## **Background simulation for ProtoDUNE**

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### **Experimental set-up: T2 target**



P	T2 production yield $(1/PoT)$	channel	$BR(P \rightarrow channel)$	$\nu_{\alpha}$ crossing ND (1/PoT)	
$K_S^0$	1	$\pi^- e^+ \nu_e$	$7.04  imes 10^{-4}$	$3.90  imes 10^{-6}$	
	$2.3 \times 10^{-1}$	$\pi^- e^+ \nu_e$	40.55%	$4.44 \times 10^{-5}$	
$\Gamma_L$	1	$\pi^-\mu^+ u_\mu$	27.04%	$2.94  imes 10^{-5}$	
$D^+$	$3.4 \times 10^{-4}$	$e^+\nu_e$	$9.49  imes 10^{-9}$	$4.78 \times 10^{-14}$	
		$\overline{K^0}e^+\overline{\nu_e}$	8.72%	$4.41 \times 10^{-7}$	
		$\mu^+ \nu_{\mu}$	$3.74 \times 10^{-4}$	$1.92 \times 10^{-9}$	
		$\overline{K^0}\mu^+\overline{\nu_{\mu}}$	8.76%	$4.48 \times 10^{-7}$	
		$\tau^+ \nu_{\tau}$	$1.20 \times 10^{-3}$	$6.09 \times 10^{-9}$	
$D_s^+$	$9.8 \times 10^{-5}$	$e^+\nu_e$	$1.25 \times 10^{-7}$	$1.48 \times 10^{-13}$	
		$\mu^+ \nu_\mu$	$5.43 \times 10^{-3}$	$6.54 \times 10^{-9}$	
		$\tau^+ \nu_{\tau}$	5.32%	$6.33 \times 10^{-8}$	
$\tau^+$ $\tau^-$	$5.2 \times 10^{-6}$	$\overline{\nu_{\tau}}e^{+}\nu_{e}$	17.82%	$1.13  imes 10^{-8}$	
		$\overline{\nu_{ au}}\mu^+ u_{\mu}$	17.39%	$1.11 \times 10^{-8}$	
		$\overline{\nu_{\alpha}}\ell_{\alpha}^{-}\nu_{\tau}$	35.21%	$2.22 \times 10^{-8}$	
		$\pi^0\pi^- u_{ au}$	25.49%	$1.60 \times 10^{-8}$	
		$\pi^- \nu_{ au}$	10.82%	$6.80 \times 10^{-9}$	

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**Table 1**. **PRELIMINARY**. List of parents P produced in the primary target T2, and decaying into SM neutrinos of flavor  $\alpha$ . The numbers are normalized per PoT.

#### Neutrinos entering the detector



# HNL

### **HNL:** Production

$$\mathcal{L} \supset -\frac{m_W}{v} \bar{N} U^*_{\alpha 4} \gamma^{\mu} l_{L\alpha} W^+_{\mu} - \frac{m_Z}{\sqrt{2}v} \bar{N} U^*_{\alpha 4} \gamma^{\mu} \nu_{L\alpha} Z_{\mu}$$

#### We consider the simplified phenomenological benchmarks of one HNL mixing with one SM neutrino of a given flavour



Parent	2-body decay	3-body decay	P	arent	2-body decay	3-body d
$\pi^+ \rightarrow$	$e^+N_4$		I	$D^+ \rightarrow$	$e^+N_4$	$e^+\overline{K^0}$
	$\mu^+ N_4$				$\mu^+ N_4$	$\mu^+\overline{K^0}$
$K^+ \rightarrow$	$e^+N_4$	$\pi^0 e^+ N_4$			$\tau^+ N_4$	
	$\mu^+ N_4$	$\pi^0 \mu^+ N_4$	I	$D_s^+ \rightarrow$	$e^+N_4$	
$\tau^- \rightarrow$	$\pi^- N_4$	$e^-\overline{\nu}N_4$			$\mu^+ N_4$	
	$ ho^- N_4$	$\mu^-\overline{\nu}N_4$			$\tau^+ N_4$	

#### (normalised per PoT)

D $4.8 \cdot 10^{-4}$ 



# Backgrounds

# $\pi^{\pm}\mu^{\mp}$ or $\mu^{\pm}\mu^{\mp}$

We have 1278 background events with the following cuts:

- We keep events with only two  $\mu$ -like ( $\pi^{\pm}$ ,  $\mu^{\pm}$ ) particles, above an energy threshold of 30 MeV.
- We reject events with other detectable particles in the final state.

We can reduce the events to 15 events with the following kinematical cuts:

- $p_T < 0.35 \text{ GeV}.$
- $\theta_{\mu\pi} < 0.18$  rad.

These 15 events of background are  $\mu^{\pm}\pi^{\mp}$ , for the channel  $\mu^{+}\mu^{-}$ , we could reduce this background further by noting that pions are more likely to interact in the TPC, producing noticeable differences in their tracks with respect to the muons.



 $e^{\pm}\mu^{\mp}$  or  $e^{\pm}\pi^{\mp}$ 

We have 1982 background events with the following cuts:

- We keep events with only one  $\mu$ -like ( $\pi^{\pm}$ ,  $\mu^{\pm}$ ) particle and one ( $e^{\pm}$ ), above an energy threshold of 30 MeV.
- We reject events with other detectable particles in the final state.

We can reduce the events to 24 events with the following kinematical cuts:

- $p_T < 0.35 \text{ GeV}.$
- $\theta_{e\mu} < 0.180$  rad.

These 24 events of background are  $e^{\pm}\pi^{\mp}$ , for the channel  $e^{-}\mu^{+}$ , we could reduce this background further by noting that pions are more likely to interact in the TPC, producing noticeable differences in their tracks with respect to the muons.



 $\pi^0$  or  $e^+e^-$ 

We have **721** background events with one single  $\pi^0$  in the final state. These  $\pi^0$  will promptly decay into 2 photons and it will be a background for  $e^+e^-$  channels when only one of the  $\gamma$  convert into the TPC. This occurs for around 1% of the events leaving a background of  $\sim$ 7 background events for  $e^+e^-$ . Further kinematical cuts need to be explored.

## Take home message

• Our preliminary studies tell us that some of the golden channels for the HNL will be close to background free.

# What is next?

- We need to compute the background using the Geant4 simulation for the neutrino flux.
- Explore other sources of background like cosmics
- Compute the background for other types of searches, like Milicharged particles where the signal comes from scattering.

### Thank you



# Back-up



### HNL: Decays into visible channels



#### New Physics: Decay in flight inside the detector

### **Detector(NP02) Liquid Argon TPC**



 $N_{dec}^{M} = N_{PoT} Y_{M} BR(M \to \Psi) \int dS \int dE_{\Psi} \mathcal{P} \left( c\tau_{\Psi} / m_{\Psi}, E_{\Psi}, \Omega_{\Psi} \right) \frac{dn^{M \to \Psi}}{dE_{\Psi} dS}$ 

 $N_{det} = N^M_{dec} \cdot BR(\Psi \rightarrow \text{ visible }) \cdot \epsilon_{det}$ 



### **HNL: Fluxes**

#### **HNL intersecting the detector**



- Wide HNL beam
- Small changes in the geometry will not significantly change the results
- Any of the two ProtoDUNE detectors can be used

• Quite energetic HNL beam



### HNL: Decays into visible channels (combination)

We consider the following channels  $N \to \nu ee, \nu \mu \mu, \nu e \mu, e \pi, \mu \pi$  and  $\nu \pi^0$ 

