The General AntiParticle Spectrometer

JENAA
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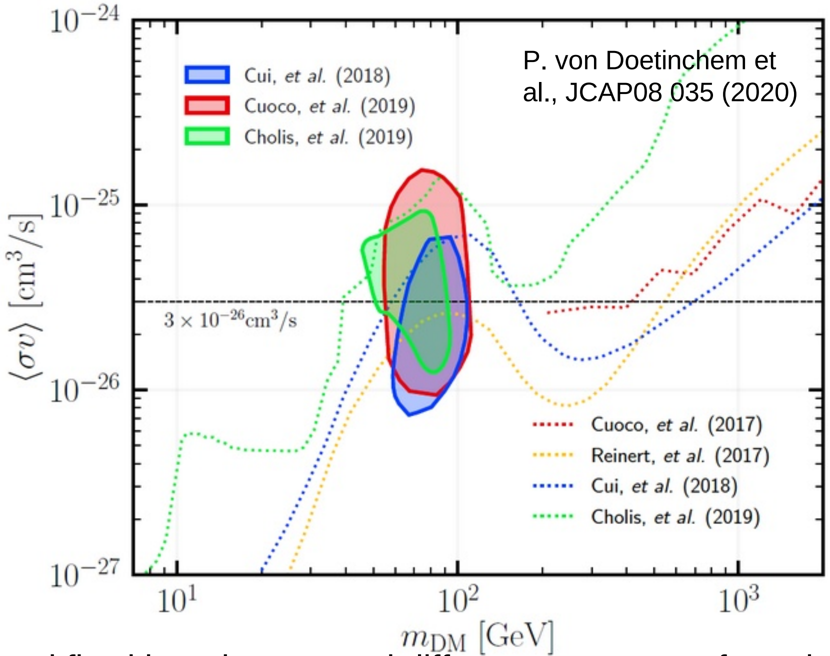
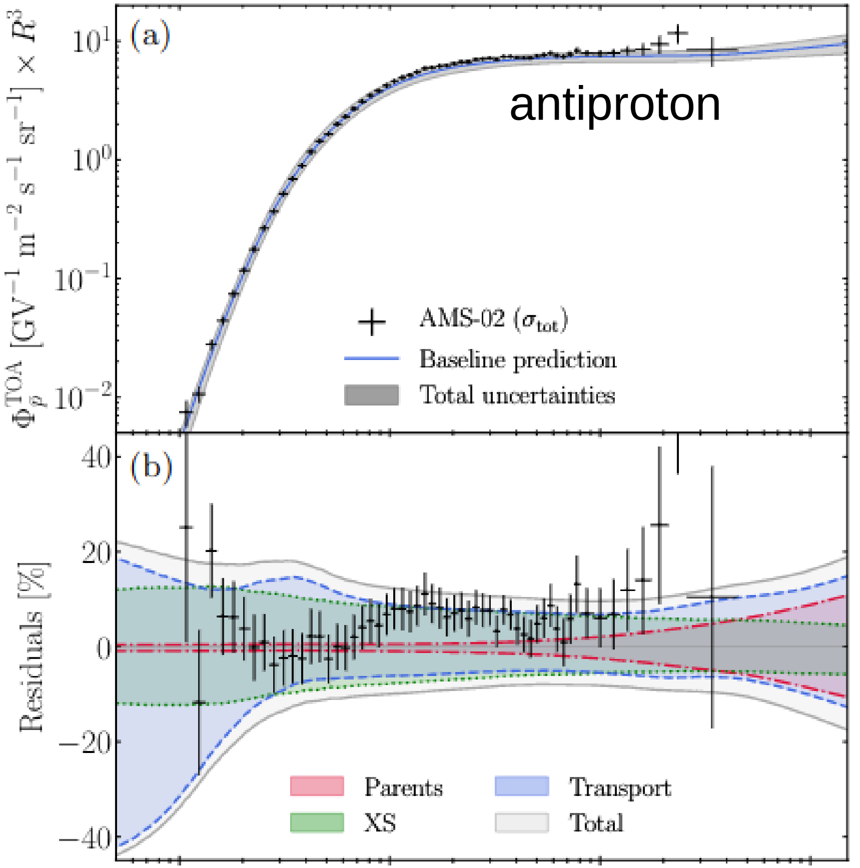
Unexplained features in cosmic antiparticles?

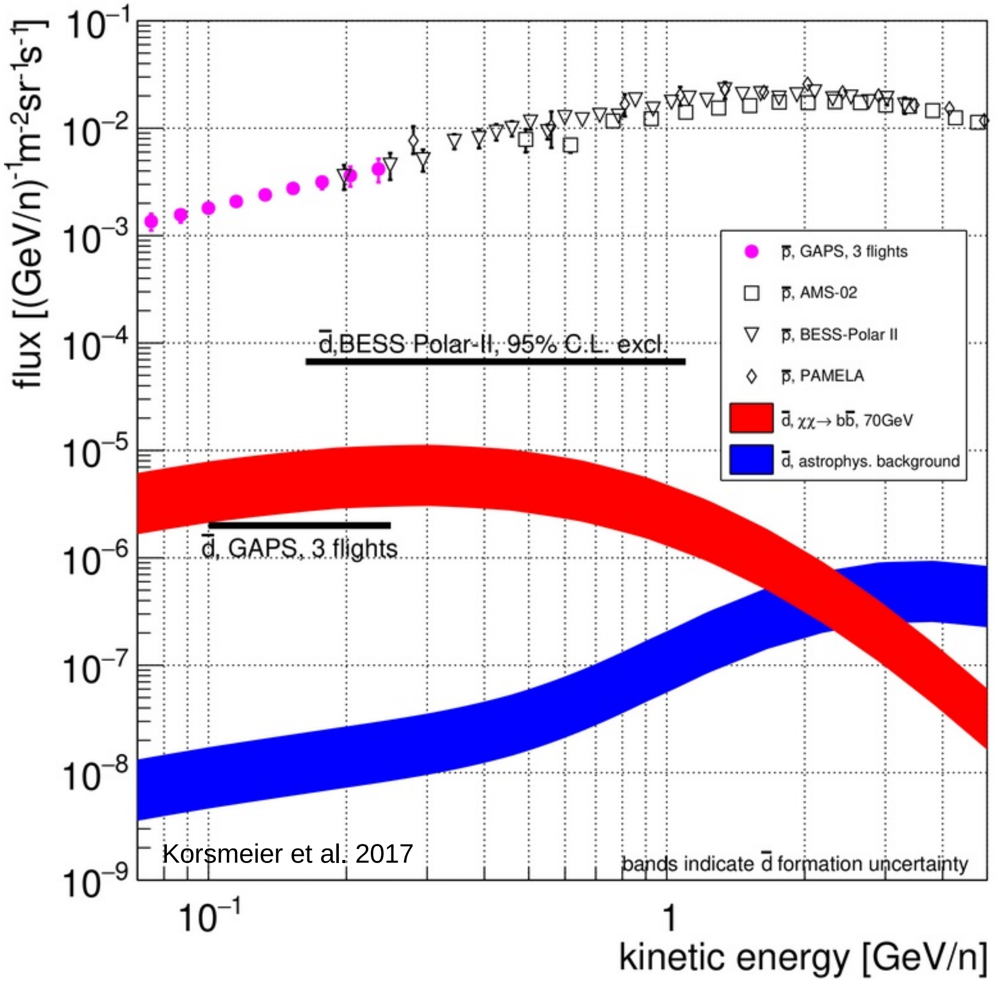




- combined fit with antiproton and diffuse gamma-rays from the Galactic Center → 70-80GeV DM particle? (ongoing debate)
- unexplained feature in positrons:
  - astrophysical origin → pulsars
  - SNR acceleration
  - dark matter annihilation
- **understanding astrophysics background is a challenge** → better constraints on cosmic-ray propagation and production needed

Status cosmic-ray antinuclei searches

- **Potential $\bar{p}$ excess** in AMS-02 data above secondary background predictions at R~10 GV was found in various studies → significance level unclear
- AMS-02 reported at conferences the observation of **antihelium candidates (~1/year)** → interpretations are actively ongoing
- **Discussed physics models that explain antihelium candidates include:**
  - Secondary astrophysical background
  - Dark matter annihilation or decay
  - Nearby antistar: at distance of ~1pc
- **No explanation of antiproton nor antihelium should overproduce antideuterons relative to existing limits**
- **Search for antinuclei with independent technique is critical**
- Review based on 2nd Cosmic-ray Antideuteron Workshop: JCAP08(2020)035, arXiv:2002.04163

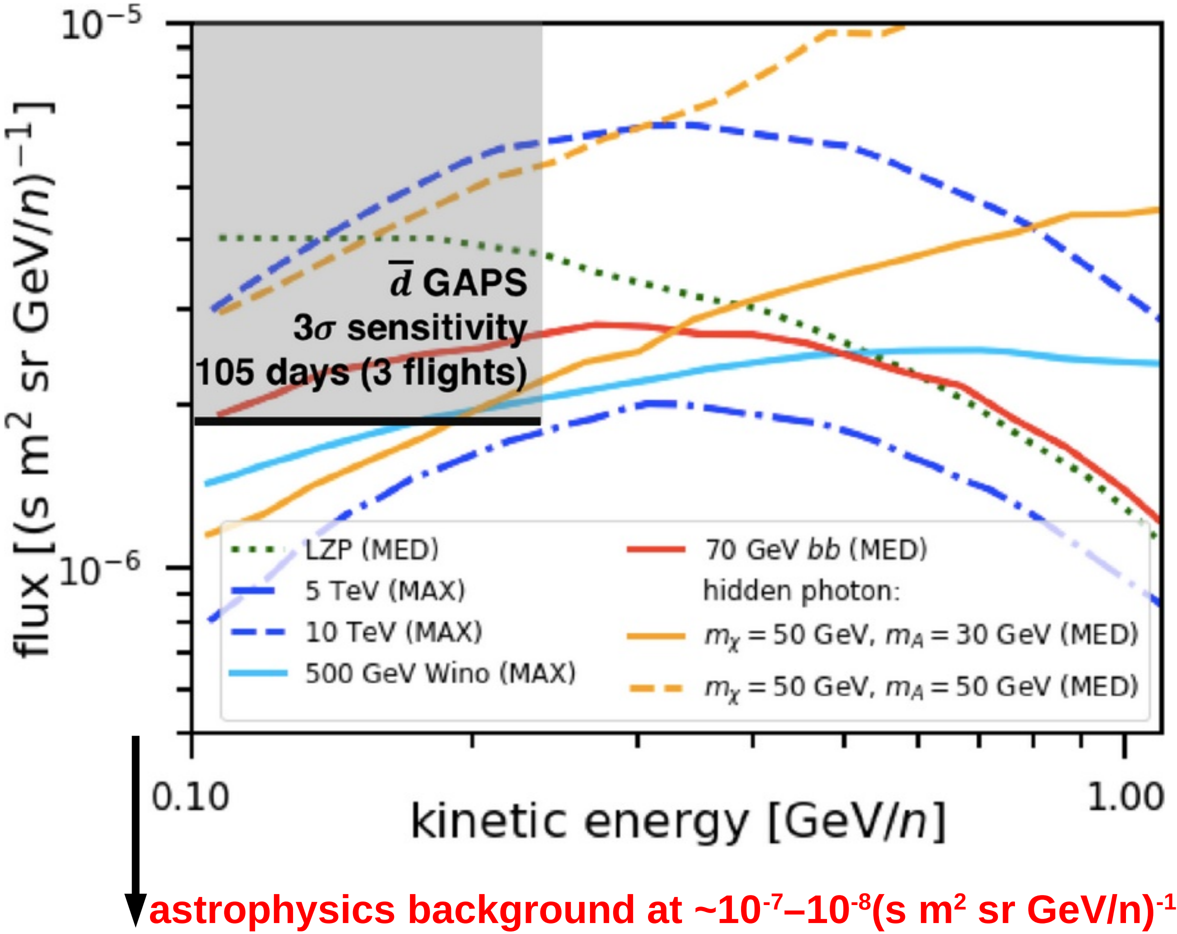


Antideuteron model sensitivity

T. Aramaki et al., Astropart. Phys. 74, 6 (2016)

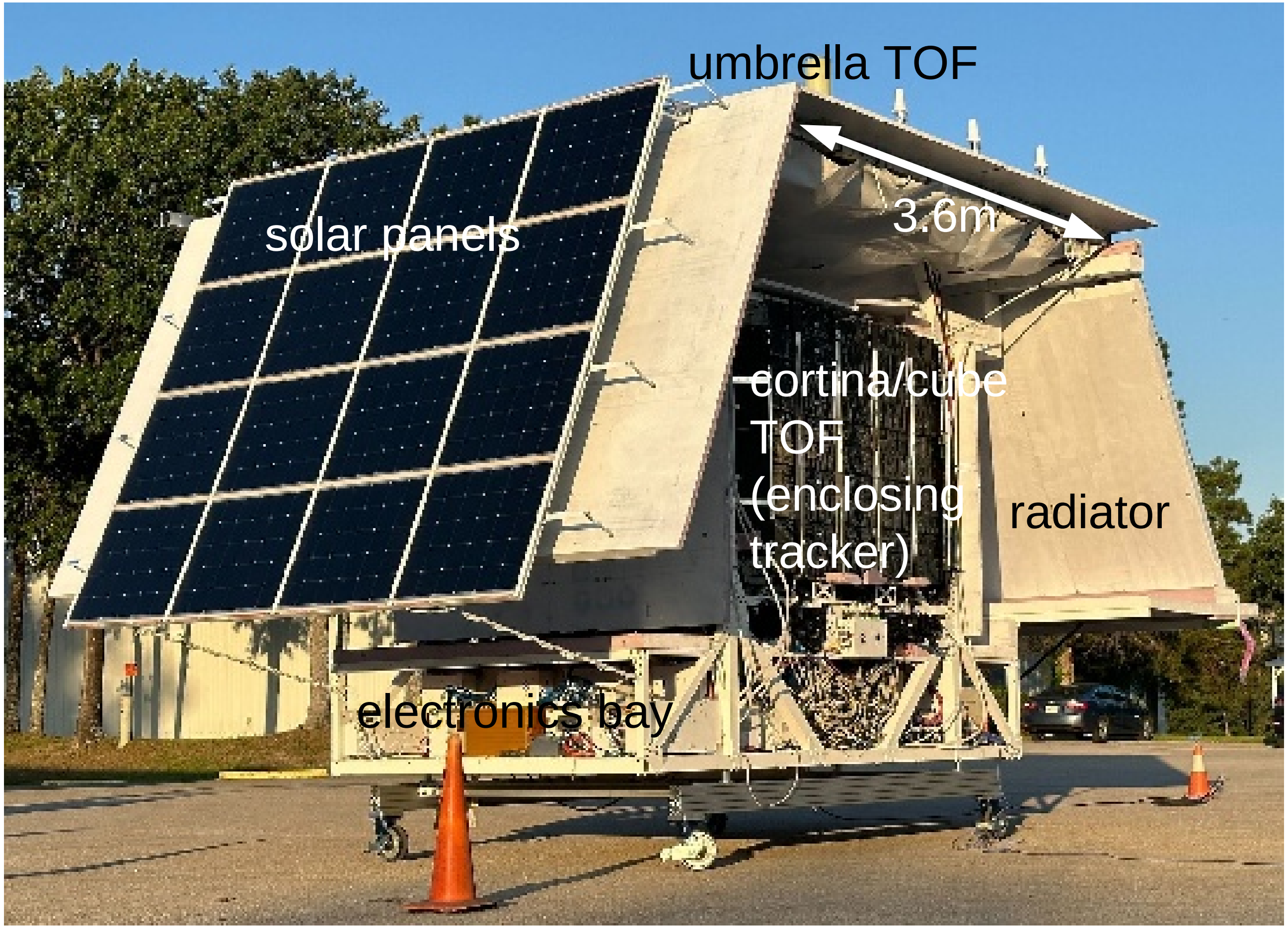
- **Low-energy antideuterons are essentially free of astrophysics background**

- **Wide range of dark matter models**, e.g.:
  - Generic 70GeV WIMP annihilation model that explains antiproton excess and γ-rays from Galactic center
  - Dark matter gravitino decay
  - Extra dimensions
  - Heavy DM models with Sommerfeld enhancement
  - Dark photons (inaccessible to other techniques)
  - Selection of publications:
    Braeuninger et al. Physics Letters B 678, 20–31 (2009)
    Cui et al, JHEP 1011, 017 (2010)
    Hryczuk et al., JCAP 1407, 031 (2014).
    Korsmeier et al., Physical Review D 97, 103011 (2018)
    Randall & Xu, JHEP (2020)



astrophysics background at ~10^{-7}–10^{-8}(s m^2 sr GeV/n)$^{-1}$

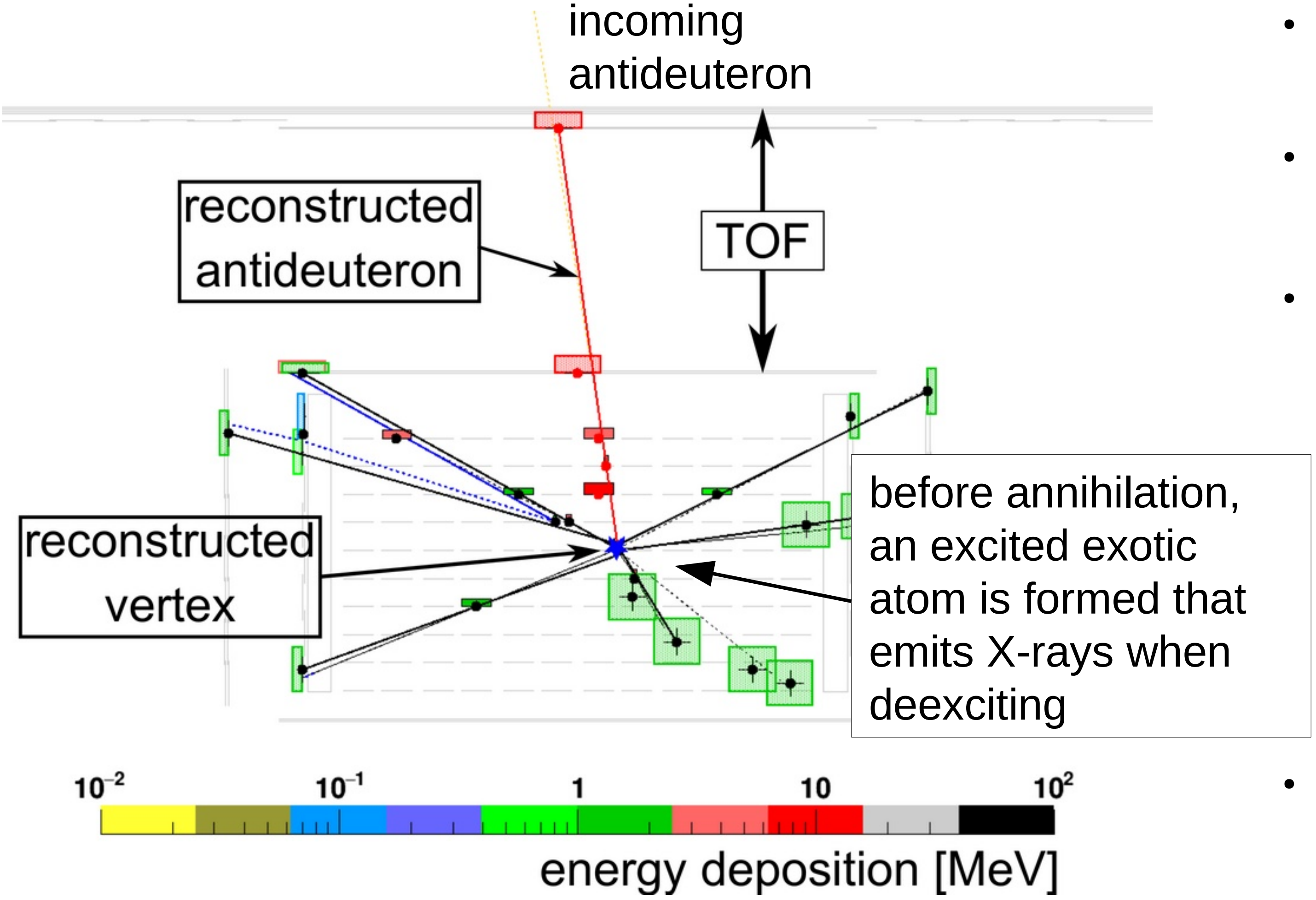
The GAPS experiment



mass: ~2,500kg
power: 1.3kW

- The **G**eneral **A**nti**P**article **S**pectrometer is the first experiment dedicated and optimized for low-energy cosmic-ray antinuclei search
- Requirements: long flight time, large acceptance, large identification power, flight at low-geomagnetic cutoff location
- **GAPS will deliver:**
  - a precision antiproton measurement in an unexplored energy range <0.25 GeV/n
  - antideuteron sensitivity 2 orders of magnitude below the current best limits, probing a variety of DM models across a wide mass range
  - leading sensitivity to low-energy cosmic antihelium nuclei
- **GAPS is under construction, preparing for first Antarctic Long Duration Balloon flight in December 2024**

GAPS principle

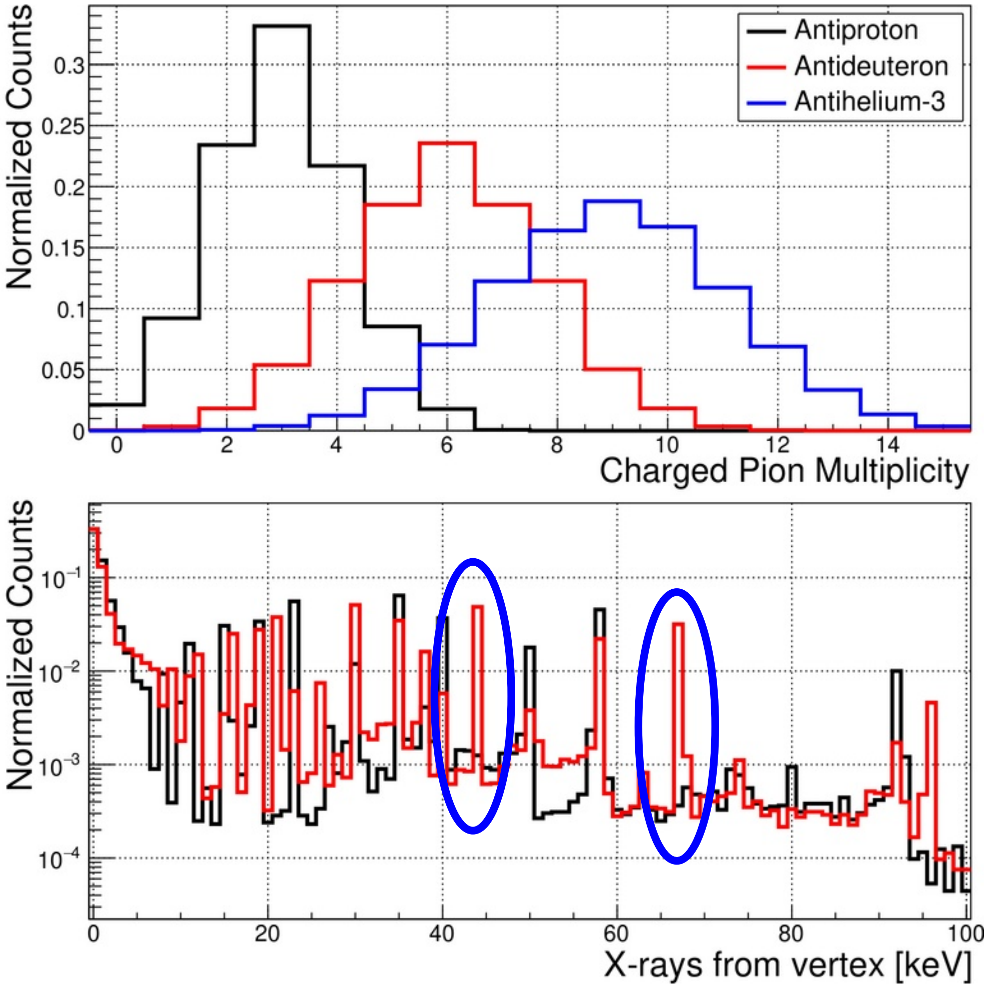


- antiparticle slows down and stops in material
- near-unity chance for creation of an excited exotic atom (E_{kin}~E_{I})
- deexcitation:
  fast ionization of bound electrons (Auger)
  → complete depletion of bound electrons
  Hydrogen-like exotic atom (nucleus+antideuteron) deexcites via characteristic X-ray transitions depending on antiparticle mass
- Nuclear annihilation with characteristic number of annihilation products

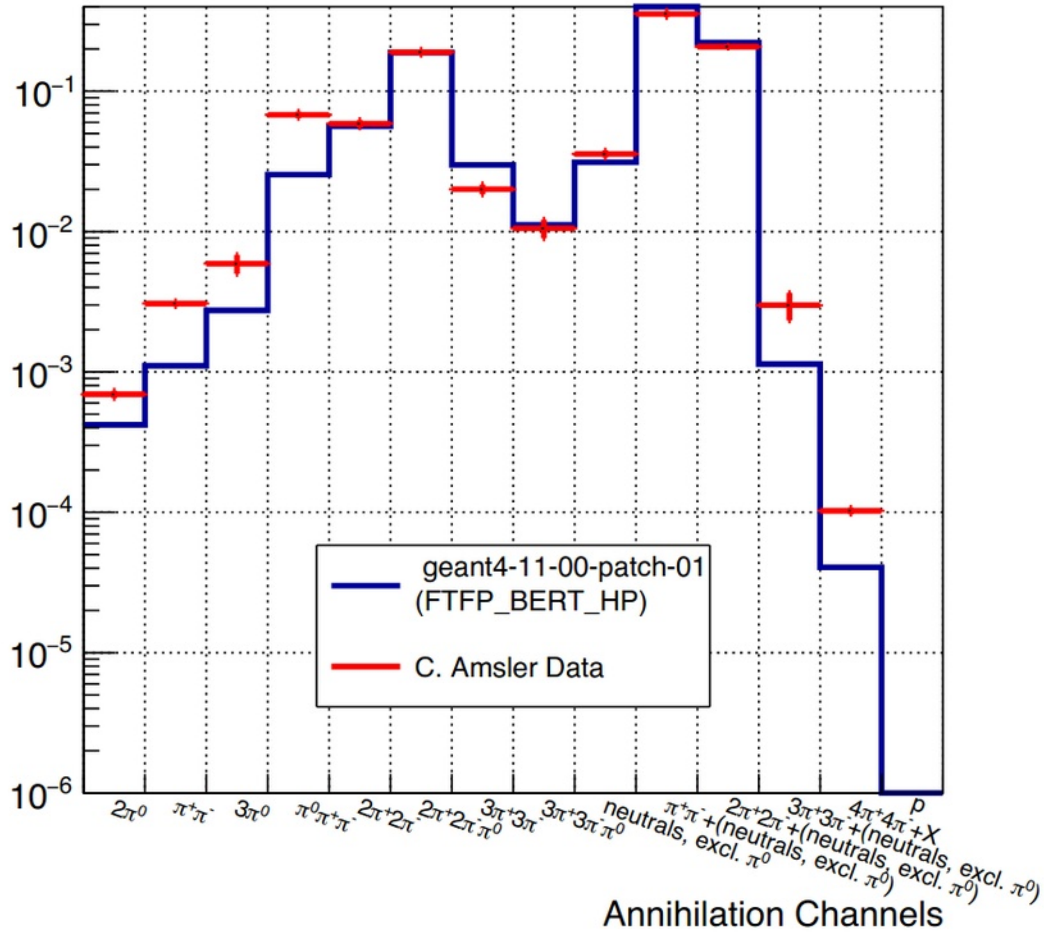
GAPS identification technique

GAPS identification technique uses:

- Energy loss in the detector of the antinucleus (depends on *Z* and β)
- Deexcitation X-rays from exotic atom
- Multiplicity of charged annihilation products



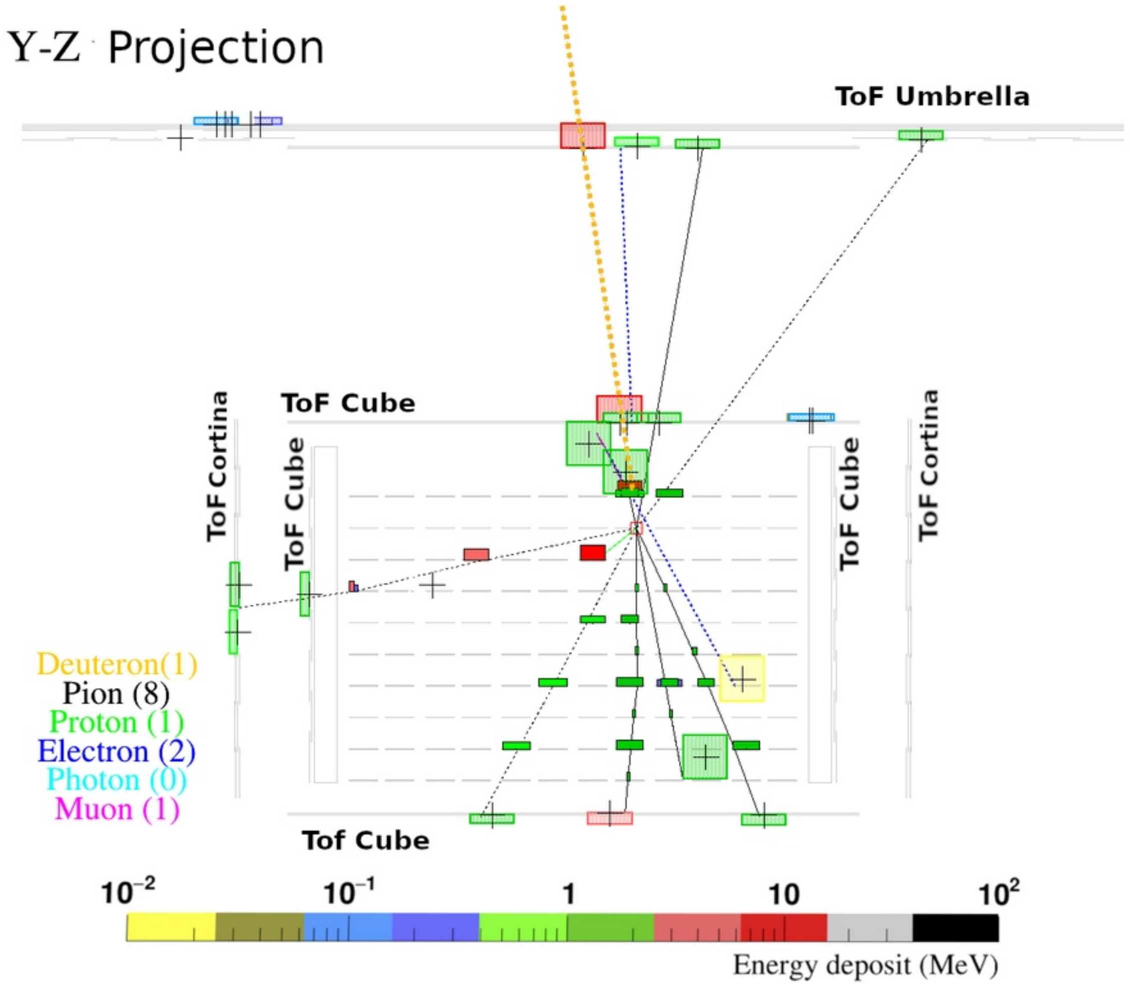
$\bar{p}$+p annihilation at rest



- test of annihilation physics in Geant4 is ongoing
- use antiproton data for validation
- work with Geant4 developers

Event reconstruction

R. Munini et al., Astropart. Phys. 133, 102640 (2021)



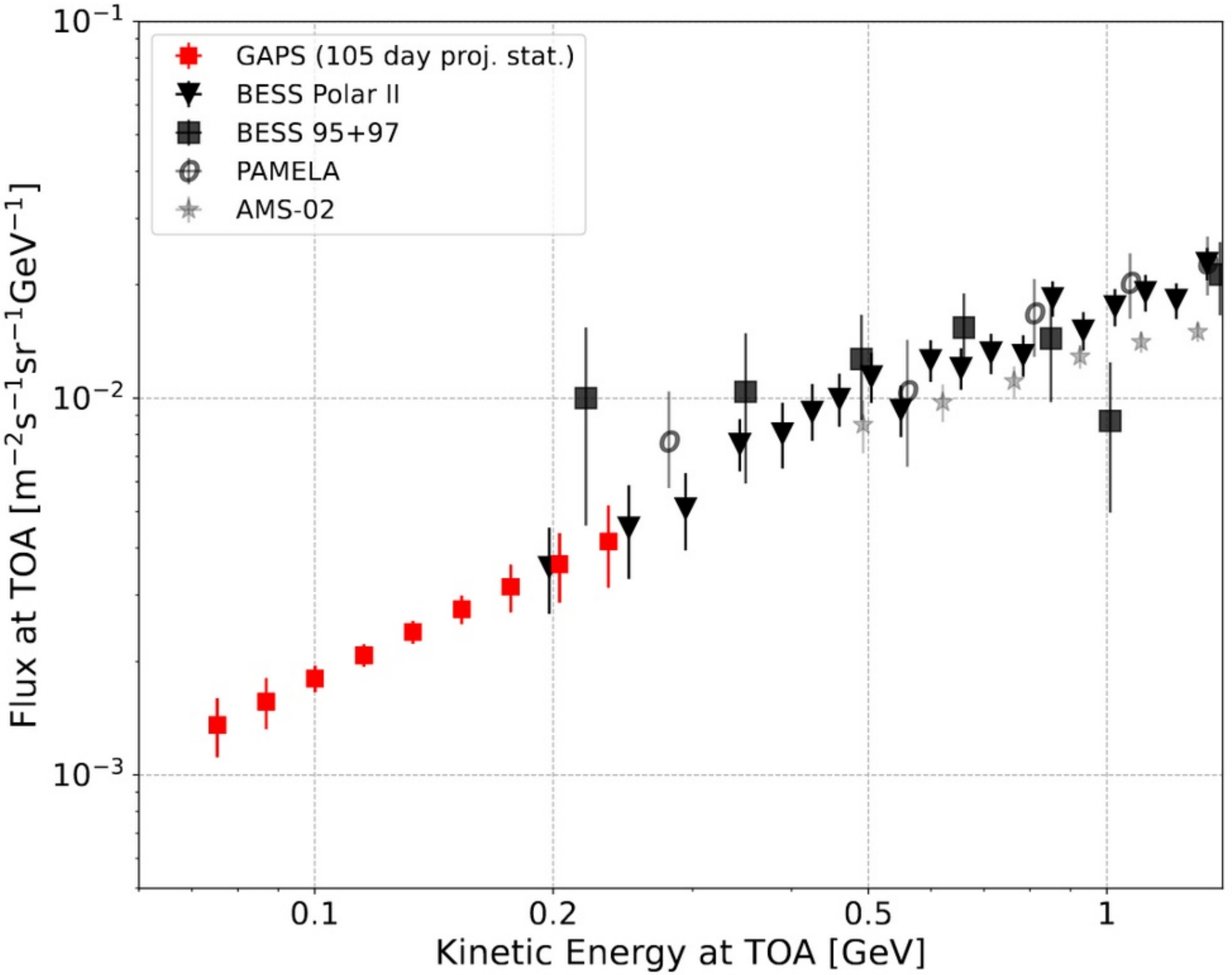
- For the event reconstruction it is critical to identify a well defined primary track
  → β measurement, energy deposition, column density
- The primary track is used as a seed for the determination of the stopping vertex with the corresponding secondary tracks

Antiproton sensitivity

F. Rogers et al., Astropart. Phys. 145, 102791 (2022)

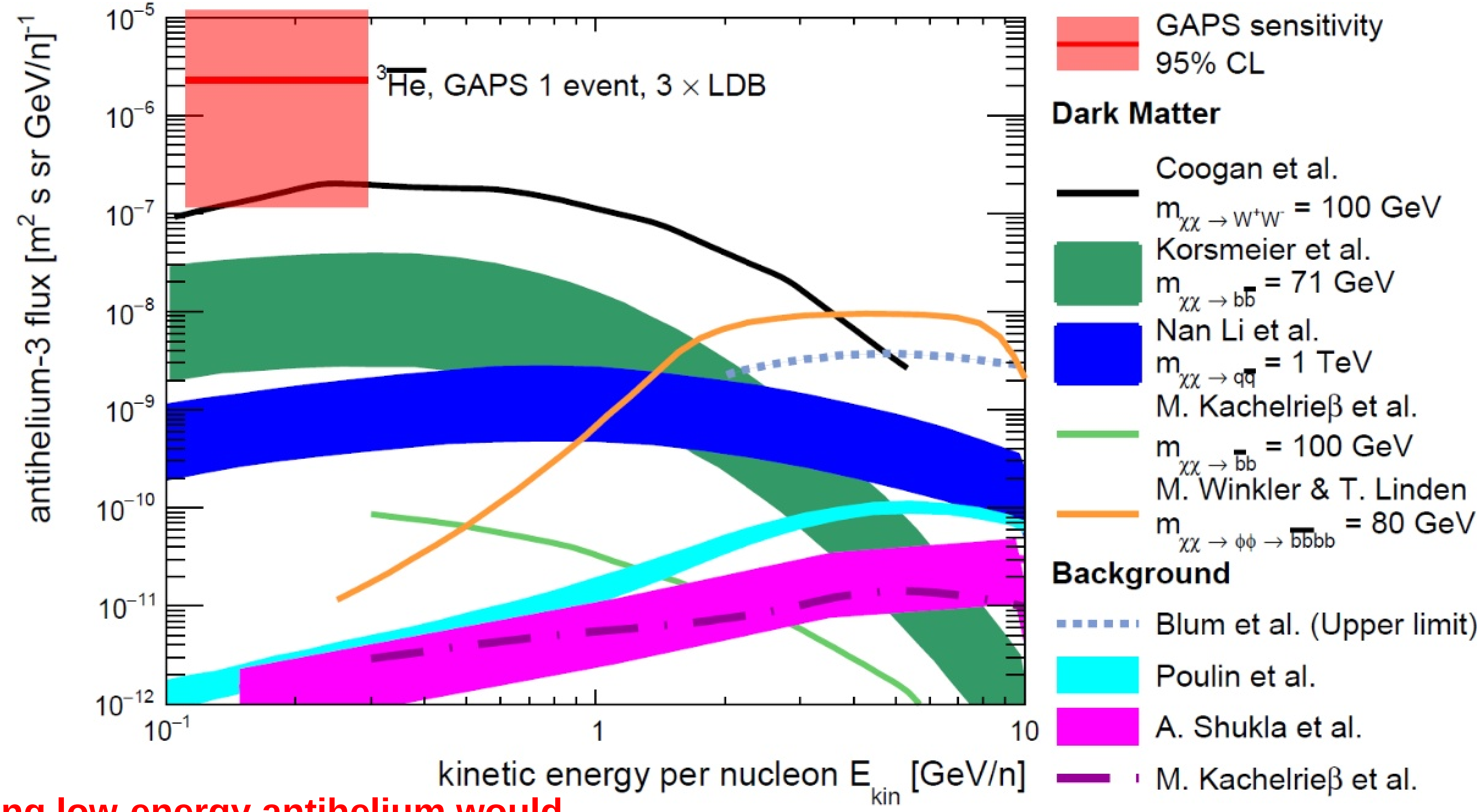
- Precision antiproton spectrum in unexplored low-energy range (<0.25 GeV/n): ~500 antiprotons for each long-duration balloon flight
- Validation of technique:
    - First cosmic rays detected with the exotic atom method
    - Reconstruction of annihilation signature
    - X-rays from exotic atom deexcitation
    - Test models for atmospheric effects

  → Reduces the systematic uncertainties for antideuteron search
- Probe light dark matter models and primordial black hole evaporation



Antihelium-3 sensitivity

N. Saffold et al., Astropart. Phys. 130, 102580 (2021)

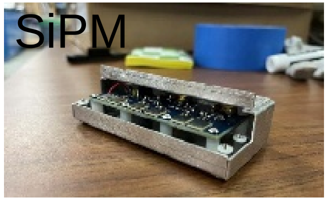


Finding low-energy antihelium would be truly revolutionary new physics

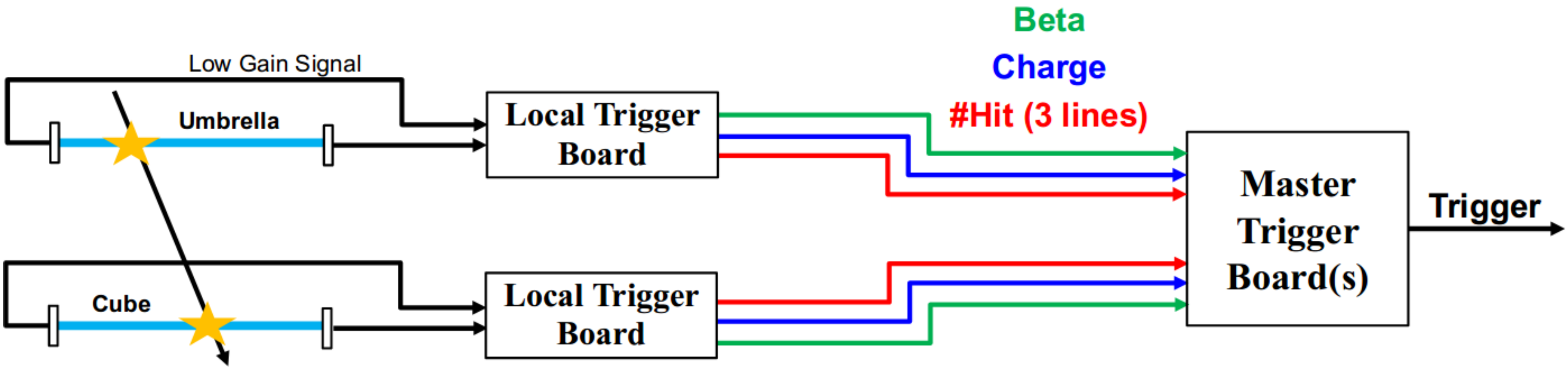
Time-of-Flight

- Tasks:
  - Main trigger system, special antinuclei trigger achieves a manageable rate of ~500 Hz
    (down from 200 kHz individual TOF paddle rate)
  - Tracking of incoming (anti)particles and outgoing secondary particles
  - Particle velocity measurement

- Plastic scintillator (Eljen EJ-200: 160-180cm long, 0.6 cm thick) with 6 SiPMs per end (Hamamatsu S13360-6050VE)

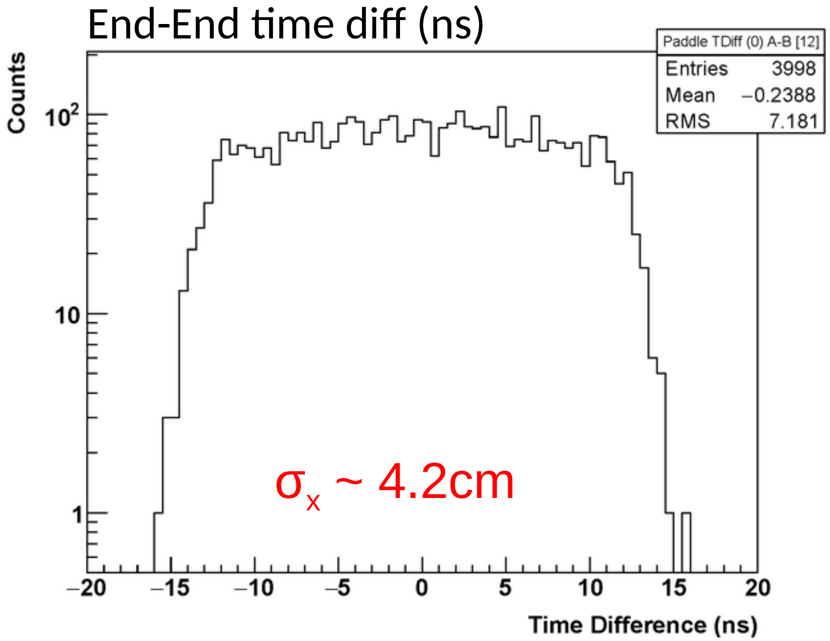


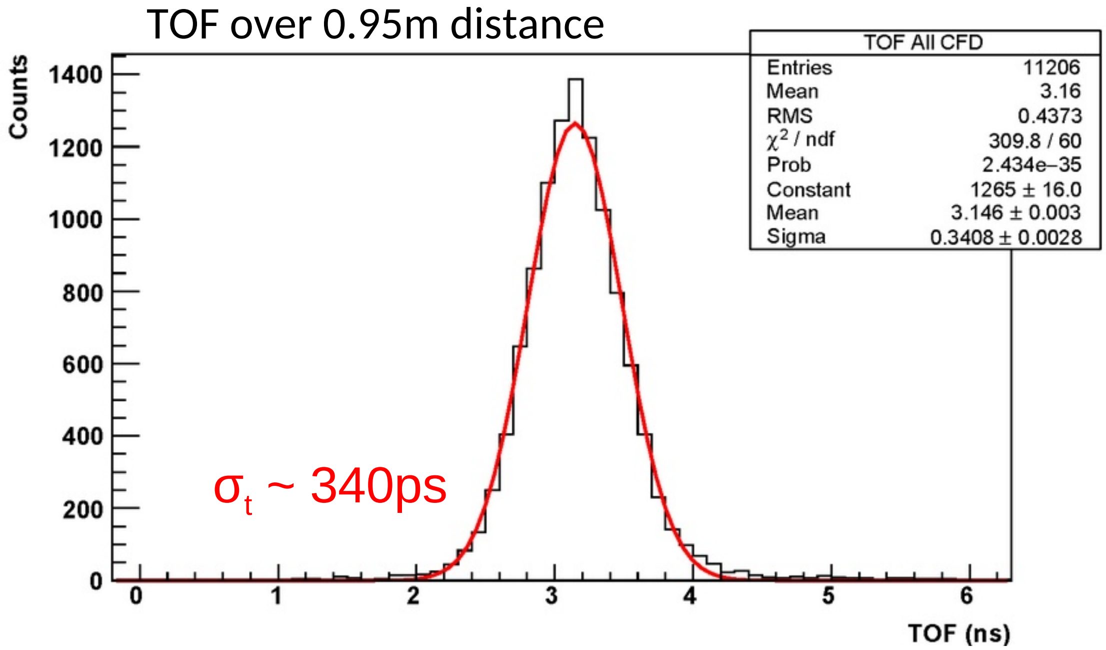
Trigger design



- main background: protons, alpha, carbon
- High-speed trigger and veto:
    - stopping events deposit more energy (lower beta)
    - annihilation events produce more TOF hits
- **smart combination reduces trigger rate to be below 500Hz**

TOF performance

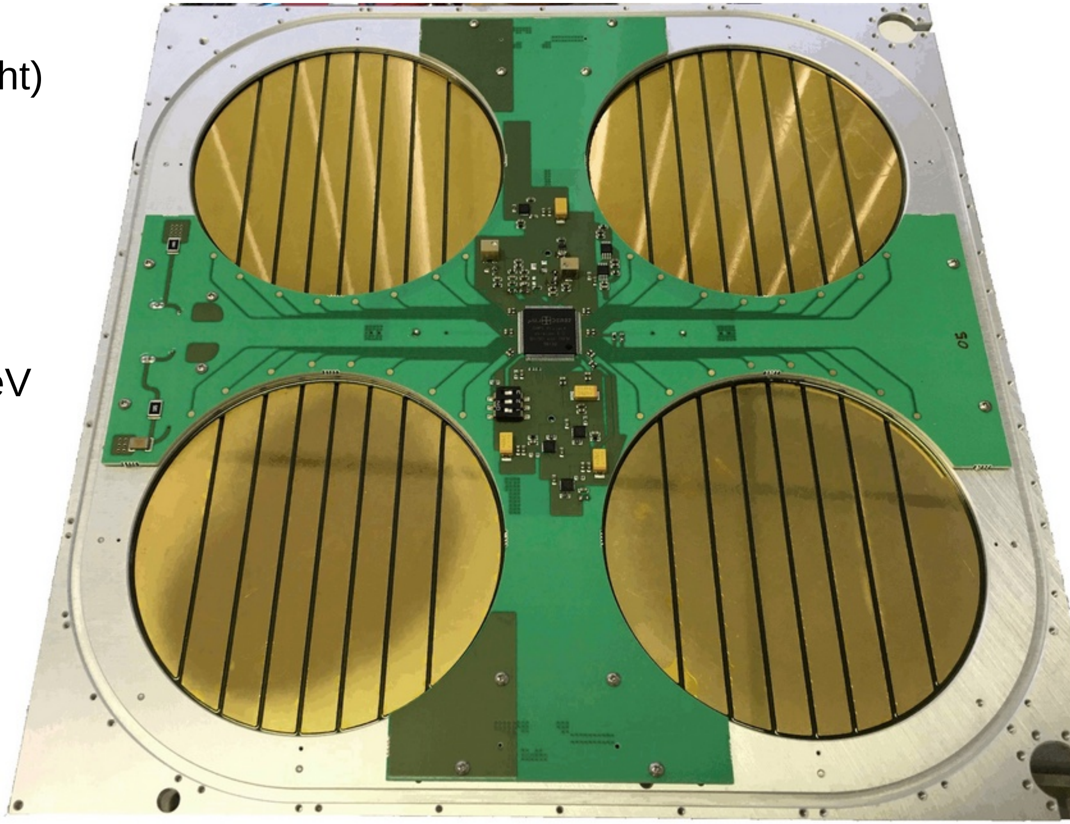




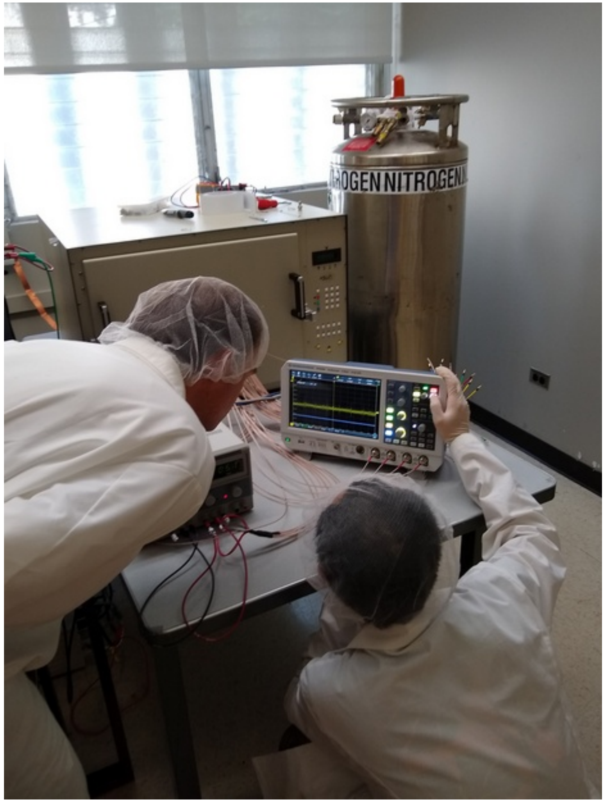
- fast sampling with custom-made readout board, based on the DRS-4 ASIC: <400ps timing resolution achieved

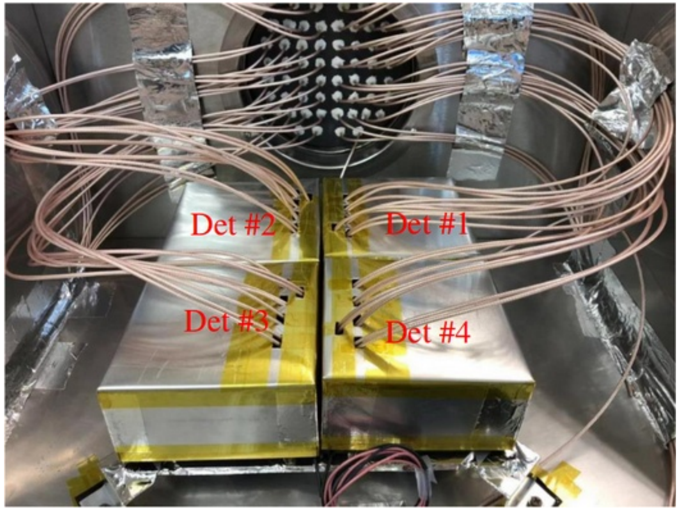
Tracker

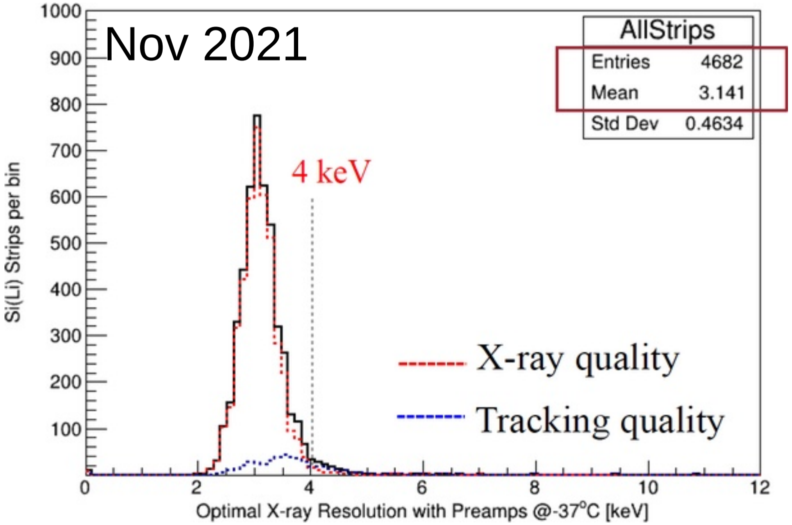
- Tracker acts as target and tracking device
- GAPS can accommodate 1,440 4” Si(Li) detectors, 2.5mm thickness (1109 detectors calibrated for first flight)
- Operation at temperature of –35C to –45C, cooling system will use novel OHP approach
- Readout via custom ASIC: integrated low-noise preamplifier with large dynamic range: 10keV to 100MeV
- Publications:
  - Perez et al., NIM A 905, 12 (2018)
  - Kozai et al., NIM A 947, 162695 (2019)
  - Rogers et al., JINST 14, P10009 (2019)
  - Saffold et al., NIM A 997, 165015 (2021)
  - Manghisoni et al., IEEE 68 (11), 2661 (2021)
  - Kozai et al., NIM A 1034, 166820 (2022)
  - Roach et al., IEEE 70 (8) (2023)
  - Xiao et al., in preparation



Tracker qualification



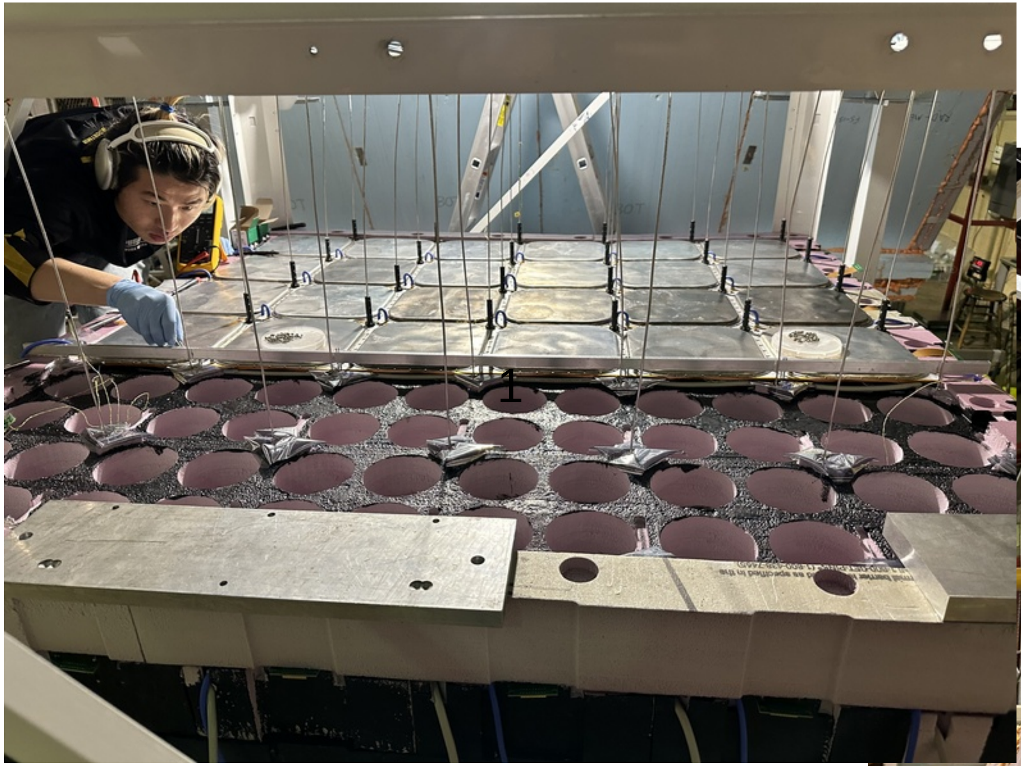


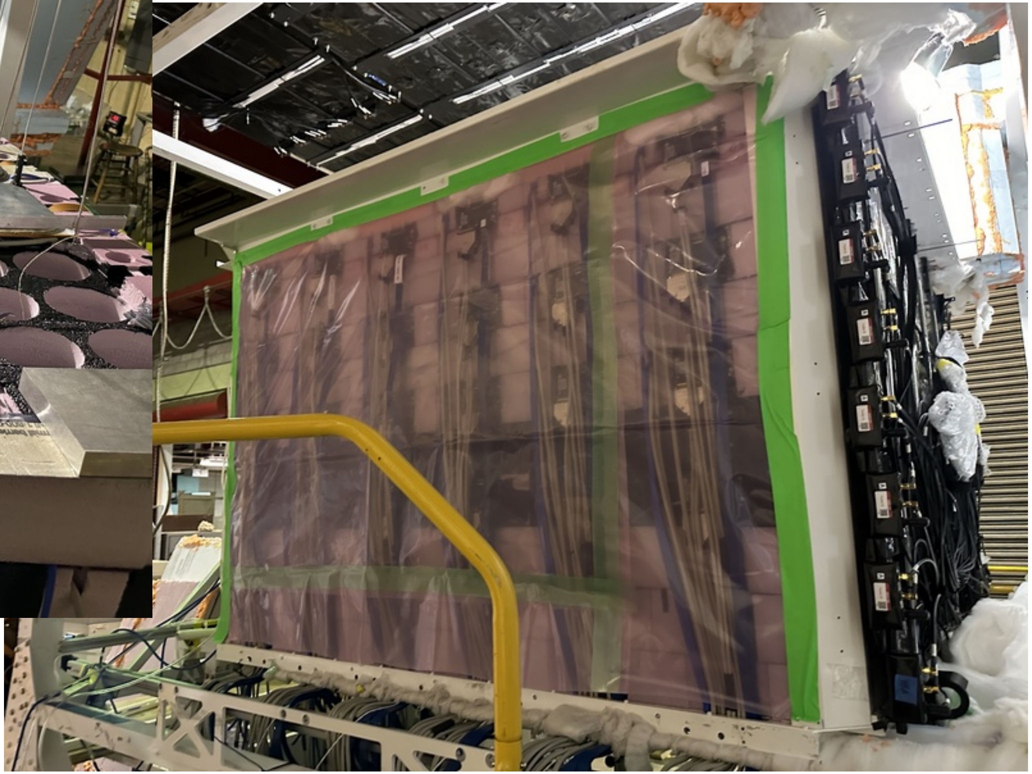


- Single detector test shows the required resolution of the detectors

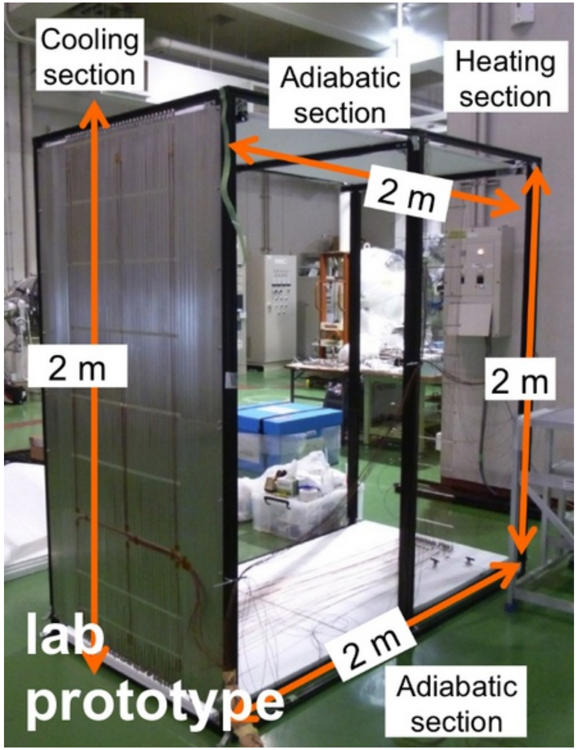
- All detector modules were calibrated at MIT and UHM

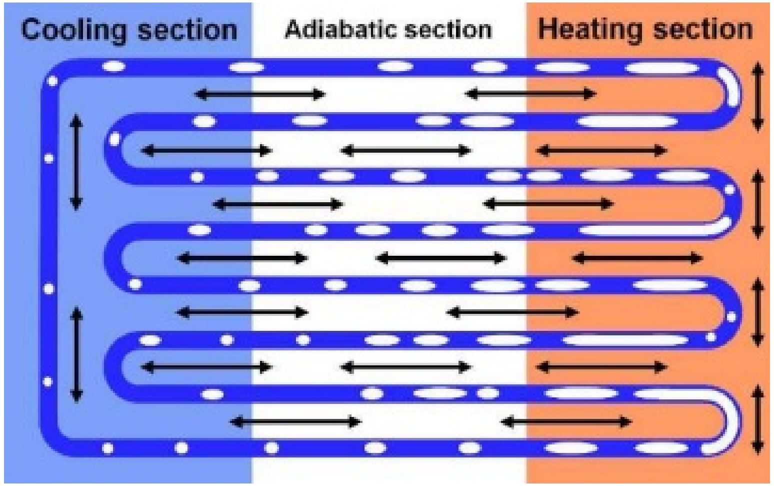
Tracker integration





Oscillating heat pipe cooling system



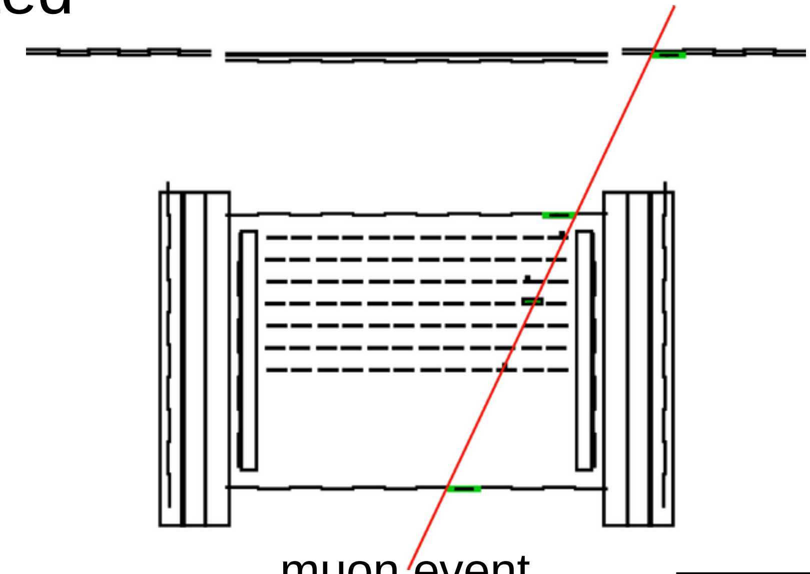


- passive cooling approach developed at JAXA/ISAS:
- small capillary metal tubes filled with a phase-changing refrigeration liquid
- small vapor bubbles form in the fluid
  → expand in warm sections/contract in cool sections
- rapid expansion and contraction of these bubbles create thermo-contraction hydrodynamic waves that transport heat
- no active pump system is required
- First prototype was flown in 2012 and another prototype was flown from Ft. Sumner in 2019

Okazaki et al., J. Astr.. Instr. 3 (2014)
Fuke et al., J. Astron. Instrum. (2017)
Okazaki et al., Appl. Therm. Eng. (2018)
Fuke et al., NIM A 1049, 168102 (2023)

GAPS in Texas

- GAPS passed compatibility at CSBF (Palestine, TX)
- Data taking period in Texas just completed and packing for Antarctica started



muon event

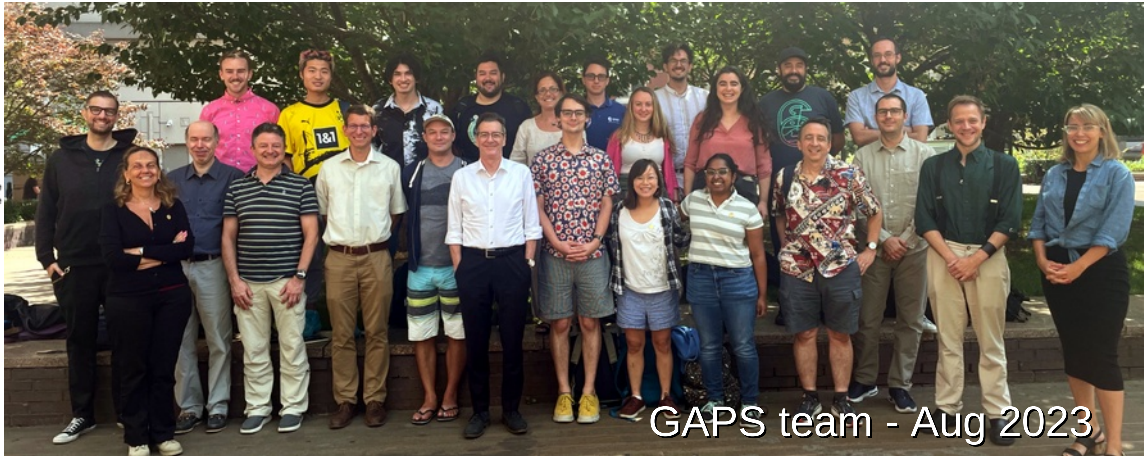
hang test

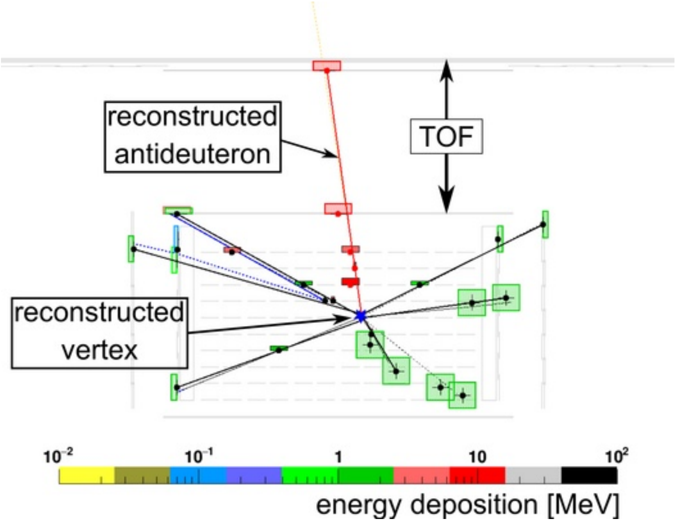
Timeline

- Integration and systems testing in spring 2023
- Thermal vacuum testing summer 2023
- Compatibility testing summer 2024
- **Continued testing before first flight in Dec. 2024 from McMurdo, Antarctica**

Image credit: NASA (cropped)

GAPS path forward





- **GAPS will deliver:**
  - a precision antiproton measurement in an unexplored energy range <0.25 GeV/n
  - antideuteron sensitivity 2 orders of magnitude below the current best limits, probing a variety of DM models across a wide mass range
  - the only complementary probe of the AMS-02 antinuclei signal

- GAPS instrument integration is ongoing → **first flight in austral summer 2024**