

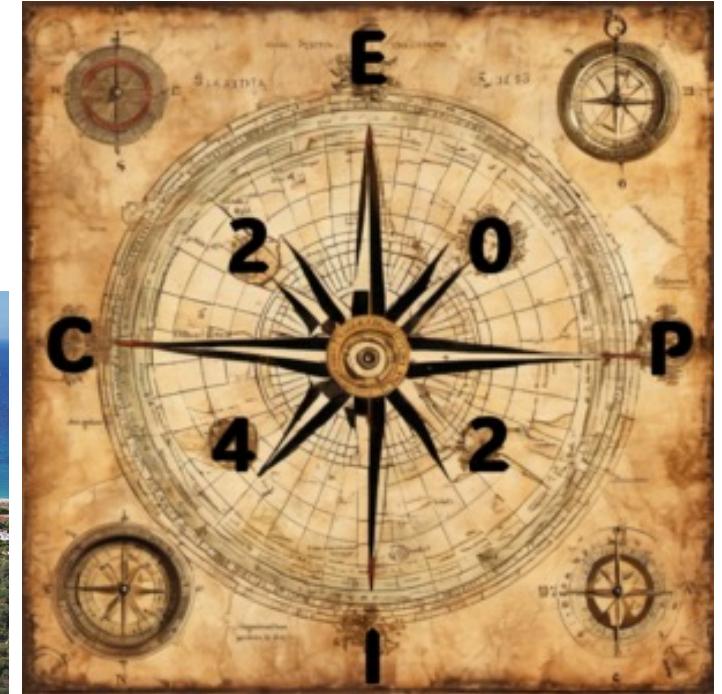
# EPIC 2024

Electroweak Physics Intersections



EPIC 2024

Electroweak Physics InterseCtions  
SEPT  
22-27, 2024  
Calaserena  
Geremeas, Italy



Simultaneous extraction of the weak radius  
and the weak mixing angle from EW probes

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EPIC 2024 - 23 September 2024

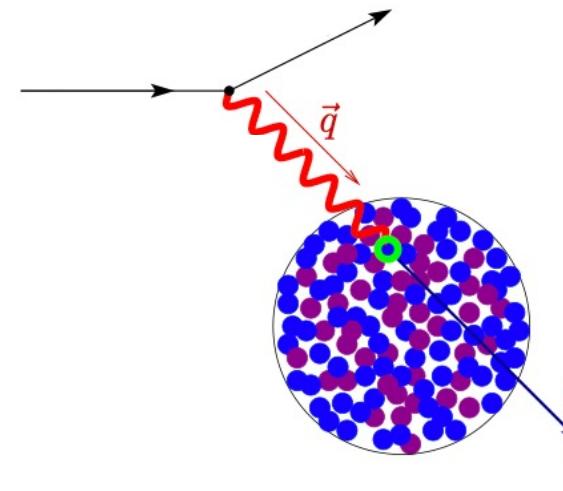


The spatial dimension probed in a scattering experiment depends on the wavelength of the mediator

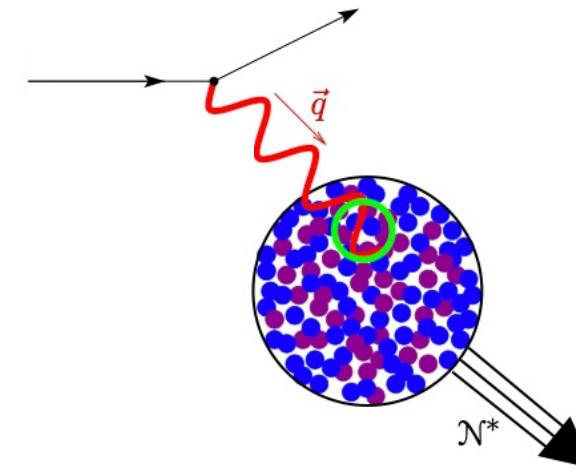
$$\text{De Broglie wavelength: } \lambda = \frac{h}{p}$$

The momentum of the mediator is of the order of the momentum transfer  $p \sim q$

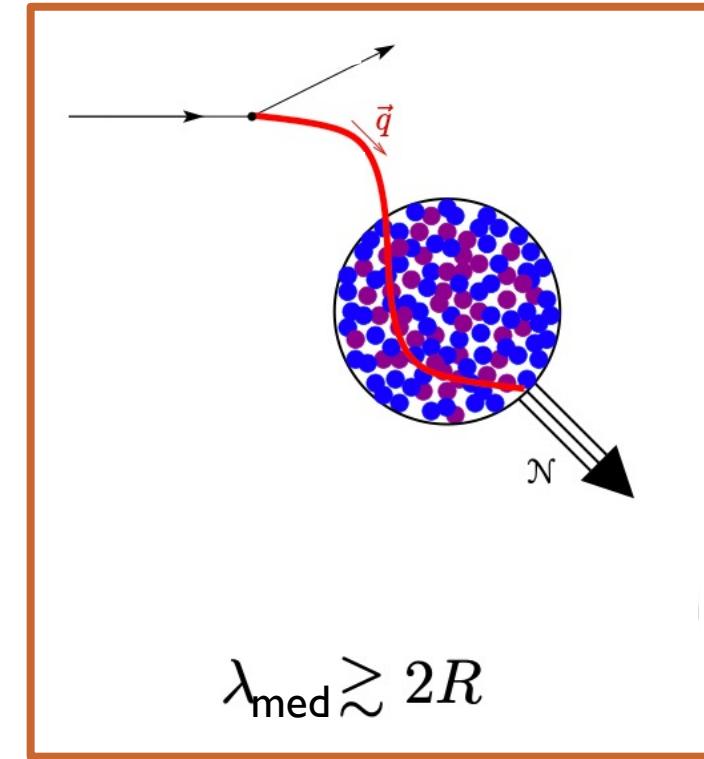
M. Cadeddu et al.,  
EPL 143 (2023) 3, 34001



$$\lambda_{\text{med}} \ll 2R$$



$$\lambda_{\text{med}} \lesssim 2R$$



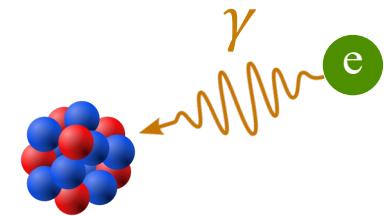
$$\lambda_{\text{med}} \gtrsim 2R$$

$$\lambda \sim \frac{200 \text{ MeV fm}}{q} \text{ for } q = 20 \text{ MeV} \rightarrow \lambda \cong 10 \text{ fm} \longrightarrow \text{MeV scale neutrinos and MeV-GeV scale electrons}$$



The dominant process is electromagnetic scattering mediated by the photon

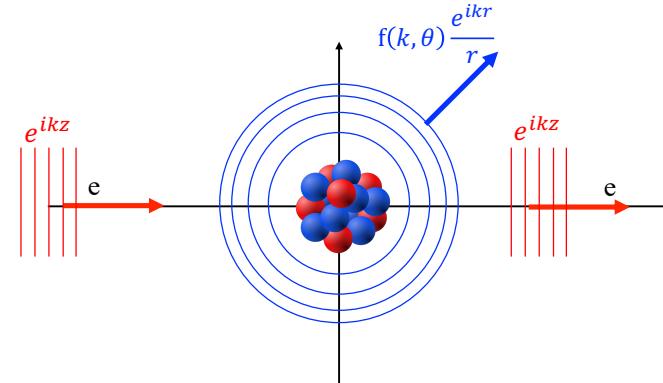
- **Rutherford scattering:**  $\left(\frac{d\sigma}{d\Omega}\right)_{\text{Ruth}} \propto \frac{Z\alpha^2}{4E_e^2} \frac{1}{\sin^4\frac{\theta}{2}}$  Point-like nucleus and nonrelativistic scattering



- **Mott Scattering:**  $\left(\frac{d\sigma}{d\Omega}\right)_{\text{Mott}} \propto \frac{Z\alpha^2}{4E_e^2} \frac{\cos^2\frac{\theta}{2}}{\sin^4\frac{\theta}{2}}$  Point-like nucleus and relativistic scattering

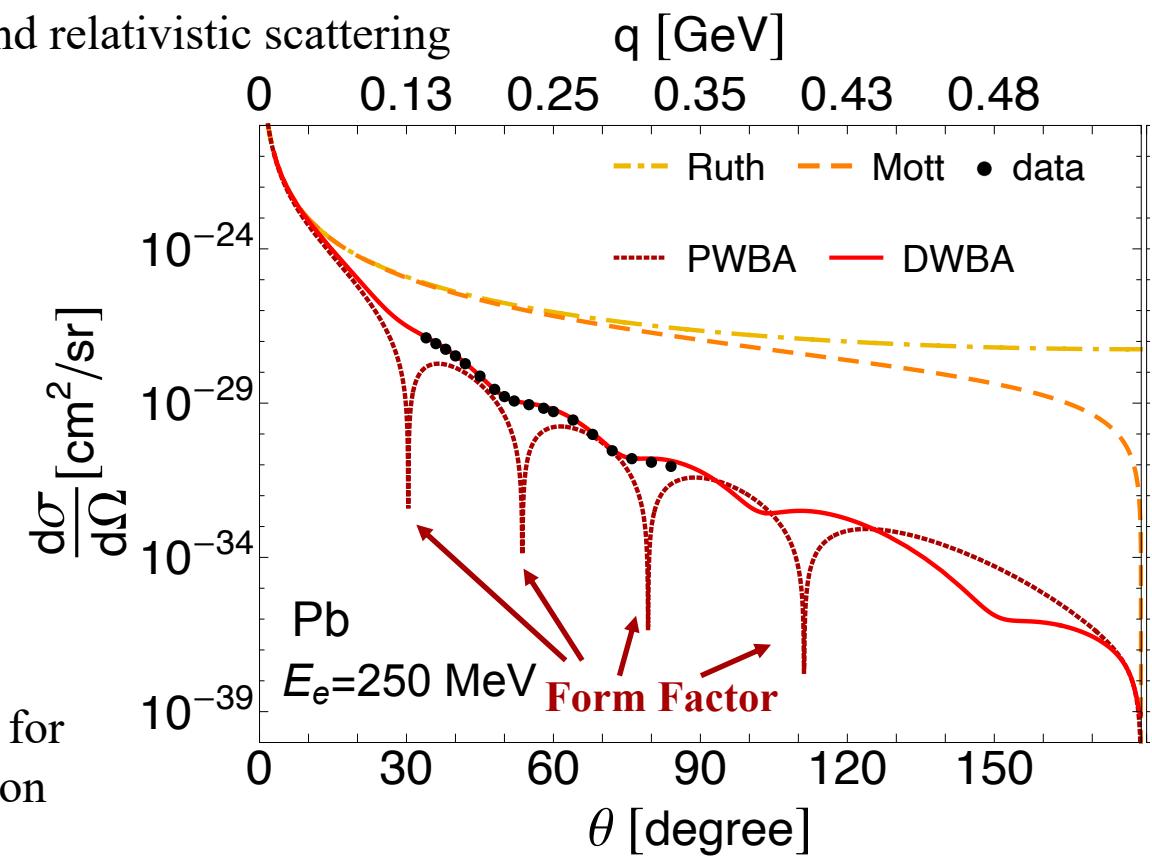
- **PWBA:**  $\left(\frac{d\sigma}{d\Omega}\right)_{\text{PWBA}} = \left(\frac{d\sigma}{d\Omega}\right)_{\text{Mott}} |F_{\text{ch}}(q^2)|^2$

Extended nucleus and relativistic scattering

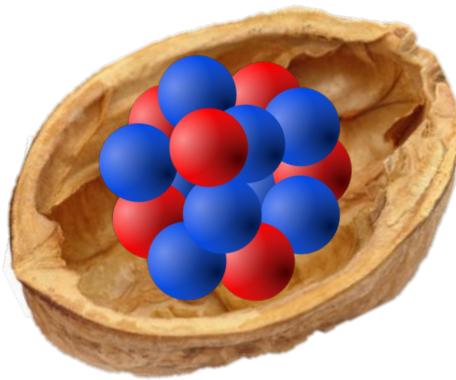


- **DWBA:** Distorted Wave Born Approximation which accounts for the effect of the Coulomb potential on the electron wavefunction

Numerical code to solve the Dirac Equation  
in a central potential



Only CHARGE and PROTON distributions



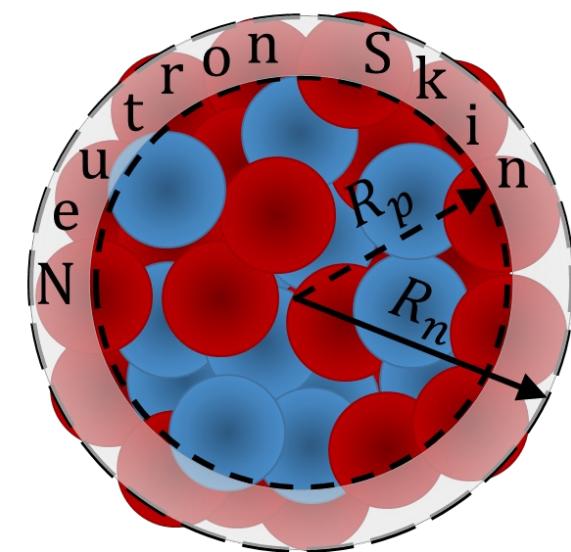
Nucleus: compact object made of nucleons and held up together from the strong force

Distribution of nucleons inside the nucleus: mass, charge, **weak** etc.

In heavy nuclei ( $N > Z$ ) protons are expected to be distributed over a smaller radius than neutrons: NEUTRON SKIN  $\Delta R_{np} = R_n - R_p$

**Neutron skin: result of the competition between the Coulomb repulsion, the surface tension and the symmetry energy**

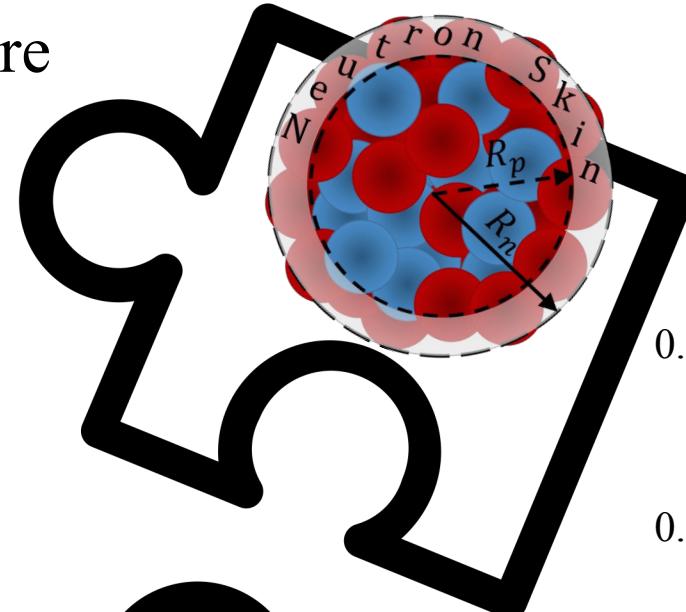
- Surface tension: decreases when the excess neutrons are pushed to the surface
- The neutron skin has implications not only in nuclear theory
- Behaviour of neutron-rich matter such as the one inside the inner core of neutron stars
- EOS of neutron stars: symmetry energy  $S$  and slope parameter  $L$  (derivative of  $S$  wrt density)



Larger neutron skins  $\longrightarrow$  Larger L parameters  $\longrightarrow$  Larger size of neutron stars



Nuclear Structure  
through the  
neutron skin

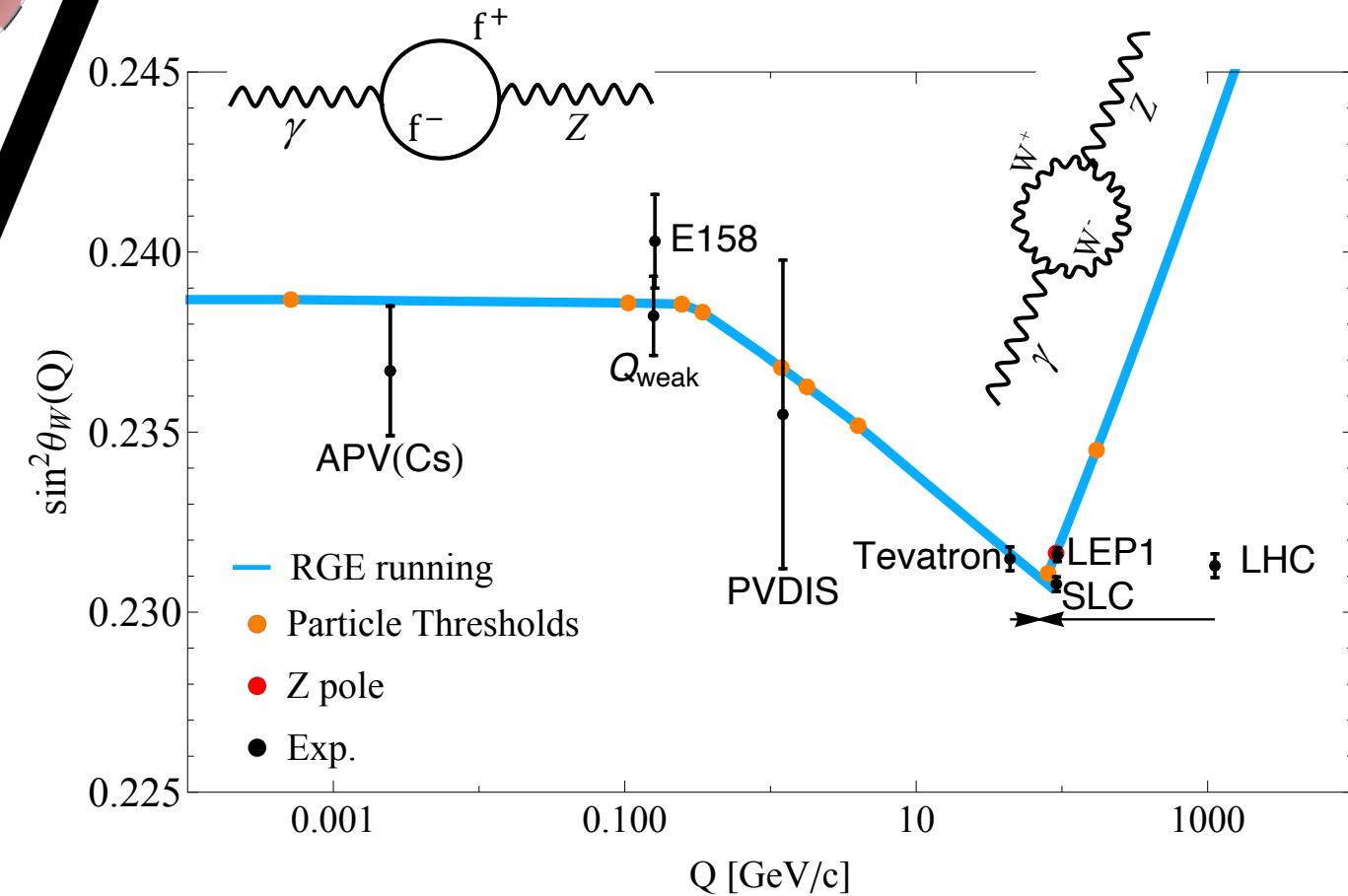


Neutrino  
Scattering  
on Nuclei (CEvNS)

Electron  
Scattering  
on Nuclei (PVES)

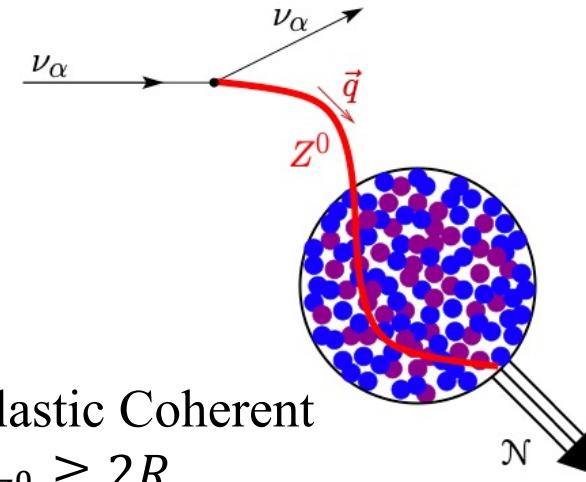
e  
Atomic  
systems  
(APV)

Weak mixing angle: key  
parameter of the EW theory





**CEvNS:** coherent elastic interaction of the neutrino with the nucleus via a purely NC process



Elastic Coherent  
 $\lambda_{Z^0} \gtrsim 2R$

$$\frac{d\sigma_{\nu_\ell - N}(E_\nu, T_{nr})}{dT_{nr}} \cong \frac{G_F^2 m_N}{\pi} \left(1 - \frac{m_N T_{nr}}{2E_\nu^2}\right) [Q_W F_W(q^2)]^2$$

Nuclear weak charge  $Q_W$   
(depends on the weak mixing angle)

Weak Form Factor (describes  
the nuclear structure effects)

$$Q_W F_W(q^2) = g_V^p (\sin^2(\vartheta_W)) Z F_Z(|\vec{q}|^2) + g_V^n N F_N(|\vec{q}|^2)$$

$$g_V^p(\nu_e) = 0.0382$$

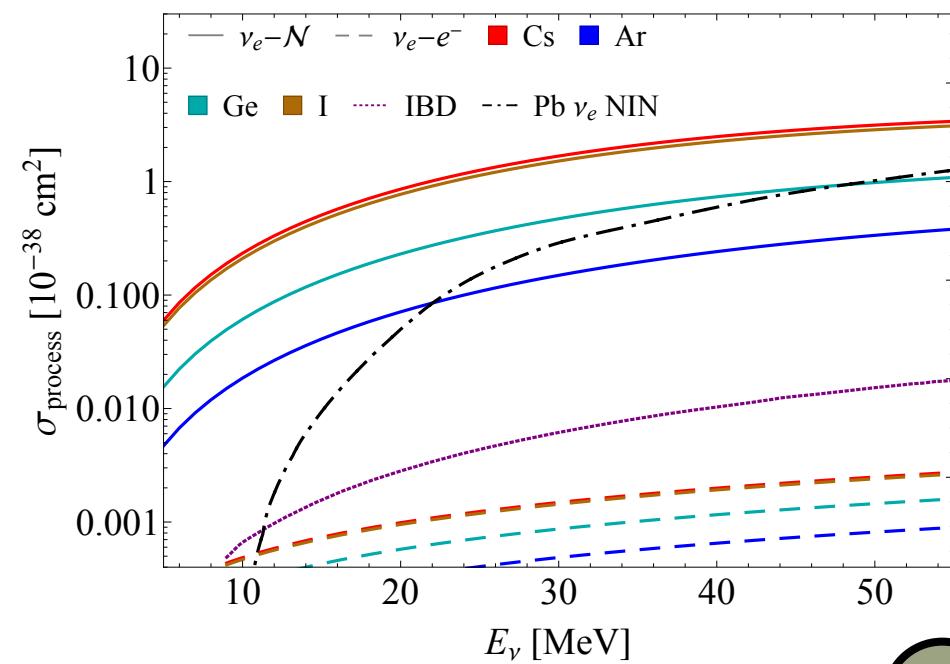
$$g_V^p(\nu_\mu) = 0.0300$$

$$g_V^n = -0.5117$$

Neutrinos couple  
mainly to neutrons!  
(Large cross  
section:  $\propto N^2$ )

Neutron nuclear  
distribution

Weak nuclear  
distribution





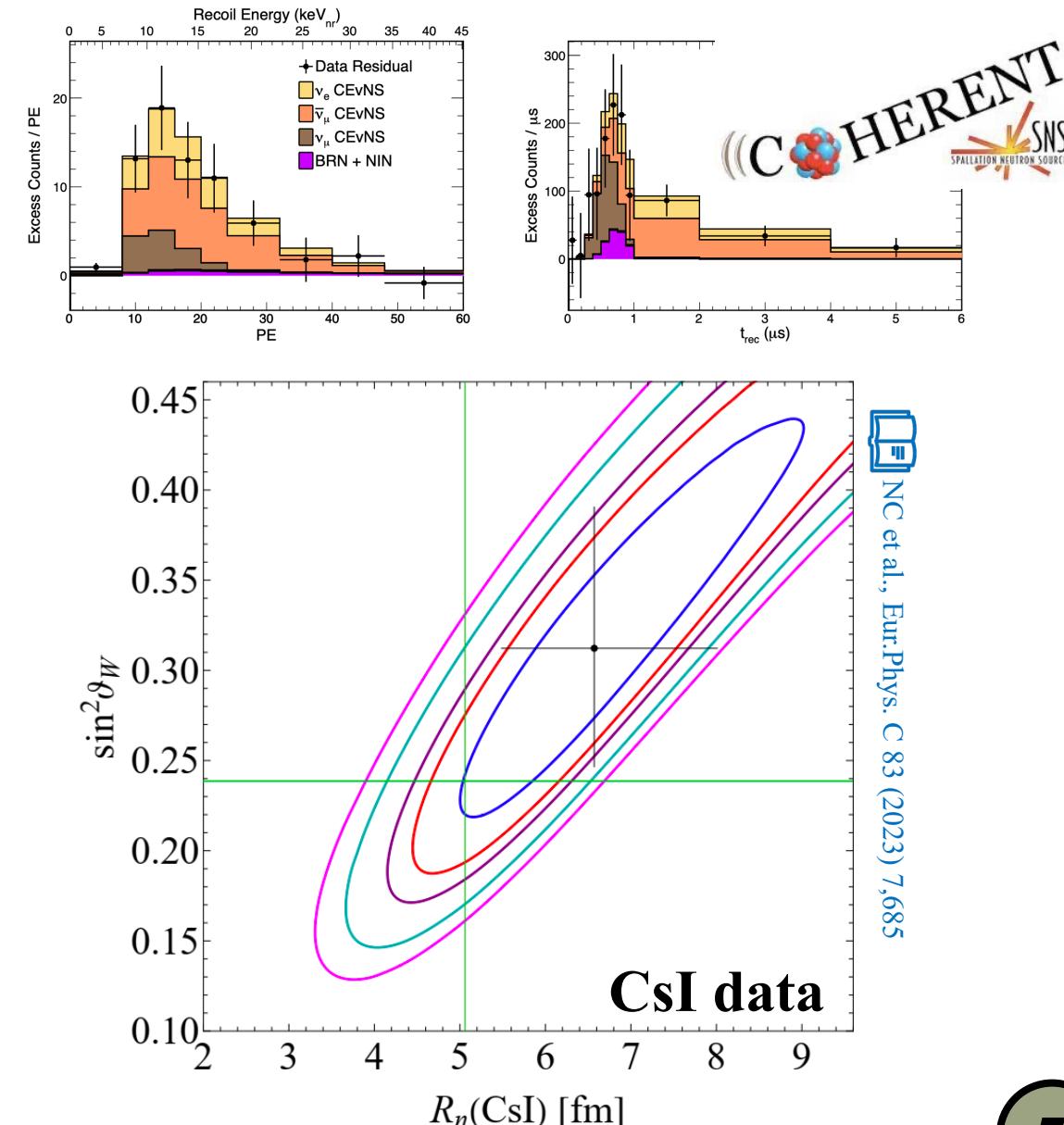
## COHERENT experimental program

- CsI crystal: D. Akimov et al. Phys.Rev.Lett. 129 (2022) 8,081801
  - 14.6 kg scintillating crystal
  - 19.3 m away from the SNS target
  - $q \sim 50$  MeV
  - $\sim 11\%$  on  $\sin^2 \theta_W$
  - $\sim 7\%$  on  $R_n(Cs)$  NC et al., Eur.Phys. C 83 (2023) 7,685
  - Future detector  $\sim 0.5\%$  on  $R_n(Cs)$

**COHERENT cannot provide (yet) a simultaneous determination of  $\sin^2 \theta_W$  and  $R_n(Cs)$  with competitive precision**

**Need for a complementary probe to disentangle the two parameters!**

- Ar single phase: D. Akimov et al. Phys.Rev.Lett. 126 (2021) 1, 012002
  - 24 kg of atmospheric argon
  - 27.5 m away from the SNS target
  - $q \sim 50$  MeV
  - Worse precision





Electrons normally interact with nuclear matter via photon exchange (EM interaction)

But there is also a subdominant weak contribution to the interaction

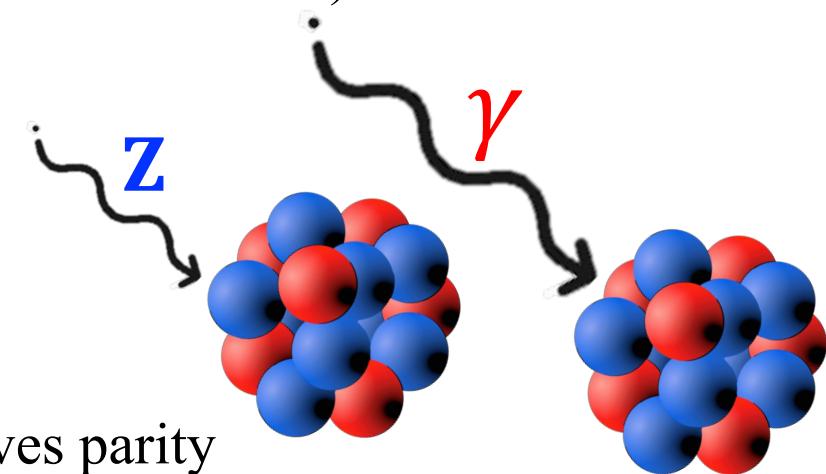
We need to select the weak part of the interaction :

## Parity Violation



- The electromagnetic interaction preserves parity
- The weak interaction violates parity

We need to isolate parity-violating contributions



### Atomic transitions forbidden by selection rules (APV)

- For atomic experiments

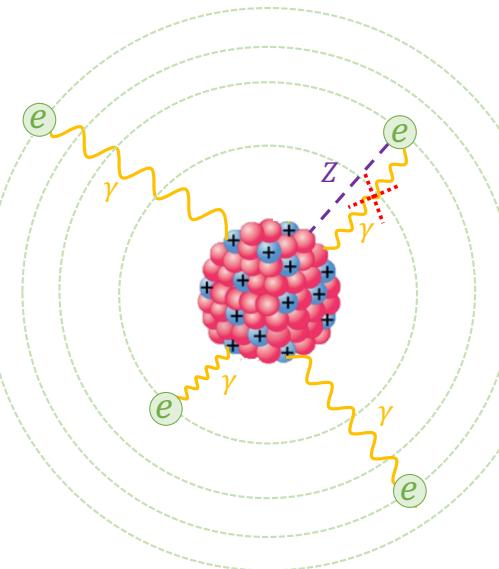
### Polarized electron scattering (PVES)

- For scattering experiments



APV experiments measure the weak mixing angle at very low momentum transfer via electronic transitions which are forbidden by selection rules: they violate parity

We consider the APV measurement on **cesium-133** and on **lead-208**



Interaction Hamiltonian

$$\hat{h}_{SI} = \frac{G_F}{2\sqrt{2}} Q_W \rho_{wk}(r) \gamma_5$$

Nuclear weak charge

↓ Evaluated between the involved atomic states  
Parity-violating amplitude\*  $E_{PNC}$

$$Q_W^{APV}(R_n) = N \left( \frac{\text{Im } E_{PNC}(R_n)}{Q_{weak}} \right)_{exp} \left( \frac{Q_{weak}}{N \text{Im } E_{PNC}(R_n)} \right)_{th} \beta_{exp+th}$$

↓ To be compared to the value of the nuclear weak charge calculated in the SM:  
dependence on  $\sin^2(\vartheta_W)$

\*There are different atomic calculations available for Cs, we adopt

B.K. Sahoo et al. Phys Rev D 103, L111303 (2021) or  
V. Dzuba et al., Phys. Rev. Letter 109, 203003 (2012)

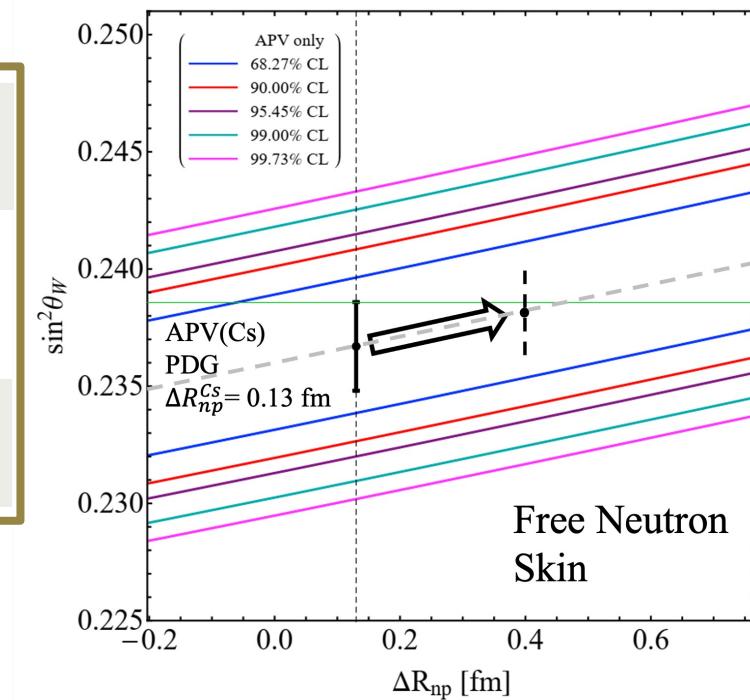
Similar, but less precise, measurement on lead-208.  
There are two experimental determinations available

$$R_{exp} = -9.80(33) \times 10^{-8}$$

$$R_{exp} = -9.86(12) \times 10^{-8}$$

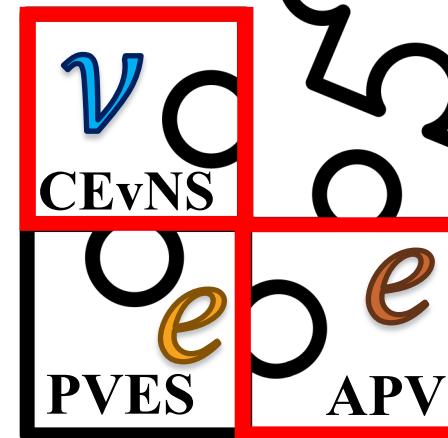
S.J. Phipp et al. Journal of Physics B 29, 1861 (1996)  
D.M. Meekhof et al. Phys Rev Lett 71, 3442 (1993)

Very sensitive to  $\sin^2 \theta_W$ ,  
not much on  $R_n(Cs)$



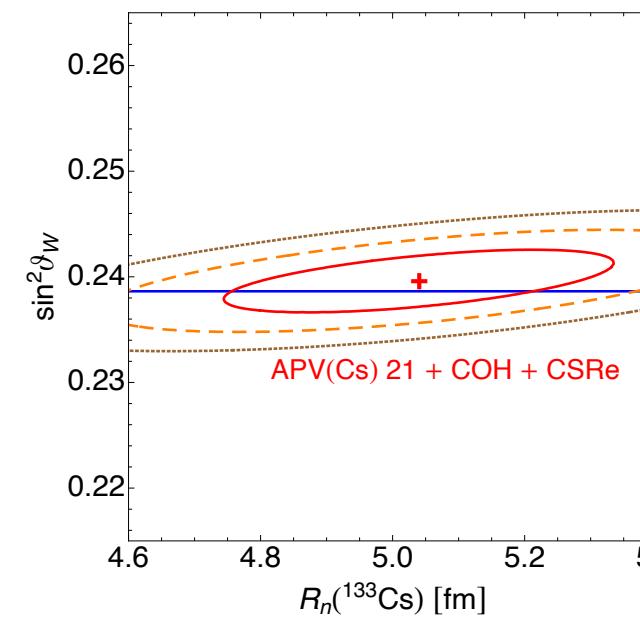
Extremely low momentum transfer  
 $Q_{Cs} \sim 2.4 \text{ MeV}$  and  $Q_{Pb} \sim 8 \text{ MeV}$





We can combine different probes to exploit different sensitivities to  $\sin^2 \theta_W$  and  $R_n(Cs)$  (different slopes)

New direct measurement of the cesium-133 neutron skin,  $\Delta R_{np}(Cs) = 0.12 \pm 0.21$  fm available from proton-cesium elastic scattering at the Cooler Storage Ring (CSRe)

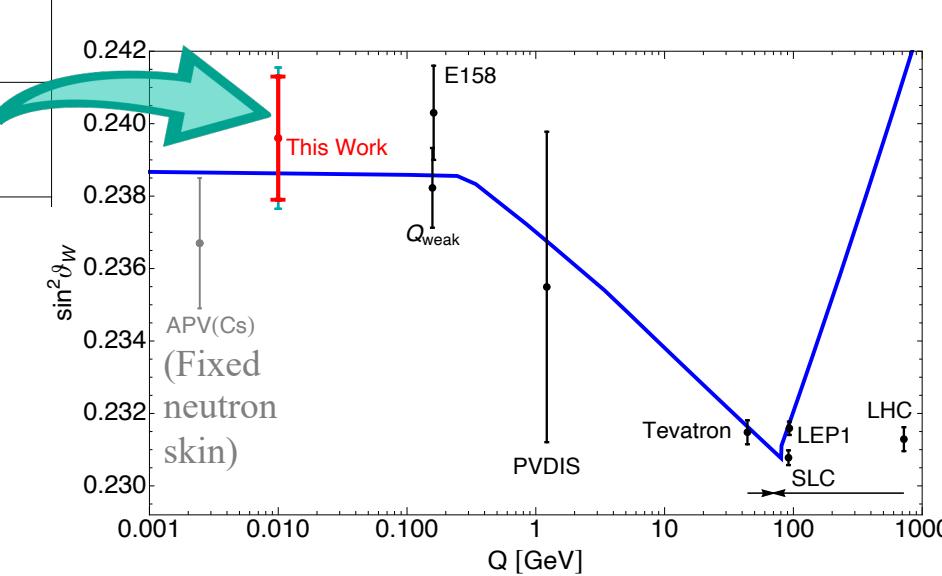
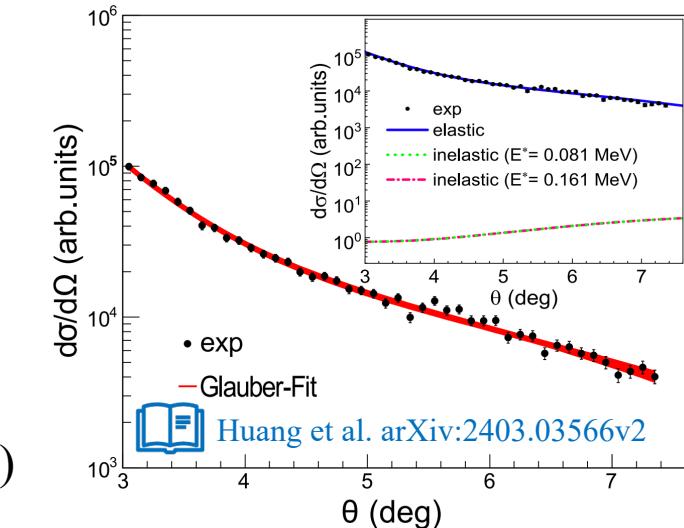


Meas.	$R_n(Cs)$ [fm]	$\sin^2 \theta_W$
APV(Cs) 21 + COH + CSRe	$5.04^{+0.19}_{-0.19}$	$0.2396^{+0.0020}_{-0.0019}$

The precision improves significantly by combining the available probes on Cs

- $\sim 3.8\%$  on  $R_n(Cs)$
- $\sim 8\%$  on  $\sin^2 \theta_W$

See M. Atzori  
Corona's poster!





In the case of electron-nucleon scattering:

$$\sigma_{\text{TOT}} \propto |\mathcal{A}_{\text{EM}} + \mathcal{A}_{\text{wk}}|^2 \propto |\mathcal{A}_{\text{EM}}|^2 + |\mathcal{A}_{\text{wk}}|^2 + 2|\mathcal{A}_{\text{EM}}| \cdot |\mathcal{A}_{\text{wk}}|$$

negligible

We can isolate the weak contribution by polarizing the electron beam

- **Parity-violating asymmetry in the case of scattering**

$$A_{\text{pv}} = \frac{\sigma_{\text{TOT}}^{\uparrow} - \sigma_{\text{TOT}}^{\downarrow}}{\sigma_{\text{TOT}}^{\uparrow} + \sigma_{\text{TOT}}^{\downarrow}} \propto \frac{|\mathcal{A}_{\text{EM}}^{\uparrow}| \cdot |\mathcal{A}_{\text{wk}}^{\uparrow}|}{|\mathcal{A}_{\text{EM}}^{\uparrow}|^2} = \frac{|\mathcal{A}_{\text{wk}}^{\uparrow}|}{|\mathcal{A}_{\text{EM}}^{\uparrow}|}$$



$$A_{\text{pv}} = -\frac{G_F Q^2}{4\sqrt{2}\pi\alpha} \frac{Q_W}{Z} \frac{F_{\text{wk}}(q^2)}{F_{\text{ch}}(q^2)}$$

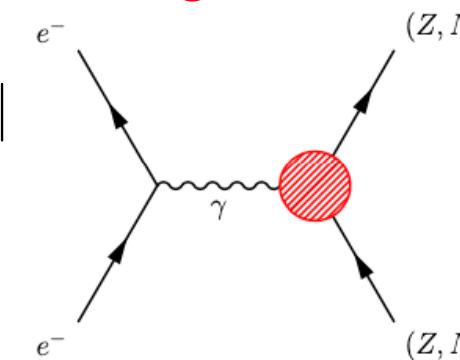
Nuclear weak charge  
Weak form factor  
Charge form factor

Known through electromagnetic measurements!

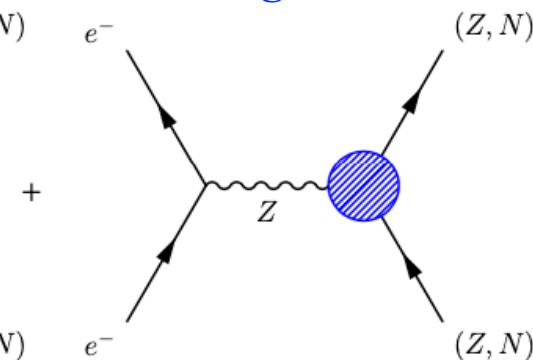
This definition works in the plane wave Born approximation: we need to account for Coulomb distortions

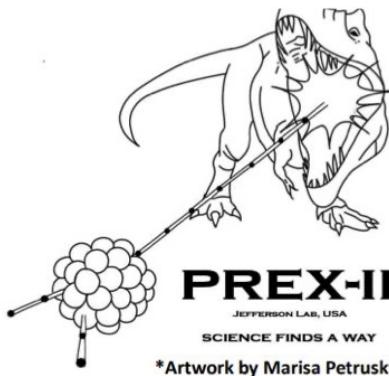
Similarly to CEvNS, the nuclear weak charge and the weak form factor enter as a product in the observable

### Electromagnetic Diagram



### Weak Diagram





We can consider the PREX measurement of the lead-208 neutron skin  
 1 GeV polarized electrons scattered at  $\theta \sim 5^\circ$  ( $q \sim 80$  MeV)  
 First measurement in 2012, updated one in 2021



S. Abrahamyan et al., Phys. Rev. Lett. 108.112502 (2012)  
 D. Adhikari et al., Phys. Rev. Lett. 126.172502 (2021)

Measured the neutron skin of lead with unprecedented precision:

$$\Delta R_{np}(\text{Pb}) = 0.278 \pm 0.078(\text{exp}) \pm 0.012(\text{th}) \text{ fm}$$

PREX found a rather thick neutron skin compared with EDF predicted values

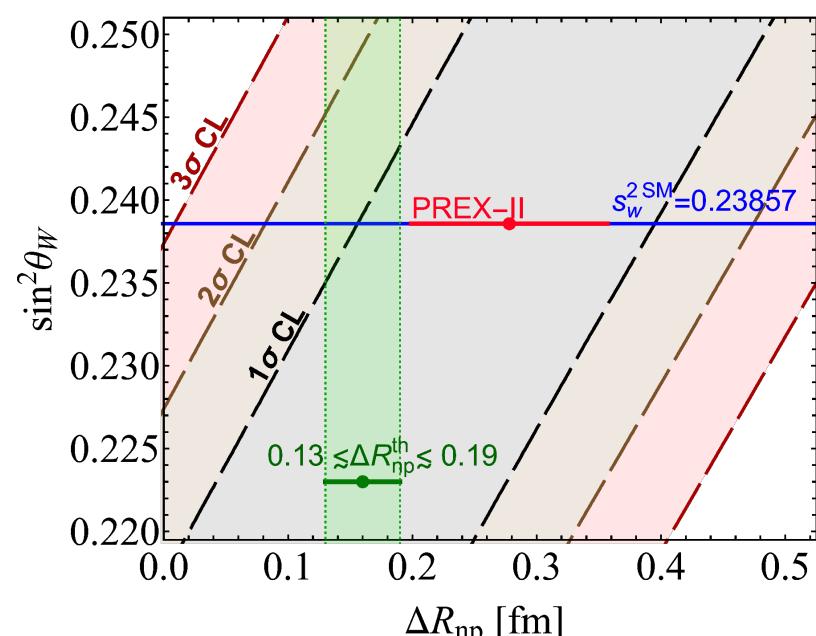
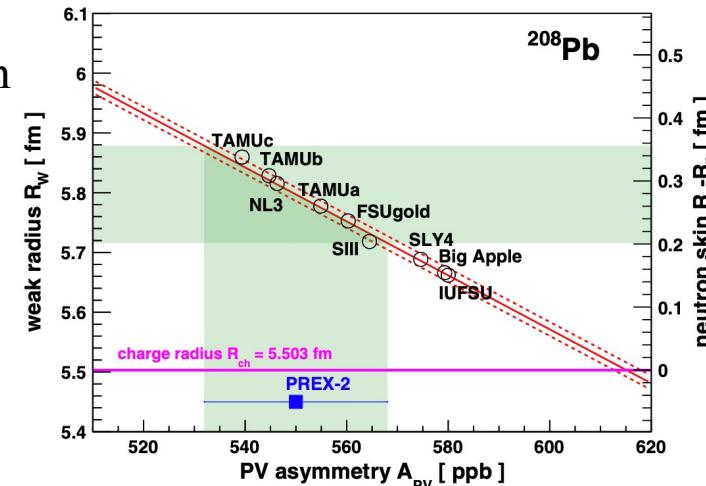
$$\Delta R_{np}^{\text{th}}(\text{Pb}) \approx 0.13 - 0.19 \text{ fm}$$

$$A_{\text{pv}} = -\frac{G_F Q^2}{4\sqrt{2}\pi\alpha} \frac{Q_W}{Z} \frac{F_{\text{wk}}(q^2)}{F_{\text{ch}}(q^2)}$$

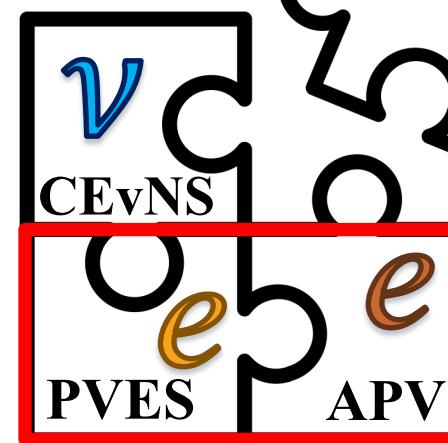
**What is the impact of the weak mixing angle?**

The PREX measurement is degenerate in the plane  $\Delta R_{np}(\text{Pb})$  vs  $\sin^2 \theta_W$

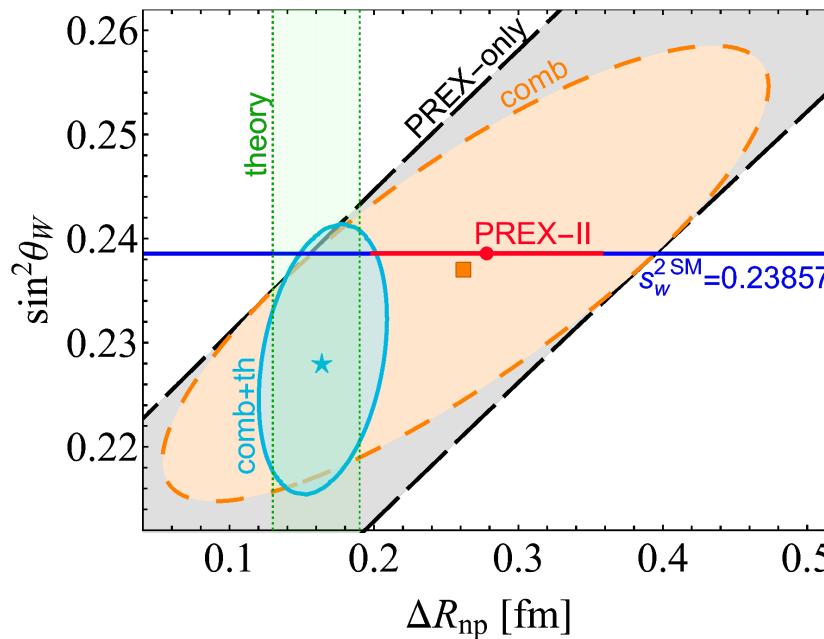
Lower values of the weak mixing angle result in larger values of the nuclear weak charge, and thus the fit gives lower neutron skin values



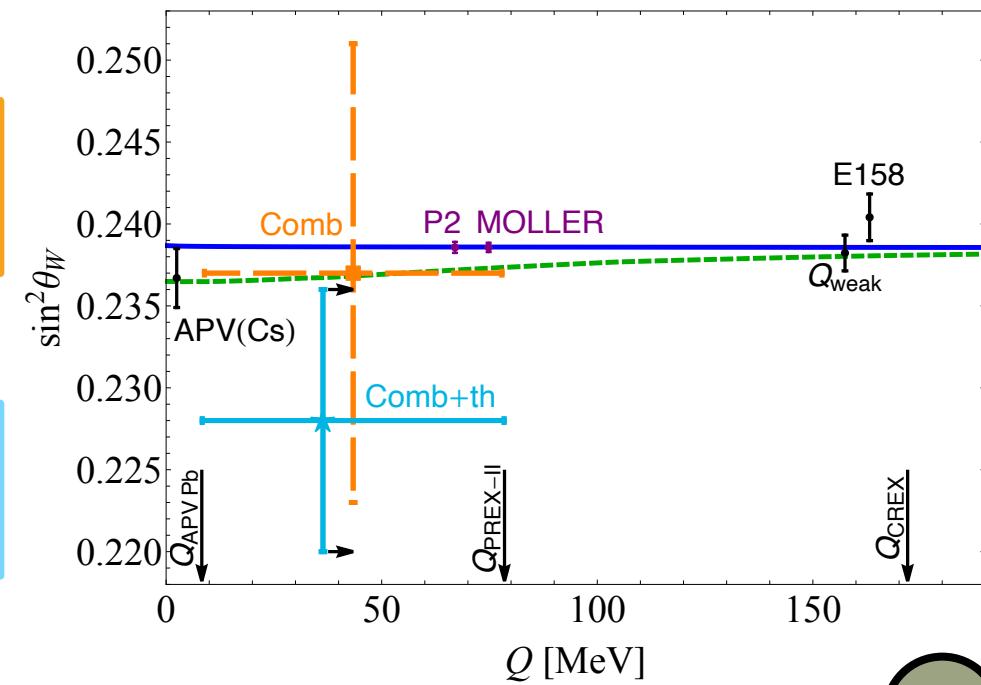
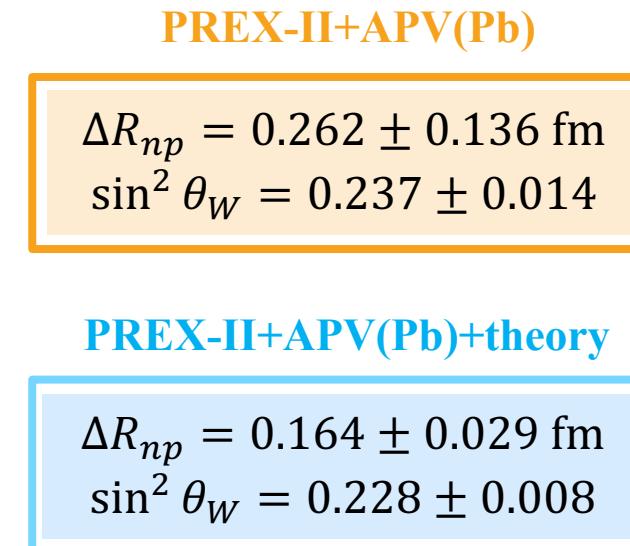
NC et al., Phys.Rev.C 105 (2022) 5, 055503

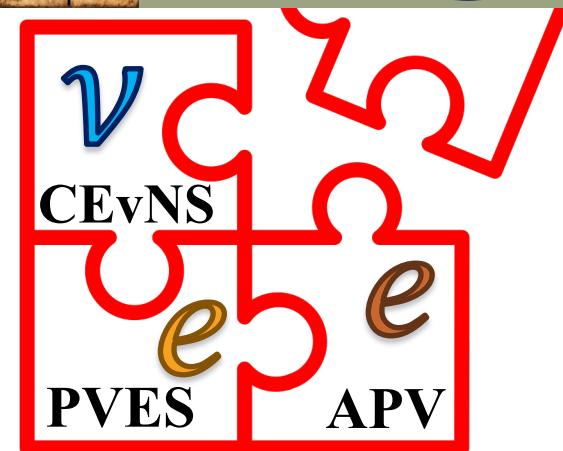


- PREX alone is degenerate in the plane  $\Delta R_{np}(\text{Pb})$  vs  $\sin^2 \theta_W$
- We exploit the very different dependence on the two free parameters for the different probes: PVES and APV(Pb)
- PVES is sensitive to the neutron skin more than to the weak mixing angle
- APV is sensitive to the weak mixing angle more than to the neutron skin
- It is possible to extract the two with a data-driven approach, but with poorer precision



NC et al., Phys.Rev.C 105 (2022) 5, 055503





## Toward a Global Analysis:

By exploiting a trend in the predictions from mean field models of  $R_n(\text{Pb})$  and  $R_n(\text{Cs})$ , we can translate a measurement on lead into a measurement on cesium

Relying on the input  
from theoretical models

Meas.	$R_n(\text{Cs}) [\text{fm}]$	$\sin^2 \theta_W$
APV(Cs) 21 + COH + CSRe	$5.04^{+0.19}_{-0.19}$	$0.2396^{+0.0020}_{-0.0019}$
APV(Cs) 21 + COH + PREX + APV Pb	$5.041^{+0.058}_{-0.057}$	$0.2396^{+0.0017}_{-0.0017}$

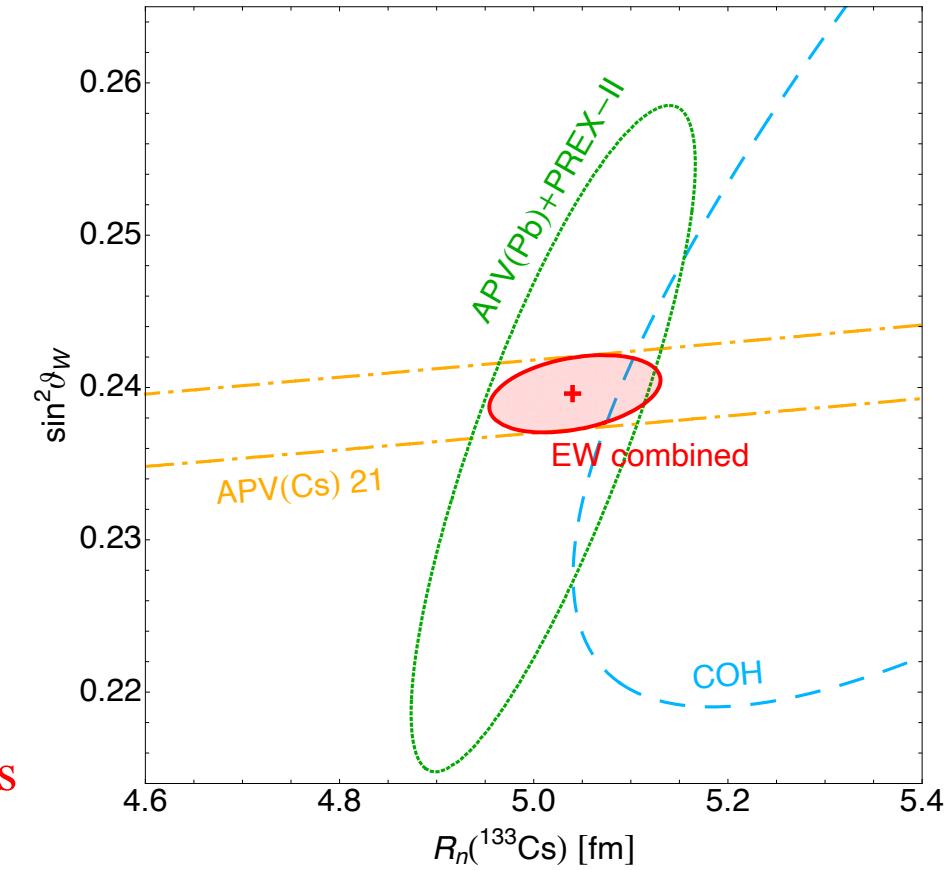
NC et al., Phys. Rev. D 110 (2024) 3, 033005

The precision improves slightly by combining the available probes on Cs with the ones on Pb

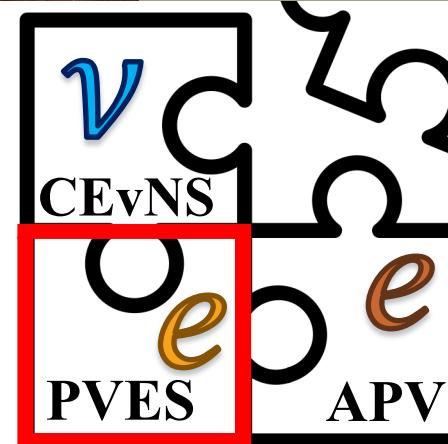
- ~1.2% on  $R_n(\text{Cs})$
- ~7% on  $\sin^2 \theta_W$



The EW combined analysis gives practically the same results of the Cs only fit



NC et al., Phys. Rev. D 110 (2024) 3, 033005



**CREX+Qweak**

$$\Delta R_{np} = 0.119 \pm 0.028 \text{ fm}$$

$$\sin^2 \theta_W = 0.2386 \pm 0.0013$$

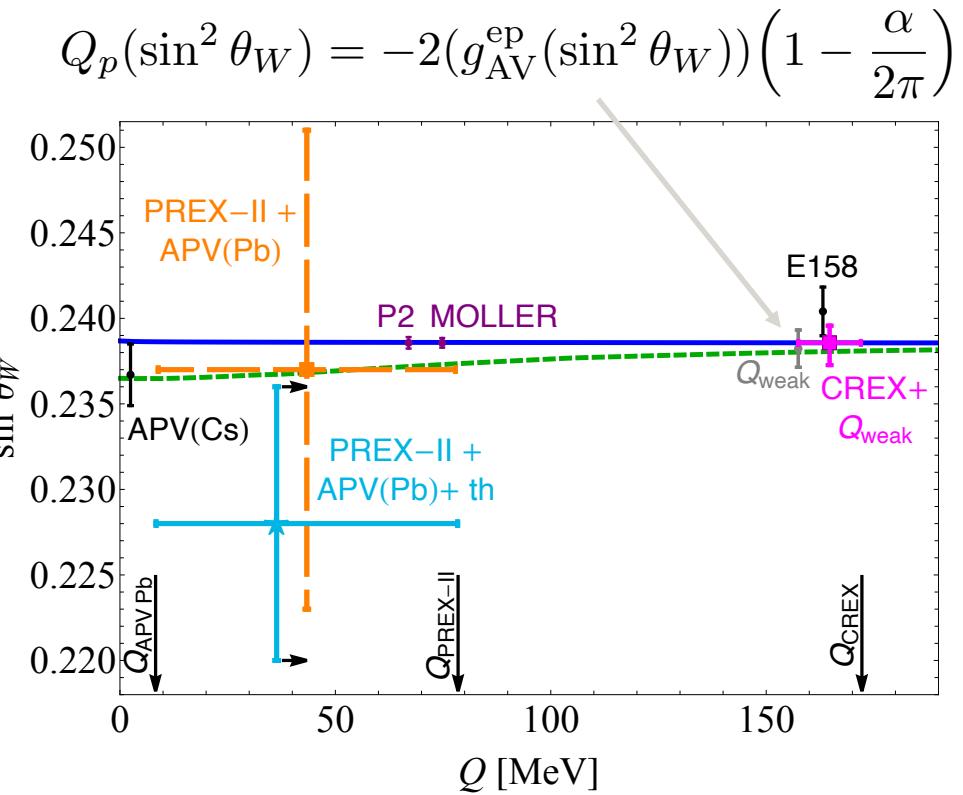
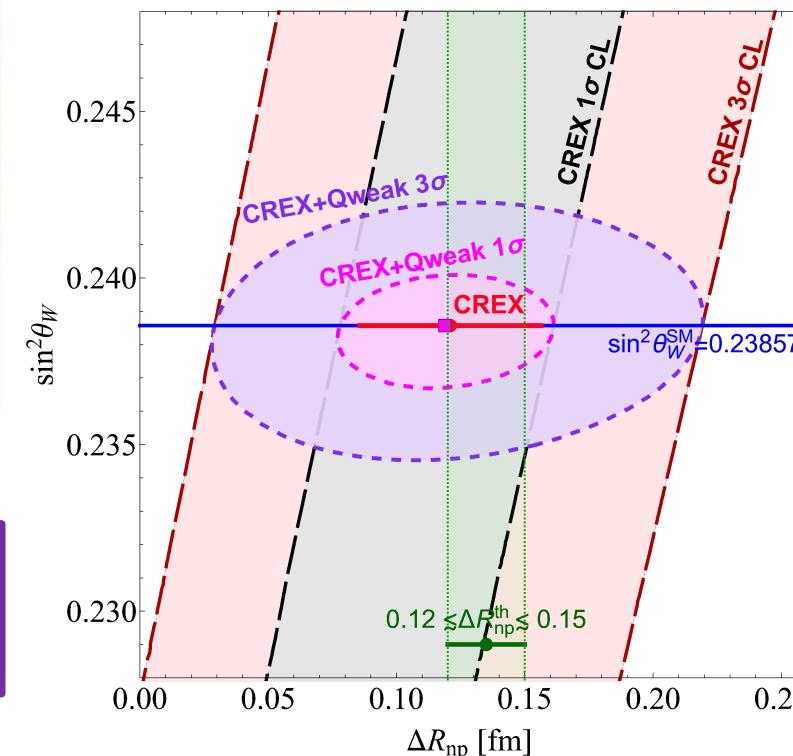
CREX: PVES on calcium-48



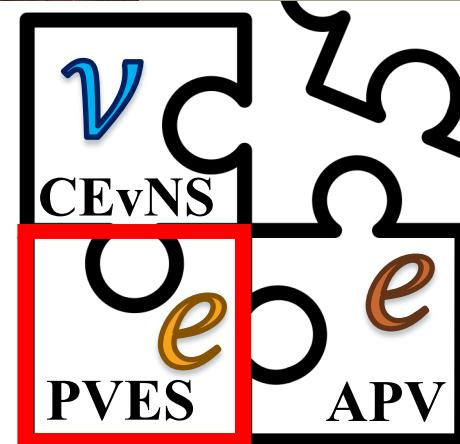
D. Adhikari et al., Phys. Rev. Lett. 129.042501(2022)

2.1 GeV polarized electrons on a calcium-48 target scattered at  $\theta \sim 5^\circ$  (higher momentum transfer  $q \sim 170$  MeV)

Only one measurement released in 2021

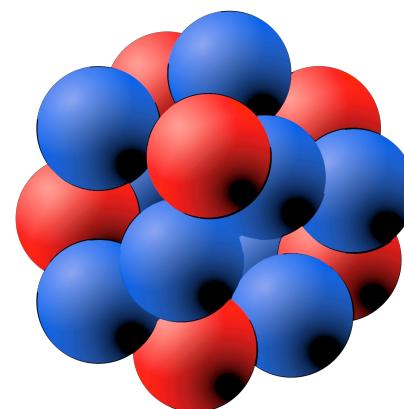


- Qweak sets a horizontal band which breaks the degeneracy of CREX (preserving the precision)
- CREX and PREX point toward different directions!



- PVES on heavy nuclei is used to extract information on the neutron skin (i.e. PREX and CREX experiments)
- PVES on light objects, such as the proton or the electrons has been used to measure the weak mixing angle (i.e. Qweak and E158 experiments)

### What happens by performing a PVES measurement on a light nucleus?

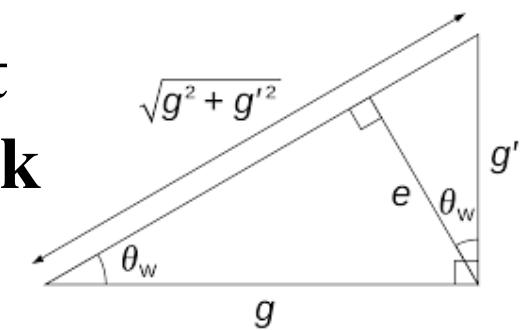


Nuclear Structure  
test through the  
neutron skin

OR/AND



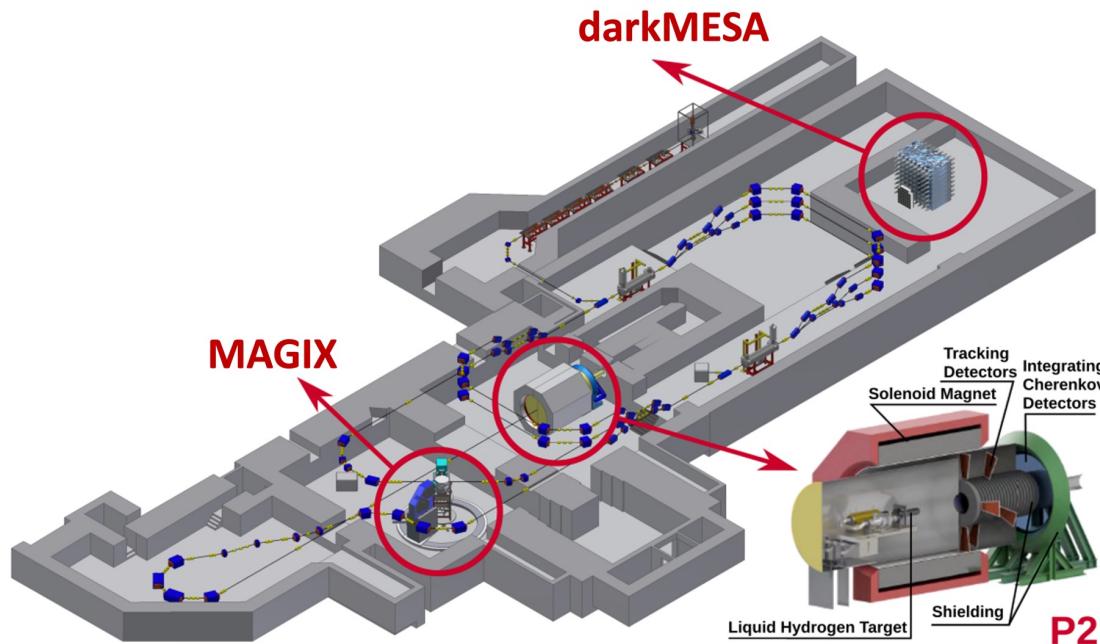
Electroweak test  
through the weak  
mixing angle



Using light nuclei, there is no need to reach high electron energies (hundreds of MeV instead of GeV electrons): MESA Facility represent a unique opportunity



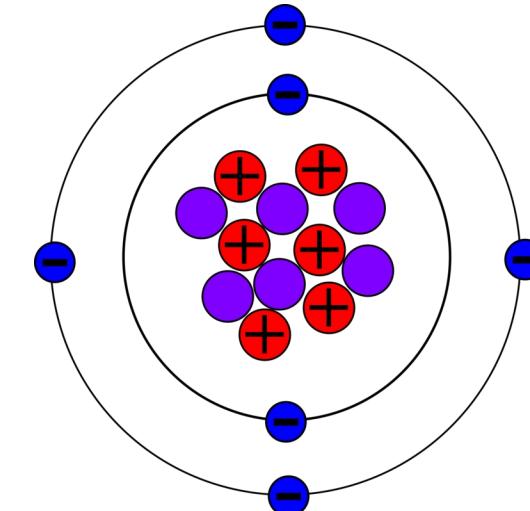
Work in Collaboration with M. Cadeddu, J. Erler, M. Gorchtein, J. Piekarewicz, X. Roca-Maza and H. Spiesberger



D. Becker et al., Eur.Phys.J. A 54, 208 (2018)

- Particular system: much **lighter** than lead and calcium
- The neutron skin is expected to be close to zero and negative
- Coulomb repulsion as dominant effect (**proton skin**)

**Z=6      N=6**



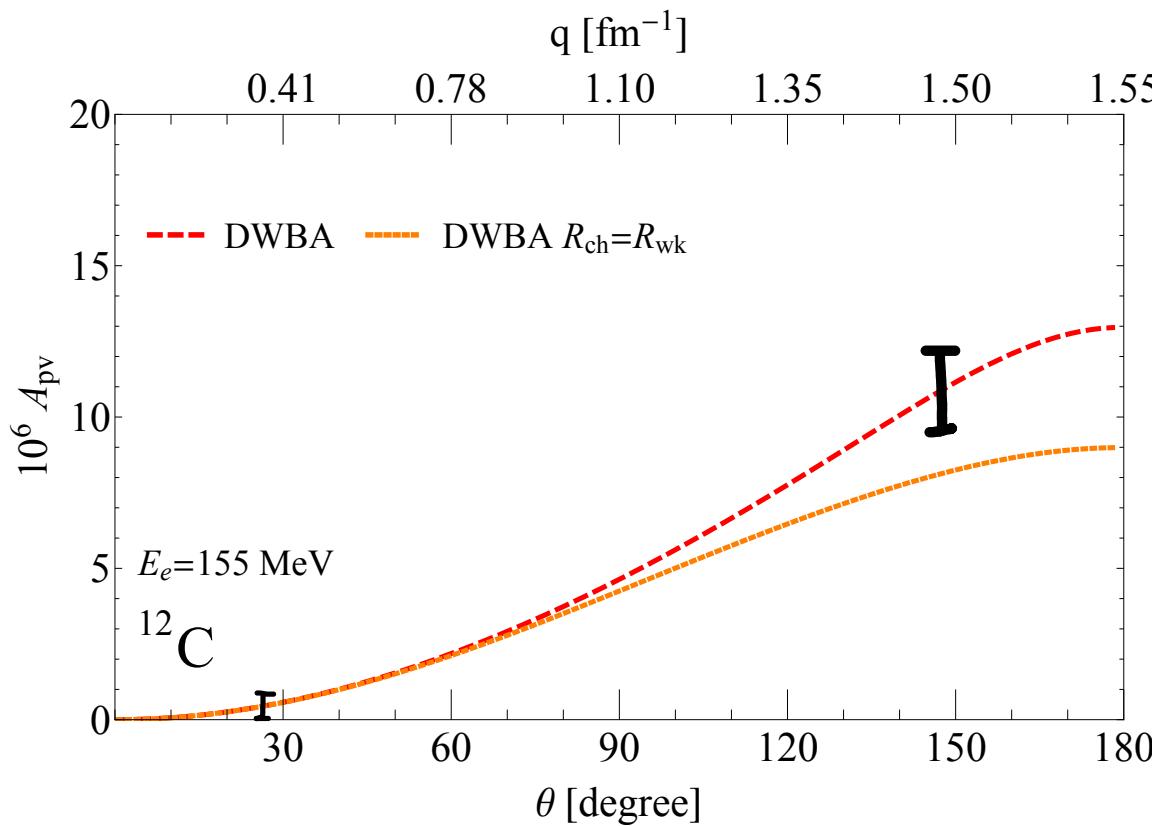


A preliminary study was discussed in a previous publication



O. Koshchii et al., Phys. Rev. C 102, 022501 (2020)

Feasibility study for the **simultaneous** extraction of the weak mixing angle and the weak skin of carbon-12



$$\text{WEAK SKIN } \Delta R_{wk,skin} = R_{wk} - R_{ch}$$

At MESA, the form factor will be measured at two **different energies** within the same experiment

- Forward is more sensitive to the weak mixing angle
- Backward is more sensitive to the weak skin

PHYSICAL REVIEW C 110, 035501 (2024)

NC et al., Phys. Rev. C 110 (2024) 3, 035501

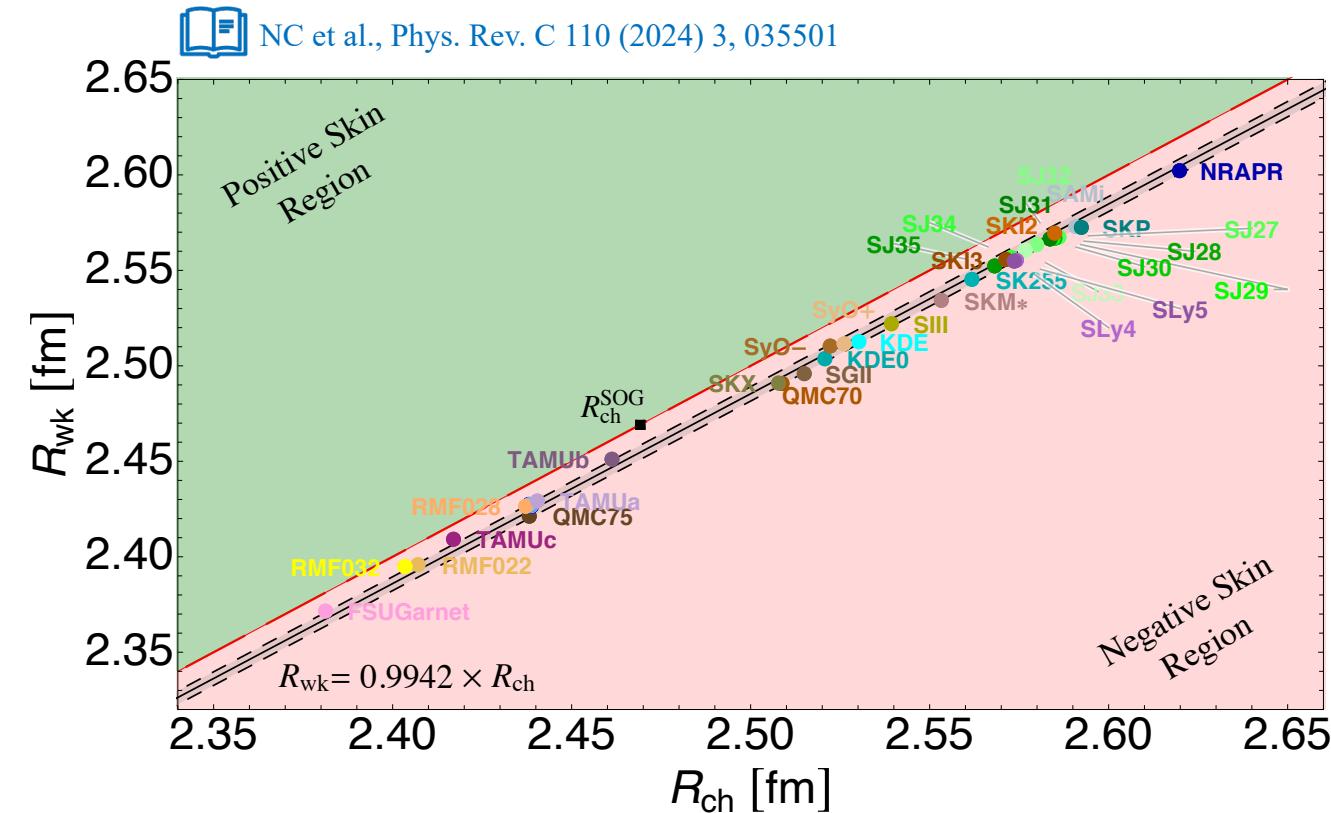
Simultaneous extraction of the weak radius and the weak mixing angle from parity-violating electron scattering on  $^{12}\text{C}$

M. Cadeddu <sup>1,\*</sup>, N. Cargioli <sup>2,1,†</sup>, J. Erler <sup>3,‡</sup>, M. Gorchtein <sup>3,§</sup>, J. Piekarewicz <sup>4,||</sup>, Xavier Roca-Maza <sup>5,6,7,8,¶</sup>, and H. Spiesberger <sup>3,#</sup>



## What does the theory say about C-12?

- The charge radius of C-12 is known to a certain extent
- $R_{ch}^{SOG} = 2.469(5) \text{ fm}$
- H. De Vries et al., At. Data Nucl. Data Tables 36, 495 (1987)
- Mean Field Models:
  - Predict both the weak and charge radius in a consistent way
  - We should not use only part of the information (weak radius) but account also for correlations
  - Predictions for charge and weak radii show a common trend



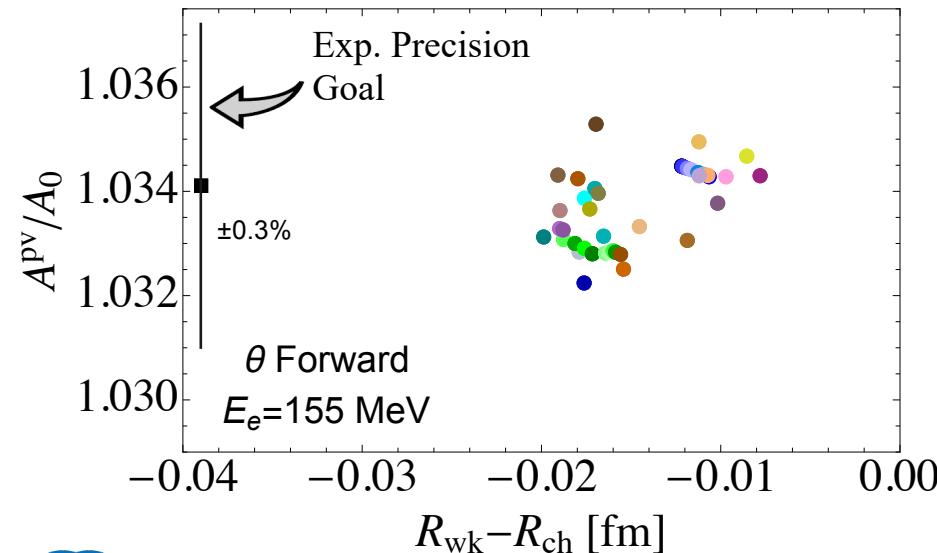
A rather large spread in  $R_{ch}$  (and  $R_{wk}$ ) is observed, while the skin remains rather stable!



## However, the experimental observable is the asymmetry

At forward angle:

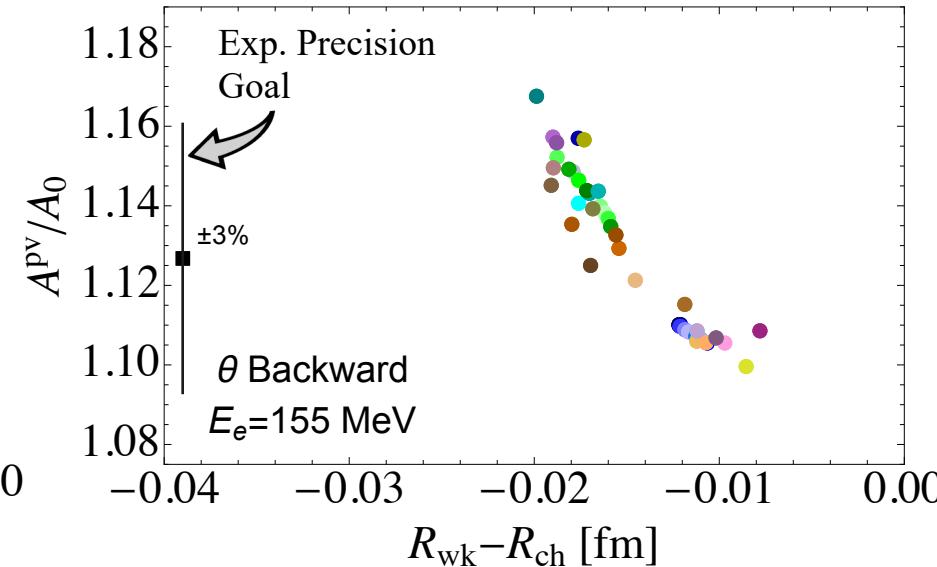
- The predictions are “randomly” scattered
- The overall vertical spread is small



NC et al., Phys. Rev. C 110 (2024) 3, 035501

At Backward angle:

- The predictions are scattered showing a general trend as a function of the weak skin
- The overall vertical spread is non-negligible

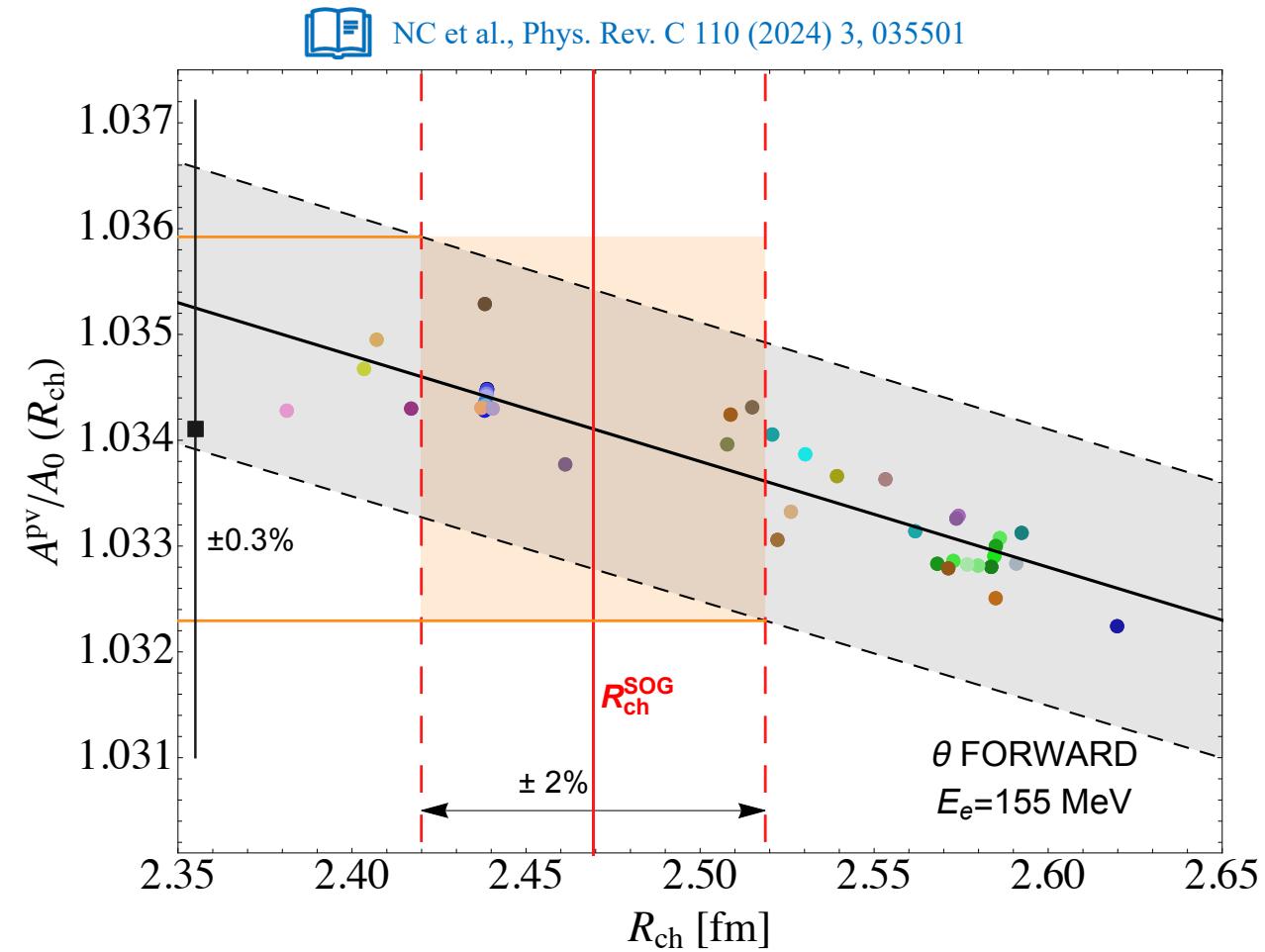


● KDE	● KDE0	● NRAPR	● QMC70	● QMC75	● SAMi	● SGII	● SIII	● SJ27	● SJ28	● SJ29
● SJ30	● SJ31	● SJ32	● SJ33	● SJ34	● SJ35	● SK255	● SKI2	● SKI3	● SKM*	● SKP
● SKX	● SLy4	● SLy5	● SyO-	● SyO+	● FSUGarnet	● FSUGold2	● FSUGold2 L047	● FSUGold2 L050	● FSUGold2 L054	● FSUGold2 L058
● FSUGold2 L069	● FSUGold2 L076	● FSUGold2 L090	● FSUGold2 L100	● RMF022	● RMF028	● RMF032	● TAMUa	● TAMUb	● TAMUc	



# The Forward measurement is practically independent of the nuclear structure contribution, but how small is the contribution?

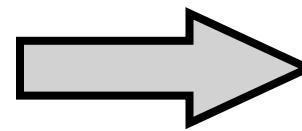
- The prediction for the asymmetry at forward angle shows a trend as a function of the charge radius
- We can assume a simple linear function and fit the predictions to obtain a correlation function
- The charge radius is measured
- We can define a confidence region around the measured charge radius to project it toward a range of values for the asymmetries
- The estimated uncertainty due to nuclear structure affecting the forward measurement is  $\sim 0.2\%$  (below the precision goal)



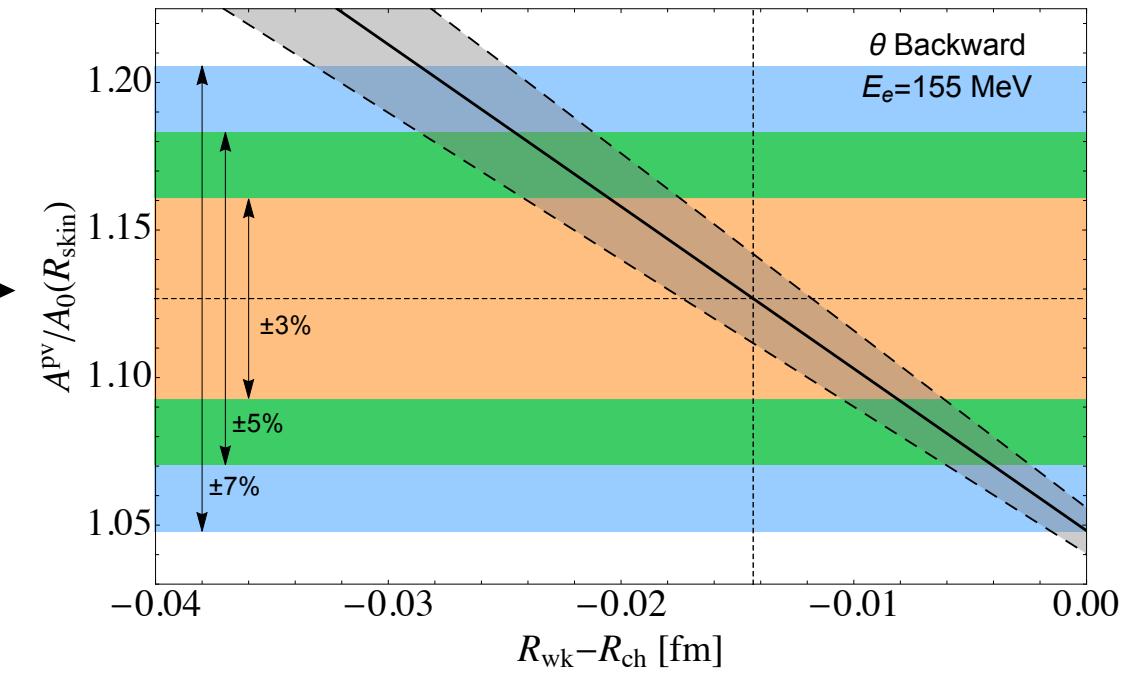
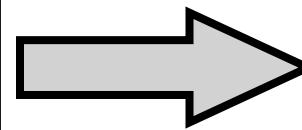
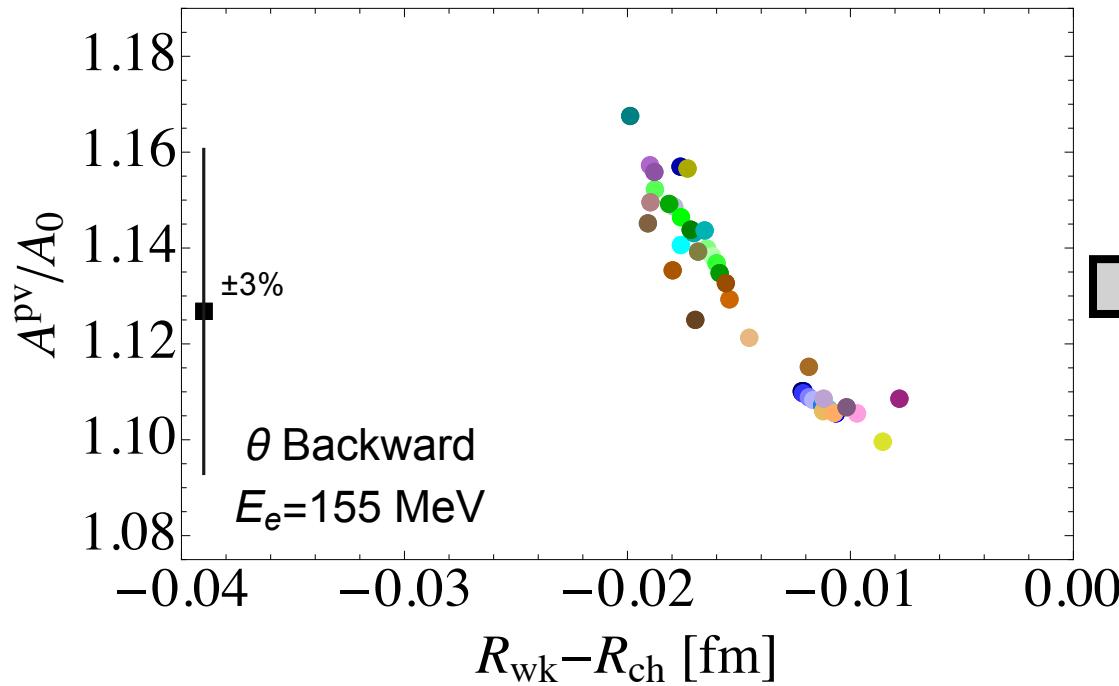


## The Backward measurement depends on the nuclear structure contribution

- We fit the prediction using a linear function:
  - $\mathcal{A}_b^{pv}(R_{wskin}) = \mathcal{A}R_{wskin} + \mathcal{B}$



We can use this function to estimate the prediction for the backward asymmetry as a function of the weak skin





To explore the sensitivity we built the following chi-square function

Experimental precision goal  
for the forward measurement

Nuisance parameter introduced to  
account for the uncertainty due to nuclear  
structure affecting the forward prediction

$$\chi^2 = \left( \frac{A_f^{\text{PV,ref}} - \eta_1 A_f^{\text{PV}}(\sin^2 \theta_W)}{\epsilon_f} \right)^2 + \left( \frac{A_b^{\text{PV,ref}} - A_b^{\text{PV}}(\eta_2, \sin^2 \theta_W, R_{\text{wk}})}{\sigma_b} \right)^2 + \left( \frac{\eta_1 - 1}{\sigma_{\eta_1}} \right)^2 + \eta_2^2,$$

Forward asymmetry prediction, the  
dependence on  $\sin^2 \theta_W$  is explicitly  
inserted by rescaling the weak charge

Backward asymmetry prediction  
which depends on both the weak  
mixing angle and the weak skin

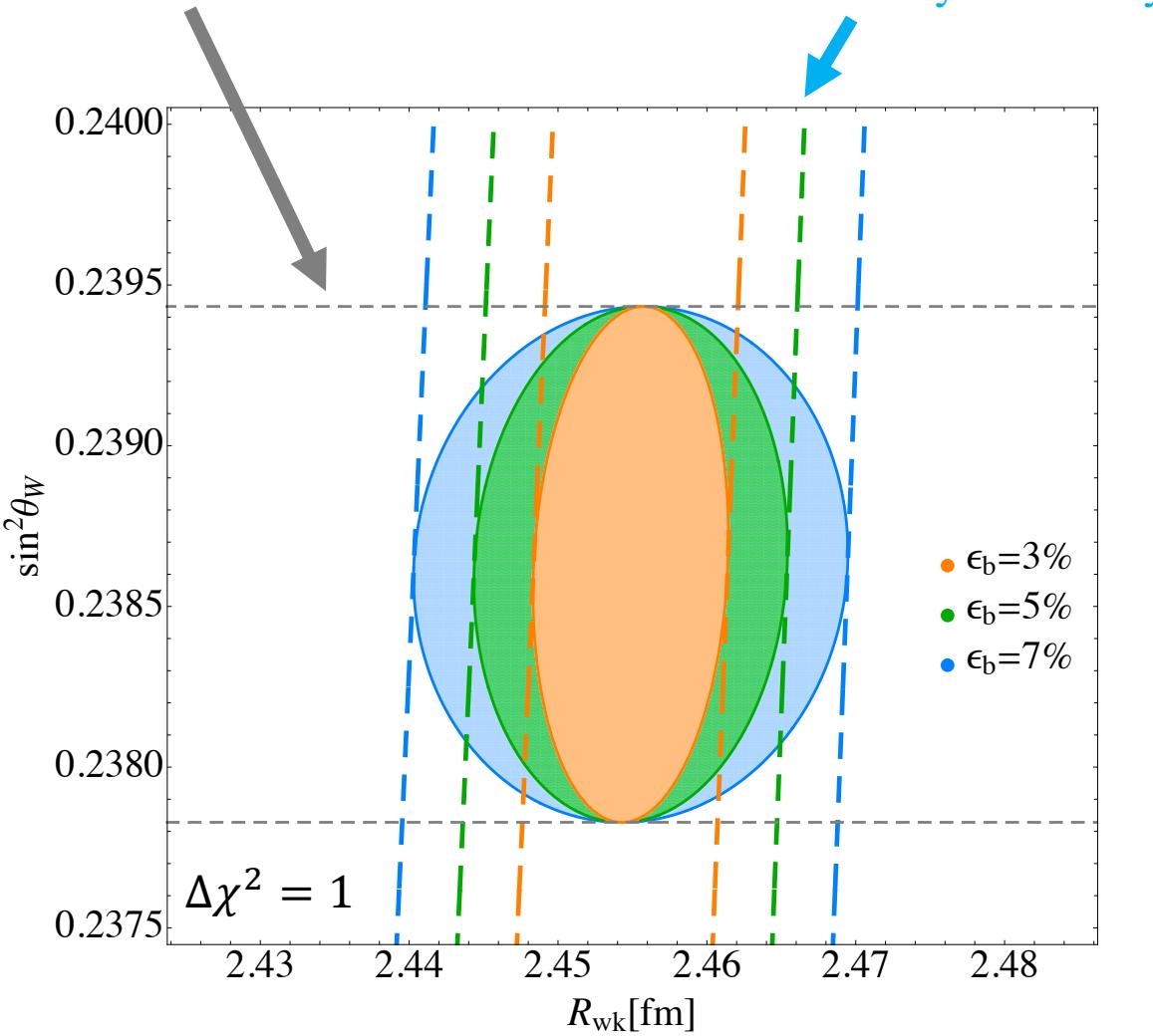
$$A_b^{\text{PV}}(\eta_2, \sin^2 \theta_W, R_{\text{wk}}) = \frac{Q_W(\sin^2 \theta_W)}{Q_W^{\text{SM}}} \times [(\mathcal{A} + \eta_2 \sigma_{\mathcal{A}}) R_{\text{wskin}} + \mathcal{B}]$$

Accounts for the experimental precision for the backward  
measurement and the uncertainty on the intercept  $\mathcal{B}$

Nuisance parameter relative to the slope of the linear  
function  $\mathcal{A}$  describing the backward prediction



Forward only sensitivity



- Forward: Horizontal band, establishes the sensitivity to the **weak mixing angle**
- Backward: Almost vertical band, establishes the sensitivity to the **weak radius** but it is marginally dependent on the weak mixing angle
- We consider different backward precision scenarios

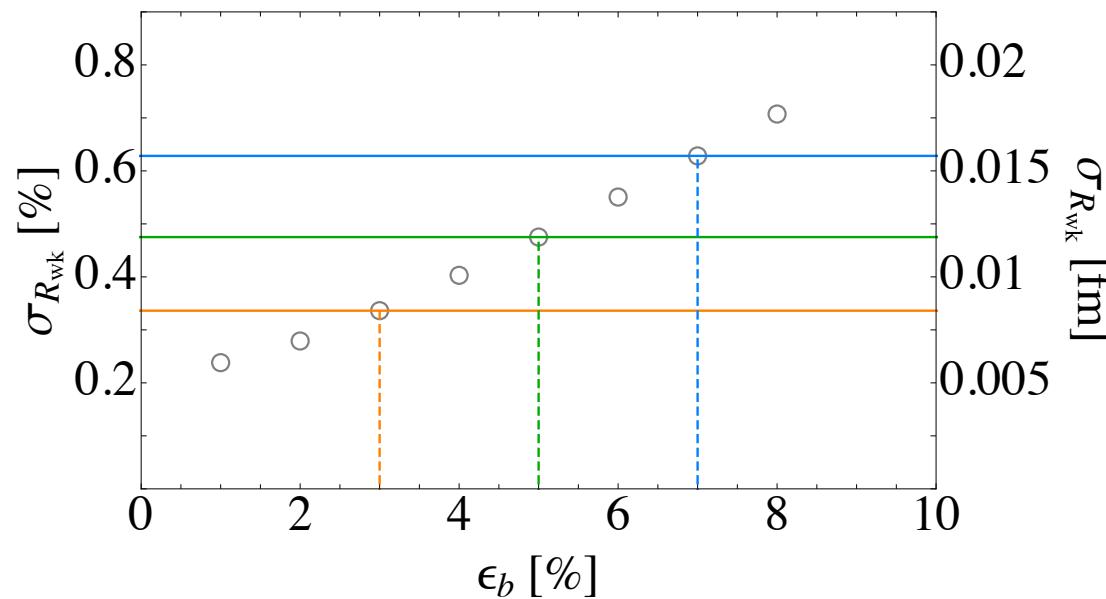
$\epsilon_f$	$\epsilon_b$	$R_{\text{wk}}$ [fm]	$\sigma_{R_{\text{wk}}}$ [%]	$\sin^2 \theta_W$
0.3%	3%	$2.455 \pm 0.008$	0.34%	$0.2386 \pm 0.0008$
0.3%	5%	$2.455 \pm 0.012$	0.48%	$0.2386 \pm 0.0008$
0.3%	7%	$2.455 \pm 0.016$	0.63%	$0.2386 \pm 0.0008$





$\epsilon_f$	$\epsilon_b$	$R_{\text{wk}}$ [fm]	$\sigma_{R_{\text{wk}}}$ [%]	$\sin^2 \theta_W$
0.3%	3%	$2.455 \pm 0.008$	0.34%	$0.2386 \pm 0.0008$
0.3%	5%	$2.455 \pm 0.012$	0.48%	$0.2386 \pm 0.0008$
0.3%	7%	$2.455 \pm 0.016$	0.63%	$0.2386 \pm 0.0008$

It would represent the first **electroweak determination** of the weak radius and potentially first «proton skin» determination

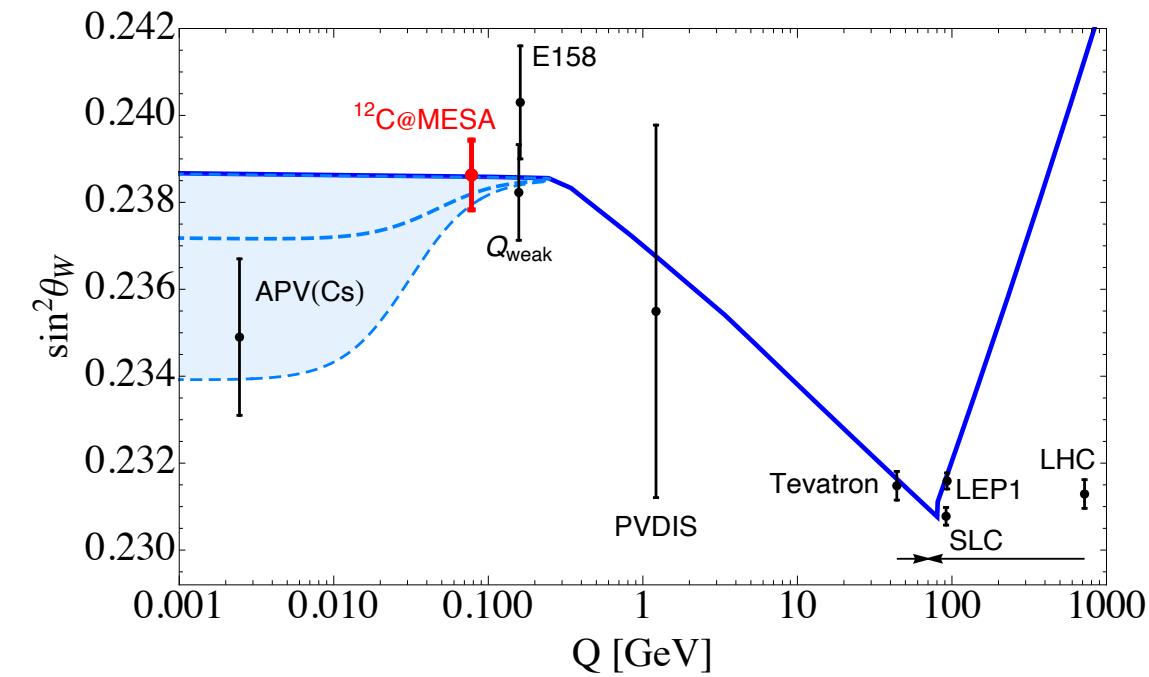


NC et al., Phys. Rev. C 110 (2024) 3, 035501



- NC et al., Phys. Rev. D 104, L011701 (2021)
- H. Davoudiasl et al., Phys. Rev. D 85, 115019 (2012)
- H. Davoudiasl et al., Phys. Rev. D 86, 095009 (2012)
- H. Davoudiasl et al., Phys. Rev. D 88, 015022 (2013)

Unique sensitivity to the weak mixing angle  $\sim 0.3\%$  at a crucial momentum transfer: where light new physics could modify the weak mixing angle running





There is intent to extend the study performed on C-12 to different nuclei, in particular to the stable daughter nuclei participating in the superallowed nuclear  $\beta$  decay.



B. Ohayon, arXiv:2409.08193v2

C.Y. Seng and M. Gorchtein, Phys. Rev. C 109 (2024) 4, 045501

C.Y. Seng, Phys. Rev. Lett. 130 (2023) 15, 152501

C.Y. Seng and M. Gorchtein, Phys. Lett. B 838 (2023) 137654

Precise measurements of neutron skins

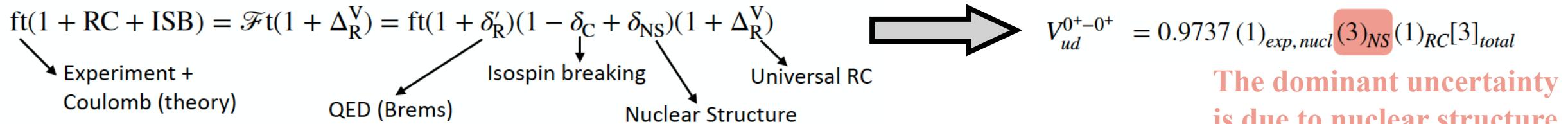


Impact on the extraction of the CKM quark mixing matrix element  $V_{ud}$

$$V_{CKM} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix}$$

There is connection between isospin-symmetry breaking corrections to the rates of superallowed nuclear  $\beta$  decay,  $\delta_C$  and isospin-breaking sensitive combinations of electroweak nuclear radii (experimentally accessible!)

$$V_{ud} \text{ can be extracted from superallowed decays: } V_{ud}^2 = \frac{2984.43s}{\mathcal{F}t(1 + \Delta_R^V)}$$





If isospin symmetry were exact

$$M_F \rightarrow M_0 = \sqrt{2}$$

Isospin symmetry is broken in nuclear states, so that in presence of ISB

$$M_F^2 = M_0^2(1 - \delta_C)$$
MacDonald 1958

$\delta_C \sim 0.17\% - 1.6\%$  Crucial for  $V_{ud}$  extraction

Nuclear community embarked on ab-initio  $\delta_C$  calculations

$\delta_C$  can be extracted from nuclear radii: we can build ISB-sensitive combinations of radii

$$\Delta M_B^{(1)} \equiv \frac{1}{2} \left( Z_1 R_{p,1}^2 + Z_{-1} R_{p,-1}^2 \right) - Z_0 R_{p,0}^2$$

$\Delta M_B^{(1)} = 0$  used for ft-value in isospin limit

$$\Delta M_A^{(1)} \equiv - \langle r_{CW}^2 \rangle + \left( \frac{N_1}{2} \langle r_{n,1}^2 \rangle - \frac{Z_1}{2} \langle r_{p,1}^2 \rangle \right)$$

Neutron radius: measurable with PV e<sup>-</sup> scattering!



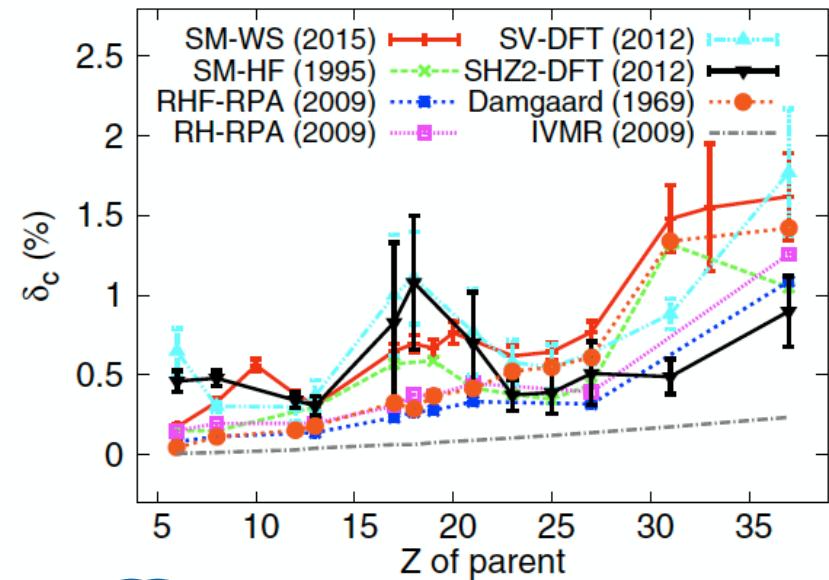
B. Ohayon, arXiv:2409.08193v2

C.Y. Seng and M. Gorchtein, Phys. Rev. C 109 (2024) 4, 045501

C.Y. Seng, Phys. Rev. Lett. 130 (2023) 15, 152501

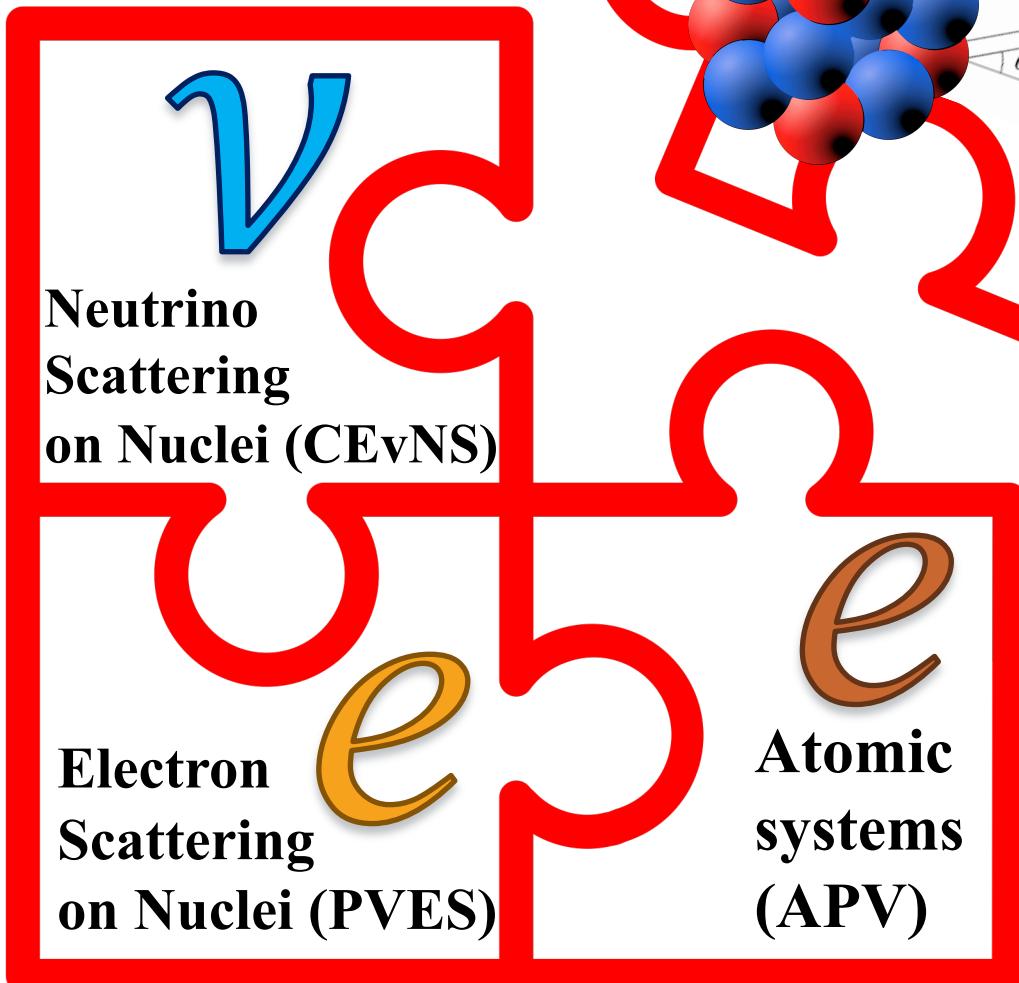
C.Y. Seng and M. Gorchtein, Phys. Lett. B 838 (2023) 137654

$\delta_C$  from shell models, density functionals and random phase approximation theories



L. Xayavong et al., Phys. Rev. C 97 (2018), 024324

The upcoming exp. Facility at Mainz (MESA) can be used to measure the neutron skins of stable daughters (e.g. Mg-26, Ca-42, Fe-54)



- Electroweak probes represent a **powerful tool** to test the standard model theory and explore the world of nuclear matter
- The **combination** of different probes enhances the sensitivity and allows one to extract **simultaneously** the **weak mixing angle** and the **neutron skin** from available data
- A **light system** such as C-12 can provide interesting information on the weak mixing angle (**world leading precision**) and on nuclear structure in light systems (and **symmetrical** matter)
- Take away message: **combine different probes and perform measurements at more energies to gain more sensitivity**

# THANK YOU FOR YOUR ATTENTION

## QUESTIONS TIME

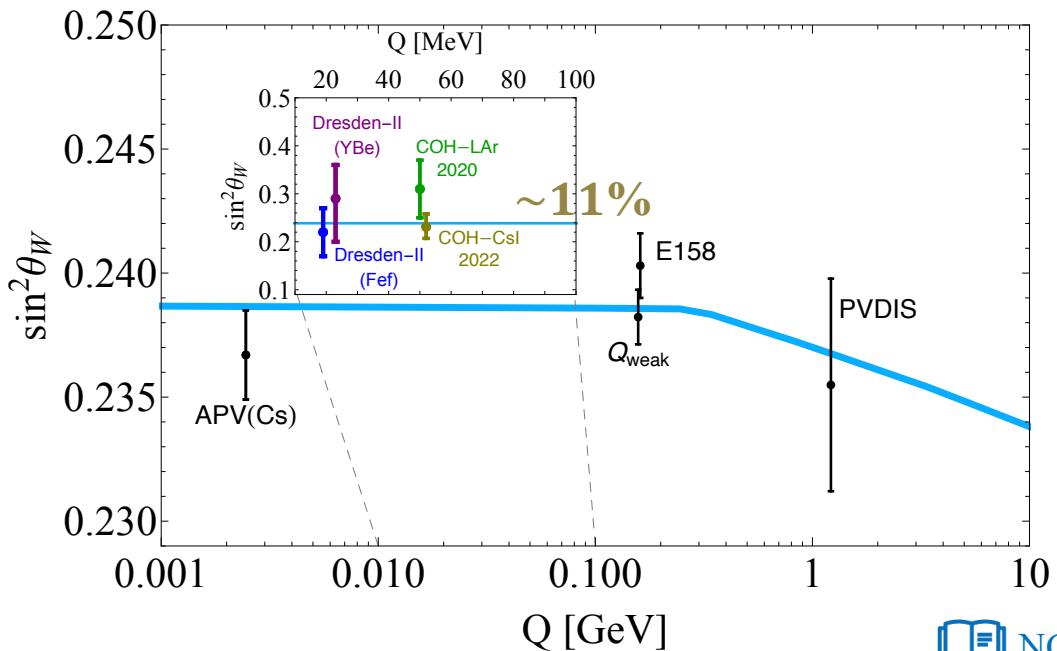
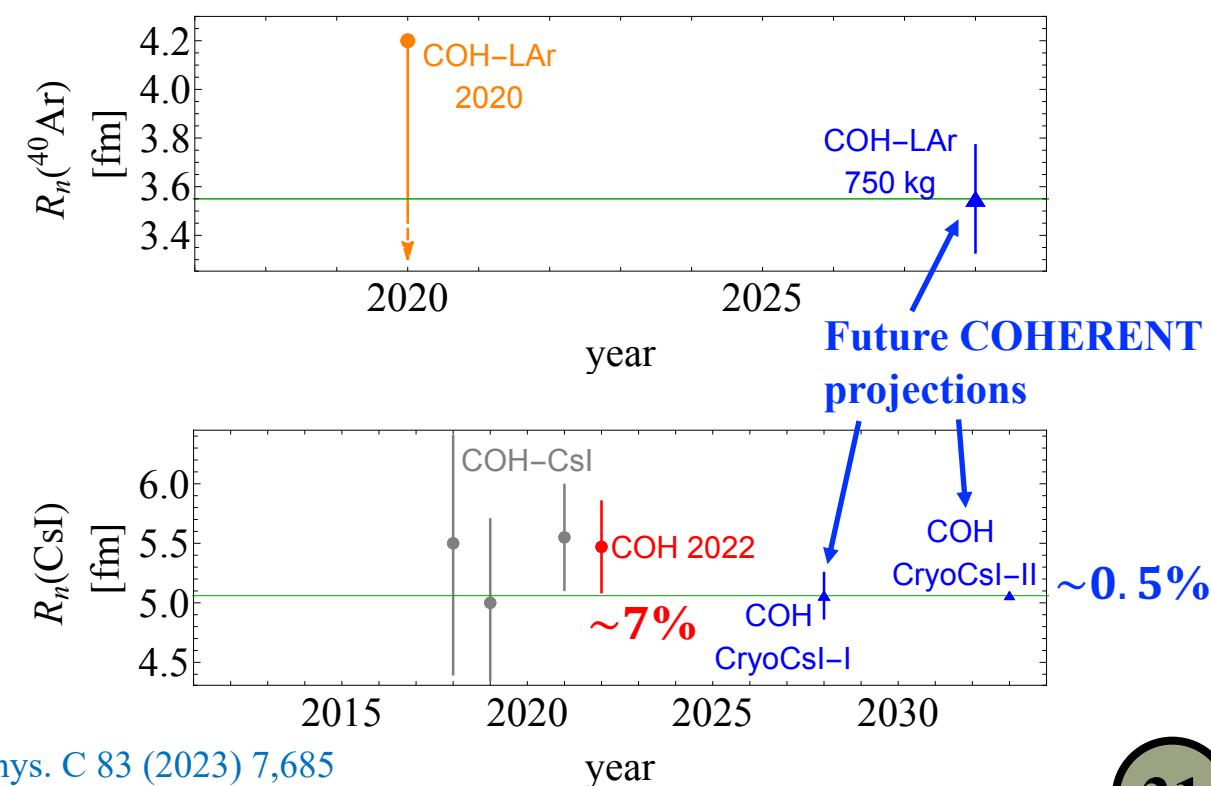
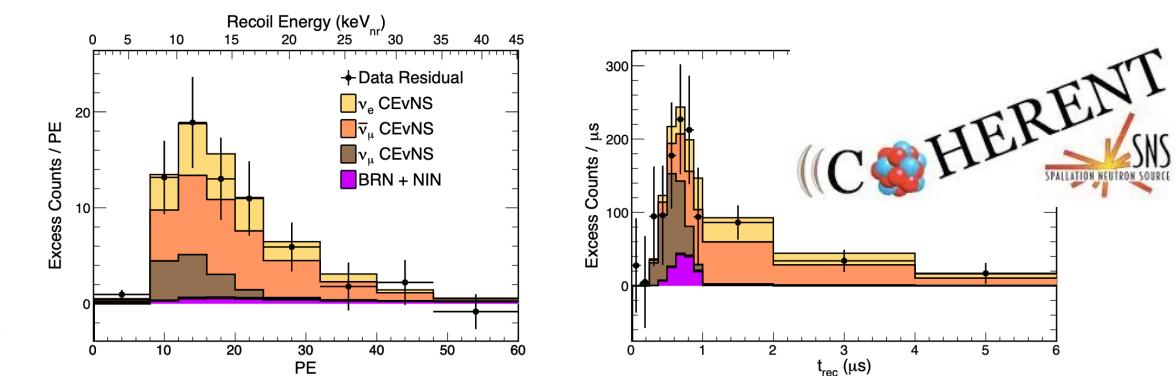


# BACK UP SLIDES



## COHERENT experimental program

- CsI crystal: D. Akimov et al. *Phys.Rev.Lett.* 129 (2022) 8,081801
  - 14.6 kg scintillating crystal
  - 19.3 m away from the SNS target
  - $q \sim 50$  MeV
- Ar single phase: D. Akimov et al. *Phys.Rev.Lett.* 126 (2021) 1, 012002
  - 24 kg of atmospheric argon
  - 27.5 m away from the SNS target
  - $q \sim 50$  MeV

NC et al., *Eur.Phys. C* 83 (2023) 7,685



APV experiments measure the weak mixing angle at very low momentum transfer (few MeVs) via electronic transitions which are forbidden by selection rules: they violate parity

We consider the APV measurement on **cesium-133**

Interaction Hamiltonian

$$\hat{h}_{SI} = \frac{G_F}{2\sqrt{2}} Q_W \rho_{wk}(r) \gamma_5$$

involved atomic states:  
 $6S \rightarrow 7S$

Parity-violating amplitude

$E_{PNC}$

$E_{PNC}$ : includes the electric-dipole operator

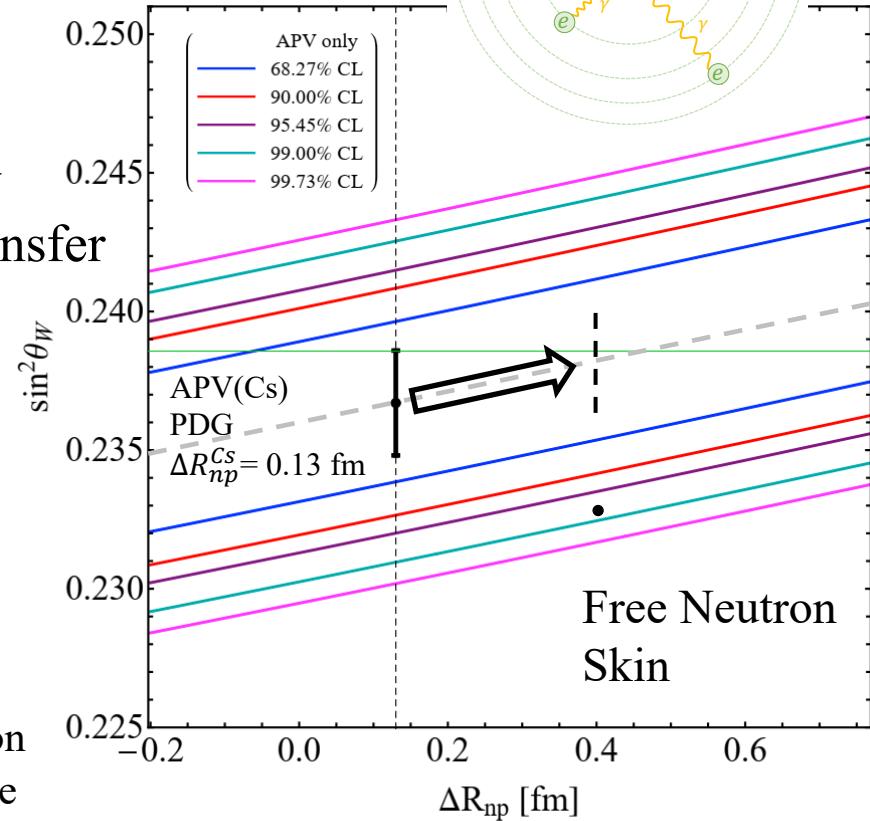
Nuclear weak charge

$$Q_W^{APV}(R_n) = N \left( \frac{\text{Im } E_{PNC}(R_n)}{Q_{weak}} \right)_{exp} \left( \frac{Q_{weak}}{N \text{Im } E_{PNC}(R_n)} \right)_{th} \beta_{exp+th}$$

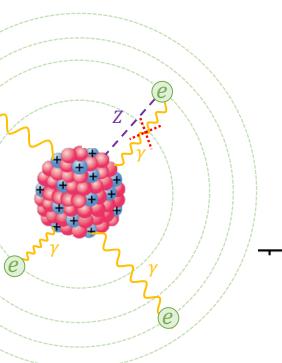
Extremely low momentum transfer  
 $Q \sim 2.4 \text{ MeV}$

$\beta$ : tensor transition polarizability: size of the Stark mixing induced electric dipole amplitude

There are different atomic calculations available, we adopt  
B.K. Sahoo et al. Phys Rev D 103, L111303 (2021)



Very sensitive to  $\sin^2 \theta_W$ , not much on  $R_n(Cs)$





APV experiments measure the weak mixing angle at very low momentum transfer (few MeVs) via electronic transitions which are forbidden by selection rules: they violate parity

We consider the APV measurement on **lead-208**

Interaction Hamiltonian

$$\hat{h}_{SI} = \frac{G_F}{2\sqrt{2}} Q_W \rho_{wk}(r) \gamma_5$$

involved atomic states:  
 $6p^2 \ ^3P_0 \rightarrow 6p^2 \ ^3P_1$

Parity-violating amplitude

$E_{PNC}$

inside the theoretical ratio  
 $R_{th} = \left( \frac{\text{Im } E_{PNC}}{M_1} \right)_{th}$

Nuclear weak charge

$$Q_W^{APV}(R_n) = -N R_{exp} \left( \frac{M_1}{\text{Im } E_{PNC}(R_n)} \right)_{th}$$

$$R_{exp} = -9.80(33) \times 10^{-8}$$
$$R_{exp} = -9.86(12) \times 10^{-8}$$

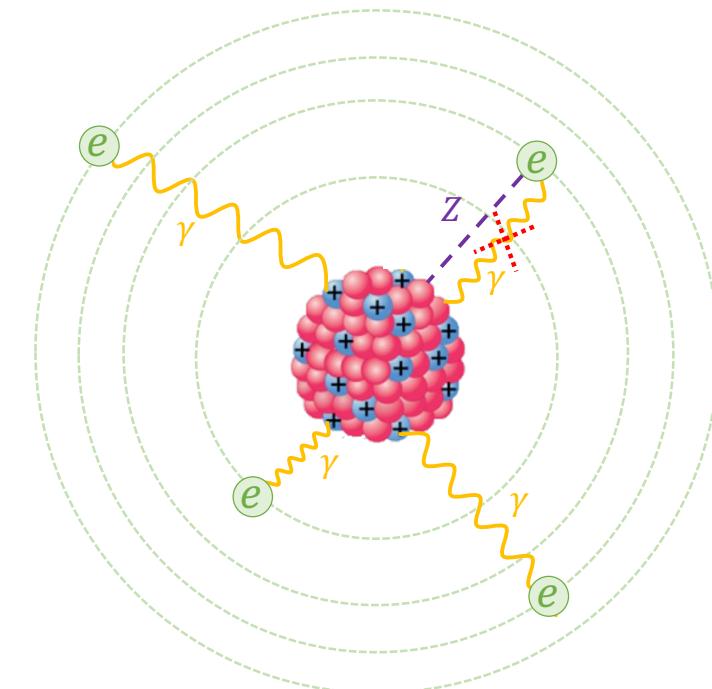
S.J. Phipp et al. Journal of Physics B 29, 1861 (1996)

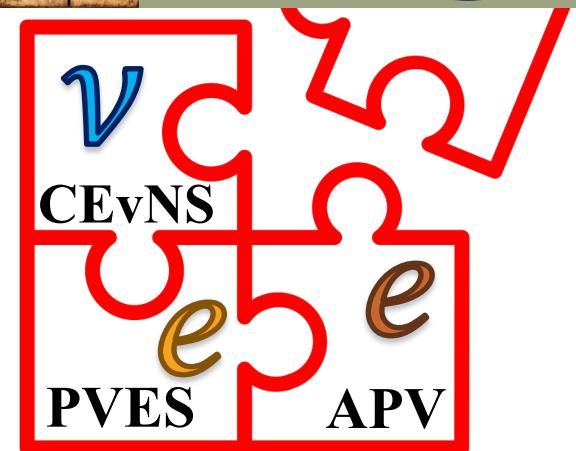
D.M. Meekhof et al. Phys Rev Lett 71, 3442 (1993)

Extremely low momentum transfer  
 **$Q \sim 8 \text{ MeV}$**

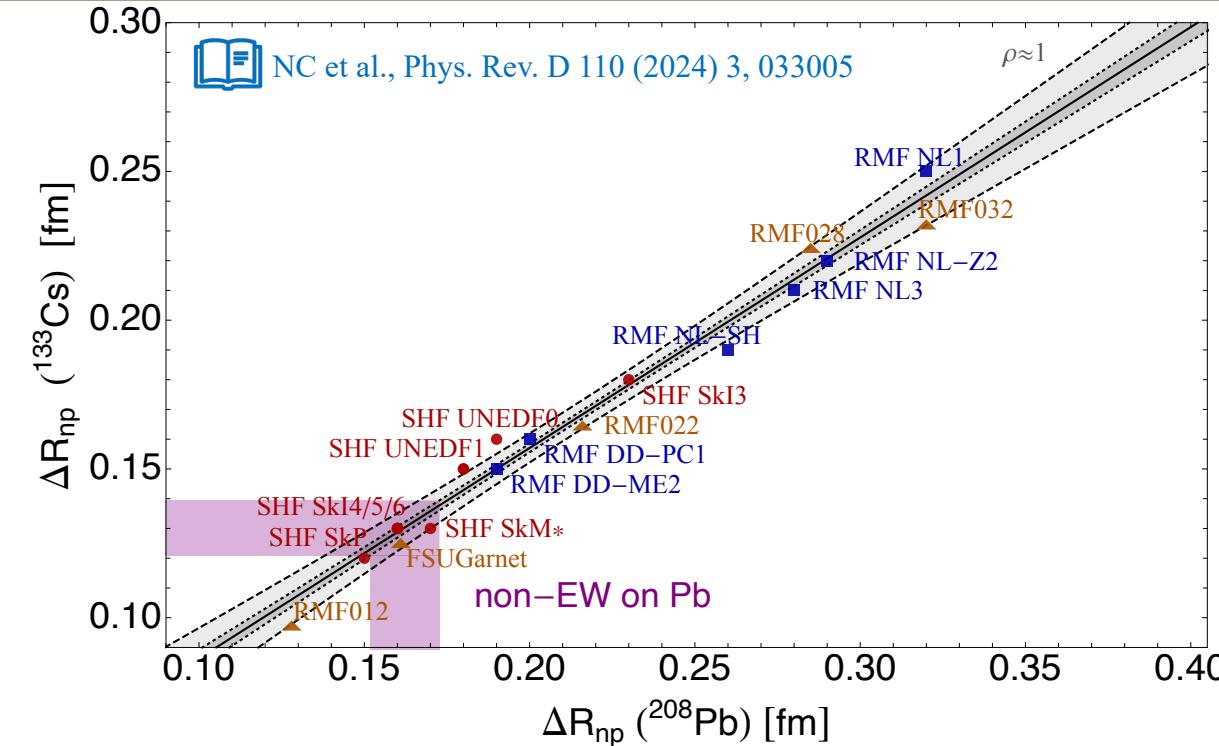
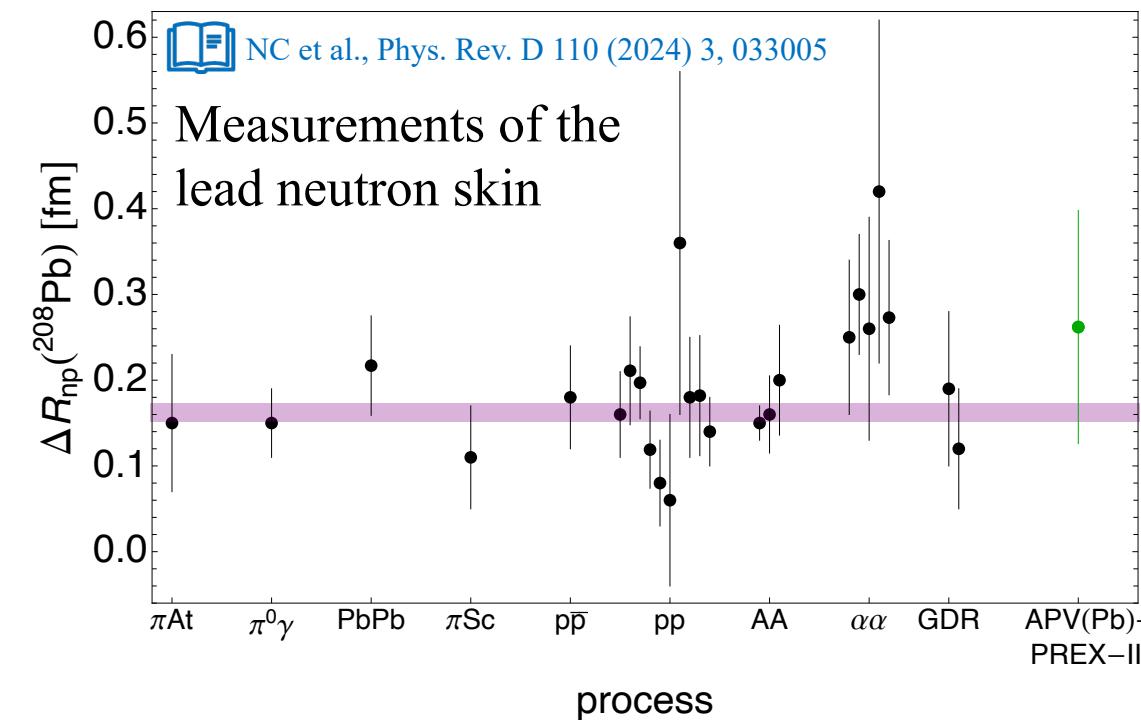
$M_1$ : reduced electric-dipole transition of the magnetic-dipole operator

To be compared to the value of the nuclear weak charge calculated in the SM:  
dependence on  $\sin^2(\vartheta_W)$

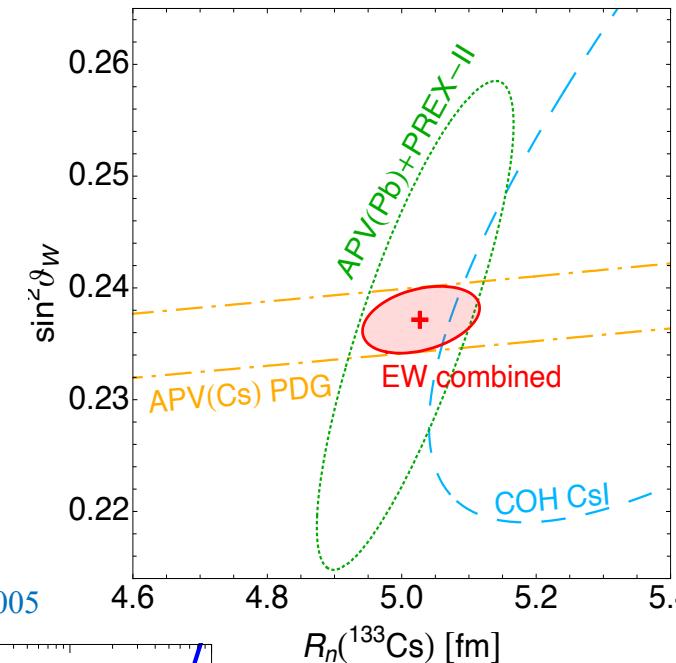
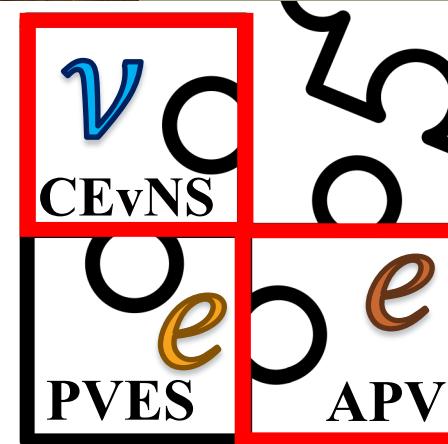




Corelation between the predictions of the lead and cesium neutron skin according to a selection of mean field models



Meas.	$R_n(\text{Cs})$ [fm]	$\sin^2 \theta_W$
APV(Cs) 21 + COH + CSRe	$5.04^{+0.19}_{-0.19}$	$0.2396^{+0.0020}_{-0.0019}$
APV(Cs) 21 + COH + PREX + APV Pb	$5.041^{+0.058}_{-0.057}$	$0.2396^{+0.0017}_{-0.0017}$
APV(Cs) 21 + COH + PREX + APV Pb + non-EW Pb + CSRe	$4.952^{+0.009}_{-0.009}$	$0.2387^{+0.0016}_{-0.0016}$

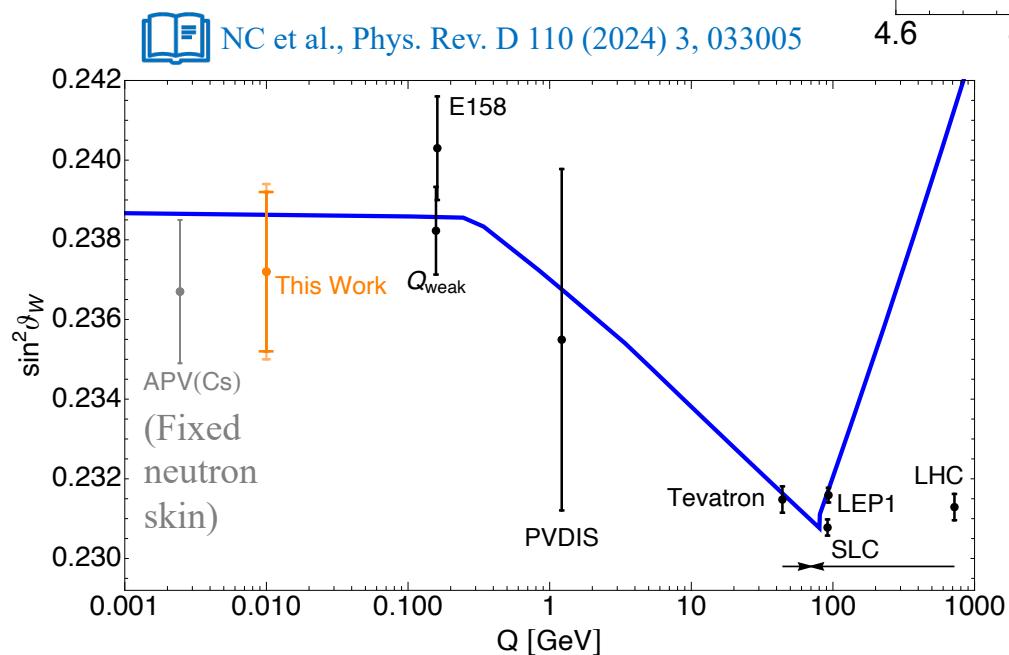


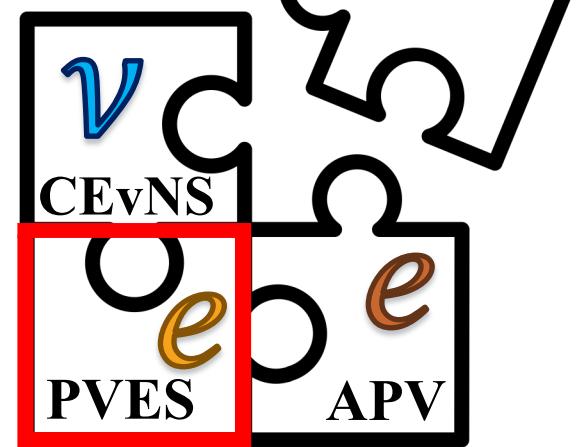
B.K. Sahoo et al. Phys Rev D 103, L111303 (2021)

Meas.	$R_n(\text{Cs}) [\text{fm}]$	$\sin^2 \theta_W$
APV(Cs) 21 + COH + CSRe	$5.04^{+0.19}_{-0.19}$	$0.2396^{+0.0020}_{-0.0019}$
APV(Cs) 21 + COH + PREX + APV Pb	$5.041^{+0.058}_{-0.057}$	$0.2396^{+0.0017}_{-0.0017}$

V. Dzuba et al., Phys. Rev. Letter 109, 203003 (2012)

Meas.	$R_n(\text{Cs}) [\text{fm}]$	$\sin^2 \theta_W$
APV(Cs) PDG + COH + CSRe	$5.04^{+0.19}_{-0.19}$	$0.2372^{+0.0022}_{-0.0022}$
APV(Cs) PDG + COH + PREX + APV Pb	$5.027^{+0.058}_{-0.057}$	$0.2372^{+0.0020}_{-0.0020}$



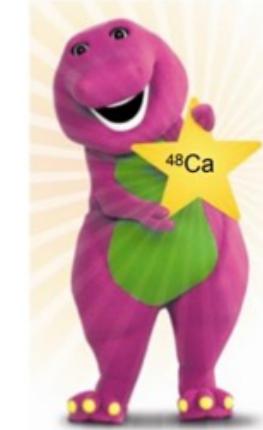


We can consider the PREX twin measurement, CREX, performed with a calcium-48 target

2.1 GeV polarized electrons on a calcium-48 target scattered at  $\theta \sim 5^\circ$  (higher momentum transfer  $q \sim 170$  MeV)

Only one measurement released in 2021

D. Adhikari et al., Phys. Rev. Lett. 129.042501(2022)



Measured the neutron skin of calcium with unprecedented precision:

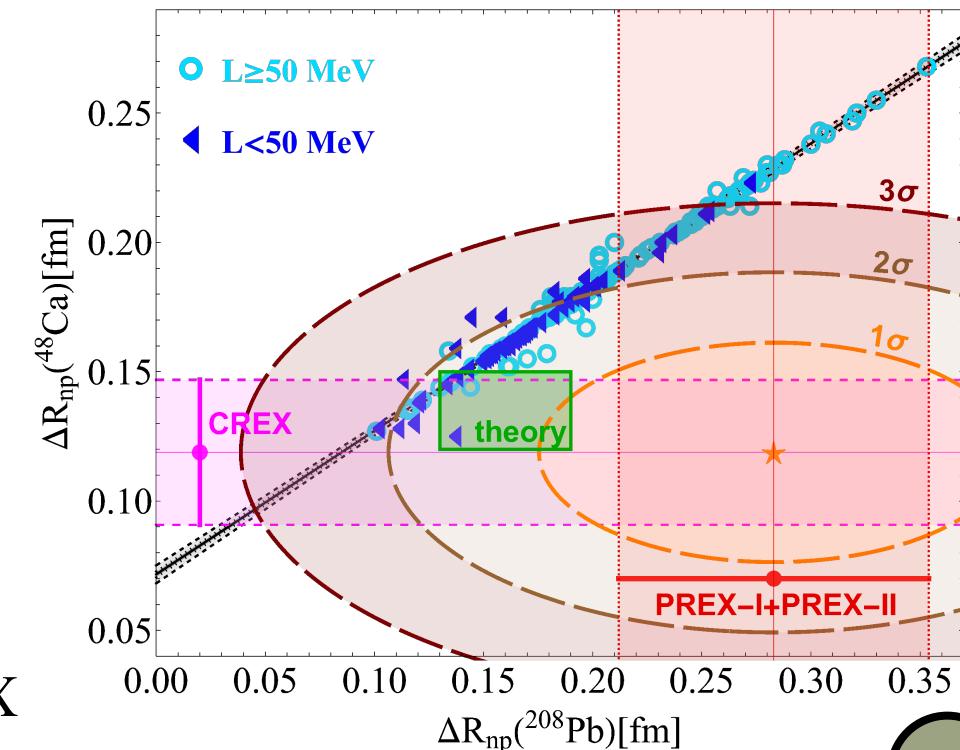
$$\Delta R_{np} = 0.121 \pm 0.026(\text{exp}) \pm 0.024(\text{mod}) \text{ fm}$$

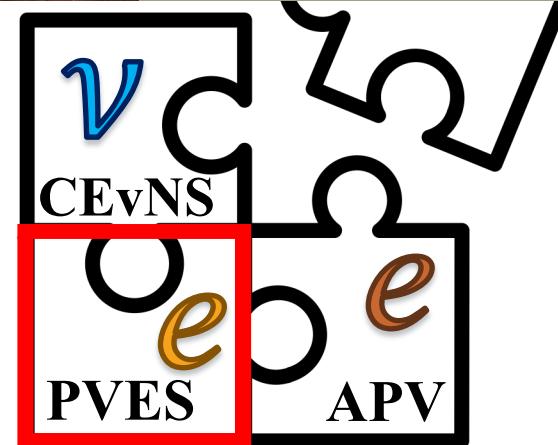
$$\Delta R_{wk} = 0.159 \pm 0.026(\text{exp}) \pm 0.023(\text{mod}) \text{ fm}$$

CREX found a neutron skin perfectly compatible with EDF and ab-initio predicted values

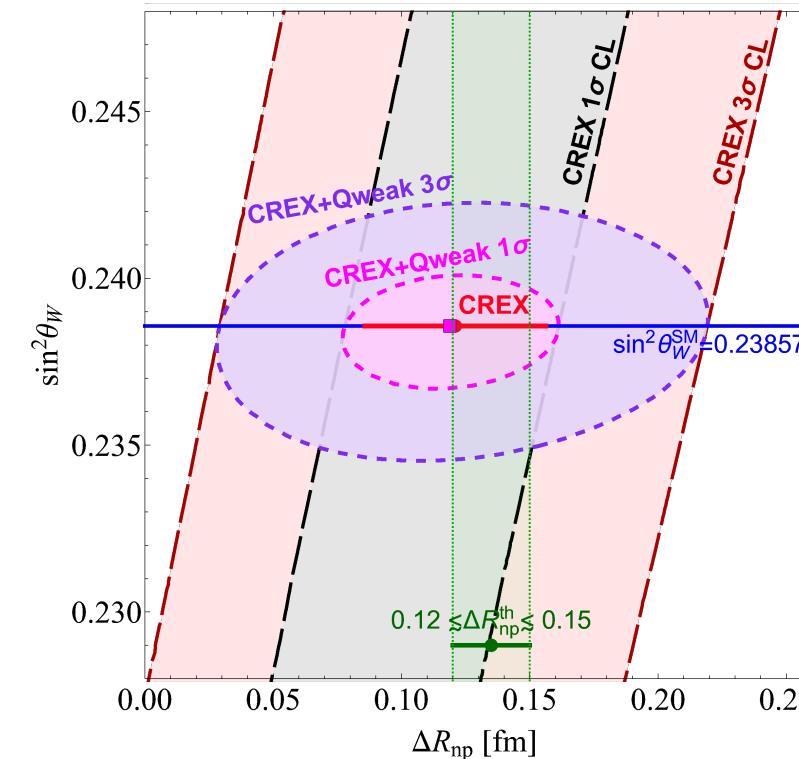
$$\Delta R_{np}^{th} \approx 0.12 - 0.15 \text{ fm}$$

We performed the same analysis applied to PREX also for CREX



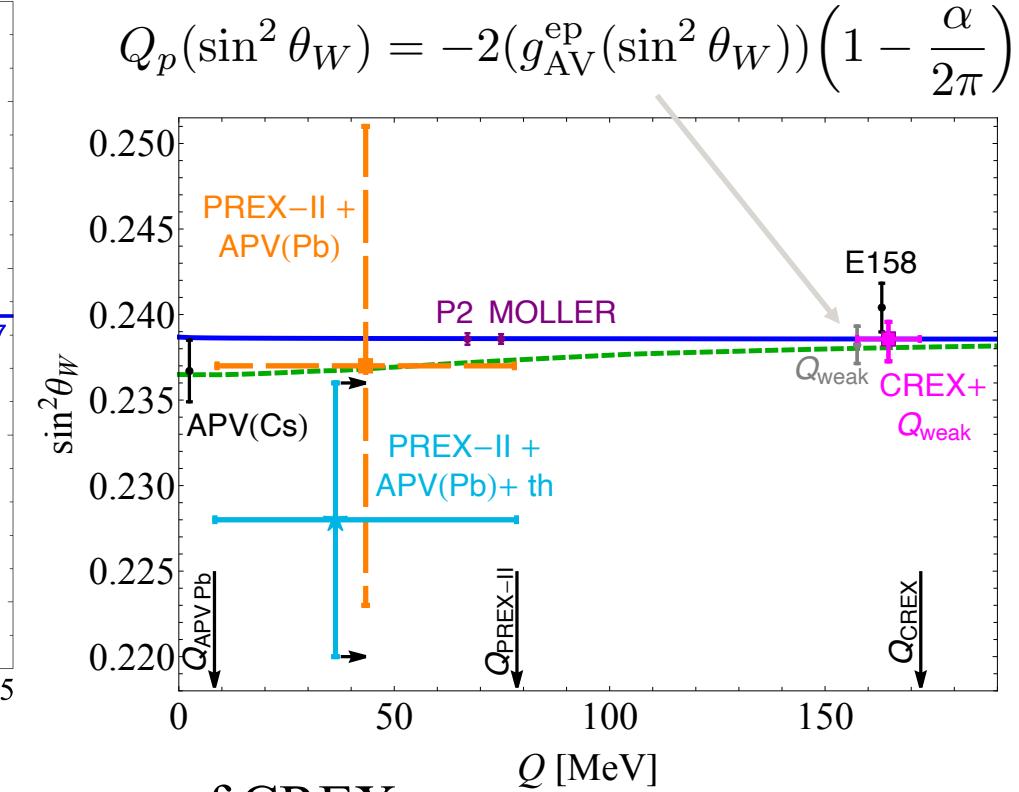


We can use the weak mixing angle measurement from the Qweak measurement of the proton weak charge (no dependence on the neutron skin)



### CREX+Qweak

$$\Delta R_{np} = 0.119 \pm 0.028 \text{ fm}$$
$$\sin^2 \theta_W = 0.2386 \pm 0.0013$$



- Qweak sets a horizontal band which breaks the degeneracy of CREX
- The precision is preserved
- CREX and PREX point toward different directions!



A preliminary study on the possibility to use a pixel-like backward detector

At MESA, the form factor could be measured at more **energies** by using a pixelized detector for the backward measurement

A scan of the form factor in a certain range of momenta could be achievable

