

2024 DISCRETE

## **Penguin Decays of B mesons**

Based on JHEP 06 (2023) 108 and JHEP 08 (2024) 030. In collaboration with Joaquim Matias, Sebastien Descotes-Genon and Gilberto Tetlalmatzi-Xolocotzi.

### Non leptonics: Motivation and Introduction

- Expectation: tensions in rare  $b \rightarrow s$  (maybe  $b \rightarrow d$ ?) if tensions in semileptonics are due to NP.
- FCNC Non leptonic decays : loop suppressed in the SM : satisfactory amount of Experimental data.
- However, increased difficulty in controlling hadronic uncertainties w.r.t semileptonics.
- Theoretical approaches available: Phenomenological extraction using flavor symmetries (GTX, TH, *Eur.Phys.J.C* 82 (2022) 3, 210). Relate to other modes using symmetry (U-spin, SU(3)) (MG, DL, etal., Nucl.Phys. B675 (2003) 333-415 etc). Compute hadronic matrix elements (QCD Factorization) (MB, MN, etal, Phys. Lett. B 514 (2001) 315, etc).
- Work with penguin dominated modes with  $B_{s,d}$  decaying to same final states:  $K^{(*)} \overline{K}^{(*)}, K^* \phi$ .
- Use them to construct observables (ratios of (longitudinal for vector-vector) branching ratios ).
   with reduced sensitivities to hadronic uncertainties (endpoint divergences).
- Use these observables to look for effects that might potentially be beyond SM: New Physics.

#### Amplitude and " $\Delta$ "

- $\bar{A}_f = A(\bar{B}_q \to F_1F_2) = \lambda_u^{(q)}T_q + \lambda_c^{(q)}P_q = \lambda_u^{(q)}\Delta_q \lambda_t^{(q)}P_q$  (unitarity).
- $\Delta_q$  is free of endpoint divergences (PRL 97 (2006) 061801: SDG, JM, JV). Because:  $T_q = A_{K^*K^*}^q \left( \alpha_4^u - \frac{1}{2} \alpha_{4,EW}^u + \beta_3^u + 2\beta_4^u - \frac{1}{2} \beta_{3,EW}^u - \beta_{4,EW}^u \right)$

$$P_q = A_{K^*K^*}^q \left( \alpha_4^c - \frac{1}{2} \alpha_{4,EW}^c + \beta_3^c + 2\beta_4^c - \frac{1}{2} \beta_{3,EW}^c - \beta_{4,EW}^c \right)$$

• Where,  

$$\alpha_i^p(M_1M_2) \propto \left[V_i(M_2) + \frac{4\pi^2}{N_c}H_i(M_1M_2)\right] + P_i^p(M_2)$$
  
 $\propto X_H^{M_1} \sim \ln(\frac{m_b}{\Lambda_{OCD}})$ 

(soft gluon spectator int, divergent, power suppressed, universal)

•  $\beta_i^p$ : Penguin annihilation,  $\beta_{i,EW}^p$ : Electroweak penguin annihilation

 $\propto X_A^{M_1} \sim \text{Endpoint divergence} \sim \ln(\frac{m_b}{\Lambda_{QCD}})$  (universal) Enter the same way in T and P (at LO in QCD).

These divergences are responsible for the model dependence of the analysis.

#### The "L" observable: definition and features



- Relative uncertainty less than relative uncertainties in branching ratios.
- Generally asymmetric SM distribution since ratio. Degree of asymmetry depends on the relative uncertainty on the denominator.
- Dominant contribution to the uncertainties from form factors and not annihilation which are dominant sources for branching ratio uncertainties (use of u-spin symmetry).
- Renders value of ratio 'L' robust: independent of dynamical model/symmetry considerations used to calculate its value.

## Diagrams: $K^{(*)}\overline{K}^{(*)}$





## Diagrams: $K^{(*)}\phi$





#### **Theory vs experiment: Current status**

Observable	SM (QCDF)	Experiment	Deviation
$10^6 BR(\bar{B}_d \to K^0 \ \bar{K}^0)$	$1.09^{+0.29}_{-0.20}$	$1.21 \pm 0.16$	$0.4\sigma$
$10^7 BR(\bar{B}_d \to K^{*0} \ \bar{K}^{*0})_L$	$2.27^{+0.99}_{-0.74}$	$6.04^{+1.81}_{-1.78}$	$1.8\sigma$
$10^5 BR(\bar{B}_s \to K^0 \ \bar{K}^0)$	$2.80^{+0.89}_{-0.62}$	$1.76 \pm 0.33$	1.6σ
$10^6 BR(\bar{B}_s \to K^{*0} \ \bar{K}^{*0})_{\rm L}$	$4.36^{+2.23}_{-1.65}$	$2.62^{+0.85}_{-0.75}$	0.9σ
$10^6 BR(\bar{B}_d \rightarrow \bar{K}^{*0}\phi)_L$	$4.53^{+2.16}_{-1.80}$	$4.96^{+0.31}_{-0.30}$	$0.3\sigma$
$10^7 BR(\bar{B}_s \to K^{*0} \phi)_L$	$2.19^{+1.05}_{-0.94}$	$5.56^{+2.78}_{-2.27}$	1.3σ
$\mathbf{L}_{\mathbf{K}^*\overline{\mathbf{K}}^*}$	$19.53^{+9.14}_{-6.64}$	$4.43 \pm 0.92$	2.6σ
$\mathbf{L}_{\mathbf{K}\overline{\mathbf{K}}}$	$26.00^{+3.88}_{-3.59}$	14.58 ± 3.37	2.4σ
$L_{K^* \phi}$	$22.04_{-4.88}^{+7.06}$	$8.80^{+6.07}_{-2.97}$	1.5σ
$10^5 (BR(\bar{B}_s \to K^{*0} \bar{K}^0) + \text{c.c.})$	$0.83^{+0.50}_{-0.25}$	$1.98 \pm 0.28 \pm 0.50$	$1.4\sigma$
$10^6 BR(\bar{B}_d \rightarrow \bar{K}^0 \phi)$	$4.28^{+2.71}_{-1.50}$	$7.3 \pm 0.7$	$1.3\sigma$
$10^6 BR(B^- \to K^- \phi)$	$4.67^{+2.98}_{-1.63}$	$8.8^{+0.7}_{-0.6}$	$1.5\sigma$
$10^6 BR(B^- \rightarrow K^{*-}\phi)$	$4.94^{+2.34}_{-1.91}$	$4.96^{+1.16}_{-1.08}$	0.05 <i>o</i>

#### **Operator basis and SM Wilson Coefficients**

SM Wilson Coefficients (at $\mu = 4.18 \text{ GeV}$ )								
$C_1$	$C_2$	$\mathcal{C}_3$	$C_4$	$C_5$	$\mathcal{C}_6$			
1.082	-0.191 0.014		-0.036	0.009	-0.042			
$C_7/\alpha_{em}$	$C_8/\alpha_{em}$	$C_9/\alpha_{em}$	$C_{10}/\alpha_{em}$	$\mathcal{C}_{7\gamma}^{\mathrm{eff}}$	$\mathcal{C}_{8g}^{\mathrm{eff}}$			
-0.011	0.060 -1.254		0.224	-0.318	-0.151			

#### Lessons from one operator scenarios

- NP in Q<sub>6d,s</sub> does not work: because of the vector modes.
- Assuming NP affects either  $Q_{4d,s}$  or  $Q_{8,gd,s}$  we find common overlaps for PP and VV modes.
- Result of including  $K^*\phi$  modes with  $K^{(*)}K^{(*)}$  modes: "allowed" range of NP values is greater for  $b \rightarrow d$  as compared to  $b \rightarrow s$ .
  - ......pattern broken when pseudoscalar vector modes included.
- NP affects  $Q_{4d,s}$ : mutual overlap among  $K^{(*)}\phi$ . Also among  $K^{(*)}K^{(*)}$ ,  $K^*K$ . But not together.
- NP affects  $Q_{8gd,s}$ : mutual overlap among  $K^{(*)}\phi$ . No mutual overlap among  $K^{(*)}K^{(*)}$ ,  $K^*K$ .
- No common one operator explanation is possible. Two operators (involving  $Q_6$ )?!
- <u>Appeal to Experimentalists</u>: Updated Measurement of BR( $\overline{B}_d \to \overline{K}\phi$ ) required to confirm or dismiss this picture.

#### Two operator scenarios: $Q_4 - Q_6$



#### Two operator scenarios: $Q_4 - Q_6$



Zoomed in version showing the common region that explains all the seven b to s branching ratios for the  $Q_4 - Q_6$  scenario.

#### Comparison: SM



#### Comparison: $Q_4 - Q_6$



#### Conclusions

Proposed optimized "L" observables which are ratios involving penguin dominated decay modes related by d to s interchange: only used while modelling the divergent annihilation and hard spectators.

Robust observables in terms of universal annihilation, given current (rather simplistic) model:	Observable	Universal	U spin broken	
	$L_{K^*\overline{K}^*}$	$19.53^{+9.14}_{-6.64}$	$19.04^{+10.20}_{-7.14}$	
	L <sub>KK</sub>	$26.00^{+3.88}_{-3.59}$	$25.79^{+5.27}_{-4.47}$	

- Dominant sources of uncertainties for theoretical SM estimates of the L's are form factors.
- The simplest NP scenario that results in common overlap among all the VV, PP and PV charged and neutral branching ratios as per the current data along with the three L's are 2 operator scenarios  $Q_{4f} Q_{6f}$ .  $Q_{6d,s}$  is important!
- <u>Appeal to Experimentalists</u>: Most recent measurements on  $BR(\overline{B}_d(B^-) \to \overline{K}^0(K^-)\phi)$  more than a decade old. PDG average involves measurements by CDF, Babar, CLEO, Belle with rather different central values. No LHCb measurement.  $1.5\sigma$  deviation between these measurements surprising because they are related by isospin. Maybe updated measurement can change this scenario. In particular, these two measurements being consistent within  $1\sigma$  with the current measurement for  $BR(\overline{B}_d \to \overline{K}^0\phi)$  will make  $Q_{6f} Q_{8gf}$  a viable scenario.

#### **Future directions and discussions**

- Correlated form factors (LCSR, Lattice)?
- Correlated measurement of  $K^{*0}\phi$  Branching fractions. LHCb is already working on these modes.
- Annihilations beyond Beneke etal. CP asymmetry measurements.
- $L_{K^*\phi}^{exp}$  has asymmetric errors. However, a correlated measurement in the future, as well as an increase in the precision of  $f_L(\overline{B}_S \to K^{*0}\phi)$  and  $BR(\overline{B}_S \to K^{*0}\phi)$  might help decrease the asymmetry. Measurement on  $b \to d BR(\overline{B}_S \to K^0\phi)$  and  $BR(\overline{B}_d \to K^{*0}\overline{K}^0 + c.c.)$  will permit construction of L's for mixed modes.
- First exploratory works. Working on rigorous statistical analysis taking asymmetric distributions into account. Possibility of three operator scenarios, complex Wilson coefficients etc.: Stay tuned!





# Backup



 Note: Dominant uncertainties from form factors and NOT divergences. (Somewhat) reduced model dependence.

#### *L<sub>K<sup>\*</sup>K<sup>\*</sup>*: Error Budget</sub>

-	Relative Error							
Input	$L_{K^*\bar{K}^*}$	$ P_s ^2$	$ P_d ^2$					
$f_{K^*}$	(-0.1%, +0.1%)	(-6.8%, +7.1%)	(-6.8%,+7%)					
$A_0^{B_d}$	(-22%, +32%)		(-24%, +28%)					
$A_0^{B_s}$	(-28%, +33%)	(-28%, +33%)						
$\lambda_{B_d}$	(-0.6%, +0.2%)	(-4.6%, +2.1%)	(-4.1%, +1.9%)					
$\alpha_2^{K^*}$	(-0.1%, +0.1%)	(-3.6%, +3.7%)	(-3.6%, +3.6%)					
$X_H$	(-0.2%, +0.2%)	(-1.8%, +1.8%)	(-1.6%, +1.6%)					
$X_A$	(-4.3%, +4.4%)	(-17%, +19%)	(-13%, +14%)					
$\kappa$	(-1.4%, +2.2%)							
Others	(-1.3%, +1.1%)	(-2.7%, +2.5%)	(-1.6%, +1.6%)					

**Table 2**. Error budget of  $L_{K^*\bar{K}^*}$  and  $|P_{d,s}|^2$ . The relative error of each theoretical input is obtained by varying them individually. The main sources of uncertainty are the form factors, followed by weak annihilation at a significantly smaller level.

					1	R. C. S. C.	
	$B_{d,s}$ Distr	ibution Amp	litudes (at	at $\mu = 1$ GeV	7) [34, 35]		
$\lambda_{B_{i}}$	[GeV]	1	$\lambda_{B_s}/\lambda_B$	34	/[/]	$\sigma_B$	
0.383	$3 \pm 0.153$		$1.19 \pm 0.$	.14	1.4	$\pm 0.4$	
	$K^*$ Dis	tribution Am	plitudes (	(at $\mu = 2$ Ge	eV) [36]		
$\alpha_1^{K^*}$		$\alpha_{1}^{K^*}$	<u> </u>	$\alpha_2^{K^*}$		$\alpha_{2}^{K^*}$	
$0.02 \pm 0.02$	.02	$0.03 \pm 0.03$		$0.08 \pm 0.06$	0	$.08 \pm 0.06$	;
	φ Dist	ribution Am	plitudes (a	at $\mu = 2$ Ge	V) [36]		
$\alpha_1^{\phi}$	$\alpha_1^{\phi}$		α	$\chi_2^{\phi}$		$\alpha_{2\perp}^{\phi}$	
0	(	)	0.13 ±	± 0.06	0.1	$1 \pm 0.05$	
Deca	y Constants i	for B mesons	s (at $\mu = 2$	2 GeV) [37]	and K meso	on [28]	
f	B <sub>d</sub>		$f_{B_s}/f_{B_d}$		f	K	
0.190 =	$\pm 0.0013$	1.2	$009 \pm 0.005$	5	0.1557 ±	$\pm 0.0003$	
	Decay Con	stants for $K$	*, $\phi, \rho, \omega$ (i	at $\mu = 2$ Ge	V) [26, 38]		
$f_{K^*}$	$f_{K^*}^{\perp}/f_{K^*}$	$f_{\phi}$	$f_{i}$	$f_{\phi}^{\perp}/f_{\phi}$	$f_{ ho}$	$f_{\omega}$	
$0.204 \pm 0.007$	$0.712 \pm 0.01$	$2 0.233 \pm 0.$	.004 0.750	$0 \pm 0.008$ 0.	$213 \pm 0.005$	$0.197 \pm 0$	0.008
1	$B_{d,s} \to K^*, \phi$	form factors	[26] and B	B-meson life	times (ps) [3	39]	
$A_0^{B_s \to K^*}(q^2 =$	$= m_{\phi}^2 A_0^{B_d \rightarrow K}$	$(q^2 = m_{\phi}^2)$	$A_0^{B_s \to \phi}(q^2$	$^{2} = m_{K^{*}}^{2})$	$ au_{B_d}$	$\tau_{B_s}$	
$0.380 \pm 0.0$	0.39	$3 \pm 0.039$	$0.438 \pm$	± 0.024 1	$.519 \pm 0.004$	$1.520 \pm$	0.005
	Mas	s and decay	widths for	$r \rho, \omega \text{ (GeV)}$	[28]		
$m_{ ho}$		$\Gamma_{ ho}$		$m_\omega$		$\Gamma_{\omega}$	
0.7745		0.1484		0.7827		0.0087	
	$B_d \rightarrow K$	$[25], B_s \to K$	[40] and	$B_s \to \phi$ for	m factors		
$f_0^{B_s}(q^2$	$= m_{\phi}^{2})$	$f_0^{B_d}($	$(q^2 = m_\phi^2)$		$A_0^{B_s \to \phi}(q^2)$	$^{2} = m_{K}^{2}$	
0.336 ±	± 0.023	0.34	$0 \pm 0.011$		$0.426 \pm$	: 0.024	
		Wolfenste	ein parame	eters [41]			
A		λ		$\bar{\rho}$		$\bar{\eta}$	
$0.8132^{+0.0}_{-0.0}$	0000 0	$.22500^{+0.0002}_{-0.0002}$	4 2	$0.1566^{+0.008}_{-0.004}$	8 0.	$3475^{+0.011}_{-0.005}$	18 54
		QCD scale a	nd masses	s [GeV] [28]			
$\bar{m}_b(\bar{m}_b)$	$m_b/m_c$	$m_{B_d}$	$m_{B_s}$	$m_{K^*}$	$m_{\phi}$ n	$n_K = \Lambda$	QCD
4.18 4.	$577 \pm 0.008$	5.27966 5.	.36692 0	0.89555 1.	01946 0.49	97611 0	.225
	SM	Wilson Coeff	ficients (at	t $\mu = 4.18$ G	leV)	_	
$\mathcal{C}_1$	$\mathcal{C}_2$	$\mathcal{C}_3$		$\mathcal{C}_4$	$C_5$	$C_6$	
1.082	-0.191	0.014		-0.036	0.009	-0.04	42 Ŧ
$C_7/\alpha_{em}$	$C_8/\alpha_{em}$	$C_9/\alpha_{er}$	m C	$c_{10}/\alpha_{em}$	$C_{7\gamma}^{en}$	C <sub>8g</sub>	51
-0.011	0.060	-1.254	¥	0.224	-0.318	-0.13	51 I G

	MLR	CDF
$L_{K^*\bar{K}^*}$	$17.2^{+8.3}_{-5.9}$	$19.5_{-6.7}^{+9.1}$
$L_{K\bar{K}}$	$25.5^{+4.0}_{-3.3}$	$26.0^{+3.9}_{-3.6}$
$\hat{L}_{K^*}$	$20.5^{+6.8}_{-6.2}$	$21.3^{+7.2}_{-6.3}$
$\hat{L}_K$	$25.3^{+3.7}_{-4.5}$	$25.0^{+4.2}_{-4.1}$
$L_{K^*}$	$16.6\substack{+6.9\\-6.0}$	$17.4_{-5.8}^{+6.6}$
$L_K$	$28.8^{+5.2}_{-4.6}$	$29.2^{+5.5}_{-5.3}$
$L_{\rm total}$	$23.5^{+3.8}_{-4.0}$	$23.5^{+4.0}_{-3.8}$
$R_d$	$0.67\substack{+0.23\\-0.24}$	$0.70\substack{+0.30\\-0.22}$
$\mathcal{B}(B_d \to K^{*0} \bar{K}^{*0}) \times 10^6$	$0.22\substack{+0.08\\-0.08}$	$0.23\substack{+0.10 \\ -0.08}$
$\mathcal{B}(B_s \to K^{*0} \bar{K}^{*0}) \times 10^6$	$3.95^{+1.88}_{-1.54}$	$4.36^{+2.23}_{-1.65}$
$\mathcal{B}(B_d \to K^0 \bar{K}^0) \times 10^6$	$1.01\substack{+0.24\\-0.16}$	$1.09\substack{+0.29\\-0.20}$
$\mathcal{B}(B_s \to K^0 \bar{K}^0) \times 10^6$	$25.6^{+7.5}_{-5.2}$	$28.0^{+8.9}_{-6.2}$



Figure 3: Hard spectator diagrams.





Figure 4: Annihilation diagrams.

#### Main caveat:

(Existence of some) **Power suppressed** but **IR divergent** spectator scattering and weak annihilation that affects amplitudes:

