

Walking in the Hidden Valley

- Exploring near-conformal dark sectors
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(Paper to appear!)





arXiv:0604261, M.J. Strassler et al. arXiv:1502.05409, P. Schwaller et al. arXiv:1503.00009, T. Cohen et al. arXiv:0712.2041, T. Han et al.

Standard Model $SU(3)_C \times SU(2)_L$ $\times U(1)_{V}$

- representation of $SU(N_C)$.
- largely undeveloped area of theory space and could give rise to distinct signatures.
- large N_F/N_C theories; can be used to build up near-conformal dark sector models in the future.

arXiv:2305.03665, T. Appelquist et al. arXiv:2306.07236, A. Hasenfratz et al. arXiv:2312.13761, R. Zwicky

arXiv:1610.01752, M. Golterman et al.

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Confining dark sector models

Heavy Z' mediator U(1)'

 $SU(N_C)$ dark sector gauge group with N_F dark quarks.

Hidden Valley models that extend the SM with a new confining dark sector resembling QCD present exciting opportunities for new physics discovery. We will focus on light dark quarks; Dirac fermions in the fundamental

• At colliders, these dark quarks undergo parton showering, hadronisation and decay giving rise to anomalous jet signatures (e.g. emerging or semi-visible jets) that are well-known at small N_F/N_C . Large N_F/N_C dark sectors are a

• There is a plethora of work focused on the non-perturbative structure and the low-energy effective descriptions of

arXiv:2312.08332, A. Pomarol et al. arXiv:2404.07601, T. Appelquist et al.



Dark parton showering



lpha runs, controlled by N_F/N_C and Λ

- (RGE). G. 't Hooft Nucl. Phys. B ('74)

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$$\mu^2 \frac{d\alpha}{d\mu^2} = \beta(\alpha) = -\alpha^2 \left(\beta_0 + \beta_1 \alpha\right) \quad \text{(at 2-loo}$$

$$N_C \alpha = f(N_F/N_C, \mu/\Lambda) + \mathcal{O}(N_F/N_C^3) \text{ correction}$$
Non-trivial
$$\alpha_* = -\frac{\beta_0}{\beta_1} ; > 0 \text{ for } N_F/N_C \gtrsim 2.7$$
fixed point:

First, we'll focus on dark parton showering; hadronisation details not yet fully understood. Dark parton showering is governed by the t' Hooft gauge coupling, $N_C \alpha$ with α in turn being governed by the Renormalisation Group Equations

• Parton shower ends near scale Λ , this scale characterises the breakdown of the perturbative expansion of α and is a good proxy for the scale of the theory. Up to corrections, the t' Hooft gauge coupling is governed solely by N_F/N_C and μ/Λ .

At two-loop, for $N_F/N_C \gtrsim 2.7$, α flows to a non-trivial infra-red fixed point (IRFP); as N_F/N_C increases α begins to slow down. New procedures are needed to understand parton showering within this region. T. Banks., A. Zaks, Nucl. Phys. B 196 ('82)



Near-conformal dark sector models

arXiv:2008.12223, J.W. Lee

Two-loop perturbative description



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Non-perturbative description

- mass (m_q) additionally deforms the IR conformal symmetry. These ensure any signatures are not purely missing energy.





Modelling of dark parton showers



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• The current approximation used within event generators (the PDG) formula) is insufficient to describe two-loop α for high N_F/N_C since it neglects effects of the IRFP.

By taking this IRFP into account, we establish a framework of two solutions to the RGE that allow for parton showering to be simulated across a wide range of parameter space.

$$\alpha_* \left[W_{-1} \left(-\frac{1}{e} \left(\frac{\mu^2}{\Lambda^2} \right)^{\beta_0 \alpha_*} \right) + 1 \right]^{-1} \quad ; \qquad \alpha = \alpha_* \left[W_0 \left(\frac{1}{e} \left(\frac{\mu^2}{\Lambda^2} \right)^{\beta_0 \alpha_*} \right) + 1 \right]^{-1} \quad ; \qquad \alpha = \alpha_* \left[W_0 \left(\frac{1}{e} \left(\frac{\mu^2}{\Lambda^2} \right)^{\beta_0 \alpha_*} \right) + 1 \right]^{-1} \quad ; \qquad \alpha = \alpha_* \left[W_0 \left(\frac{1}{e} \left(\frac{\mu^2}{\Lambda^2} \right)^{\beta_0 \alpha_*} \right) + 1 \right]^{-1} \quad ; \qquad \alpha = \alpha_* \left[W_0 \left(\frac{1}{e} \left(\frac{\mu^2}{\Lambda^2} \right)^{\beta_0 \alpha_*} \right) + 1 \right]^{-1} \quad ; \qquad \alpha = \alpha_* \left[W_0 \left(\frac{1}{e} \left(\frac{\mu^2}{\Lambda^2} \right)^{\beta_0 \alpha_*} \right) + 1 \right]^{-1} \quad ; \qquad \alpha = \alpha_* \left[W_0 \left(\frac{1}{e} \left(\frac{\mu^2}{\Lambda^2} \right)^{\beta_0 \alpha_*} \right) + 1 \right]^{-1} \right]^{-1} \quad ; \qquad \alpha = \alpha_* \left[W_0 \left(\frac{1}{e} \left(\frac{\mu^2}{\Lambda^2} \right)^{\beta_0 \alpha_*} \right) + 1 \right]^{-1} \left[W_0 \left(\frac{1}{e} \left(\frac{\mu^2}{\Lambda^2} \right)^{\beta_0 \alpha_*} \right) + 1 \right]^{-1} \right]^{-1} \quad ; \qquad \alpha = \alpha_* \left[W_0 \left(\frac{1}{e} \left(\frac{\mu^2}{\Lambda^2} \right)^{\beta_0 \alpha_*} \right) + 1 \right]^{-1} \left[W_0 \left(\frac{1}{e} \left(\frac{\mu^2}{\Lambda^2} \right)^{\beta_0 \alpha_*} \right) + 1 \right]^{-1} \right]^{-1} \quad ; \qquad \alpha = \alpha_* \left[W_0 \left(\frac{1}{e} \left(\frac{\mu^2}{\Lambda^2} \right)^{\beta_0 \alpha_*} \right) + 1 \right]^{-1} \left[W_0 \left(\frac{1}{e} \left(\frac{\mu^2}{\Lambda^2} \right)^{\beta_0 \alpha_*} \right) + 1 \right]^{-1} \right]^{-1} \quad ; \qquad \alpha = \alpha_* \left[W_0 \left(\frac{1}{e} \left(\frac{\mu^2}{\Lambda^2} \right)^{\beta_0 \alpha_*} \right) + 1 \right]^{-1} \left[W_0 \left(\frac{1}{e} \left(\frac{\mu^2}{\Lambda^2} \right)^{\beta_0 \alpha_*} \right) + 1 \right]^{-1} \left[W_0 \left(\frac{1}{e} \left(\frac{\mu^2}{\Lambda^2} \right)^{\beta_0 \alpha_*} \right) + 1 \right]^{-1} \right]^{-1} \left[W_0 \left(\frac{1}{e} \left(\frac{\mu^2}{\Lambda^2} \right)^{\beta_0 \alpha_*} \right]^{-1} \right]^{-1} \left[W_0 \left(\frac{1}{e} \left(\frac{\mu^2}{\Lambda^2} \right)^{\beta_0 \alpha_*} \right) + 1 \right]^{-1} \left[W_0 \left(\frac{\mu^2}{\Lambda^2} \right)^{\beta_0 \alpha_*} \right]^{-1} \left[W_0 \left(\frac{\mu^2}{\Lambda^2} \right)^{-1} \left[W_0 \left(\frac{\mu^2}{\Lambda^2} \right)^{\beta_0 \alpha_*} \right]^{-1} \left[W_0 \left(\frac{\mu^2}{\Lambda^2} \right)^{-1} \left[W_0 \left(\frac{\mu^2}{\Lambda^2} \right)^{-1} \right]^{-1} \left[W_0 \left(\frac{\mu^2}{\Lambda^2} \right]^{-1} \left[W_0 \left(\frac{\mu^2}{\Lambda^2} \right)^{-1} \left[W_0 \left(\frac{\mu^2}{\Lambda^2} \right)^{-1} \left[W_0 \left(\frac{\mu^2}{\Lambda^2} \right)^{-1} \left[W_0 \left(\frac{\mu^2}{\Lambda^2} \right$$

QL region (no IRFPs)

CW region (IRFPs)

This procedure defines Λ within the IRFP region as the scale at which power-law running takes over from logarithmic running.

arXiv:9602385, T. Appelquist et al. arXiv:1406.2337, D. Litim et al.

arXiv:9810192 - E. Gardi et al.





The Sudakov veto algorithm at two-loop

- arXiv:0603175 T. Sjöstrand et al.

$$Q_{i}^{2} = \Lambda^{2} \left(\frac{Q_{i-1}^{2}}{\Lambda^{2}} \right)^{\tilde{\Delta}^{2\pi\beta_{0}/e}} \left[\tilde{\Delta}^{2\pi\beta_{0}/e} \left(\mp e W_{n} \left(\mp \frac{1}{e} \left(\frac{Q_{i-1}^{2}}{\Lambda^{2}} \right)^{\beta_{0}\alpha_{*}} \right) \right)^{1-\tilde{\Delta}^{2\pi\beta_{0}/e}} \right]^{1/\beta_{0}\alpha_{*}}; \quad \epsilon = \int_{\tilde{\xi}_{min}}^{\tilde{\xi}_{max}} \tilde{P}_{a \to bc}(\xi') d\xi' \quad ; \quad \pi, n = \begin{cases} -, -1 & (\text{QL regions of } \xi)^{1/2} \\ +, 0 & (\text{CW regions of } \xi)^{1/2} \end{cases}$$

- even provides a way to simulate below $\mu/\Lambda = 1$ in the CW region.
- The Lambert W function is put within Pythia through a combination of approximation and interpolation.

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• To generate some scale Q_i^2 given some initial scale Q_{i-1}^2 , Pythia generates a random number R_1 and solves for $\Delta(Q_i^2, Q_{i-1}^2) = R_1$. Δ is the Sudakov form factor, the probability of no parton emissions between Q_{i-1}^2 and Q_i^2 .

arXiv:1102.2126 - W. Giele et al. arXiv:1101.2599 - A. Buckley et al. arXiv:1211.7204 - L. Lonnblad et al.

• This inversion is complicated, so Δ is usually overestimated with a much simpler $\tilde{\Delta}$. The overestimate is then corrected for through the Sudakov veto algorithm. At two-loop, the inverse can be performed exactly as,

Previous two-loop efforts relied on a separate veto algorithm that corrected one-loop showering with two-loop effects - a method which did not converge for the entire $N_F/N_C - \mu/\Lambda$ space. The new implementation now







Simulation of dark parton showers



- Theories with large IRFPs ($\alpha_* \gg 1$) (around $N_F/N_C \sim 3$) have similar parton showering behaviour to those without have both hard and soft partons; having an overall small multiplicity of hard partons.
- splitting at $N_F/N_C \sim 5$ and average parton multiplicity is almost 2 the 2 initial dark quarks.

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Simulated with a custom Pythia 8.307 with benchmark: $e^+e^- \rightarrow Z' \rightarrow q_D \overline{q_D}$, $\sqrt{s} = M_{Z'} = 1$ TeV , hadronisation off , $\Lambda=5$ GeV , $N_C = 3$.

IRFPs, having a overall large multiplicity of soft partons. Theories with small IRFPs ($\alpha_* \ll 1$) (around $N_F/N_C \sim 5$)

Since parton splitting probability is proportional to α , it thus vanishes as $N_F/N_C \rightarrow 5.5$. Hence there is very little









Simulation of dark parton showers

Showered dark parton $\textbf{p}_{\!_{\!\!\!\!\!\!\!\!\!\!\!\!\!}}$ and η distribution



- states, as this could have an additionally large influence on the dark shower phenomenology.

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Showered dark parton p_{τ} and η distribution

At around $N_F/N_C \sim 5$, there is a transition in the $p_T - \eta$ plane from the majority of partons being soft to a majority being hard, reflecting how branching probability is negligible and the majority of partons are initial dark quarks.

This new procedure allows for the simulation of the anomalous jets signatures of near-conformal Hidden Valley theories. Motivates further investigations into the hadronisation and subsequent decay of near-conformal bound







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Thank you! Questions?

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