

# ⊙ Thermalization Of Jets In QCD Plasma

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Based on:

- » IS [arXiv:2409.04806](https://arxiv.org/abs/2409.04806)
- » Y. Mehtar-Tani, S. Schlichting, and IS [JHEP 05 \(2023\) 091](https://arxiv.org/abs/2305.091)
- » S. Schlichting, IS [JHEP 07 \(2021\), 077](https://arxiv.org/abs/2107.077)



## ◉ Table of contents

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- ◉ Introduction
- ◉ Effective Kinetic Description
- ◉ Longitudinal Energy Loss
- ◉ Thermalization To Large Angles
- ◉ Beyond AMY

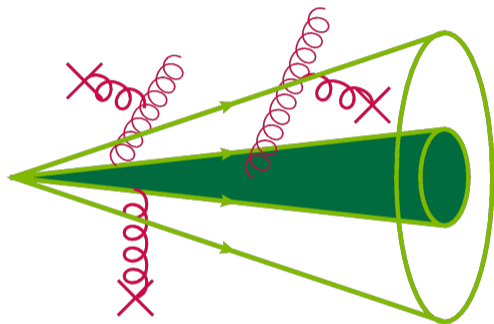
## ◉ Outline

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- ◉ Thermalization To Large Angles
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## Introduction I

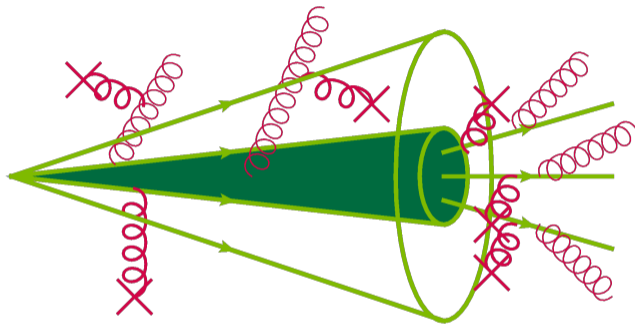
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- Hard parton evolution in heavy ion collisions

## Introduction I

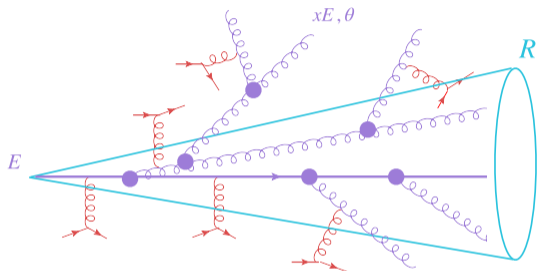
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- Hard parton evolution in heavy ion collisions
- Focus on energy loss and equilibration of hard partons  $\Rightarrow$  resolved by the medium

## ⊙ In-Medium Shower

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- Main focus: Hard parton traversing a QCD plasma
- Understand: Energy cascade, out-of-cone energy loss, medium response and full thermalization of the shower  $\Rightarrow$  Important for low energy jets

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## Effective Kinetic Description

- Leading Order Effective Kinetic Theory:

$$p^\mu \partial_\mu f_i(\vec{x}, \vec{p}, t) = C_i[\{f_i\}] , \quad (1)$$

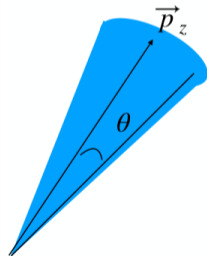
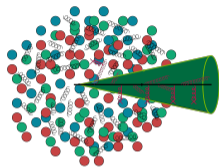
- Hard partons  $\Rightarrow$  Linearized Fluctuation on top of Static Equilibrium

$$f(p, t) = n_{\text{eq}}(p, t) + \delta f_{\text{jet}}(p, t) , \quad (2)$$

- Energy distribution

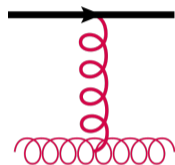
$$D(x, \theta, t) \equiv x \frac{dN_a}{dx d\cos\theta} \sim \frac{\nu_a(N_f)}{E_j} \delta f_{\text{jet}}(p, \theta, t) . \quad (3)$$

- Exact conservation of energy, momentum and valence charge  $\rightarrow$  Study evolution from  $\sim E$  to  $\sim T$  including full thermalization

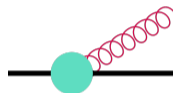




- Elastic Scatterings using HTL screened  $2 \leftrightarrow 2$



- Collinear Radiation effective  $1 \leftrightarrow 2$



$$C[\{f_i\}] = C^{2 \leftrightarrow 2}[\{f_i\}] + C^{1 \leftrightarrow 2}[\{f_i\}]. \quad (4)$$

- Analogous to QGP thermalization in pre-Hydro phase
- Using AMY rates for in-medium radiation:  $t_{\text{form}} \sim \sqrt{\frac{z(1-z)P}{\hat{q}}} \ll L$

## ◉ Elastic Scattering

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Elastic scattering collision integral with HTL screened matrix elements

$$C_a^{2\leftrightarrow 2}[\{f_i\}] = \frac{1}{2|p_1|v_a} \sum_{bcd} \int d\Omega^{2\leftrightarrow 2} |\mathcal{M}_{cd}^{ab}(p_1, p_2, p_3, p_4)|^2 \delta\mathcal{F}(p, p_1, p_2, p_3, p_4) \quad (5)$$

Full detailed balance terms  $\Rightarrow$  Medium response

$$\begin{aligned} \delta F(p_1, p_2, p_3, p_4) = & \delta f_a(p_1)[\pm_a n_c(p_3) n_d(p_4) - n_b(p_2)(1 \pm n_c(p_3) \pm n_d(p_4))] \\ & + \delta f_b(p_2)[\pm_b n_c(p_3) n_d(p_4) - n_a(p_1)(1 \pm n_c(p_3) \pm n_d(p_4))] \\ & - \delta f_c(p_3)[\pm_c n_a(p_1) n_b(p_2) - n_d(p_4)(1 \pm n_a(p_1) \pm n_b(p_2))] \\ & - \delta f_d(p_4)[\pm_d n_a(p_1) n_b(p_2) - n_c(p_3)(1 \pm n_a(p_1) \pm n_b(p_2))] . \end{aligned} \quad (6)$$

## Collinear Radiation

Collinear radiation (including BH & LPM effects)

$$\begin{aligned}
 C_g^{g \leftrightarrow gg}[\{D_i\}] = & \int_0^1 dz \frac{d\Gamma_{gg}^g\left(\left(\frac{xE}{z}\right), z\right)}{dz} \left[ D_g\left(\frac{x}{z}\right) \left(1 + n_B(xE) + n_B\left(\frac{\bar{z}xE}{z}\right)\right) \right. \\
 & + \frac{D_g(x)}{z^3} \left( n_B\left(\frac{xE}{z}\right) - n_B\left(\frac{\bar{z}xE}{z}\right) \right) + \frac{D_g\left(\frac{\bar{z}xE}{z}\right)}{\bar{z}^3} \left( n_B\left(\frac{xE}{z}\right) - n_B(xE) \right) \left. \right] \\
 & - \frac{1}{2} \int_0^1 dz \frac{d\Gamma_{gg}^g(xE, z)}{dz} \left[ D_g(x) (1 + n_B(zxE) + n_B(\bar{z}xE)) \right. \\
 & \left. + \frac{D_g(zx)}{z^3} (n_B(xE) - n_B(\bar{z}xE)) + \frac{D_g(\bar{z}x)}{\bar{z}^3} (n_B(xE) - n_B(zxE)) \right],
 \end{aligned}$$

- AMY Radiation Rate

$$\frac{d\Gamma_{bc}^a}{dz}(P, z, \infty) = \frac{g^2 P_{bc}^a(z)}{4\pi P^2 z^2 (1-z)^2} \text{Re} \int_0^\infty dt_1 \int_{p,q} \frac{i\mathbf{q}\cdot\mathbf{p}}{\delta E(\mathbf{q})} \Gamma_3(t) \circ G(\infty, \mathbf{q}; t_1, \mathbf{p}). \quad (7)$$

- Merging/Splitting  $\Rightarrow$  Thermalization of soft sector

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- ◉ Beyond AMY

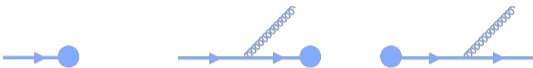
## ◉ Longitudinal Energy Loss

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- Evolution can be divided into three regimes:
  1. Initial energy loss: mediated by single gluon radiation
  2. Energy cascade: successive emissions lead to an energy cascade
  3. Equilibration: Late time decay to soft sector

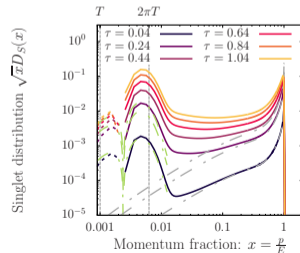
## Single Emission

- Initial distribution determined from direct emissions



$$D_q(x, \tau) \simeq \delta(1-x) + \left[ x \frac{d\Gamma_{gq}^q(x)}{dz} - \int_0^1 dz z \frac{d\Gamma_{gq}^q(x, z)}{dz} \delta(1-x) \right] \tau \quad (8)$$

- Short lived stage  $\Rightarrow$  Subsequent emissions lead to energy cascade



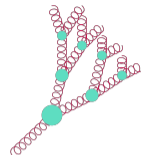
# Energy Cascade

- Stationary turbulent solution in intermediate energy range  $T/E \ll x \ll 1$ :

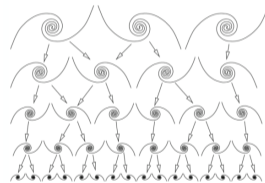
$$D_g(x) = \frac{G}{\sqrt{x}}, \quad D_S(x) = \frac{S}{\sqrt{x}}. \quad (9)$$

- Fixed point of the differential equation

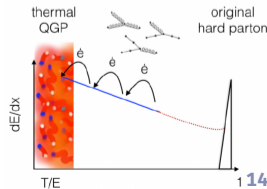
$$C_g[\{D_i\}] = \int_0^1 dz \frac{d\Gamma_{gg}^g(xE/z, z)}{dz} D_g\left(\frac{x}{z}\right) - \frac{1}{2} \frac{d\Gamma_{gg}^g(xE, z)}{dz} D_g(x) \\ + \int_0^1 dz \frac{d\Gamma_{gq}^q(xE/z, z)}{dz} D_S\left(\frac{x}{z}\right) - N_f \int_0^1 dz \frac{d\Gamma_{q\bar{q}}^g(xE, z)}{dz} D_g(x), \quad (10)$$



Energy injection



Viscous dissipation



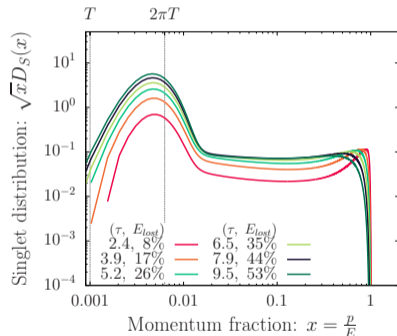
## Energy Cascade

- Stationary turbulent solution in intermediate energy range

$T/E \ll x \ll 1$ :

$$D_g(x) = \frac{G}{\sqrt{x}}, \quad D_S(x) = \frac{S}{\sqrt{x}}. \quad (9)$$

- Existence of a fixed point of the differential equation comes from the fact that the rate behaves as  $1/\sqrt{xE}$





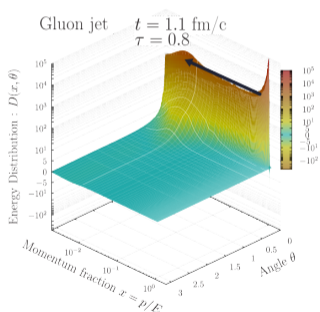
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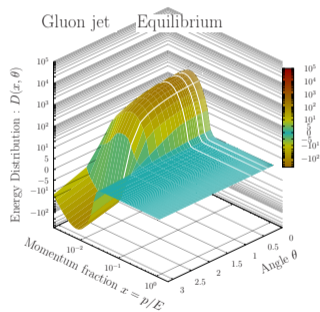
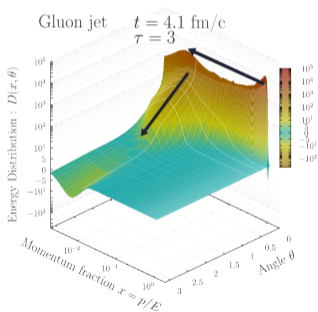
- ◉ Introduction
- ◉ Effective Kinetic Description
- ◉ Longitudinal Energy Loss
- ◉ Thermalization To Large Angles
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## Thermalization To Large Angles

- Energy cascade to soft sector
- Broadening of soft partons to large angle
- Thermalization of soft sector



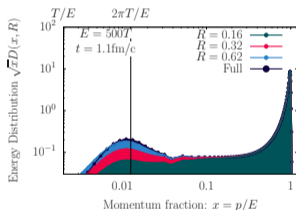
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<sup>1</sup> $E/T = 500, g = 2.$

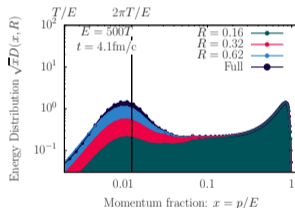
## Thermalization To Large Angles

- Energy cascade to soft sector

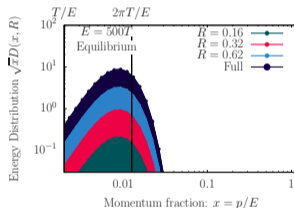


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- Broadening of soft partons to large angle



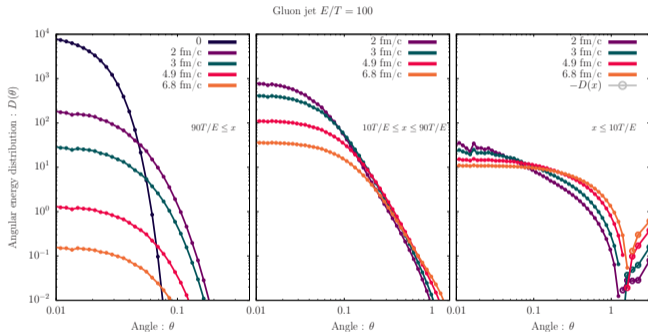
- Thermalization of soft sector



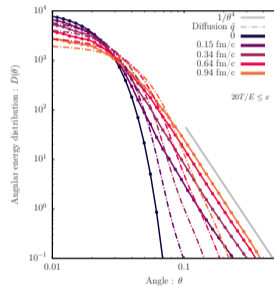
<sup>1</sup> $E/T = 500, g = 2.$

# ⊙ Molière Scattering

- Angular distribution in different energy regimes



- Hard sector



$$\partial_t D(x, \mathbf{p}_\perp, t) \simeq \int_{\mathbf{q}_\perp} C(\mathbf{q}_\perp) [D(x, \mathbf{p}_\perp - \mathbf{q}_\perp, t) - D(x, \mathbf{p}_\perp, t)] + \sum_{bc} \int_0^1 dz \left[ \frac{1}{z^2} \frac{d\Gamma_{ac}^b}{dz} \left( \frac{xE}{z}, z \right) D\left( \frac{x}{z}, \frac{\mathbf{p}_\perp}{z}, t \right) - \frac{1}{2} \frac{d\Gamma_{bc}^a}{dz} (xE, z) D(x, \mathbf{p}_\perp, t) \right], \quad (10)$$

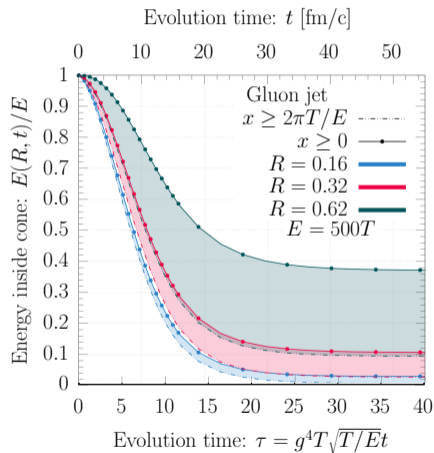
$$C(\mathbf{q}_\perp) \propto \frac{1}{\mathbf{q}_\perp^2 (\mathbf{q}_\perp^2 + m_D^2)} \quad (11)$$

## ○ Cone-Size Dependence

$$\text{---} \bullet \text{---} : E(R, \tau) = \int_0^\infty dx \int_{\cos R}^1 d\cos\theta D(x, \cos\theta, \tau) \quad (12)$$

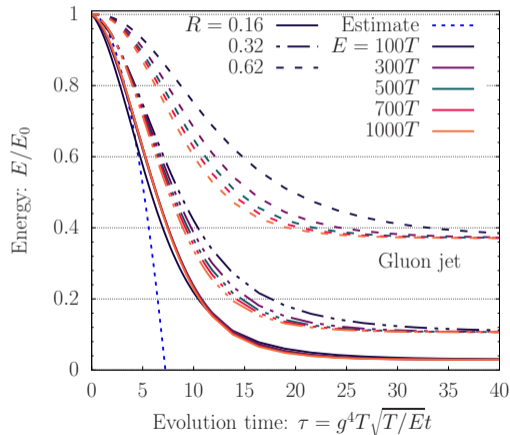
- Small cone-sizes: Soft sector does not play major role
- Large cone-sizes: Energy loss display different behavior  $\Rightarrow$  Dominated by thermalization

$$\text{---} \bullet \text{---} : E_{2\pi}(R, \tau) = \int_{2\pi T/E}^\infty dx \int_{\cos R}^1 d\cos\theta D(x, \cos\theta, \tau) \quad (13)$$



## ○ Sensitivity To Initial Parton's Energy

- Characteristic time of the turbulent cascade  $t_{th} = \frac{1}{\alpha_s} \sqrt{\frac{E}{\hat{q}}}$
- Scaling between different initial parton's energy for small cone-sizes
- Broken for large cone-sizes



## ◉ Modeling Jet Quenching

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- In the presence of QGP, the jet spectrum factorizes

$$\frac{d\sigma}{dp_T} = \int_0^\infty d\epsilon \, P(\epsilon, R) \frac{d\sigma_{\text{vac}}}{dp_T^{\text{in}}} (p_T^{\text{in}} \equiv p_T + \epsilon), \quad (14)$$

- The energy loss probability  $P(\epsilon, R)$  is obtained using the BDMPS rate

$$P(\epsilon, R) = \sum_{n=0}^{\infty} \frac{1}{n!} \left[ \prod_{i=1}^n \omega_i \frac{dI}{\omega_i} d\omega_i \right] \delta \left( \epsilon - \sum_{i=1}^n \omega_i \right) \exp \left[ - \int_0^\infty d\omega \frac{dI}{d\omega} \right] \quad (15)$$

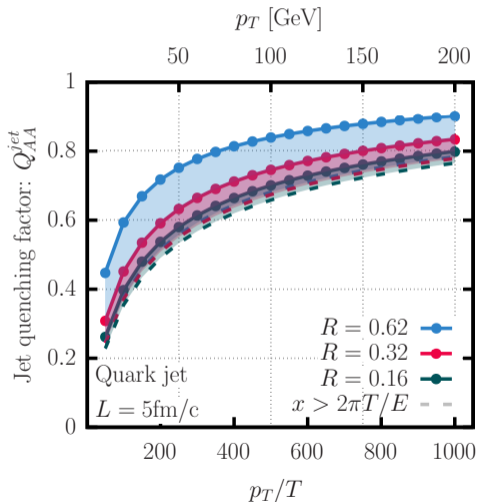
- Using a Milne transform  $\Rightarrow$  Exponential
- Taking vacuum spectrum  $\frac{d\sigma_{\text{vac}}}{dp_T^{\text{in}}} (p_T^{\text{in}}) \propto (p_T^{\text{in}})^{-\alpha}$

$$\frac{d\sigma}{dp_T} = Q(p_T, R) \frac{d\sigma_{\text{vac}}}{dp_T}, \quad \Rightarrow \quad Q(p_T, R) \approx \exp \left[ - \int d\omega \frac{dI}{d\omega} \left( 1 - e^{-n\omega/p_T} \right) \right] \quad (16)$$

## Modeling Jet Quenching

- The first emission is modeled using BDMPS finite medium rate  $\frac{d\Gamma}{d\omega}(P, \omega, t)$  at time  $t$
- Medium energy loss computed by modeling the energy remaining inside the cone  $E(\omega, R, L - t)$  after a time  $(L - t)$  in the medium

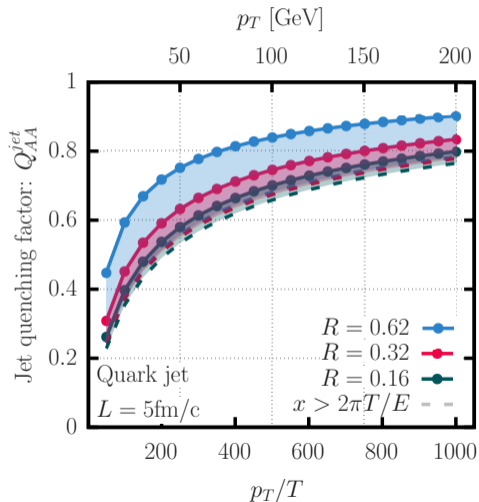
$$Q(p_T) = \exp \left[ \int_0^L dt \int d\omega \times \frac{d\Gamma}{d\omega} \left( 1 - e^{-n \frac{\omega}{p_T} \left[ 1 - \frac{E(\omega | R, \tau = \frac{L-t}{t_{th}})}{\omega} \right]} \right) \right]. \quad (17)$$





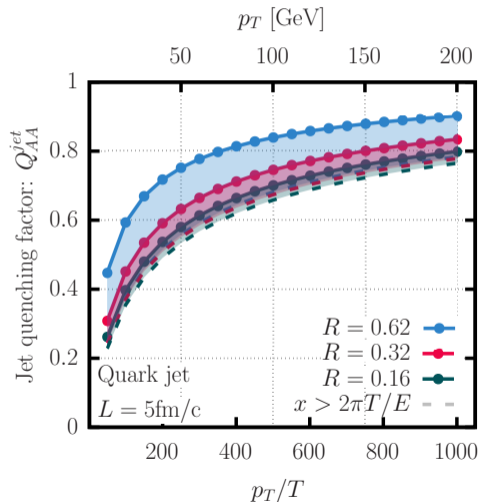
## Modeling Jet Quenching

- Jets with energies  $p_T \leq 25\text{GeV}$  lose significant energy in length  $L = 5\text{fm}/c$   
 $\Rightarrow$  large suppression at low  $p_T$ , milder for  $p_T \leq 200 - 400\text{GeV}$
- Negligible contribution of soft fragments to narrow cone
- Large cone size ( $\geq 0.3$ )  $\Rightarrow$  Recover energy from the soft sector  
 $\triangleright$  **Medium response**



## Modeling Jet Quenching

- ⚠ Over-estimate energy loss since we neglect finite size effects, Work in progress
- ⚠ Requires more refined studies of near-equilibrium physics and jet recoil onto the medium

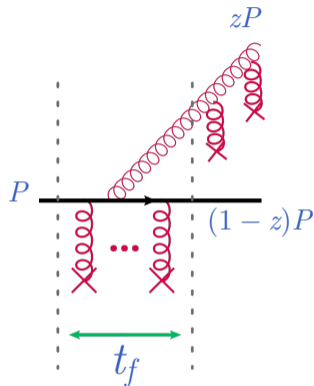


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## Collinear Radiation



- Multiple scatterings  $\Rightarrow$  induced radiation
- Emission controlled by the formation time

$$t_{\text{form}} \sim \frac{z(1-z)xE}{k_T^2}, \quad k_T^2 \sim \hat{q}t_{\text{form}} \Rightarrow t_{\text{form}} \sim \sqrt{\frac{z(1-z)xE}{\hat{q}}}, \quad (17)$$

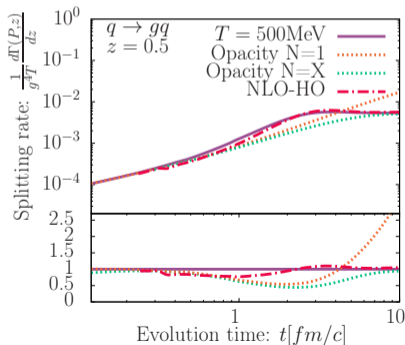
- $t_{\text{form}} \ll \lambda_{\text{mfp}}$ : Medium cannot resolve the quanta until it is formed
- $t_{\text{form}} \gg \lambda_{\text{mfp}}$ : Multiple scatterings act coherently
- Quantum interference  $\Rightarrow$  suppression of high energy radiation  $\Rightarrow$  LPM effect

<sup>1</sup>(Baier, Dokshitzer, Mueller, Peigné, Schiff, Zakharov, Wiedemann, Arnold, Moore, Yaffe..)

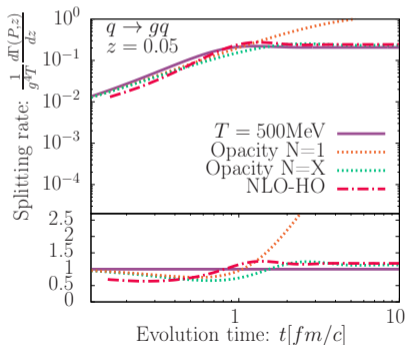
## Finite Size Effects: Rates

$$\frac{d\Gamma_{bc}^a}{dz}(P, z, t) = \frac{g^2 P_{bc}^a(z)}{4\pi P^2 z^2 (1-z)^2} \text{Re} \int_0^t dt_1 \int_{p,q} \frac{i\mathbf{q}\cdot\mathbf{p}}{\delta E(\mathbf{q})} \Gamma_3(t) \circ G(\infty, \mathbf{q}; t_1, \mathbf{p}). \quad (18)$$

- Long formation time:  $z \simeq 0.5$



- Short formation time:  $z(1-z) \simeq 1$



<sup>1</sup>S. Caron-Huot, C. Gale PRC(2010), S. Schlichting IS PRD(2022)

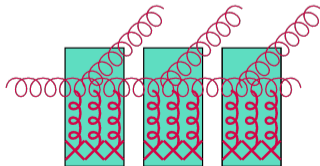
## Finite Size Effects: Evolution I

- Including finite size effects in the evolution
- Global formation time:

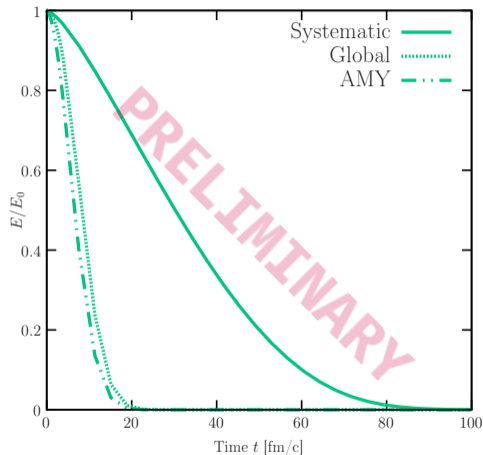
$$\partial_t D(x) = \int_0^1 dz \left[ \frac{d\Gamma_{bc}^a}{dz} \left( \frac{x}{z}, z, t \right) zD \left( \frac{x}{z} \right) - \frac{d\Gamma_{bc}^a}{dz} (x, z, t) zD(x) \right]$$

- Systematic formation time:

$$\partial_t D(x) = \int_0^t ds \int_0^1 dz \left[ \frac{d\Gamma_{bc}^a}{dz} \left( \frac{x}{z}, z, t-s \right) zD \left( \frac{x}{z} \right) - \frac{d\Gamma_{bc}^a}{dz} (x, z, t-s) zD(x) \right]$$

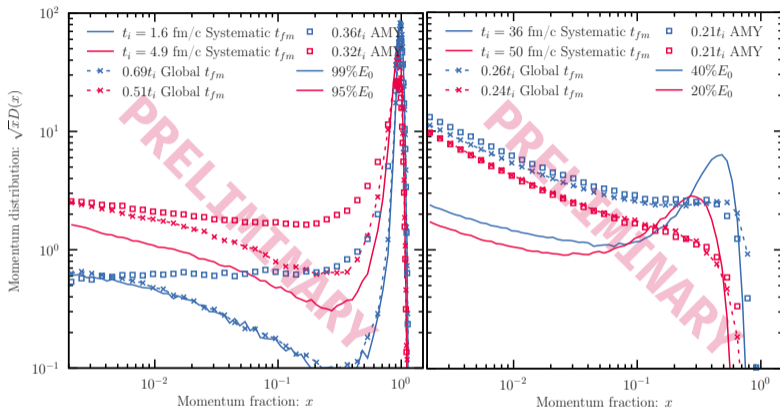


## Finite Size Effects: Evolution II



- Delay of energy loss due to formation time

## Finite Size Effects: Evolution III



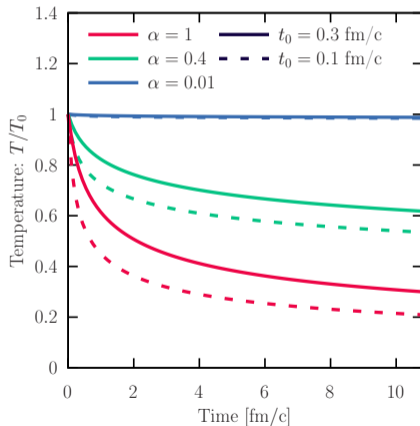
- Markedly different evolution for Systematic treatment of formation time

<sup>1</sup>Barata, Domínguez, Salgado, Vila JHEP (2021)



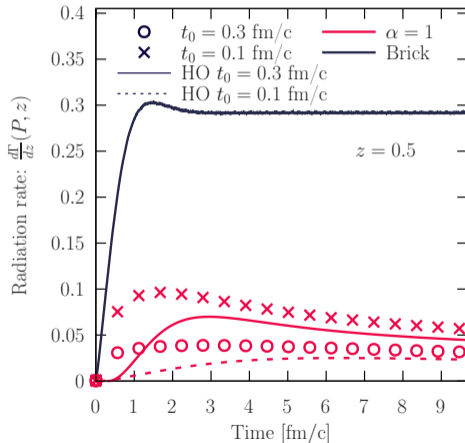
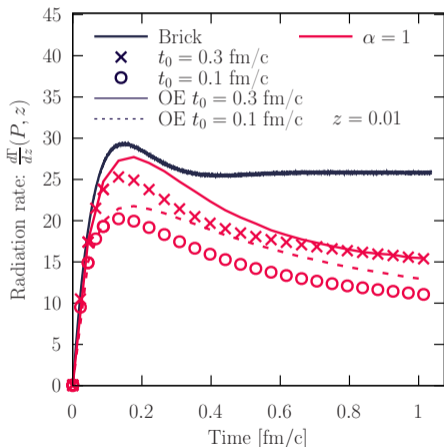
## ⊙ Bjorken Expansion: Medium-Induced Radiation I

- Bjorken expansion:  $T(t) = T_0 \left( \frac{t_0}{t+t_0} \right)^{1/3}$



## ⊙ Bjorken Expansion: Medium-Induced Radiation II

- Bjorken expansion:  $T(t) = T_0 \left( \frac{t_0}{t+t_0} \right)^{1/3}$



## ⊙ Conclusion

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- Energy loss is governed by an inverse energy cascade  
⇒ driven by successive collinear radiation
- Energy deposited collinearly at the soft scales rapidly broadens to large angles
- Formation time lead to dramatic effect on medium-cascade ⇒ Delay of energy loss, survival of high energy partons

Thank you for your attention \o/

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