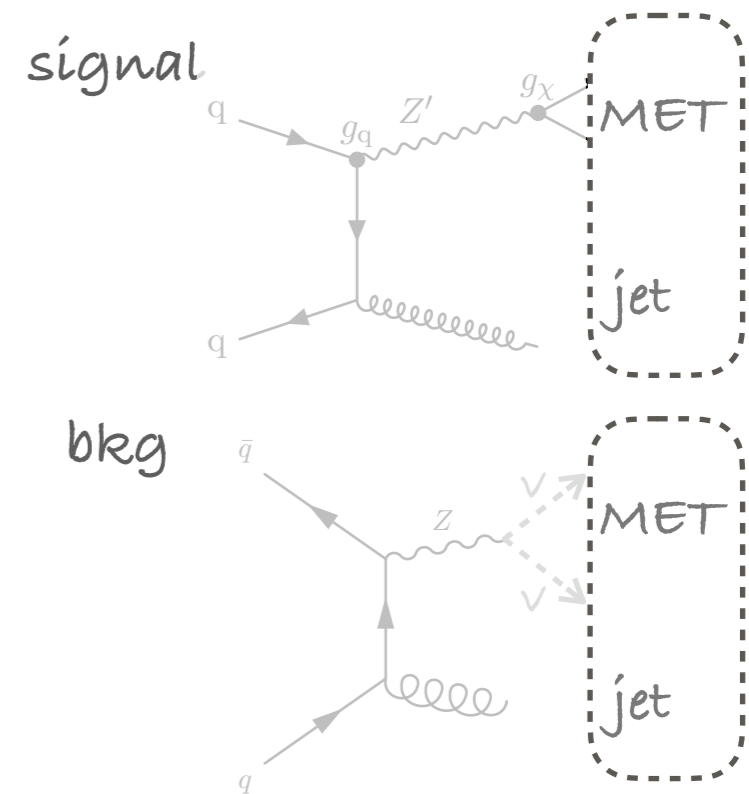
How do we search for DM at colliders?

Recap

- DM could be produced at colliders, rare process
- long lived and neutral, will appear as MET

► 1- Selection

- many SM processes can have similar characteristics (or fake them) as the signal - SM background 📄🔍
- these SM processes are much more probable than signal
- require additional criteria to enhance the signal vs background - signal region (SR) 📄



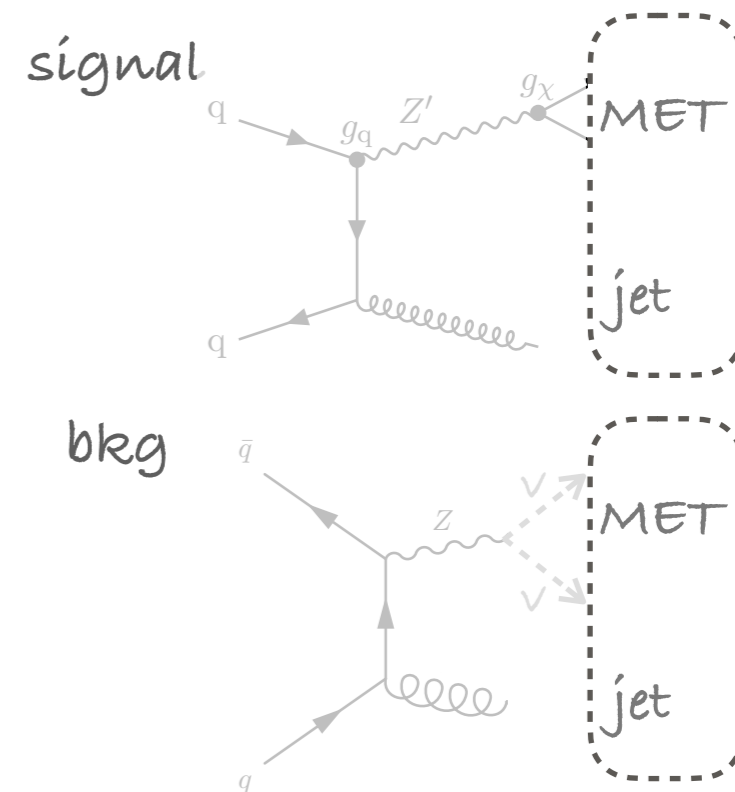
How do we search for DM at colliders?

Recap

- DM could be produced at colliders, rare process
- long lived and neutral, will appear as MET

► 1- Selection

- many SM processes can have similar characteristics (or fake them) as the signal - *SM background* 📖🔍
- these SM processes are much more probable than signal
- require additional criteria to enhance the signal vs background - *signal region (SR)* 📖🔍



SR
----->



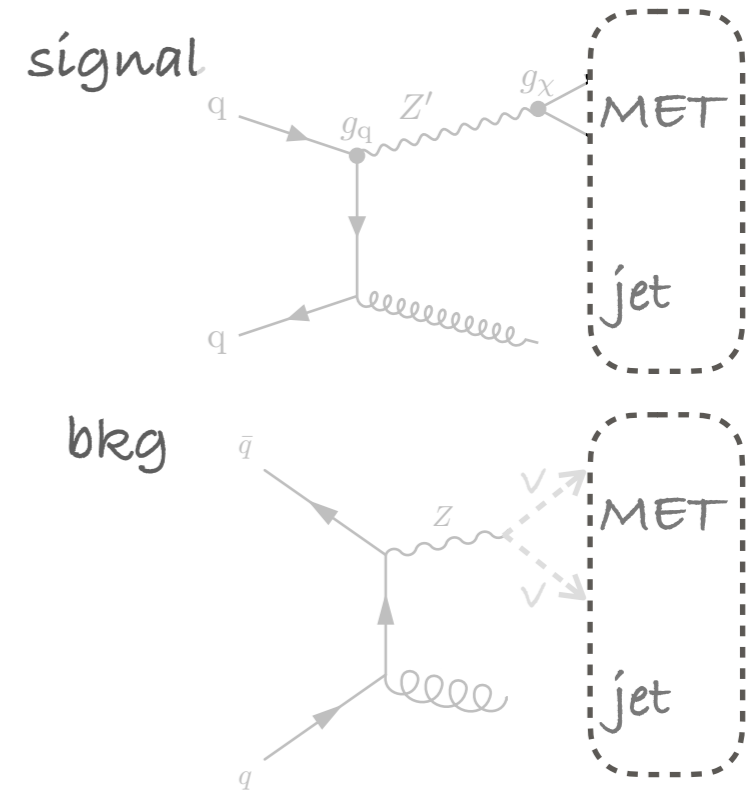
How do we search for DM at colliders?

Recap

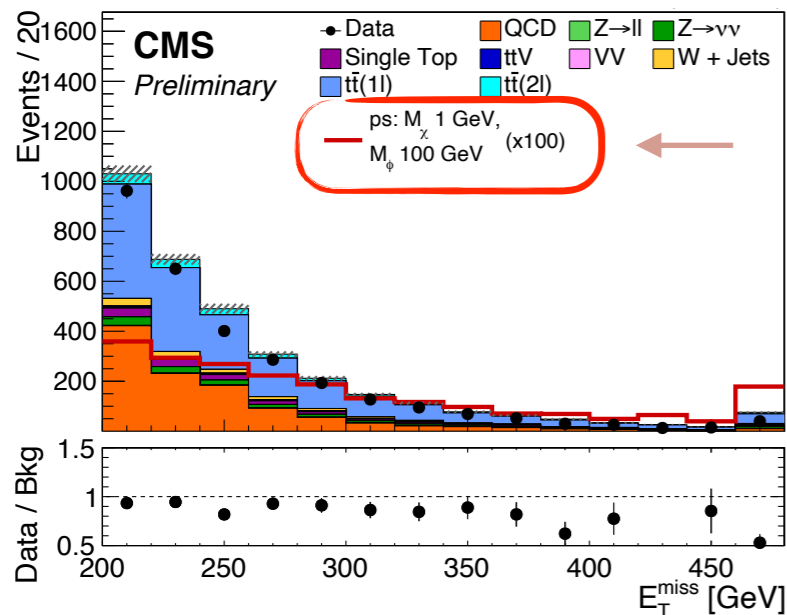
- DM could be produced at colliders, rare process
- long lived and neutral, will appear as MET

► 1- Selection

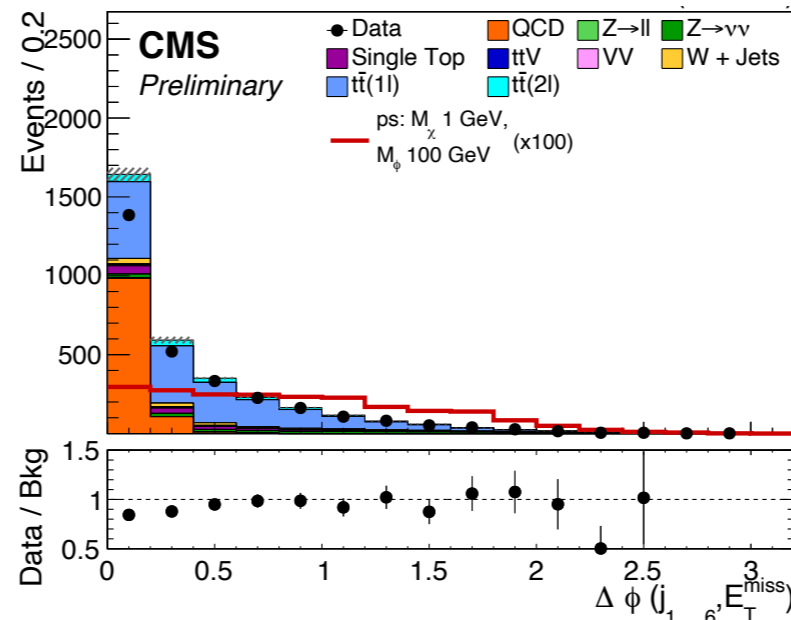
- many SM processes can have similar characteristics (or fake them) as the signal - SM background
- these SM processes are much more probable than signal
- require additional criteria to enhance the signal vs background - signal region (SR)



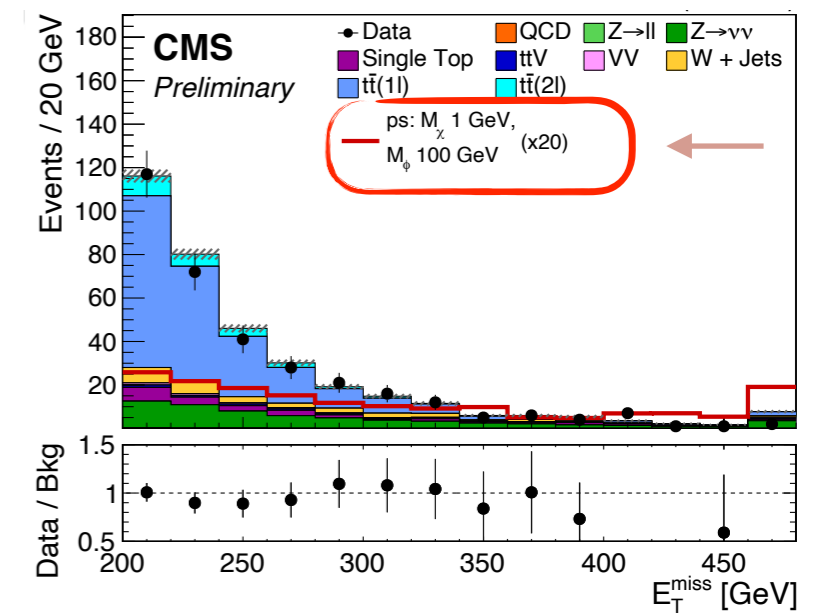
pre-selection



additional criteria



signal region



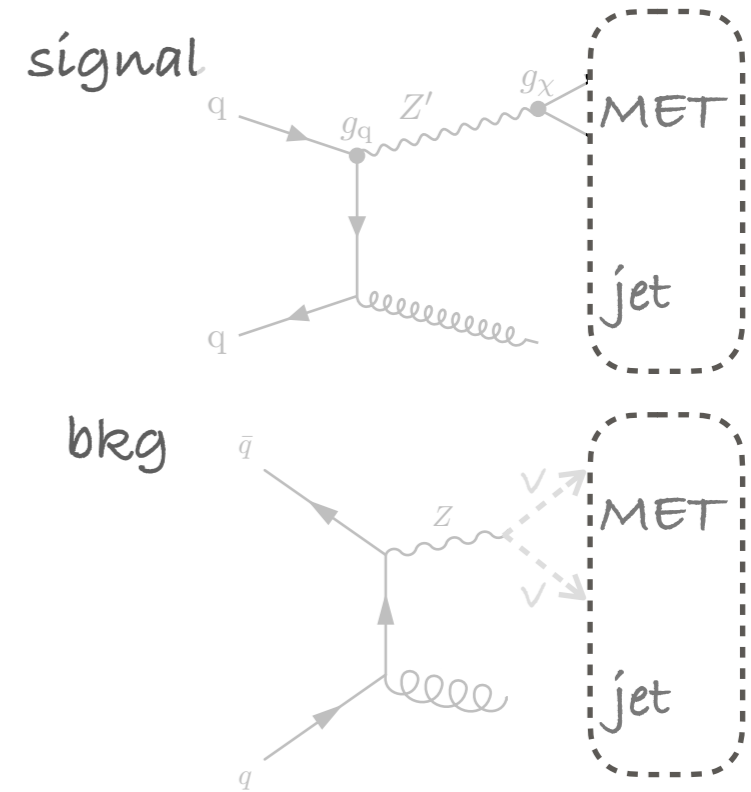
How do we search for DM at colliders?

Recap

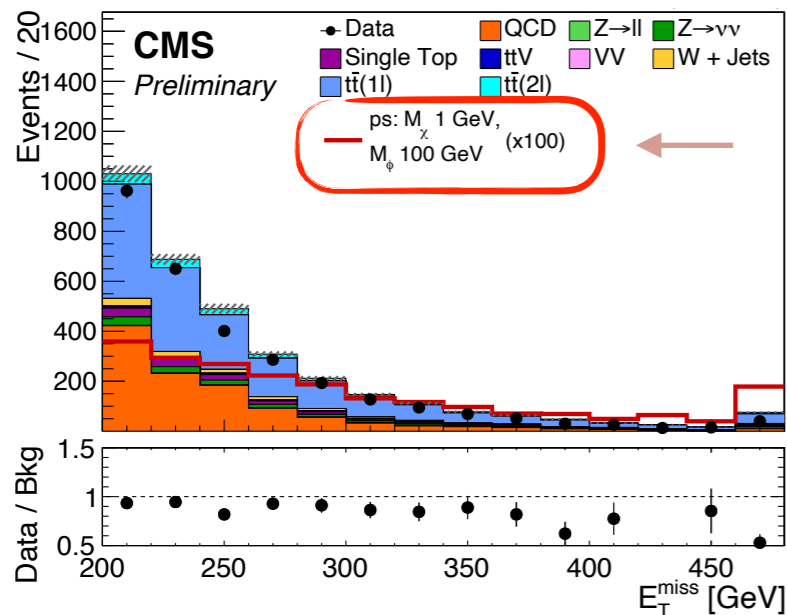
- DM could be produced at colliders, rare process
- long lived and neutral, will appear as MET

► 1- Selection

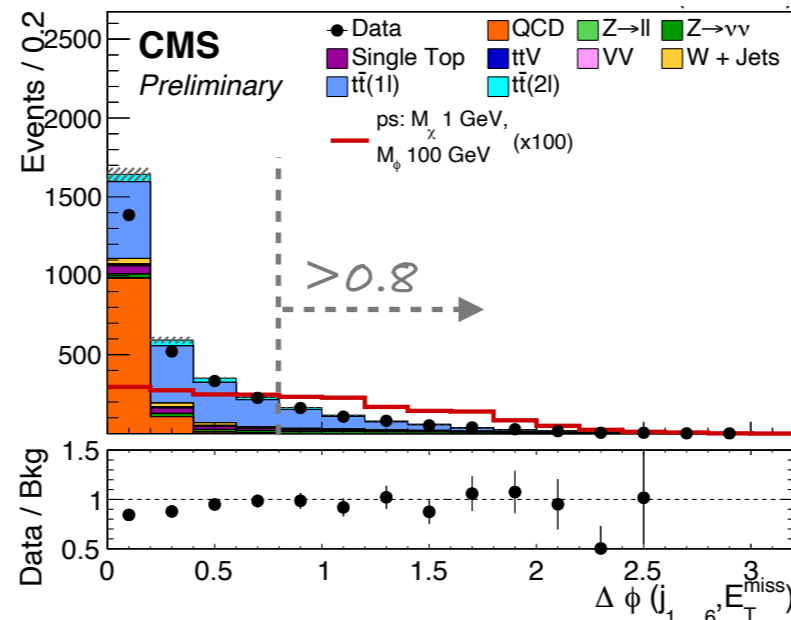
- many SM processes can have similar characteristics (or fake them) as the signal - SM background
- these SM processes are much more probable than signal
- require additional criteria to enhance the signal vs background - signal region (SR)



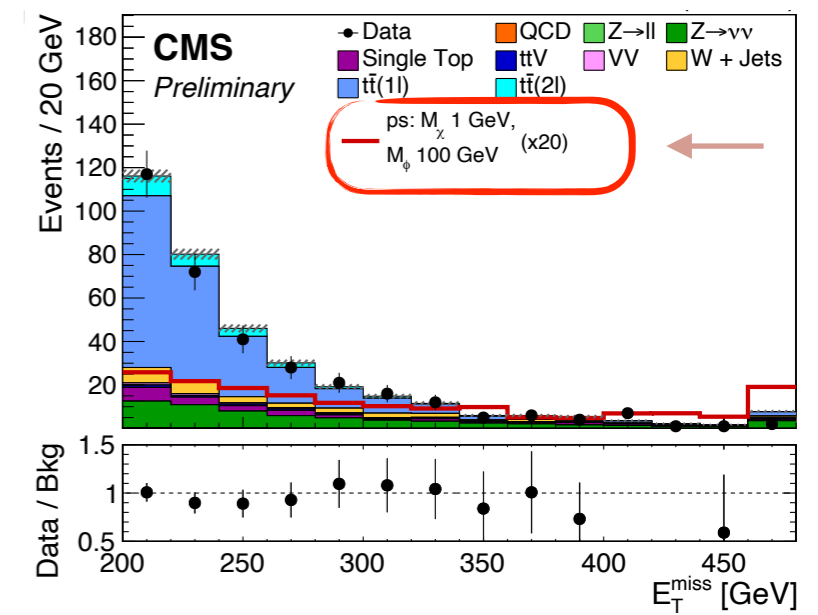
pre-selection



additional criteria



signal region



How do we search for DM at colliders?

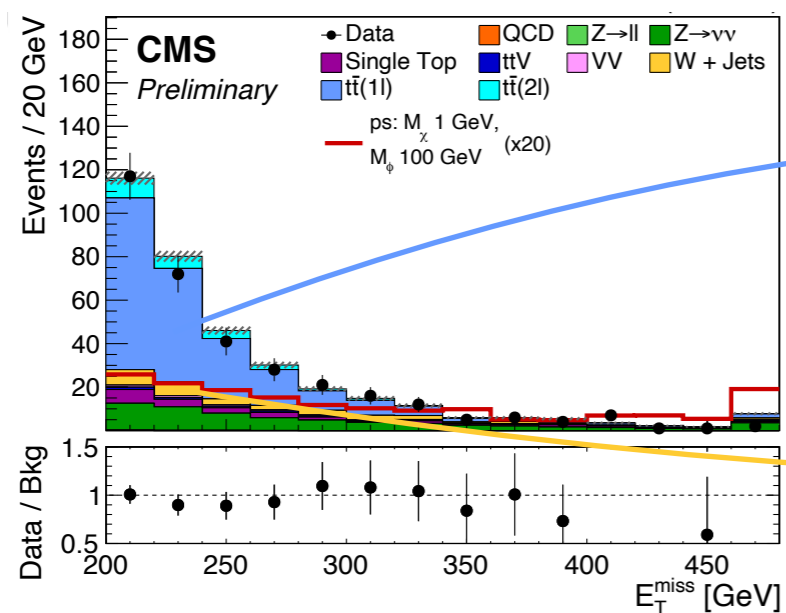
Recap

- DM could be produced at colliders, rare process
- long lived and neutral, will appear as MET

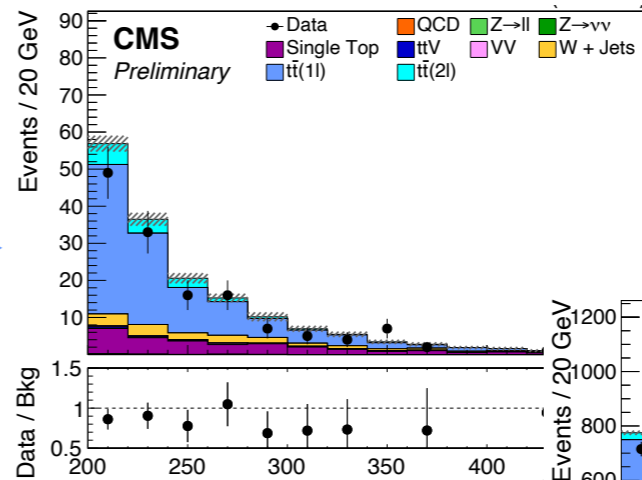
► 2- Background

- DM production is a rare process. We need a precise modeling and evaluation of SM bkg in SR essential to “see” the signal
- Achieved through use of multiple control regions (CRs)
- CR definition: similar to the SR, good purity in bkg we want to check, no signal

signal region

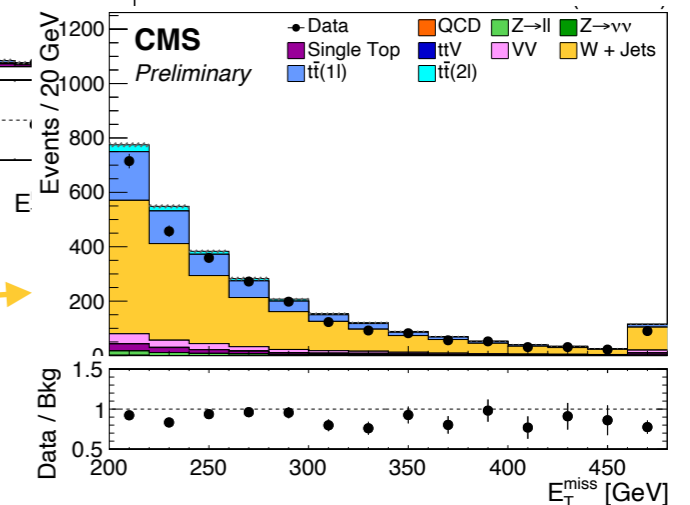


tt(1l)



control regions

W+jets



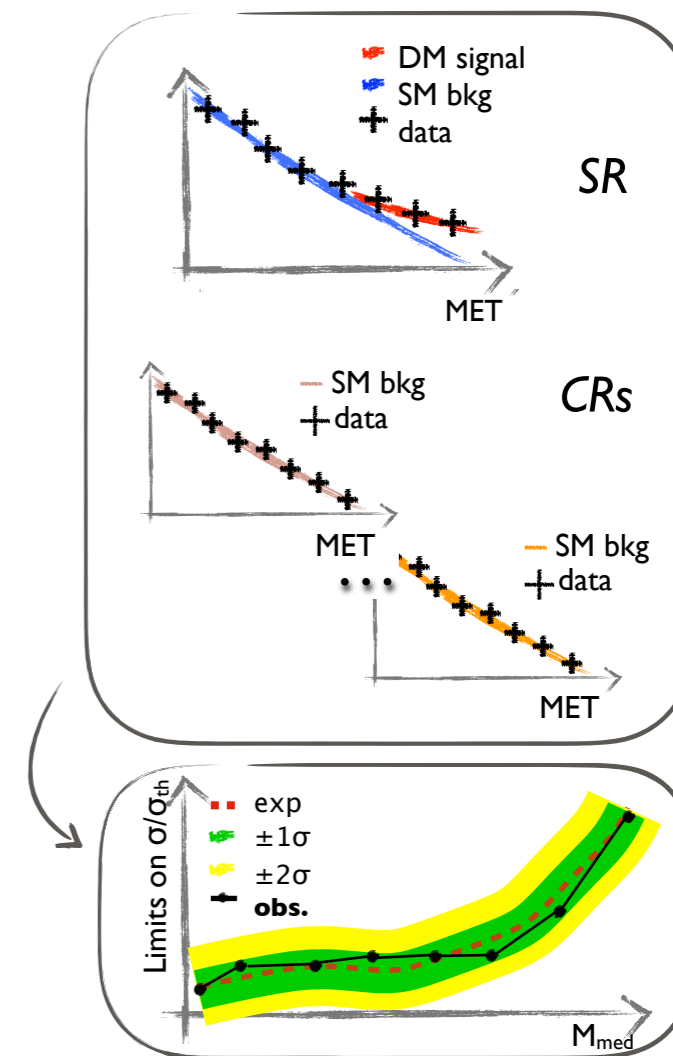
How do we search for DM at colliders?

Recap

- DM could be produced at colliders, rare process
- long lived and neutral, will appear as MET

3- Results

- DM appears as excess of events in MET tail in SR wrt SM background
 - no very striking signature, eg. mass peak, m_T kinematic endpoint
- excess of events in data. Did we find DM?
- no excess, interpret result in terms of model parameters



Experimental challenges

- * accurate E calibration/resolution of visible objects (*"fake" MET from mis-measured jets*)
- * precise particle reconstruction and identification
- * mitigate effects from additional pp collisions (pile-up)
- * MET thresholds affected by trigger (very high collision rates)

DARK MATTER

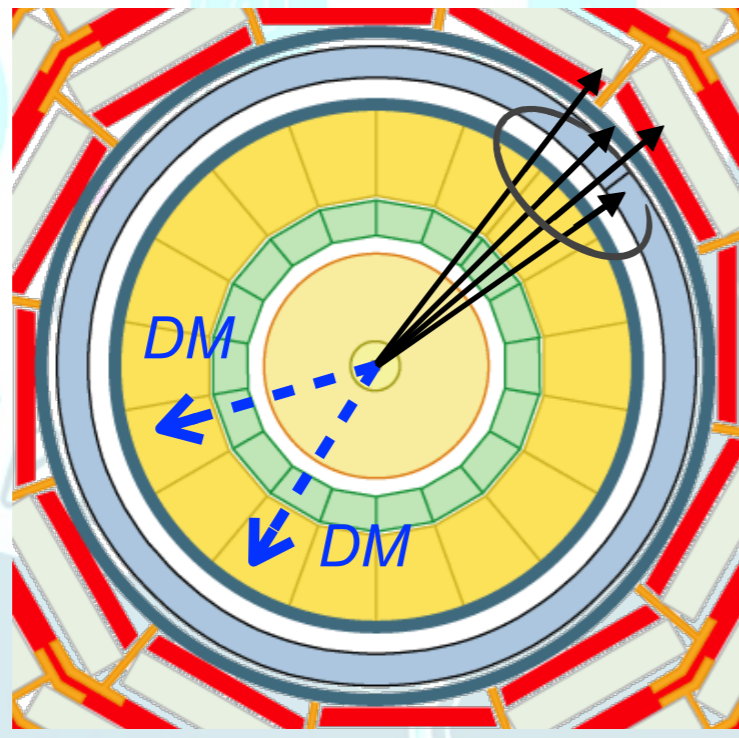
Simplified models

Reminder:

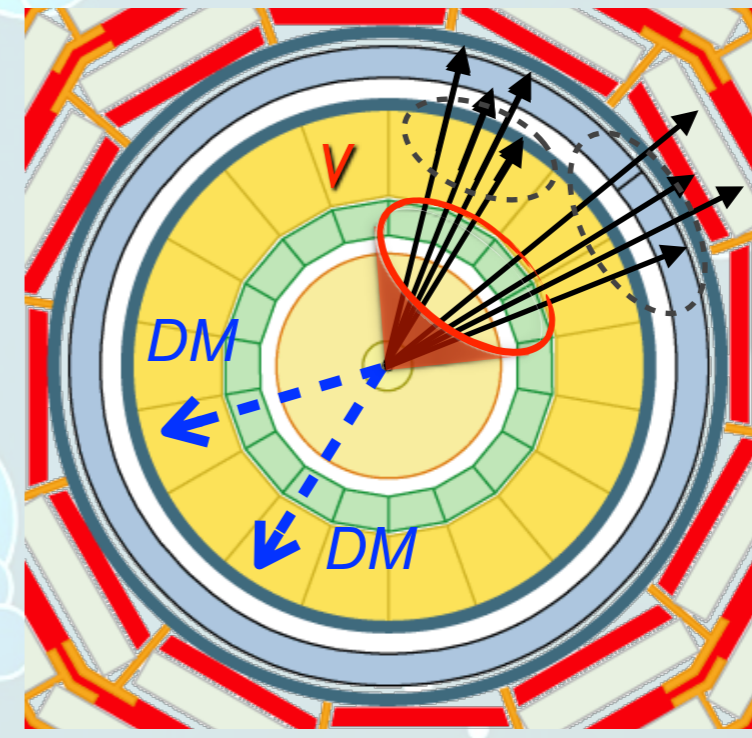
<p>vector</p> $g_q \sum_q V_\mu \bar{q} \gamma^\mu q$ <p>* choose X to increase xsec or bkg rejection</p>	<p>axial-vector</p> $g_q \sum_q A_\mu \bar{q} \gamma^\mu \gamma^5 q$
<p>scalar</p> $g_q \frac{\phi}{\sqrt{2}} \sum_f y_f \bar{f} f$ <p>* choose X to exploit coupling \propto to quark mass (or increase xsec)</p>	<p>pseudoscalar</p> $g_q \frac{iA}{\sqrt{2}} \sum_f y_f \bar{f} \gamma^5 f$

Signature: large MET and ≥ 1 high- p_T jet/vector boson

mono-jet



mono-V(=W,Z)

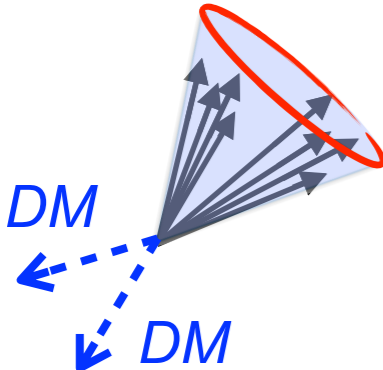


* CMS: EXO-20-004 (2016-2018)

DM+jet/V search

► 1 - Selection: events categorized based on jet nature

mono-V




DM

DM

j [large-cone jet, eg. $R=0.8$]

- * MET > 250 GeV
- * ≥ 1 jets, $p_T(j_1) > 250$ GeV
- * machine learning technique to identify V hadronic decays
- * categories based on MVA score
- * jet mass consistent with V

mono-jet



DM

DM

j [small-cone jet, eg. $R=0.4$]

- * not selected as mono-V
- * ≥ 1 jets, $p_T(j) > 100$ GeV
- * b-tagged jets veto

► 2- Bkg:

- Z(vv)+jets and W(lv)+jets from CRs

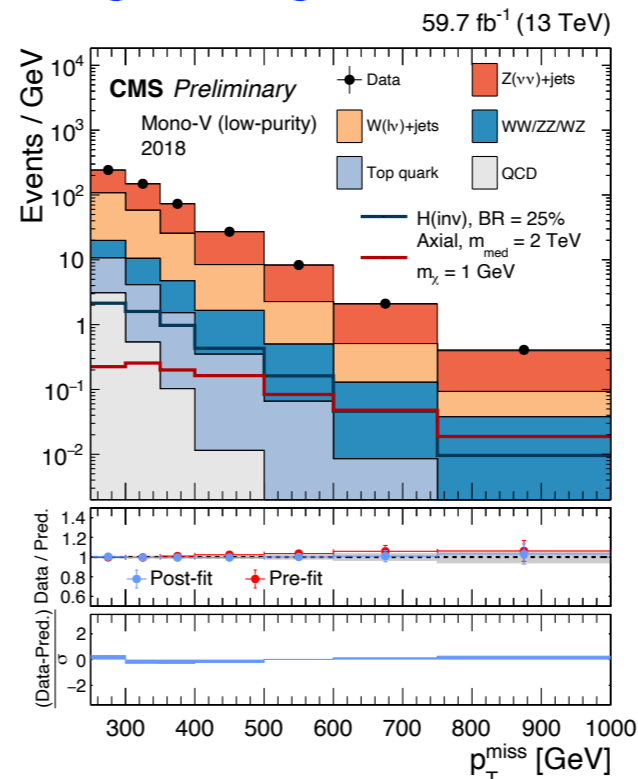
► 3- Results: combined fit of SRs, CRs

- systematic unc. included as nuisance parameters

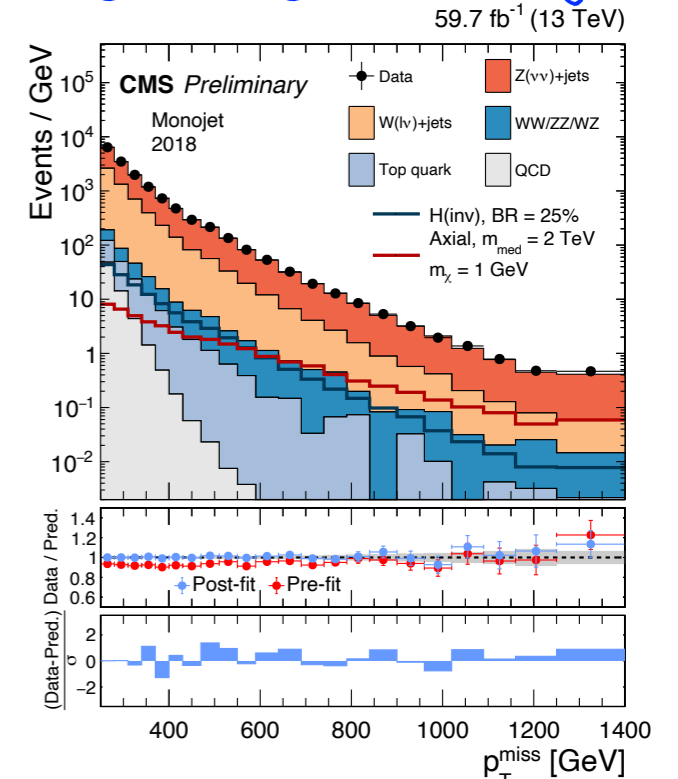
* prefit: as from simulation

* post-fit: after allowing simulation to vary within unc and the scaling factors from CRs

signal region mono-V

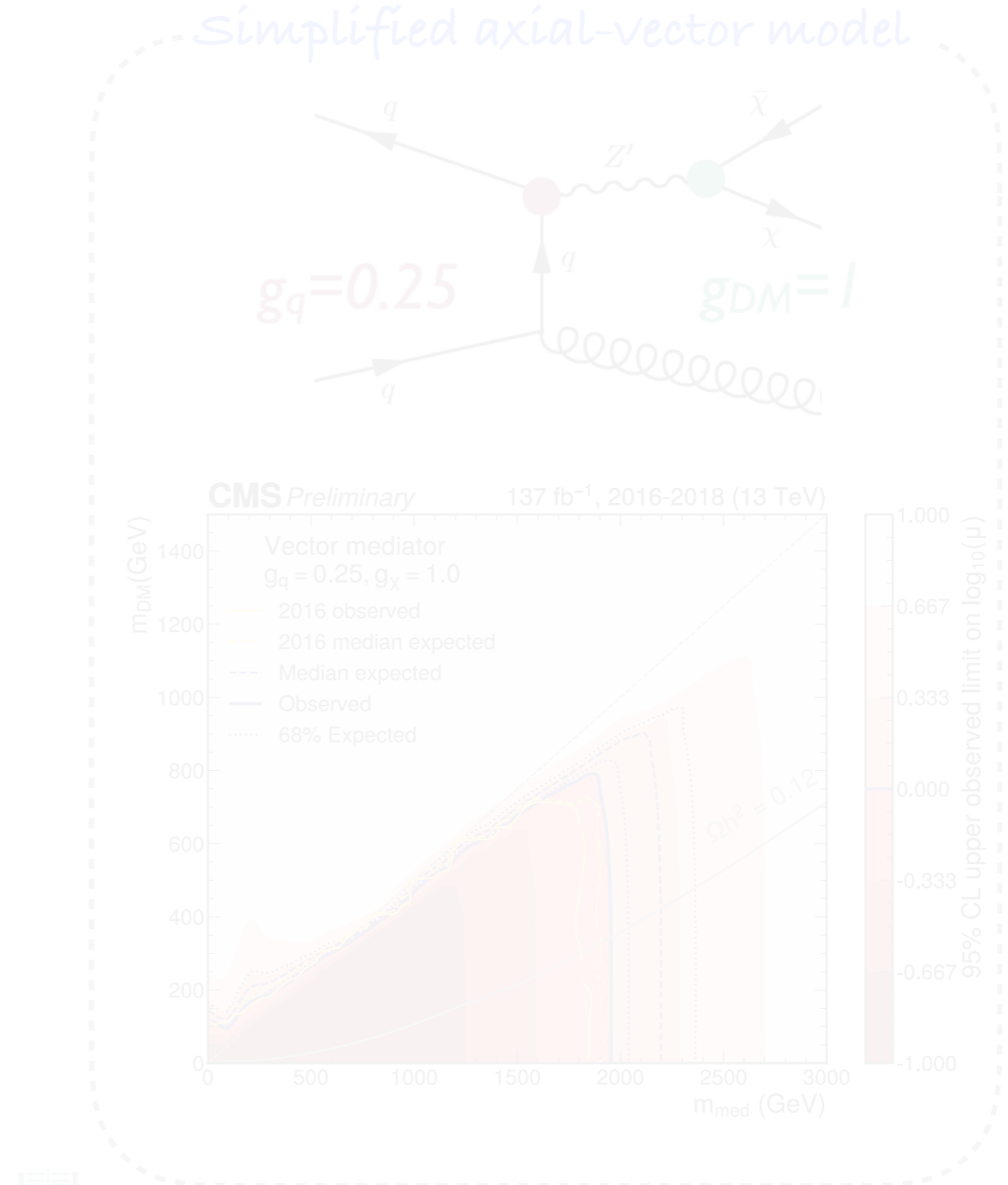
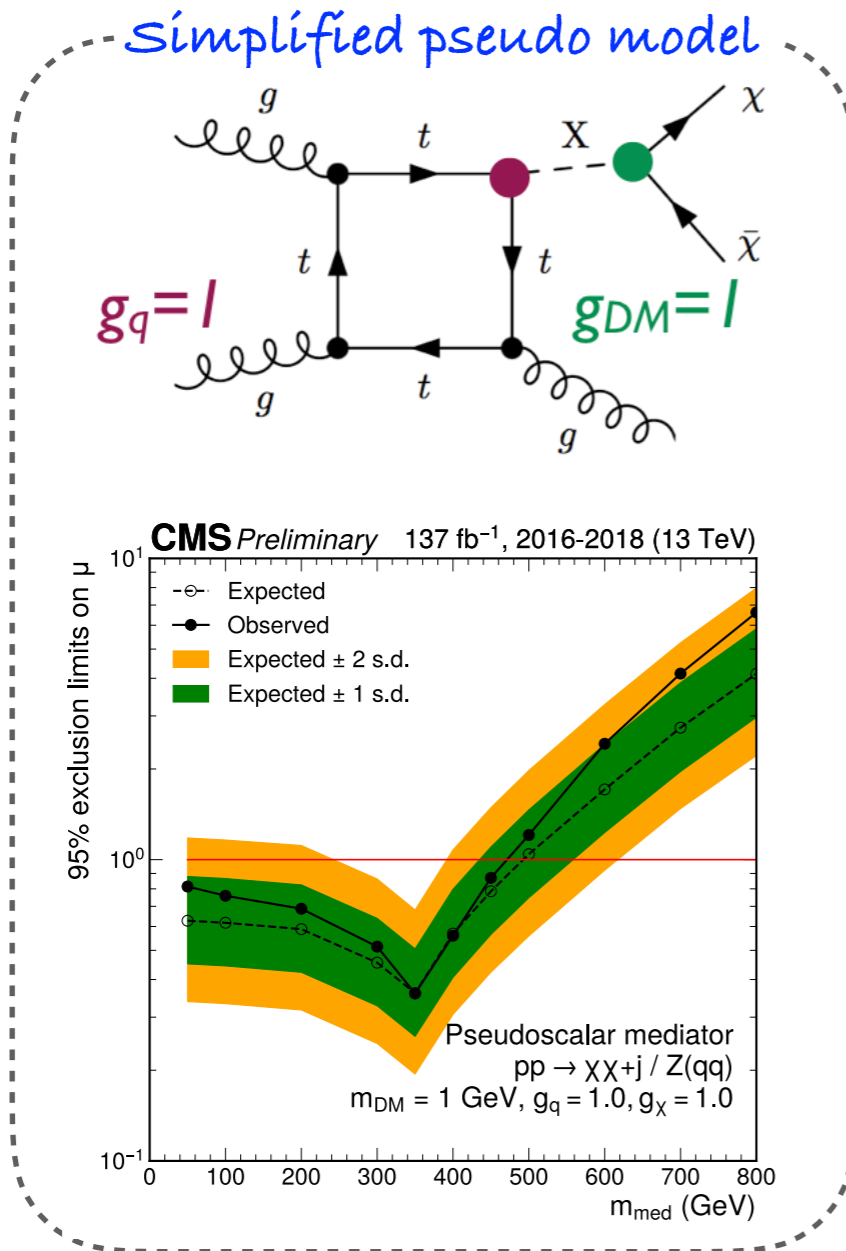


signal region mono-jet



DM+jet/V search

► 3- Results: interpretation in terms of DM model, upper limits at 95% CL on cross section

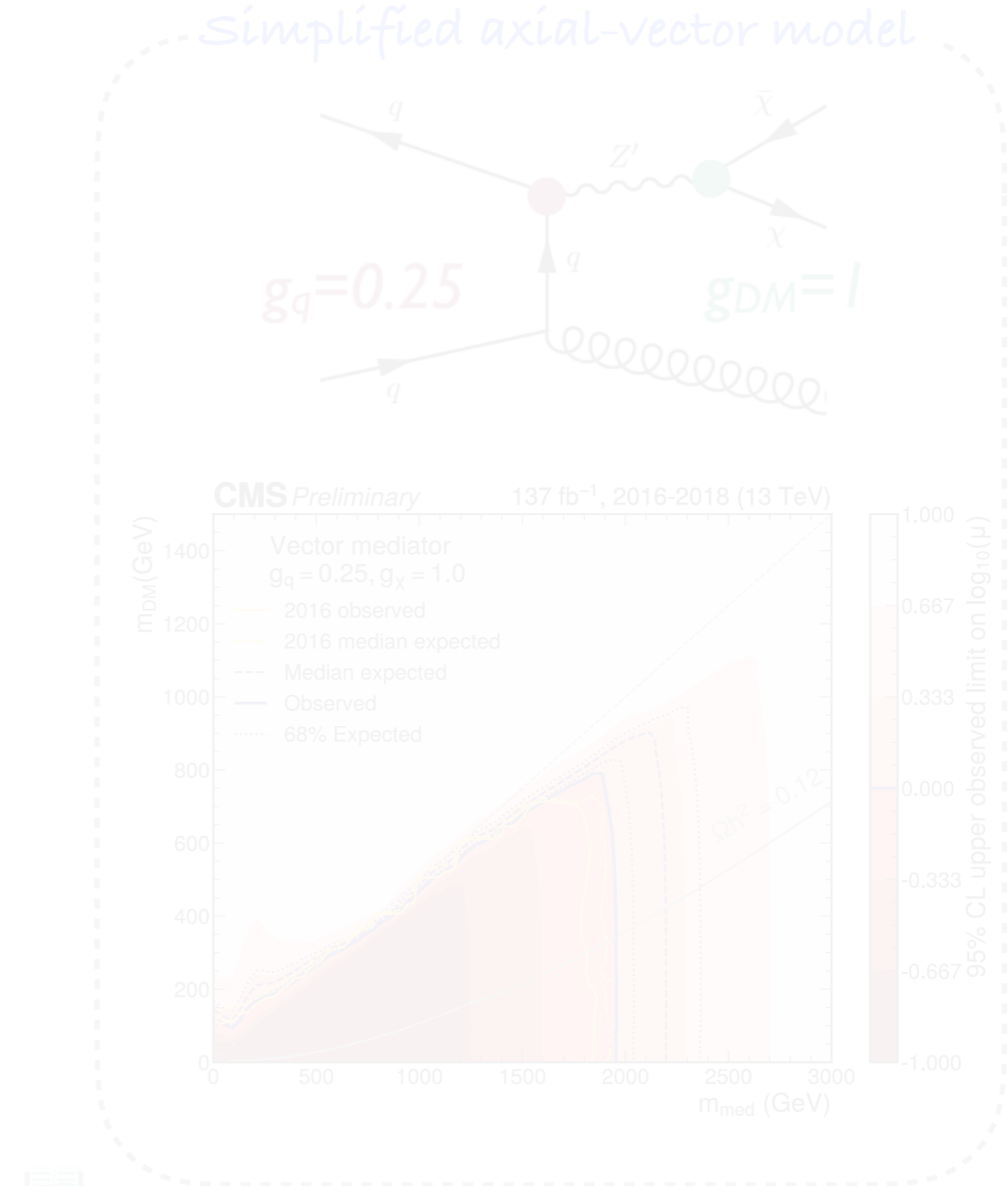
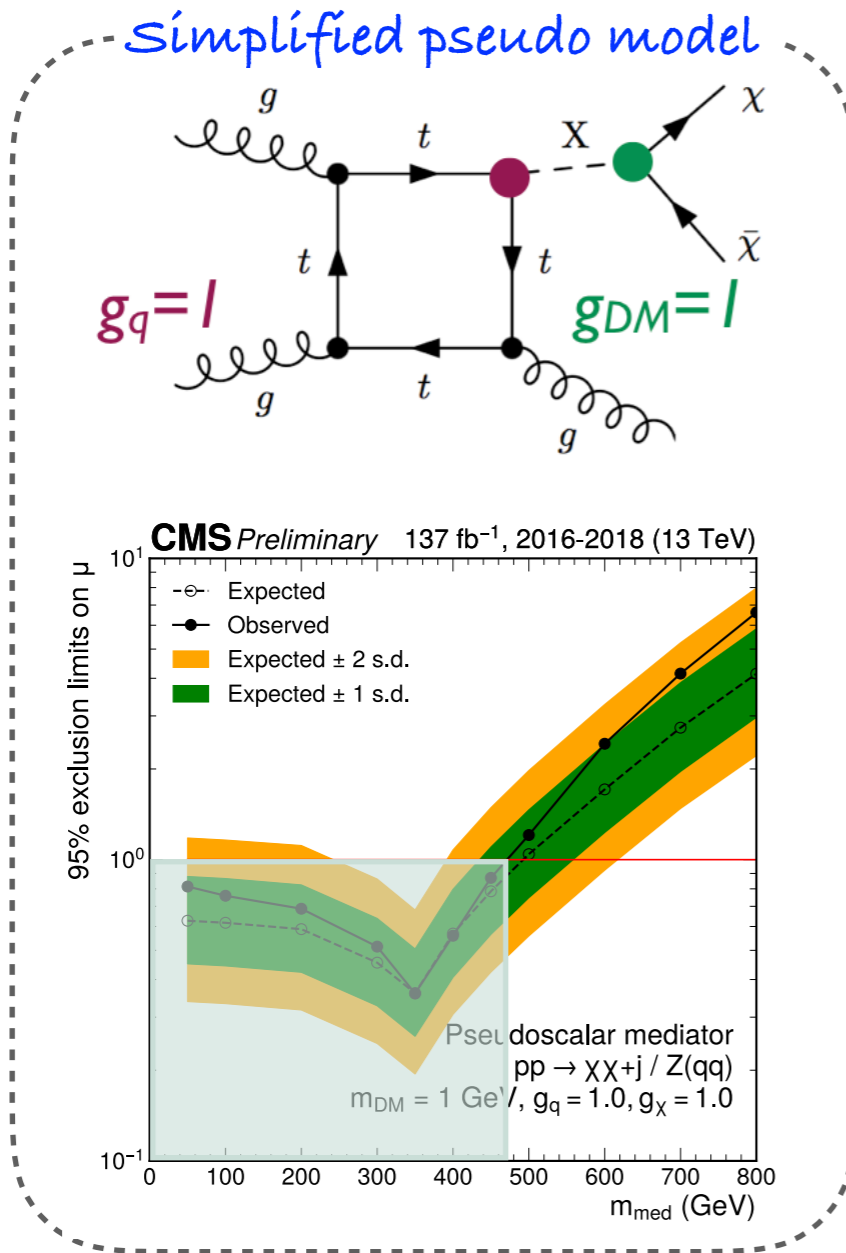


- * $\mu = \sigma / \sigma_{th}$, $\mu = 1$ exclude the theory value, $\mu < 1$ exclude below theory value, $\mu > 1$ does not exclude theory value
- * parameter: cannot scan all parameters at once. Fixed ones only affect σ_{sec} but not kinematic (selection)

- * here $\mu = \sigma / \sigma_{th}$ in on z axis (notice here $\log_{10}(\mu)$)
 - * parameter: here we scan 2 parameters at the time
- see also N. Strobbe's talk

DM+jet/V search

► 3- Results: interpretation in terms of DM model, upper limits at 95% CL on cross section

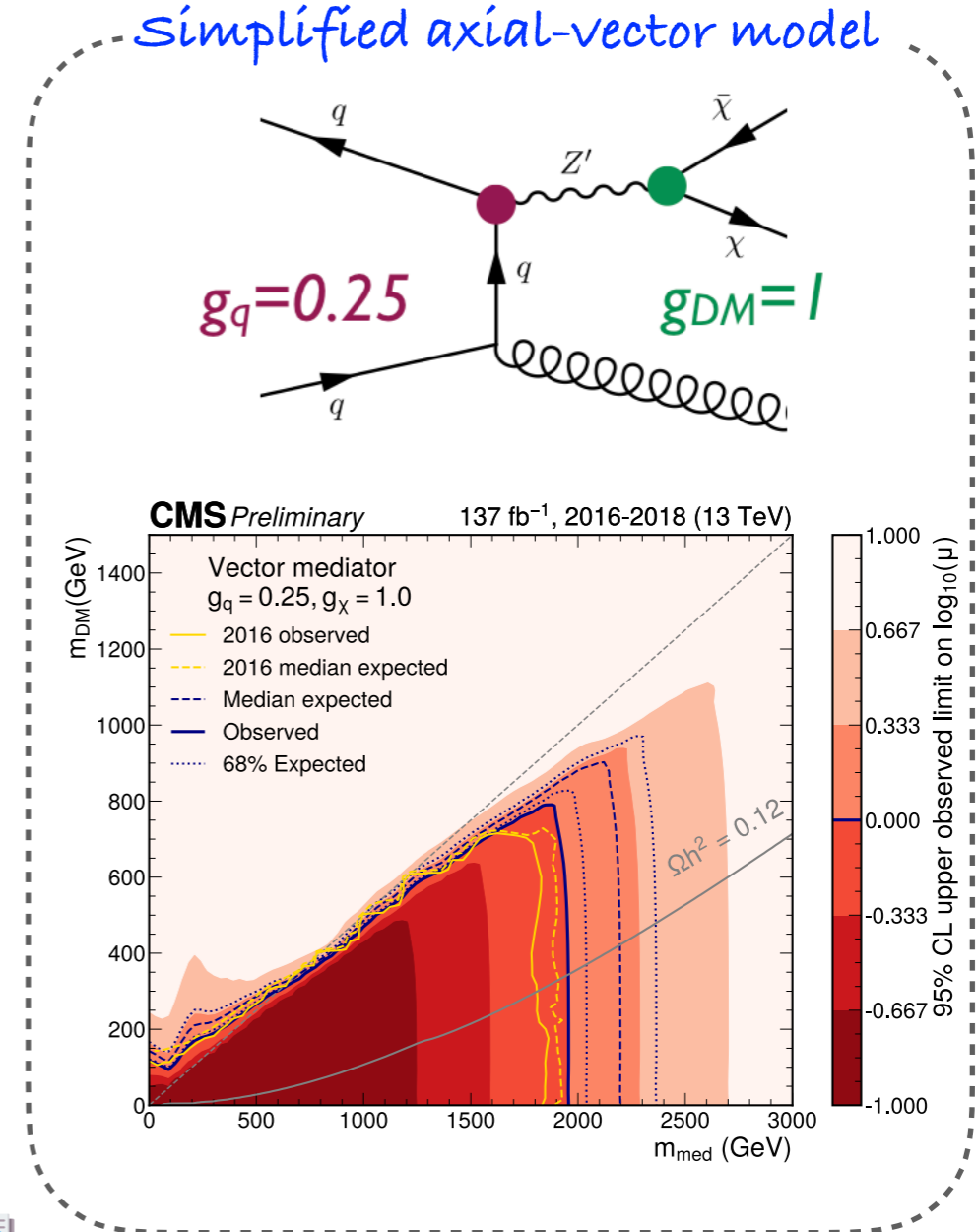
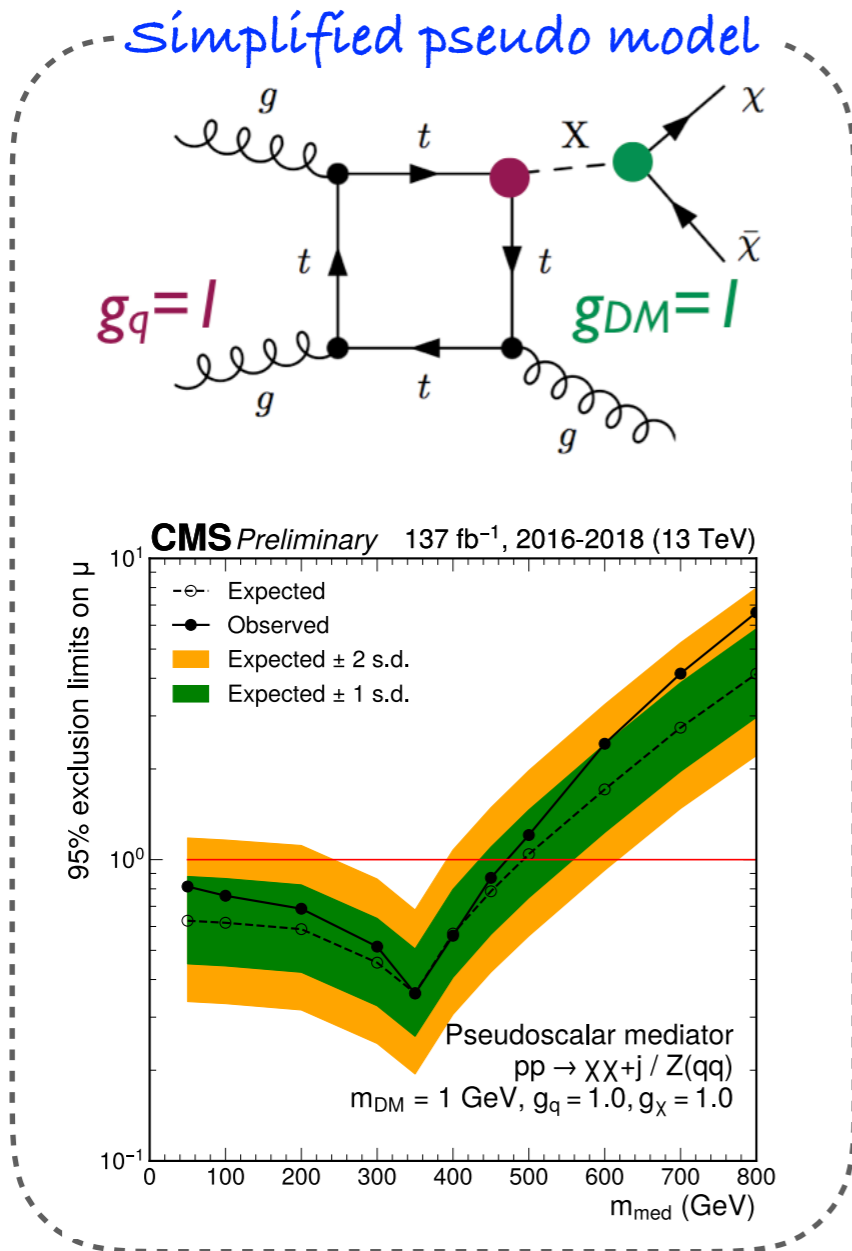


- * $\mu = \sigma / \sigma_{th}$, $\mu = 1$ exclude the theory value, $\mu < 1$ exclude below theory value, $\mu > 1$ does not exclude theory value
- * parameter: cannot scan all parameters at once. Fixed ones only affect xsec but not kinematic (selection)

- * here $\mu = \sigma / \sigma_{th}$ in on z axis (notice here $\log_{10}(\mu)$)
 - * parameter: here we scan 2 parameters at the time
- see also N. Strobbe's talk

DM+jet/V search

► 3- Results: interpretation in terms of DM model, upper limits at 95% CL on cross section

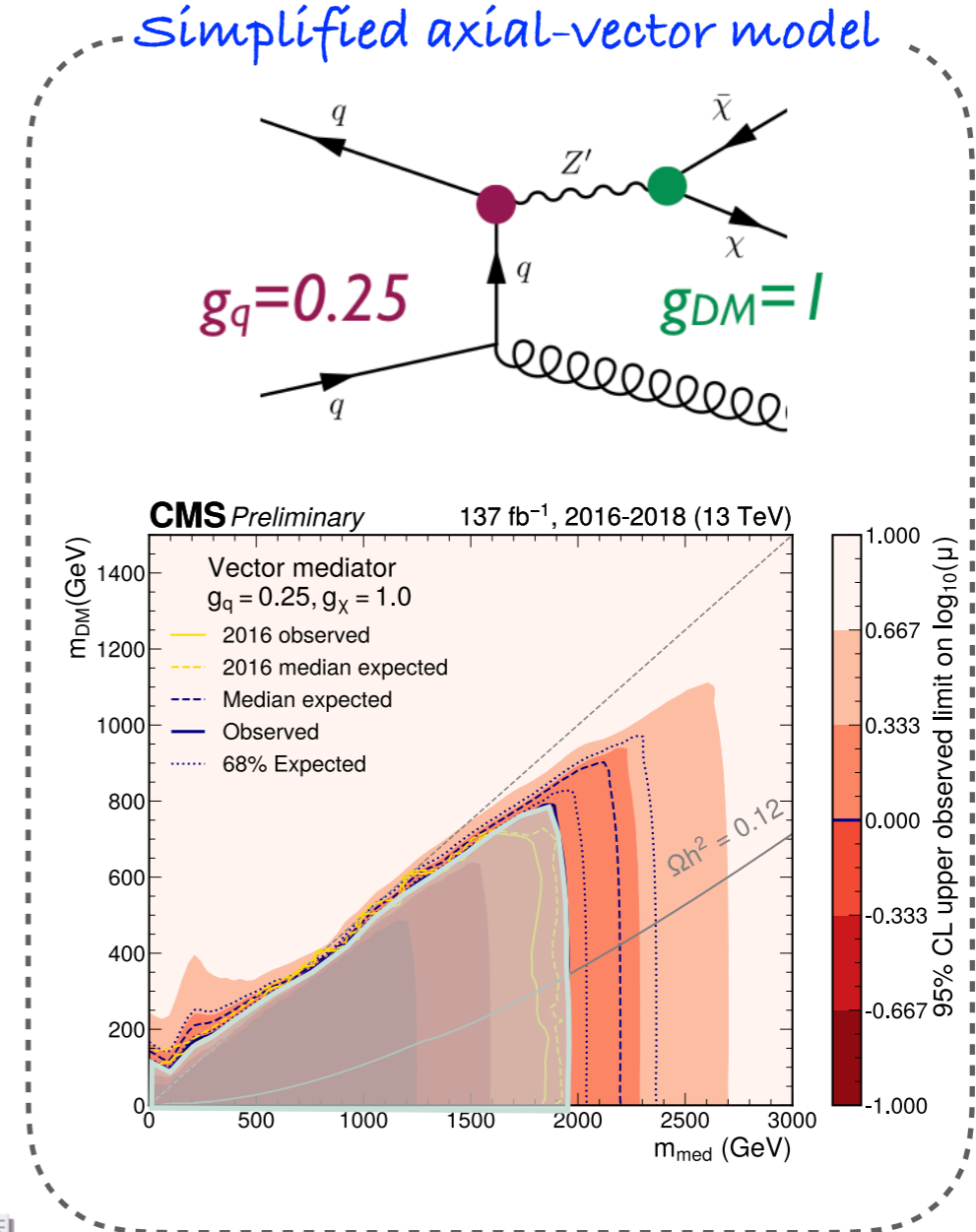
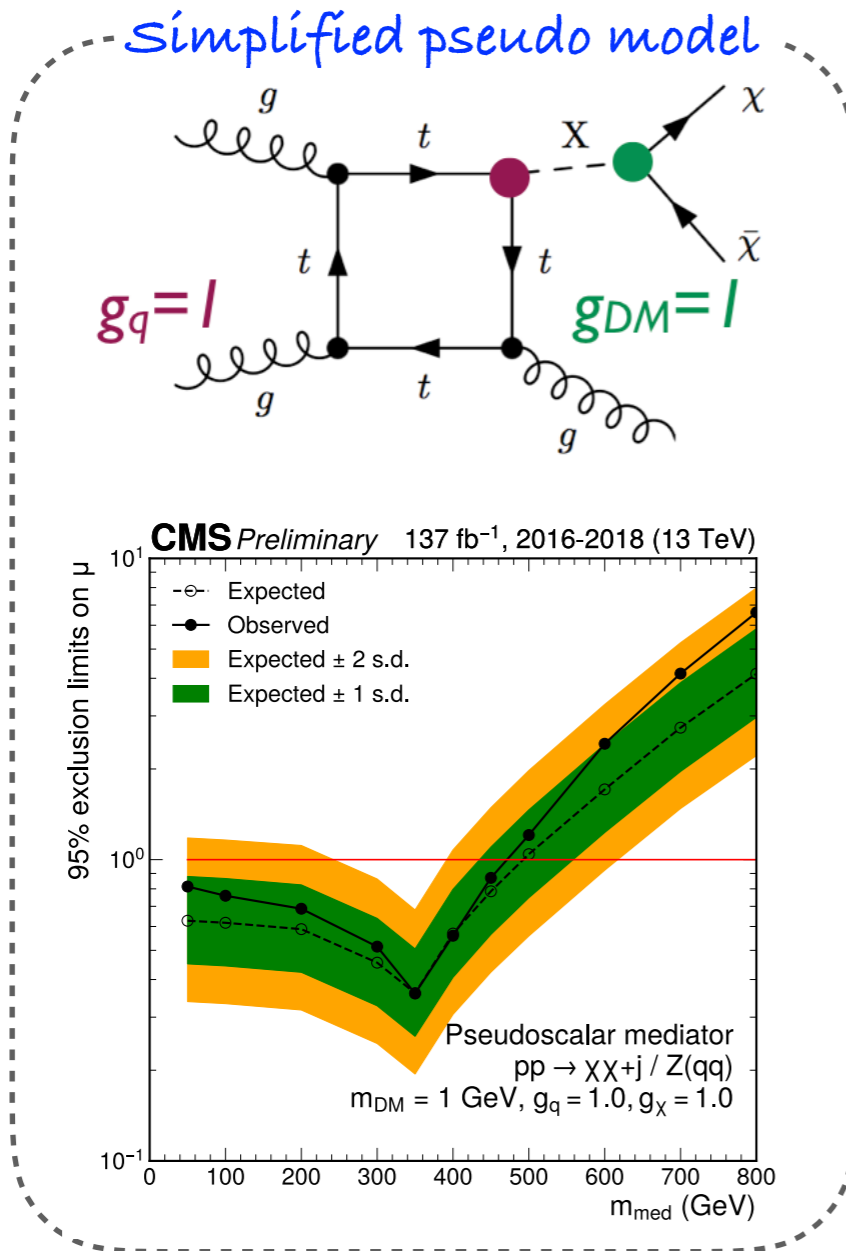


- * $\mu = \sigma/\sigma_{th}$, $\mu = 1$ exclude the theory value, $\mu < 1$ exclude below theory value, $\mu > 1$ does not exclude theory value
- * parameter: cannot scan all parameters at once. Fixed ones only affect xsec but not kinematic (selection)

- * here $\mu = \sigma/\sigma_{th}$, in on z axis (notice here $\log_{10}(\mu)$)
- * parameter: here we scan 2 parameters at the time

DM+jet/V search

► 3- Results: interpretation in terms of DM model, upper limits at 95% CL on cross section

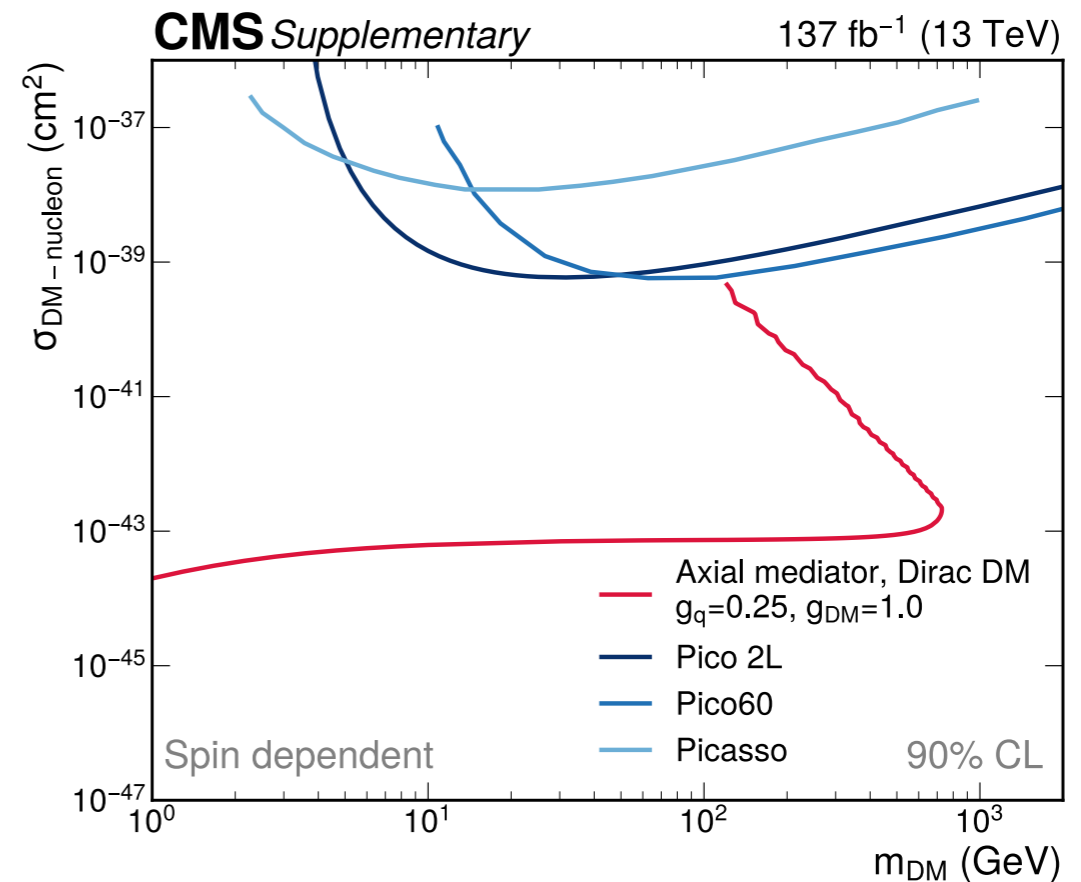
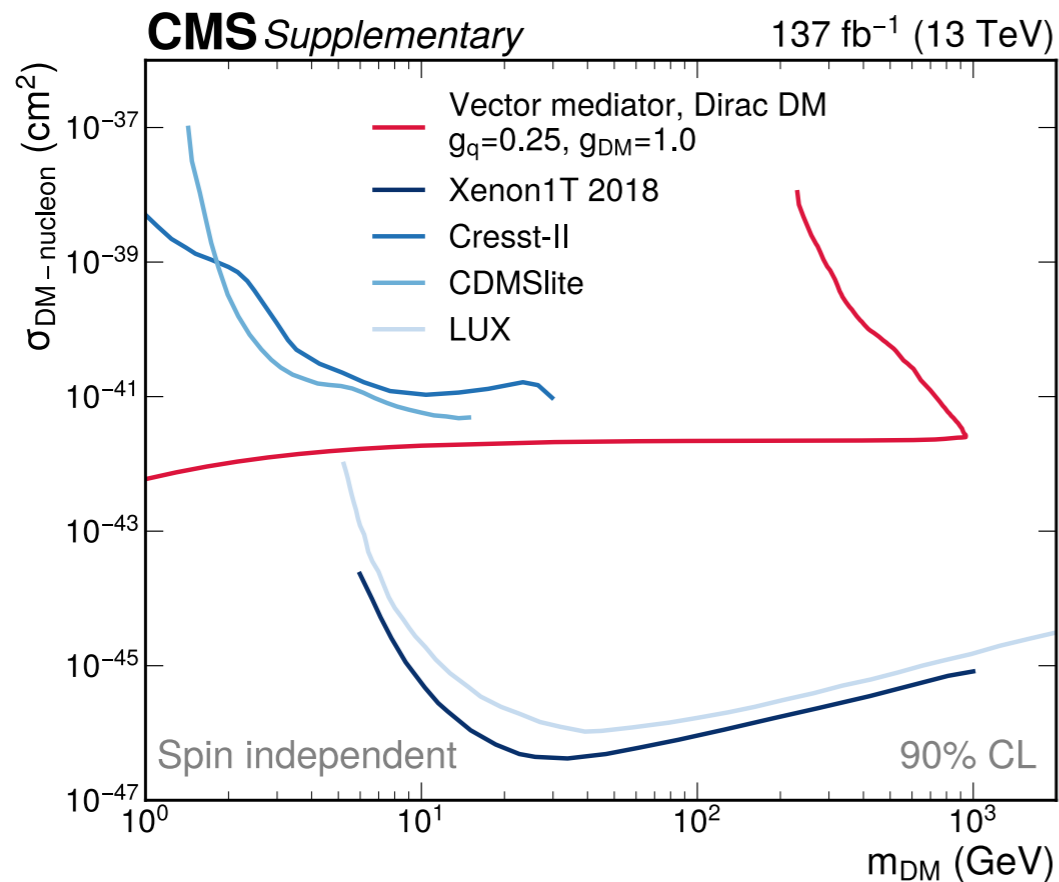


- * $\mu = \sigma / \sigma_{th}$, $\mu = 1$ exclude the theory value, $\mu < 1$ exclude below theory value, $\mu > 1$ does not exclude theory value
- * parameter: cannot scan all parameters at once. Fixed ones only affect xsec but not kinematic (selection)

- * here $\mu = \sigma / \sigma_{th}$, in on z axis (notice here $\log_{10}(\mu)$)
- * parameter: here we scan 2 parameters at the time

DM+jet/V: interplay with direct detection

► **3- Results:** lower limits at 90% CL on interaction cross section between DM candidates and the nuclei



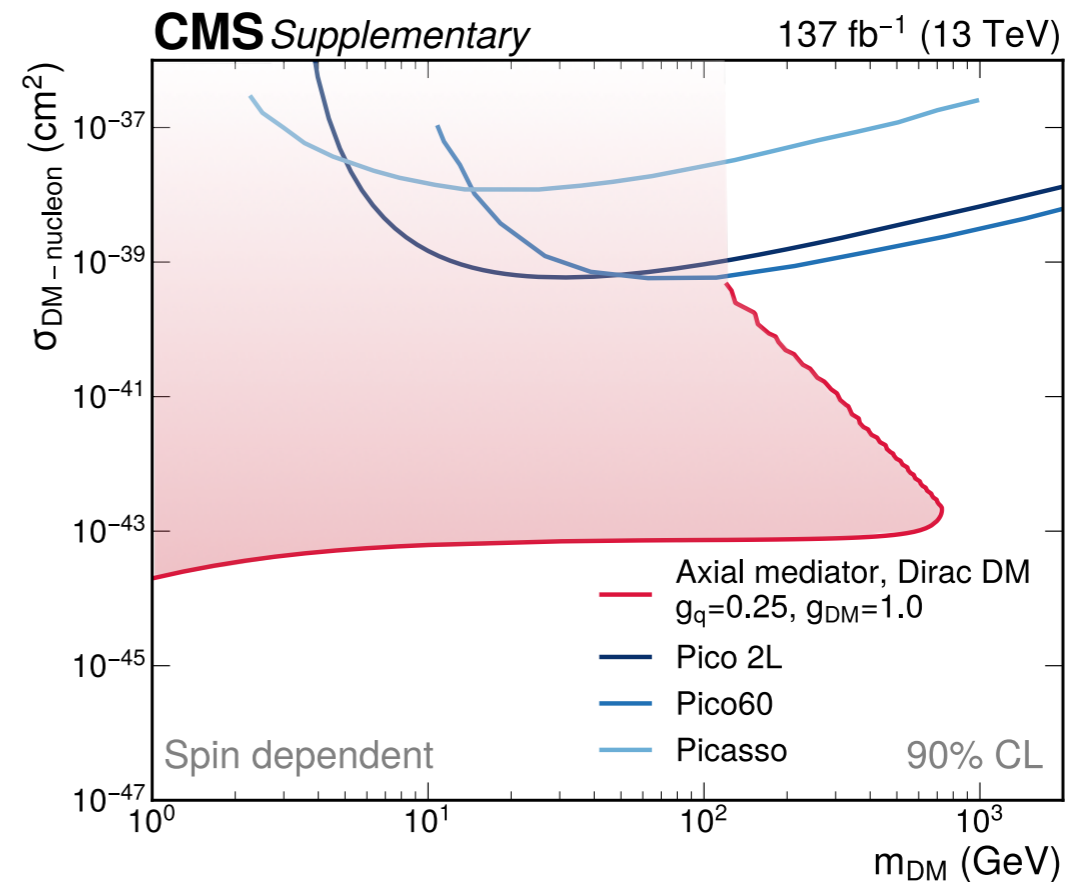
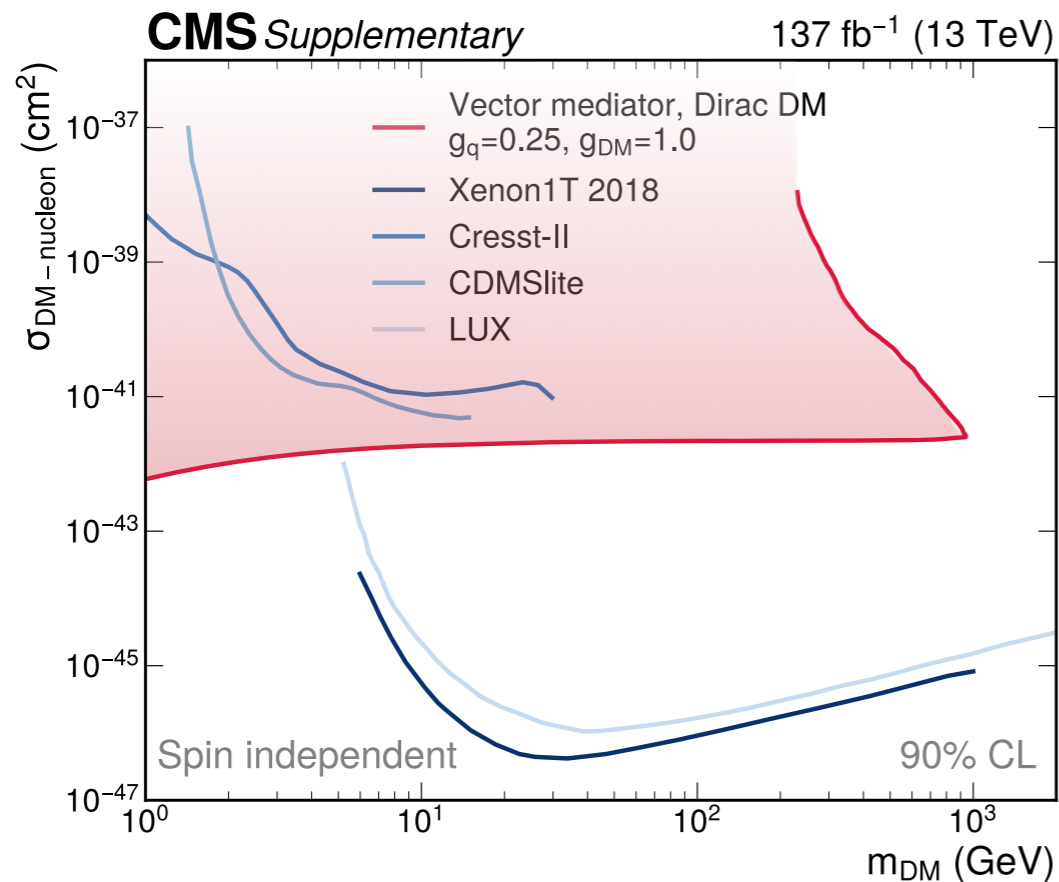
Comparison

- * DM particle non-relativistic: dominant DM-nuclei interactions described by spin-independent and spin-dependent scattering cross section
 - * *vector/scalar* mediator lead to a *SI* interaction
 - * *axial-vector/pseudo-scalar* lead to *SD* interaction
- * comparison is very model dependent
 - * DD bounds may be valid for multiple models, LHC limits hold exclusively for considered simpl. model

comparisons
recommendations
[arXiv:1603.04156]

DM+jet/V: interplay with direct detection

► **3- Results:** lower limits at 90% CL on interaction cross section between DM candidates and the nuclei



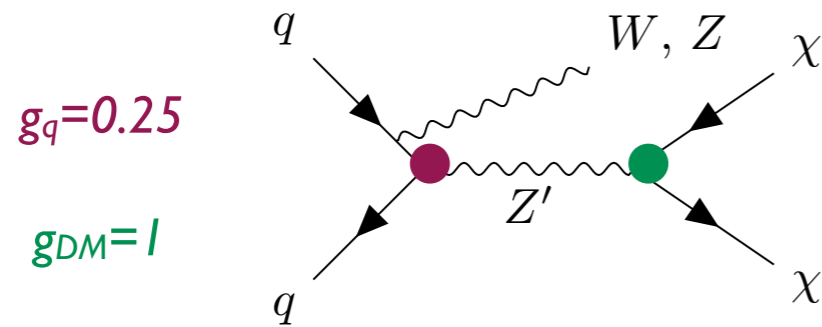
Comparison

- * DM particle non-relativistic: dominant DM-nuclei interactions described by spin-independent and spin-dependent scattering cross section
 - * *vector/scalar* mediator lead to a *SI* interaction
 - * *axial-vector/pseudo-scalar* lead to *SD* interaction
- * comparison is very model dependent
 - * DD bounds may be valid for multiple models, LHC limits hold exclusively for considered simpl. model

comparisons
recommendations
[arXiv:1603.04156]

Simplified models

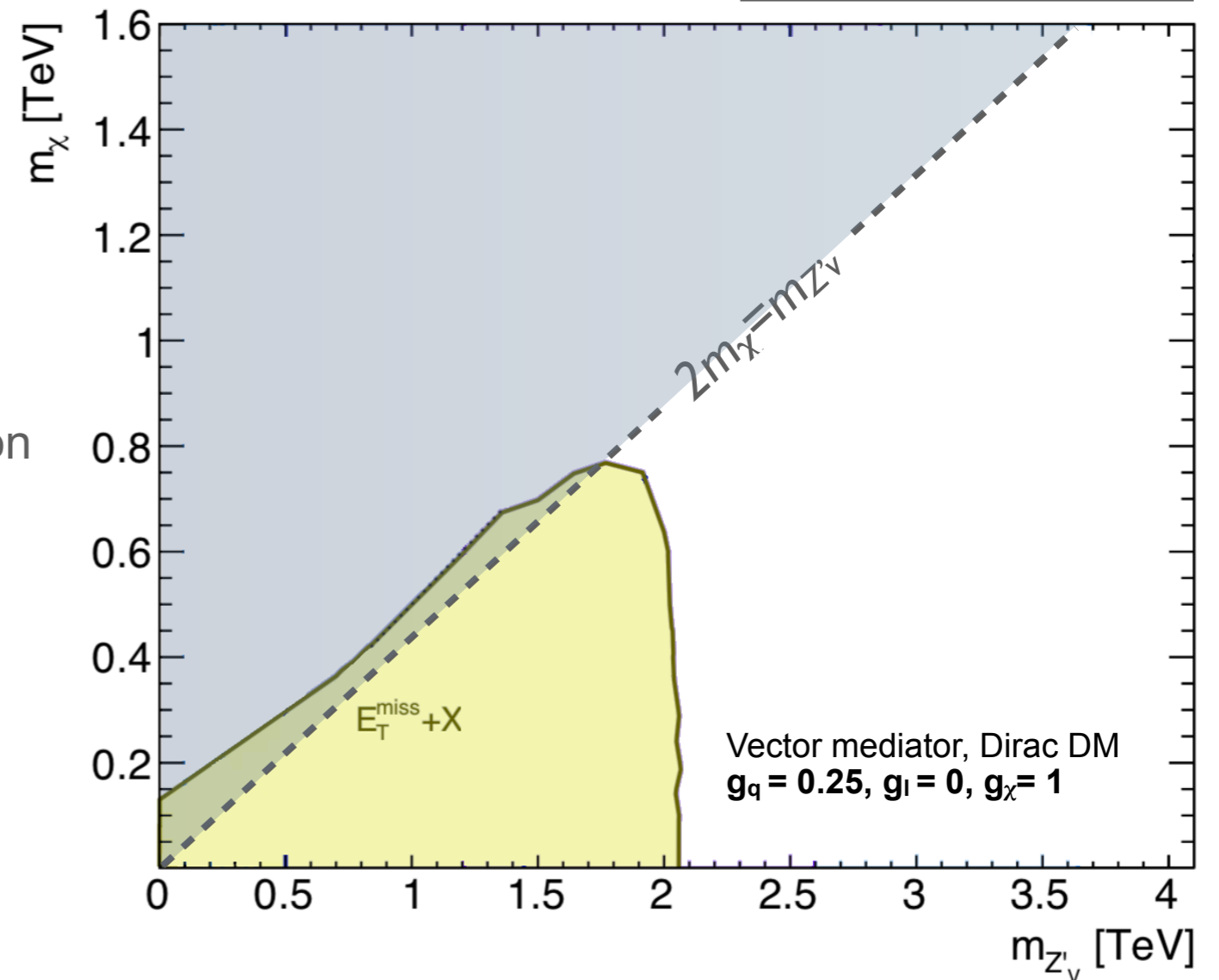
Simplified vector model



► low sensitivity to off-shell region due to strong reduction of production cross-section

► Can we recover the sensitivity?

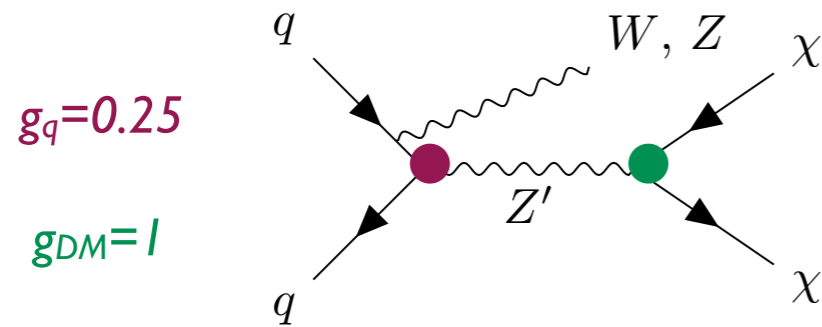
* ATLAS:ATL-PHYS-PUB-2021-006



Simplified models

Simplified vector model

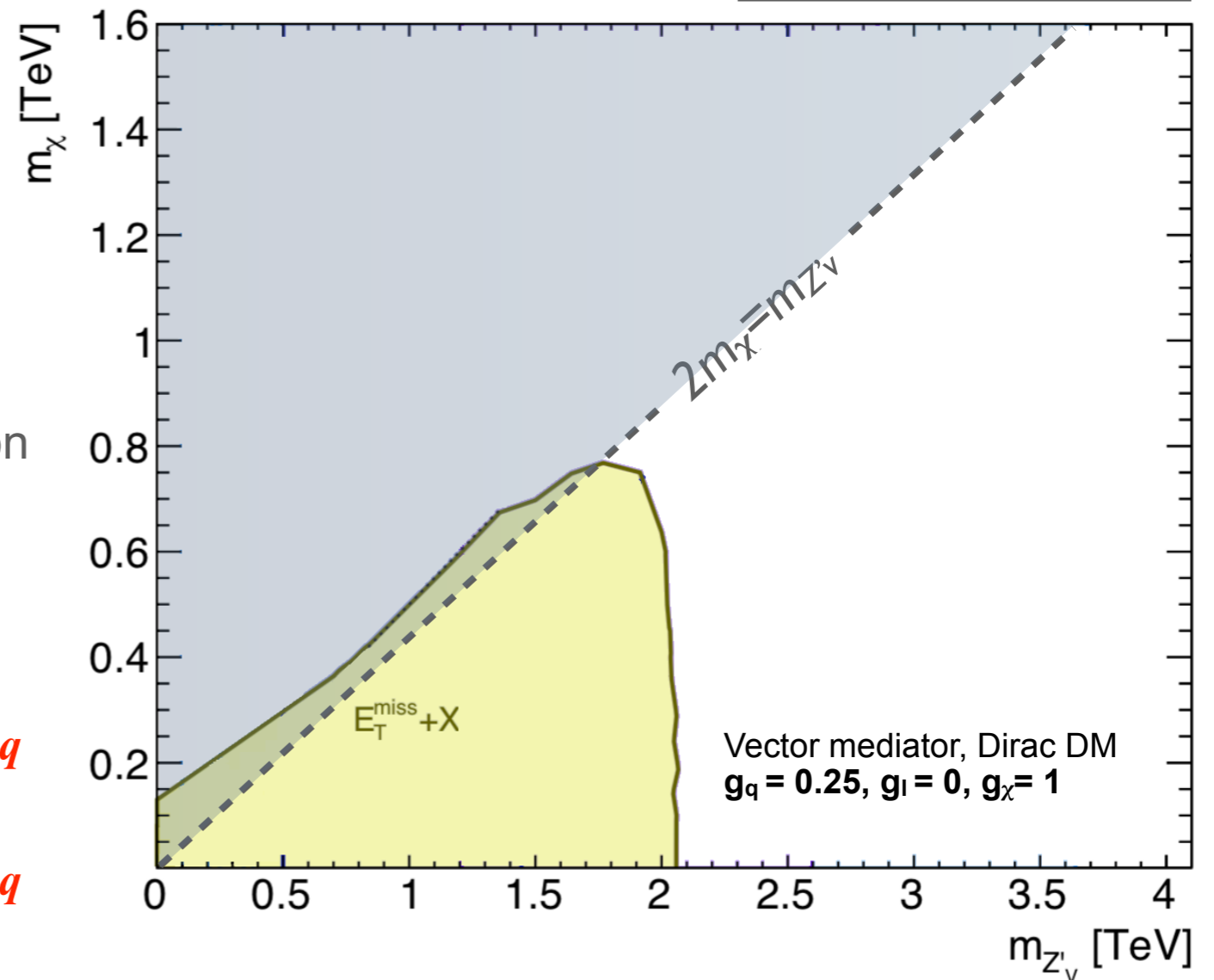
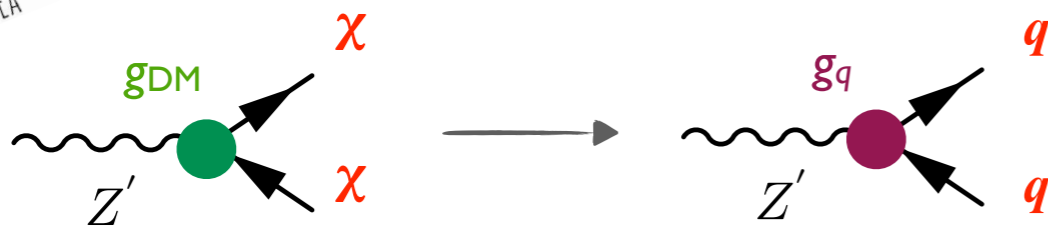
* ATLAS:ATL-PHYS-PUB-2021-006



► low sensitivity to **off-shell region** due to strong reduction of production cross-section

► *Can we recover the sensitivity?*

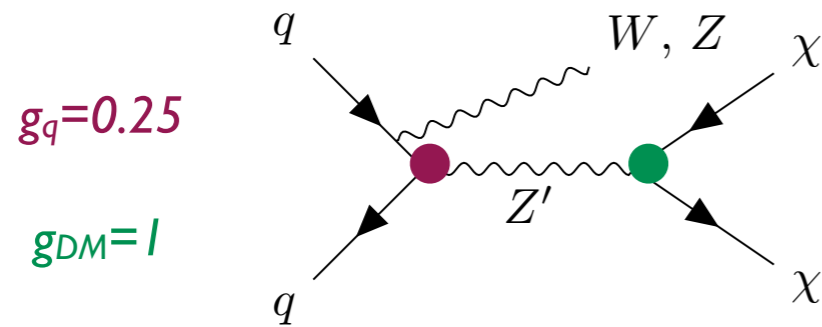
 visible decays



Simplified models

Simplified vector model

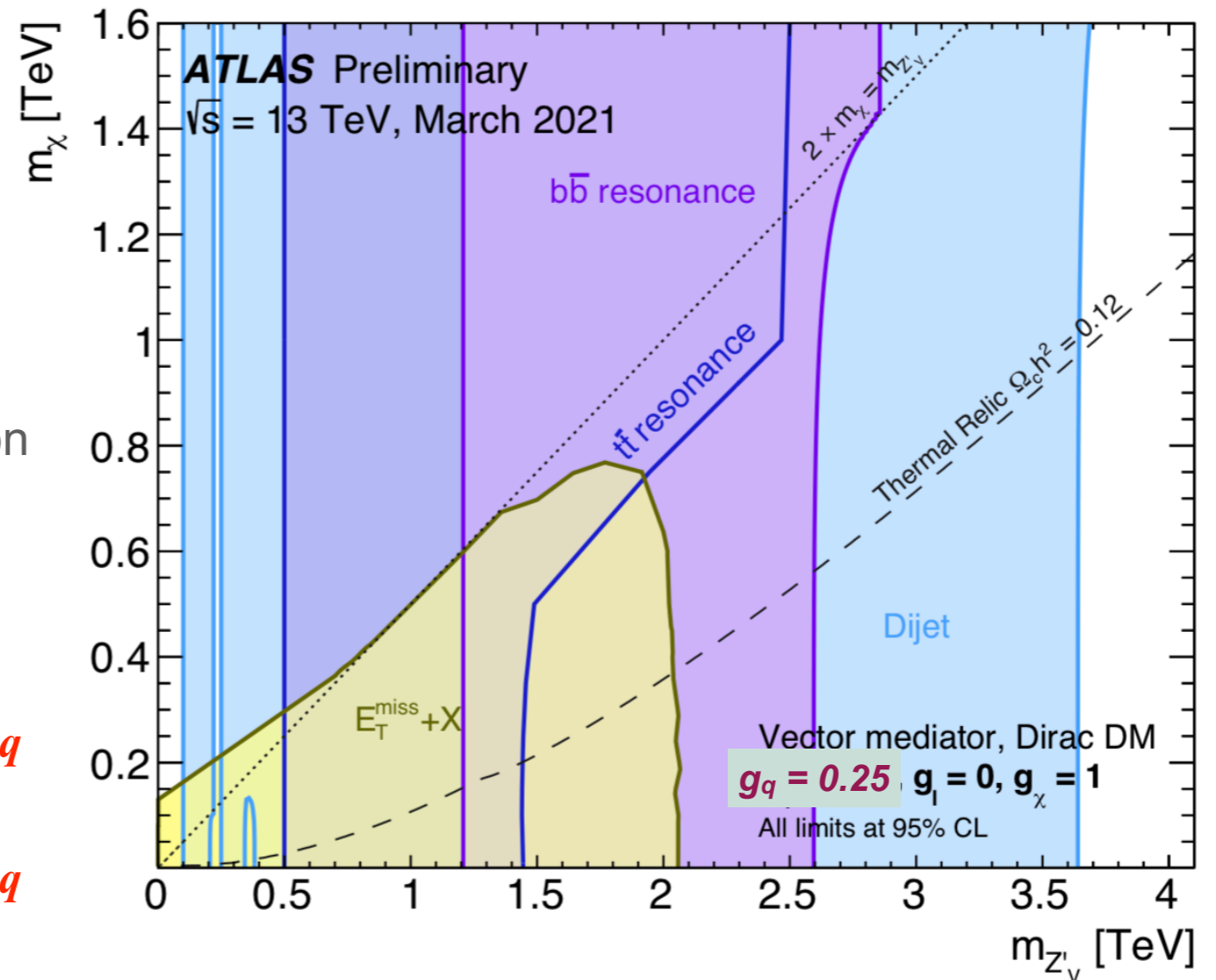
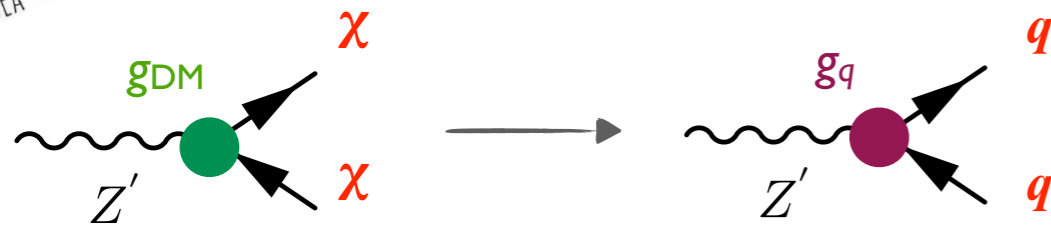
* ATLAS:ATL-PHYS-PUB-2021-006



► low sensitivity to **off-shell region** due to strong reduction of production cross-section

► Can we recover the sensitivity?

IDEA visible decays



* mediator

$g_q =$

- narrow resonance
- wide resonance

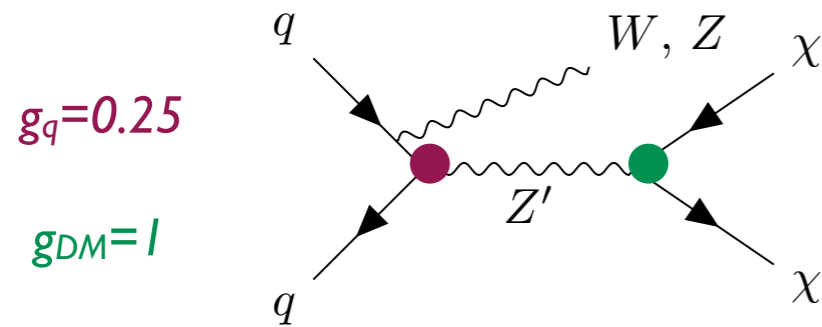
Did we exclude already everything?

Dijet	$t\bar{t}$ resonance	$E_T^{\text{miss}} + X$
Dijet, 139 fb ⁻¹ JHEP 03 (2020) 145	$t\bar{t}$ resonance, 36.1 fb ⁻¹ EPJC 78 (2018) 565	$E_T^{\text{miss}} + \text{jet}$, 139 fb ⁻¹ arXiv:2102.10874
Dijet TLA, 29.3 fb ⁻¹ PRL 121 (2018) 081801	$b\bar{b}$ resonance, 139 fb ⁻¹ JHEP 03 (2020) 145	$E_T^{\text{miss}} + \gamma$, 139 fb ⁻¹ arXiv:2011.05259
Dijet+ISR, 79.8 fb ⁻¹ PLB 795 (2019) 56		$E_T^{\text{miss}} + V(\text{had})$, 36.1 fb ⁻¹ JHEP 10 (2018) 180
Boosted dijet+ISR, 36.1 fb ⁻¹ PLB 788 (2019) 316		$E_T^{\text{miss}} + Z(\text{ll})$, 36.1 fb ⁻¹ PLB 776 (2017) 318
Boosted di- b +ISR, 80.5 fb ⁻¹ ATLAS-CONF-2018-052		

Simplified models

Simplified vector model

* ATLAS:ATL-PHYS-PUB-2021-006

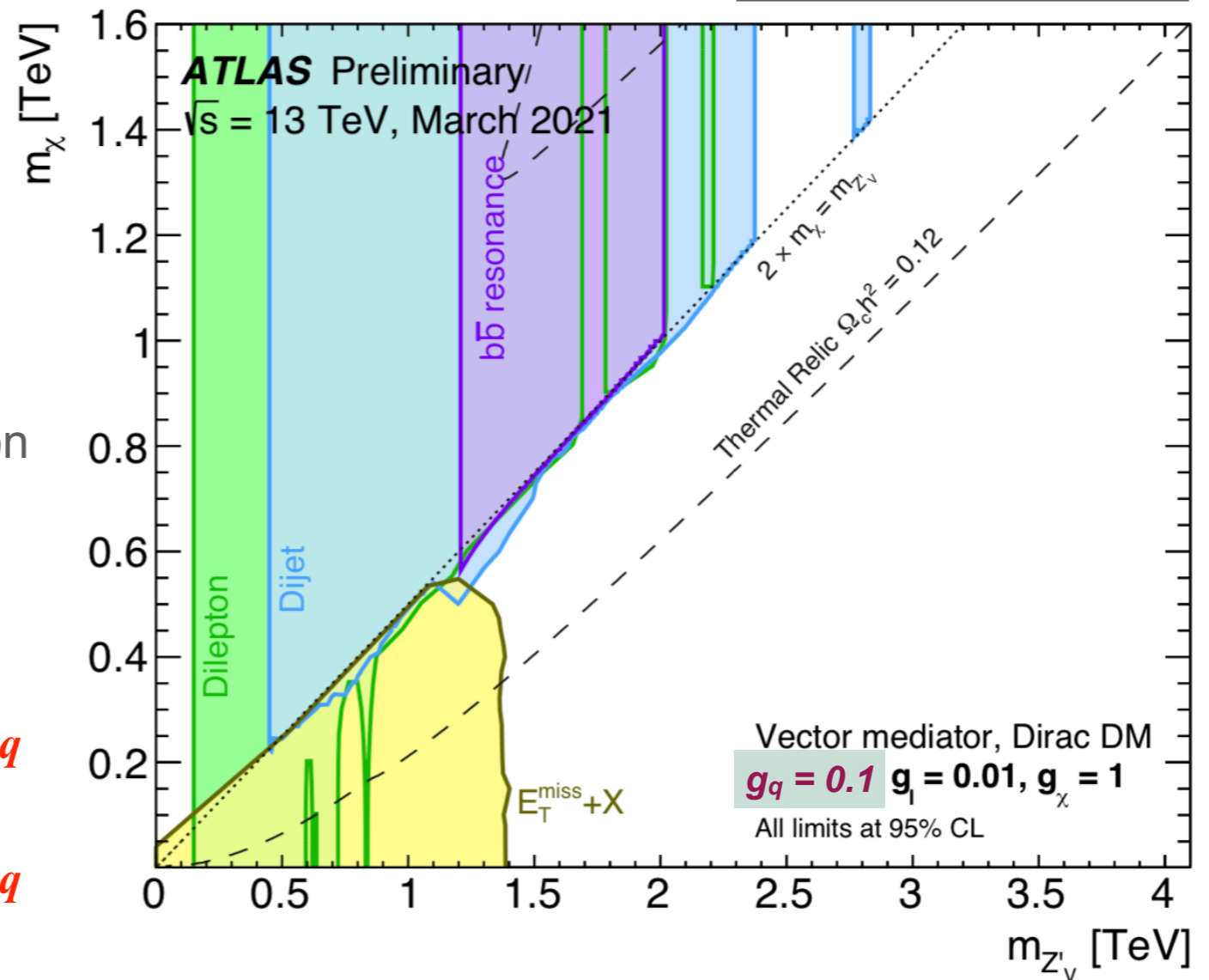
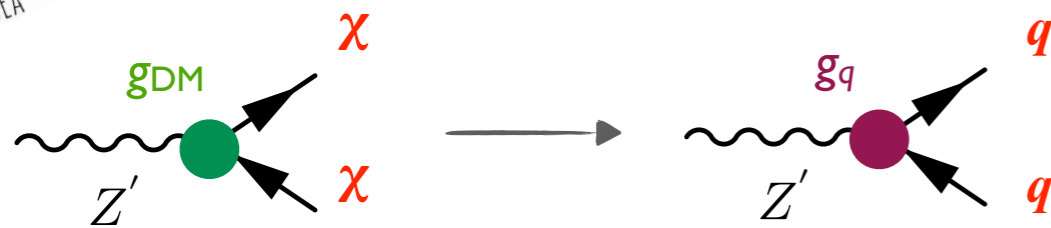


► low sensitivity to **off-shell region** due to strong reduction of production cross-section

► *Can we recover the sensitivity?*



visible decays



■ Dilepton 139 fb ⁻¹ PLB 796 (2019) 68	■ $b\bar{b}$ resonance 139 fb ⁻¹ JHEP 03 (2020) 145
■ Dijet Dijet, 139 fb ⁻¹ JHEP 03 (2020) 145 Dijet TLA, 29.3 fb ⁻¹ PRL 121 (2018) 081801	■ $E_T^{\text{miss}} + X$ $E_T^{\text{miss}} + \gamma$, 139 fb ⁻¹ arXiv:2011.05259 $E_T^{\text{miss}} + \text{jet}$, 139 fb ⁻¹ arXiv:2102.10874

* mediator
- narrow resonance
- wide resonance
* interplay changes

$g_q =$

Did we exclude
already
everything? NO!

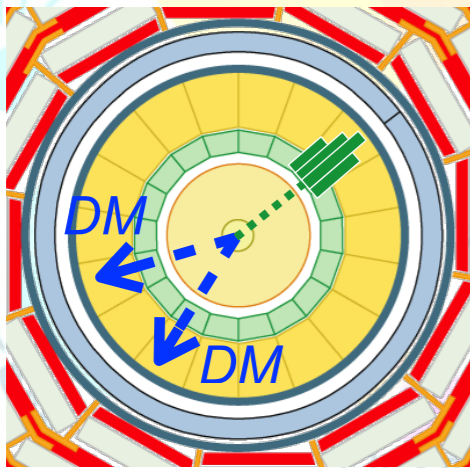
Simplified models ...

... many more mono-X

Reminder:

vector	axial-vector
$g_q \sum_q V_\mu \bar{q} \gamma^\mu q$	$g_q \sum_q A_\mu \bar{q} \gamma^\mu \gamma^5 q$
* choose X to increase xsec or bkg rejection	
scalar	pseudoscalar
$g_q \frac{\phi}{\sqrt{2}} \sum_f y_f \bar{f} f$	$g_q \frac{iA}{\sqrt{2}} \sum_f y_f \bar{f} \gamma^5 f$
* choose X to exploit coupling \propto to quark mass (or increase xsec)	

mono- γ



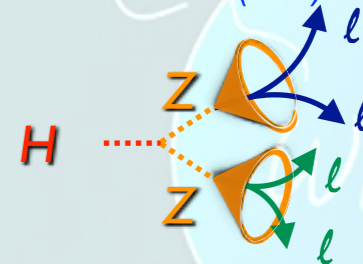
DM+tt



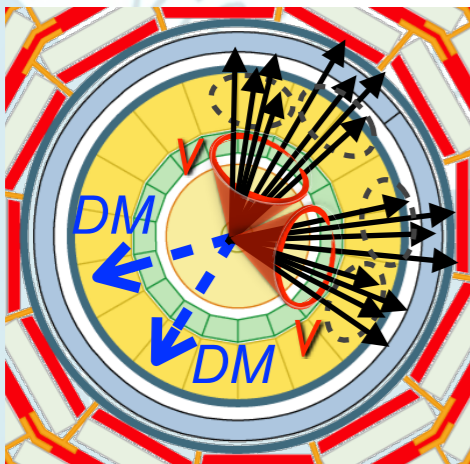
mono-H(WW)



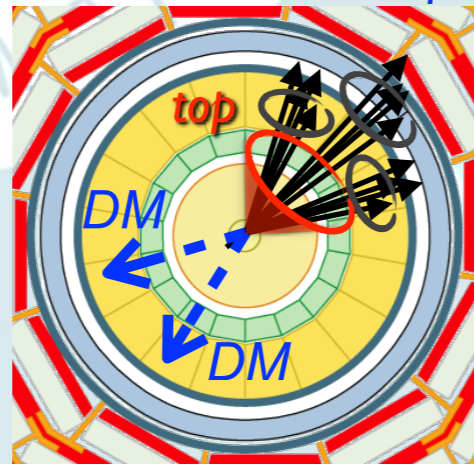
mono-H(ZZ)



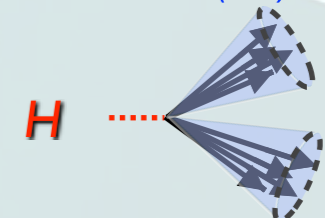
mono-VV(=WW,ZZ)



DM+top: t/tW-channel



mono-H(bb)



Simplified models: Higgs boson portal DM

Reminder:

- ▶ Higgs decay branching fractions not yet sufficiently constrained
 - in SM, $H \rightarrow \text{inv} \sim 0.1\%$
 - direct coupling H-DM will enhance H invisible decays

Higgs

DM

DM

Wimp

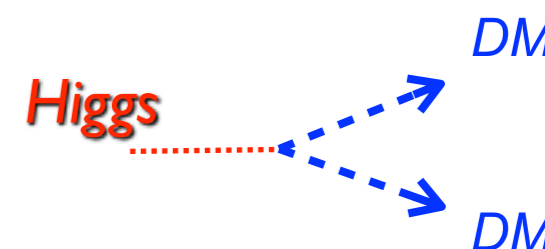
Higgs: a portal to the invisible?

DM-SM interactions mediated by Higgs boson

- direct coupling to DM enhance H invisible decays (SM $\sim 0.1\%$)

Higgs production as in SM

- gluon fusion (MET+j)
- associated VH (MET+v)
- * vector-boson fusion (MET+2jets)



1 - Selection:

- 2 jets (large $|\Delta\eta_{jj}|$, small $|\Delta\Phi_{jj}|$), MET > 180-250 GeV

2- Bkg:

- V+jets main bkg from CRs



Experimental challenges

- * precise estimation of bkg m_{jj} shape distribution, signal as excess of events at large m_{jj}
- * excellent calorimetry in forward region to measure jets

Higgs: a portal to the invisible?

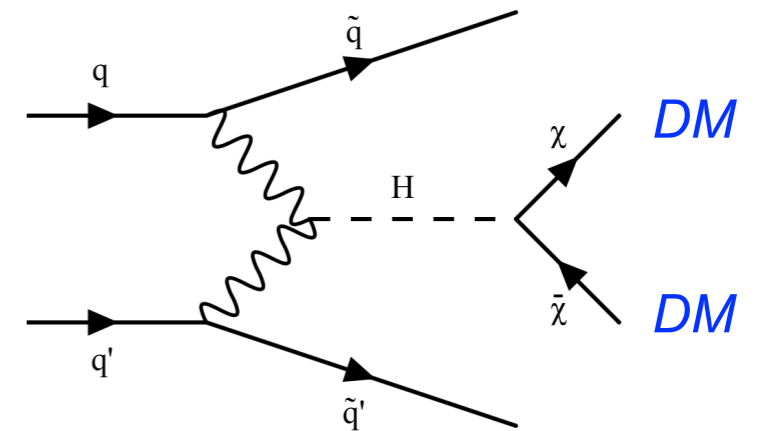
DM-SM interactions mediated by Higgs boson

- direct coupling to DM enhance H invisible decays (SM $\sim 0.1\%$)

Higgs production as in SM

- gluon fusion ($MET+j$)
- associated VH ($MET+v$)

* vector-boson fusion ($MET+2jets$)

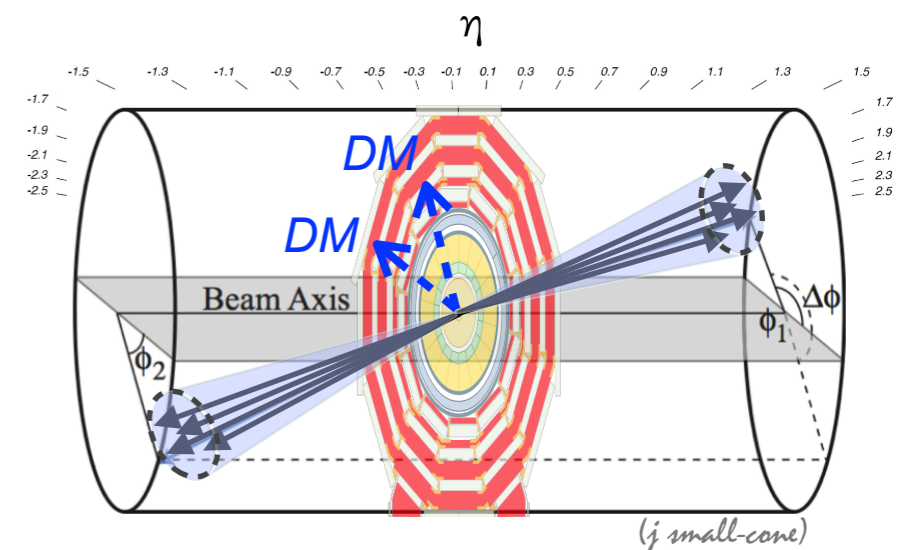


1- Selection:

- 2 jets (large $|\Delta\eta_{jj}|$, small $|\Delta\Phi_{jj}|$), $MET > 250$ GeV

2- Bkg:

- **V+jets** main bkg from CRs



Experimental challenges

- * precise estimation of bkg m_{jj} shape distribution, signal as excess of events at large m_{jj}
- * excellent calorimetry in forward region to measure jets

Higgs: a portal to the invisible?

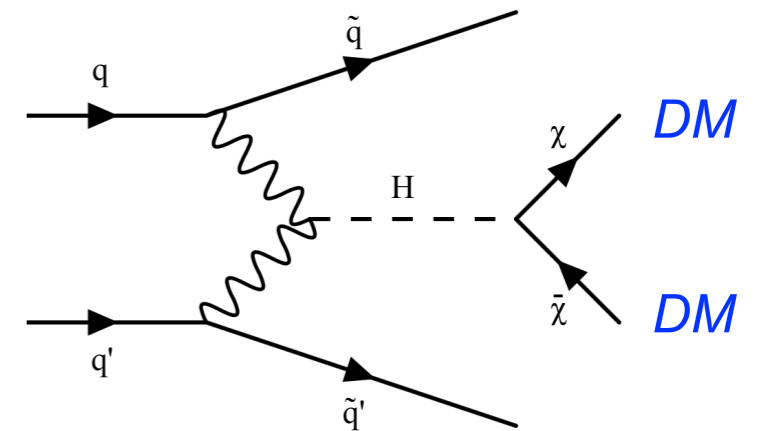
DM-SM interactions mediated by Higgs boson

- direct coupling to DM enhance H invisible decays (SM $\sim 0.1\%$)

Higgs production as in SM

- gluon fusion ($MET+j$)
- associated VH ($MET+v$)

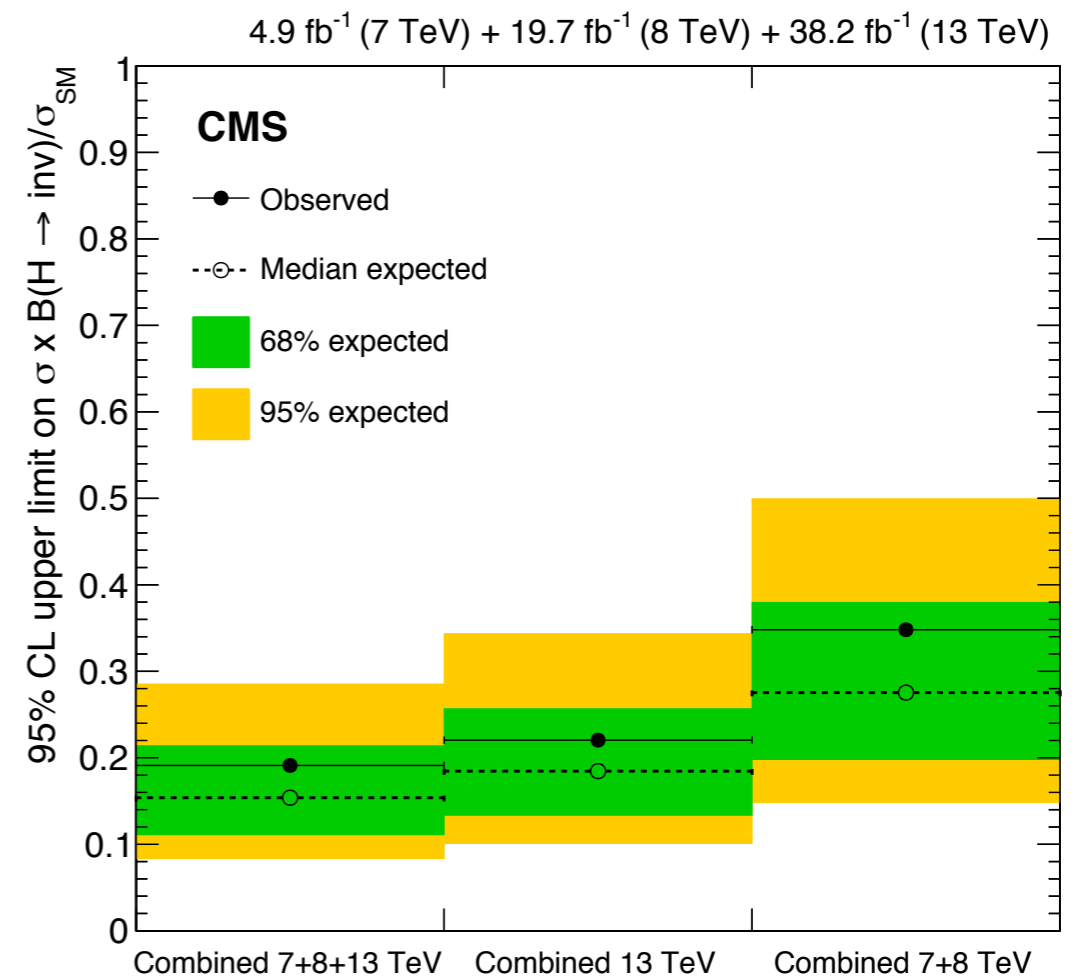
* vector-boson fusion ($MET+2jets$)



3- Results:

- combined fit of SRs and CRs to m_{jj}
- upper limits on cross section \times probability to decay to invisible particles (branching ratio - BR)

$$BR(H \rightarrow inv) < 0.19 (0.15) \text{ obs (exp.)}$$



Higgs: a portal to the invisible?

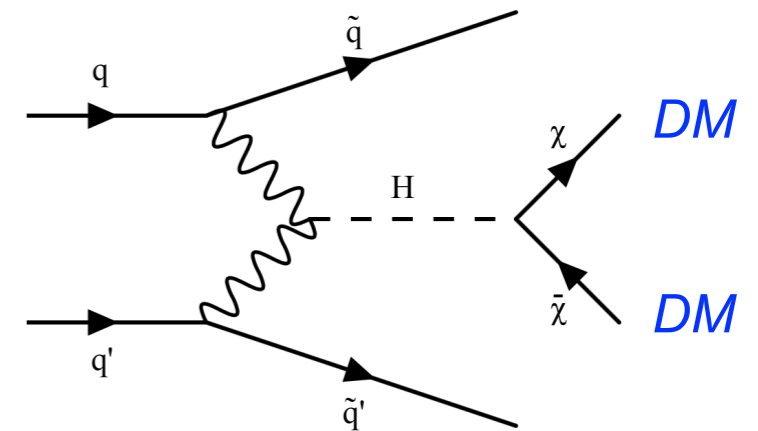
DM-SM interactions mediated by Higgs boson

- direct coupling to DM enhance H invisible decays (SM $\sim 0.1\%$)

Higgs production as in SM

- gluon fusion (MET+j)
- associated VH (MET+v)

* vector-boson fusion (MET+2jets)

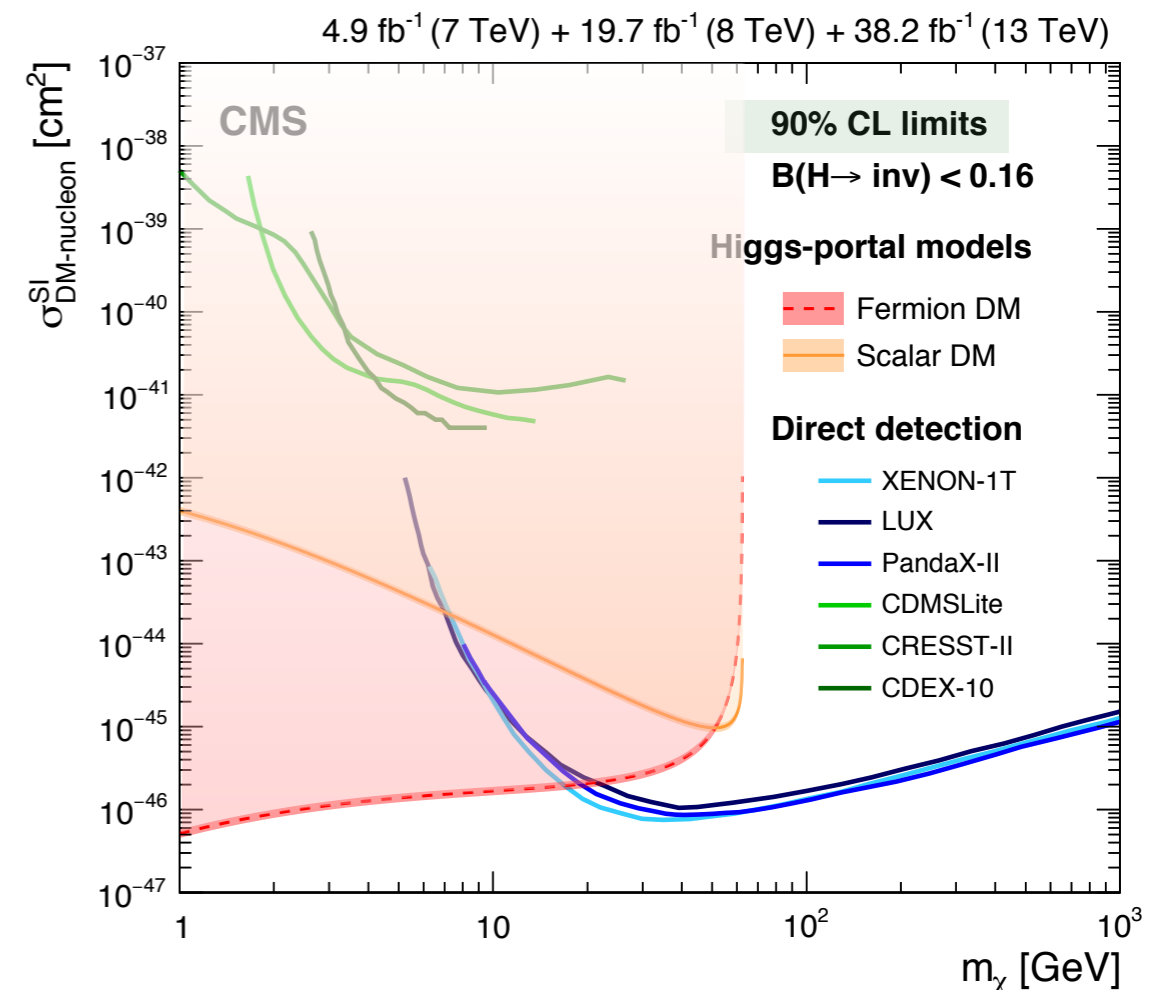


3- Results:

- upper limits on on cross section \times probability to decay to invisible particles (branching ratio - BR)

- translated into a spin-independent DM-nucleon elastic scattering σ_{SI} limit

- * m_{DM} smaller than half of m_{H}
- * interaction between DM and nucleus mediated by H exchange



Summary

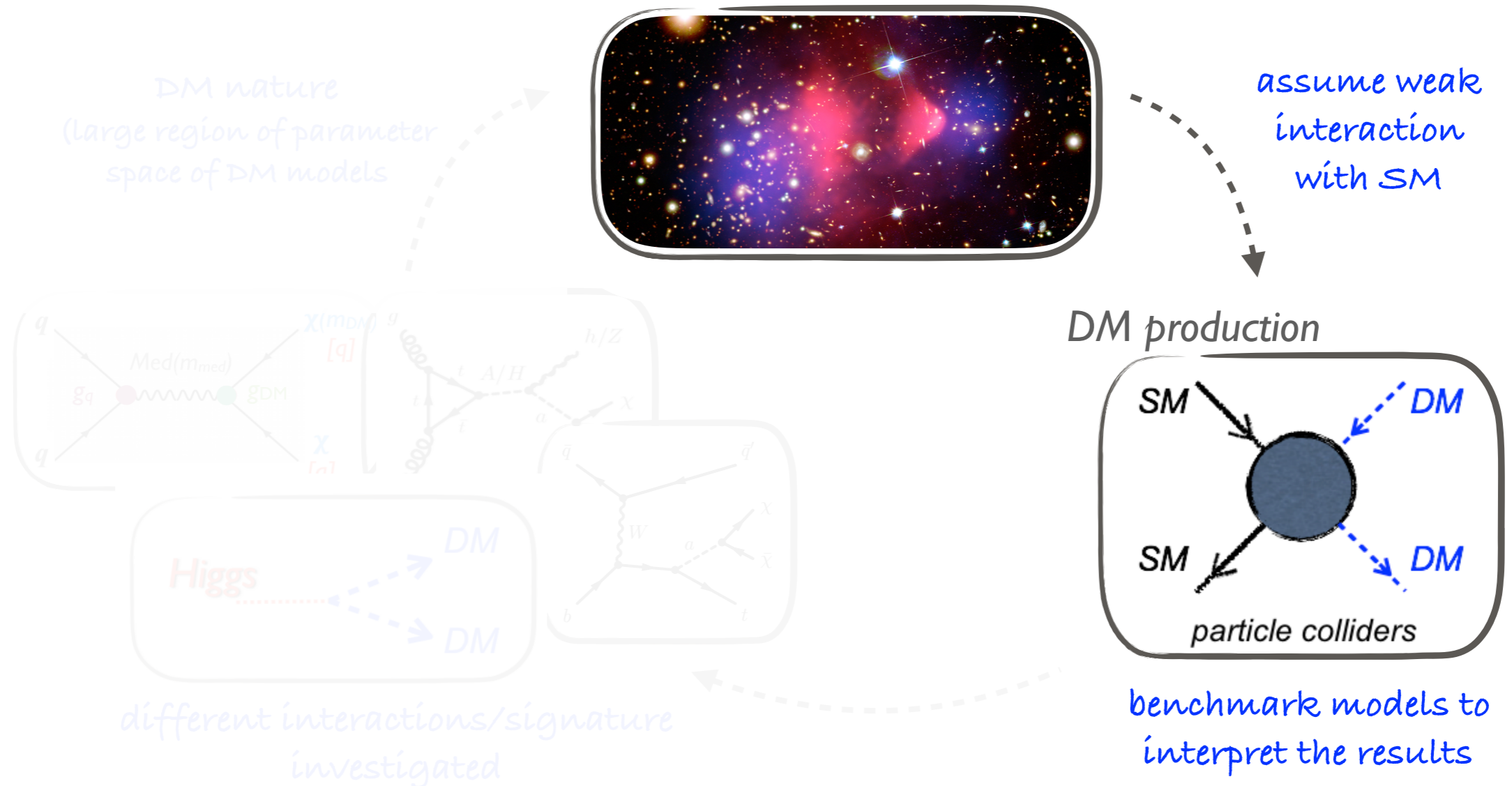


*Rich DM physics program at colliders

- essential complementarity with non-collider searches
- inputs from various signatures (mono- X , resonances, $H \rightarrow \text{inv}$, ...)
- supersymmetry can be a dark matter model too *see more in N. Strobbe's talk*

*Many new results with Run-3 data are expected!

Summary

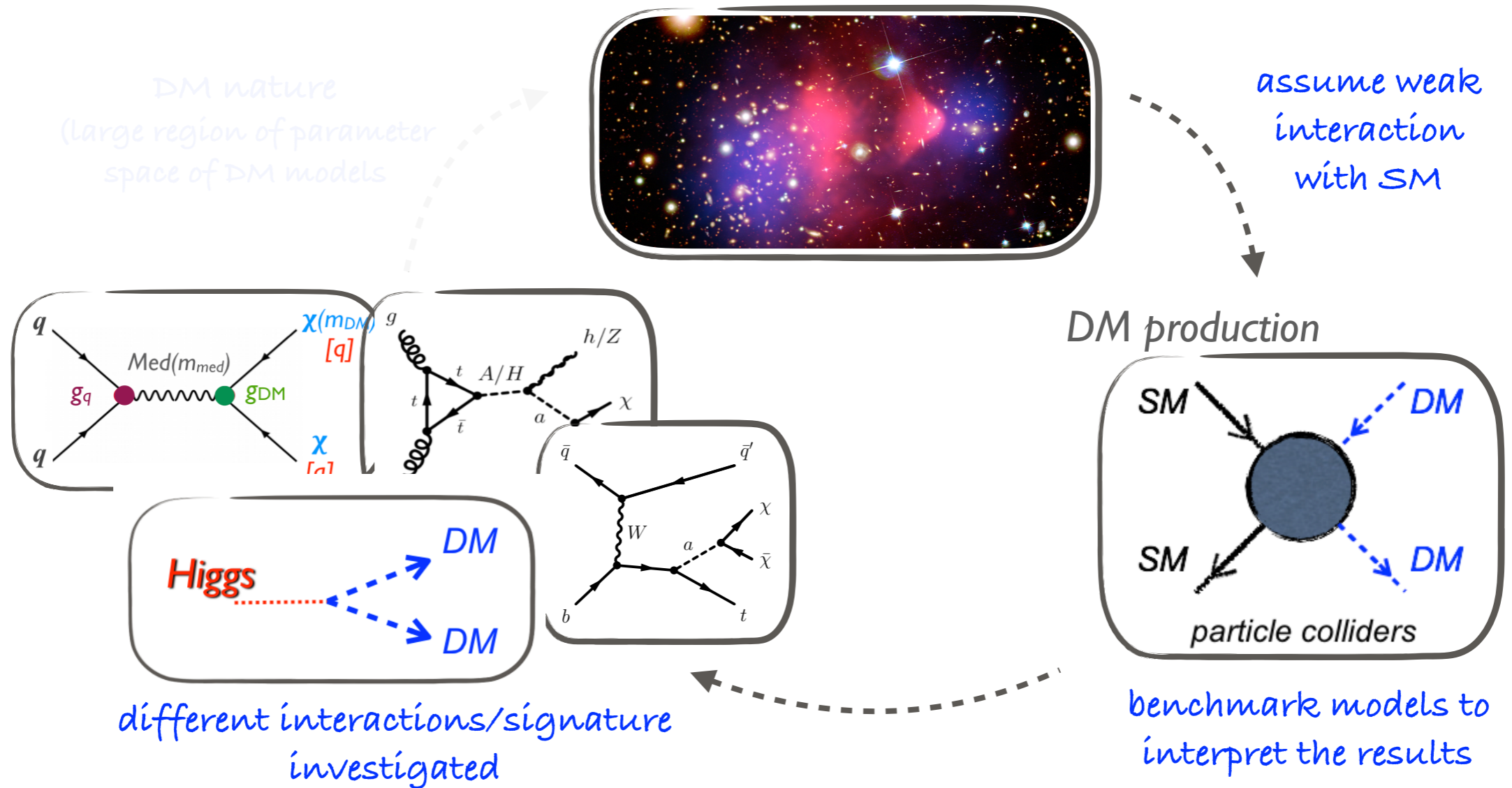


*Rich DM physics program at colliders

- essential complementarity with non-collider searches
- inputs from various signatures (mono- X , resonances, $H \rightarrow \text{inv}$, ...)
- supersymmetry can be a dark matter model too *see more in N. Strobbe's talk*

*Many new results with Run-3 data are expected!

Summary

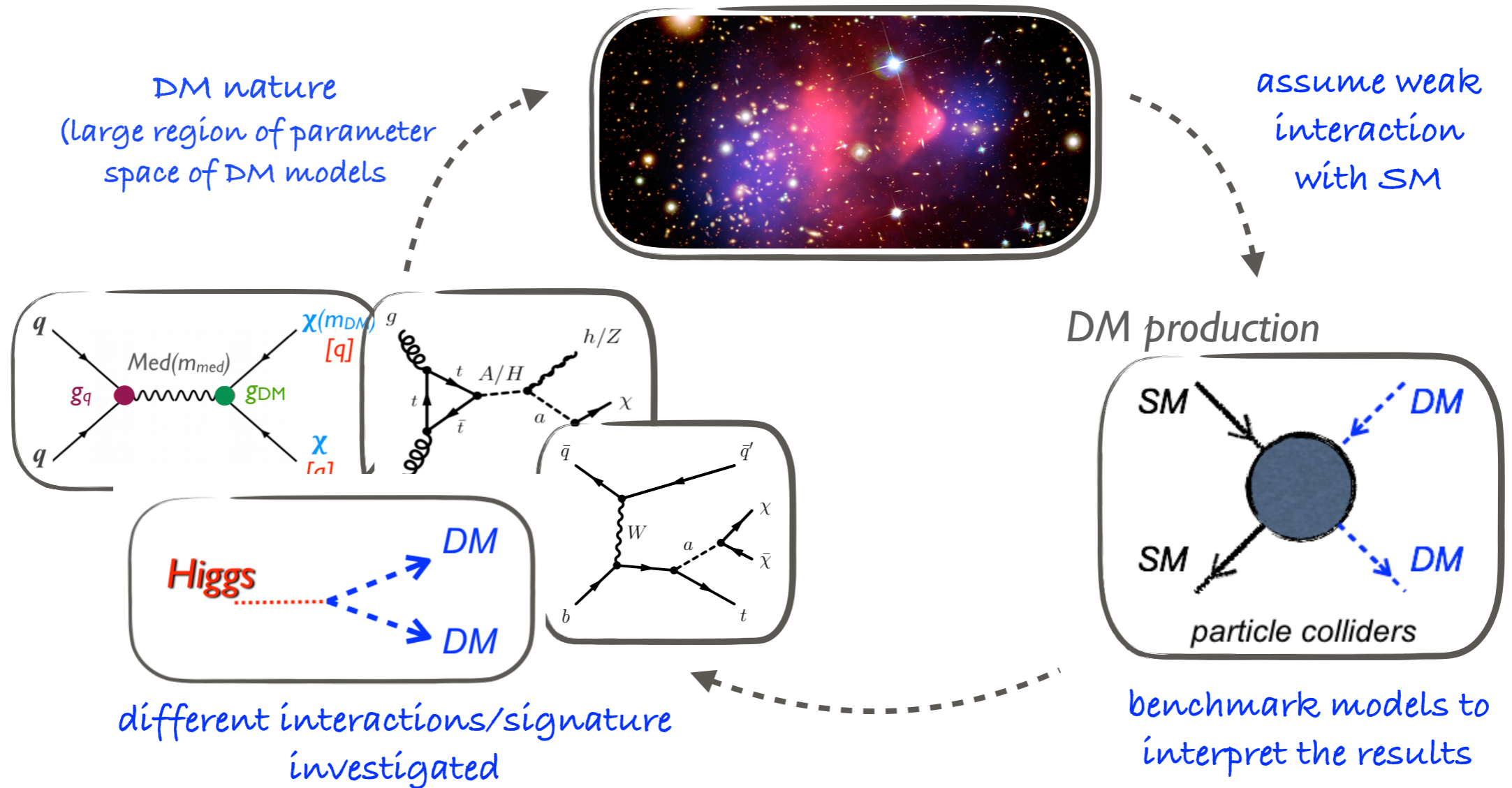


* Rich DM physics program at colliders

- essential complementarity with non-collider searches
- inputs from various signatures (mono- χ , resonances, $H \rightarrow inv$, ...)
- supersymmetry can be a dark matter model too *see more in N. Strobbe's talk*

* Many new results with Run-3 data are expected!

Summary

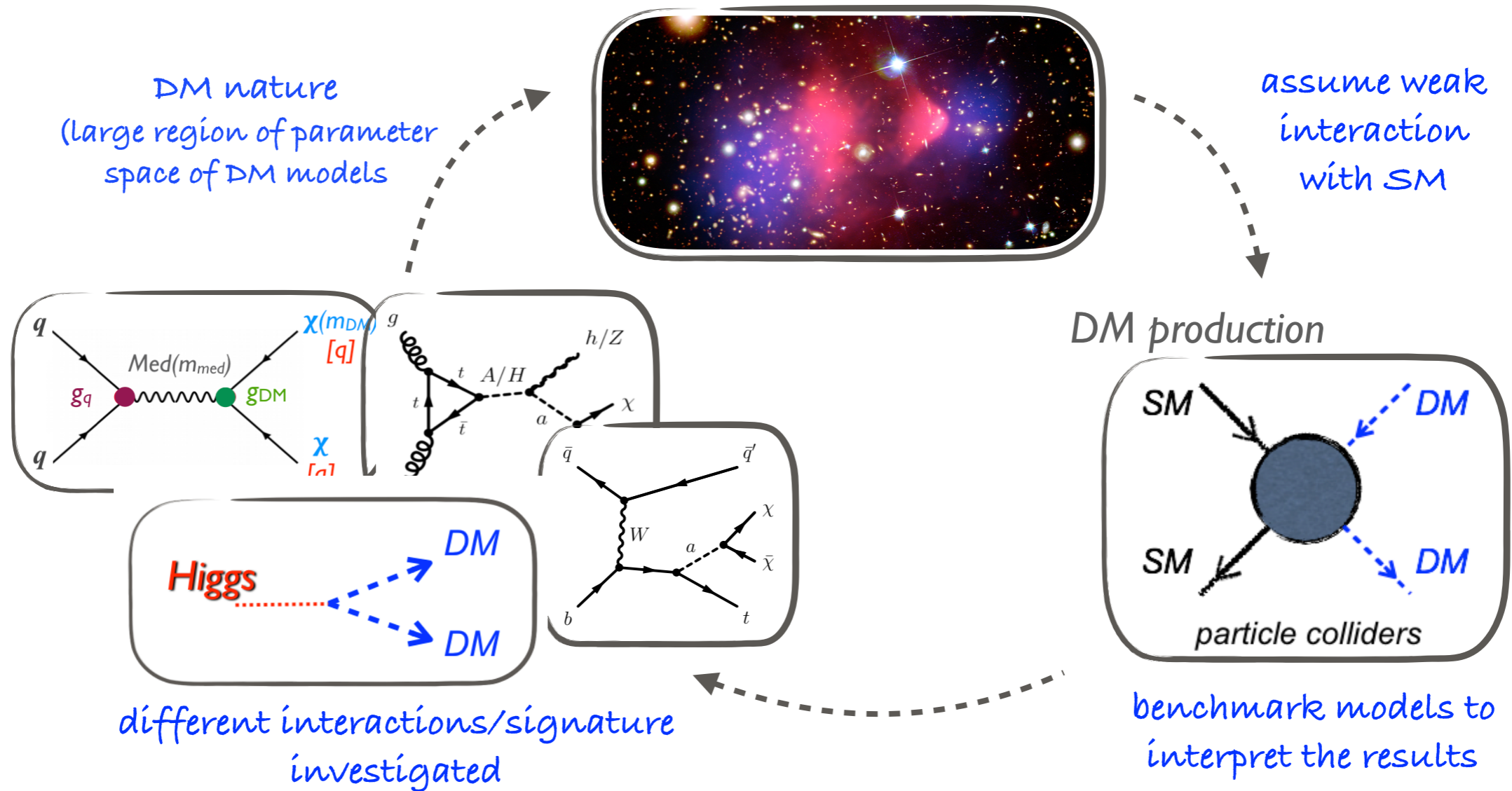


* Rich DM physics program at colliders

- essential complementarity with non-collider searches
- inputs from various signatures (mono- χ , resonances, $H \rightarrow \text{inv}$, ...)
- supersymmetry can be a dark matter model too *see more in N. Strobbe's talk*

* Many new results with Run-3 data are expected!

Summary



* Rich DM physics program at colliders

- essential complementarity with non-collider searches
- inputs from various signatures (mono-X, resonances, $H \rightarrow inv$, ...)
- supersymmetry can be a dark matter model too

* Many new results with Run-3 data are expected!

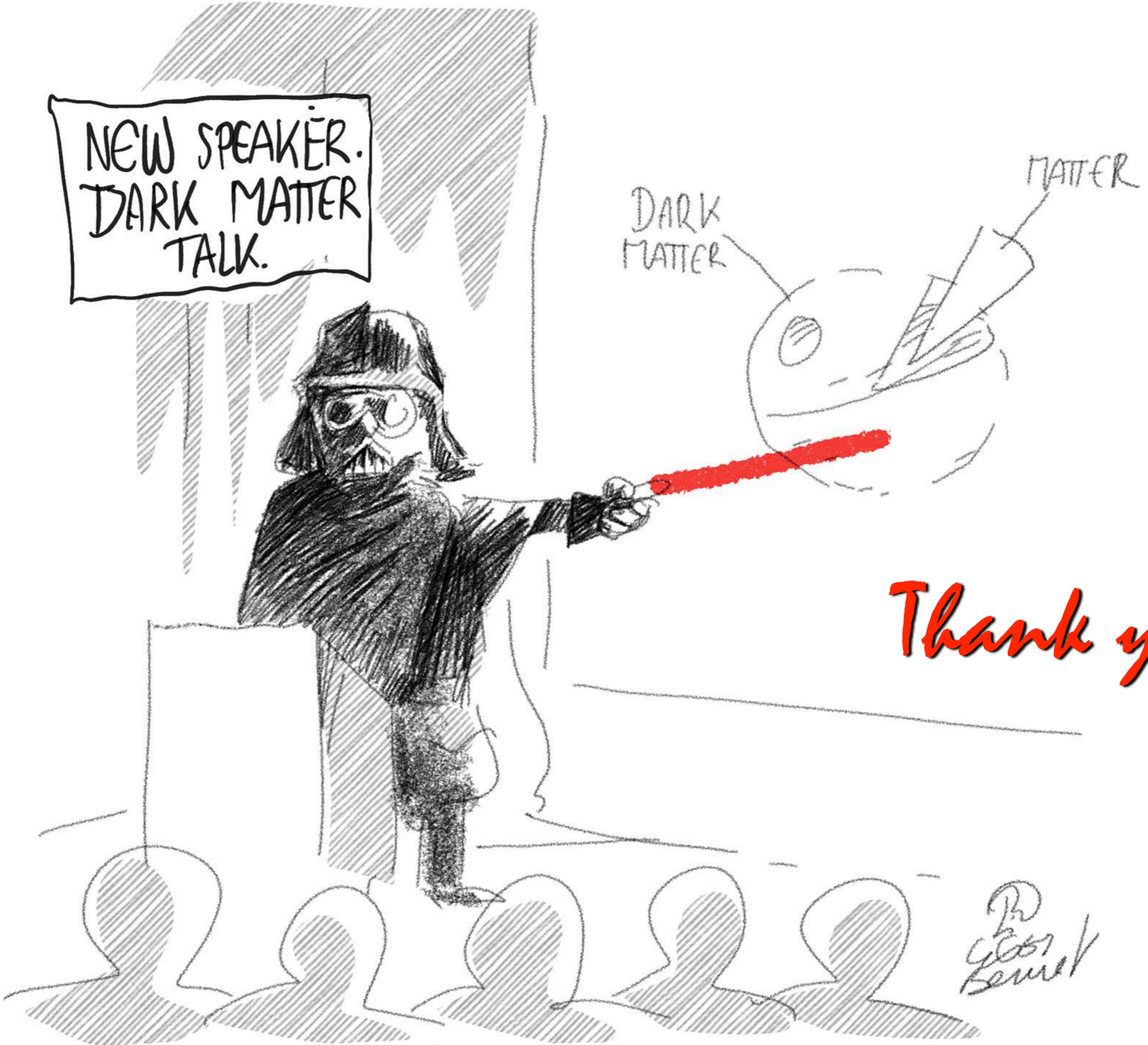
NEW SPEAKER.
DARK MATTER
TALK.

DARK
MATTER

MATTER

Thank you!

*Dr
Gib
Bauer*



DARK MATTER

Backup

PROTON

PHOTON

NEUTRINO

MUON

QUARK

VISIBLE MATTER

NEUTRALINO

AXION

AXION

GRAVITINO

Wimp



Di-lepton high-mass resonances

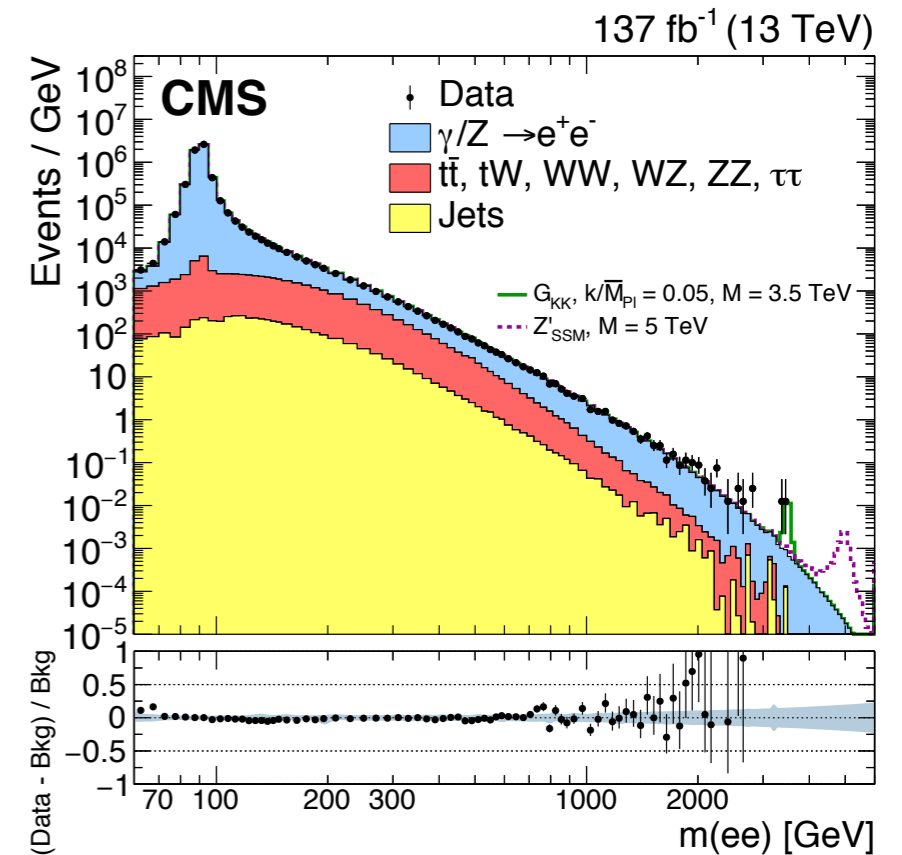
1 - *Selection*: resonance appears as peak wrt SM invariant mass spectrum

- * 2 electrons or 2 opp-sign muons

2 - *Bkg*:

- **Z(II)** main bkg, normalized from CR
- **QCD multi-jet, W+jets** with mis-identified leptons from CR

3 - *Results*: compare SM predictions with data, fit to dilepton invariant mass (systematic unc. included as nuisance parameters)



Di-lepton high-mass resonances

- * lower masses wrt searches for di-jet resonances (lower thresholds for lepton triggers)
- * can investigate very low mass mediators

