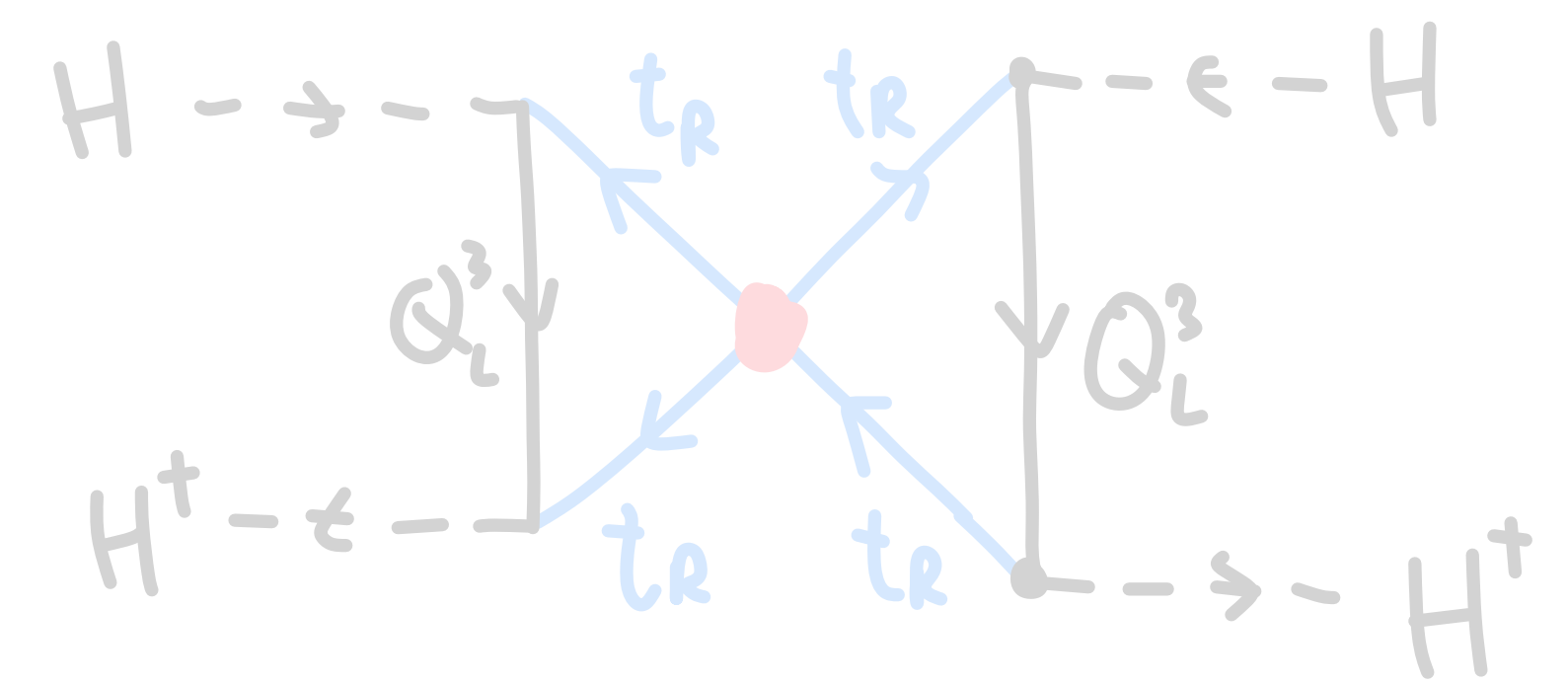
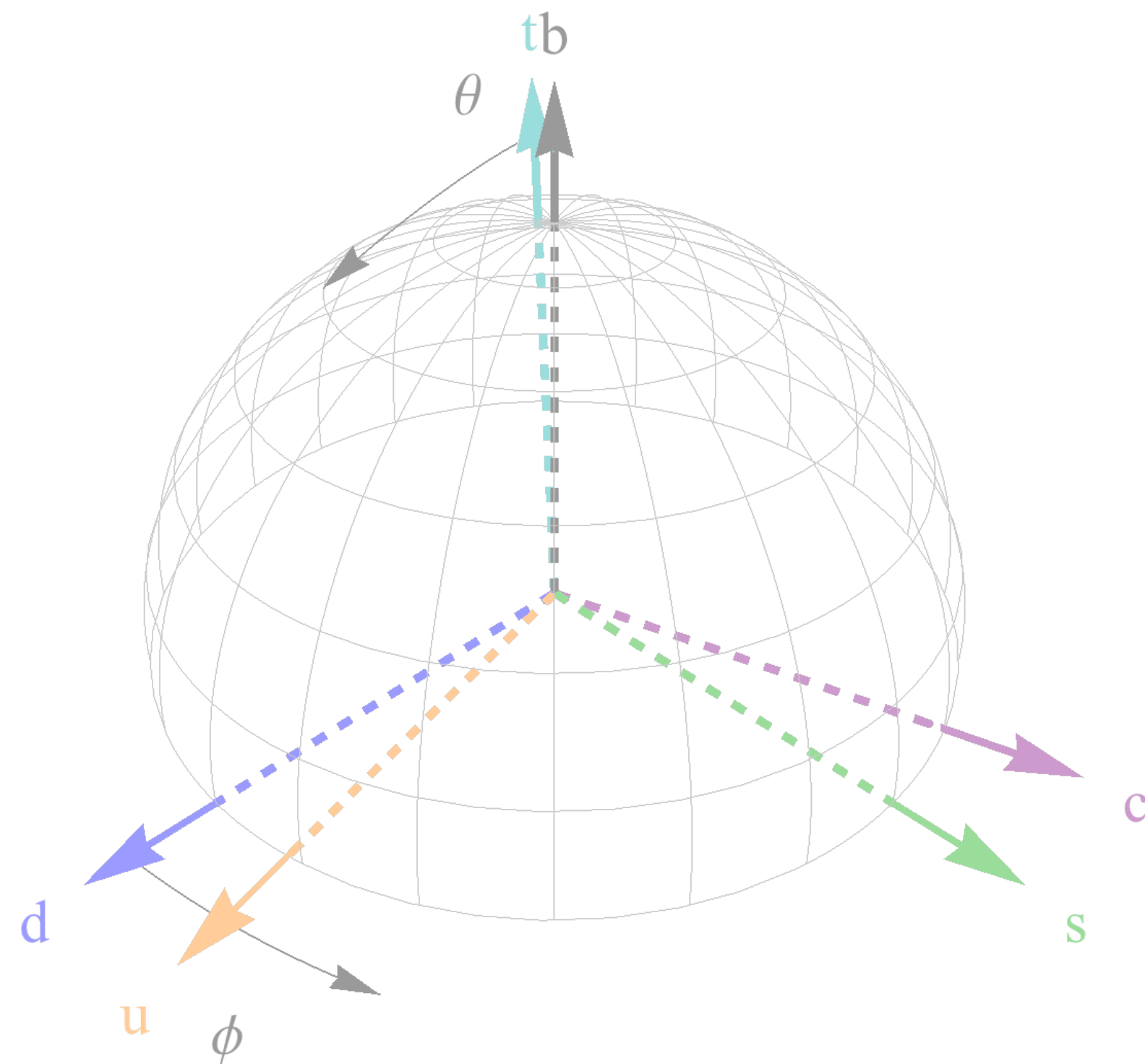


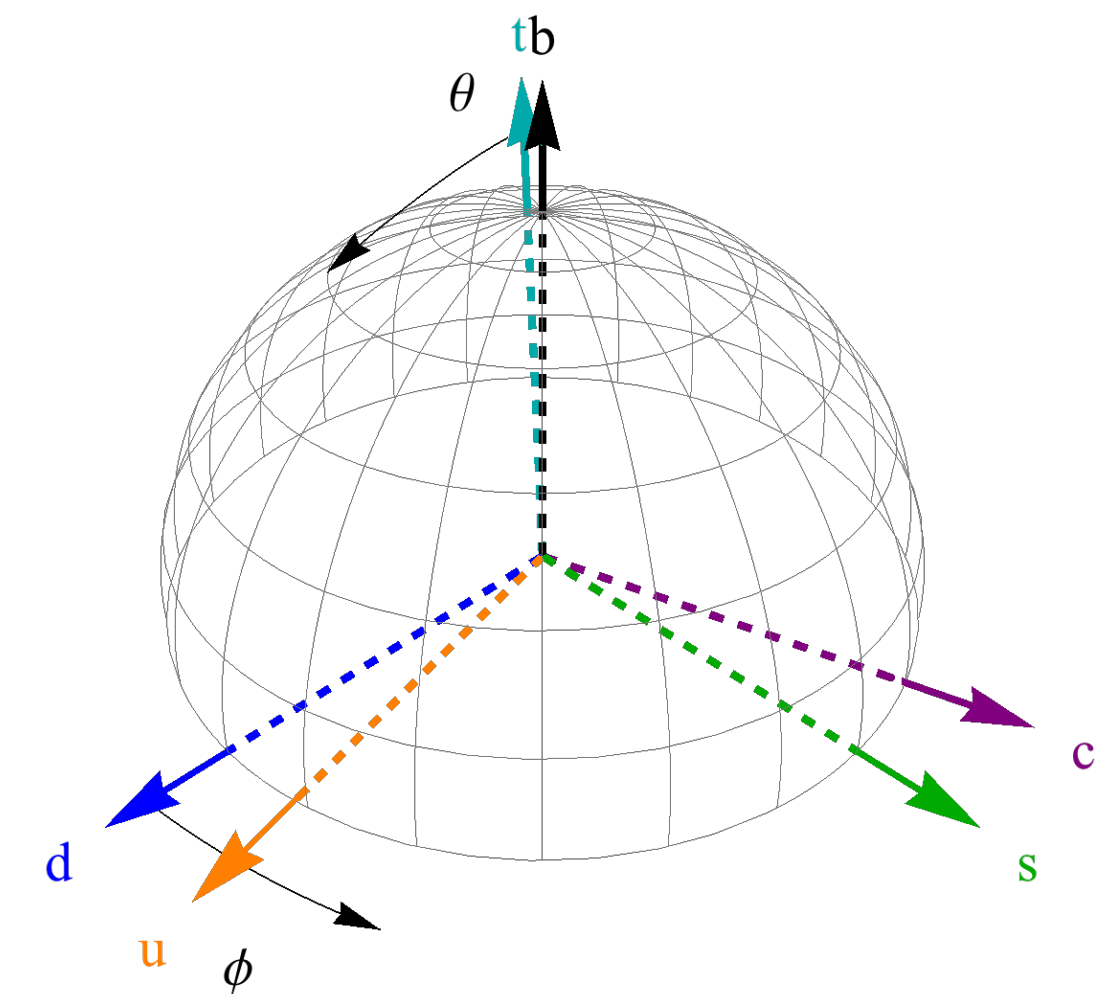
New Physics in Top Operators

David Marzocca



DISCRETE 2024 - 02/12/2024

Outline



1) Introduction: the New Physics flavour problem & EFT approach

2) How much aligned to the third generation should New Physics couplings be?

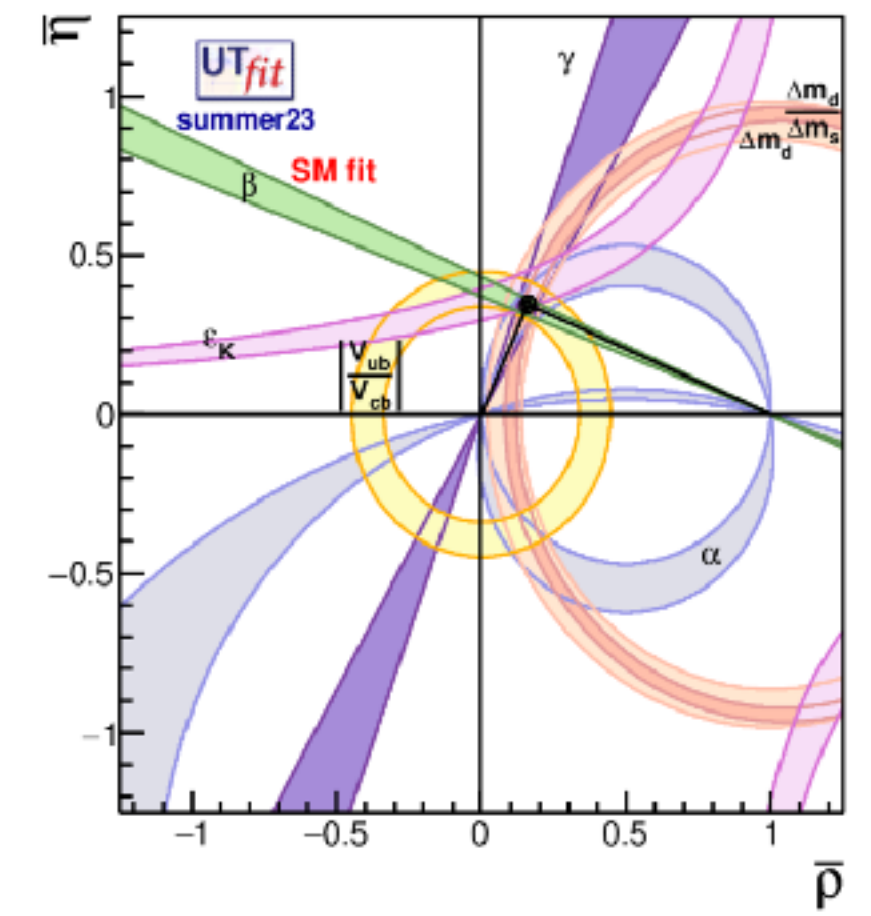
- Take the $B \rightarrow K\nu\nu$ **excess** as a guide to set an overall scale of New Physics
- Consider a **specific flavour structure** that allows to deviate arbitrarily from the third generation, while keeping only a few parameters (unlike a general SMEFT approach).
- Evaluate **quantitatively** the **allowed misalignment** from third generation.

3) What are the constraints on heavy New Physics coupled to the top quark?

Indirect (Flavour + EW) vs. **Direct** (LHC)

The BSM Flavour Problem

Flavour in the **SM** has a **rigid structure**: accidental symmetries and suppression (**FCNC**, **CP violation**, **LFV**, etc).

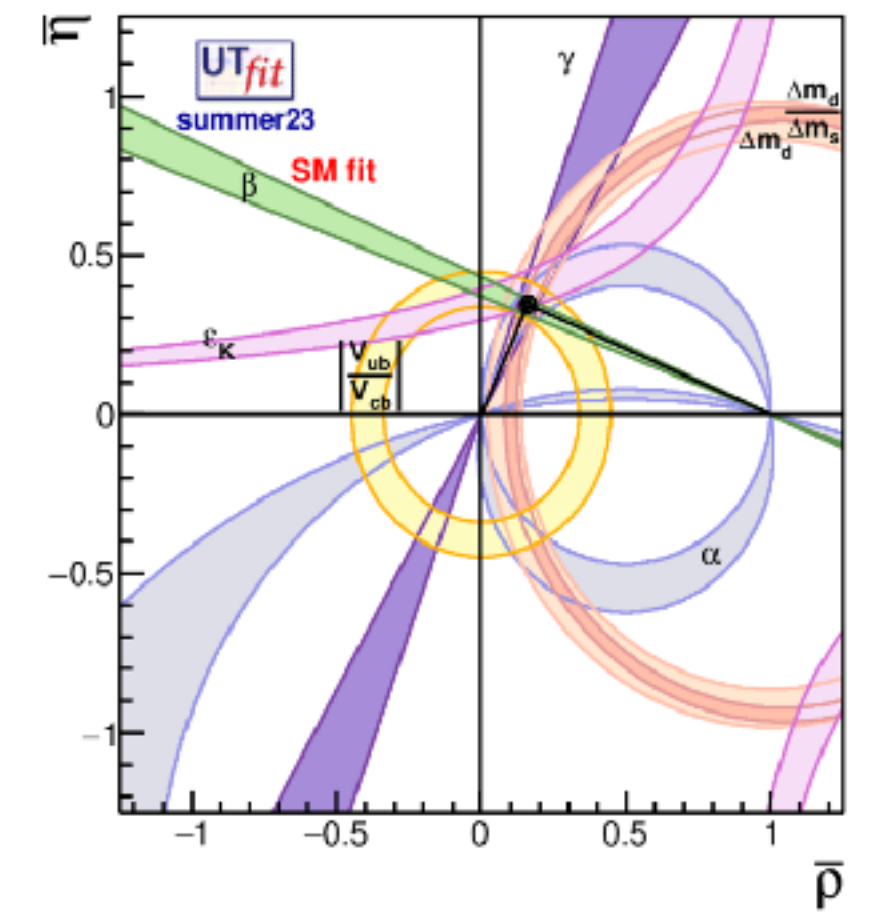
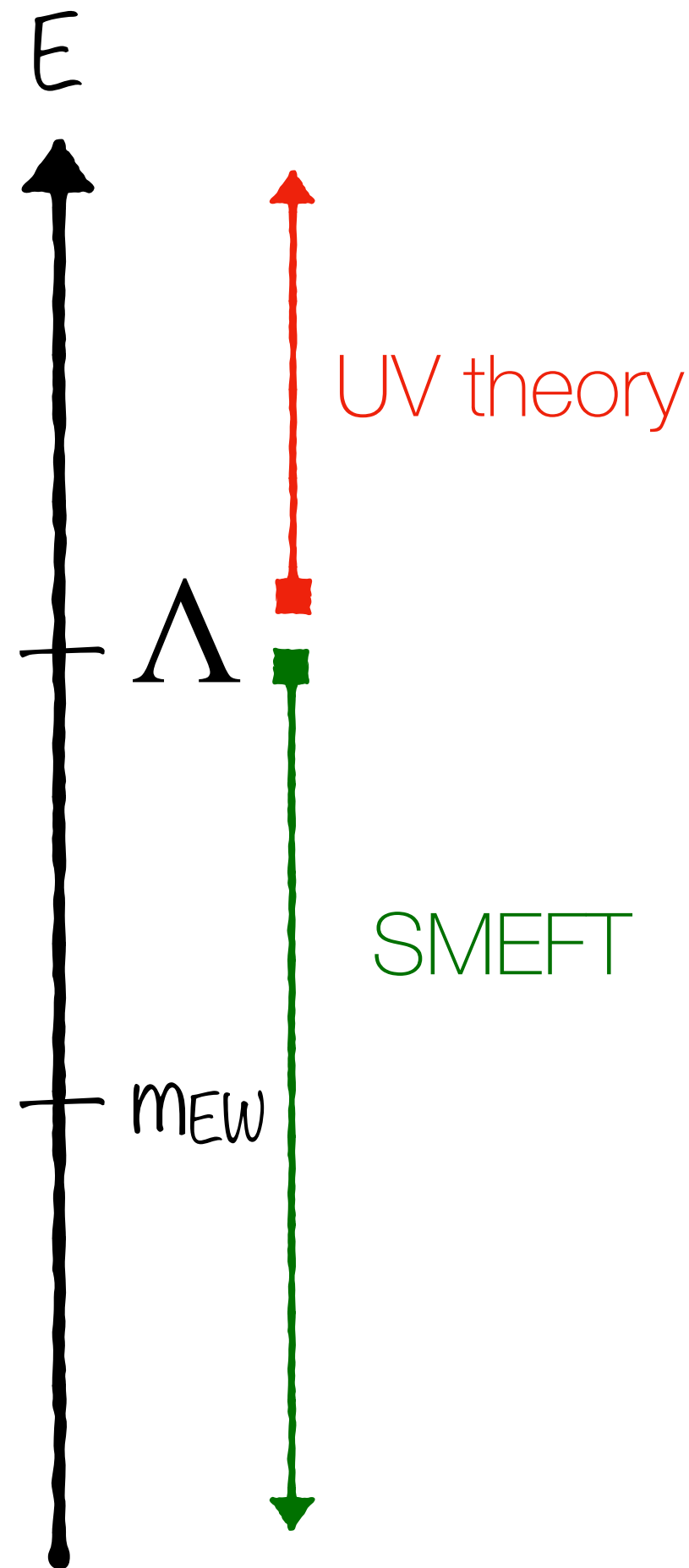


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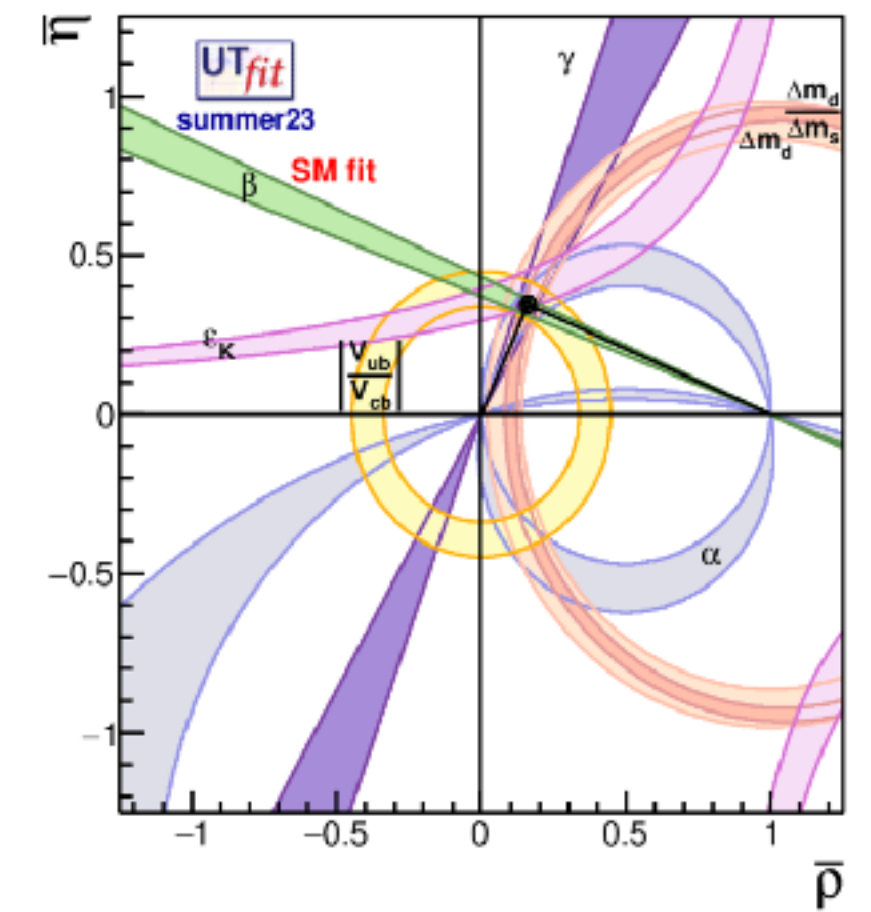
Generic **heavy new physics**, parametrised in SMEFT:

$$\mathcal{L}_{\text{SMEFT}}^{(d=6)} = \sum_i \frac{C_i^{(6)}}{\Lambda^2} \mathcal{O}_i^{(6)}[\psi_{\text{SM}}]$$



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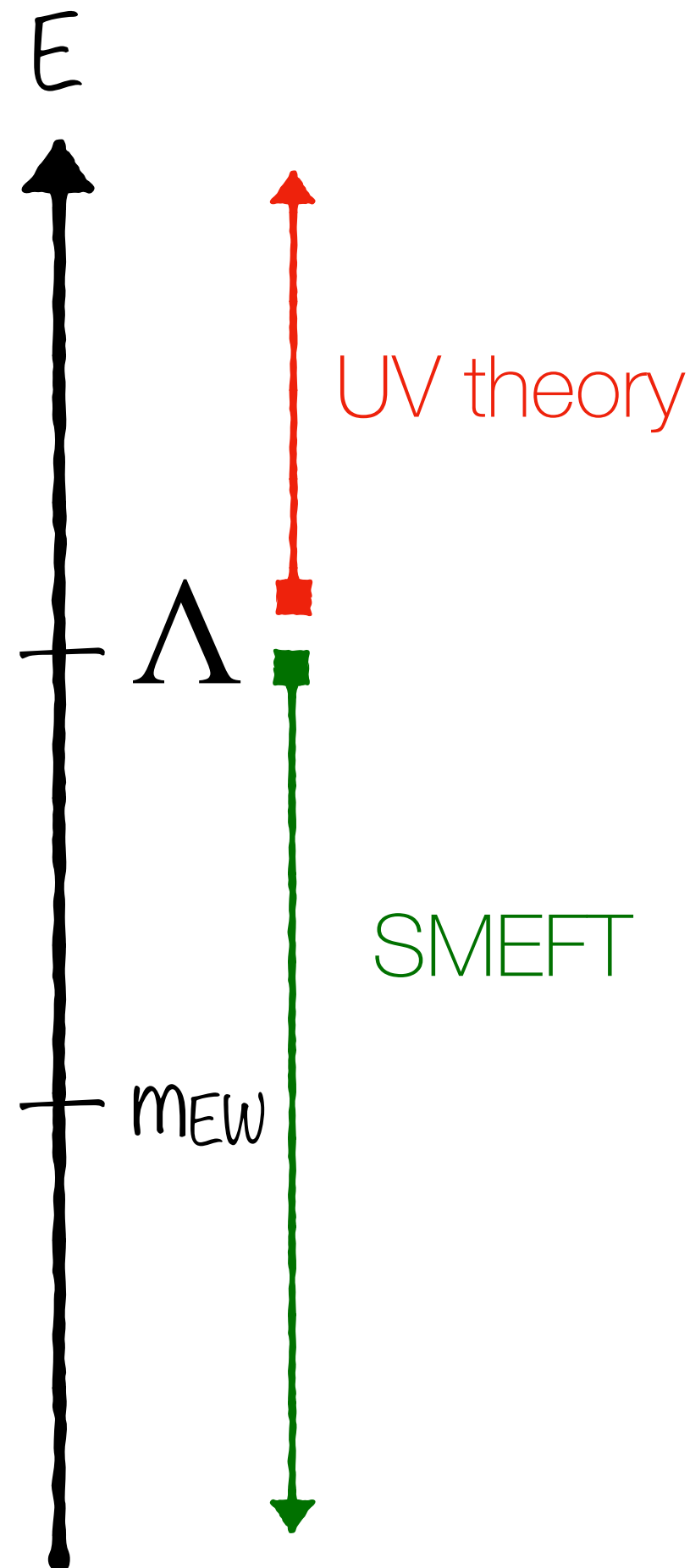


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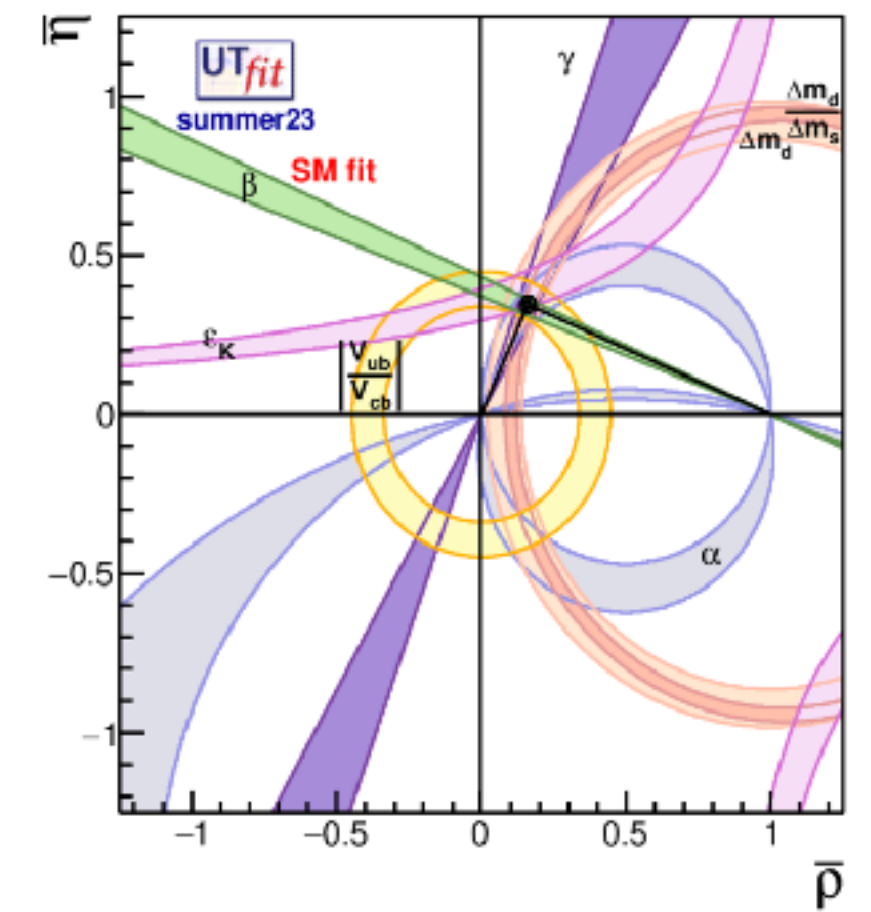
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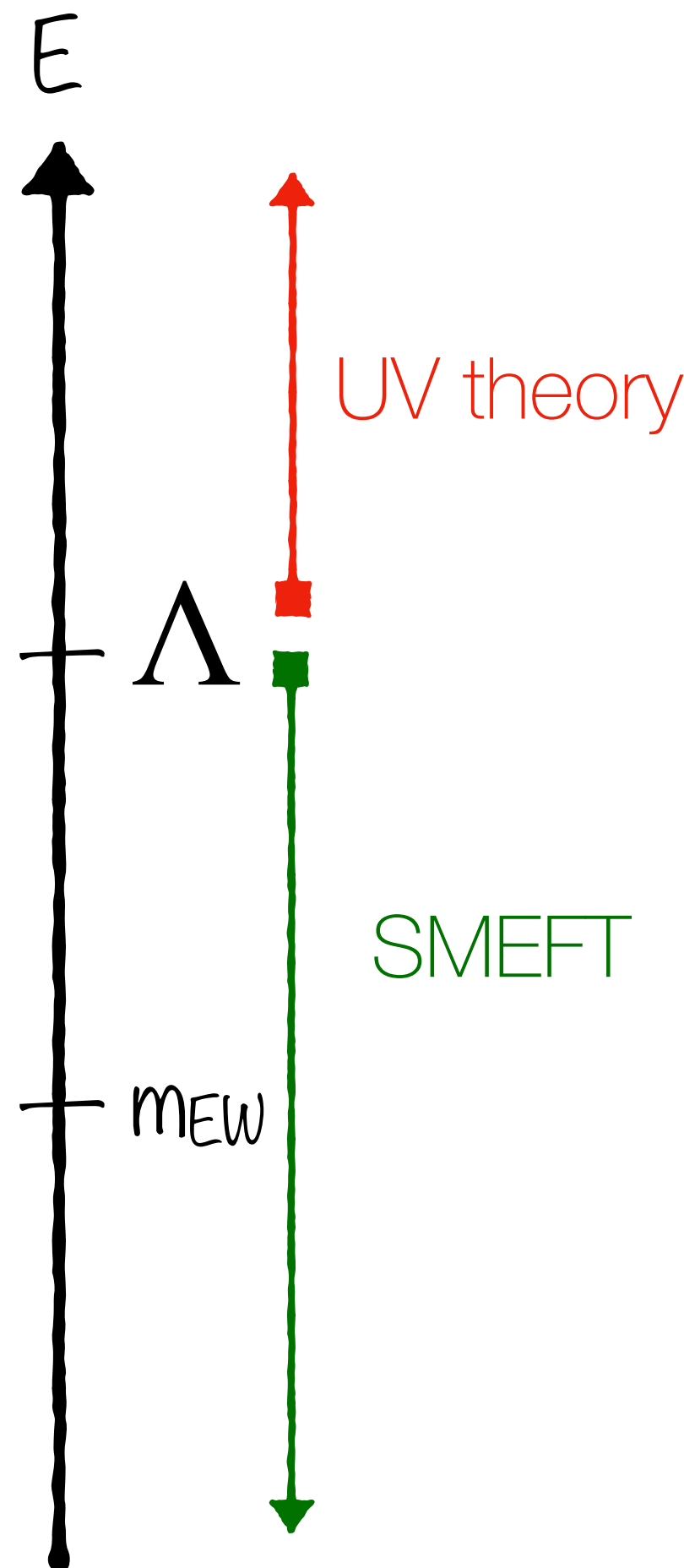
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Precision tests of forbidden or suppressed processes in the SM are powerful probes of physics **Beyond the Standard Model.**

>> Flavour Physics ! <<

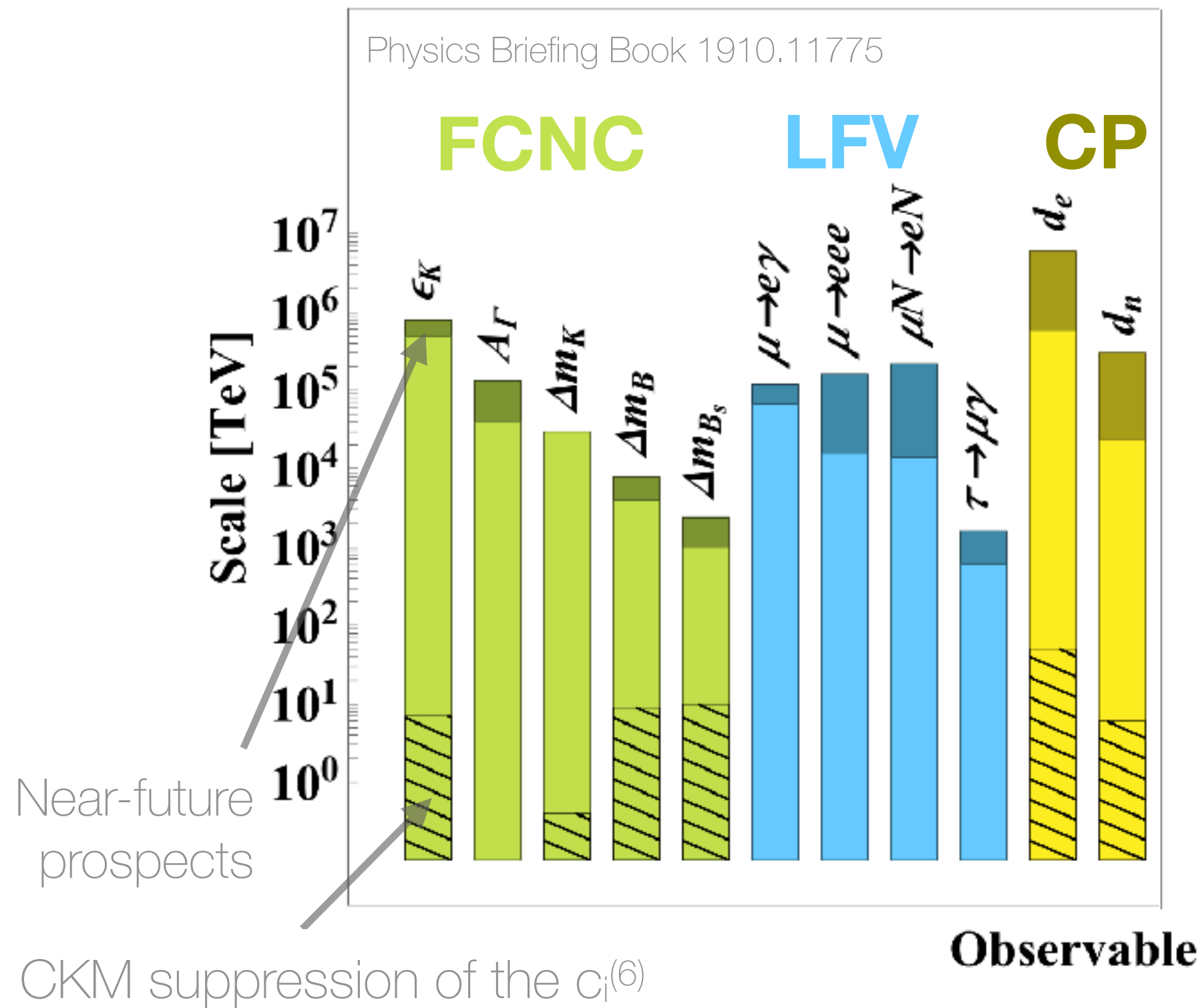


The BSM Flavour Problem

Precision tests push Λ to be very high

Bounds on Λ (taking $c_i^{(6)} = 1$) from various processes

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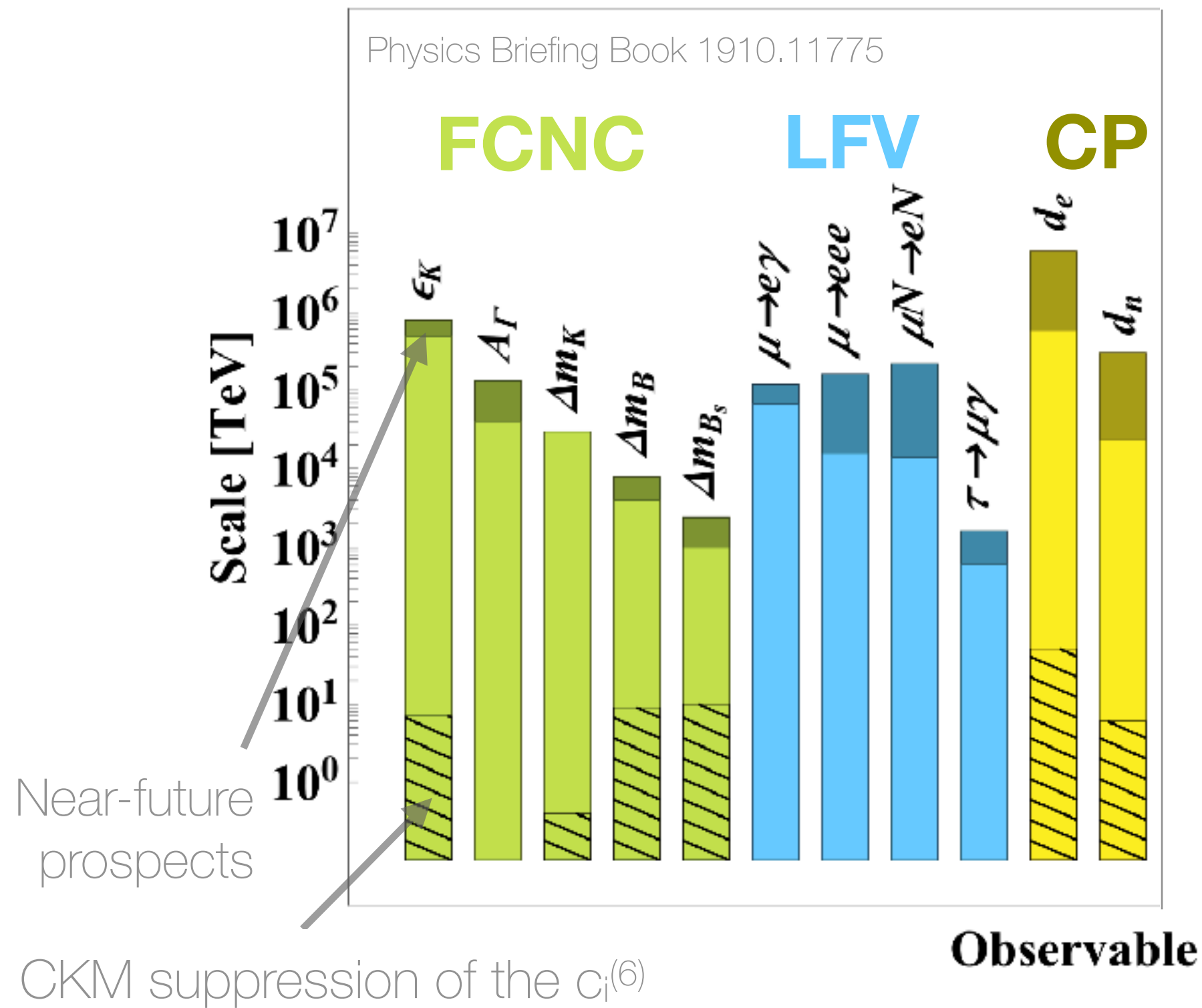
If $c_{FV}^{(6)} = 1$: $\Lambda_{FV} \gtrsim 10^6$ TeV

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Smaller values of the NP scale are instead **motivated**

$$\Lambda_{\text{NP}} \sim 1 - 10 \text{ TeV}$$

- Solutions to the Hierarchy Problem
- Reach of present/future colliders
- Experimental anomalies

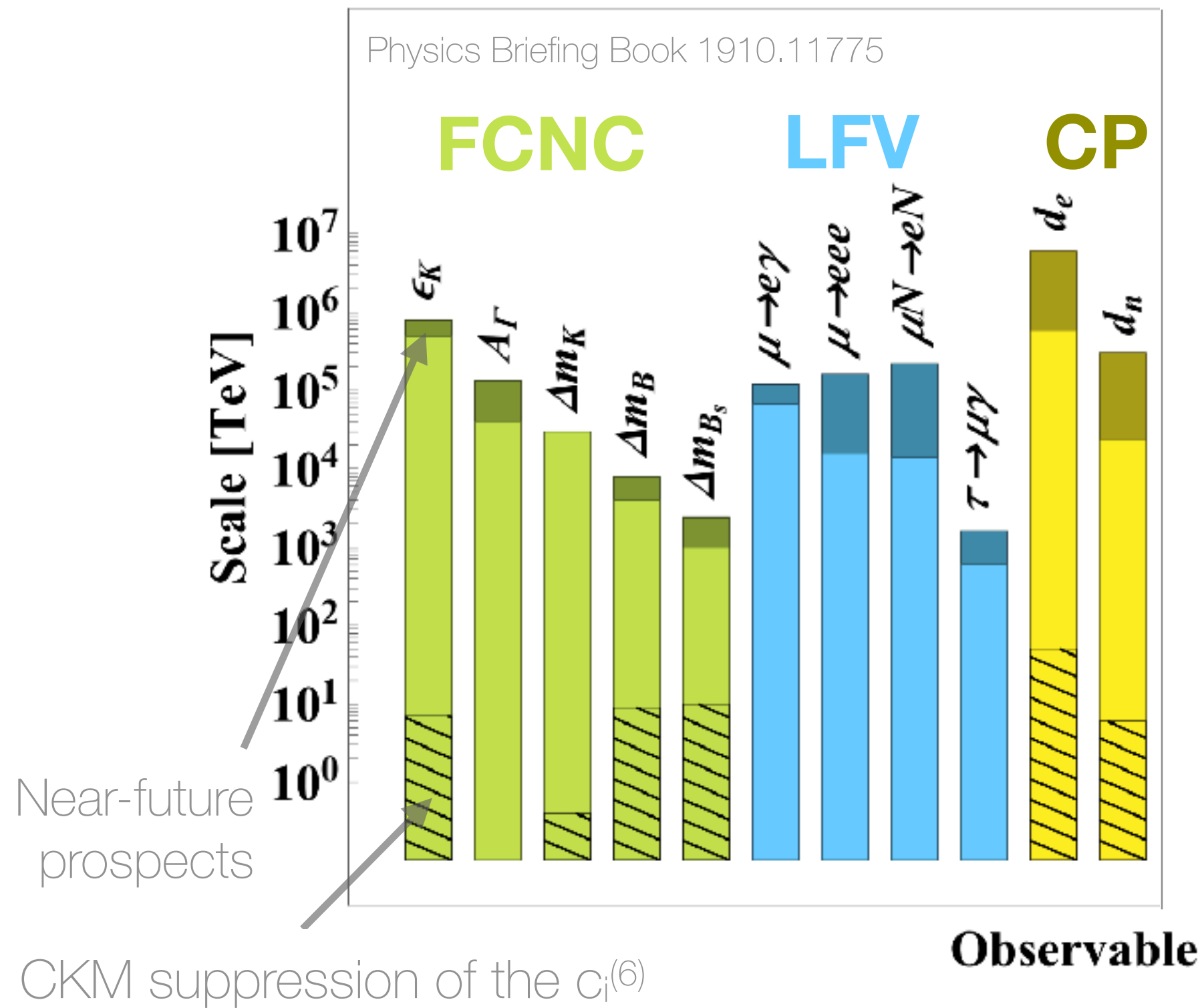
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★ Need some Flavour Protection

E.g. CKM-like suppression of FCNC

If $c_{\text{FV}}^{(6)} = 1$: $\Lambda_{\text{FV}} \gtrsim 10^6 \text{ TeV}$

$$C_{ij} \sim \begin{pmatrix} \epsilon_1 & \lambda^5 & \lambda^3 \\ \lambda^5 & \epsilon_2 & \lambda^2 \\ \lambda^3 & \lambda^2 & 1 \end{pmatrix} \quad \lambda \sim \sin \vartheta_c$$

U(2)-like: $\epsilon_{1,2} \ll 1$

MFV-like: $\epsilon_{1,2} \sim 1$

New physics likes the Top

(1)

Bounds from direct searches @ LHC are **stronger for light fermions than for third generation** ones.

E.g. squark: $M_{\tilde{q}_{(1,2)}} \gtrsim 2 \text{ TeV}$ $M_{\tilde{t},\tilde{b}} \gtrsim 1.4 \text{ TeV}$

E.g. Scalar LQ: $M_{LQ(\mu,e)} \gtrsim 1.8 \text{ TeV}$ $M_{LQ(\tau)} \gtrsim 1.1 \text{ TeV}$

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New Physics coupled to the top should be **lighter** in order to address the **Higgs hierarchy problem**, could also be related to the SM flavour puzzle (lighter NP gives larger Yukawas in some models).

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★ Non-universal couplings preferred

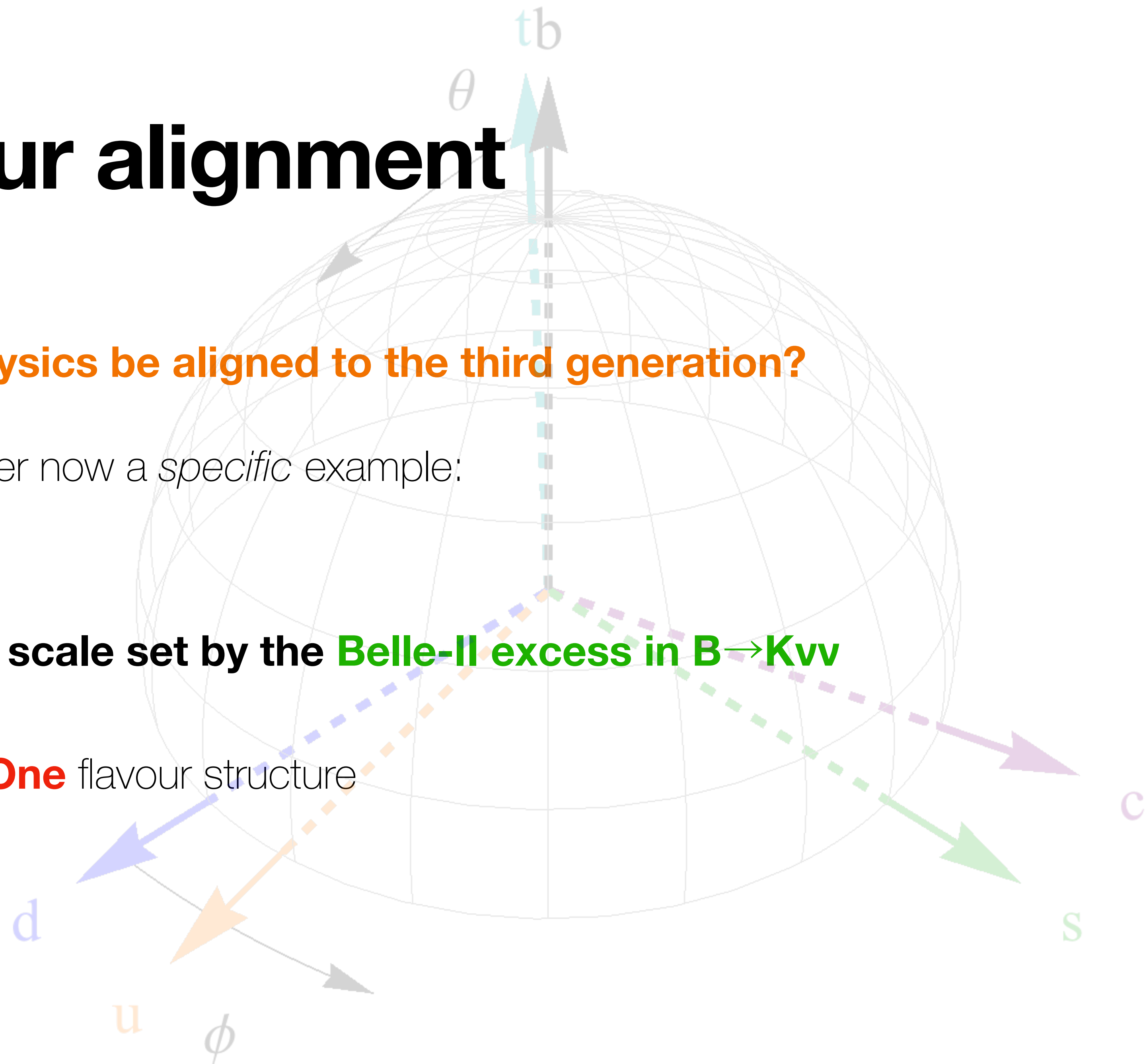
U(2)-like: $\xi_{1,2} \ll 1$

Flavour alignment

How much should New Physics be aligned to the third generation?

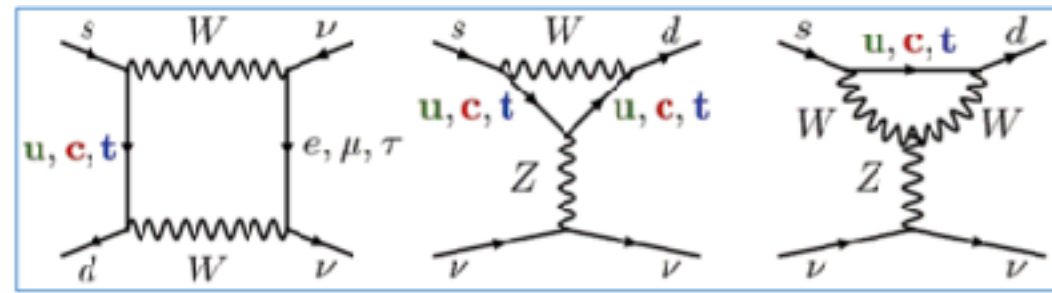
We consider now a *specific* example:

- Overall **New Physics scale** set by the **Belle-II excess in $B \rightarrow K_{\nu\nu}$**
- We assume a **Rank-One** flavour structure



Golden channel decay $B \rightarrow K^{(*)} \nu \bar{\nu}$ $b \rightarrow s \nu \bar{\nu}$

Precise SM predictions possible due to absence of long-distance QCD effects:
neutrinos do not couple to the electromagnetic current.



see 1409.4557, 1503.02693, 2109.11032, 2301.06990, ...

$$\text{BR}(B^+ \rightarrow K^+ \nu \bar{\nu})_{\text{SM}} = (0.444 \pm 0.030) \times 10^{-5}$$

Becirevic et al. 2301.06990

$$\text{BR}(B^0 \rightarrow K^{*0} \nu \bar{\nu})_{\text{SM}} = (9.05 \pm 1.4) \times 10^{-6}$$

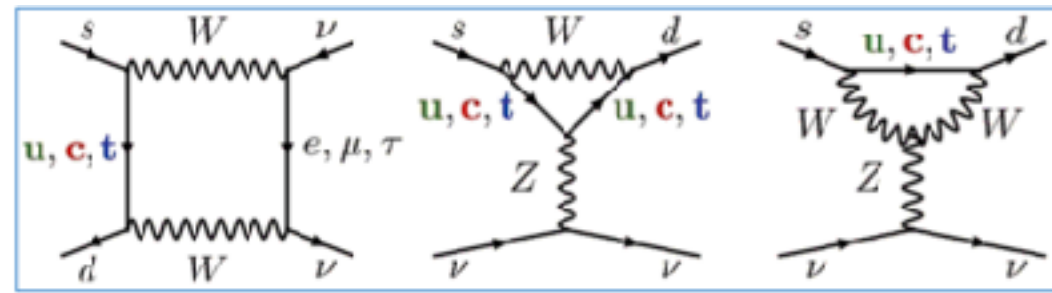
$$\text{Belle-II}_{2023}: \text{BR}(B^+ \rightarrow K^+ \nu \bar{\nu}) = (2.3 \pm 0.6) \times 10^{-5}$$

$$\text{Belle}_{2017}: \text{BR}(B \rightarrow K^* \nu \bar{\nu}) < 2.7 \times 10^{-5} \quad @ 90\% \text{CL}$$

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$$R_K^\nu = \frac{\text{BR}(B \rightarrow K \nu \bar{\nu})}{\text{BR}(B \rightarrow K \nu \bar{\nu})_{\text{SM}}} = 2.93 \pm 0.90 \quad 2.1\sigma$$

$$R_{K^*}^\nu = \frac{\text{BR}(B \rightarrow K^* \nu \bar{\nu})}{\text{BR}(B \rightarrow K^* \nu \bar{\nu})_{\text{SM}}} < 3.2 \quad @ 95\% \text{CL}$$

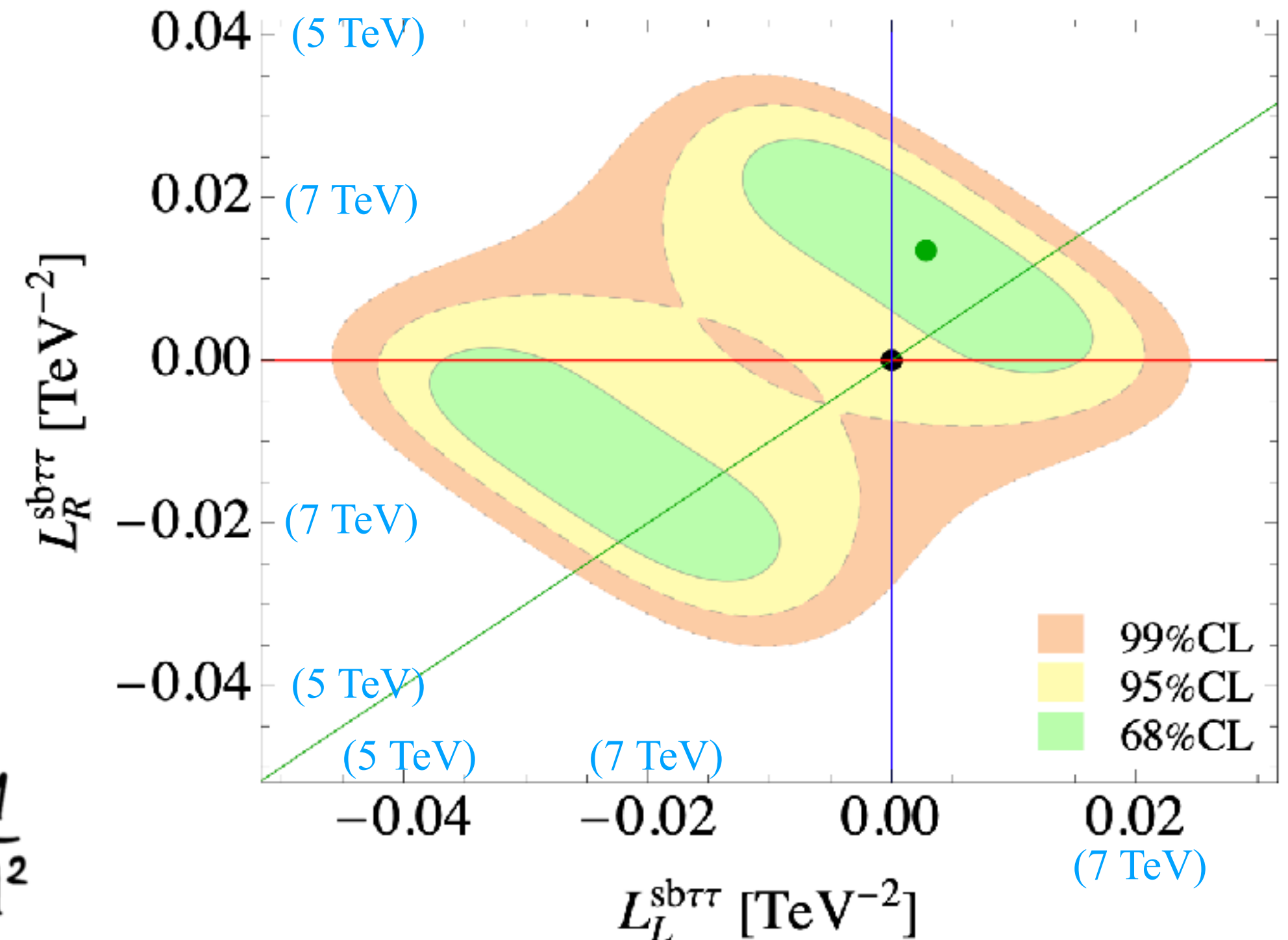
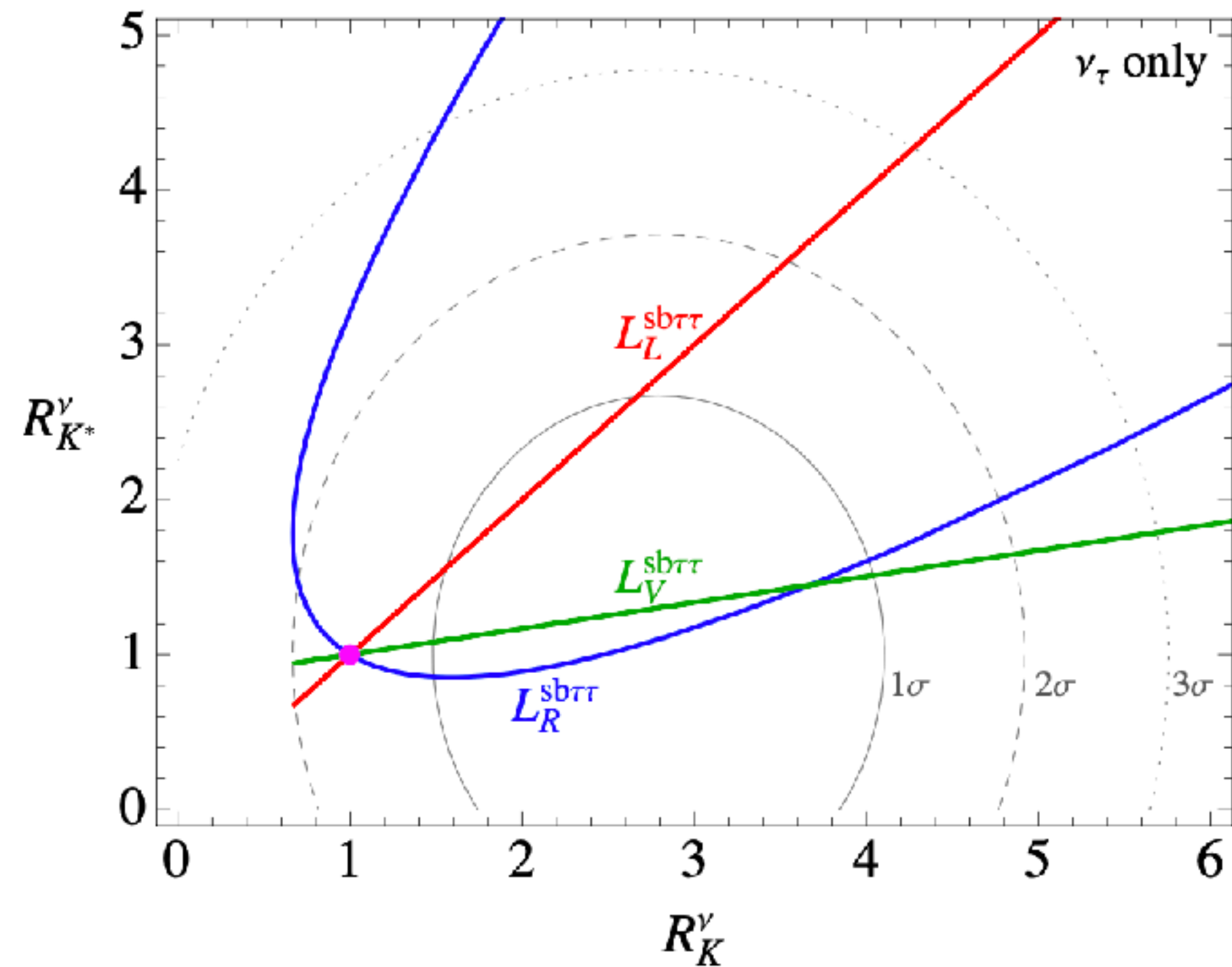
Golden channel decay $B \rightarrow K^{(*)} \nu \bar{\nu}$

Assuming **only NP in tau**
(see paper for other cases)

$$\mathcal{L}_{\text{EFT}} \supset L_{L,R}^{i j \tau \tau} \left(\bar{d}_{iL,R} \gamma_\mu d_{jL,R} \right) \left(\bar{\nu}_\tau \gamma^\mu \nu_\tau \right)$$

DM, M. Nardecchia, A. Stanzione, C. Toni [2404.06533]

$$L_{V,A}^{sb\alpha\beta} \equiv L_R^{sb\alpha\beta} \pm L_L^{sb\alpha\beta}$$



$$L \sim \frac{1}{\Lambda^2}$$

$$L^{sb\nu\nu} \sim \frac{1}{(8 \text{ TeV})^2}$$

They probe scales of about 8 TeV,

with a slight excess from the SM preferring either a RH or vector-like quark current. Future Belle II results (in particular from the K^* mode) will help to clarify the situation.



NA62 (CERN)

$$\text{BR}(K^+ \rightarrow \pi^+ \nu \bar{\nu})_{\text{SM}} = (8.09 \pm 0.63) \times 10^{-11}$$

Allwicher et al. [2410.21444] (see also Buras et al. 1503.02693, 2109.11032, etc.)

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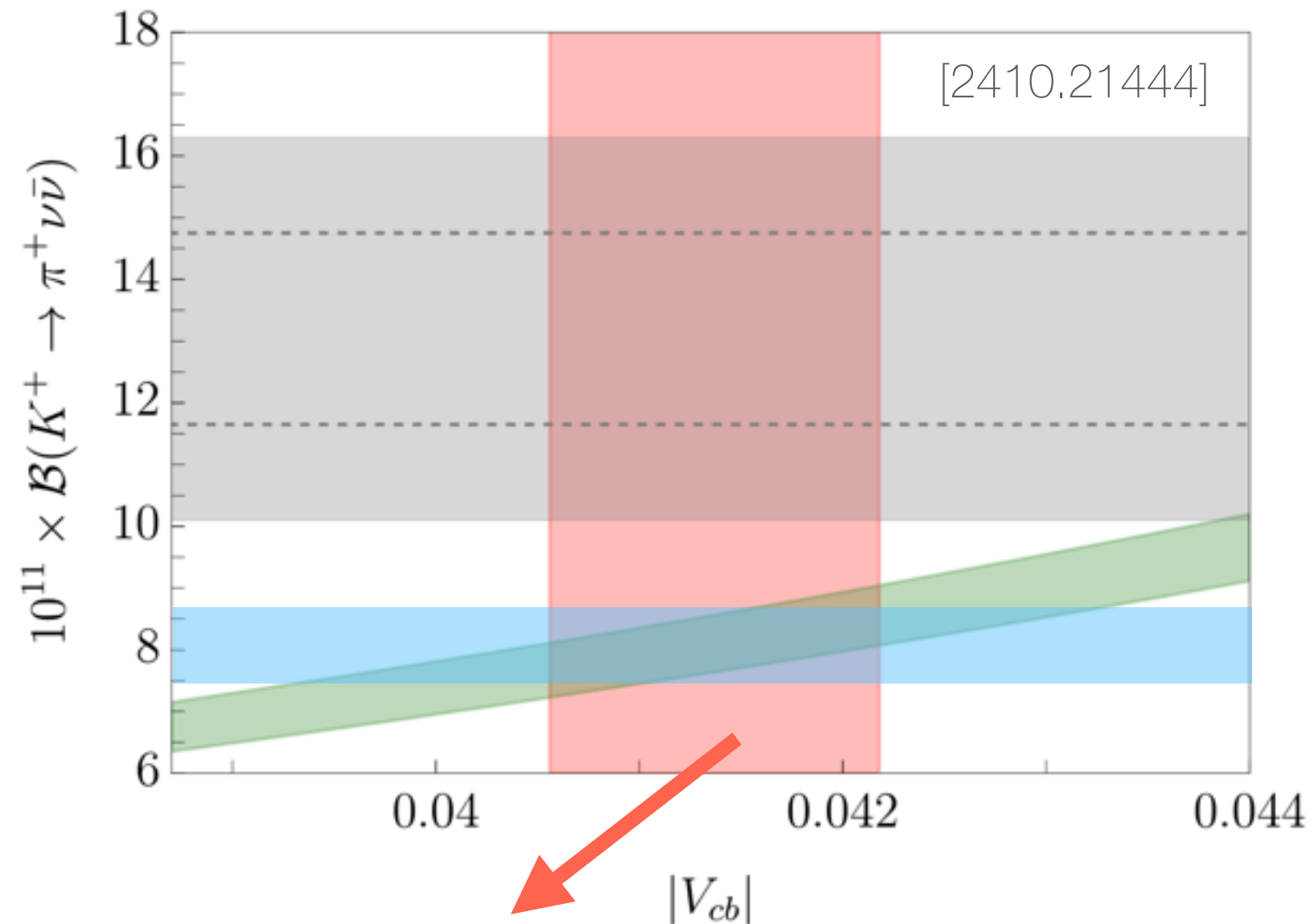
KOTO (JPARC)

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$$|V_{cb}| = (41.37 \pm 0.81) \times 10^{-3}$$

Derived by combining exclusive and inclusive determinations. [2310.20324, 2406.10074]



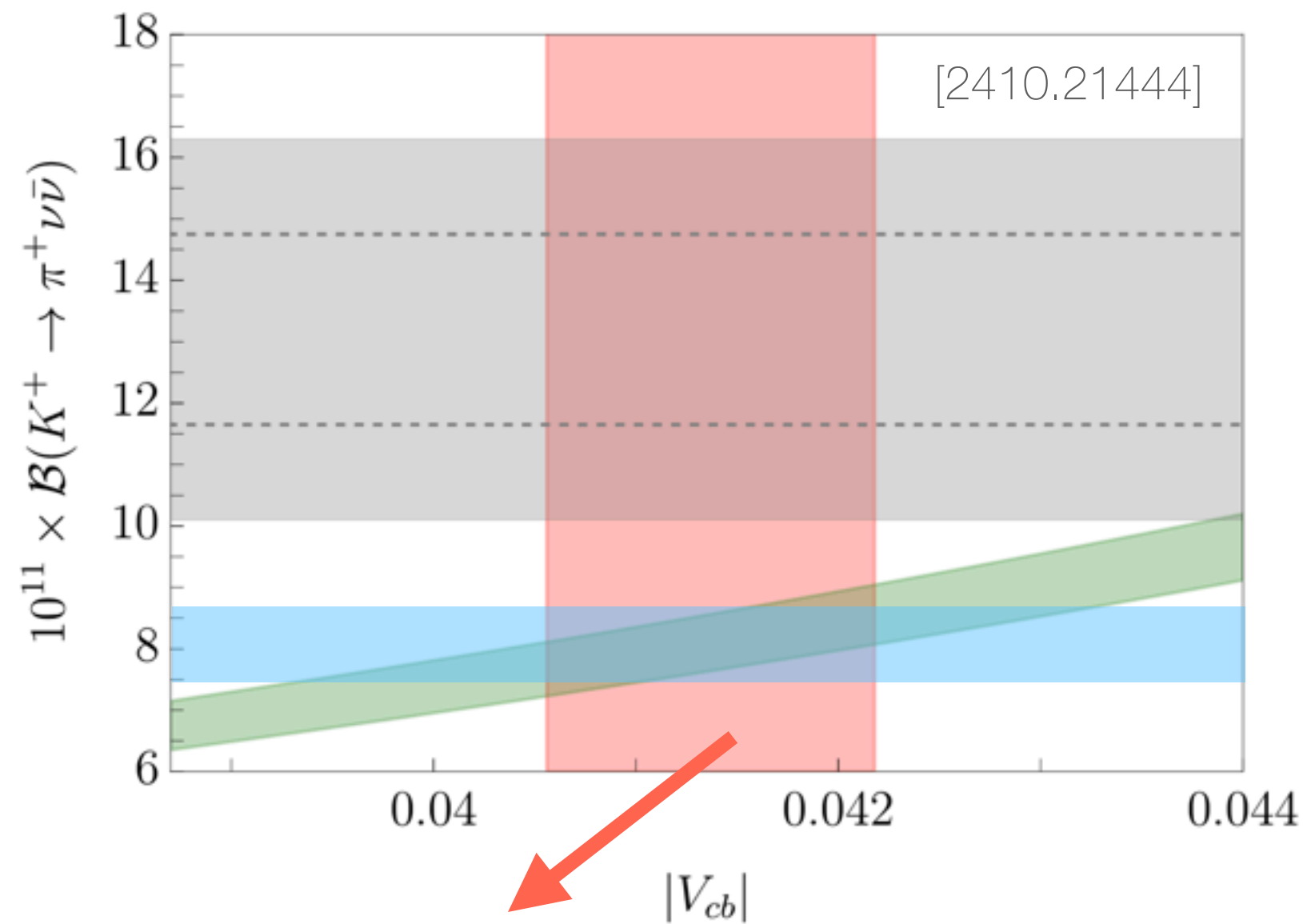
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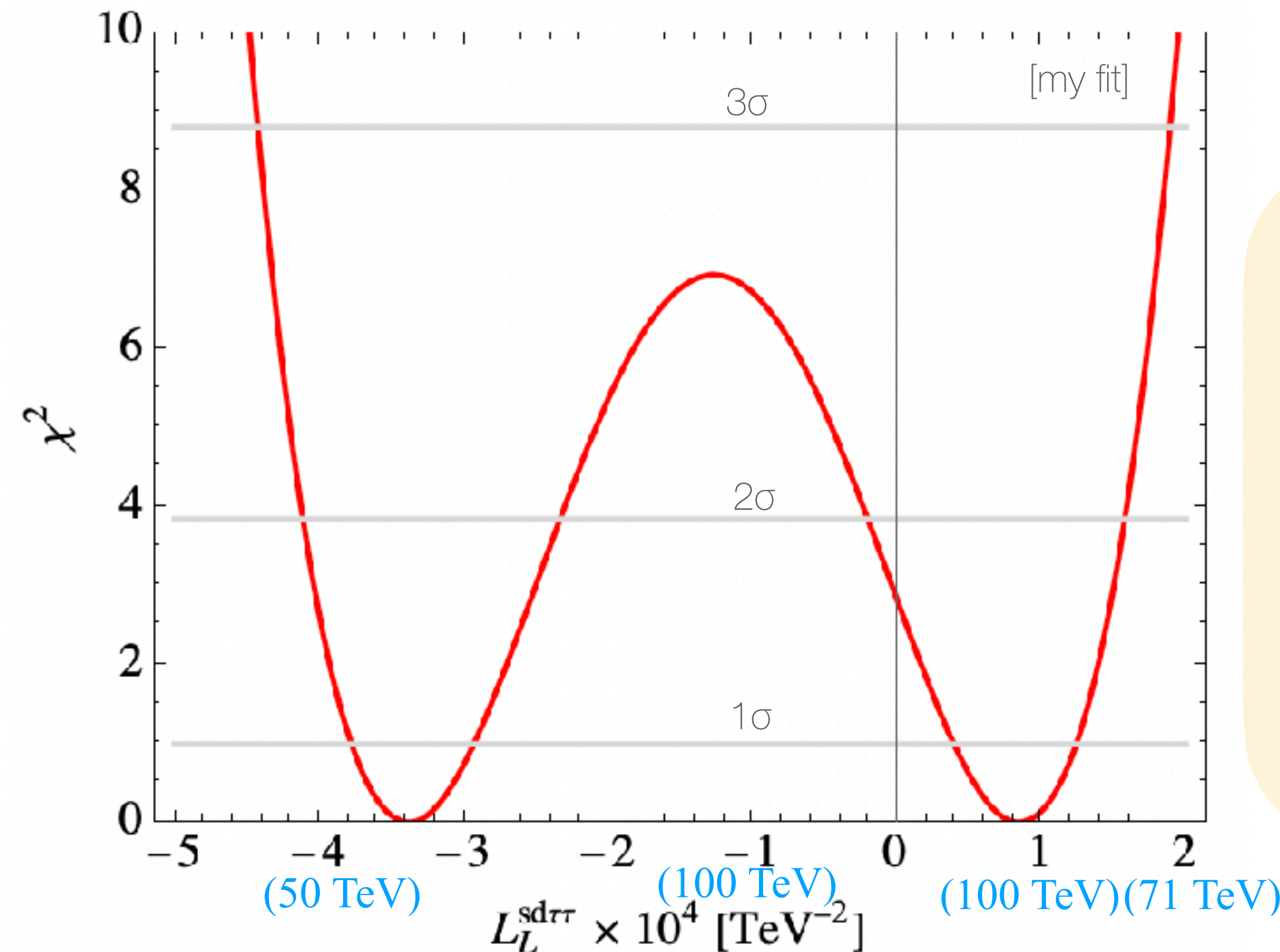
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$$L \sim \frac{1}{\Lambda^2}$$

The **slight $<2\sigma$ excess** points to new physics scales

$$\Lambda_{\text{sd}\nu\nu} \sim 100 \text{TeV}$$

Directions in Flavour Space

Gherardi, DM, Nardecchia, Romanino [1903.10954](#)
DM, Nardecchia, Stanzione, Toni [2404.06533](#)

Consider the vector space spanned by the
3 generations of down quarks, $SU(3)_q$:

$$\hat{d} \sim \begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix} \quad \hat{s} \sim \begin{pmatrix} 0 \\ 1 \\ 0 \end{pmatrix} \quad \hat{b} \sim \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix}$$

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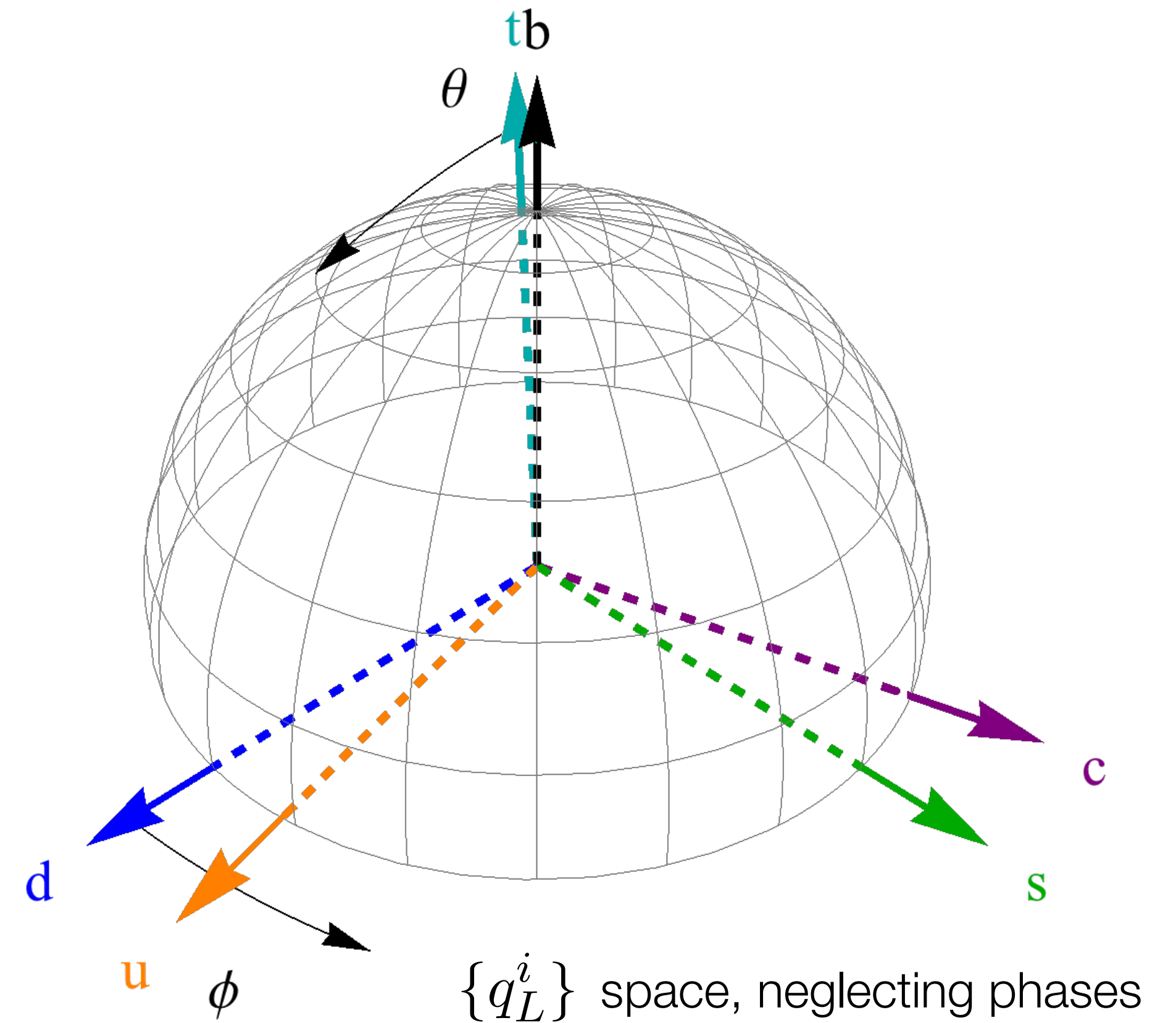
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We can parametrise a generic directions as:

$$\hat{n} = \begin{pmatrix} \sin \theta \cos \phi e^{i\alpha_{bd}} \\ \sin \theta \sin \phi e^{i\alpha_{bs}} \\ \cos \theta \end{pmatrix} \quad \begin{array}{l} \text{neglecting phases,} \\ \text{it is a unit-vector} \\ \text{on a semi-sphere} \end{array}$$

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The overall phase is unphysical: $U(1)_B$



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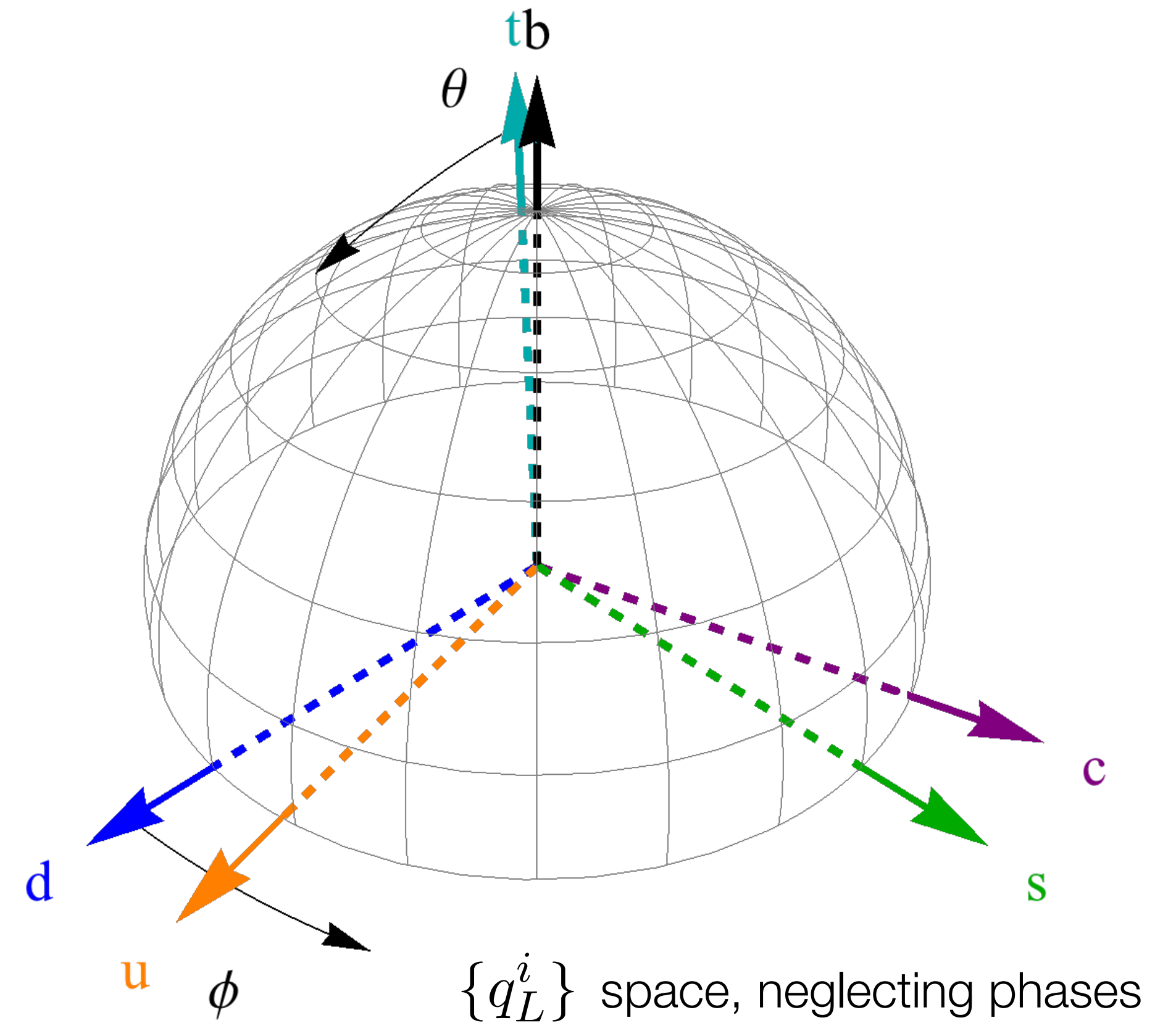
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We show also up quarks using:

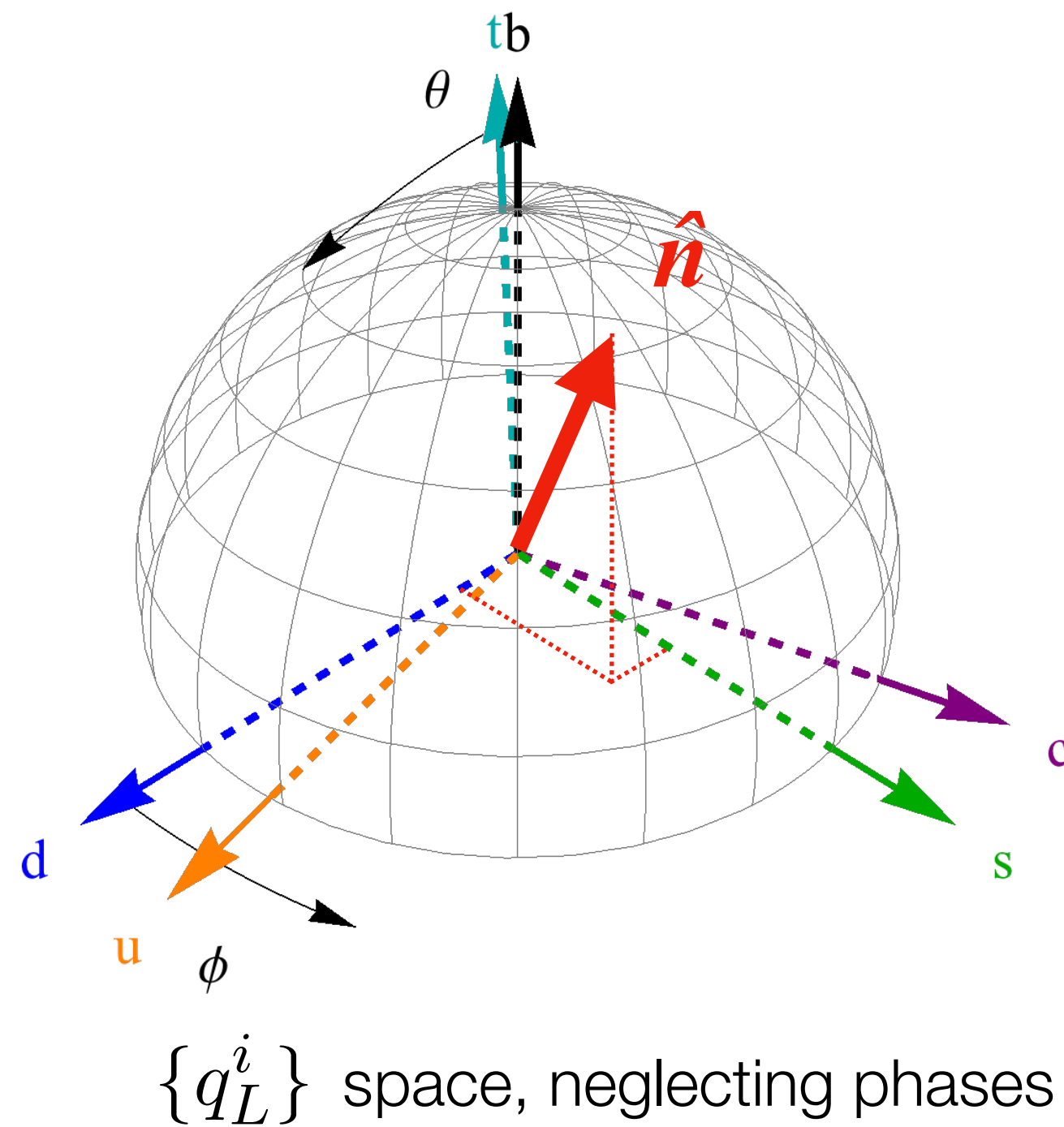
$$q_L^i = \begin{pmatrix} V_{ji}^* u_L^i \\ d_L^i \end{pmatrix}$$

Rank-One Flavour Violation

Gherardi, DM, Nardecchia, Romanino [1903.10954](#)

DM, Nardecchia, Stanzione, Toni [2404.06533](#)

$$\mathcal{L}_{\text{LEFT}}^{\text{NP}} = \sum_{ij\alpha\beta} \left[L_L^{ij\alpha\beta} (\bar{d}_L^i \gamma_\mu d_L^j) (\bar{\nu}_L^\alpha \gamma^\mu \nu_L^\beta) + L_R^{ij\alpha\beta} (\bar{d}_R^i \gamma_\mu d_R^j) (\bar{\nu}_L^\alpha \gamma^\mu \nu_L^\beta) \right]$$



We assume that **New Physics is aligned to a specific direction \hat{n}** .

> the **EFT coefficients** are given by an overall scale times the **projection of \hat{n} on the specific flavour direction**

$$L^{ij\nu\nu} = C \hat{n}_i \hat{n}_j^*$$

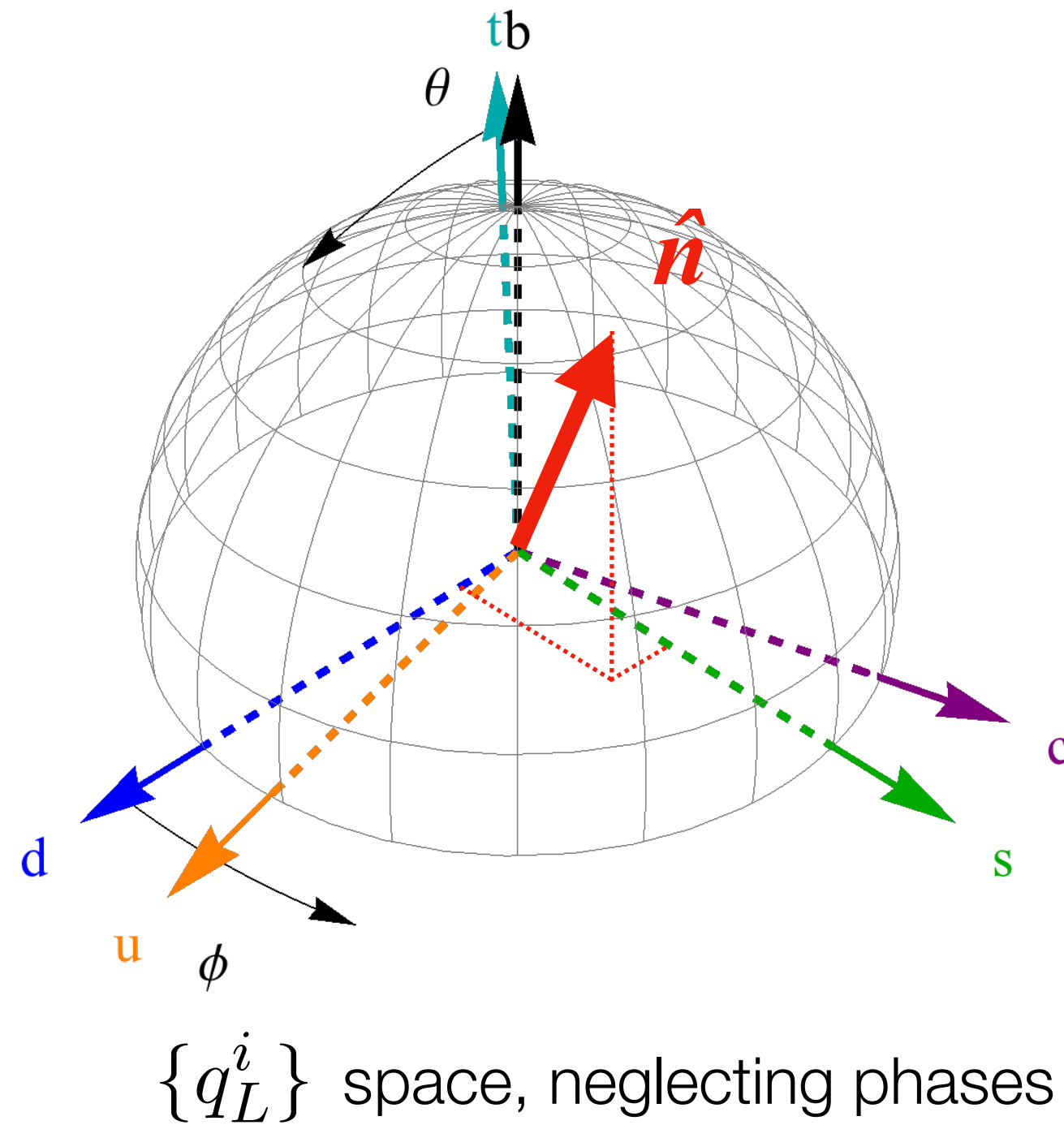
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This structure is **automatic** if New Physics couples linearly to a single combination of quarks:

$$\mathcal{L} \supset \lambda_i \bar{q}^i \mathcal{O}_{\text{NP}} + \text{h.c.}$$

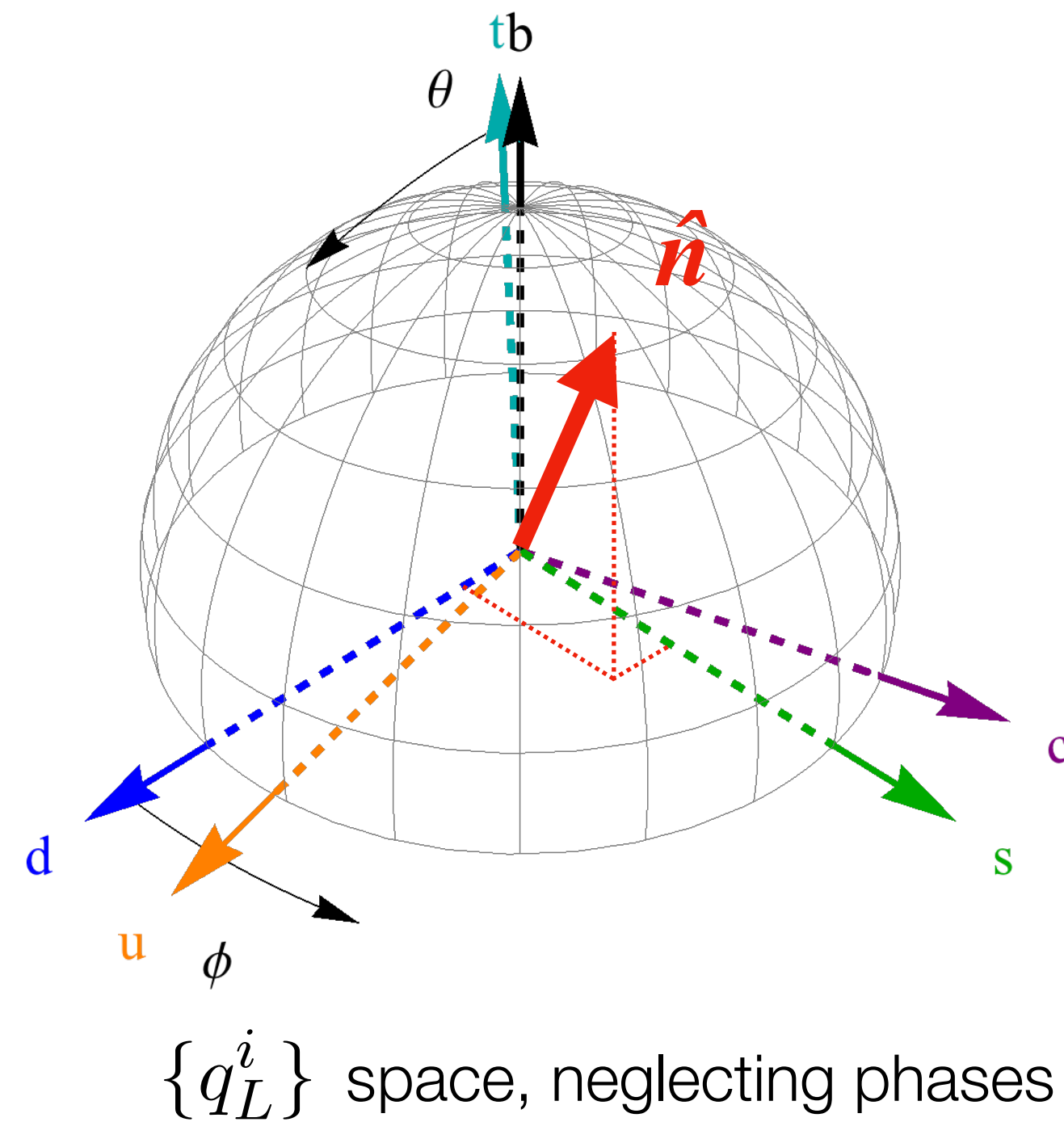
- e.g.
- leptoquarks coupled mainly to 1 lepton family
 - Vector coupled via the mixing of a single vector-like quark

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Gherardi, DM, Nardecchia, Romanino [1903.10954](#)

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At any value of (ϕ, θ) we can fix the overall scale **C** by imposing the **best-fit of $B \rightarrow K^{(*)} \nu\nu$** .

$$L^{sb\nu\nu} = C \cos \theta \sin \theta \sin \phi \equiv (8 \text{ TeV})^{-2}$$

For the best-fit of $R_{\nu K}$ and for simplicity we fix: $\alpha_{bs} = \alpha_{bd} = 0$ (fit in backup slides)

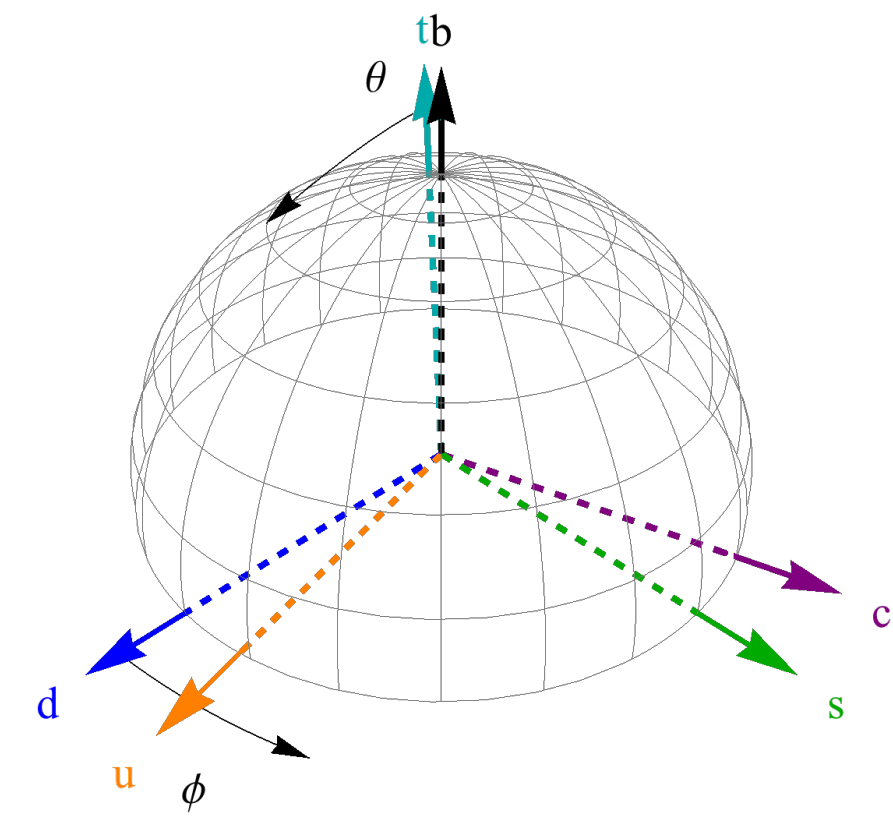
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Once **C** is fixed as function of (θ, ϕ) , all parameters are set and we can check the **constraints from other observables**

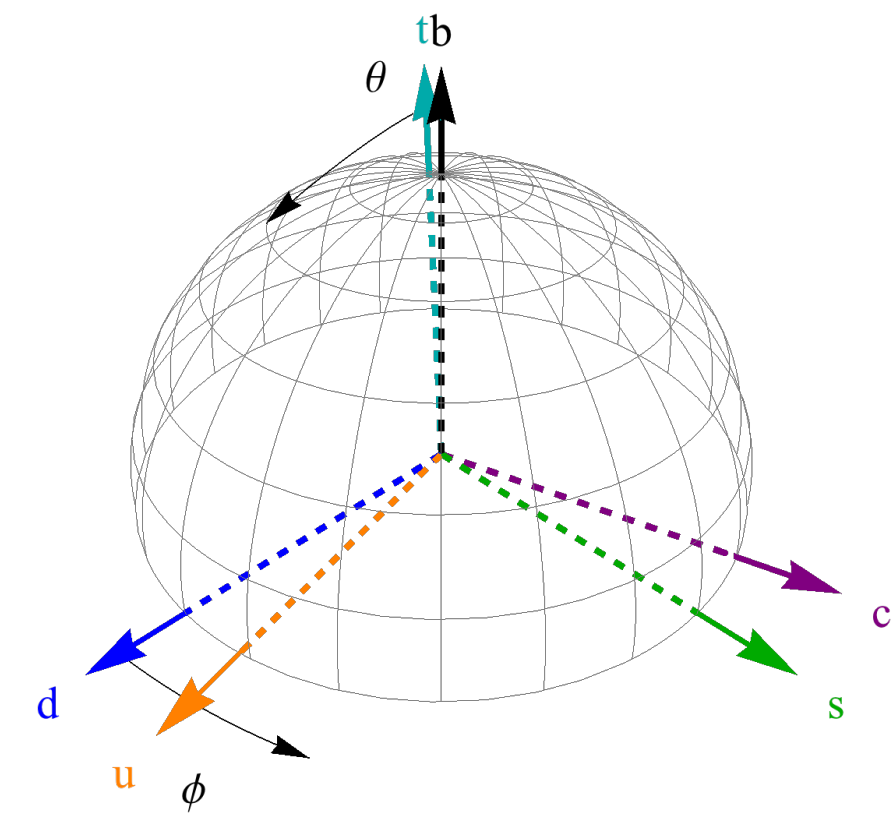
[2404.06533](#)



Rank-One Flavour Violation

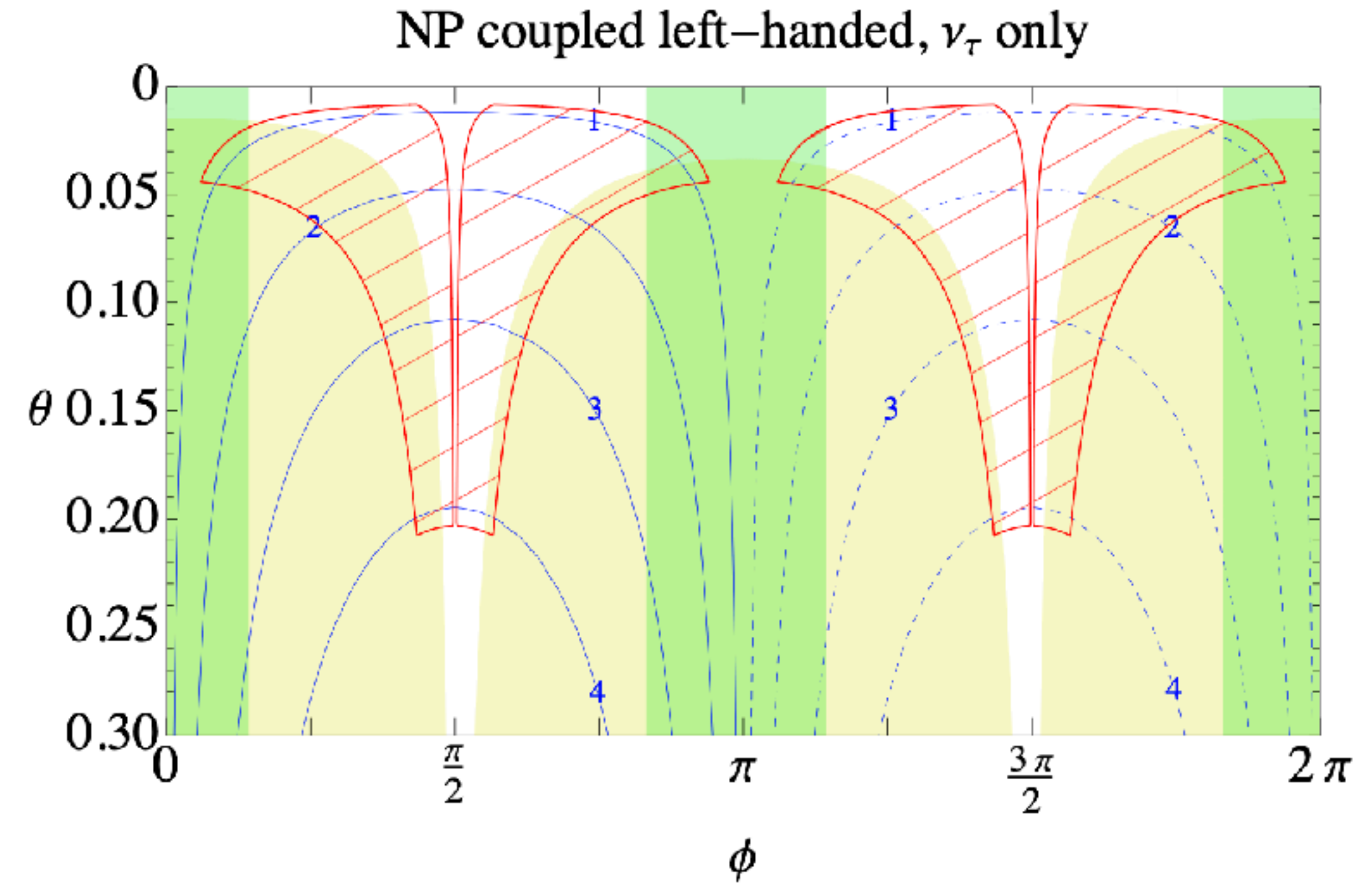
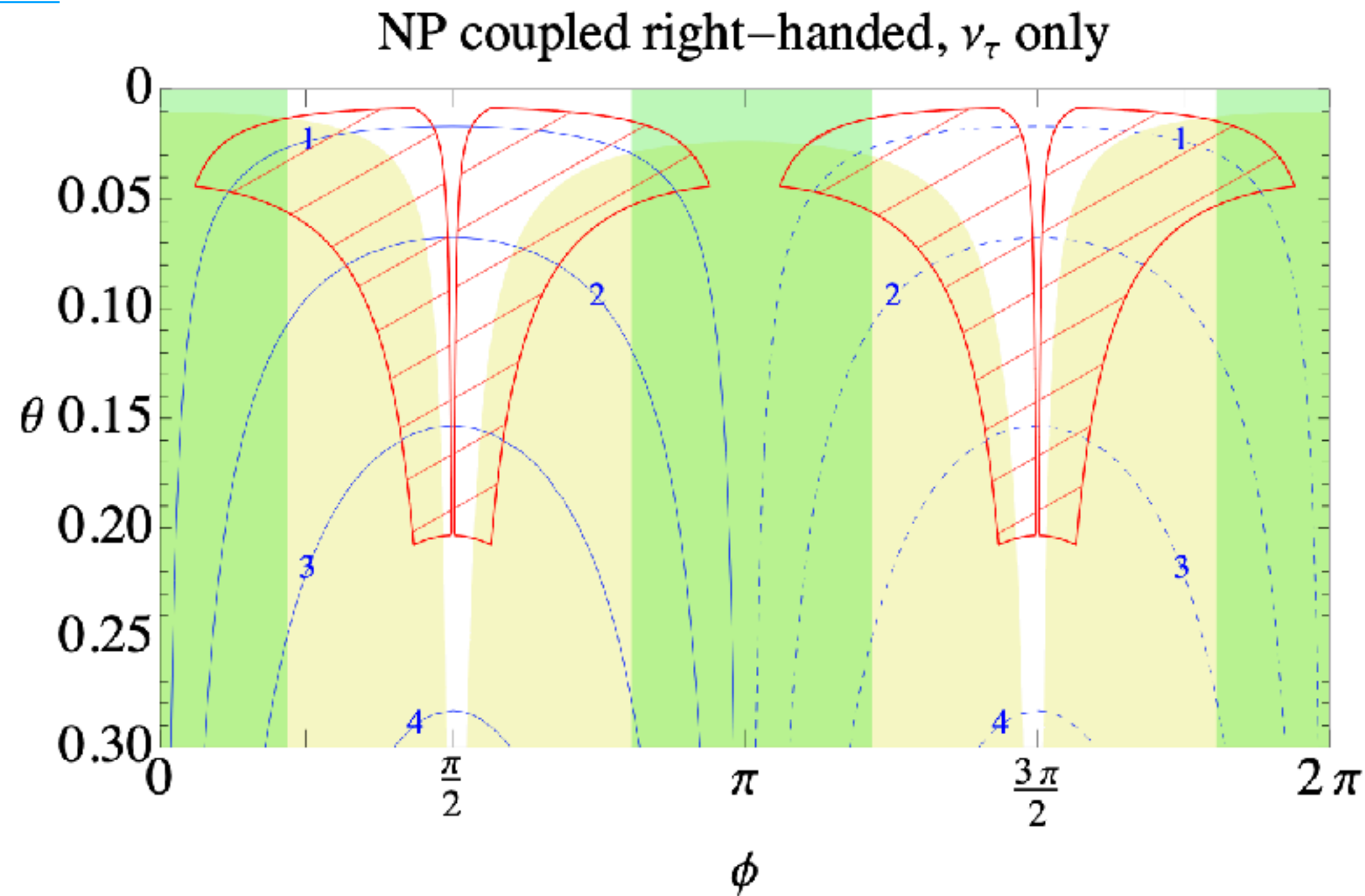
$$L^{ij\nu\nu} = C \hat{n}_i \hat{n}_j^*$$

$$L^{sb\nu\nu} = C \cos\theta \sin\theta \sin\phi \equiv (8\text{TeV})^{-2}$$



Once **C is fixed as function of (θ, ϕ)**, all parameters are set and we can check the **constraints from other observables**

[2404.06533](#)



$b \rightarrow d\nu\bar{\nu}$

$s \rightarrow d\nu\bar{\nu}$

\hat{n}_{3rd}

$|C|^{-1/2}$ [TeV]

$$\hat{n}_{3rd} \sim (V_{td}, V_{ts}, \gamma)$$

The allowed region (white) is close to the third generation, with a misalignment of $O(\text{CKM})$.

Top-philic New Physics

We just saw how the **preferred region** to address the Belle-II excess compatibly with other constraints is **close to the third generation quarks** (up to $O(\text{CKM})$ deviations).

This trend is expected and well known, can be generalised:

Both experimental and theory arguments motivate having
TeV-scale New Physics coupled mostly to the top quark.

[e.g. review by Franceschini 2301.04407]

If, in the quark sector, **New Physics couples indeed mostly to the top quark:**



What are the strongest constraints we can put?



Where do they come from?



How do indirect bounds compare with direct ones from LHC?

[Garosi, DM, Rodriguez-Sanchez, Stanzione 2310.00047]

[see also these, on similar spirit: 0704.1482, 0802.1413, 1109.2357, 1408.0792, 1909.13632, 2012.10456]

Top SMEFT

Let us **assume heavy NP couples, among quarks, mostly to the top.**

We leave **arbitrary lepton flavour and gauge structures**

These are the **SMEFT dim-6** operators satisfying these conditions:

Since we are assuming that the top quark is somehow "special" from the UV point of view, we work in the **up quark mass basis**:

$$q^i = (u_L^i, V_{ij} d_L^j)$$

[see Isidori et al. 2024 for an analysis varying continuously from up to down basis]

Semi-leptonic		Four quarks	
$\mathcal{O}_{lq}^{(1),\alpha\beta}$	$(\bar{\ell}^a \gamma_\mu \ell^\beta)(\bar{q}^3 \gamma^\mu q^3)$	$\mathcal{O}_{qq}^{(1)}$	$(\bar{q}^3 \gamma^\mu q^3)(\bar{q}^3 \gamma_\mu q^3)$
$\mathcal{O}_{lq}^{(3),\alpha\beta}$	$(\bar{\ell}^a \gamma_\mu \tau^a \ell^\beta)(\bar{q}^3 \gamma^\mu \tau^a q^3)$	$\mathcal{O}_{qq}^{(3)}$	$(\bar{q}^3 \gamma^\mu \tau^a q^3)(\bar{q}^3 \gamma_\mu \tau^a q^3)$
$\mathcal{O}_{lu}^{\alpha\beta}$	$(\bar{\ell}^\alpha \gamma_\mu \ell^\beta)(\bar{u}^3 \gamma_\mu u^3)$	\mathcal{O}_{uu}	$(\bar{u}^3 \gamma^\mu u^3)(\bar{u}^3 \gamma_\mu u^3)$
$\mathcal{O}_{qe}^{\alpha\beta}$	$(\bar{q}^3 \gamma^\mu q^3)(\bar{e}^\alpha \gamma_\mu e^\beta)$	$\mathcal{O}_{qu}^{(1)}$	$(\bar{q}^3 \gamma^\mu q^3)(\bar{u}^3 \gamma_\mu u^3)$
$\mathcal{O}_{eu}^{\alpha\beta}$	$(\bar{e}^\alpha \gamma_\mu e^\beta)(\bar{u}^3 \gamma_\mu u^3)$	$\mathcal{O}_{qu}^{(8)}$	$(\bar{q}^3 \gamma^\mu T^A q^3)(\bar{u}^3 \gamma_\mu T^A u^3)$
$\mathcal{O}_{lequ}^{(1),\alpha\beta}$	$(\bar{\ell}^\alpha e^\beta) \epsilon (\bar{q}^3 u^3)$	Higgs-Top	
$\mathcal{O}_{lequ}^{(3),\alpha\beta}$	$(\bar{\ell}^\alpha \sigma_{\mu\nu} e^\beta) \epsilon (\bar{q}^3 \sigma^{\mu\nu} u^3)$	$\mathcal{O}_{Hq}^{(1)}$	$(H^\dagger i \overleftrightarrow{\mathcal{D}}_\mu H)(\bar{q}^3 \gamma^\mu q^3)$
Dipoles		$\mathcal{O}_{Hq}^{(3)}$	$(H^\dagger i \overleftrightarrow{\mathcal{D}}_\mu^a H)(\bar{q}^3 \gamma^\mu \tau^a q^3)$
\mathcal{O}_{uG}	$(\bar{q}^3 \sigma^{\mu\nu} T^A u^3) \tilde{H} G_{\mu\nu}^A$	\mathcal{O}_{Hu}	$(H^\dagger i \overleftrightarrow{\mathcal{D}}_\mu H)(\bar{u}^3 \gamma^\mu u^3)$
\mathcal{O}_{uW}	$(\bar{q}^3 \sigma^{\mu\nu} u^3) \tau^a \tilde{H} W_{\mu\nu}^a$	\mathcal{O}_{uH}	$(H^\dagger H)(\bar{q}^3 u^3 \tilde{H})$
\mathcal{O}_{uB}	$(\bar{q}^3 \sigma^{\mu\nu} u^3) \tilde{H} B_{\mu\nu}$		

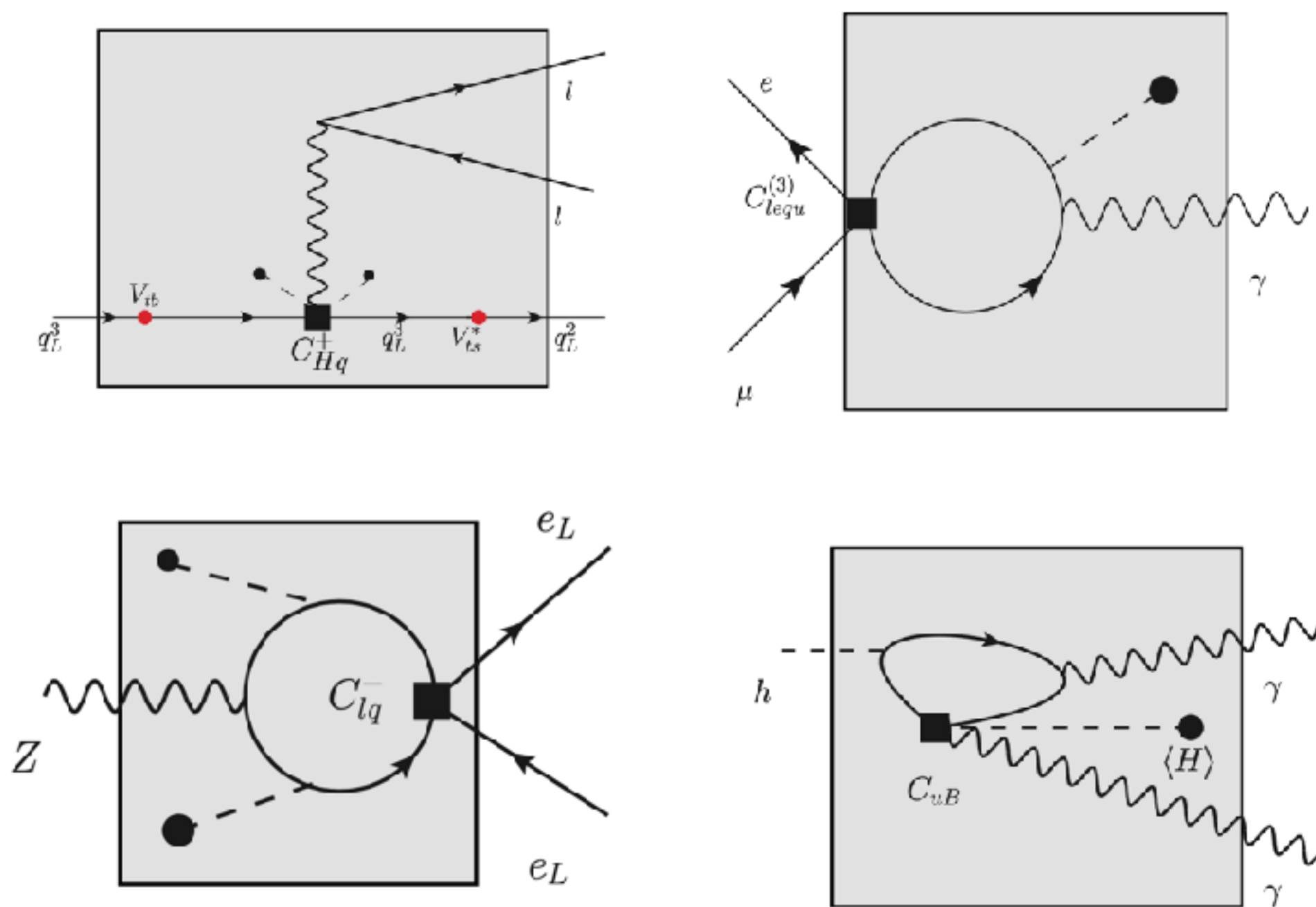
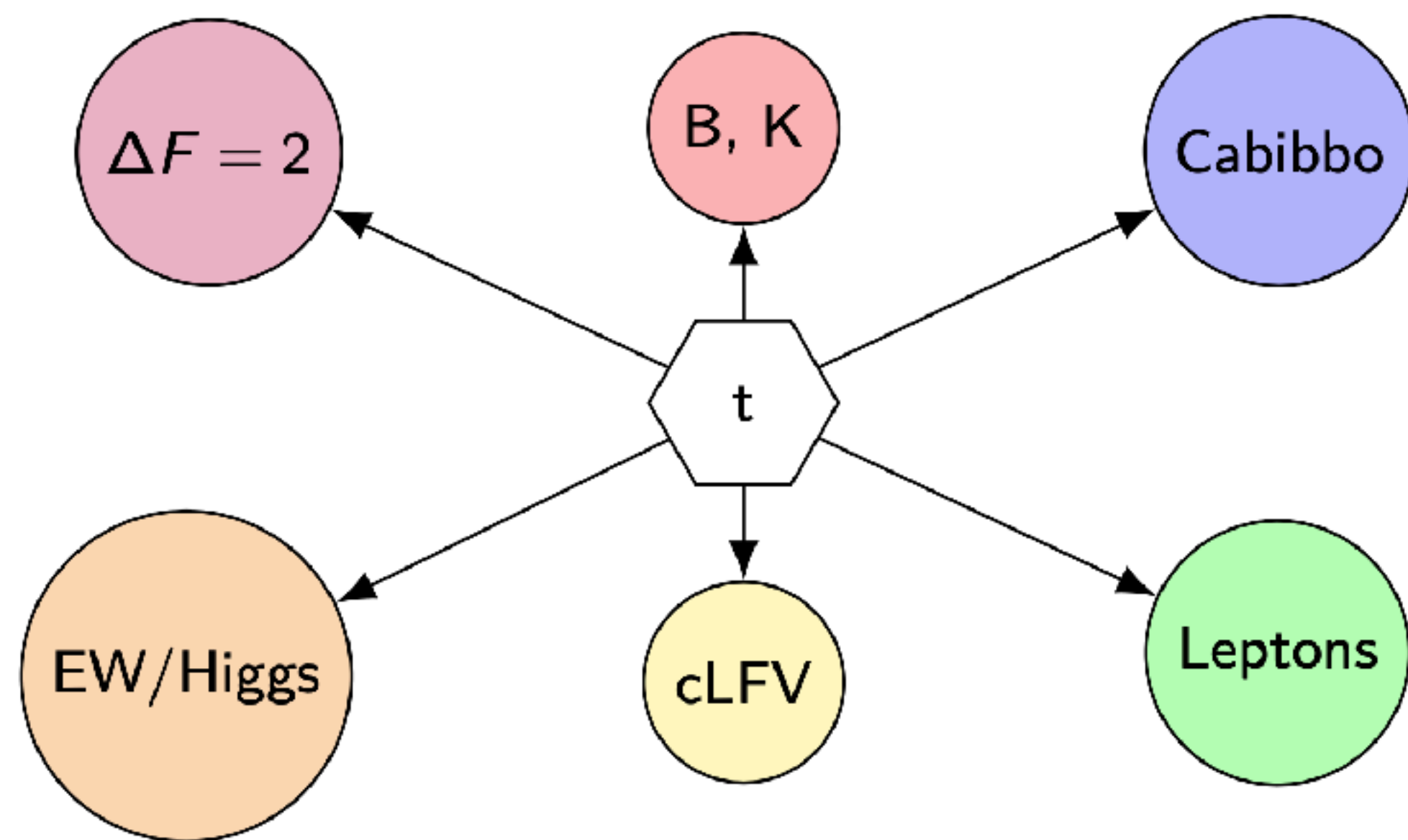
[Garosi, DM, Rodriguez-Sanchez, Stanzione 2310.00047]

Indirect constraints



[Jenkins, Manohar and Trott 2013; Dekens and Stoffer [1908.05295]; Jenkins, Manohar and Stoffer [1711.05270], DSixTools ...]

RG evolution to low scales induce effects in a wide range of observables



Observables included

B physics

Observable	Experimental value
$B \rightarrow X_s \gamma$	$(3.49 \pm 0.19) \times 10^{-4}$ [38]
R_K^ν	2.93 ± 0.90 [35, 36]
$R_{K^*}^\nu$	< 3.21 [35, 37]
$R_K[1.1, 6]$	0.949 ± 0.047 [34]
$R_{K^*}[1.1, 6]$	1.027 ± 0.077 [34]
$\mathcal{B}(B \rightarrow K e \mu)$	$< 4.5 \times 10^{-8}$ [39]
$\mathcal{B}(B \rightarrow K e \tau)$	$< 3.6 \times 10^{-5}$ [40]
$\mathcal{B}(B \rightarrow K \mu \tau)$	$< 4.5 \times 10^{-5}$ [41]

Observable	Experimental value
$\mathcal{B}(B_s \rightarrow ee)$	$< 11.2 \times 10^{-9}$ [42]
$\mathcal{B}(B_s \rightarrow \mu\mu)$	$(3.01 \pm 0.35) \times 10^{-9}$ [43]
$\mathcal{B}(B_s \rightarrow \tau\tau)$	$< 6.8 \times 10^{-3}$ [44]
$\mathcal{B}(B_s \rightarrow e\mu)$	$< 6.3 \times 10^{-9}$ [45]
$\mathcal{B}(B_s \rightarrow \mu\tau)$	$< 4.2 \times 10^{-5}$ [46]
$\mathcal{B}(B_d \rightarrow ee)$	$< 3.0 \times 10^{-9}$ [42]
$\mathcal{B}(B_d \rightarrow \mu\mu)$	$< 2.6 \times 10^{-10}$ [43]
$\mathcal{B}(B_d \rightarrow \tau\tau)$	$< 2.1 \times 10^{-3}$ [44]
$\mathcal{B}(B_d \rightarrow e\mu)$	$< 1.3 \times 10^{-9}$ [45]
$\mathcal{B}(B_d \rightarrow \mu\tau)$	$< 1.4 \times 10^{-5}$ [46]

$\Delta F = 2$

Observable	Experimental value	SM prediction
ϵ_K	$(2.228 \pm 0.011) \times 10^{-3}$	$(2.14 \pm 0.12) \times 10^{-3}$
ΔM_s	$(17.765 \pm 0.006) \text{ ps}^{-1}$	$(17.35 \pm 0.94) \text{ ps}^{-1}$
ΔM_d	$(0.5065 \pm 0.0019) \text{ ps}^{-1}$	$(0.502 \pm 0.031) \text{ ps}^{-1}$

Leptonic

Observable	Experimental value	
	$\ell = e$	$\ell = \mu$
Δa_ℓ	$(2.8 \pm 7.4) \times 10^{-13}$	$(20.0 \pm 8.4) \times 10^{-10}$
$g_\tau/g_\ell - 1$	$(2.7 \pm 1.4) \times 10^{-3}$	$(0.9 \pm 1.4) \times 10^{-3}$

Kaon physics

Observable	Experimental value
$\mathcal{B}(K^+ \rightarrow \pi^+ \nu \bar{\nu})$	$(1.14_{-0.33}^{+0.4}) \times 10^{-10}$ [47] [48]
$\mathcal{B}(K_L \rightarrow \pi^0 \nu \bar{\nu})$	$< 3.6 \times 10^{-9}$ [49]
$\mathcal{B}(K_S \rightarrow \mu^+ \mu^-)$	$< 2.5 \times 10^{-10}$ [50]
$\mathcal{B}(K_L \rightarrow \mu^+ \mu^-)_{SD}$	$< 2.5 \times 10^{-9}$ [51]
$\mathcal{B}(K_L \rightarrow \mu^\pm e^\mp)$	$< 5.6 \times 10^{-12}$ [52]
$\mathcal{B}(K_L \rightarrow \pi^0 \mu^+ \mu^-)$	$< 4.5 \times 10^{-10}$ [53]
$\mathcal{B}(K_L \rightarrow \pi^0 e^+ e^-)$	$< 3.3 \times 10^{-10}$ [54]
$\mathcal{B}(K_L \rightarrow \pi^0 e^+ \mu^-)$	$< 9.1 \times 10^{-11}$ [55]
$\mathcal{B}(K^+ \rightarrow \pi^+ e^+ \mu^-)$	$< 7.9 \times 10^{-11}$ [56]

LFV

Observable	Experimental limit
$\mathcal{B}(\mu \rightarrow e \gamma)$	5.0×10^{-13} [102]
$\mathcal{B}(\mu \rightarrow 3e)$	1.2×10^{-12} [103]
$\mathcal{B}(\mu \text{ Au} \rightarrow e \text{ Au})$	8.3×10^{-13} [104]
$\mathcal{B}(\tau \rightarrow e \gamma)$	3.9×10^{-8} [105]
$\mathcal{B}(\tau \rightarrow 3e)$	3.2×10^{-8} [106]
$\mathcal{B}(\tau \rightarrow e \bar{\mu} \mu)$	3.2×10^{-8} [106]
$\mathcal{B}(\tau \rightarrow e \pi^0)$	9.5×10^{-8} [107]
$\mathcal{B}(\tau \rightarrow e \eta)$	1.1×10^{-7} [107]
$\mathcal{B}(\tau \rightarrow e \eta')$	1.9×10^{-7} [107]

Observable	Experimental limit
$\mathcal{B}(\tau \rightarrow e \pi^+ \pi^-)$	2.7×10^{-8} [108]
$\mathcal{B}(\tau \rightarrow e K^+ K^-)$	4.1×10^{-8} [108]
$\mathcal{B}(\tau \rightarrow \mu \gamma)$	5.0×10^{-8} [109]
$\mathcal{B}(\tau \rightarrow 3\mu)$	2.5×10^{-8} [106]
$\mathcal{B}(\tau \rightarrow \mu \bar{e} e)$	2.1×10^{-8} [106]
$\mathcal{B}(\tau \rightarrow \mu \pi^0)$	1.3×10^{-7} [110]
$\mathcal{B}(\tau \rightarrow \mu \eta)$	7.7×10^{-8} [107]
$\mathcal{B}(\tau \rightarrow \mu \eta')$	1.5×10^{-7} [107]
$\mathcal{B}(\tau \rightarrow \mu \pi^+ \pi^-)$	2.5×10^{-8} [108]
$\mathcal{B}(\tau \rightarrow \mu K^+ K^-)$	5.2×10^{-8} [108]

Cabibbo angle-related

Global analysis of Cabibbo-related observables by [Cirigliano et al. 2112.02087]

EW precision obs. + Higgs

[Falkowski et al. [1503.07872, 1911.07866]

One-parameter fits

What **scale** are we probing with **indirect** probes?

One-parameter fits from our global analysis of indirect constraints on top quark operators. In the third column we report the observable giving the dominant constraint in each case.

Wilson	Global fit [TeV ⁻²]	Dominant
$C_{qq}^{(+)}$	$(-1.9 \pm 2.3) \times 10^{-3}$	ΔM_s
$C_{qq}^{(-)}$	$(-2.0 \pm 1.0) \times 10^{-1}$	$B_s \rightarrow \mu\mu$
$C_{qu}^{(1)}$	$(1.3 \pm 1.0) \times 10^{-1}$	ΔM_s
$C_{qu}^{(8)}$	$(-1.7 \pm 4.4) \times 10^{-1}$	ΔM_s
C_{uu}	$(-3.0 \pm 1.7) \times 10^{-1}$	$\delta g_{L,11}^{Ze}$
$C_{Hq}^{(+)}$	$(18.7 \pm 8.8) \times 10^{-3}$	$B_s \rightarrow \mu\mu$
$C_{Hq}^{(-)}$	$(5.8 \pm 4.5) \times 10^{-2}$	$\delta g_{L,11}^{Ze}$
C_{Hu}	$(-4.3 \pm 2.3) \times 10^{-2}$	$\delta g_{L,11}^{Ze}$
C_{uB}	$(-0.6 \pm 2.0) \times 10^{-2}$	$c_{\gamma\gamma}$
C_{uG}	$(-0.1 \pm 2.0) \times 10^{-2}$	c_{gg}
C_{uH}	$(-0.3 \pm 5.2) \times 10^{-1}$	$C_{uH,33}$
C_{uW}	$(-0.1 \pm 3.1) \times 10^{-2}$	$c_{\gamma\gamma}$

Wilson	Global fit [TeV ⁻²]	Dominant
$C_{lq}^{(+),11}$	$(2.4 \pm 3.5) \times 10^{-3}$	R_K
$C_{lq}^{(+),22}$	$(-4.0 \pm 3.4) \times 10^{-3}$	R_K
$C_{lq}^{(+),33}$	$(7.2 \pm 4.4) \times 10^{-1}$	g_τ/g_i
$C_{lq}^{(-),11}$	$(10.9 \pm 7.6) \times 10^{-2}$	$R_{K^{(*)}}^\nu$
$C_{lq}^{(-),22}$	$(-6.0 \pm 7.0) \times 10^{-2}$	$R_{K^{(*)}}^\nu$
$C_{lq}^{(-),33}$	$(-1.8 \pm 1.0) \times 10^{-1}$	$R_{K^{(*)}}^\nu$
C_{lu}^{11}	$(-1.7 \pm 7.0) \times 10^{-2}$	$\delta g_{L,11}^{Ze}$
C_{lu}^{22}	$(-4.3 \pm 1.8) \times 10^{-1}$	$\delta g_{L,22}^{Ze}, R_K$
C_{lu}^{33}	$(0.5 \pm 2.4) \times 10^{-1}$	$\Delta g_{L,33}^{Ze}$
C_{qe}^{11}	$(-0.7 \pm 3.9) \times 10^{-2}$	R_{K^*}
C_{qe}^{22}	$(12.1 \pm 9.2) \times 10^{-3}$	$B_s \rightarrow \mu\mu$
C_{qe}^{33}	$(2.2 \pm 2.4) \times 10^{-1}$	$\delta g_{R,33}^{Ze}$

Wilson	Global fit [TeV ⁻²]	Dominant
C_{cu}^{11}	$(5.0 \pm 8.1) \times 10^{-2}$	$\Delta g_{R,11}^{Ze}$
C_{cu}^{22}	$(4.8 \pm 2.1) \times 10^{-1}$	$\Delta g_{R,22}^{Ze}$
C_{cu}^{33}	$(-2.3 \pm 2.5) \times 10^{-1}$	$\Delta g_{R,33}^{Ze}$
$C_{lequ}^{(1),11}$	$(0.4 \pm 1.0) \times 10^{-2}$	$(g-2)_e$
$C_{lequ}^{(1),22}$	$(1.8 \pm 1.6) \times 10^{-2}$	C_{eH22}
$C_{lequ}^{(1),33}$	$(8.0 \pm 9.1) \times 10^{-2}$	C_{eH33}
$C_{lequ}^{(3),11}$	$(-0.6 \pm 1.5) \times 10^{-5}$	$(g-2)_e$
$C_{lequ}^{(3),22}$	$(-19.3 \pm 8.1) \times 10^{-5}$	$(g-2)_\mu$
$C_{lequ}^{(3),33}$	$(-7.0 \pm 7.8) \times 10^{-1}$	C_{eH33}

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C_{uH}	$(-0.3 \pm 5.2) \times 10^{-1}$	$C_{uH,33}$
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$C_{lq}^{(+),33}$	$(7.2 \pm 4.4) \times 10^{-1}$	g_τ/g_i
$C_{lq}^{(-),11}$	$(10.9 \pm 7.6) \times 10^{-2}$	$R_{K^{(*)}}^\nu$
$C_{lq}^{(-),22}$	$(-6.0 \pm 7.0) \times 10^{-2}$	$R_{K^{(*)}}^\nu$
$C_{lq}^{(-),33}$	$(-1.8 \pm 1.0) \times 10^{-1}$	$R_{K^{(*)}}^\nu$
C_{lu}^{11}	$(-1.7 \pm 7.0) \times 10^{-2}$	$\delta g_{L,11}^{Ze}$
C_{lu}^{22}	$(-4.3 \pm 1.8) \times 10^{-1}$	$\delta g_{L,22}^{Ze}, R_K$
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$C_{lequ}^{(3),33}$	$(-7.0 \pm 7.8) \times 10^{-1}$	C_{eH33}

Indirect bounds are in the **few TeV range**.

Exception for tree-level contributions to **Bs-mixing**, **R_K**, **B_s→μμ**, and top-loop **from dipoles to (g-2)_{e,μ}**.

Λ ≳ 10 TeV

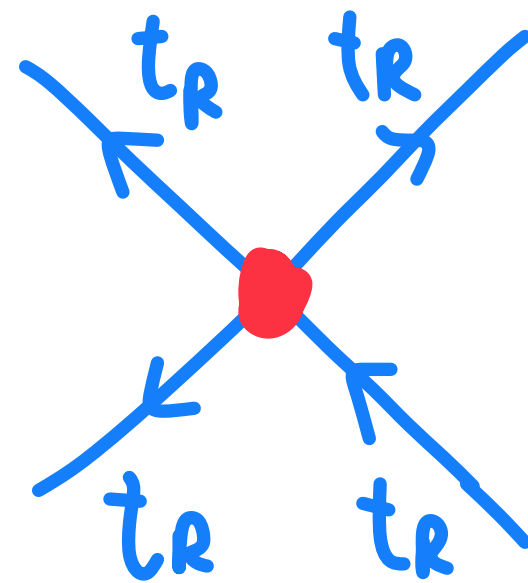
Λ ≳ 180 - 80 TeV

Example

Let us take for example this 4-top operator.
Its strongest bound is from LEP (Z-pole).

C_{uu}	$(-3.0 \pm 1.7) \times 10^{-1}$	$\delta g_{L,11}^{Ze}$
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How does it generate a contribution?



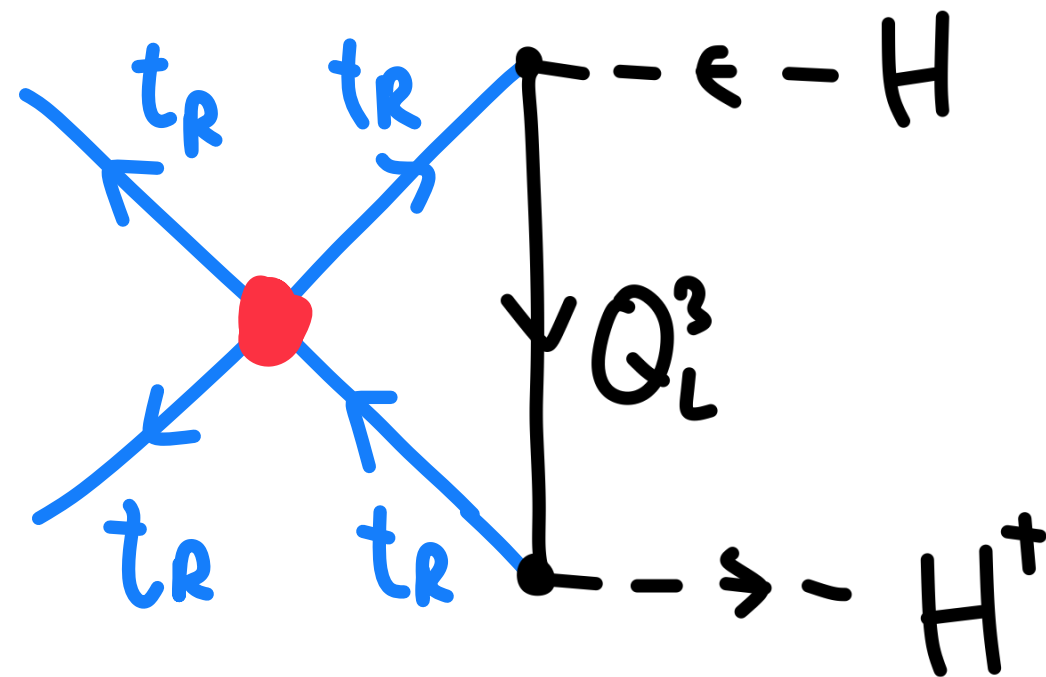
$$\mathcal{L}_{\text{SMEFT}} = C_{uu} (\bar{t}_R \gamma_\mu t_R) (\bar{t}_R \gamma^\mu t_R)$$

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How does it generate a contribution?



$$\mathcal{L}_{\text{SM\text{E}FT}} = C_{uu} (\bar{t}_R \gamma_\mu t_R) (\bar{t}_R \gamma^\mu t_R)$$

1-loop



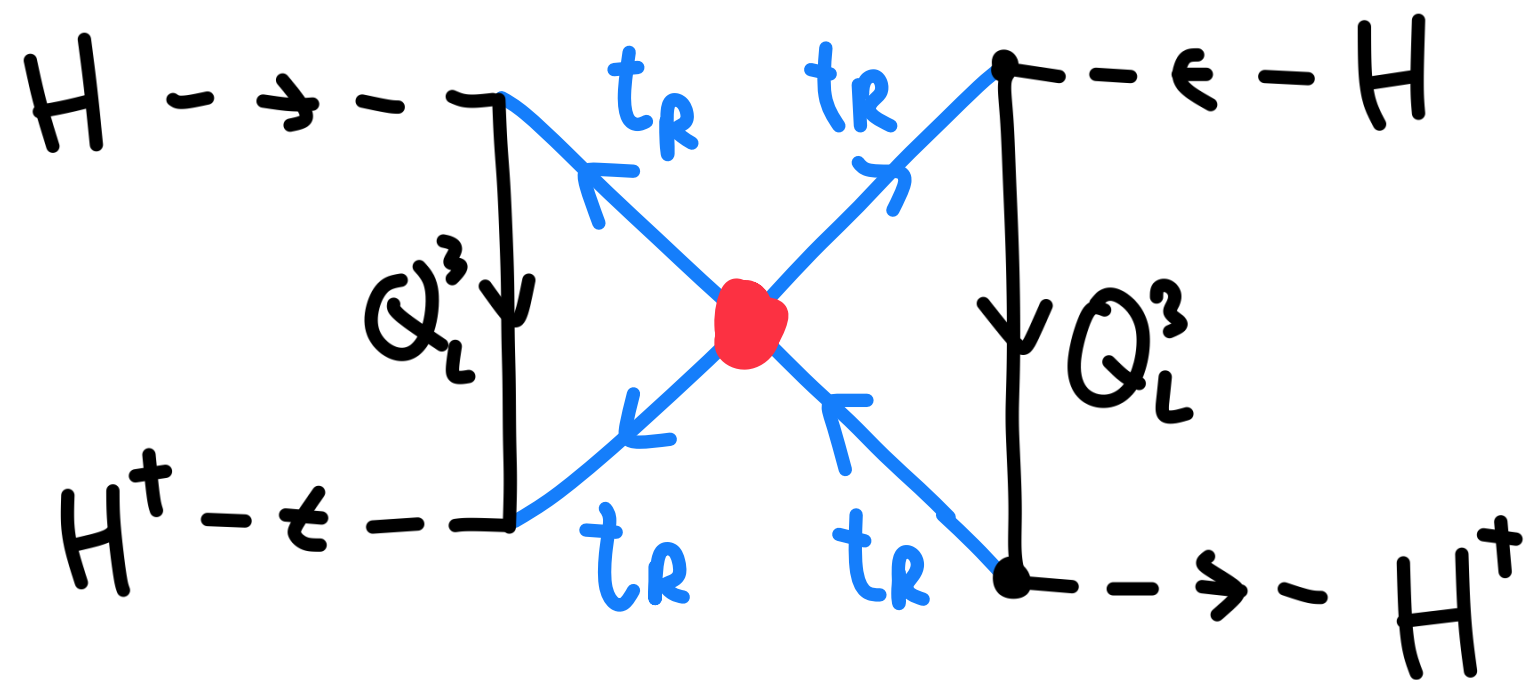
$$C_{uu} \frac{N_c \gamma_t^2}{(4\pi)^2} \log \frac{\Lambda^2}{m_t^2} (\bar{t}_R \gamma_\mu t_R) (H^\dagger \overleftrightarrow{D}^\mu H)$$

Example

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1-loop

$$C_{uu} \frac{N_c \gamma_t^2}{(4\pi)^2} \log \frac{\Lambda^2}{m_t^2} (\bar{t}_R \gamma_\mu t_R) (H^\dagger \overleftrightarrow{D}^\mu H)$$

2-loop (Leading Log)

$$C_{uu} \left(\frac{N_c \gamma_t^2}{(4\pi)^2} \log \frac{\Lambda^2}{m_t^2} \right)^2 (H^\dagger \overleftrightarrow{D}^\mu H)^2$$

This is the operator contributing to the **EW T-parameter** (Z-mass).
In the fit we used it shows up as a Z-coupling contribution.

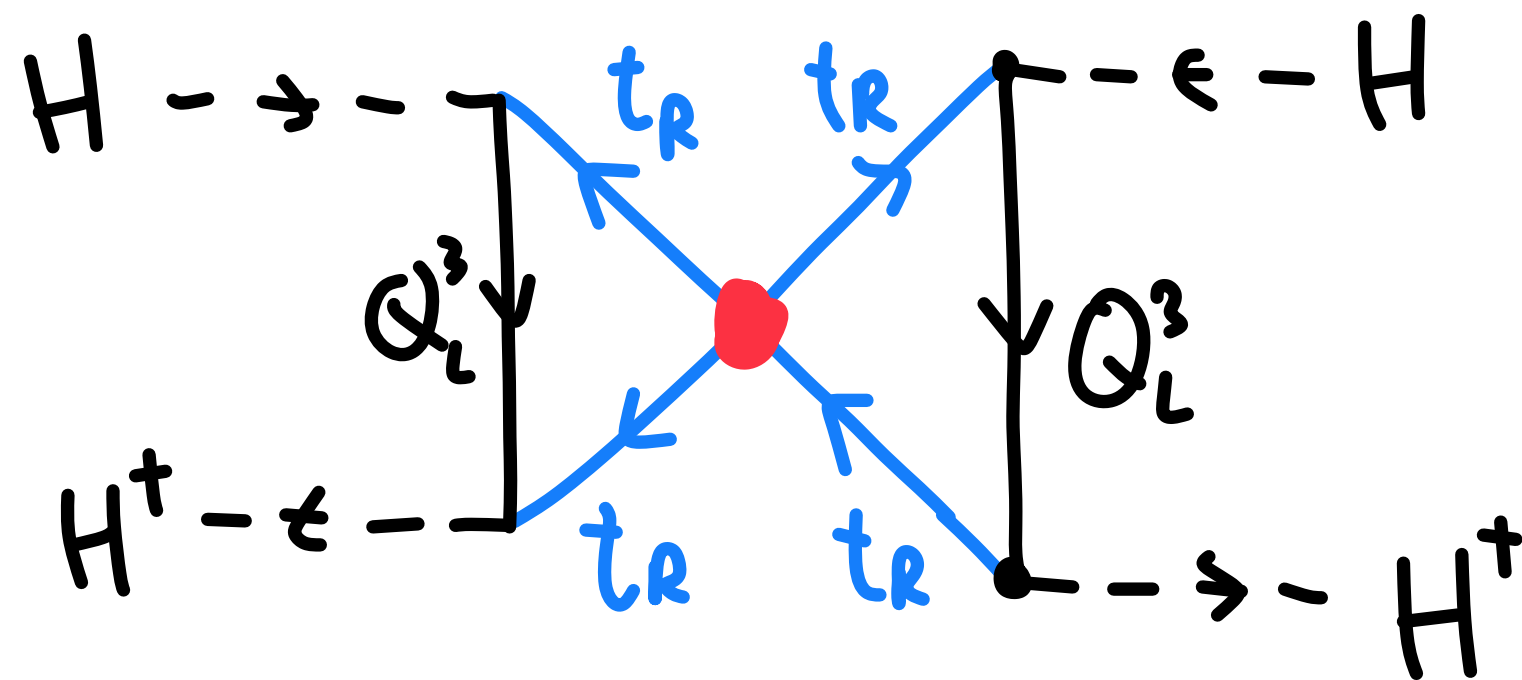
[1911.07866]

Example

Let us take for example this 4-top operator.
Its strongest bound is from LEP (Z-pole).

C_{uu}	$(-3.0 \pm 1.7) \times 10^{-1}$	$\delta g_{L,11}^{Ze}$
----------	---------------------------------	------------------------

How does it generate a contribution?



$$\mathcal{L}_{\text{SM\textt{EFT}}} = C_{uu} (\bar{t}_R \gamma_\mu t_R) (\bar{t}_R \gamma^\mu t_R)$$

1-loop

$$C_{uu} \frac{N_c Y_t^2}{(4\pi)^2} \log \frac{\Lambda^2}{m_t^2} (\bar{t}_R \gamma_\mu t_R) (H^\dagger \overleftrightarrow{D}^\mu H)$$

2-loop (Leading Log)

$$C_{uu} \left(\frac{N_c Y_t^2}{(4\pi)^2} \log \frac{\Lambda^2}{m_t^2} \right)^2 (H^\dagger \overleftrightarrow{D}^\mu H)^2$$

This is the operator contributing to the **EW T-parameter** (Z-mass).
In the fit we used it shows up as a Z-coupling contribution.

[1911.07866]

Direct constraints from LHC

Top operators can be constrained directly from LHC measurements of top quark processes:

SMEFIT 2105.00006

$$pp \rightarrow t\bar{t}$$

$$pp \rightarrow t\bar{t}t\bar{t}$$

$$pp \rightarrow t\bar{t}b\bar{b}$$

$$pp \rightarrow t\bar{t}z$$

$$pp \rightarrow t\bar{t}W$$

$$pp \rightarrow t\bar{t}H$$

$$pp \rightarrow t + X$$

$$pp \rightarrow t\bar{t} + X$$

$$pp \rightarrow tW + X$$

$$pp \rightarrow tH + X$$

Higgs physics
ggF, VBF, Vh, etc..

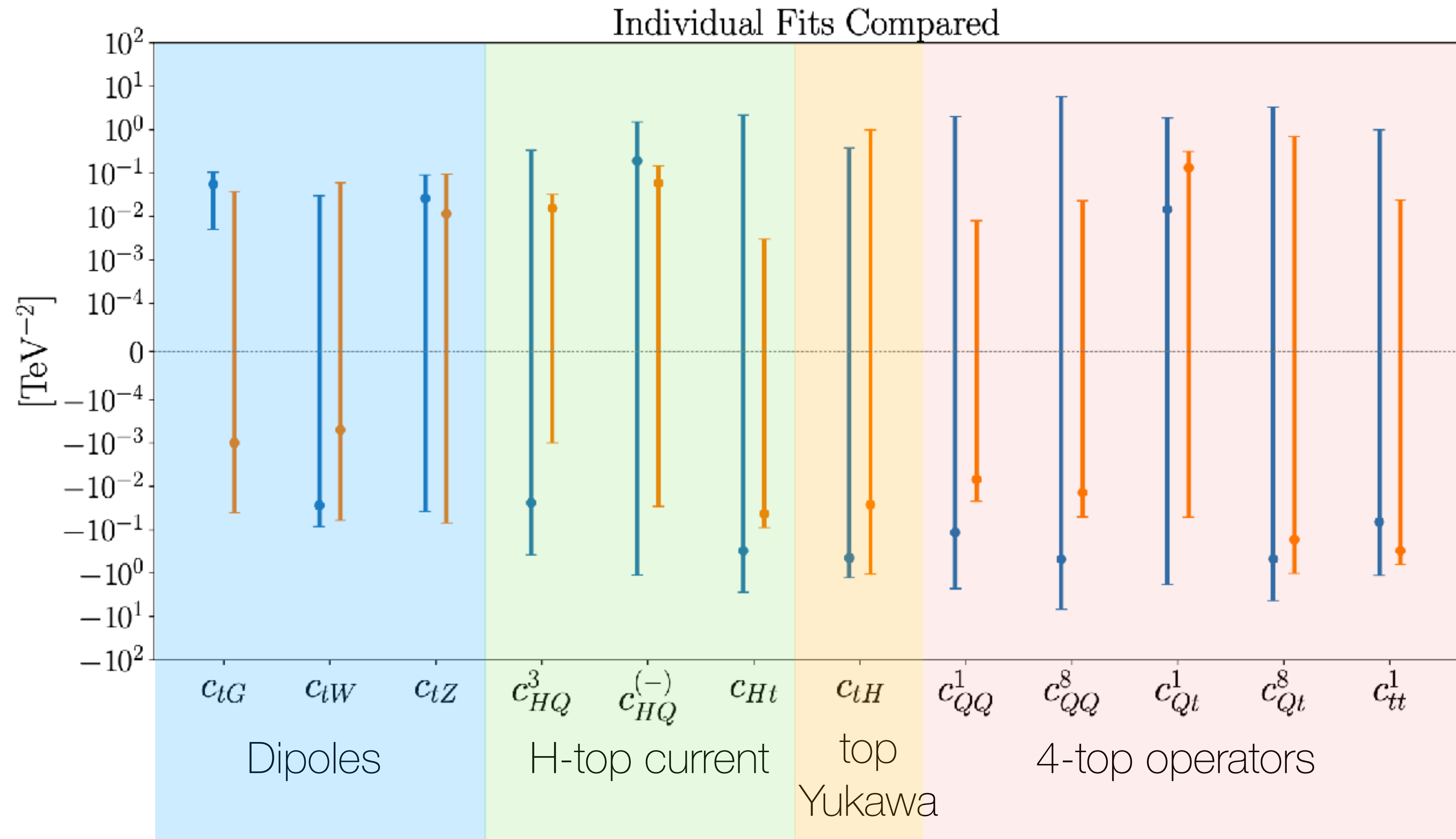
Indirect vs. Direct

How **direct bounds** compare with **indirect** ones? **Indirect are typically much stronger.**

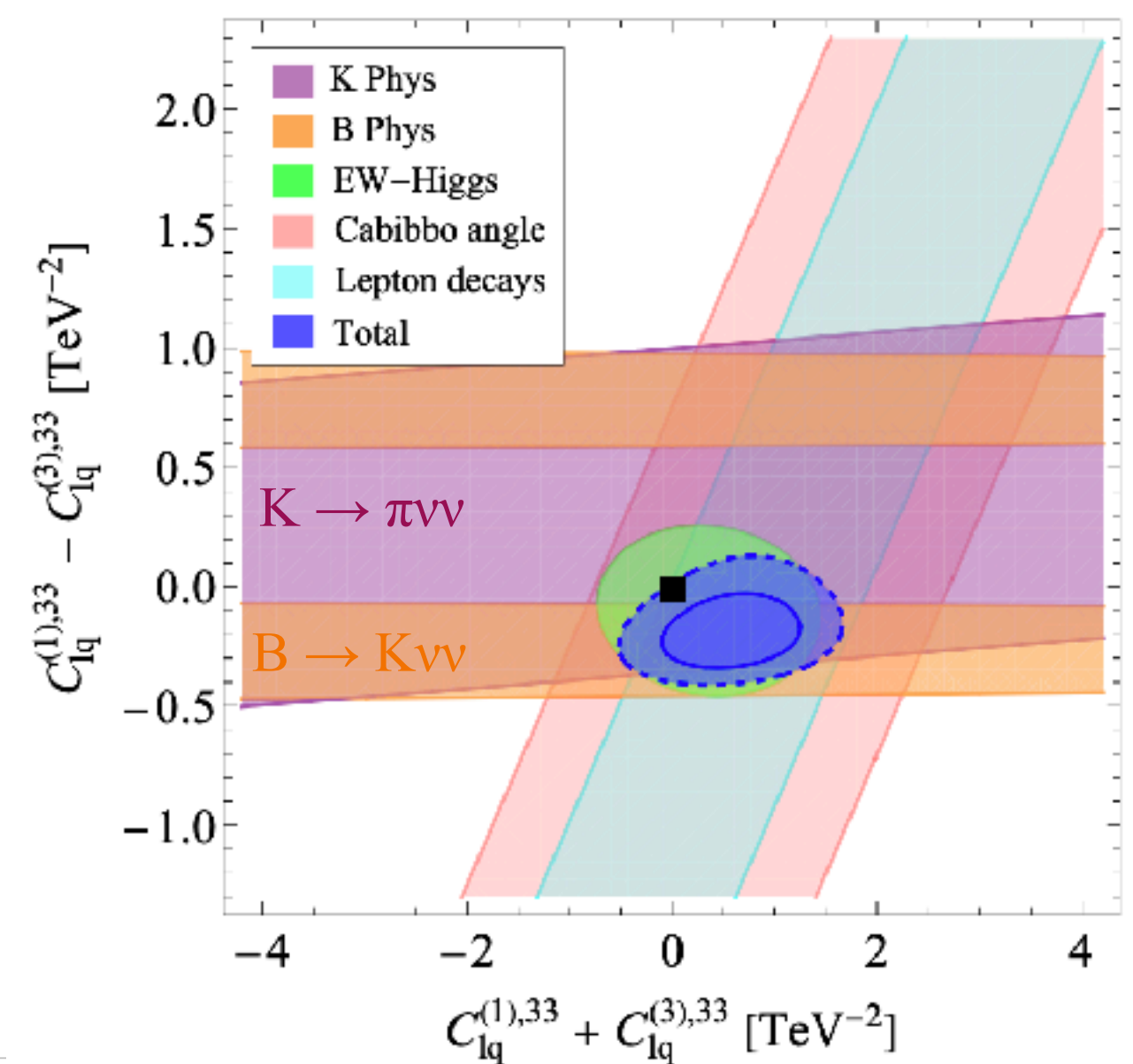
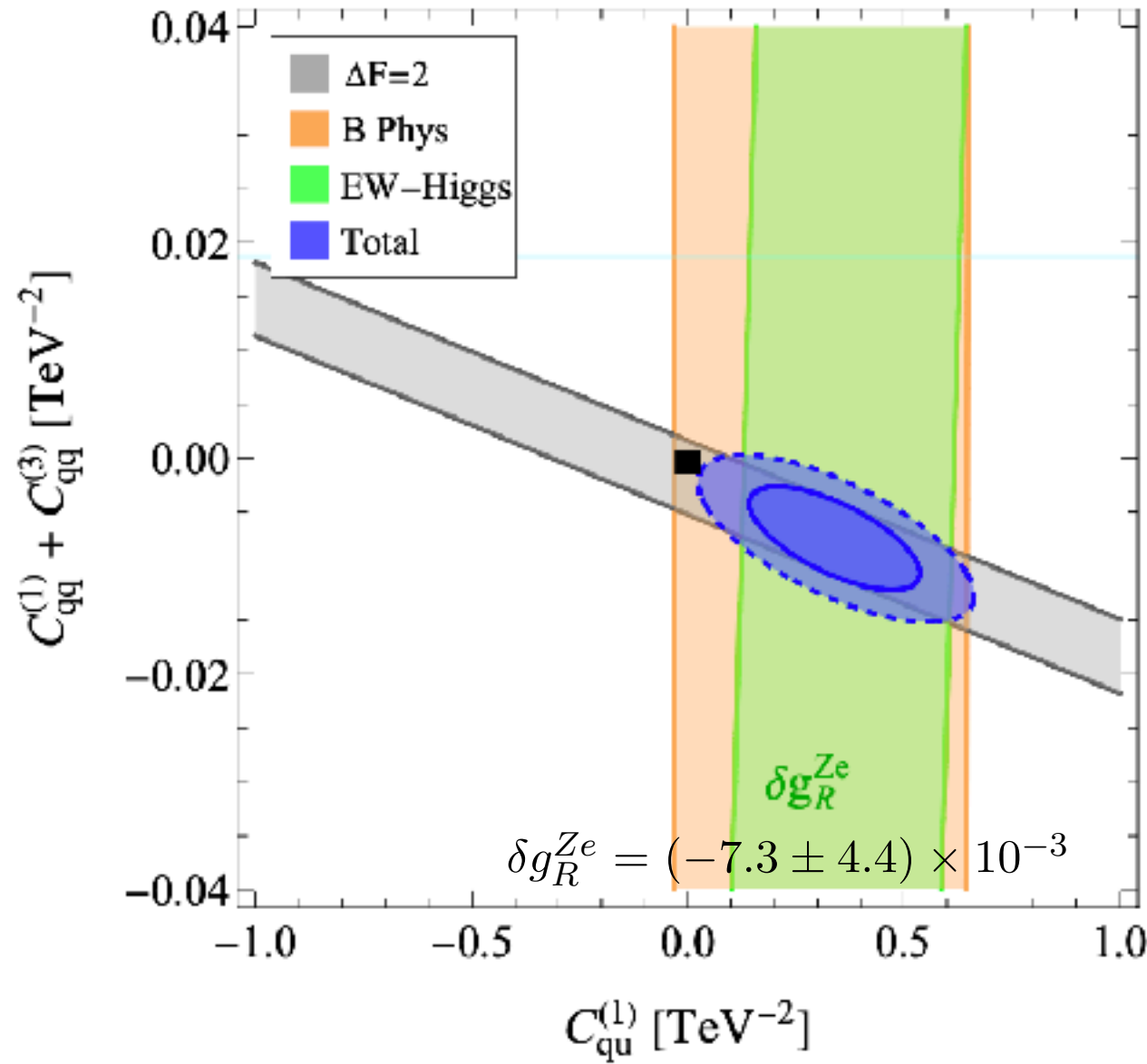
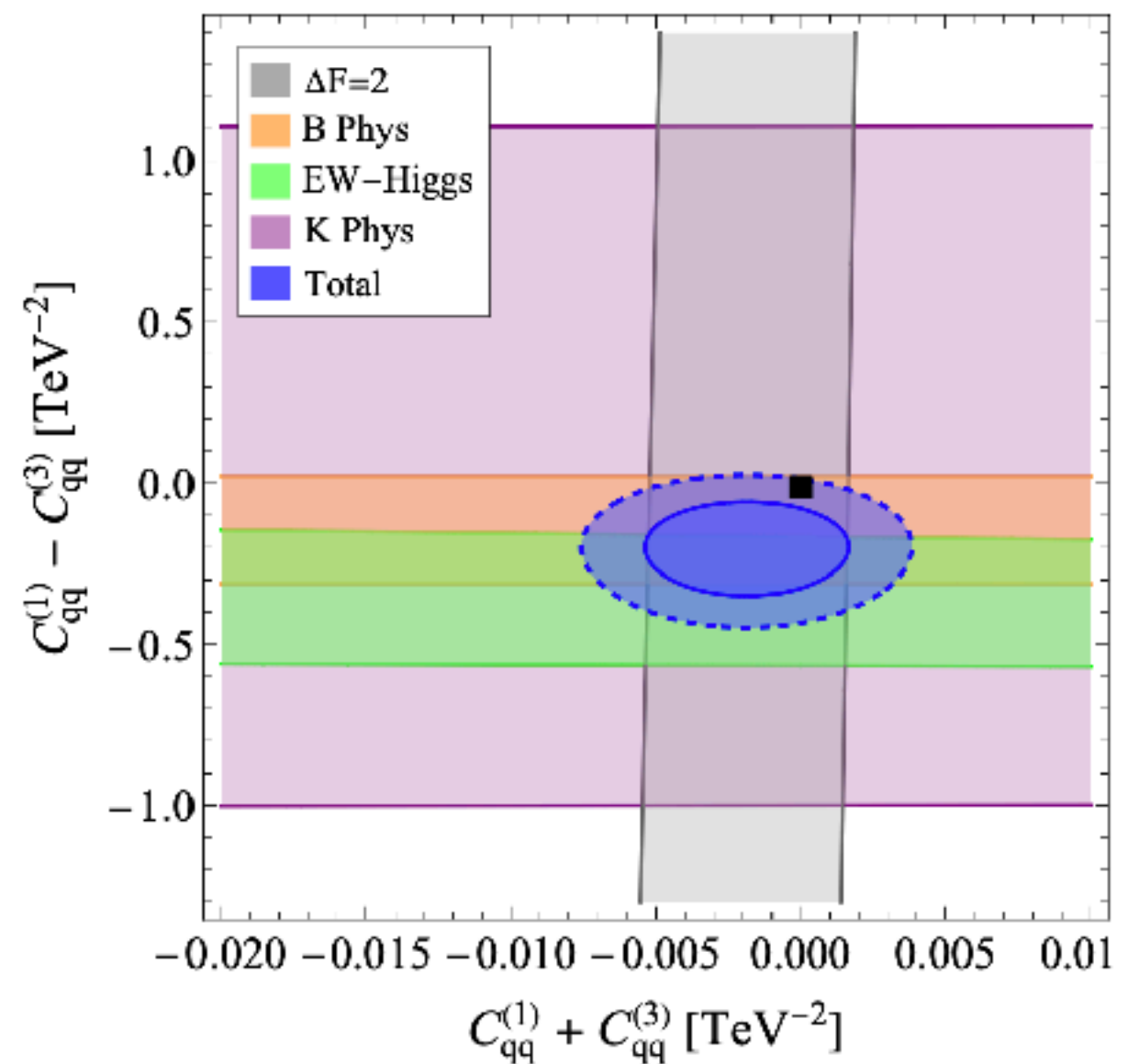
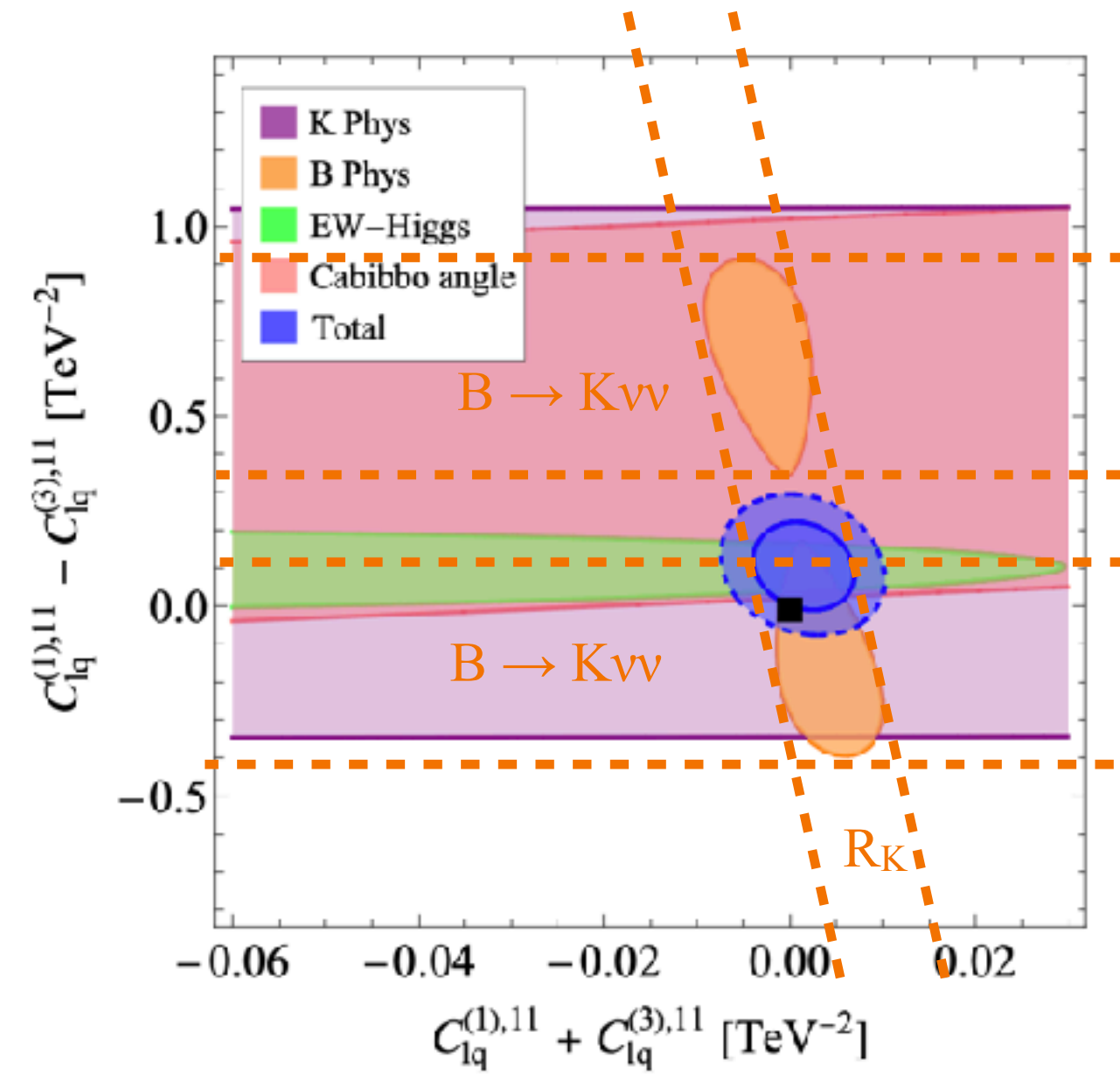
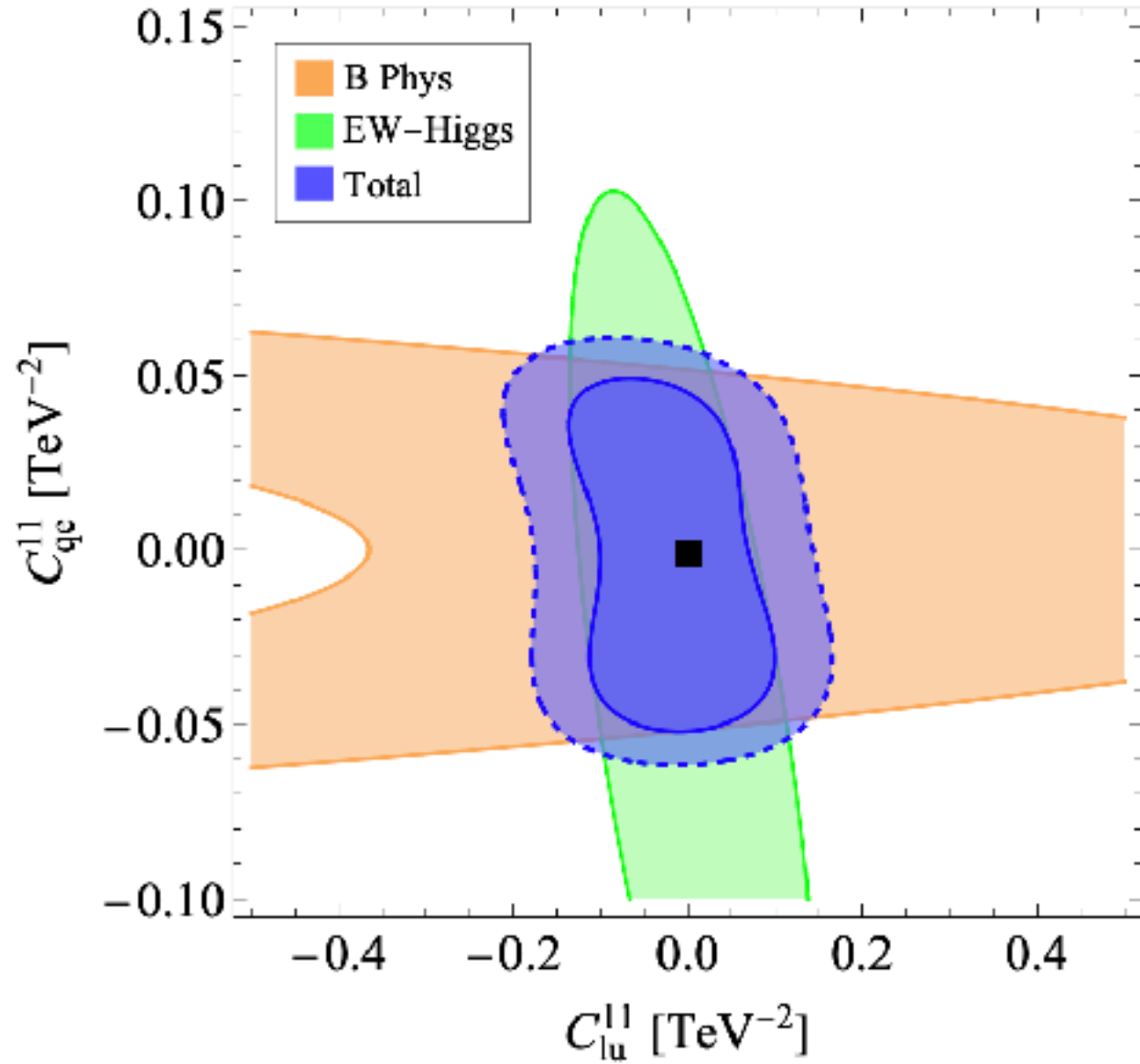
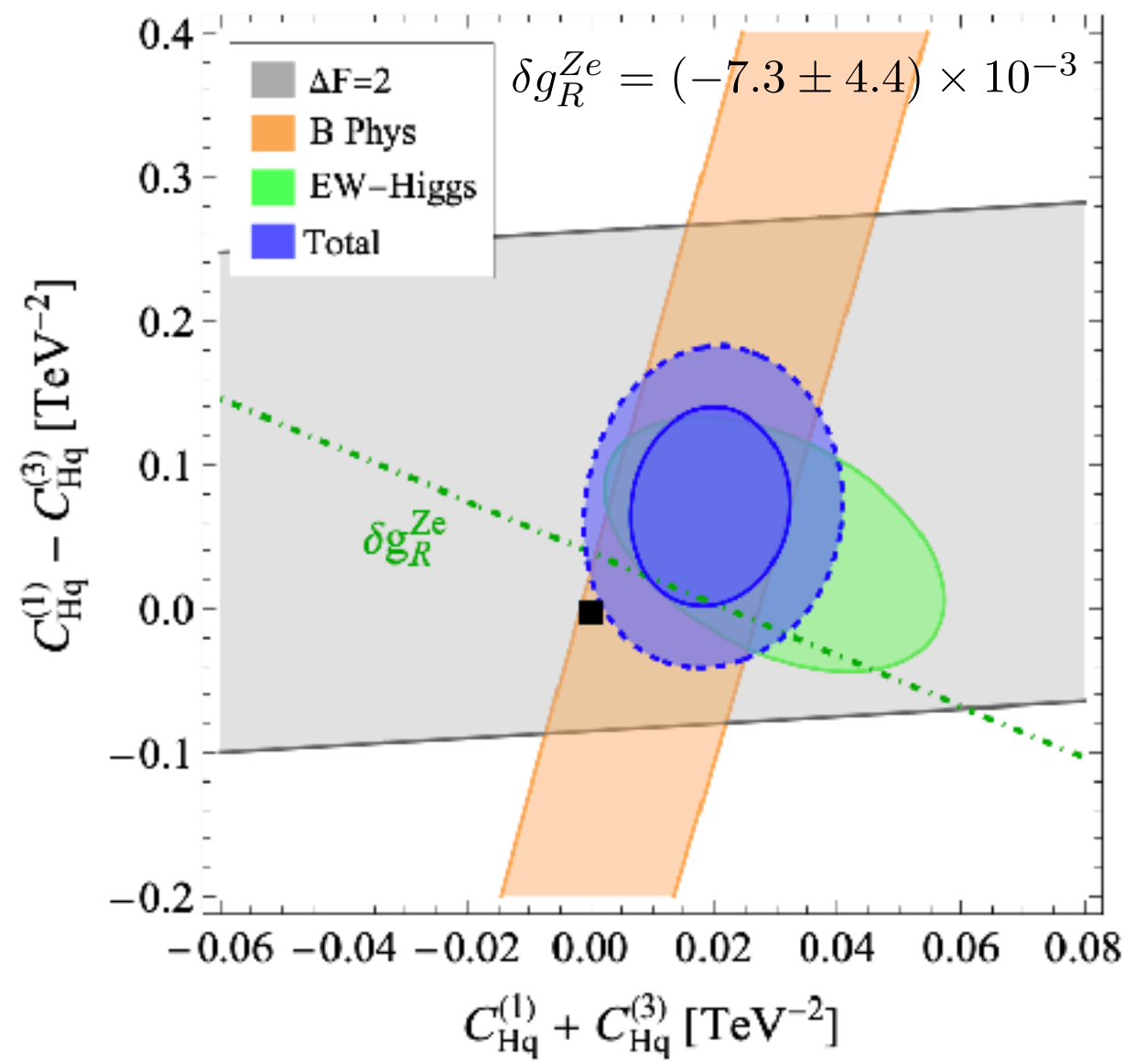
[Garosi, DM, Rodriguez-Sanchez, Stanzione 2310.00047]


 Direct bounds from LHC
**top production,
 Higgs physics,
 diboson production**
SMEFiT 2105.00006


 Indirect bounds from
 EW+Flavour
 2310.00047



2D fits



Combining bounds from different datasets allows to derive **much stronger constraints**.

Conclusions

If **New Physics** is present at a scale reachable by present or (~near) future colliders, then it must enjoy some **non-trivial flavour structure** that suppresses large FCNC effects.

Typically this tends to **align it close it to the 3rd generation**.

With a **Rank-One Flavour Violation** setup we show how close to the top direction this should be, in the case of fitting the $B \rightarrow K_w$ excess: **only deviations of $\approx O(\text{CKM})$** are allowed.

Correlations between different observables are crucial to identify the flavour structure.

Assuming **New Physics couples mostly to the top quark**, we show that **indirect bounds provide almost always stronger constraints than direct bounds** from LHC: also here it is crucial to combine different datasets.

Conclusions

If **New Physics** is present at a scale reachable by present or (~near) future colliders, then it must enjoy some **non-trivial flavour structure** that suppresses large FCNC effects.

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With a **Rank-One Flavour Violation** setup we show how close to the top direction this should be, in the case of fitting the $B \rightarrow K_w$ excess: **only deviations of $\approx O(\text{CKM})$** are allowed.

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Assuming **New Physics couples mostly to the top quark**, we show that **indirect bounds provide almost always stronger constraints than direct bounds** from LHC: also here it is crucial to combine different datasets.

Thank you!

Backup

Model	Signature	$\int \mathcal{L} dt$ [fb ⁻¹]	Mass limit	Reference	
Inclusive Searches	$q\bar{q}, \bar{q} \rightarrow q\tilde{\chi}_1^0$	0 e, μ mono-jet	2-6 jets 1-3 jets E_T^{miss} 140	\tilde{q} [1x, 8x Degen.] 1.0 \tilde{q} [8x Degen.] 0.9	$m(\tilde{\chi}_1^0) < 400$ GeV $m(\tilde{q}) - m(\tilde{\chi}_1^0) = 5$ GeV 2010.14293 2102.10874
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}\tilde{\chi}_1^0$	0 e, μ	2-6 jets E_T^{miss} 140	\tilde{g} 2.3 Forbidden 1.15-1.95	$m(\tilde{\chi}_1^0) = 0$ GeV $m(\tilde{\chi}_1^0) = 1000$ GeV 2010.14293 2010.14293
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}W\tilde{\chi}_1^0$	1 e, μ	2-6 jets E_T^{miss} 140	\tilde{g} 2.2	$m(\tilde{\chi}_1^0) < 600$ GeV 2101.01629
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}(\ell\ell)\tilde{\chi}_1^0$	$ee, \mu\mu$	2 jets E_T^{miss} 140	\tilde{g} 2.2	$m(\tilde{\chi}_1^0) < 700$ GeV 2204.13072
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow qqWZ\tilde{\chi}_1^0$	0 e, μ	7-11 jets E_T^{miss} 140	\tilde{g} 1.97	$m(\tilde{\chi}_1^0) < 600$ GeV 2008.06032
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow qqWZ\tilde{\chi}_1^0$	SS e, μ	6 jets E_T^{miss} 140	\tilde{g} 1.15	$m(\tilde{g}) - m(\tilde{\chi}_1^0) = 200$ GeV 2307.01094
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\bar{t}\tilde{\chi}_1^0$	0-1 e, μ	3 b E_T^{miss} 140	\tilde{g} 2.45	$m(\tilde{\chi}_1^0) < 500$ GeV 2211.08028
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\bar{t}\tilde{\chi}_1^0$	SS e, μ	6 jets E_T^{miss} 140	\tilde{g} 1.25	$m(\tilde{g}) - m(\tilde{\chi}_1^0) = 300$ GeV 1909.08457
3 rd gen. squarks direct production	$\tilde{b}_1\tilde{b}_1$	0 e, μ	2 b E_T^{miss} 140	\tilde{b}_1 1.255 \tilde{b}_1 0.68	$m(\tilde{\chi}_1^0) < 400$ GeV 10 GeV $< \Delta m(\tilde{b}_1, \tilde{\chi}_1^0) < 20$ GeV 2101.12527 2101.12527
	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow b\tilde{\chi}_2^0 \rightarrow bh\tilde{\chi}_1^0$	0 e, μ 2 τ	6 b 2 b E_T^{miss} 140	Forbidden 0.23-1.35 \tilde{b}_1 0.13-0.85	$\Delta m(\tilde{\chi}_2^0, \tilde{\chi}_1^0) = 130$ GeV, $m(\tilde{\chi}_1^0) = 100$ GeV $\Delta m(\tilde{\chi}_2^0, \tilde{\chi}_1^0) = 130$ GeV, $m(\tilde{\chi}_1^0) = 0$ GeV 1908.03122 2103.08189
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow t\tilde{\chi}_1^0$	0-1 e, μ	≥ 1 jet E_T^{miss} 140	\tilde{t}_1 1.25	$m(\tilde{\chi}_1^0) = 1$ GeV 2004.14050, 2012.03799
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow Wb\tilde{\chi}_1^0$	1 e, μ	3 jets/1 b E_T^{miss} 140	Forbidden 1.05	$m(\tilde{\chi}_1^0) = 500$ GeV 2012.03799, 2401.13430
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow \tilde{\tau}_1 b\nu, \tilde{\tau}_1 \rightarrow \tau\tilde{G}$	1-2 τ	2 jets/1 b E_T^{miss} 140	Forbidden 1.4	$m(\tilde{\tau}_1) = 800$ GeV 2108.07665
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow c\tilde{\chi}_1^0 / \tilde{c}\tilde{c}, \tilde{c} \rightarrow c\tilde{\chi}_1^0$	0 e, μ	2 c E_T^{miss} 36.1	\tilde{c} 0.85	$m(\tilde{\chi}_1^0) = 0$ GeV 1805.01649
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow c\tilde{\chi}_1^0$	0 e, μ	mono-jet E_T^{miss} 140	\tilde{t}_1 0.55	$m(\tilde{t}_1, \tilde{c}) - m(\tilde{\chi}_1^0) = 5$ GeV 2102.10874
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow t\tilde{\chi}_2^0, \tilde{\chi}_2^0 \rightarrow Z/h\tilde{\chi}_1^0$	1-2 e, μ	1-4 b E_T^{miss} 140	\tilde{t}_1 0.067-1.18	$m(\tilde{\chi}_2^0) = 500$ GeV 2006.05880
$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + Z$	3 e, μ	1 b E_T^{miss} 140	Forbidden 0.86	$m(\tilde{\chi}_1^0) = 360$ GeV, $m(\tilde{t}_1) - m(\tilde{\chi}_1^0) = 40$ GeV 2006.05880	
EW direct	$\tilde{\chi}_1^+\tilde{\chi}_2^0$ via WZ	Multiple l /jets $ee, \mu\mu$	≥ 1 jet E_T^{miss} 140	$\tilde{\chi}_1^+/\tilde{\chi}_2^0$ 0.96 $\tilde{\chi}_1^+/\tilde{\chi}_2^0$ 0.205	$m(\tilde{\chi}_1^0) = 0$, wino-bino $m(\tilde{\chi}_1^+) - m(\tilde{\chi}_2^0) = 5$ GeV, wino-bino 2106.01676, 2103.07586 1911.12606
	$\tilde{\chi}_1^+\tilde{\chi}_1^0$ via WW	2 e, μ	E_T^{miss} 140	$\tilde{\chi}_1^+$ 0.42	$m(\tilde{\chi}_1^0) = 0$, wino-bino 1908.08215
	$\tilde{\chi}_1^+\tilde{\chi}_2^0$ via Wh	Multiple l /jets	E_T^{miss} 140	Forbidden 1.06	$m(\tilde{\chi}_1^0) = 70$ GeV, wino-bino 2004.10894, 2103.07586
	$\tilde{\chi}_1^+\tilde{\chi}_1^0$ via $\tilde{l}_L/\tilde{\nu}$	2 e, μ	E_T^{miss} 140	$\tilde{\chi}_1^+$ 1.0	$m(\tilde{l}, \tilde{\nu}) = 0.5(m(\tilde{\chi}_1^+) + m(\tilde{\chi}_1^0))$ 1908.08215
	$\tilde{\tau}\tilde{\tau}, \tilde{\tau} \rightarrow \tau\tilde{\chi}_1^0$	2 τ	E_T^{miss} 140	$\tilde{\tau}$ ($\tilde{\tau}_R, \tilde{\tau}_L$) 0.35 0.5	$m(\tilde{\chi}_1^0) = 0$ 2402.00603
	$\tilde{\ell}_{L,R}, \tilde{\ell}_{L,R}, \tilde{\ell} \rightarrow \ell\tilde{\chi}_1^0$	2 e, μ $ee, \mu\mu$	0 jets ≥ 1 jet E_T^{miss} 140	$\tilde{\ell}$ 0.7 $\tilde{\ell}$ 0.26	$m(\tilde{\chi}_1^0) = 0$ $m(\tilde{\ell}) - m(\tilde{\chi}_1^0) = 10$ GeV 1908.08215 1911.12606
	$\tilde{H}\tilde{H}, \tilde{H} \rightarrow h\tilde{G}/Z\tilde{G}$	0 e, μ 4 e, μ 0 e, μ	≥ 3 b 0 jets ≥ 2 large jets E_T^{miss} 140	\tilde{H} 0.94 \tilde{H} 0.55 \tilde{H} 0.45-0.93	$\text{BR}(\tilde{\chi}_1^0 \rightarrow h\tilde{G}) = 1$ $\text{BR}(\tilde{\chi}_1^0 \rightarrow Z\tilde{G}) = 1$ $\text{BR}(\tilde{\chi}_1^0 \rightarrow Z\tilde{G}) = 1$ 2401.14922 2103.11684 2108.07506
	$\tilde{H}\tilde{H}, \tilde{H} \rightarrow h\tilde{G}/Z\tilde{G}$	2 e, μ	≥ 2 jets E_T^{miss} 140	\tilde{H} 0.77	$\text{BR}(\tilde{\chi}_1^0 \rightarrow Z\tilde{G}) = \text{BR}(\tilde{\chi}_1^0 \rightarrow h\tilde{G}) = 0.5$ 2204.13072
Long-lived particles	Direct $\tilde{\chi}_1^+\tilde{\chi}_1^-$ prod., long-lived $\tilde{\chi}_1^\pm$	Disapp. trk	1 jet E_T^{miss} 140	$\tilde{\chi}_1^\pm$ 0.66 $\tilde{\chi}_1^\pm$ 0.21	Pure Wino Pure higgsino 2201.02472 2201.02472
	Stable \tilde{g} R-hadron	pixel dE/dx	E_T^{miss} 140	\tilde{g} 2.05	2205.06013
	Metastable \tilde{g} R-hadron, $\tilde{g} \rightarrow qq\tilde{\chi}_1^0$	pixel dE/dx	E_T^{miss} 140	\tilde{g} [$\tau(\tilde{g}) = 10$ ns] 2.2	$m(\tilde{\chi}_1^0) = 100$ GeV 2205.06013
	$\tilde{\ell}, \tilde{\ell} \rightarrow \ell\tilde{G}$	Displ. lep	E_T^{miss} 140	$\tilde{\ell}, \tilde{\mu}$ 0.74 $\tilde{\tau}$ 0.36 $\tilde{\tau}$ 0.36	$\tau(\tilde{\ell}) = 0.1$ ns $\tau(\tilde{\ell}) = 0.1$ ns $\tau(\tilde{\ell}) = 10$ ns ATLAS-CONF-2024-011 ATLAS-CONF-2024-011 2205.05013
RPV	$\tilde{\chi}_1^+\tilde{\chi}_1^-/\tilde{\chi}_1^0, \tilde{\chi}_1^\pm \rightarrow Zl \rightarrow \ell\ell\ell$	3 e, μ	140	$\tilde{\chi}_1^\pm/\tilde{\chi}_1^0$ [BR(Z τ)=1, BR(Z e)=1] 0.625 1.05	Pure Wino 2011.10543
	$\tilde{\chi}_1^+\tilde{\chi}_1^-/\tilde{\chi}_2^0 \rightarrow WW/Z\ell\ell\nu\nu$	4 e, μ	0 jets E_T^{miss} 140	$\tilde{\chi}_1^\pm/\tilde{\chi}_2^0$ [$\lambda_{33} \neq 0, \lambda_{12} \neq 0$] 0.95 1.55	$m(\tilde{\chi}_1^0) = 200$ GeV 2103.11684
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow qq\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow qqq$	≥ 8 jets	140	\tilde{g} [$m(\tilde{\chi}_1^0) = 50$ GeV, 1250 GeV] 1.6 2.34	Large λ'_{112} 2401.16333
	$\tilde{u}, \tilde{t} \rightarrow t\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow tbs$	Multiple	36.1	\tilde{t} [$\lambda'_{323} = 2e-4, 1e-2$] 0.55 1.05	$m(\tilde{\chi}_1^0) = 200$ GeV, bino-like ATLAS-CONF-2018-003
	$\tilde{u}, \tilde{t} \rightarrow b\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow bbs$	$\geq 4b$	140	\tilde{t} Forbidden 0.95	$m(\tilde{\chi}_1^+) = 500$ GeV 2010.01015
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow bs$	2 jets + 2 b	36.7	\tilde{t}_1 [qq, bs] 0.42 0.61	1710.07171
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow q\ell$	2 e, μ	2 b 140	\tilde{t}_1 0.4-1.85	$\text{BR}(\tilde{t}_1 \rightarrow b\ell/b\mu) > 20\%$ 2406.18367
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow q\ell$	1 μ	DV 136	\tilde{t}_1 [$1e-10 < \lambda'_{234} < 1e-8, 3e-10 < \lambda'_{234} < 3e-5$] 1.0 1.6	$\text{BR}(\tilde{t}_1 \rightarrow q\mu) = 100\%$, $\cos\theta_1 = 1$ 2003.11956
$\tilde{\chi}_1^+/\tilde{\chi}_2^0/\tilde{\chi}_1^0, \tilde{\chi}_{1,2}^0 \rightarrow tbs, \tilde{\chi}_1^+ \rightarrow bbs$	1-2 e, μ	≥ 6 jets 140	$\tilde{\chi}_1^0$ 0.2-0.32	Pure higgsino 2106.09609	

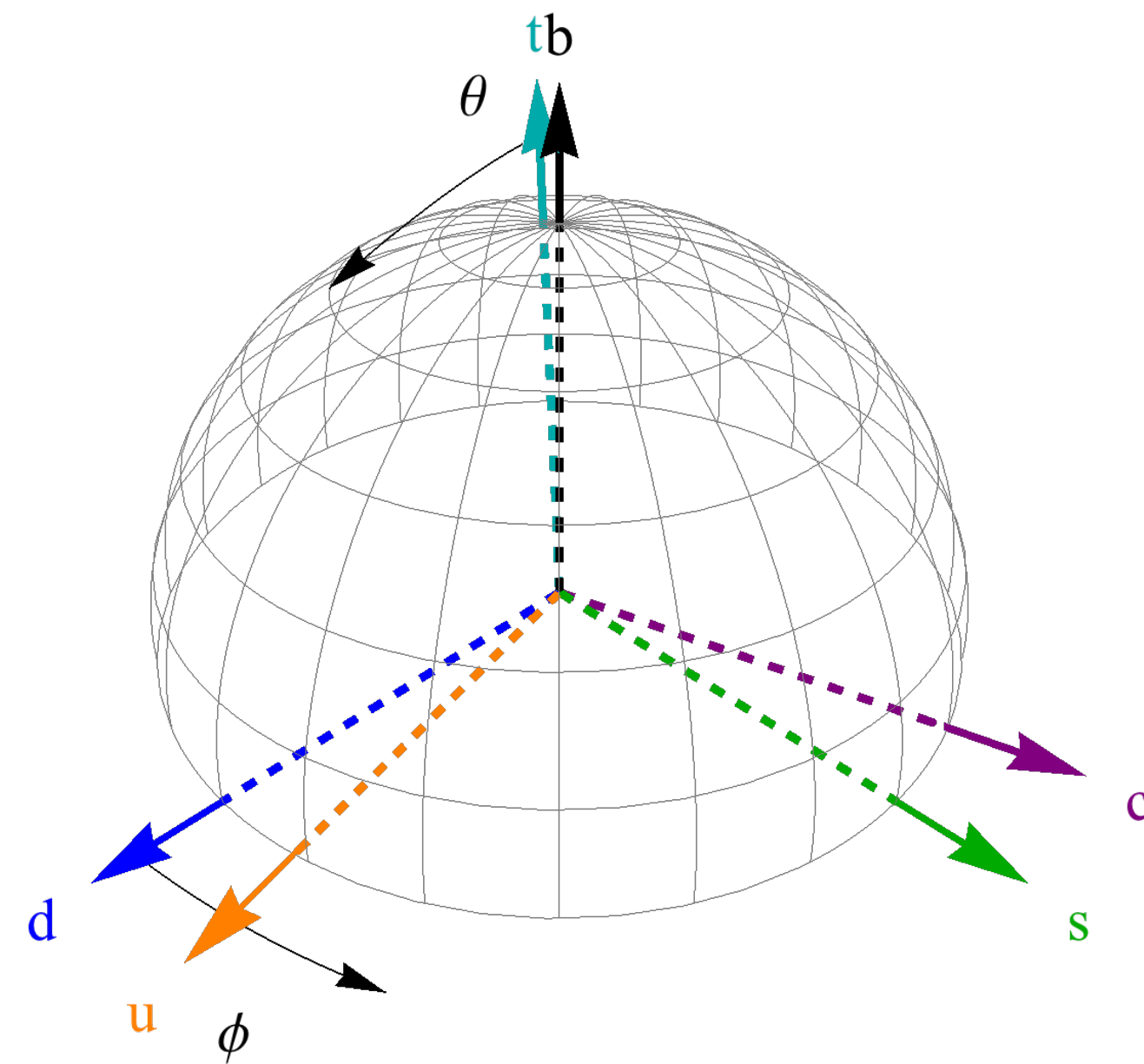
*Only a selection of the available mass limits on new states or phenomena is shown. Many of the limits are based on simplified models, c.f. refs. for the assumptions made.

10⁻¹ 1 Mass scale [TeV]

Directions in Flavour Space

Gherardi, DM, Nardecchia, Romanino [1903.10954](#)

DM, Nardecchia, Stanzione, Toni [2404.06533](#)



$\{q_L^i\}$ space, neglecting phases

$$\hat{n} = \begin{pmatrix} \sin \theta \cos \phi e^{i\alpha_{bd}} \\ \sin \theta \sin \phi e^{i\alpha_{bs}} \\ \cos \theta \end{pmatrix}$$

$$\theta \in \left[0, \frac{\pi}{2}\right], \quad \phi \in [0, 2\pi), \quad \alpha_{bd} \in \left[-\frac{\pi}{2}, \frac{\pi}{2}\right], \quad \alpha_{bs} \in \left[-\frac{\pi}{2}, \frac{\pi}{2}\right]$$

$$q_L^i = \begin{pmatrix} V_{ji}^* u_L^i \\ d_L^i \end{pmatrix}$$

quark	\hat{n}	ϕ	θ	α_{bd}	α_{bs}
down	(1, 0, 0)	0	$\pi/2$	0	0
strange	(0, 1, 0)	$\pi/2$	$\pi/2$	0	0
bottom	(0, 0, 1)	0	0	0	0
up	$e^{i \arg(V_{ub})} (V_{ud}^*, V_{us}^*, V_{ub}^*)$	0.23	1.57	-1.17	-1.17
charm	$e^{i \arg(V_{cb})} (V_{cd}^*, V_{cs}^*, V_{cb}^*)$	1.80	1.53	-6.2×10^{-4}	-3.3×10^{-5}
top	$e^{i \arg(V_{tb})} (V_{td}^*, V_{ts}^*, V_{tb}^*)$	4.92	0.042	-0.018	0.39

The **misalignment** between **down-** and **up-quarks** is described by the **CKM matrix**.

SMEFT

Since the **scale of New Physics is ~TeV**, the contribution could come from heavy New Physics: **SMEFT**.

Possible tree-level contributions from the following SMEFT dim-6 operators:

$$\begin{aligned}
 \mathcal{O}_{lq}^{(1)\alpha\beta ij} &= \left(\bar{l}_L^\alpha \gamma_\mu l_L^\beta \right) \left(\bar{q}_L^i \gamma^\mu q_L^j \right), & \mathcal{O}_{Hq}^{(1)ij} &= \left(H^\dagger \overleftrightarrow{D}_\mu H \right) \left(\bar{q}_L^i \gamma^\mu q_L^j \right), \\
 \mathcal{O}_{lq}^{(3)\alpha\beta ij} &= \left(\bar{l}_L^\alpha \gamma_\mu \sigma_a l_L^\beta \right) \left(\bar{q}_L^i \gamma^\mu \sigma_a q_L^j \right), & \mathcal{O}_{Hq}^{(3)ij} &= \left(H^\dagger \sigma_a \overleftrightarrow{D}_\mu H \right) \left(\bar{q}_L^i \gamma^\mu \sigma_a q_L^j \right), \\
 \mathcal{O}_{ld}^{\alpha\beta ij} &= \left(\bar{l}_L^\alpha \gamma_\mu l_L^\beta \right) \left(\bar{d}_R^i \gamma^\mu d_R^j \right), & \mathcal{O}_{Hd}^{ij} &= \left(H^\dagger \overleftrightarrow{D}_\mu H \right) \left(\bar{d}_R^i \gamma^\mu d_R^j \right).
 \end{aligned}$$

$$\begin{aligned}
 L_L^{ij\alpha\beta} &= C_{lq}^{(1)\alpha\beta ij} - C_{lq}^{(3)\alpha\beta ij} + C_{Hq}^{(1)ij} \delta_{\alpha\beta} + C_{Hq}^{(3)ij} \delta_{\alpha\beta} \\
 L_R^{ij\alpha\beta} &= C_{ld}^{\alpha\beta ij} + C_{Hd}^{ij} \delta_{\alpha\beta}.
 \end{aligned}$$



SMEFT: which combinations of coefficients to study?

We assume they are **induce by specific heavy UV states** and study those simplified models instead.

SMEFT

$$\begin{aligned}\mathcal{O}_{lq}^{(1)\alpha\beta ij} &= (\bar{l}_L^\alpha \gamma_\mu l_L^\beta) (\bar{q}_L^i \gamma^\mu q_L^j) \\ \mathcal{O}_{lq}^{(3)\alpha\beta ij} &= (\bar{l}_L^\alpha \gamma_\mu \sigma_a l_L^\beta) (\bar{q}_L^i \gamma^\mu \sigma_a q_L^j) \\ \mathcal{O}_{ld}^{\alpha\beta ij} &= (\bar{l}_L^\alpha \gamma_\mu l_L^\beta) (\bar{d}_R^i \gamma^\mu d_R^j)\end{aligned}$$

Colorless vectors & Leptoquarks

	Spin	G_{SM}	Interaction term	SMEFT coeff.
V'	1	$(\mathbf{1}, \mathbf{3}, 0)$	$[g_q^{ij} (\bar{q}_L^i \gamma^\mu \sigma_a q_L^j) + g_\ell^{\alpha\beta} (\bar{l}_L^\alpha \gamma^\mu \sigma_a l_L^\beta)] V'_{a\mu}$	$C_{lq}^{(3)}$
Z'_L	1	$(\mathbf{1}, \mathbf{1}, 0)$	$[g_q^{ij} (\bar{q}_L^i \gamma^\mu q_L^j) + g_\ell^{\alpha\beta} (\bar{l}_L^\alpha \gamma^\mu \sigma_a l_L^\beta)] Z'_{L\mu}$	$C_{lq}^{(1)}$
Z'_R	1	$(\mathbf{1}, \mathbf{1}, 0)$	$[g_q^{ij} (\bar{d}_R^i \gamma^\mu d_R^j) + g_\ell^{\alpha\beta} (\bar{l}_L^\alpha \gamma^\mu \sigma_a l_L^\beta)] Z'_{R\mu}$	C_{ld}
S_1	0	$(\bar{\mathbf{3}}, \mathbf{1}, 1/3)$	$\lambda_{i\alpha}^* (\bar{q}_L^{i,c} \epsilon l_L^\alpha) S_1$	$C_{lq}^{(1)} = -C_{lq}^{(3)}$
S_3	0	$(\bar{\mathbf{3}}, \mathbf{3}, 1/3)$	$\lambda_{i\alpha}^* (\bar{q}_L^{i,c} \epsilon \sigma_a l_L^\alpha) (S_3)_a$	$C_{lq}^{(1)} = 3C_{lq}^{(3)}$
U_3	1	$(\mathbf{3}, \mathbf{3}, 2/3)$	$\lambda_{i\alpha} (\bar{q}_L^i \gamma_\mu \sigma_a l_L^\alpha) (U_3^\mu)_a$	$C_{lq}^{(1)} = -3C_{lq}^{(3)}$
\tilde{R}_2	0	$(\mathbf{3}, \mathbf{2}, 1/6)$	$\lambda_{i\alpha} \bar{d}_R^i (l_L^\alpha \epsilon \tilde{R}_2)$	C_{ld}
V_2	1	$(\bar{\mathbf{3}}, \mathbf{2}, 5/6)$	$\lambda_{i\alpha}^* \bar{d}_R^{i,c} \gamma_\mu (l_L^\alpha \epsilon V_2^\mu)$	C_{ld}

U_1 LQ does not contribute to bsw: we don't consider it

$$\begin{aligned}\mathcal{O}_{Hq}^{(1)ij} &= (H^\dagger \overleftrightarrow{D}_\mu H) (\bar{q}_L^i \gamma^\mu q_L^j) , \\ \mathcal{O}_{Hq}^{(3)ij} &= (H^\dagger \sigma_a \overleftrightarrow{D}_\mu H) (\bar{q}_L^i \gamma^\mu \sigma_a q_L^j) , \\ \mathcal{O}_{Hd}^{ij} &= (H^\dagger \overleftrightarrow{D}_\mu H) (\bar{d}_R^i \gamma^\mu d_R^j) .\end{aligned}$$

vectorlike quarks

Simplified model	Spin	SM irrep	SMEFT couplings
D	1/2	$(\mathbf{3}, \mathbf{1}, -1/3)$	$C_{Hq}^{(1)} = C_{Hq}^{(3)}$
T_1	1/2	$(\mathbf{3}, \mathbf{3}, -1/3)$	$C_{Hq}^{(1)} = -3C_{Hq}^{(3)}$
T_2	1/2	$(\mathbf{3}, \mathbf{3}, 2/6)$	$C_{Hq}^{(1)} = 3C_{Hq}^{(3)}$
Q_1	1/2	$(\mathbf{3}, \mathbf{2}, 1/6)$	C_{Hd}
Q_5	1/2	$(\mathbf{3}, \mathbf{2}, -5/6)$	C_{Hd}

These give too large contributions to B_s mixing and $B_s \rightarrow \mu\mu$:

A good fit of the $R_{\nu K}$ excess is never allowed. (see backup slide)

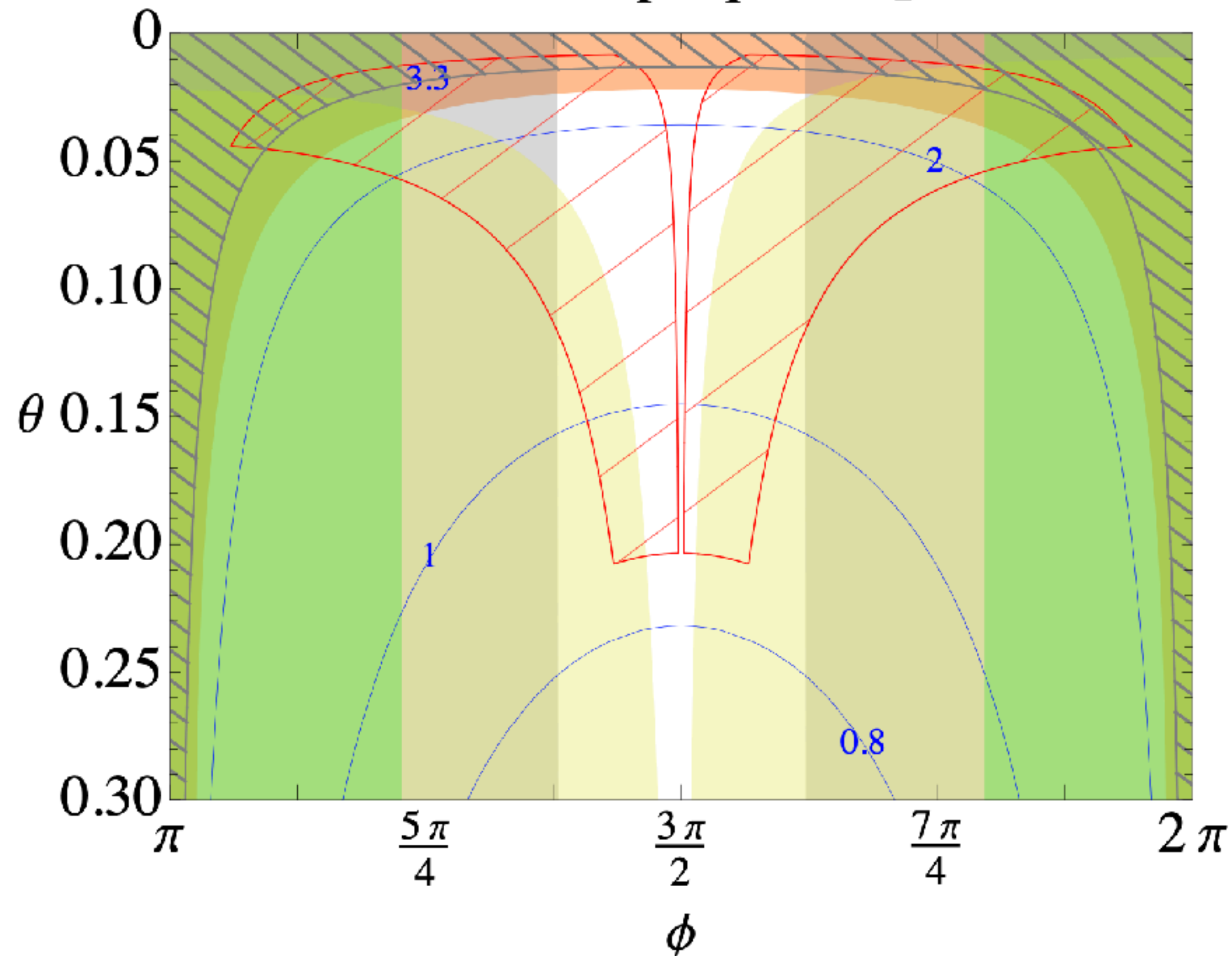
UV mediators - \tilde{R}_2 leptoquark

(similar for S_1)

LQ	Spin	G_{SM}	Interaction term	SMEFT coeff.	C
\tilde{R}_2	0	$(\mathbf{3}, \mathbf{2}, 1/6)$	$\lambda_{i\tau} \bar{d}_R^i (l_L^\tau \epsilon \tilde{R}_2)$	C_{ld}	$-\frac{1}{2} \frac{ \lambda ^2}{m_{LQ}^2}$

We fix the **LQ mass at 2 TeV** to avoid direct-searches bounds.

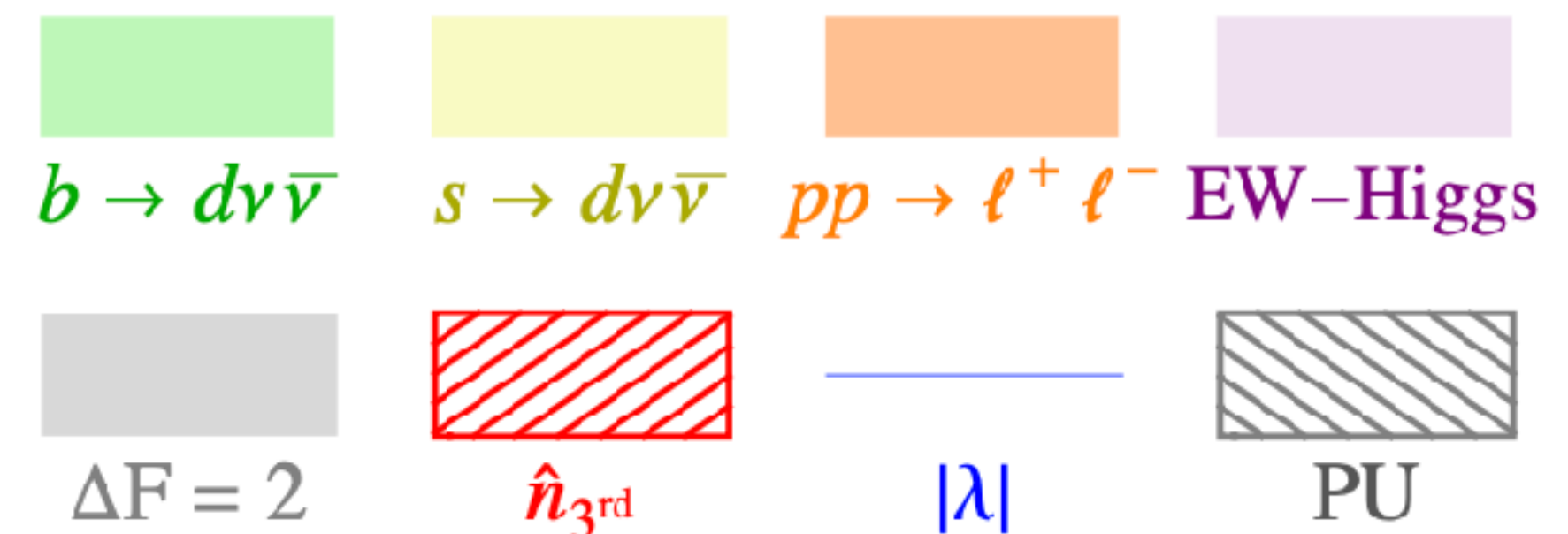
scalar leptoquark \tilde{R}_2



At each parameter space point we fix the best-fit:

$$C_{ld}^{\tau\tau sb} \Big|_{\tilde{R}_2, \text{best-fit}} \approx (7.5 \text{ TeV})^{-2}$$

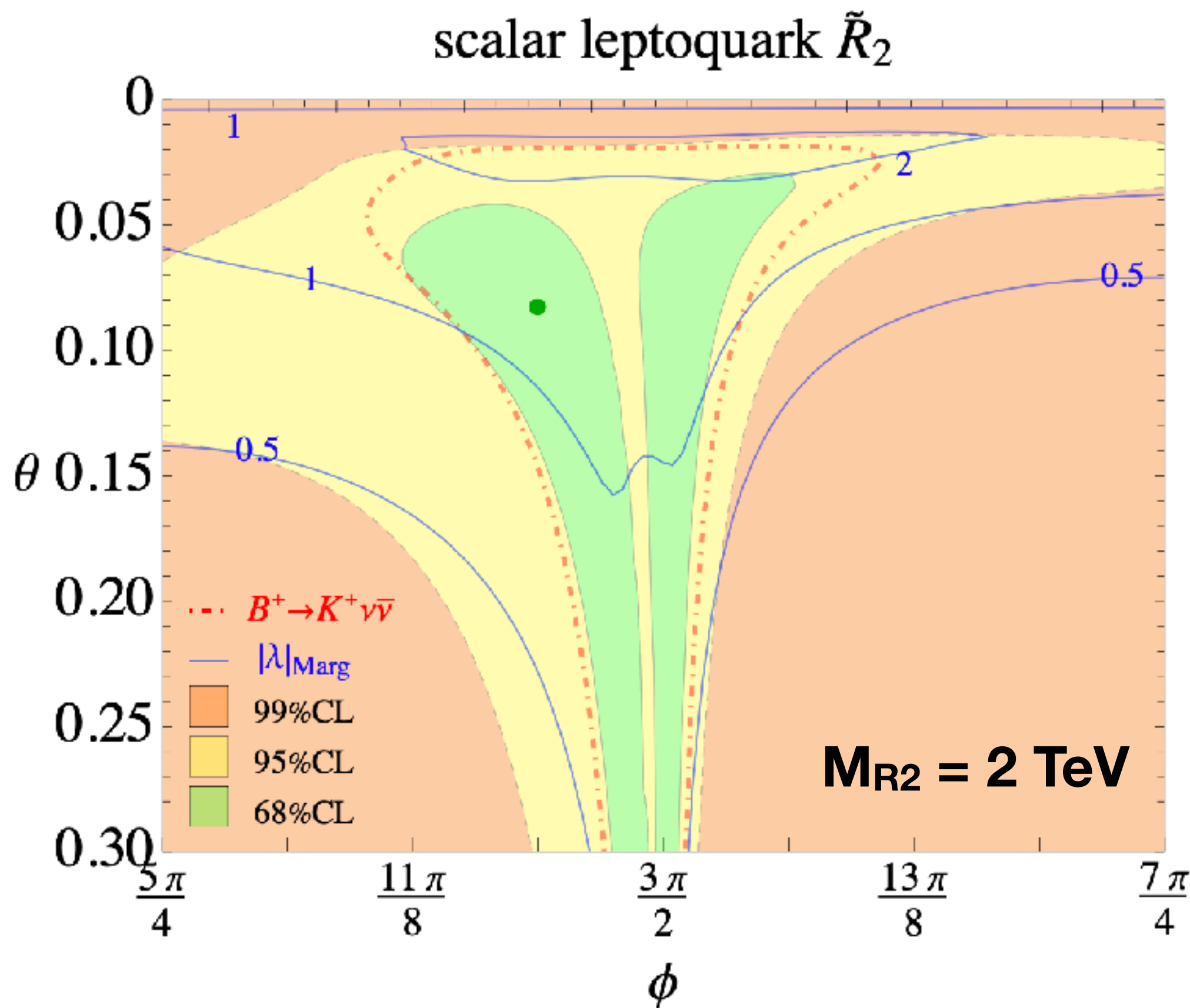
Show regions excluded by:



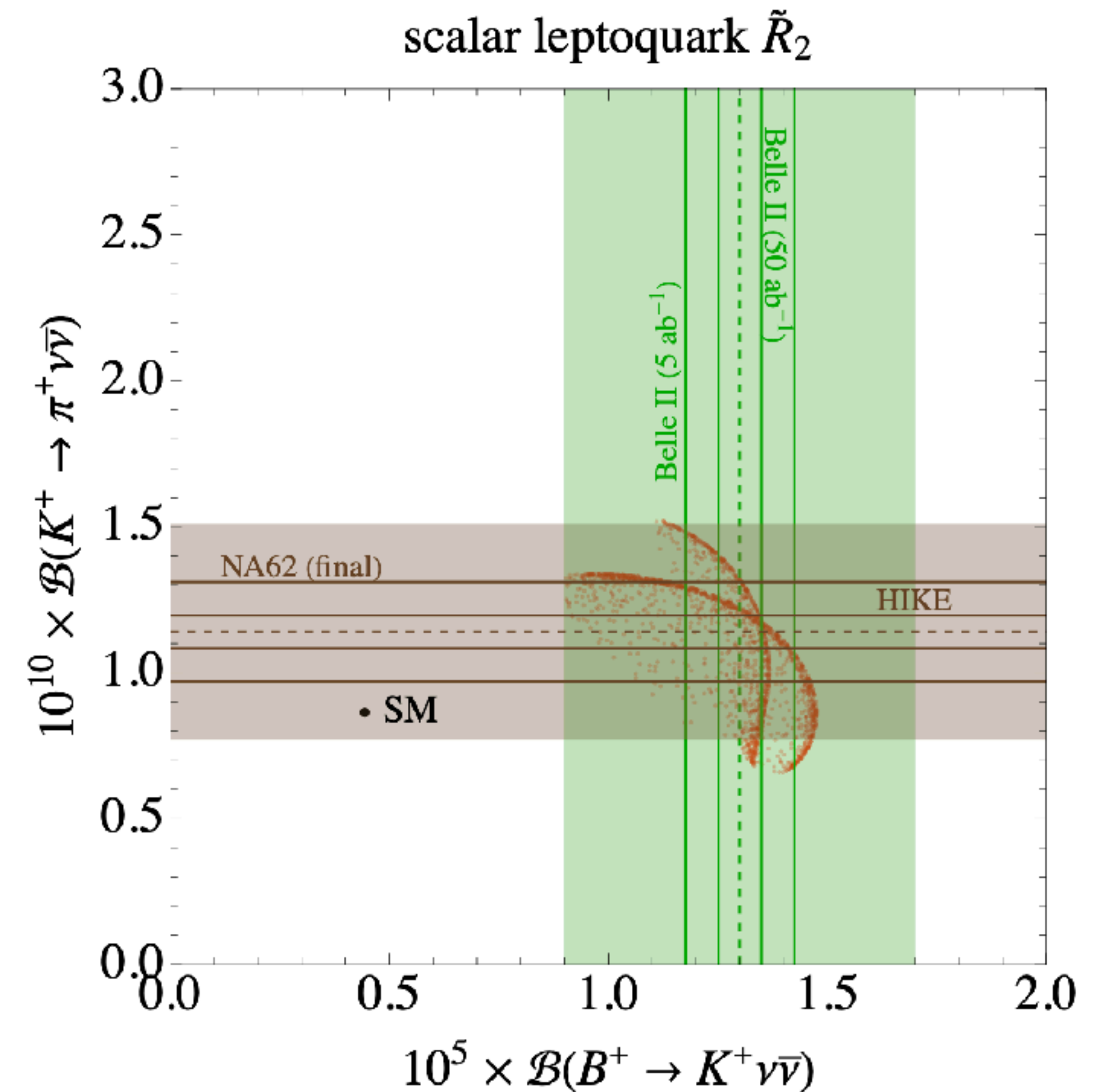
UV mediators - \tilde{R}_2 leptoquark

(similar for S_1)

Favoured region from the global fit of all observables, marginalising over $|\lambda|$.



Correlation between $BK\nu\nu$ and $K\pi\nu\nu$ for the points within 1σ region.



UV mediators - Z'

$$\mathcal{L} \supset \left[\sum_{ij} g_L^{ij} (\bar{q}_L^i \gamma^\mu q_L^j) + \sum_{ij} g_R^{ij} (\bar{d}_R^i \gamma^\mu d_R^j) + \sum_{\alpha\beta} g_\ell^{\alpha\beta} \bar{l}_L^\alpha \gamma^\mu l_L^\beta \right] Z'_\mu$$

vector singlet Z'_R , l^τ only $g_\ell^{\alpha\beta} = g_\ell \delta^{\tau\alpha} \delta^{\tau\beta}$

(similar for a coupling to LH quarks or for a SU(2) triplet V')

$$C_{ld}^{\tau\tau sb} \Big|_{Z_{R, \text{best-fit}}} \approx (7.7 \text{ TeV})^{-2}$$

Meson mixing puts an upper bound on $|g_q/g_\ell|$.

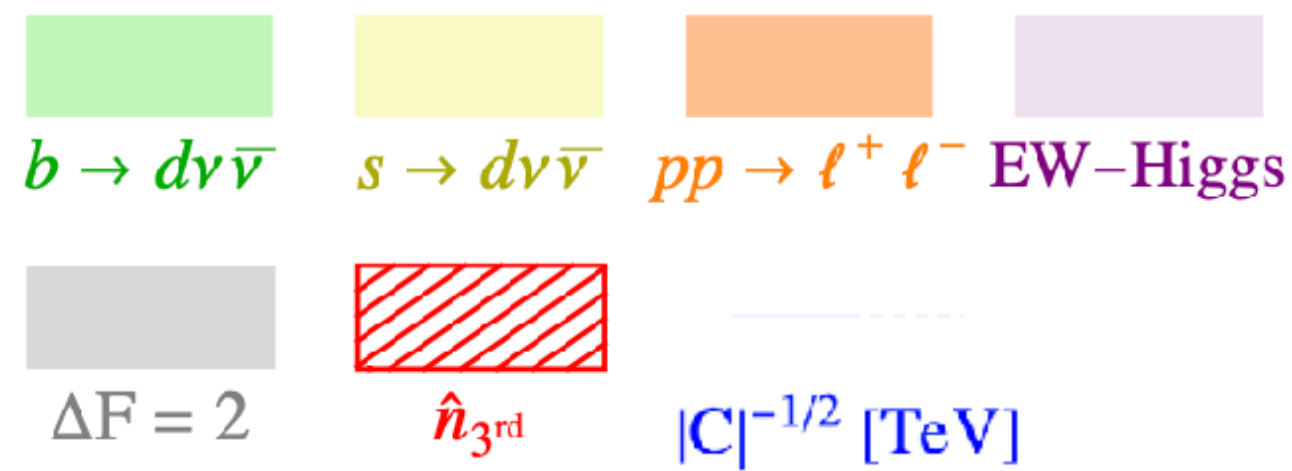
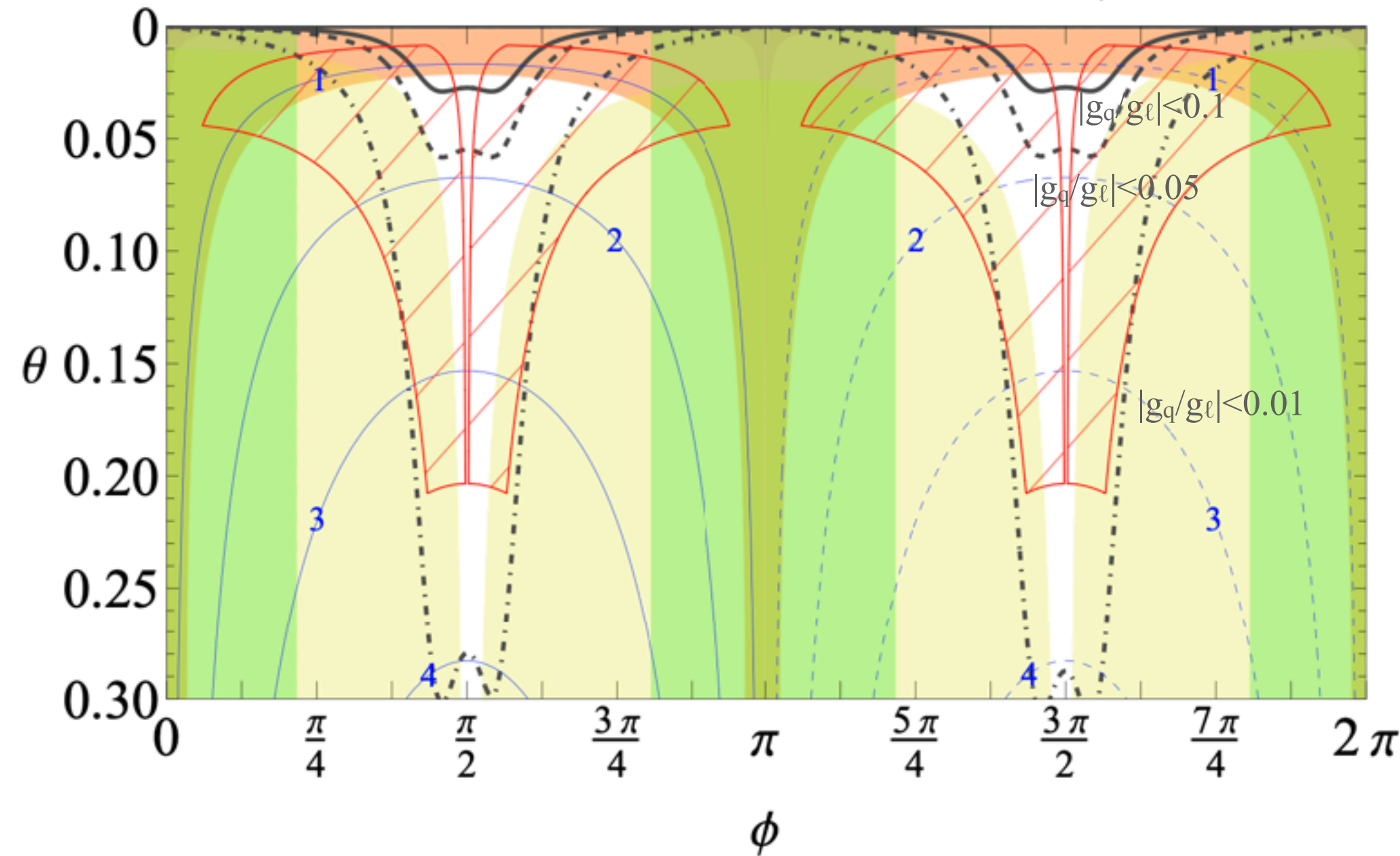
Combined with the **perturbative unitarity** constraint $g_\ell < 2.3$ we get an **upper limit on the vector mass** for each point in the plane, when imposing a fit to the anomaly:

$$M_{V'} \lesssim 1391 \text{ GeV} \left(\frac{|g_q/g_\ell|^{\text{max}}}{0.05} \right)^{1/2} |\sin \theta \cos \theta \sin \phi|^{\frac{1}{2}}$$

$$\approx 762 \text{ GeV} \left(\frac{|g_q/g_\ell|^{\text{max}}}{0.05} \right)^{1/2} \left| \frac{\theta}{0.3} \right|^{1/2}, \quad (\text{for } \theta \ll 1 \text{ and } \phi \sim \pi/2)$$

The **vector must be rather light** in the allowed region!

Need to **check di-tau bounds** without assuming EFT.



Region excluded by $\Delta F=2$ for different max values of $|g_q/g_\ell|$.

$$C = -g_q g_\ell / M_{Z'}^2$$

UV mediators - Z'

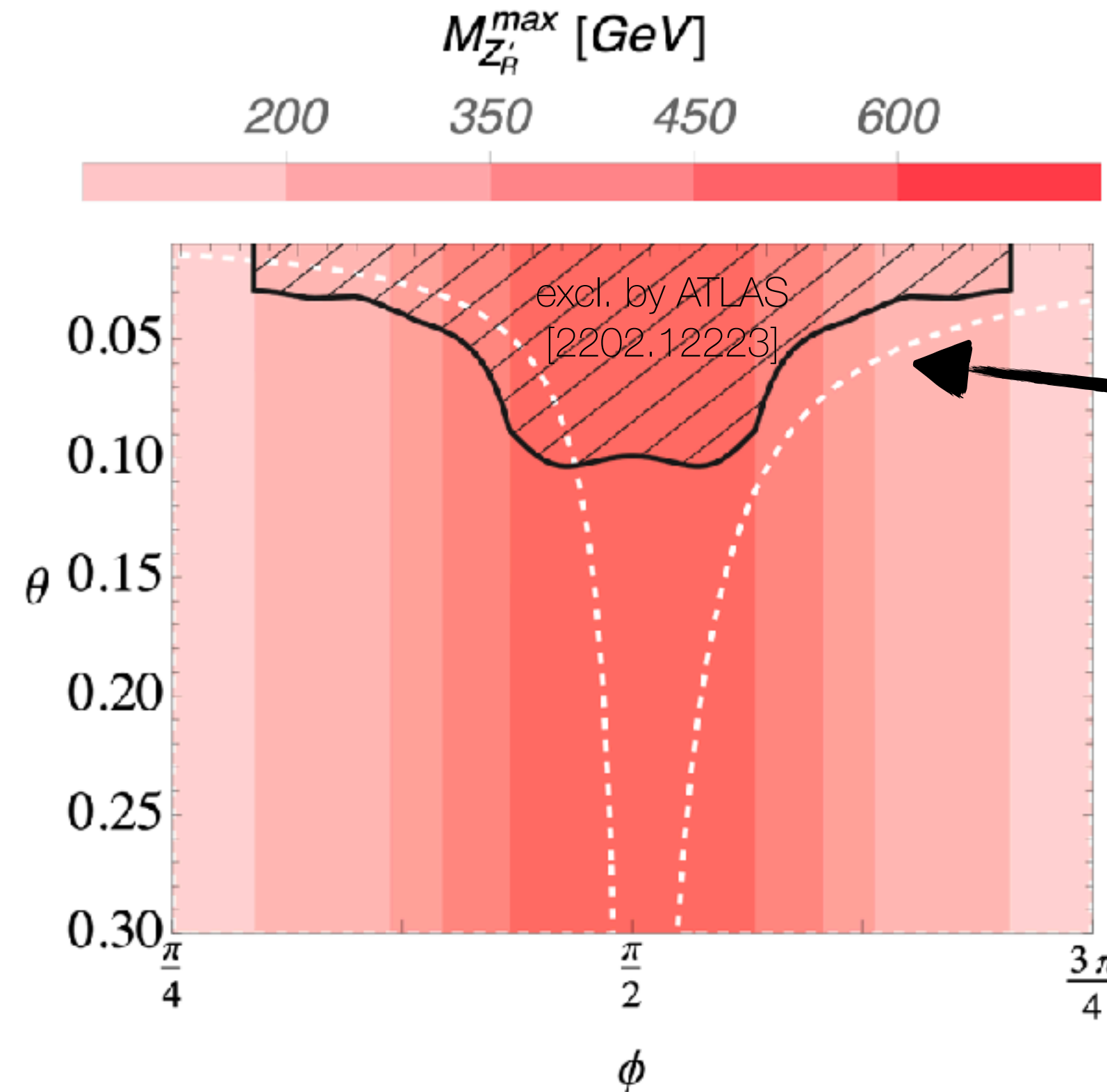
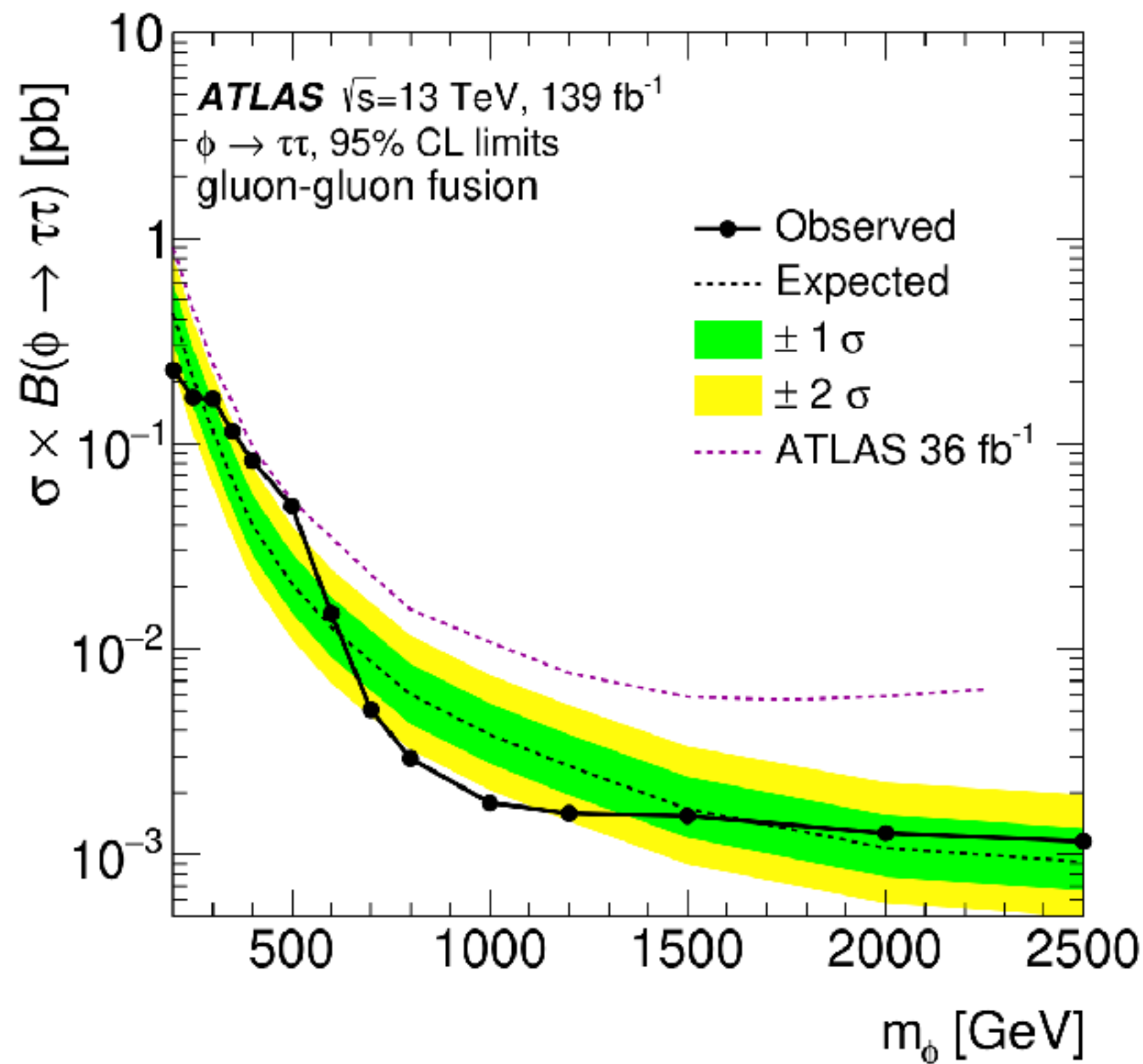
$$\mathcal{L} \supset \left[\sum_{ij} g_L^{ij} (\bar{q}_L^i \gamma^\mu q_L^j) + \sum_{ij} g_R^{ij} (\bar{d}_R^i \gamma^\mu d_R^j) + \sum_{\alpha\beta} g_\ell^{\alpha\beta} \bar{l}_L^\alpha \gamma^\mu l_L^\beta \right] Z'_\mu$$

Since g_ℓ is large, the Z' decays mostly to $\tau\tau$ or $\nu\nu$, with $\text{Br} \sim 1/2$.

$$\Gamma_{\text{tot}}/M_{V'} \approx 14\%$$

$$\sigma(pp \rightarrow V'^0 \rightarrow \tau^+\tau^-) \approx \frac{4\pi^2}{3} \mathcal{B}(V'^0 \rightarrow \tau^+\tau^-) \sum_{i,j=u,d,s,c,b} \frac{\Gamma(V'^0 \rightarrow q^i\bar{q}^j)}{M_{V'}} \frac{2}{s_0} \mathcal{L}_{q^i\bar{q}^j}(M_{V'})$$

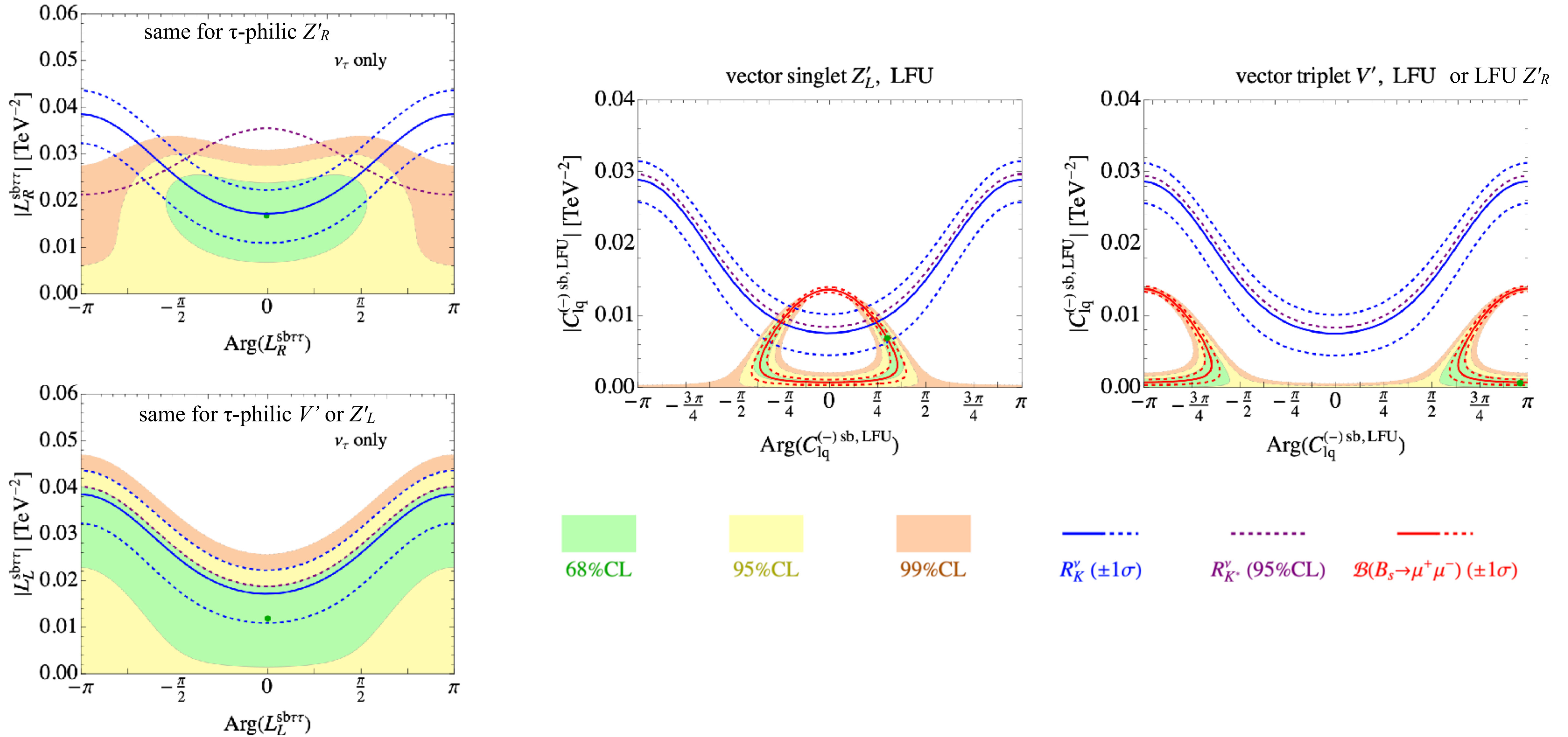
$$\Gamma(V'^0 \rightarrow q^i\bar{q}^j) = \frac{M_{V'} N_c}{24\pi} |g_q^{ij}|^2$$



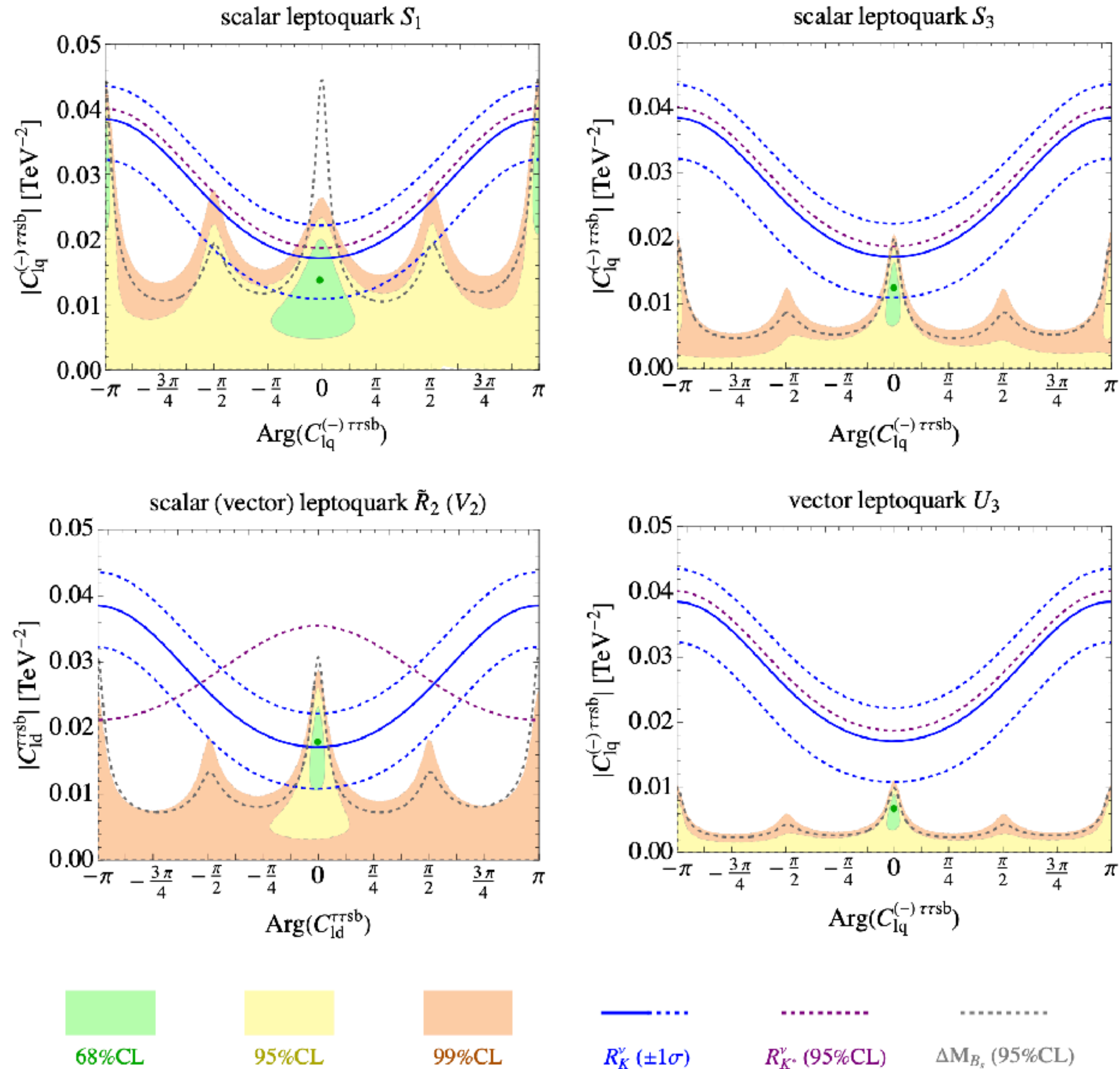
Outside the white dashed line is excluded by Flavour + EW

(this is a rough constraint, we neglect effect of different acceptances between scalar and vector resonances)

Fits of the s-b couplings - LEFT & vectors



Fits of the s-b couplings - LQ

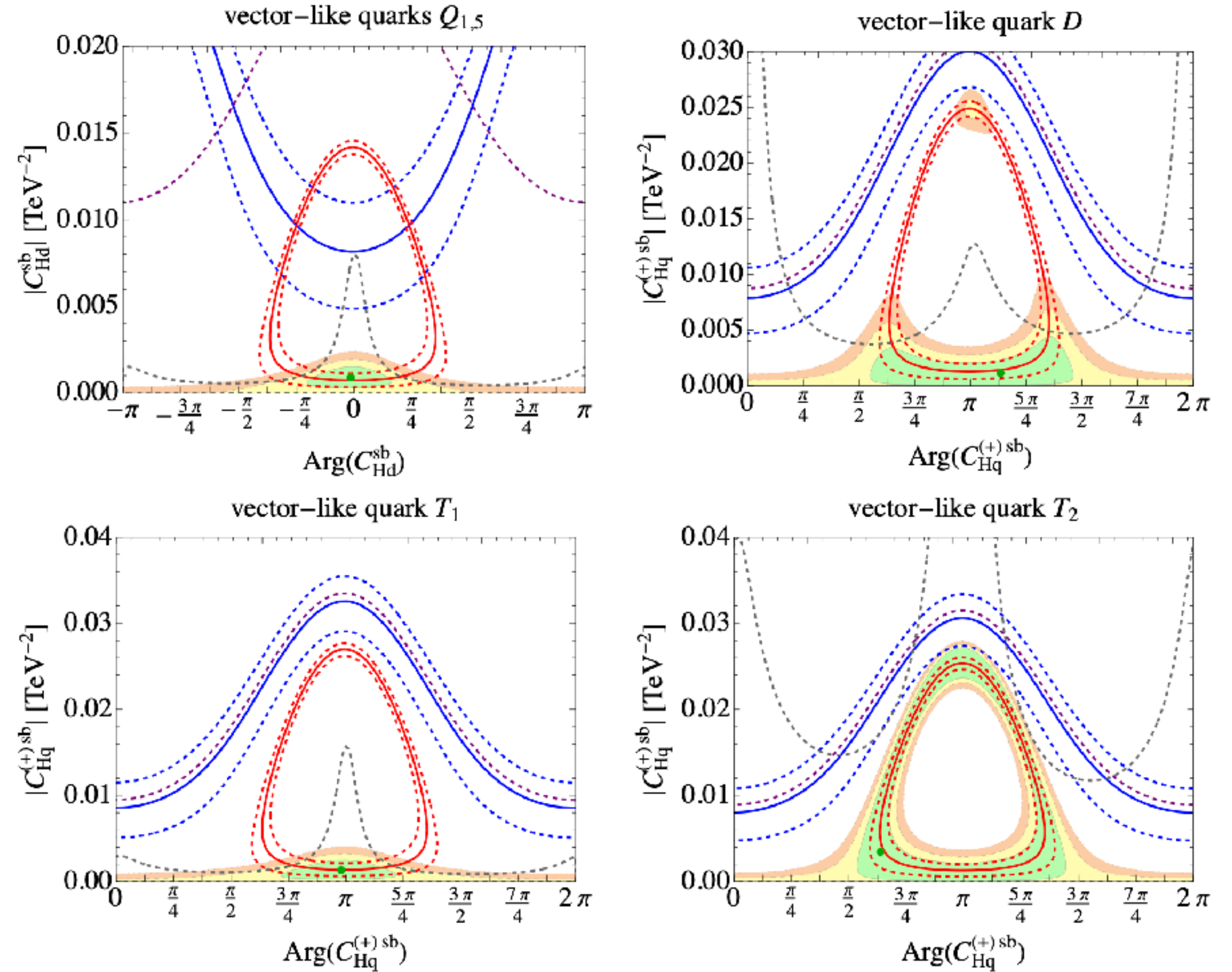


Fits of the s-b couplings - VLQ

Given the multiple possible combinations of coefficients, we assume they are induced by specific heavy UV states.

In this case: vector-like quarks:

Simplified model	Spin	SM irrep	SMEFT couplings
D	1/2	$(\mathbf{3}, \mathbf{1}, -1/3)$	$C_{Hq}^{(1)} = C_{Hq}^{(3)}$
T_1	1/2	$(\mathbf{3}, \mathbf{3}, -1/3)$	$C_{Hq}^{(1)} = -3C_{Hq}^{(3)}$
T_2	1/2	$(\mathbf{3}, \mathbf{3}, 2/6)$	$C_{Hq}^{(1)} = 3C_{Hq}^{(3)}$
Q_1	1/2	$(\mathbf{3}, \mathbf{2}, 1/6)$	C_{Hd}
Q_5	1/2	$(\mathbf{3}, \mathbf{2}, -5/6)$	C_{Hd}



A good fit of the R_K^ν excess is never allowed.

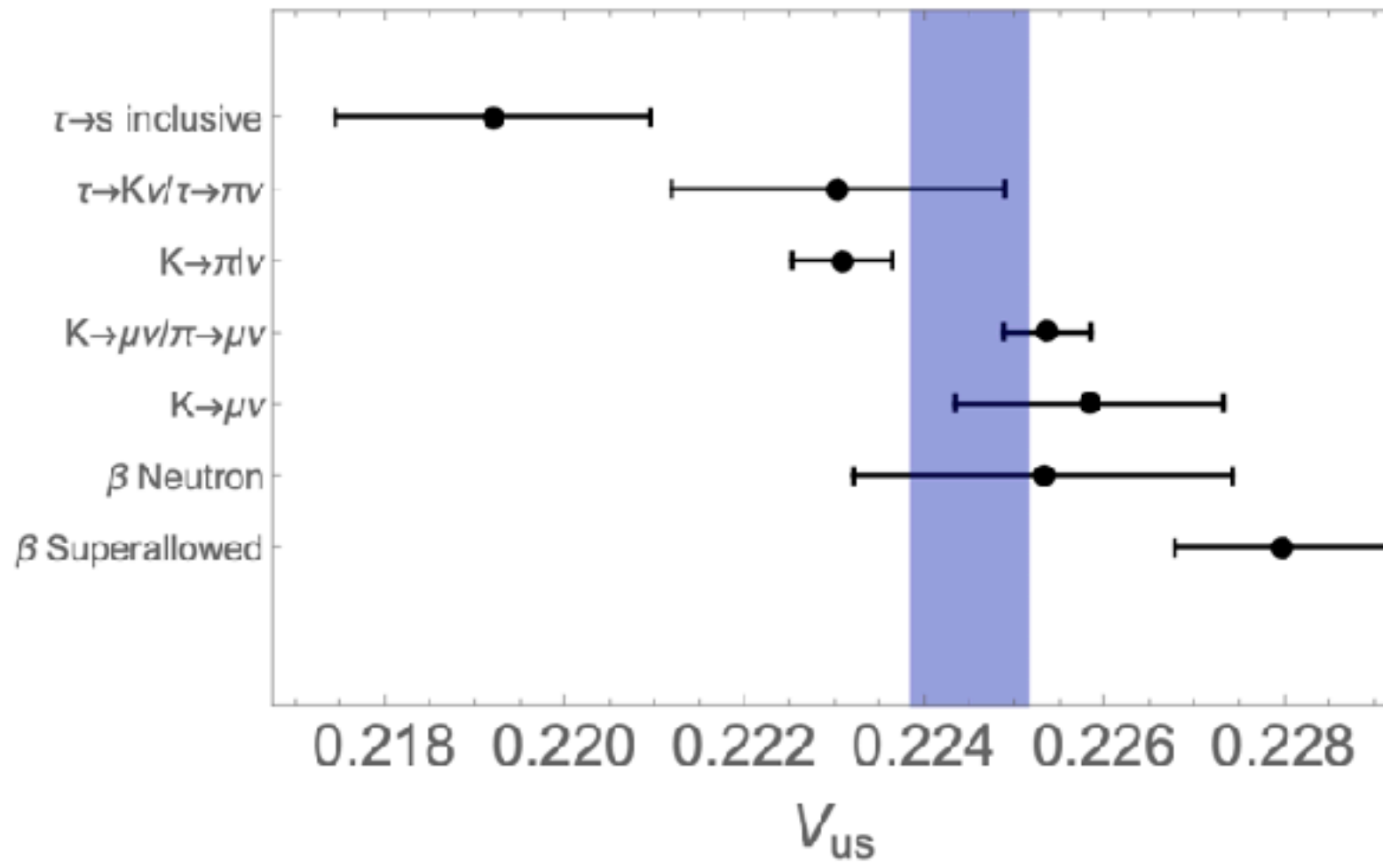
L. Allwicher, D. Becirevic, G. Piazza, S. Rosauero-Alcaraz and O. Sumensari [2309.02246]

Observables included

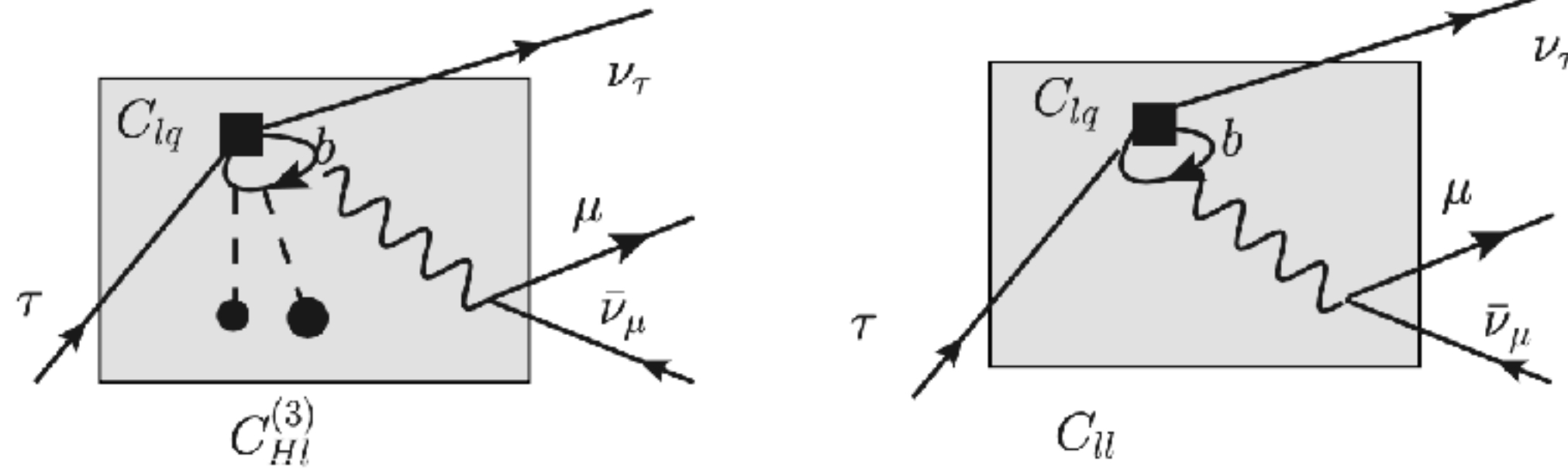
Cabibbo angle

unitarity relation (V_{ub} negligible):

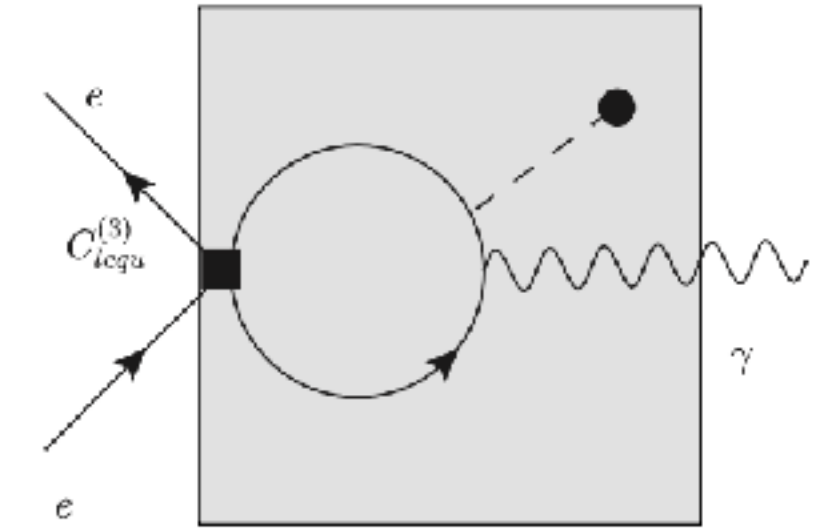
$$|V_{ud}|^2 + |V_{us}|^2 = 1$$



LFU



Magnetic Moments



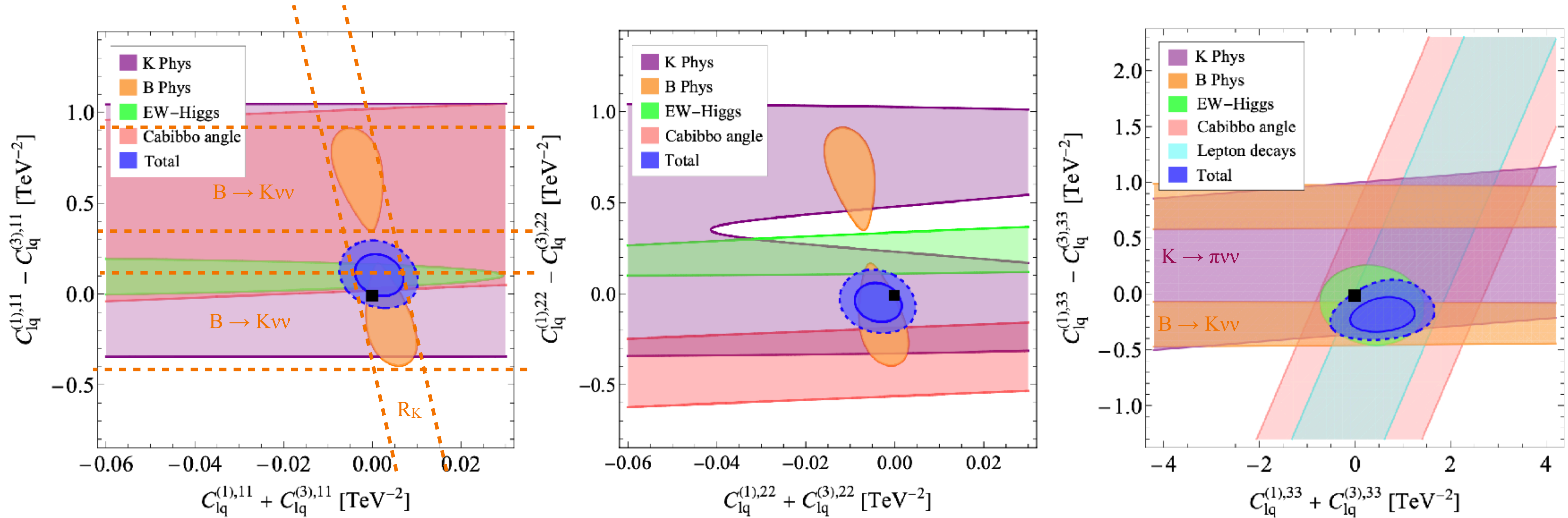
Higgs and EW

- δg_L^{Zl} $\leftarrow C_{uB}, C_{uW}, C_{Hu}, C_{Hq}^{(1,3)}, C_{lq}^{(1,3),ll}, C_{lu}^{ll}, \dots$
- δg_L^{Wl} $\leftarrow C_{uB}, C_{uW}, C_{Hu}, C_{Hq}^{(1,3)}, C_{lq}^{(3),ll}, \dots$
- δg_R^{Zl} $\leftarrow C_{uB}, C_{uW}, C_{Hu}, C_{Hq}^{(1,3)}, C_{eu}^{ll}, C_{qe}^{ll}, \dots$
- δg_L^{Zb} $\leftarrow C_{Hq}^{(1,3)}, C_{Hu}, C_{qq}^{(1,3)}, \dots$
- δg_R^{Zb} $\leftarrow C_{Hq}^{(1)}, C_{Hu}, C_{qq}^{(1,3)}, C_{uB}, C_{uW}, \dots$
- $c_{\gamma\gamma}$ $\leftarrow C_{uB}, C_{uW}, C_{uG}$
- c_{gg} $\leftarrow C_{uG}$
- $[C_{eH}]_{\alpha\alpha}$ $\leftarrow C_{lequ}^{(1),\alpha\alpha}$
- $[C_{uH}]_{33}$ $\leftarrow C_{uH}, C_{uG}, C_{Hq}^{(1,3)}, C_{qu}^{(1,8)}, \dots$

B physics

	Tree level matching	RG and 1-loop matching
$R_{K^{(*)}}^\nu$ $K \rightarrow \pi \nu \bar{\nu}$	$C_{Hq}^{(1,3)}, C_{lq}^{(1,3),\alpha\beta}$	$C_{Hu}, C_{qq}^{(1,3)}, C_{lu}^{\alpha\beta}, C_{qe}^{\alpha\beta}$ $C_{qu}^{(1,8)}, C_{uu}, C_{uW}$
$B \rightarrow K^{(*)} l \alpha \beta$ $B_{s,d} \rightarrow l \alpha \beta$ $K \rightarrow \pi l \alpha \beta$ $K \rightarrow l \alpha \beta$	$C_{Hq}^{(1,3)}, C_{lq}^{(1,3),\alpha\beta}, C_{qe}^{\alpha\beta}$	$C_{qq}^{(1,3)}, C_{lu}^{\alpha\beta}, C_{eu}^{\alpha\beta}$
$R_{K^{(*)}}$	$C_{lq}^{(1,3),ll}, C_{qe}^{ll}$	C_{lu}^{ll}
$B \rightarrow X_s \gamma$		$C_{Hq}^{(1,3)}, C_{uB}, C_{uW}, C_{uG}$

2D fits



With electrons and muons: **EW bounds** do **not allow** a combined explanations of **Cabibbo anomaly** and **$B \rightarrow K\nu\nu$**

Gaussian Fit ex. semileptonic

$$\vec{C} = (C_{qq}^{(+)}, C_{qq}^{(-)}, C_{uu}, C_{qu}^{(1)}, C_{qu}^{(8)}, C_{Hq}^{(+)}, C_{Hq}^{(-)}, C_{Hu}, C_{uH}, C_{uG}, C_{uW}, C_{uB})$$

$$\Delta\chi^2 \equiv \chi_{\text{SM}}^2 - \chi_{\text{best-fit}}^2 \approx 10.9 \quad (\text{only mild improvements in several observables, not a single "anomaly"})$$

Due to flat directions, we report the result in terms of the eigenvectors of the Hessian matrix around the minimum

$$\chi^2 = \chi_{\text{best-fit}}^2 + (C_i - \mu_{C_i})(\sigma^2)_{ij}^{-1}(C_j - \mu_{C_j}) = \chi_{\text{best-fit}}^2 + \frac{(K_i - \mu_{K_i})^2}{\sigma_{K_i}^2} \quad \vec{K} = U_{KC}\vec{C}$$

Coefficient	Gaussian fit [TeV ⁻²]	Coefficient	Gaussian fit [TeV ⁻²]
K_1	0.0019 ± 0.0023	K_7	0.54 ± 0.79
K_2	0.0179 ± 0.0083	K_8	0.74 ± 0.88
K_3	-0.002 ± 0.015	K_9	-0.8 ± 1.3
K_4	-0.016 ± 0.021	K_{10}	-0.7 ± 1.8
K_5	0.044 ± 0.029	K_{11}	12 ± 13
K_6	-0.30 ± 0.38	K_{12}	-11 ± 16

Flat directions:

$$K_{11} \approx -0.80C_{qq}^{(-)} + 0.45C_{uu} - 0.36C_{qu}^{(1)} - 0.12C_{Hu} + \dots ,$$

$$K_{12} \approx +0.40C_{qq}^{(-)} + 0.88C_{uu} + 0.24C_{qu}^{(1)} - 0.09C_{Hu} + \dots .$$